



Derivation of diffusion coefficients for the transport of charged particles in magnetic fields

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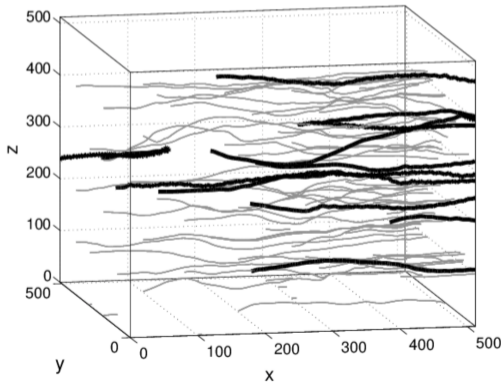
Cosmic-ray diffusion in turbulent magnetic field



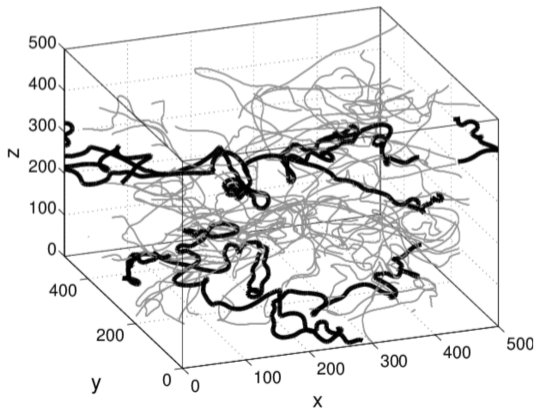
Cosmic-ray diffusion in turbulent magnetic field

- $D_{ij}(t) \simeq \langle x_i x_j \rangle / \Delta t$
- MC simulations

Credits: Xu and Yan, ApJ 2013



(a)



(b)

- Spatial diffusion-tensor coefficients related to the velocity correlation function through a time integration [Taylor (1922); Green (1951); Kubo(1957)]:

$$D_{ij}(t) = \int_0^t dt' \langle v_{0i} v_j(t') \rangle$$

- Aim: calculate $\langle v_{0i} v_j(t) \rangle$ given any turbulence $\delta \mathbf{B}$ on top of a mean field \mathbf{B}_0
- Equation of motion: $\partial v_i(t)/\partial t = \delta \Omega \epsilon_{ijk} v_j(t) \delta b_k(t) + \Omega_0 \epsilon_{ijk} v_j(t) \mathbf{b}_{0k}$
 - $\delta \Omega = c^2 Z |e| \delta B / E$: turbulent-related gyrofrequency
 - $\Omega_0 = c^2 Z |e| B_0 / E$: mean field-related gyrofrequency
 - $\delta b_k(t) \equiv \delta b_k(\mathbf{x}(t))$: k -th component of the magnetic field, expressed in units of δB , at the spatial coordinate $\mathbf{x}(t)$ of the particle at time t

Equation of motion – Auxiliary variable

- Mean field oriented in the z direction, $\mathbf{B}_0 = B_0 \mathbf{u}_z$
- Equation of motion: $\partial v_i(t)/\partial t = \delta\Omega \epsilon_{ijk} v_j(t) \delta b_k(t) + \Omega_0 \epsilon_{ijk} v_j(t) \delta_{kz}$
- Expansion: infinite number of Dyson series mixing powers of δB and B_0
- Alternative: auxiliary variable $v_i(t) = R_{ij}^{-1}(\Omega_0 t) w_j(t)$, with R_{ij} the matrix elements of the rotation matrix around Z [Plotnikov, Pelletier, & Lemoine, A&A 532 (2011) A68]
- Given that $\partial R_{ij}^{-1}(\Omega_0 t)/\partial t = \Omega_0 \epsilon_{ikz} R_{kj}^{-1}(\Omega_0 t)$,

$$\frac{\partial w_i(t)}{\partial t} = \delta\Omega R_{im}(\Omega_0 t) \epsilon_{mnk} \delta b_k R_{nj}^{-1}(\Omega_0 t) w_j(t)$$

