



Apparatus for Meson and Baryon Experimental Research

AMBER (CSN1)

P. Zuccon preCSN2 TIFPA 2023

Introduction

- New accurate measurements have brought Cosmic Rays to a precision era.
- A similar precision is needed in the knowledge of peculiar cross sections is now requested
- Three kind, maybe four, class of problems
 - For all elements: Isotopic Nuclear fragmentation cross sections on ISM (p and He nuclei)
 - Production of pbar and positrons in dominant CR-ISM interactions (p-p and p-He, He-p)
 - Disappearance cross sections of Nuclei on the CR detectors material (mostly C and Al).
 - Cross section for producing anti-nuclei (Dbar, ³Hebar, ⁴Hebar) in dominant CR-ISM interactions (p-p and p-He, He-p)



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Solve the "cross-pollination" of CR spectra measured in the Heliosphere



Obtain a better sensitivity to signals from exotic physics



Reduce systematic errors on the measurements



Improve sensitivity to primordial antimatter and to DM **Properties of Cosmic Antiprotons** The antiproton-to-proton flux ratio shows that above 60 GV the ratio is energy independent.



Antiproton and proton spectra



A sample of recent papers on AMS antiproton data

P. Mertsch *et al.*, Phys. Rev. D 104 (2021) 103029
M. Boudaud et al., Phys. Rev. Research 2, 023022 (2020)
V. Bresci *et al.*, Mon. Not. R. Astron. Soc., 488 (2019), p. 2068
M. Korsmeier *et al.*, Phys. Rev. D 97 (2018), 103019
P. Lipari, Phys. Rev. D, 95 (2017), 063009
I. Cholis *et al.*, Phys. Rev. D 95(2017), 123007
M. Winkler, JCAP, 2017(02), 048

J. Heisig, Modern Physics Letters A, (2021), 36, 05 Y. Genolini *et al.*, arXiv:2103.04108 (2021) I. Cholis *et al.*, Phys. Rev. D, 99 (2019), 103026 A. Cuoco *et al.*, Phys. Rev. D, 99 (2019), 103014 M. Carena *et al.*, Phys. Rev. D, 100 (2019), 055002 A. Reinert *et al.*, JCAP, 01 (2018), p. 055 A. Cuoco *et al.*, Phys. Rev. Lett., 118 (2017), 191102 M. Cui *et al.*, Phys. Rev. Lett., 118 (2017), 191101 Y. Chen *et al.*, Phys. Rev. D, 93 (2016), p. 015015 P. Zuccon - UNITY & INFN HEPA

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Antiproton production and propagation

Antiprotons from Dark Matter

ΦÃ³ [m⁻²sr¹s⁻¹GV²] 1GV² ⊅Ř³ [m⁻² AMS AMS • when the MNRAS 488,2,2019 JCAP 01, 055 (2018) p from collision of cosmic rays p from collision of cosmic rays IRigidityl [GV] Rigidityl [GV] 10 10^{2} 10 10² 0.3×10⁻³ ×10⁻³ p/p flux ratio p/p flux ratio AMS AMS • 0.2 0.15 0.15 0. - Phys. Rev. D 99, 103026 (2019) Phys. Rev. Lett. 118, 191102 (2017) 0.05 0.05

IRigidityl AMPP 2023

p from collision of cosmic rays

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10²

Example: AMS Antiproton Results compared with Cosmic Ray Models Based on AMS Data

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Rigidityl [GV]

p from collision of cosmic rays

10²

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The AMBER experiment at CERN



Apparatus for Meson and Baryon Experimental Research

- The AMBER experiment (NA66) is a fixed target experiment at CERN SPS.
- It has been approved for three measurements
 - pbar production cross section in p-He interaction
 - Proton radius measurement though muonhydrogen scattering
 - Drell Yan processes to study form factors and emergence of the hadron mass
- A proposal for a AMBER phase2 to be done after the LHC long shutdown is in preparation



(the AMBER/COMPASS spectrometer)

Existing anti-p cross section

- anti-p production cross section from p-p and p-He interactions is poorly measured and cannot simply constrained from available measurements.
- an accurate prediction of the expected anti-p flux in cosmic rays in the rigidity range from few GeV to several hundreds of GeVs, is interesting to understand cosmic ray and possibly search for signals of new physics
- LHC-b collaboration reported a measurement the anti-p XS from 8 TeV p-He, and foresee a similar measurement with 4TeV protons.
- NA61 published p-p to anti-p at 20, 31, 40, 80, and 158 GeV/c
- we want to investigate the possibility to perform a measurement with the SPS protons between 50 and 280 GeV/c on fixed LH2 and LHe targets, and a magnetic spectrometer

