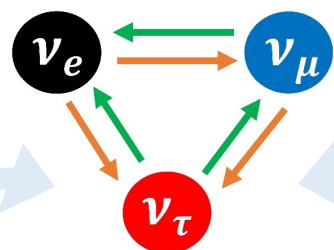


la Fisica  
“senza gli acceleratori”

## Ricerca di Materia Oscura



Neutrini

 VIRGO



Esperimenti per onde  
gravitazionali

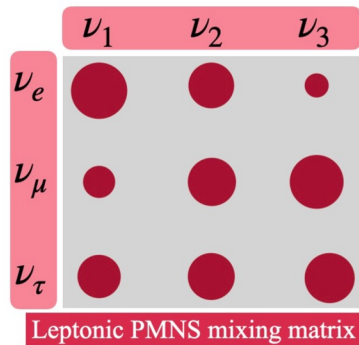
# The Flavor Problem: a question of numbers

## Mass hierarchies

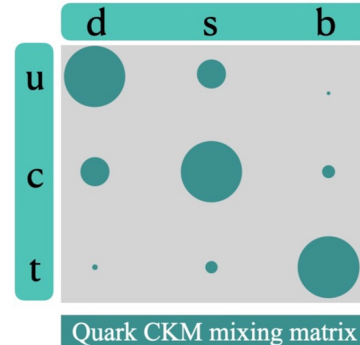
$$m_d \ll m_s \ll m_b, \quad \frac{m_d}{m_s} = 5.02 \times 10^{-2}, \quad \frac{m_s}{m_b} = 2.22 \times 10^{-2}, \quad m_b = 4.18 \text{ GeV};$$

$$m_u \ll m_c \ll m_t, \quad \frac{m_u}{m_c} = 1.7 \times 10^{-3}, \quad \frac{m_c}{m_t} = 7.3 \times 10^{-3}, \quad m_t = 172.9 \text{ GeV};$$

## Fermion mixing

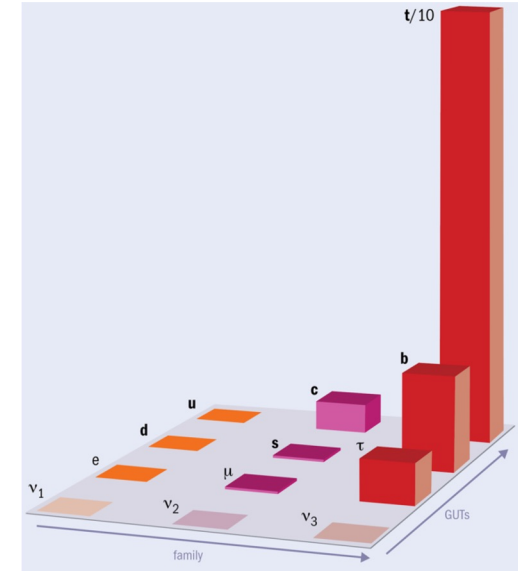


all mixing are large  
but the 1,3 element



almost a diagonal  
matrix

very small neutrino masses



Numbers NOT predicted by the SM:

***New ingredients are needed***

# The Flavor Problem: some popular suggested solutions

## - (Old) Good ideas

Smallness of neutrino masses:

**See-saw** mechanism:

SM  $\oplus$  heavy sterile neutrinos with mass  $M_R$

$$m_{light} \sim \frac{(vev_{SM})^2}{M_R}$$

**BUT no clue  
on mixing**

Mixing angles:

SM  $\otimes$  flavour symmetries

$$\text{PMNS} = \begin{bmatrix} 0.799 \dots 0.844 & 0.516 \dots 0.582 & 0.141 \dots 0.156 \\ 0.242 \dots 0.494 & 0.467 \dots 0.678 & 0.639 \dots 0.774 \\ 0.284 \dots 0.521 & 0.490 \dots 0.695 & 0.615 \dots 0.754 \end{bmatrix}$$

**BUT complicated  
scalar sector**

# The Flavor Problem: some popular suggested solutions

## - (Old) Good ideas

Smallness of neutrino masses:

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**BUT complicated scalar sector**

## - (New) Promising idea

Modular Symmetry: theory invariant under transformations induced by suitable 2x2 matrices acting on a complex variable  $\tau$

**Properties:**

# small number of operators (few free parameters)  $\rightarrow$  **predictability**

# no new matter fields  $\rightarrow$  **minimality**

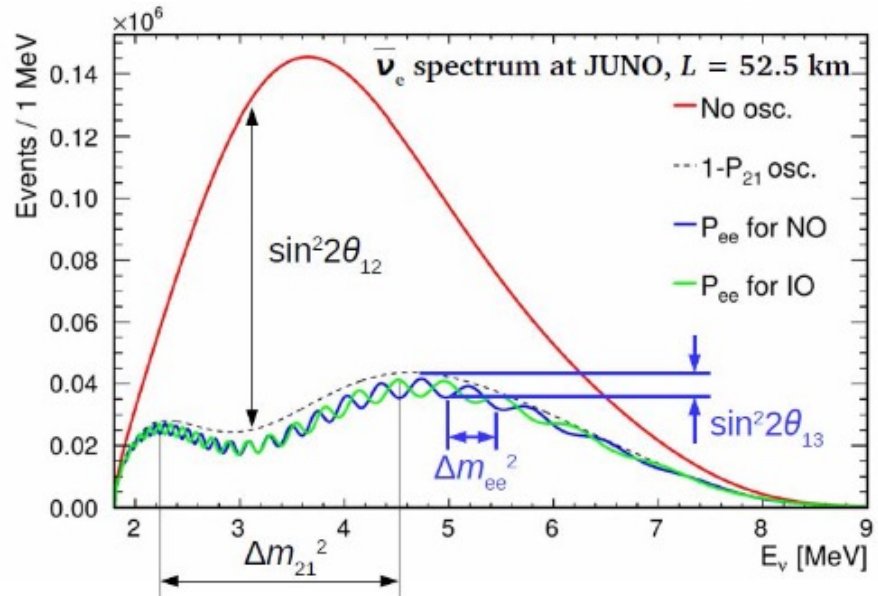
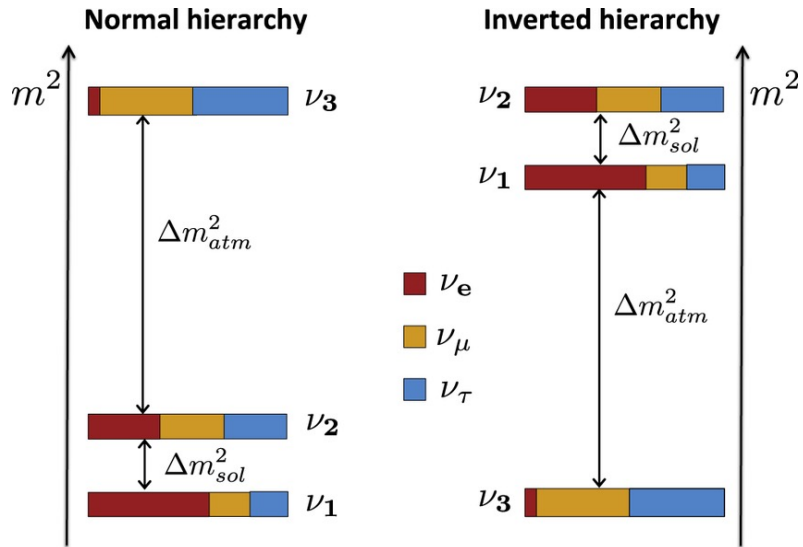
# no new scalar fields beside Higgs(es)  $\rightarrow$  **symmetry breaking**

