### <u>Workshop Italiano sulla Fisica delle Alte Intensità - 2023</u>

# B rare decays: theory overview

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## Introduction

one is the absence of tree-level Flavour Changing Neutral Currents (FCNC)

hence fundamental probe of heavy NP effects

Among the several accidental symmetries of the Standard Model, a particularly interesting

These decays occur at loop-level, and are both GIM- and CKM-suppressed: very rare,

Indeed, since no NP has been (so far) directly observed at colliders, is fundamental to have input from indirect searches where BSM appears through virtual, intermediate states











•  $B \rightarrow \mu \mu$ 

•  $B \to K^{(*)} \nu \nu$ 

•  $B \to K^{(*)}\ell\ell, B_s \to \phi\ell\ell$ 

•  $b \rightarrow s\gamma$ 

### <u>Overview</u>



Helicity suppressed, tree-level decay

Main uncertainties come from CKM elements (UTA) and decay constants (Lattice)

$$\mathcal{B}(B_q^+ \to \tau^+ \nu_\tau)^{\text{SM}} = \tau_{B_q^+} \frac{G_F^2 |V_{qb}|^2 f_{B_q^+}^2 m_{B_q^+} m_\tau^2}{8\pi} \left(1 - \frac{m_\tau^2}{m_{B_q^+}^2}\right)^2, \quad q = u, c$$

 $|V_{cb}|^{\text{UTA}} = 42.22(51) \times 10^{-3}, f_{B_c} = 427(6) \text{ MeV}$  $|V_{\mu b}|^{\text{UTA}} = 3.70(11) \times 10^{-3}, f_{B^+} = 190.0(1.3) \text{ MeV}$ 

According to present Lattice estimates, decay constants errors could be halved in the next decade!

### $B \rightarrow \tau \nu$ : the SM status

$$\Rightarrow \quad \mathcal{B}(B_c^+ \to \tau^+ \nu_\tau)^{\text{SM}} = 2.29(9) \times 10^{\circ}$$

$$\mathcal{B}(B^+ \to \tau^+ \nu_\tau)^{\rm SM} = 0.87(5) \times 10$$





$$\mathcal{B}(B_{q}^{+} \to \tau^{+}\nu_{\tau}) = \mathcal{B}(B_{q}^{+} \to \tau^{+}\nu_{\tau})^{\mathrm{SM}} \times \left| 1 - \left(C_{V_{R}}^{q} - C_{V_{L}}^{q}\right) + \left(C_{S_{R}}^{q} - C_{S_{L}}^{q}\right) \frac{m_{B_{q}}^{2}}{m_{\tau}(m_{b} + m_{q})} \right|^{2}$$

$$g_{(R)} = (\bar{q}_{L(R)}\gamma_{\mu}b_{L(R)})(\bar{\tau}_{L}\gamma_{\mu}\nu_{L})$$

$$g_{L(R)} = (\bar{q}_{R(L)}b_{L(R)})(\bar{\tau}_{R}\nu_{L})$$

$$g_{L(R)} = (\bar{q}_{R(L)}b_{L(R)})(\bar{\tau}_{R$$

$$\mathcal{B}(B_{q}^{+} \to \tau^{+} \nu_{\tau}) = \mathcal{B}(B_{q}^{+} \to \tau^{+} \nu_{\tau})^{\text{SM}} \times \left| 1 - (C_{V_{R}}^{q} - C_{V_{L}}^{q}) + (C_{S_{R}}^{q} - C_{S_{L}}^{q}) \frac{m_{B_{q}}^{2}}{m_{\tau}(m_{b} + m_{q})} \right|^{2}$$

$$O_{V_{L(R)}} = (\bar{q}_{L(R)}\gamma_{\mu}b_{L(R)})(\bar{\tau}_{L}\gamma_{\mu}\nu_{L})$$

$$O_{S_{L(R)}} = (\bar{q}_{R(L)}b_{L(R)})(\bar{\tau}_{R}\nu_{L})$$

$$\int_{1.5}^{0.0} \frac{1.5}{-1.0} - 0.5 - 0.0 - 0.5 - 1.0 - 1.5} \int_{1.0}^{0.0} \frac{1.5}{-0.5} \frac{1.0}{-0.5} - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0.0 - 0.5 - 1.0 - 0.5 - 0$$

## <u> $B \rightarrow \tau \nu$ : NP implications</u>

Extremely sensitive to scalar BSM extensions (2HDM, LQ), which lift helicity suppression





• Helicity suppressed, loop-level decay dominated by short-distance effects ( $C_{10}$ )

