

ARXIV: [2403.07984](https://arxiv.org/abs/2403.07984)

Final State Radiation And (Ultra)High Energy Neutrinos

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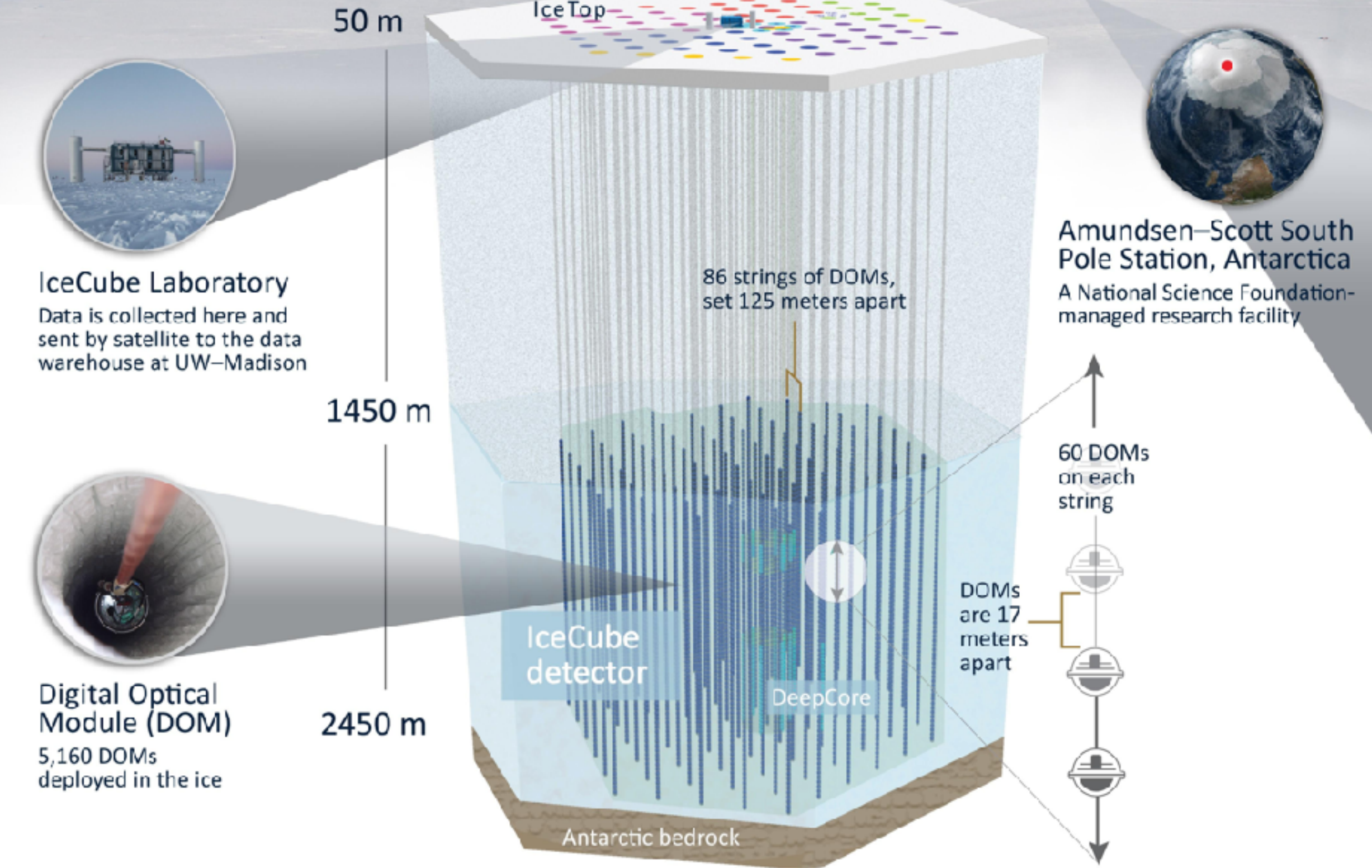
COLLABORATORS

BEI ZHOU | FERMILAB

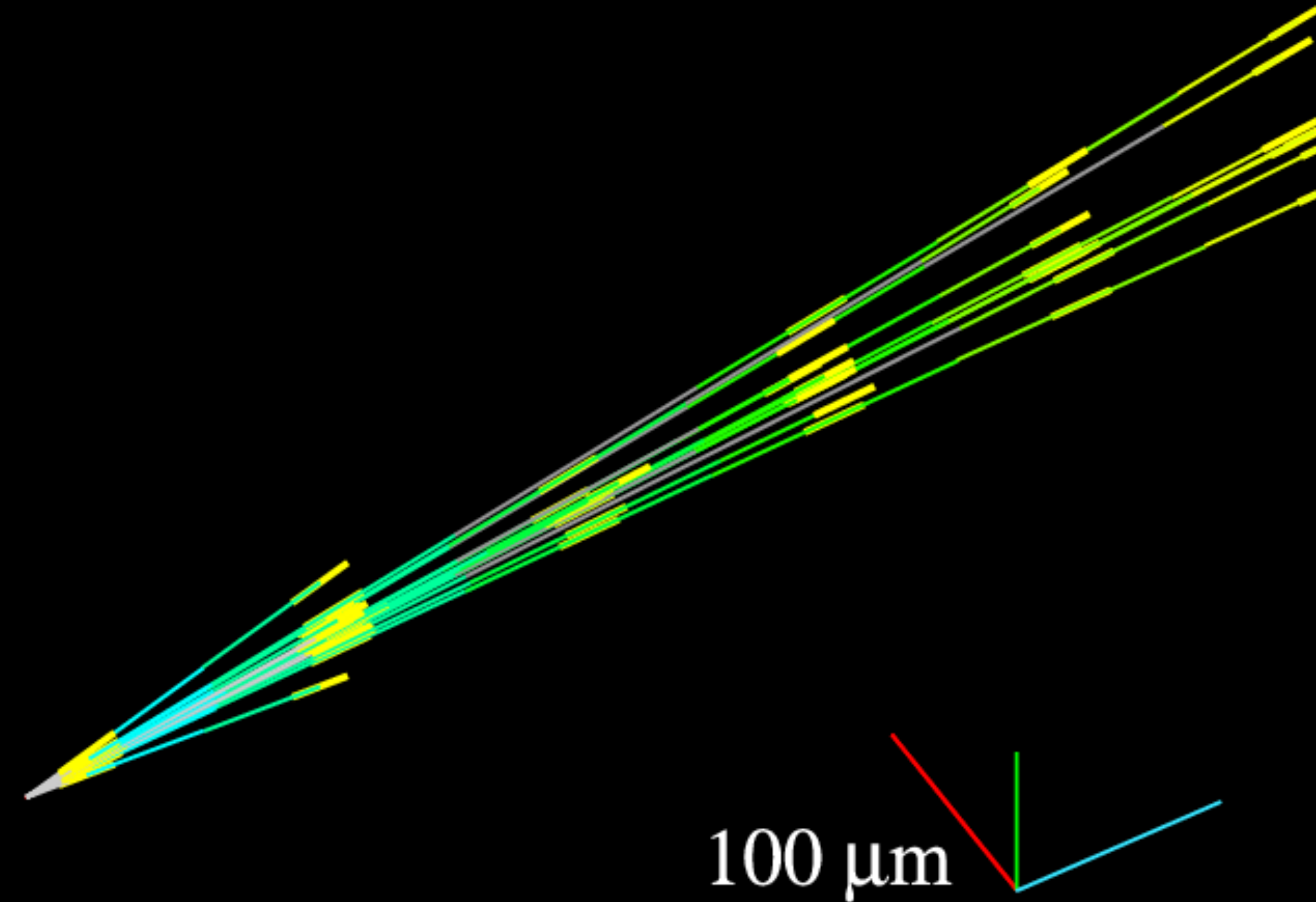
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Caltech

Neutrino Theory Network



FASER



Motivation & Context

MULTI-MESSENGER ASTRONOMY



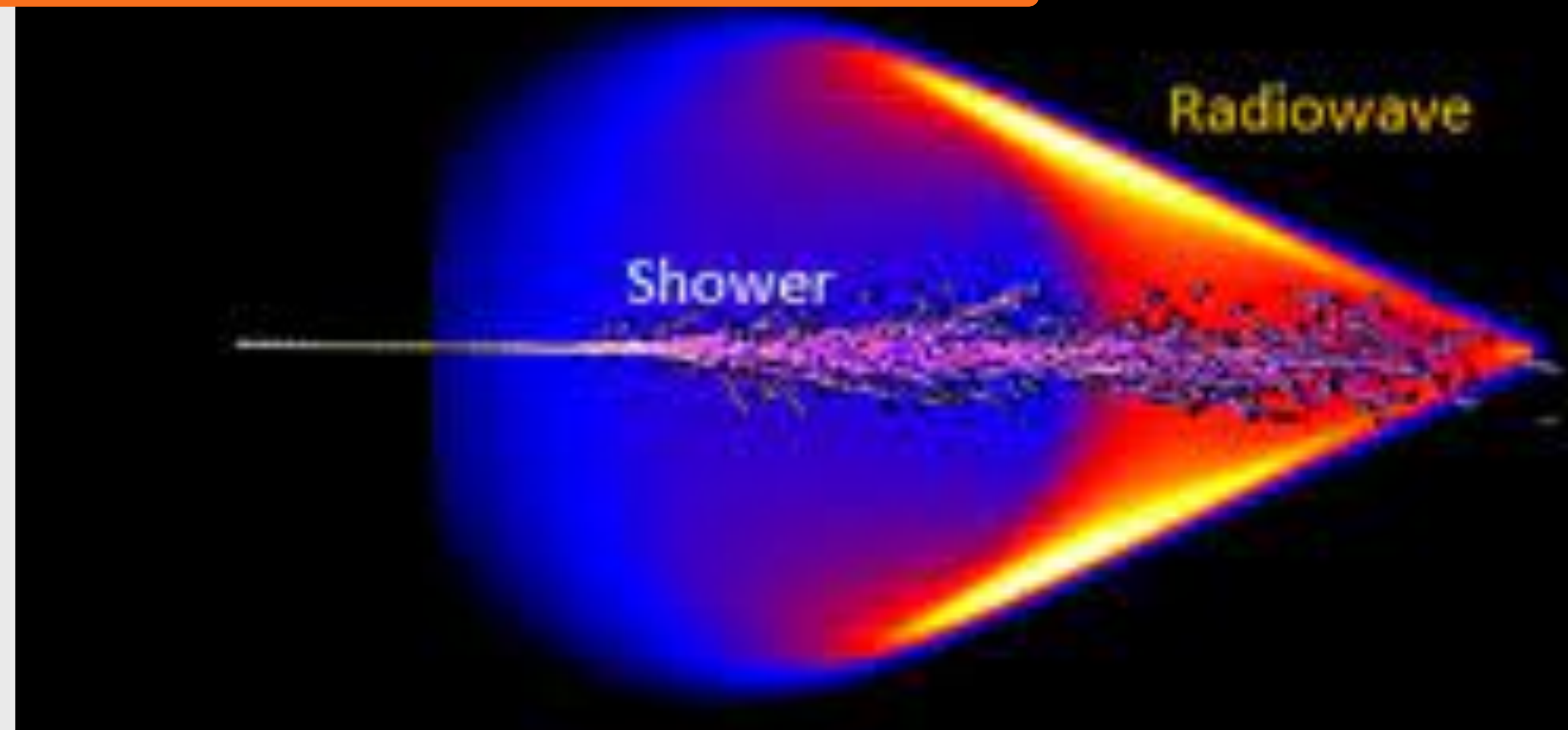
- Neutrinos are an increasingly important for astrophysics.
- Neutrino interactions are crucial to interpret data.

HIGH ENERGY NEUTRINO TELESCOPES



- Neutrinos offer a "sterile" messenger of astrophysical phenomena.
- IceCube Gen-2 will instrument a very large area,
- Few singlet operators available.
Focus on "portals".

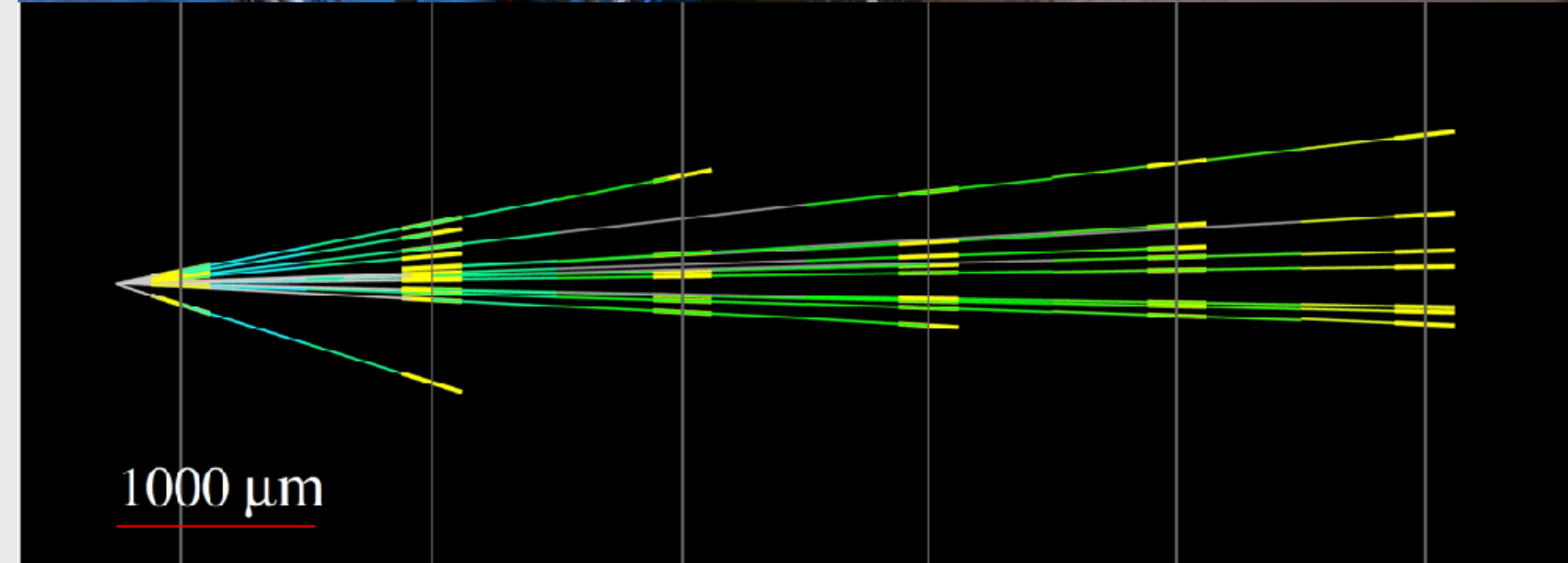
ULTRAHIGH ENERGY NEUTRINO DETECTION



- Ultra high energy neutrinos (above 100 PeV) are difficult to study due to their much smaller flux.
- Near-term telescopes/facilities planned.
- Use very different detection strategies.