**dictated by the vev of  $\tau$**



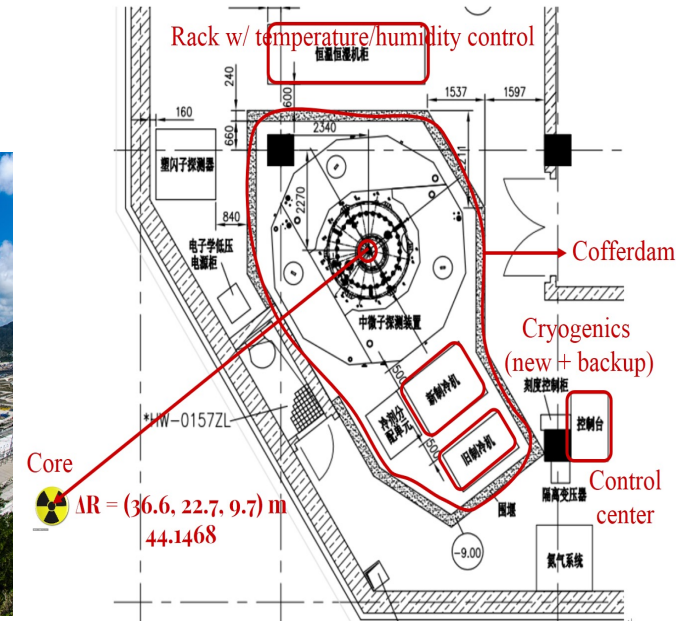
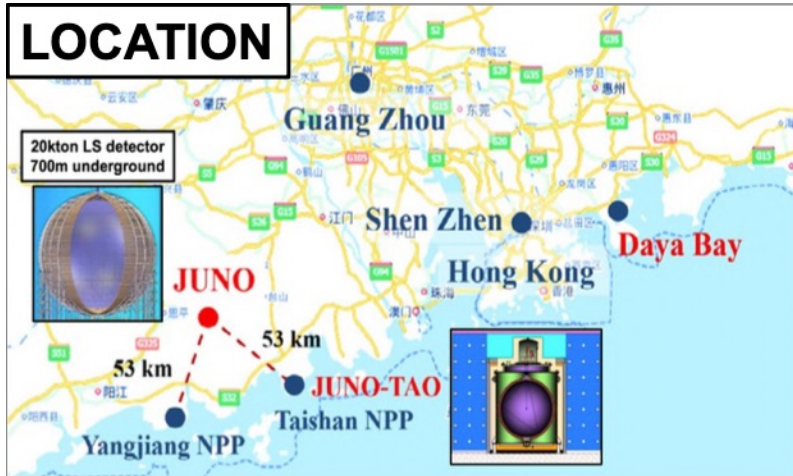


# The JUNO experiment





# The JUNO-TAO satellite detector

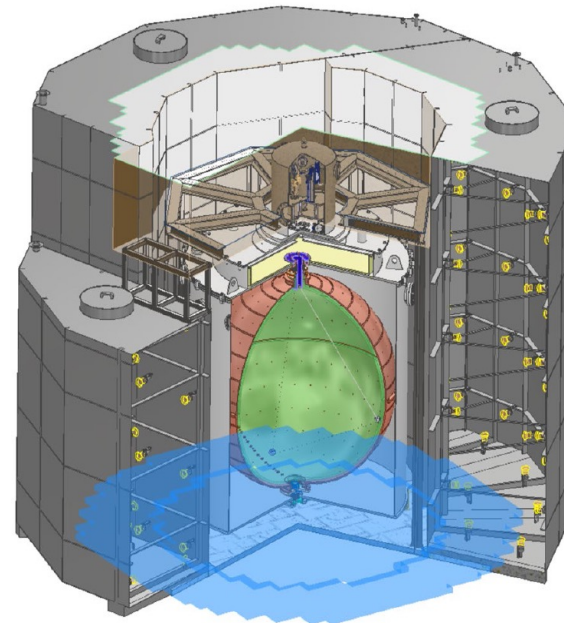


Taishan Antineutrino Observatory:  
a ton-level, liquid scintillator (LS) detector.

~10 m<sup>2</sup> of SiPMs at -50°C

Measure reactor neutrino spectrum  
with sub-percent E resolution.

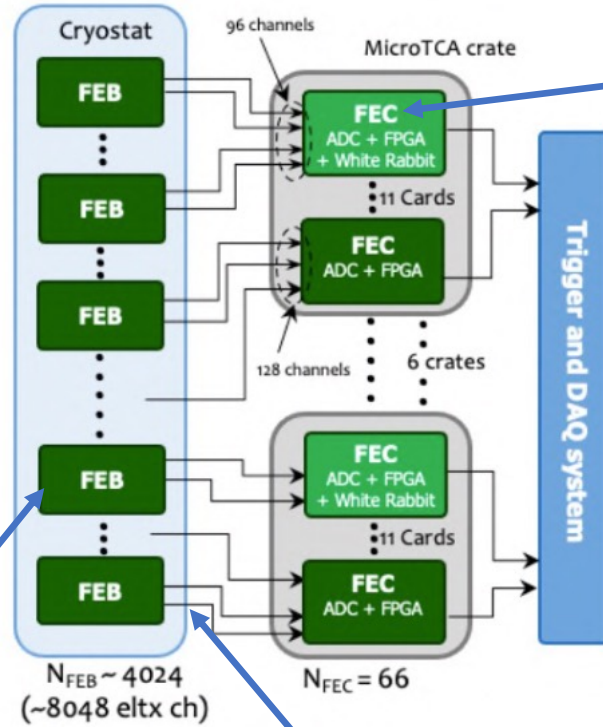
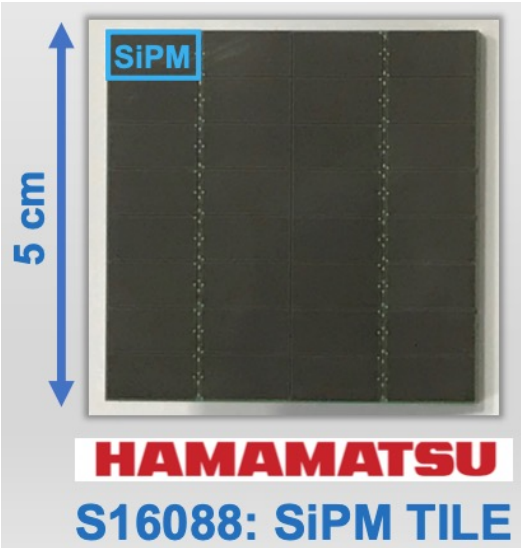
Primary physics goal: **Model-independent  
reference spectrum for JUNO**



