# Flavour News from my Homeoffice





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**QCD@Work 2024** 

(Trani, 17-21 June, 2024)





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# Overture

#### News from the Homeoffice in Ottobrunn



# Trani Symphony



**Standard Model Predictions for** 

- Rare K and B Decays without New Physics Infection
- : Z´at Work
  - Disentangling New Physics in  $K \to \pi \nu \overline{\nu}$  and  $B \to K(K^*) \overline{\nu} \nu$  Decays

4<sup>th</sup> Movement

More Flavour News

# 1st Movement: Standard Model Predictions for Rare K and B Decays without New Physics Infection



AJB: 2209.03968

#### Problems with SM Predictions for TH "clean" Rare K and B Decays

(AJB 2209.03968)



In a global fit New Physics can infect them through CKM parameters.



Tensions in the determination of  $|V_{cb}|$  from inclusive vs exclusive tree level decays. (Lower the precision and should be presently avoided)



Hadronic uncertainties in some observables included in the fit are much larger than in many rare K and B decays. (Lower the precision and should be presently avoided)

#### **Suggested Strategy**

AJB 0303060 AJB+E.Venturini 2109.11032 " 2203.11960 AJB 2209.03968

**No New** 

**Physics!** 



Remove CKM dependence by calculating suitable ratios of branching ratios to  $\Delta M_d$ ,  $\Delta M_s$ ,  $|\epsilon_k|$ 

CKM can be fully eliminated for all rare B decays. For K decays only the dependence on β remains. (γ dependence irrelevant!!)



Very precise predictions for rare decays branching ratios independent of CKM parameters!



Rapid test of New Physics infection in the  $\Delta$ F=2 sector using  $|V_{cb}|$  -  $\gamma$  plots

BV1 + BV2 + AJB 2204.10337



Determination of CKM parameters from  $\Delta$ F=2 only.

Advantages over full global fits



 $\Delta F = 2$  sector appears to be free of NP infection: NP is not required.



The remaining observables outside the " $\Delta F = 2$  archipelago" that could be infected by NP can be predicted within the SM and the pulls can be better estimated.



 $|V_{cb}|$  and  $|V_{ub}|$  tensions can be avoided.

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# Standard Model Predictions for Rare K and B Decays without |V<sub>cb</sub>| uncertainties and New Physics Infection

but with



E. Venturini

V<sub>cb</sub>| Tension

$$|V_{cb}|_{inclusive} = (41.97 \pm 0.48) \cdot 10^{-3}$$
 Finauri & Gambino (2310.20234)  
 $|V_{cb}|_{exclusive} = (39.21 \pm 0.62) \cdot 10^{-3}$  (FLAG)  
(2022)

(see also Bordone, Gubernari, van Dyk, Jung (1912.09335) Bordone, Capdevilla, Gambino (2107.00604)

Note: Changing 
$$|\mathbf{V}_{cb}|$$
 :  $39 \cdot 10^{-3} \Rightarrow 42 \cdot 10^{-3}$   
changes  $|\mathbf{V}_{cb}|^2$ : by 16%  $(\mathbf{B}_{s,d} \rightarrow \mu^+ \mu^-, \Delta \mathbf{M}_{s,d})$   
 $|\mathbf{V}_{cb}|^3$ : by 25%  $(\mathbf{K}^+ \rightarrow \pi^+ \nu \overline{\nu}, \varepsilon_{\mathbf{K}})$   
 $|\mathbf{V}_{cb}|^4$ : by 35%  $(\mathbf{K}_{\mathbf{L}} \rightarrow \pi^0 \nu \overline{\nu}, \mathbf{K}_{\mathbf{S}} \rightarrow \mu^+ \mu^-)$ 

 $|V_{cb}|$  tension is a disaster for those who spent decades to calculate NLO and NNLO QCD Corrections to basically all important rare K and B decays.

Achieving the reduction of TH uncertainties to 1% - 2% level.

Similar disaster for Lattice QCD which for  $\Delta M_s$ ,  $\Delta M_d$ ,  $\varepsilon_K$  and weak decay constants achieved accuracy below 5%. Moreover experimental data are very precise for them.

