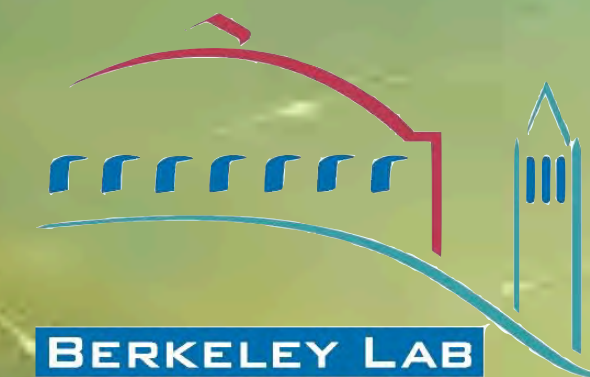


Let There be Light

The development of “Hybrid”
Cherenkov / scintillation
neutrino detectors

Next-generation neutrino detection
with EOS & THEIA



*Gabriel D. Orebi Gann
UC Berkeley & LBNL
Pisa Meeting 2024*

THEIA

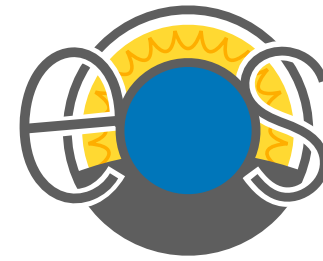


- **Hybrid Cherenkov/ scintillation detector**
- **Multi-messenger astrophysics**
- **Probe the fundamental nature of matter: CPV and Majorana ν**
- **Unique opportunity to engage a broad community in world-leading “big science”**

Disclaimer: calls out a subset of the critical team involved in this effort; all citations at the end

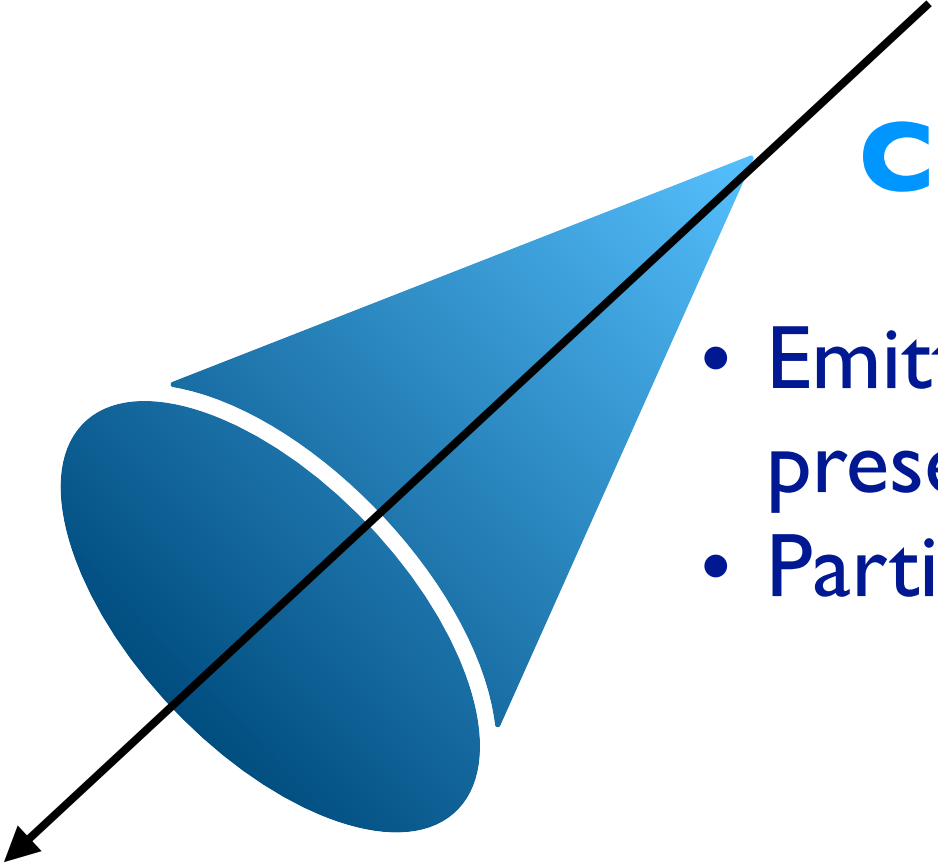


Innovative neutrino detection



Combine two well-tested methods for neutrino detection, for enhanced precision:

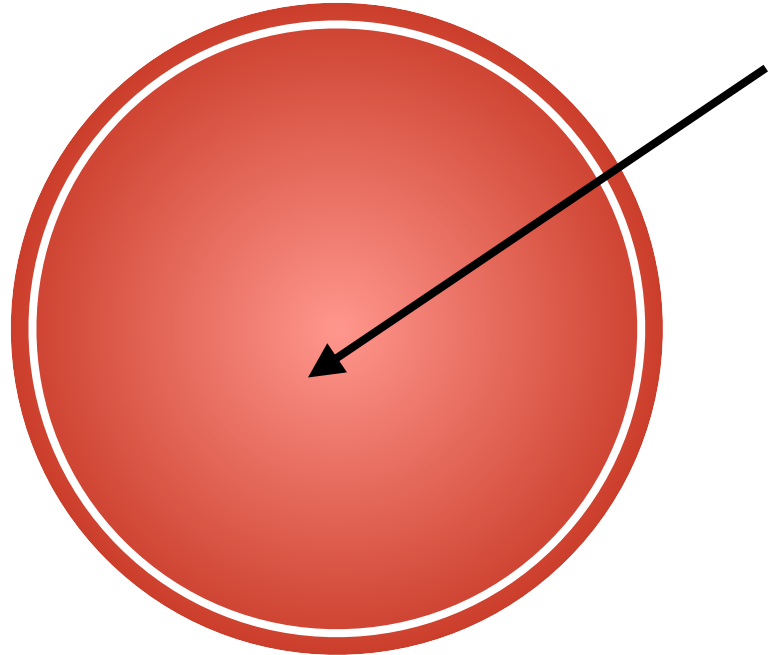
Cherenkov light



- Emitted with a unique topology that preserves directional information
- Particle-dependent threshold

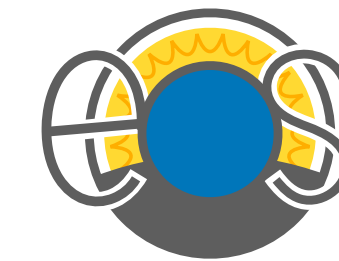
Scintillation light

- Extremely high light yield: high efficiency for detection
- Particle-dependent response offers particle identification



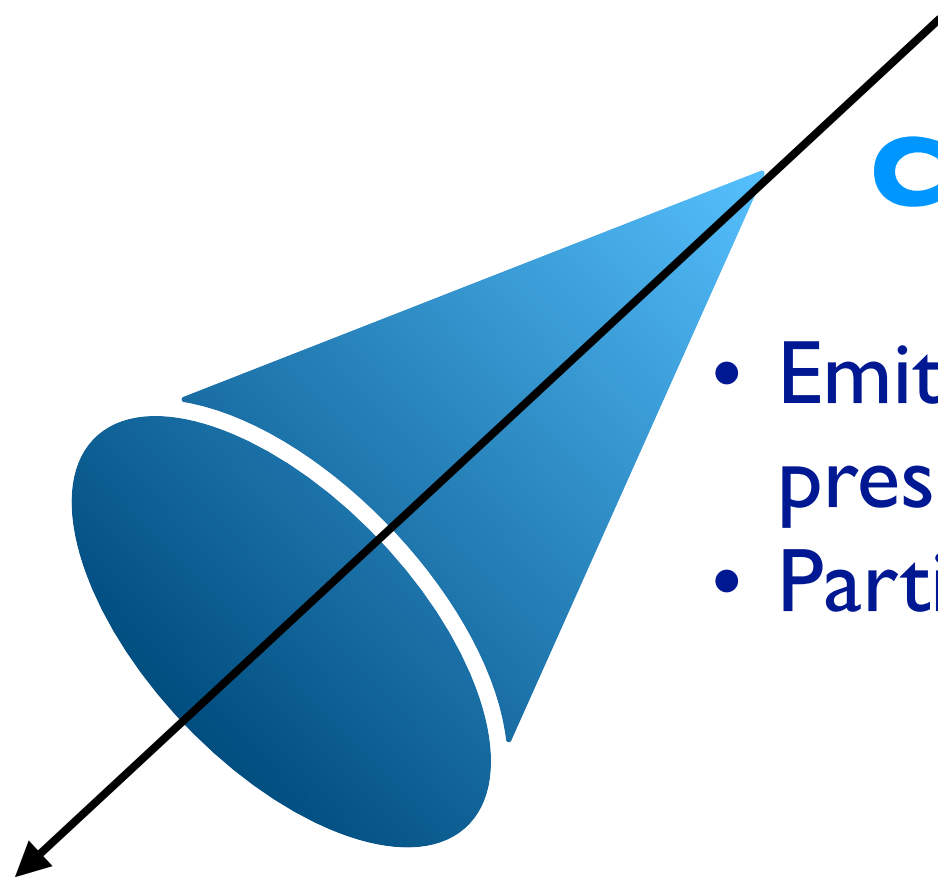


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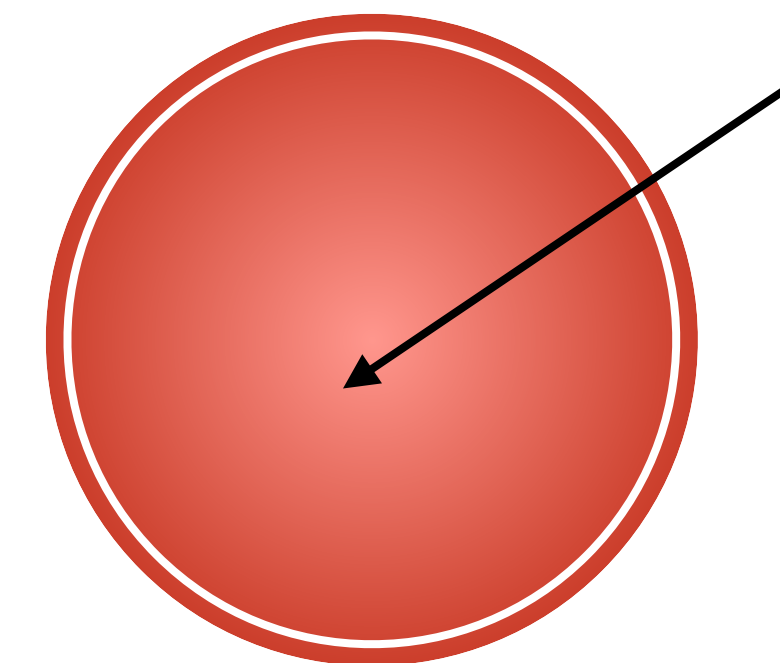
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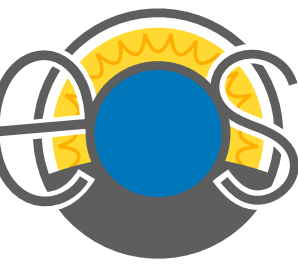


The whole is greater than the sum of the parts

*The ratio of the two signals provides additional information on the type of particle interacting:
Improved background rejection for precision measurements*

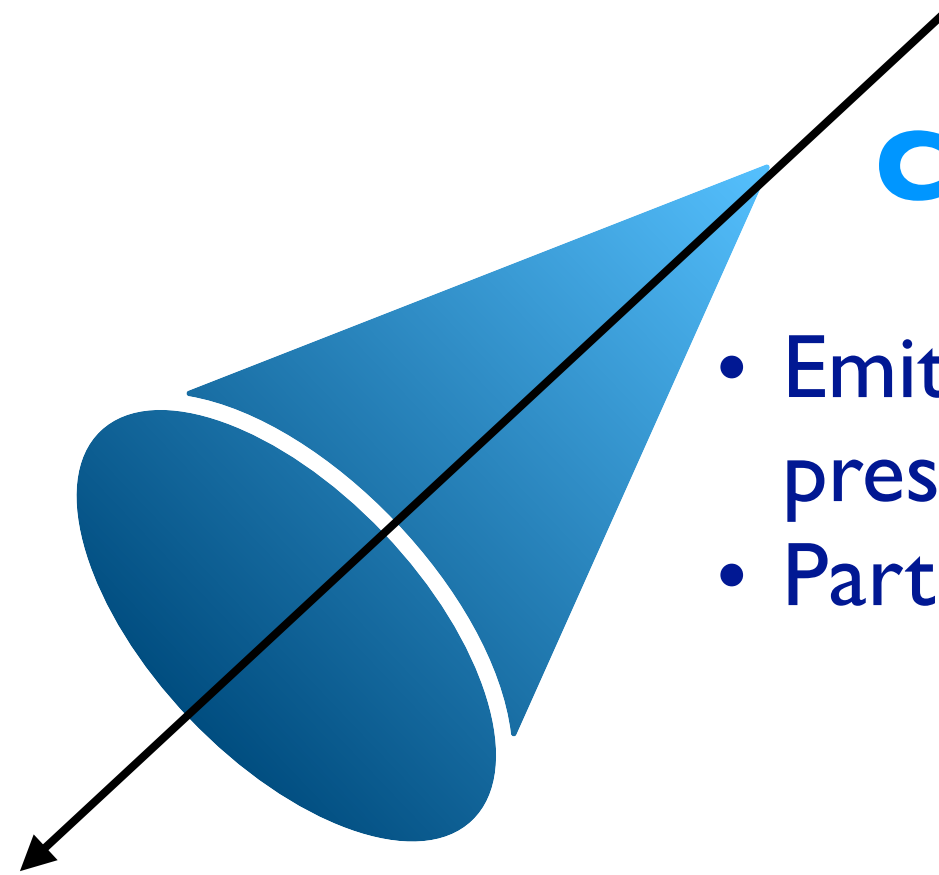


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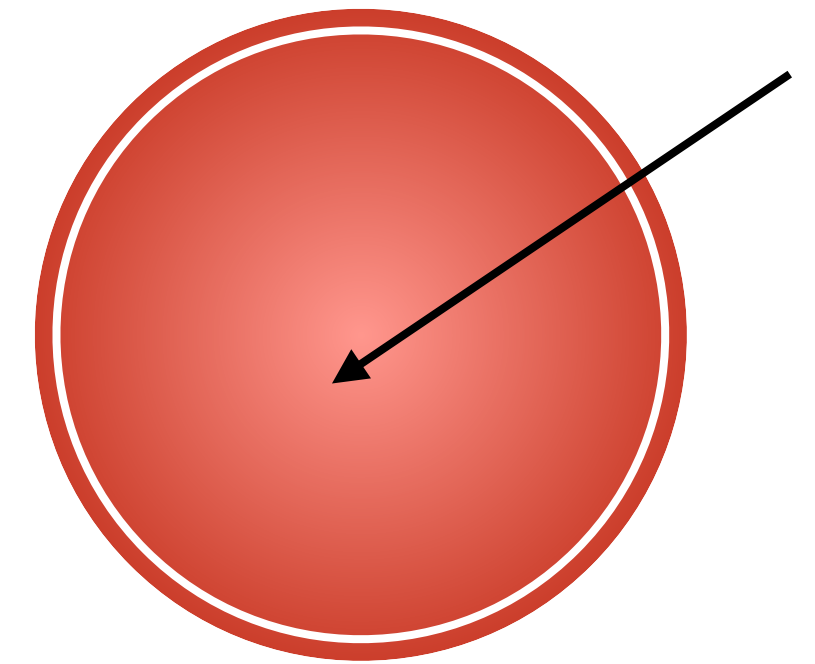
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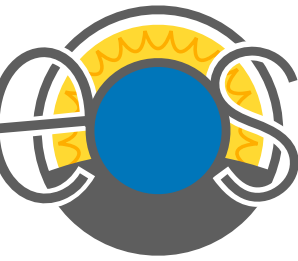
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Challenging technical goal: preserve directional Cherenkov signature against more abundant scintillation yield

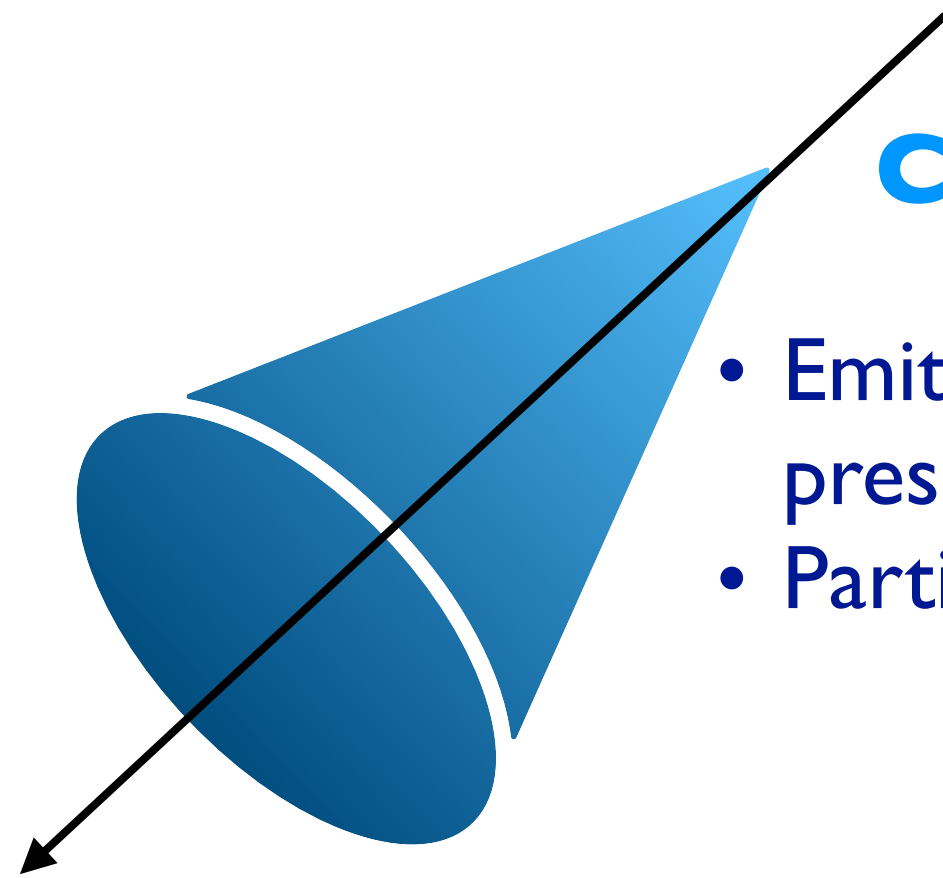


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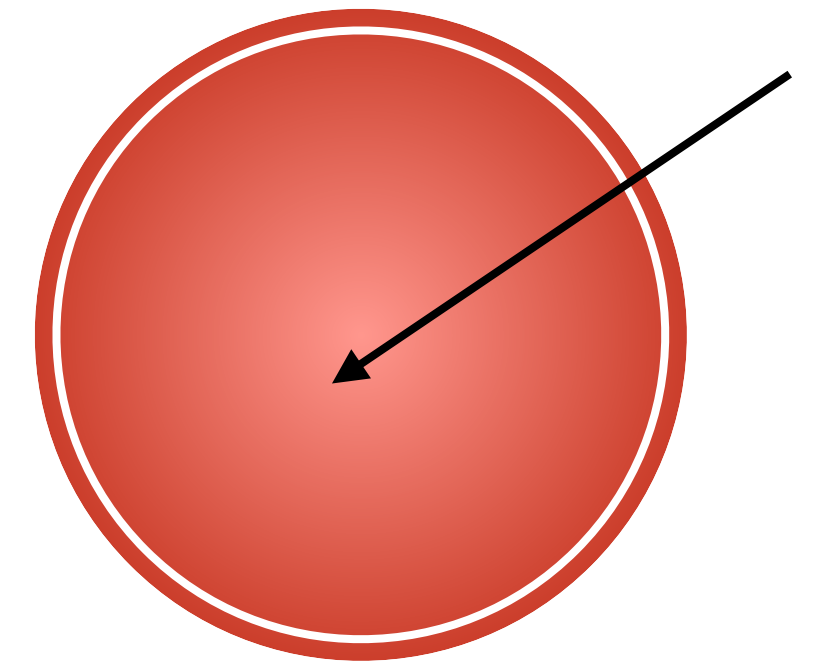
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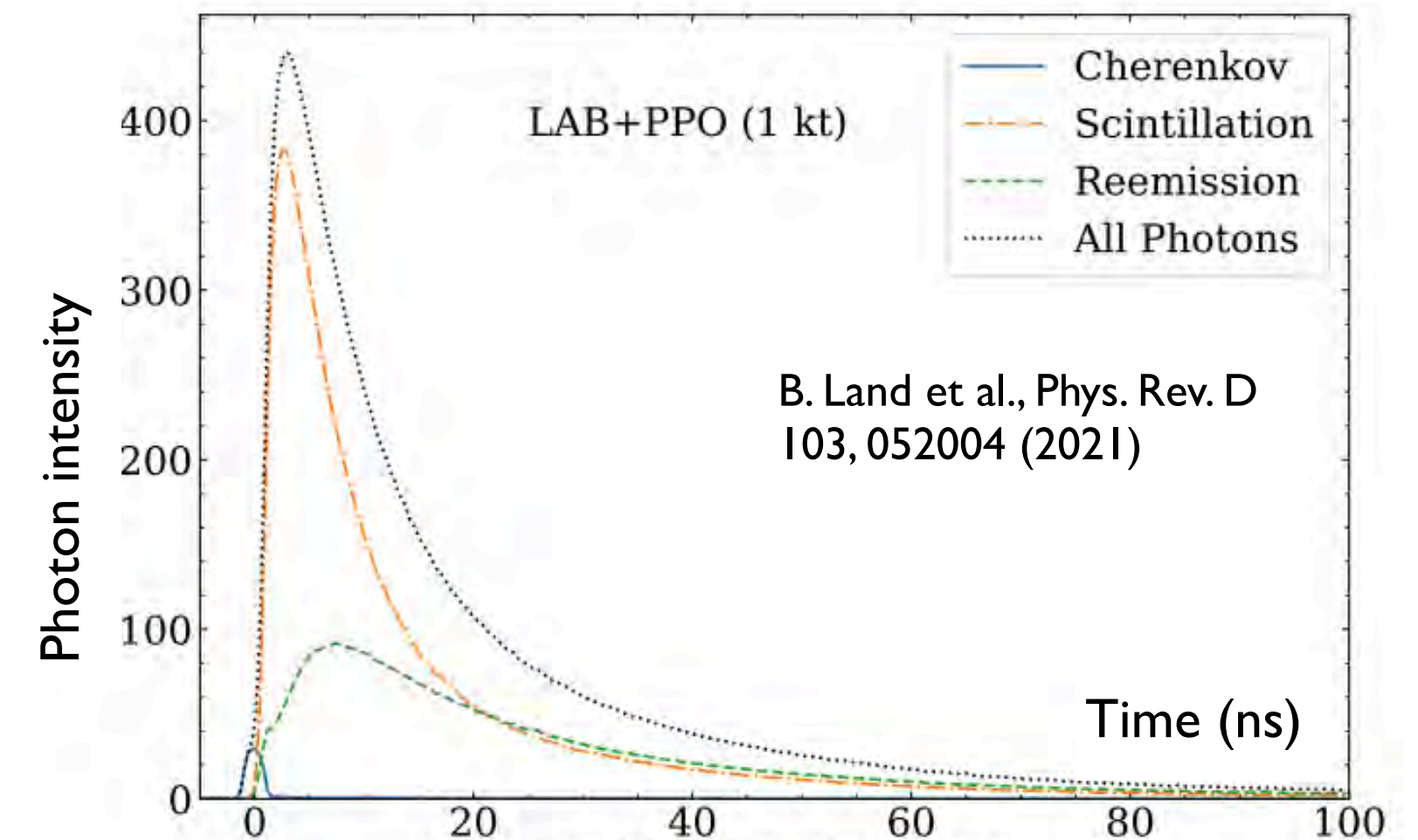
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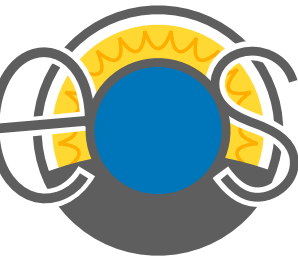
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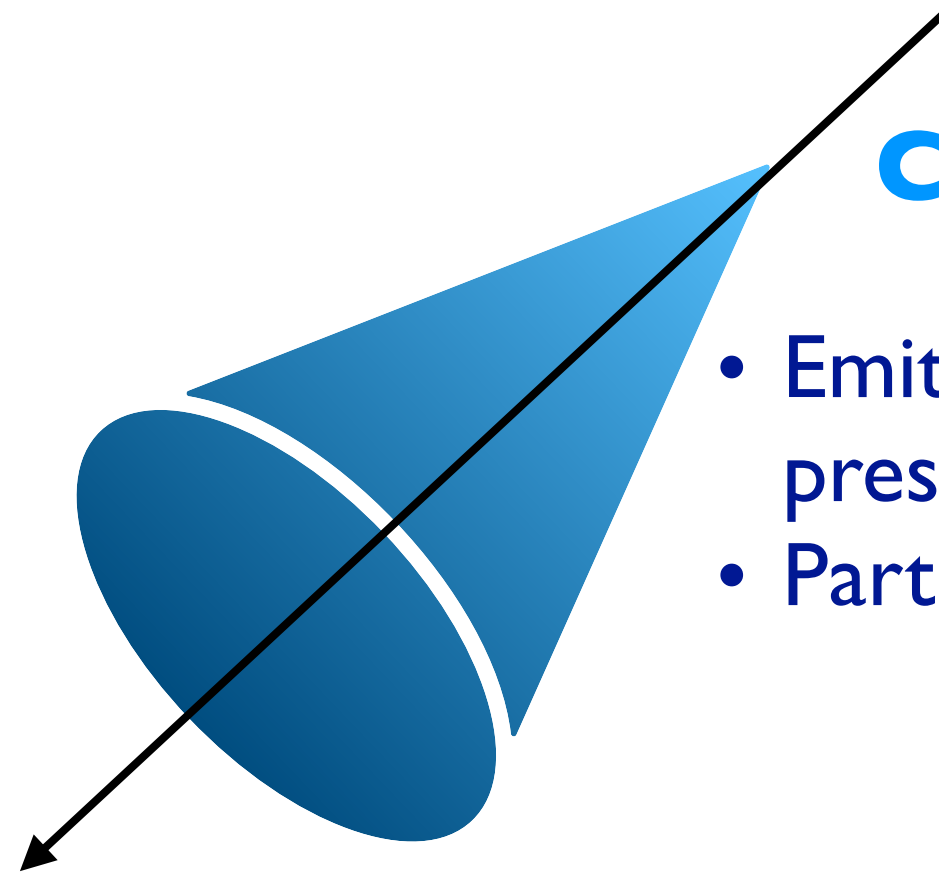


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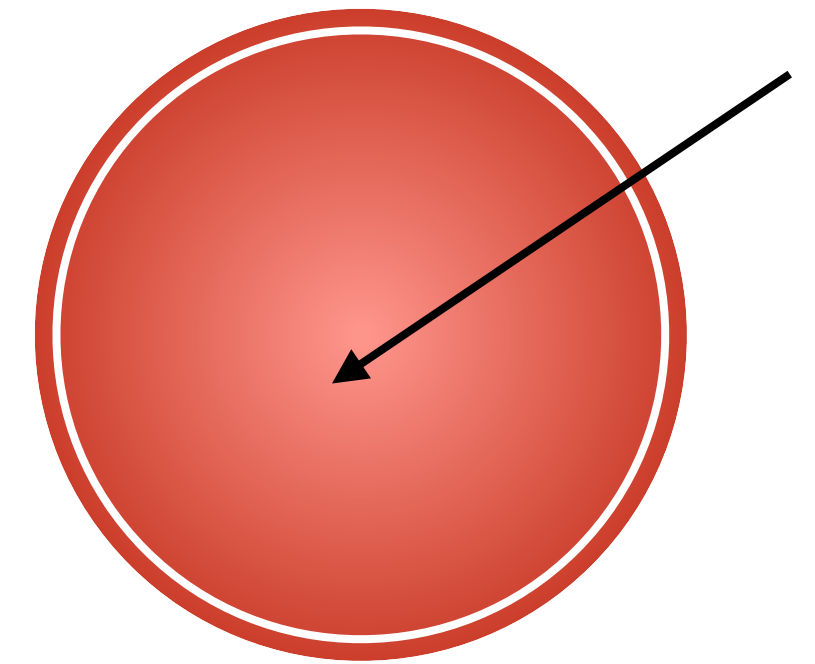
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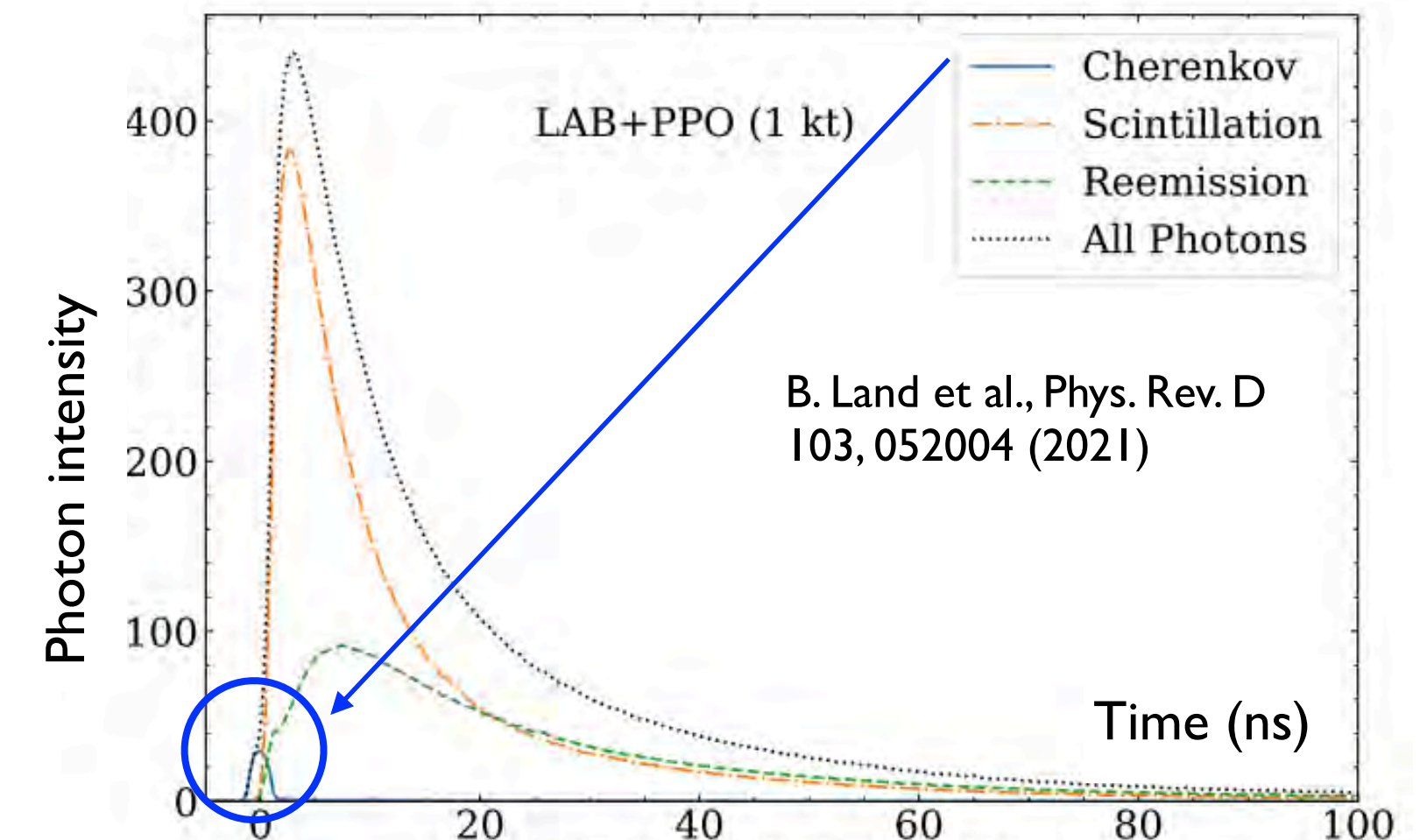
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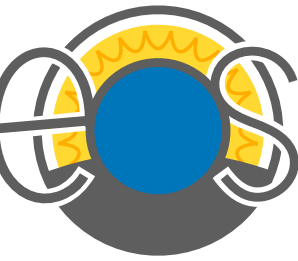
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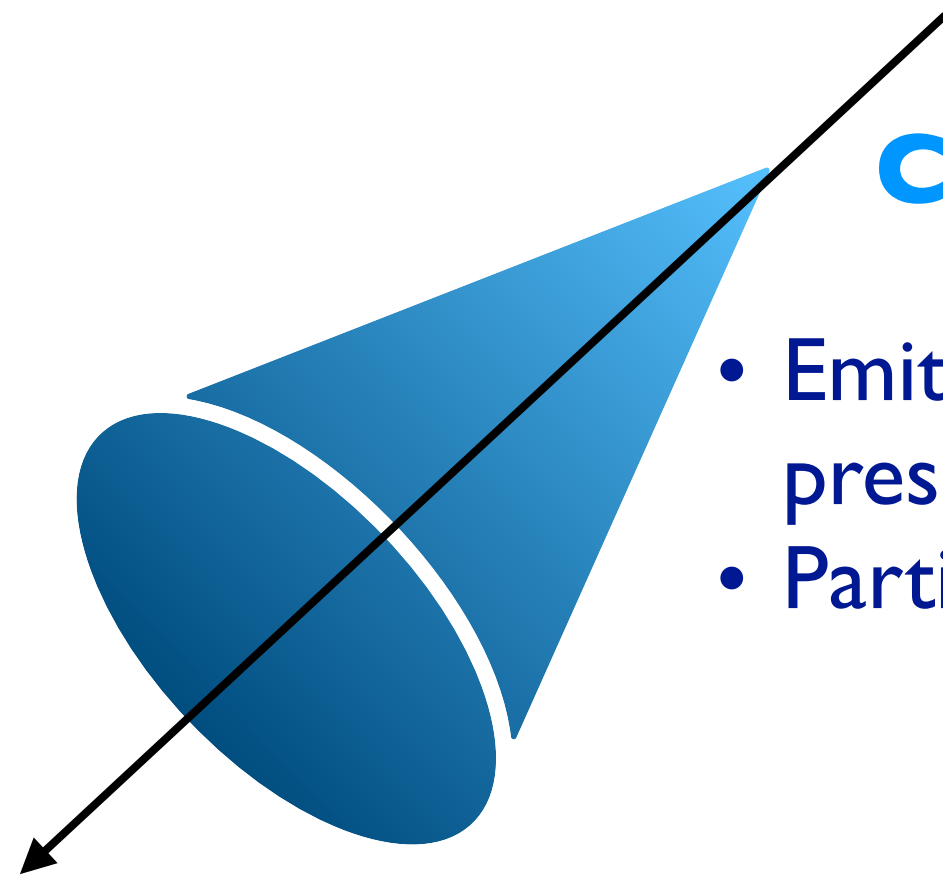


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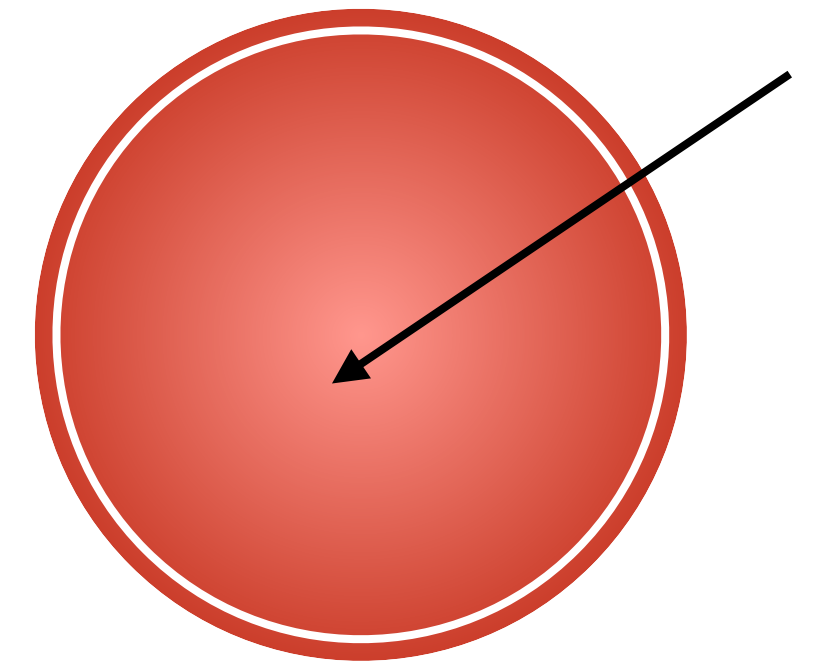
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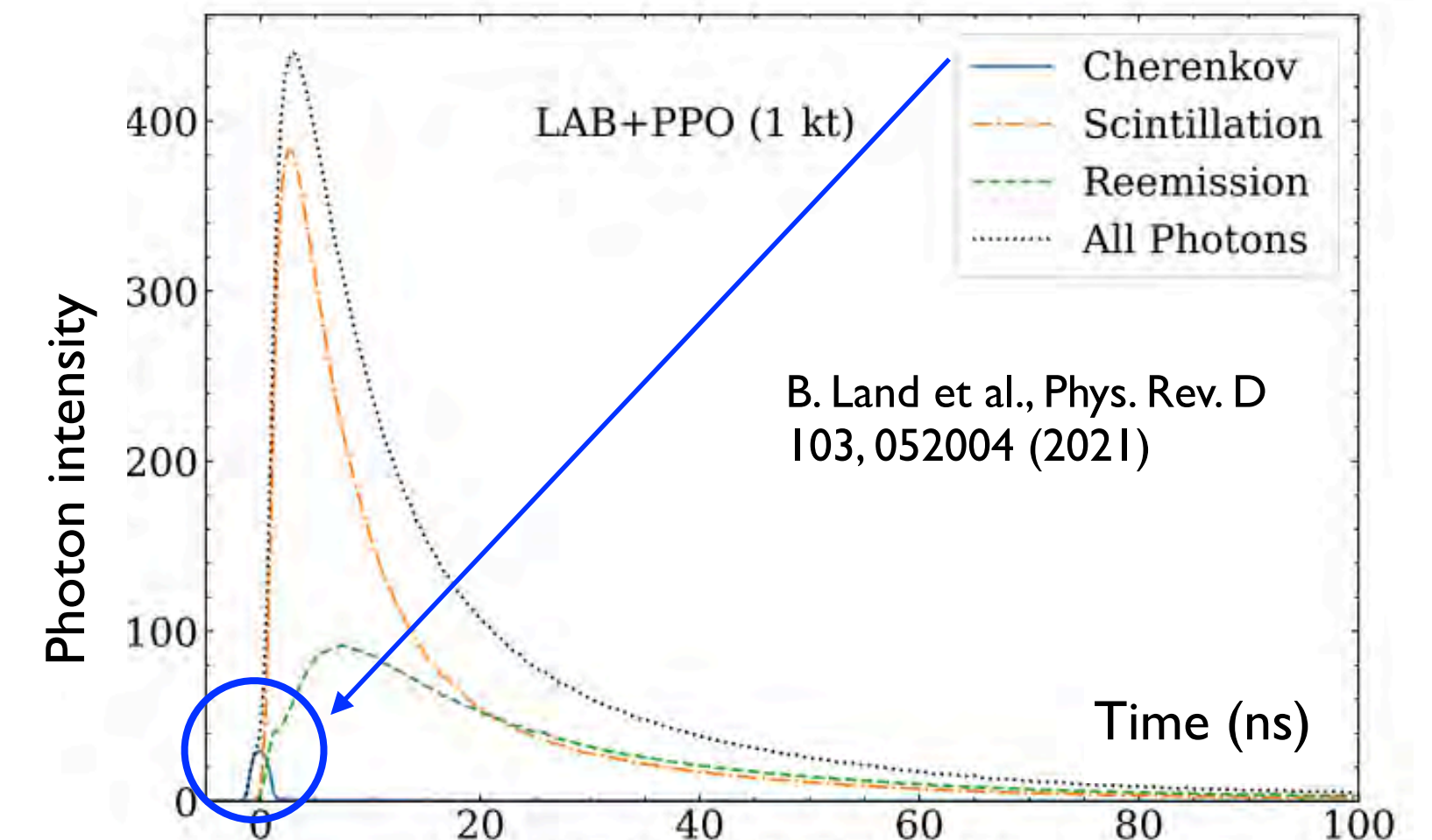
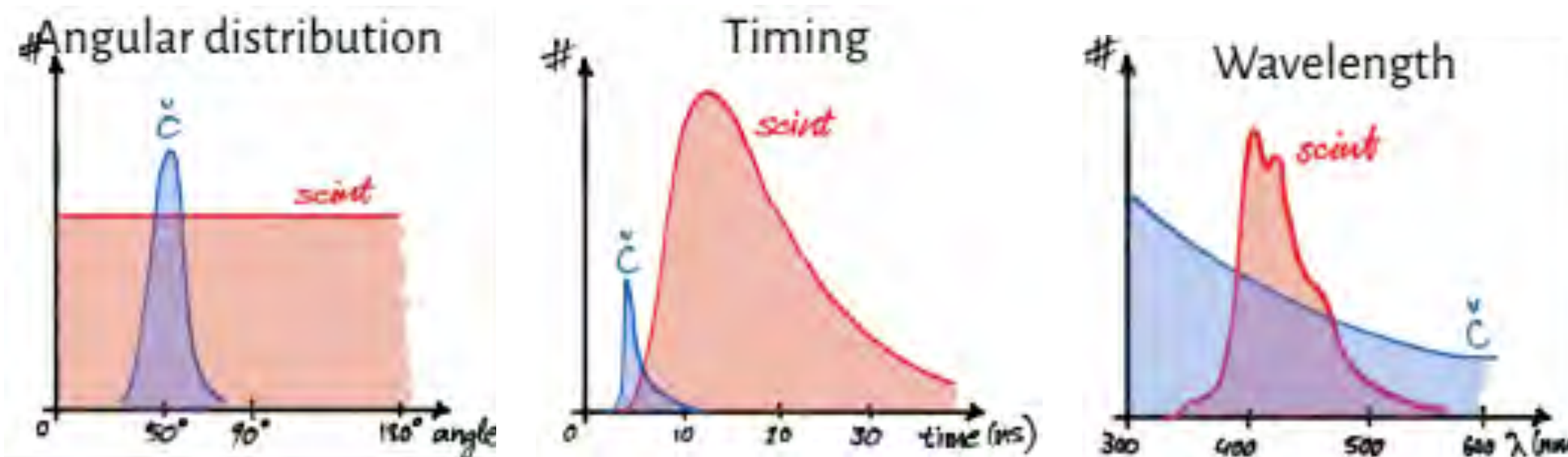
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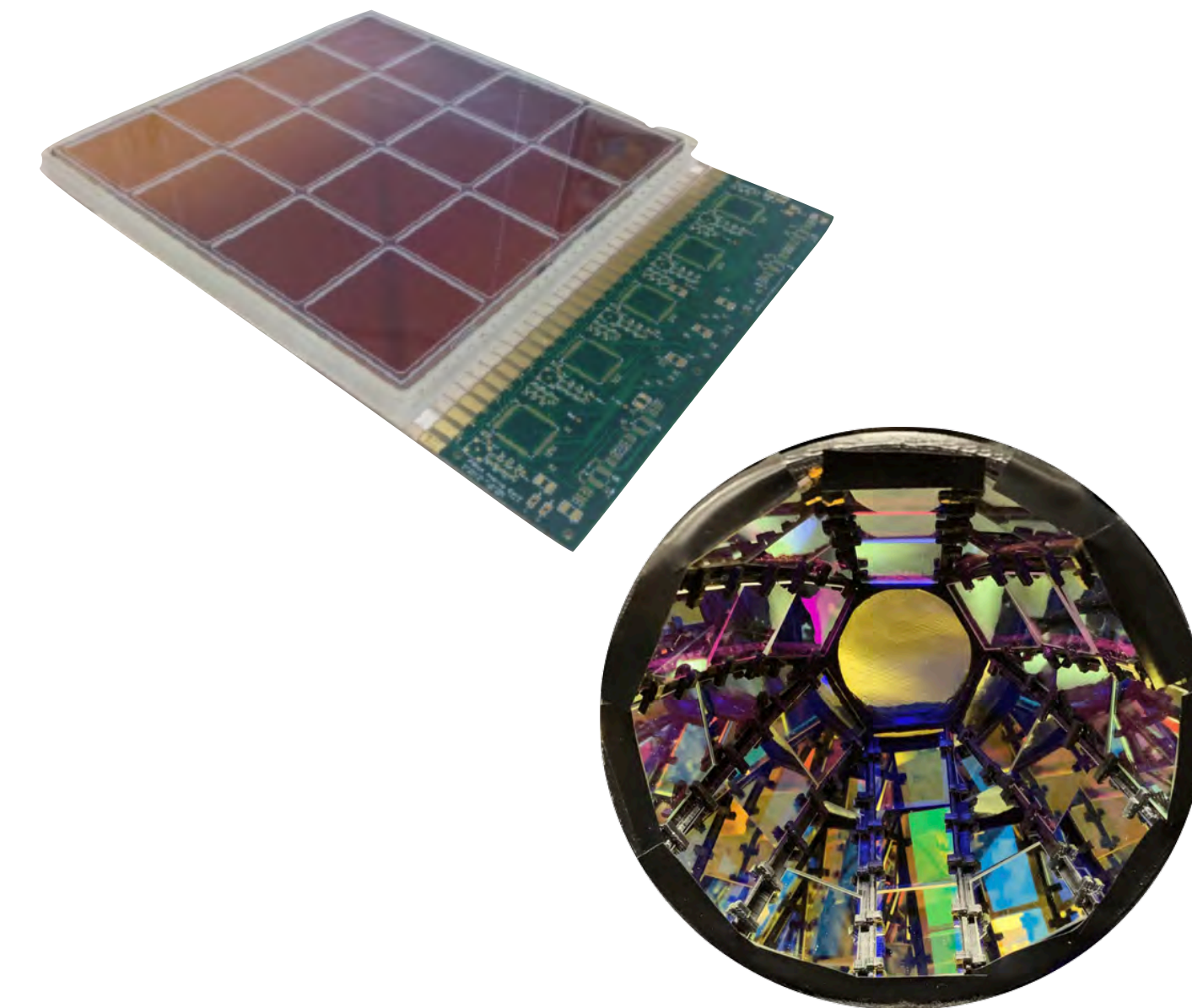
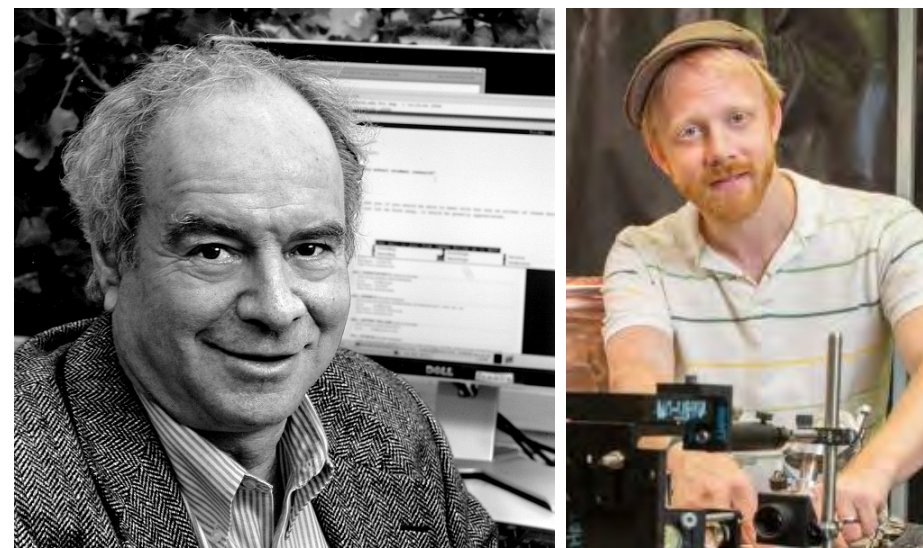
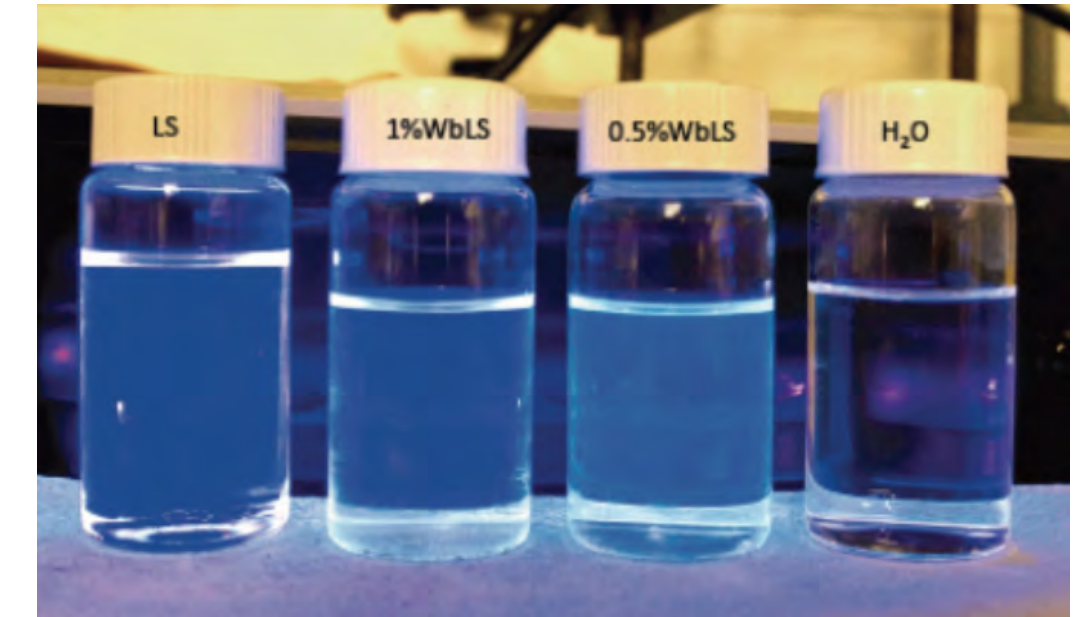
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Enabling technology

We focus our studies on three technologies that optimize hybrid Cherenkov/scintillation detection:

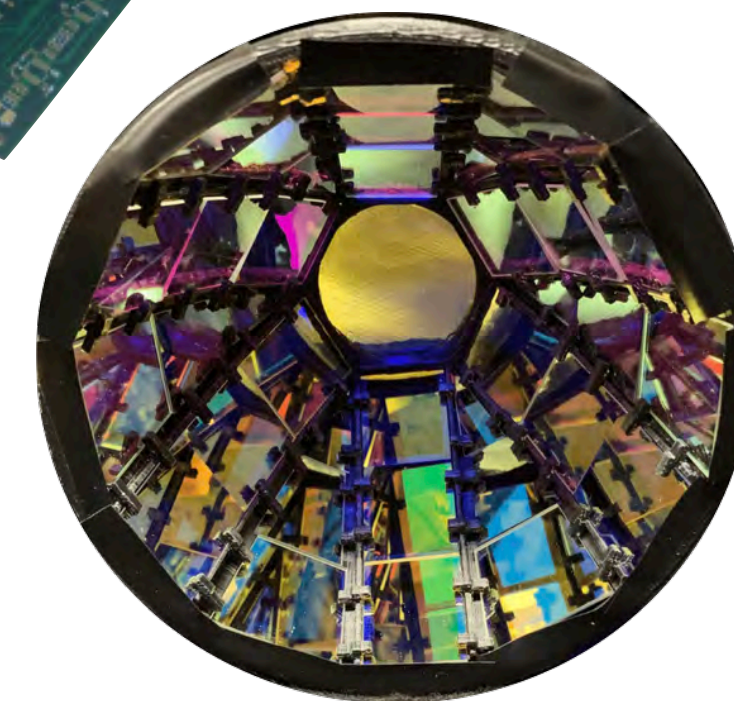
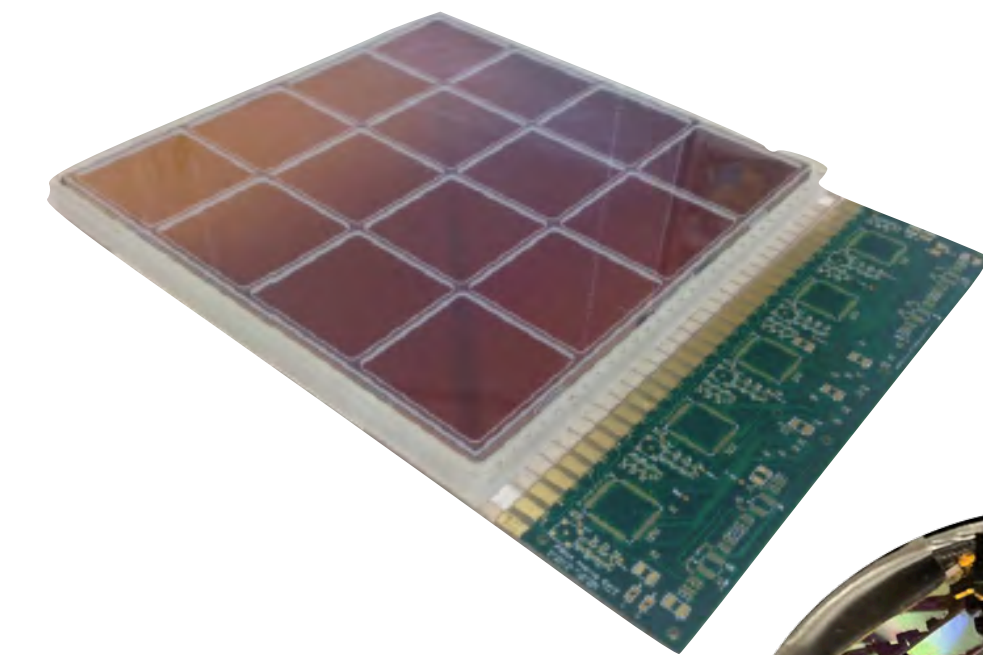
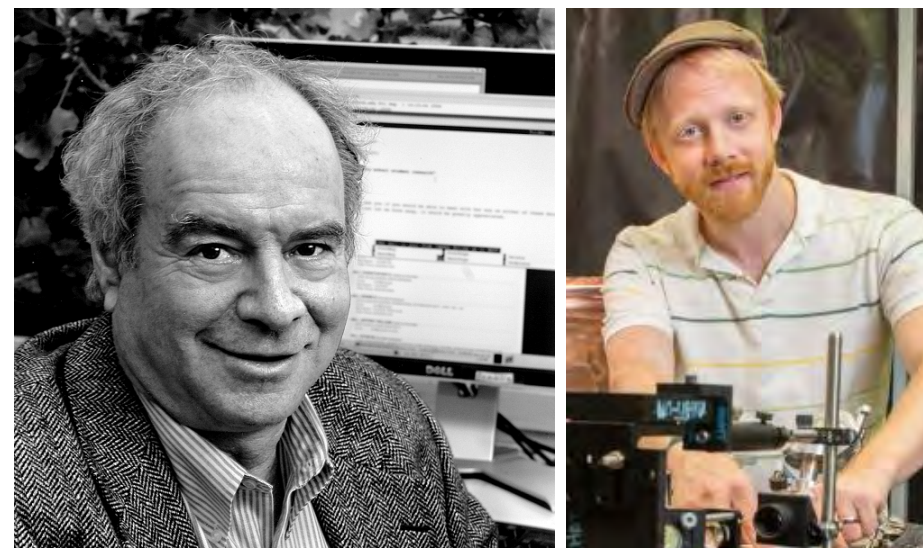
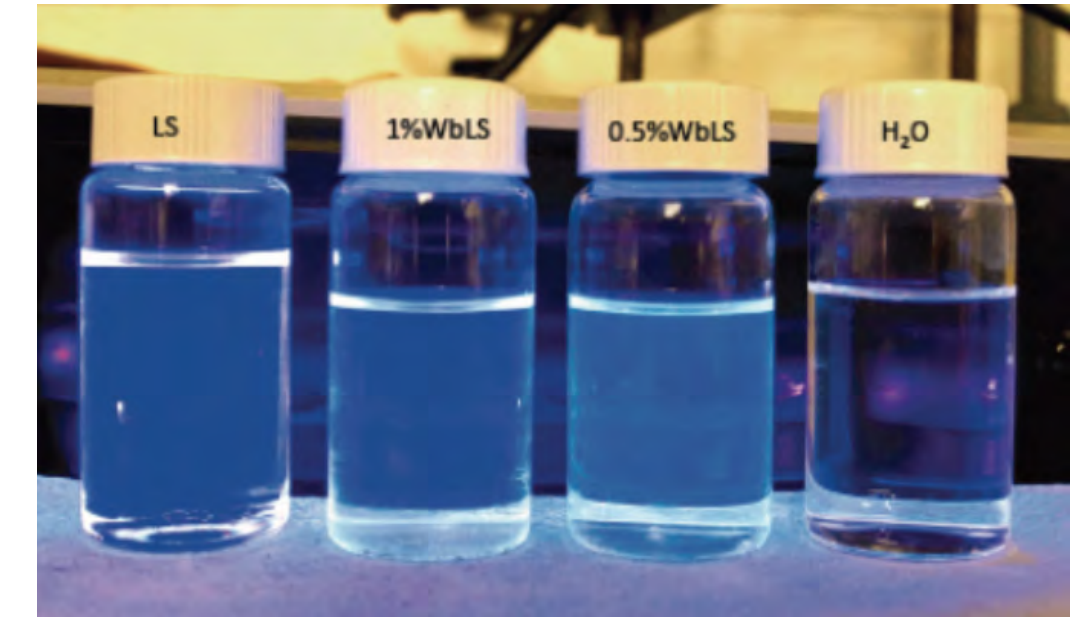


We seek to characterize behavior, understand and model performance at a microphysical level, and use results to extrapolate performance to kton scales.