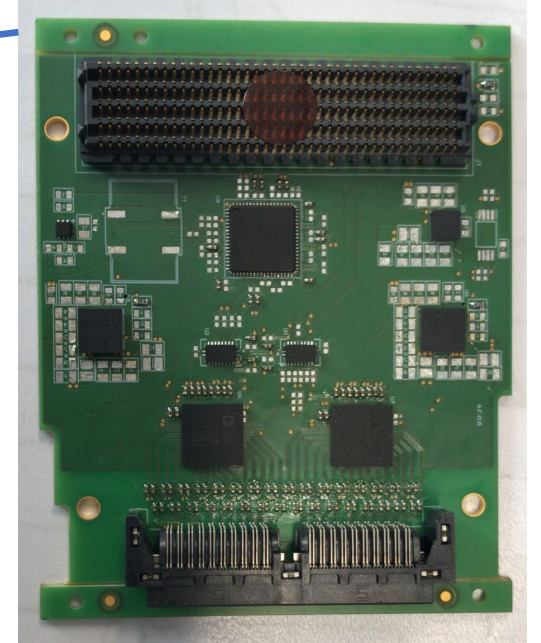


# The JUNO-TAO satellite detector

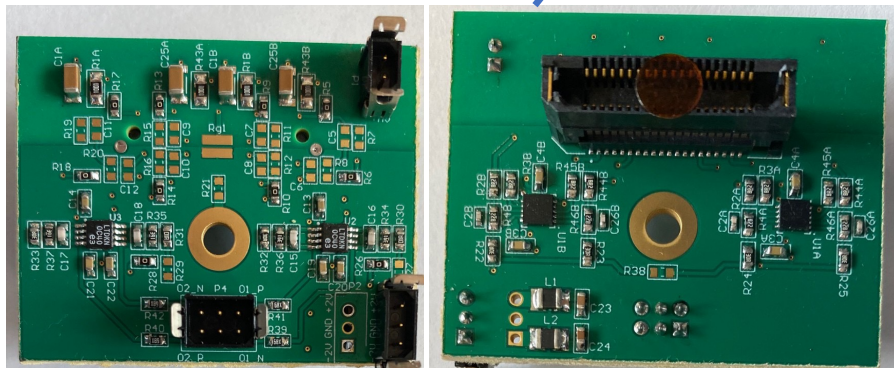
~4000 tiles → ~ 10 m<sup>2</sup>



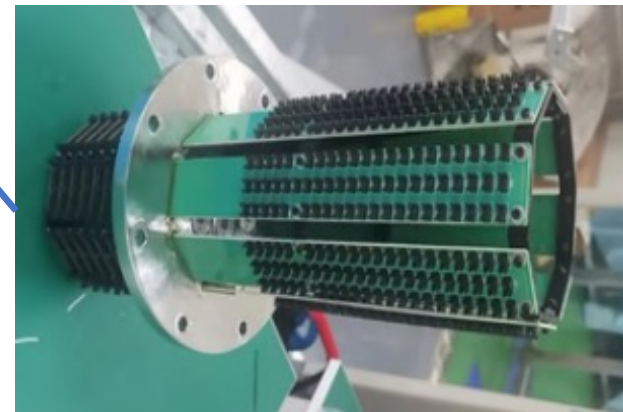
ADC



FEB



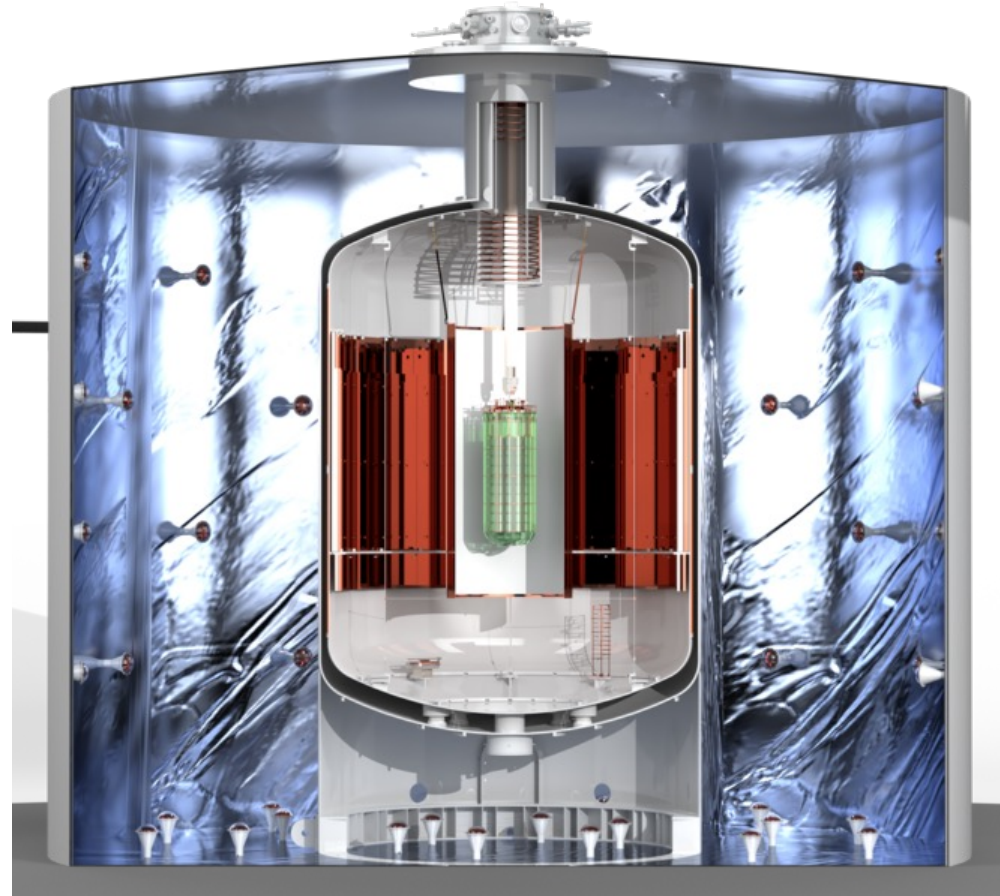
Flanges





# LEGEND

Large Enriched  
Germanium Experiment  
for Neutrinoless  $\beta\beta$  Decay





## Gruppo giovane (4 anni), ma in crescita

- 5 FTE (4 staff + 1 post-doc + 1 dottoranda + alcuni laureandi)
- **1 infrastruttura criogenica locale “OLAF”**
- Vari articoli e proceedings

### Ruoli di responsabilità scientifica:

- Run coordination (L200)
- E-scale analysis convener (L200)
- L3 FE electronics SIPM (L200 e L1000)



Photo: M. Willers / LEGEND Collaboration

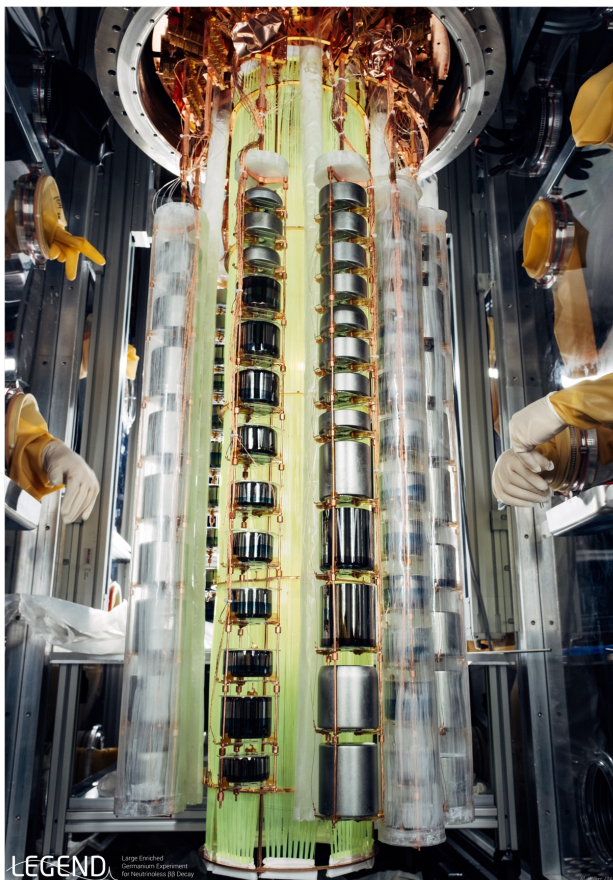
## Attività HW e SW

### LEGEND-200:

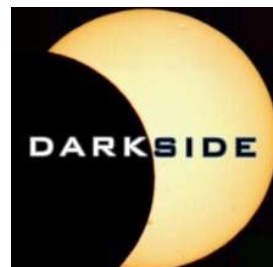
- Contributo al progetto e costruzione del **sistema di lettura in argon liquido**
- Guida delle **analisi** di scala e calibrazione **energia dei Germani**

### LEGEND-1000:

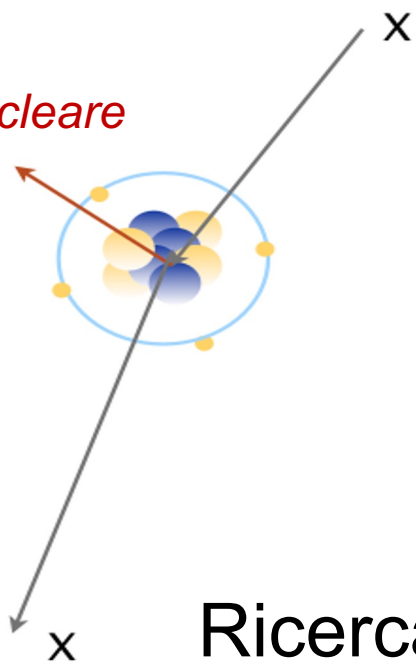
- Partecipazione all'R&D per il sistema di veto in argon atmosferico contro neutroni cosmogenici
- Infrastruttura per i test finali del FE dei Germani



