

# Presentazione sigla VIRGO

M. Bawaj, Virgo Perugia group

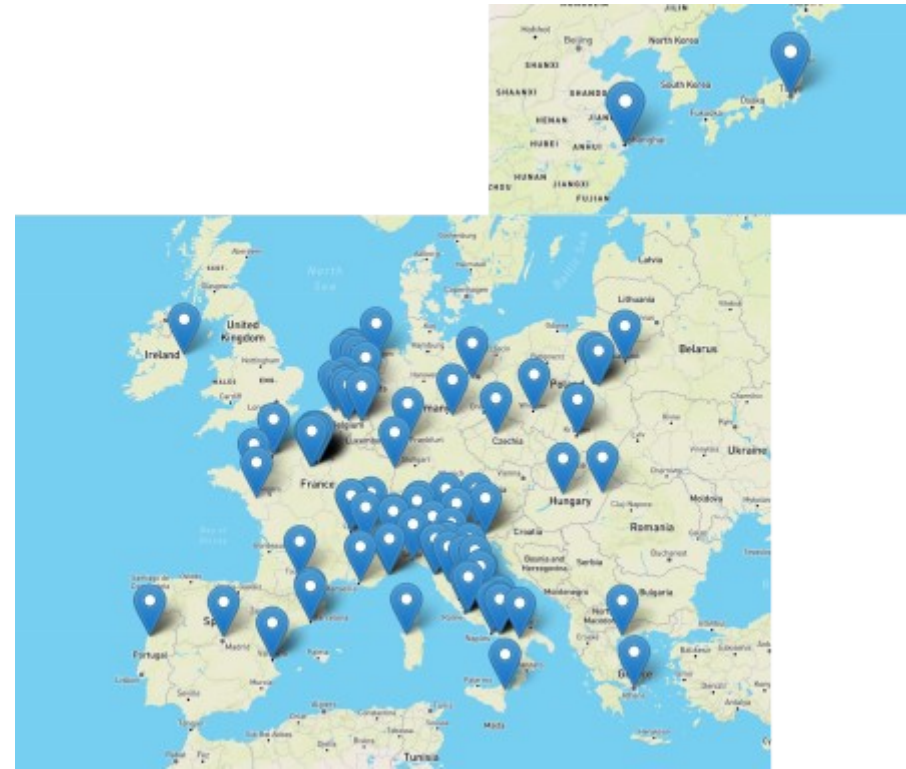
Consiglio di Sezione – 14.07.2022





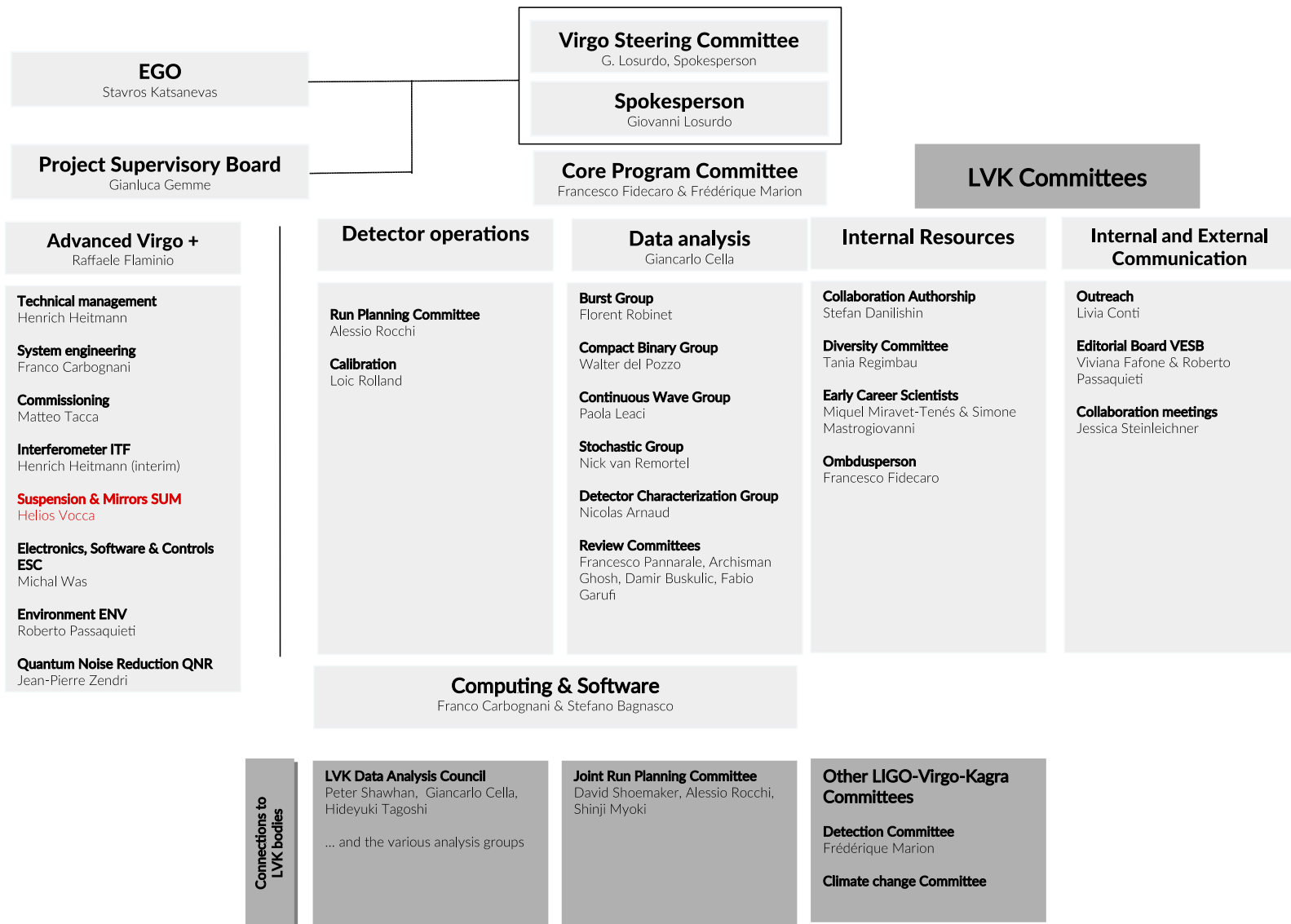
# Virgo Collaboration

- ~840 members, ~580 authors, 146 institutions from 16 countries  
+70 members, +130 authors, +15 institutions w.r.t. to last year
- 37 Groups:
- 9 countries represented in the VSC