## **Basic Strategy for Rare B and K Decays**

AJB + E. Venturini (2109.11032)



Use as basic parameters

 $\lambda$ ,  $|\mathbf{V}_{cb}|$ ,  $\beta$ ,  $\gamma$ 



Construct  $|V_{cb}|$  independent Ratios R<sub>i</sub> ( $\beta$ , $\gamma$ )



#### **16 Ratios involving**

 $|\varepsilon_{\kappa}|, \Delta M_{d}, \Delta M_{s}$ 

$$B_{s} \rightarrow \mu^{+}\mu^{-}, B_{d} \rightarrow \mu^{+}\mu^{-}$$

$$B^{+} \rightarrow K^{+}\nu\overline{\nu}, B^{0} \rightarrow K^{0^{*}}\nu\overline{\nu}$$

$$K^{+} \rightarrow \pi^{+}\nu\overline{\nu}, K_{L} \rightarrow \pi^{0}\nu\overline{\nu}, K_{s} \rightarrow \mu^{+}\mu^{-}$$



Once  $\gamma$  and  $\beta$  will be precisely measured very good test of SM

Additional ratios with  $B \rightarrow K(K^*)\mu^+\mu^-$ ,  $B_s \rightarrow \phi\mu^+\mu^-$  in 2209.03968



#### **"Critical Exponents" of Flavour Physics**

AJB + Venturini (2109.11032) (All decays TH clean)

$$\mathbf{Br}\left(\mathbf{K}^{+} \to \pi^{+} \nu \overline{\nu}\right) \sim \left|\mathbf{V}_{cb}\right|^{2.8} \left[\sin\gamma\right]^{1.4} \mathbf{Br}\left(\mathbf{K}_{L} \to \pi^{0} \nu \overline{\nu}\right) \sim \left|\mathbf{V}_{cb}\right|^{4} \left[\sin\gamma\right]^{2} \left[\sin\beta\right]^{2}$$

$$\mathbf{Br}\left(\mathbf{K}_{s} \to \mu^{+}\mu^{-}\right)_{sD} \sim \left|\mathbf{V}_{cb}\right|^{4} \left[\sin\gamma\right]^{2} \left[\sin\beta\right]^{2} \qquad \left|\epsilon_{\kappa}\right| \sim \left|\mathbf{V}_{cb}\right|^{3.4} \left[\sin\gamma\right]^{1.67} \left[\sin\beta\right]^{0.87}$$

$$\mathbf{Br} \left( \mathbf{B}_{s} \to \mu^{+} \mu^{-} \right) \sim \left| \mathbf{V}_{cb} \right|^{2} \quad \mathbf{Br} \left( \mathbf{B}_{d} \to \mu^{+} \mu^{-} \right) \sim \left| \mathbf{V}_{cb} \right|^{2} \left[ \sin \gamma \right]^{2}$$

$$\mathsf{Br}(\mathsf{B}^{\scriptscriptstyle +}\to\mathsf{K}^{\scriptscriptstyle +}\nu\overline{\nu})\sim \left|\mathsf{V}_{_{\mathsf{c}\mathsf{b}}}\right|^2 \quad \mathsf{Br}(\mathsf{B}^{\scriptscriptstyle 0}\to\mathsf{K}^{\scriptscriptstyle 0^{\scriptscriptstyle *}}\nu\overline{\nu})\sim \left|\mathsf{V}_{_{\mathsf{c}\mathsf{b}}}\right|^2$$

$$\Delta \mathbf{M}_{s} \sim \left| \mathbf{V}_{cb} \right|^{2} \quad \Delta \mathbf{M}_{d} \sim \left| \mathbf{V}_{cb} \right|^{2} \left[ \sin \gamma \right]^{2}$$

$$\bm{S}_{\psi\bm{K}_{s}}=\bm{sin2\beta}$$

### $|V_{cb}|$ Independent Ratios in the SM

AJB + E. Venturini (B-K Correlations)

$$\mathbf{R}_{1}(\boldsymbol{\beta},\boldsymbol{\gamma}) = \frac{\mathbf{Br}\left(\mathbf{K}^{+} \to \pi^{+} \nu \overline{\nu}\right)}{\left[\overline{\mathbf{B}r}\left(\mathbf{B}_{s} \to \mu^{+} \mu^{-}\right)\right]^{1.4}} = \mathbf{C}_{1}\left(\sin \gamma\right)^{1.4}\left(\mathbf{F}_{\mathbf{B}_{s}}\right)^{-2.8}$$

$$\mathbf{R}_{2}(\beta,\gamma) = \frac{\mathbf{Br}\left(\mathbf{K}^{+} \to \pi^{+}\nu\overline{\nu}\right)}{\left[\overline{\mathbf{B}r}\left(\mathbf{B}_{d} \to \mu^{+}\mu^{-}\right)\right]^{1.4}} = \mathbf{C}_{2}\left(\sin\gamma\right)^{-1.4}\left(\mathbf{F}_{\mathbf{B}_{d}}\right)^{-2.8}$$

