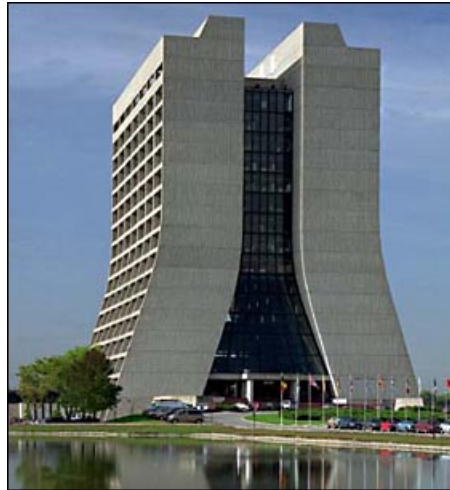
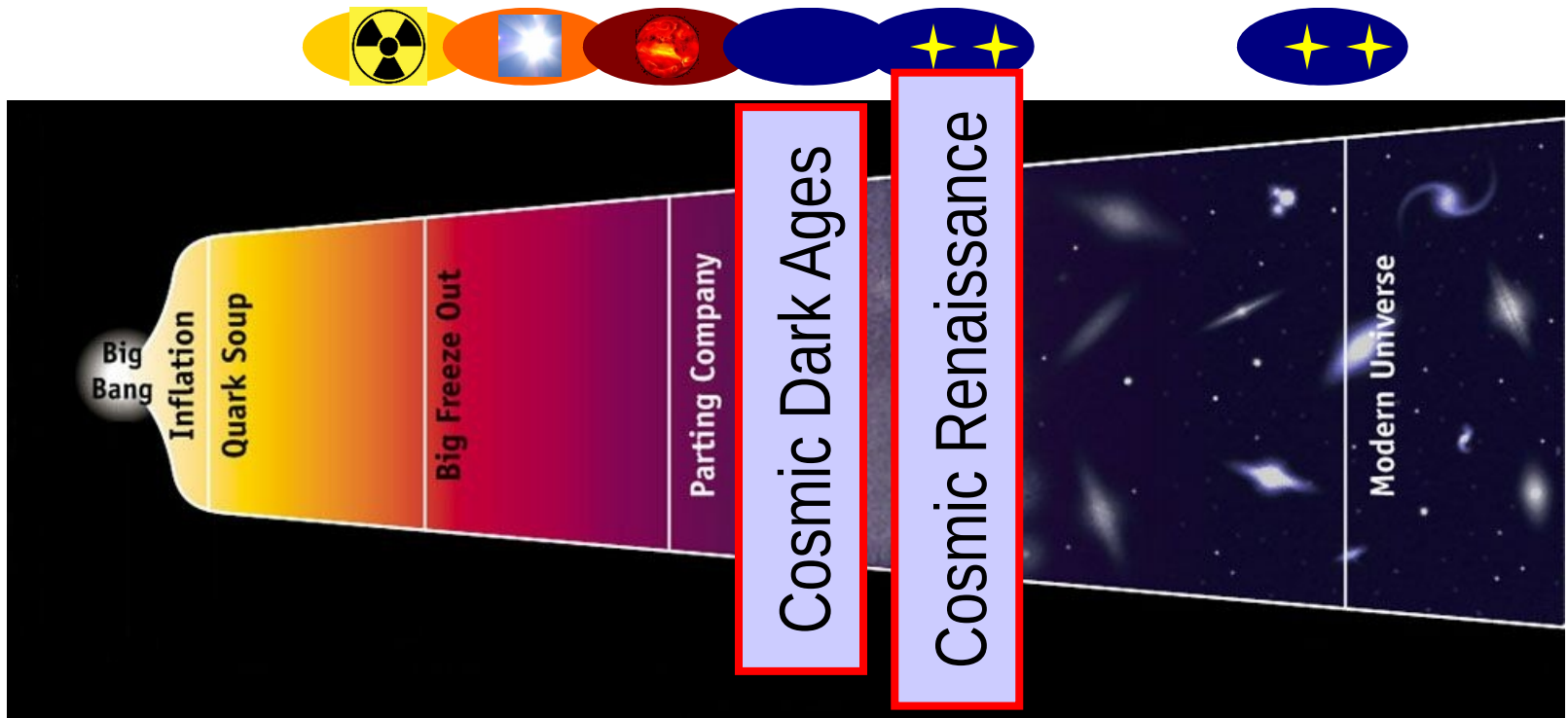


The End of the Dark Ages

Nick Gnedin



The Brief History of Time



Time →

← Redshift z

← Temperature

What You See Is Not What You Get

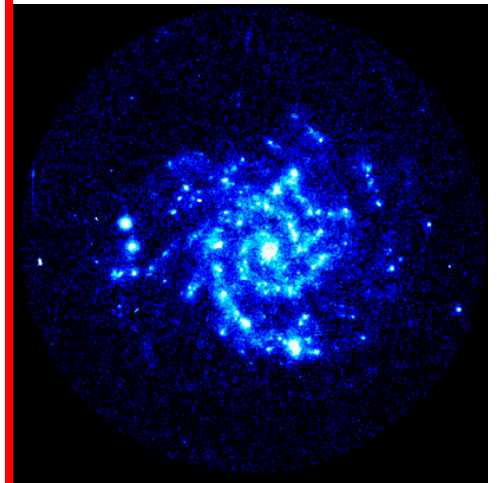
Infra-Red (IR)



Visible



Ultra-Violet (UV)

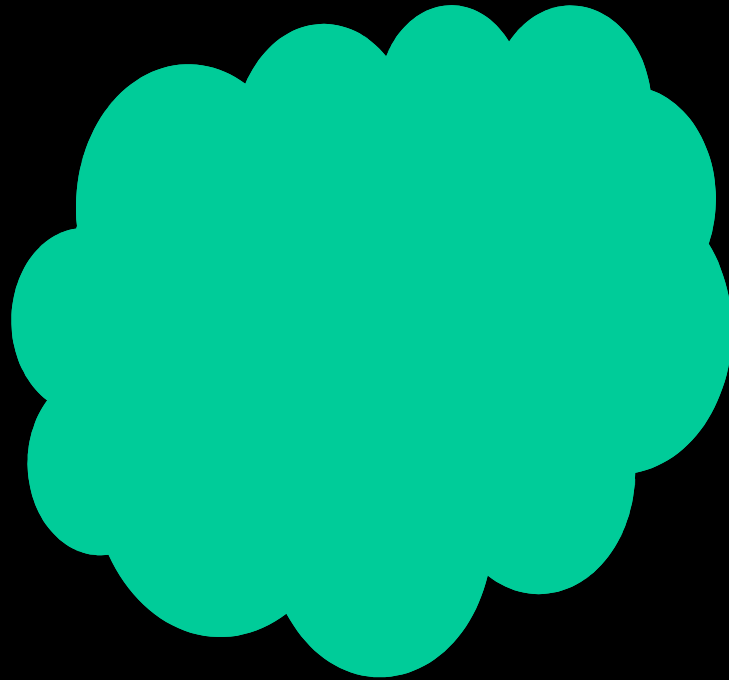


Now

High z

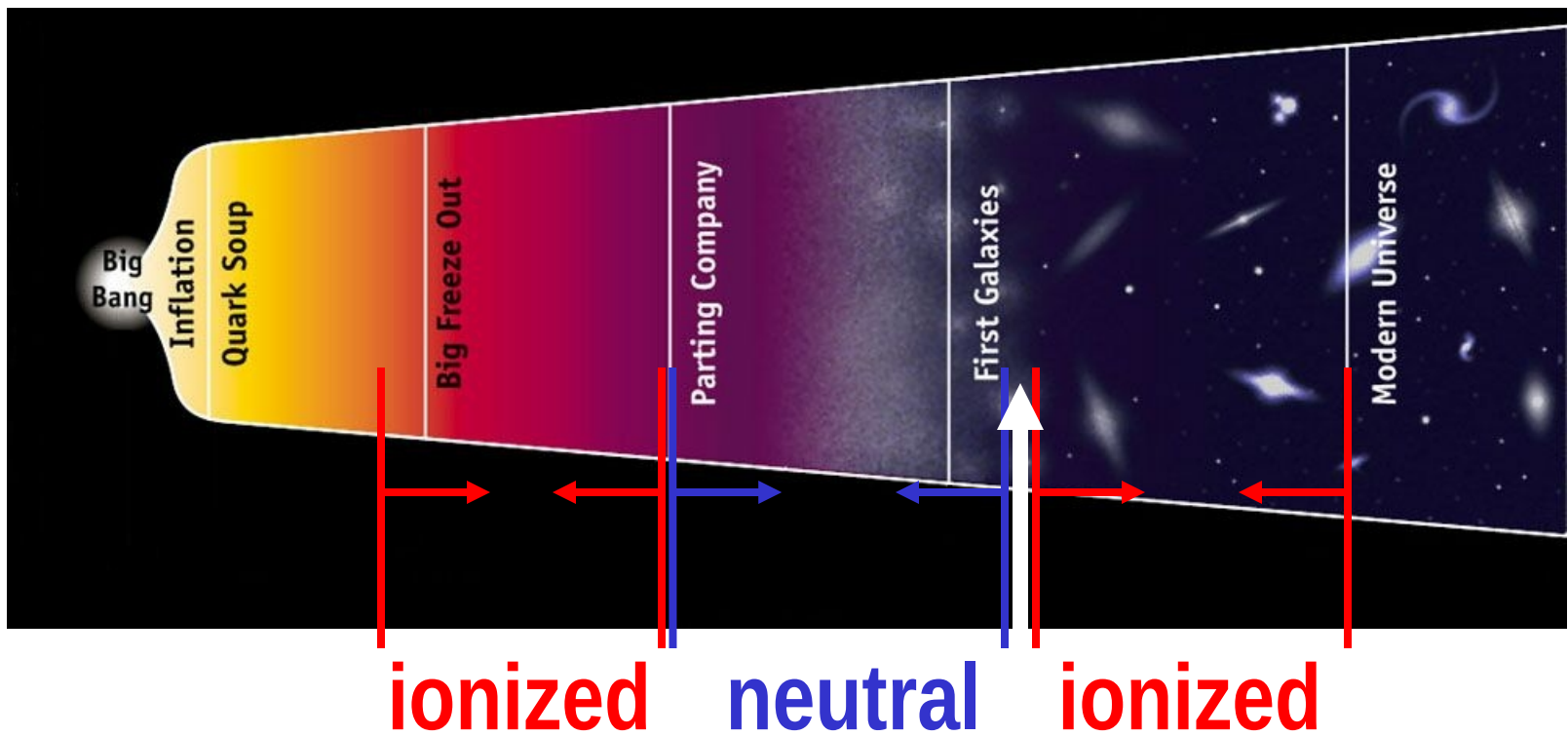
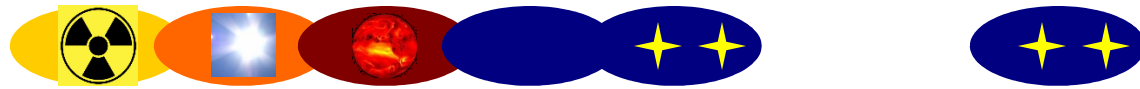


Transparent & Opaque



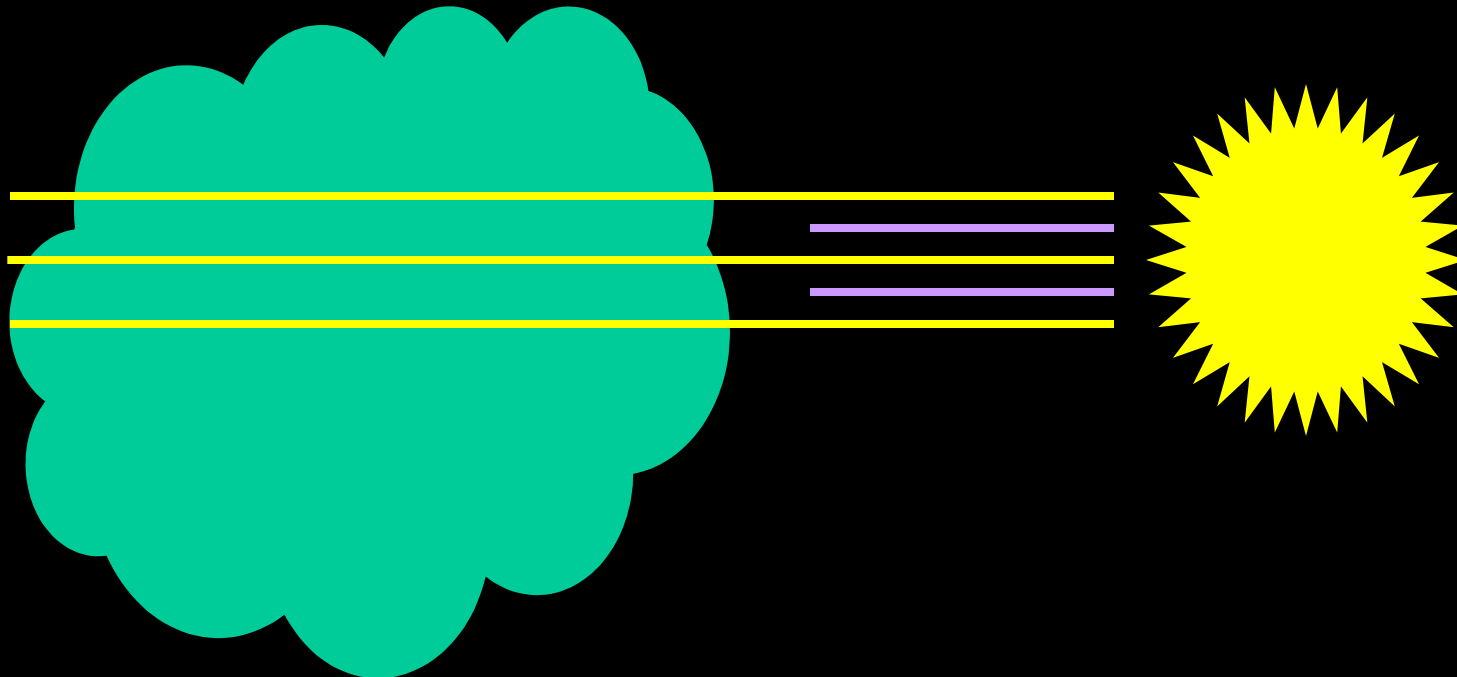
Hydrogen

The Brief History of Hydrogen



RE-IONIZATION

The End of the “Dark Ages”