- No additional input w.r.t. the pure turbulence case

Dyson series

- First iterative solution:

$$\langle w_i(t) \rangle = w_{i_0} + \delta\Omega \epsilon_{mnk} \int_0^t dt' R_{im}(\Omega_0 t') \langle w_j(t') \delta b_k(t') \rangle R_{nj}^{-1}(\Omega_0 t')$$

- Dyson series:

$$\langle v_{i_0}(t) \rangle = v_{0i_0} + \sum_{n=1}^{\infty} \delta\Omega^n \epsilon_{i_0 i_1 j_1} \epsilon_{i_1 i_2 j_2} \dots \epsilon_{i_{n-1} i_n j_n} v_{0i_n} \int_0^t dt_1 \int_0^{t_1} dt_2 \dots \int_0^{t_{n-1}} dt_n \langle \delta b_{j_1}(t_1) \dots \delta b_{j_n}(t_n) \rangle$$

- Wick theorem: expectation value in the integrand expressed in terms of all possible permutations of products of contractions of pairs of $\langle \delta b_{i_1}(t_{j_1}) \delta b_{i_2}(t_{j_2}) \rangle$
- Using the Ansatz $\langle \delta b_{i_1}(t_{j_1}) \delta b_{i_2}(t_{j_2}) \rangle = \varphi(t_{j_1} - t_{j_2}) \delta_{i_1 i_2/3}$:

$$\langle w_{i_0}(t) \rangle = w_{0i_0} + \sum_{p=1}^{\infty} \left(\frac{\delta\Omega^2}{3} \right)^p w_{0i_{2p}} \int_0^t dt_1 \int_0^{t_1} dt_2 \dots \int_0^{t_{2p-1}} dt_{2p} \sum_{\text{pairings } j < \ell} \prod \varphi(t_j - t_\ell)$$

$$\left(R_{i_{j-1}k_j}(\Omega_0 t_j) R_{i_{\ell-1}k_\ell}(\Omega_0 t_\ell) R_{m_j j}^{-1}(\Omega_0 t_j) R_{m_\ell \ell}^{-1}(\Omega_0 t_\ell) - R_{i_{j-1}k_j}(\Omega_0 t_j) R_{i_{\ell-1}k_\ell}(\Omega_0 t_\ell) R_{k_\ell j}^{-1}(\Omega_0 t_j) R_{k_j \ell}^{-1}(\Omega_0 t_\ell) \right)$$

Partial summation of Dyson series

- Include all possible ways to have one nested loop in unconnected diagrams
(Kraichnan propagator [Kraichnan, JMP 2 (1961) 124]):

$$\overset{\text{wavy}}{\text{---}}_{0,j} \quad t,i = \text{---}_{0,j} \quad t,i + \overset{\text{dashed}}{\text{---}}_{0,j} \quad t_2 \quad t_1 \quad t,i$$

- Include all possible nested and crossed diagrams:

$$\text{====}_{0,j} \quad t,i \approx \text{---}_{0,j} \quad t,i + \overset{\text{dashed}}{\text{---}}_{0,j} \quad t_2 \quad t_1 \quad t,i + \overset{\text{dashed}}{\text{---}}_{0,j} \quad t_4 \quad t_3 \quad t_2 \quad t_1 \quad t,i,$$

- Approximation:

$$\text{====}_{0,j} \quad t,i \approx \text{---}_{0,j} \quad t,i + \overset{\text{dashed}}{\text{---}}_{0,j} \quad t_2 \quad t_1 \quad t,i + \overset{\text{dashed}}{\text{---}}_{0,j} \quad t_4 \quad t_3 \quad t_2 \quad t_1 \quad t,i$$

2-pt correlation function of the experienced magnetic field

- Formal expression:

$$\langle \delta b_i(t) \delta b_j(0) \rangle \simeq \iint dk dk' \langle \delta b_i(k) \delta b_j(k') \rangle \langle e^{ik \cdot x(t)} \rangle$$

- Expansion of $\langle e^{ik \cdot x(t)} \rangle$ as a Dyson series:

$$\langle e^{ik \cdot x(t)} \rangle = \sum_{n \geq 0} \frac{i^n}{n!} \int_0^t dt_1 \cdots \int_0^t dt_n \langle (k \cdot v(t_1)) \cdots (k \cdot v(t_n)) \rangle$$