Main uncertainties come from CKM elements (UTA) and decay constants (Lattice)

$$\mathcal{B}(B_q^0 \to \mu^+ \mu^-)^{\rm SM} = \tau_{B_q^0} \frac{G_F^4 |V_{tb}^* V_{tq}|^2 f_{B_q}^2 m_W^4 m_{B_q^0} m_\mu^2}{2\pi^5} \sqrt{1 - \frac{4m_\mu^2}{m_{B_q^0}^2}} |C_{10}^{\rm q,SM}|^2, \quad q = d, s$$

 $|V_{td}|^{\text{UTA}} = 8.59(11) \times 10^{-3}, f_{B_d} = 190.5(1.3) \text{ MeV}$  $|V_{ts}|^{\text{UTA}} = 41.28(46) \times 10^{-3}, f_{B_s} = 230.1(1.2) \text{ MeV}$ 

According to present Lattice estimates, decay constants errors could be halved in the next decade!

### <u> $B \rightarrow \mu\mu$ : the SM status</u>

$$\Rightarrow \begin{array}{l} \mathcal{B}(B_d \to \mu^+ \mu^-)^{\text{SM}} = 9.48(36) \times 10 \\ \mathcal{B}(B_s \to \mu^+ \mu^-)^{\text{SM}} = 3.47(14) \times 10 \end{array}$$







Sensitive to BSM effect on axial and (pseudo)scalar operators, which again lift helicity suppression

$$\mathcal{B}(B_q \to \mu^+ \mu^-) = \mathcal{B}^{\rm SM} \times \left( \left| \frac{C_{10}^{\rm q,NP} - C_{10}^{\prime q,NP}}{C_{10}^{\rm q,SM}} + \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_P^{\rm q,NP} - C_P^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{C_S^{\rm q,NP} - C_S^{\prime q,NP}}{C_{10}^{\rm q,SM}} \right|^2 + \left| \sqrt{1 - \frac{4m_\mu^2}{m_{B_q}^2}} \frac{m_{B_q}^2}{2m_\mu m_b} \frac{m_{B_q}^2}{2m_\mu$$



## <u> $B \rightarrow \mu \mu$ : NP implications</u>

Current results are (now) in perfect agreement with SM prediction, NP strongly constrained Fundamental player in global fit to  $b \rightarrow s\ell\ell$  transitions (in a few slides)





## $B \rightarrow K^{(*)}\nu\nu$ : the SM status

igcolor Loop-level decay dominated by short-distance effects ( $C_L$ ), negligible long-distance

• Main uncertainties as the ones from  $B_s \rightarrow \mu \mu$ , plus additional ones from Form Factors (Lattice)

$$\langle \bar{K}(k) | \bar{s} \gamma^{\mu} b | \bar{B}(p) \rangle = \left[ (p+k)^{\mu} - \frac{m_B^2 - m_K^2}{q^2} q^{\mu} \right] f_+(q^2) + \frac{m_B^2 - m_K^2}{q^2} q^{\mu} f_0(q^2)$$

$$\langle \bar{K}^*(k) | \bar{s}\gamma_\mu (1-\gamma_5) b | \bar{B}(p) \rangle = \varepsilon_{\mu\nu\rho\sigma} \varepsilon^{*\nu} p^\rho k^\sigma \frac{2V(q^2)}{m_B + m_{K^*}} - i\varepsilon^*_\mu (m_B + m_B) + i\varepsilon^*_\mu (m_B + m_B)$$











 $\frac{\mathrm{d}\mathcal{B}}{\mathrm{d}q^2}(B \to K\nu\bar{\nu}) = \mathcal{N}_K(q^2) |C_L^{\mathrm{SM}}|^2 |\lambda_t|^2 \left[f_+(q^2)\right]$ 

$$\mathcal{O}_{L}^{\nu_{i}\nu_{j}} = \frac{e^{2}}{(4\pi)^{2}} (\bar{s}_{L}\gamma_{\mu}b_{L})(\bar{\nu}_{i}\gamma^{\mu}(1-\gamma_{5})\nu_{j})$$

$$\begin{aligned} \mathcal{B}(B^+ \to K^+ \nu \bar{\nu}) \times 10^6 \ \sigma_{\mathcal{B}_{K^+}} / \mathcal{B}_{K^+} \ \mathcal{B}(B^0 \to K_S \nu \bar{\nu}) \times 10^6 \ \sigma_{\mathcal{B}_{K_S}} / \mathcal{B}_{K_S} \end{aligned} \\ (5.06 \pm 0.14 \pm 0.28) \ 0.06 \ (2.05 \pm 0.07 \pm 0.12) \ 0.07 \end{aligned}$$

$$\begin{aligned} \mathcal{B}(B^+ \to K^{*+} \nu \bar{\nu}) \times 10^6 \ \sigma_{\mathcal{B}_{K^{*+}}} / \mathcal{B}_{K^{*+}} \ \mathcal{B}(B^0 \to K^{*0} \nu \bar{\nu}) \times 10^6 \ \sigma_{\mathcal{B}_{K^{*0}}} / \mathcal{B}_{K^{*0}} \\ (10.86 \pm 1.30 \pm 0.59) \ 0.12 \ (9.05 \pm 1.25 \pm 0.55) \ 0.15 \end{aligned}$$