## Ricerca di Materia Oscura



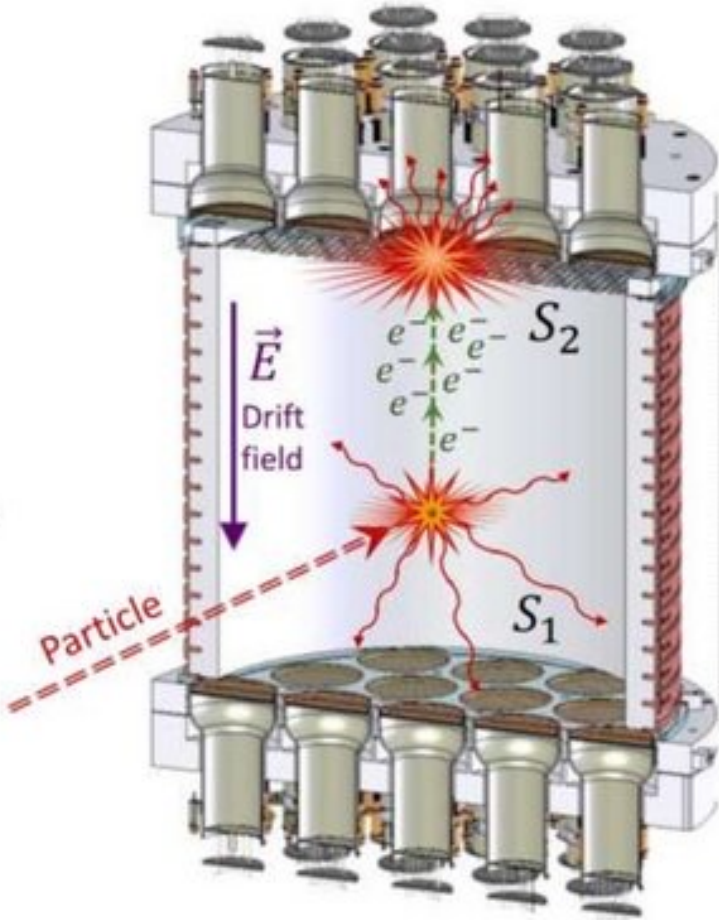
*Rincolo nucleare*



Ricerca Diretta



# DarkSide



DarkSide-10

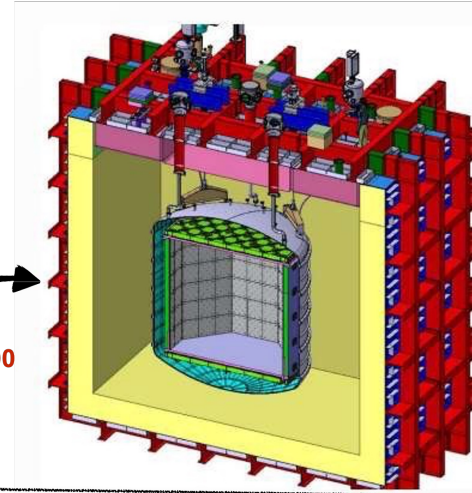


DarkSide-50

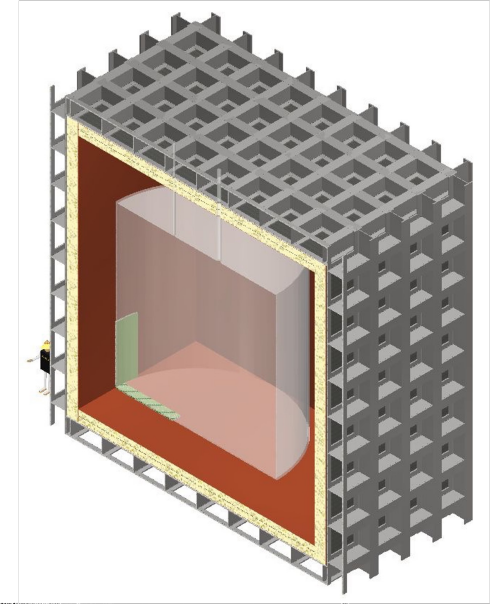


x 1000

DarkSide-20k



ARGO



2011

2013-2018 at LNGS  
0.03 t.yr exposure

2025+

200 t.yr target exposure

2030+

3000 t.yr target exposure



Installation started  
in Hall C at LNGS



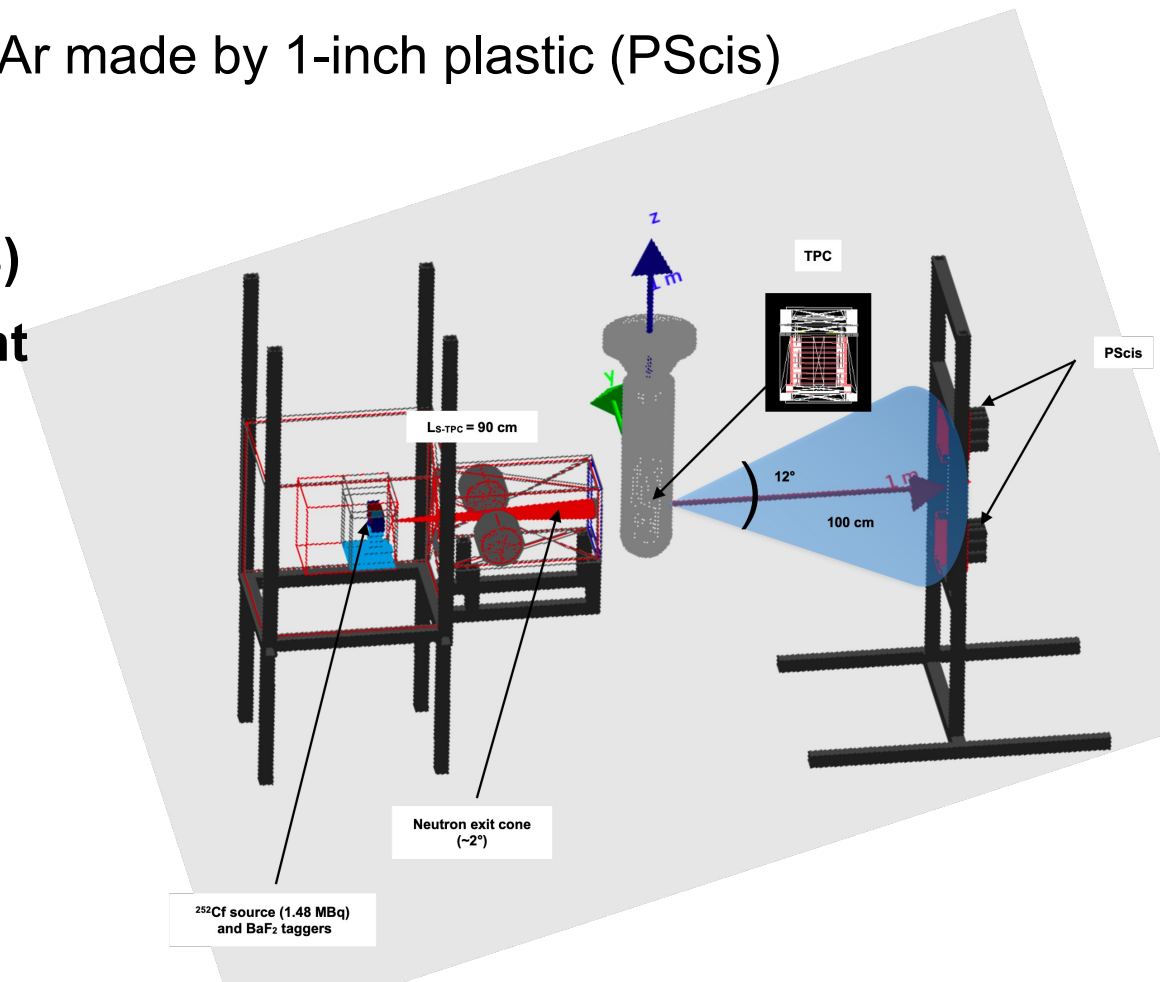
# DarkSide: Low-Energy Calibration at INFN - CT