V<sub>cb</sub>-independent correlations between K and B Decays

$$\mathbf{R}_{3}(\beta,\gamma) = \frac{\mathbf{Br}\left(\mathbf{K}_{L} \to \pi^{0}\nu\overline{\nu}\right)}{\left[\overline{\mathbf{B}r}\left(\mathbf{B}_{s} \to \mu^{+}\mu^{-}\right)\right]^{2}} = \mathbf{C}_{3}\left[\sin\beta\sin\gamma\right]^{2}\left(\mathbf{F}_{\mathbf{B}_{s}}\right)^{-4}$$

$$\mathbf{R}_{4}(\beta,\gamma) = \frac{\mathbf{Br}\left(\mathbf{K}_{L} \to \pi^{0}\nu\overline{\nu}\right)}{\left[\overline{\mathbf{Br}\left(\mathbf{B}_{d} \to \mu^{+}\mu^{-}\right)}\right]^{2}} = \mathbf{C}_{4}\left[\frac{\sin\beta}{\sin\gamma}\right]^{2}\left(\mathbf{F}_{\mathbf{B}_{d}}\right)^{-4}$$

C<sub>i</sub> = CKM independent known factors

### Important V<sub>cb</sub> – Independent Formulae

AJB + E. Venturini (2109.11032)

$$\frac{\text{Br}\left(\text{K}^{+} \rightarrow \pi^{+}\nu\overline{\nu}\right)}{\left|\epsilon_{\text{K}}\right|^{0.82}} = \left(1.31 \pm 0.05\right) \cdot 10^{-8} \left[\frac{\sin 22.2}{\sin \beta}\right]^{0.71} \left[\frac{\sin \gamma}{\sin 67^{\circ}}\right]^{0.015}$$

$$\frac{\text{Br}\left(\text{K}_{\text{L}} \rightarrow \pi^{0} \nu \overline{\nu}\right)}{\left|\epsilon_{\text{K}}\right|^{1.18}} = \left(3.87 \pm 0.06\right) \cdot 10^{-8} \left[\frac{\sin\beta}{\sin 22.2}\right]^{0.98} \left[\frac{\sin\gamma}{\sin 67^{\circ}}\right]^{0.030}$$

$$\left\{ \left| \epsilon_{\kappa} \right|_{exp}, S_{\psi \kappa_{s}}^{exp} = sin 2\beta \right\} \Rightarrow \begin{cases} Most \ accurate \\ Predictions \ to \\ date \end{cases} \right\}$$

Note: practically γ-independent

> 14 additional ratios in 2109.11032

Important reduction of TH uncertainties in  $\epsilon_{K}$  (Brod, Gorbahn, Stamou, 1911.06822)



$$\begin{array}{c} \textbf{Standard Model} \qquad (2024) \\ \textbf{SM:} \\ \hline \textbf{Br}(K^+ \to \pi^+ \nu \bar{\nu}) = (8.6 \pm 0.4) \cdot 10^{-11} \\ \textbf{Br}(K_L \to \pi^0 \nu \bar{\nu}) = (2.94 \pm 0.15) \cdot 10^{-11} \\ \textbf{AJB} + \textbf{Venturini} (2109.11032) \\ \hline \textbf{AJB} + \textbf{Venturini} (2109.11032) \\ \hline \textbf{Relativ to} \\ 1503.02693 \\ \textbf{(AJB, Buttazzo, Girrbach-Noe, Knegiens)} \\ \hline \textbf{Reduction of uncertainties:} \\ \textbf{In } K^+ \to \pi^+ \nu \bar{\nu} \textbf{ by factor } 2.4 \\ \textbf{K}_L \to \pi^0 \nu \bar{\nu} \textbf{ by factor } 4.0 \\ \hline \textbf{News from NA62 and KOTO} \\ \hline \textbf{Br}(K^+ \to \pi^+ \nu \bar{\nu}) = (10.6 \pm 3.8) \cdot 10^{-11} \\ \textbf{Br}(K_L \to \pi^0 \nu \bar{\nu}) \leq 2.0 \cdot 10^{-9} \\ \hline \textbf{(KOTO)} \\ \hline \textbf{V}_{cb} = 42.6(4) \cdot 10^{-3} \quad |\textbf{V}_{cb}|_{inl} = 42.0(5) \cdot 10^{-3} \\ \hline \textbf{News from Nach and State of the state$$