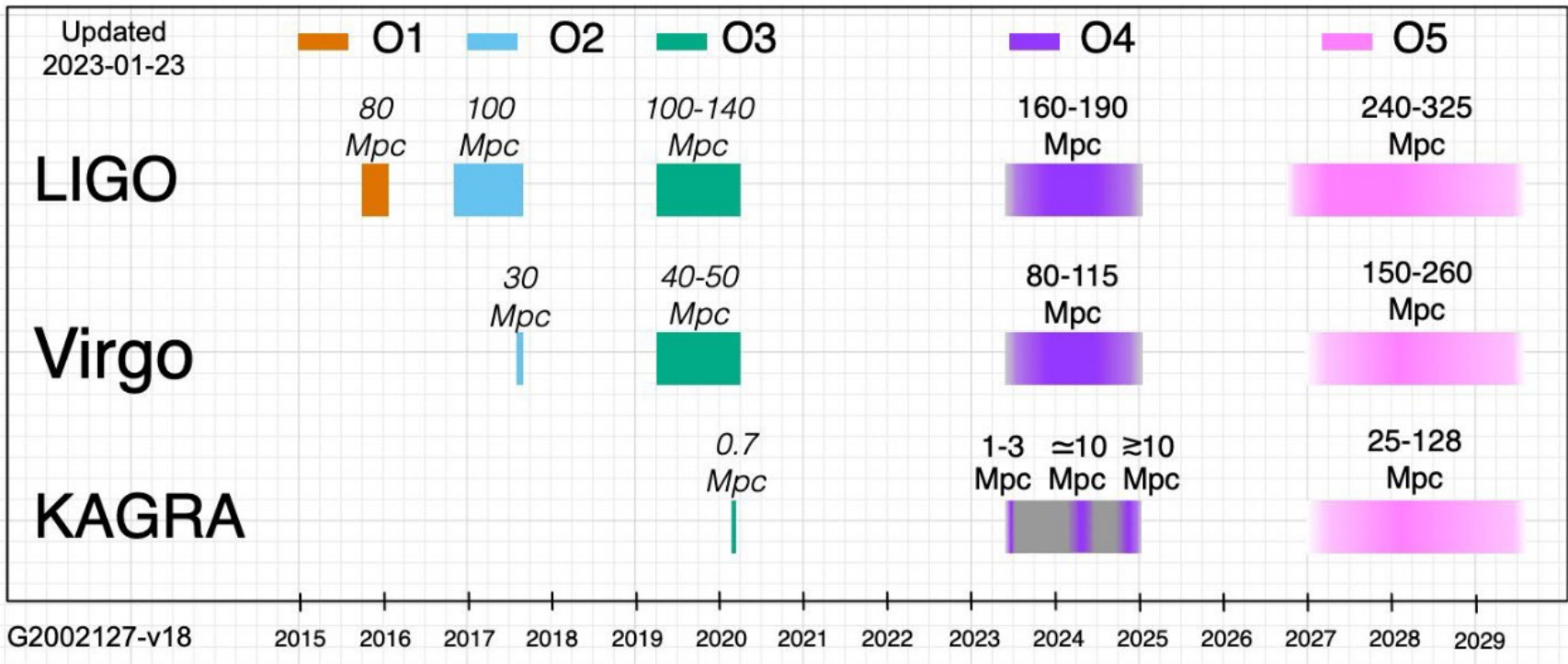
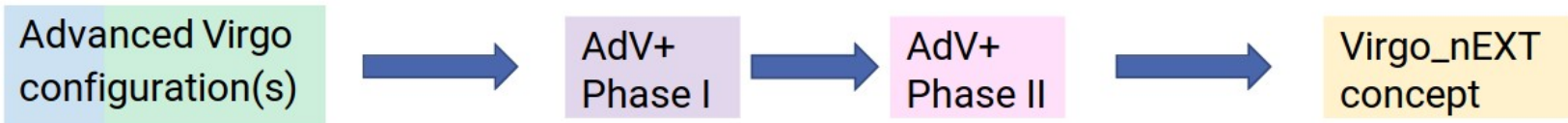


# Organization chart

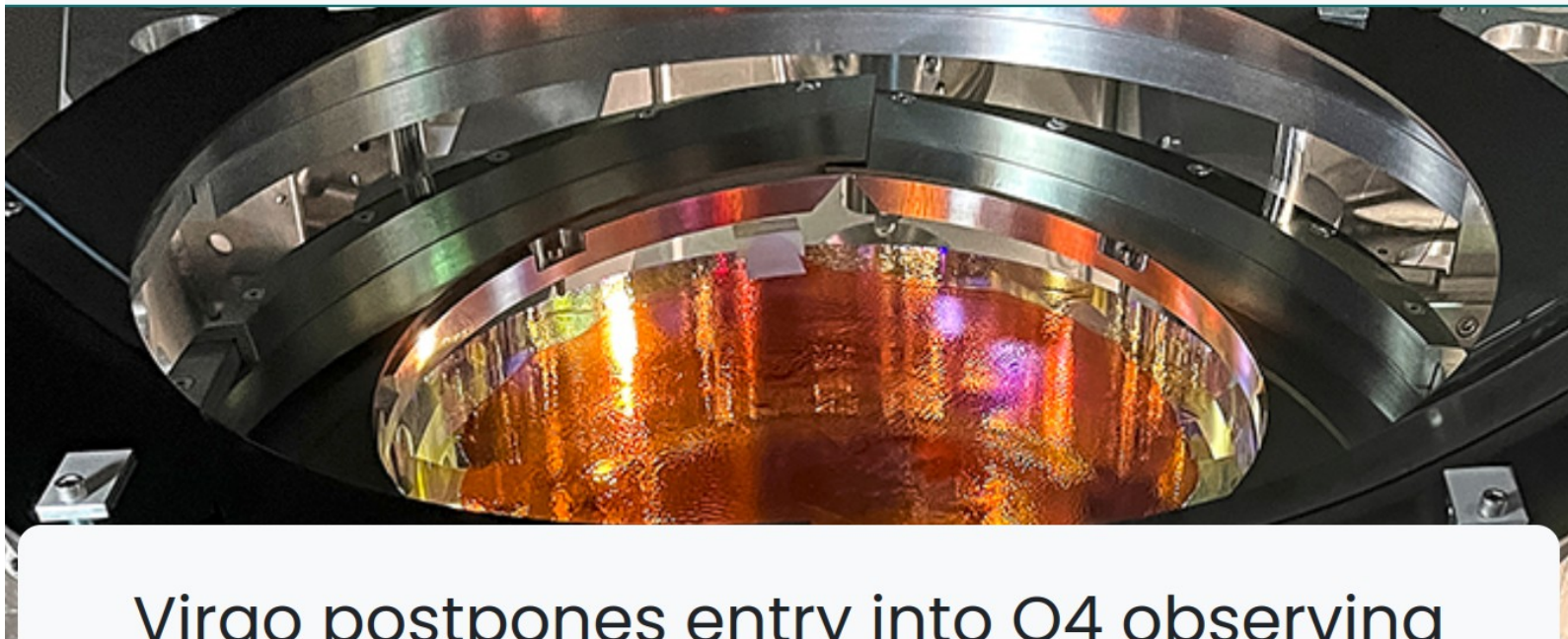




# Virgo upgrades and Observing runs



G2002127-v18



# Virgo postpones entry into O4 observing run

May 11, 2023

The Virgo Collaboration has decided to postpone the entry of Virgo in the next observing run (O4), scheduled for May 24th, in order to continue the detector commissioning activities and further increase its sensitivity.

Since the end of the O3 observing run in 2020, the Virgo interferometer has undergone a major upgrade to improve its sensitivity for the new joint observing run with the LIGO and KAGRA interferometers. This upgrade has nevertheless required several months for the team to make the detector stable again.