**ReD: Low-energy recoil measurements (few keV) by using a  $^{252}\text{Cf}$  neutron source**

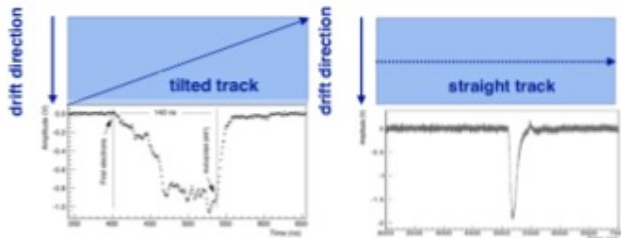
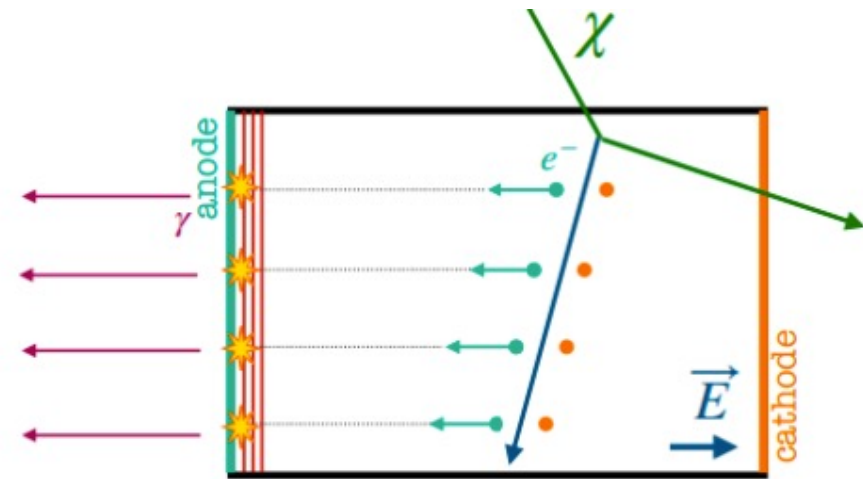
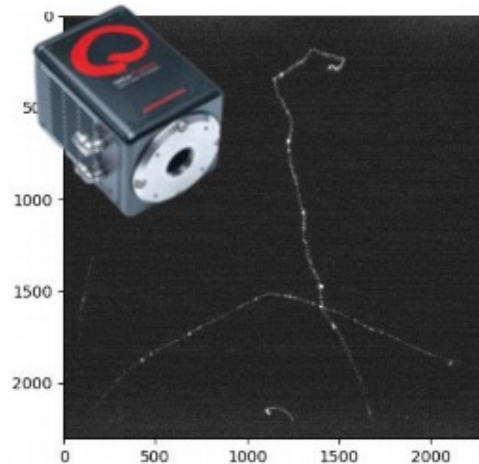
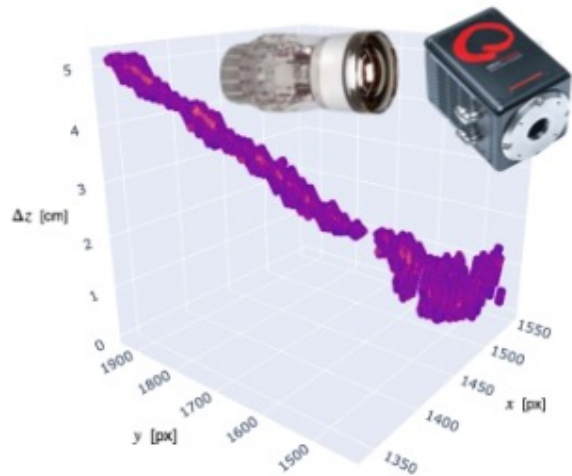
- Neutrons of  $\sim 2$  MeV kinetic energy more appropriate for  $E_{\text{rec}} \sim \text{few keV}$
- Neutron spectrometer to detect neutrons scattered off-Ar made by 1-inch plastic (PScis)
- Use  $\text{BaF}_2$  as close fission tagger

**Strategy: to tag the fission events (neutrons and  $\gamma$ s) with  $\text{BaF}_2$  detectors and measuring the time-of-flight in the “far” PScis**

- Use the PSD and time of flight for the n/ $\gamma$  rejection
- Sensitivity down to 2-5 keV<sub>NR</sub>

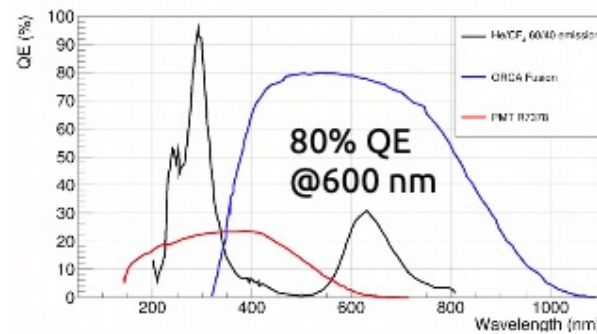


# CYGNO: a gaseous TPC with optical readout for directional DM searches



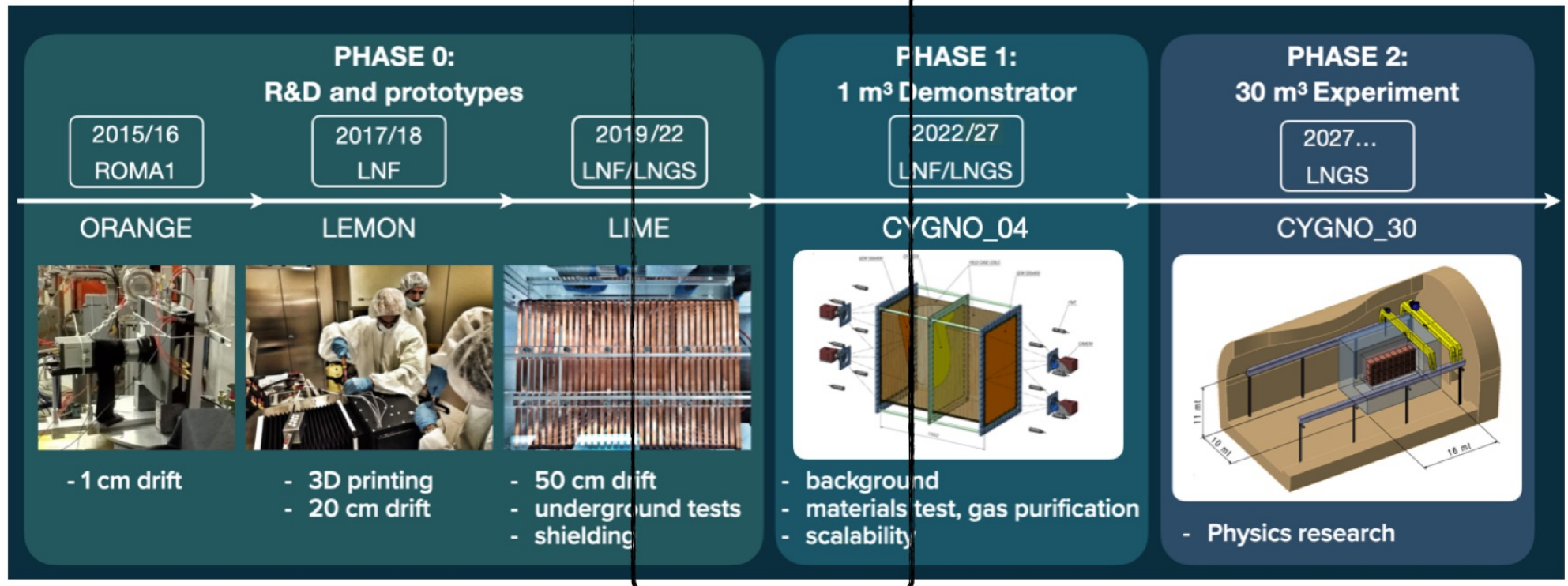
## PMTs:

Energy, z component of track from time structure of signal

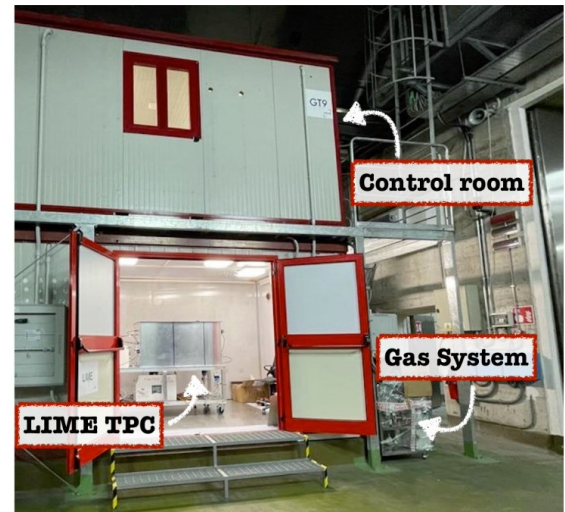


**sCMOS camera:**  
Energy, xy track projection

- He:CF<sub>4</sub> (60:40) at **atmospheric pressure**, GEM amplification, **optical readout** (sCMOS+PMT)
  - With suitable lenses we can **image large areas** O(1 m<sup>2</sup>) with single sensor, with effective pixel area O(100 μm)
- **3D reconstruction** capability: directionality, particle ID, background rejection, fiducialization
- **Low threshold** (< 1 keV)



