

I.S. GSS-MI - INFN, 05 July 2023

Gauge theories, Supergravity and String theory

Mi-Bicocca, Mi-Statale, Genova,
Padova, Lecce, Pisa, Torino

Coord. naz.: Anna Teresa Ceresole (TO) → Davide Cassani (PD)

String theory, M-theory, Supergravity

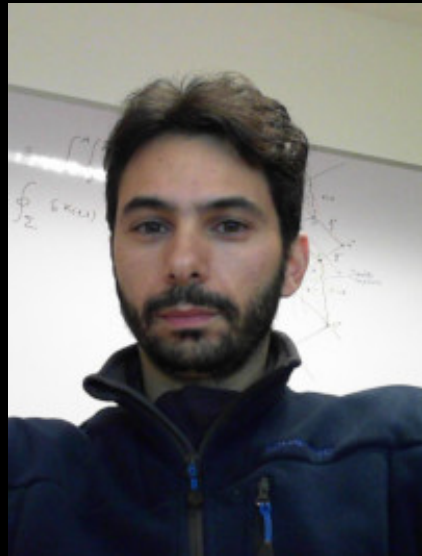
Perturbative and non perturbative gauge theories

Supersymmetric Black Holes, holography, micro state counting

Models of SUSY breaking in cosmology and particle physics



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<https://www0.mi.infn.it/~strings/index.php>

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A covariant description of space-times in Einstein and extended gravity

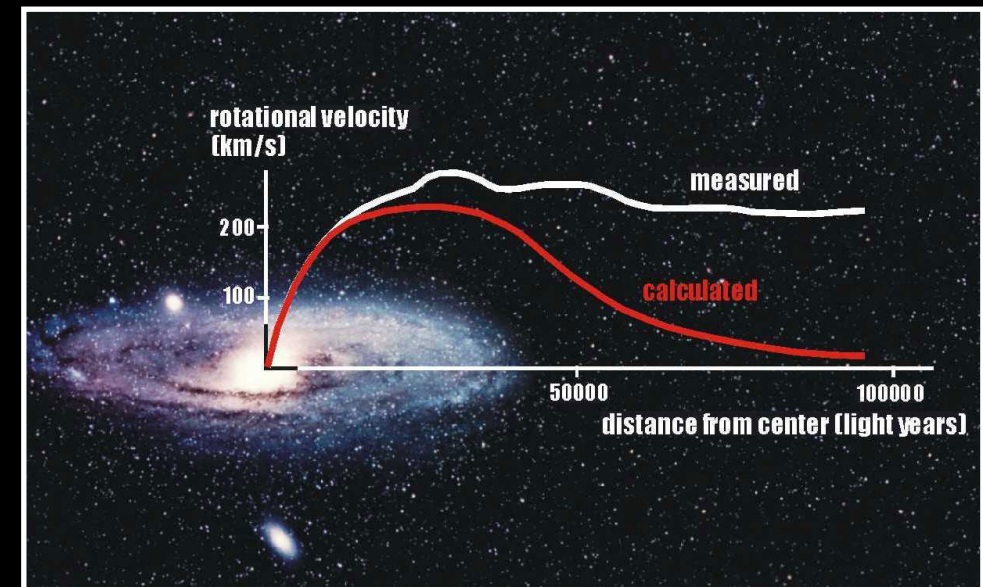
The field equations of gravitational theories are covariant, and equate a geometric tensor (e.g. Einstein, Cotton, Bach tensor) to a tensor describing matter. Solutions are usually found in coordinates that exploit the symmetries. However, there are advantages in keeping the coordinate-free tensor description as far as possible. Besides the formal elegance, it naturally addresses scalar identities.

A selection of publications by C.A.Mantica & L.G.Molinari (& S. Capozziello, NA):

- The covariant approach to static space-times in Einstein and extended gravity theories.
- Codazzi tensors and their space-times, and Cotton gravity (GERG 2023)
- Spherical doubly warped spacetimes for radiating stars and cosmology (GERG 2022)
- Geometric perfect fluids from Extended Gravity (EPL perspectives 2022)
- Doubly torqued vectors and a classification of doubly twisted and Kundt spacetimes (GERG 2021)
- $w = 1/3$ to $w = -1$ evolution in a Robertson–Walker space-time with constant scalar curvature (IJGMMP 2019)
- Cosmological perfect-fluids in $f(R)$ gravity (IJGMMP 2019)
- ...
- Extended Derdziński-Shen theorem for curvature tensors (Coll Math 2012)

Col gruppo abbiamo condiviso tesi di laurea

Fritz Zwicky (Mt. Wilson 1933): missing mass in the motion of galaxies in galactic clusters (virial) - Vera Rubin ('70): the rotation curve of galaxies (Newt)



THE QUEST FOR DARK MATTER

New geometry ?

$$G_{\mu\nu} = \kappa T_{\mu\nu}$$

New particles ?

For cosmology we prove:

In many extended theories of gravity in RW background (ex: $f(R)$, string-inspired corrections to Einstein-Hilbert action) the curvature corrections to Einstein equations have the perfect fluid form as required by the COSMOLOGICAL PRINCIPLE.

Cosmological perfect-fluids in $f(R)$ gravity (IJGMMP 2019), ...

Geometric perfect fluids from Extended Gravity (EPL 2022)

They modify the perfect-fluid parameters of the matter source and may correspond to observed effects of unobserved DM (S. Capozziello).



Doubly torqued vectors and a covariant classification of a (large) class of space-times

$$ds^2 = - e^{-\beta(x,t)} dt^2 + e^{\alpha(x,t)} g_{\mu\nu}^{\star}(x) dx^{\mu} dx^{\nu}$$

$$\begin{aligned}\nabla_i \tau_j &= \kappa g_{ij} + \alpha_i \tau_j + \tau_i \beta_j \\ \alpha_j \tau^j &= 0, \quad \beta_j \tau^j = 0\end{aligned}$$

A covariant approach to static space-times in Einstein and extended gravity theories

$$ds^2 = - e^{-\beta(x)} dt^2 + g_{\mu\nu}^{\star}(x) dx^{\mu} dx^{\nu}$$

$$\begin{aligned}\nabla_i u_j &= - u_i \dot{u}_j \\ \nabla_i \dot{u}_j &= \nabla_j \dot{u}_i\end{aligned}$$

Geometry dictates the covariant expressions for the Ricci, Weyl, Cotton, Bach tensors

A gravity theory (Einstein, Cotton, f(R), conformal - each one with its own variational principle) implies the form of energy-momentum tensor.

In spherical symmetry, we reobtain in simple manner several known solutions (Schwarzschild, Reissner-Nordstrom, Harada, Mannheim-Kazanas, some BH ... and new ones), and properties of non-lin electrodynamics

$$F_{jk} = \frac{E}{\sqrt{\eta}} (u_j \dot{u}_k + u_k \dot{u}_j) + B (y_j z_k + y_k z_j) \quad (\text{Faraday tensor})$$

(The M-K and Harada solutions both have a log term in beta that mimics DM)