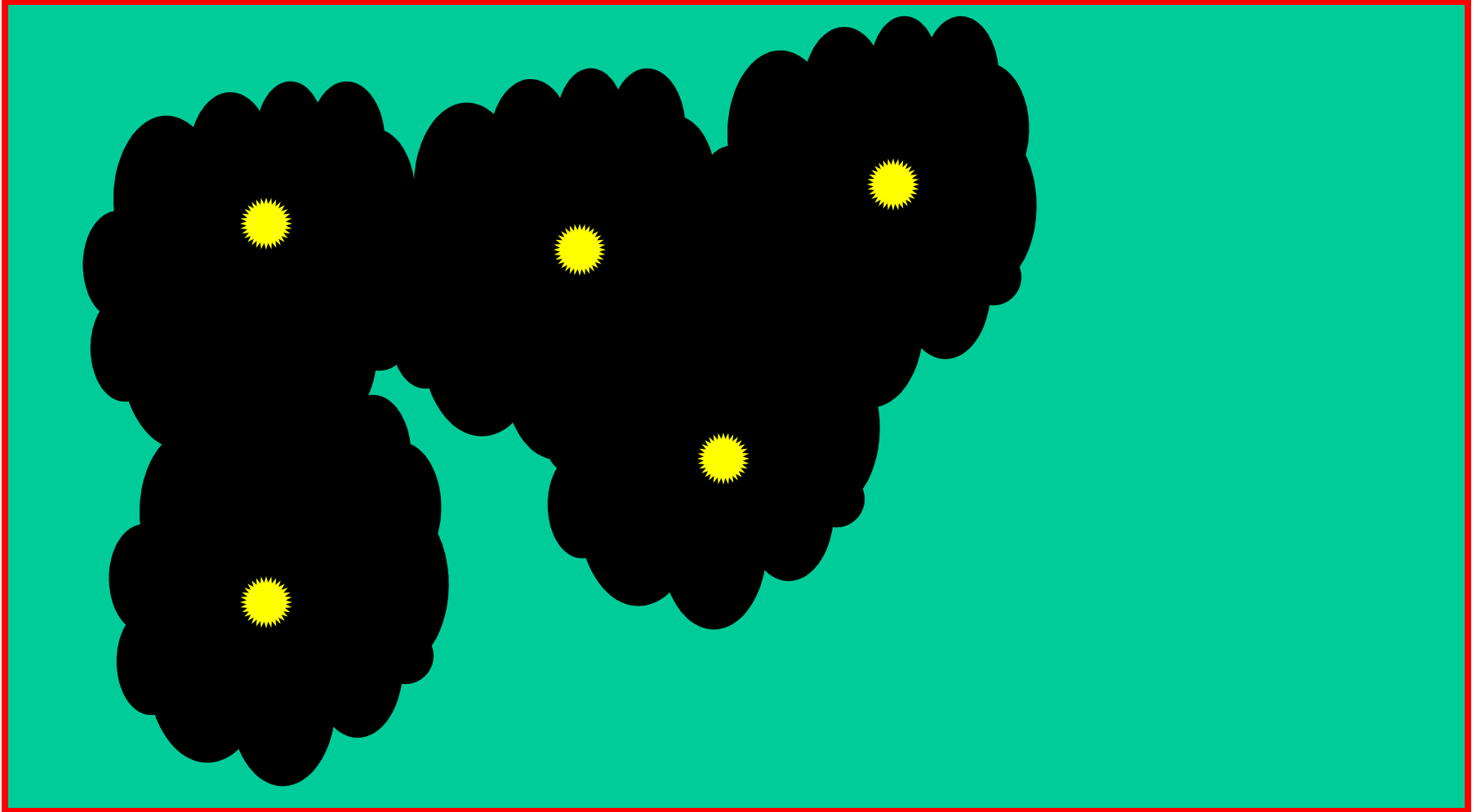
Hydrogen

**The End of the “Dark
Ages”**

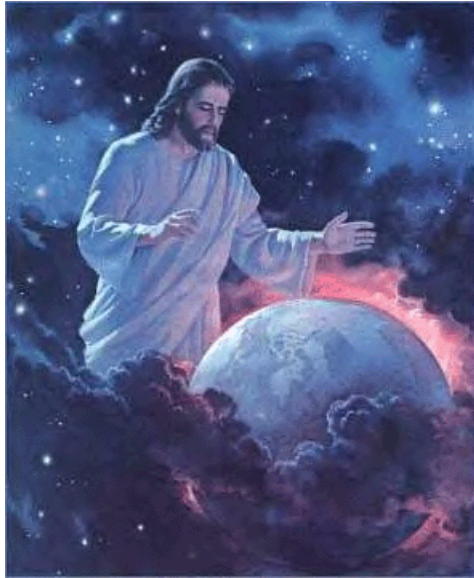
IS

Re-ionization

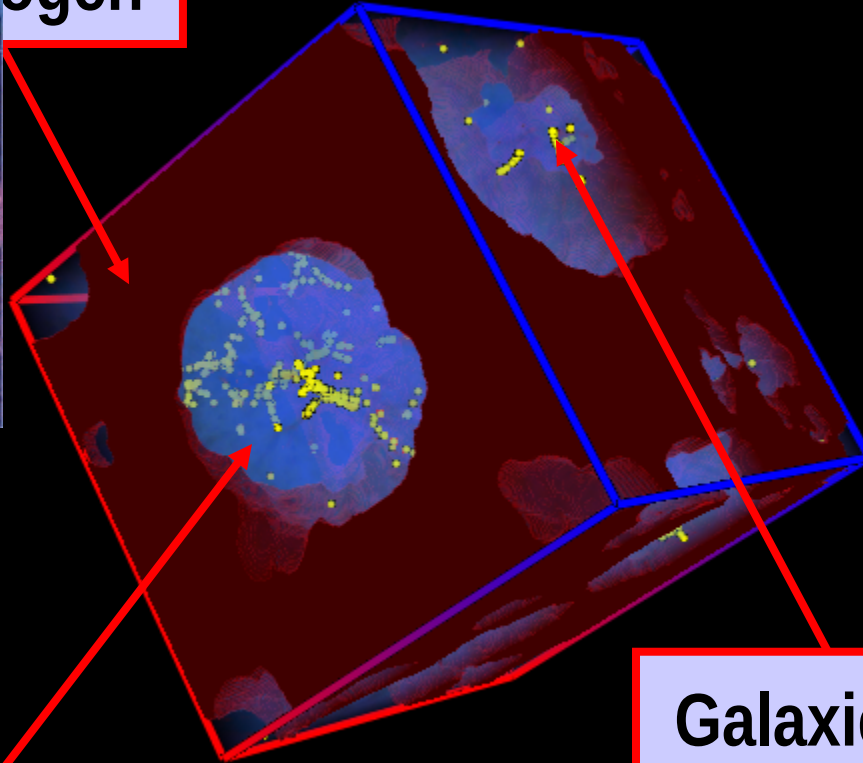
Bubbly Story



Let Us Be Gods...

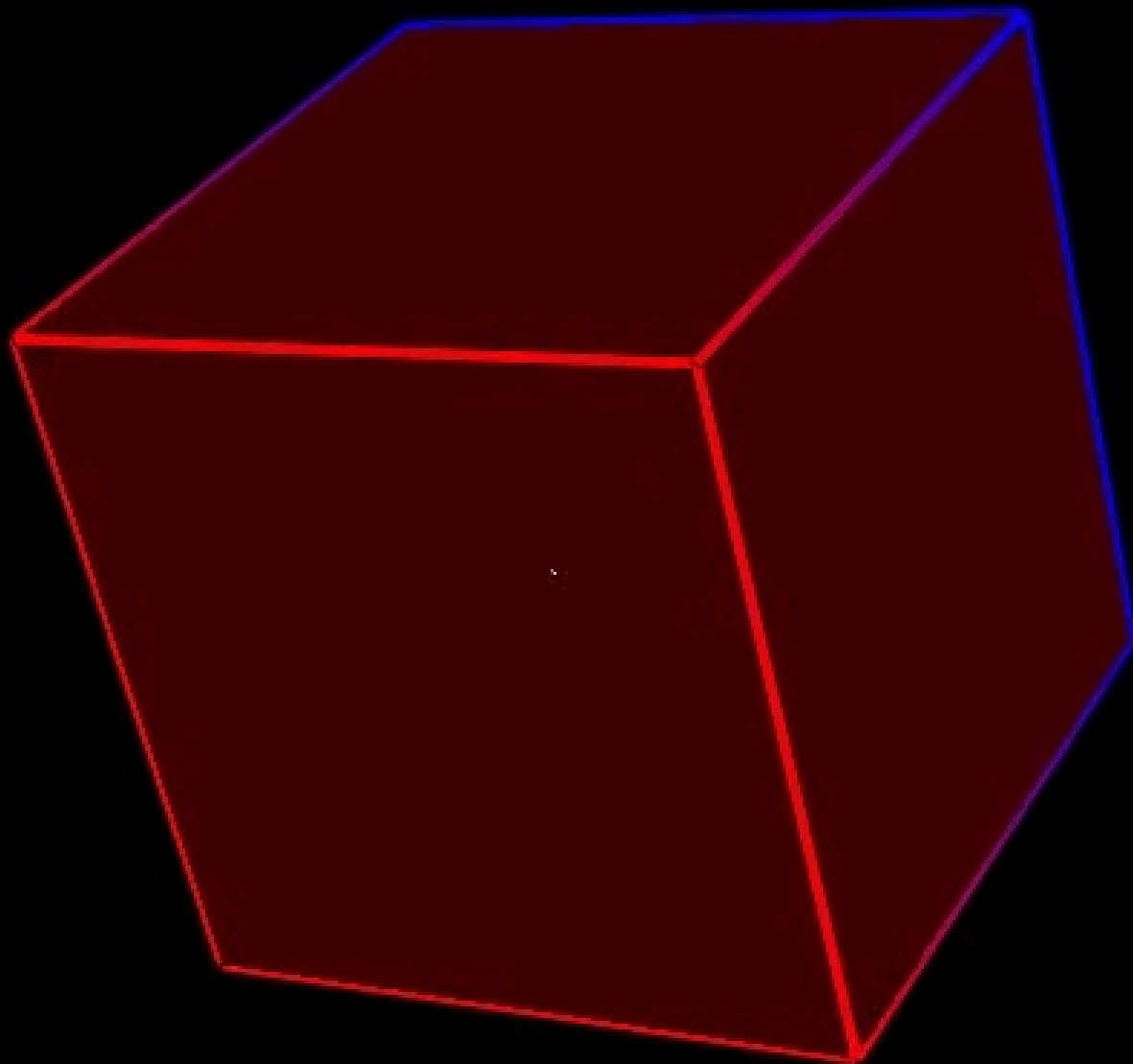


Hydrogen

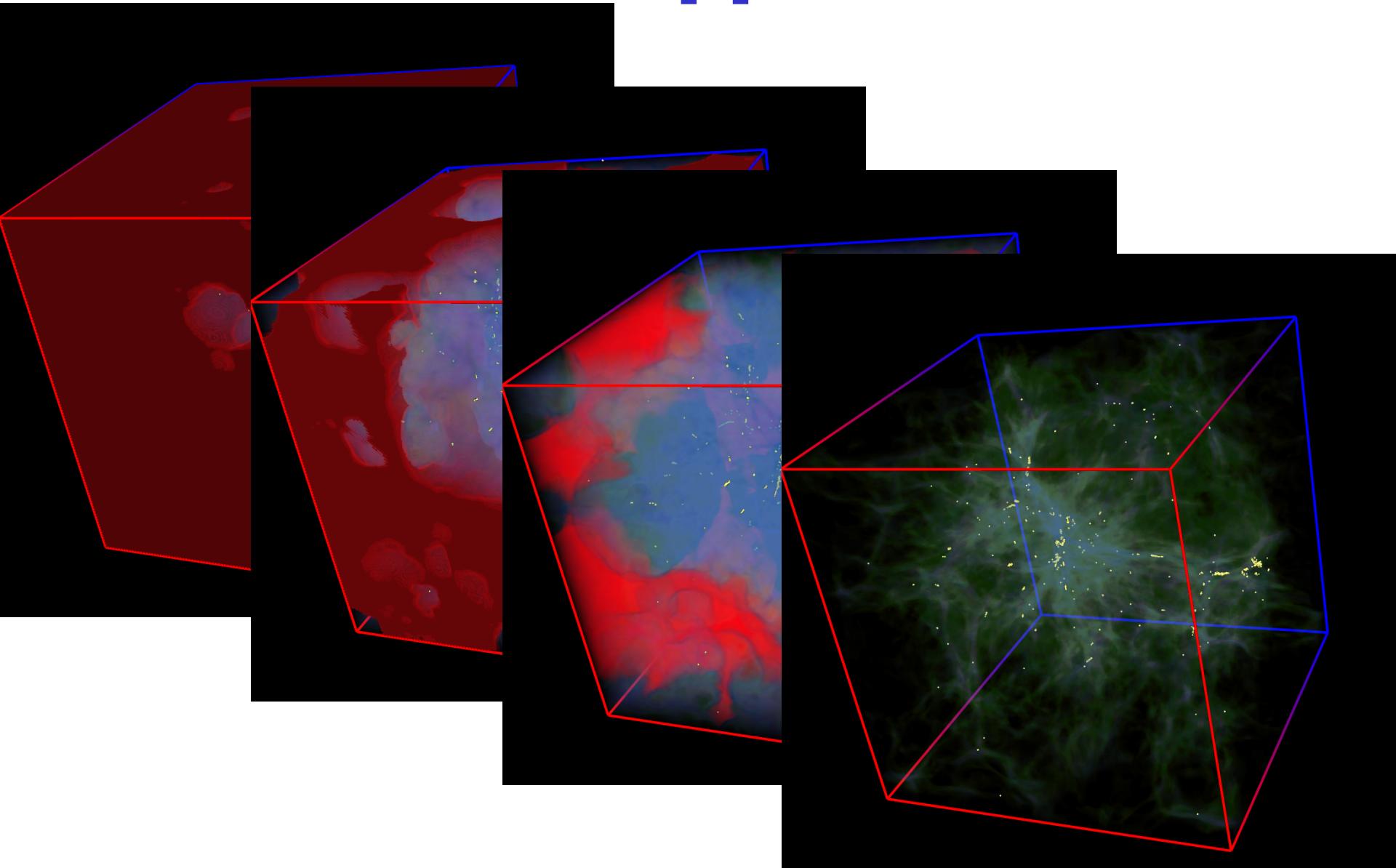


Galaxies

Ionized bubbles



How It All Happens...

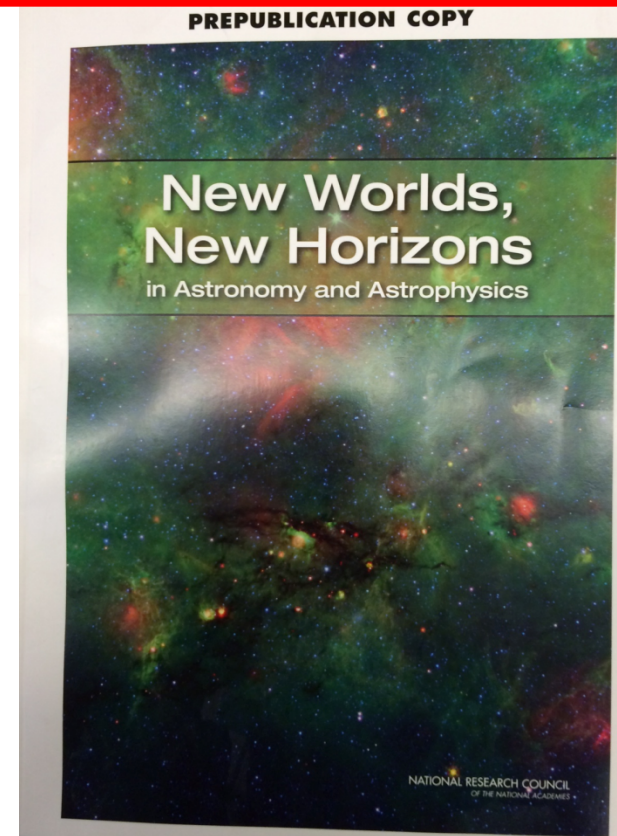


Not Just Fun: Priority Science

- Astro2010:

The priority science objectives chosen by the survey committee for the decade 2012-2021 are searching for the first stars, galaxies, and black holes;

- Also identified as a likely discovery area.