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- I. Novel targets, such as water-based liquid scintillator (WbLS). Enhances Cherenkov detection by “dialling down” or otherwise modifying the scintillation signal

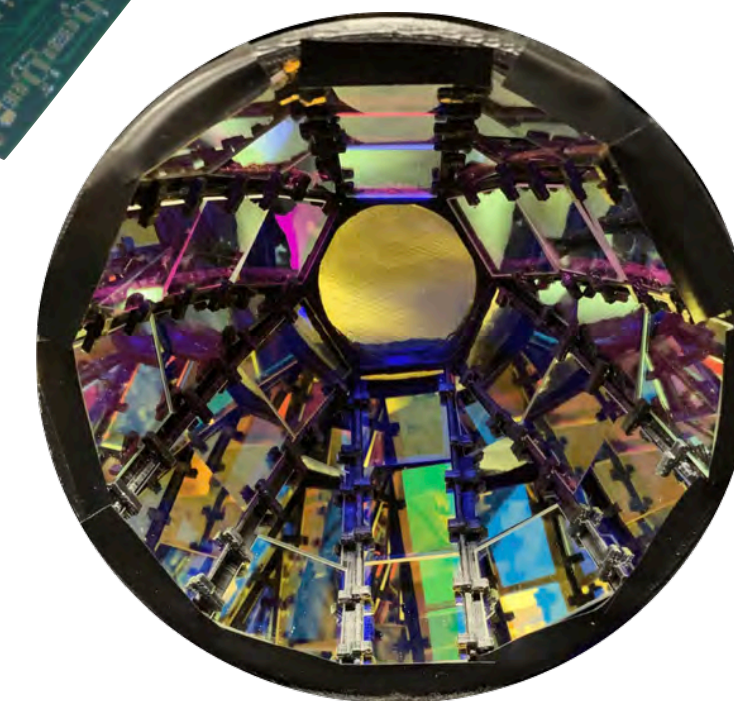
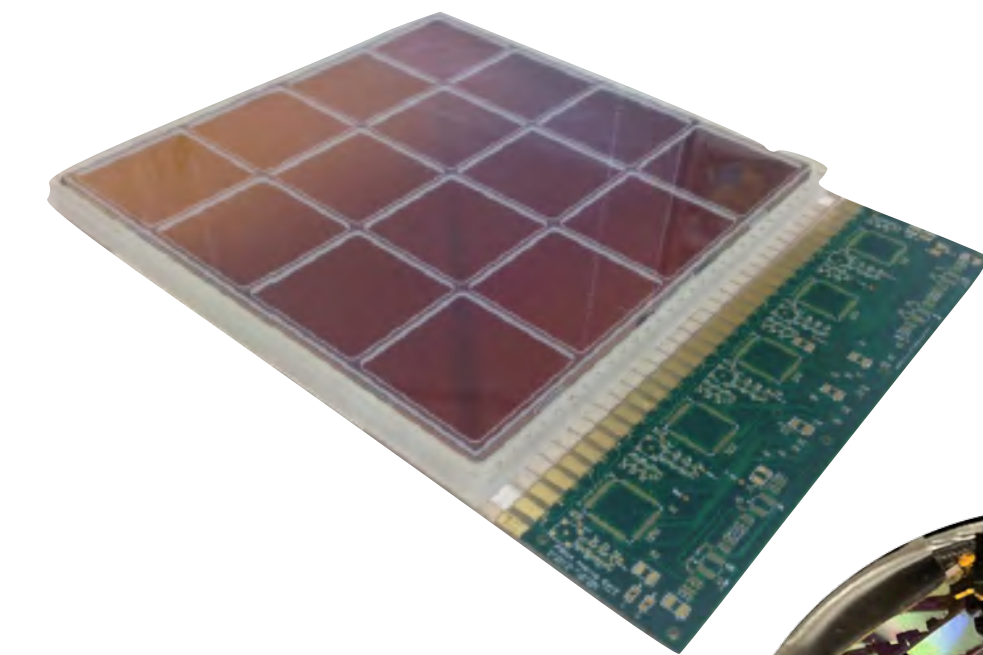
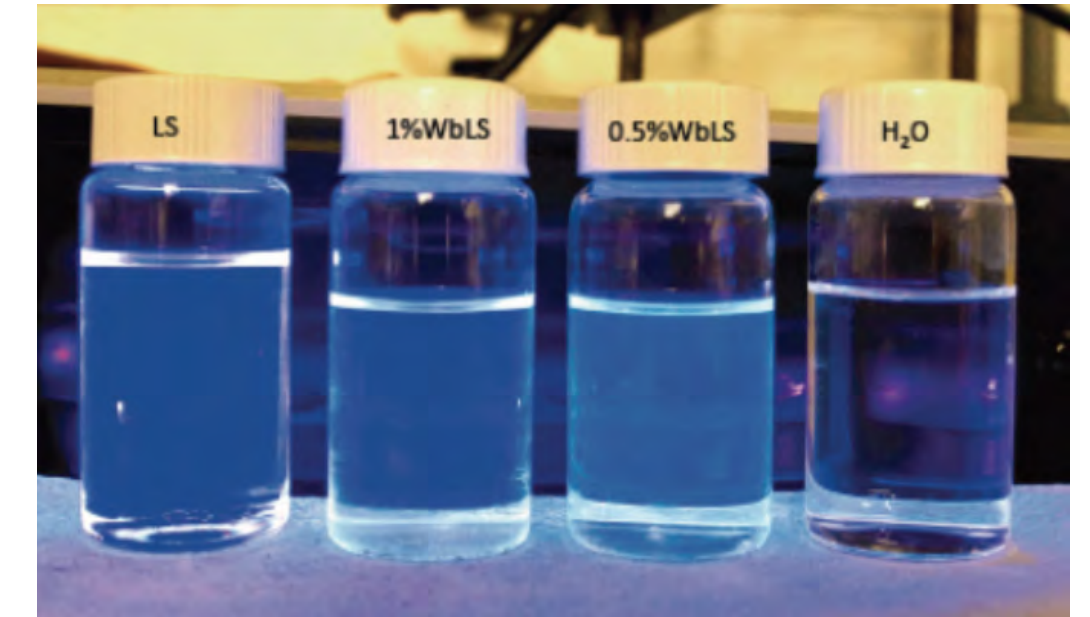


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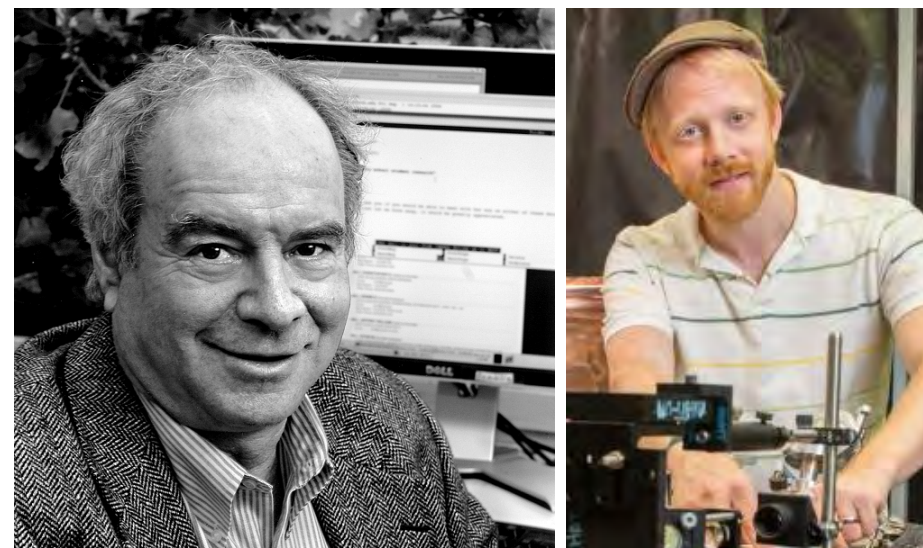
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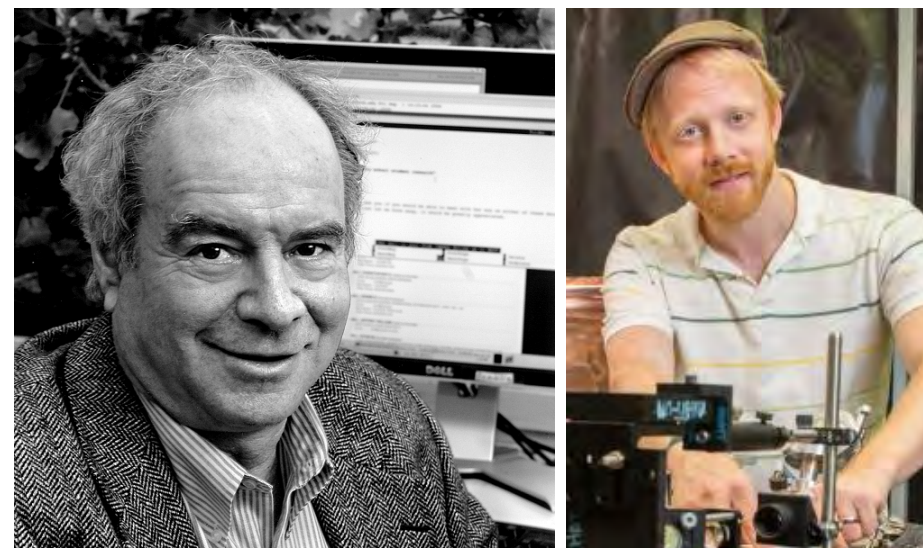
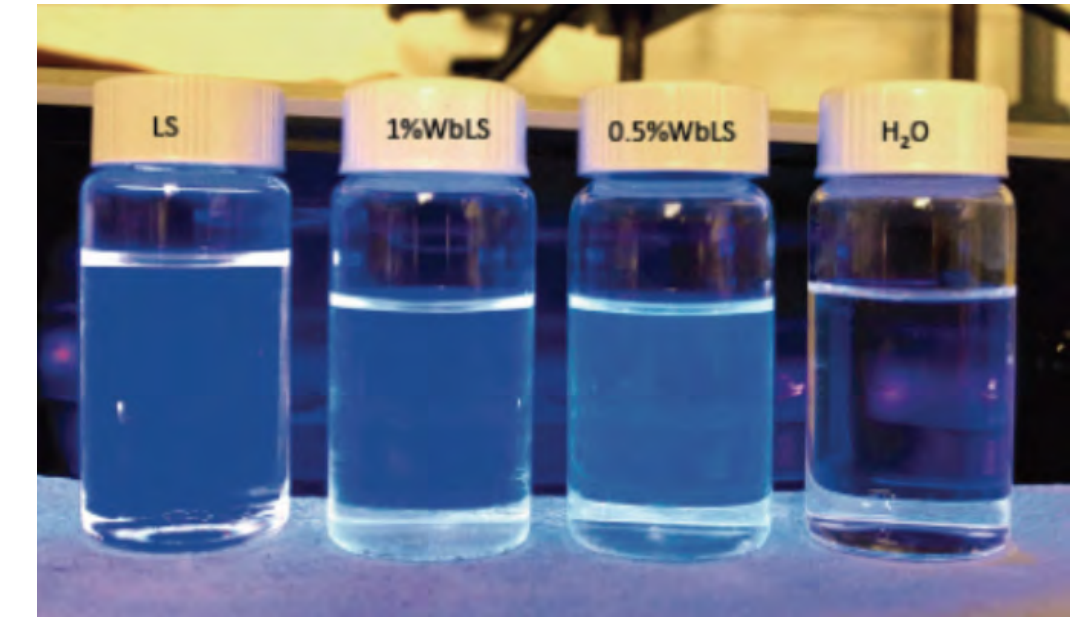


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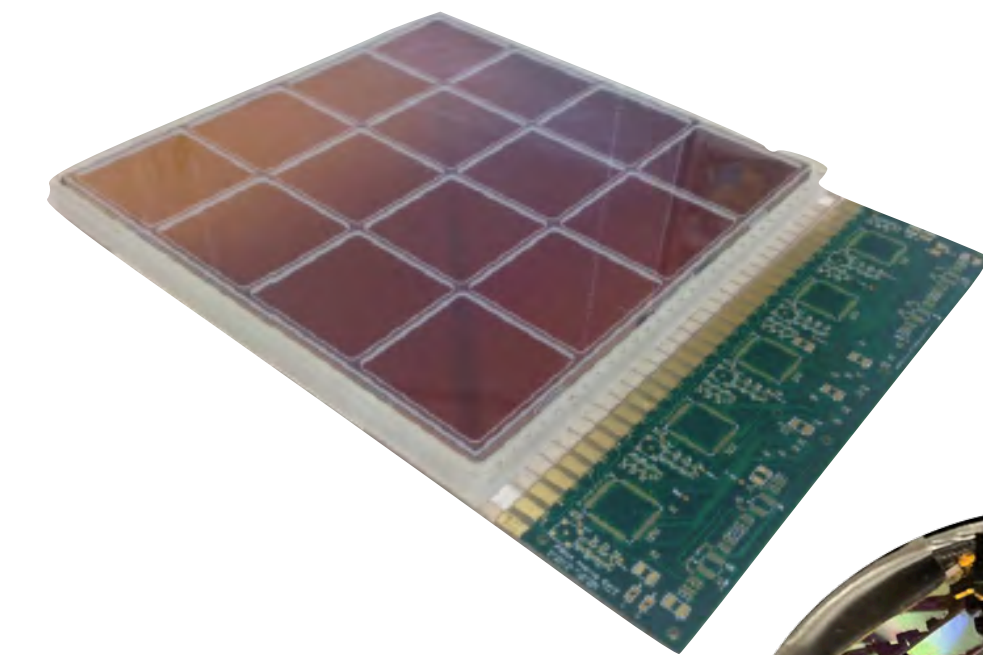
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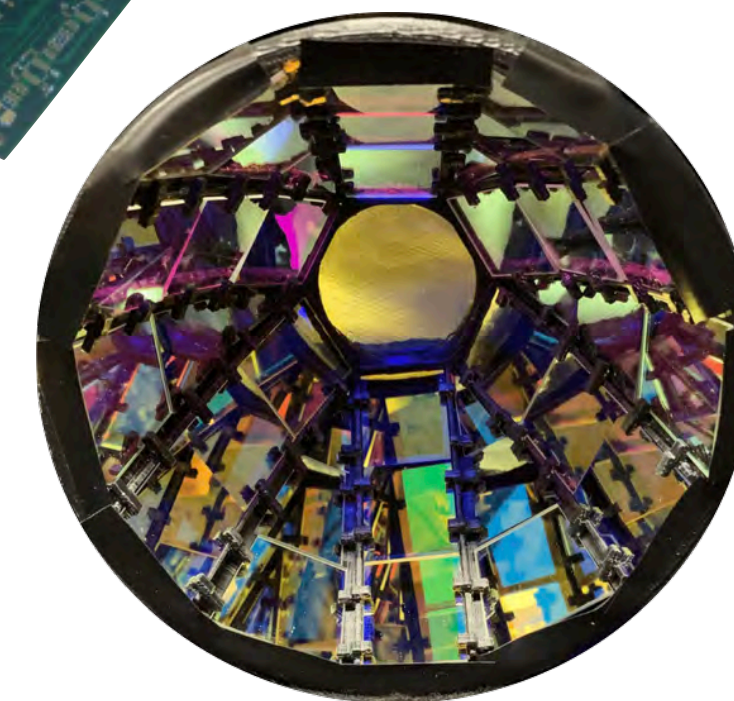
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2. Large-Area Picosecond Photon Detectors (LAPPDs). Fast-timing discrimination for vertex resolution and Cherenkov/scintillation separation



3. Dichroicons (“chromatic quantum sensing”). Cherenkov/scintillation separation via spectral sorting



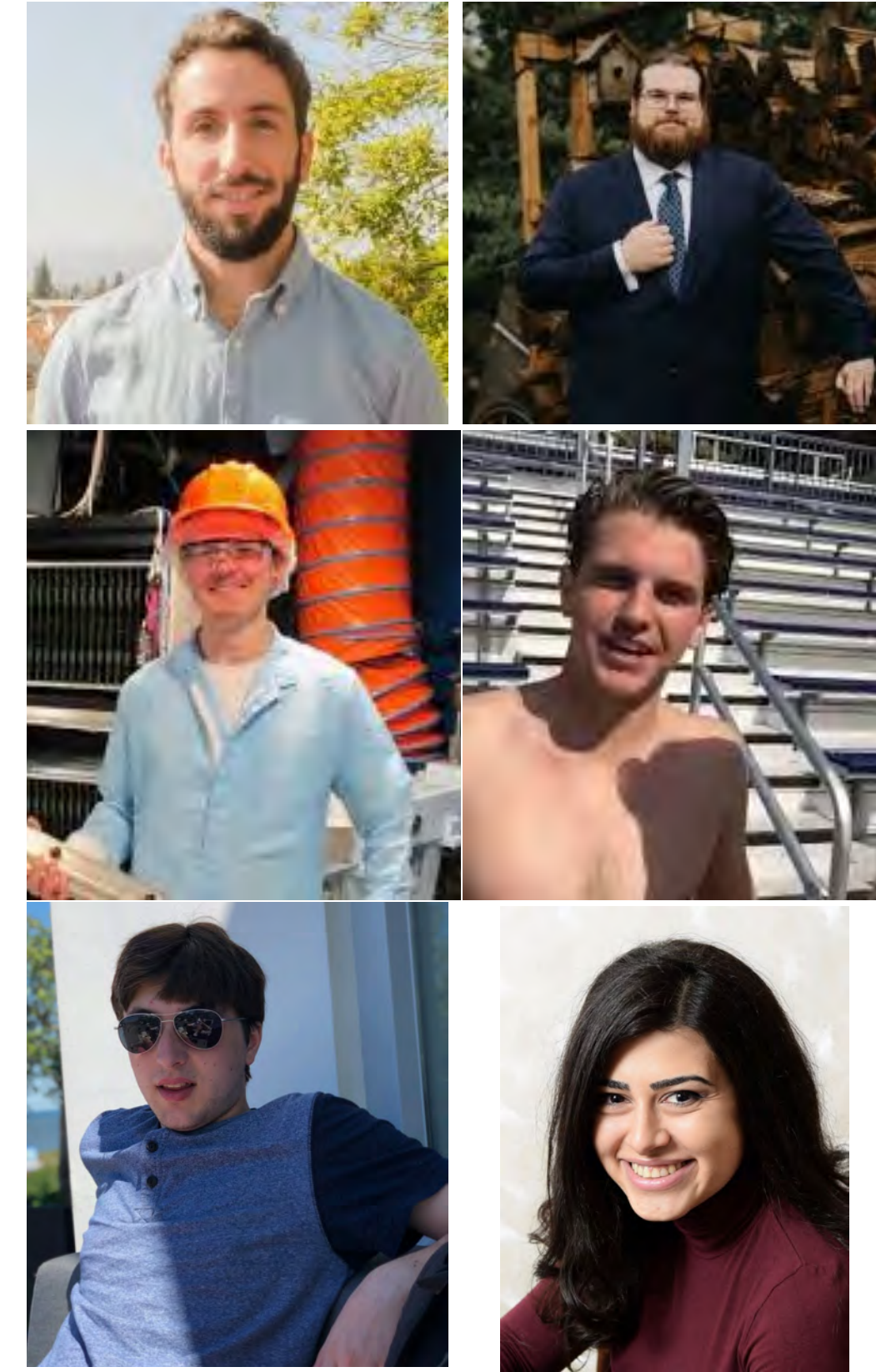
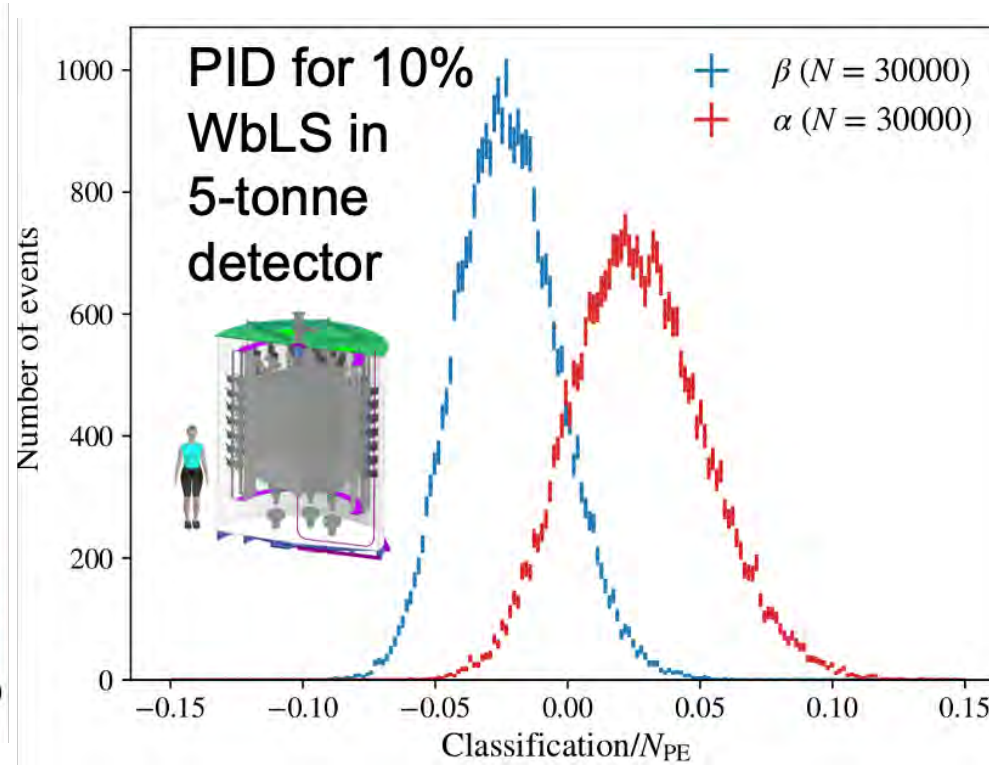
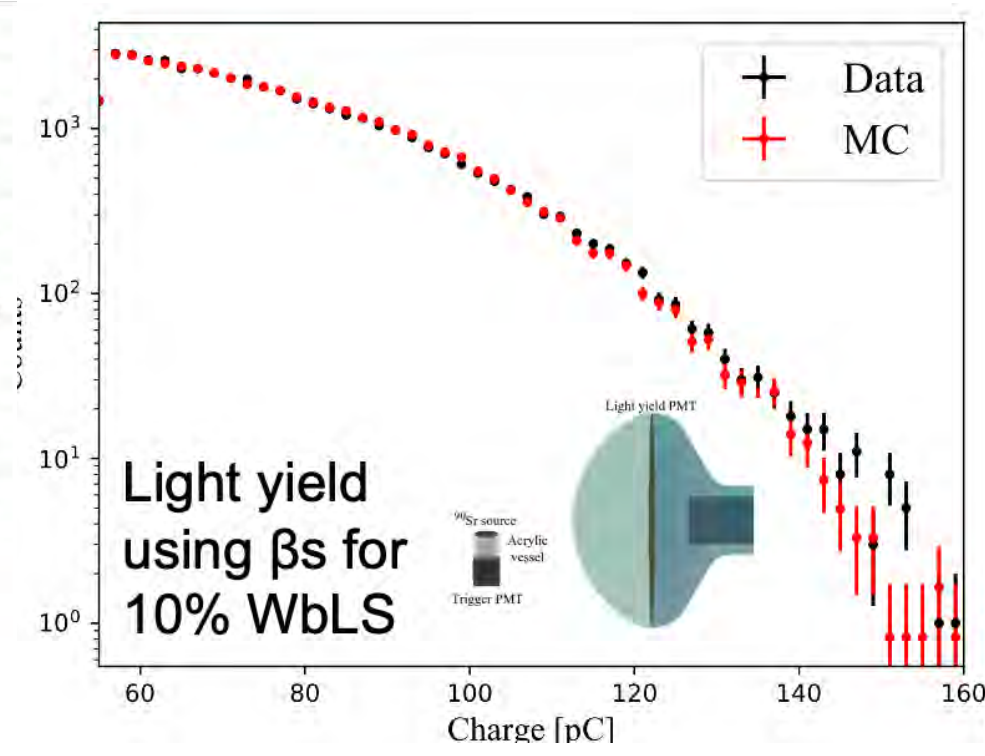
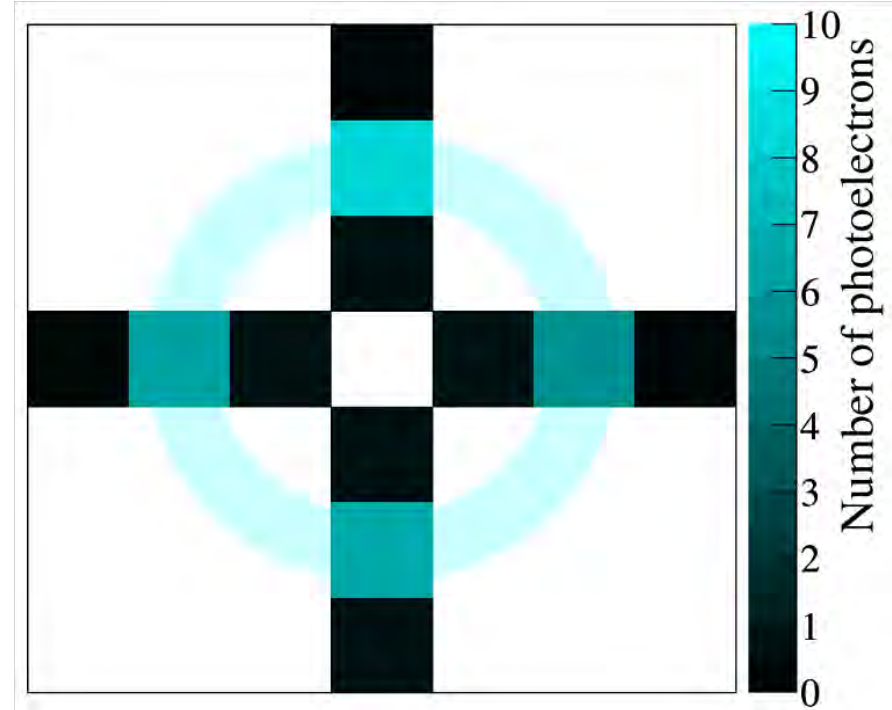
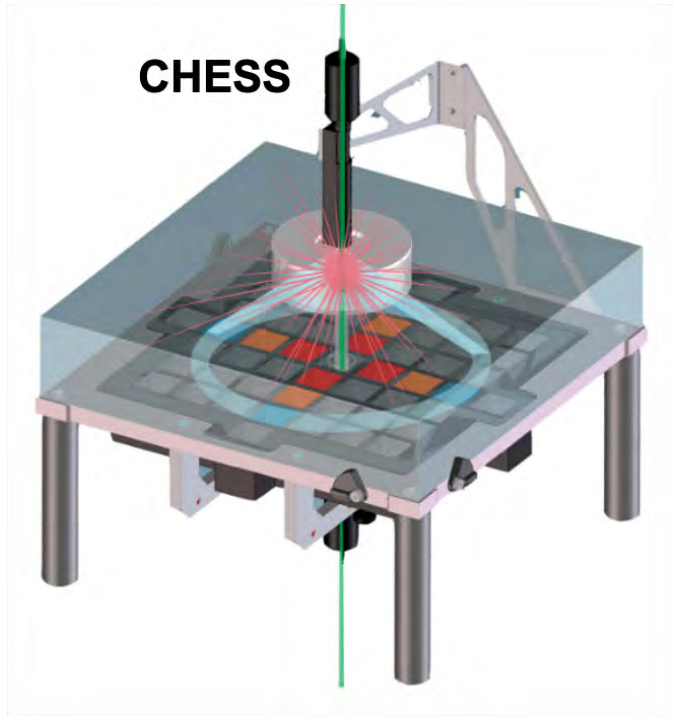
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Builds on core (Wb)LS development at BNL (Yeh et al.)

Technical accomplishments

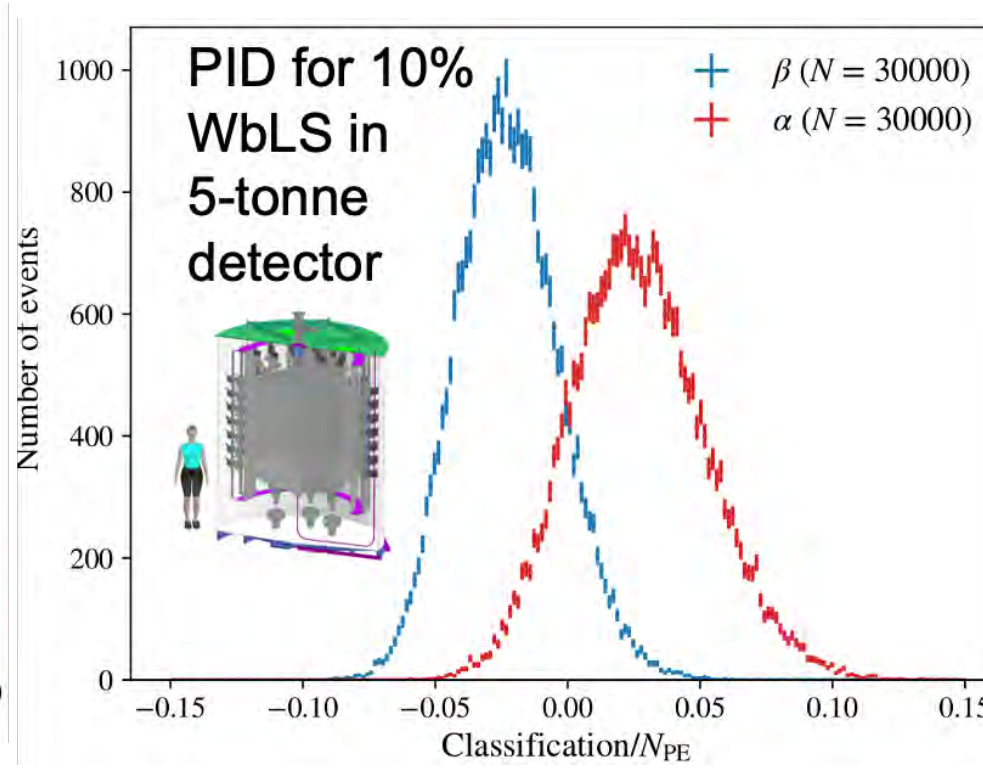
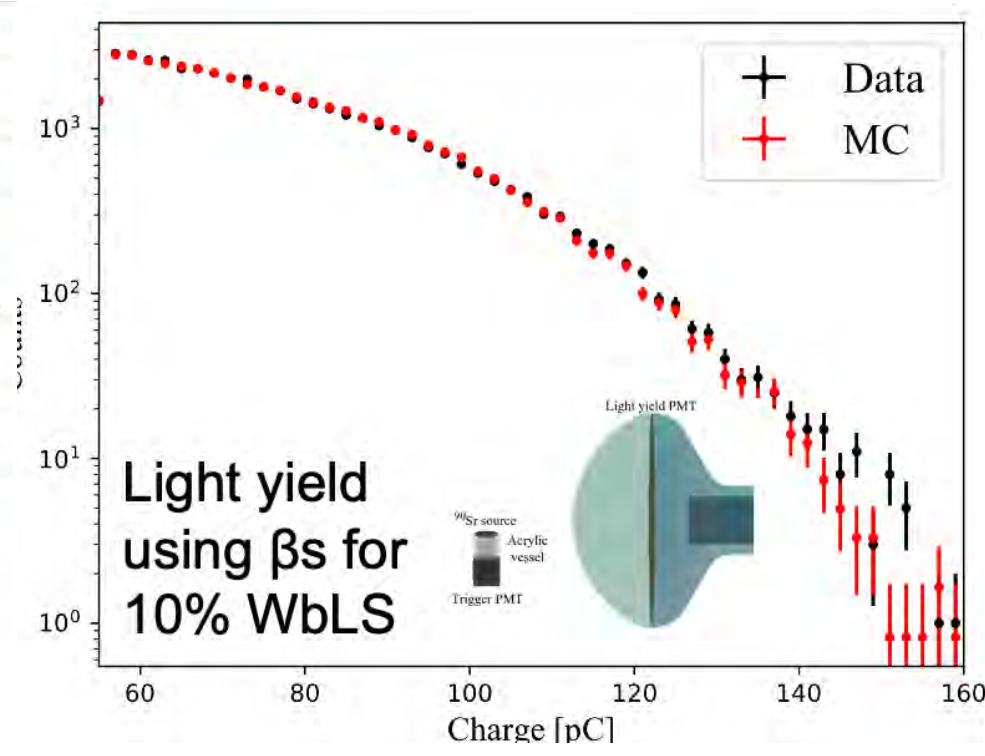
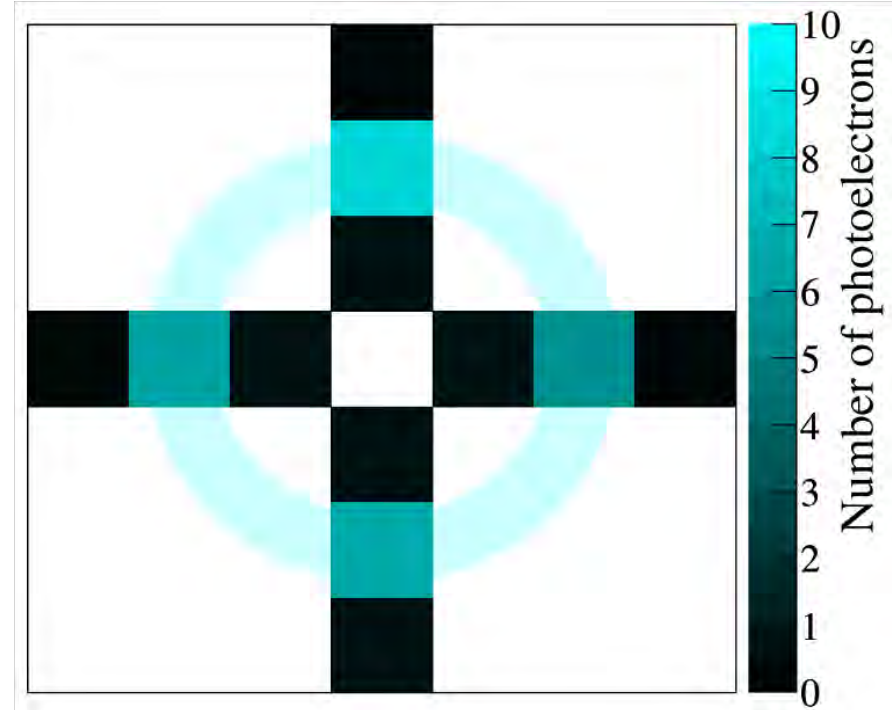
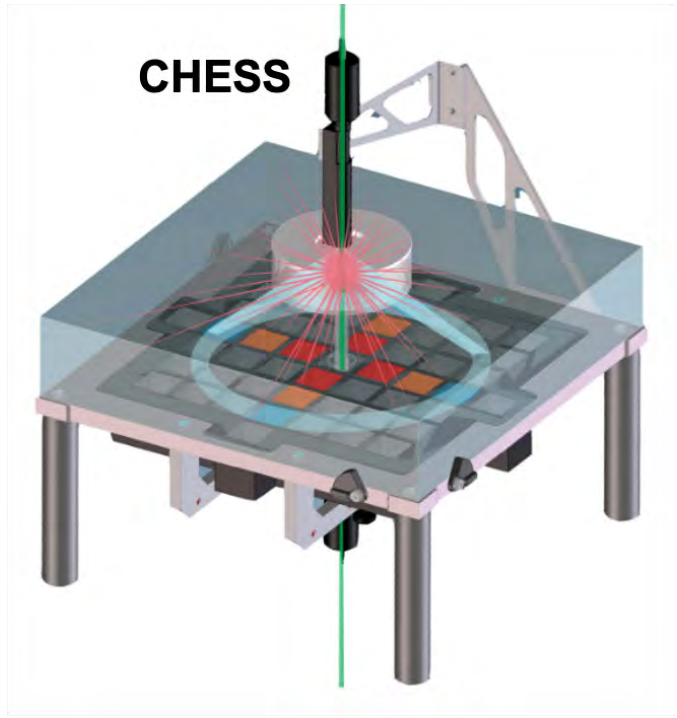
WbLS characterization



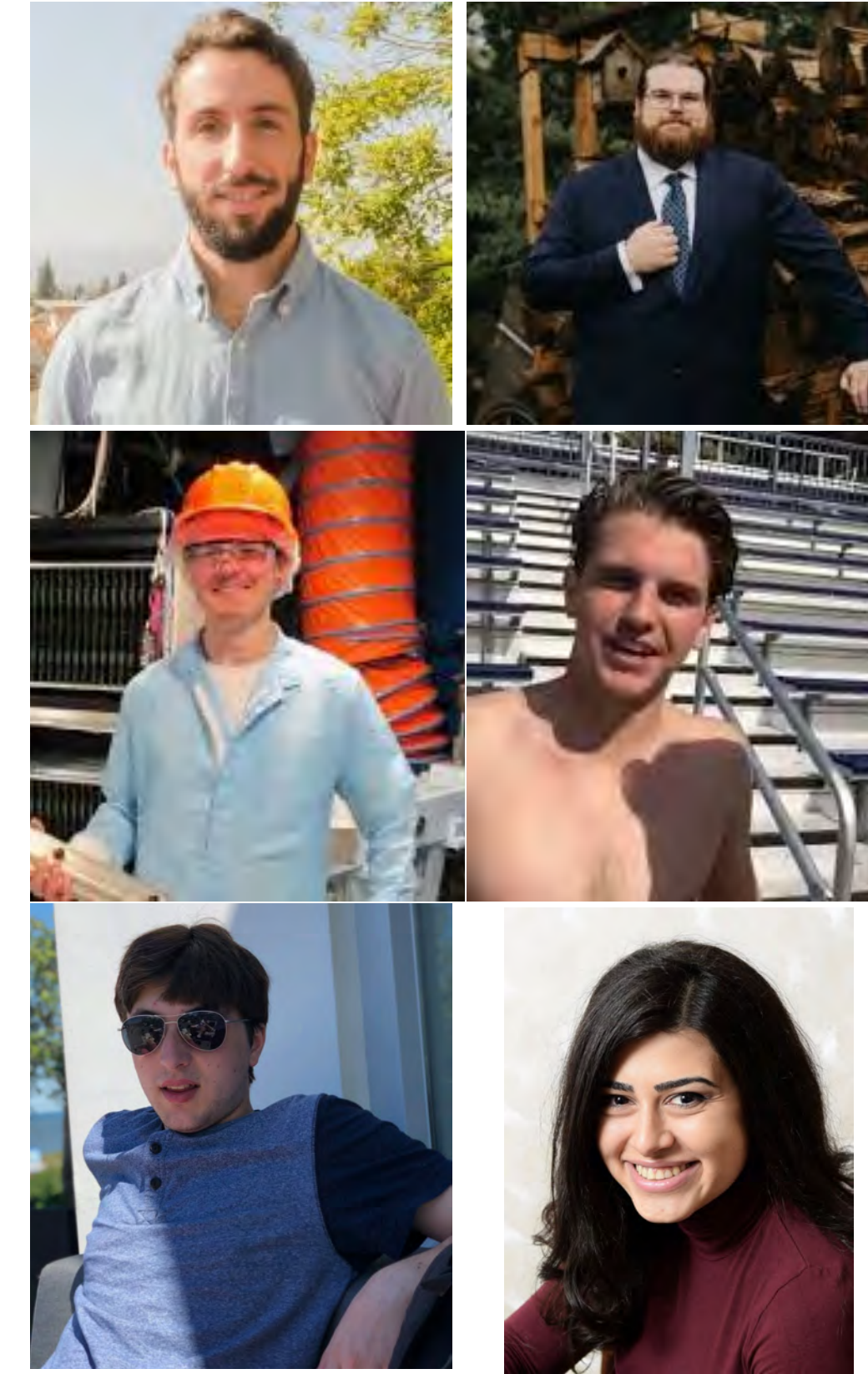
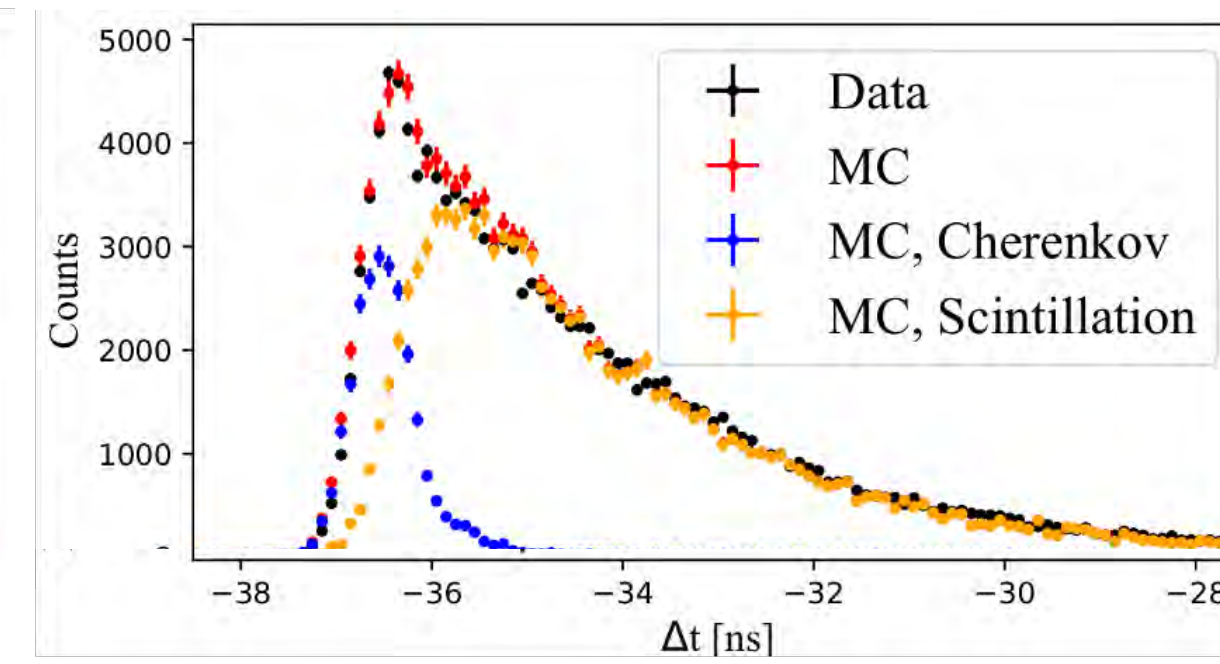
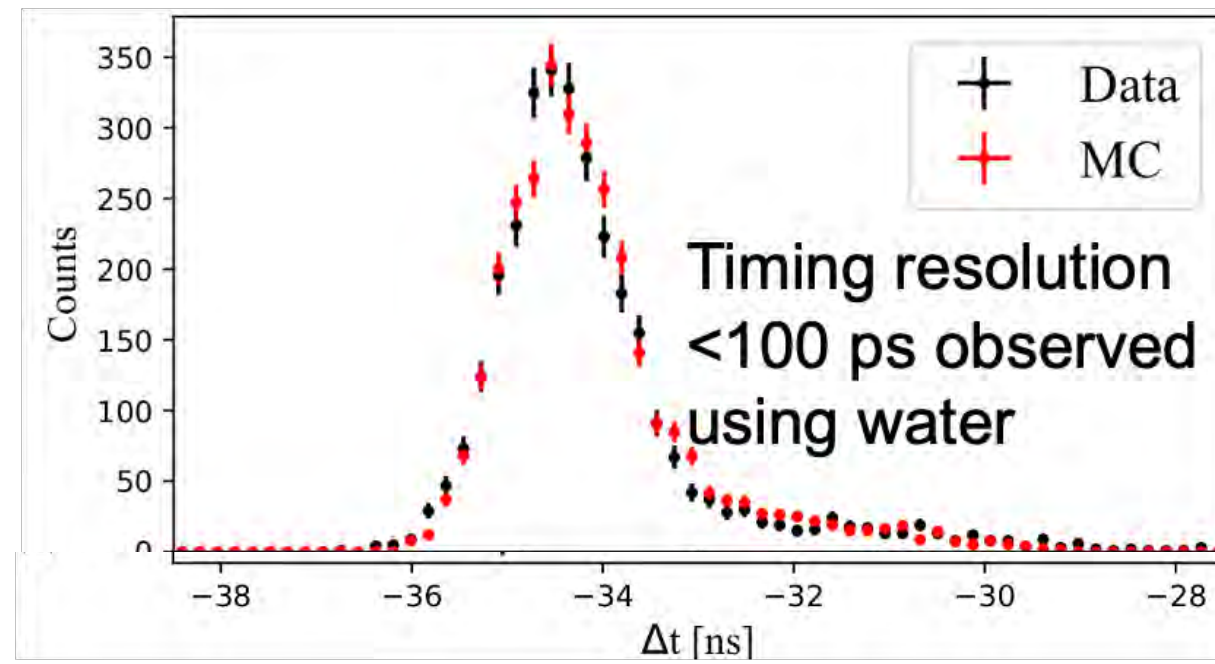
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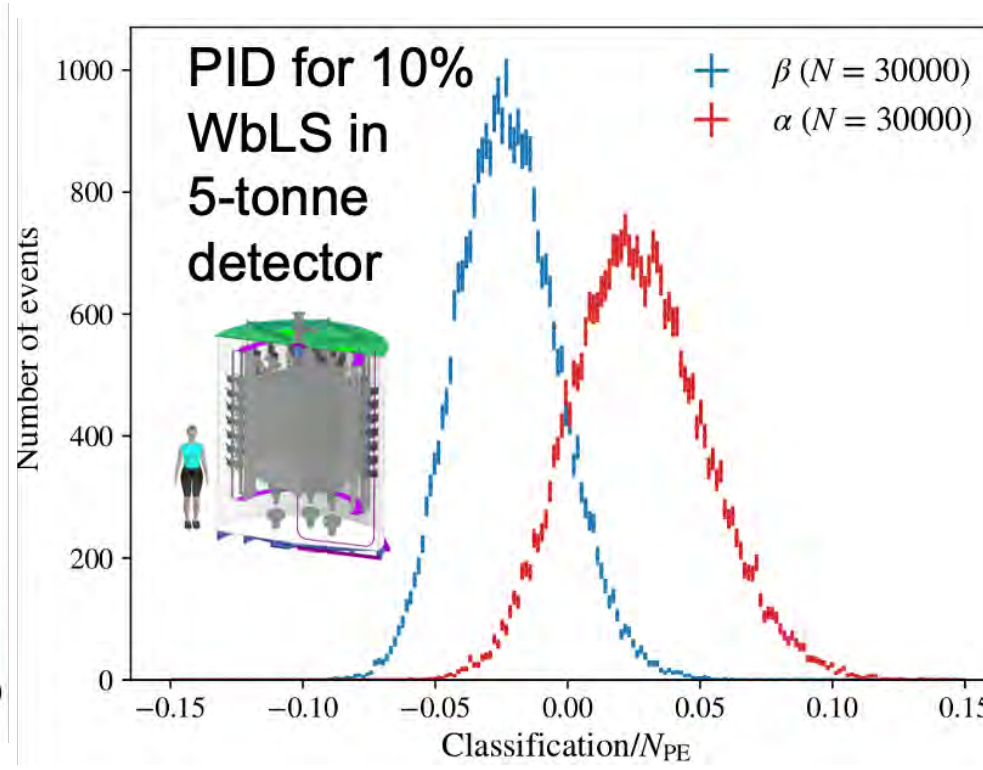
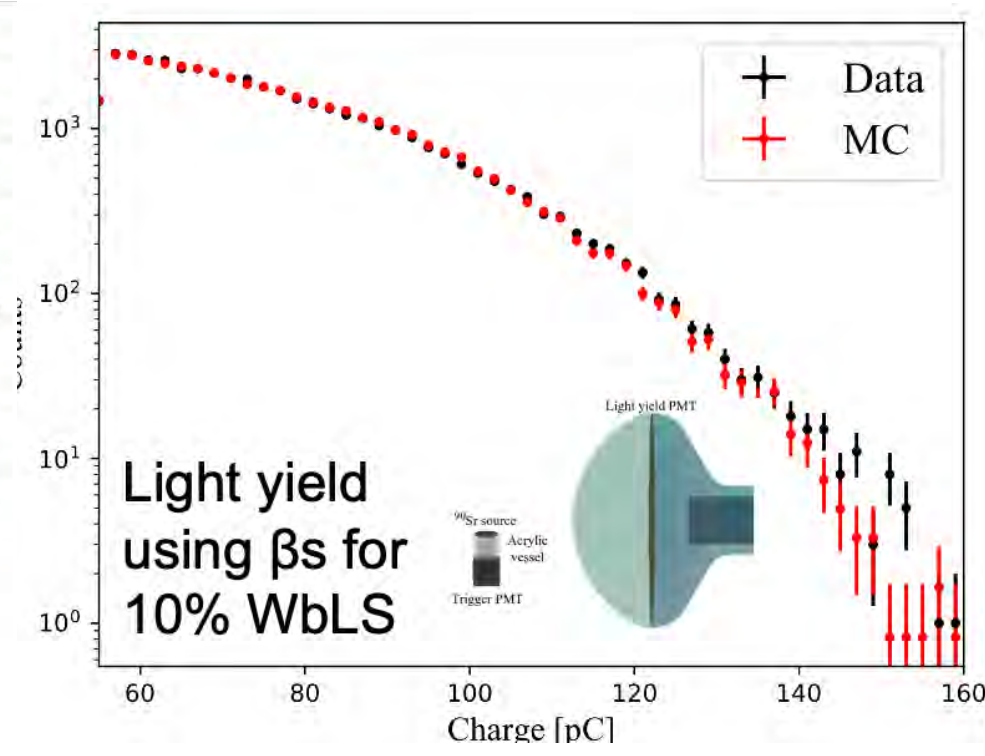
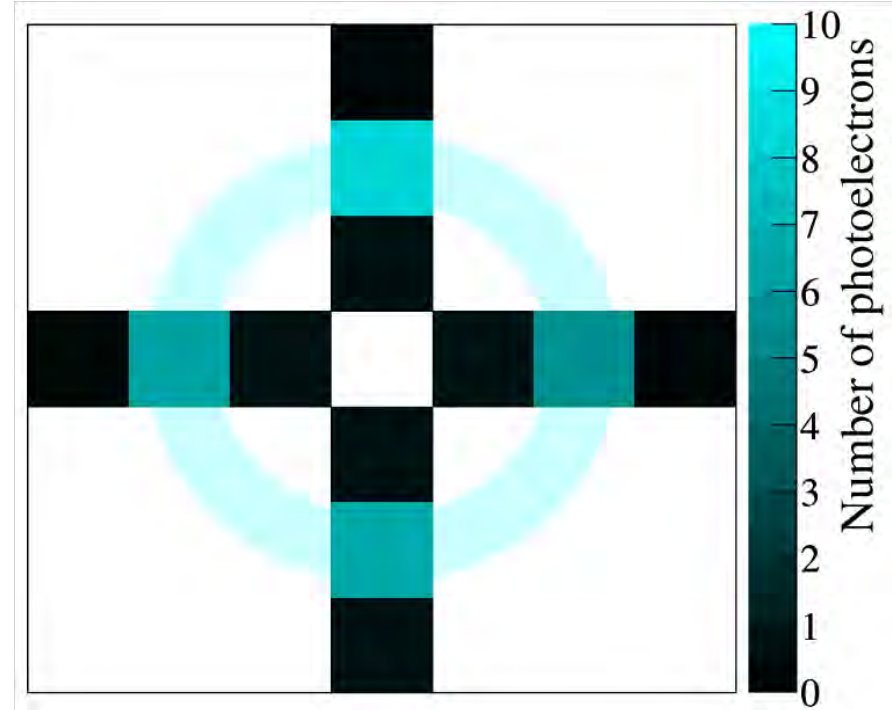
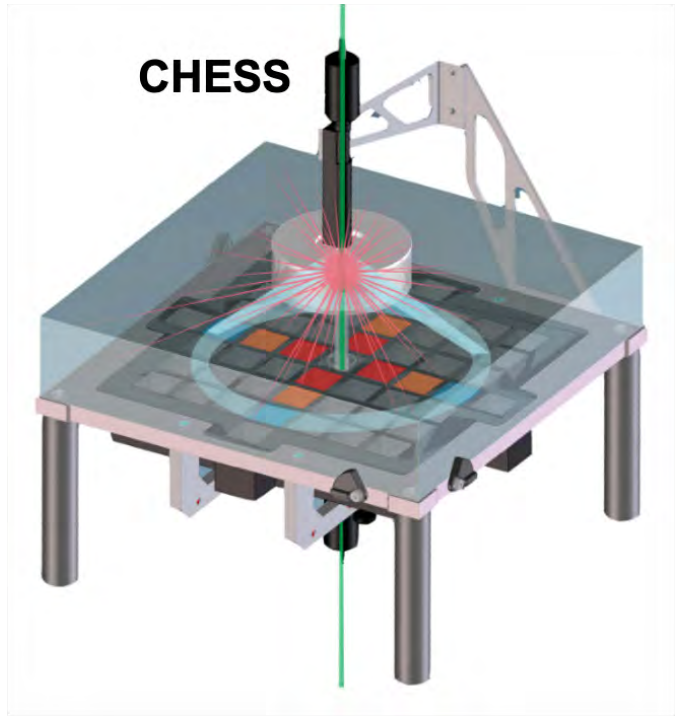
Fast timing photon detection



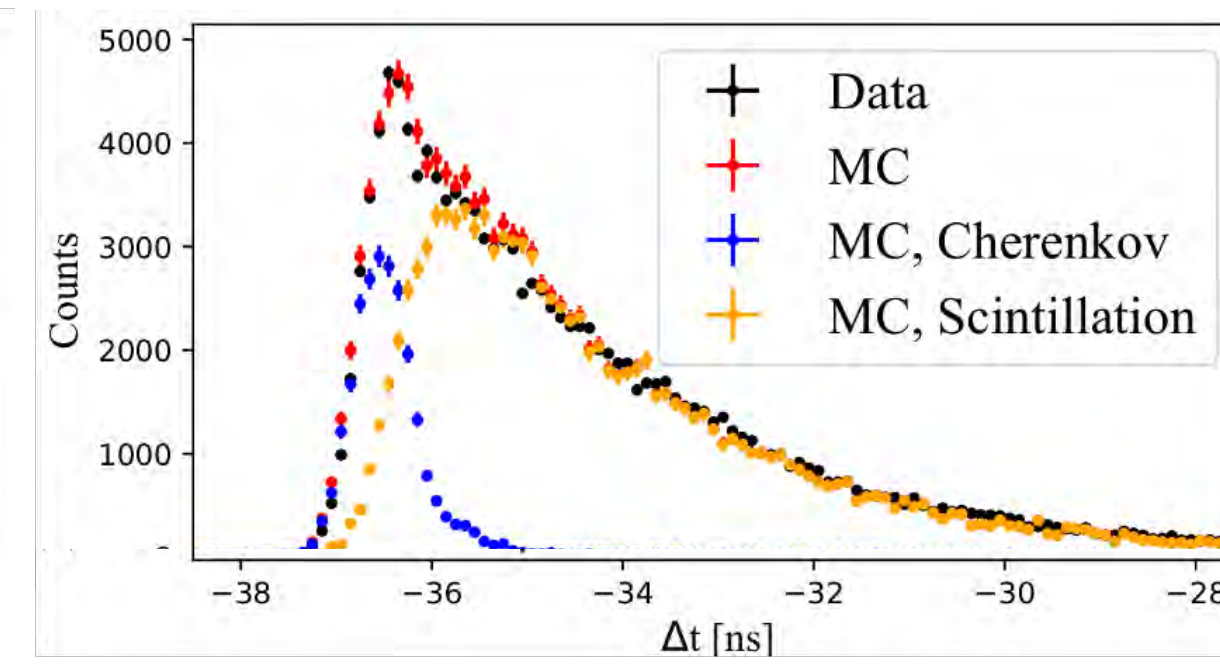
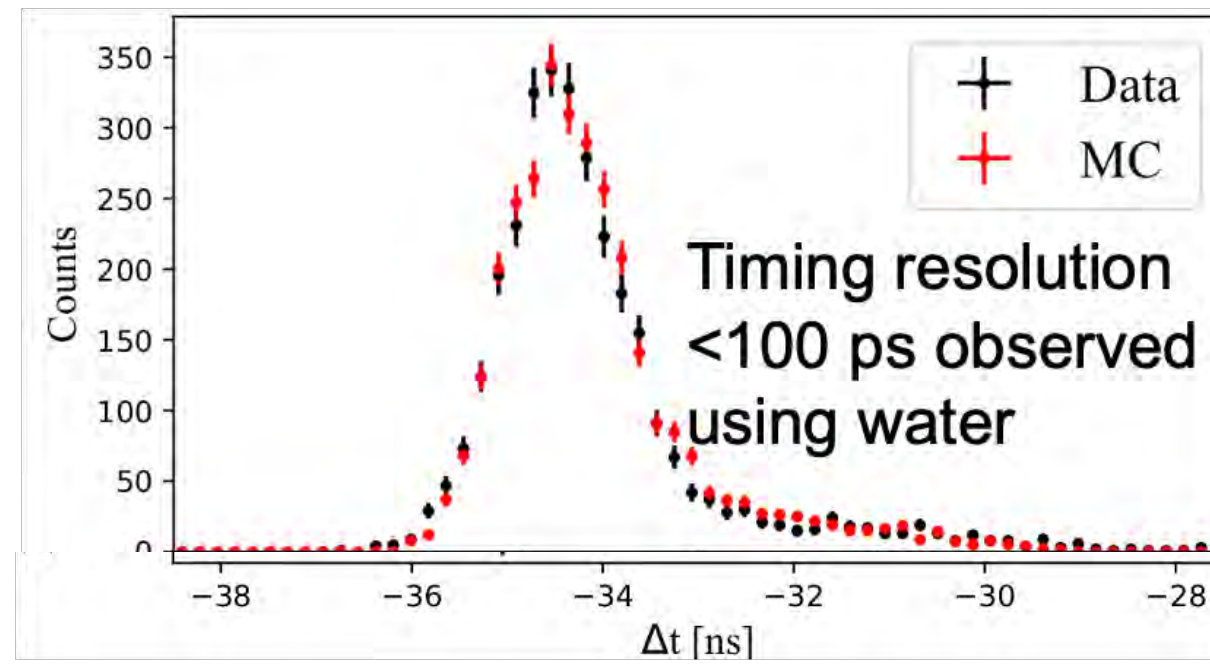
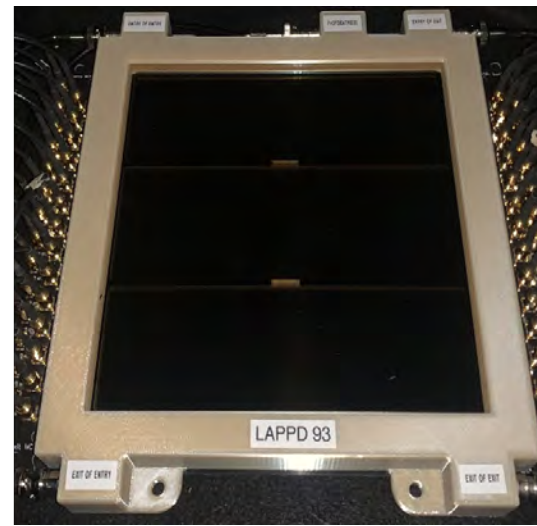
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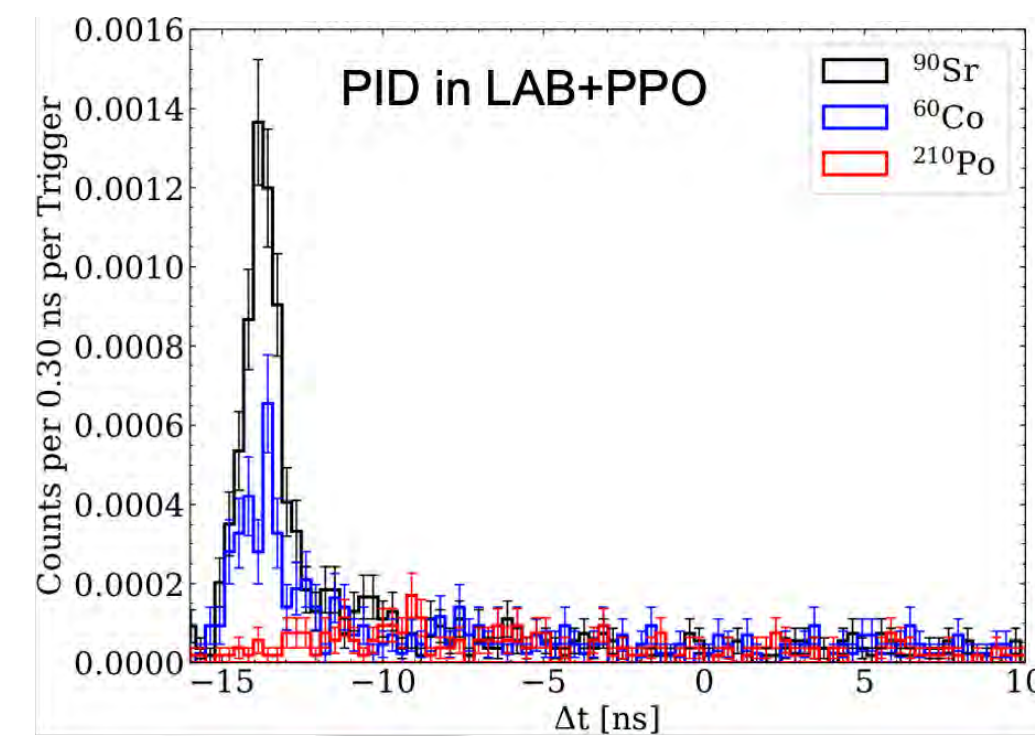
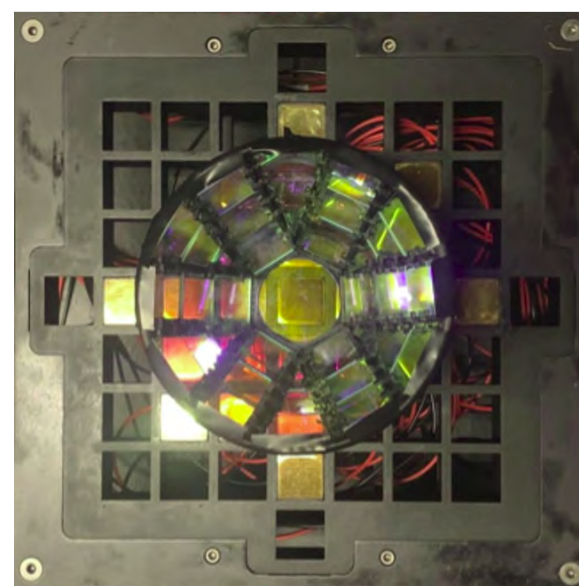
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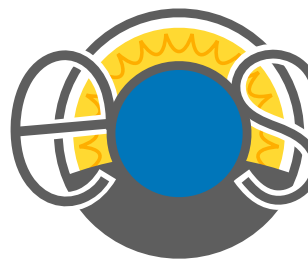


Fast timing photon detection



Quantum chromatic sorting

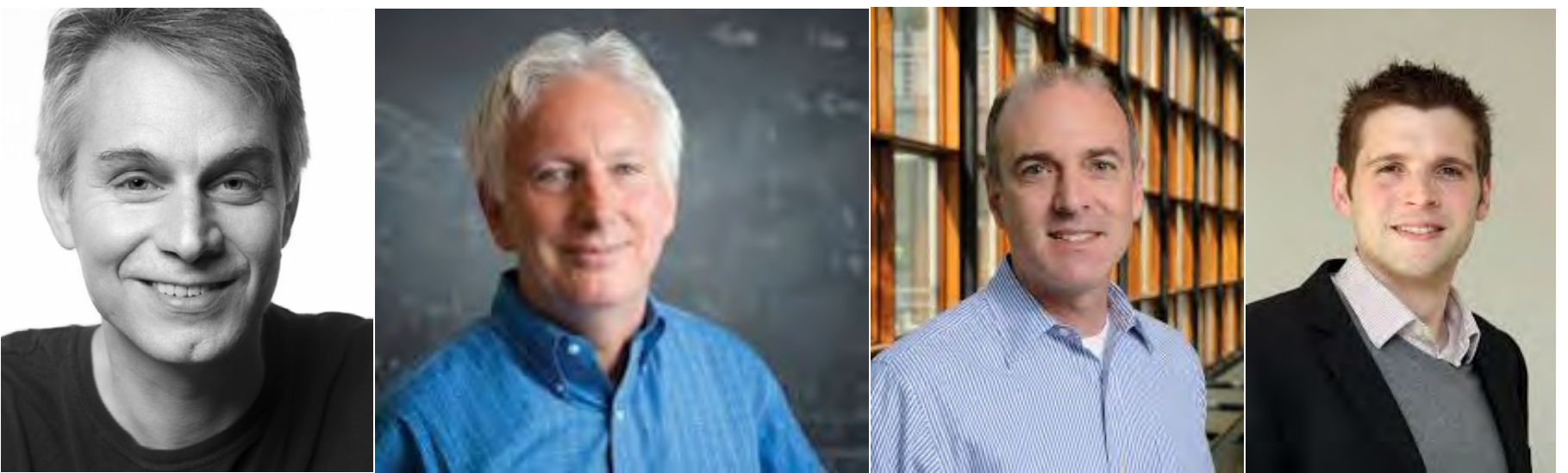
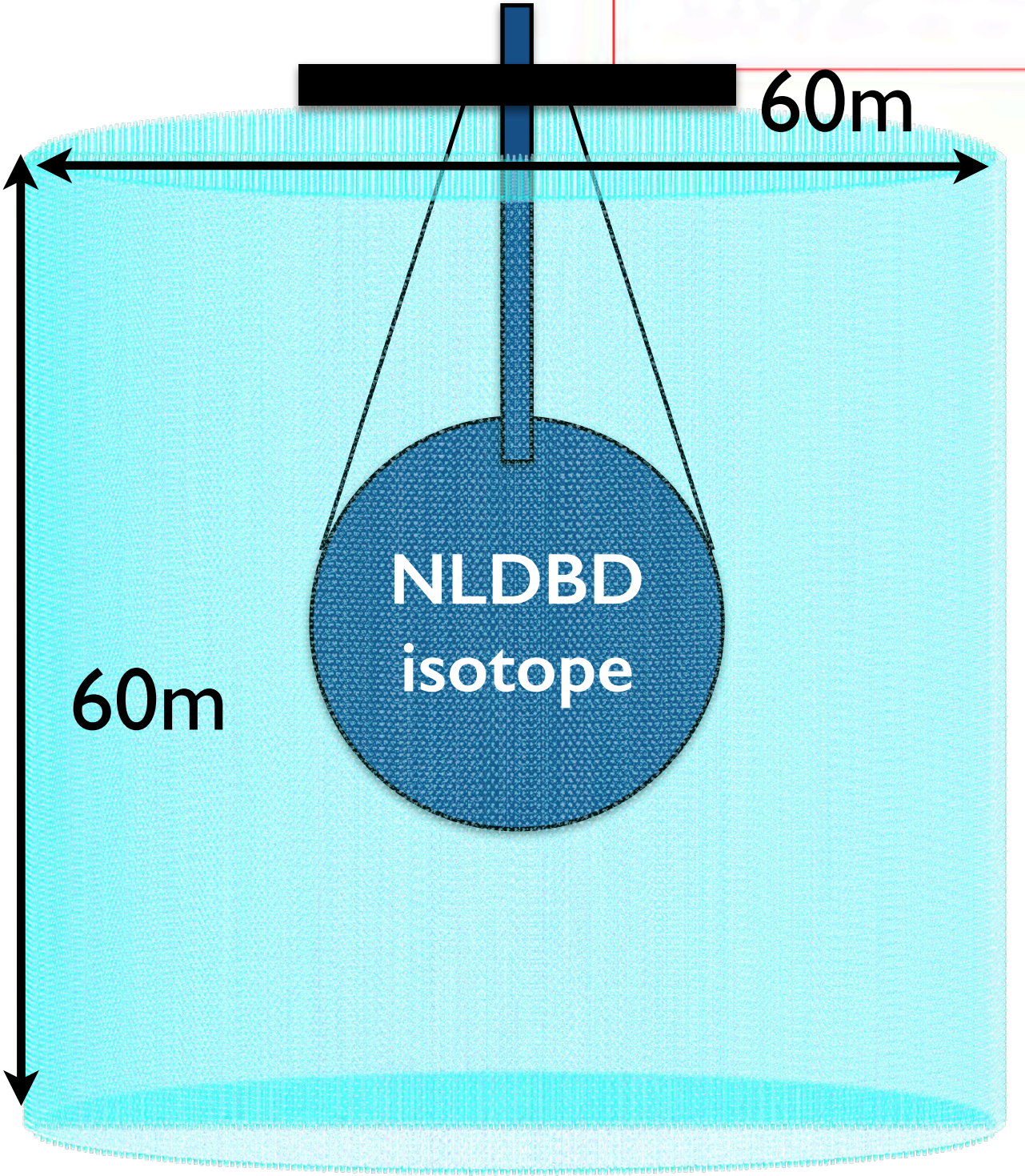
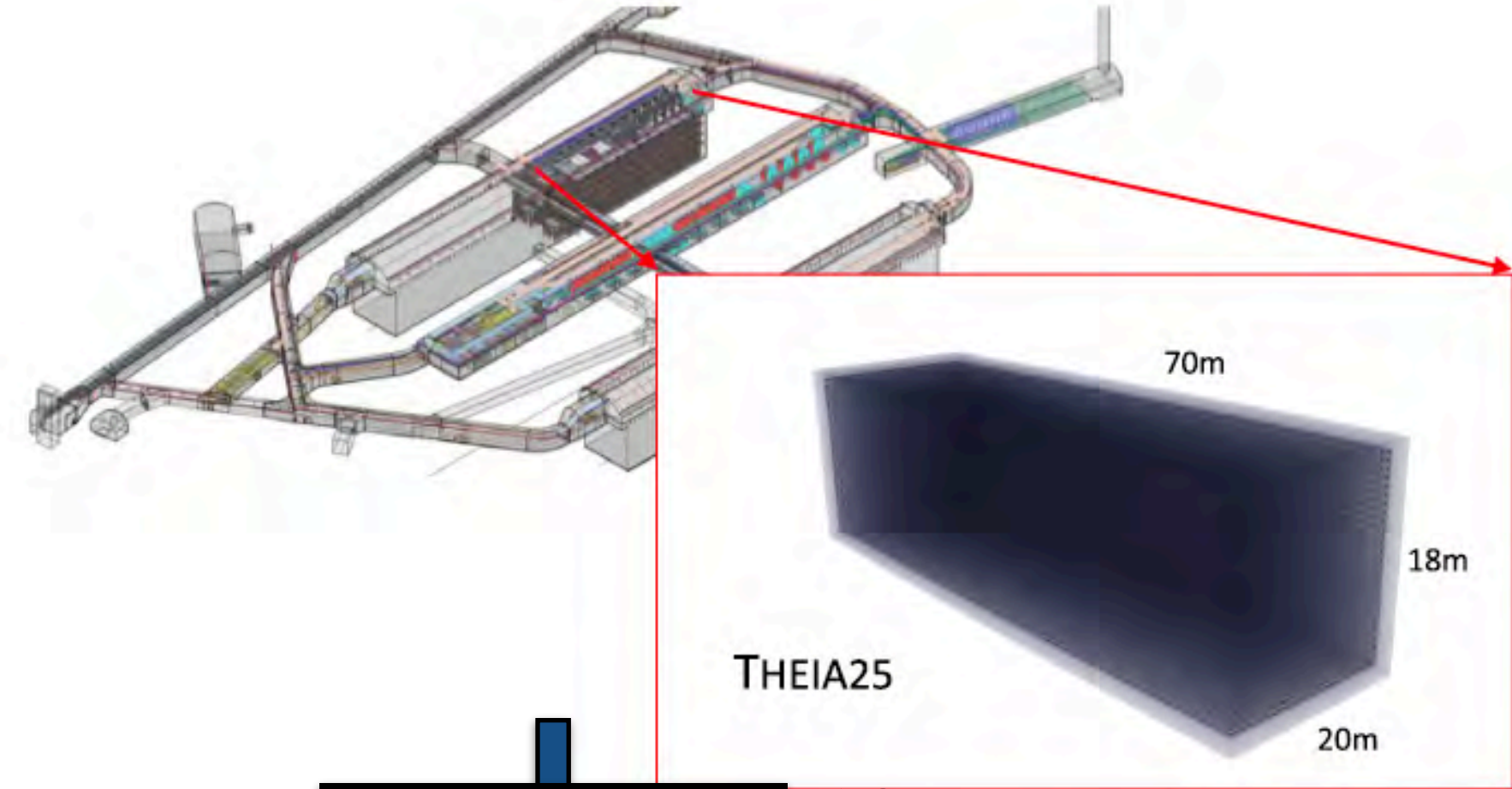
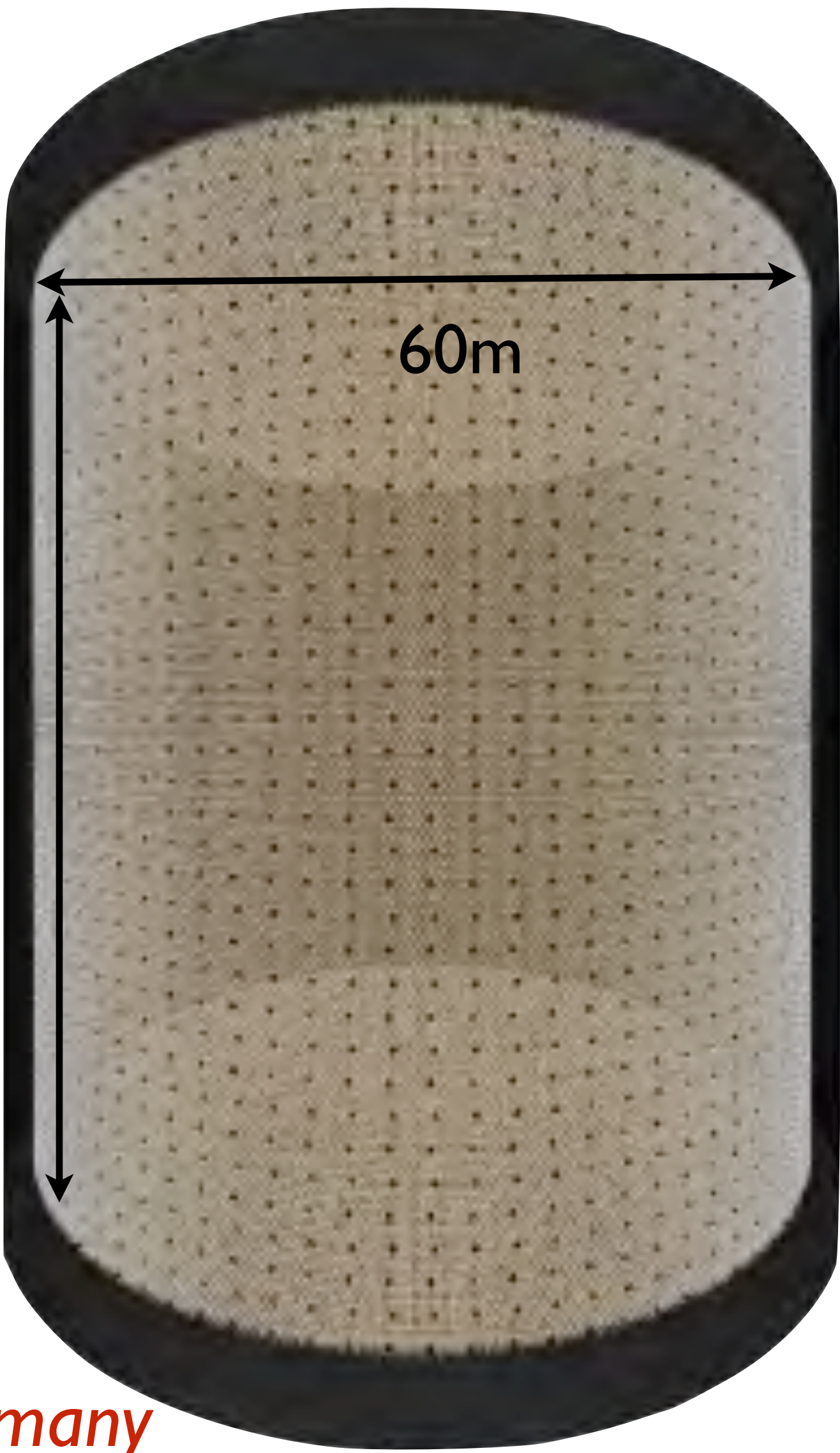