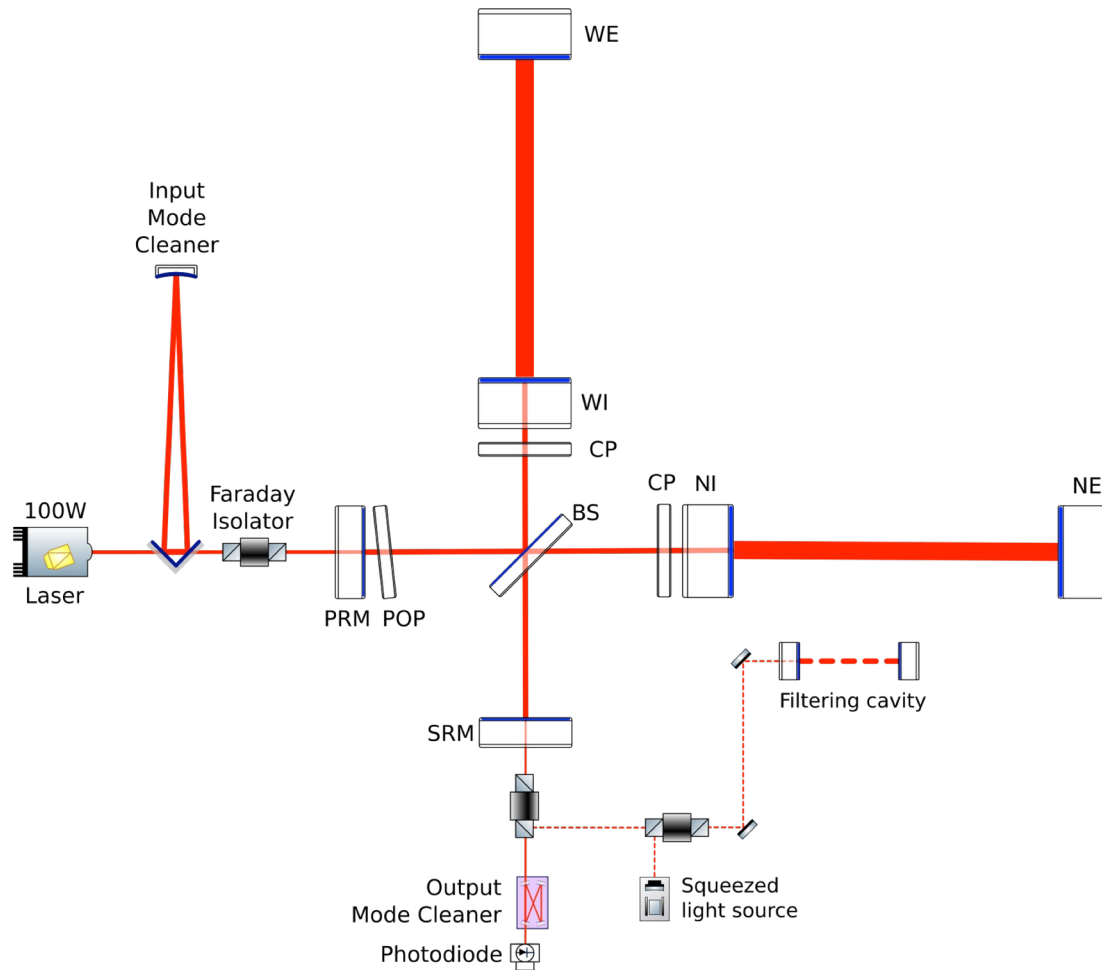
To date Virgo would be able to observe events similar to some already detected in previous observing runs. To go further, a long and intense hunt for the multiple sources of noise that could limit the sensitivity of the interferometer is now necessary. Additionally, subtle technical