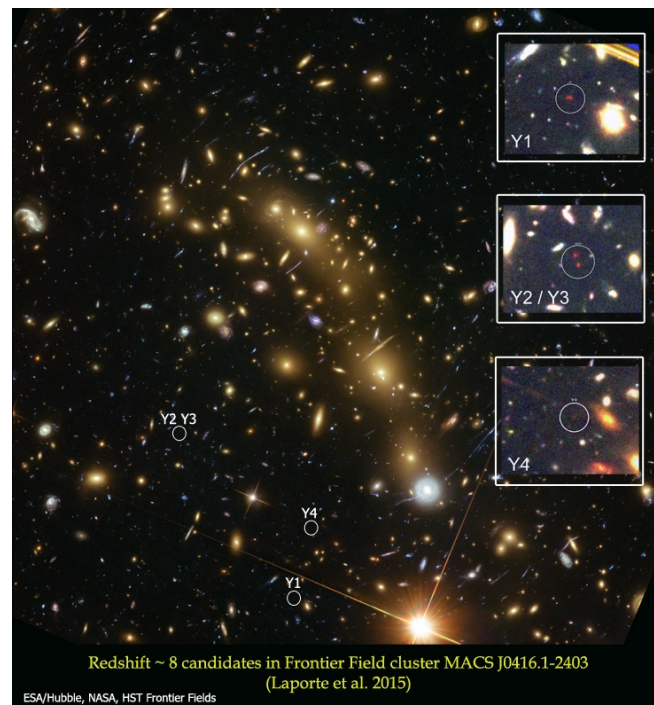
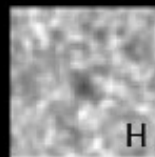
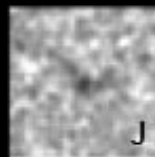
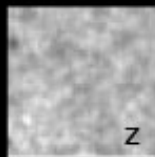
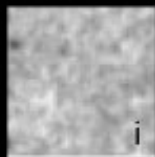
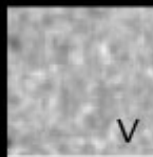
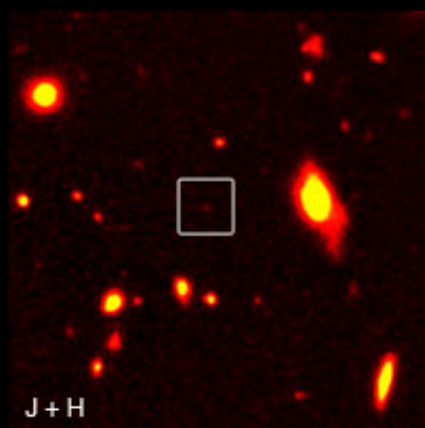
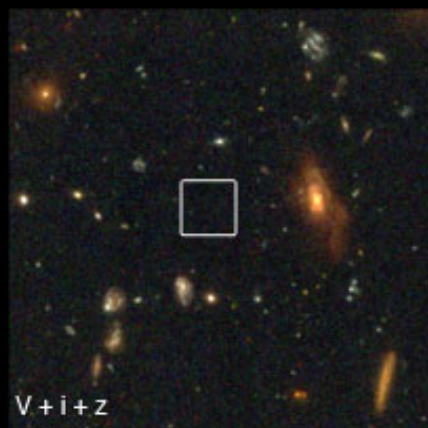
The Flood Is Coming

- Atacama Large Millimeter Array (ALMA) 2017+
- James Webb Space Telescope (JWST) 2021?
- Hydrogen Epoch of Reionization Array (HERA) 2021-22?
- Next generation optical telescopes 2021-25? (GMT, TMT, E-ELT)



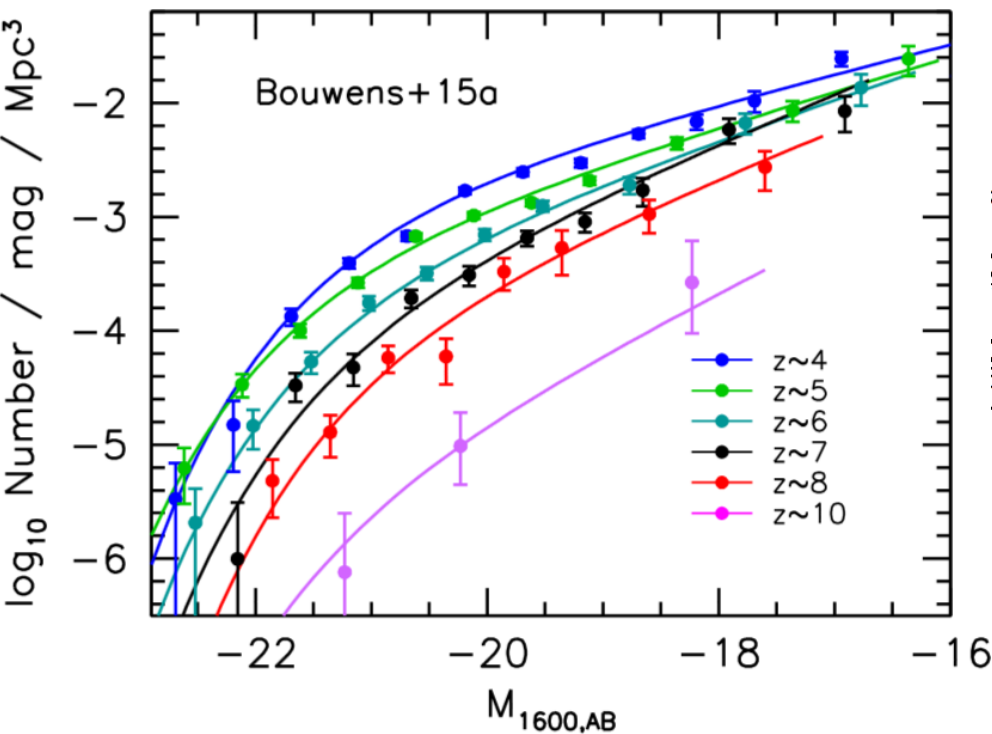
What We Know Now: Galaxy Luminosity Functions

- High redshift galaxies are not particularly glaring.
- The faintest ones are detected with the help of gravitational lensing in HST *Frontier Fields*.

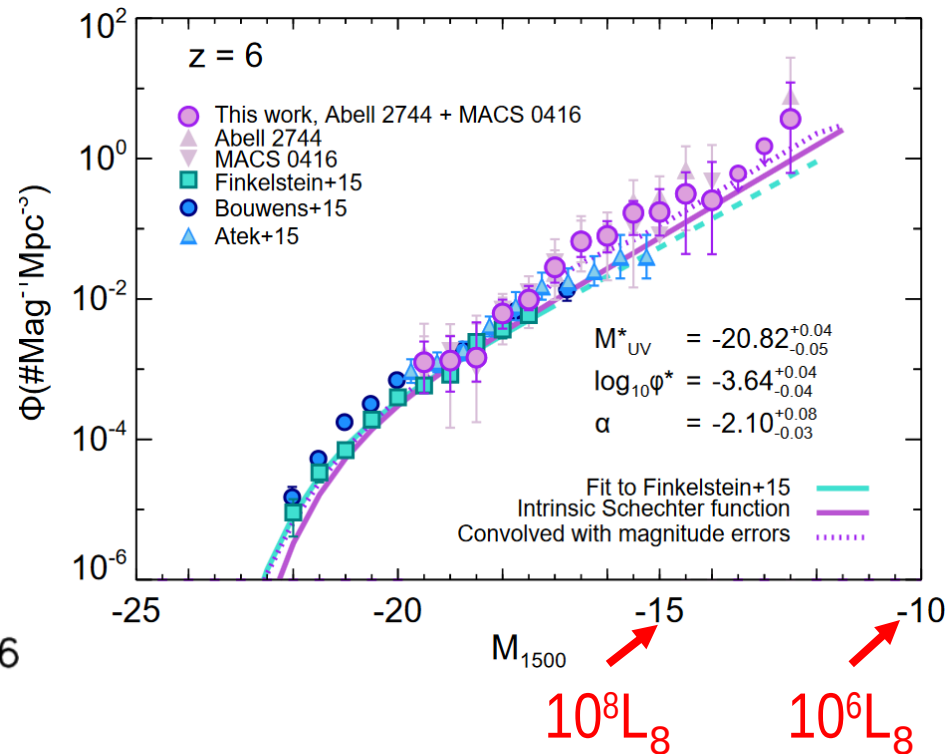


What We Know Now: Galaxy Luminosity Functions

- Data (may) exist all the way to $z=10$ and $M_{UV}=-13$.



Bouwens+2015



Livermore+ 2016

About “Magnitudes”

- Astronomers use magnitudes to maintain self-esteem. Without them astrophysics is too easy.

$$\begin{aligned}
 1 \quad L = & -\frac{1}{2}\partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4}g_s^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e + \\
 & \frac{1}{2}ig_s^2 (\bar{q}_i^\sigma \gamma^\mu q_j^\sigma) g_\mu^a + \bar{G}^a \partial^2 G^a + g_s f^{abc} \partial_\mu \bar{G}^a G^b g_\mu^c - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- - \\
 2 \quad M^2 W_\mu^+ W_\mu^- & - \frac{1}{2}\partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2} M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - \frac{1}{2}\partial_\mu H \partial_\mu H - \\
 & \frac{1}{2}m_h^2 H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - M^2 \phi^+ \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \frac{1}{2c_w^2} M \phi^0 \phi^0 - \beta_h \left[\frac{2M^2}{g^2} + \right. \\
 & \left. \frac{2M}{g} H + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-) \right] + \frac{2M^4}{g^2} \alpha_h - igc_w [\partial_\nu Z_\mu^0 (W_\mu^+ W_\nu^- - \\
 & W_\nu^+ W_\mu^-) - Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\mu^0 (W_\nu^+ \partial_\nu W_\mu^- - \\
 & W_\nu^- \partial_\nu W_\mu^+)] - ig_s w [\partial_\nu A_\mu (W_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - \\
 & W_\nu^- \partial_\nu W_\mu^+) + A_\mu (W_\nu^+ \partial_\nu W_\mu^- - W_\nu^- \partial_\nu W_\mu^+)] - \frac{1}{2}g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \\
 & \frac{1}{2}g^2 W_\mu^+ W_\nu^- W_\mu^- W_\nu^+ + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - Z_\mu^0 Z_\nu^0 W_\mu^+ W_\nu^-) + \\
 & g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w [A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\
 & W_\nu^+ W_\mu^-) - 2A_\mu Z_\mu^0 W_\nu^+ W_\nu^-] - g\alpha [H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-] - \\
 & \frac{1}{8}g^2 \alpha_h [H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2] - \\
 & gM W_\mu^+ W_\mu^- H - \frac{1}{2}g \frac{M}{c_w^2} Z_\mu^0 Z_\mu^0 H - \frac{1}{2}ig [W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - \\
 & W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)] + \frac{1}{2}g [W_\mu^+ (H \partial_\mu \phi^- - \phi^- \partial_\mu H) - W_\mu^- (H \partial_\mu \phi^+ - \\
 & \phi^+ \partial_\mu H)] + \frac{1}{2}g \frac{1}{c_w} (Z_\mu^0 (H \partial_\mu \phi^0 - \phi^0 \partial_\mu H) - ig \frac{s_w^2}{c_w} M Z_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \\
 & ig_s w M A_\mu (W_\mu^+ \phi^- - W_\mu^- \phi^+) - ig \frac{1-2c_w^2}{2c_w} Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + \\
 & ig_s w A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \frac{1}{4}g^2 W_\mu^+ W_\mu^- [H^2 + (\phi^0)^2 + 2\phi^+ \phi^-] - \\
 & \frac{1}{4}g^2 \frac{1}{c_w^2} Z_\mu^0 Z_\mu^0 [H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2 \phi^+ \phi^-] - \frac{1}{2}g^2 \frac{s_w^2}{c_w} Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + \\
 & W_\mu^- \phi^+) - \frac{1}{2}ig^2 \frac{s_w^2}{c_w} Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2}g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\
 & W_\mu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{s_w}{c_w} (2c_w^2 - 1) Z_\mu^0 A_\mu \phi^+ \phi^- - \\
 & g^1 s_w^2 A_\mu A_\mu \phi^+ \phi^- - \bar{e}^\lambda (\gamma \partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda \gamma \partial \nu^\lambda - \bar{u}_j^\lambda (\gamma \partial + m_u^\lambda) u_j^\lambda -
 \end{aligned}$$

Particle Physics

$$\begin{aligned}
 3 \quad g^1 s_w^2 A_\mu A_\mu \phi^+ \phi^- & - \bar{e}^\lambda (\gamma \partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda \gamma \partial \nu^\lambda - \bar{u}_j^\lambda (\gamma \partial + m_u^\lambda) u_j^\lambda - \\
 & \bar{d}_j^\lambda (\gamma \partial + m_d^\lambda) d_j^\lambda + ig_s w A_\mu [-\bar{e}^\lambda \gamma^\mu e^\lambda + \frac{2}{3}(\bar{u}_j^\lambda \gamma^\mu u_j^\lambda) - \frac{1}{3}(\bar{d}_j^\lambda \gamma^\mu d_j^\lambda)] + \\
 & \frac{ig}{4c_w} Z_\mu^0 [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{u}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - \\
 & 1 - \gamma^5) u_j^\lambda) + (\bar{d}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 - \gamma^5) d_j^\lambda)] + \frac{ig}{2\sqrt{2}} W_\mu^+ [(\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) e^\lambda) + \\
 & (\bar{u}_j^\lambda \gamma^\mu (1 + \gamma^5) C_{\lambda\kappa} d_j^\kappa)] + \frac{ig}{2\sqrt{2}} W_\mu^- [(\bar{e}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{d}_j^\kappa C_{\lambda\kappa}^\dagger \gamma^\mu (1 + \\
 & \gamma^5) u_j^\lambda)] + \frac{ig}{2\sqrt{2}} \frac{m_e^\lambda}{M} [-\phi^+ (\bar{\nu}^\lambda (1 - \gamma^5) e^\lambda) + \phi^- (\bar{e}^\lambda (1 + \gamma^5) \nu^\lambda)] - \\
 4 \quad \frac{g}{2} \frac{m_e^\lambda}{M} [H (\bar{e}^\lambda e^\lambda) + i\phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda)] & + \frac{ig}{2M\sqrt{2}} \phi^+ [-m_d^\kappa (\bar{u}_j^\lambda C_{\lambda\kappa} (1 - \gamma^5) d_j^\kappa) + \\
 & m_u^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 + \gamma^5) d_j^\kappa) + \frac{ig}{2M\sqrt{2}} \phi^- [m_d^\lambda (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 + \gamma^5) u_j^\kappa) - m_u^\lambda (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 - \\
 & \gamma^5) u_j^\kappa) - \frac{g}{2} \frac{m_u^\lambda}{M} H (\bar{u}_j^\lambda u_j^\lambda) - \frac{g}{2} \frac{m_d^\lambda}{M} H (\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2} \frac{m_u^\lambda}{M} \phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \\
 & \frac{ig}{2} \frac{m_d^\lambda}{M} \phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \\
 5 \quad \frac{M^2}{2c_w} X^0 + \bar{Y} \partial^2 Y & + igc_w W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \partial_\mu \bar{X}^+ X^0) + ig_s w W_\mu^+ (\partial_\mu \bar{Y} X^- - \\
 & \partial_\mu \bar{X}^+ Y) + igc_w W_\mu^- (\partial_\mu \bar{X}^- X^0 - \partial_\mu \bar{X}^0 X^+) + ig_s w W_\mu^- (\partial_\mu \bar{X}^- Y - \\
 & \partial_\mu \bar{Y} X^+) + igc_w Z_\mu^0 (\partial_\mu \bar{X}^+ X^+ - \partial_\mu \bar{X}^- X^-) + ig_s w A_\mu (\partial_\mu \bar{X}^+ X^+ - \\
 & \partial_\mu \bar{X}^- X^-) - \frac{1}{2}gM [\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w} \bar{X}^0 X^0 H] + \\
 & \frac{1-2c_w^2}{2c_w} igM [\bar{X}^+ X^0 \phi^+ - \bar{X}^- X^0 \phi^-] + \frac{1}{2c_w} igM [\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + \\
 & igM s_w [\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-] + \frac{1}{2}igM [\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0]
 \end{aligned}$$

About “Magnitudes”

- Astronomers use magnitudes to maintain self-esteem. Without them astrophysics is too easy.

