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International Workshop on QCD Theory and Experiment

The effect of electric and chiral magnetic conductivities on azimuthally fluctuating electromagnetic fields and observables in isobar collisions.

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Date: 20th-June-2024





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Outline

- Introduction & Motivation
- EM fields with and without medium feedback
- Numerical Results
- Summary

SIDDIQUE IRFAN. DOI: 10.1103/PhysRevC.109.034905 (2024)

Magnetic field in HICs



Extremely Strong magnetic fields

- x-axis impact parameter z-axis beam direction
- **B** Fields perpendicular to reaction plane
- Influencing the dynamics of QGP
- CME etc
- Time evolution and spatial distribution etc.



0.6	Earth		
100	Hand held magnet		
8.3 * 10 ⁴	Superconducting Magnet LHC		
4.5 * 10 ⁵	Strongest steady magnetic field		
10 ¹²	Neutron Stars		
10 ¹⁵	Magnetars		
10 ¹⁸	PHIC at RHIC		
10 ¹⁹	PHIC at LHC		
Units of Gauss			

- W.-T. Deng and X.-G. Huang, Phys. Rev. C 85, 044907 (2012), 1201.5108.
- J. Bloczynski, X.-G. Huang, X. Zhang, and J. Liao, Phys. Lett. B 718, 1529 (2013), 1209.6594.
- ✤ K. Hattori and X.-G. Huang, Nucl. Sci. Tech. 28, 26 (2017), 1609.00747.
- ✤ K. Tuchin, Phys. Rev. C 88, 024911 (2013), 1305.5806.

Electromagnetic Fields in HICs

- Without medium feedback:
 - Use of Event generator/ transport model
 - Use Lienard-Wiechert potential

$$E = \frac{e}{4\pi} \sum_{n} \frac{(1 - v_n^2) R_n}{(R_n^2 - (R_n \times v_n)^2)^{3/2}}$$
$$B = \frac{e}{4\pi} \sum_{n} \frac{(1 - v_n^2) (v_n \times R_n)}{(R_n^2 - (R_n \times v_n)^2)^{3/2}}$$

Here $\mathbf{R}_n = \mathbf{x} - \mathbf{x}_n$ is the relative position vector between the field point \mathbf{x} and the source point \mathbf{x}_n

- Over estimate or underestimate
- W.-T. Deng and X.-G. Huang, Phys. Rev. C 85, 044907 (2012), 1201.5108.
 J. Bloczynski, X.-G. Huang, X. Zhang, and J. Liao, Phys. Lett. B 718, 1529 (2013).
- J. BIOCZYNSKI, X.-G. Huang, X. Zhang, and J. Liao, Phys. Lett. B 718, 1529 (201.
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- K. Tuchin, Phys. Rev. C 88, 024911 (2013), 1305.5806.
- LI H, Li Sheng X, Wang Q. DOI:10.1103/physrevc.94.044903
- Irfan Siddique et. al Phys. Rev. C 105, 054909 (2022)
- Irfan Siddique et. al., Phys. Rev. C 104, 034907(2021)

$$\therefore \Gamma(a,z) = \int_{z}^{\infty} dt \ t^{a-1} exp(-t)$$

• With medium feedback ($\sigma \& \sigma_{\gamma}$):



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Isobar collisions $(Ru_{44}^{96} + Ru_{44}^{96} \& Zr_{40}^{96} + Zr_{40}^{96})$

The difference in number of protons can generate different magnitudes of |Deformed Nuclei Case electromagnetic fields and related induced effects, but the same mass number in two isobar systems can generate the same background effect.

Woods-Saxon distribution for Ru and Zr

$$\rho = \frac{\rho_0}{1 + exp\left[\frac{r - R(1 + \beta_2 Y_{20} + \beta_4 Y_{40})}{a}\right]}$$

 β_i deformation parameter

 $Y_i(\theta)$ spherical harmonic functions

f surface thickness parameter

	R_0	a	eta_2	
Ru	5.085	0.46	0.158	
Zr	5.020	0.46	0.08	
Halotype Nuclei Case				
Ru, n	5.085	0.523	0	
Ru, p	5.085	0.523	0	
Zr, n	5.021	0.592	0	
Zr, p	5.021	0.523	0	

Woods-Saxon parameters for Ru and Zr

MCGlauber model

- ◆ Q. Y. Shou *et al.*, arXiv:1409.8375.
- H.-j. Xu, H. Li, X. Wang, C. Shen, and F. Wang, arXiv:2103.05595.
- X.-L. Zhao and G.-L. Ma, arXiv:2203.15214

B. Pritychenko, M. Birch, B. Singh, and M. Horoi, arXiv:1312.5975.

Time-Evolution



Spatial Distribution



Effects on correlation

According to the expectations from CME, the difference between the correlation of opposite charge pairs and same charge pairs is expected to be directly proportional to the strength of the squared magnetic field and $\cos 2(\Psi_B - \Psi_2)$,

$$\Delta \gamma = \gamma_{opposite} - \gamma_{same} \propto (eB)^2 \cos 2(\Psi_B - \Psi_2)$$

Quantitative contribution to B-induced effect

Where Ψ_B represents the azimuthal angle of the magnetic field and Ψ_2 represents the second harmonic participant plane

$$\Psi_n = \frac{\operatorname{atan2}(\langle r_p^2 \sin(n\phi_p) \rangle, \langle r_p^2 \cos(n\phi_p) \rangle + \pi)}{n}$$
$$X_c = 2\frac{c^{Ru} - c^{Zr}}{c^{Ru} + c^{Zr}} \text{ , Relative Ratios}$$

For similarity or dissimilarity

- J. Bloczynski, X.-G. Huang, X. Zhang, and J. Liao, Phys. Lett. B 718, 1529 (2013), arXiv:1209.6594.
- J. Bloczynski, X.-G. Huang, X. Zhang, and J. Liao, Nucl. Phys. A 939, 85 (2015), arXiv:1311.5451.
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- X.-L. Zhao, G.-L. Ma, and Y.-G. Ma, Phys. Rev. C 99, 034903 (2019), arXiv:1901.04151.

Correlations between magnetic field and participant plane





Correlations

3.
$$\langle (\mathbf{eB})^2 \mathbf{cos2}(\Psi_B - \Psi_2) \rangle_t$$

$$\langle \mathbf{G} \rangle_t(x) \equiv \frac{\int \mathbf{G}(t,x) dt}{\int dt} \quad \therefore \mathbf{G} \equiv (\mathbf{e}\mathbf{F})^2 \cos 2(\boldsymbol{\Psi}_F - \boldsymbol{\Psi}_2)$$



Time-averaged correlation

EM fields behavior varies with respect to both time and space, so their impact on physical observables should be at average level in lifespan of quark and nuclear matter. To quantify the average effects of correlators on physical observables time-averaged correlation can be defined

$$\langle \mathbf{G} \rangle_t(x) \equiv \frac{\sum_i \mathbf{G}(t_i, x) \Delta t_i}{\sum_i \Delta t_i}$$



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Summary

Effects of the electric (σ) and chiral magnetic (σ) conductivities on the space and time evolution of the electromagnetic fields. Partially asymmetric spatial distribution as compared to zeroconductivity system. Decay in the presence of conductivities is much slower as compared to zero conductivity system. Studied effect on magnetic field related correlations which reflect the importance of taking into account medium feedback.