# Advanced Virgo+ phase I

**+ Signal Recycling Cavity**

**+ Frequency Dependent Squeezing**

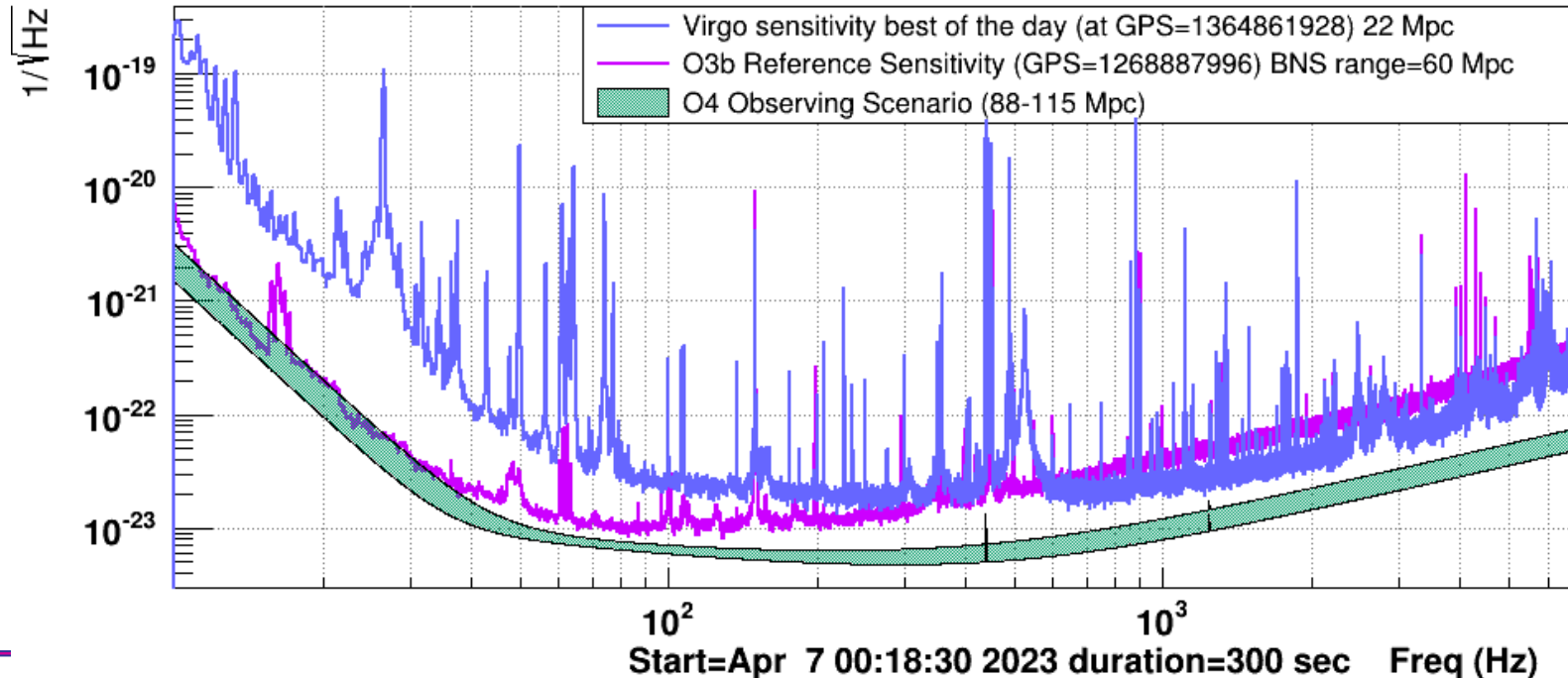




# Virgo sensitivity at high frequency

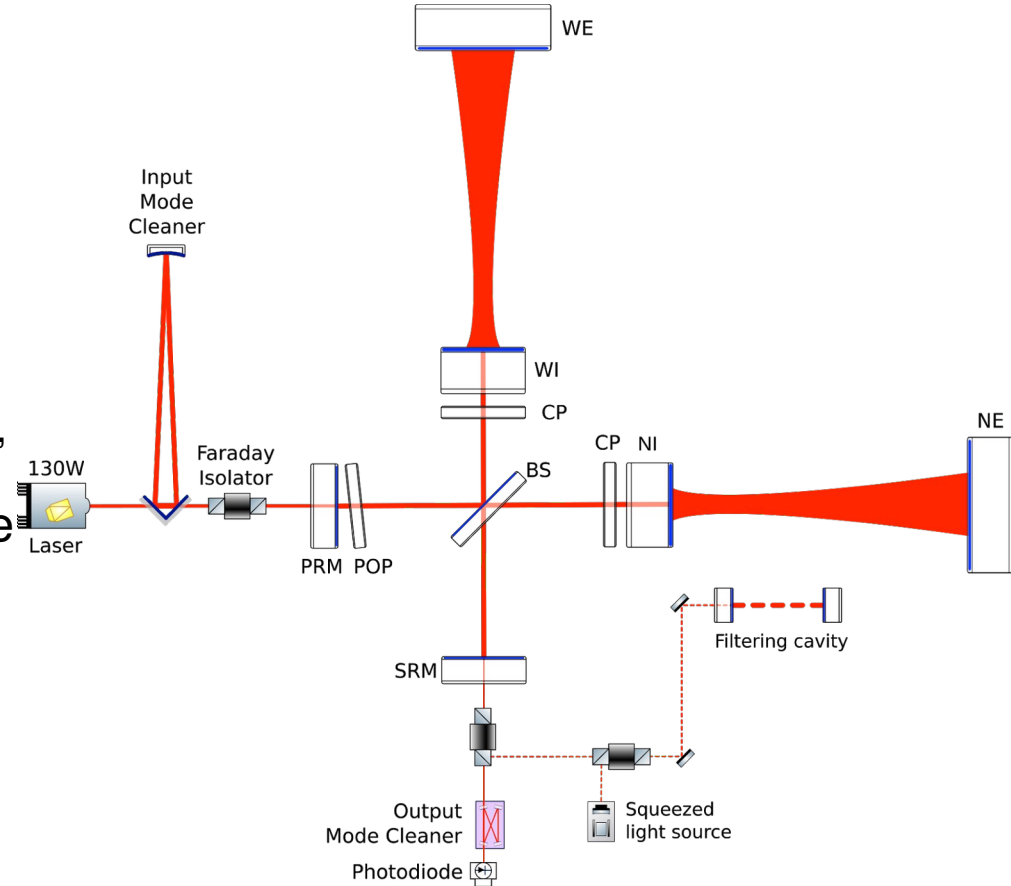
- Sensitivity at high frequency better than in O3
  - ◆ The signal recycling cavity does enlarge the detector bandwidth

Sensitivity for best BNS range of the day (22 Mpc)



## ● Main changes

- ◆ Larger beams on end test masses
  - » 6 cm radius  $\rightarrow$  10 cm radius
- ◆ Larger end mirrors
  - » 35 cm diameter  $\rightarrow$  55 cm diameter
  - » 40 kg  $\rightarrow$  100 kg
- ◆ Better mirror coatings
  - » Lower mechanical losses, less point defects, better uniformity
- ◆ New suspensions/seismic isolators for large
- ◆ Further increase of laser power
  - » 40W  $\rightarrow$  60W  $\rightarrow$  80 W

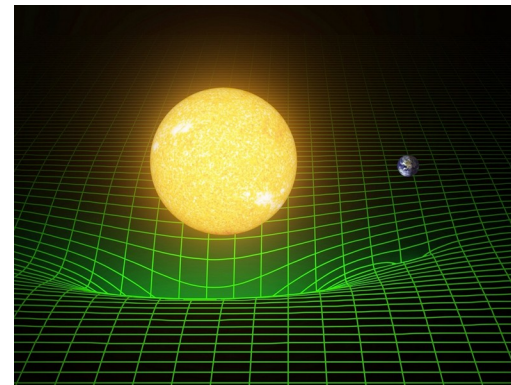






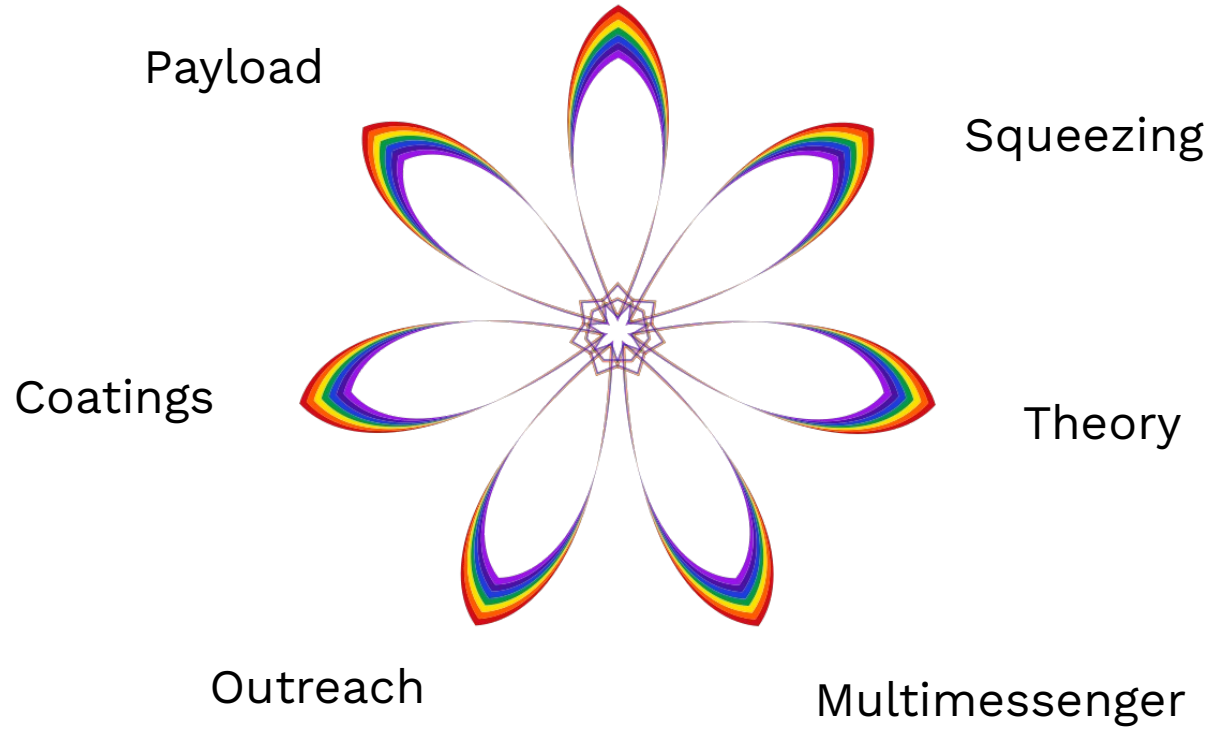
# Framework

- ET is the future of the EU GW community
- ET is becoming real 😊
- ET is a great attractor but also a great perturbation
- Virgo community may have an important role in the ET success by contributing to ET
- **We need to balance between the participation in ET and the continuation of the Virgo program**