# The CYGNO timeline



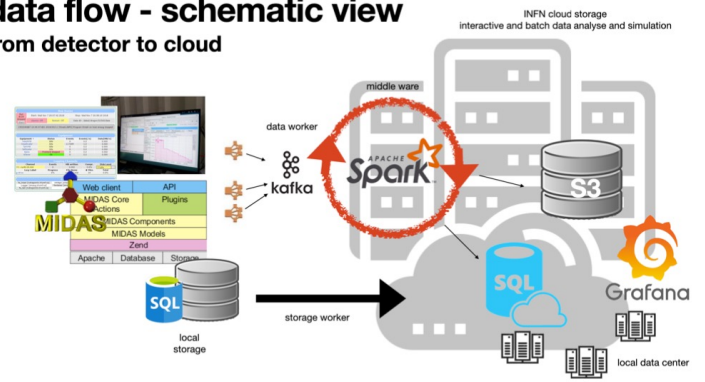




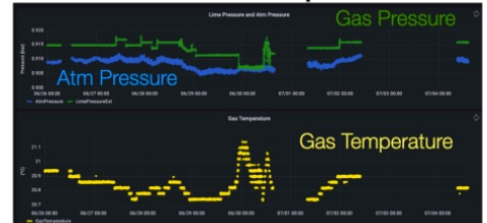
# Attività in CYGNO a Roma Tre



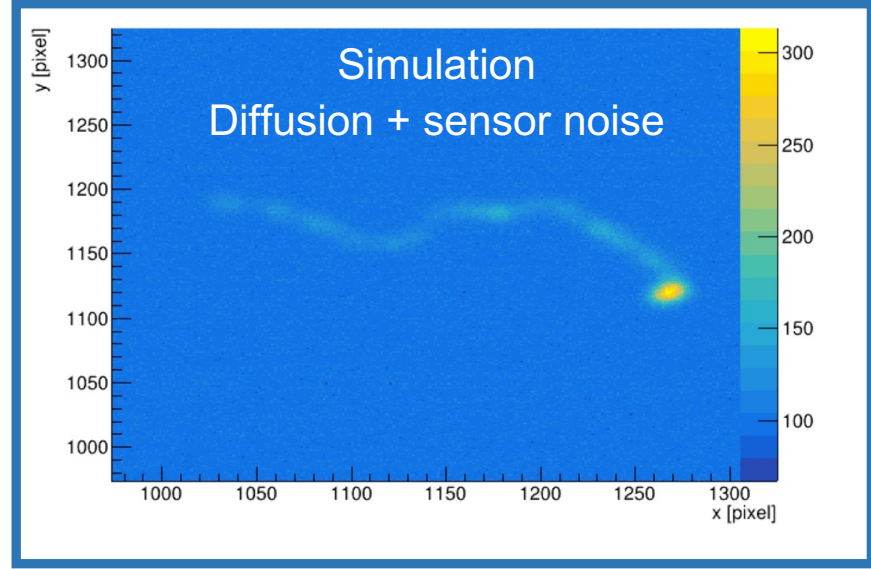
## data flow - schematic view from detector to cloud



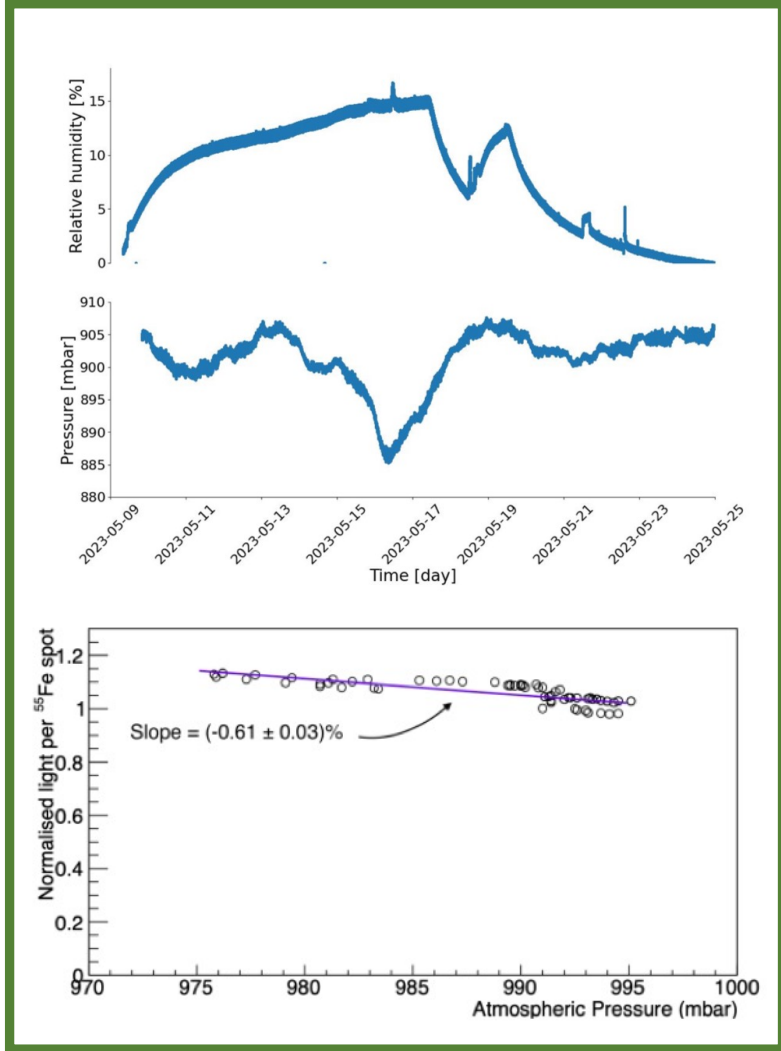
### Gas and environmental parameters

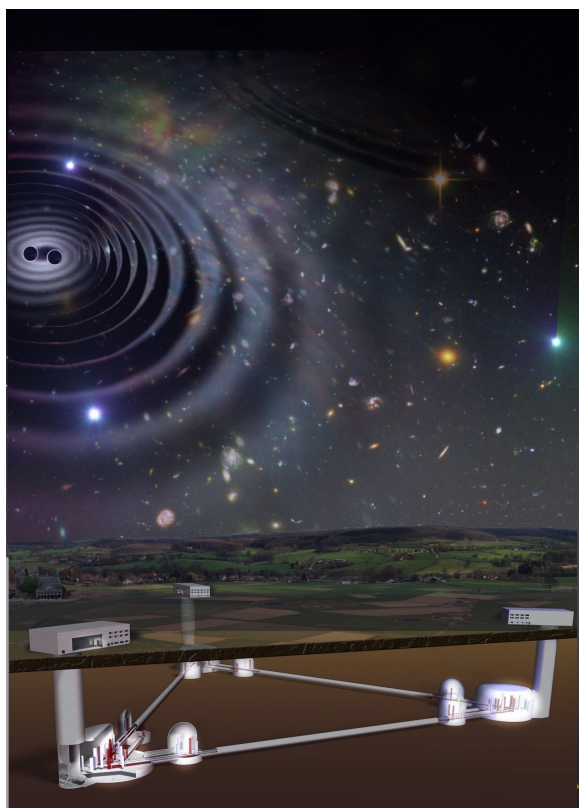


### Detector performance



- “Middleware”
  - Simulazione/Digitizzazione
  - Ricostruzione
  - Slow control
- Detector stability



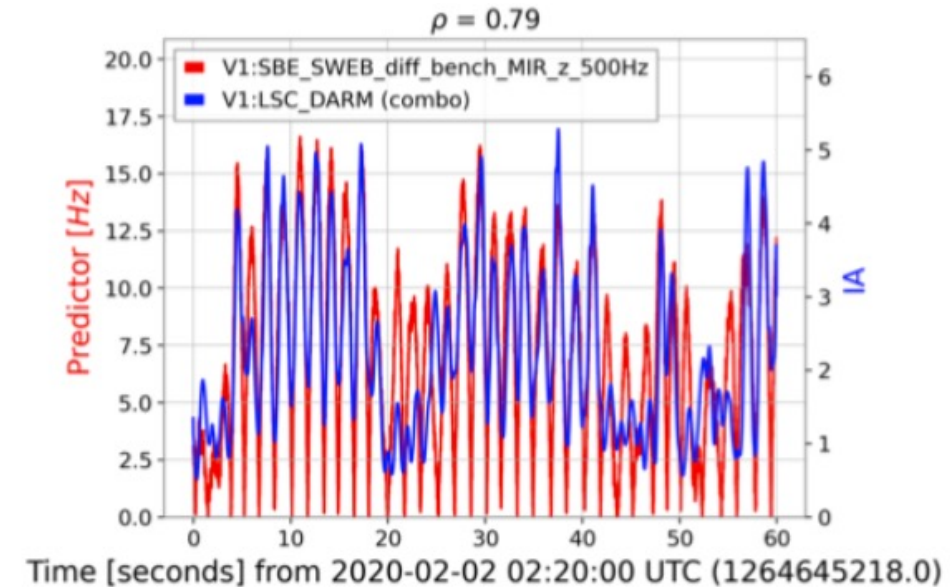


Esperimenti per onde  
gravitazionali

Method for the characterization of the noise from Scattering Light that can lower the sensitivity, based on adaptive algorithms