$$\gamma = 64.6(16)^{\circ}$$
  $\gamma = 63.8(36)^{\circ}$  LHCb

$$|V_{cb}| - \gamma Plot = Rapid Test$$

#### Perfect consistency between $\Delta M_s$ , $\Delta M_d$ , $\epsilon_K$ , $S_{\psi K}$

AJB + Venturini 2203.11960



## **Positive Tests**

#### AJB + Venturini 2203.11960



Precise Lattice QCD and higher order QCD calculations are necessary to make the rapid tests reliable!

Rapid Test: cover picture of EPJC Vol. 83 number 1, January 2023



#### **SM** without uncertainties

#### **Impact of New Physics**

 $V_{cb}$  -  $\gamma$  Plot



## $R_i(\beta, \gamma)$ can now be predicted in the SM

AJB 2209.03968

$$\frac{Br(K^+ \rightarrow \pi^+ \nu \overline{\nu})}{[\overline{Br}(B_s \rightarrow \mu^+ \mu^-)]^{1.4}} = 53.69 \pm 2.75$$

$$\frac{Br(K^+ \to \pi^+ \nu \overline{\nu})}{[Br(B^+ \to K^+ \nu \overline{\nu})]^{1.4}} = (1.90 \pm 0.13) \cdot 10^{-3}$$

Many other results in 2209.03968

# 2nd Movement: Z´at Work

10 Years Anniversary (Z', 331)

#### AJB – Fulvia de Fazio – Jennifer Girrbach-Noe Collaboration







AJB

Fulvia

Jennifer

1211.1896	
1211.1237	
1303.3723	
1311.6729	
1404.3824	
1405.3850	



#### **Peculiar Pattern of Flavour Data**



 $\Rightarrow$  Impact on  $\Delta M_{s, S_{\psi\varphi}, \dots}$ 

#### Which NP scenario can reproduce this pattern ?

 $\epsilon_{\rm K}, \epsilon'/\epsilon, \Delta M_{\rm K}, {\rm K}^+ \rightarrow \pi^+ \nu \overline{\nu}, {\rm K}_{\rm L} \rightarrow \pi^0 \nu \overline{\nu}$ 

New heavy gauge boson Z':  $\Delta_L^{sd}(Z') = \left|\Delta_L^{sd}(Z')\right| e^{i\phi}$ 

$$\begin{split} & \epsilon_{K}^{NP} \backsim Im \left( \Delta_{L}^{sd}(Z') \right)^{2} \backsim \left[ Re \Delta_{L}^{sd}(Z') \right] \left[ Im \Delta_{L}^{sd}(Z') \right] \\ & (\epsilon'/\epsilon)^{NP} \backsim Im \Delta_{L}^{sd}(Z') \\ & \Delta M_{K}^{NP} \backsim \left( Re \Delta_{L}^{sd}(Z') \right)^{2} - \left( Im \Delta_{L}^{sd}(Z') \right)^{2} \qquad \left( K^{0} - \overline{K}^{0} \right) \end{split}$$

With  $Re\Delta_L^{sd}(Z') \ll Im\Delta_L^{sd}(Z')$ 

(Imaginary coupling)

 $\begin{array}{ll} \epsilon_{K}^{NP} \simeq 0 & (\epsilon'/\epsilon)^{NP} \text{ can be enhanced} \\ \Delta M_{K} \text{ can be suppressed + Interesting implications} & Aebischer \\ \text{AJB} \\ \text{(possibly required by} & \text{for } K \rightarrow \pi \nu \overline{\nu} \\ \text{Lattice QCD} & & & & & & & & & \\ \end{array}$ 



Based on the insights from Monika Blanke (0904.1545)