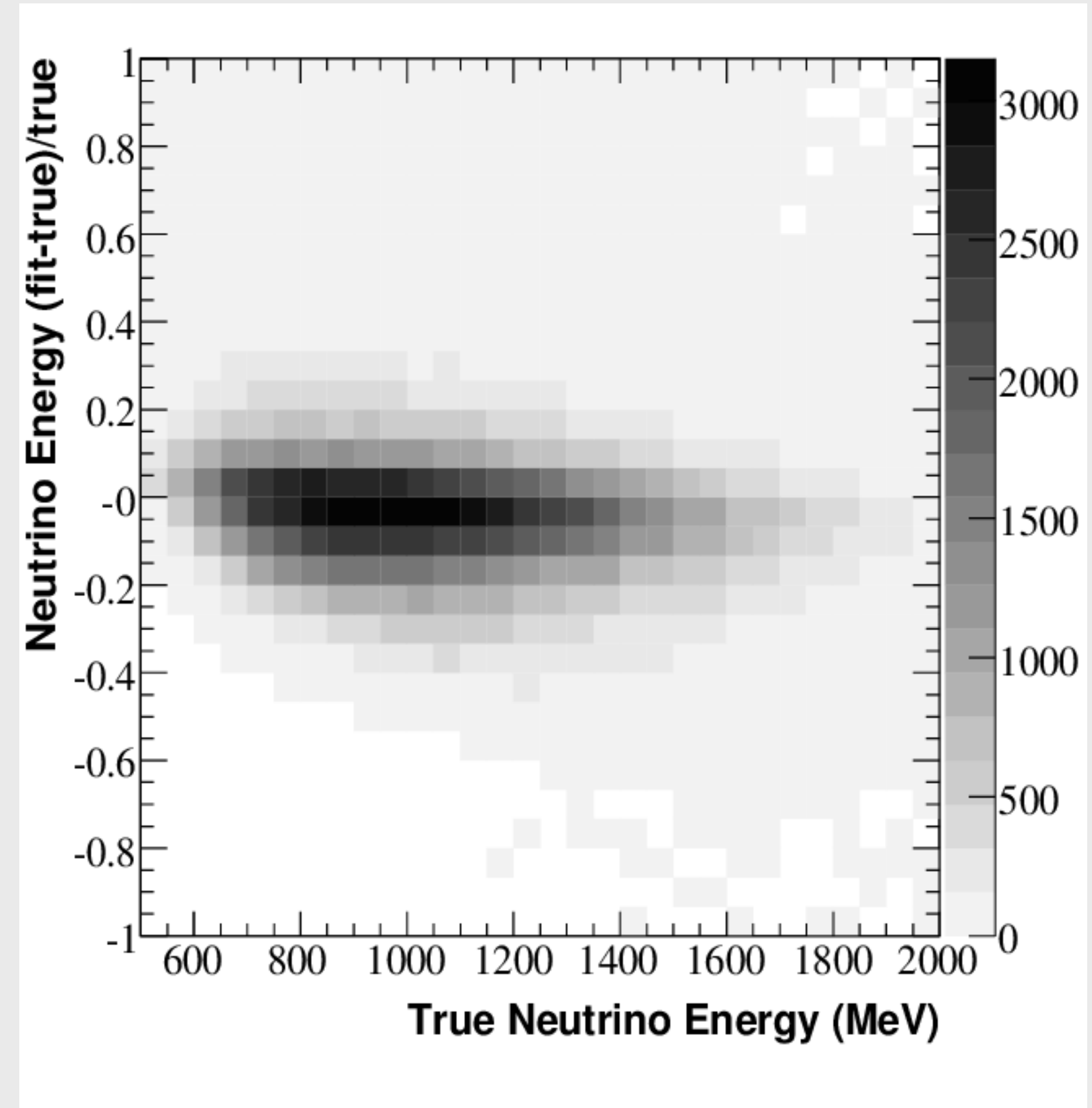
COLLIDER NEUTRINOS

- New collider based experiments at the LHC.
- Lab-measurements of \sim TeV neutrino energies.
- Granular detectors with high vertex resolution.



ENERGY ESTIMATORS

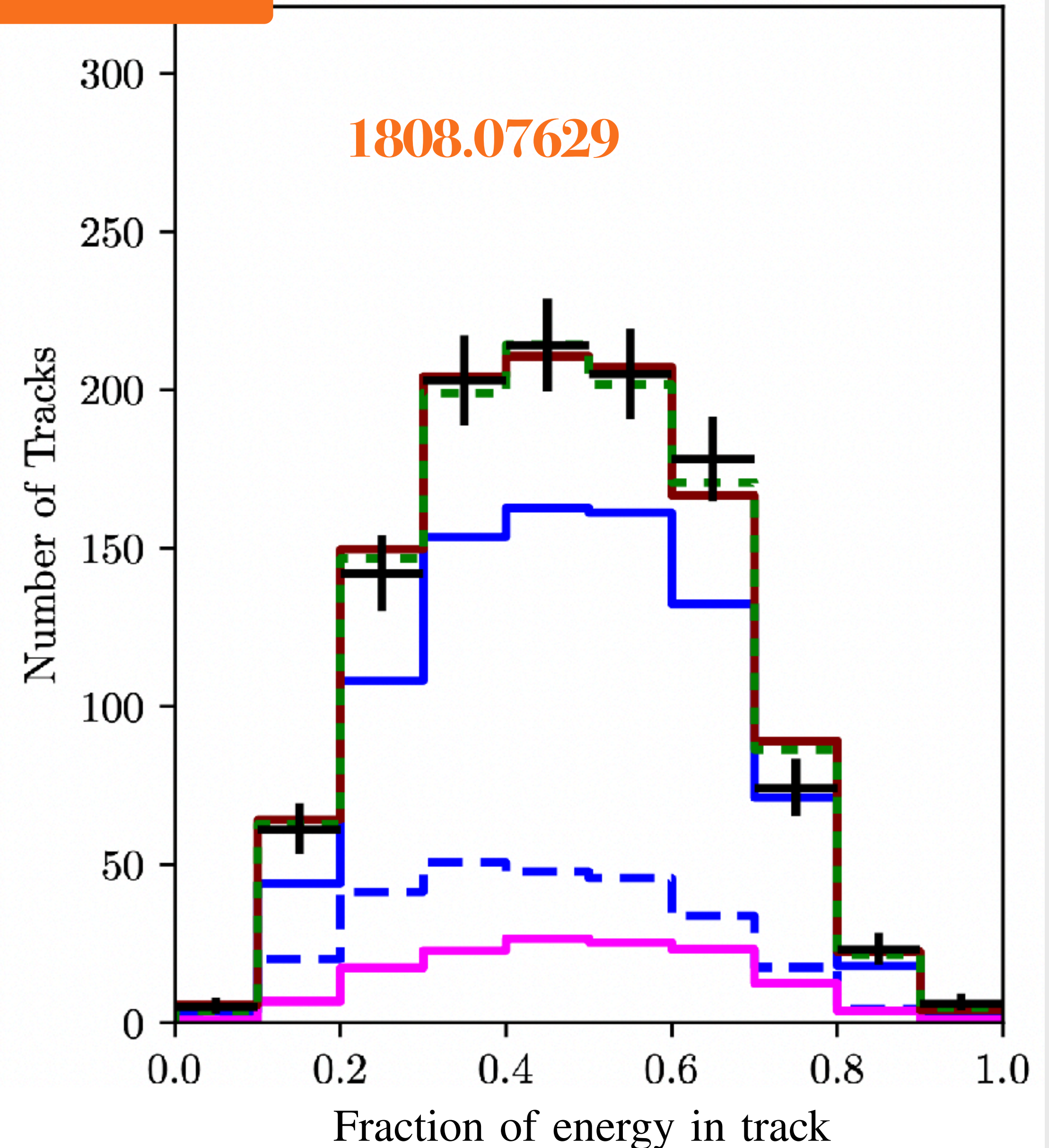
- Estimating the parent neutrino energy is important for all experiments.
- Any effect which causes a systematic bias in energy reconstruction is important.
- Photons and muons "look" very different in a detector.

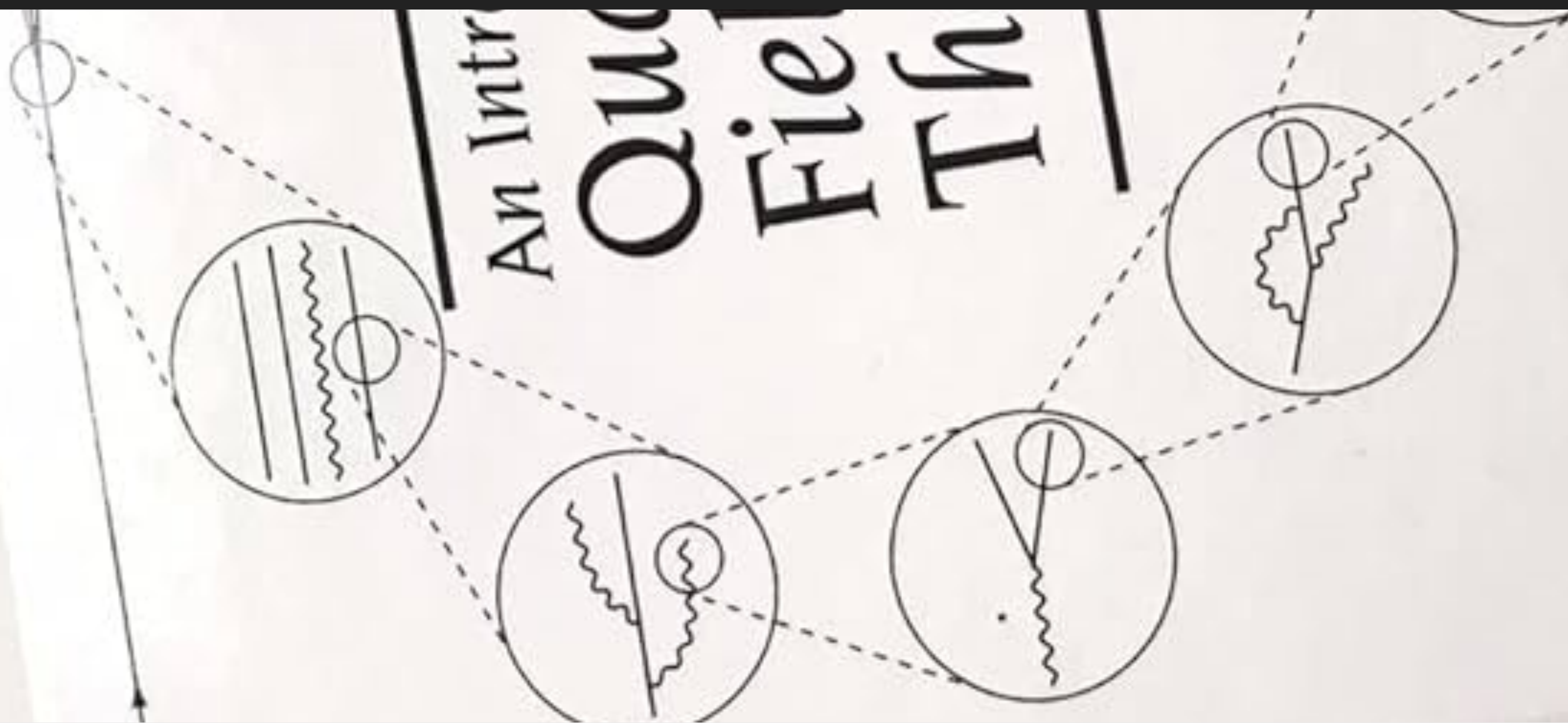


INCLUSIVE VS EXCLUSIVE OBSERVABLES

- Data sets are growing and can bin events.
- Differential distributions with respect to e.g. lepton energy.
- Goal of $\sim 10\%$ accuracy as a heuristic benchmark.

$$10^{3.0} \text{ GeV} < E_{\text{vis.}} < 10^{3.5} \text{ GeV}$$





An Intro
QED
Field
Th

Michael E. Peskin

Radiative Corrections : The Basics



OLD IDEA

Nuclear Physics B154 (1979) 394–426
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USE SPLITTING
FUNCTION

**RADIATIVE CORRECTIONS TO HIGH-ENERGY NEUTRINO
SCATTERING**

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Received 19 January 1979

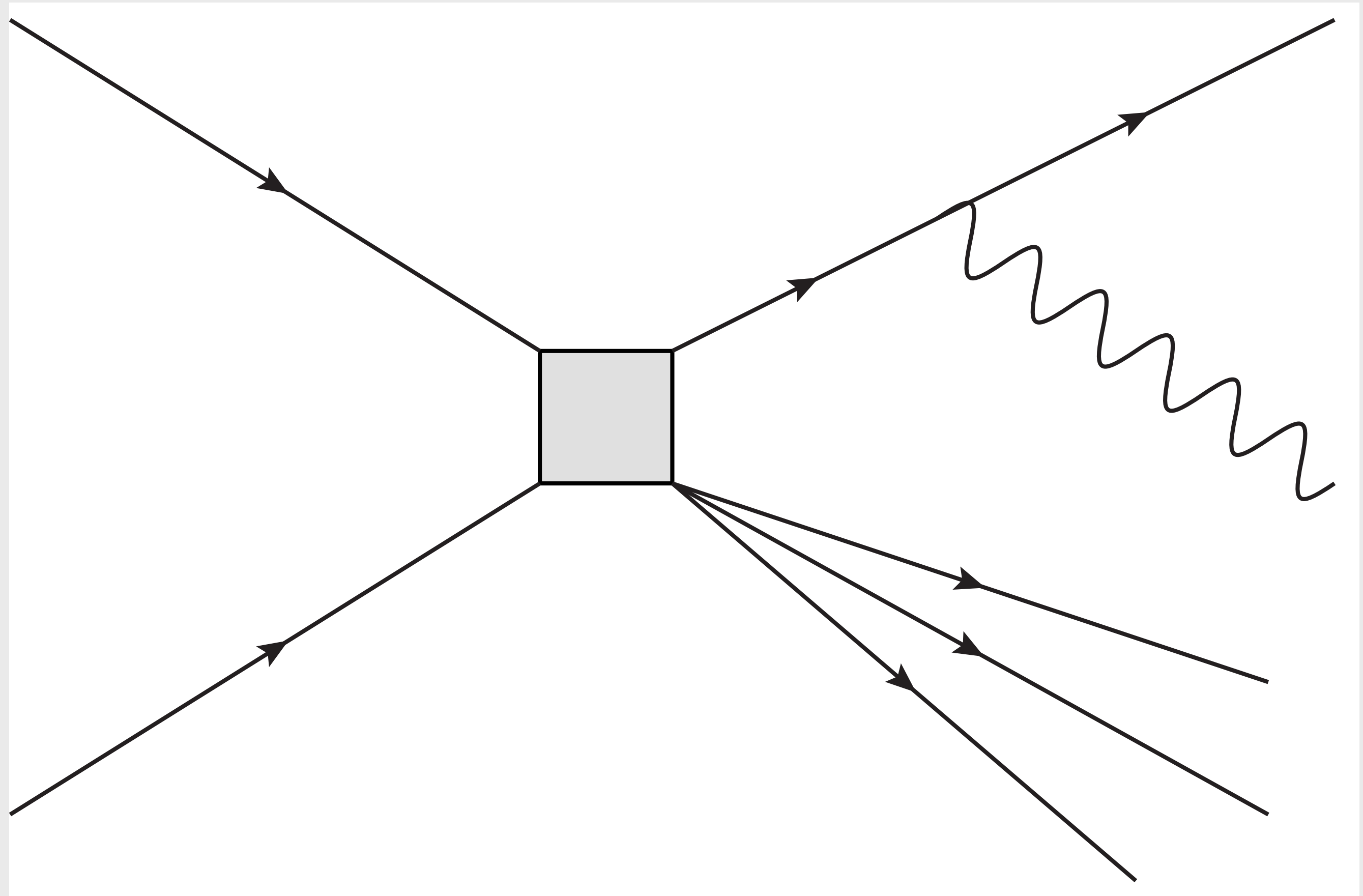
NEW
APPLICATIONS

REAL PHOTON EMISSION

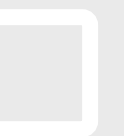
- Lepton radiates a photon.