## Methodology

- Lowpass filter differential arm motion DoF (DARM) time series,
- Apply pytvfemd adaptive algorithm to obtain DARM's oscillatory modes (IMF)
- Mode's envelope is the **instantaneous amplitude IA(t) of the arm oscillatory modes**
- Compute **correlation  $\rho$  between IA(t) and the predictor  $f_{arch}$  [Hz]** for a given optic
- **Predictor:** obtained from position data  $x(t)$  of optical component
- Considering 60s interval: a time series of 1440 values of  $\rho$  is obtained each day
- **Daily monitoring: Compare  $\rho$  time series with seismometer data** in the microseismic band (0.1-1Hz)

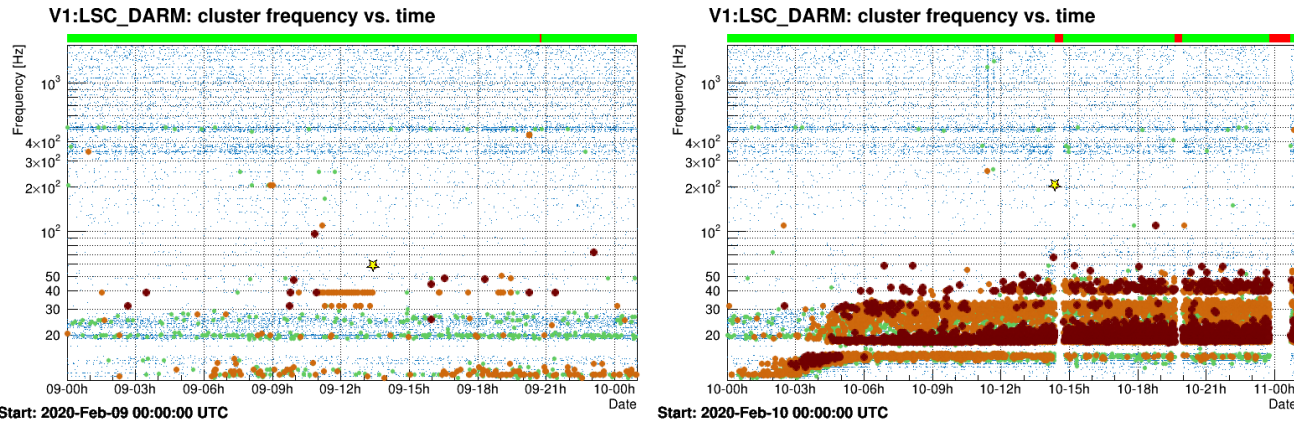


Predictor: tracks scattered light arches appearing in DARM spectrogram

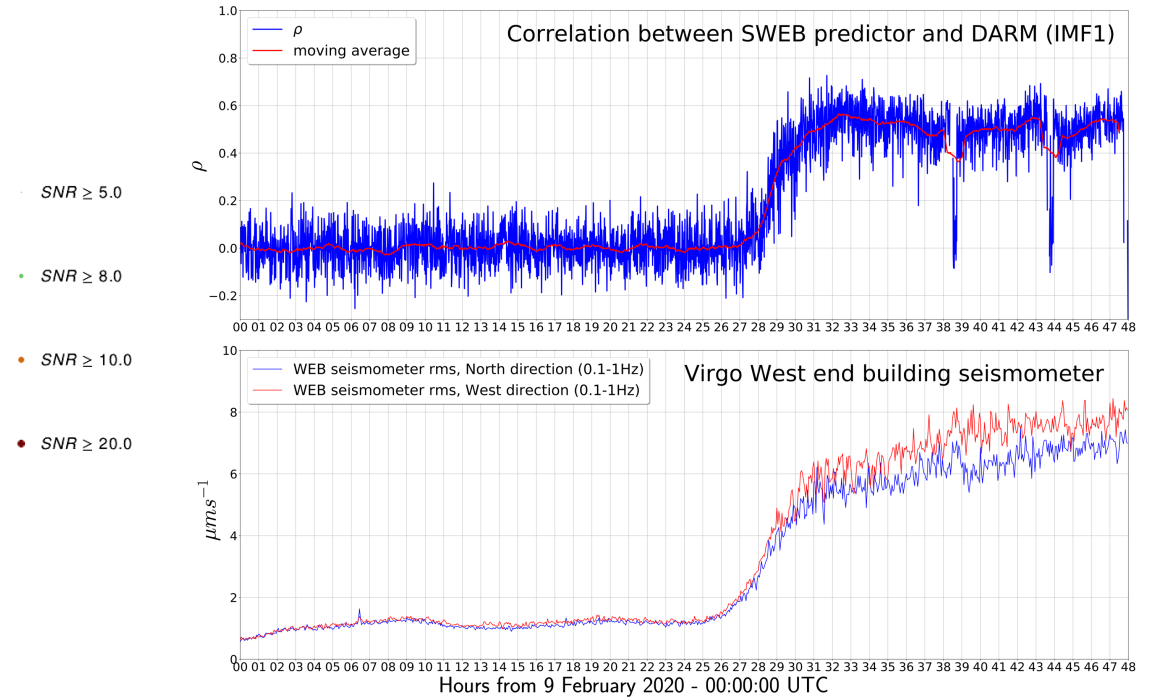
$$f_{arch}(t) = \frac{2|v(t)|}{\lambda}$$

# Example: Daily monitoring of Scattered light from Virgo Suspended West End Building (SWEB) during O3

Noise in Virgo DARM increases in a specific moment (2020/02/10 03h)

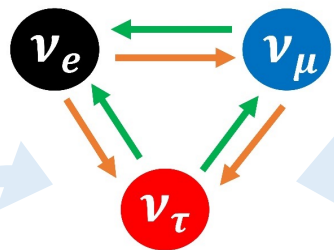


Source identified and monitored



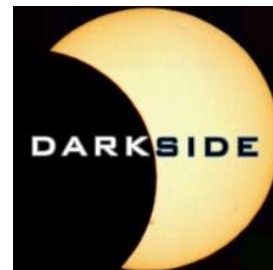
The monitoring system allows to continuously monitor multiple optics, possible sources of SL. Will run during O4 on Condor.

For each optic to be monitored —> 1440 parallel jobs per daily analysis  
*Pytvfemd* is used. Is available at <https://pypi.org/project/pytvfemd/> (Stefano Bianchi, PhD)