Monika Blanke

#### Kaon Physics without New Physics in $\epsilon_{K}$

$$R_{\nu\bar{\nu}}^{+} = \frac{\mathcal{B}(K^{+} \to \pi^{+}\nu\bar{\nu})}{\mathcal{B}(K^{+} \to \pi^{+}\nu\bar{\nu})_{SM}}, \quad R_{\nu\bar{\nu}}^{0} = \frac{\mathcal{B}(K_{L} \to \pi^{0}\nu\bar{\nu})}{\mathcal{B}(K_{L} \to \pi^{0}\nu\bar{\nu})_{SM}},$$
$$R_{\mu^{+}\mu^{-}}^{S} = \frac{\mathcal{B}(K_{S} \to \mu^{+}\mu^{-})_{SD}}{\mathcal{B}(K_{S} \to \mu^{+}\mu^{-})_{SM}^{SD}}, \quad R_{\pi\ell^{+}\ell^{-}}^{0} = \frac{\mathcal{B}(K_{L} \to \pi^{0}\ell^{+}\ell^{-})}{\mathcal{B}(K_{L} \to \pi^{0}\ell^{+}\ell^{-})_{SM}},$$
$$R_{\Delta M_{K}} = \frac{\Delta M_{K}^{BSM}}{\Delta M_{K}^{exp}}, \quad \Delta \left(\frac{\varepsilon'}{\varepsilon}\right) = \kappa_{\varepsilon'} \cdot 10^{-3}, \quad \Delta(\varepsilon_{K}) = \kappa_{\varepsilon} \cdot 10^{-3}$$



#### **B** Physics without NP in Quark Mixing

Fine tuning in 
$$\Delta M_q$$
 q=d,s suppression factor  
 $M_{12}(Z') \sim \left[ 1 + \left( \frac{\Delta_R^{bq}(Z')}{\Delta_L^{bq}(Z')} \right)^2 + 2K_{bq} \frac{\Delta_R^{bq}}{\Delta_L^{bq}} \right] \frac{\Delta_L^{bq}(Z')}{M_{Z'}^2}$   
 $K_{bq} = \frac{\langle \widehat{Q}_1^{LR}(M_{Z'}) \rangle^{bq}}{\langle \widehat{Q}_1^{VLL}(M_{Z'}) \rangle^{bq}} \approx -5$   $\Delta_R^{bq}(Z') \approx 0.1 \Delta_L^{bq}(Z')$ 

AJB, De Fazio, Girrbach-Noe 1404.3824 AJB, Buttazzo, Girrbach-Noe 1408.0728 Crivellin, Hofer, Matias, Nierste, Pokorski, Rosiek 1504.07928

## Strong Suppression of Z' to $B_s - \overline{B}_s$ Mixing

**Requires** 

(+)

$$\Delta_{\mathbf{R}}^{\mathbf{bs}}(\mathbf{Z}') \approx \mathbf{0.1} \Delta_{\mathbf{L}}^{\mathbf{bs}}(\mathbf{Z}')$$

Non-negligible RH couplings

#### **Implications for rare B-Decays**

Suppression of  $B^+ \rightarrow K^+ \mu^+ \mu^-$ ,  $B_s \rightarrow \phi \mu^+ \mu^-$ ,  $B \rightarrow K^* \mu^+ \mu^-$ 

Enhancement of  $B^+ \rightarrow K^+ \nu \overline{\nu}$ ,  $B^0 \rightarrow K^0 \nu \overline{\nu}$  up to 20%

AJB +



**Peter Stangl** 

(2407.xxx)

**News from Belle II** 

(2311.14647)

\*)

 $Br(B^+ \to K^+ \nu \bar{\nu}) = (13 \pm 4) \cdot 10^{-6}$  $Br(B^+ \to K^+ \nu \bar{\nu})_{SM} = (4.92 \pm 0.30) \cdot 10^{-6}$ 

AJB + Stangl (2024)

News from CERN (LHCb, CMS, ATLAS)

$$\begin{split} \overline{Br}(B_s \to \mu^+ \mu^-) &= (3.45 \pm 0.29) \cdot 10^{-9} \\ \overline{Br}(B_s \to \mu^+ \mu^-)_{SM} &= (3.78 \pm 0.12) \cdot 10^{-9} \\ \end{split} \label{eq:Br}$$