$$\frac{1}{2p \cdot q + q^2} \rightarrow \frac{1}{2p \cdot q}$$

$$\int \frac{d^3q}{2E_q} |\mathcal{M}|^2 \sim \int \frac{d\omega}{\omega}$$



- Infrared divergence

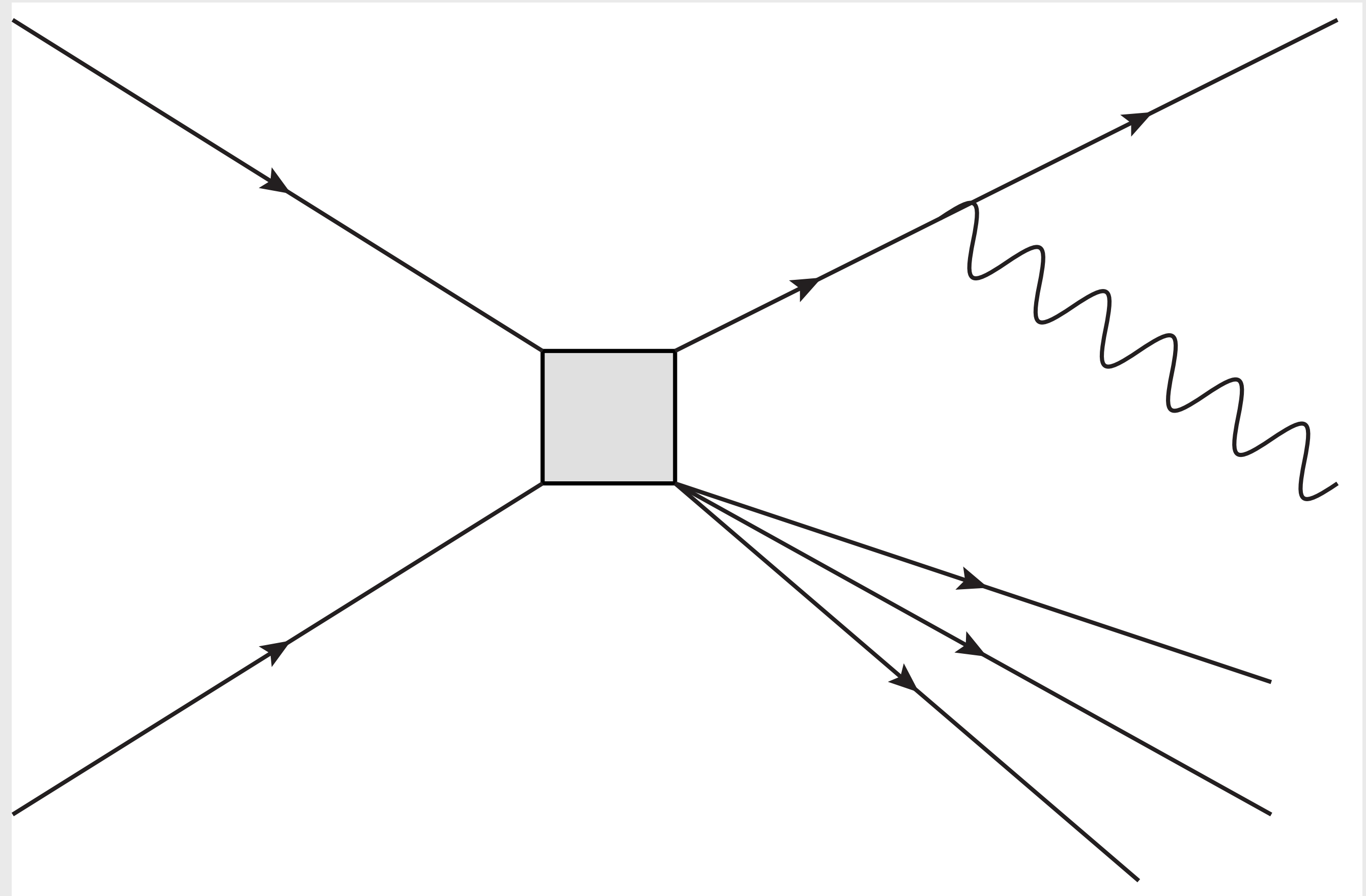


REAL PHOTON EMISSION

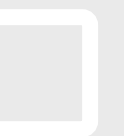
- Lepton radiates a photon.

$$\frac{1}{2p \cdot q + q^2} \rightarrow \frac{1}{2p \cdot q}$$

$$\int \frac{d^3q}{2E_q} |\mathcal{M}|^2 \sim \int \frac{d\theta^2}{\theta^2 + m_\ell^2}$$



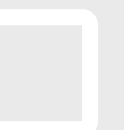
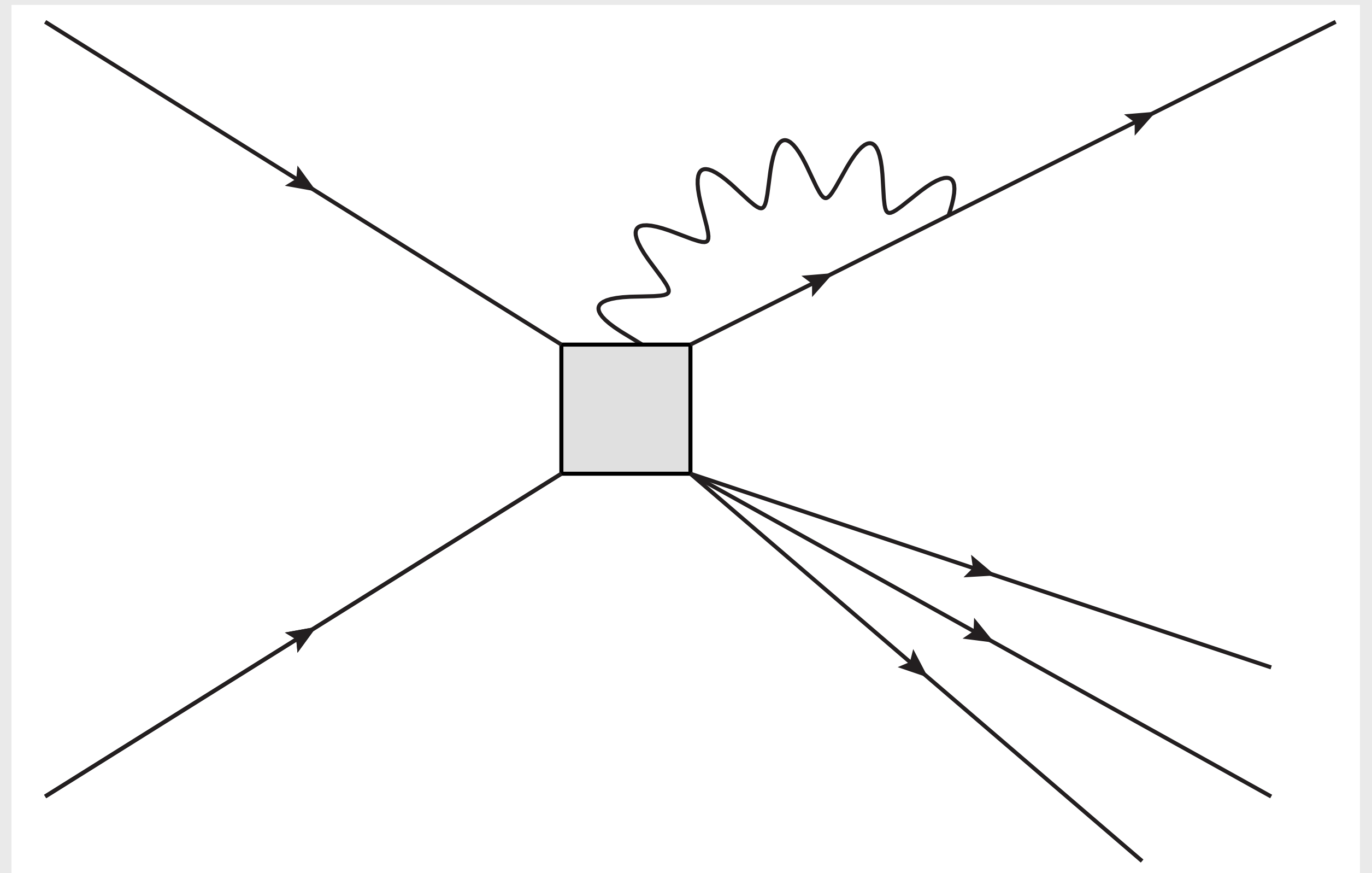
- Collinear divergence



VIRTUAL CORRECTIONS

- In addition to diagrams where a photon is explicitly radiated, we have to include corrections to the process where no photon is radiated.
- These diagrams **decrease** the cross section.

CANCELS AGAINST
REAL RADIATION
CONTRIBUTION



REAL + VIRTUAL CORRECTIONS

- KLN Theorem implies that IR divergences must cancel between real and virtual corrections.
- Implies the same for collinear divergences.

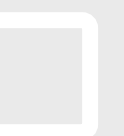
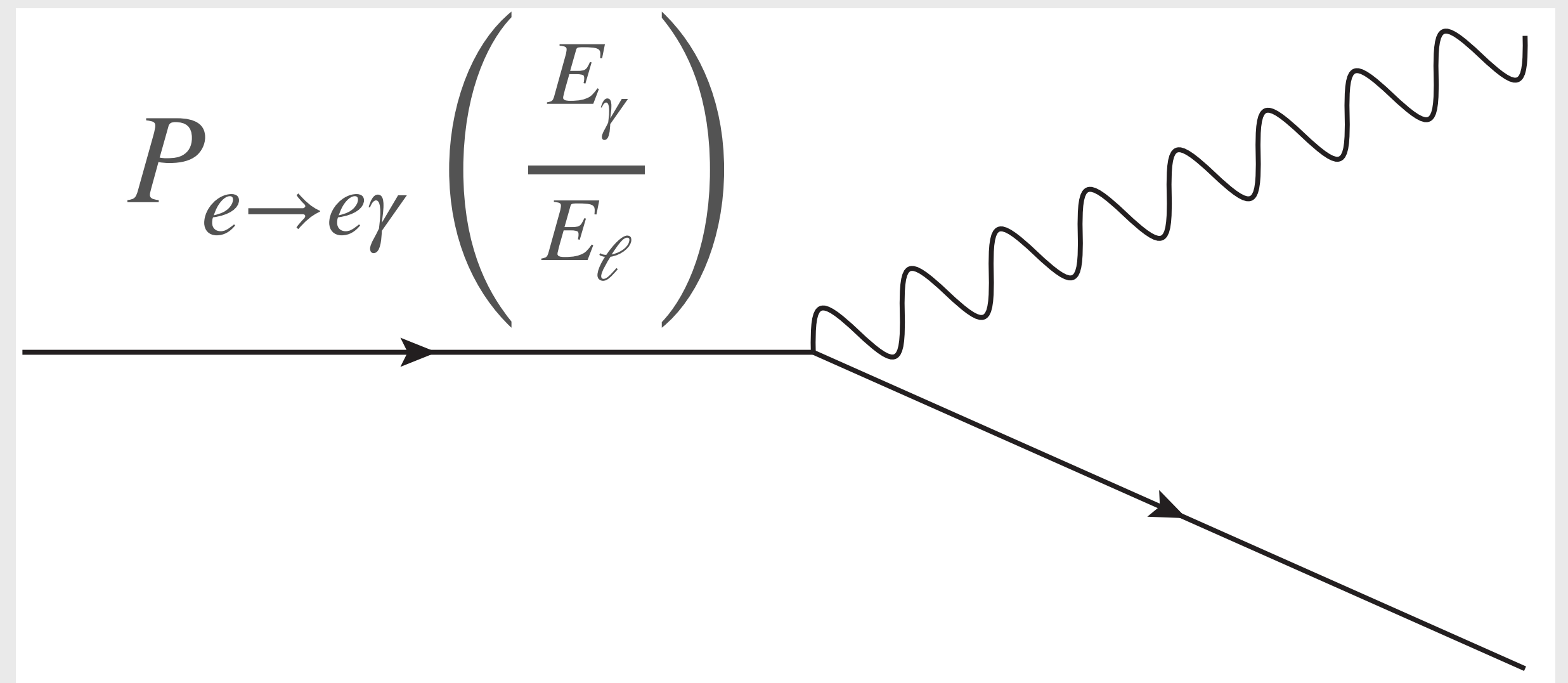
$$d\sigma = d\sigma^{(0)} + \left(d\sigma_R^{(1)} + d\sigma_V^{(1)} \right) + \dots$$

$$\int d\Pi \left(d\sigma_R^{(1)} + d\sigma_V^{(1)} \right) = \mathcal{O} \left(\frac{\alpha}{4\pi} \right)$$

IR AND COLLINEAR
DIVERGENCES ARE
RELATED

SPLITTING FUNCTIONS

- We can use a calculation of real radiation to predict logarithmically enhanced parts of virtual correction.
- Result is effectively classical.
- Think of electron as having a distribution of QED-partons.