THEIA

- Hybrid Cherenkov/scintillation detector (25-100 kton)
- Novel LS target e.g. WbLS, slow LS
- Fast, high-efficiency spectrally sensitive photon detection with high coverage
- Isotope loading (Gd, Te, Li...)
- Flexible! Target, loading, configuration
- Broad physics program

60m



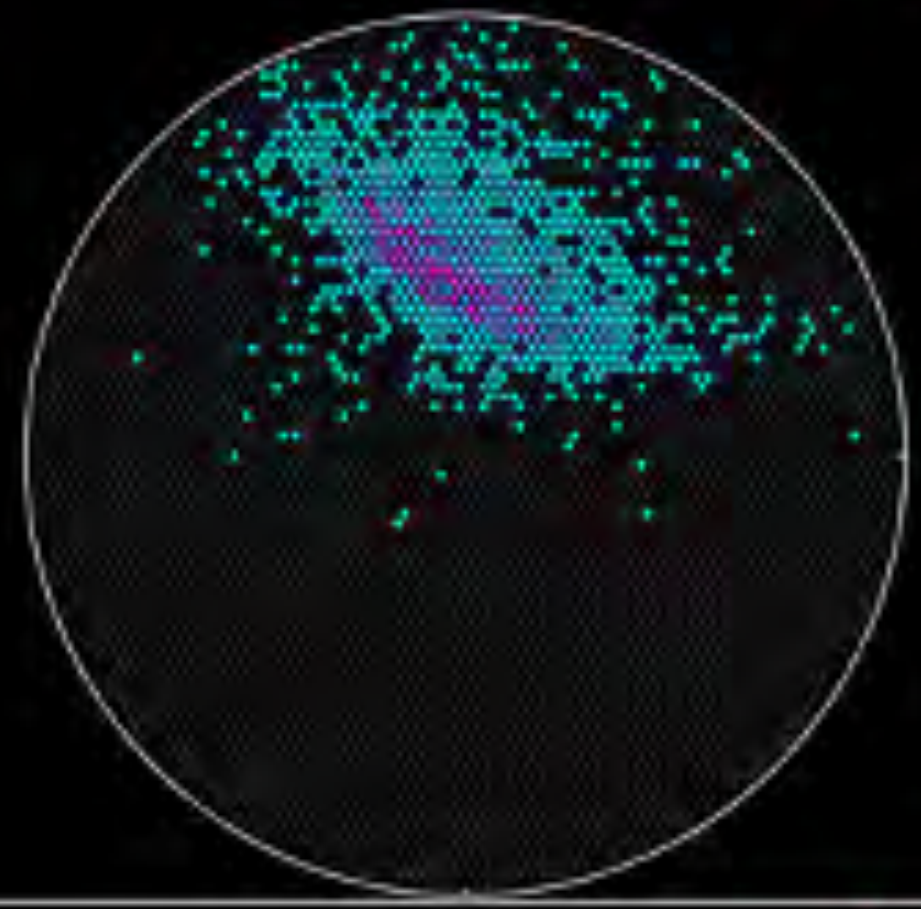
+ many

1 GeV
 e^-

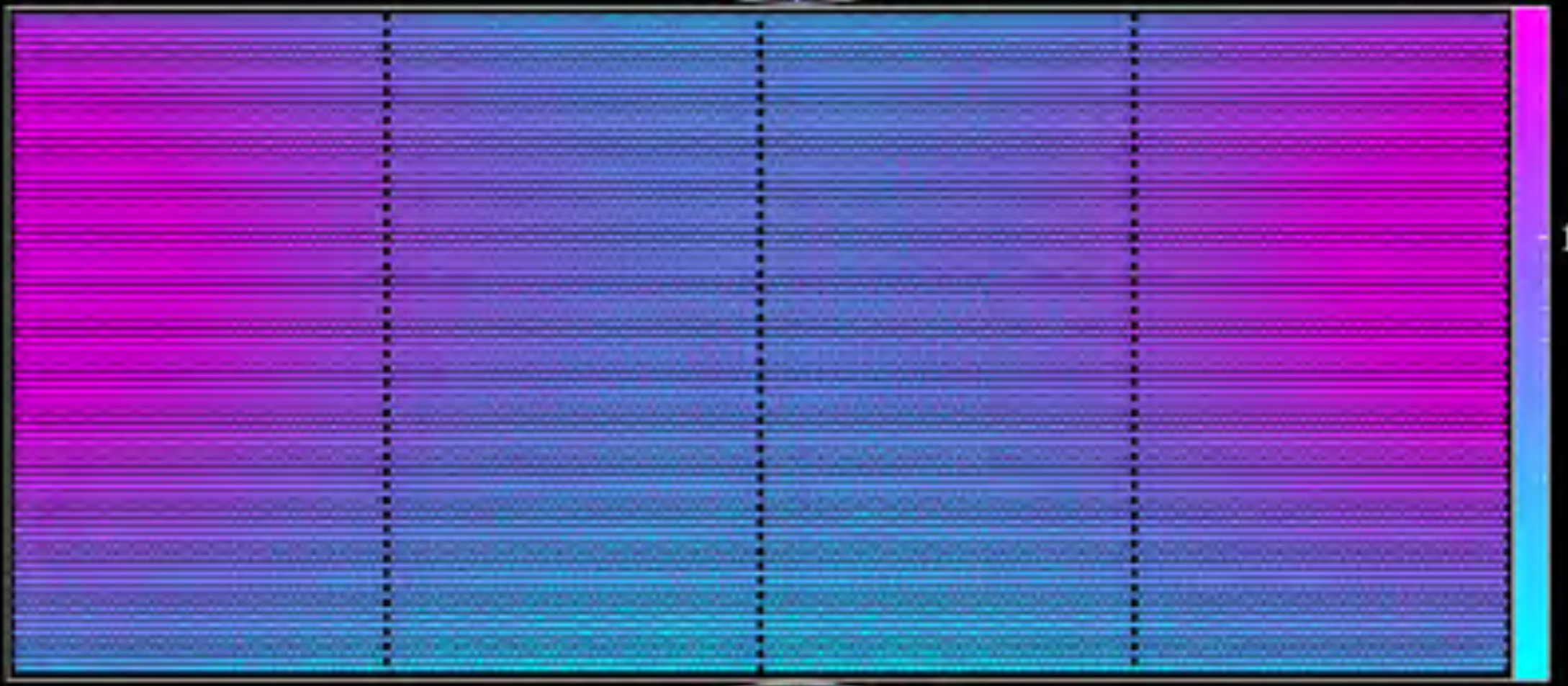
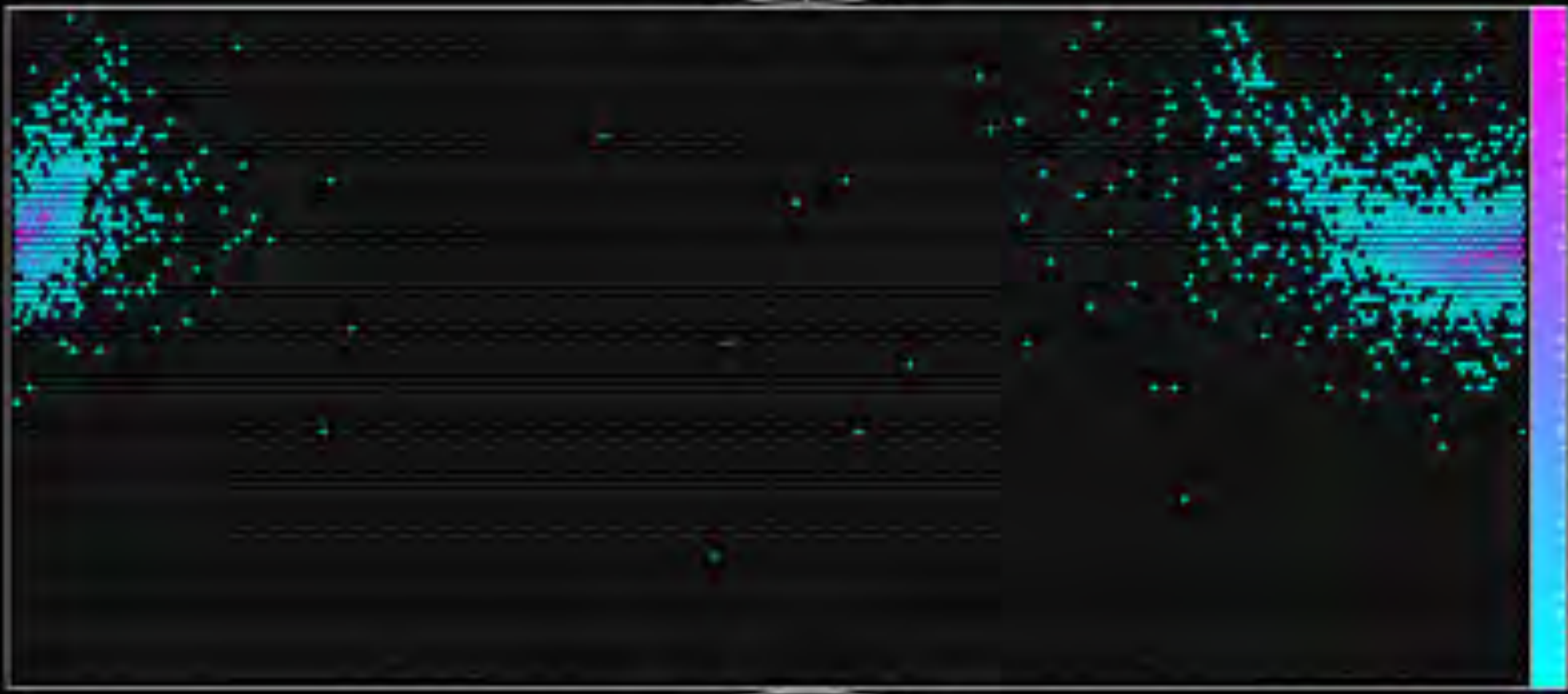
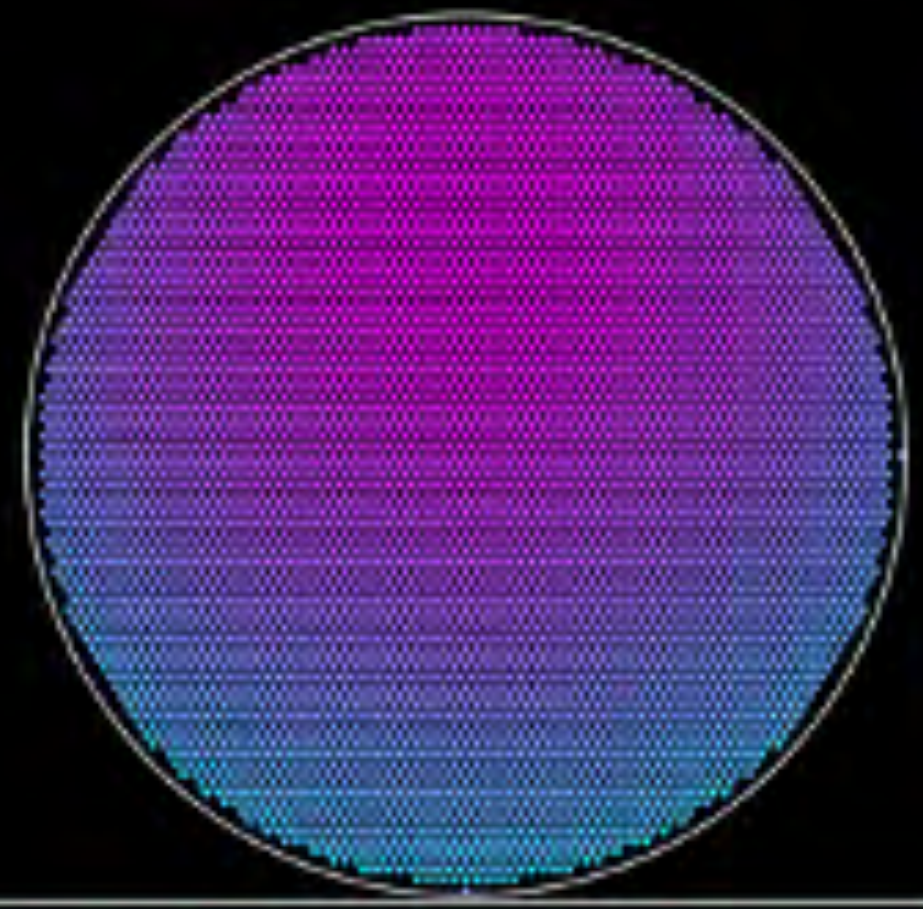
Long Wavelengths ($t_{resid} < 3.0$ ns)

1 GeV
 e^-

Short Wavelengths (All)



Pure LS +
dichroicons

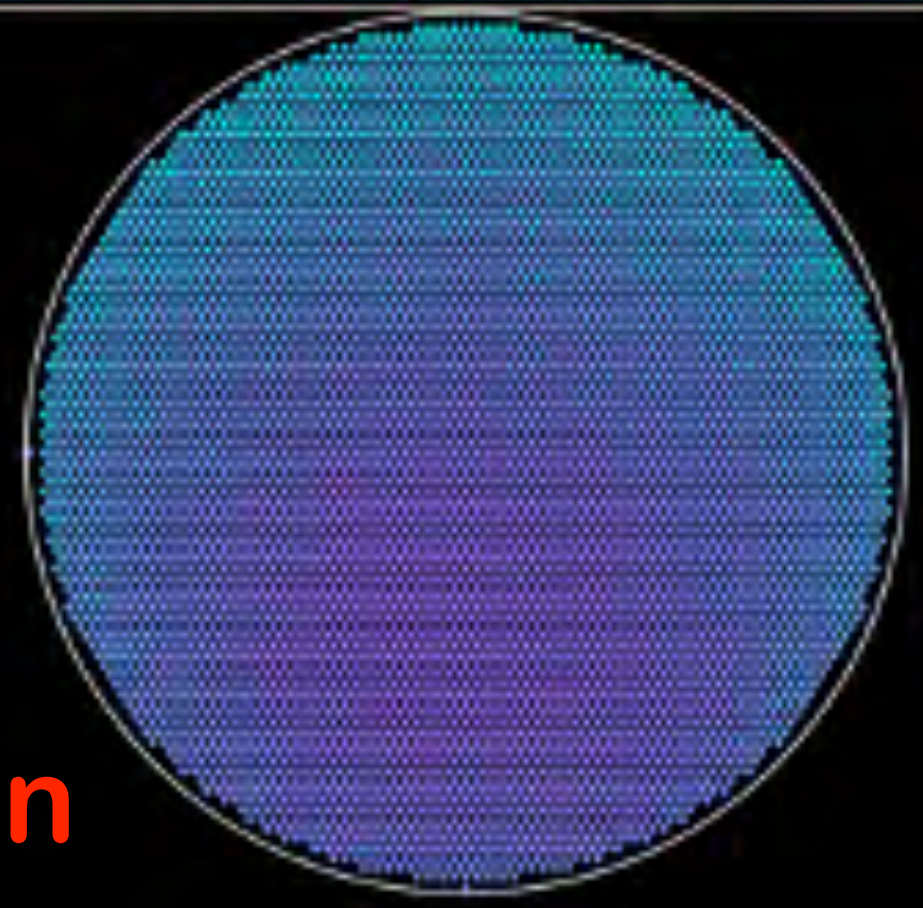
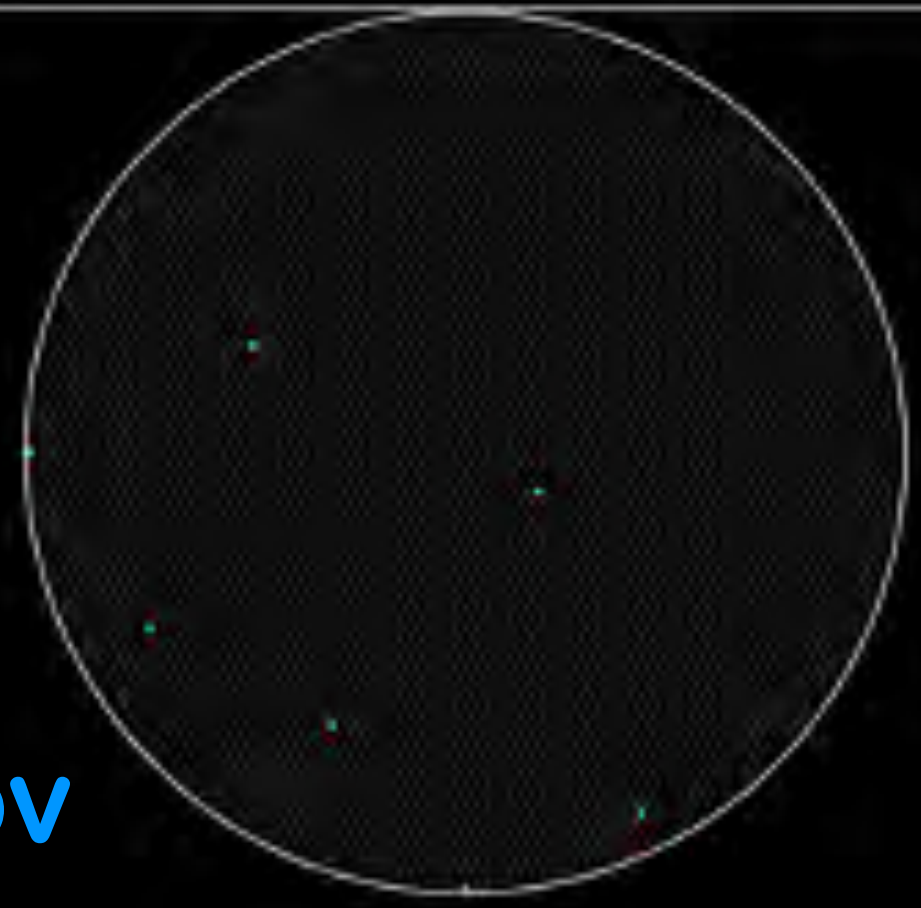


2-for-1

Credit: Sam Young, Penn

Cherenkov

Scintillation



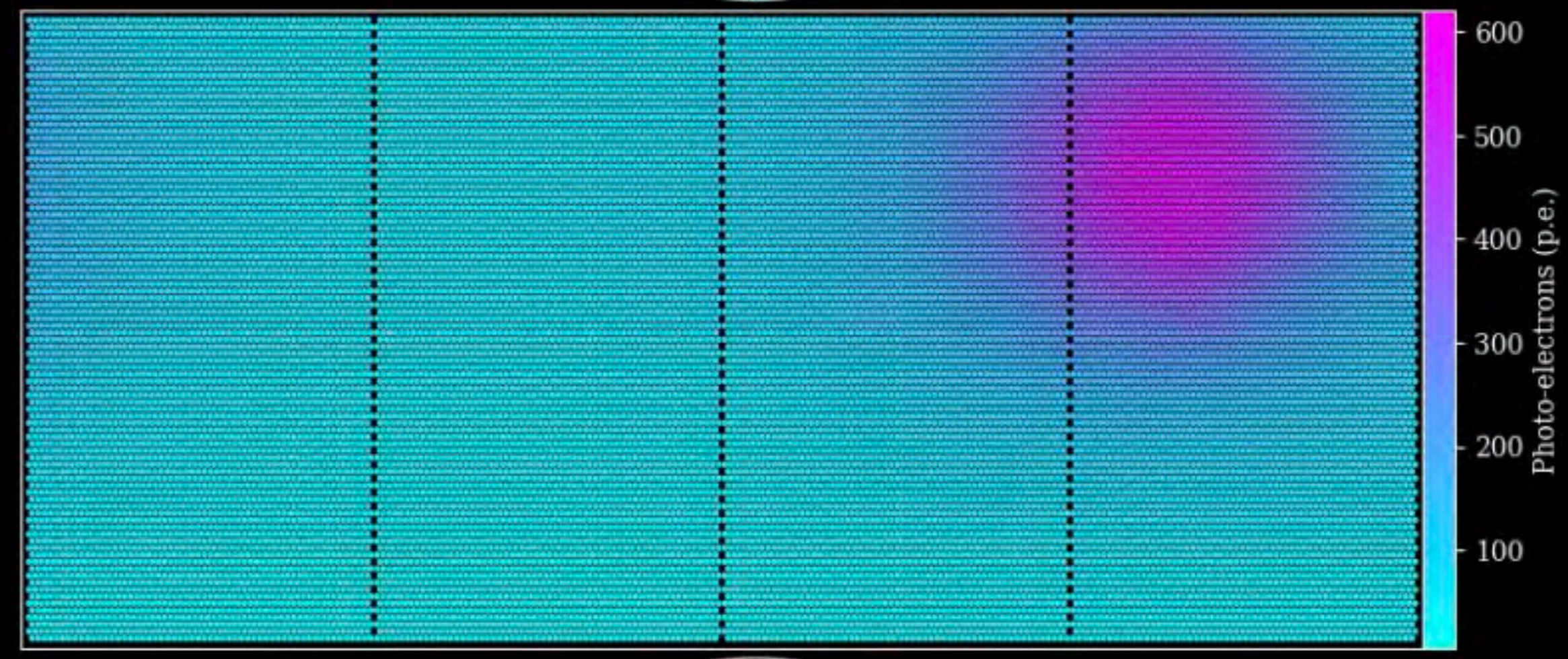
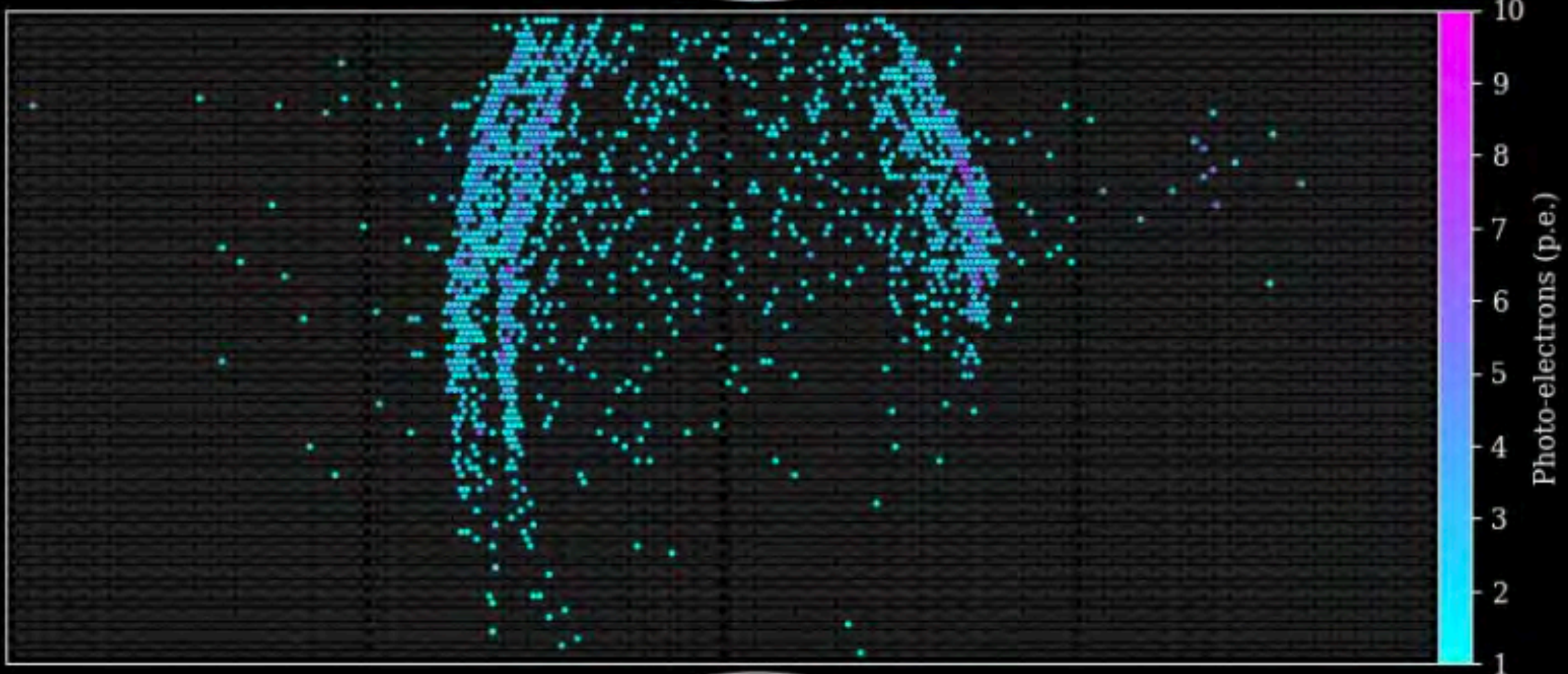
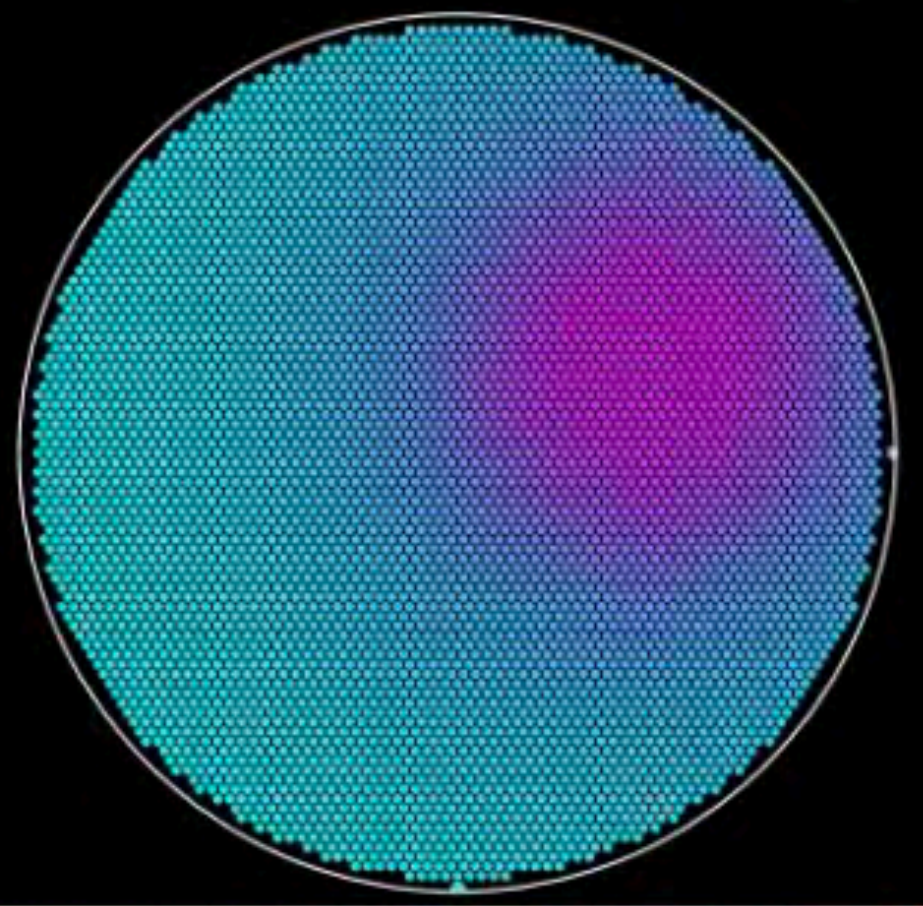
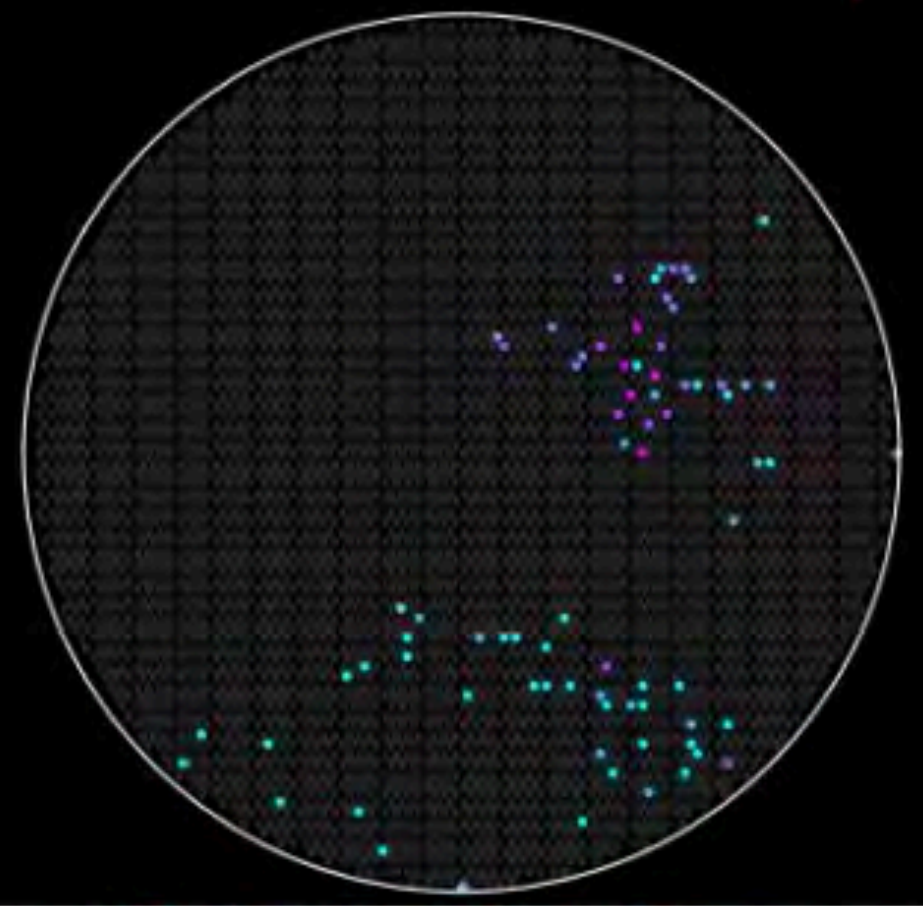
1 GeV
 μ^-

Long Wavelengths ($t_{\text{resid}} < 3.0$ ns)

1 GeV
 μ^-

Short Wavelengths (All)

2-for-1



Credit: Sam Young, Penn

Cherenkov

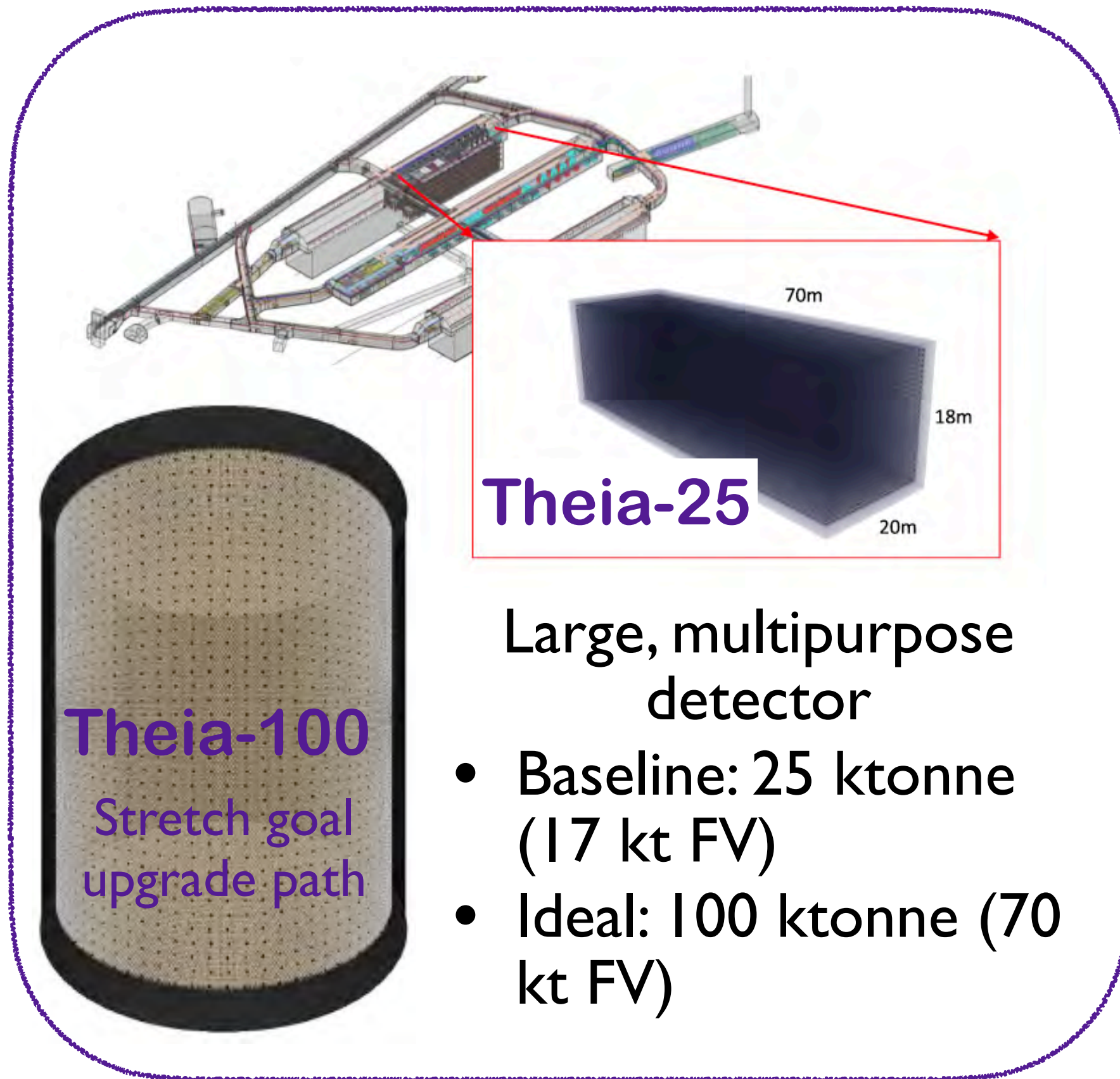
Scintillation

Theia as a DUNE module in Phase II

Long-baseline sensitivity comparable to a LAr DUNE module

Complementary supernova sensitivity (primarily anti- ν , fast response: can act as trigger)

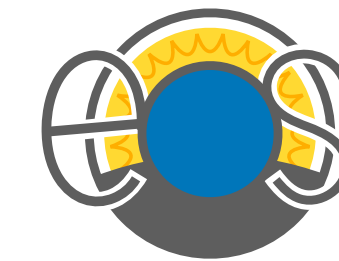
+ broad (new!) additional physics program



- DUNE Phase II formal process includes Theia as 1 of 3 options
- Theia is technically mature, and brings a broad physics program beyond any alternative (LAr) tech.
- Strong international team actively engaged
- Current R&D support from HEP, NNSA; LDRD at BNL to study ND requirements
- Technical demonstrators underway (BNL 30t, Eos @ LBL, ANNIE, BUTTON)

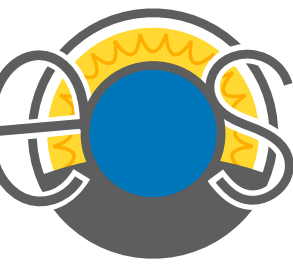


THEIA physics program





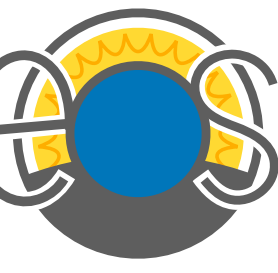
THEIA physics program



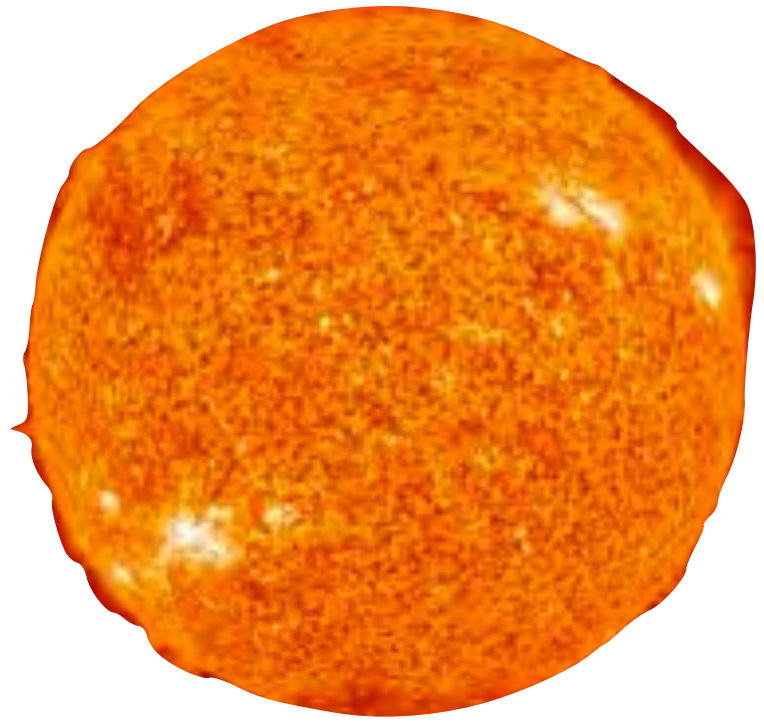
Neutrinos as a probe of nature



THEIA physics program

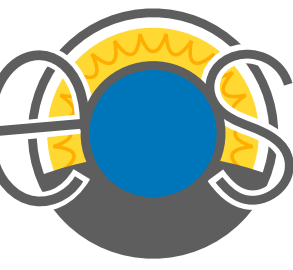


Neutrinos as a probe of nature

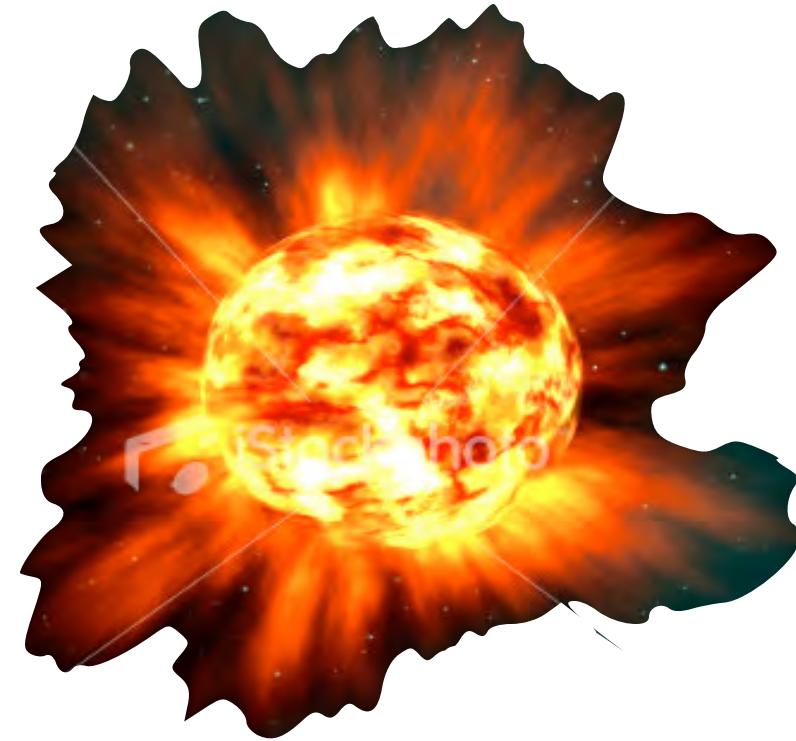
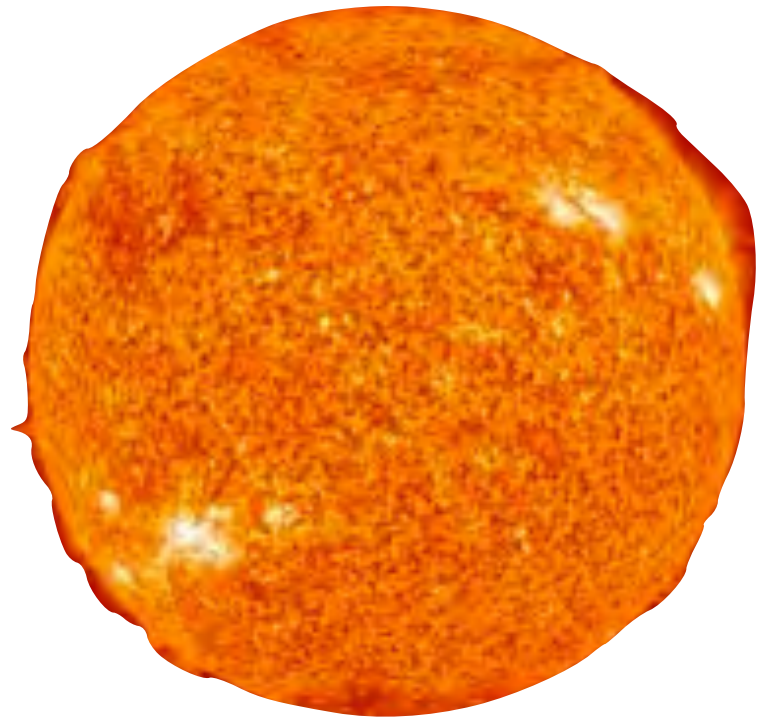




THEIA physics program

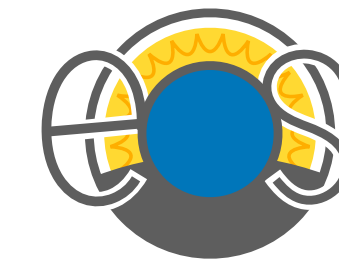


Neutrinos as a probe of nature

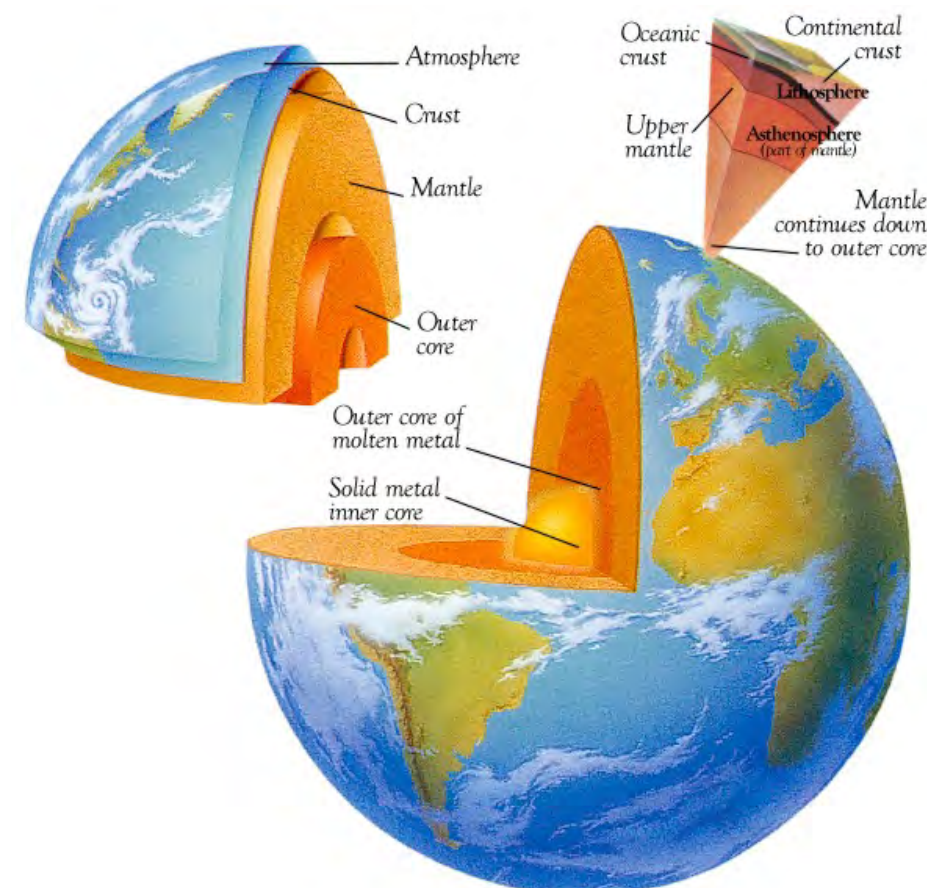
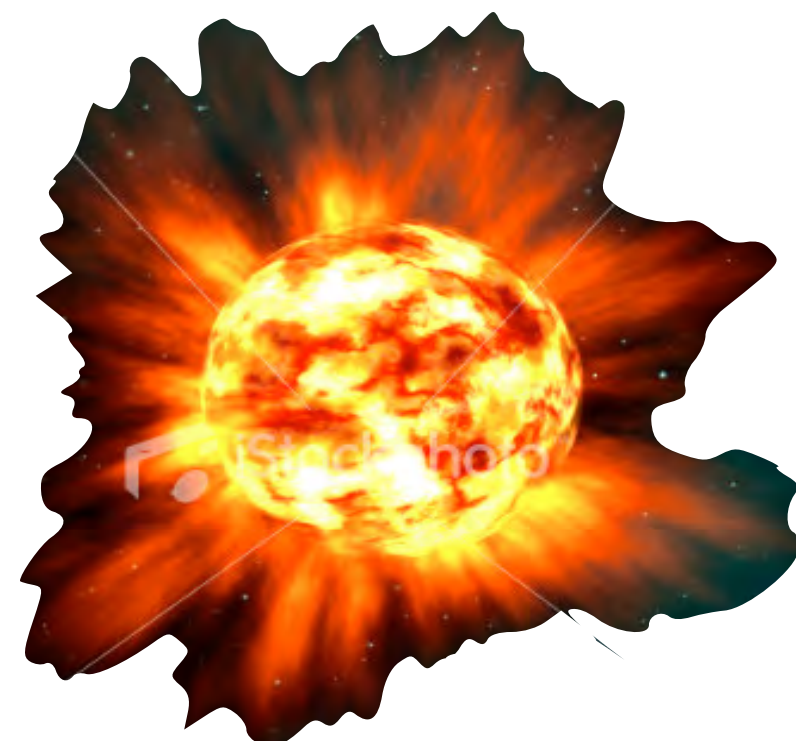
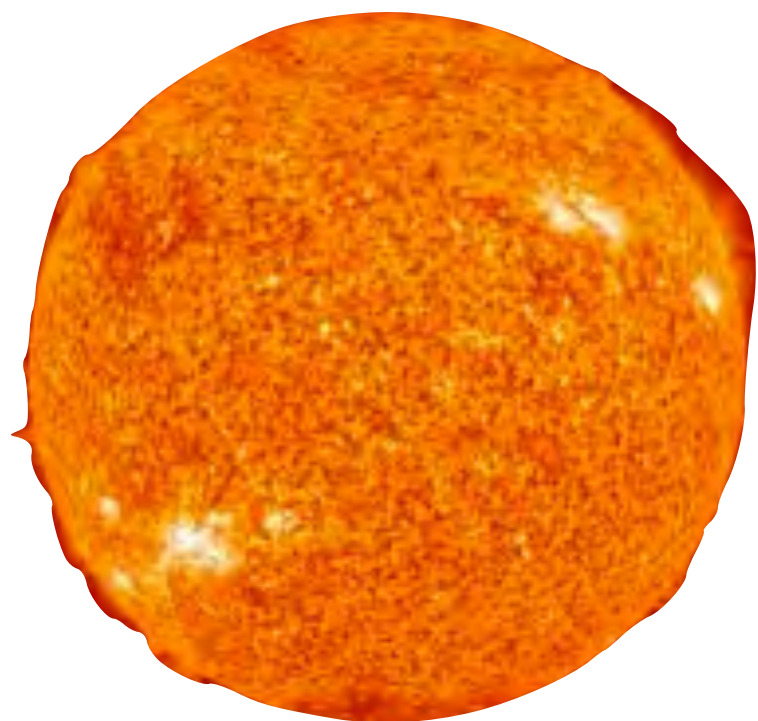




THEIA physics program

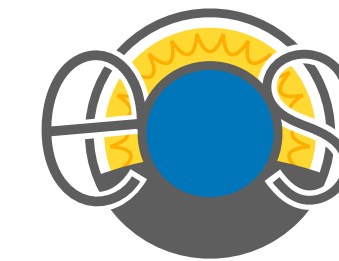


Neutrinos as a probe of nature

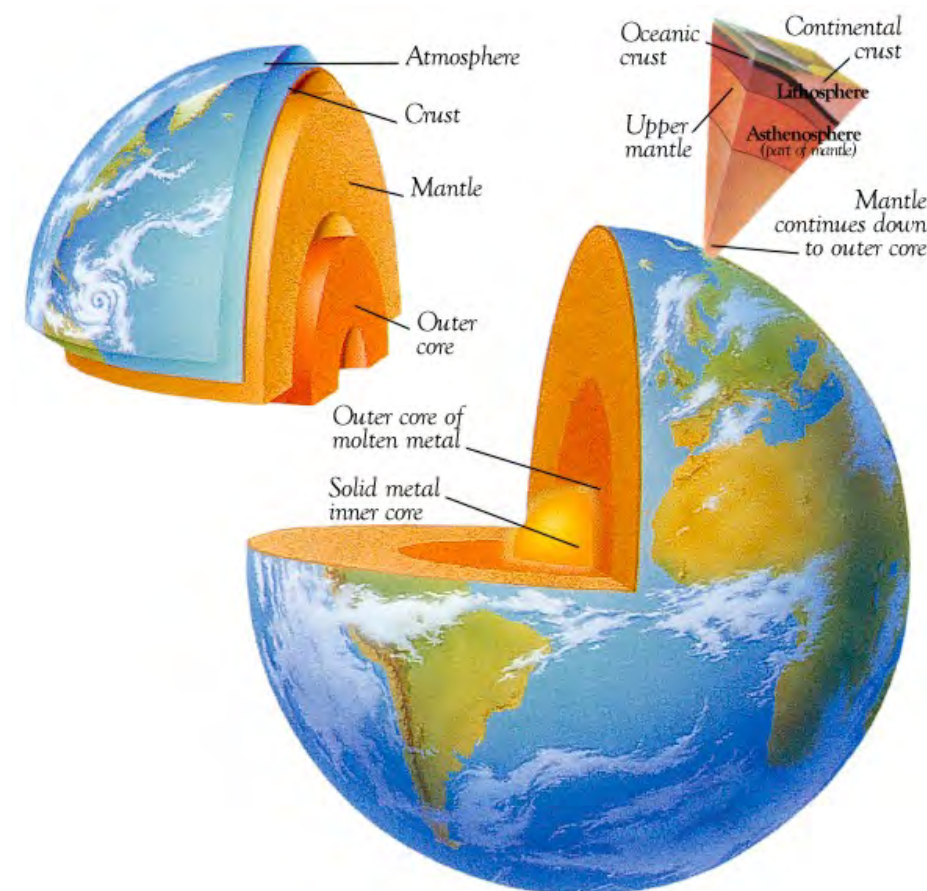
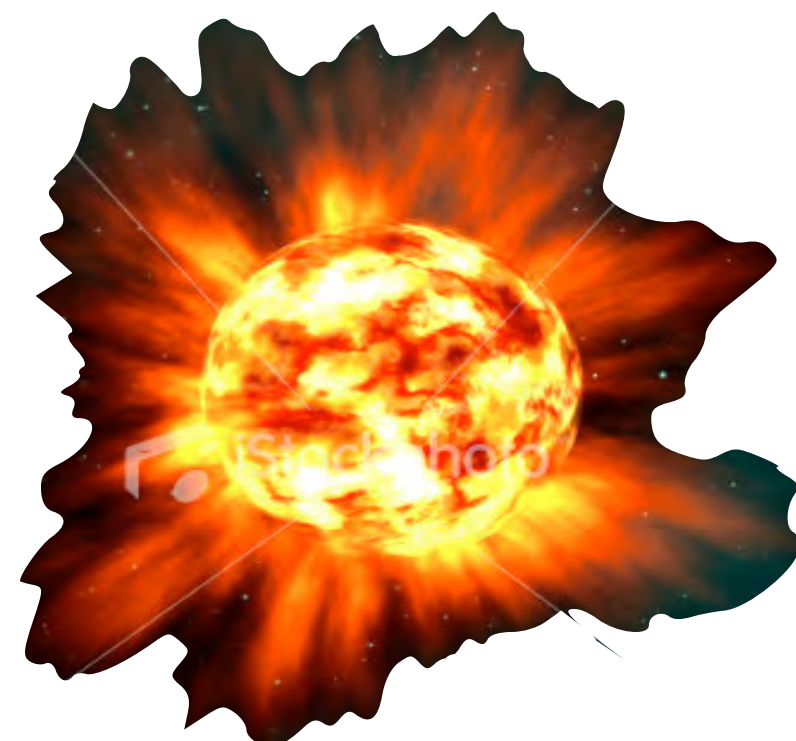
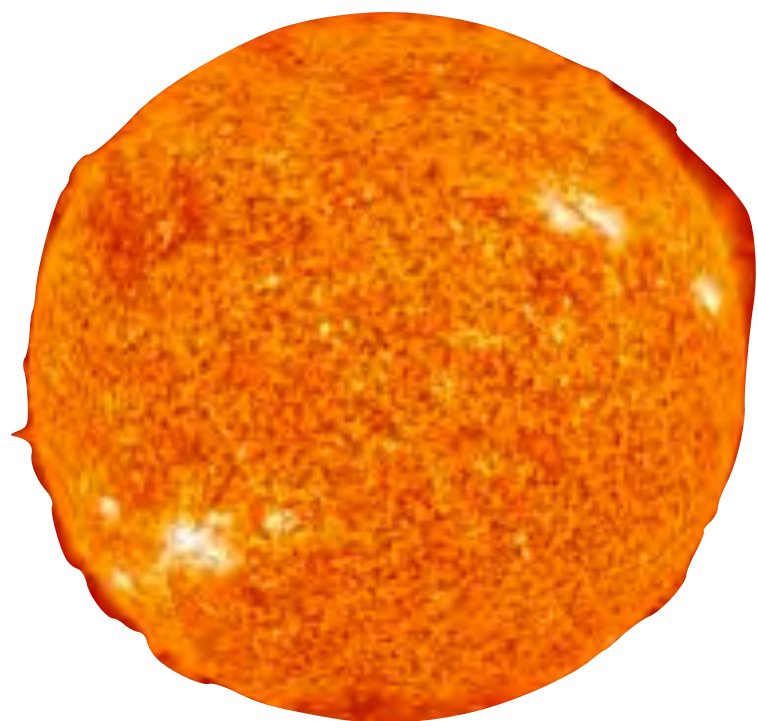




THEIA physics program



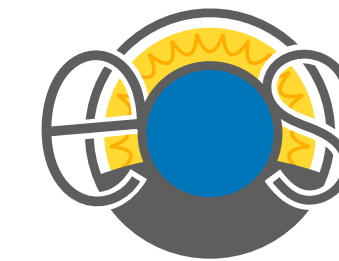
Neutrinos as a probe of nature



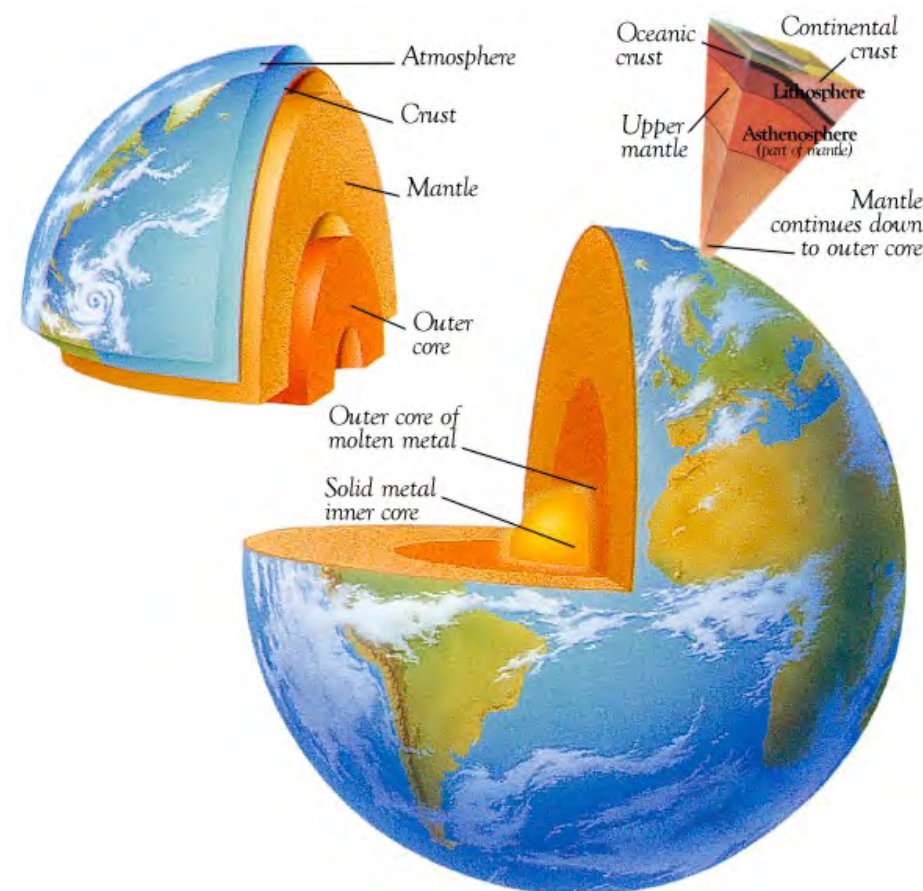
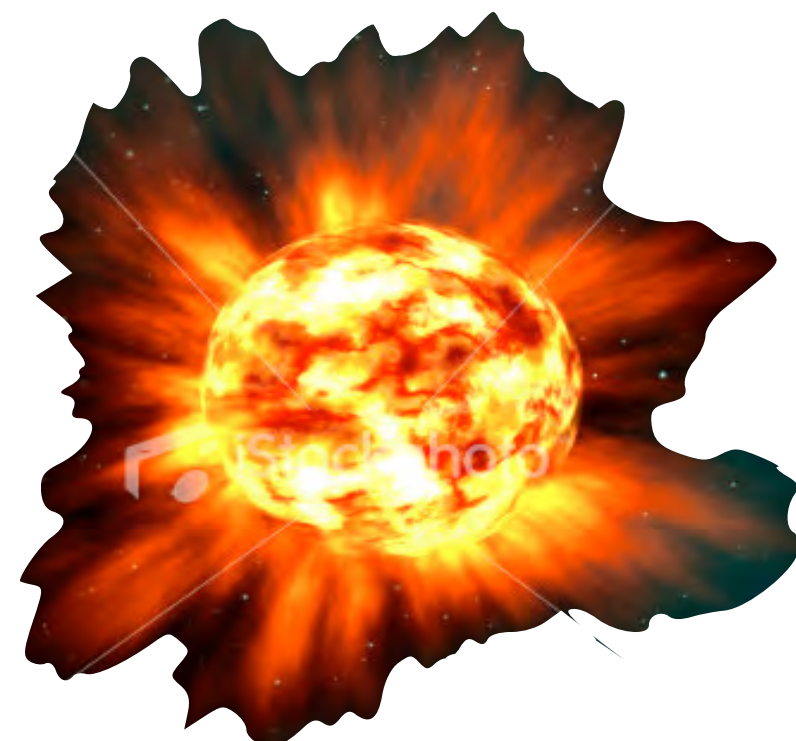
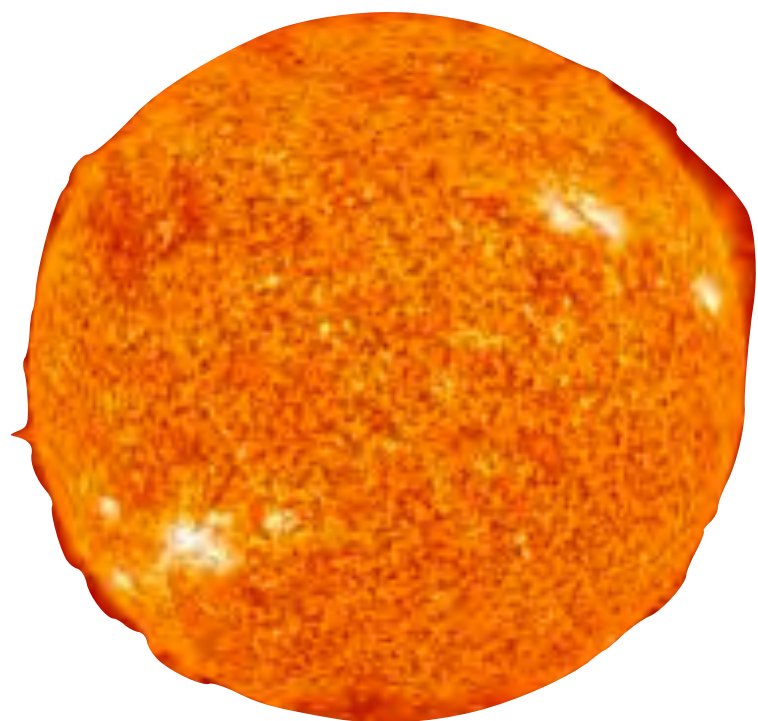
Studying the fundamental nature of matter



THEIA physics program



Neutrinos as a probe of nature



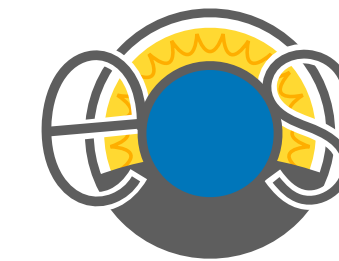
Studying the fundamental nature of matter

$0\nu\beta\beta$

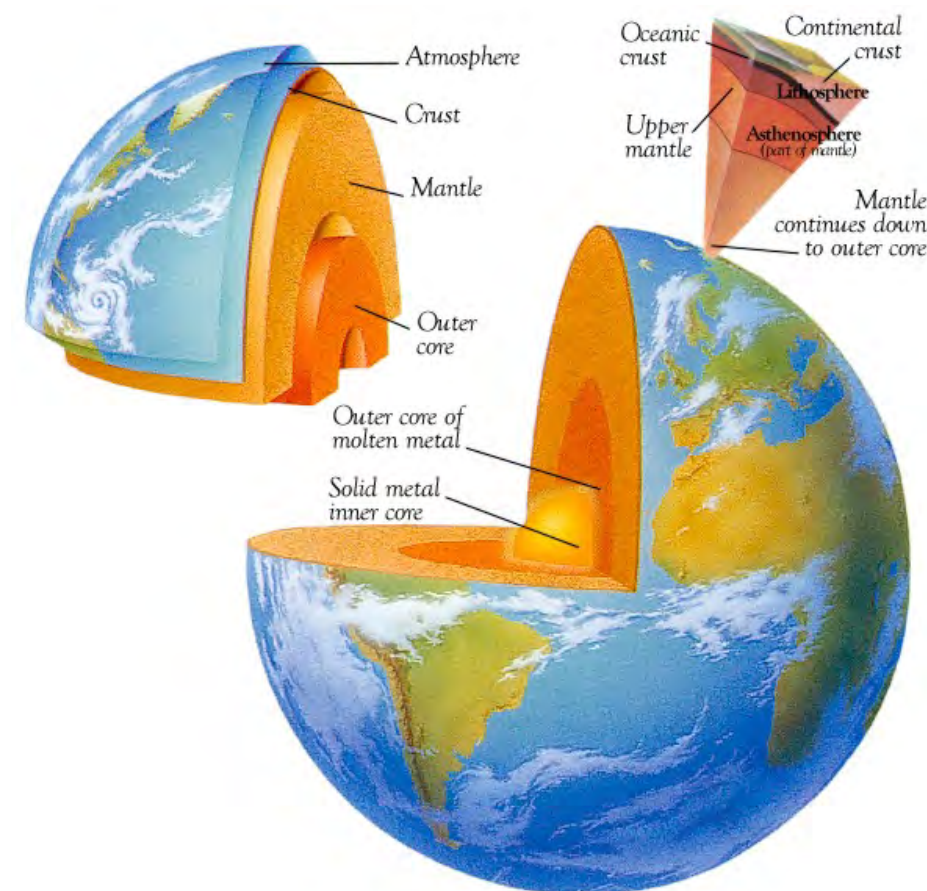
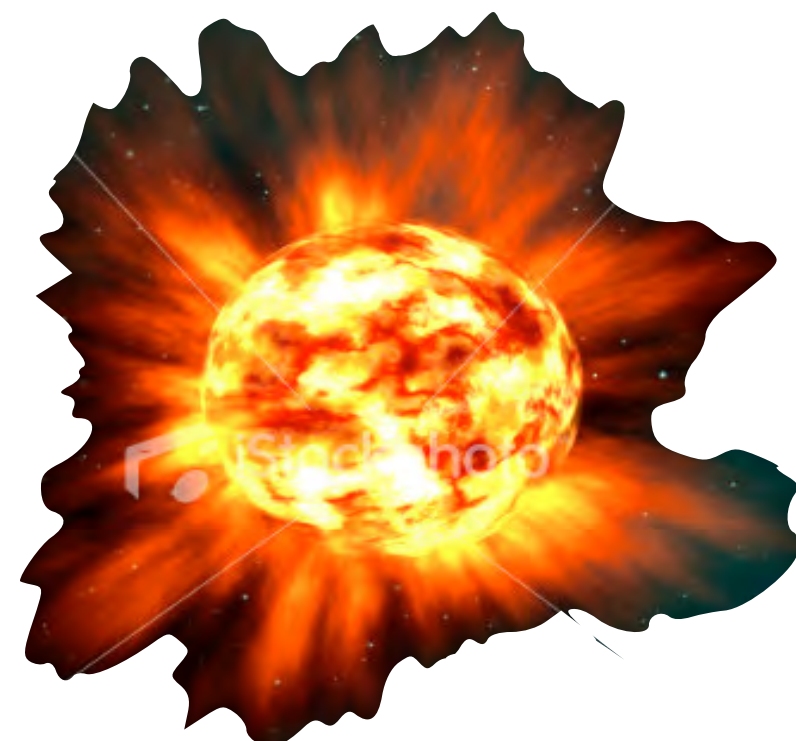
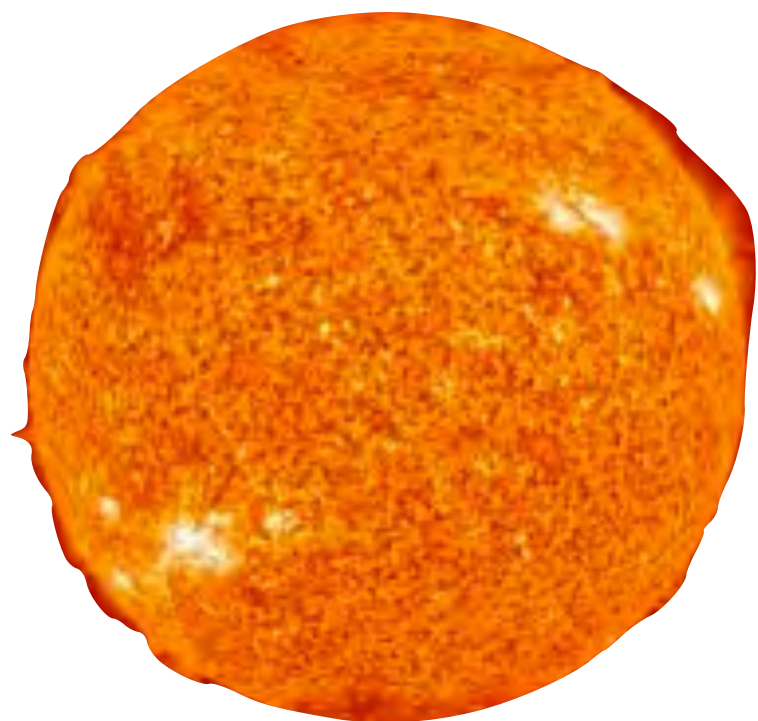
$${}^A_Z X \rightarrow {}^A_{Z+2} X + 2e^-$$