\*) Many analyses: Bause et al. (2309.00075) Becirevic et al. (2301.06990, 2309.02246) Dreiner et al. (2309.03727) He et al. (2309.12741)



# **3rd Movement:**

# Distangling New Physics in $K \to \pi \nu \overline{\nu}$ and $B \to K(K^*) \overline{\nu} \nu$ Decays

M. Mojahed

AJB +



J. Harz

2405.06742

35 Trani0624

Goal: Disentangling different New Physics contributions to the rare decays  $K \rightarrow \pi + E$  and  $B \rightarrow K(K^*) + E$ through kinematic distributions in the missing energy E

Step 1: WET with active or sterile neutrinos including Lepton Number violating operators with scalar and tensor currents

Step 2: Dark WET: new invisible particles in the final state: two dark scalars, two dark fermions, two dark vectors

Main Results

A. Vector, scalar and tensor quark currents can be uniquely determined from experimental data of kinematic distributions

- B. Measurements of kinematic distributions make it possible to disentangle the contributions of WET operators from most of the dark-sector operators
- C. Sum Rules for vector currents in WET are also satisfied in some new dark-physics scenarios that mimic WET







# 4th Movement More Flavour News

# **Dual QCD Approach for Weak Decays**

Successful low energy approximation of QCD for  $K \rightarrow \pi\pi$  K<sup>0</sup>-K<sup>0</sup> mixing (Large N framework)



W. Bardeen





AJB





J.-M. Gérard



1986

2024





QCD dynamics dominate this rule but New Physics could still contribute AJB F. de Fazio J. Girrbach-Noe (1404.3824)



Reviews AJB: 2101.00020, 2203.12632 2307.15737 **Good News on**  $\varepsilon'/\varepsilon$ 

 $\epsilon'/\epsilon = QCD$  Penguins – Electroweak Penguin

$$\left(\frac{\varepsilon}{\varepsilon}\right)_{SM}^{EWP} = -(7\pm1)\cdot10^{-4} \quad (RBC - UKQCD \text{ and } DQCD) \qquad Perfect \\ Agreement!$$

Chiral Pert Th:  $\approx$  (-3.5 ± 2.0)  $\cdot$  10<sup>-4</sup>

**Disagreements on QCD Penguin contribution.** 

#### Main Activities in the Homeoffice in Ottobrunn



#### **NLO QCD in WET and SMEFT** (in Homeoffice)

$\textbf{WET} \qquad \mu \leq EW$	SU(3) <sub>c</sub> ⊗ U(1) <sub>QED</sub> SM + New Physics Operators
$\Delta F = 1$ (Non-Leptonic)	J. Aebischer, C. Bobeth, AJB, J. Kumar, M. Misiak (2107.10262) (2107.12391) (ε΄/ε)
$\Delta F = 2$ (Non-Leptonic)	J. Aebischer, C. Bobeth, AJB, J. Kumar (2009.07276)
$\begin{array}{ c c } \textbf{SMEFT} & \mu \geq EW \\ \Delta \textbf{F} = \textbf{2} & \textbf{(Non-Leptonic)} \end{array}$	J. Aebischer, AJB, J. Kumar (2203.11224) (2202.01225)



J. Aebischer



AJB



J. Kumar



**Christoph Bobeth** 



M. Misiak



Messages to take to your Homeoffice

 $V_{cb}$  – independent ratios and  $V_{cb}$  -  $\gamma$  plots will play important roles in the search for New Physics



can reveal NP easier than  $B \rightarrow K\mu^+\mu^-, B \rightarrow K^*\mu^+\mu^-$ (smaller long-distance uncertainties)



It is crucial that several lattice QCD groups calculate  $\Delta M_d$ ,  $\Delta M_s$ ,  $\epsilon'/\epsilon$ ,  $\Delta I = \frac{1}{2}$  rule with 2 + 1 + 1 flavours

**Coming Years** : Flavour Precision Era

LHC Upgrade E = 14 TeV (CERN)



 $K^+ 
ightarrow \pi^+ 
u \overline{
u} ig( 10^{-10} ig)$  (CERN)  $K_L \rightarrow \pi^0 \nu \tilde{\nu} (3 \cdot 10^{-11})$  J-PARC (Japan)



# $\begin{array}{l} 2024\mathchar`-2046: Expedition \\ Attouniverse \rightarrow Zeptouniverse \\ 10^{\mathchar`-18}m \mathchar`-10^{\mathchar`-21}m \end{array}$

# Hopefully meeting Z<sup>´</sup>, Leptoquarks, Vector-Like Quarks and Leptons

#### Zeptouniverse Guide

Published July 2020



Exciting Years !