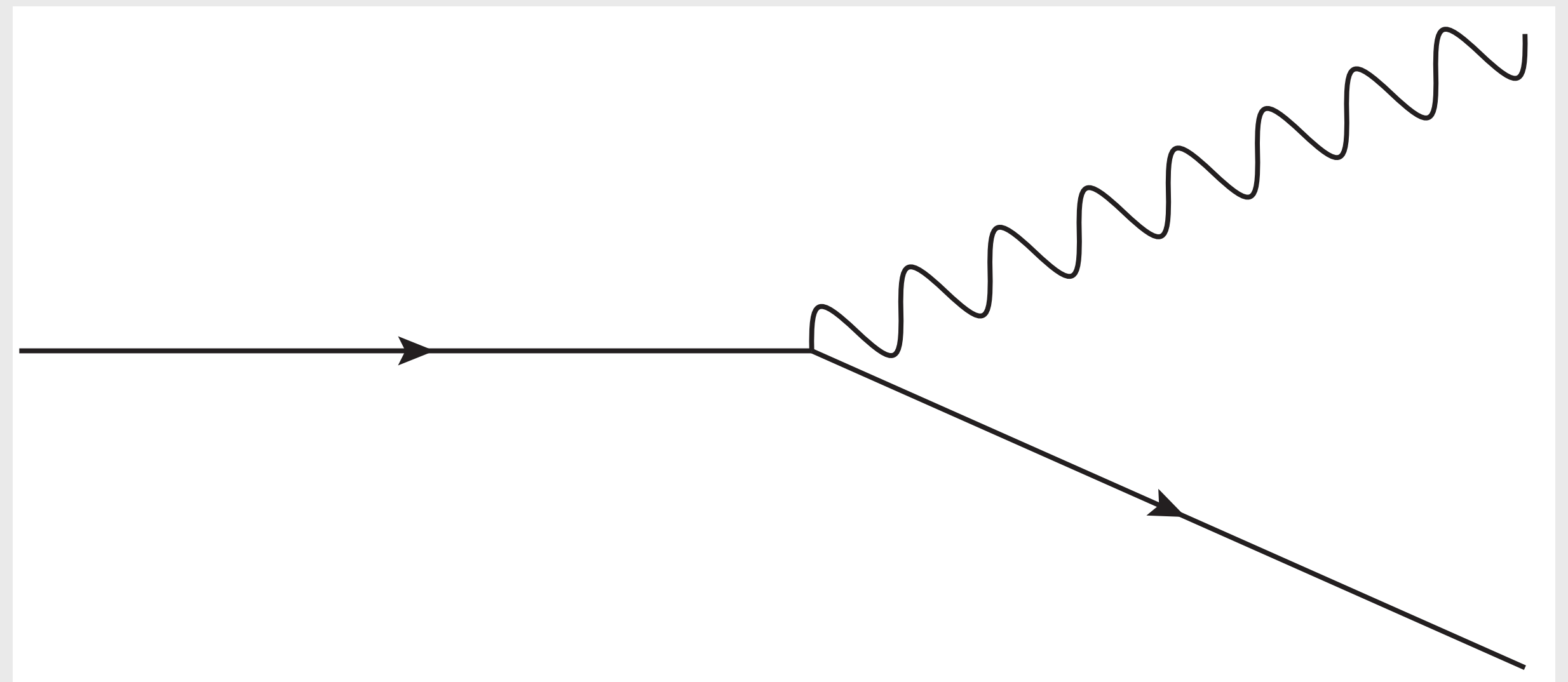


SPLITTING FUNCTIONS

- We can use a calculation of real radiation to predict logarithmically enhanced parts of virtual correction.

$$d\sigma_V^{(1)} \simeq - \int d\Pi_\gamma d\sigma_R^{(1)}$$

- This works for the largest parts of the corrections (double logs)

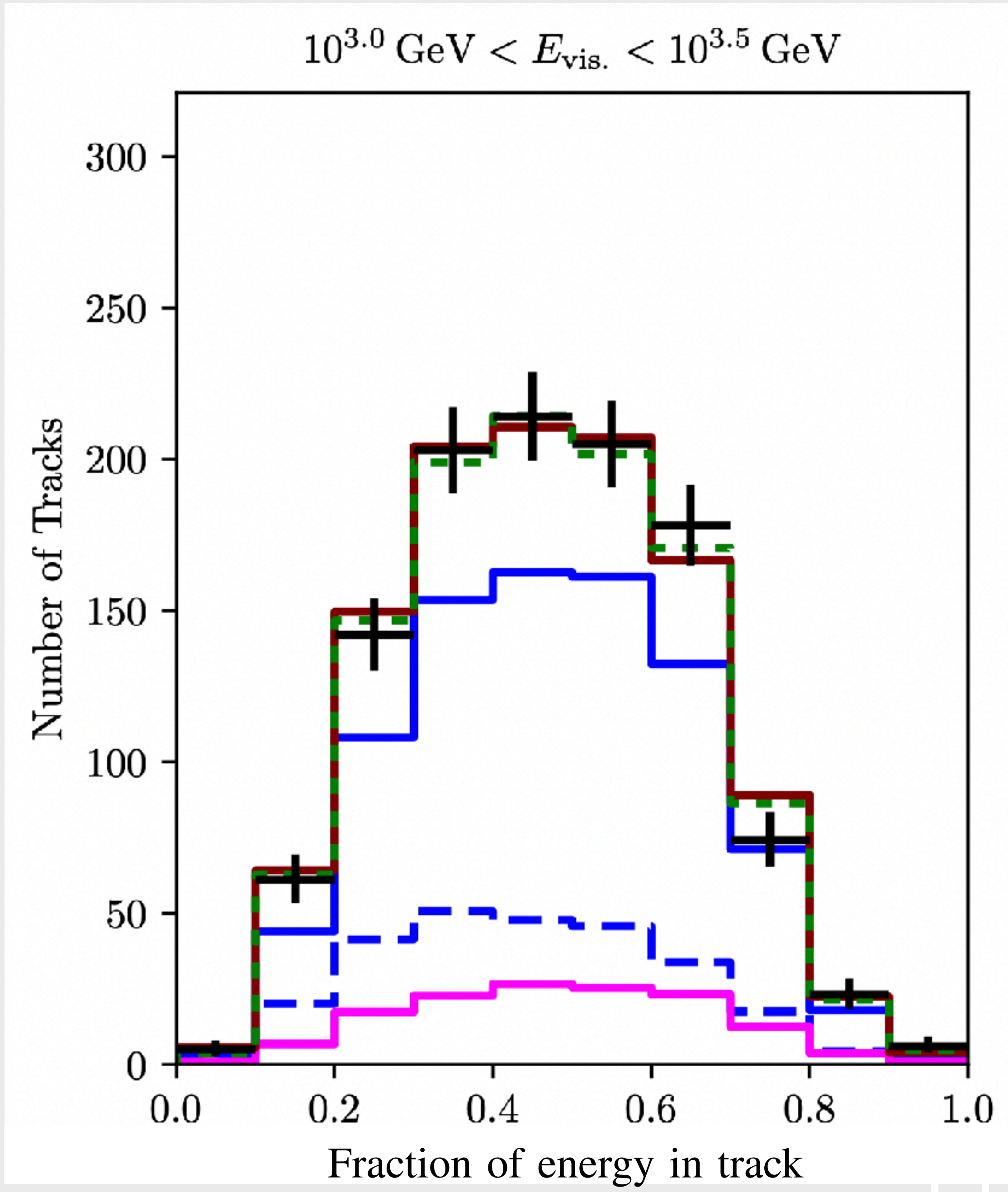
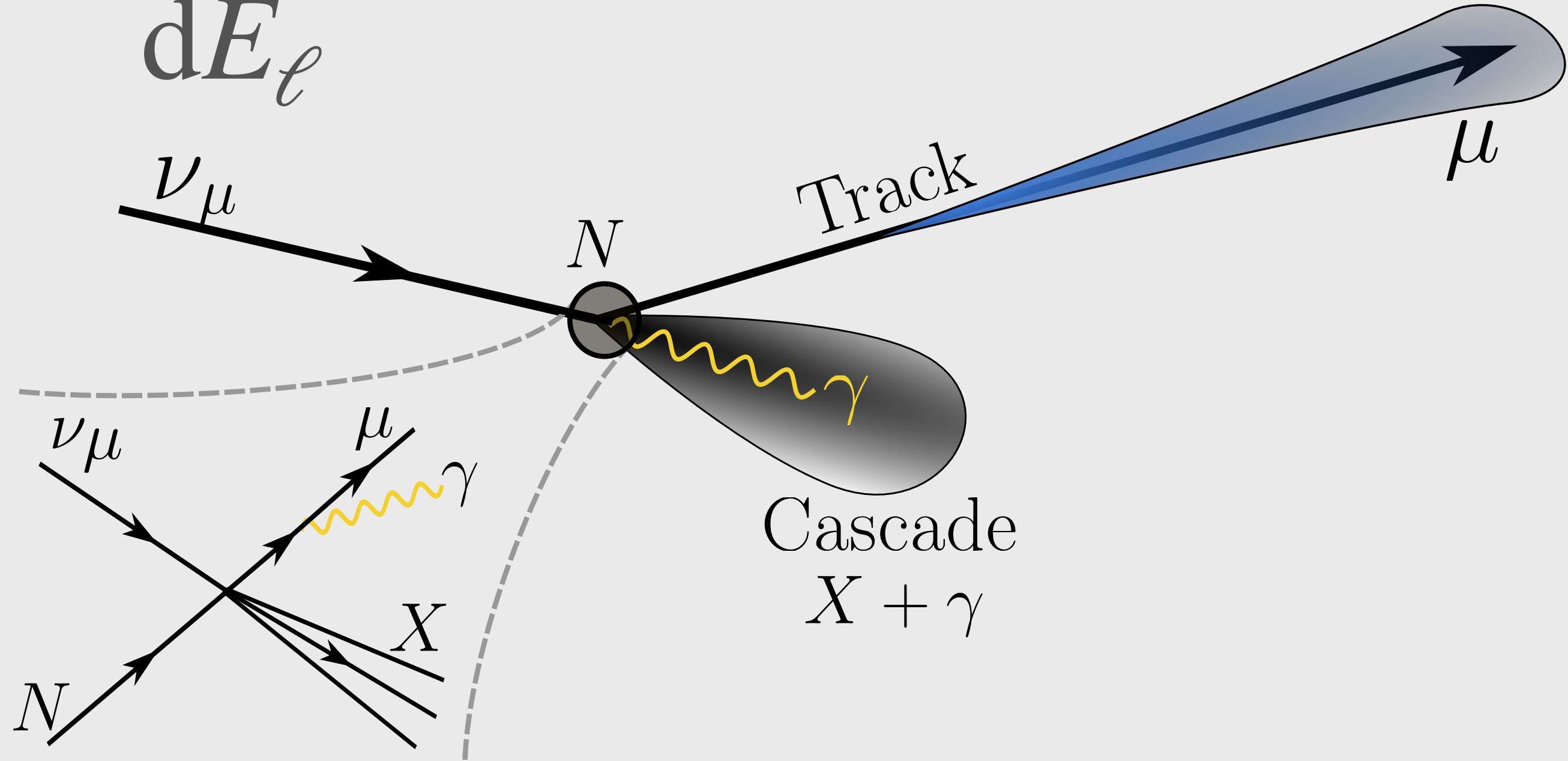


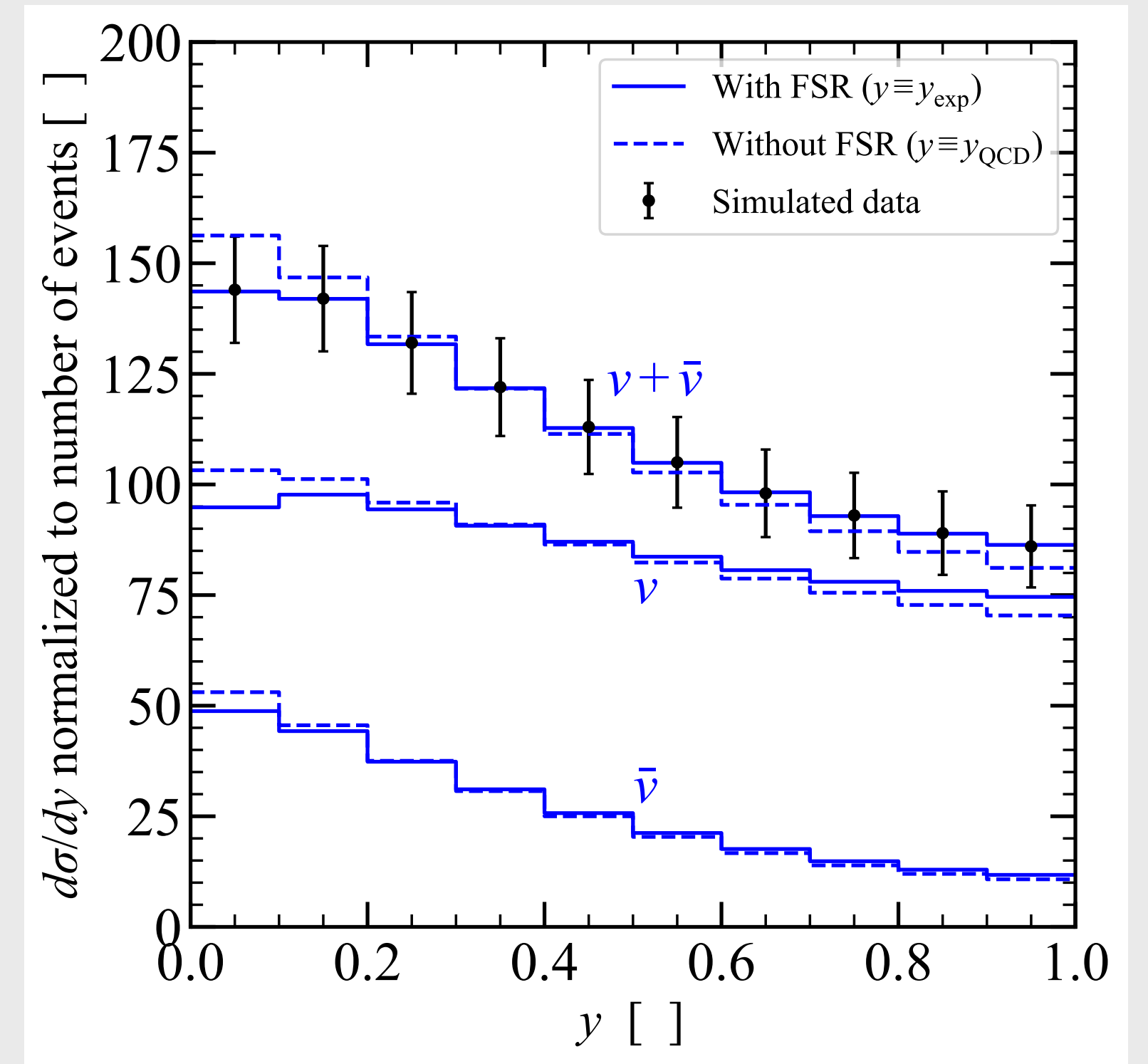
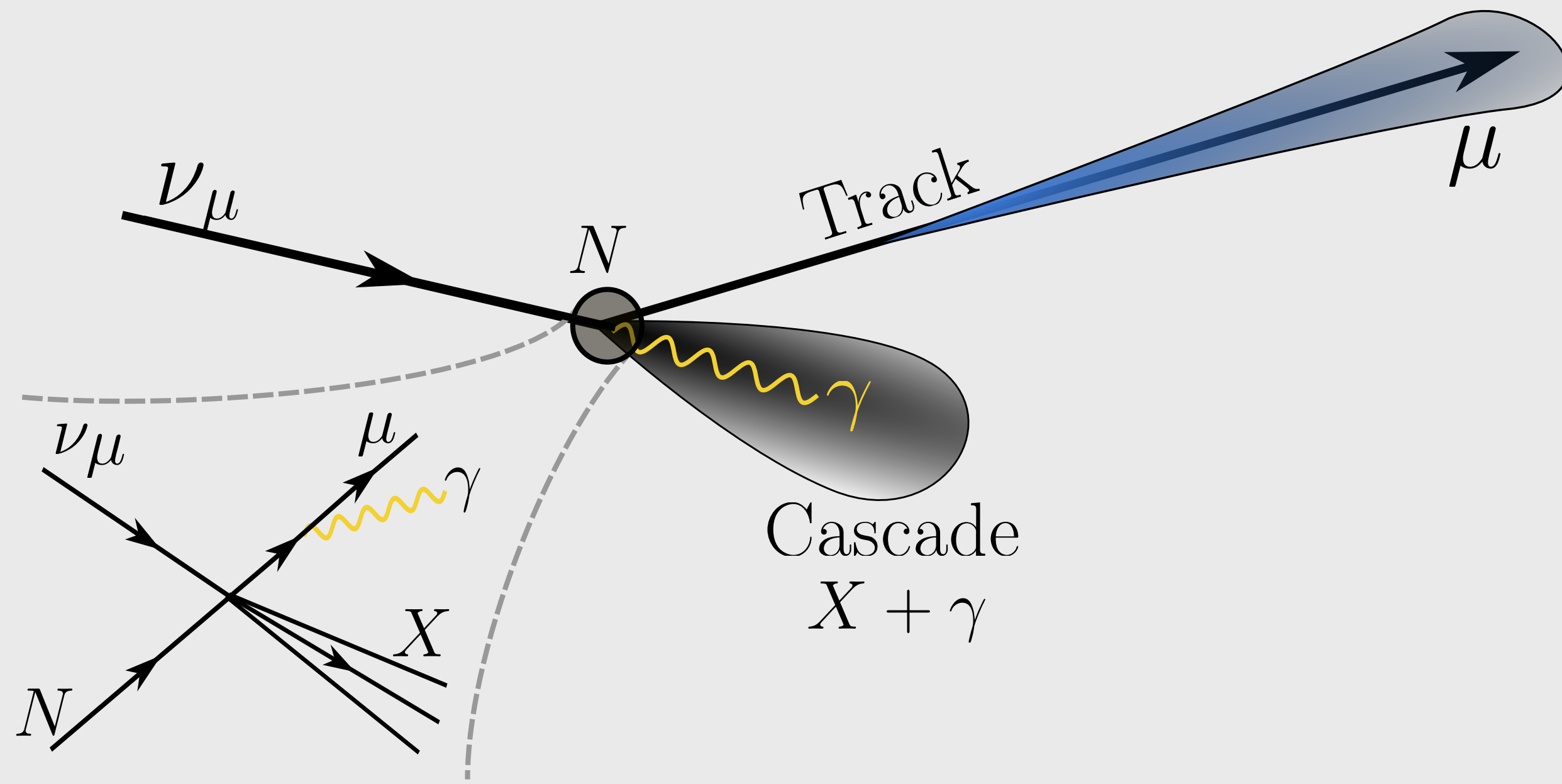
"TRUE" INELASTICITY: INCLUSIVE