## <u> $B \rightarrow K^{(*)}\nu\nu$ </u>: the SM status

$$]^2$$

$$\frac{\mathrm{d}\mathcal{B}}{\mathrm{d}q^2}(B \to K^* \nu \bar{\nu}) = \mathcal{N}_{K^*}(q^2) |C_L^{\mathrm{SM}}|^2 |\lambda_t|^2 \mathcal{F}(Q^2) |C_L^{\mathrm{SM}}|^2 |\lambda_t|^2 |\lambda_t|^2 \mathcal{F}(Q^2) |C_L^{\mathrm{SM}}|^2 |\lambda_t|^2 |\lambda_t|^2$$

$$\mathcal{O}_{R}^{\nu_{i}\nu_{j}} = \frac{e^{2}}{(4\pi)^{2}} (\bar{s}_{R}\gamma_{\mu}b_{R})(\bar{\nu}_{i}\gamma^{\mu}(1-\gamma_{5})\nu_{R})$$







## <u> $B \rightarrow K^{(*)} \nu \nu$ : NP implications</u>



Possible interpretation also in terms of weakly interacting light NP (axions)

Sensitive to BSM effect on both left-handed and right-handed operator





ullet Loop-level decays dominated by short-distance effects ( $C_{9,10}$ ), important long-distance

Additional uncertainties coming from non-perturbative charming penguins



Potential, unaccounted for, pollution in computation of BRs and angular obs!

## <u> $B \to K^{(*)}\ell\ell, B_{c} \to \phi\ell\ell$ : the SM status</u>



Estimates performed with LCSR, but lacking potential effects from  $D_s$ -D rescattering 





 $\underline{B} \to K^{(*)}\ell\ell, B_{s} \to \phi\ell\ell$ : the SM status









## <u> $B \to K^{(*)}\ell\ell, B_{c} \to \phi\ell\ell$ </u>: NP implications



Combined global fit to all  $b \rightarrow s\ell\ell$  data, connection with  $b \rightarrow s \nu \nu$  to be inspected

• Loop-level decay dominated by short-distance effects ( $C_7$ )

$$BR(B \to X_s \gamma)_{E_{\gamma} > E_0} = BR(B \to X_c \ell \nu) \left| \frac{\lambda_t}{V_{cb}} \right|^2 \frac{6\alpha_{em}}{\pi C} \left[ |C_7^{eff}|^2 + |C_7'|^2 + \delta_{\text{nonp.}} \right]$$

$$BR(B_q \to V\gamma) = \tau_{B_q} \frac{G_F^2 \alpha_{em} m_{B_q}^3 m_b^2}{32\pi^3} \left(1 - \frac{m_V^2}{m_B^2}\right)^3 |\lambda^t|^2 \left(|\mathcal{C}_7|^2 + |\mathcal{C}_7'|^2\right) T_1(0)$$

$$A_{\rm CP}(B_q(t) \to V\gamma) = \frac{\Gamma(\bar{B}_q(t) \to \bar{V}\gamma) - \Gamma(B_q(t) \to V\gamma)}{\Gamma(\bar{B}_q(t) \to \bar{V}\gamma) + \Gamma(B_q(t) \to V\gamma)}$$

### $b \rightarrow s\gamma$ : the SM status

- Inclusive: main uncertainties come from CKM elements (UTA) and non-perturbative contributions

Exclusive: main uncertainties come from CKM elements (UTA) and form factor (Lattice + LCSR)



## Very strong constraints on possible BSM contribution to the radiative operator, particularly from inclusive decay

## <u> $b \rightarrow s\gamma$ : NP implications</u>





Rare decays are a fundamental probe for the search of NP effects. Main theory uncertainties coming from CKM elements, decay constants and form factors

potential origin and connection with other sectors (light NP?)

## Conclusions

After re-analysis of LFUV ratios by LHCb, evidence of LFV NP is gone. Remaining hints of deviation in the muon sector are to be considered with care, due to charming penguins

New discrepancy recently observed in  $B \rightarrow K \nu \nu$ , still much work to do to understand its