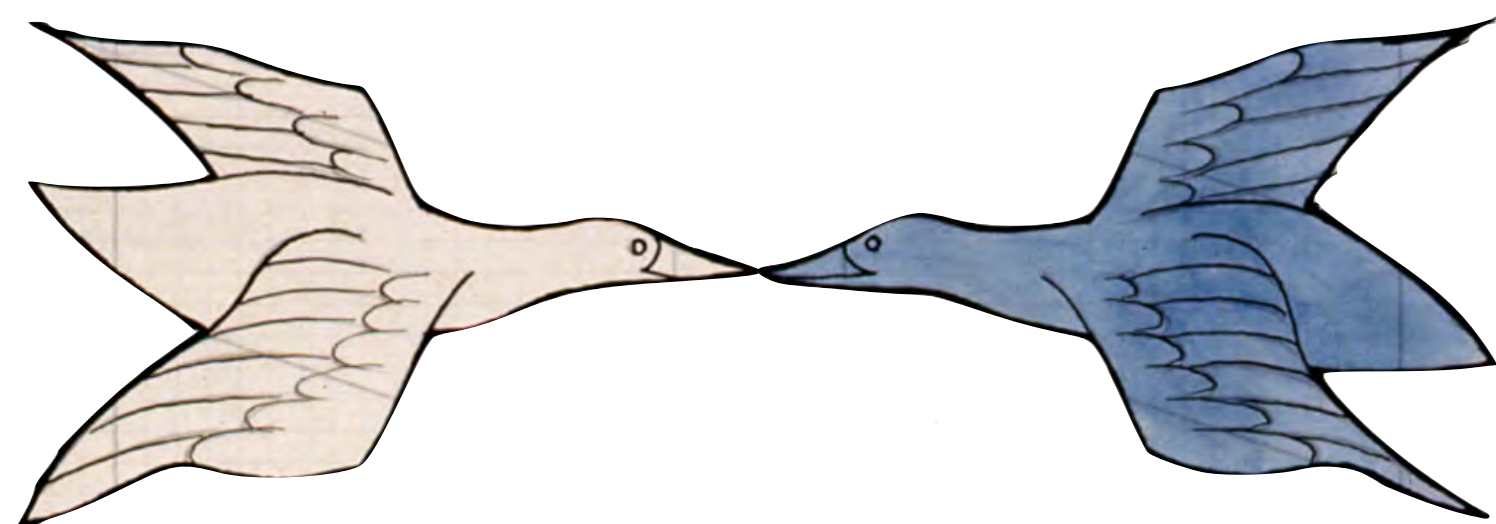
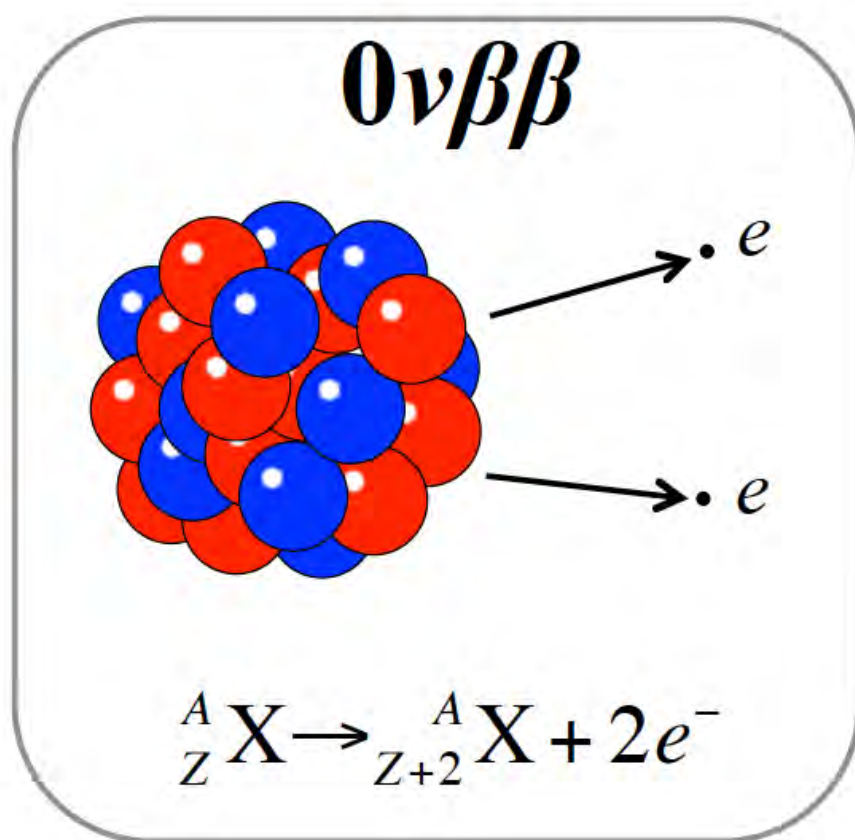
THEIA physics program



Neutrinos as a probe of nature

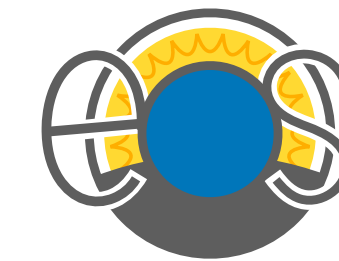


Studying the fundamental nature of matter

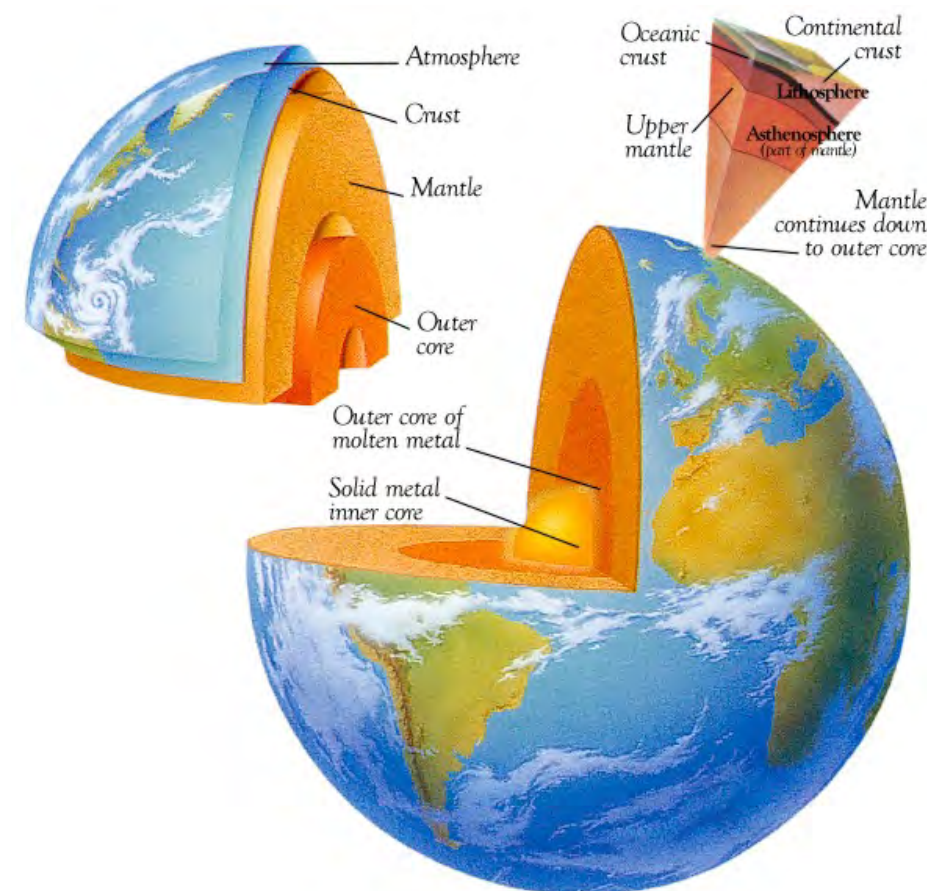
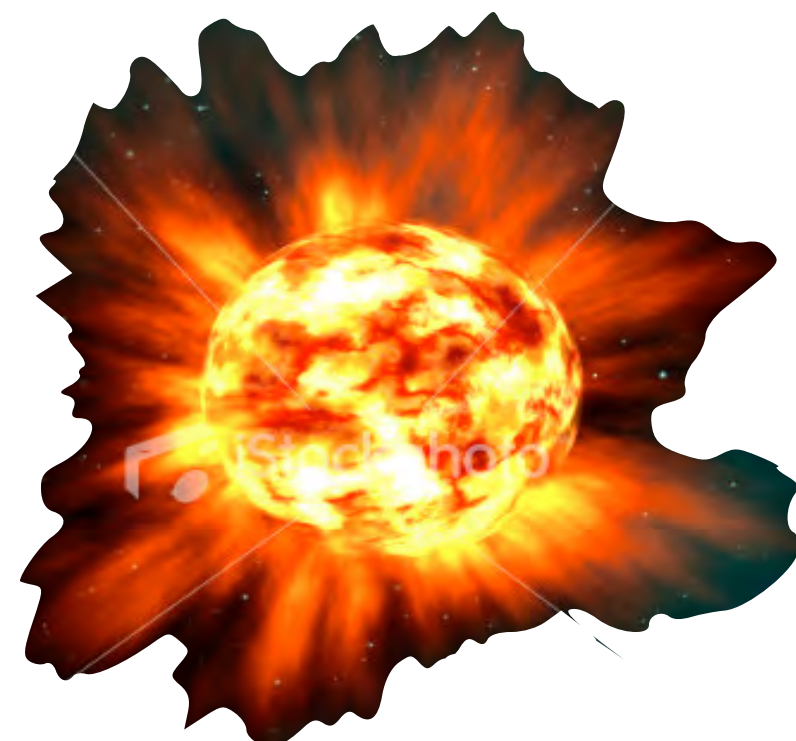
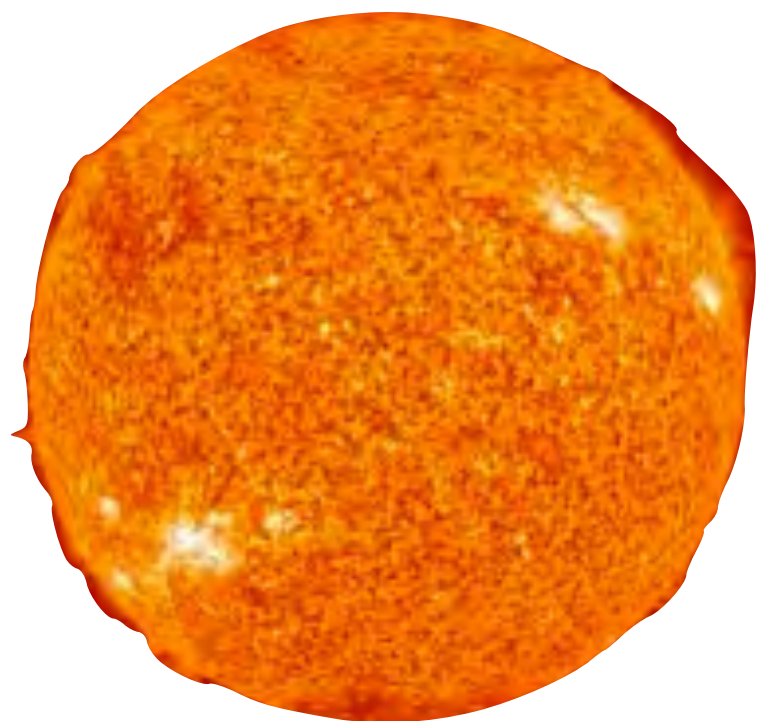




THEIA physics program

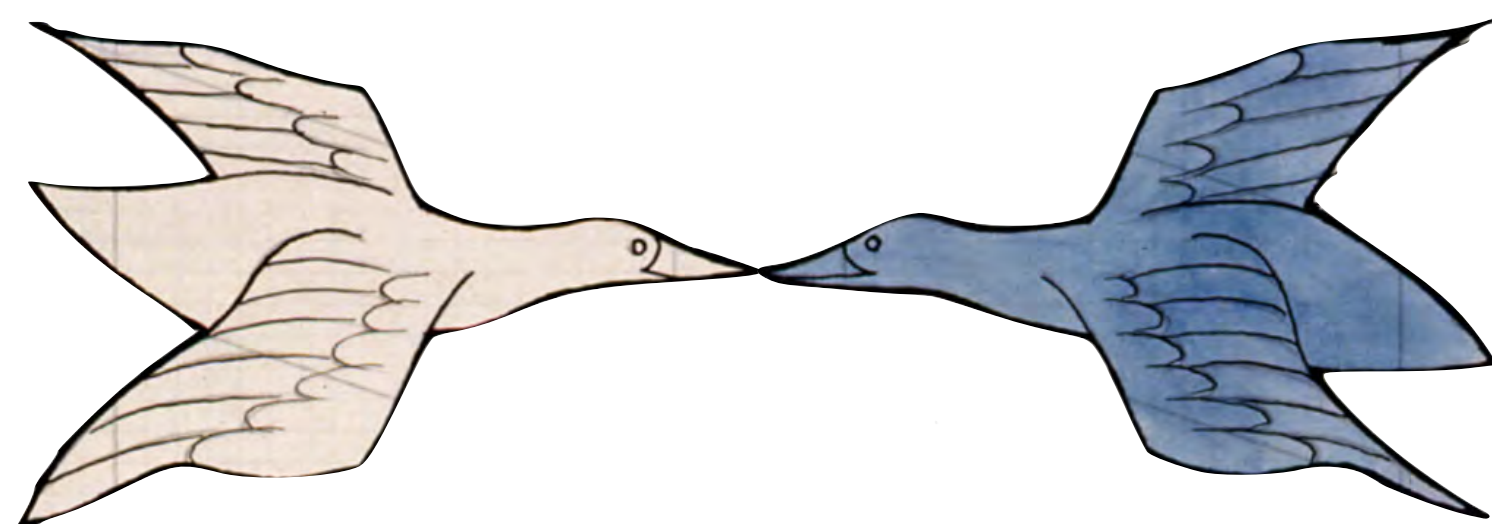
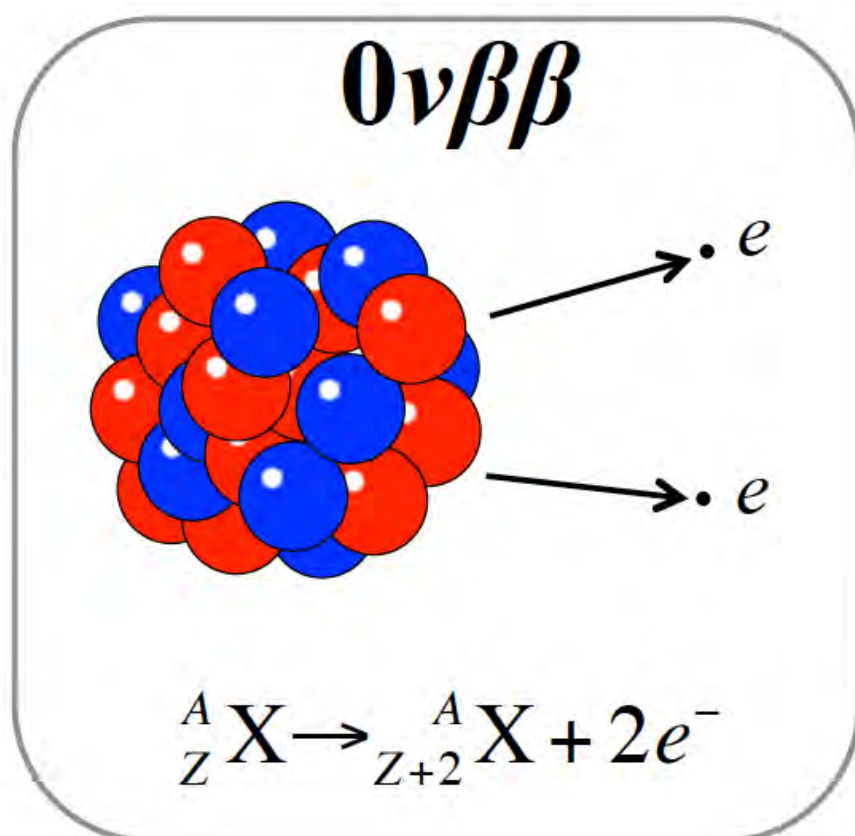


Neutrinos as a probe of nature



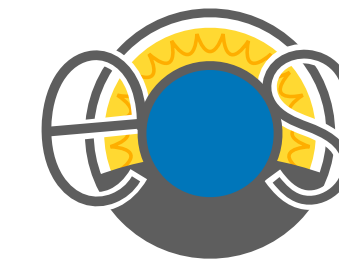
Particle astrophysics

Studying the fundamental nature of matter

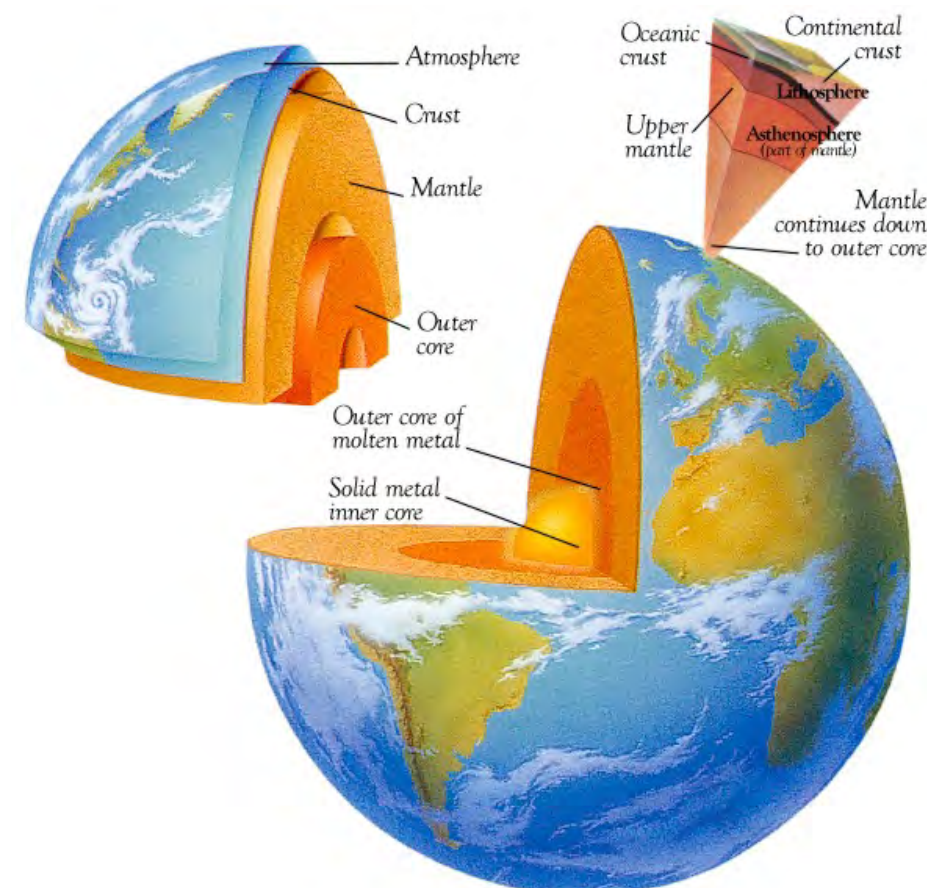
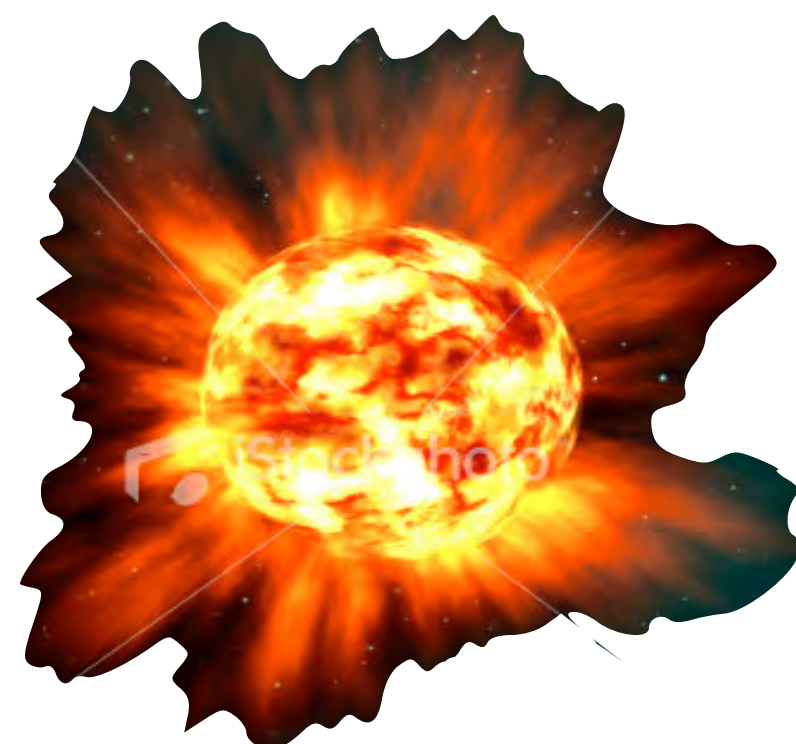
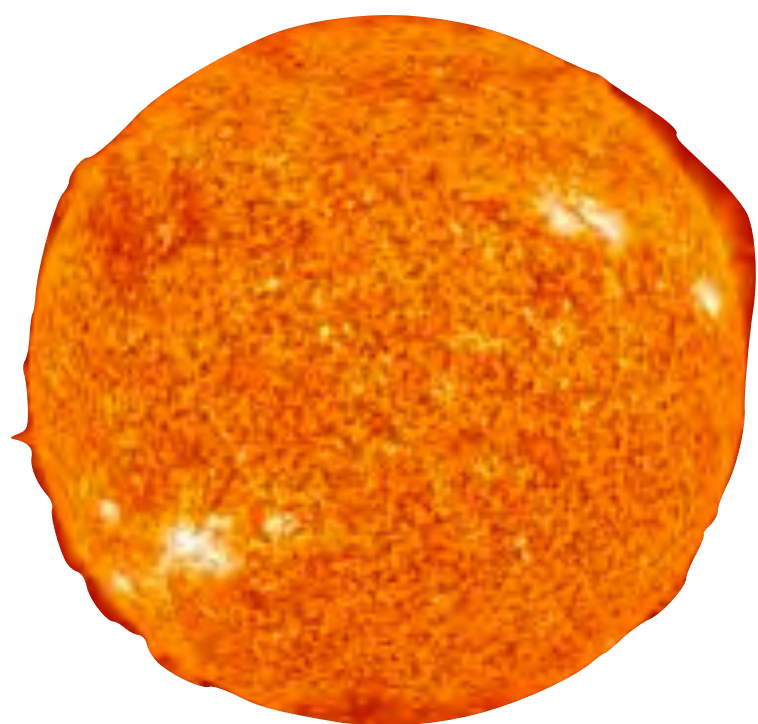




THEIA physics program



Neutrinos as a probe of nature



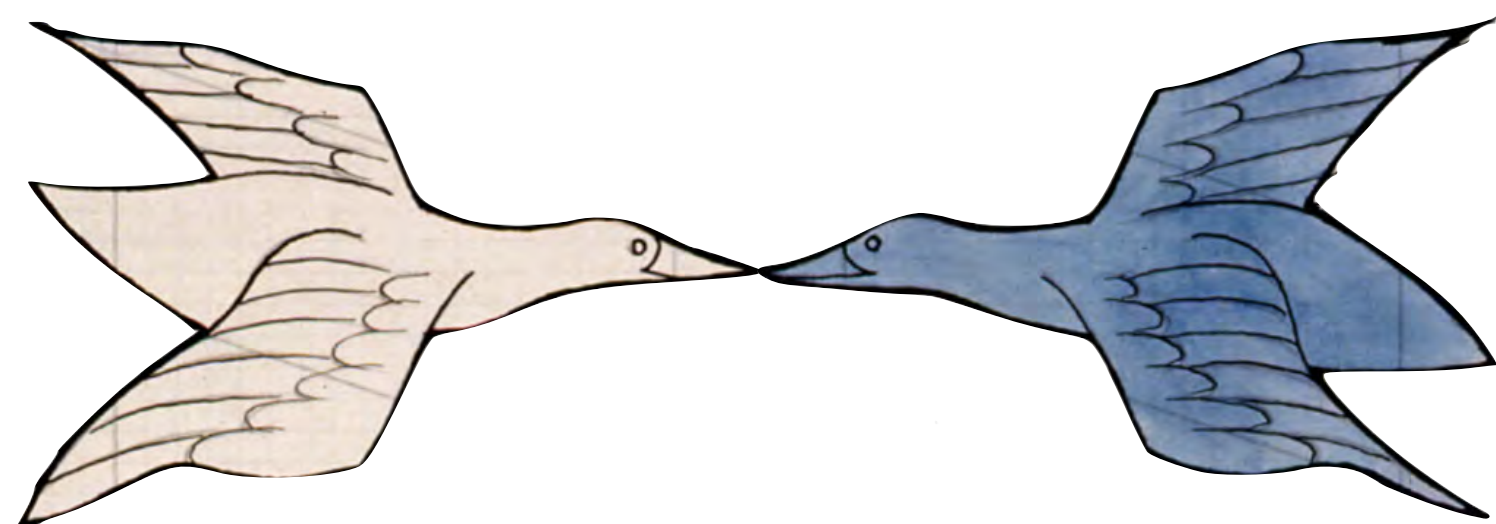
Particle astrophysics

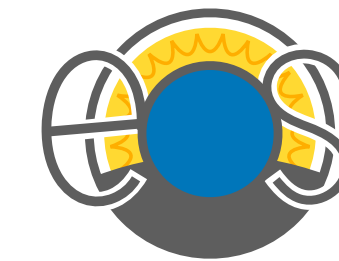
Nuclear physics

Studying the fundamental nature of matter

$0\nu\beta\beta$

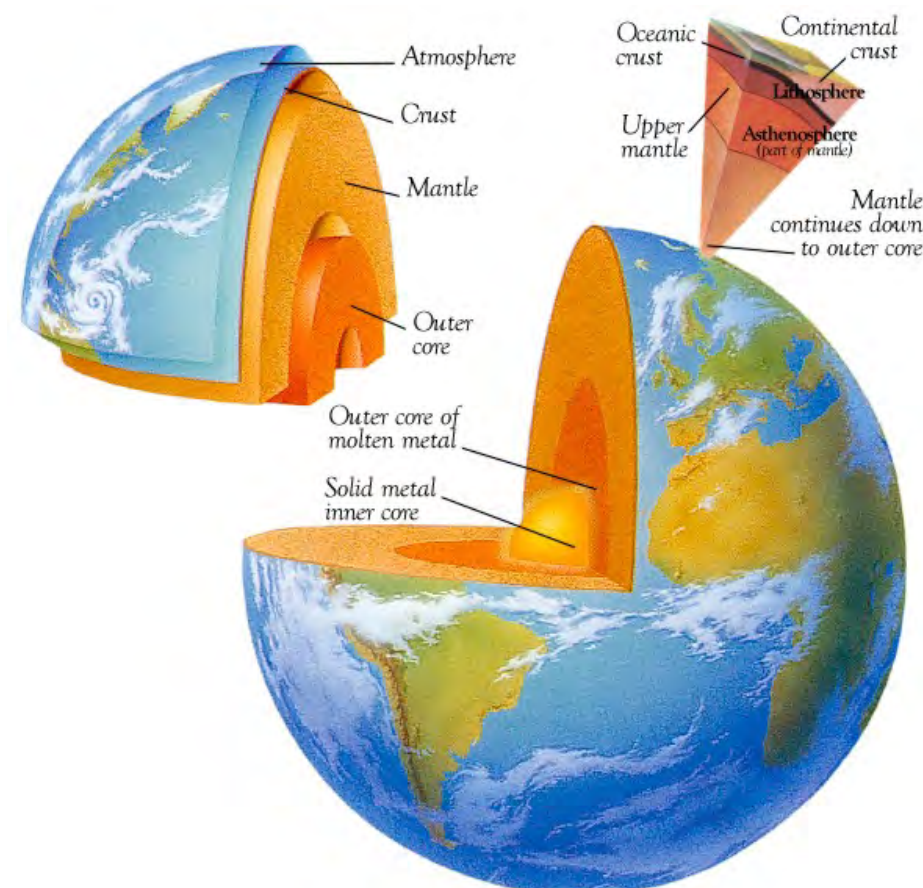
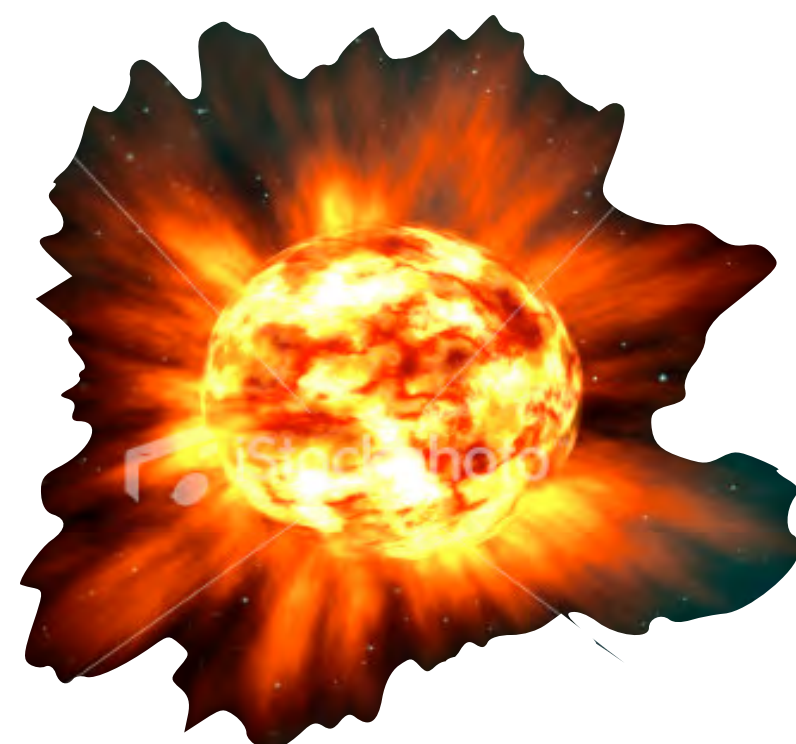
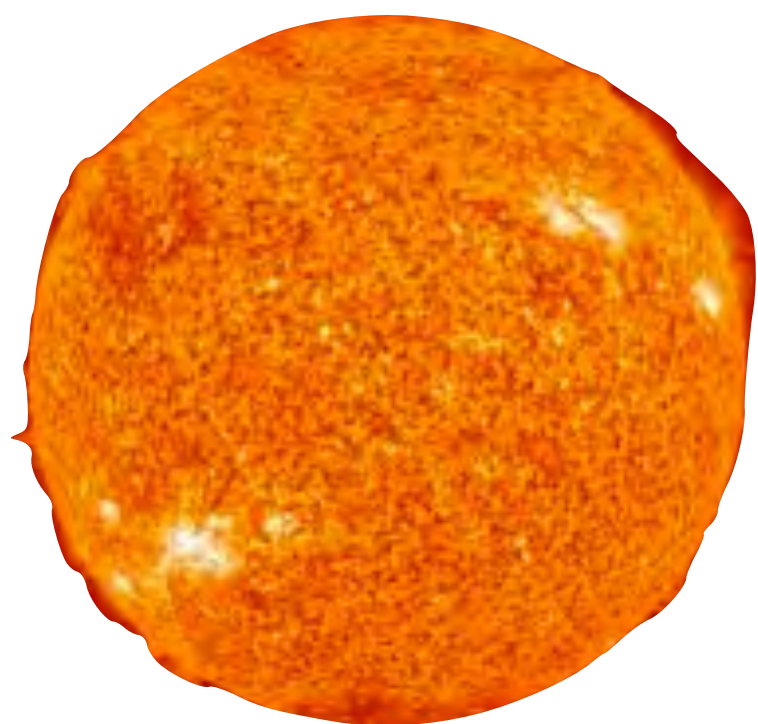
${}^A_Z X \rightarrow {}^A_{Z+2} X + 2e^-$





THEIA physics program

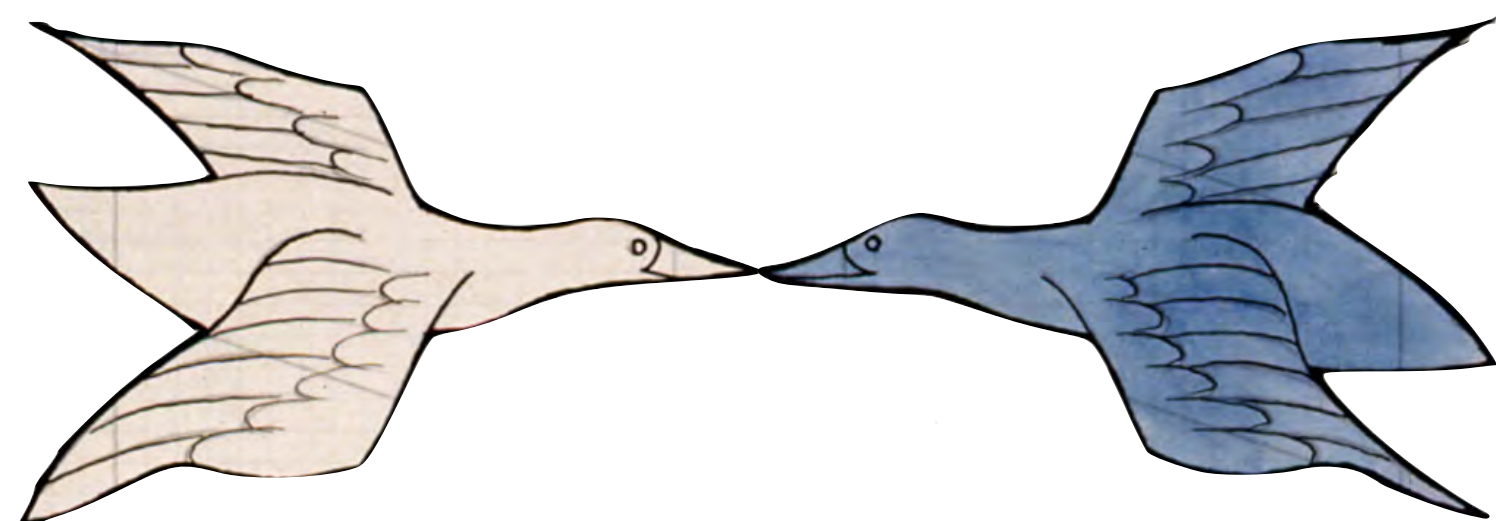
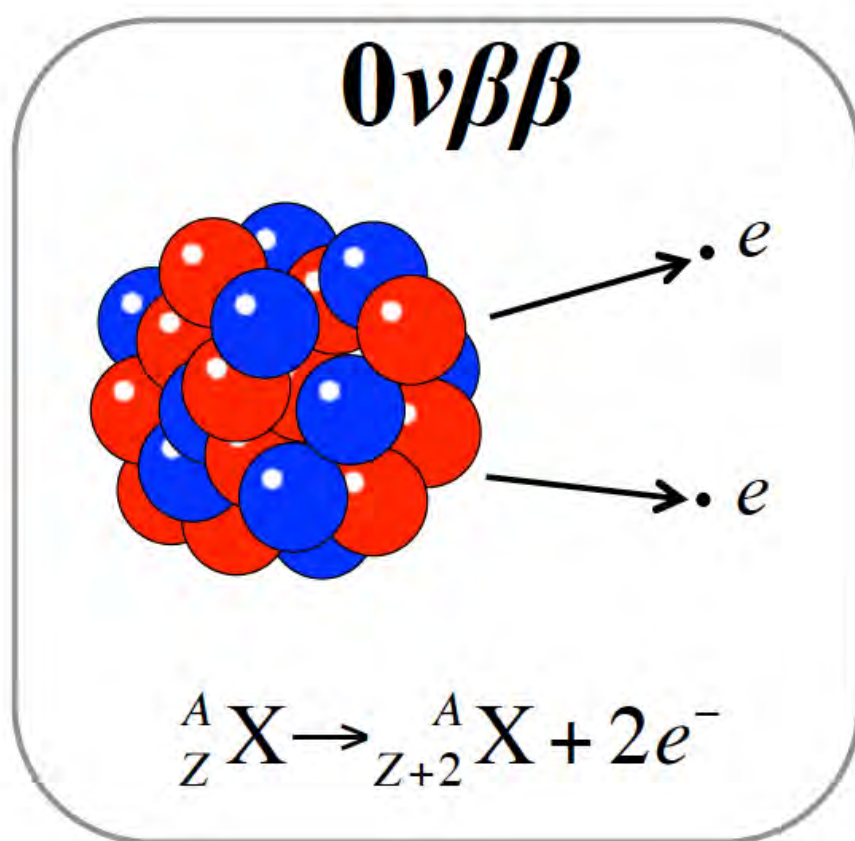
Neutrinos as a probe of nature



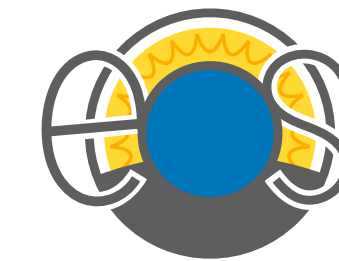
Particle astrophysics

Nuclear physics

Studying the fundamental nature of matter

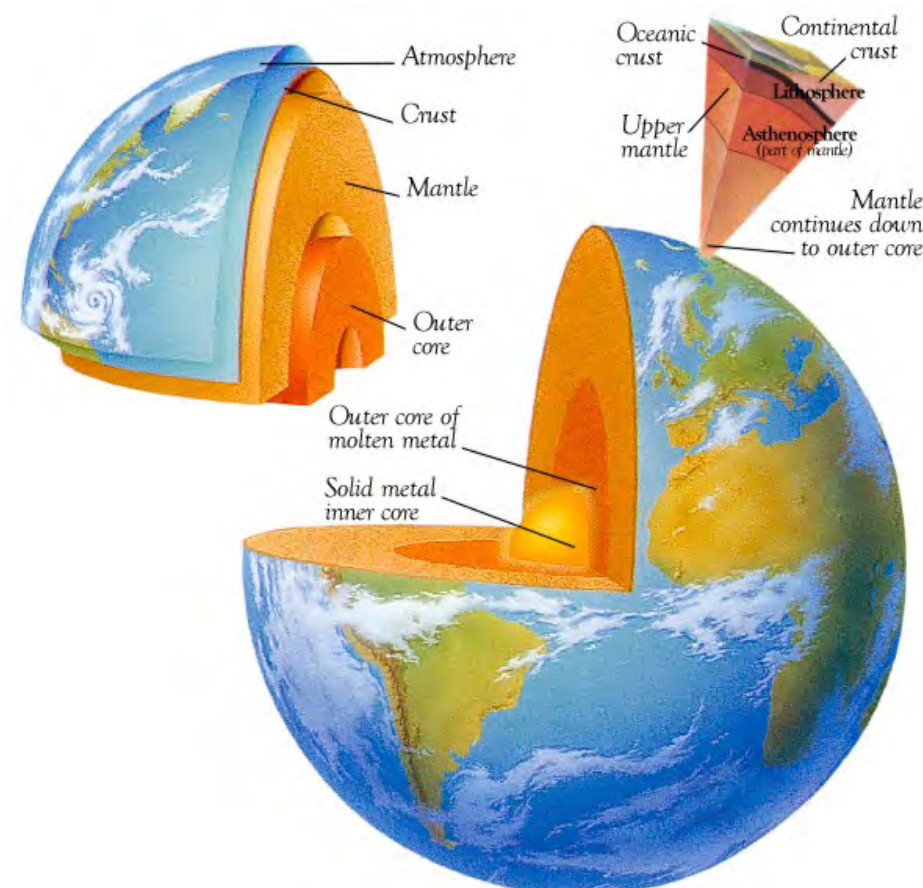
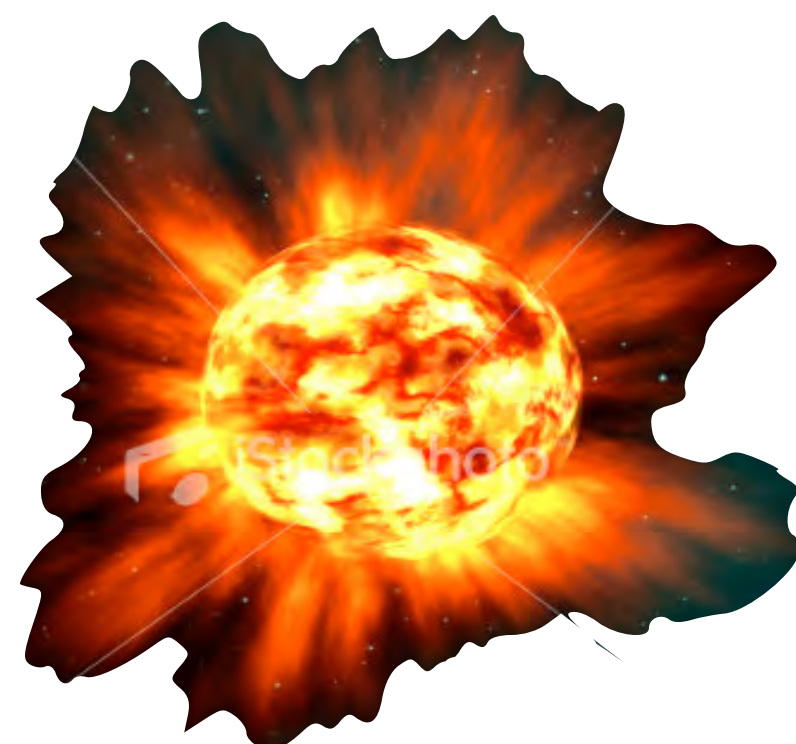
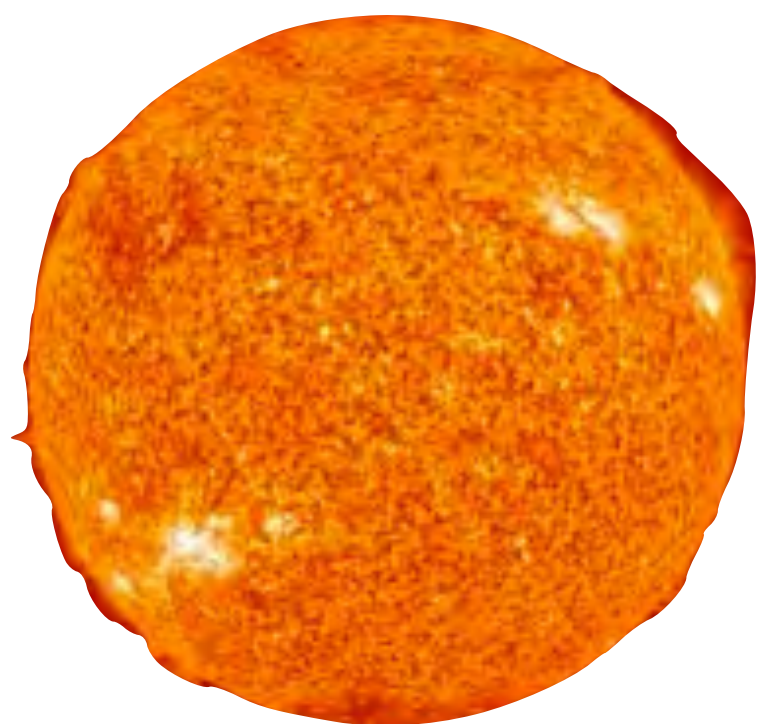


High energy physics



THEIA physics program

Neutrinos as a probe of nature

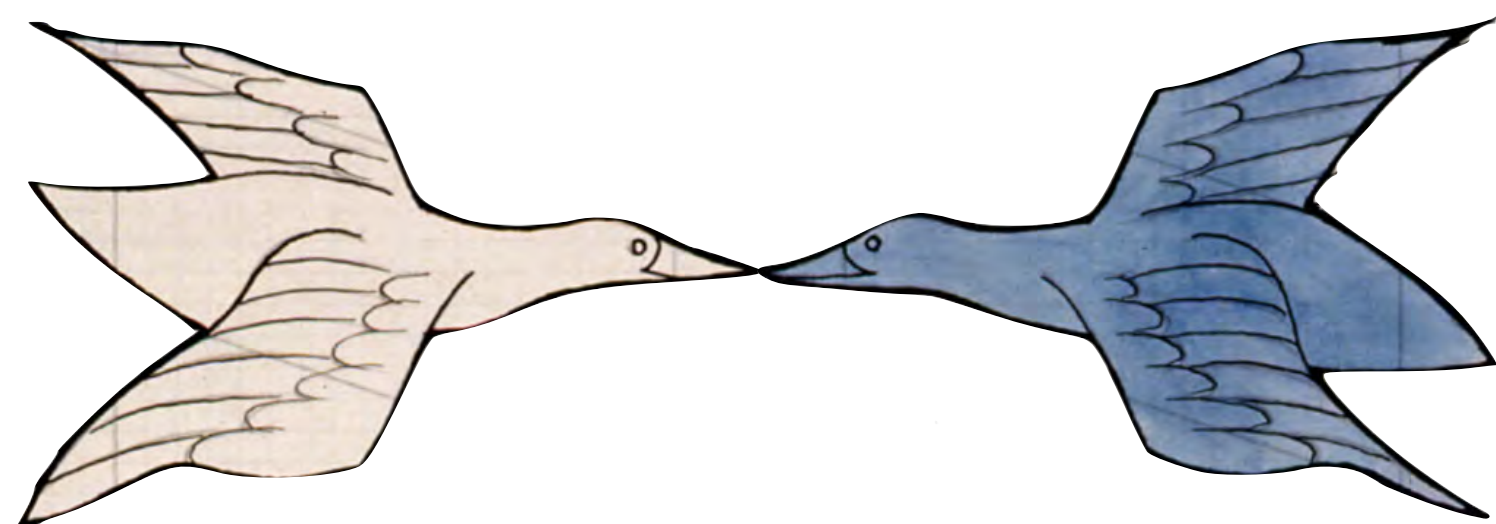
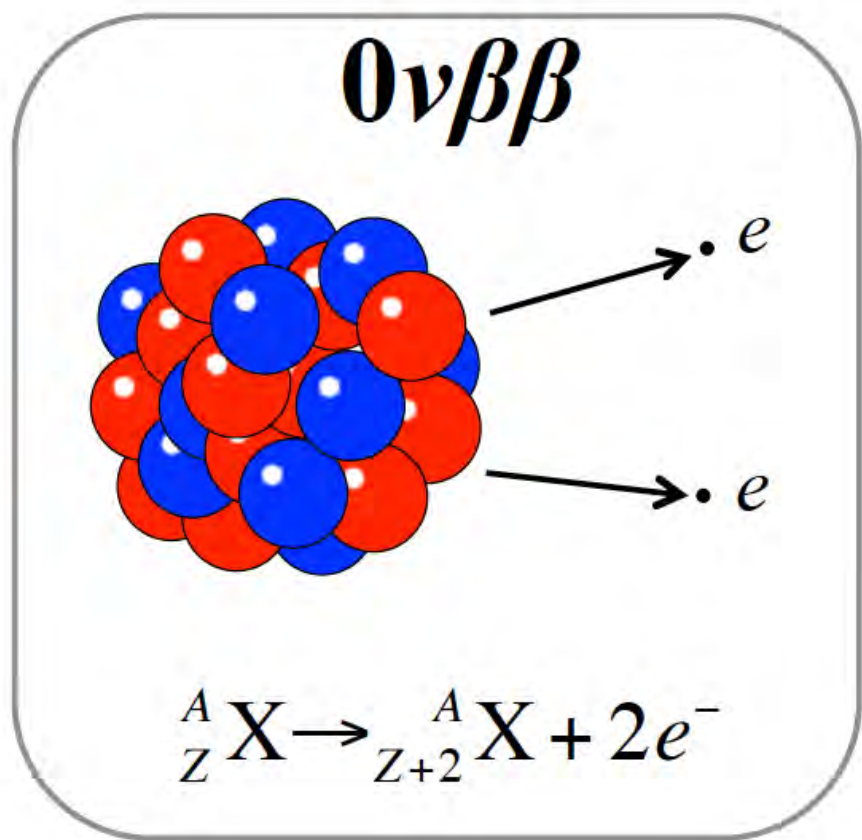


Physics over 5 orders of magnitude

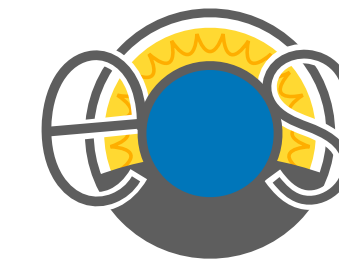
Particle astrophysics

Nuclear physics

Studying the fundamental nature of matter

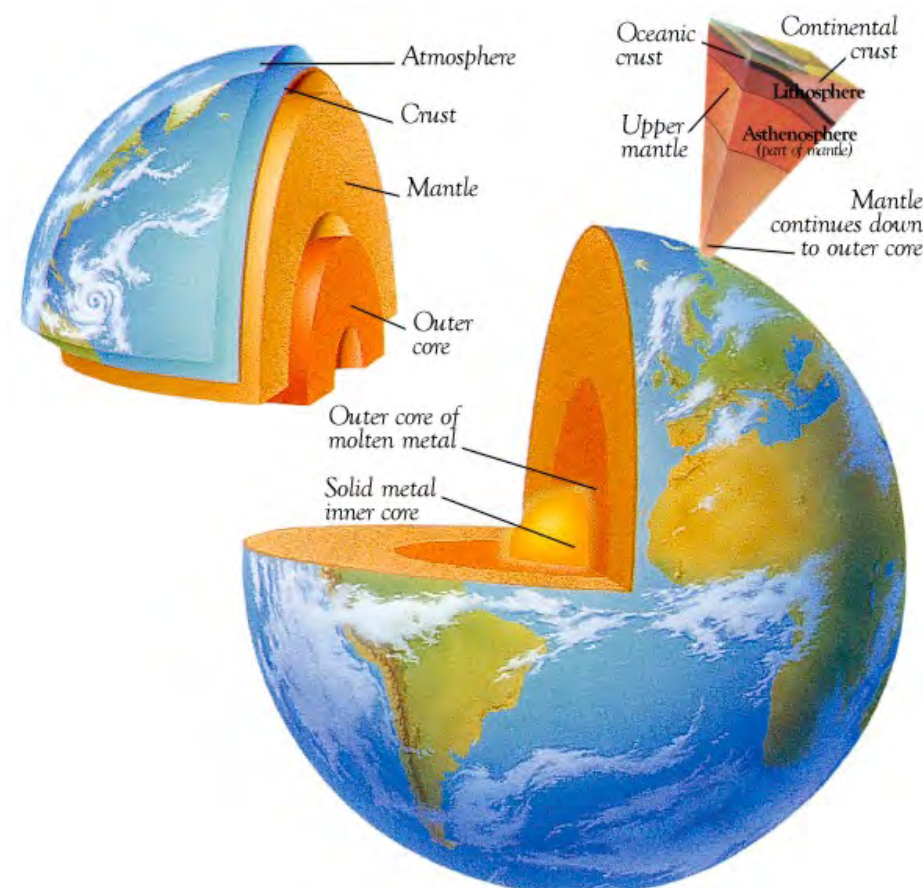
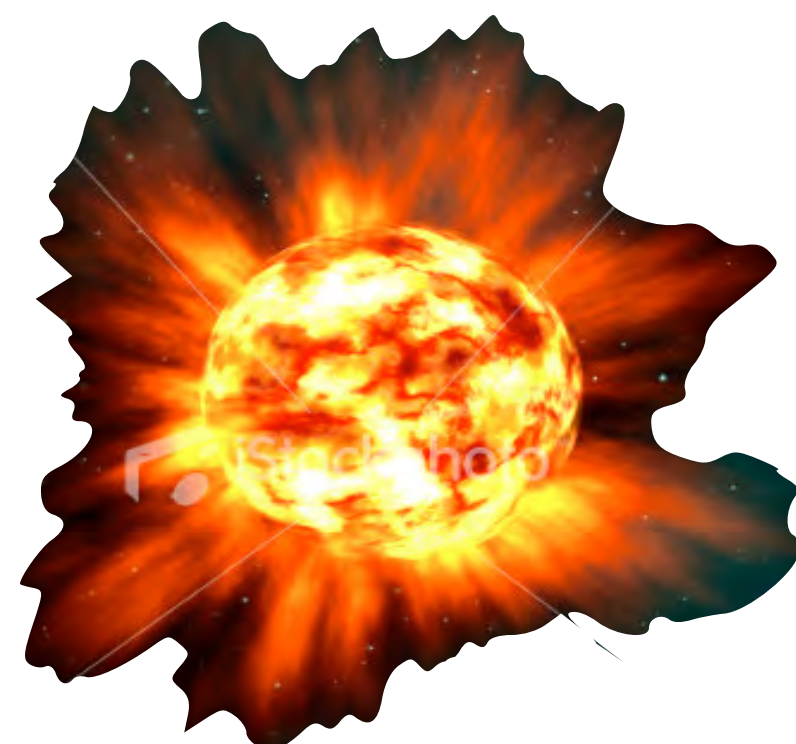
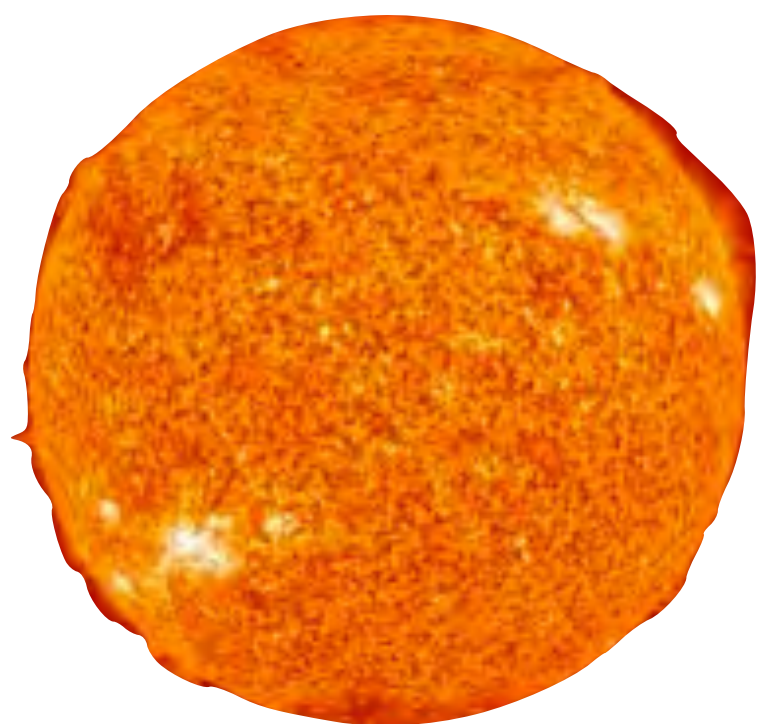


High energy physics



THEIA physics program

Neutrinos as a probe of nature



Physics over 5 orders of magnitude

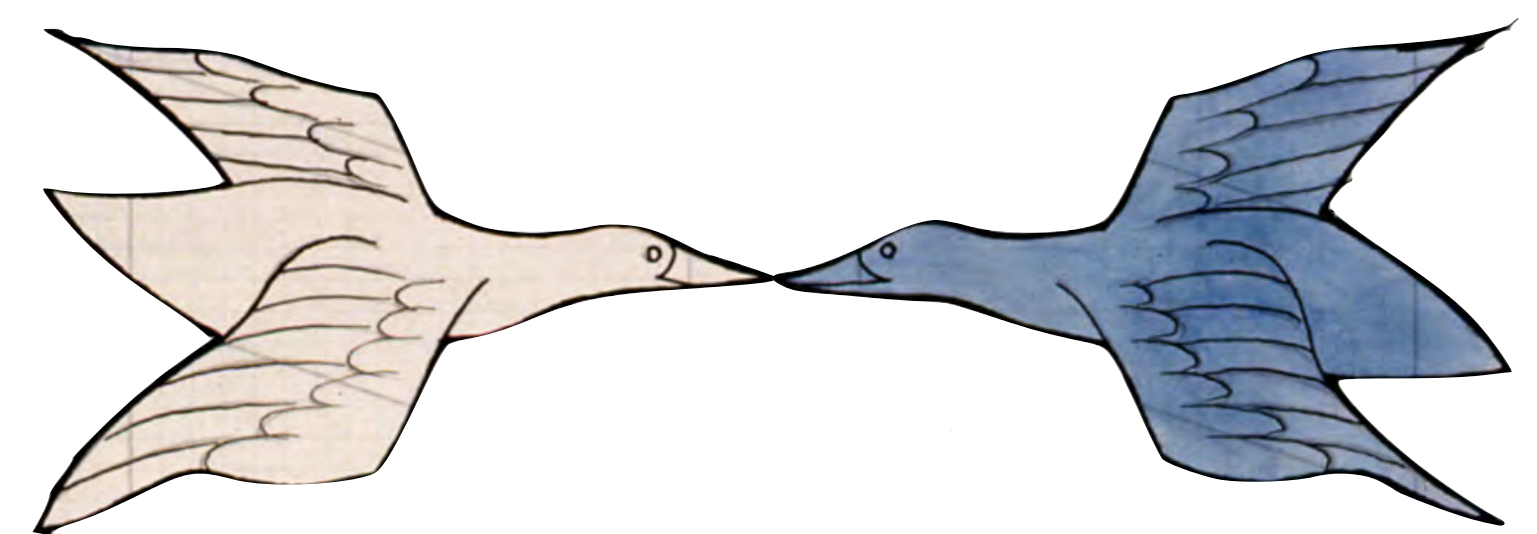
Particle astrophysics

Nuclear physics

Studying the fundamental nature of matter

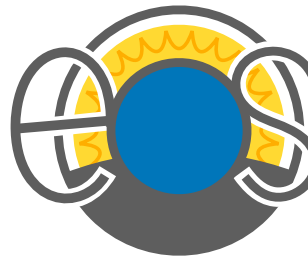
$0\nu\beta\beta$

${}^A_Z X \rightarrow {}^A_{Z+2} X + 2e^-$



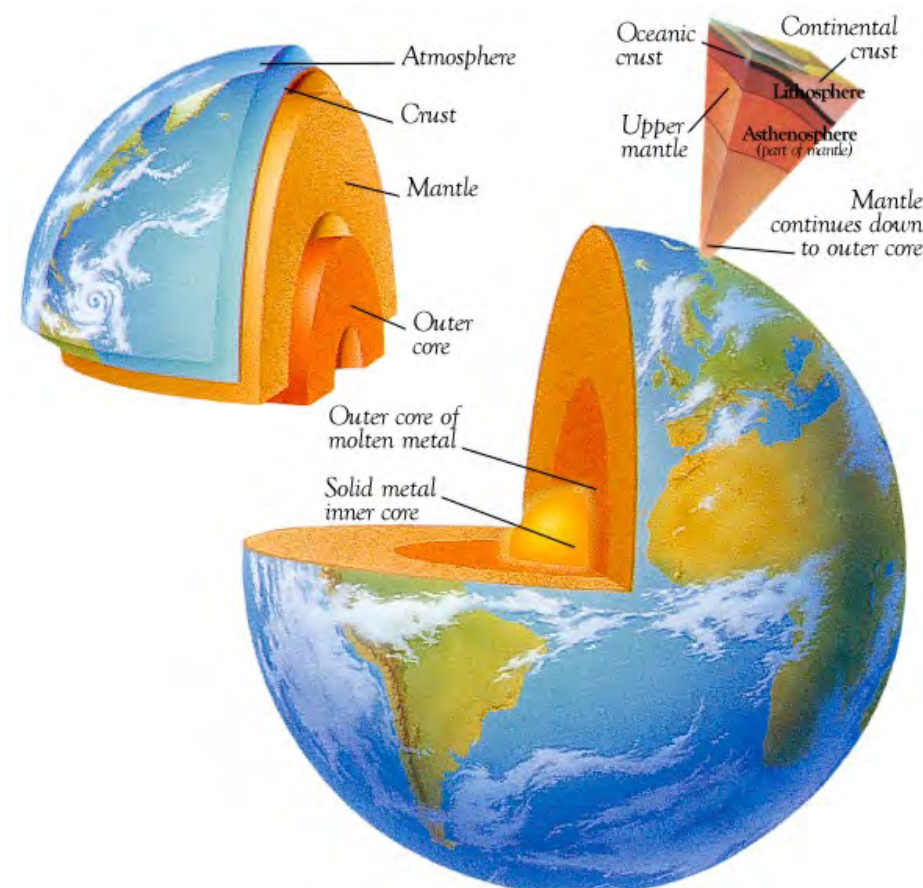
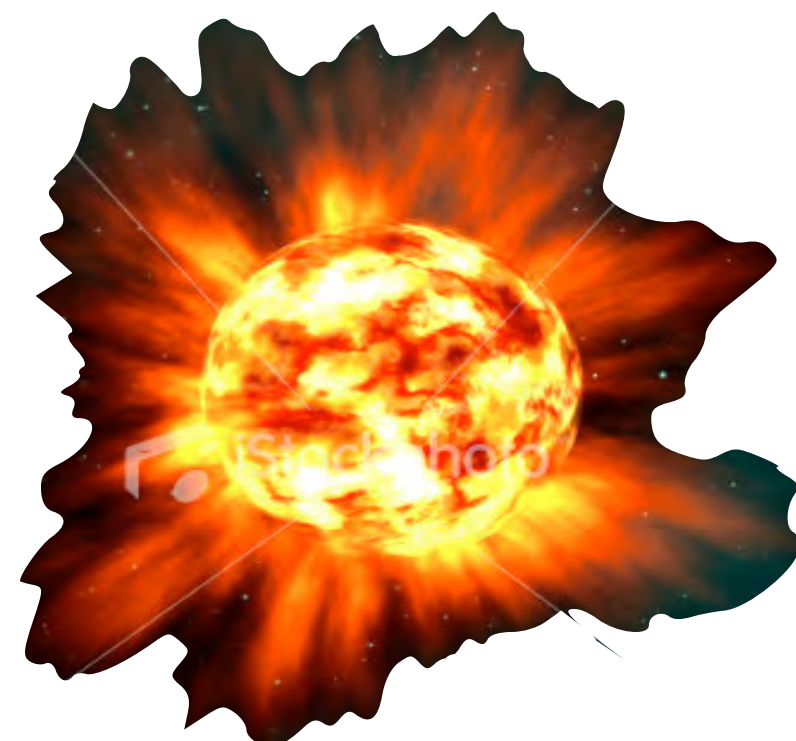
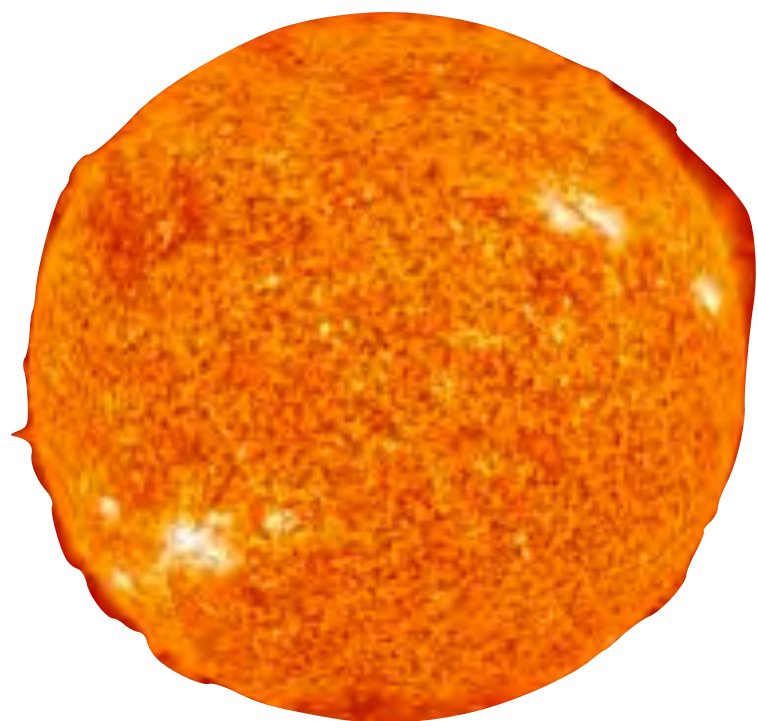
High energy physics

Remarkably, the same detector could show that neutrinos and antineutrinos are the same, **and** that “neutrinos” and “antineutrinos” oscillate differently