$$\frac{d\sigma}{dy} \quad y = E_{\text{had}} / E_\nu$$

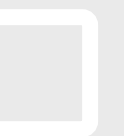
LEPTON ENERGY: EXCLUSIVE

$$\frac{d\sigma}{dE_\ell} \quad E_\nu = E_{\text{had}} + E_\ell + E_\gamma$$





Neutrino Applications

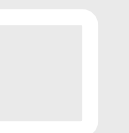


TOTAL CROSS SECTIONS

- The total cross section is by definition inclusive, up to very small bits of phase space that fall below detector threshold.
- KLN theorem guarantees that there are no kinematic logs.

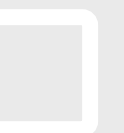
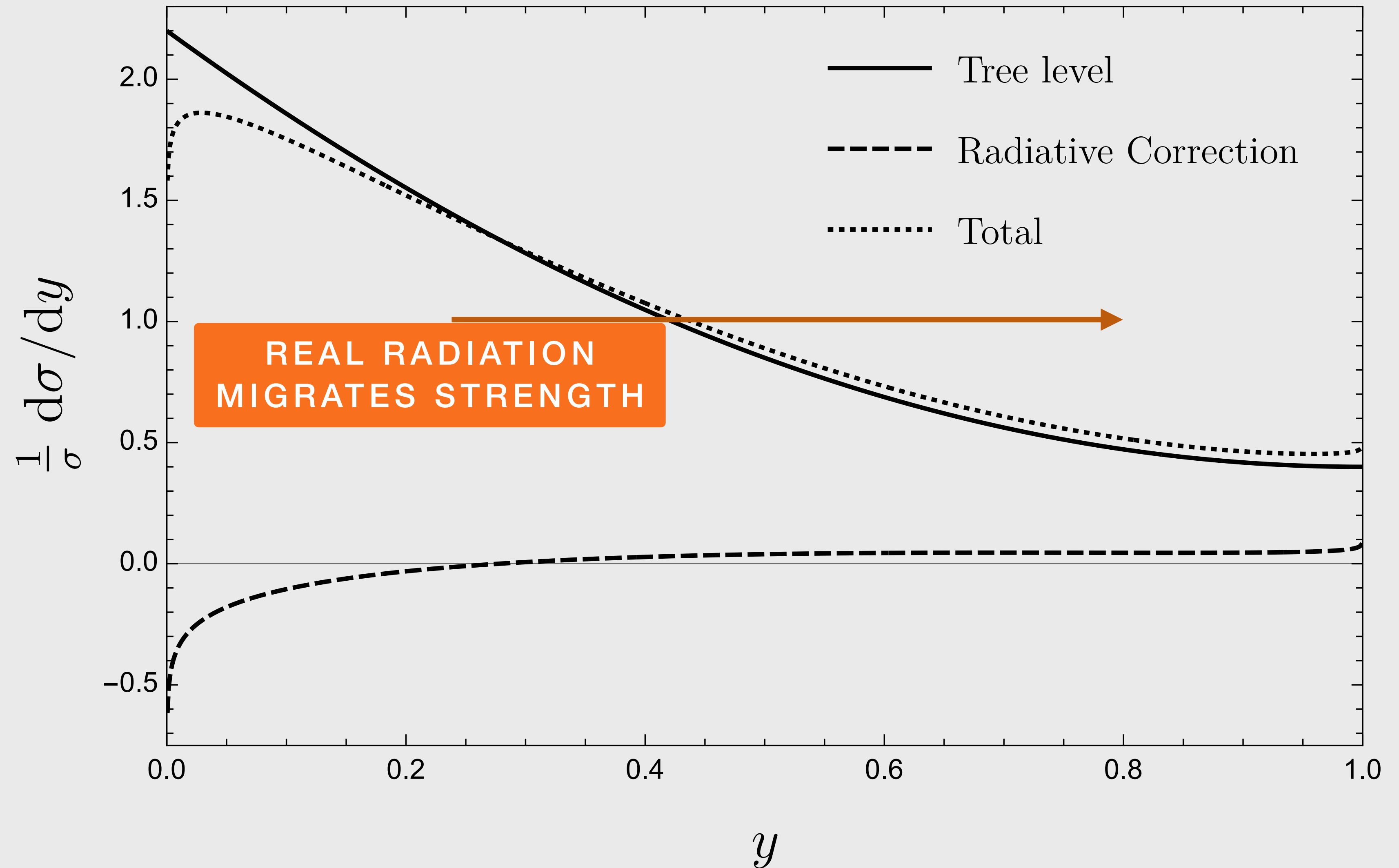
RADIATION OFF NUCLEONS

- If radiation comes from hadrons, then the "cascade" or "shower" topology is completely inclusive.
- We can therefore neglect QED radiative corrections of the nucleon lines.



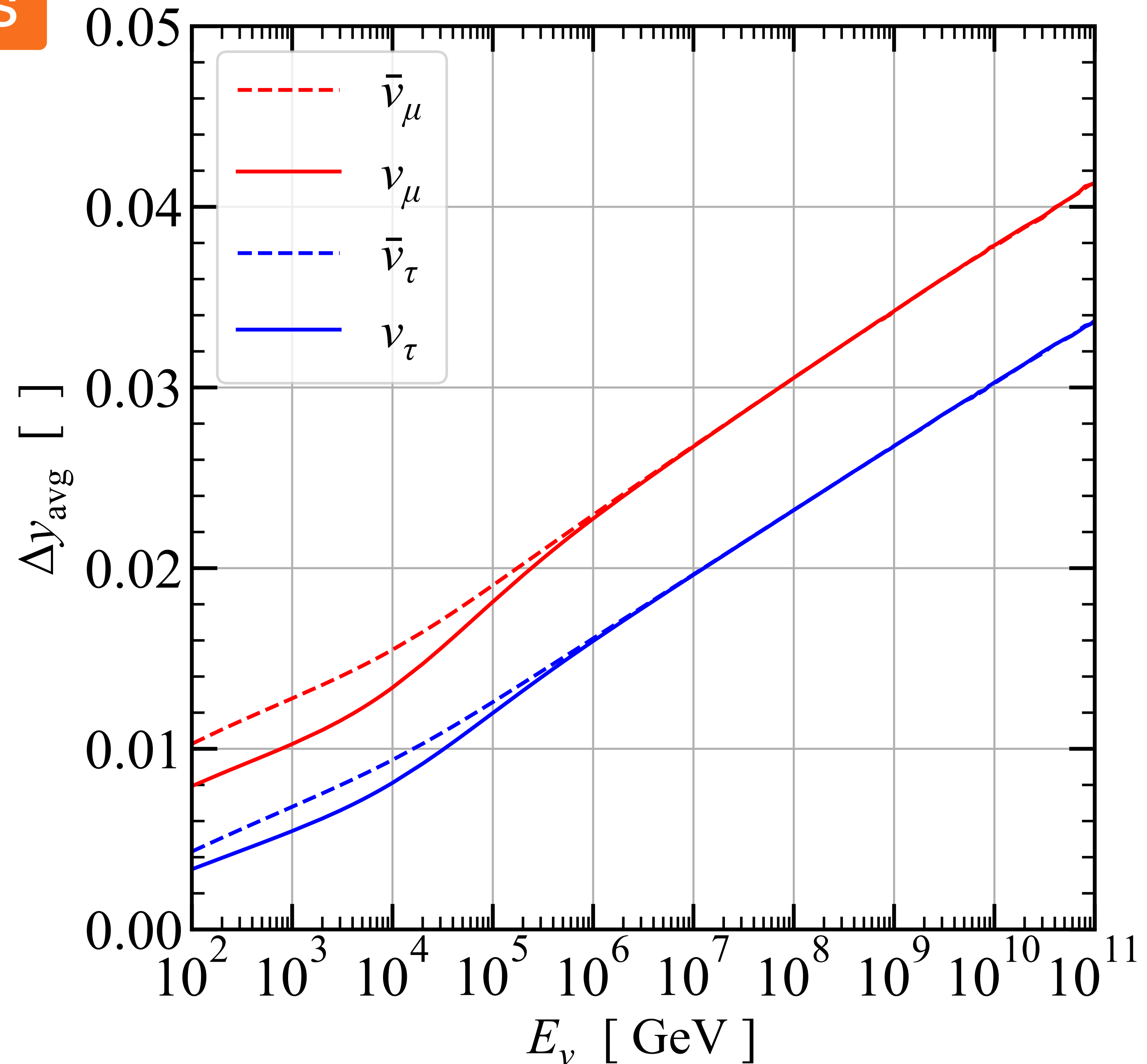
LEPTON ENERGY DISTRIBUTIONS: TOY DISTRIBUTION

- Emitting a photon shifts inelasticity strength.
- Systematically shifts $\langle y \rangle$ to larger values.



INELASTICITY DISTRIBUTIONS

- Shifts in inelasticity are modest
 $\delta\langle y \rangle \sim 0.03$
- Approximate size agrees with counting logarithms.

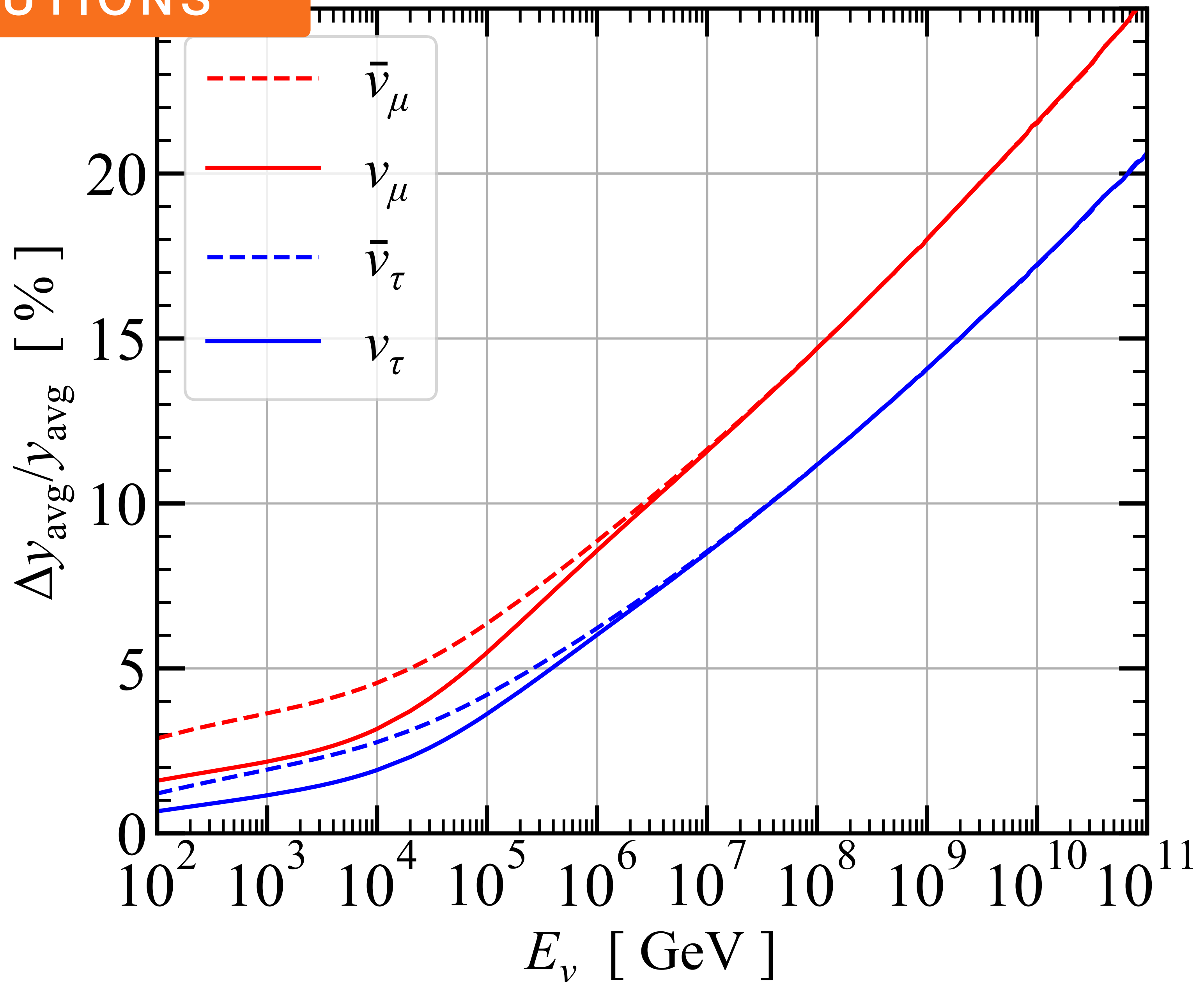


INELASTICITY DISTRIBUTIONS

- In relative terms the shift can be very large.