#### Andrzej J. Buras

## GAUGE THEORY OF Weak Decays

The Standard Model and the Expedition to New Physics Summits



739 pages 1350 references

> Cambridge University Press

#### Flavour Physics (2024-)







#### Flavour Physics (2024-)

S STREET

#### Zeptouniverse

**New Physics Summits** 

SMEFT Energy gap

SM

Crevasses

**Allan Buras** 

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#### Flavour Physics (2024-)

#### Zeptouniverse

**New Physics Summits** 

SMEFT Energy gap

Crevasses

SM

# Thank You !

**Allan Buras** 

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# Backup

## Footprints of Majorana Neutrinos in Rare K and B Decays

AJB + Julia Harz

All existing calculations of  $K^+ \rightarrow \pi^+ \nu \overline{\nu}$  and  $K_{L} \rightarrow \pi^0 \nu \overline{\nu}$ assumed until recently that neutrinos are of Dirac type.

What if neutrinos are Majorana neutrinos? First pioneering studies:



J. Harz

1912.10433	T. Li, XD. Ma, M. A. Schmidt
2009.04494	F. Deppisch, K. Fridell, J. Harz

## Main Messages from these Studies

1.



Difference between LNV and LNC seen in s-distributions, s = the invariant mass<sup>2</sup> of  $\nu\overline{\nu}$ 



Scale 
$$\Lambda_{NP}^{LNV} \approx 20 \text{ TeV}$$
 can be probed



All neutrino generations involved as opposed to neutrinoless double beta decay

#### Main Goals of AJB – JH Collaboration

AJB + Julia Harz



Closer look at the impact of Majorana neutrinos on the  $K^+ \rightarrow \pi^+ \nu \overline{\nu}$  -  $K_{L} \rightarrow \pi^0 \nu \overline{\nu}$  plane



Generalization to 
$$B \to Kv\overline{\nu}$$
,  $B \to K^*v\overline{\nu}$ ,  $B \to Xv\overline{\nu}$ 



Efficient strategies that would allow NA62, KOTO and Belle II to find possible footprints of Majorana neutrinos in their data.



Strategies valid in the presence of right-handed currents, LFUV and LFV

$$\Delta C_{v} = \left| C_{v}^{NP} \right| e^{i\phi_{v}} \qquad C_{s} = \left| C_{s} \right| e^{i\phi_{s}}$$

#### **Present Anomalies**

(2024)



## **New Particles behind Anomalies**

**Top candidates** 

2

Review: Capdevilla, Crivellin, Matias 2309.01311

Z' boson

: heavy neutral gauge boson (Spin 1)

Leptoquarks

Spin 0 or Spin 1 (provide interactions between quarks and leptons)

Dinosaurs of Flavour Physics?



Vector-like quarks :

Left and right components transform identically under SU(2)<sub>L</sub>



## **SM Relation for** $\Delta M_s$ , $\Delta M_d$ , $|\varepsilon_K|$ , $\beta$

AJB: 2209.03968

$$R \equiv \frac{|\epsilon_{\rm K}|^{1.18}}{\Delta M_d \Delta M_s} = (8.22 \pm 0.18) \cdot 10^{-5} \left(\frac{\sin\beta}{\sin 22.2^{\circ}}\right)^{1.027} {\rm K \ ps^2}$$



$$R_{exp} = (8.26 \pm 0.06) \cdot 10^{-5} ps^2$$
  $K = 1.0$ 

$$\frac{\operatorname{Br}(\mathsf{K}^{+} \to \mu^{+} \nu \overline{\nu})_{SM} \operatorname{and} \operatorname{Br}(\mathsf{K}_{L} \to \pi^{0} \nu \overline{\nu})_{SM}}{AJB + E. \operatorname{Venturini} (2109.11032)}$$

$$\overset{\mathsf{A} JB + E. \operatorname{Venturini} (2109.11032)}{\overset{\mathsf{A} JB^{+}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}}{\overset{\mathsf{A} JB^{+}}}}\overset{\mathsf{A} JB^{+}}} \overset{\mathsf{A} JB^{+}}}}{\overset{J} JB^{+}}}\overset{\mathsf{A} JB^{+}}} \overset{\mathsf{A} JB^{+}}}{\overset{J} JB^{+}}} \overset{\mathsf{A} JB^{+}}} \overset{\mathsf{A} JB^{+}}} \overset{\mathsf{A} JB^{+$$