General Relativity

$$\begin{aligned}
 & \frac{1}{2}g^{rs} \left(-\frac{\partial^2 g_{ij}}{\partial x^r \partial x^s} + \frac{\partial^2 g_{is}}{\partial x^r \partial x^j} + \frac{\partial^2 g_{rj}}{\partial x^i \partial x^s} - \frac{\partial^2 g_{rs}}{\partial x^i \partial x^j} \right) + \frac{1}{4}g^{qp} \left(-\frac{\partial g_{is}}{\partial x^p} + \frac{\partial g_{pi}}{\partial x^s} + \right. \\
 & \left. \frac{\partial g_{ps}}{\partial x^i} \right) \times \left(\frac{\partial g_{qj}}{\partial x^r} + \frac{\partial g_{qr}}{\partial x^j} - \frac{\partial g_{rj}}{\partial x^q} \right) - \frac{1}{4}g^{qp} \left(-\frac{\partial g_{ij}}{\partial x^p} + \frac{\partial g_{pi}}{\partial x^j} + \frac{\partial g_{pj}}{\partial x^i} \right) \left(\frac{\partial g_{qr}}{\partial x^s} + \right. \\
 & \left. \frac{\partial g_{qs}}{\partial x^r} - \frac{\partial g_{rs}}{\partial x^q} \right) - \frac{1}{4}g_{ij}g^{rs}g^{uv} \left(-\frac{\partial^2 g_{rs}}{\partial x^u \partial x^v} + \frac{\partial^2 g_{rv}}{\partial x^u \partial x^s} + \frac{\partial^2 g_{us}}{\partial x^r \partial x^v} - \frac{\partial^2 g_{uv}}{\partial x^r \partial x^s} \right) + \\
 & \frac{1}{8}g_{ij}g^{rs}g^{uv}g^{qp} \left(\frac{\partial g_{qr}}{\partial x^v} + \frac{\partial g_{qv}}{\partial x^r} - \frac{\partial g_{rv}}{\partial x^q} \right) \left(\frac{\partial g_{ps}}{\partial x^u} + \frac{\partial g_{pu}}{\partial x^s} - \frac{\partial g_{us}}{\partial x^p} \right) - \\
 & \frac{1}{8}g_{ij}g^{rs}g^{uv}g^{qp} \left(\frac{\partial g_{qr}}{\partial x^s} + \frac{\partial g_{qs}}{\partial x^r} - \frac{\partial g_{rs}}{\partial x^q} \right) \left(\frac{\partial g_{pu}}{\partial x^v} + \frac{\partial g_{pv}}{\partial x^u} - \frac{\partial g_{uv}}{\partial x^p} \right) = \frac{8\pi G}{c^4}T_{ij}.
 \end{aligned}$$

About “Magnitudes”

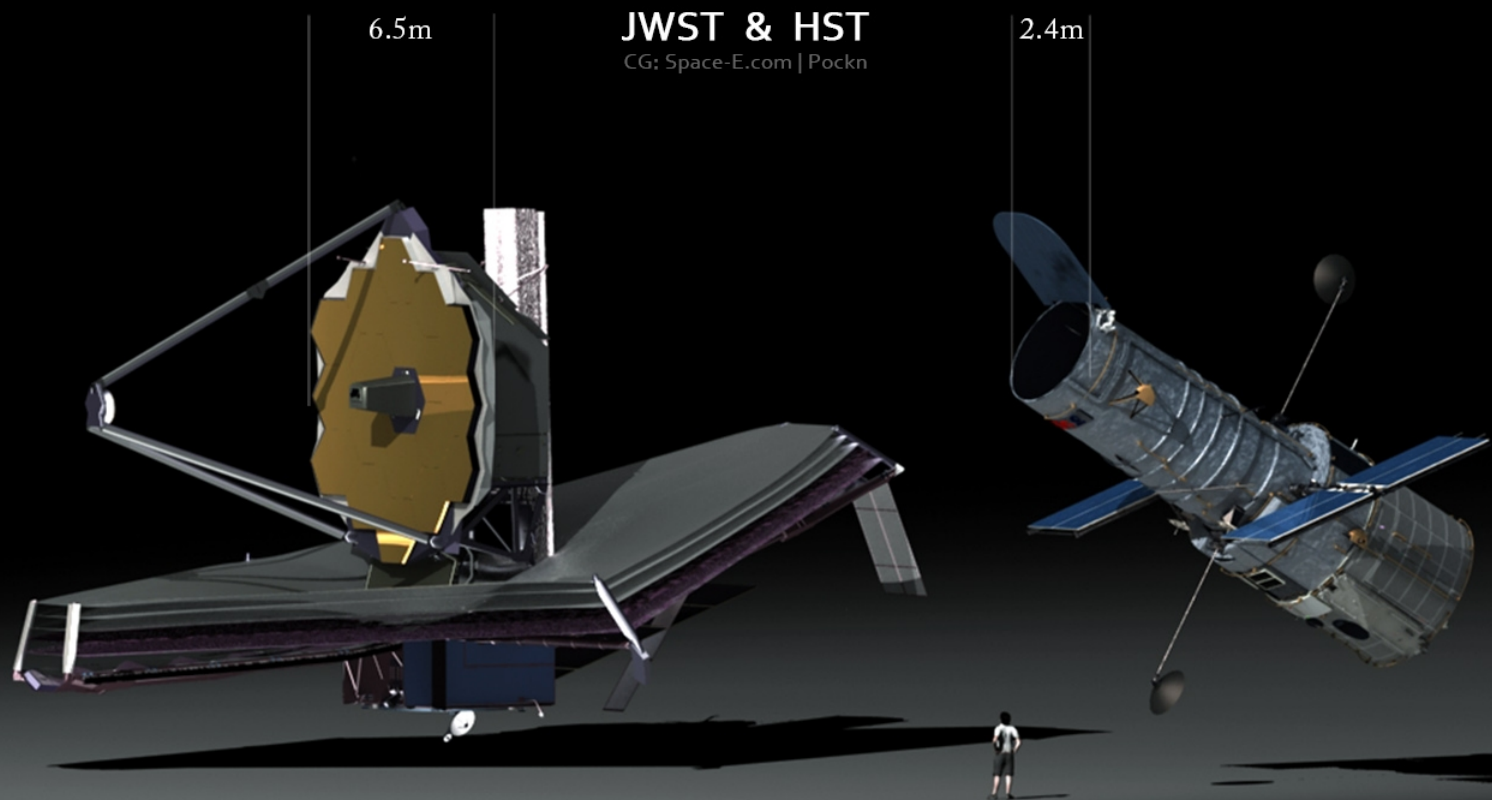
- Astronomers use magnitudes to maintain self-esteem. Without them astrophysics is too easy.

Astrophysics

$$\frac{\partial f}{\partial t} + \frac{\vec{p}}{m} \frac{\partial f}{\partial \vec{x}} + \vec{F} \frac{\partial f}{\partial \vec{p}} = C[f]$$

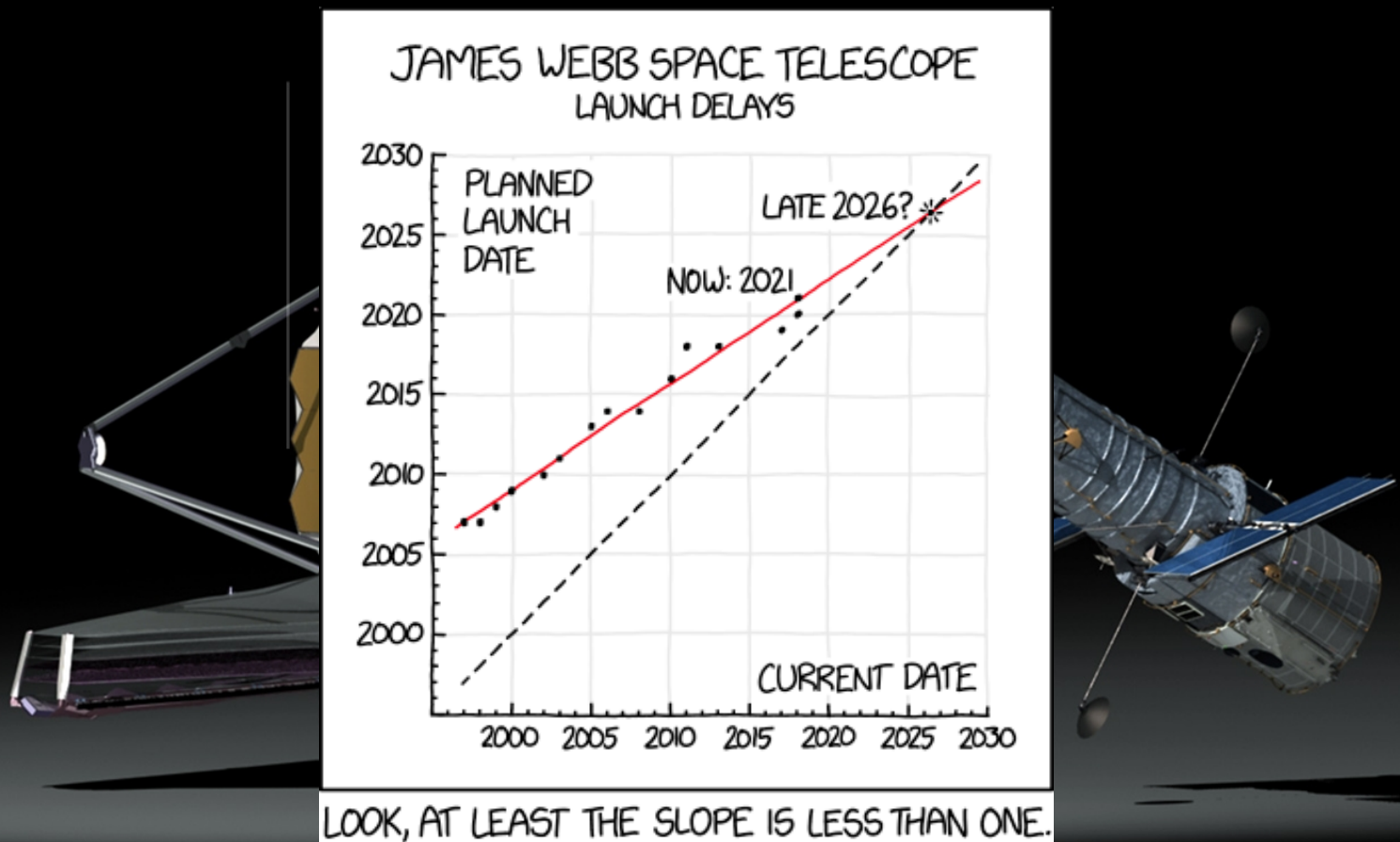
The Flood Is Coming: JWST

- James Webb Space Telescope (JWST) is the primary NASA mission for the next decade.



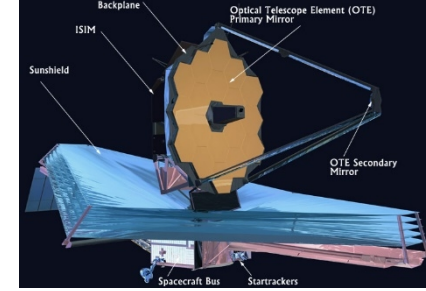
The Flood Is Coming: JWST

- James Webb Space Telescope (JWST) is the primary NASA mission for the next decade.

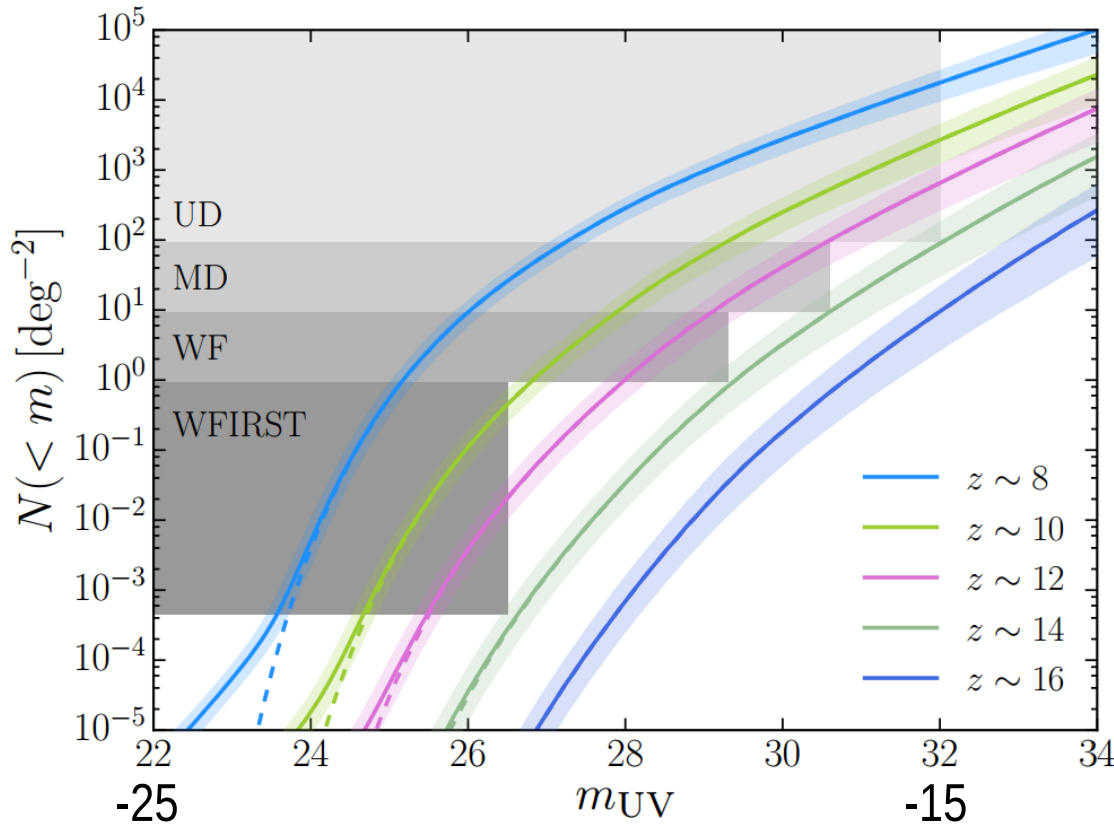


LOOK, AT LEAST THE SLOPE IS LESS THAN ONE.