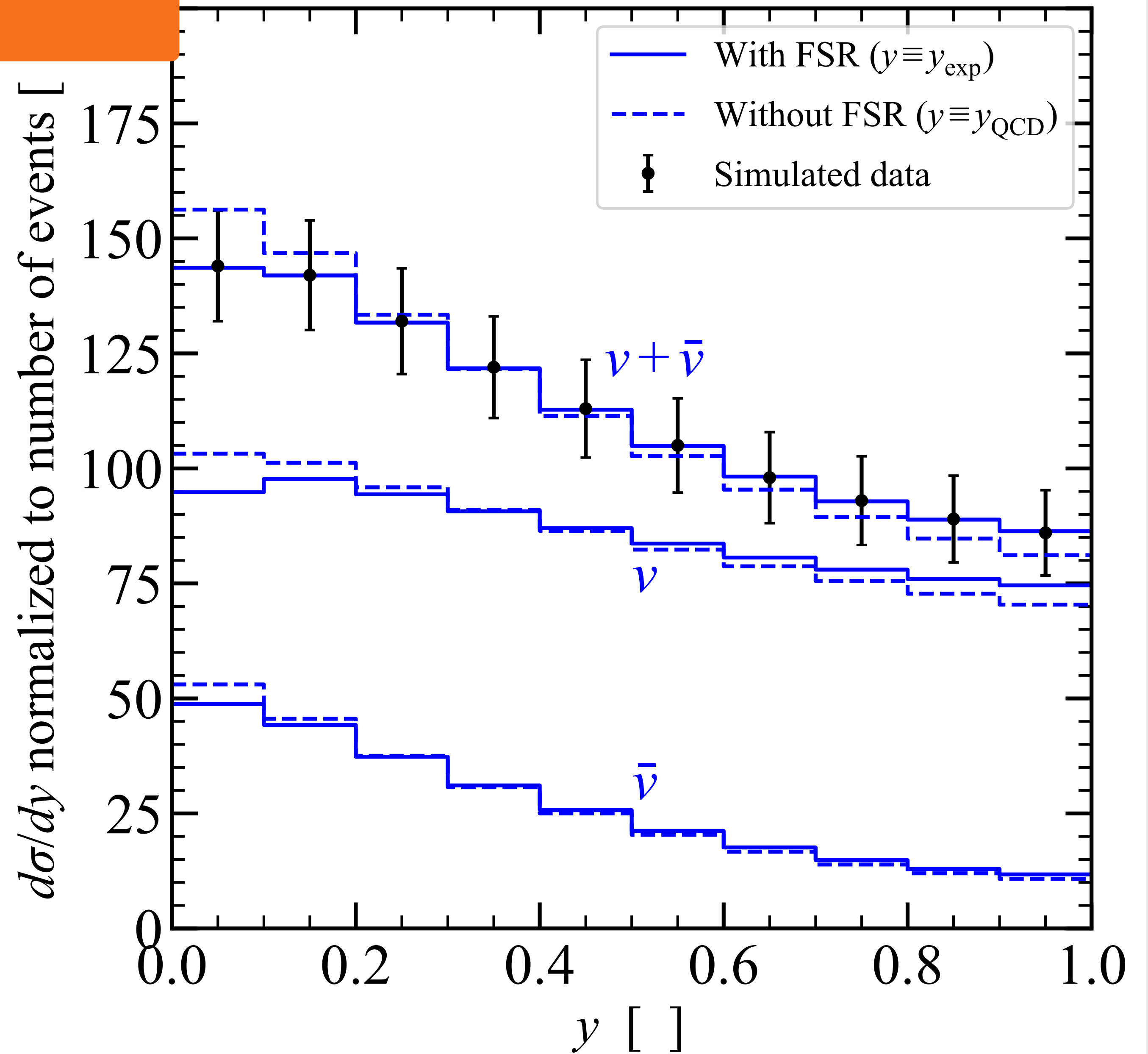
- This is because

$$E_\ell \sim 4 \times E_{\text{had}}$$



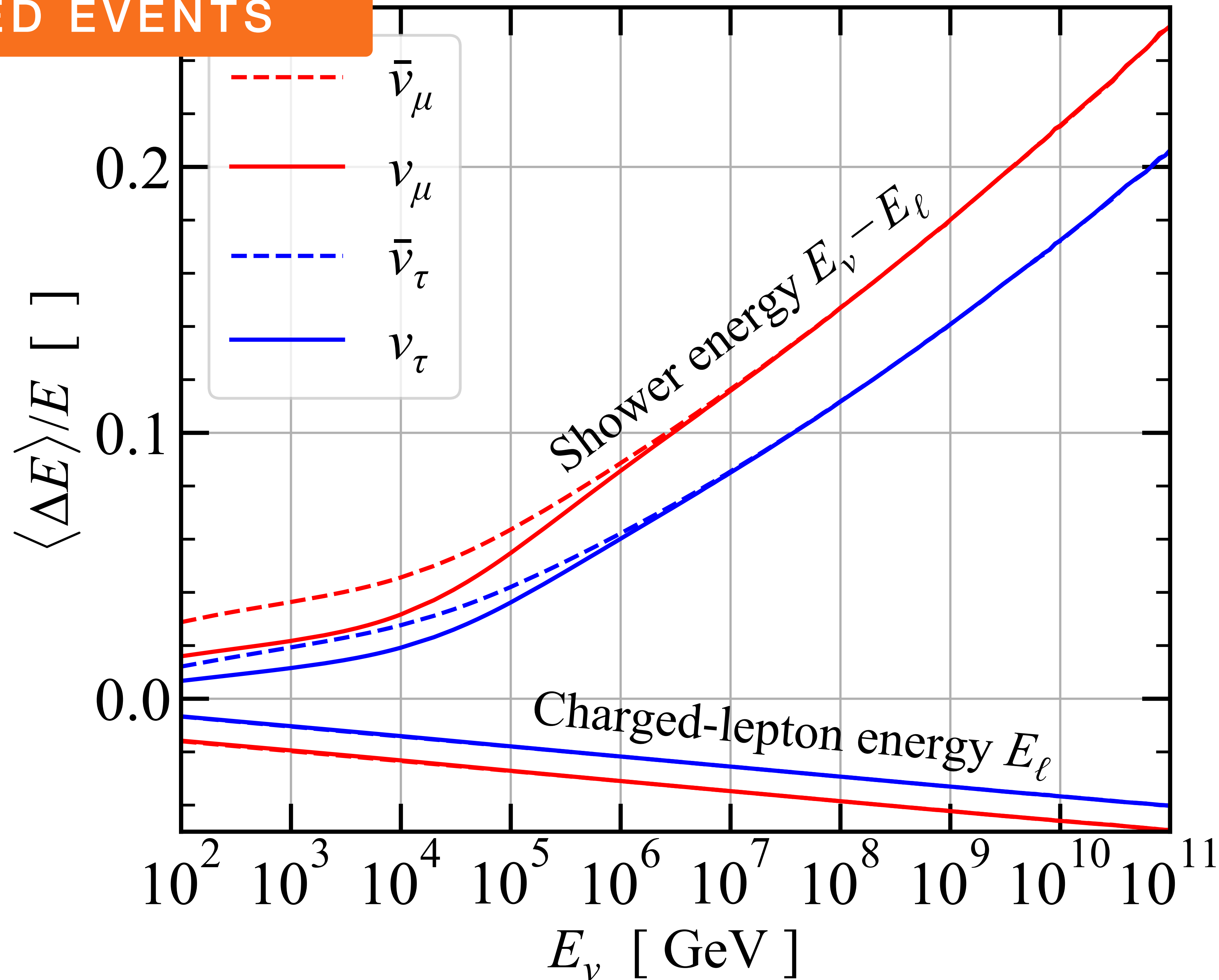
APPLICATION: $\nu/\bar{\nu}$ RATIO

- FSR distorts shape of $d\sigma/dy$
- Significant impact for some observables.
- E.g. $\nu/\bar{\nu}$ ratio.
- Sensitive to systematics in y .

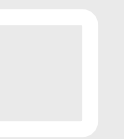
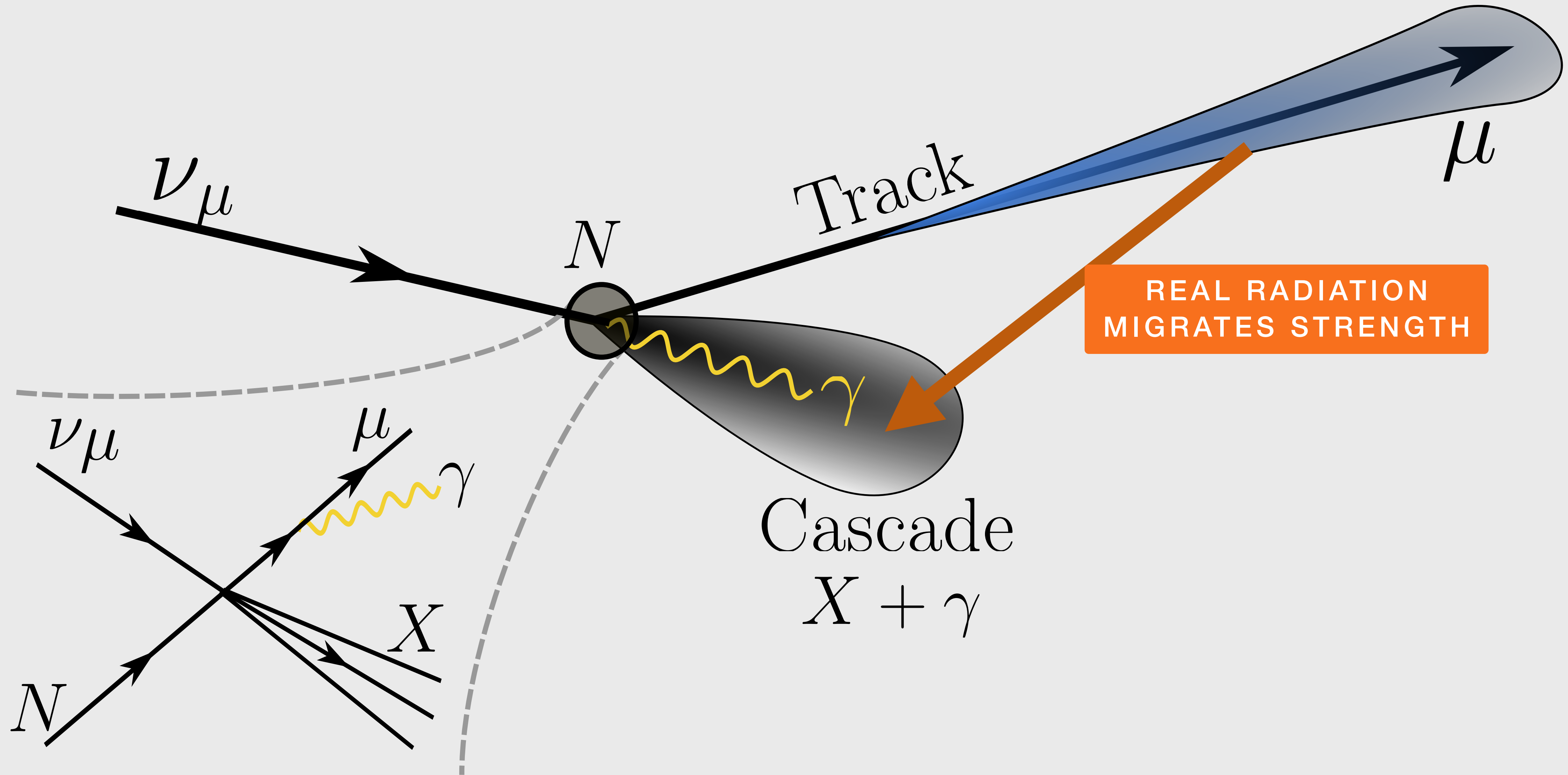


STARTING/CONTAINED EVENTS

- Re-balancing of shower vs track energy.
- Can effect energy estimators.