# Research lines





# Research on mirror suspensions

## Main research topics: characterization and installation of mirror suspensions

- Mechanical loss measurements and thermal noise modelling
- Optimization and realization of the HCB bonding of all functional parts of the main optics
- Realization of the handling boxes for all the main Virgo optics

## Experimental facilities:

- FARO-CAM2 arm
- Setup for mechanical losses measurement at room and cryogenic temperature
- Materials Testing Machine
- Microscopy (SEM-EDX)
- FTIR spectroscopy
- X-ray diffraction and x-ray fluorescence

## HCB procedure at LMA and at positioning test with FARO arm Virgo/EGO site





# Research on suspensions

## **Members:**

B. Bracco, A. Di Michele, F. Travasso, H. Vocca,  
A. Piluso, D. Aisa, S. Aisa

## **Responsibilities:**

- AdV+ SUM  
(Suspensions&Mirrors)  
System coordination
- Monolithic suspension  
assembling coordination



# Test Large Mass

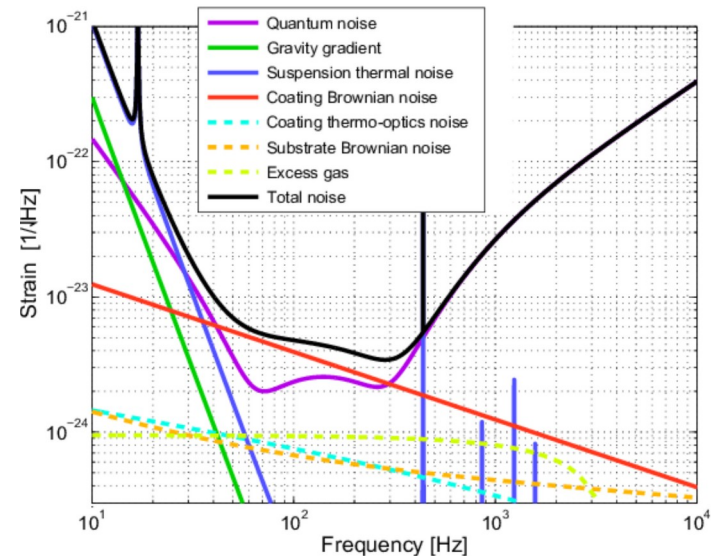
Il rumore termico degli specchi è il limite fondamentale alla sensibilità di Virgo nella regione di frequenze 40Hz-300Hz. Per poterlo ridurre vi sono 2 vie:

- Coating con dissipazioni meccaniche più basse
- Specchi più larghi

Quest'ultima via prevede l'assemblaggio delle varie parti con una precision intorno ai 10 $\mu$ m, cosa possibile tramite l'uso del braccio FARO

$$\text{CTN}(f) \propto \frac{\sqrt{\frac{T}{f} \frac{1}{w^2} \phi d}}{L}$$

Temperature (points to T)  
Mechanical loss (points to  $\phi$ )  
Coating thickness (points to d)  
Beam-size (points to w)  
Arm length (points to L)



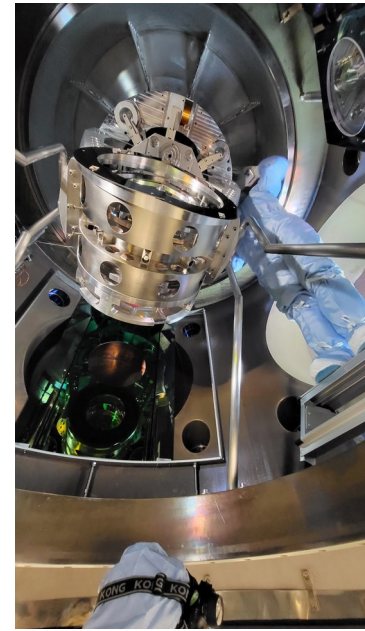
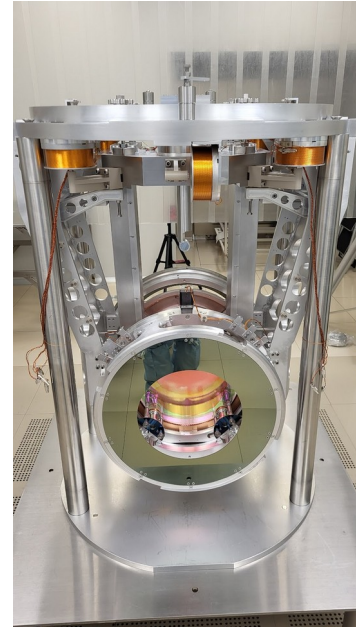




# Lavori sulle sospensioni WI e NE

La sospensione della torre WI presentava dei rumori in eccesso dovute ad un magnete rotto che è stato tolto in torre.

La spensione della torre NE presentava un comportamento anomalo per cui si è deciso di riassemblare tutta la sospensione.



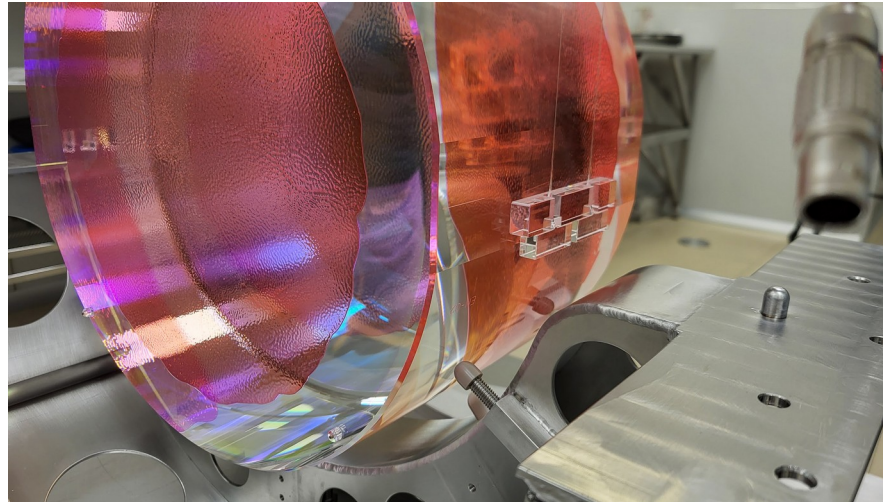




# PRIN – Incollaggi rimovibili

Le parti di supporto per la sospensione degli specchi sono costituite da elementi di vetro incollati allo specchio o tra di loro con la tecnica del HCB\*. Tale tecnica permette un incollaggio poco dissipativo e molto forte che in caso di rottura delle sospensioni porta al danneggiamento dei vari elementi.