Neutrini

Ricerca di  
Materia Oscura



VIRGO



Esperimenti per onde  
gravitazionali

***Grazie per l'attenzione!***

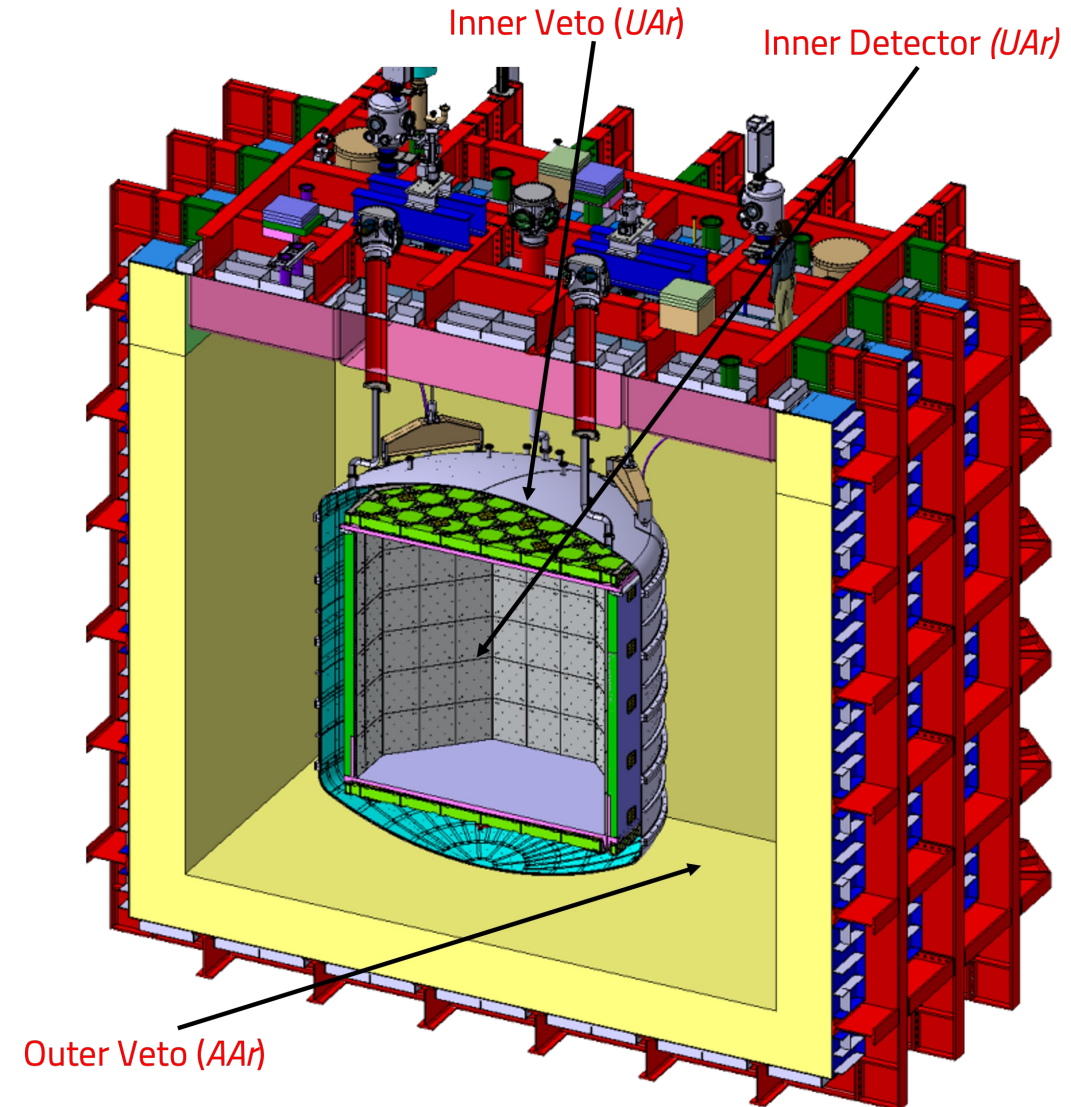
## **Additional Material**



# DarkSide-20k: overview



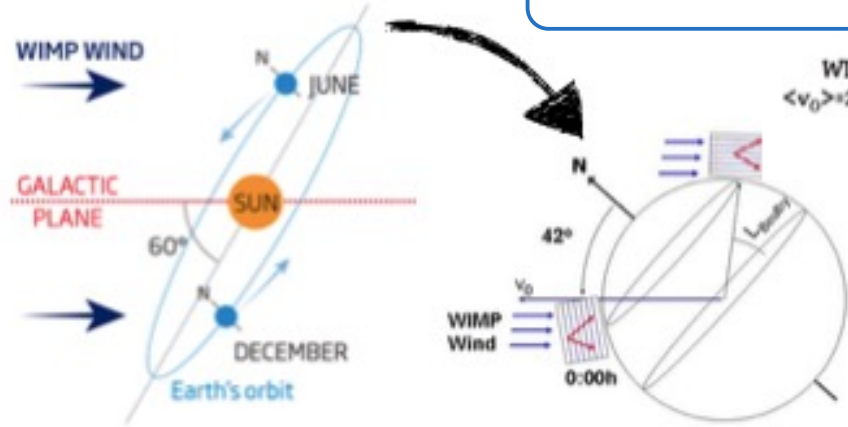
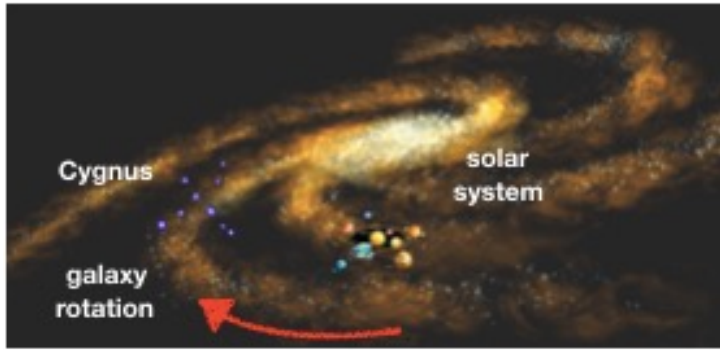
- ProtoDune-like cryostat instrumented like a muon veto ( $8 \times 8 \times 8 \text{ m}^3$ )
- Titanium vessel separating Atmospheric Argon (AAr) from Underground Argon (UAr)
- WIMP detector fiducial volume of  $\sim 20$  tonnes ( $\sim 50$  tonnes total) of UAr, depleted in  $^{39}\text{Ar}$
- Active neutron veto integrated into the TPC structure via gadolinium-loaded acrylic (PMMA)
- Silicon PhotoMultipliers (SiPMs) as photo detection devices (total area  $\sim 26 \text{ m}^2$ )



WIMP flux  
 $\langle v \rangle \sim 220 \text{ km/s}$

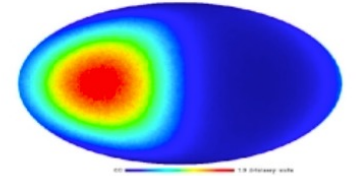
few % DM flux  
 annual  
 modulation

- Anisotropy in the angular distribution of nuclear recoils;
- **No background can mimic it.**

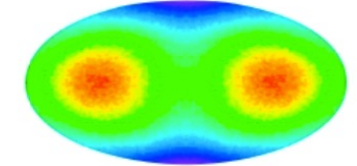


A. M. Green et. al, Astropart. Phys. 27 (2007) 14.

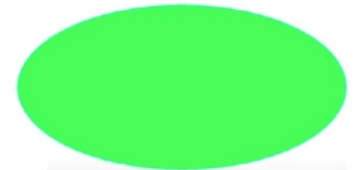
Nuclear recoil map  
 (direction+head-tail)



Nuclear recoil map  
 (only direction)

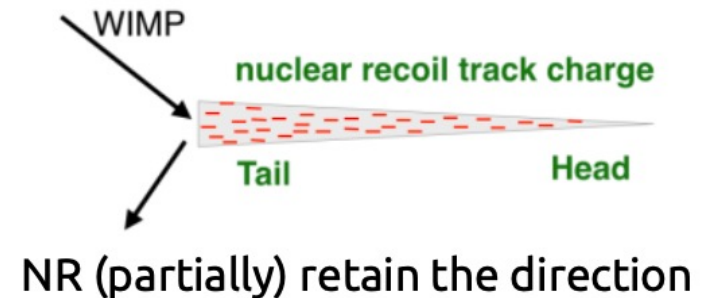


Background  
 (isotropic)



**Directional discrimination** is the only way to:

- **Unambiguously identify a DM signal**
- Searching beyond the neutrino floor
- **Moreover:**
  - Properties of the solar neutrino flux;
  - DM halo properties (DM astronomy).





# References

- **Alessandro Longo, Stefano Bianchi**, Guillermo Valdes, Nicolas Arnaud, **Wolfgang Plastino**, *Scattered light monitoring system at the Virgo interferometer: performance improvement and automation based on O3 data*. **Currently submitted to CQG.**
  - **Stefano Bianchi, Alessandro Longo**, Guillermo Valdes, Gabriela González, **Wolfgang Plastino**, *gwadaptive scattering: an automated pipeline for scattered light noise characterization*. *Classical and Quantum Gravity* 39.19 (2022): 195005.
  - **Alessandro Longo, Stefano Bianchi**, Guillermo Valdes, Nicolas Arnaud, **Wolfgang Plastino**, *Daily monitoring of scattered light noise due to microseismic variability at the Virgo interferometer*. *Classical and Quantum Gravity* 39.3 (2021), 035001.
  - **Alessandro Longo, Stefano Bianchi, Wolfgang Plastino**, Nicolas Arnaud, *Scattered light noise characterization at the Virgo interferometer with tvf-EMD adaptive algorithm*. *Class. Quantum Grav.* 37, 145011, 2020.
- 
- **Alessandro Longo, Stefano Bianchi, Wolfgang Plastino**, Kouseki Miyo, Takaaki Yokozawa, Tatsuki Washimi, and Akito Araya, *Local Hurst Exponent Computation of Data from Triaxial Seismometers Monitoring KAGRA*. *Pure and Applied Geophysics* 178.9 (2021): 3461-3470.
  - **Alessandro Longo, Stefano Bianchi, Wolfgang Plastino** et al., *Adaptive denoising of acoustic noise injections performed at Virgo Interferometer*. *Pure and Applied Geophysics* 177, 3395–3406, 2020.
  - **Alessandro Longo, Stefano Bianchi, Wolfgang Plastino** et al., *Fractal analysis of data from seismometer array monitoring Virgo Interferometer*. *Pure and Applied Geophysics* 177, 2597-2603, 2020.