THEIA physics program

Neutrinos as a probe of nature

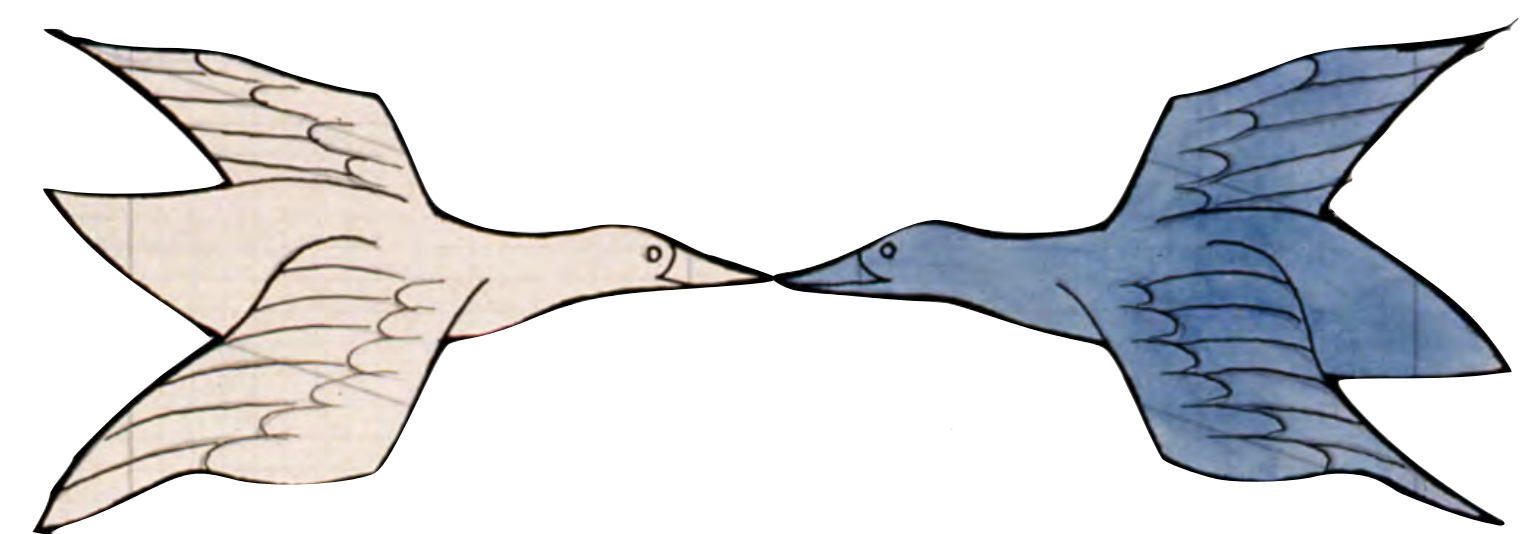
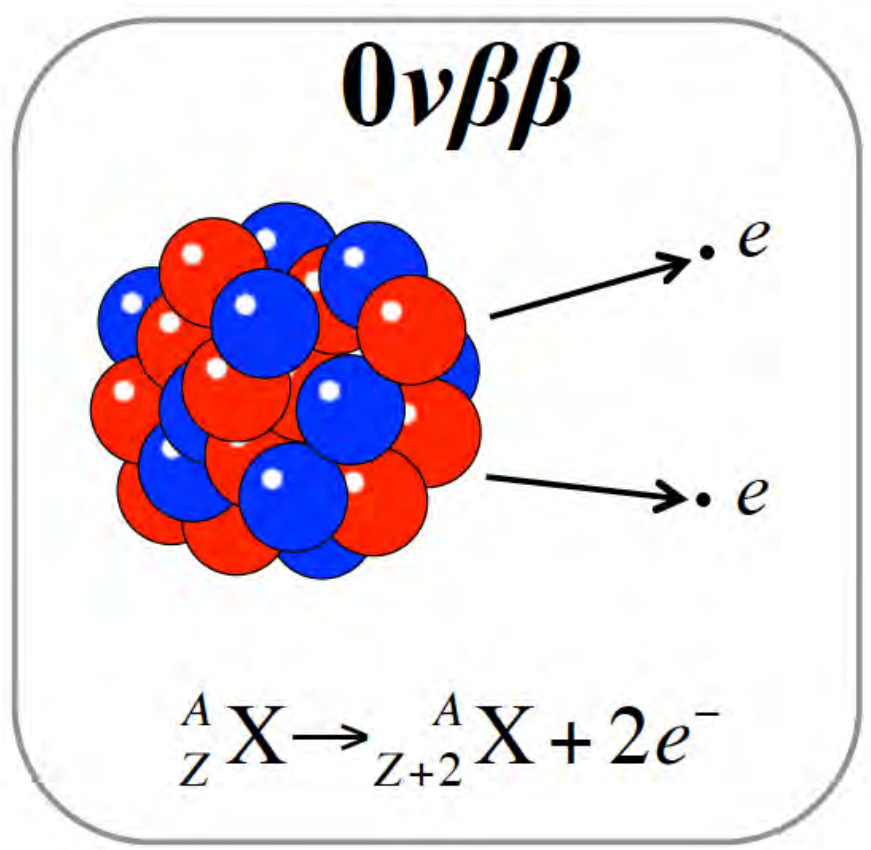


Physics over 5 orders of magnitude

Particle astrophysics

Nuclear physics

Studying the fundamental nature of matter



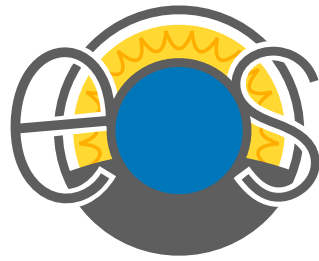
Remarkably, the same detector could show that neutrinos and antineutrinos are the same, **and** that “neutrinos” and “antineutrinos” oscillate differently

High energy physics

Matter-dominated universe



Neutrinos for nonproliferation



Neutrino properties

Produced as a by-product of fission and fusion



Provide a unique signature of nuclear fission / fusion

Weakly interacting



Signature can't be shielded

Close to massless, travel $\sim c$



Near-instantaneous detection

Interaction products preserve incident direction



Can be used to "point" back to the source

Test site transparency



Tunnel at the Ula Complex

Small modular reactors

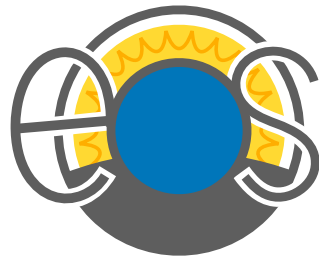


Maritime sensing





Neutrinos for nonproliferation



Neutrino properties

Produced as a by-product of fission and fusion



Provide a unique signature of nuclear fission / fusion

Weakly interacting



Signature can't be shielded

Close to massless, travel $\sim c$



Near-instantaneous detection

Interaction products preserve incident direction



Can be used to "point" back to the source

- NuTools: 2021 study (DNN R&D) "exploring practical roles for neutrinos in nuclear energy and security"
- LBNL's focus is to advance the technology to enhance the capabilities of such a detector
 - Reduce the required scale, increase standoff, and provide additional synergies with Office of Science interests
- Close partnership with related efforts at BNL & LLNL

Test site transparency



Tunnel at the U1a Complex

Small modular reactors

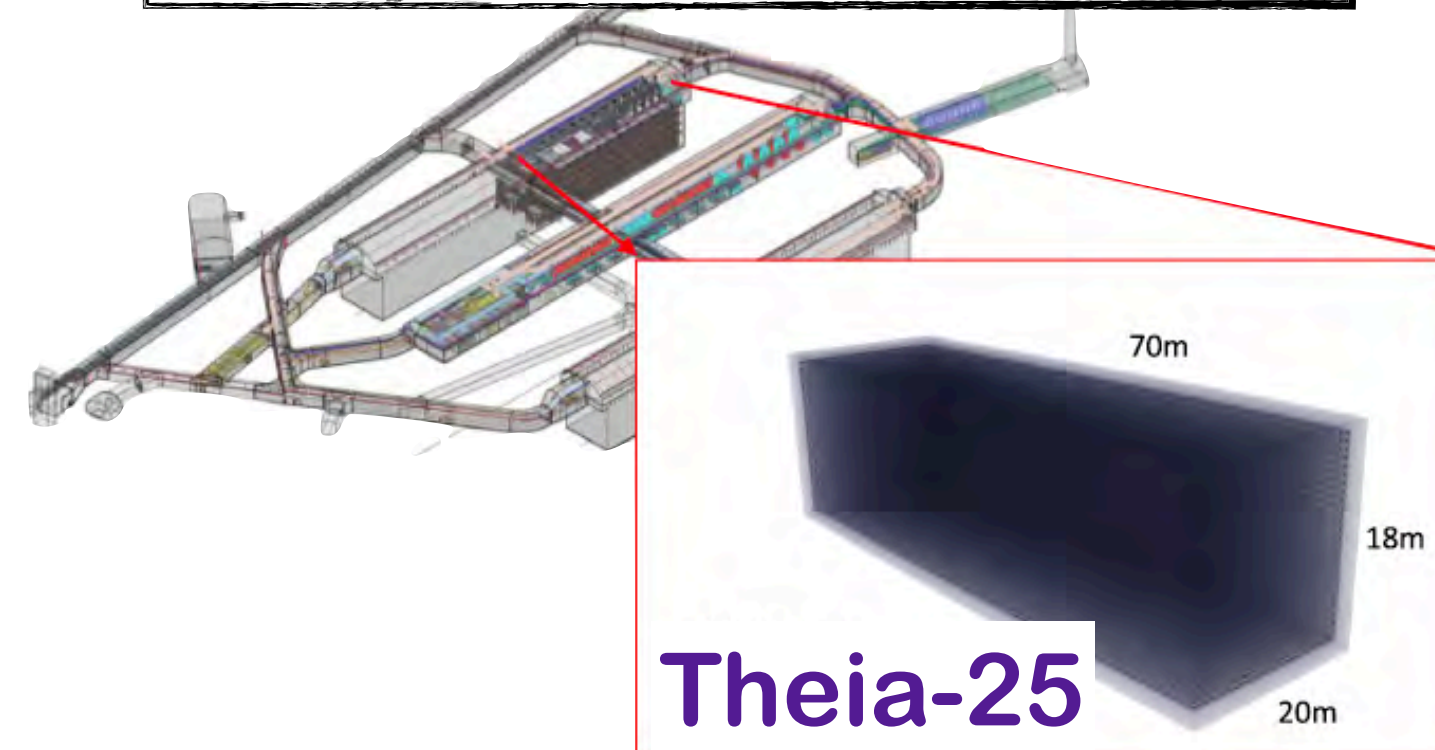


Maritime sensing

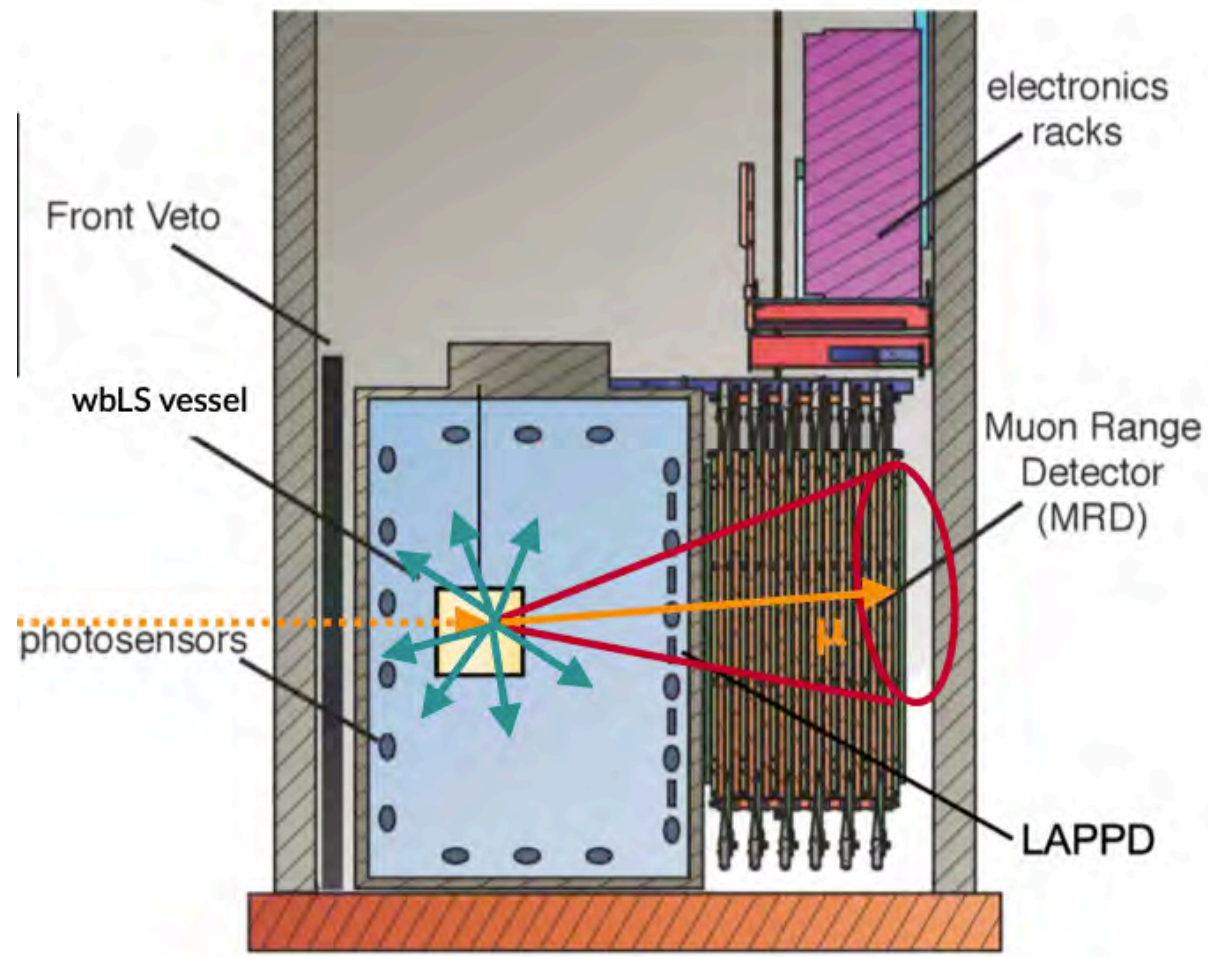


The path to THEIA

Building on a broad program of bench-top scale development



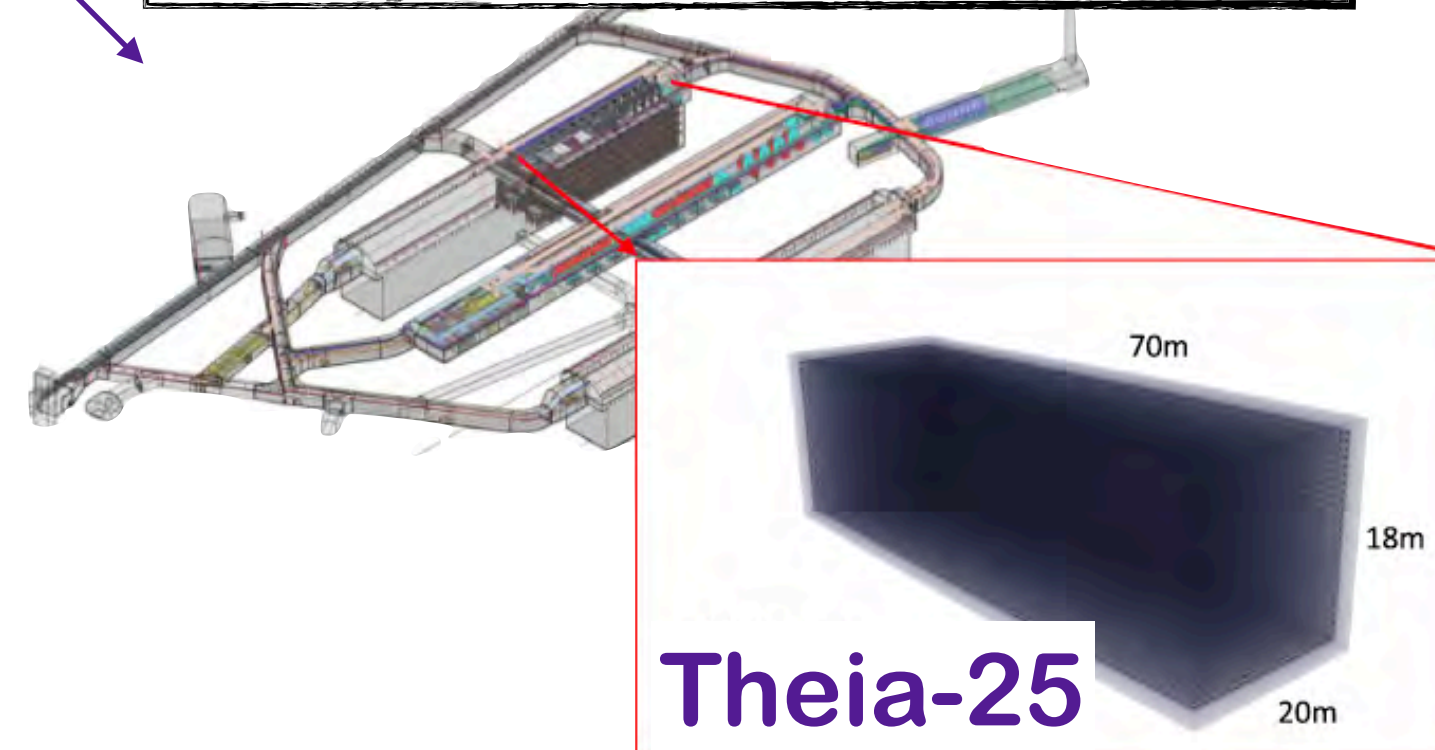
The path to THEIA



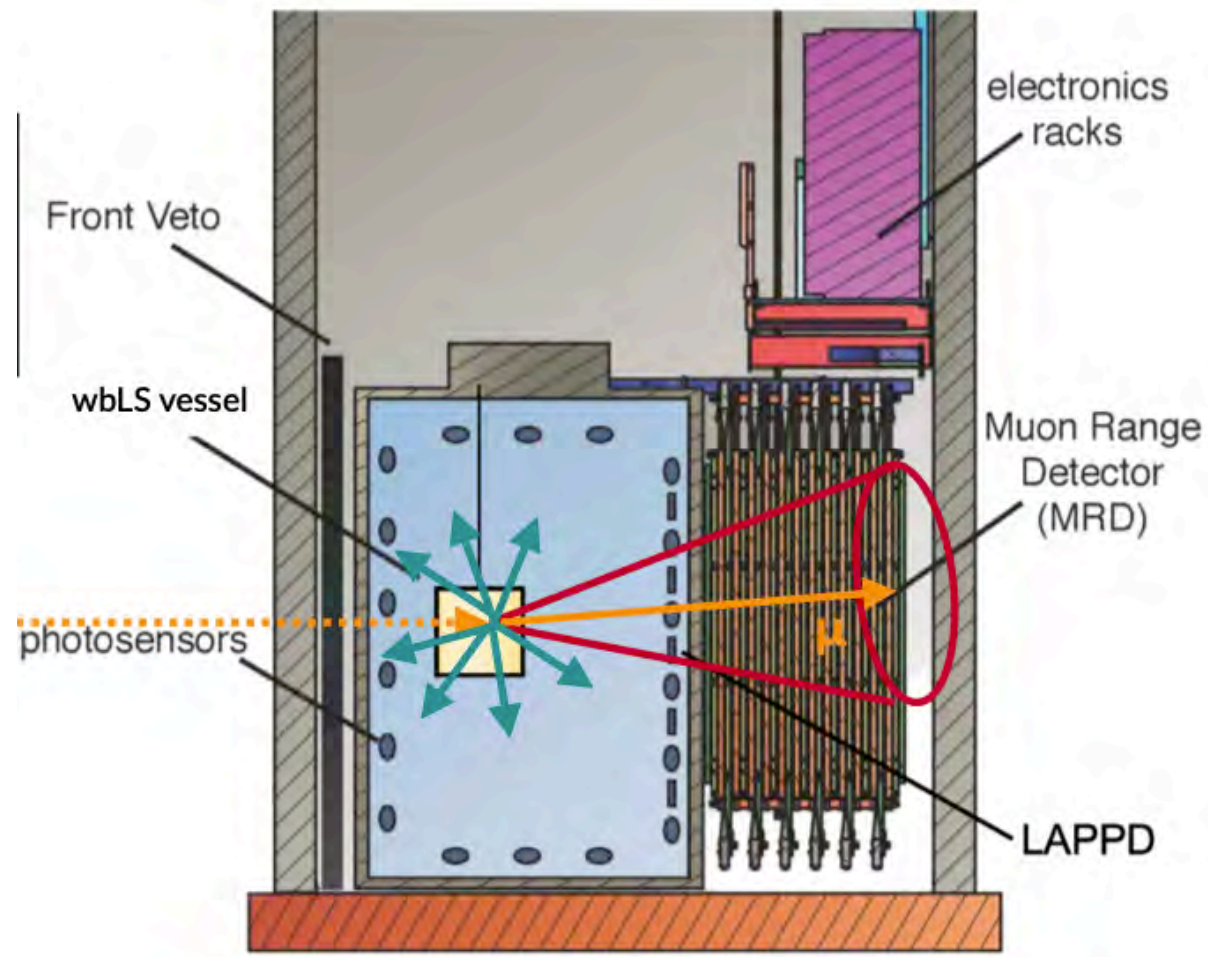
ANNIE: 365 kg

High-energy event reconstruction,
neutrino detection

Building on a broad program of
bench-top scale development



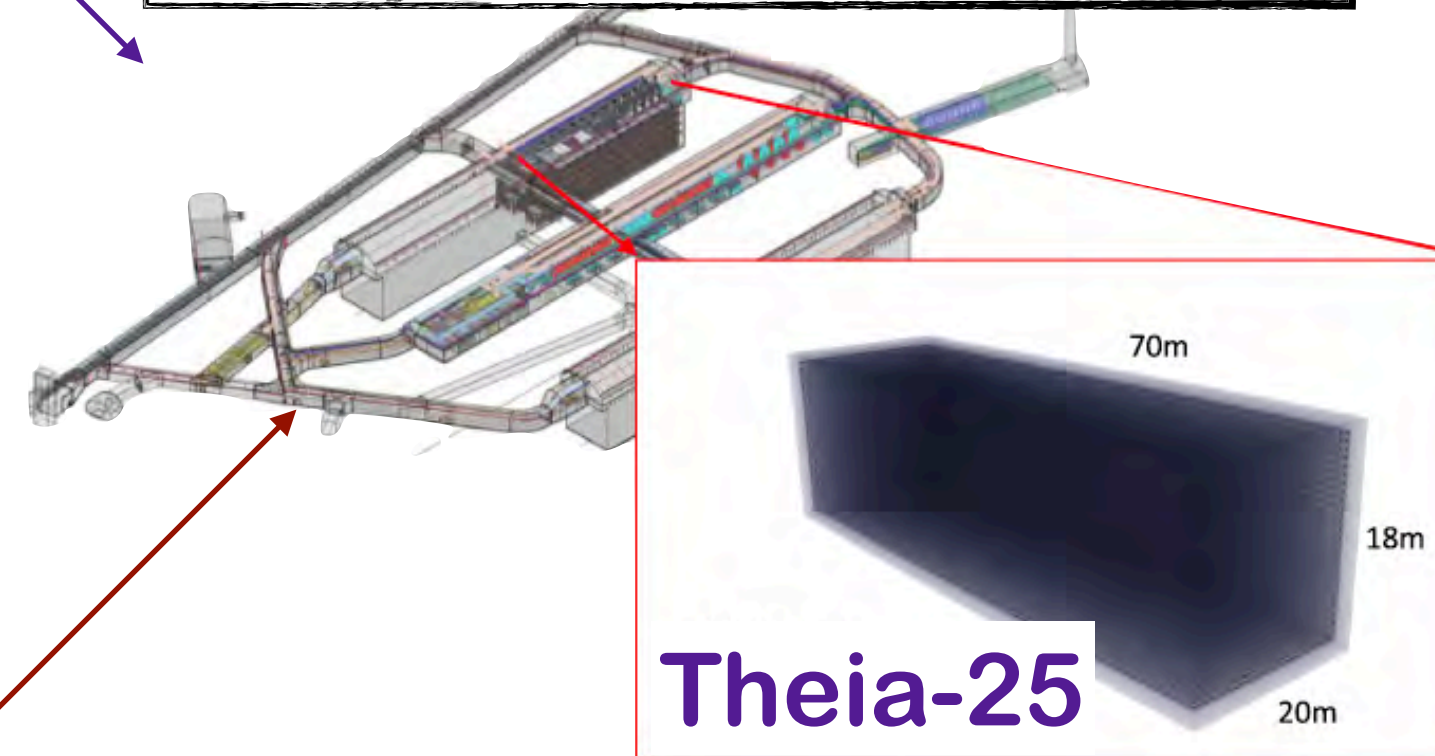
The path to THEIA



ANNIE: 365 kg

High-energy event reconstruction, neutrino detection

Building on a broad program of bench-top scale development

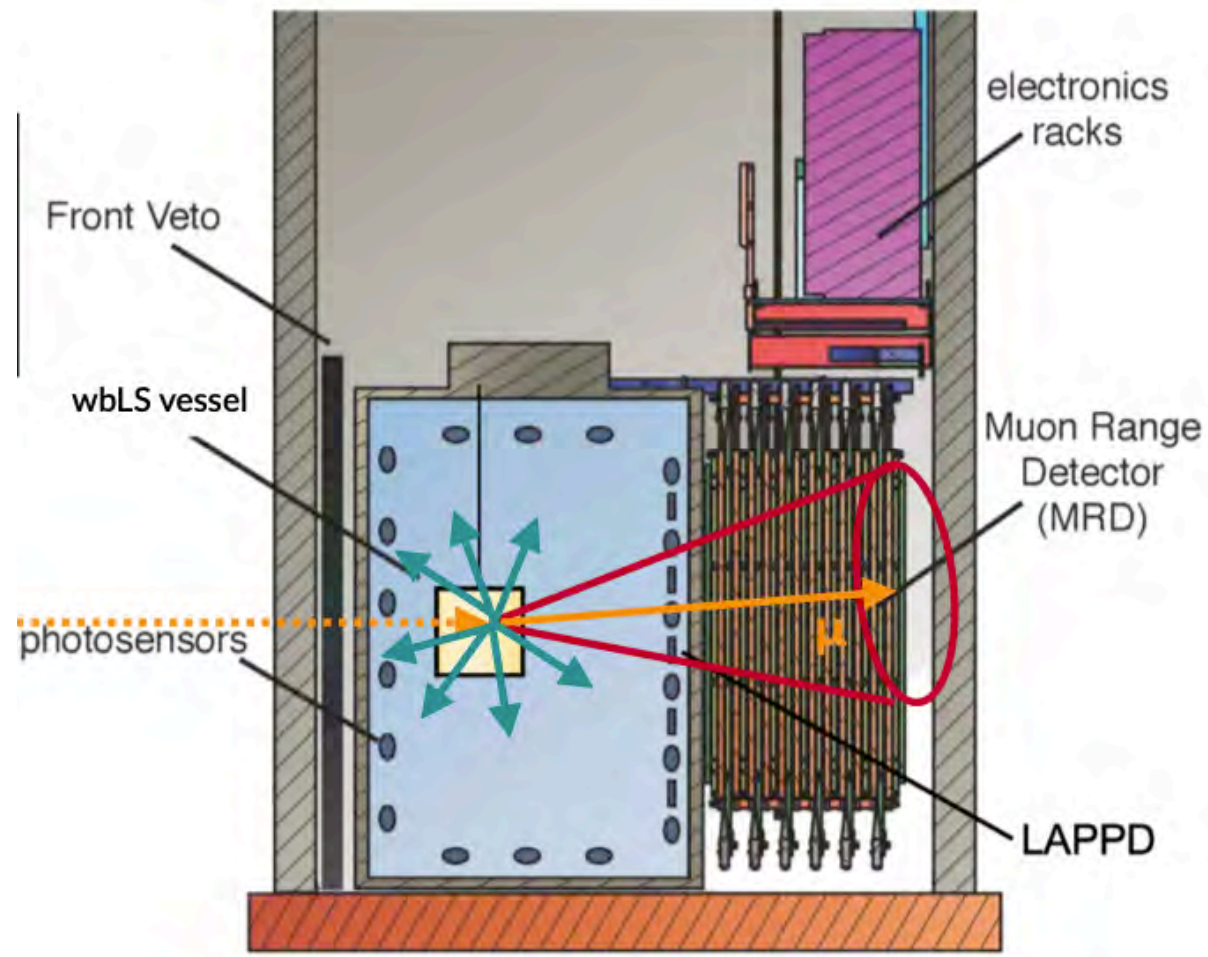


BNL: 1- and 30-ton

Deployment, purification, recirculation, transparency

Theia-25

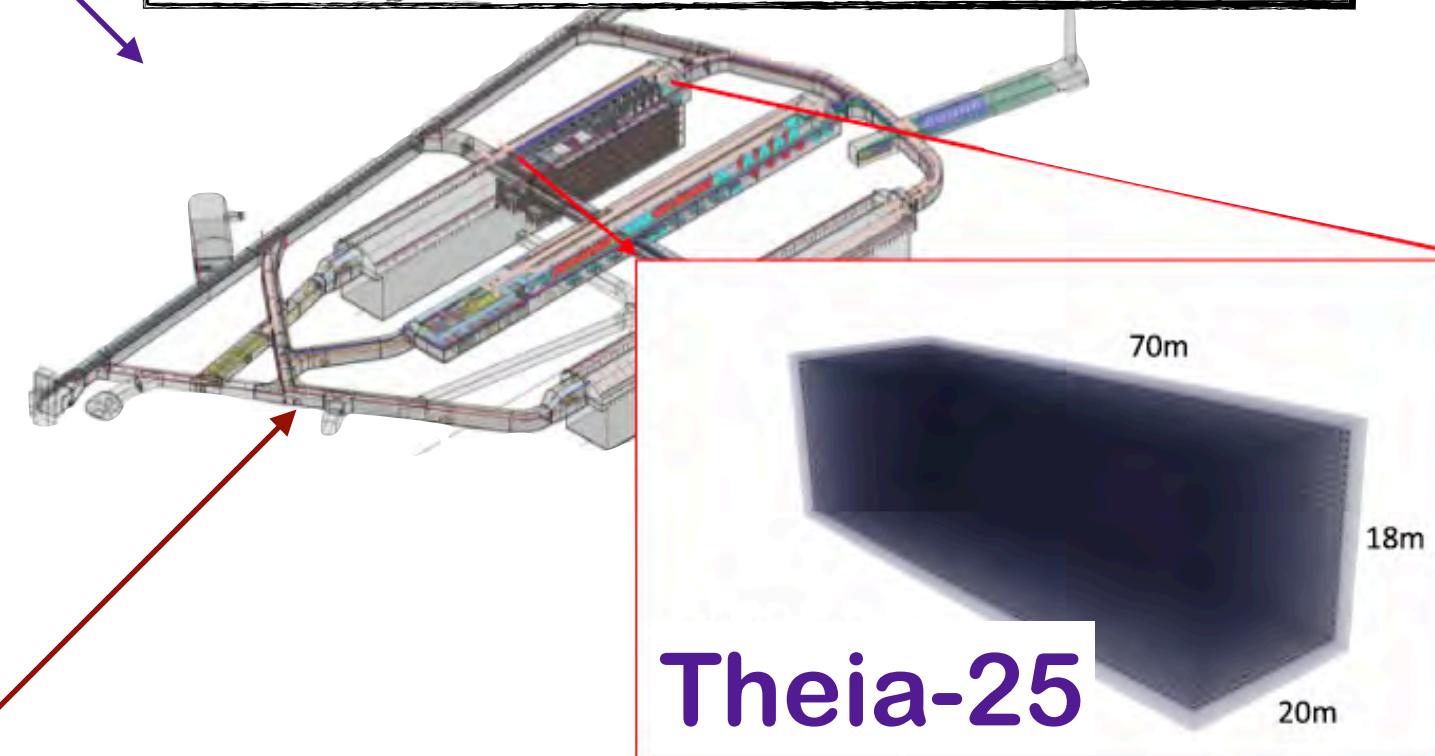
The path to THEIA



ANNIE: 365 kg

High-energy event reconstruction, neutrino detection

Building on a broad program of bench-top scale development



Theia-25



BNL: 1- and 30-ton

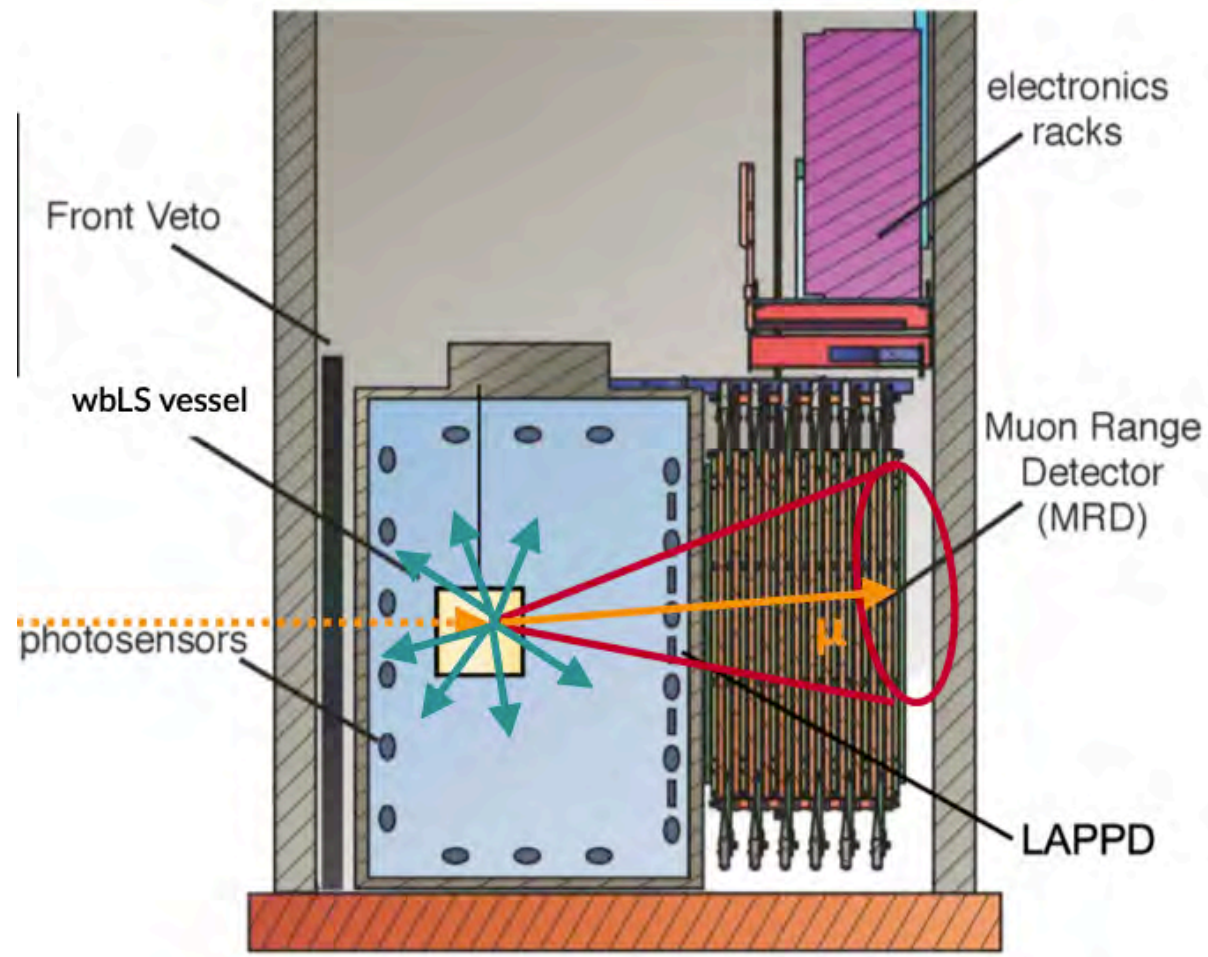
Deployment, purification, recirculation, transparency



NuDot: 1 ton

Isotope loading, NLDBD topology

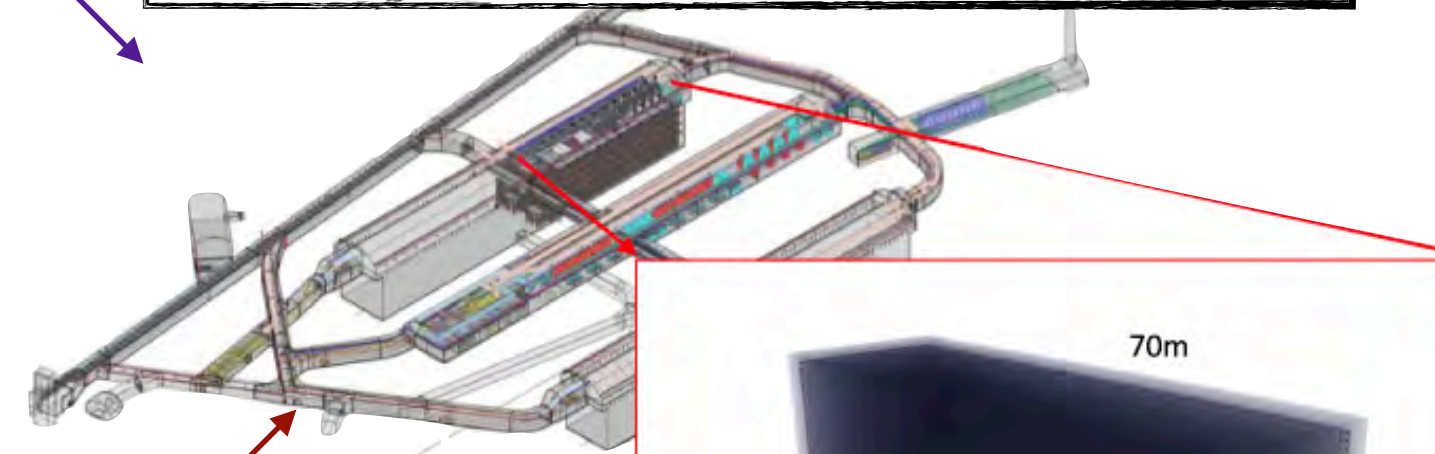
The path to THEIA



ANNIE: 365 kg

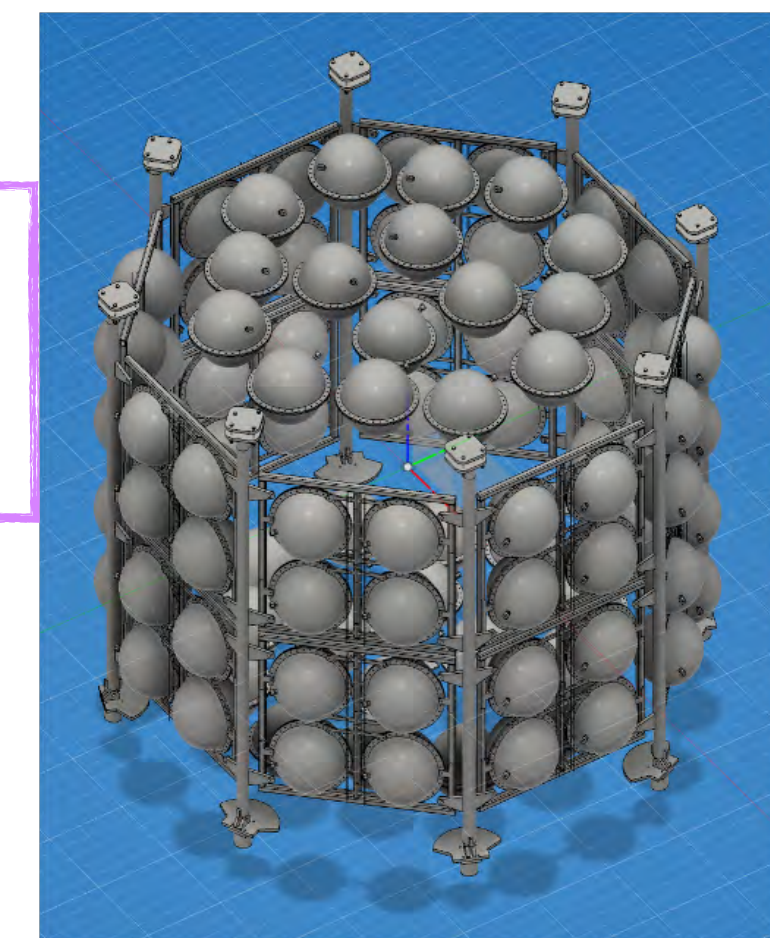
High-energy event reconstruction, neutrino detection

Building on a broad program of bench-top scale development



BUTTON: 30 ton

Underground deployment, low bkg verification



BNL: 1- and 30-ton

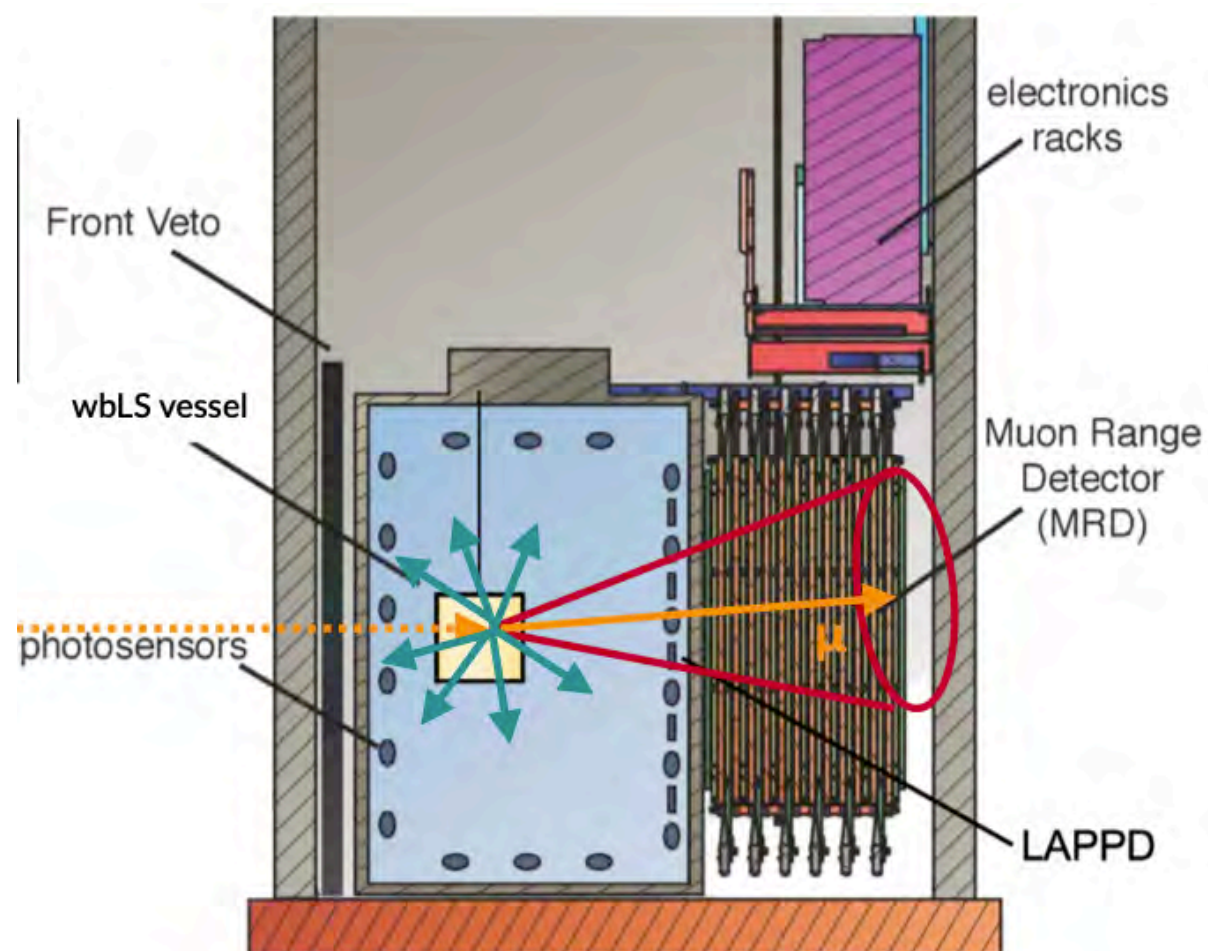
Deployment, purification, recirculation, transparency



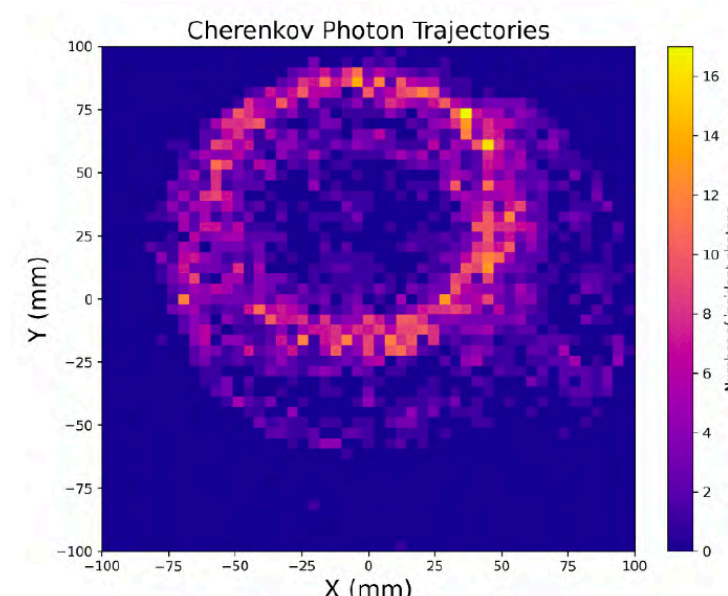
NuDot: 1 ton

Isotope loading, NLDBD topology

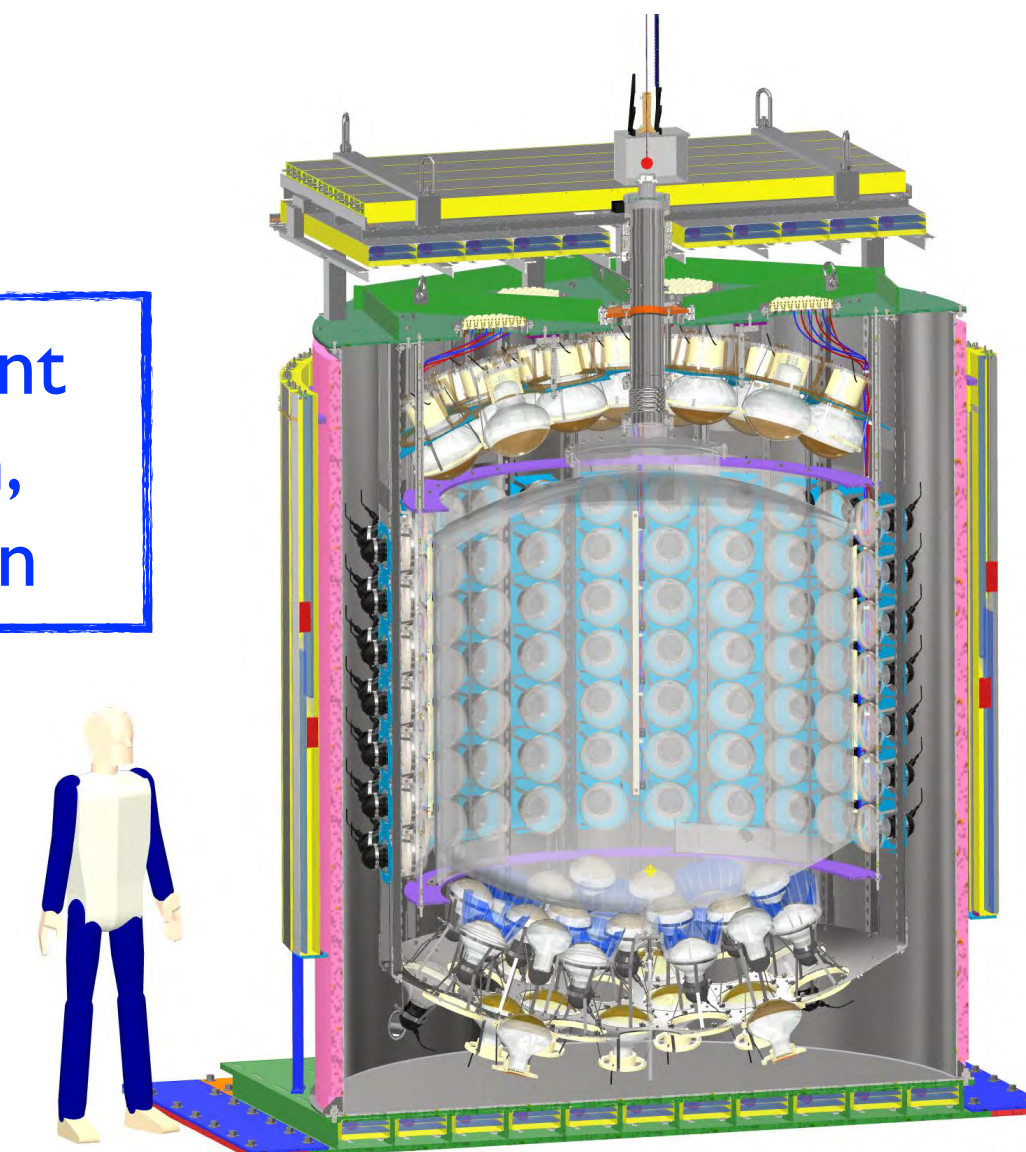
The path to THEIA



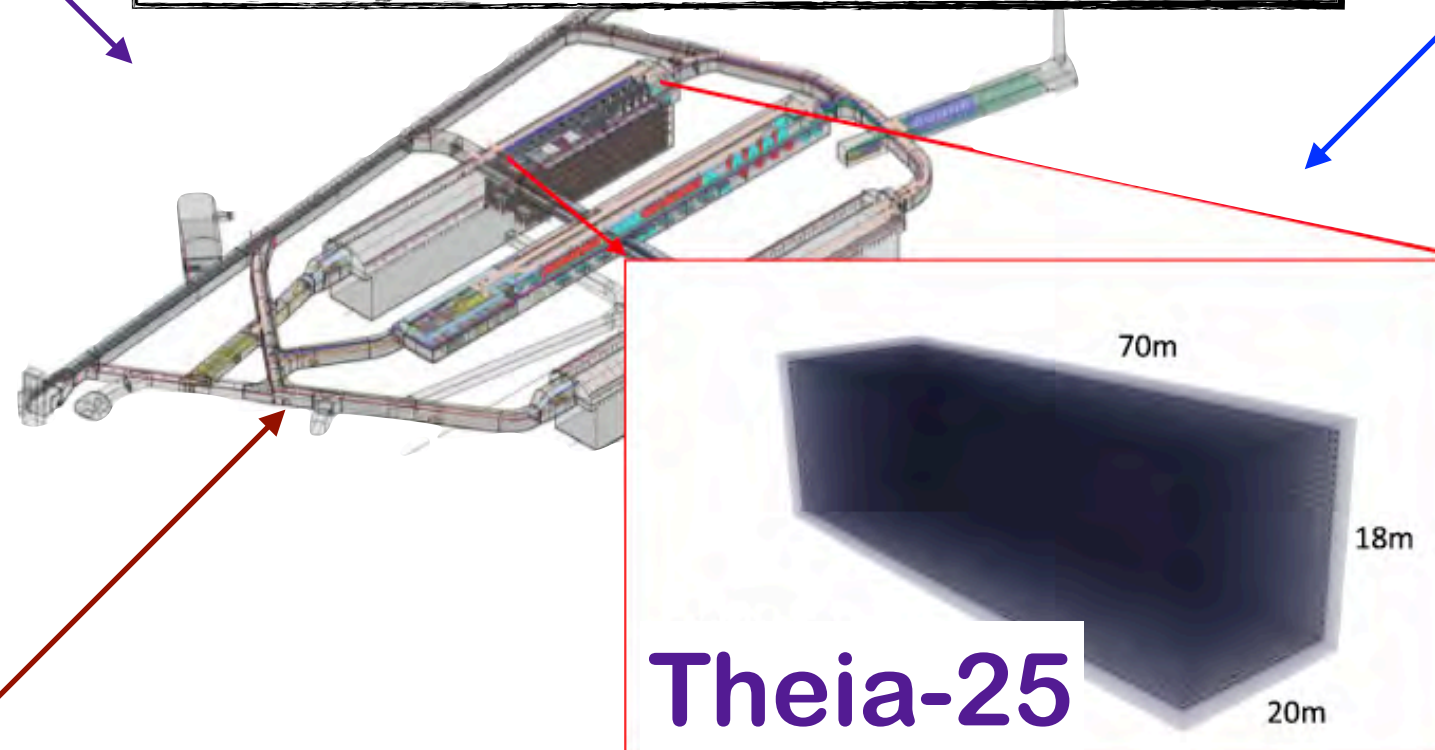
ANNIE: 365 kg
 High-energy event reconstruction, neutrino detection



Eos: 20 ton
 Low-energy event reconstruction, model validation

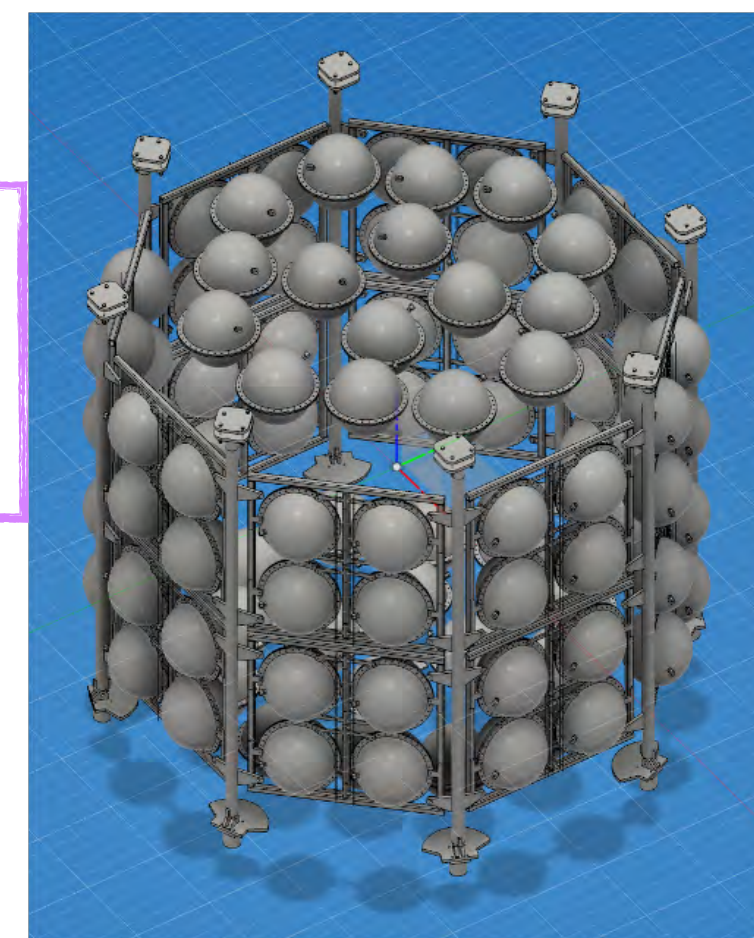


Building on a broad program of bench-top scale development



Theia-25

BUTTON: 30 ton
 Underground deployment, low bkg verification



BNL: 1- and 30-ton
 Deployment, purification, recirculation, transparency



NuDot: 1 ton
 Isotope loading, NLDBD topology

A woman with long dark hair, wearing a long, flowing white dress, stands on a bed of white clouds. She has her arms outstretched, looking upwards. In the background, a large, glowing blue planet with a ring system is visible on the left, and a bright sun or star is on the right, casting a warm glow. The sky is filled with soft, colorful clouds in shades of blue, purple, and orange.

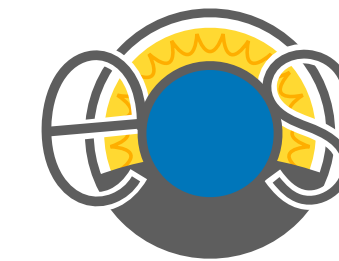
Eos (Darwin)

*Funded by NNSA,
DNN R&D
FY22-24*

Let There be Light



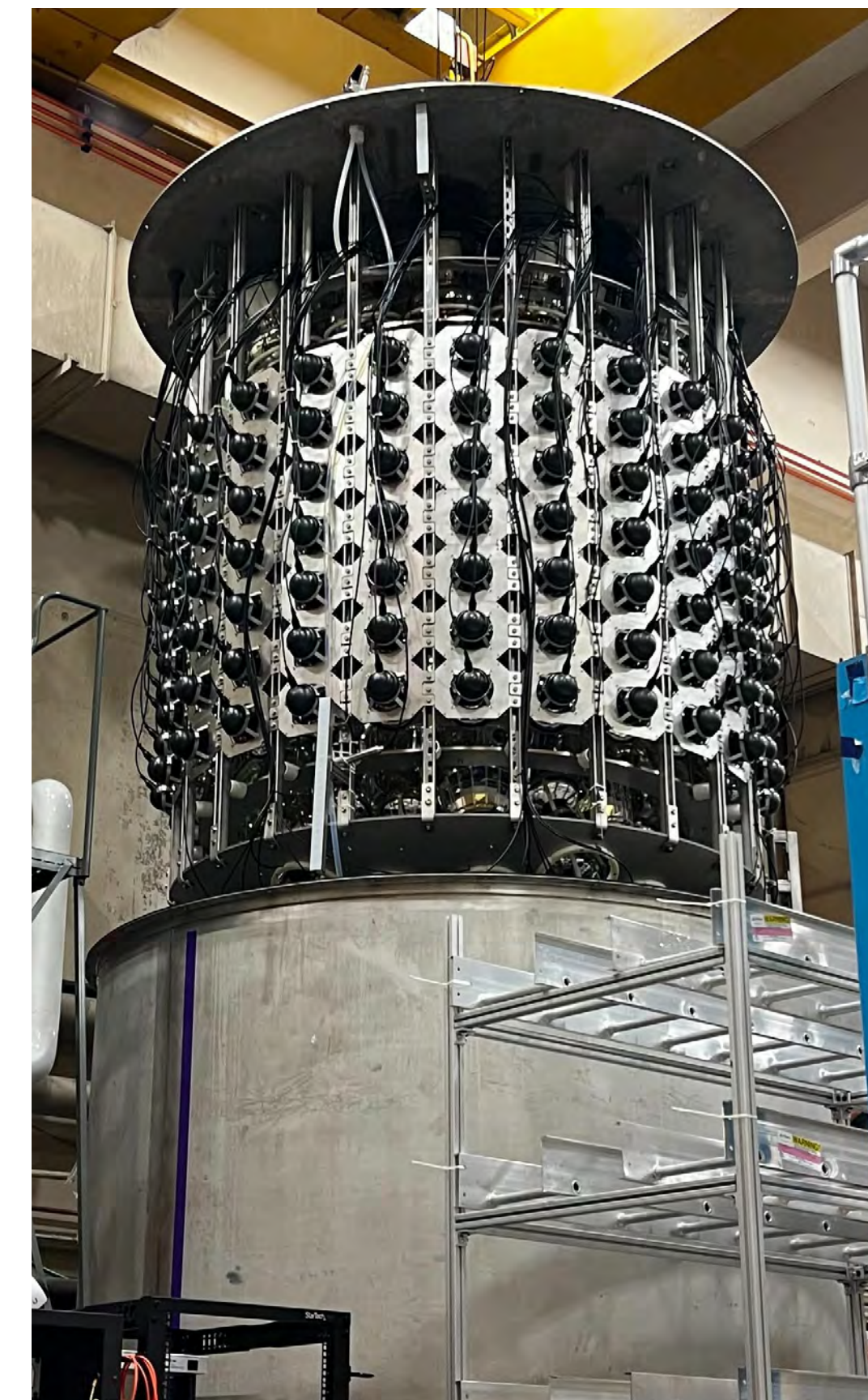
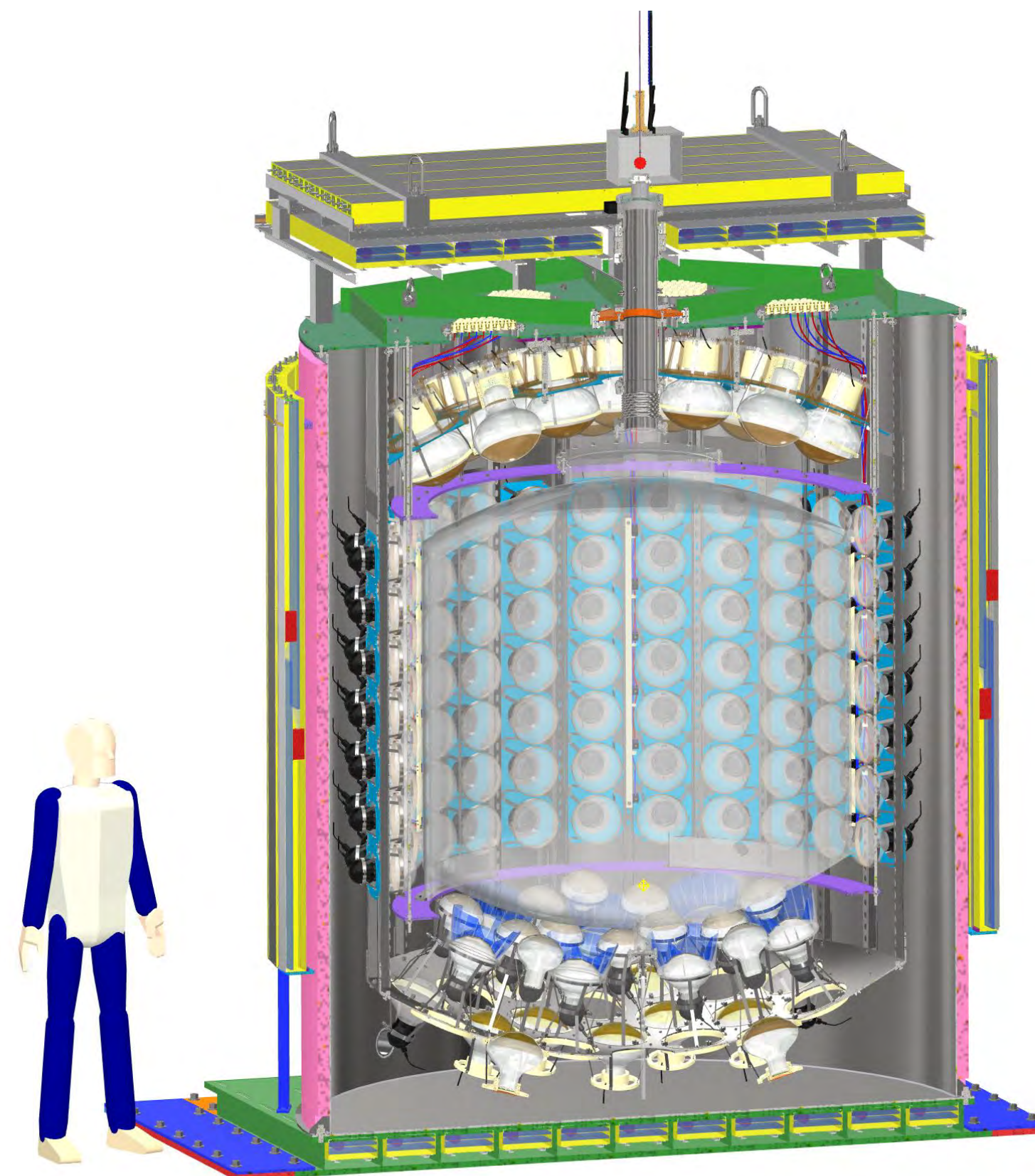
EOS: performance demonstrator



Approach: design, construct and operate an integrated testbed to demonstrate the performance of novel technology

Novelty / technology:

- Novel scintillating liquids — water-based scintillator, slow scintillator
- Ultra-fast photon detectors — novel 8" PMTs (200 8" PMTs: RI4688-100, 900ps FWHM)
- “Quantum chromatic sorting”: dichroicons for spectrally sensitive photon detection
- AI/ML-based analysis techniques
- Deployable sources for studies of vertex, energy, direction reconstruction & PID
- 36-fiber light injection system for optical calibration

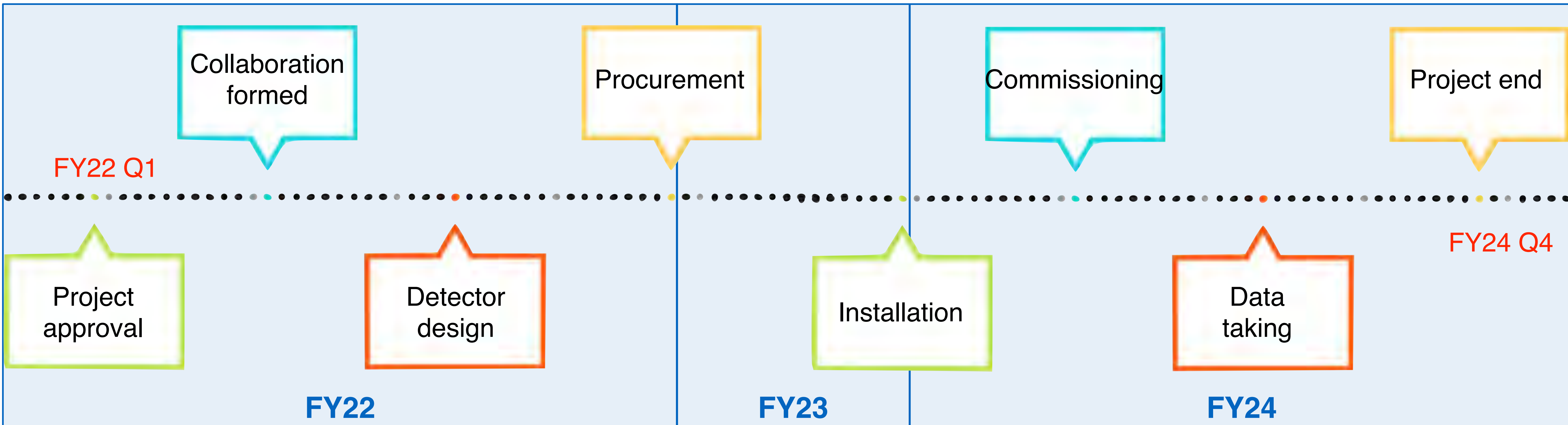


- Sited on UC Berkeley campus, in Nuclear Engineering (NE) department

*Designed for flexible upgrade paths & to be redeployed at a neutrino source
→ demonstrate viability of future applications*

EOS concept paper published: *JINST 18 P02009 (2023)*,
<https://doi.org/10.1088/1748-0221/18/02/P02009>

Eos Demonstrator timeline

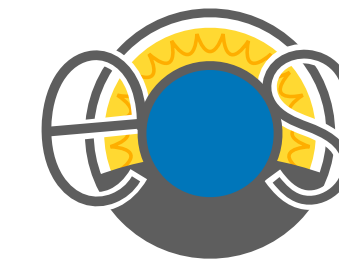


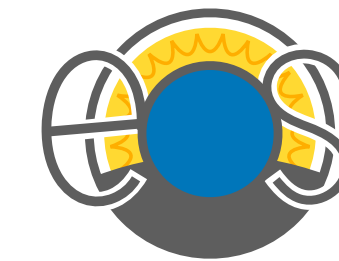
→ FY25-27?