Per questa ragione si sta cercando di sviluppare un incollaggio poco dissipativo ma rimovibile senza danneggiare le varie parti





# Research on coatings

## Main activities: characterization of amorphous oxides and semiconductors coatings

- Mechanical loss measurements (GeNS), thermal noise modelling
  - Morphological and compositional characterizations (SEM-EDX)
  - Structural characterizations (XRD-XRF, Raman, FTIR and XAS spectroscopy)
  - Elastic characterization (Brillouin spectroscopy)
- mechanical characterization of optimized silicon substrates
- first layers of GaAs deposited (collab. with CNR-IOM), structure & morphology measurements done

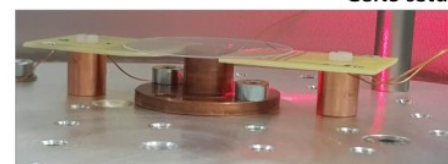
## Upgrades of experimental facilities:

- A new design for a two-chambers GeNS cryogenic has been accepted by the company
- Brillouin setup fully operational for experiments on thin films (upgrade for Q-vector selection and polarization analysis in progress)
- FTIR upgraded with sample holders for grazing incidence measurements
- SEM-EDX: new software for EDX analysis
- High temperature (1500 °C) cell for Raman/x-rays available

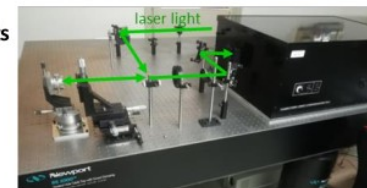
SEM analysis



GeNS setup



cell for HT Raman/x-ray expts



BLS setup

## Experiments at external laboratories & synchrotron radiation facilities:

May 2022: @University of Venice: investigation of Argon desorption from silica coatings

*on approved synchrotron radiation peer-reviewed proposals:*

Nov 2022: SESAME (Amman, JO) (proposal n. 20190078), *Structure of Ta<sub>2</sub>O<sub>5</sub> and TiO<sub>2</sub>:Ta<sub>2</sub>O<sub>5</sub> films*

*Submitted:*

May 2023: FERMI@Elettra (Trieste, IT) (proposal n. 20234083), *Nanoscale study of mechanical properties of amorphous coatings for gravitational wave detectors mirrors*

ATR-FTIR setup





# Research on coatings

## Members:

B. Bracco,  
S. Corezzi,  
A. Di Cicco,  
A. Di Michele,  
P. Sassi,  
L. Silenzi,  
A. Trapananti,  
F. Travasso,  
H. Vocca

## Presentations at conferences & workshops

- B. Bracco et al., “Structural characterization of amorphous silica coatings combining specular reflection (SR) and attenuated total reflection (ATR) infrared spectroscopic techniques”, 12<sup>th</sup> International Conference on Advanced Vibrational Spectroscopy, August 27 – September 1, 2023, Krakow (PL) - oral
- B. Bracco et al., “Structural characterization of amorphous silica coatings”, VISPEC conference, June 14-16, 2023, Perugia - oral
- L. Silenzi et al., “A structural study of properties of amorphous silica coatings for low internal friction optics”, Gravitational-Wave Advanced Detector Workshop GWADW, 21-27 May 2023, La Biodola (Isola d’Elba, Italy) – oral
- F. Travasso et al. “Cryogenic measurements of the loss angle of thin films: a geometric approach”, XIII ET Symposium, May 8-12, 2023 - oral
- B. Bracco et al., “A multi-technique approach to the chemical characterization of amorphous coatings” LVK meeting, March 13, 2023 - poster
- L. Silenzi et al, “Cryogenic measurements of crystalline substrates”, Virgo Week, 11 November 2022 - oral
- L. Silenzi et al., “Effects of thermal treatments on silica substrates”, LVK meeting, 12 September 2022 - oral
- A. Trapananti et al., “Short range atomic structure in amorphous oxides for low mechanical loss optical coatings”, JMC2022, 22-26 August 2022, Lyon (France) - invited oral
- S. Corezzi et al., “Status of surface Brillouin light spectroscopy of coatings” Virgo Week, 8 July 2022, Cascina (Pisa), Italy - oral

## Thesis/traineeships

- “Analisi di film sottili mediante riflettività di raggi X”, BSc. (laurea triennale Fisica) – Univ. Di Camerino, E. De Berardinis, supervisor A. Trapananti, October 2022
- ongoing: PhD thesis L. Silenzi (Univ. Camerino)
- “Caratterizzazione IR di film sottili di SiO<sub>2</sub>”, BSc. (laurea triennale Chimica) – Univ. di Perugia, F. Tuteri, supervisors P.Sassi & B.Bracco, June 2023

## Publications

- L. Silenzi et al, “Structure and elastic properties of annealed silica coatings for mirrors of gravitational waves detectors”, in preparation
- L. Silenzi et al, “Cryogenic measurements of the loss angle of thin films: a geometric approach”, in preparation





# Research on the Quantum Noise Reduction

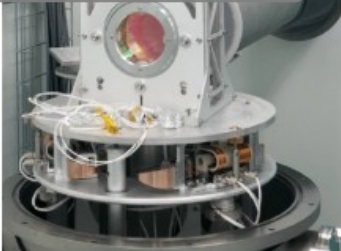
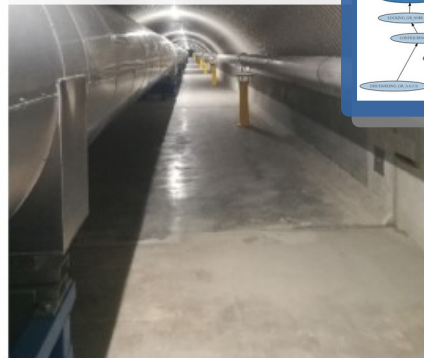
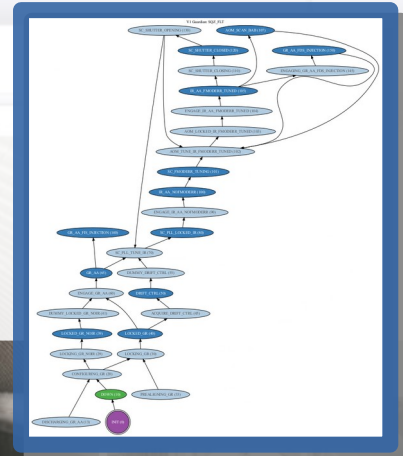
**Members:** M. Bawaj

**Main research topics:**

- Commissioning, Automation and Electronics development of the Frequency Dependent Squeezing implementation.
- Software and Electronics development for the EPR squeezing demonstrator.

**Collaborations:**

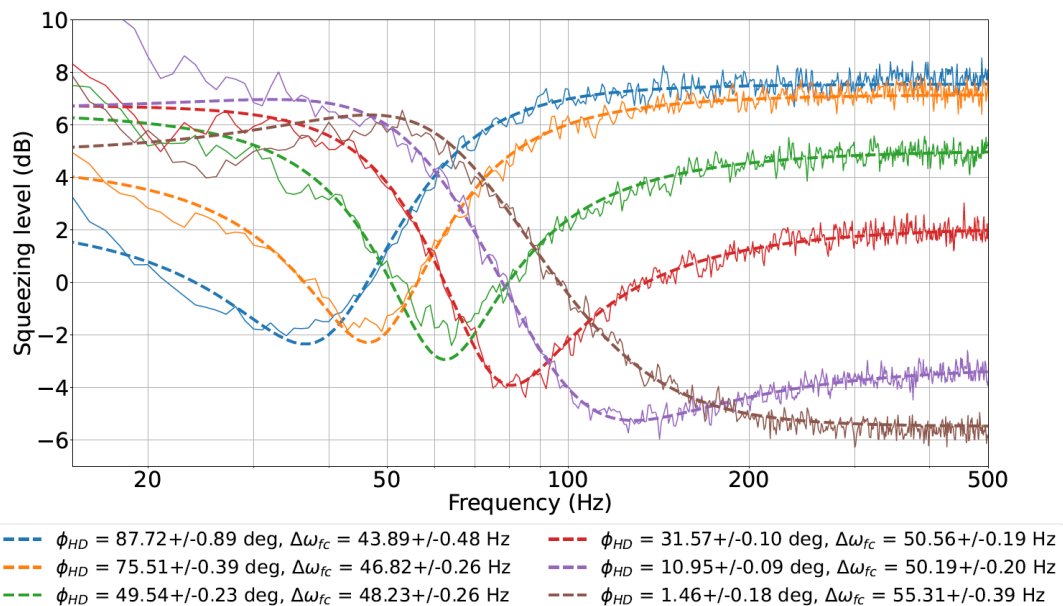
- with the Kyung Hee University, Korea Astronomy and Space Science Institute and Yonsei University on the construction of the EPR demonstrator.