Fraction origin of anti-p from CR interaction with ISM



LHCb-CONF-2017-002 Measurement performed at 7 TeV p-He -> pbar + X

NA61 p+p data beam momenta of 20, 31, 40, 80, and 158 GeV/c Eur. Phys. J. C 77, 671 (2017)



Experimental Research

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Measuring pbar cross section with AMBER

- Target: Liquid ⁴He cylinder 1.2 m long for 10 cm of diameter
- Precise beam telescope measuring the incoming particle trajectory
- Two stages spectrometer (50m long) with typical momentum resolution ranging from 1% to 0.1%
- Large gas RICH detector for Particle identification
- Use secondary hadron beams from primary SPS proton on a Be target.
- Explore 6 proton momenta: 60, 80, 100, 160, 190, 250 GeV/c
- Use 2 long CEDARs (threshold Cerenkov) to identify protons in the hadron beam

Non Tracking detectors

- Two sets of ECAL and HCAL
- One of the largest RICH detectors ever operated.









AMBER -2023 Physics run

- April 24th: Detector and beam commissioning starts
- May 20th: Data taking starts with p at 190 GeV/c
- May 25th : switch for the long run with p at 60 GeV/c
- June 6th : Switch to p at 100 GeV/c
- June 11th : switch to 250 GeV/c
- June 14^{th} : switch to 160 GeV/c
- June 18th : switch to 80 GeV/c
- June 25th : Last day with liquid He target
- June 26⁻28th: Empty target runs
- June 29th-30th: Drell Yan high intensity test
- July 1st : CEDAR PID efficiency runs

NOTE: SM1 and SM2 currents have been scaled such that the primary beam at different momenta follows the same path within the spectrometer as the 190 GeV one

We added to additional energy points (80 and 160 GeV), because p-p measurements at those energies are available and it is interesting to compare

RICH Measurement example

 θ_{CH} vs p - negative 0.06 θ_{CH} [rad] 10⁴ 0.05 10³ 0.04 hthetach n Entries 1520819 0.03 8.136 Mean x 10² Mean y 0.04113 Std Dev x 8.996 Std Dev y 0.01766 0.02 10 0.01 50 10 15 20 25 30 45 5 35 40 p [GeV/c]

Proton identification in the Secondary hadron beam

CEDARs

CErenkov Differential counters with Achromatic Ring focus





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CEDAR Selection p 190 GeV/c

CEDAR1 >= 6 pmts CEDAR2 >= 7 pmts



Effect of the CEDAR selection cutt

Before N pmt cut

After N pmt cut



Cross section measurement

- Strategy
 - Count all the p-p (or p-He) interaction in the target (R_i)
 - Identify events with one (or multiple) anti-p vs reconstructed momentum and angle $(R_s (p, \theta))$
 - Calculate the double differential cross section as

$$\frac{d\sigma_{\overline{p}}}{dp \ d\theta} = \frac{R_s(p,\theta)}{R_i}$$



Explored phase space for anti-p production

Antiproton tracking efficiency

Pseudo-rapidity vs log₁₀(momentum)

Particle Identification Threshold Ē 11.5 _11⊦ New udorepidity 0.9 10.5 10 **RICH** 0.8 9.5 9 Acc 0.7 8.5 ۵Ľ 7.5 0.6 6.5 0.5 5.5 0.4 **RICH** 0.3 Acc 0.2 2.5 0.1 1.5 1 = _ _ _ -1 -0.8 -0.6 -0.4 -0.2 0 0.2 0.4 0.6 0.8 1 1.2 1.4 1.6 1.8 2 2.2 2.4 Momentum (log₁₀ GeV)

Removal of RICH BeamPipe



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p-He He-p Source term coverage



Expected cross section measurement p-He -> bar-p + X



p 190 GeV/c

Anagrafica

		2023	2024	
Battiston Roberto	РО	0.15	0.15	
Dass Abhinandan	Dottorando	0.5	0.5	
Bisht Ashish	Dottorando	0.15	0	
Gebbia Giuseppe	Assegnista	0.15	0	
Rossi Francesco	Dottorando	0	0.5	
Nicolaidis Riccardo	Dottorando	0.35	0.15	
Nozzoli Francesco	Ricercatore	0.35	0.35	
Zuccon Paolo	PA	0.35	0.35	
TOTALE		2	2	