- \$10M supported by DOE-NNSA-DNN R&D
- Led by LBNL
- FY22-24



Challenges faced

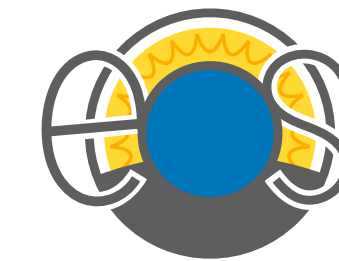




Challenges faced

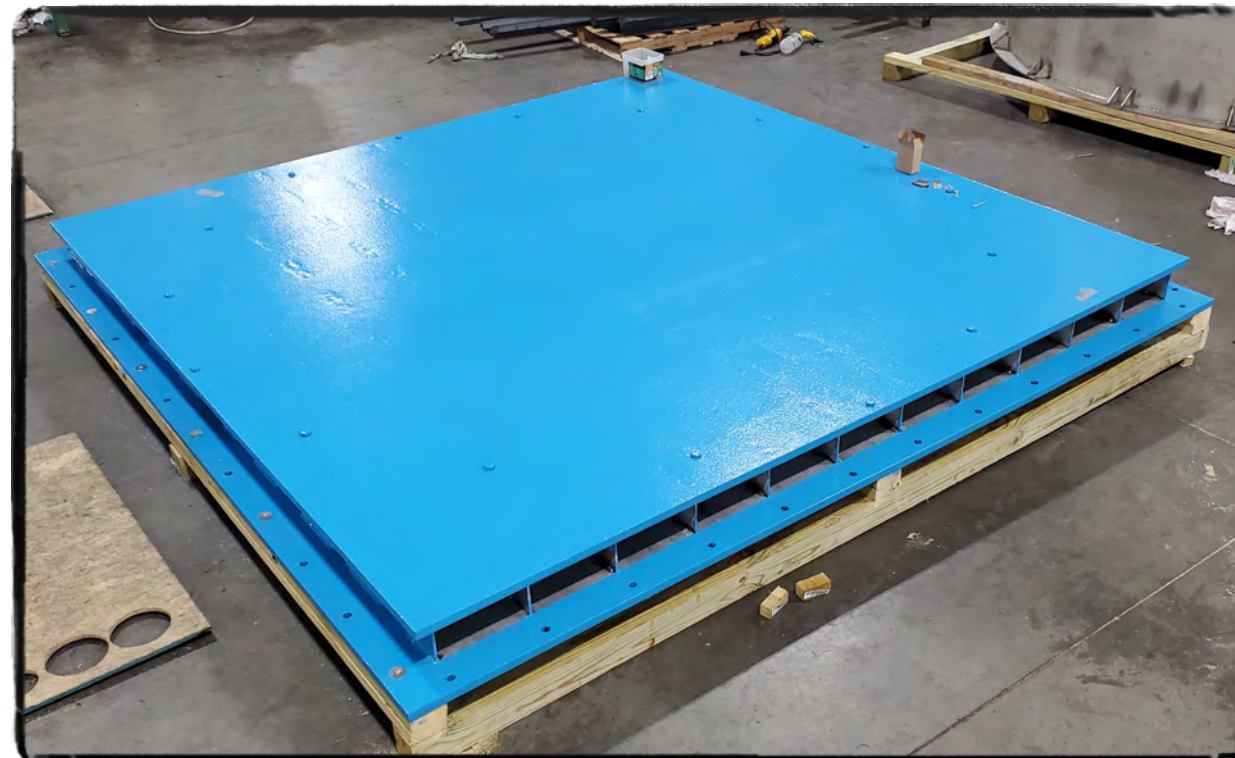
Outer vessel and assembly stand, May '23





Challenges faced

Outer vessel and assembly stand, May '23



Shipping, June 1 '23

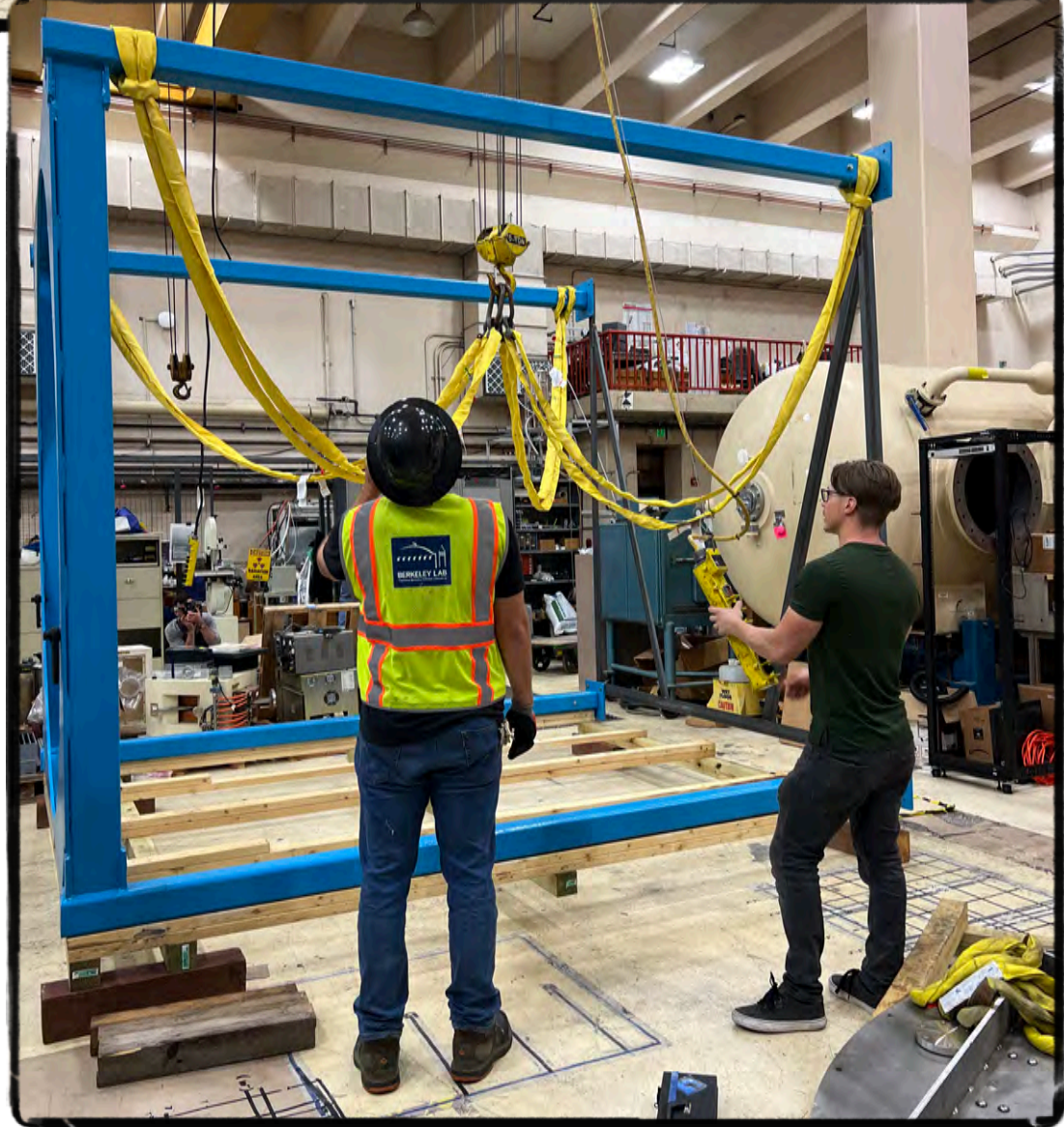
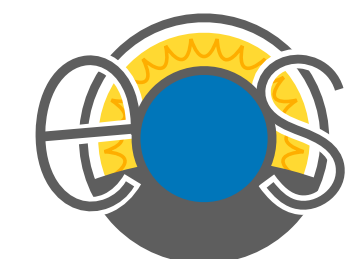


June 2, 2023



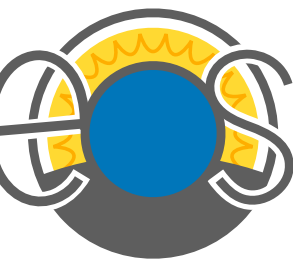


Tank install (Sept 14th)



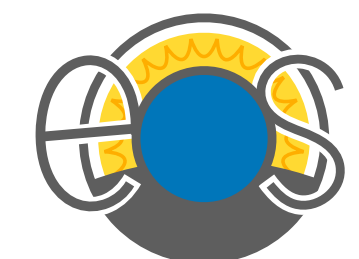


Upper PMT array (Oct 18)



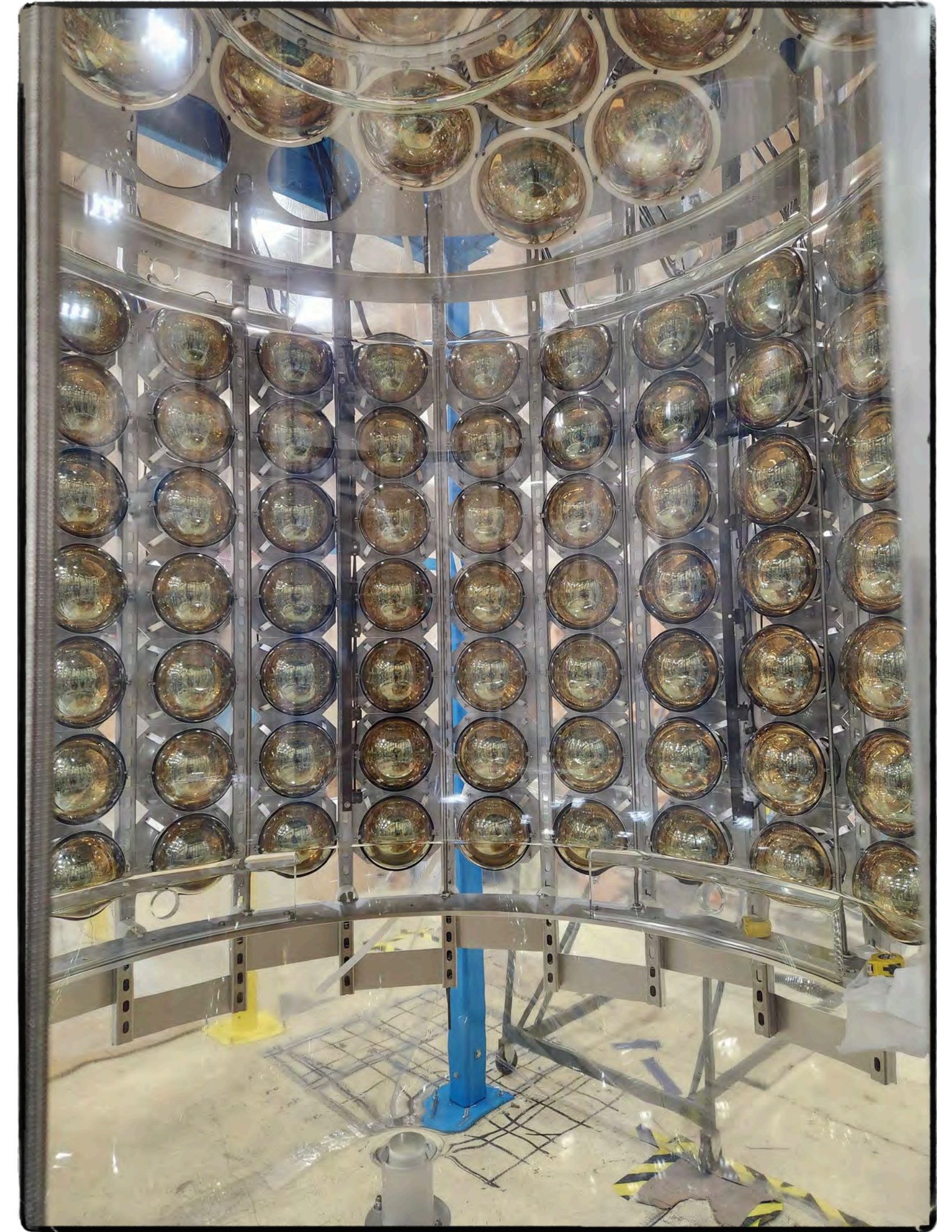
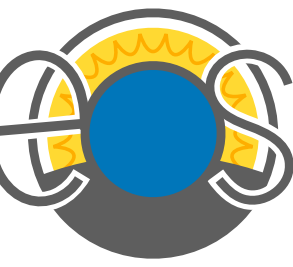


Inner vessel (Oct 25)



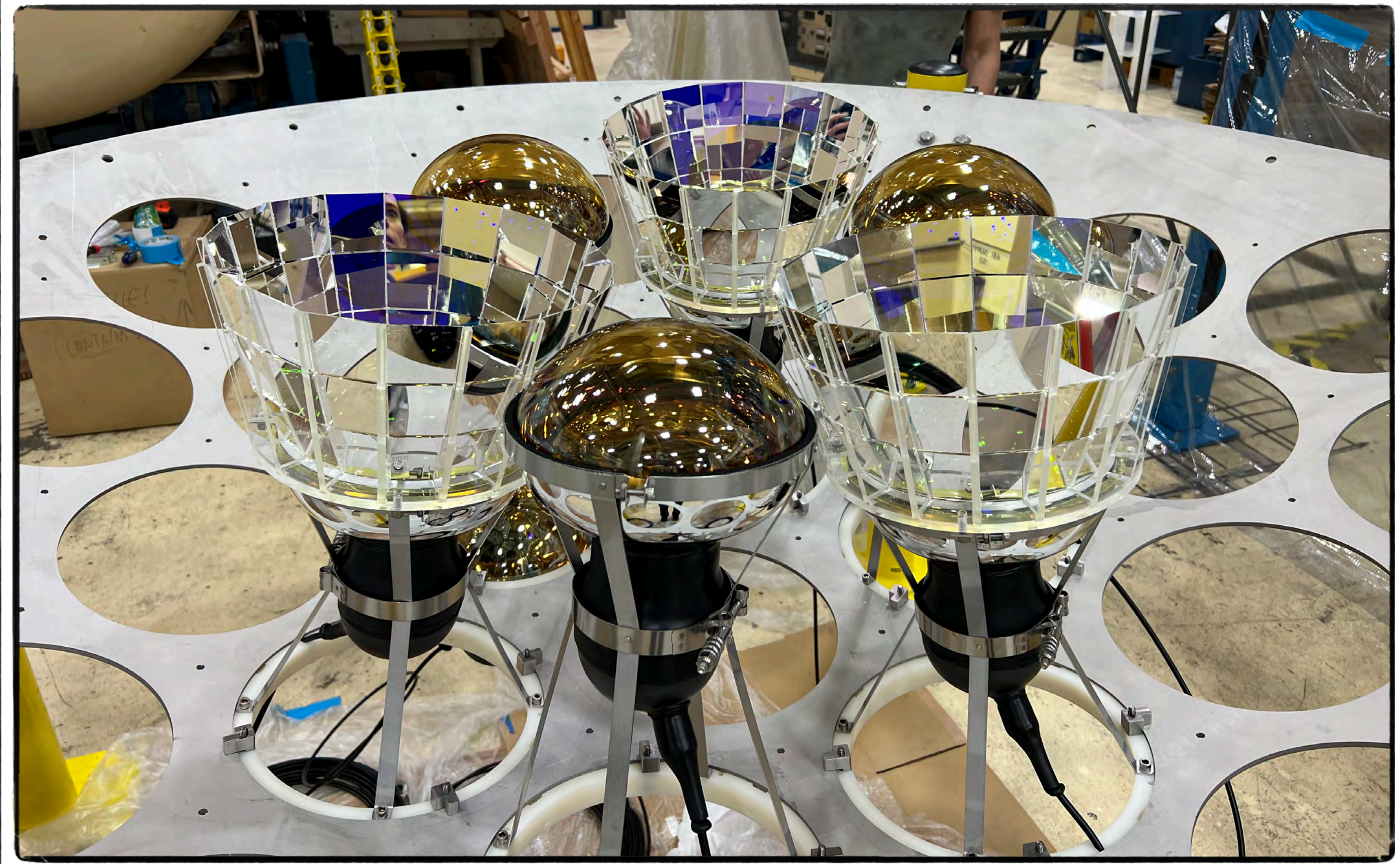
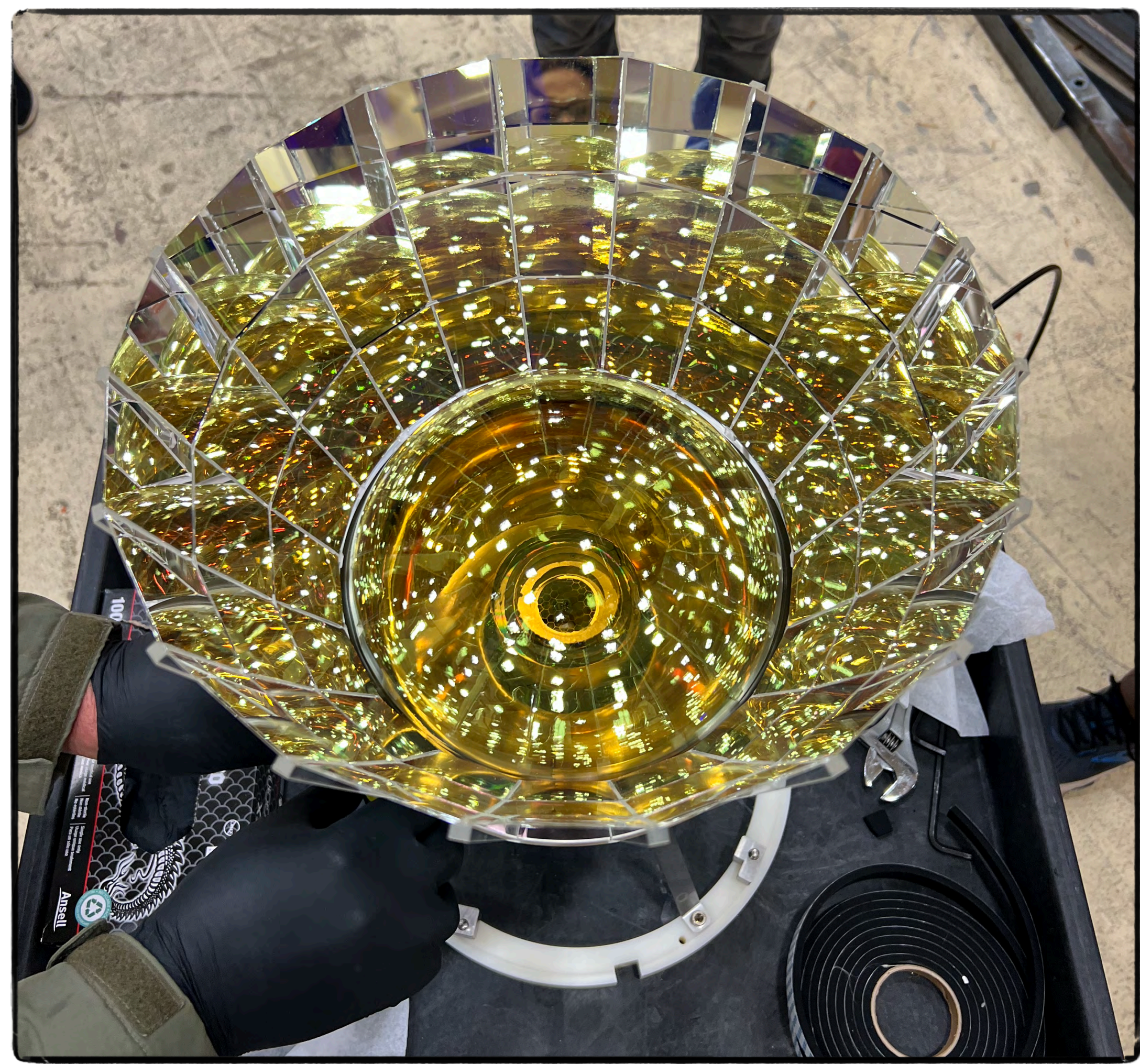
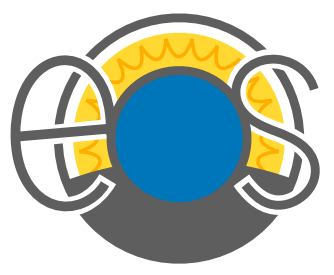


IV unveiling, Barrel PMTs (early Nov)



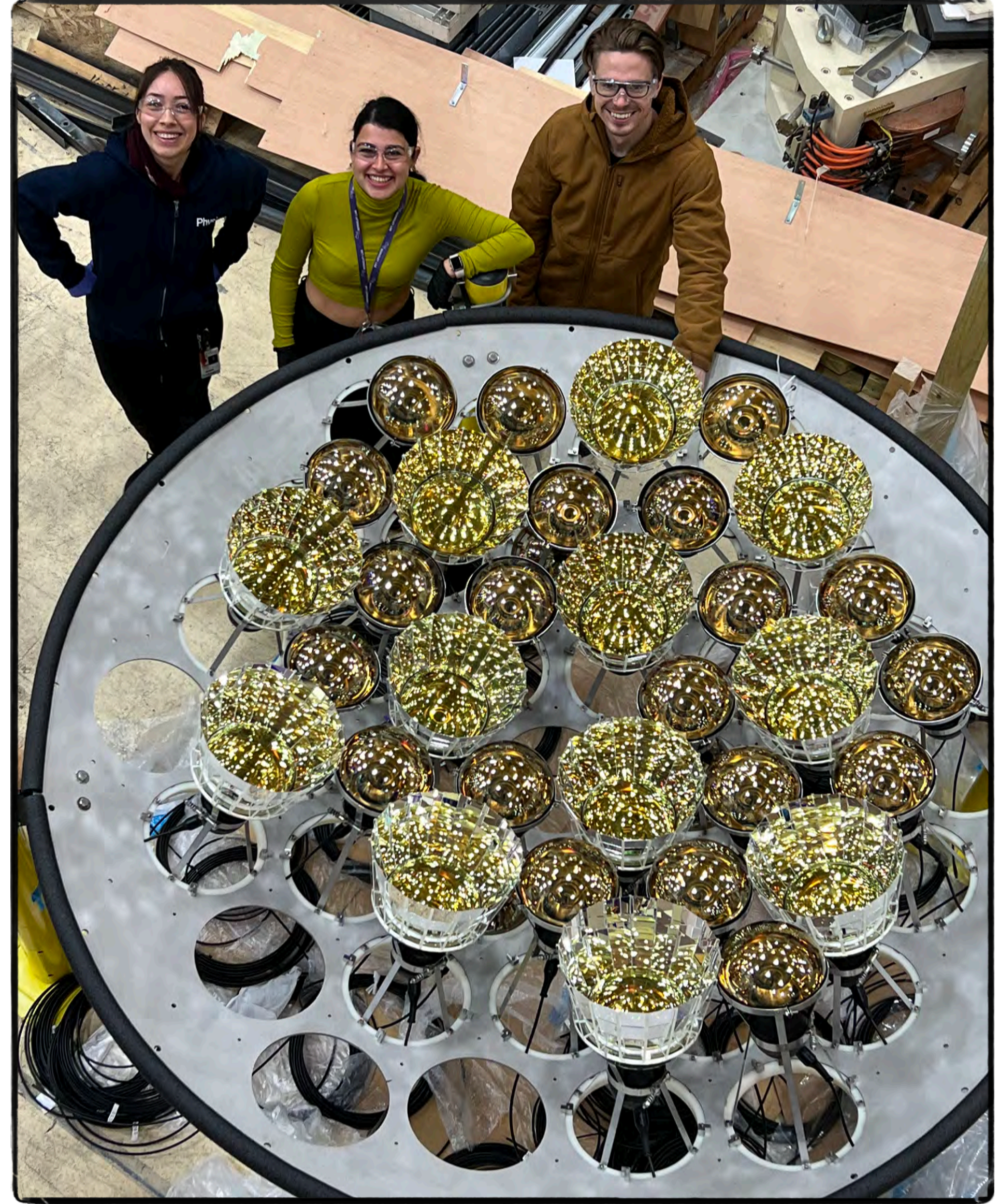
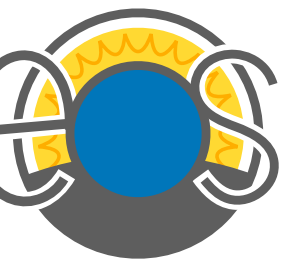


First dichroicon install (Nov 13)

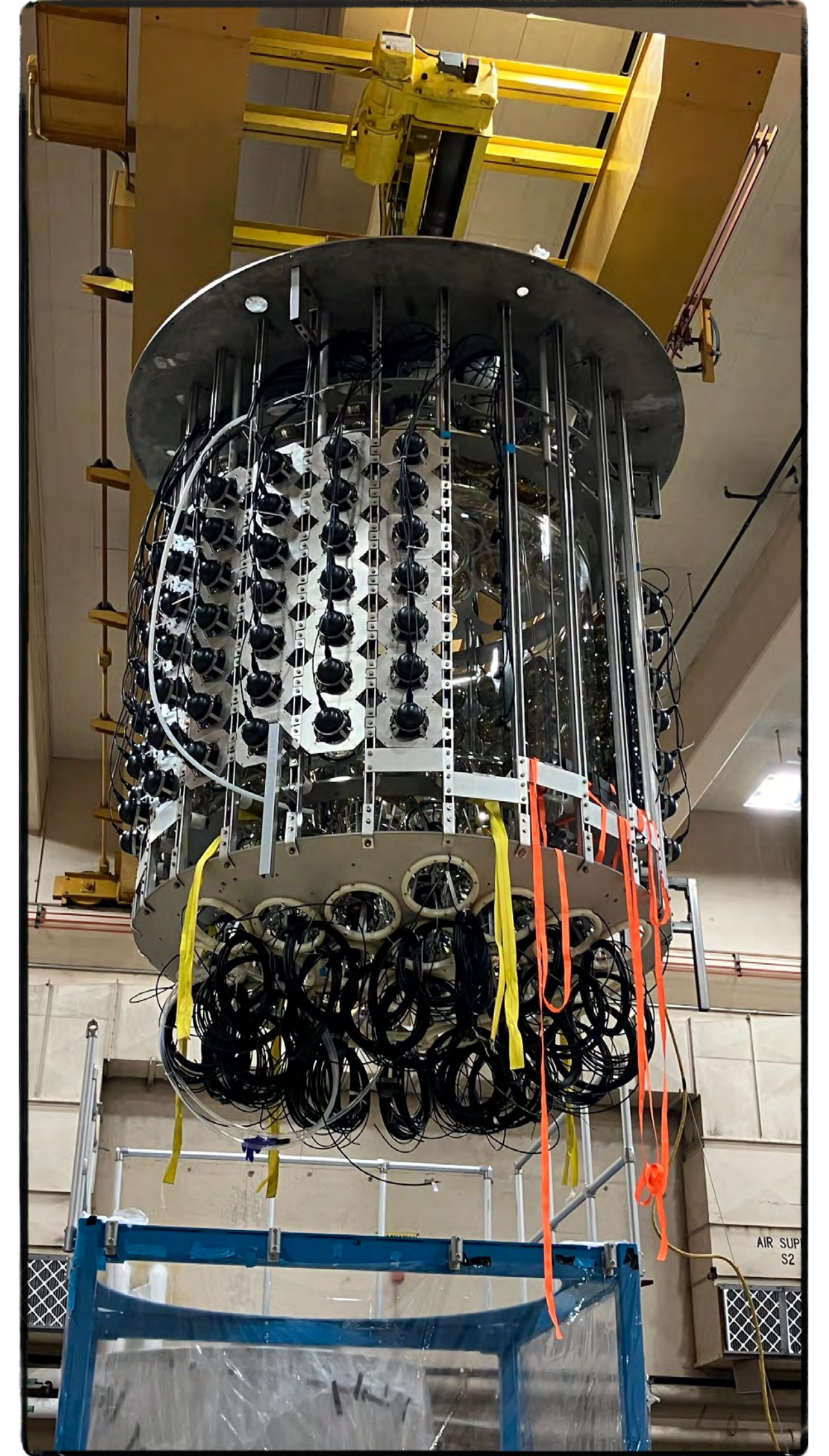
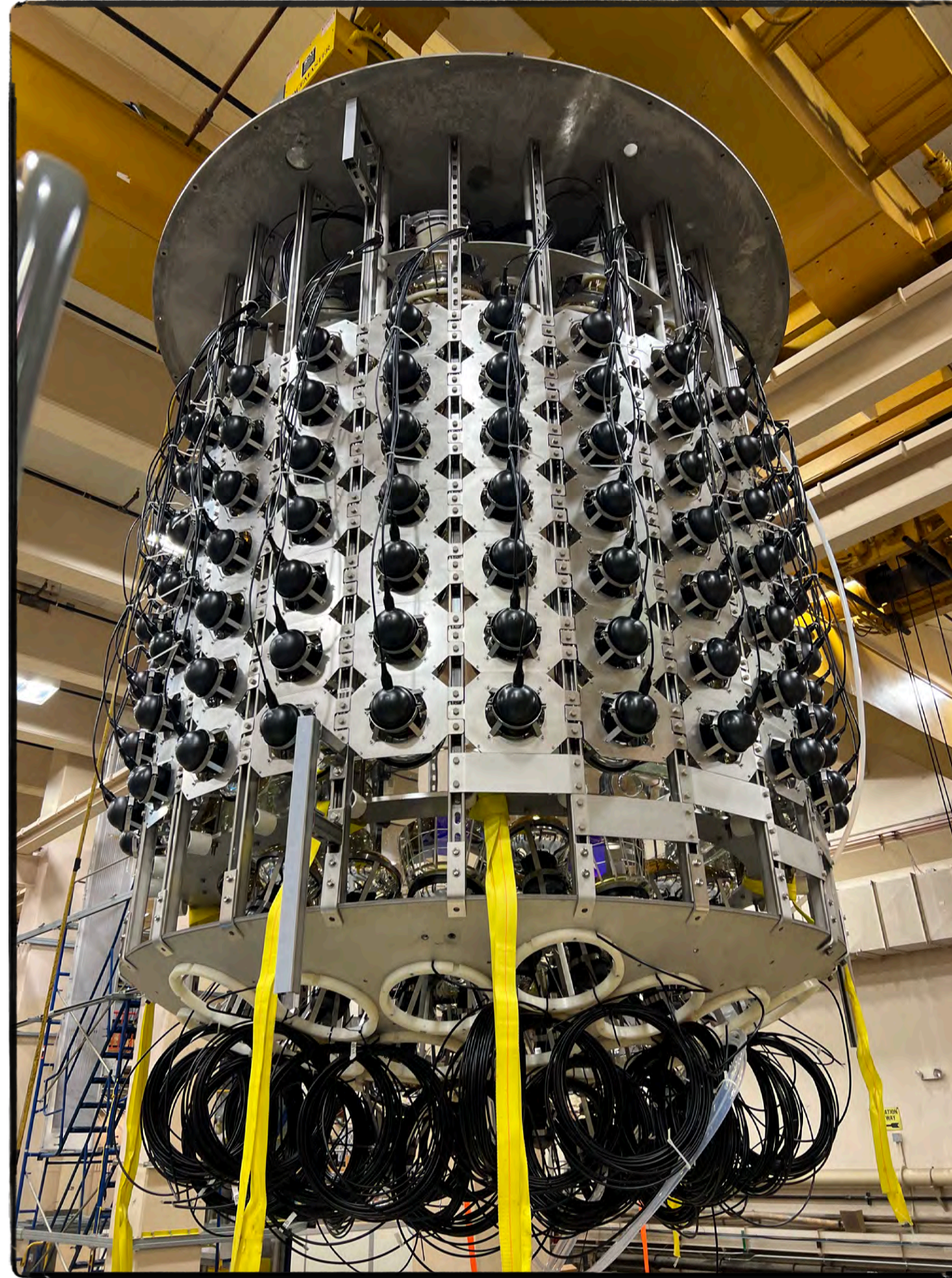




Complete dichroicon install (Nov 17)



Lower array connection (Nov 24)

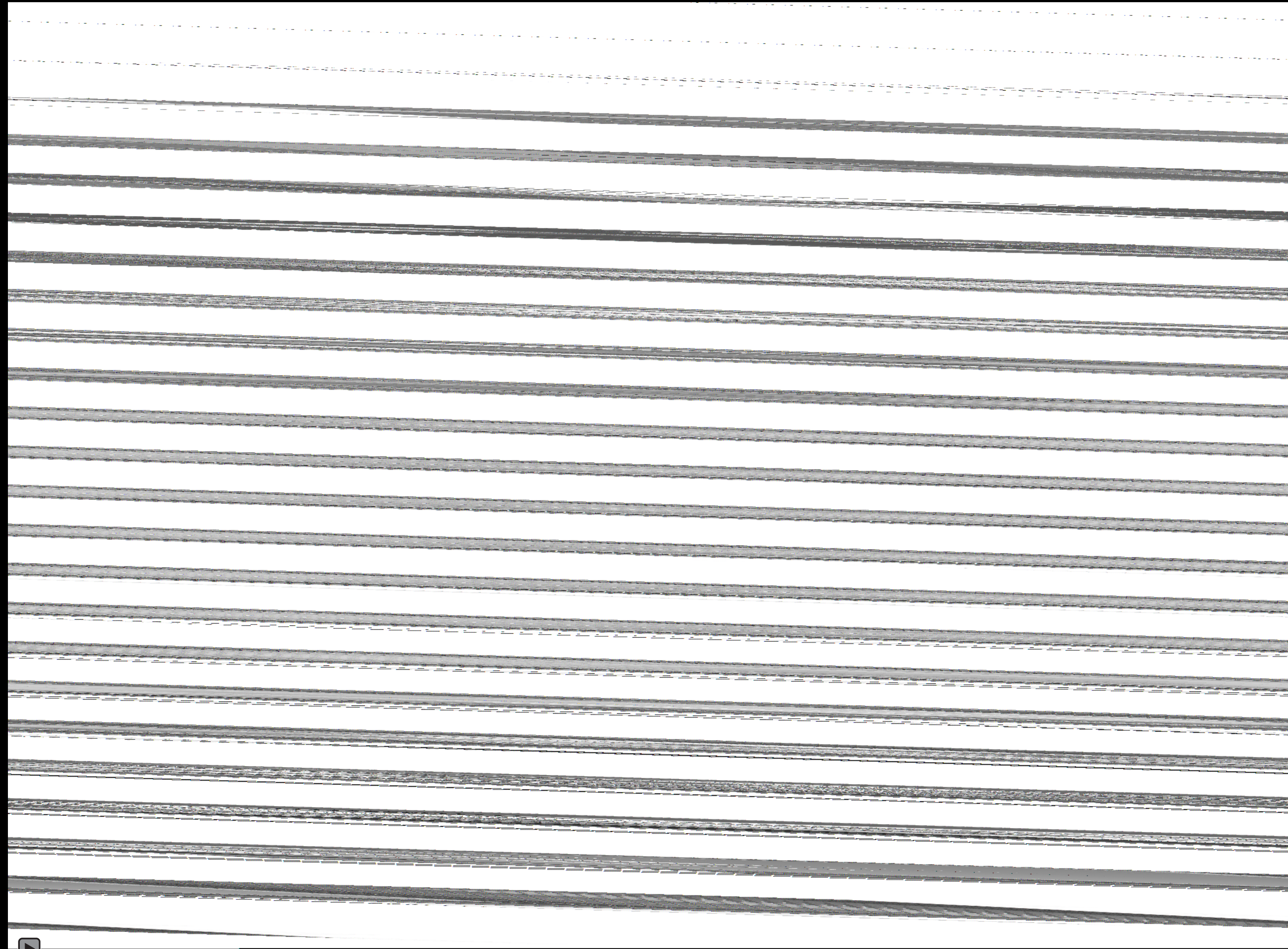




Detector lift (Jan 26)



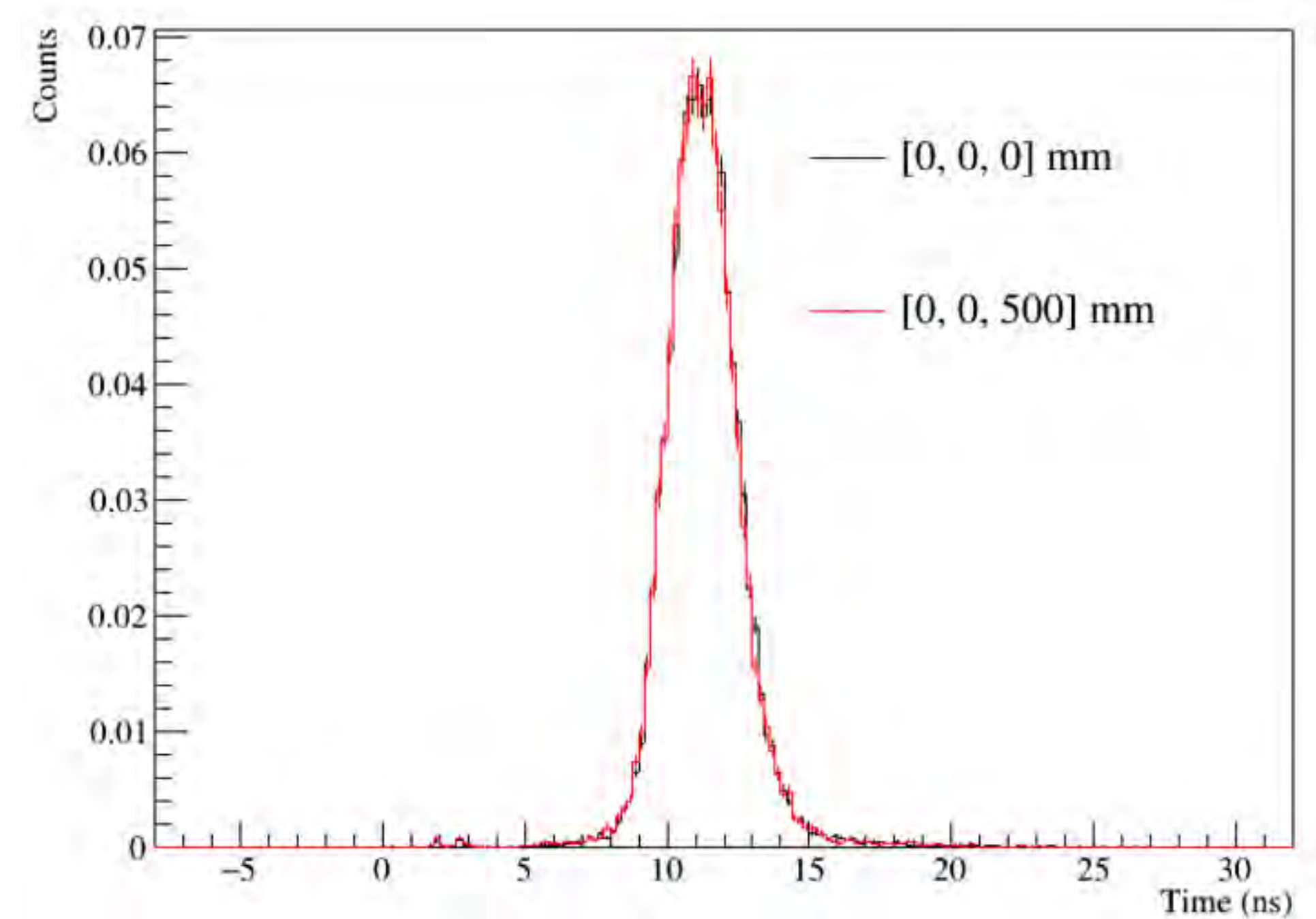
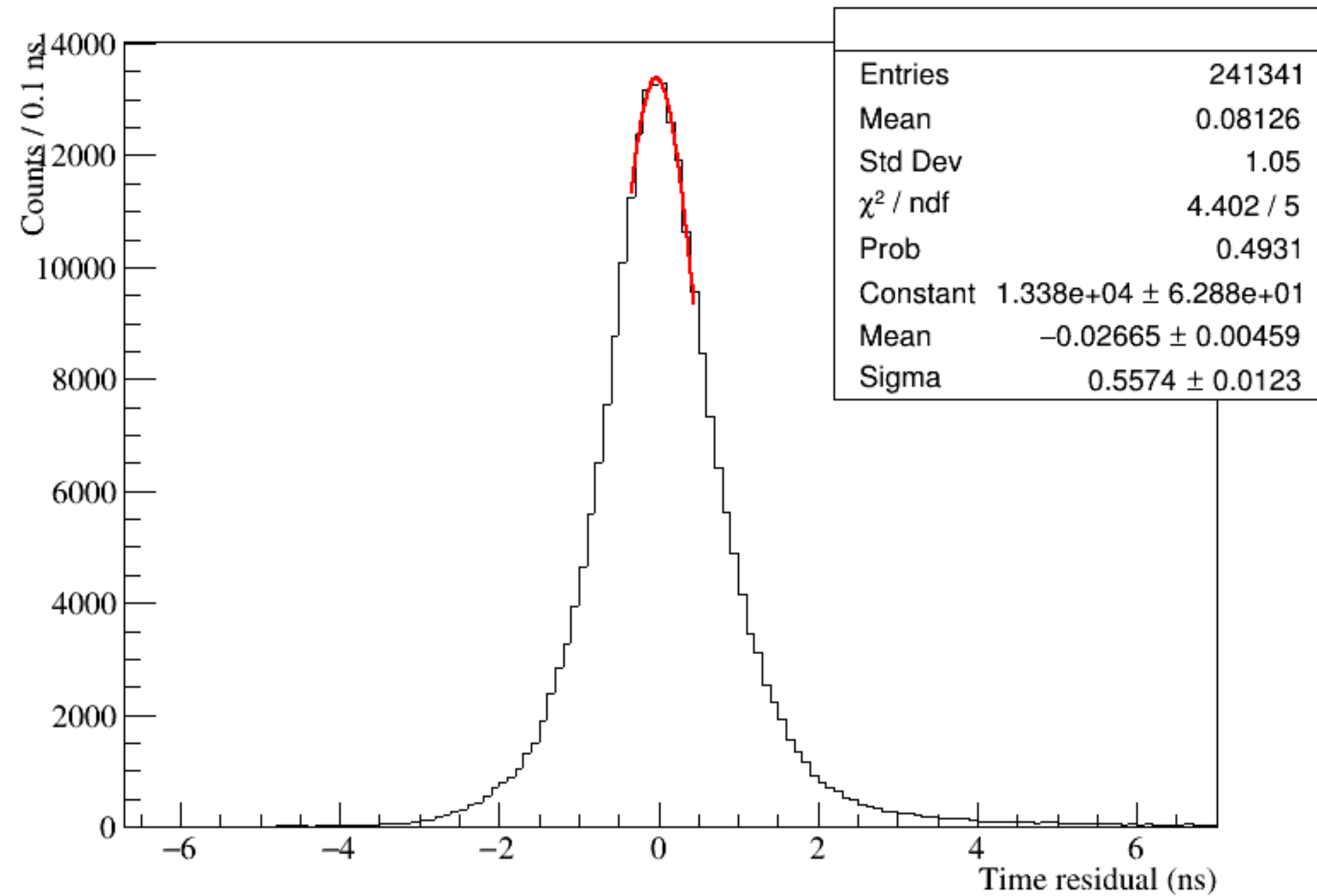
Detector lift (Jan 26)



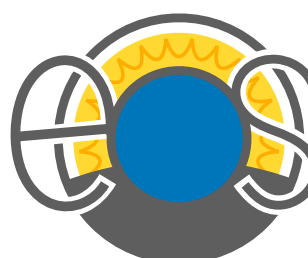
Detector lift (Jan 26)



Preliminary data



- Time precision evaluated using a pico-second laser injected into a single-mode optical fiber, terminating in a teflon diffuser ball
- $< 600\text{ps}$ time resolution across all PMTs (8", 10", 12")
- Stable and reproducible across runs
- Next step: evaluate energy / vertex / direction reconstruction



The EOS team

USA



Team breakdown:

- 3 national laboratories
- 18 top-tier US institutions
- 8 international collaborating institutions
- 10 postdoctoral scholars
- 9 graduate students
- 18 undergraduate students



Germany



Portugal



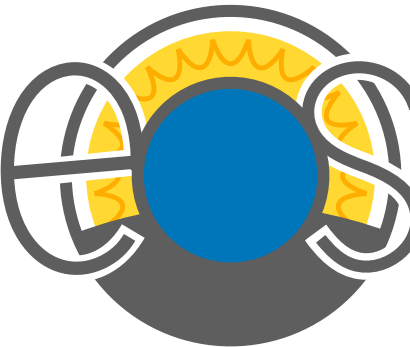
Finland



Turkey



Canada



THEIA (proto-)Collaboration



White paper - Eur. Phys. J. C 80, 416 & arXiv:2202.12839 [hep-ex]

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A. Weber,⁵ M. Wetstein,²³ M. J. Wilking,¹⁵ L. Winslow,³⁴ P. Wittich,⁴⁴ B. Wonsak,⁹
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of Physics and Astronomy, Stony Brook, New York, USA

⁴²Department of Engineering Physics, Tsinghua University, Beijing 100084, China

⁴³Department of Physics, South Dakota School of Mines and Technology, Rapid City, SD 57701, USA

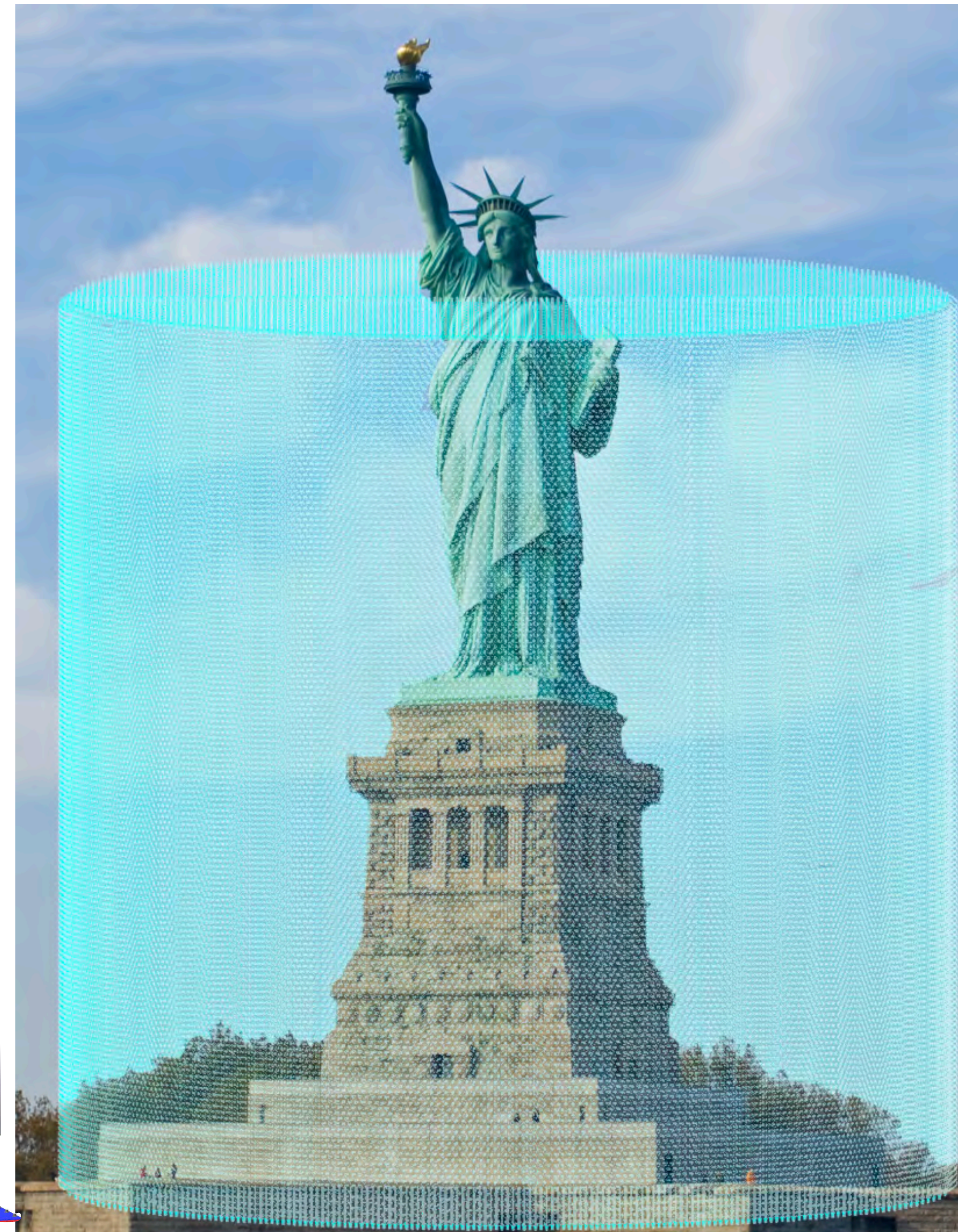
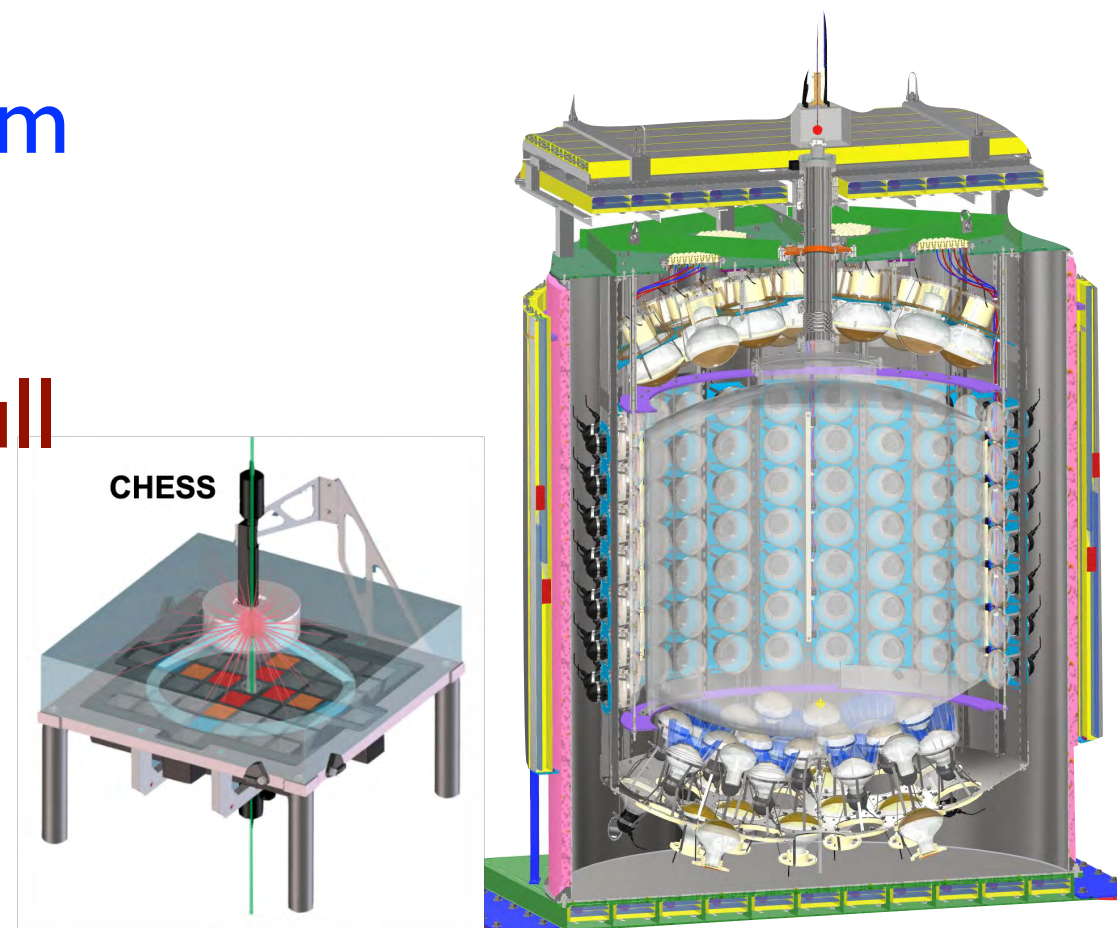
⁴⁴Cornell University, Ithaca, NY, USA

⁴⁵University of Colorado at Boulder, Department of Physics, Boulder, Colorado, USA

⁴⁶Institut für Kern und Teilchenphysik, TU Dresden, Zellescher Weg 19, 01069, Dresden, Germany

Summary

- Hybrid Cherenkov/scintillation technology can interrogate a broad program of compelling science
- Conventional neutrino physics & rare-event searches in a single, large detector
- A hybrid detector module would add to the LBL program at DUNE and bring a broad program of additional physics
- THEIA offers inspirational physics to motivate a new, broad community of scientists
- Technological developments have evolved from bench-top to fully integrated demonstrators
- Prototypes underway will demonstrate the full range of capabilities
- ***Exciting physics to come!***



Non-comprehensive list of citations

- **Physics**

- Eur. Phys. Jour. C (2018) 78: 435
<https://doi.org/10.1140/epjc/s10052-018-5925-7>
- Eur. Phys. Jour. C (2020) 80:418
<https://doi.org/10.1140/epjc/s10052-020-7977-8>
- arXiv:2202.12839 [hep-ex]
- Eur. Phys. Jour. C 82 (2022) 12, 1151
<https://doi.org/10.1140/epjc/s10052-022-11106-1>
- Phys. Rev. D 103, 052004 (2021)
<https://doi.org/10.1103/PhysRevD.103.052004>
- NIMA 943 (2019) 162420
<https://doi.org/10.1016/j.nima.2019.162420>
- JINST 9 (2014) P06012
<https://doi.org/10.1088/1748-0221/9/06/P06012>
- <https://doi.org/10.48550/arXiv.1409.5864>

- **Photon detection**

- <https://doi.org/10.48550/arXiv.2203.07479>
- <https://doi.org/10.48550/arXiv.1603.01843>
- NIMA 958 (2020) 162834
<https://doi.org/10.1016/j.nima.2019.162834>
- JINST 19 (2024) P02032
<https://doi.org/10.1088/1748-0221/19/02/P02032>

- **Scintillator development / studies**

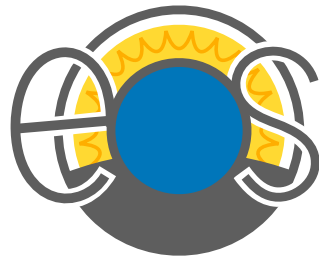
- Phys. Rev. D 105 (2022) 9
<https://doi.org/10.1103/PhysRevD.105.092006>
- NIMA 972 (2020) 164106
<https://doi.org/10.1016/j.nima.2020.164106>
- Materials Advances, 1 (2020) 71 - 76
<https://doi.org/10.1039/D0MA00055H>
- Eur. Phys. Jour. C (2023) 83:134
<https://doi.org/10.1140/epjc/s10052-023-11242-2>
- <https://doi.org/10.48550/arXiv.2405.01100>
- JINST 16 (2021) P05009
<https://doi.org/10.1088/1748-0221/16/05/P05009>
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<https://doi.org/10.1016/j.nima.2019.01.014>
- Astro.Part. Phys. 109 (2019) 33
<https://doi.org/10.1016/j.astropartphys.2019.02.00>
- J. Phys. G: Nucl. Part. Phys. 43 (2016) 093001.
<https://doi.org/10.1088/0954-3899/43/9/093001>
- NIMA 830 (2016) 303
<https://doi.org/10.1016/j.nima.2016.05.132>
- JINST 10 (2015) P12009
<https://doi.org/10.1088/1748-0221/10/12/P12009>
- <https://doi.org/10.48550/arXiv.2405.05743>

- **Technology demonstrators**

- JINST 18 (2023) P02009
<https://doi.org/10.1088/1748-0221/18/02/P02009>
- Phys. Rev. D 109 (2024) 072002
<https://doi.org/10.1103/PhysRevD.109.072002>
- JINST 19 (2024) 05, P05070.
<https://doi.org/10.1088/1748-0221/19/05/P05070>
- <https://doi.org/10.48550/arXiv.2403.13231>
- Phys. Rev. C 95, 055801 (2017)
<https://doi.org/10.1103/PhysRevC.95.055801>
- Eur. Phys. J. C (2017) 77: 811
<https://doi.org/10.1140/epjc/s10052-017-5380-x>
- Eur. Phys. J. C 80, 867 (2020)
<https://doi.org/10.1140/epjc/s10052-020-8418-4>
- Eur. Phys. J. C 82-2 (2022) 169
<https://doi.org/10.1140/epjc/s10052-022-10087-5>
- Eur. Phys. Jour. C (2023) 83:1094
<https://doi.org/10.1140/epjc/s10052-023-12278-0>
- Phys. Rev. D 105 (2022) 7
<https://doi.org/10.1103/PhysRevD.105.072003>
- <https://doi.org/10.48550/arXiv.2312.09293>

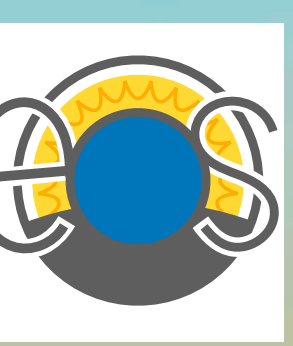


Acknowledgements



This work was performed under the auspices of the U.S. Department of Energy by Lawrence Berkeley National Laboratory under Contract DE-AC02-05CHI1231. The project was funded by the U.S. Department of Energy, National Nuclear Security Administration, Office of Defense Nuclear Nonproliferation Research and Development (DNN R&D) (FY19, FY20-24). This material is based upon work supported by the U.S. Department of Energy, Office of Science, Office of High Energy Physics, under Award Number DE-SC0018974 (FY18-20). This work was funded in-part by the Consortium for Monitoring, Technology, and Verification under Department of Energy National Nuclear Security Administration award number DE-NA0003920 (FY20-24), and the Nuclear Science and Security Consortium under Award Number DE-NA0003180 (FY19-20).





Thank you!

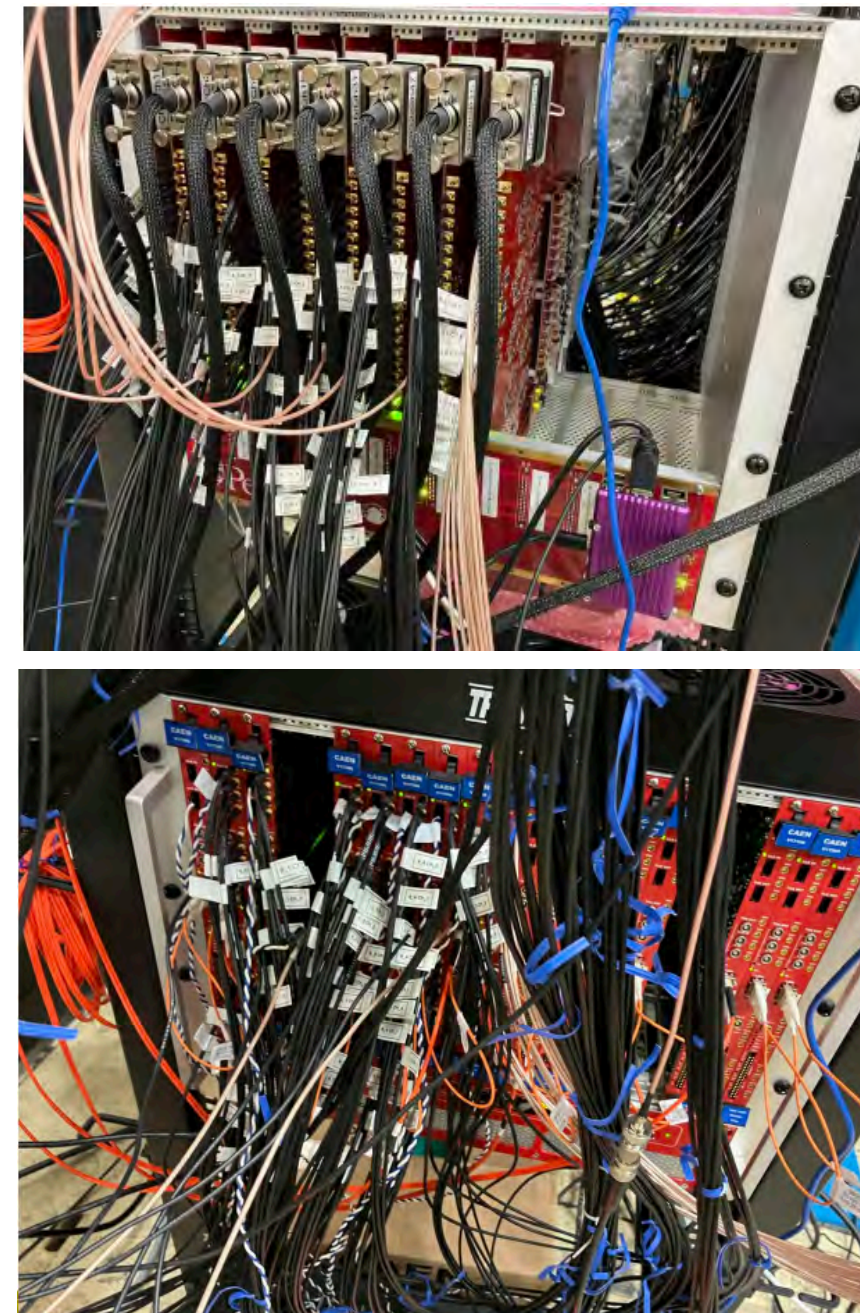
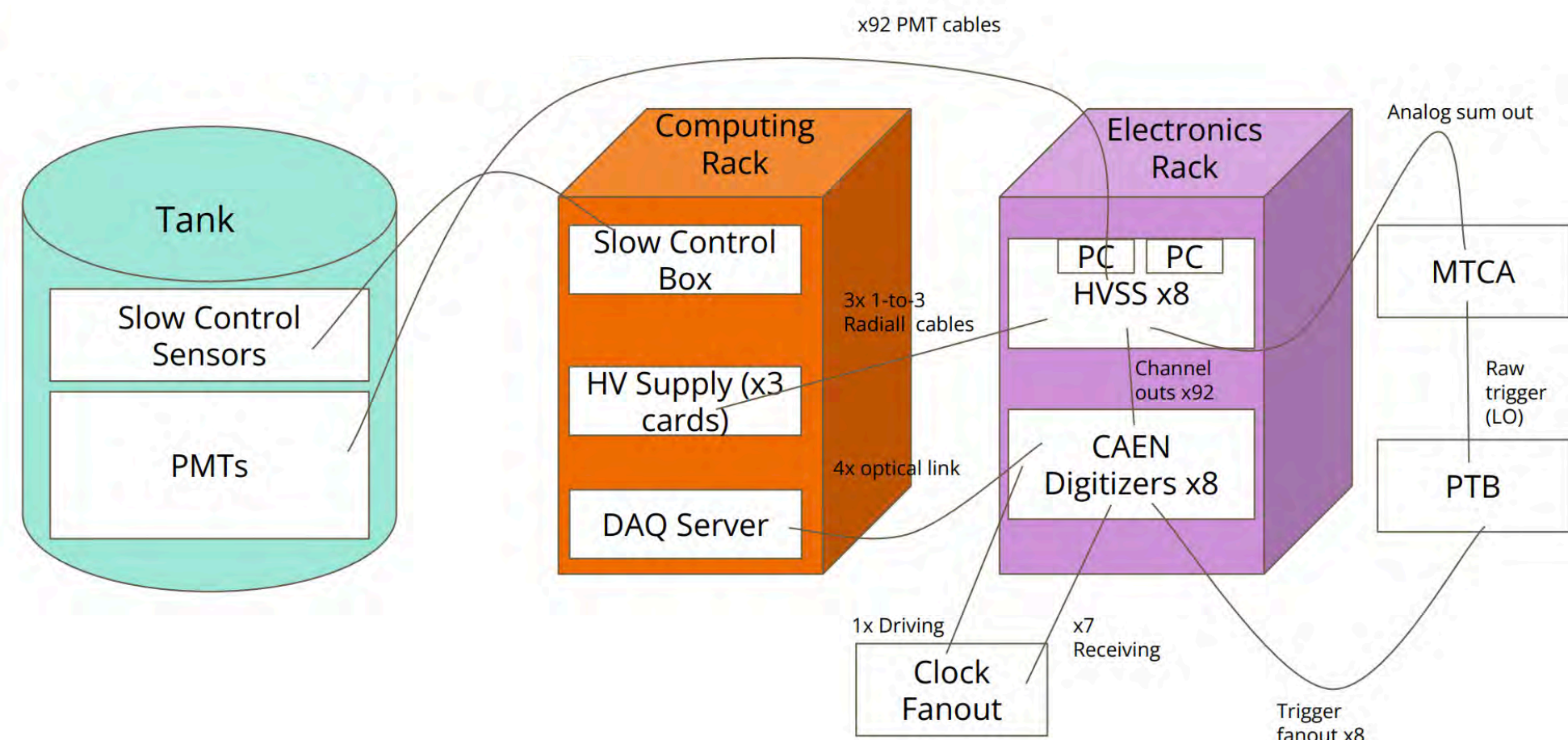
- *Thank you to DNN R&D: Chris Ramos, Brian Yoxall, Craig Sloan for their support*
- *Thank you to DOE-SC HEP: Helmut Marsiske, Brian Beckford, Alan Stone, Glen Crawford*
- *Thank you to DOE-SC NP: Paul Sorensen, Tim Hallman*
- *Thank you to this audience for your attention!*