The Flood Is Coming: JWST



- Studying reionization sources is the primary science goal of JWST.



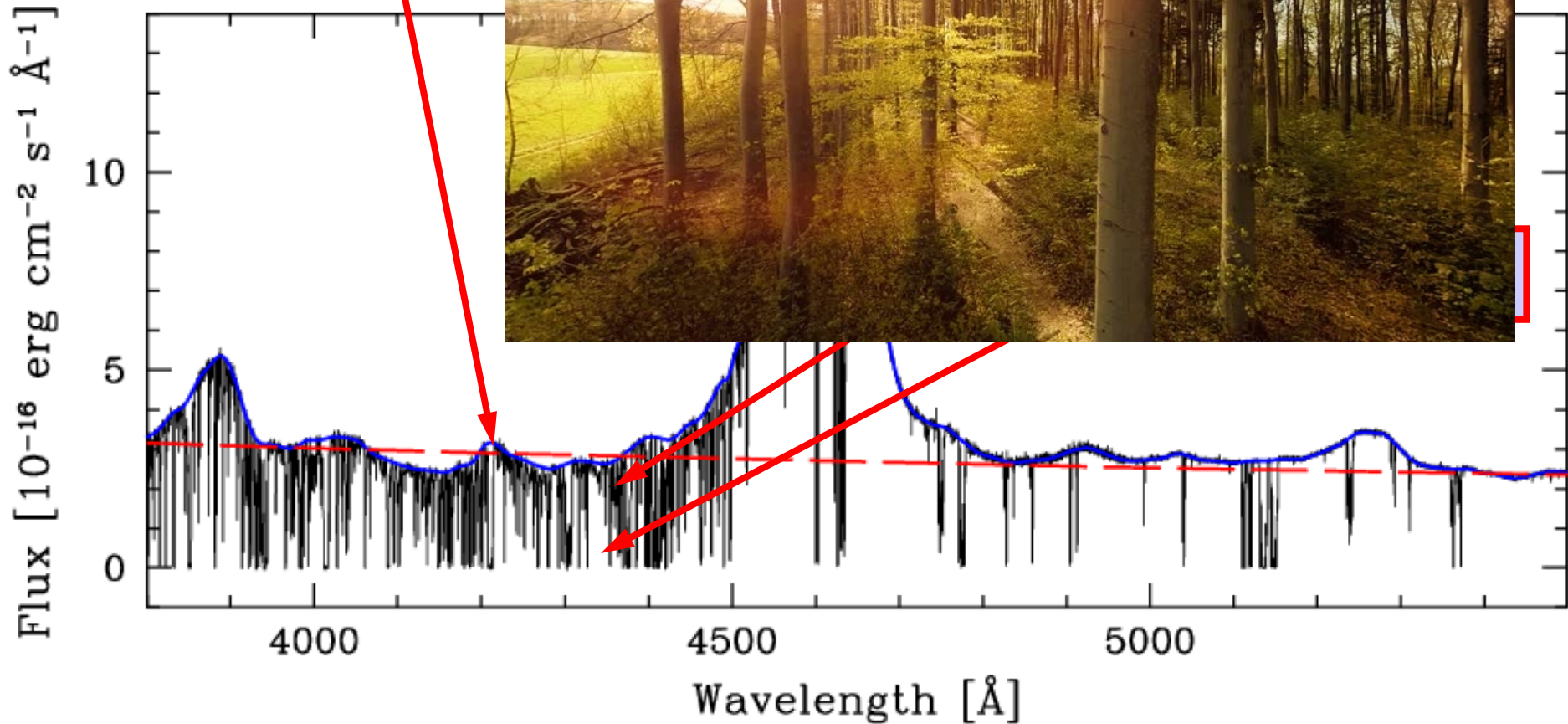
- It will perform a range of surveys of varying depth, greatly improving precision of luminosity functions, as well as studying individual galaxies.

Mason+ 2015

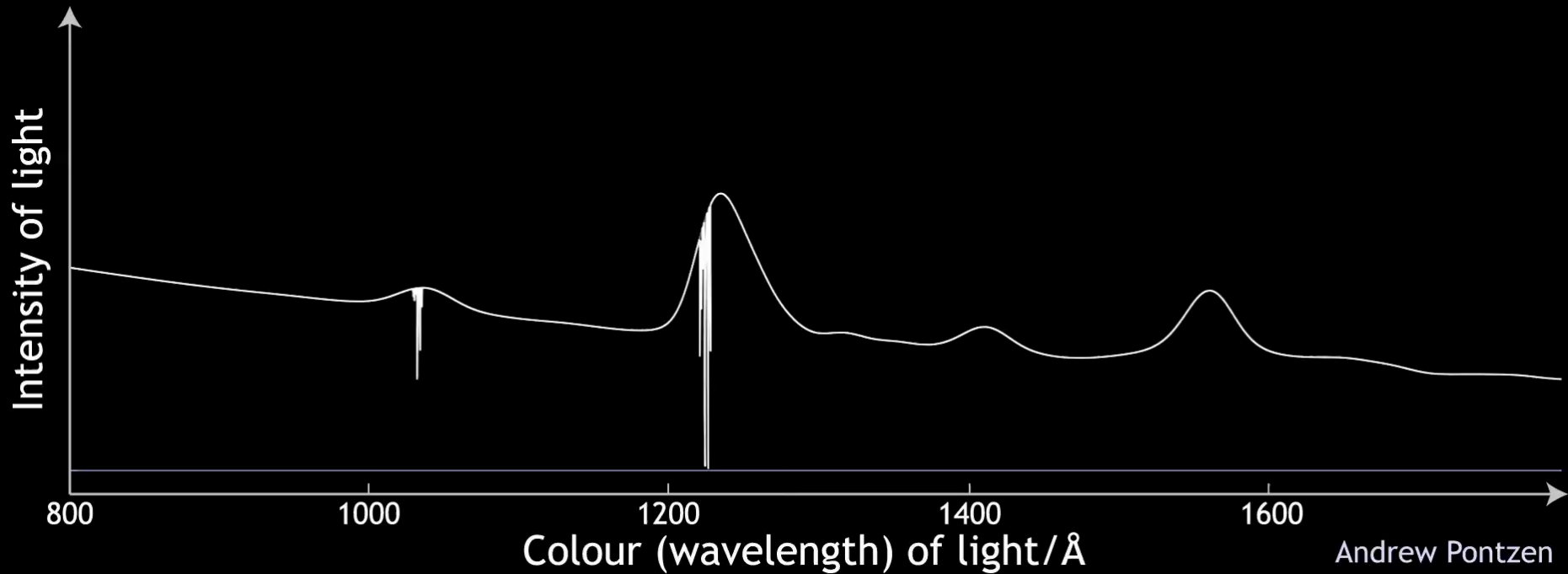
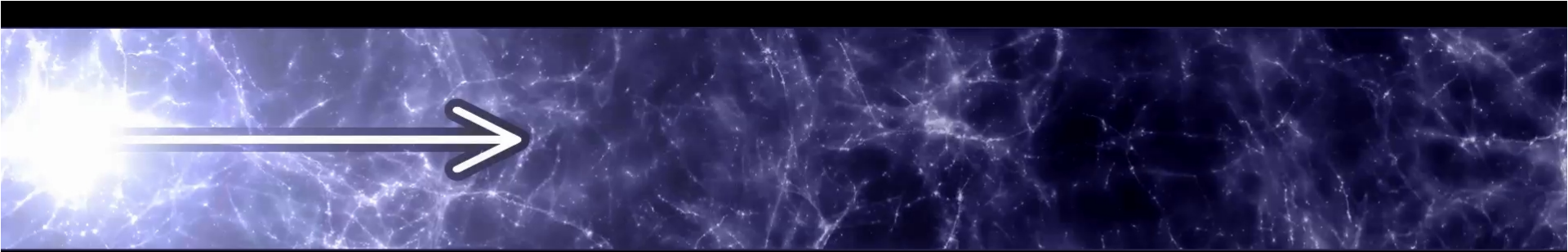
Intro: Ly- α Forest



Intrinsic spectrum



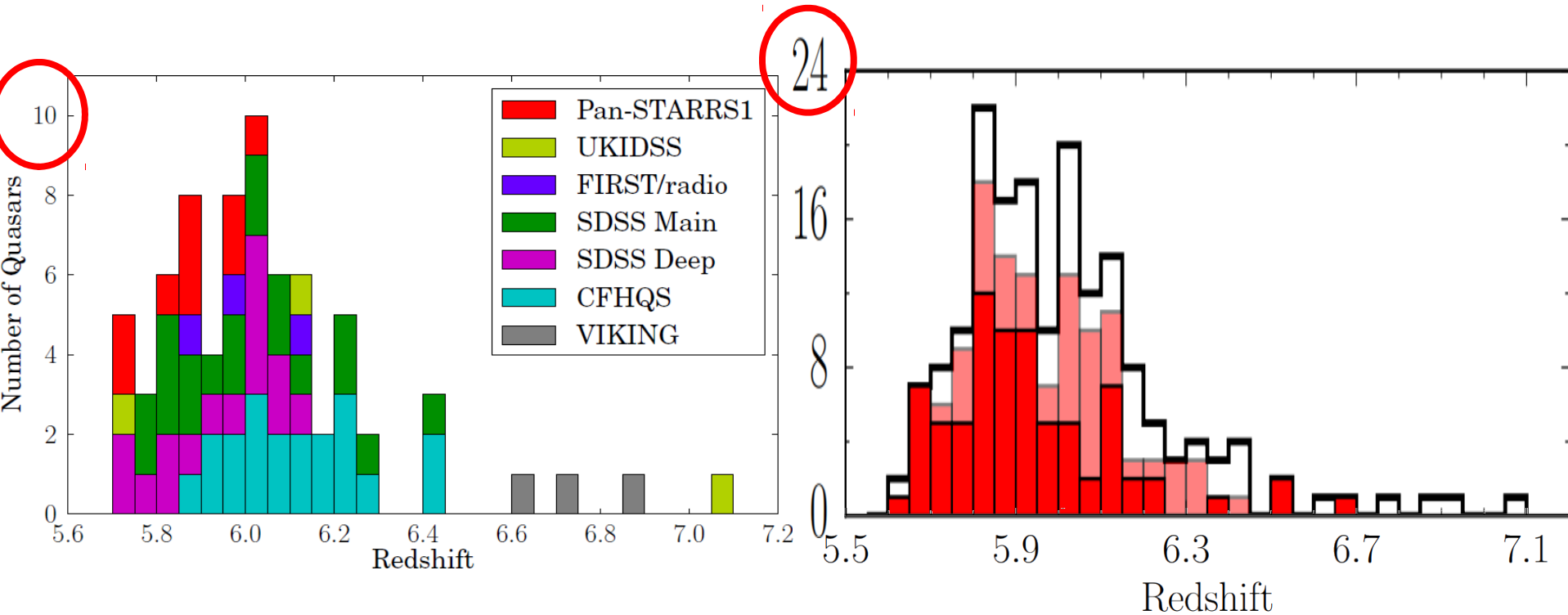
Intro: Ly- α Forest



Andrew Pontzen

What We Know Now: Post-reionization IGM

- High- z QSOs appear as mushrooms after a rain, and data become extremely constraining.

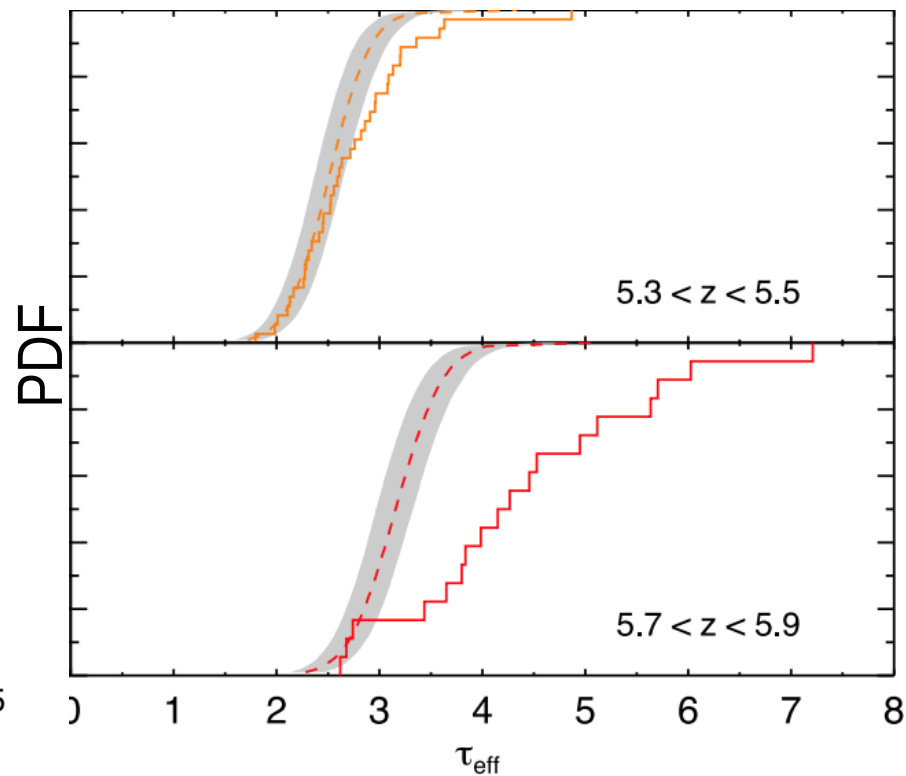
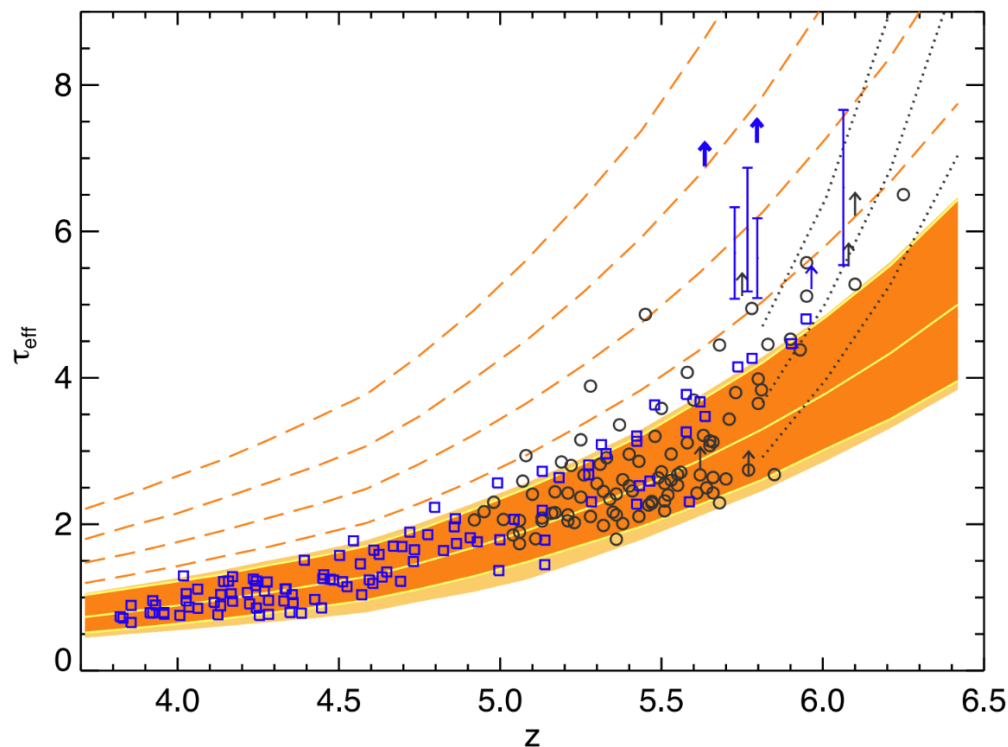


Banados+ 2014

Banados+ 2016

What We Know Now: Post-reionization IGM

- High- z QSOs appear as mushrooms after a rain, and data become extremely constraining.