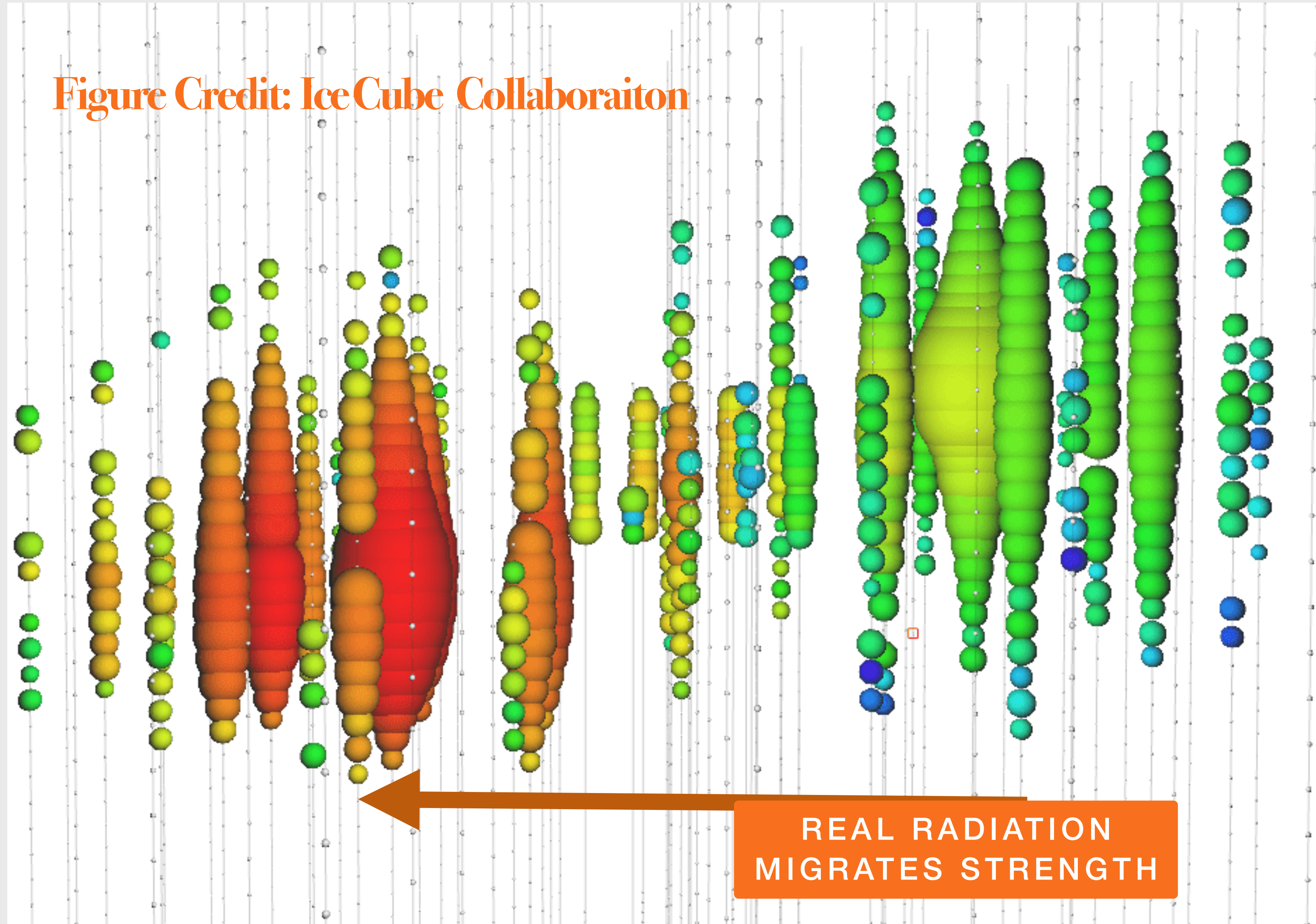


MUON TRACKS



DOUBLE BANG SIGNATURE

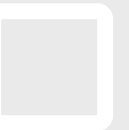
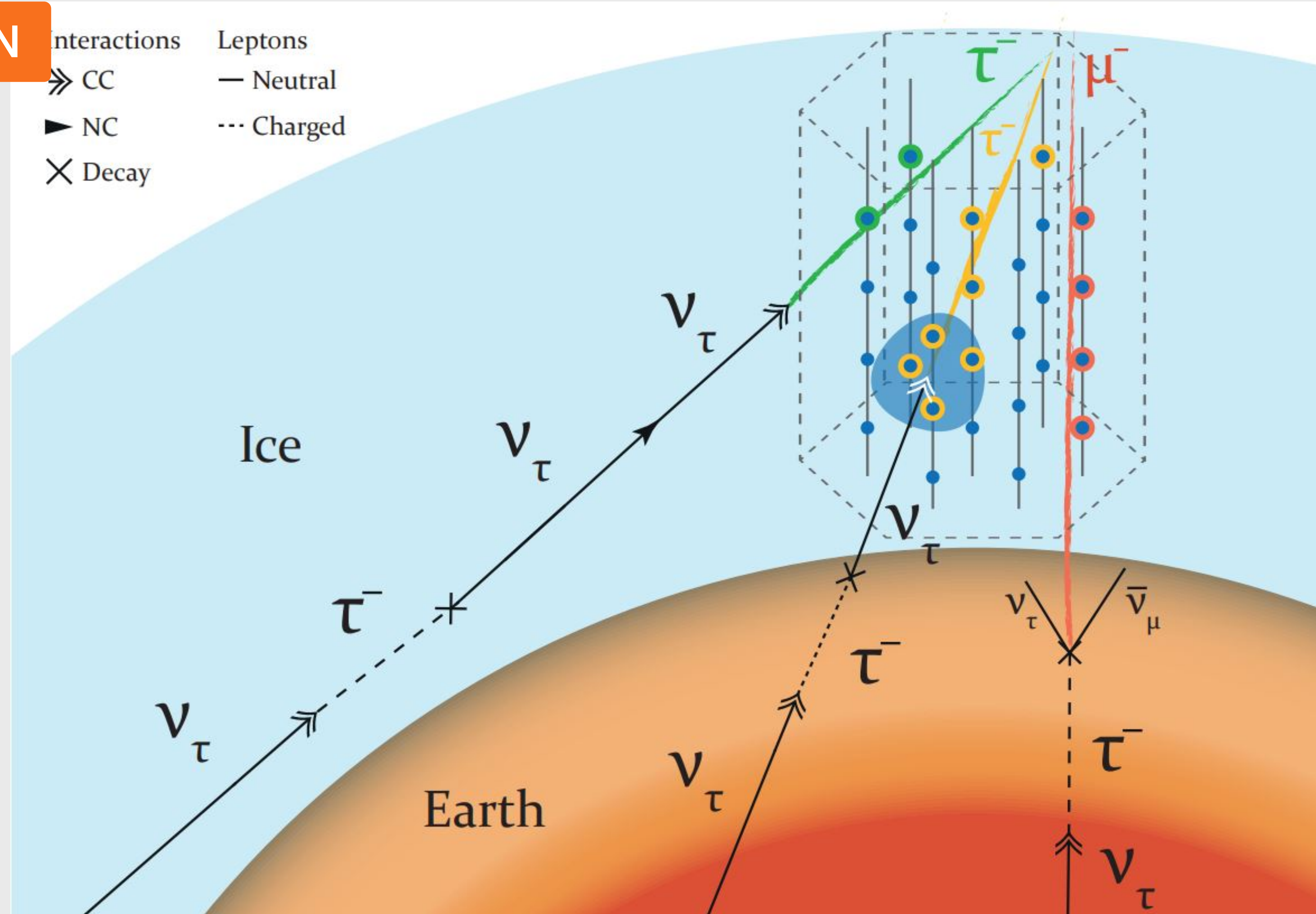
Figure Credit: IceCube Collaboraiton



- Reshuffles energy between bangs.
- First bang will have more energy.

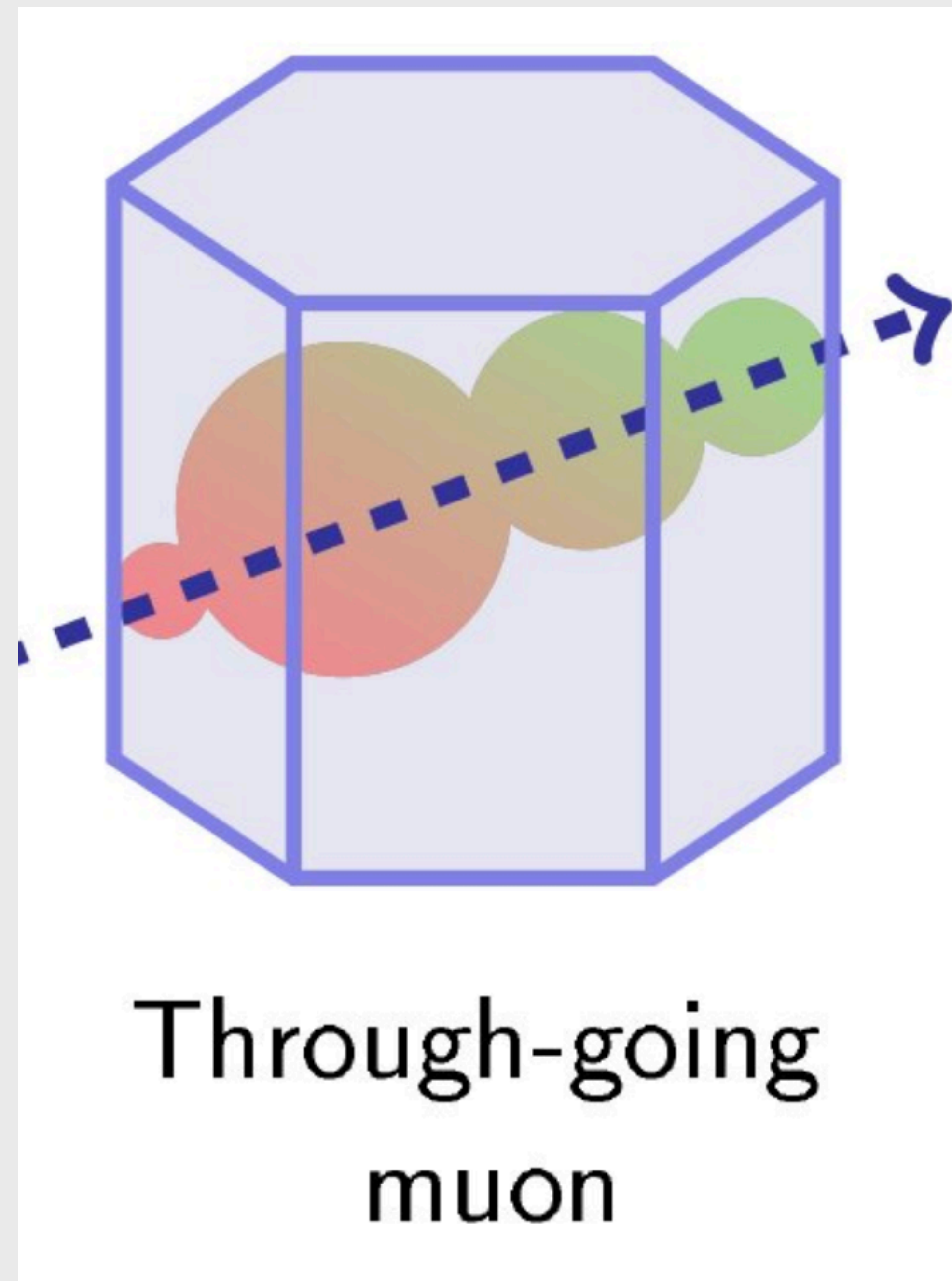
TAU REGENERATION

- For regenerated neutrinos FSR can "build up".
- Many interactions means more chances for FSR.



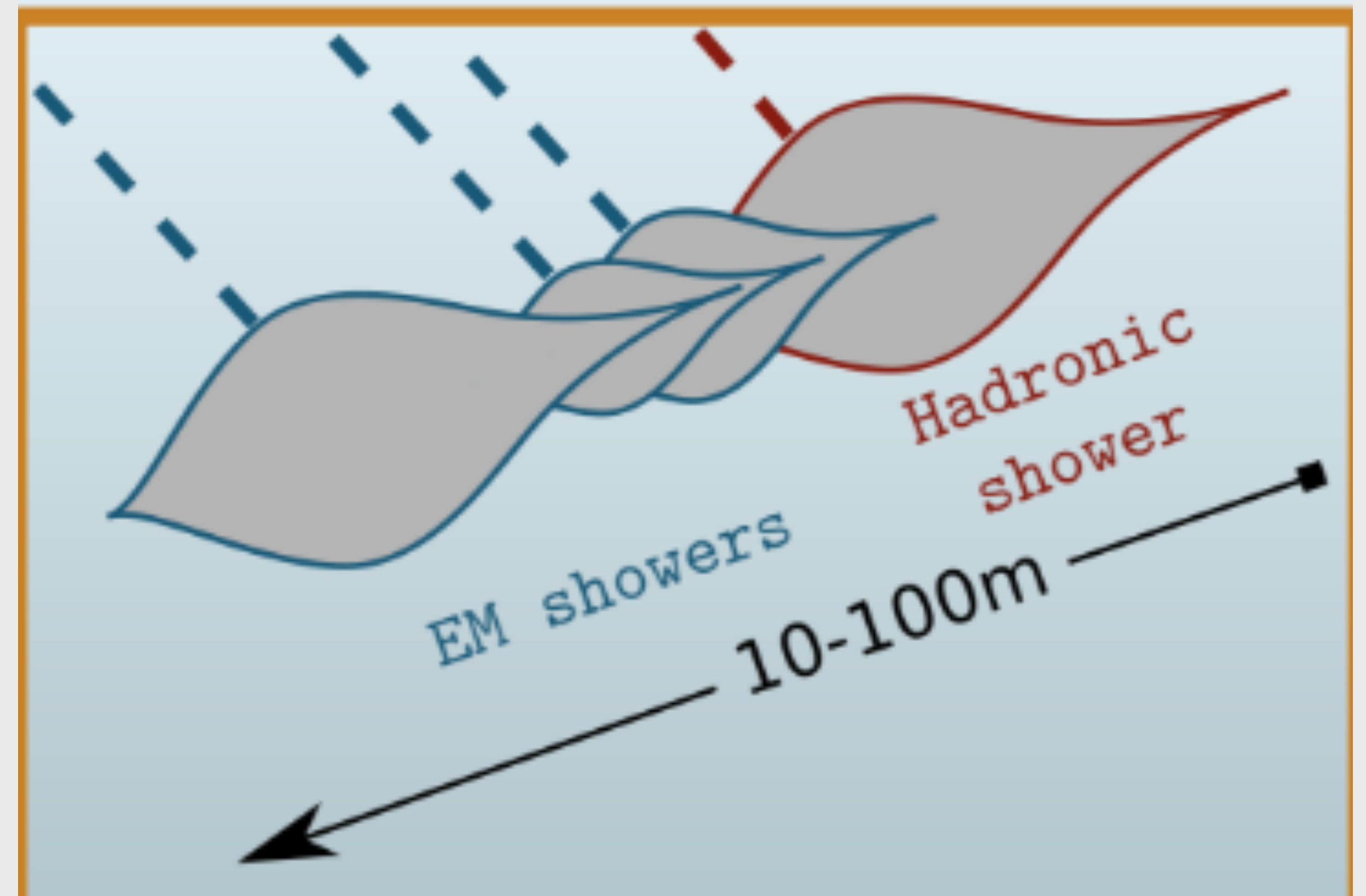
UHE NEUTRINOS

- External leptons can be necessary to increase effective area.
- FSR will be lost outside detector.
- Introduces $\sim 5\%$ bias in neutrino energy estimator.

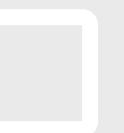


UHE NEUTRINOS

- Recent proposal to use LPM elongated showers to distinguish flavour.
- Relies on "subtle waveforms".
- FSR will distort these and should be included in templates.