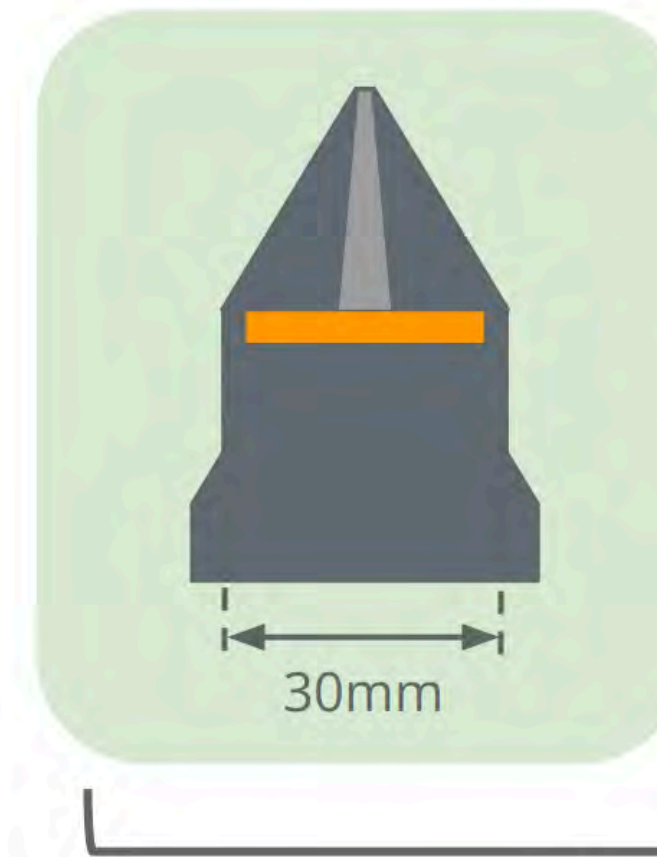
BACKUP

Subsystem status

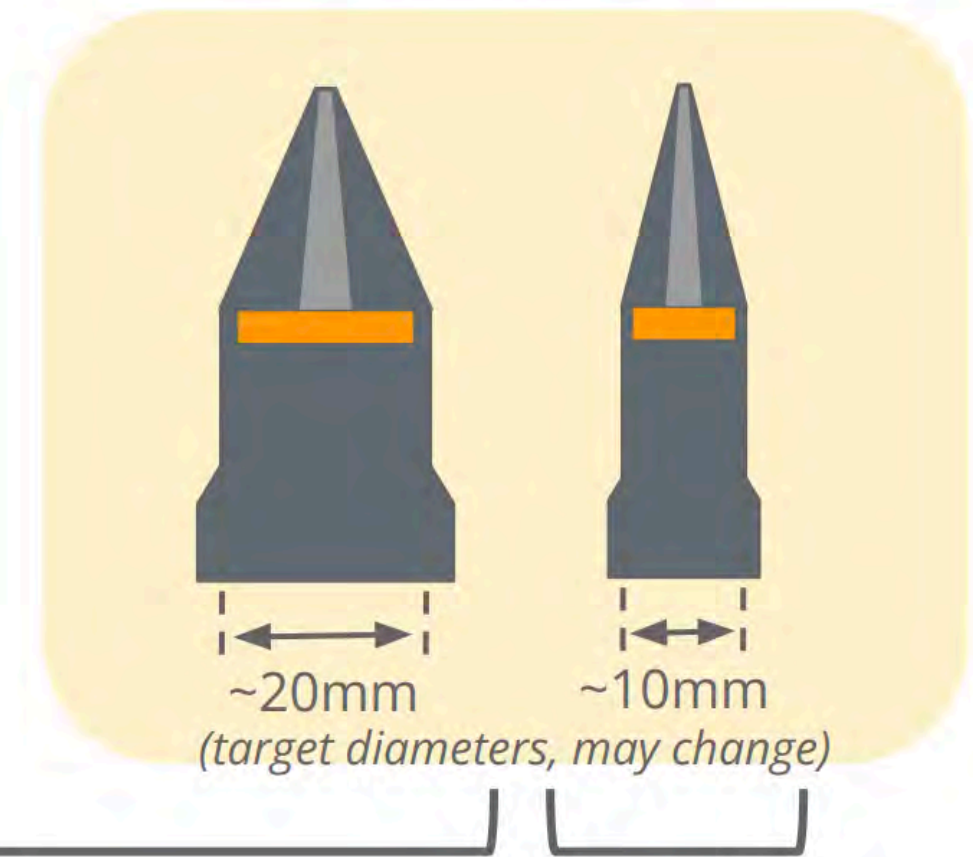
Electronics / readout



Current design



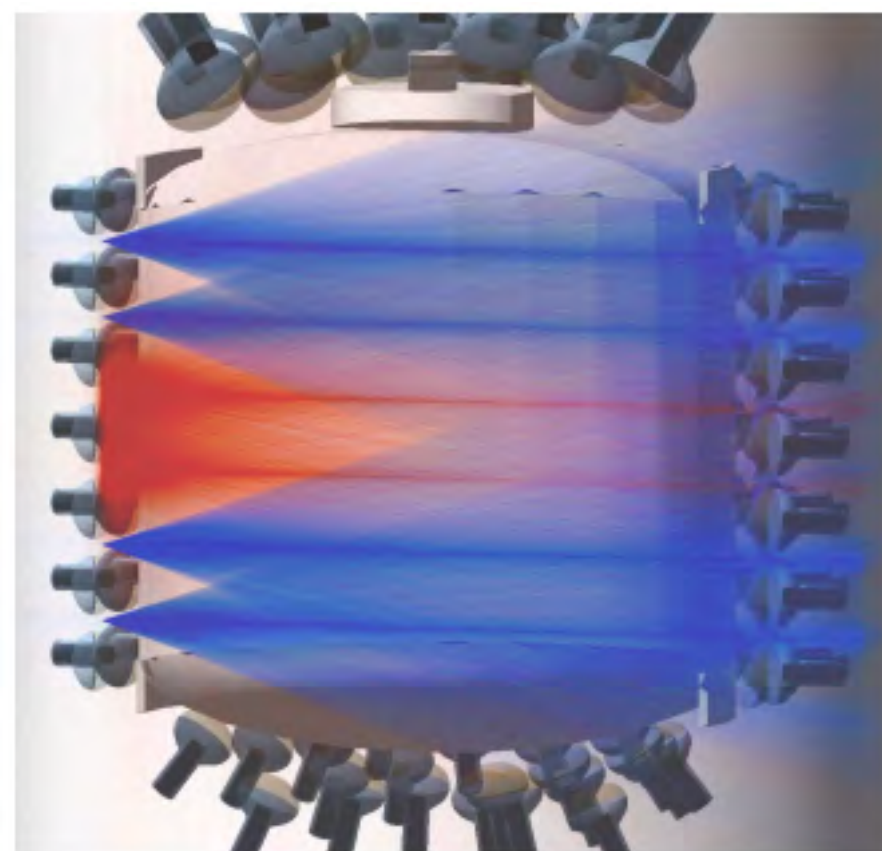
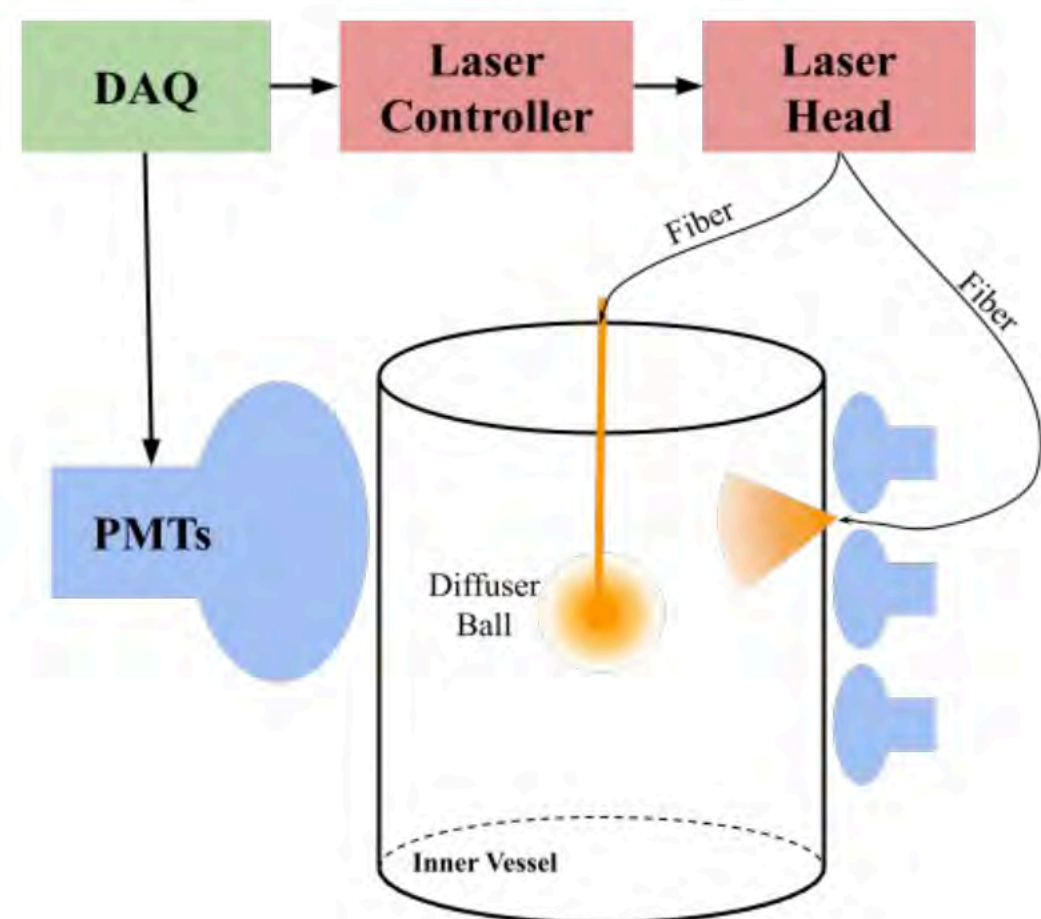
Next source designs



185kBq

37kBq

Laser light injection



Calibration

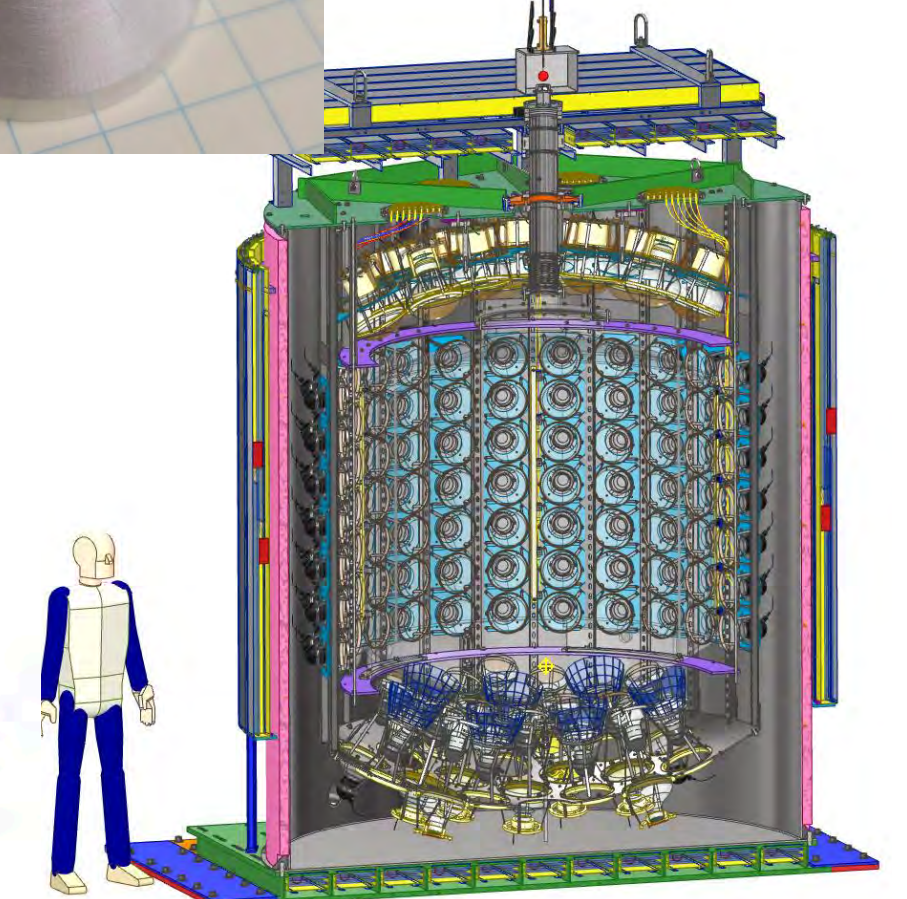
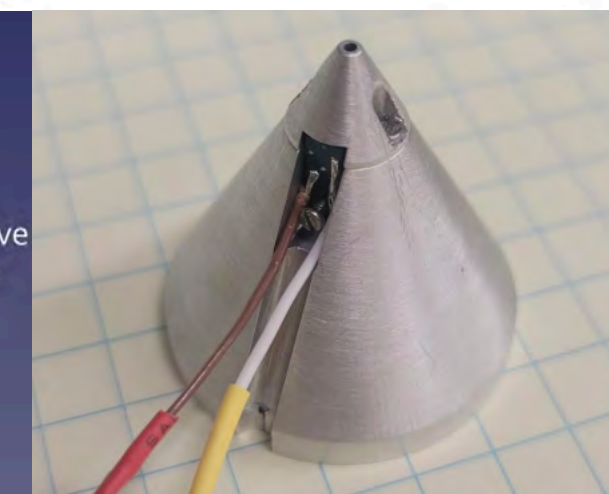
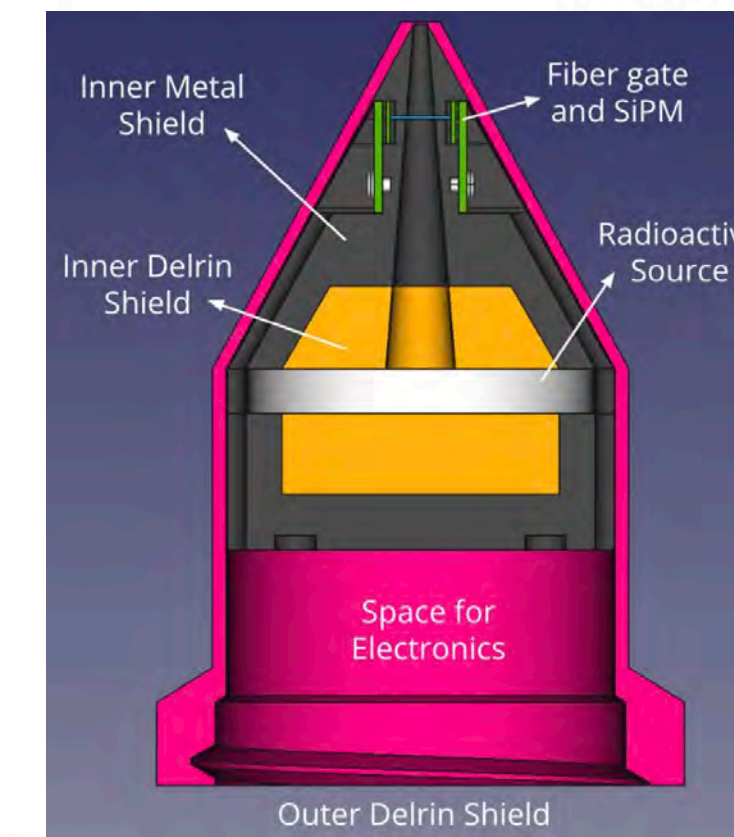
β - $^{90}\text{Sr}/^{106}\text{Ru}$

Low-energy γ - ^{137}Cs

High-energy γ 's - AmBe/PuBe

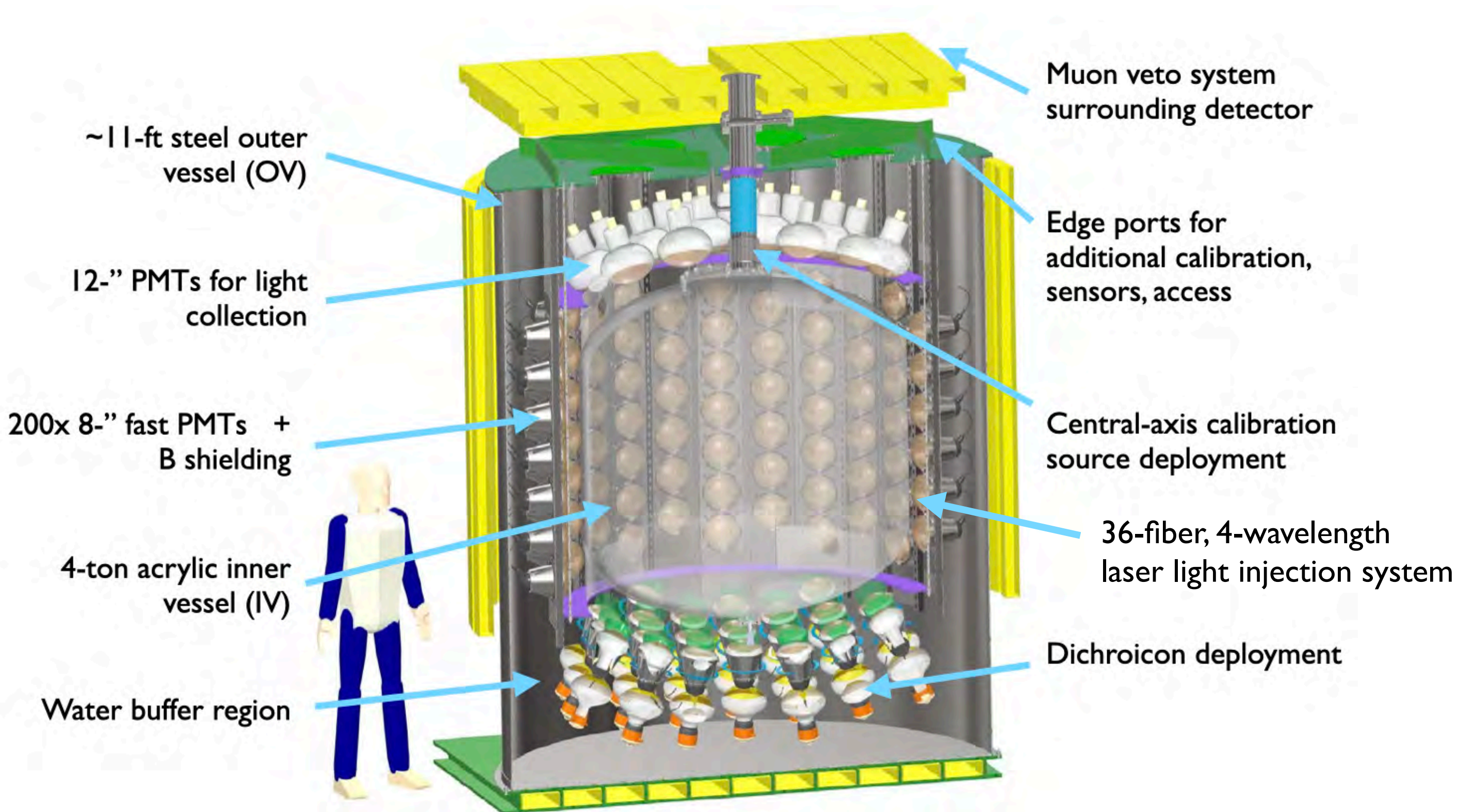
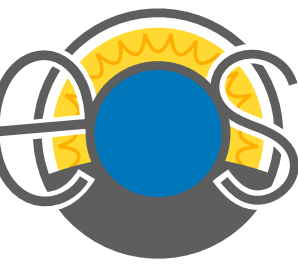
Light injection system - mounted fibers, deployed diffuser

Natural sources - cosmogenics, radioactivity

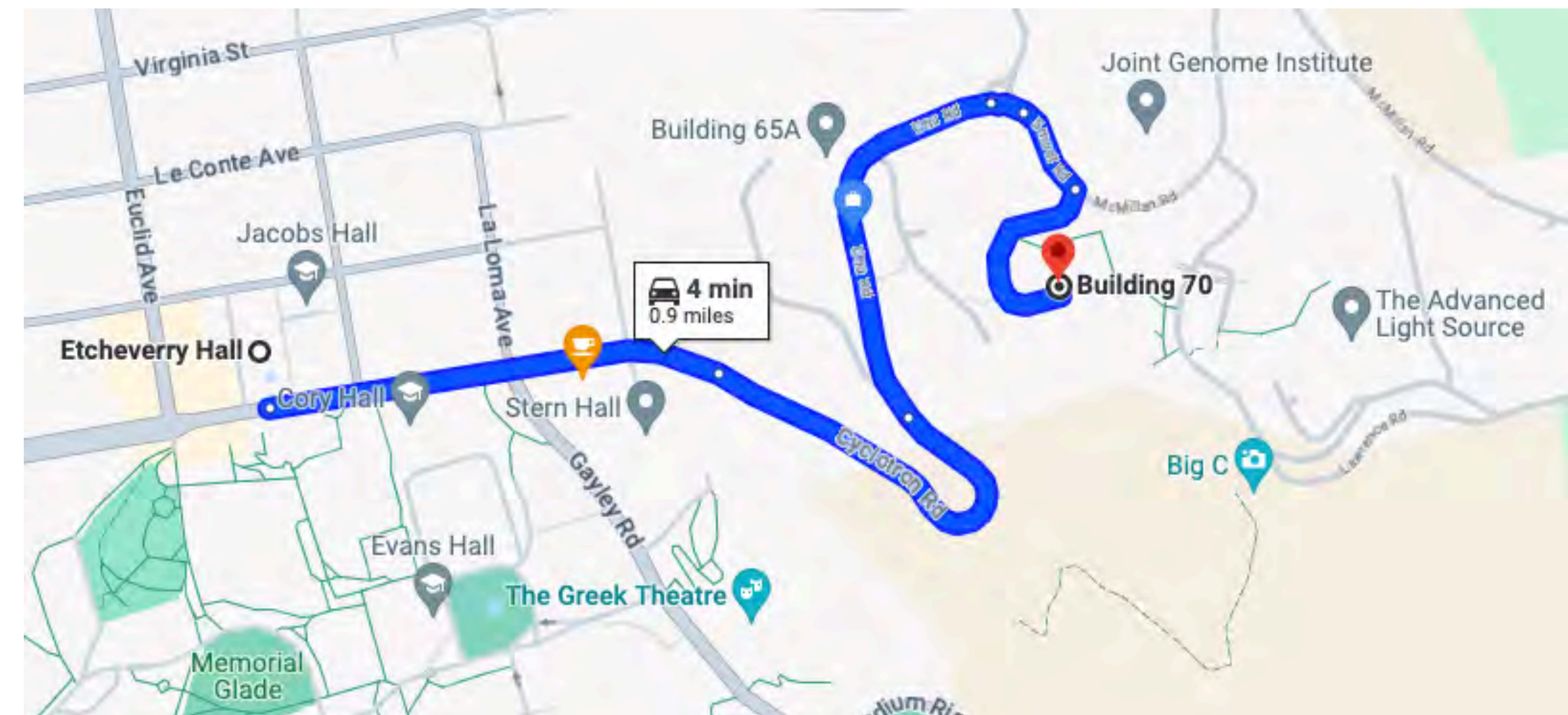


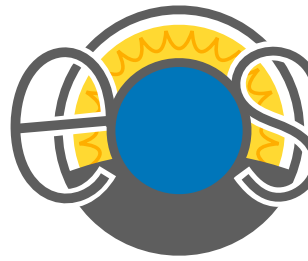


EOS detector design & site selection



- Sited on UC Berkeley campus, in Nuclear Engineering (NE) department
- Supported by NSSC & NE colleagues
- High-bay area with 5-ton capacity bridge crane





Community planning

Snowmass Neutrino Frontier

- *NF: Detectors: one of 5 priorities for development*
- *NF: Long-term outlook also BSM, terrestrial & astrophysical nu*

- **Pursuit of hybrid Cherenkov/scintillation Detectors:** Many different technologies are being developed for these, including water-based liquid scintillator, slow fluors, fast timing with LAPPDs and other devices, and spectral photon sorting with dichroicons. At very large scales like the proposed **Theia** detector, these could have very broad physics programs.

ordering region. These next-generation LAr detector ideas go by different names, such as “SLoMo”, “SoLAr”, and “LArXe.” Another idea is the proposed **Theia** detector, which is a hybrid Cherenkov/scintillation detector that could do precision measurements of very low-energy solar neutrinos, diffuse supernova neutrino detection, perform searches for sterile neutrinos, and also push well beyond DUNE in precision tests of the three-flavor mixing model (including, for example, studies of the second oscillation maximum). On the low-threshold side, much will depend on what the

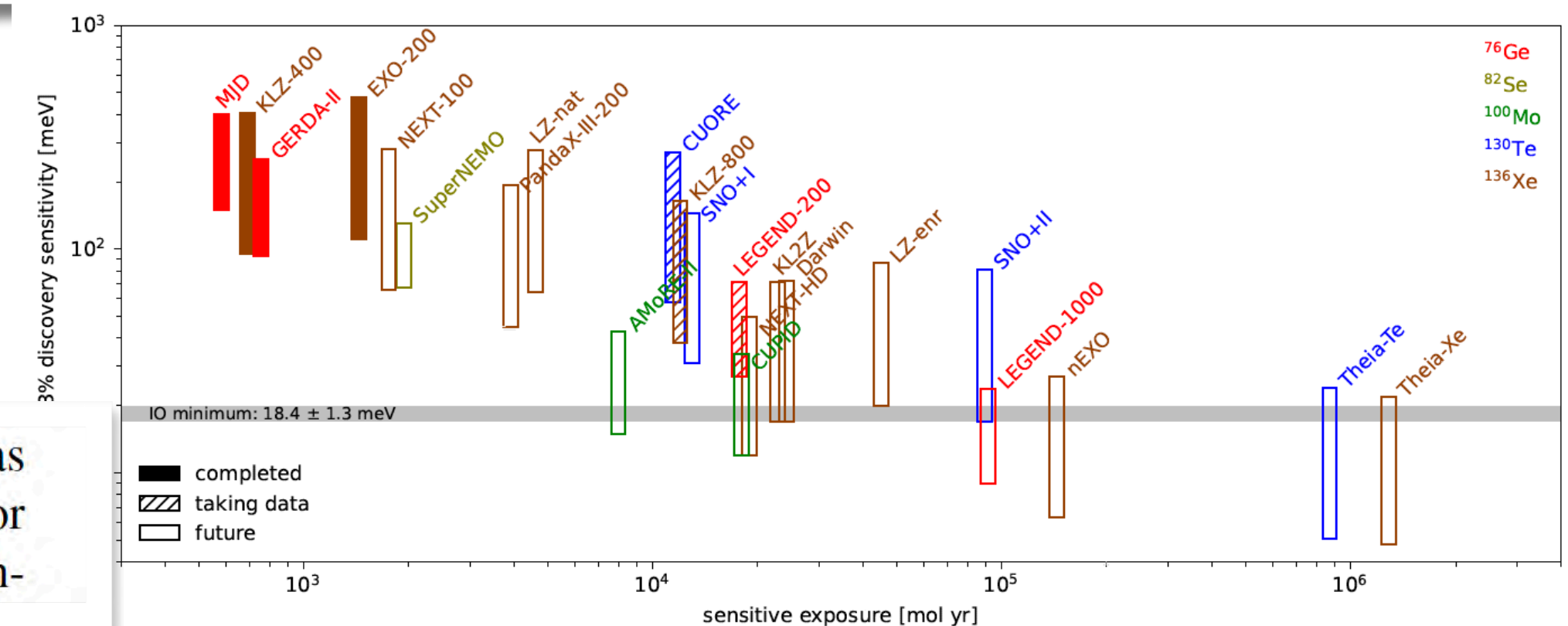
Snowmass Instrumentation Frontier

- *IF: Going beyond DUNE also photon detectors & spectral sorting*

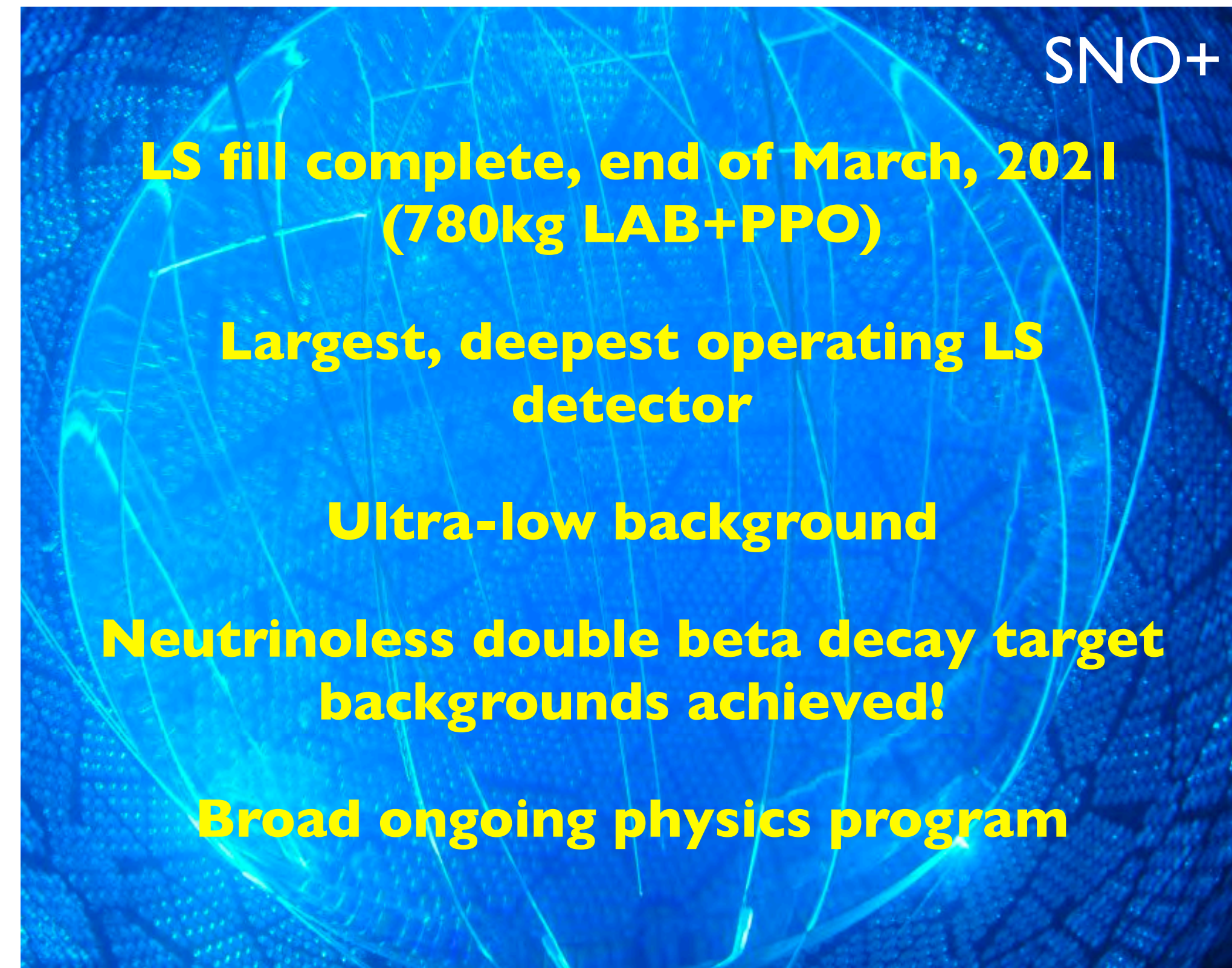
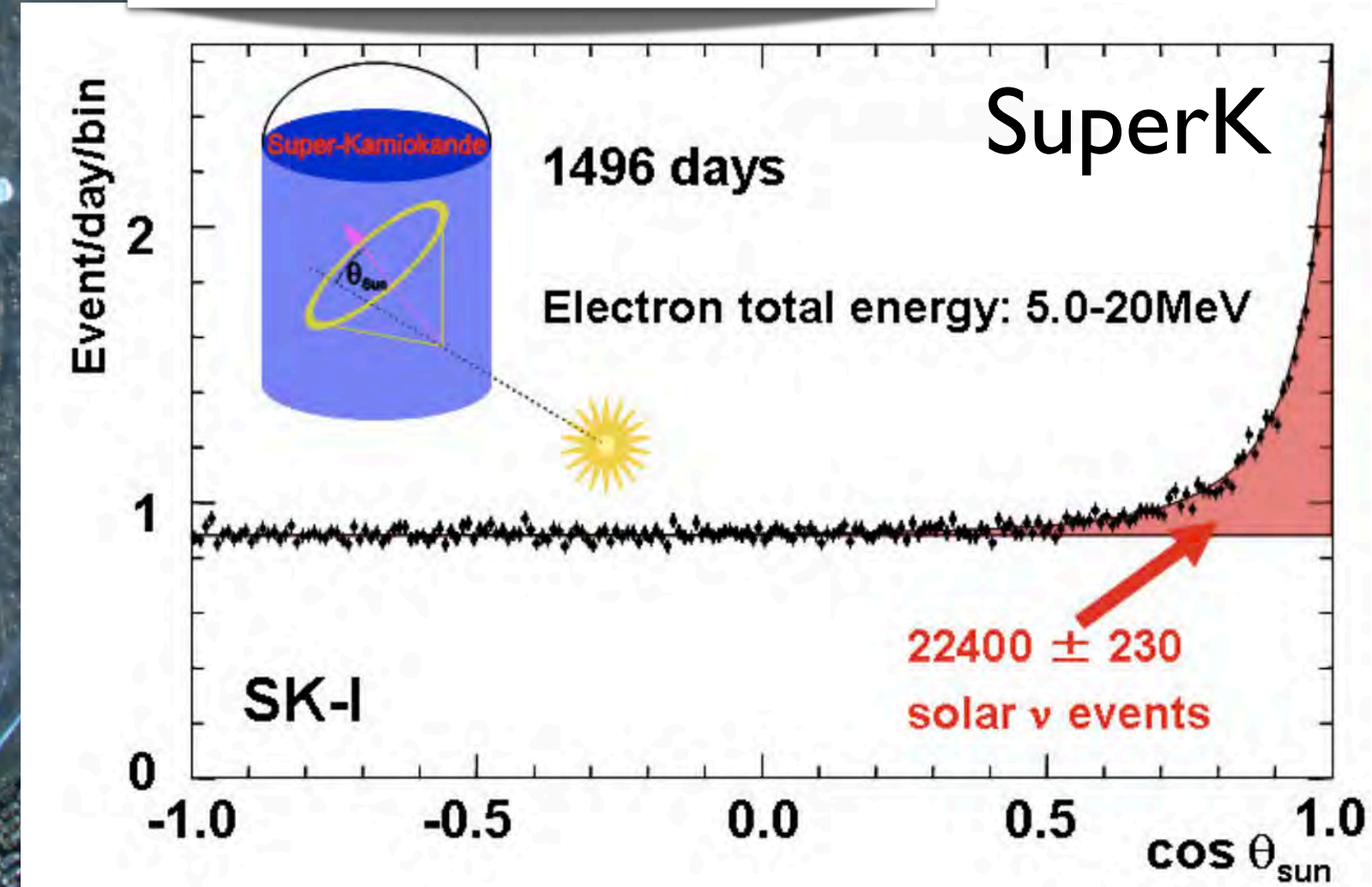
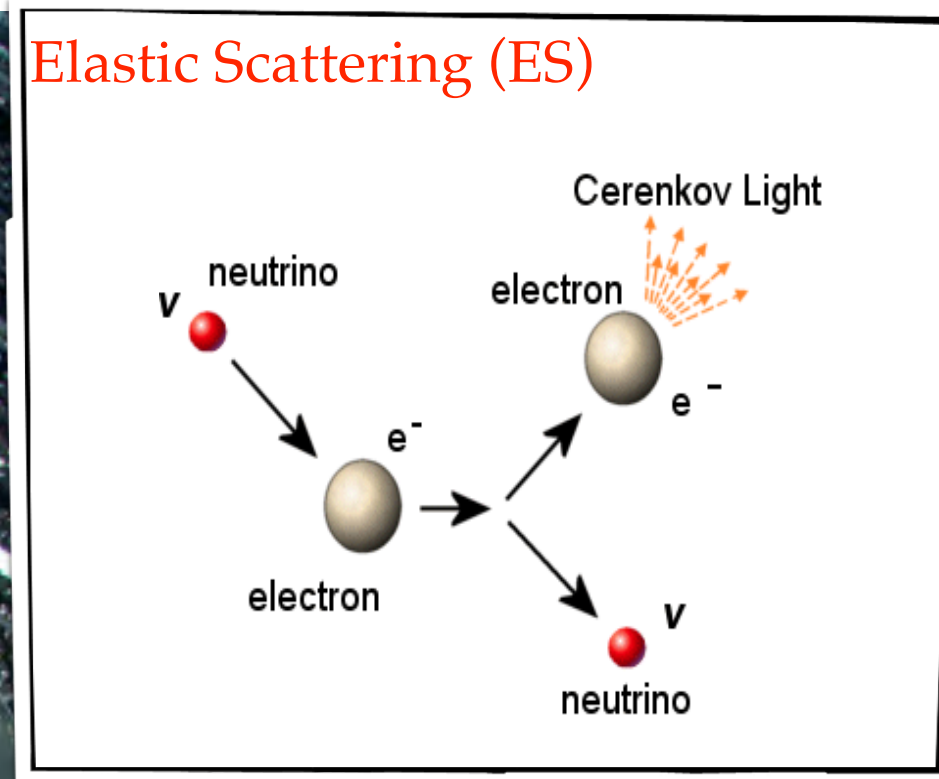
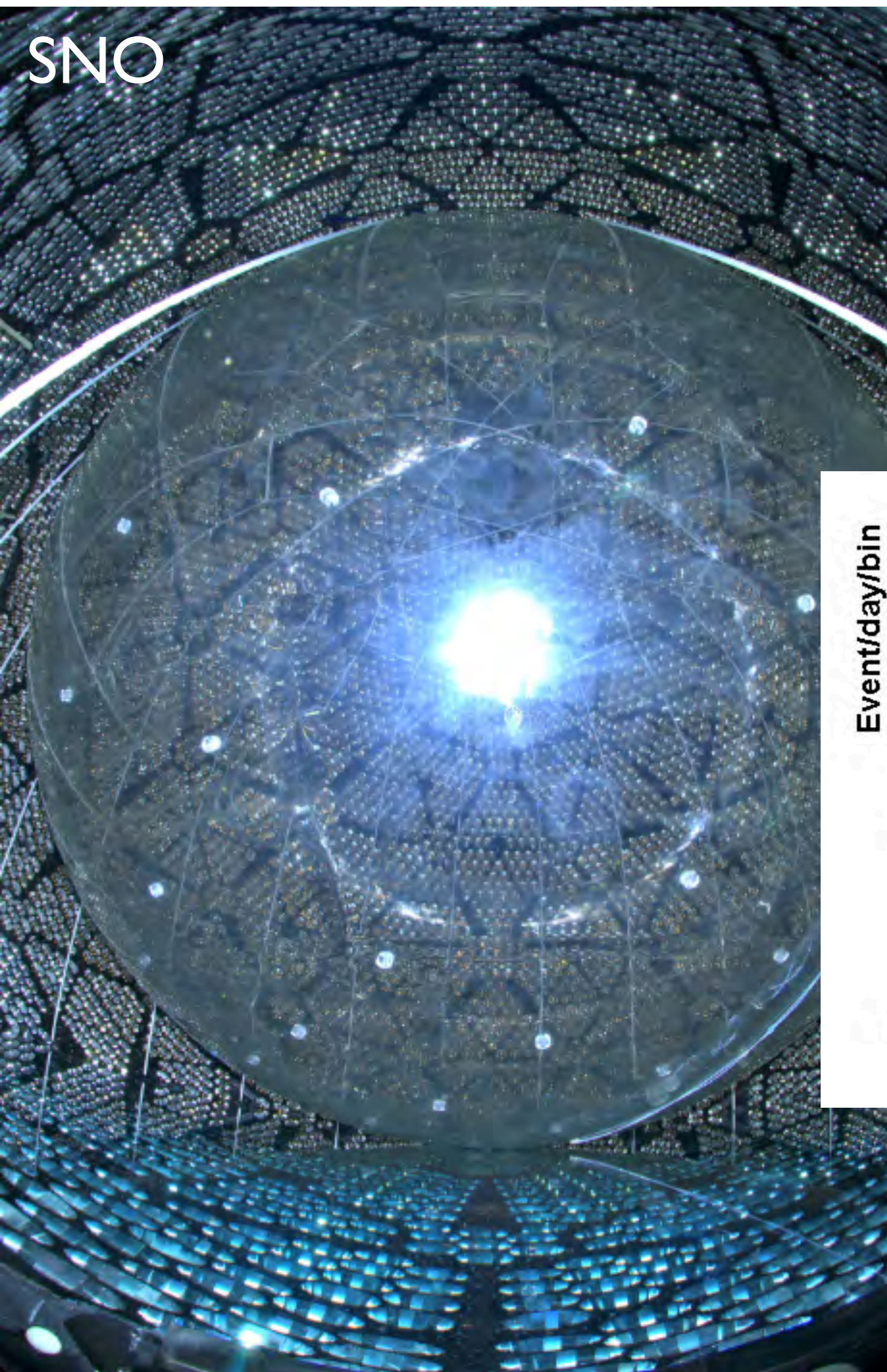
Of particular community interest is the development of hybrid Cherenkov-scintillation detectors, which can simultaneously exploit the advantages of Cherenkov light’s reconstruction of direction and related high-energy particle identification (PID) and the advantages of scintillation light, high light-yield, low-threshold detection with low-energy PID. Hybrid Cherenkov-scintillation detectors could have an exceptionally broad dynamic range in a single experiment, allowing them to have both high-energy, accelerator-based sensitivity while also achieving a broad low-energy neutrino physics and astrophysics program. Recently the Borexino

NP LRP - Fundamental Symmetries, Neutrons, and Neutrinos Whitepaper

Among possible beyond-ton-scale experiments are: NEXT, which will employ high pressure xenon gas time projection chambers with barium tagging; **THEIA**, a large-scale hybrid Cherenkov/scintillation detector that will be an outgrowth of the SNO+ and KamLAND-Zen experiments; Selena, which will employ high-



Cherenkov vs Scintillation



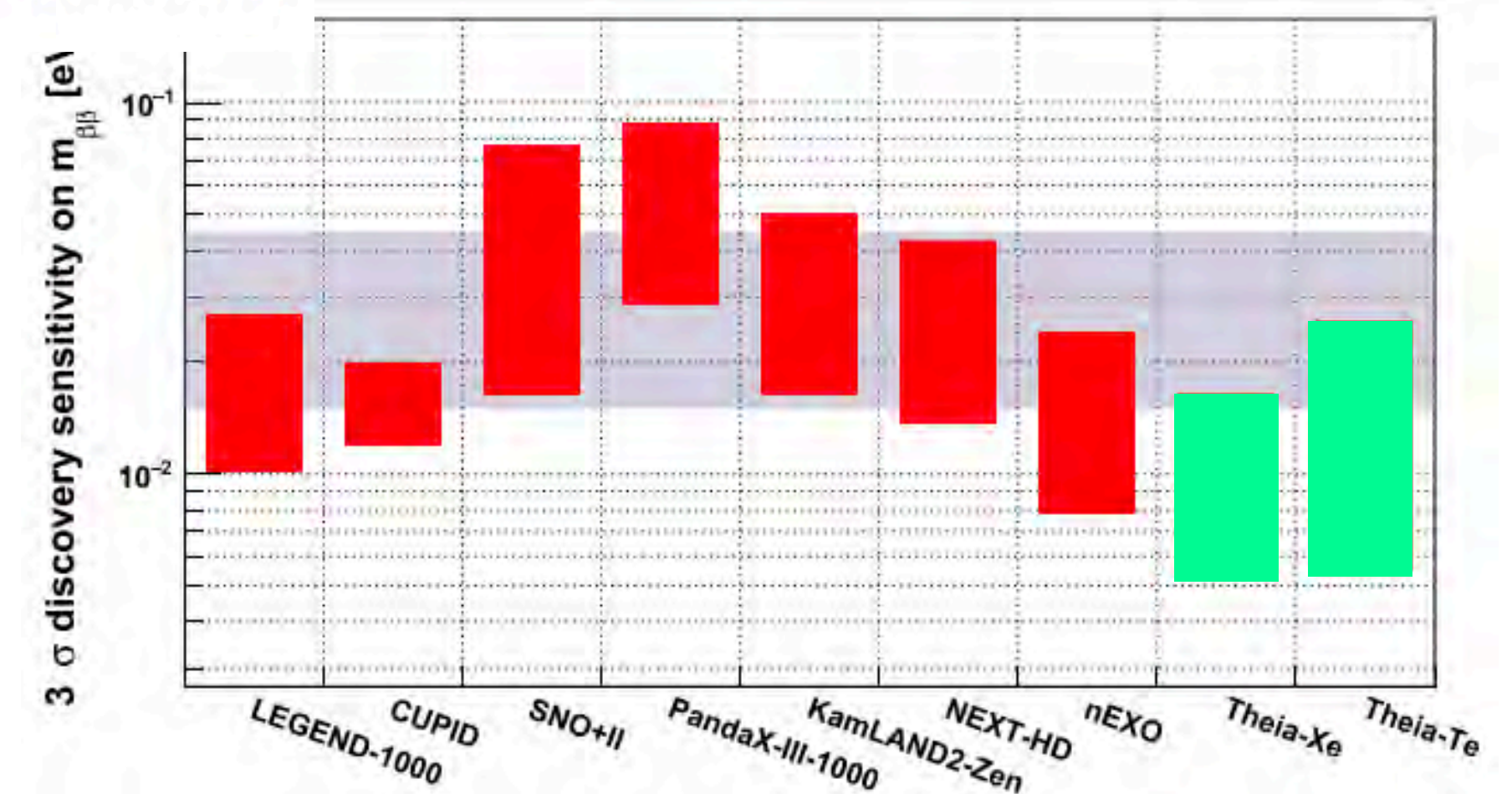
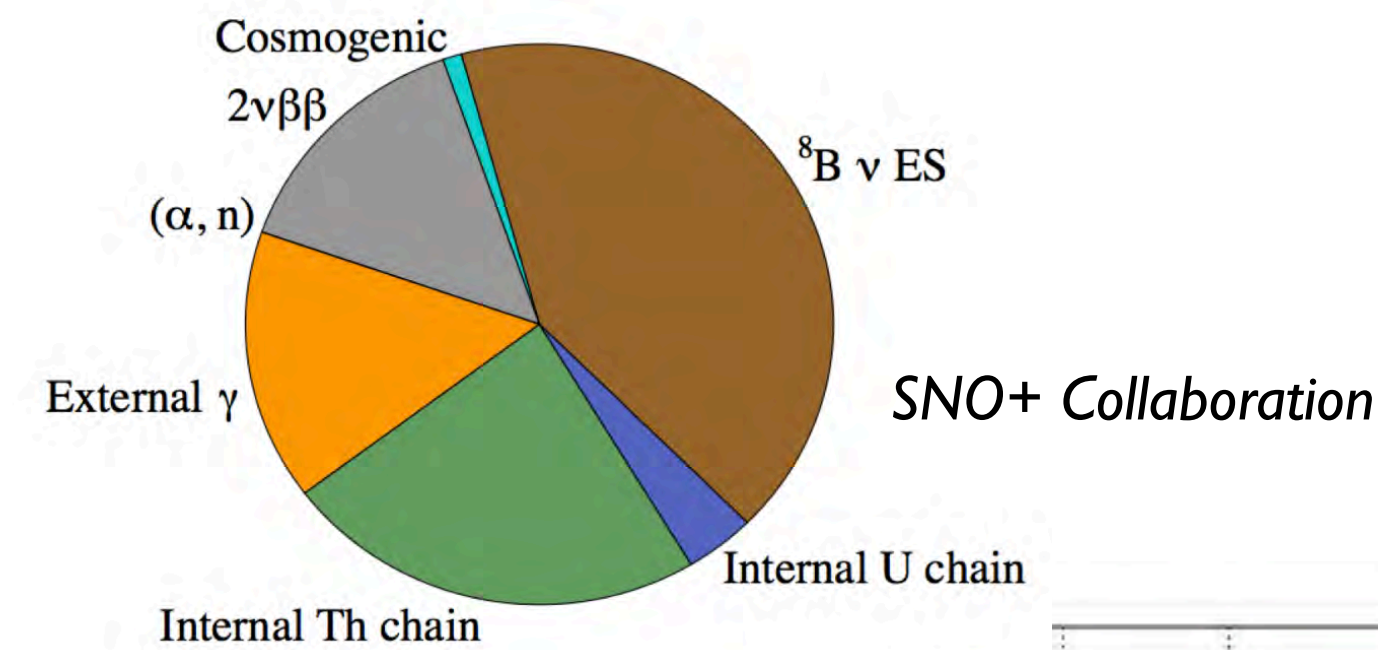
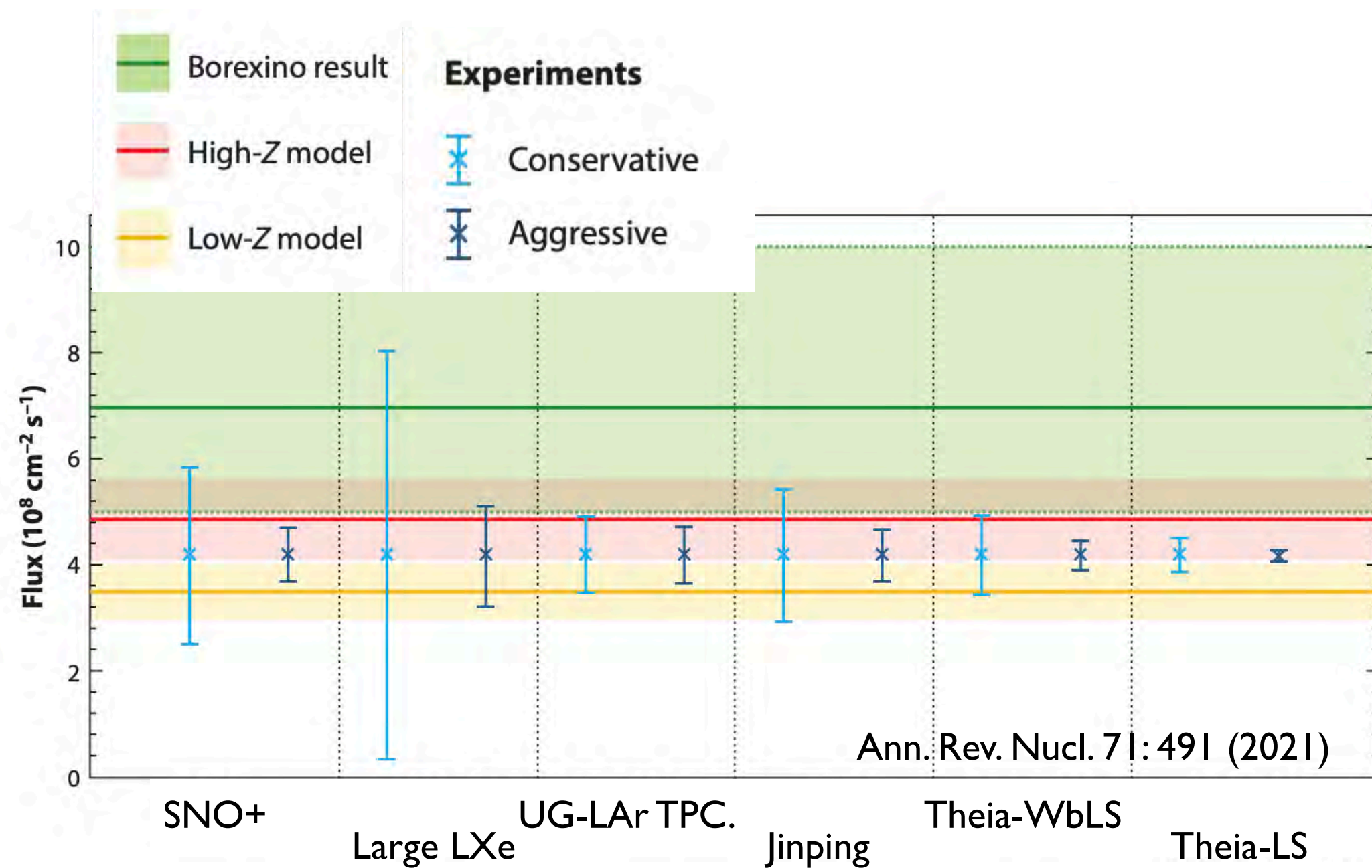
- Directional Cherenkov

- Intense scintillation

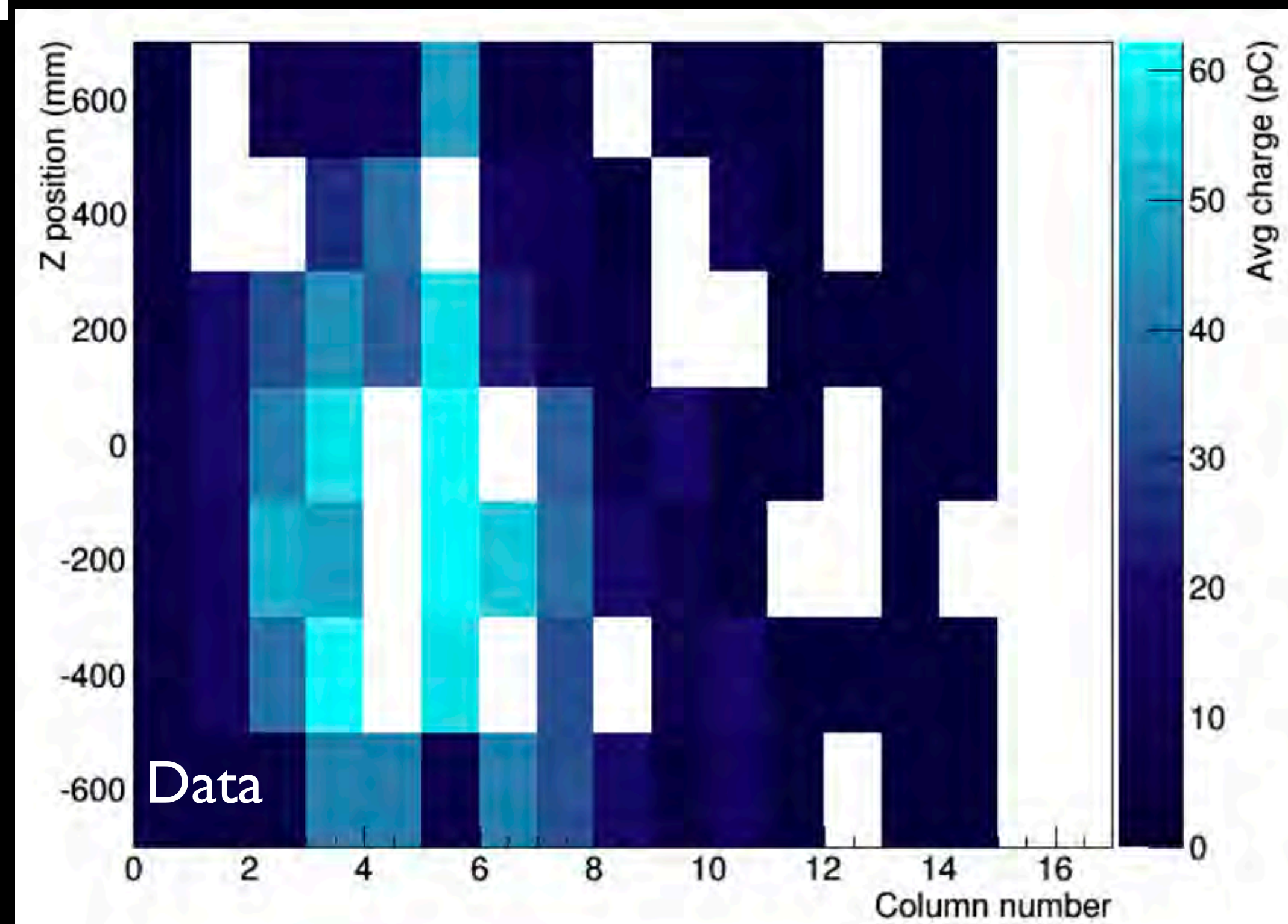
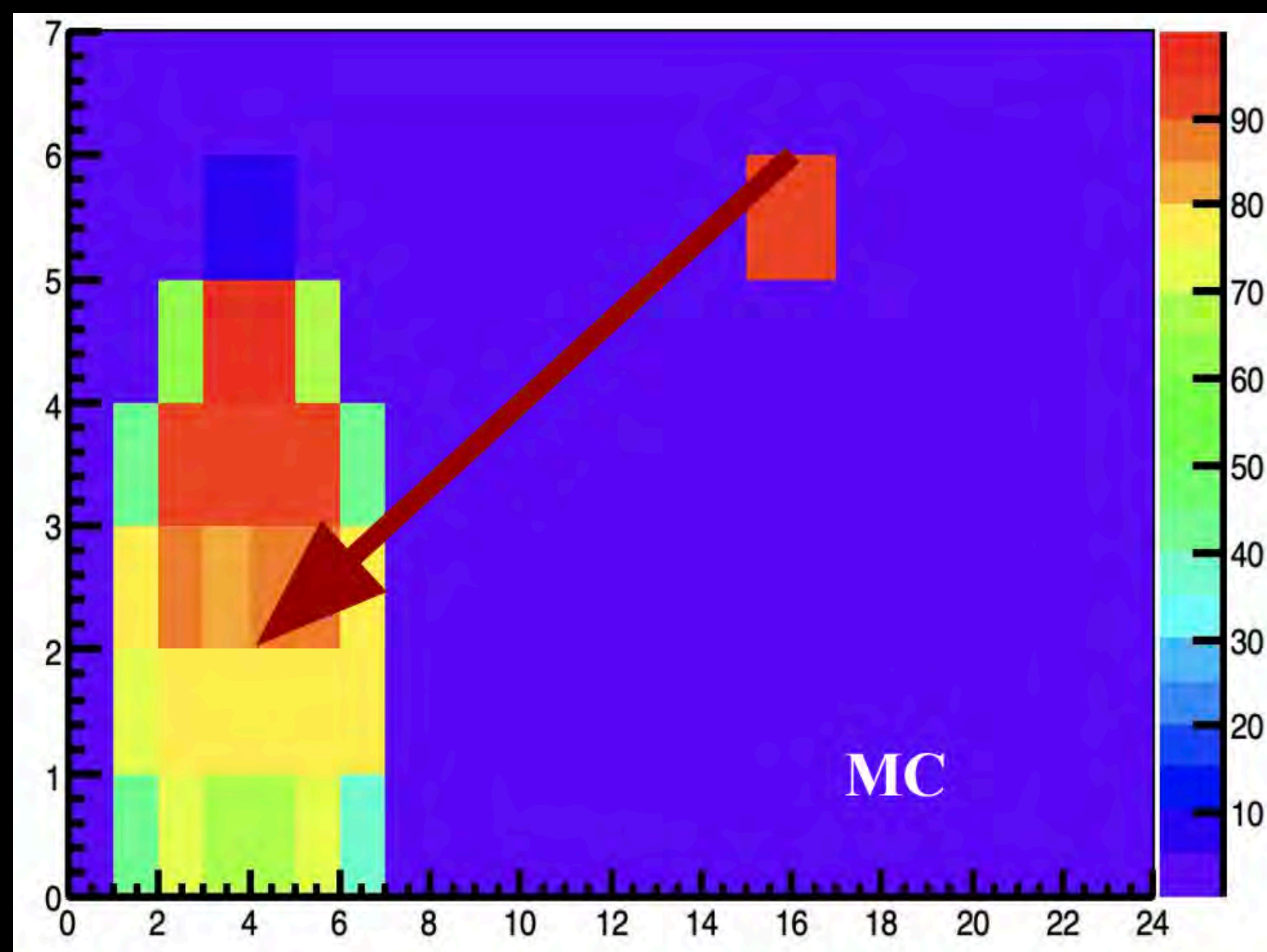
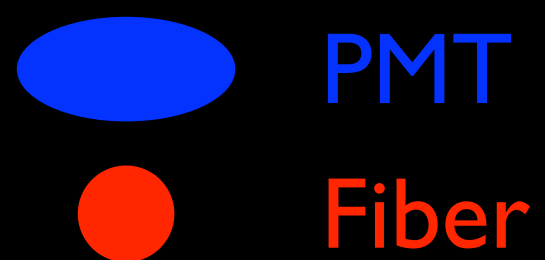
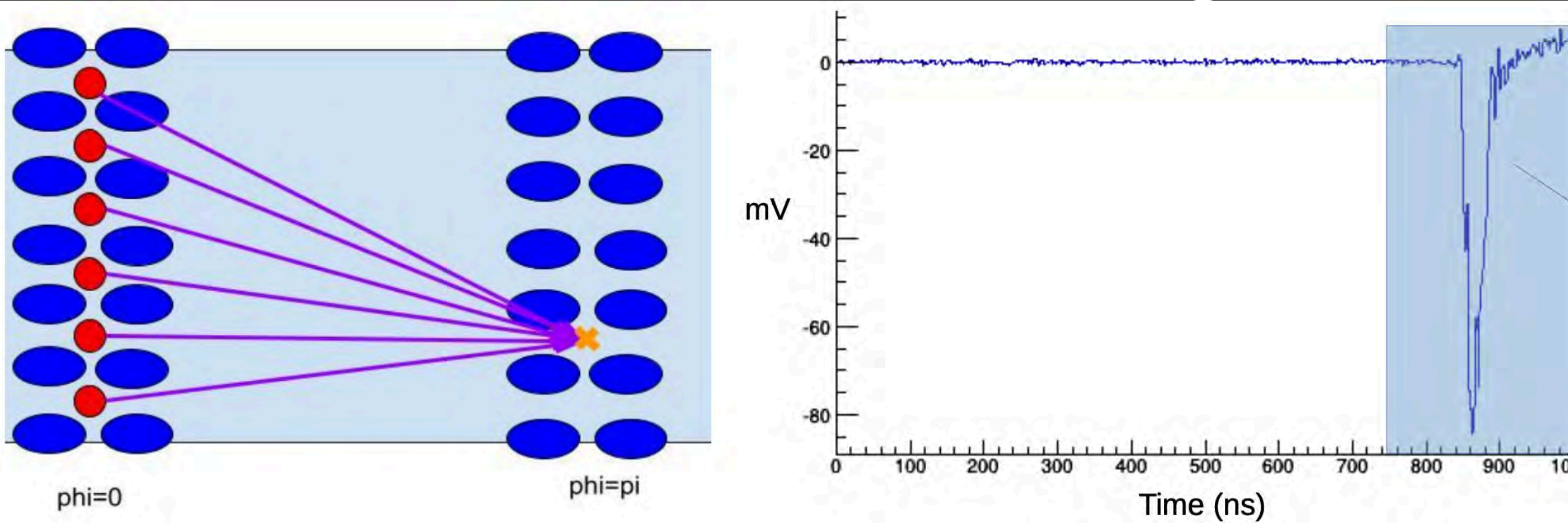
Physics Highlights

- Solar neutrinos
- Unique low-threshold, directional detection
- Particle and event ID from LS time profile, quenching, Ch/S ratio
- Few-% level sensitivity to CNO ν
- Plus precision pp, 8B shape

- Neutrinoless double beta decay
- Inner containment vessel with high-LY LS and isotope
- Background reduction via event imaging: PID, multi-site, directionality

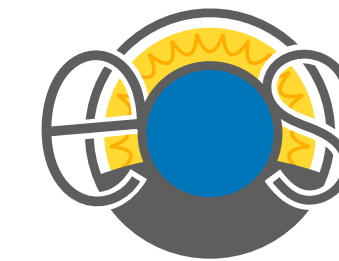


First light (Mar 8)





EOS @ SNS (ORNL)

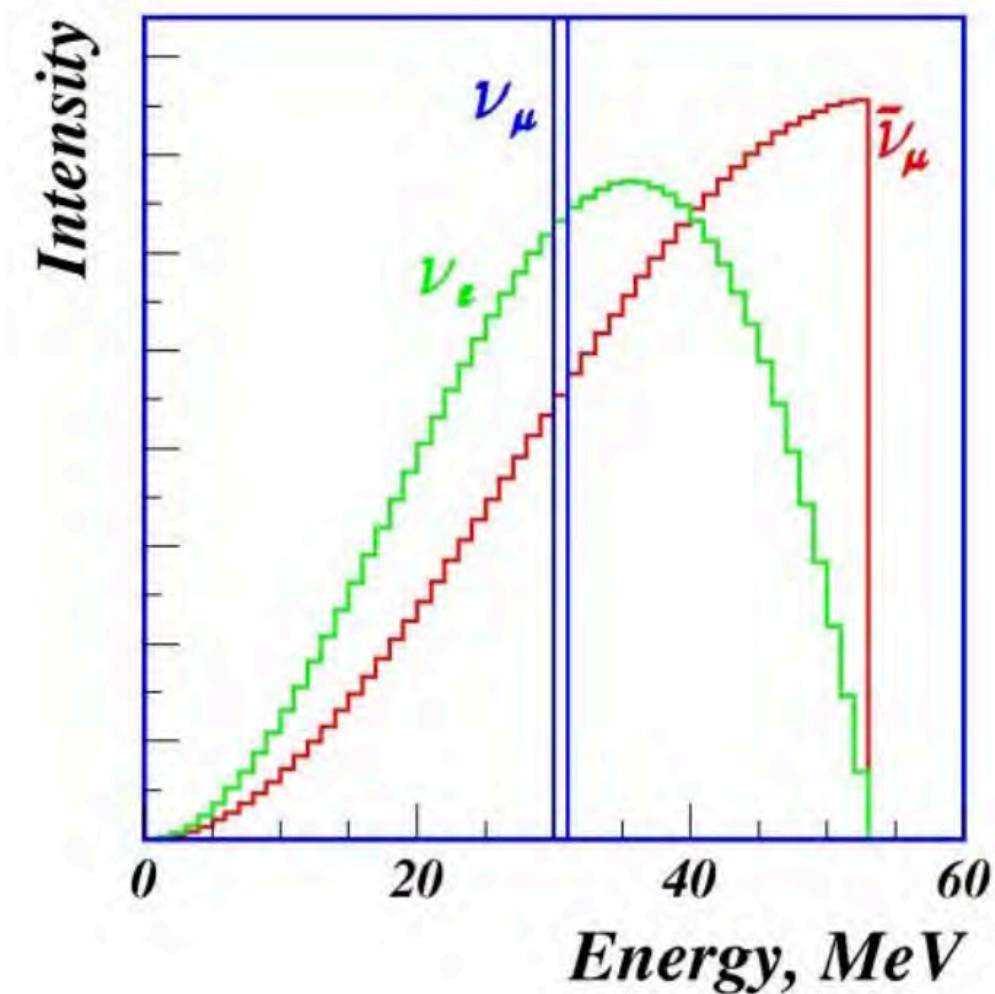


- SNS provides both neutrinos and anti neutrinos
- Detection of Inverse Beta Decay: relevant for reactors
- Detection of Elastic Scattering events: directionality “holy grail”
- Neutron studies: evaluate background rejection
- Possible space identified at ORNL
- Additional technology development opportunity:
 - ▶ Li-loaded WbLS (5% organic, 10% Li)
 - ▶ Enhanced ν_e detection : CC on ${}^7\text{Li}$, spectral precision
- Supernova-relevant demonstration
- Beyond-SM searches