# Commissioning of Frequency Dependent Squeezing

- Frequency dependent squeezing measured down to 25 Hz
- System ready, paper published on PRL



## Frequency dependent squeezed vacuum source for the Advanced Virgo gravitational wave detector

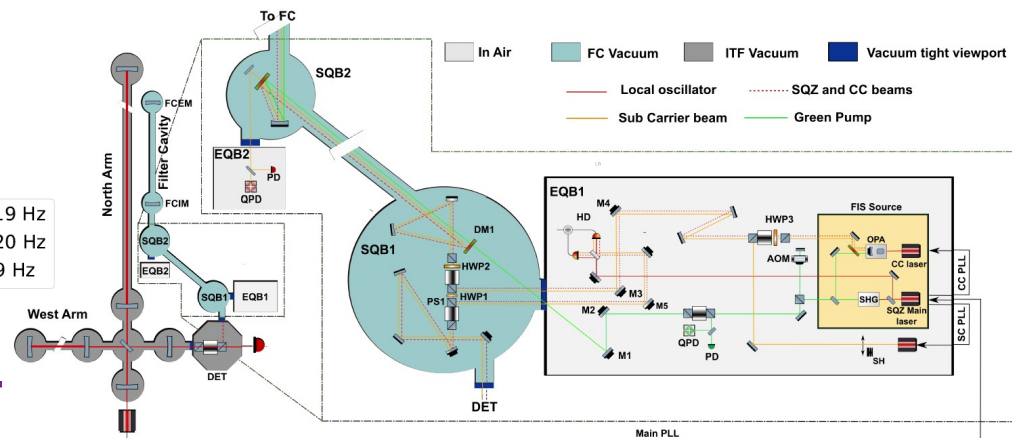
F. Acernese *et al.* (the Virgo Collaboration),

H. Vahlbruch, M. Mehmet, H. Lück, and K. Danzmann

Institut für Gravitationsphysik, Leibniz Universität Hannover and Max-Planck-Institut für Gravitationsphysik (Albert-Einstein-Institut), Callinstr. 38, 30167 Hannover, Germany

(Dated: May 4, 2023)

In this paper we present the design and performance of the frequency dependent squeezed vacuum source that will be used for the broadband quantum noise reduction of the Advanced Virgo Plus gravitational-wave detector in the upcoming observation run. The frequency dependent squeezed field is generated by a phase rotation of a frequency independent squeezed state through a 285 m long, high-finesse, near-detuned optical resonator. With about 8.5 dB of generated squeezing, up to 5.6 dB of quantum noise suppression has been measured at high frequency while close to the filter cavity resonance frequency, the intra-cavity losses limit this value to about 2 dB. Frequency dependent squeezing is produced with a rotation frequency stability of about 6 Hz RMS, which is maintained over the long term. The achieved results fulfill the frequency dependent squeezed source requirements for Advanced Virgo Plus. With the current squeezing source, considering also the estimated squeezing degradation induced by the interferometer, we expect a reduction of the quantum shot noise and radiation pressure noise of up to 4.5 dB and 2 dB, respectively.





# Multi-messenger astronomy

## Members:

M. Bawaj, M.L. Brozzetti, S. Cutini, G. Greco, T. Matcovich, M. Punturo



## Main activities:

- Preparation of tutorials for scientists from the multi-messenger field in the framework of EOSC, IVOA, VO School
- Development of progressive Web Apps for Multi-Messenger Cosmology
- Update to the LIGO-Virgo KAGRA User Guide and lowlatency in O4
- Active Rapid Response Team for immediate GW trigger handling during O4

<https://virgo.pg.infn.it/maps/>

<https://virgo.pg.infn.it/gladenet/catalogs/>



## Related presentations of the group:

- M. Bawaj, “Transient localization web service based on open gravitational-wave data for the multi-messenger community” talk at the GRAvitational - wave Science&technology Symposia (GRASS) in Padova
- M. L. Brozzetti, “GLADEnet: a progressive Web App for Multi-Messenger Cosmology” talk at the LVK collaboration meeting in Chicago
- G. Greco, Invited speaker, “GWapps: to drive multi-messenger researches based on the FAIR approach” at the European Astronomical Society Annual Meeting
- G. Greco, Invited speaker, "ESCAPE to the Future" talk for the Royal Belgium Institute of Natural Sciences
- G. Greco, LOC member & speaker, The IVOA Interoperability Meeting (Northern Spring)





# Multi-messenger Virgo group

## Members:

M. Bawaj, M.L. Brozzetti, S. Cutini, G. Greco, T. Matcovich, M. Punturo



## Related papers of the group:

- M. L. Brozzetti, et al., “GLADEnet: a progressive Web App for Multi-Messenger Cosmology” (preliminary title) paper to be submitted soon.
- Greco, G., et al. “Multi Order Coverage Data Structure to Plan Multi-Messenger Observations.” *Astronomy and Computing* 39 (April 2022): 100547. <https://doi.org/10.1016/j.ascom.2022.100547>. + Associated tutorial published in the European Open Science Cloud and promoted by the International Virtual Observatory Alliance
- Greco, G., et al., “Innovative web apps for sky regions in the multi-messenger research” (preliminary title) EAS proceeding to be submitted soon.