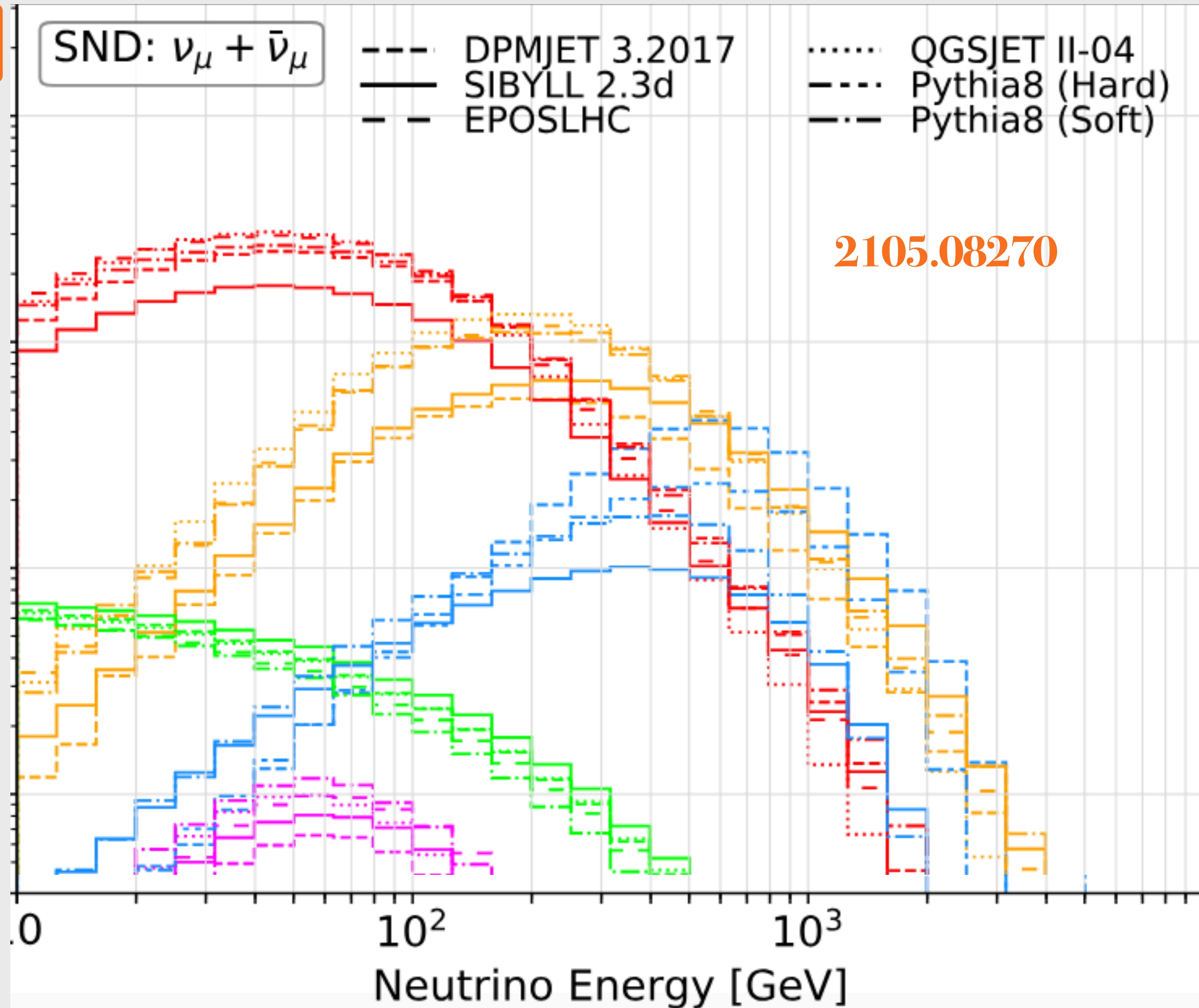


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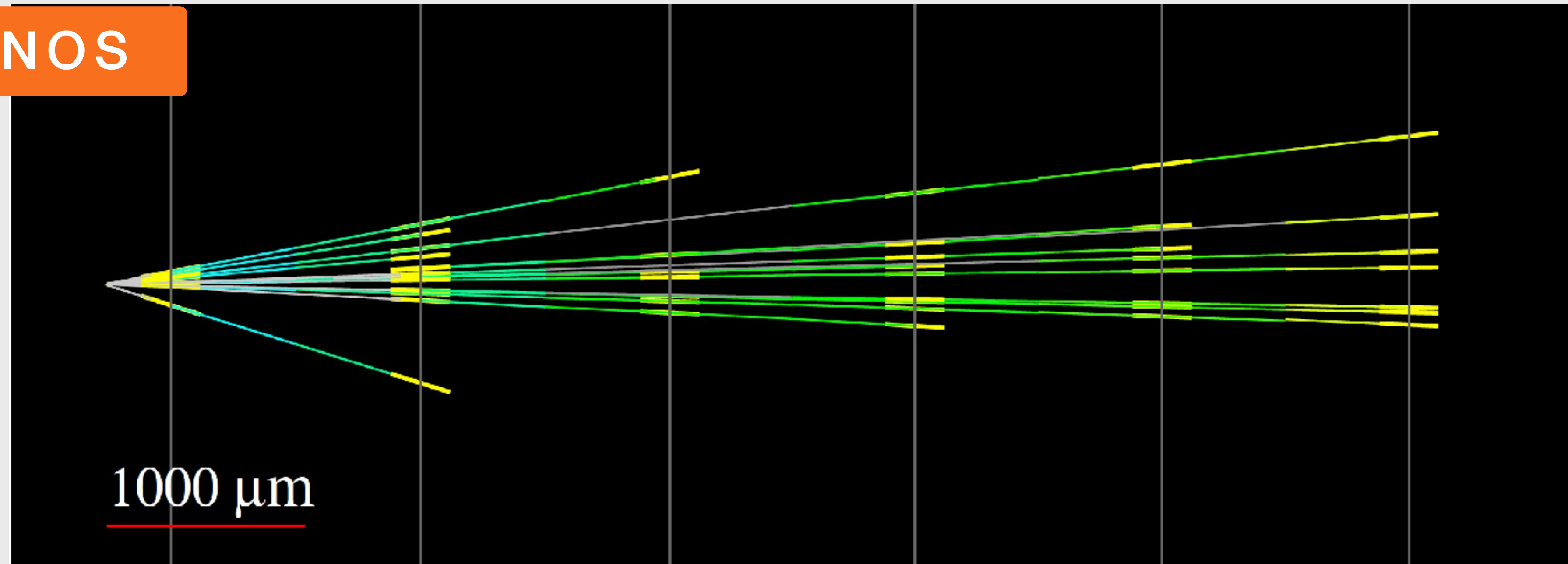


COLLIDER NEUTRINOS

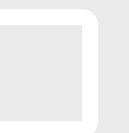
- Event rate peaks near TeV energies.
- Detectors are very different than neutrino telescopes.
- Warrants separate discussion.



COLLIDER NEUTRINOS



- Photon still gets absorbed into shower energy (hard to distinguish from $\pi^0 \rightarrow \gamma\gamma$).
- Electrons may be identifiable with primary vertex ID.

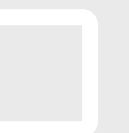


KINEMATICS

- Reconstructed DIS variables get distorted.

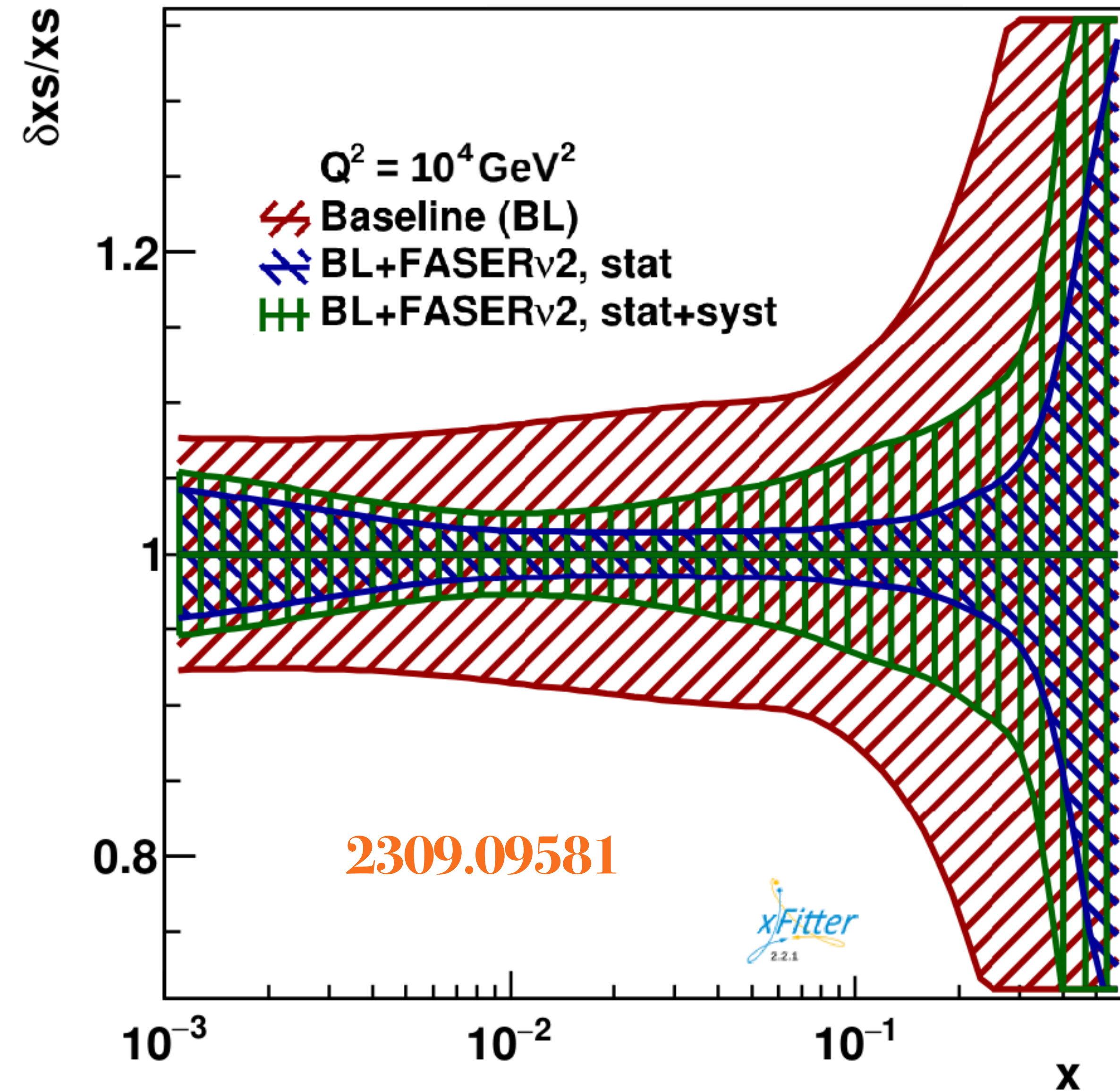
$$[\Delta Q^2]_{\text{FSR}} \simeq -4E_\nu E_\gamma \sin^2(\theta_\ell/2)$$

$$[\Delta x]_{\text{FSR}} \simeq \frac{[\Delta Q^2]_{\text{FSR}}}{2m_N E_X} - \frac{E_\gamma}{E_X} x^{(0)}$$



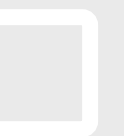
IMPACT ON PDF EXTRACTIONS

- LHC neutrinos can supply best constraints on certain PDFs.
- Specifically strange and quark singlet PDFs.
- Will be important to include radiative corrections for these.





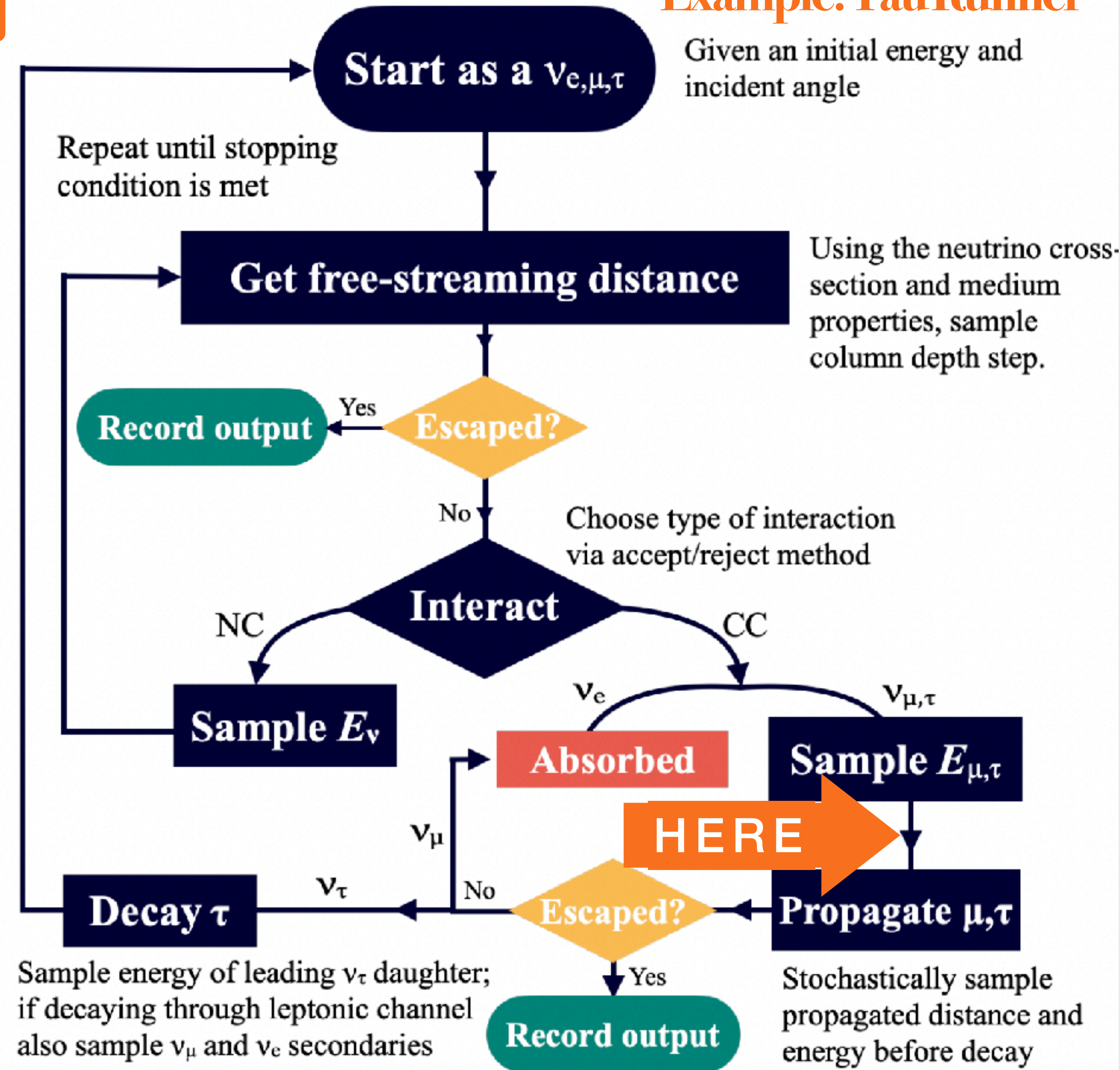
Monte Carlo Implementation



HOW TO IMPLEMENT FSR

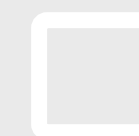
- At leading-log accuracy process is essentially classical.
- Can generate lepton "as is".
- Perform a final step where energy fraction to photon is sampled.

Example: Tau Runner





Conclusions & Outlook



Conclusions

- Final state radiation can be enhanced by large kinematic logarithms.
- Effects are $\sim 10\%$ in size when lepton energy is measured.
- Influences reconstruction of kinematic variables and estimates of neutrino energy