The Flood Is Coming: GMT,TMT,E-ELT

- Large optical telescopes will increase high-redshift quasar samples ~ 100 -fold.

E-ELT: 42m

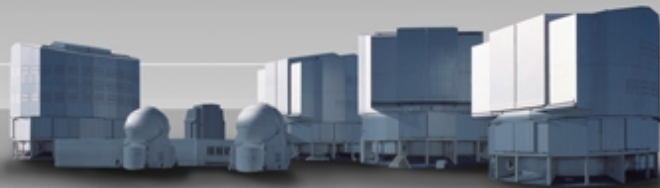
100 m

80 m

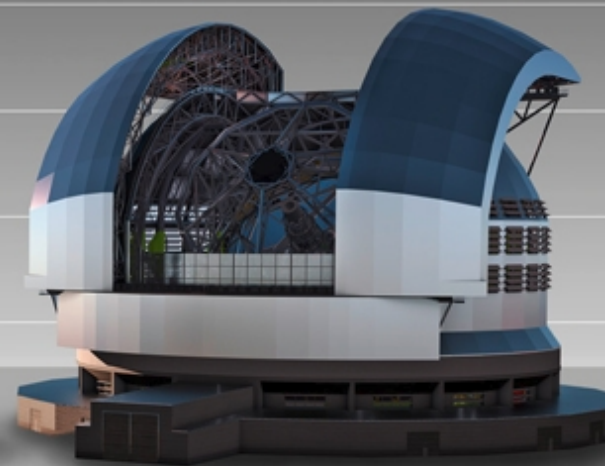
60 m

40 m

20 m



Very Large Telescope



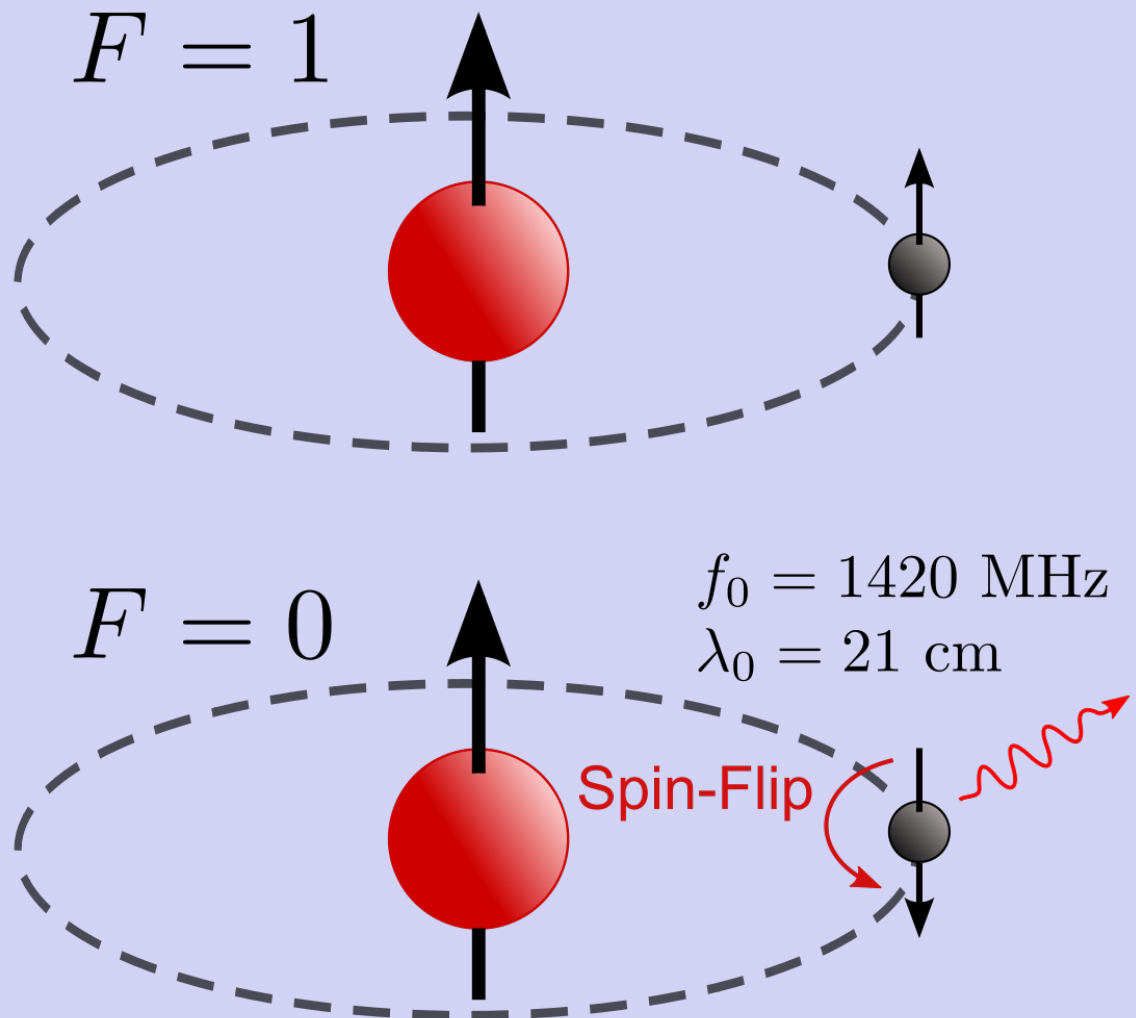
European Extremely Large Telescope



Keck Telescope

What We Know Now: Redshifted 21cm Emission

- Neutral hydrogen structure line.
- During reionization neutral, that emission
- Epoch of reionization (E is far away, the
- Foregrounds (galactic) are \sim times stronger the cosmic signal



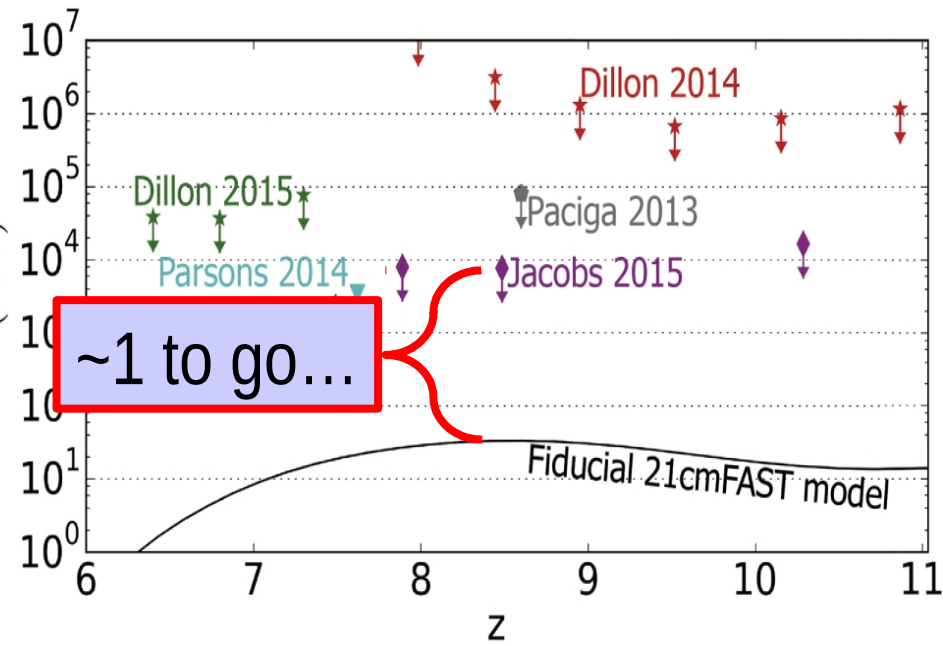
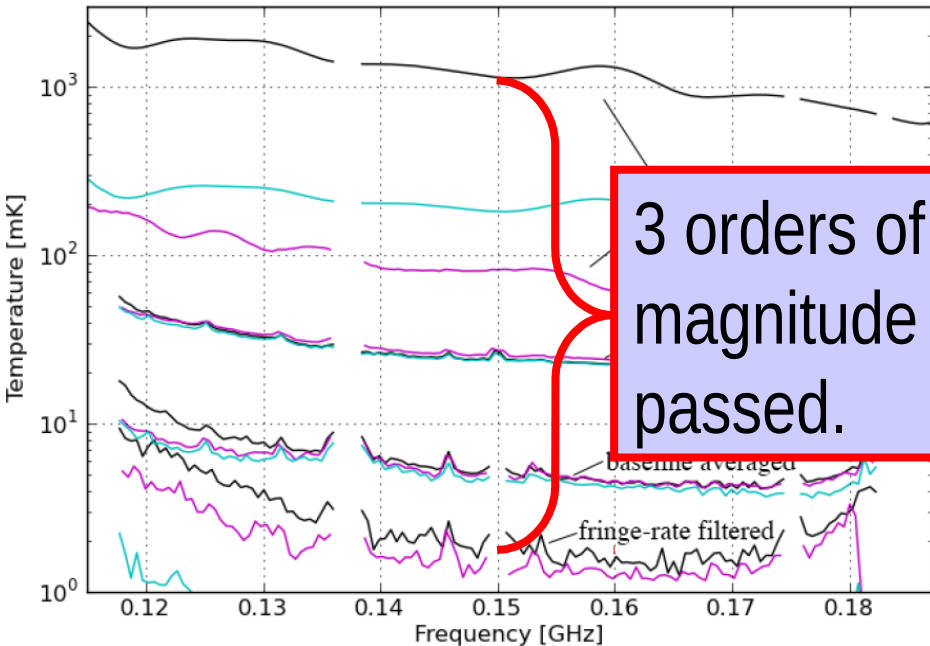
What We Know Now: Redshifted 21cm Emission

- Current experiments:
 - Precision Array for Probing the Epoch of Reionization (PAPER), South Africa, finished
 - Murchison Widefield Array (MWA), western Australia, finished
 - Low Frequency Array (LOFAR), *Netherlands*



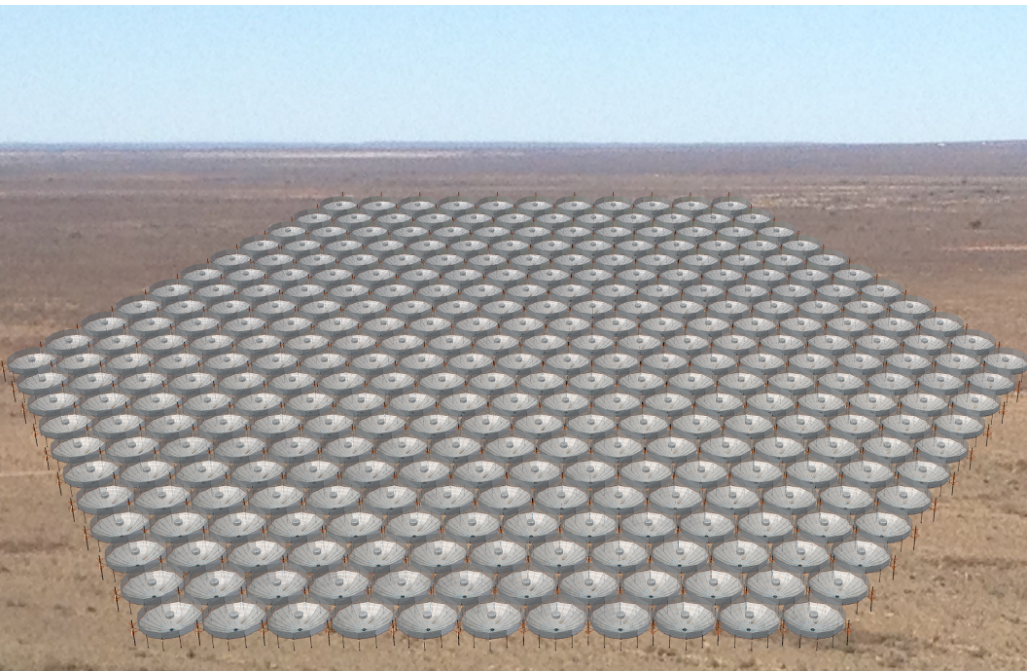
What We Know Now: Redshifted 21cm Emission

- No measurement yet.
- PAPER and MWA have placed the upper limits so far, the LOFAR limit is forthcoming.
- They are within a factor of ~ 10 ...