## Collaborations:

- GECKOrchestrator on the software development at the Seoul National University.  
Gravitational-wave Follow-up with GECKO and 7-DIMENSIONAL TELESCOPE (paper in the review process)



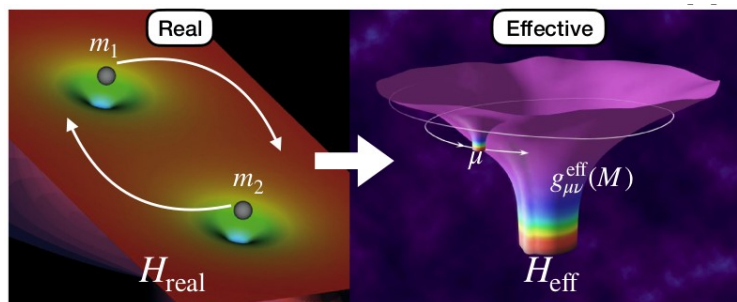


# Research on waveform modelization

## Main research topics:

- Effective-one-body approach to gravitational wave models for compact binary systems
- Inclusion of eccentricity effects
- Inclusion of memory effects

## EOB approach



[Buonanno-Damour 1999]

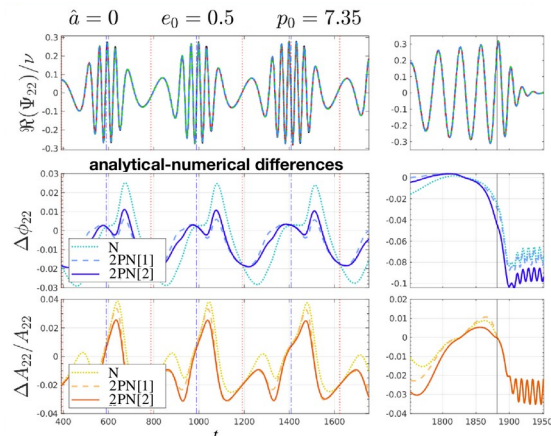
$$H_{\text{real}} = Mc^2 \sqrt{1 + 2\nu \left( \frac{H_{\text{eff}}}{\mu c^2} - 1 \right)}$$

## Upgrades for the eccentric waveform model TEOBResumS-DALI

$$h_+ - ih_\times = D_L^{-1} \sum_{\ell=2}^{\ell_{\text{max}}} \sum_{m=-\ell}^{\ell} h_{\ell m} {}_{-2}Y_{\ell m}$$

$$h_{\ell m} = h_{\ell m}^{(N, \epsilon)} \hat{h}_{\ell m}^c \hat{h}_{\ell m}^{\text{nc}}$$

Extra noncircular factors  
(inst, tail, memory)





# Theory Virgo Group

## **Members:**

G. Grignani, T. Harmark, R. Olivieri, M. Orselli and A. Placidi

## **Related papers of the group:**

- [arXiv:2112.05448](https://arxiv.org/abs/2112.05448)
- [arXiv:2202.10063](https://arxiv.org/abs/2202.10063)
- [arXiv:2203.16286](https://arxiv.org/abs/2203.16286)
- [arXiv:2305.14440](https://arxiv.org/abs/2305.14440)

Paper on PM EOB Hamiltonian in preparation

## **Collaborators therein:**

- Niels Bohr Institute (Copenhagen) - Troels Harmark
- IHES (Paris) - Alessandro Nagar, Thibault Damour



# Outreach

## Participants:

M. Bawaj, S. Corezzi, A. Di Michele, G. Greco, F. Travasso

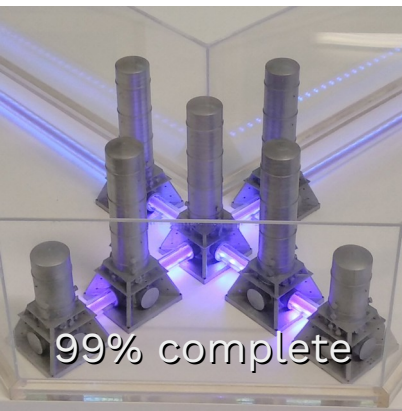
## Main group activities:

- publication of the interactive website *Soundmap*

<https://virgo.pg.infn.it/soundmap>

- proposal for SHARPER 2023:

*work in progress...*



Educational Soundmap: exploring gravitational-wave transient events

The educational soundmap shows some sky localizations among the most significant gravitational-wave events discovered by the LIGO, Virgo and KAGRA collaborations. The interactive sky map is sponsored. A specific chord is played when the cursor enters or leaves the coverage of the gravitational-wave sky localization. The binary merger of compact objects that LIGO and Virgo have detected so far are in the audio band. They can be converted to sound files, so that you can hear them. Once the map has been anchored by the cursor, the associated sound file is executed. These audio files were produced by the Gravitational Waves Data Science Center and they are intended for educational purposes. For additional information you can visit [the page Audio Files](https://www.ligo.org). The gravitational-wave sky localizations shown here are the areas of the 90% credible regions and they are built using the skymaps from the Gravitational-wave Transient Catalog (GWTC). **Note:** To enjoy the full use of headphones is recommended to hear the frequency content.

Detection	Sky localisation
GW150914 - GWTC-1 skymap	<input type="checkbox"/>
GW170814 - rapid LIGO detection	<input checked="" type="checkbox"/>
GW170814 - rapid LIGO and Virgo detection	<input checked="" type="checkbox"/>
GW170814 - GWTC-1 skymap	<input checked="" type="checkbox"/>
GW170817 - GWTC-1 skymap	<input checked="" type="checkbox"/>
GW170817 - GWTC-1 skymap	<input checked="" type="checkbox"/>
GW170817 - GWTC-1 skymap	<input checked="" type="checkbox"/>
AT2017gfo - kilonova	<input checked="" type="checkbox"/>
GW190412 - GWTC-1 skymap	<input checked="" type="checkbox"/>
GW190521 - GWTC-1 skymap	<input checked="" type="checkbox"/>
GW190814 - GWTC-1 skymap	<input checked="" type="checkbox"/>
GW200115 - GWTC-1 skymap	<input checked="" type="checkbox"/>

**Sky surveys**

- Fermi -GALEX GRB AS BOSS
- Hubble -Pantheon -ZEMOS -GRB -KAGRA

**Artistic constellations**

- Painting-Kagaya
- Painting-Col

<https://www.sharper-night.it/>

<https://www.festascienzafilosofia.it/>



**SHARPER**  
**29.09.2023**

SHARPER 30.09.2022  
SHARPER 24.09.2021

HOME | IL PROGETTO | LE CITTÀ | PROGRAMMA | MARATONA | SOCIAL WALL

CONTATTI

ONDE GRAVITAZIONALI - NUOVI INGREDIENTI





# Virgo Perugia

## **Servizi**

Il gruppo continuerà a richiedere il coinvolgimento dei servizi, per lo più meccanici ma anche di calcolo. Non si prevedono impegni continui per le sospensioni. È vitale una presenza di circa 2 mesi uomo del servizio meccanico in missione in Giappone (principalmente nei periodi estivi).



# Anagrafica

Nel 2023 e 2024 il Gruppo di Perugia mantiene il numero di FTE rispetto all'anno scorso e mantiene anche la divisione di FTE fra gli argomenti di ricerca.

9,4 FTE, distribuiti tra PG: 5,8 e Cam: 3,6

(Lo scorso anno gli FTE erano circa 9)

Richiesta 2024 @ CSN2 ~280k€

Nome	Qualifica	Impegno
E. Z. Appavuravther	PhD	100%
M. Bawaj	RTD-b	80%
M. L. Brozzetti	PhD	50%
S. Corezzi	Prof. Associato	50%
S. Cutini	Ric. INFN	20%
A. Di Cicco	Prof. Ordinario	30%
A. Di Michele	RTD-a	100%
G. Grignani	Prof. Ordinario	20%
F. Marchesoni	Prof. Ordinario	30%
M. Orselli	Prof. Associato	20%
T. Pitik	Borsista	100%
M. Punturo	Dir. di Ric. INFN	40%
P. Sassi	Prof. Associato	40%
L. Silenzi	PhD	100%
A. Trapananti	Prof. Associato	40%
F. Travasso	RTD-b	60%
H. Vocca	Prof. Associato	60%
<b>Tot.</b>		<b>9,4</b>