\_\_\_\_

7

Br(

 $1. \times 10^{-10}$ 

 $9. \times 10^{-11}$ 

 $\overset{(\widehat{A}, \widehat{A})}{\underset{+}{\uparrow}} 8. \times 10^{-11} \\ \overset{(\widehat{A}, \widehat{A})}{\underset{+}{\uparrow}} 8 \\ 7. \times 10^{-11}$ 

 $6. \times 10^{-11}$ 

## The Story of $B_s \rightarrow \mu^+\mu^-$ continues

 $\overline{Br}(B_s \to \mu^+ \mu^-) = (3.78 \pm 0.12) \cdot 10^{-9}$ 

 $\overline{Br}(B_s \to \mu^+ \mu^-) = (3.45 \pm 0.29) \cdot 10^{-9}$ 

 $\overline{Br}(B_s \to \mu^+ \mu^-) = (3.47 \pm 0.14) \cdot 10^{-9}$ 

AJB + Venturini 2203.11960

HFLAV (CMS, LHCb, ATLAS)

(SM)

UTfitter 2212.1051

Theory SM

: Buchalla + AJB (1993, 1998) Misiak + Urban (1998) NLO QCD

Bobeth, Gorbahn, Stamou (2013) NLO EW Hermann, Misiak, Steinhauser (2013) NNLO QCD Beneke, Bobeth, Szafron (2017, 2019) QED

#### Searching for Majorana Footprints through LNC Sum Rules

$$\begin{array}{l} \mbox{LNC Sum Rules} \quad \mbox{AJB, J. Girrbach-Noe, C. Niehoff, D. Straub} \\ (1409.4557) \quad & \left(r_1^{LNC} = r_2^{LNC} = 1\right) \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_{K^*}}{\left(\kappa_\eta + 2\right) R_{K^*}} \Bigg] r_1^{LNV} \quad \box{$r_1^{LNV} \neq 1$} \\ F_L = F_L^{SM} \Bigg[ \frac{\left(\kappa_\eta - 2\right) R_K + 4 R_K +$$

fraction

$$\begin{split} \mathsf{R}_{\mathsf{K}} &= \frac{\mathsf{Br}\big(\mathsf{B} \to \mathsf{K}\nu\overline{\nu}\big)}{\mathsf{B}_{\mathsf{SM}}\big(\mathsf{B} \to \mathsf{K}\nu\overline{\nu}\big)} \qquad \mathsf{R}_{\mathsf{K}^{\star}} = \frac{\mathsf{Br}\big(\mathsf{B} \to \mathsf{K}^{\star}\nu\overline{\nu}\big)}{\mathsf{Br}\big(\mathsf{B} \to \mathsf{K}^{\star}\nu\overline{\nu}\big)_{\mathsf{SM}}} \\ \kappa_{\eta} &= 1.33 \pm 0.05 \text{ (formfactor)} \qquad \mathsf{F}_{\mathsf{L}}^{\mathsf{SM}} = 0.49 \pm 0.04 \\ &\qquad \mathsf{K}^{\star} \text{ longitudinal polarization} \end{split}$$

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#### **Vector Z' Couplings to Leptons**



 $\begin{array}{l} \Delta C_9(\Lambda_{NP}) \neq \mathbf{0} \\ \Delta C_{10}(\Lambda_{NP}) = \mathbf{0} \end{array}$ 

AJB + Stangl (2407.xxx)

#### Strong Suppression of Z' to $\Delta F = 2$ Process

$$\label{eq:K-System} \hline Re\Delta_L^{sd}(Z') \ll Im\Delta_L^{sd}(Z')$$

Negligible RH couplings

$$\textbf{B}_{s,d}\textbf{-Systems} \qquad \Delta_R^{bq}(Z') \approx 0.1 \, \Delta_L^{bq}(Z')$$

Non-negligible RH couplings