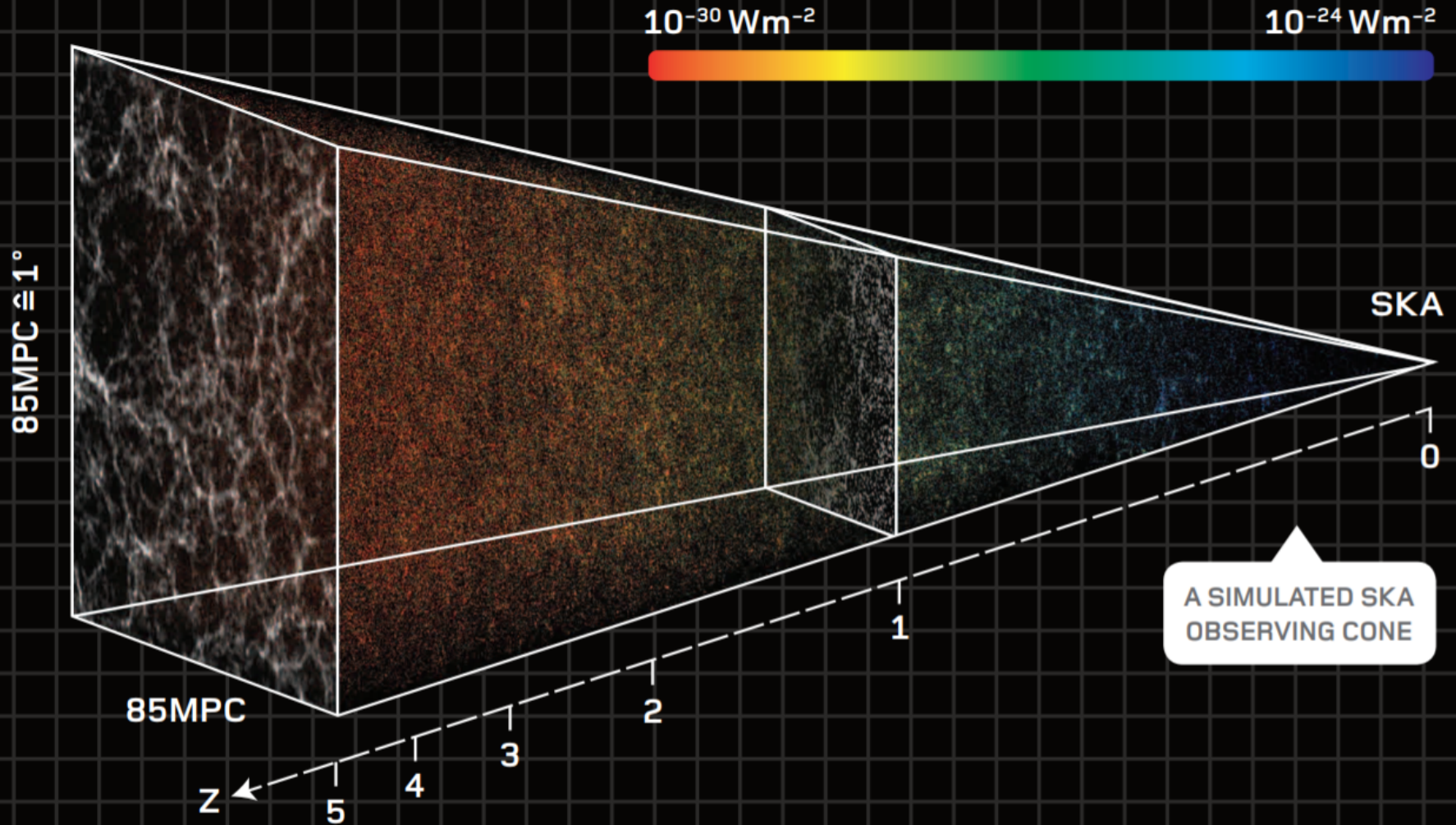


The Flood Is Coming: Redshifted 21cm Signal

- Despite 1st generation experiments not detecting any signal, NSF+MF already fully funded the 2nd generation experiment:
Hydrogen Epoch of Reionization Array (HERA).



The Flood Is Coming: Redshifted 21cm Signal



What Will We Learn From All That Data?

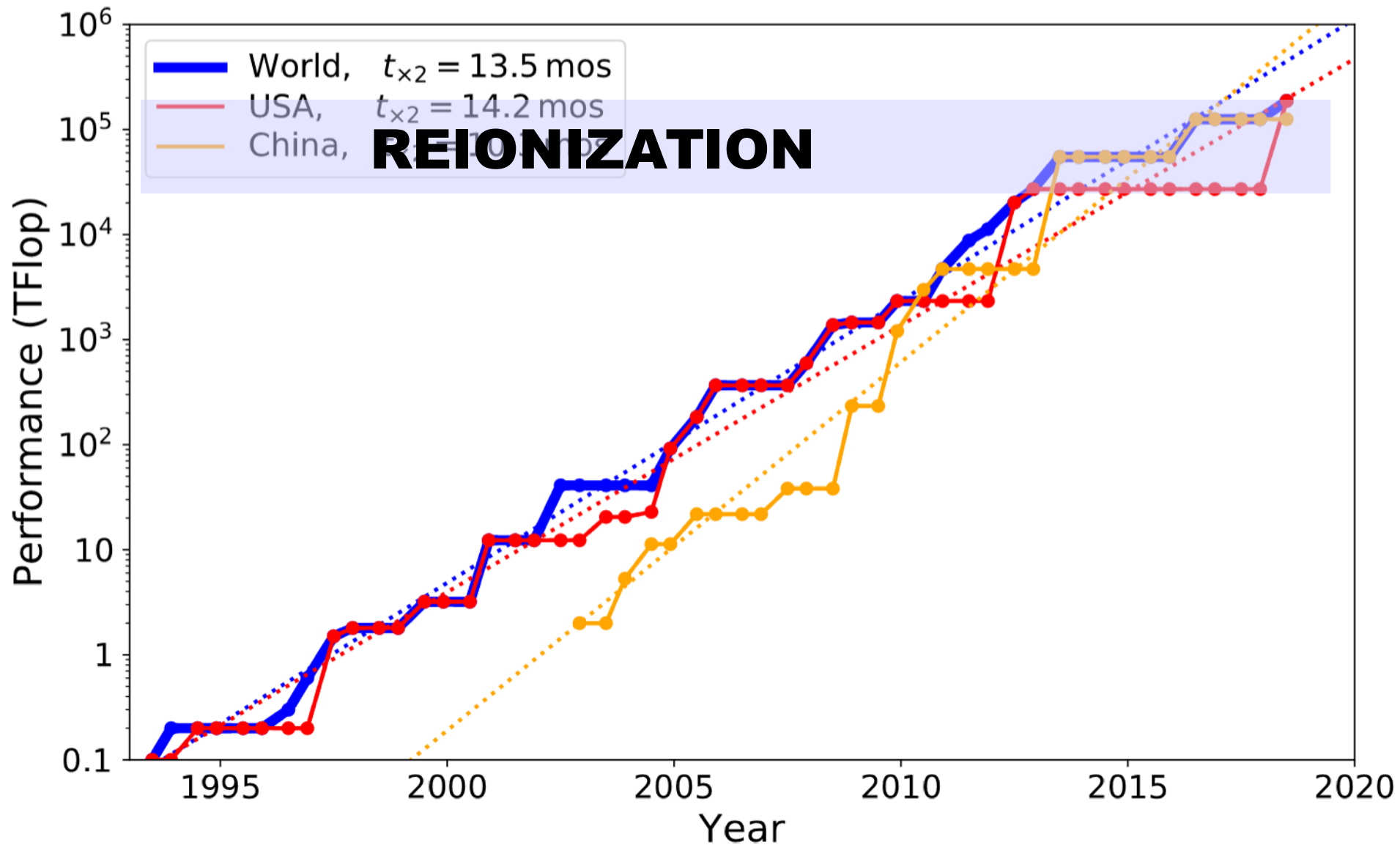
- Actually, not much, unless...
- ...we get equally precise theory.



The Flood Is Coming

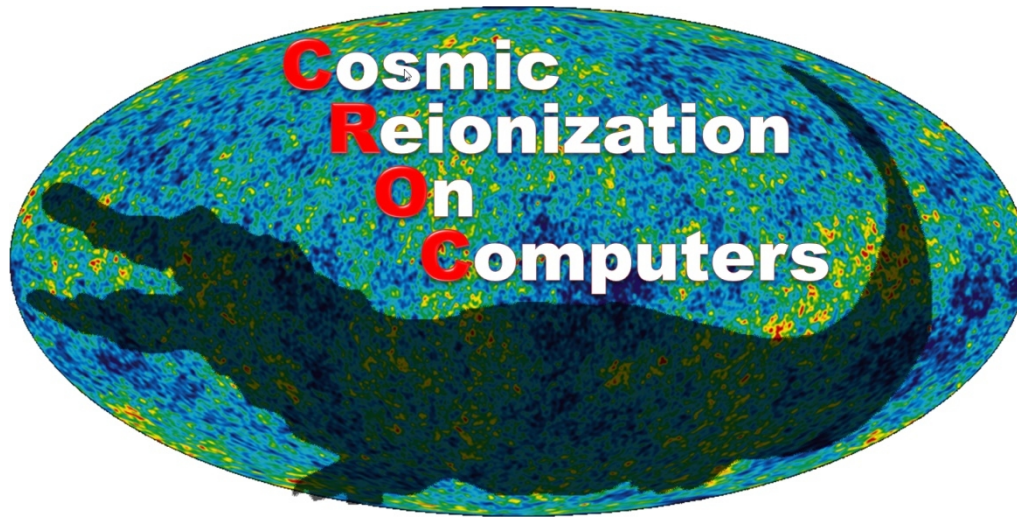
- Existing observations already make many models obsolete; forthcoming observations will make obsolete **all** of them.
- Hence, theorists' task is not to make more models now, but rather to develop new modeling technology.
- And it is happening:
 - *Aurora*
 - *Cosmic Dawn*
 - *CRASH*
 - *CROC*
 - *DRAGONS*
 - *Emma*
 - *Renaissance Simulations*
 - *SPHINX*

Moore's Law



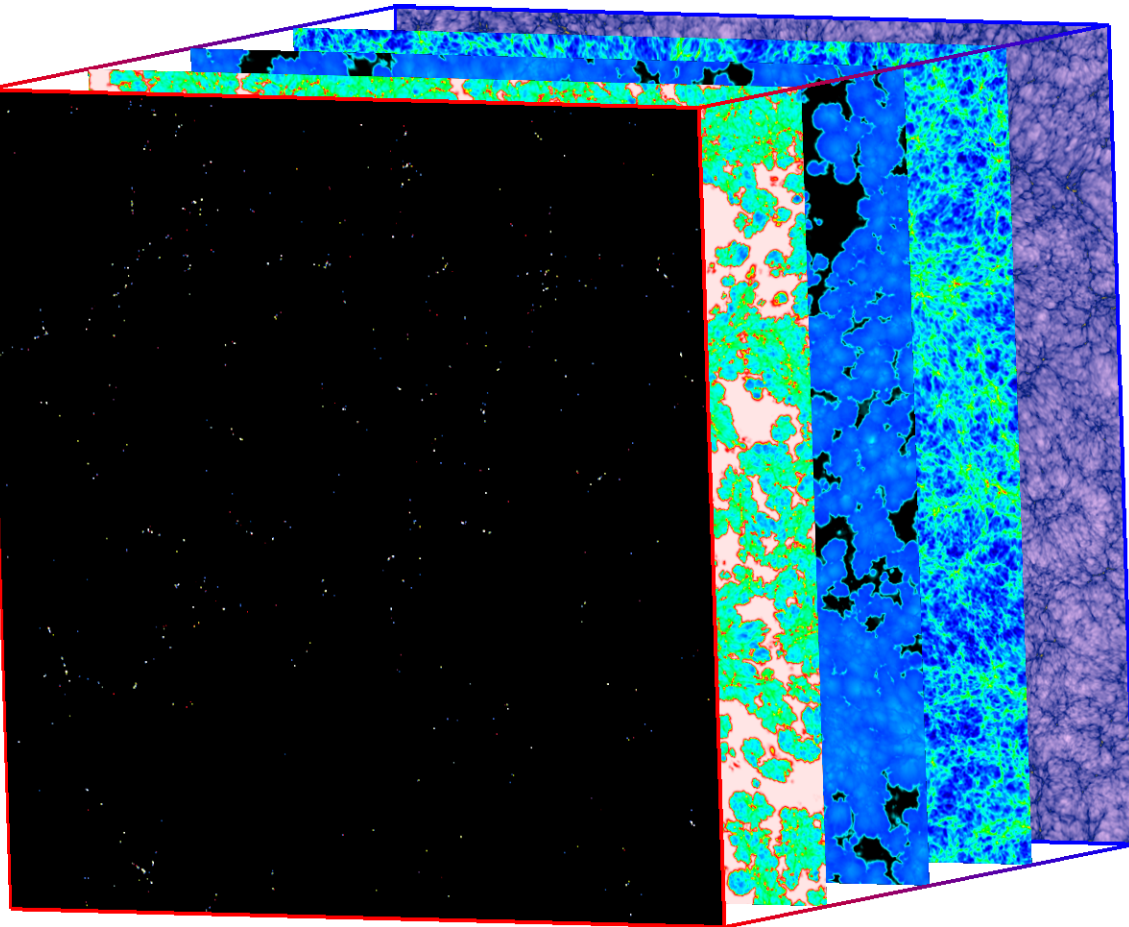
The Flood Is Coming

- One (typical) example in action:



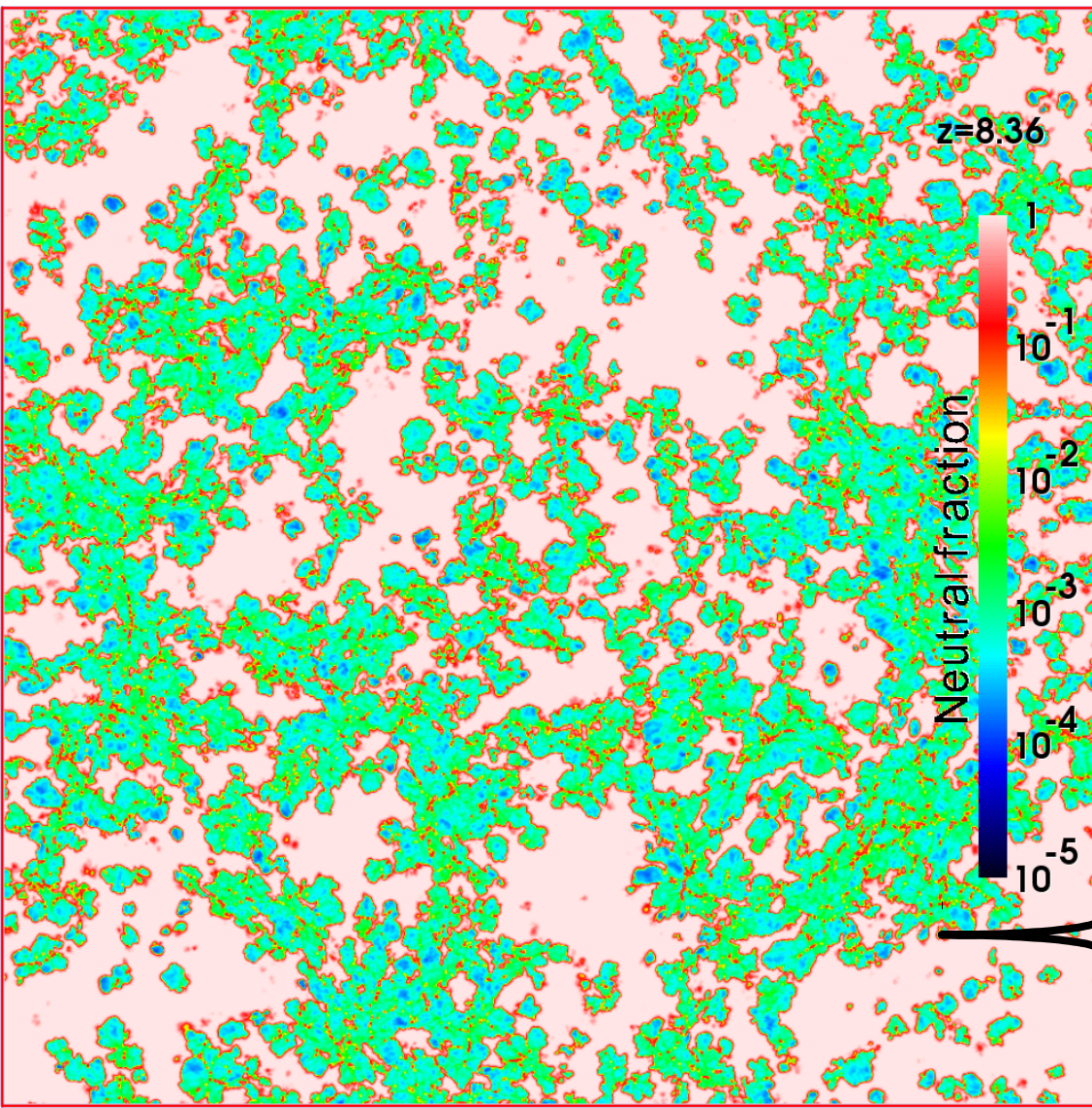
- Disclaimer: *other projects/efforts are achieving similar overall levels of computational scale, agreement with data, etc.*