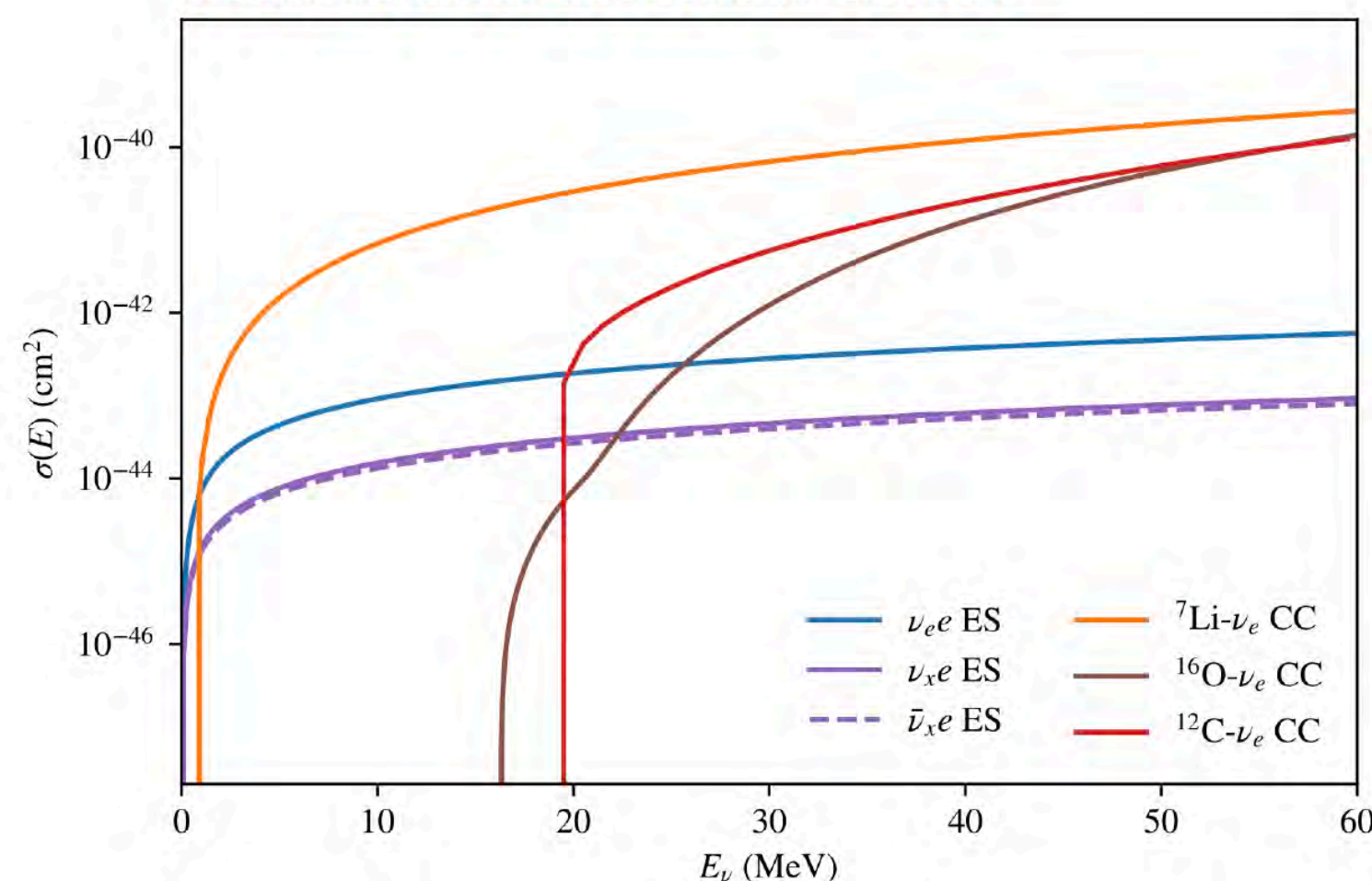
Channel	Rate at 20m Standoff (ev/yr)
$\nu_e e$ ES	136.89
$\nu_\mu e$ ES	20.89
$\bar{\nu}_\mu e$ ES	22.48
ν_e - ${}^7\text{Li}$ CC	533.30
ν_e - ${}^{16}\text{O}$ CC	459.34
ν_e - ${}^{12}\text{C}$ CC	37.08

event rates expected for 4 tons of LiWbLS

ν flux at SNS



neutrino cross-sections in LiWbLS



possible site for EOS deployment

Physics program

Primary physics goal	Reach	Context
Long-baseline oscillations	$>5\sigma$ for 30% of δ_{CP}	Comparable / complementary to DUNE
Nucleon decay $p \rightarrow \nu K^+$	$T > 3.8 \times 10^{34}$ year	Complementary to DUNE (sensitivity to different modes)
Supernova burst	$< 1(2)^\circ$ pointing 20K(5K) events	Complementary to DUNE (nu vs anti-nu)
Diffuse Supernova Neutrino	5σ	Unique background rejection (deep, Cher+scint ratio)
CNO neutrinos	$< 5(10)\%$	Unique capability, order of magnitude improvement
MSW transition	5σ	Unique capability: depth, low-bkg, low-threshold, PID
Geoneutrinos	$< 7\%$	Only plausible high stats measurement in North America
$0\nu\nu\beta$	$T_{1/2} > 1.1 \times 10^{28}$ year (90%C.L.)	Beyond ton-scale sensitivity

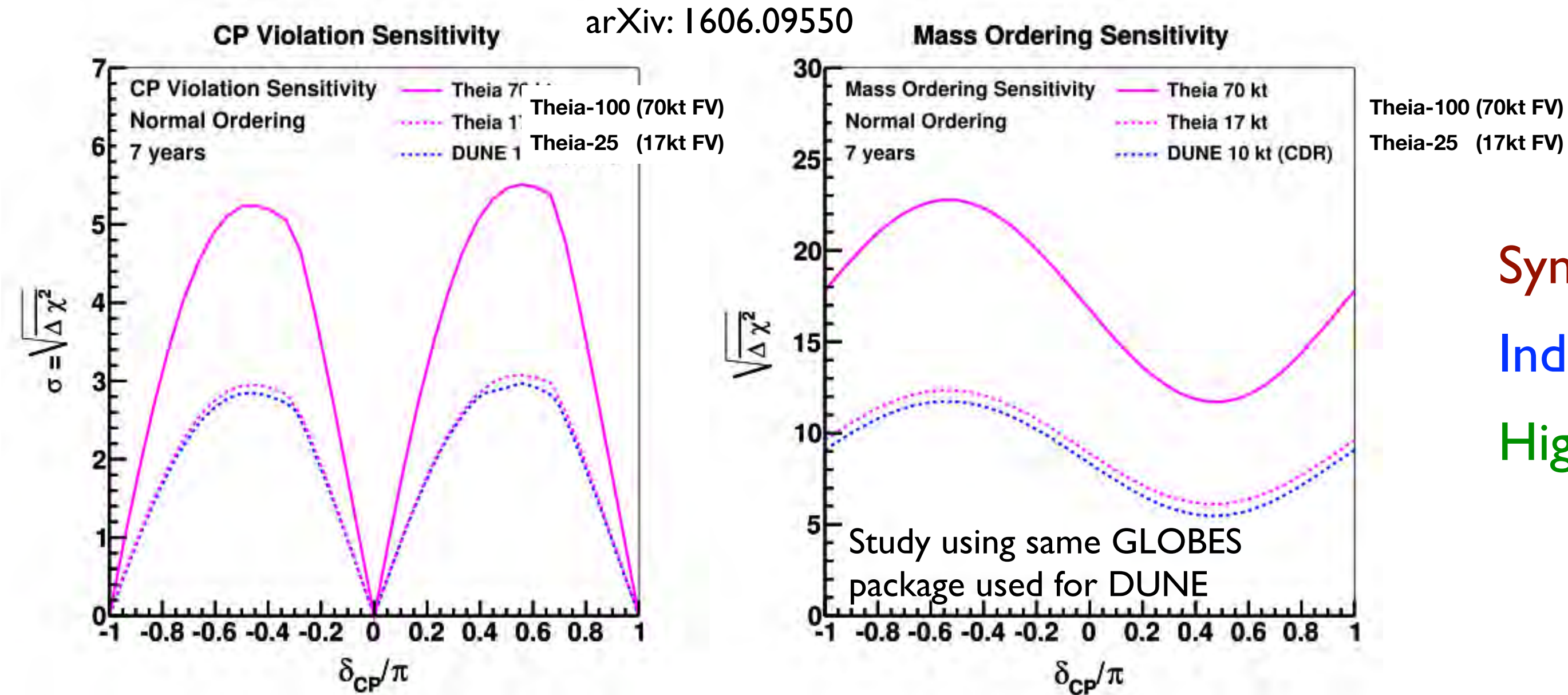
- **Low-energy (< 5 MeV) Solar vs**
 - No long-lived cosmogenics
 - Naturally neutron-shielding
 - High light-yield for good resolution/low threshold
- **SN Burst and Diffuse SN anti- ν Background**
 - Literally complementary to LAr: anti- ν vs. ν
 - Atmospheric NC rejected via Cher+scint
- **$0\nu\nu\beta$ with natural isotopic (e.g. ^{nat}Te) loading**
 - Beyond tonne-scale at low cost
- **Long-baseline oscillations**
 - Similar CP sensitivity to 10 kt LAr
 - Lighter target, different systematics
 - Very narrow energy resolution

Antinu signal \gg nu signal



Require UAr + n shield to achieve this physics in LArTPC DUNE due to intrinsic ³⁹Ar and ⁴²Ar & neutron background

Long-Baseline Sensitivity



Synergy with LAr TPC
 Independent systematics
 High-energy events

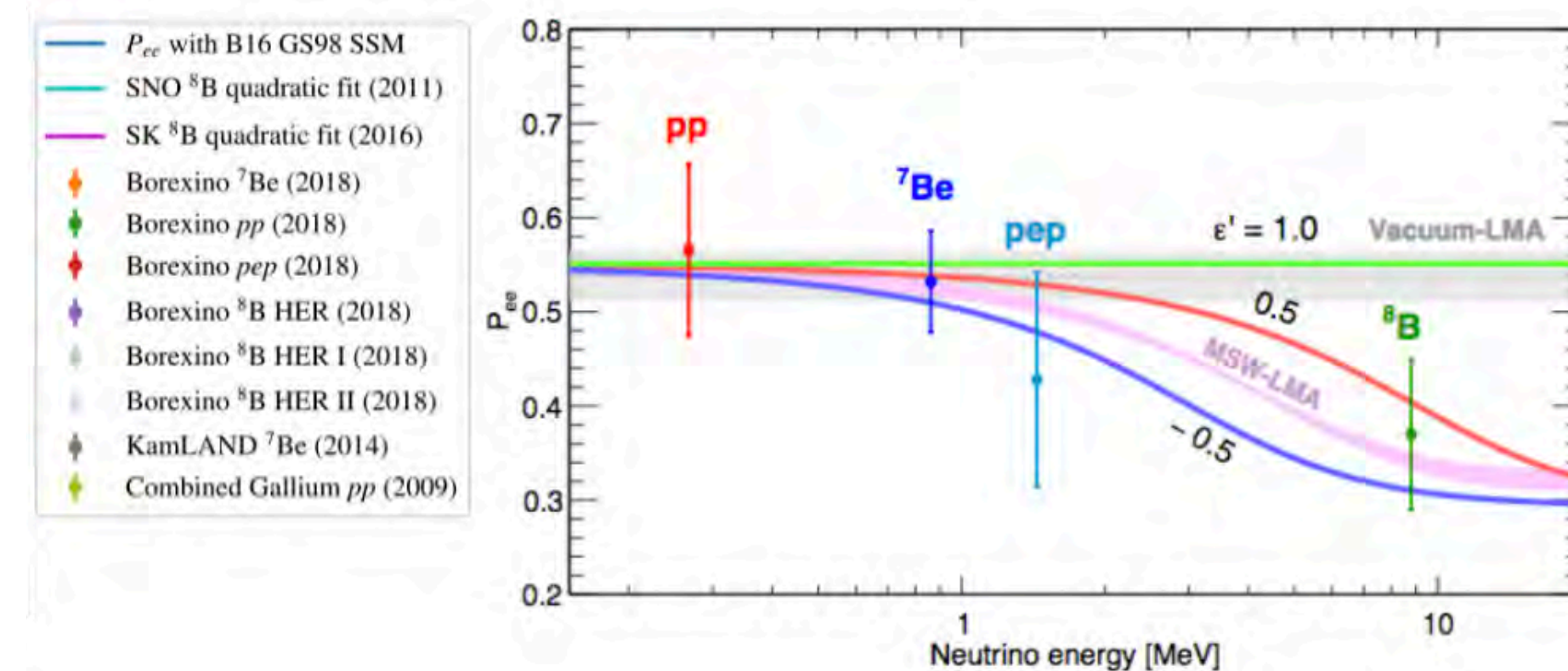
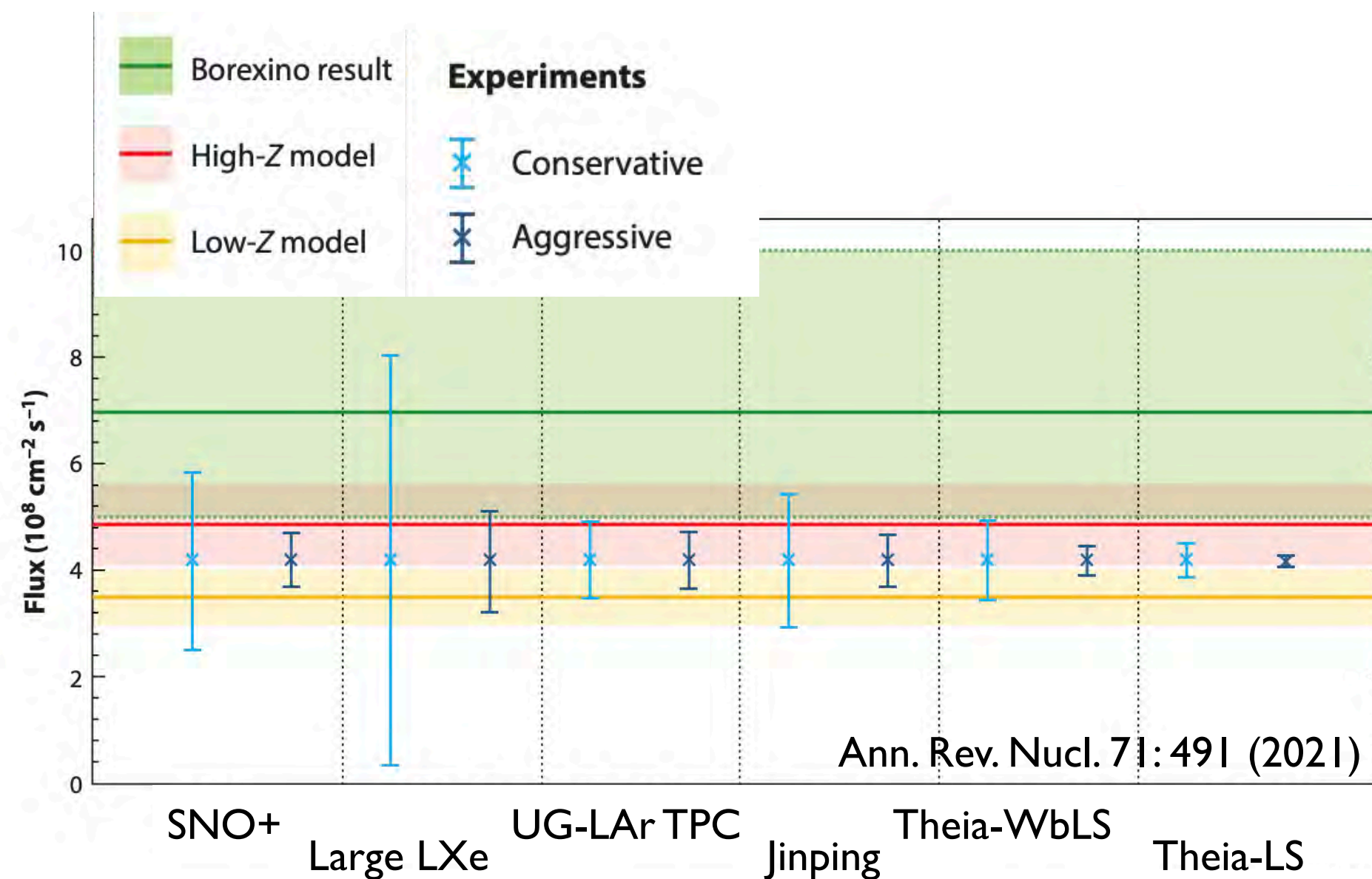
Performance of small (25kt)
 Theia module competitive
 with 10kt LAr TPC

- Ring-imaging of a water Cherenkov detector
- Particle ID from Cher/scint separation
- n and low-E hadron detection (low threshold)
 - ▶ reduce wrong-sign component (ν vs anti- ν)
 - ▶ reduce NC background by detecting $\pi^0 \rightarrow \gamma\gamma$
- THEIA100: large size \rightarrow sensitivity to 2nd oscn max

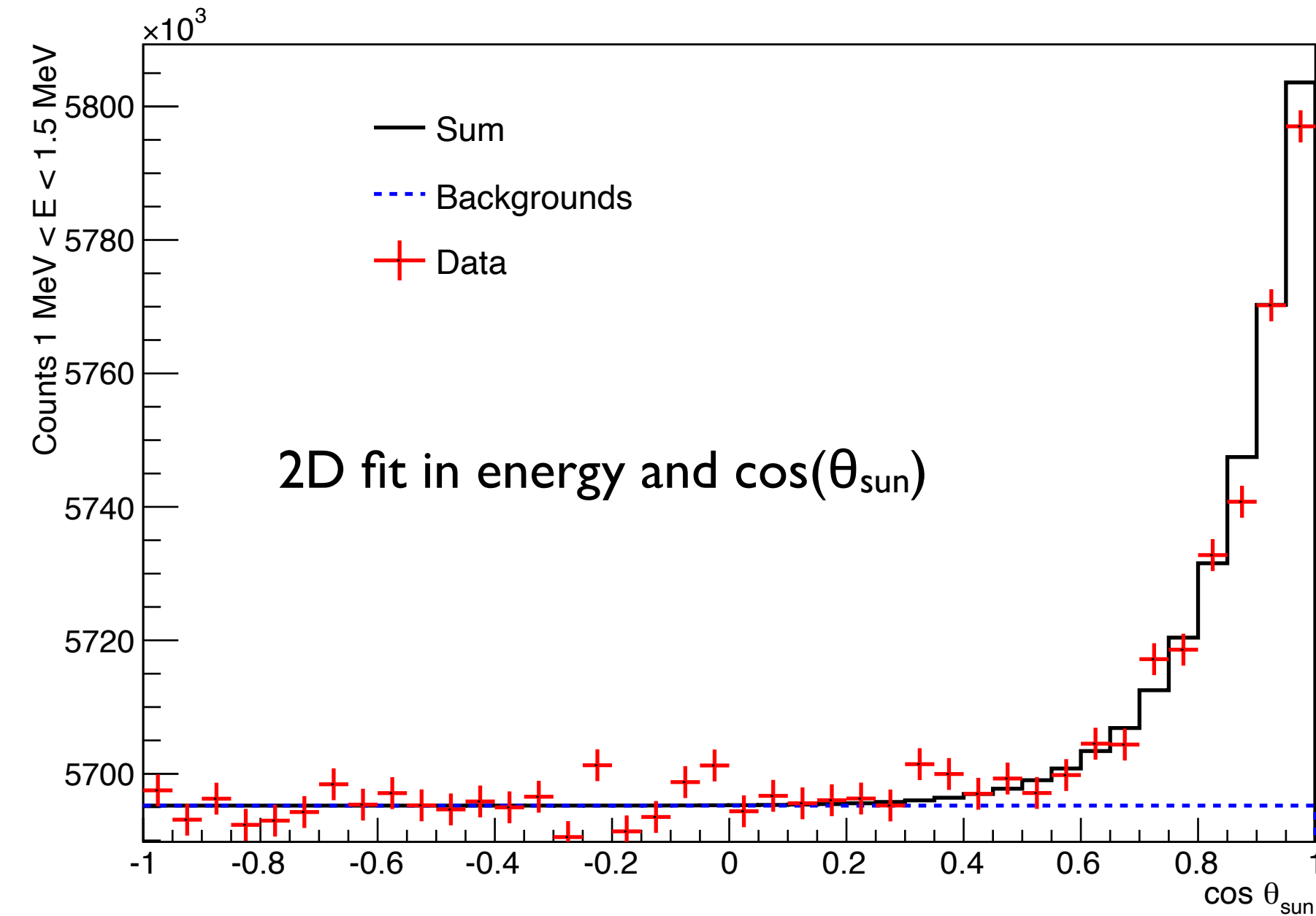
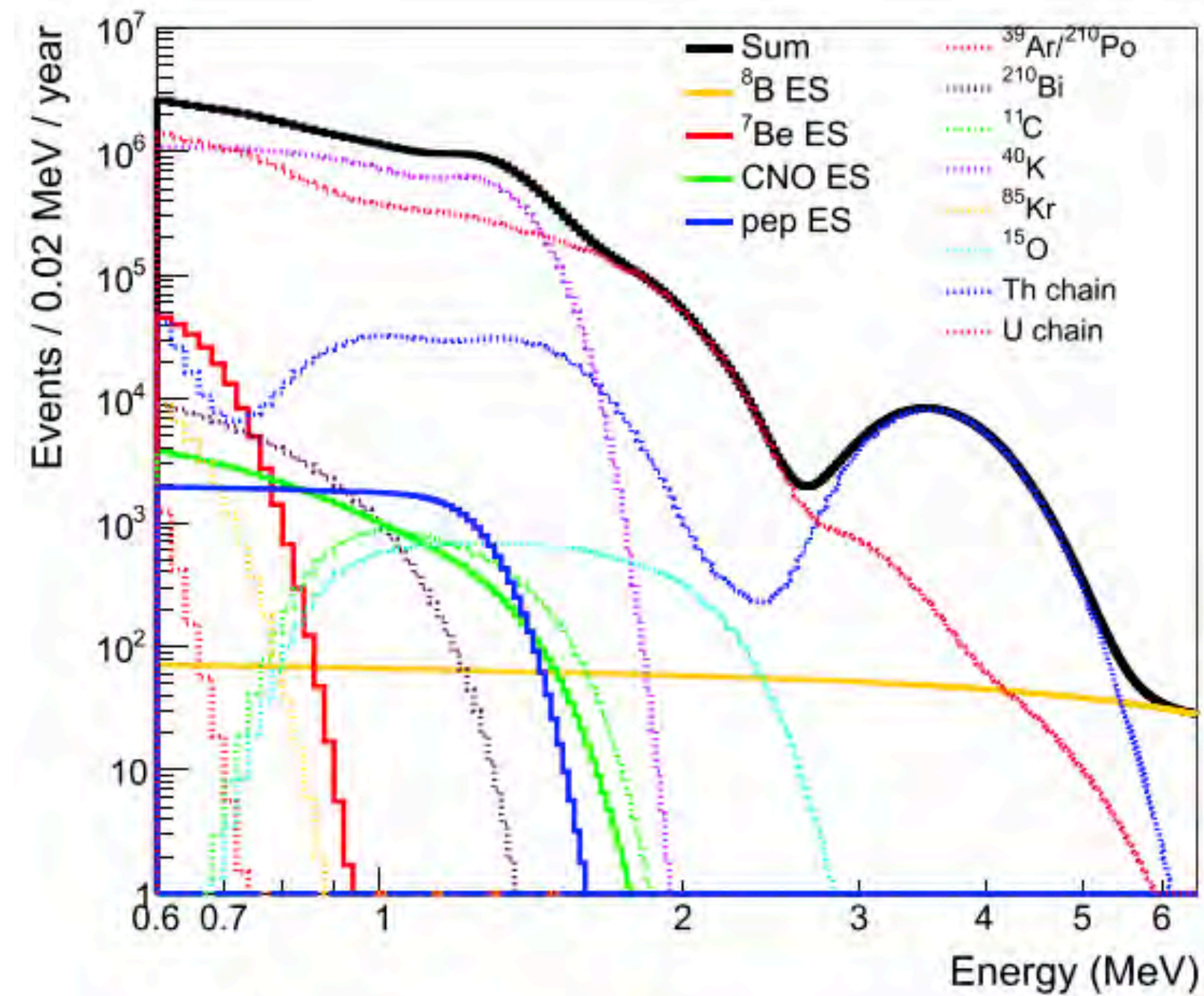
Solar Neutrinos with THEIA

- Dominant background: natural radioactivity e.g. ^{210}Bi
- Theia offers unique low-threshold, directional detection
- Particle and event ID from LS time profile, quenching, Ch/S ratio

- Unique few-% level sensitivity to CNO ν
- Precision pp: luminosity, understand solar energy production
- Unique probe of matter effect / matter-vacuum transition
- Potential Li loading for CC (Haxton)

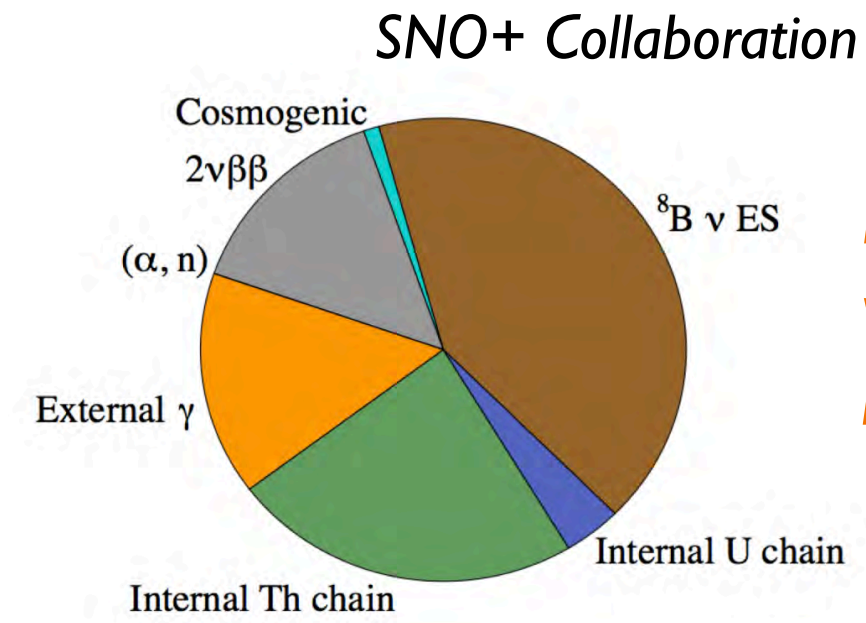
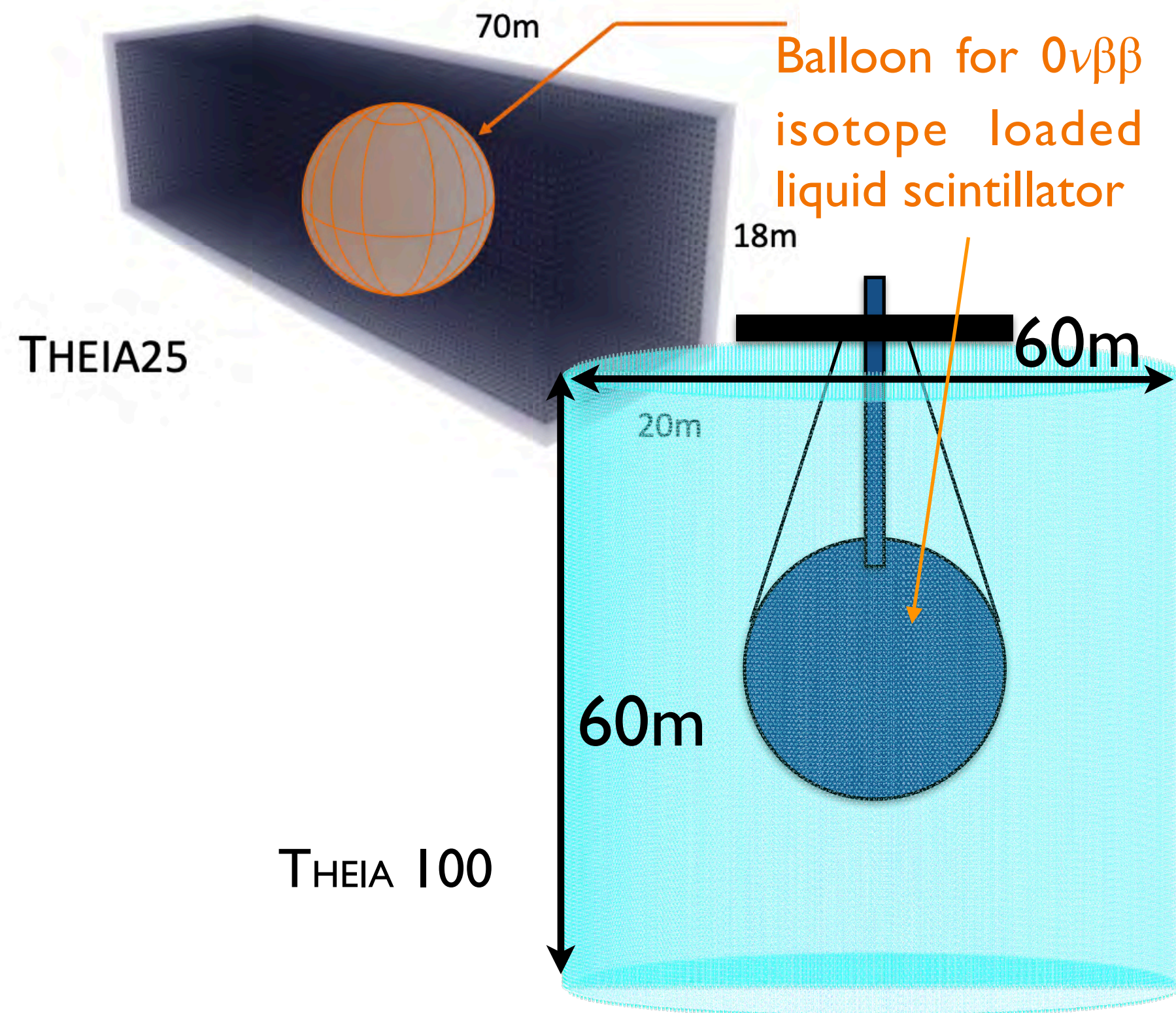


The challenge of backgrounds

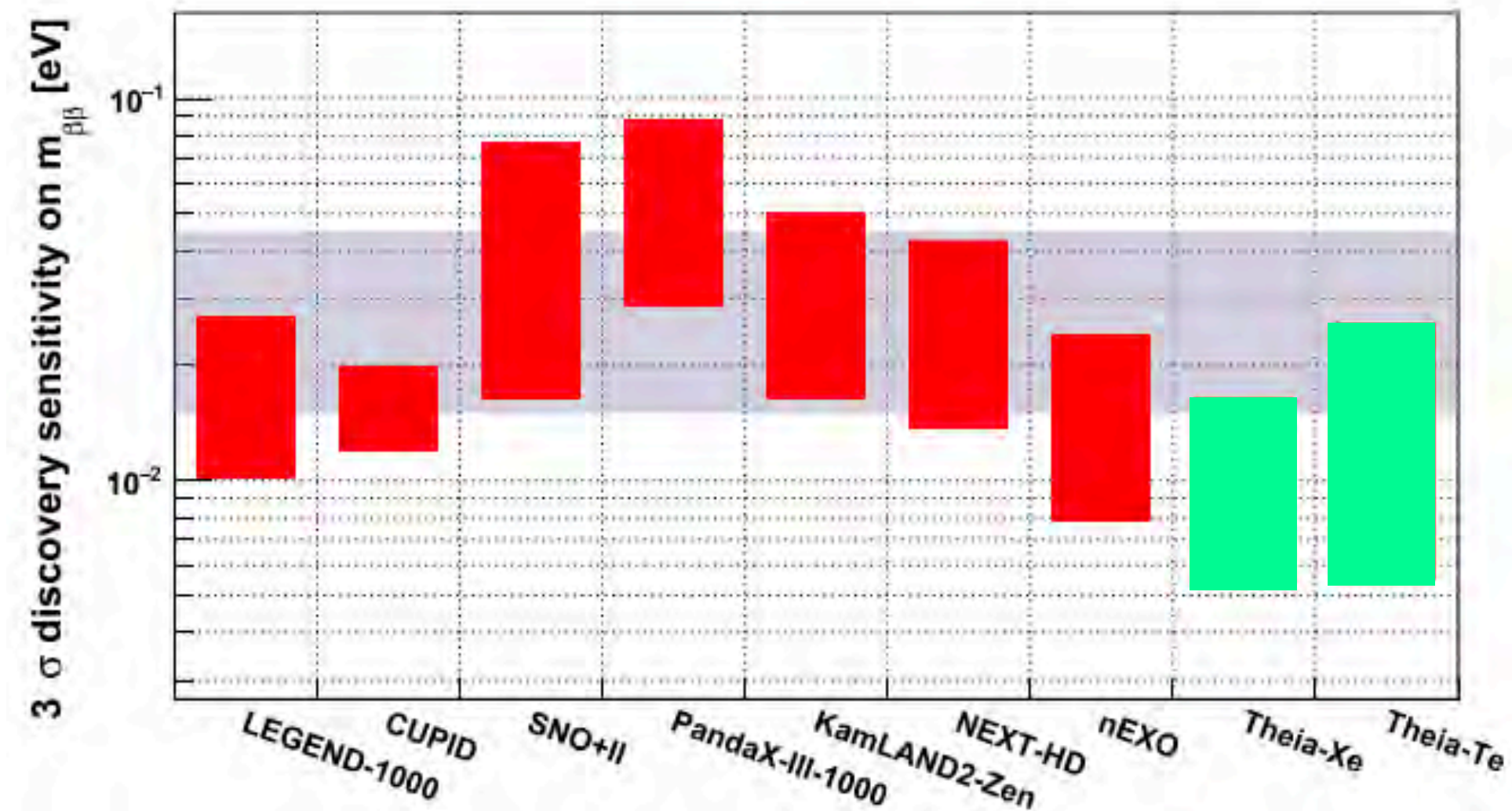


$0\nu\beta\beta$ with THEIA

25-100 kton hybrid optical neutrino detector
 8-m radius balloon with high-LY LS and isotope
 7-m fiducial, 3% ^{nat}Te (or ^{enr}Xe), 10 years



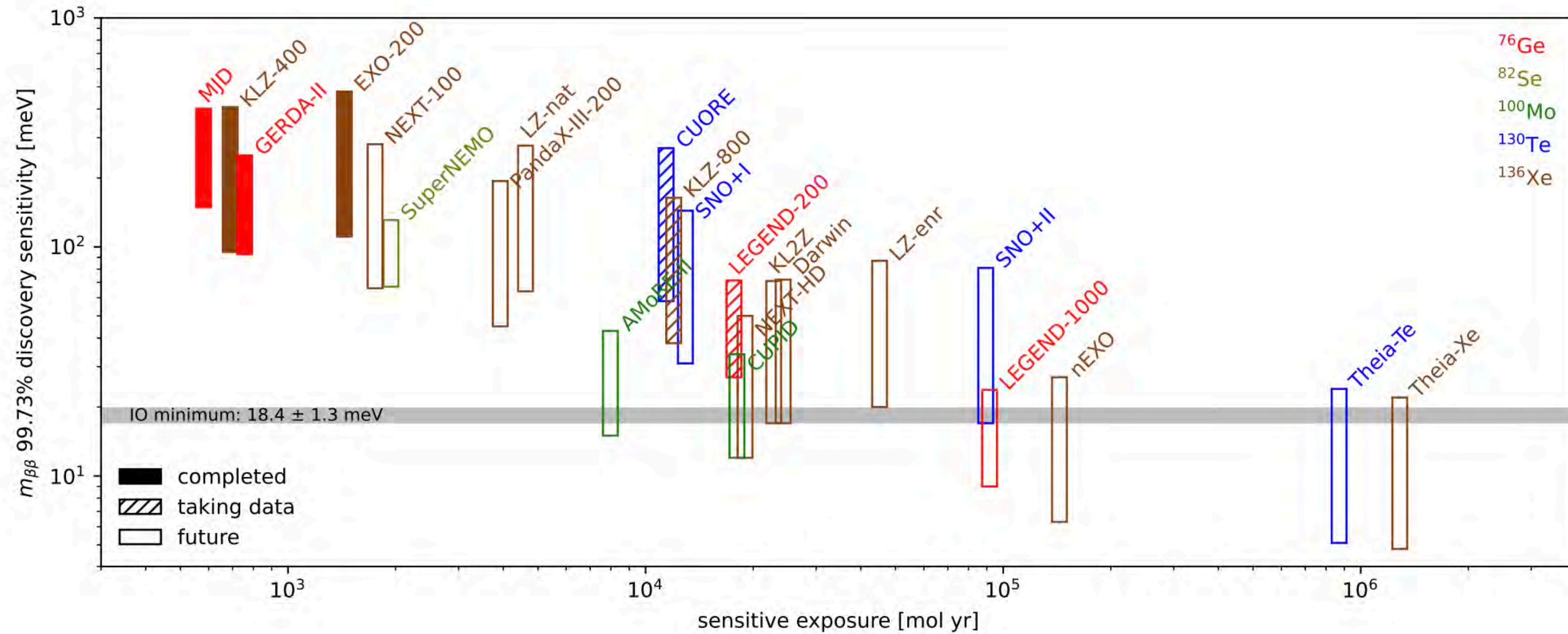
Background reduction via event imaging: PID, multi-site, directionality



Builds on critical developments by KLZ & SNO+ collaborations

Phys.Rev.Lett. 110 : 062502 (2013); Adv.High Energy Phys. 2016 (2016) 6194250; Phys. Rev. D 87 no. 7 : 071301 (2013)

Sensitivity

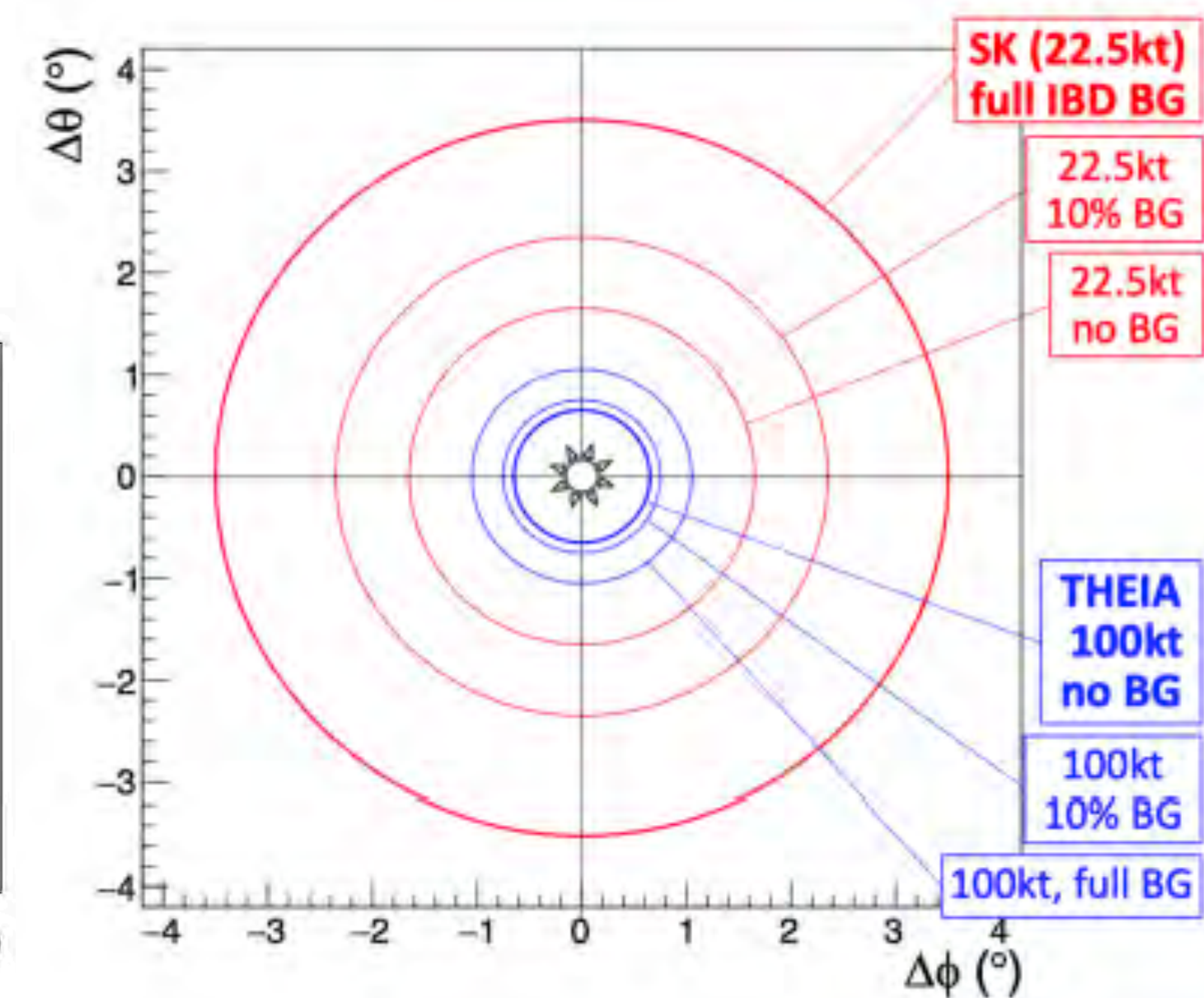
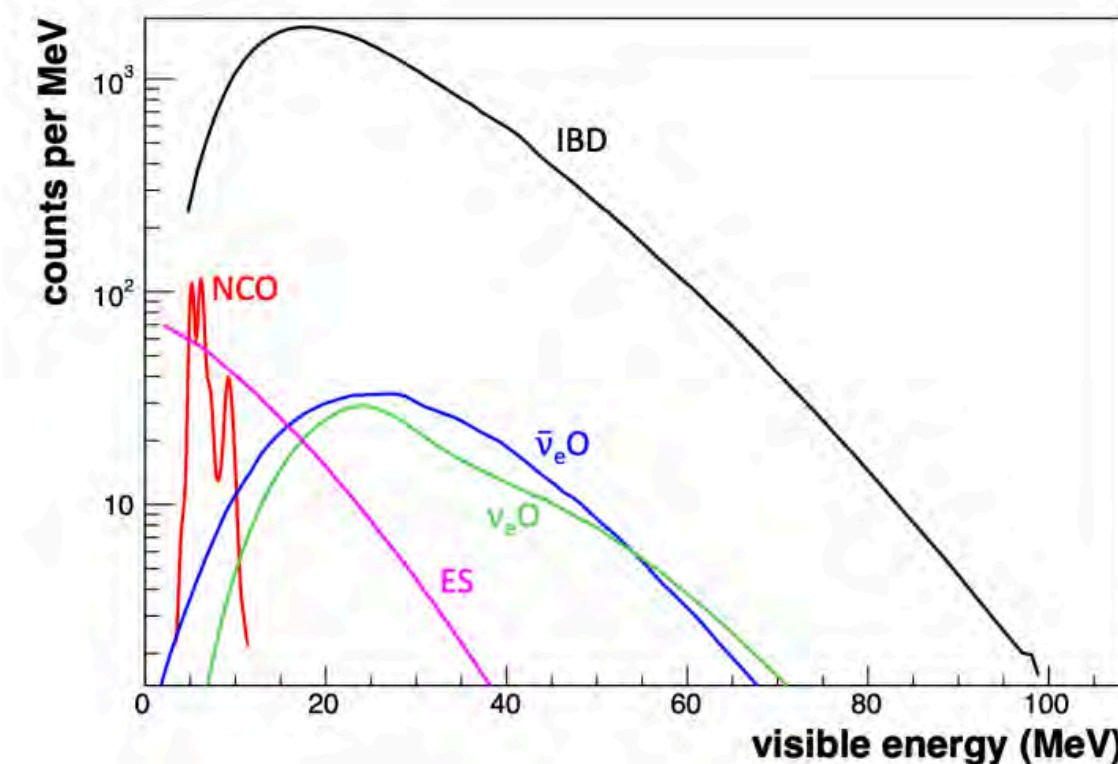


Supernova Detection

- ~90% events are IBD
Highly complementary to ν_e LAr signal
Fast, can act as trigger for DUNE
- ES \Rightarrow pointing accuracy $< 1^\circ$
- CC & monoE γ from NC \Rightarrow
burst T & subsequent mixing
- Flavour-resolved neutrino spectra
- High-stats, low-threshold signal with good resolution
- Pre-supernova ν sensitivity
- Enhanced CC with

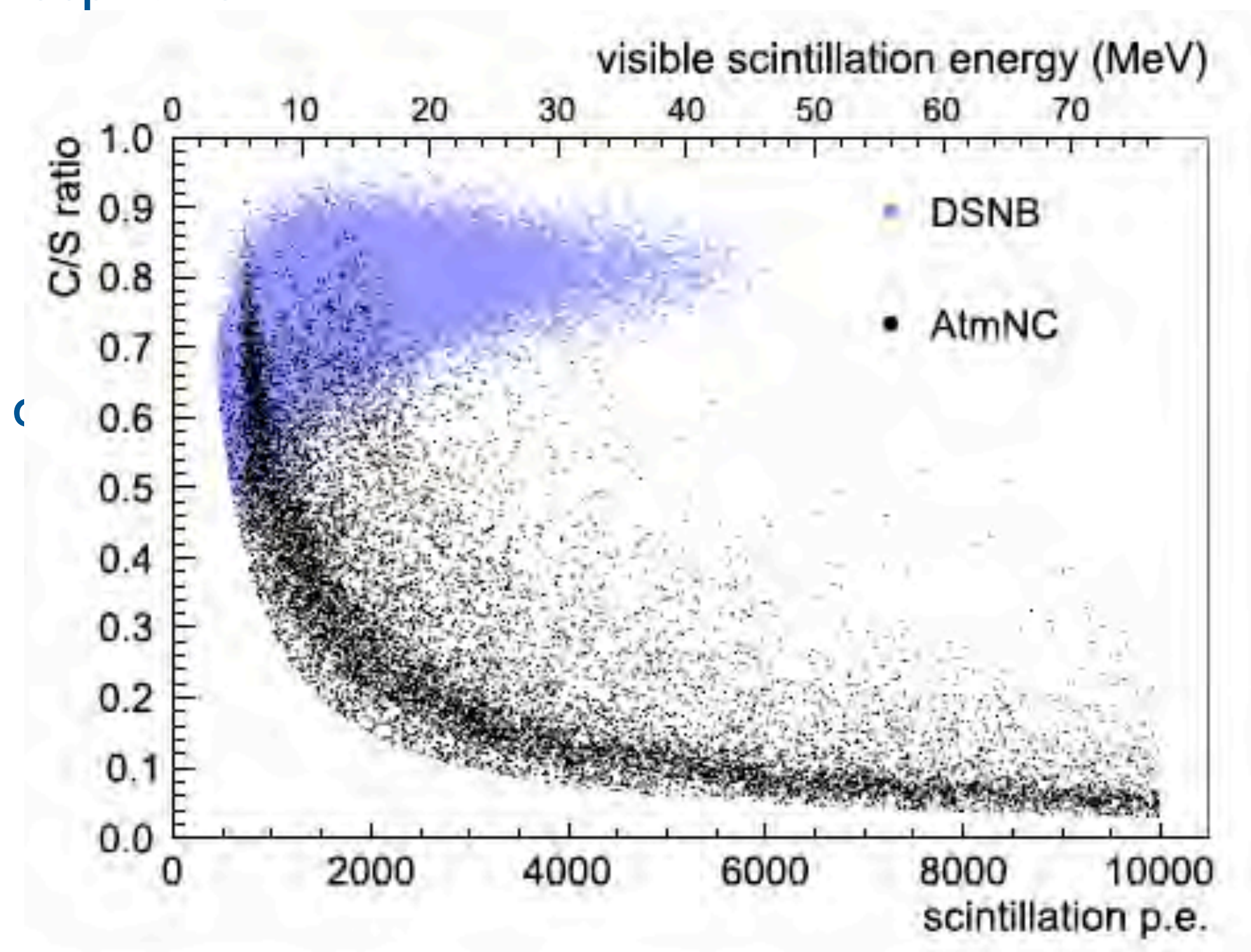
Event rate in 100-kt WbLS, SN at 10 kpc

Reaction	Rate
(IBD) $\bar{\nu}_e + p \rightarrow n + e^+$	19,800
(ES) $\nu + e \rightarrow e + \nu$	960
(ν_e O) $^{16}\text{O}(\nu_e, e^-)^{16}\text{F}$	340
($\bar{\nu}_e$ O) $^{16}\text{O}(\bar{\nu}_e, e^+)^{16}\text{N}$	440
(NCO) $^{16}\text{O}(\nu, \nu)^{16}\text{O}^*$	1100



Diffuse Supernova ν Background

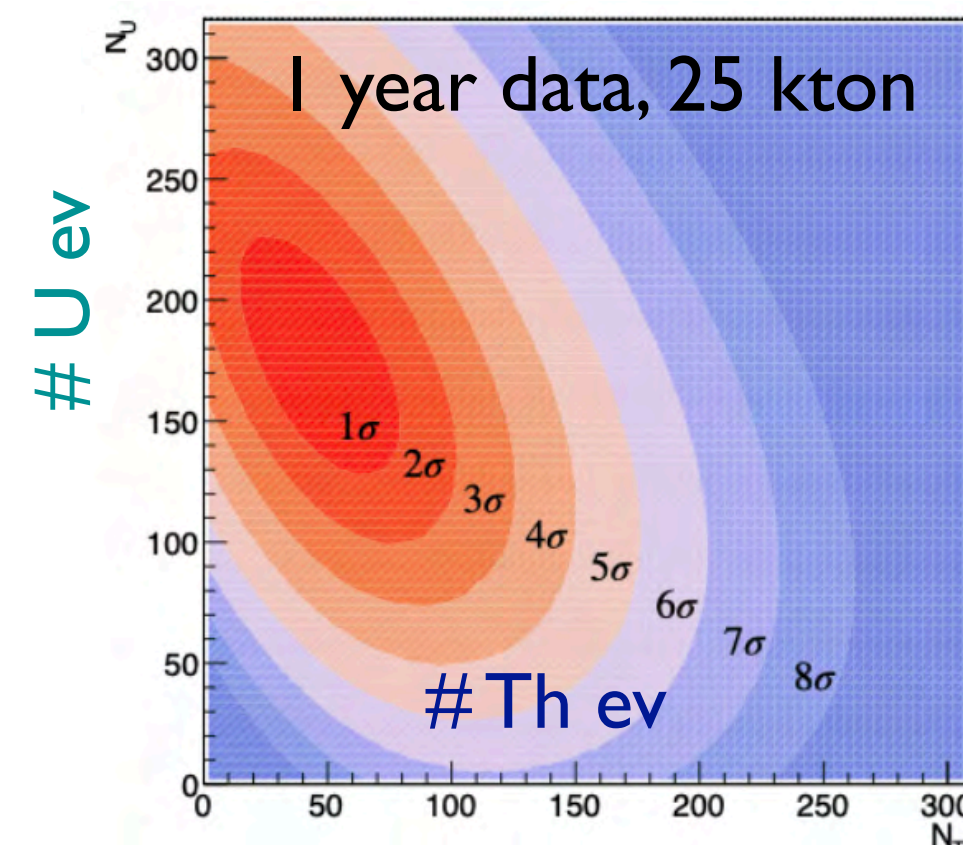
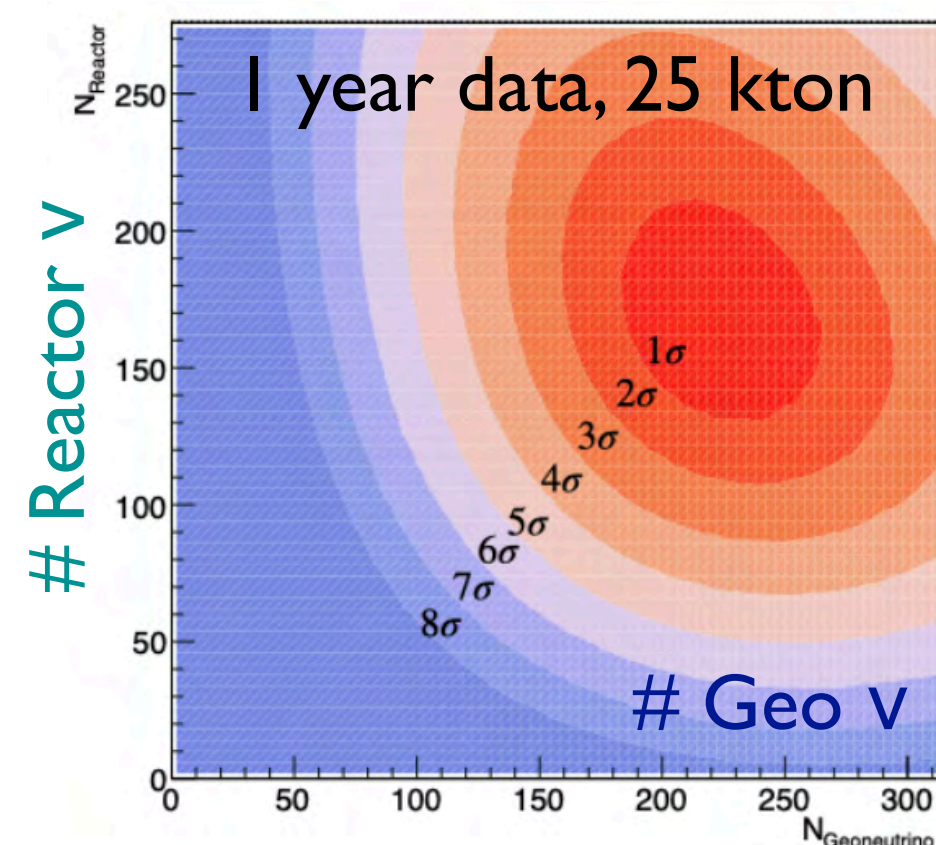
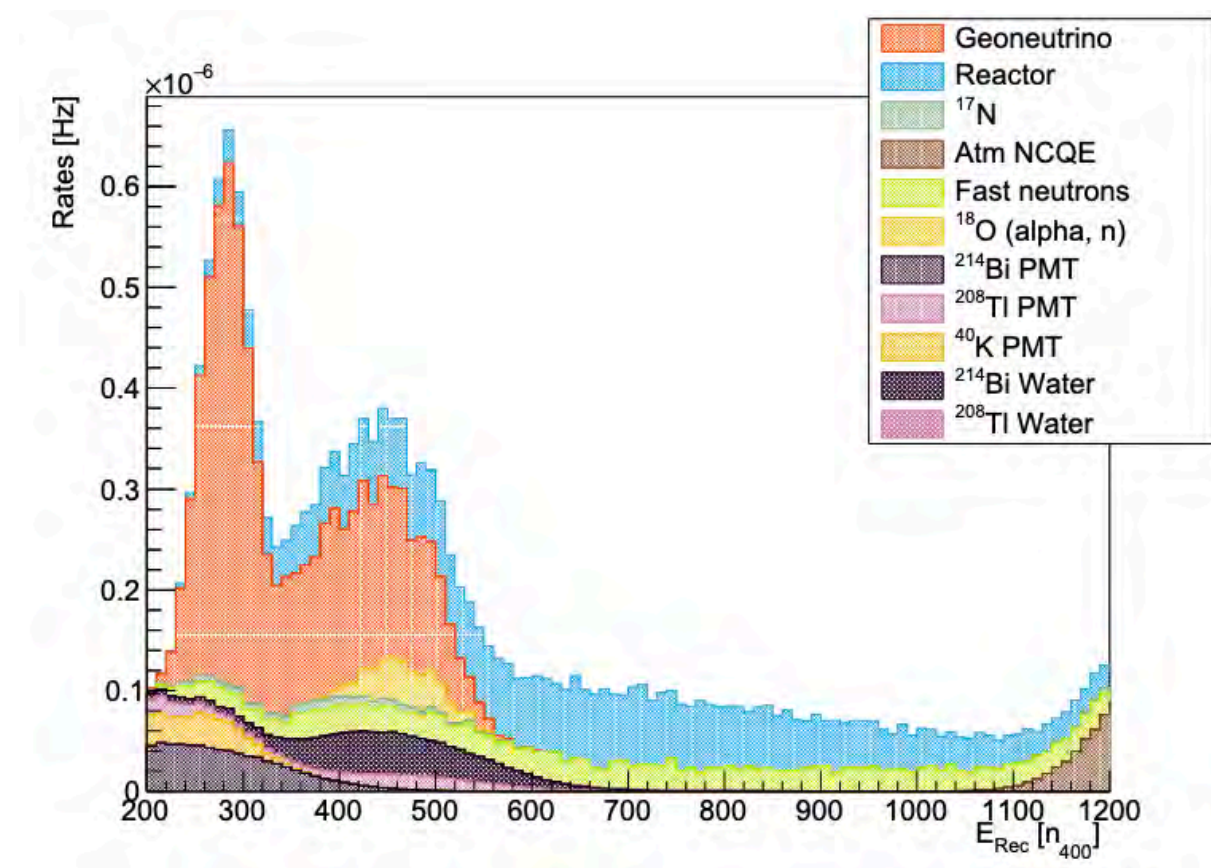
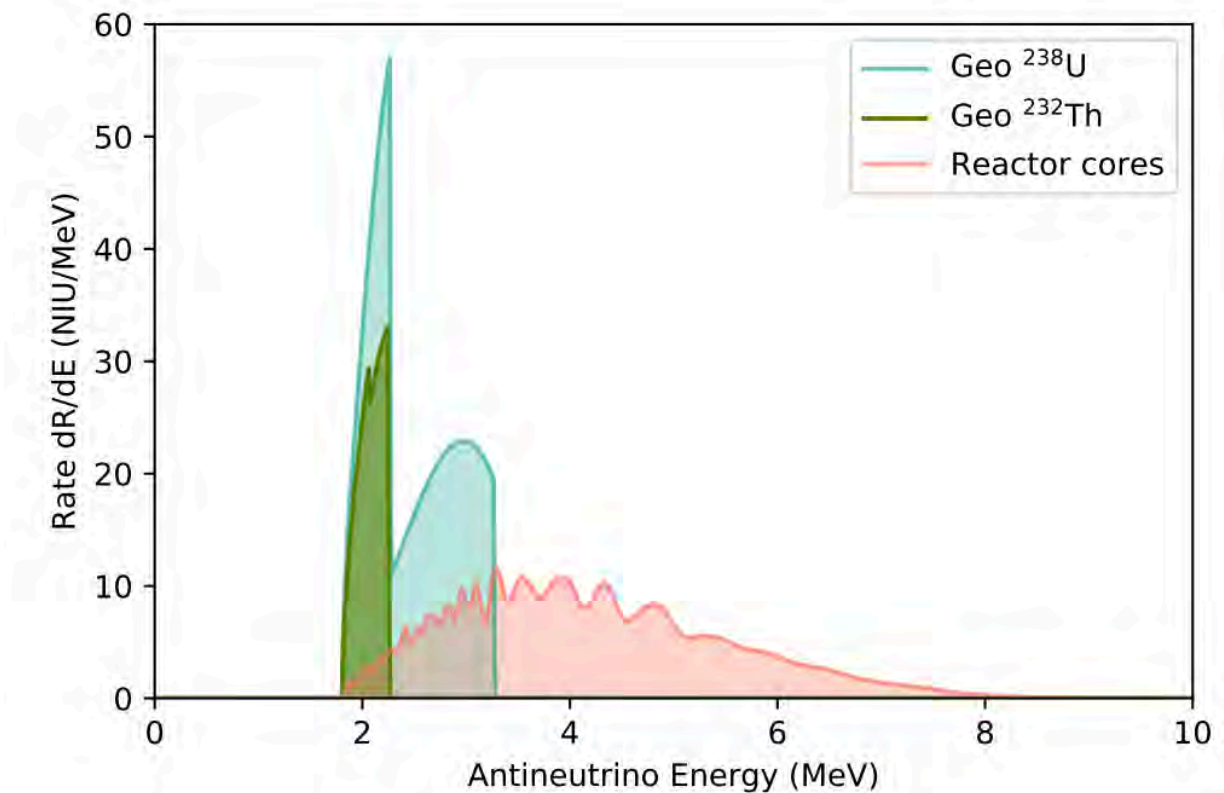
- Diffuse ν “glow” from past core-collapse supernovae
- Astrophysics of SNe
- Signature: IBD detection of antineutrino signal
 - Prompt e^+ and delayed n-capture signal
- Main background: NC interaction of atmospheric ν
 - ν hits C nucleus, causing recoil
 - n captures
 - Can mimic signal
- Cherenkov/scintillation ratio provides a powerful handle for background
- 5σ in 125 kton-yrs



Anti- ν Detection



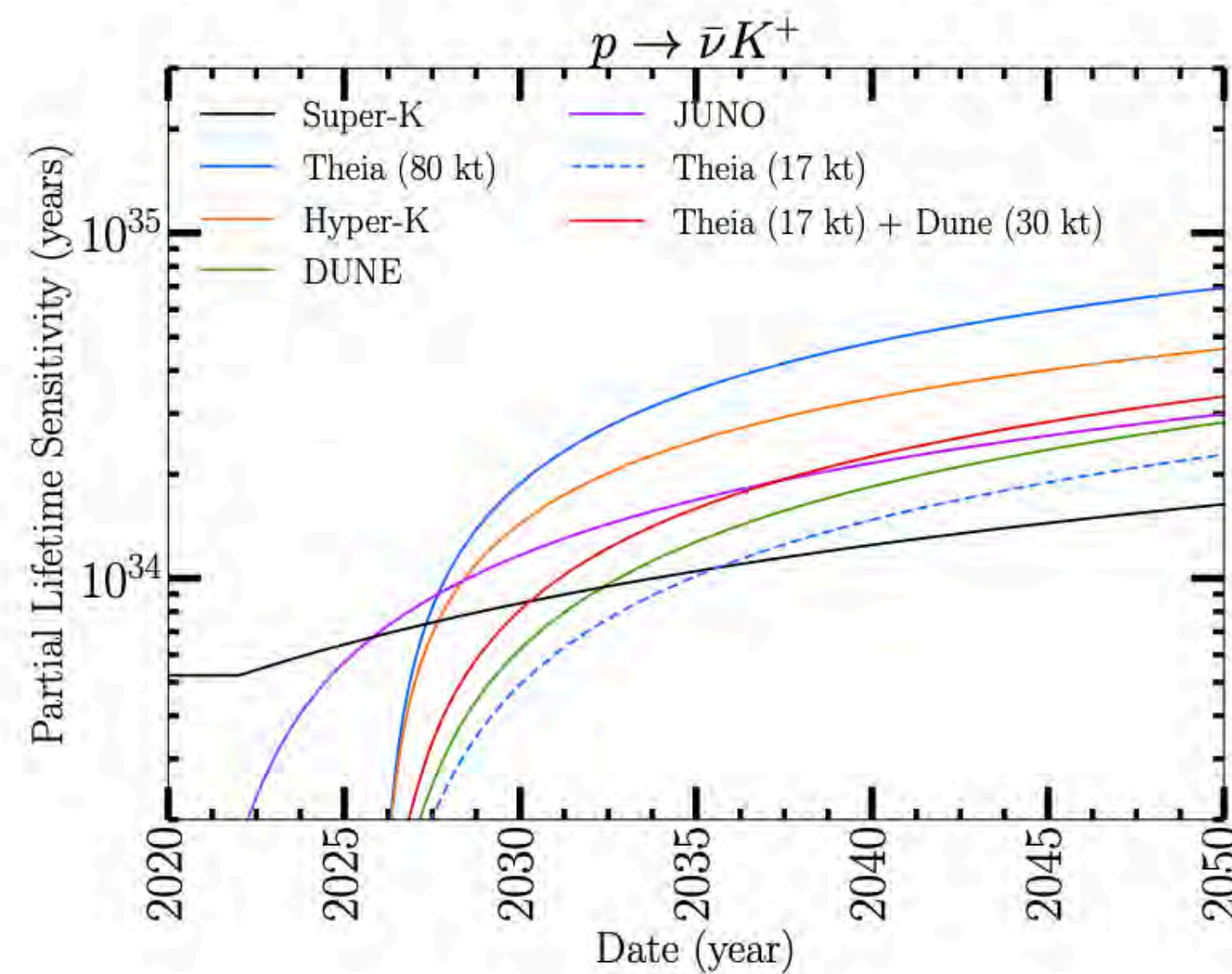
- **Geo- ν** observation by KL, Borexino (< 220 ev)
- **THEIA**: large statistics, complementary site: 218 ev/yr (25 kt)
- Full spectral analysis with BDT for bkg rejection
- Future improvements: PID (p/e^+ , e^-/e^+)
- Could offer first evidence for surface variation
- U/Th ratio to 15% precision in 10 years
- **Reactor ν** prospects: ~ 20 reactor ev/kt-yr
- Demonstrate techniques for remote reactor monitoring
- Range & direction at > 1000 km standoff



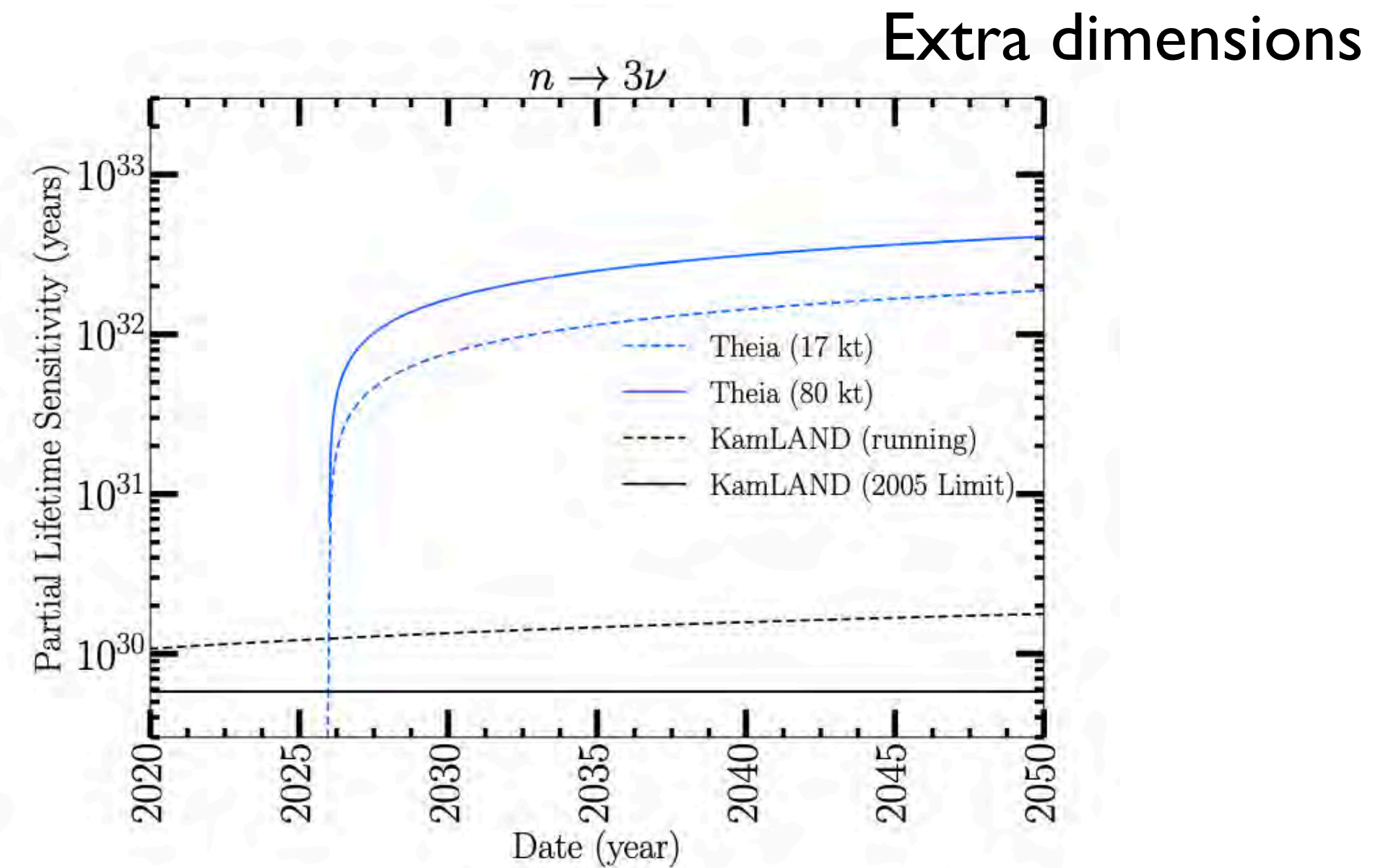
Nucleon Decay

Testing the existence of GUTs with THEIA:

- Large size (statistics), deep location, very clean
- n tagging (low threshold plus potential isotope loading)
- Sub-Cherenkov threshold detection



Sub-Chr t/h detection
 \Rightarrow Directly visible K^+



Deep, low threshold
 Directionality + n tag