Simulating Reionization: Physics



- Dark matter
- Gas dynamics
- Atomic processes
- Radiative transfer
- Star formation & stellar feedback

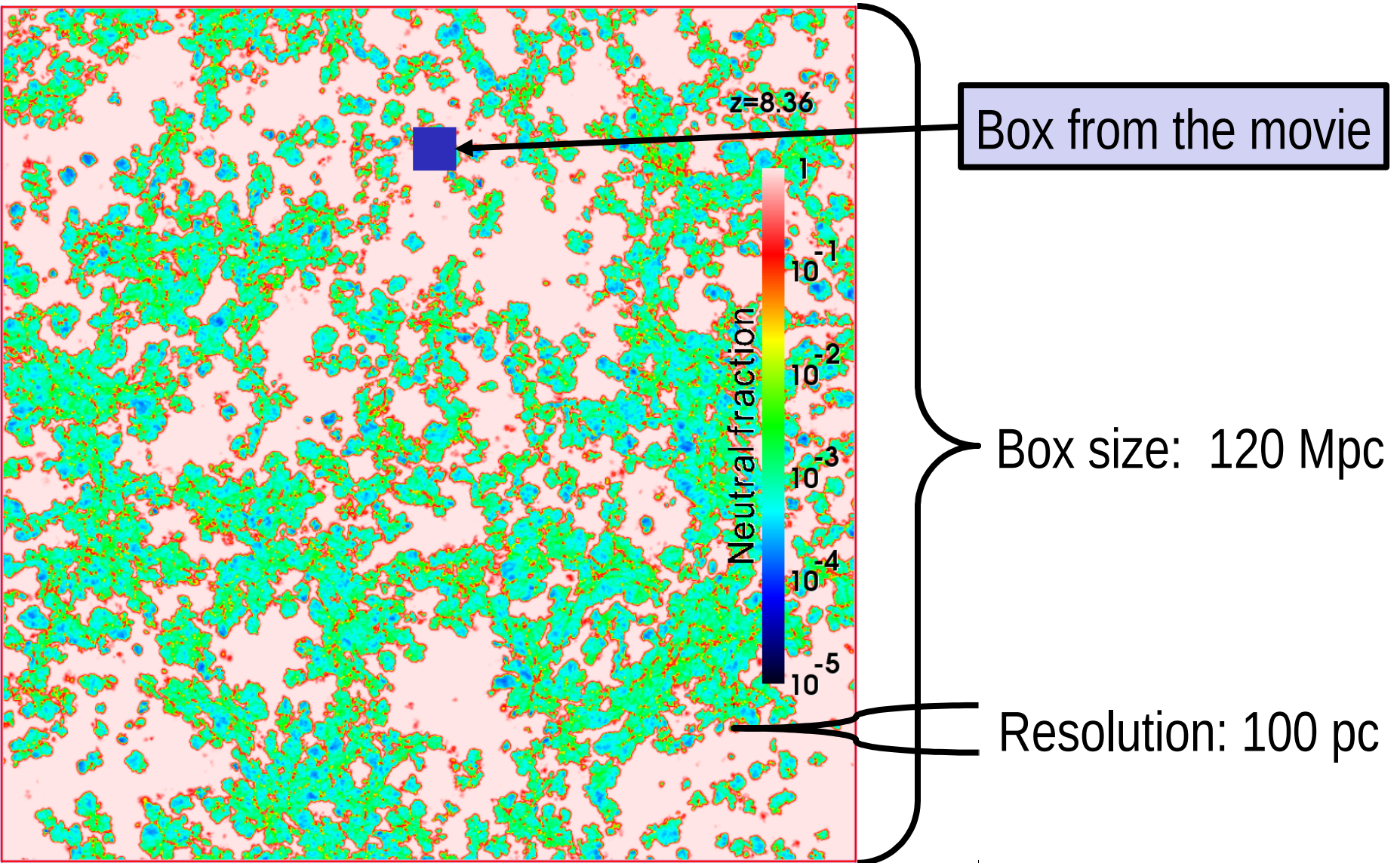
Simulating Reionization: CROC on Blue Waters

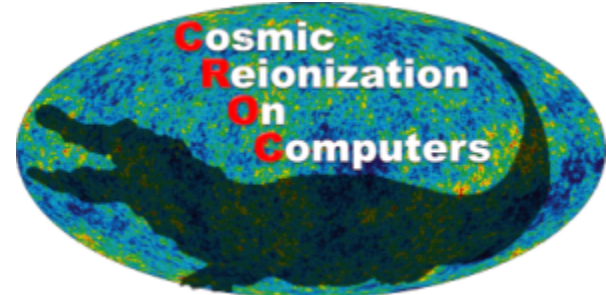


Box size: 120 Mpc

Resolution: 100 pc

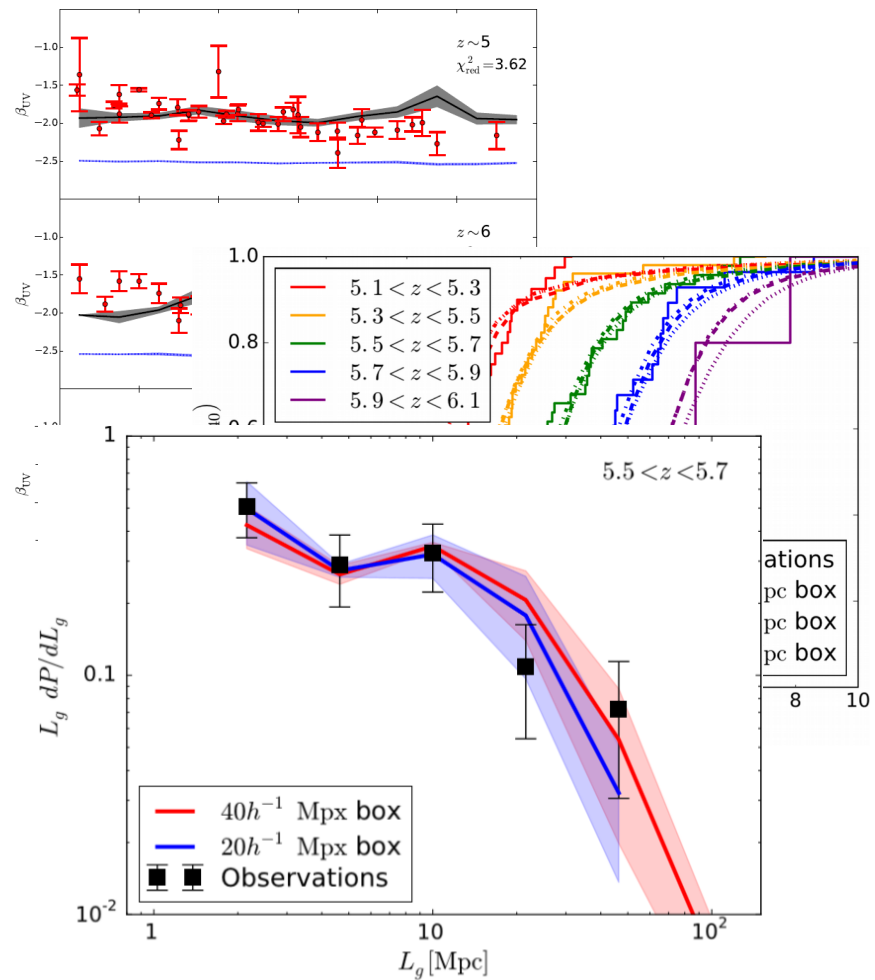
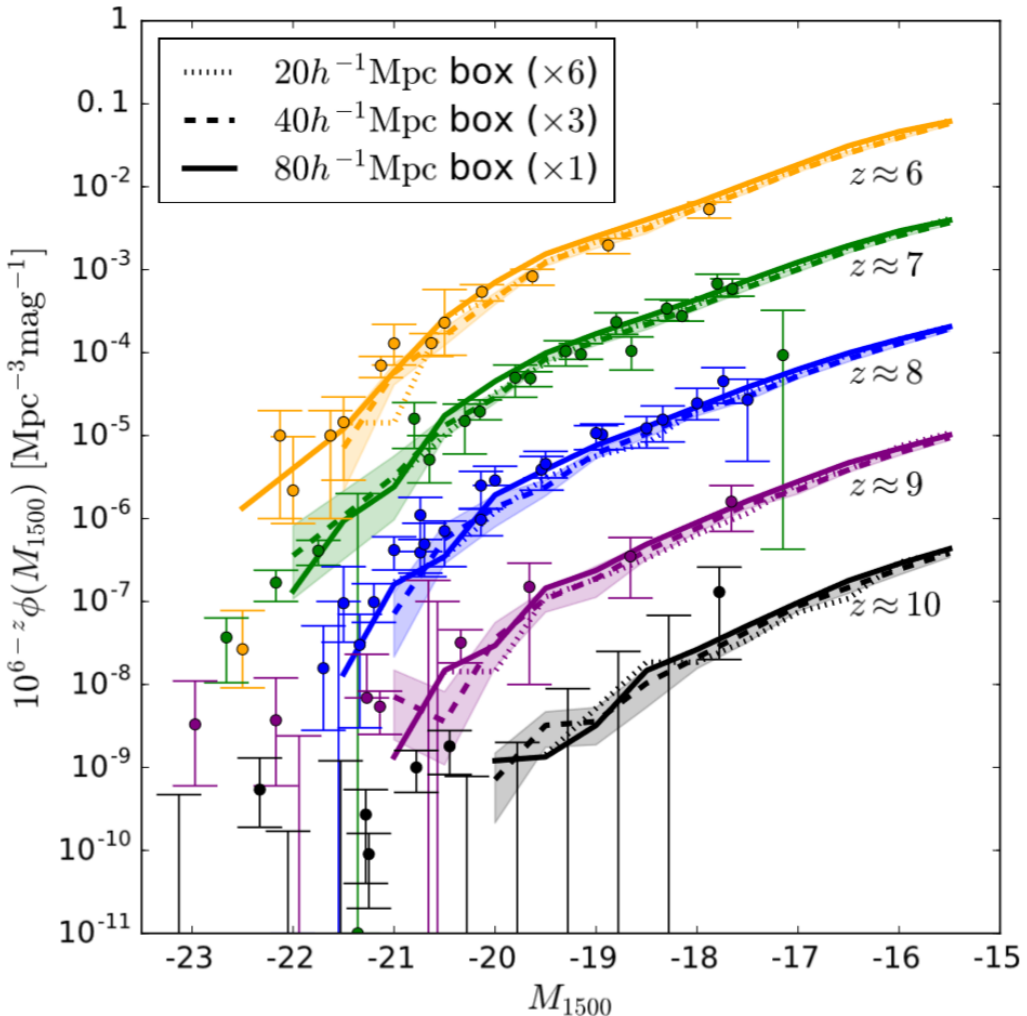
Simulating Reionization: CROC on Blue Waters

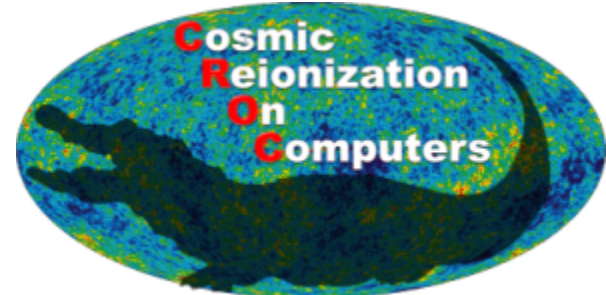




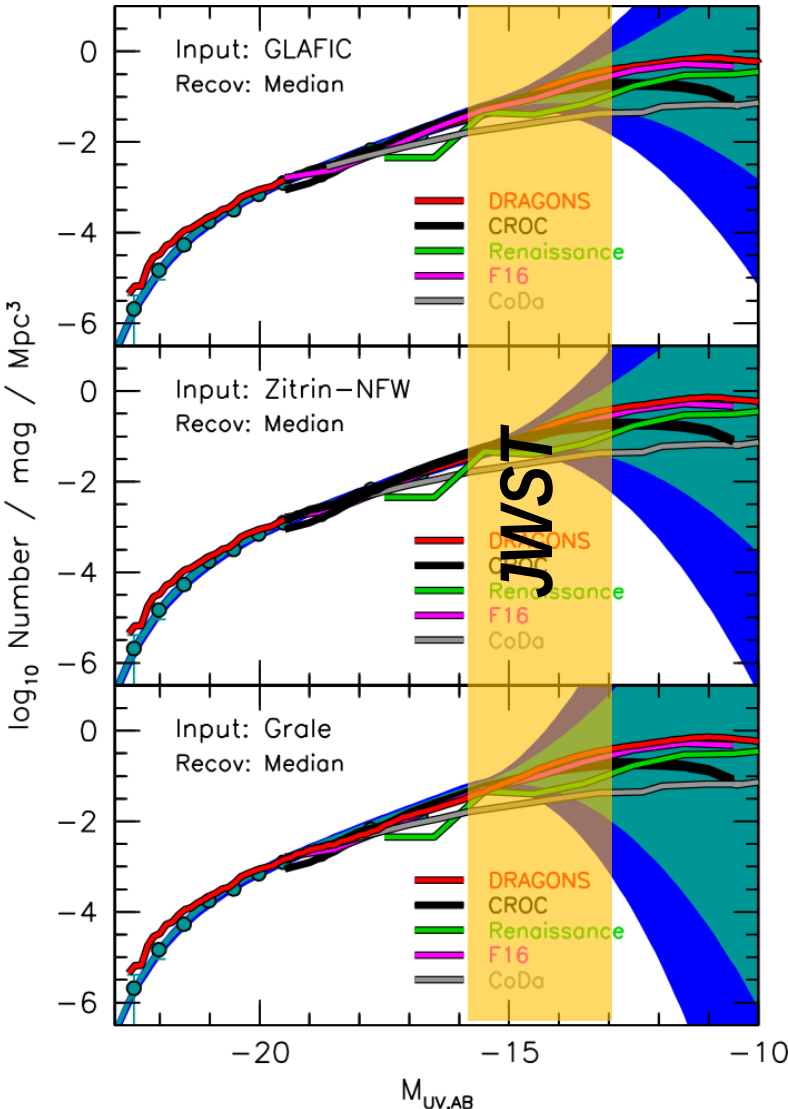
The CROC Project: Success

Galaxies (and many other things) are ok!





The Role of JWST: There Will Be Blood!



- All simulation agree with the data where the data exist.
- They all disagree (for a good reason!) strongly in the regime that JWST will probe.
- JWST will kill most (or all) existing models.

Conclusions

- We are entering a golden age of reionization studies, with observations increasing the data volume ~ 100 -fold and opening new subfields of astronomy.
- Theory is finally maturing to the level of theoretical uncertainties approaching observational ones, and Moore's law ensures that it will only get better.
- Lots of fun ahead...

The End

Galaxy Formation Large-Scale Structure



Reionization