- n -point correlation function substituted for the sum of all possible contractions of pairs:

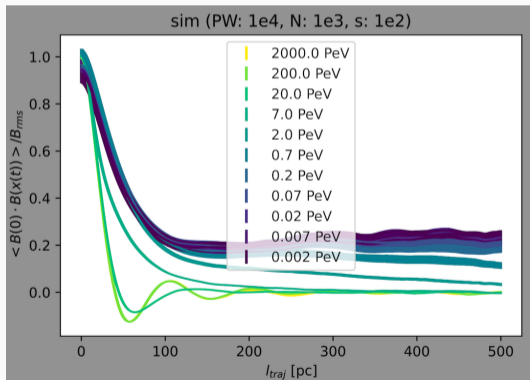
$$\langle (k \cdot v(t_1)) (k \cdot v(t_2)) \rangle \simeq \langle k^2 \rangle \langle v(t_1) \cdot v(t_2) \rangle \rightarrow k^2 c^2 e^{-(t_1 - t_2)/\xi(k)}$$

- Hence, with the heuristic expression $\xi(k) = \sqrt{\rho}/kc$,

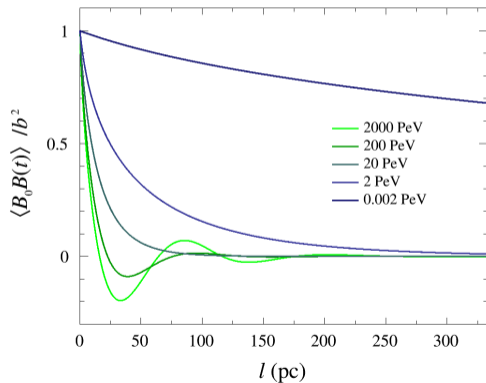
$$\langle e^{ik \cdot x(t)} \rangle_0 \simeq 1 - (kc)^2 \int_0^t dt_1 \int_0^{t_1} dt_2 e^{-(t_1 - t_2)/\xi(k)} \langle e^{ik \cdot x(t - t_1)} \rangle_0$$

2-pt correlation function of the experienced magnetic field

- MC simulations
- Numerous features not accounted for in various approaches (e.g. QLT)



- Sum of unconnected diagrams
- Main features reproduced in gyro-resonant and high-rigidity regimes



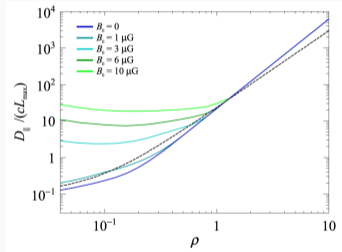
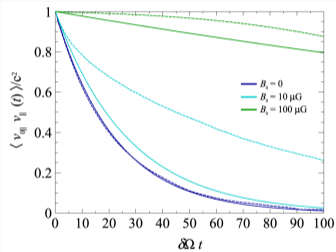
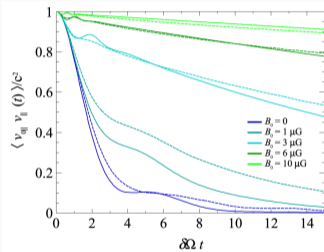
- First iteration:

$$W_{0z}(s) = \frac{1}{s} - \frac{2\delta\Omega^2}{3} \frac{W_{0z}(s)}{s} \mathcal{L} [\varphi(x) \cos \Omega_0 x] (s),$$

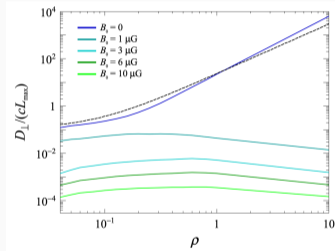
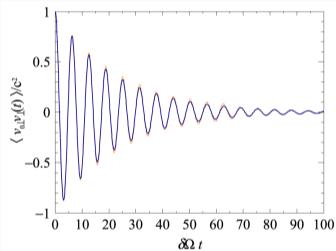
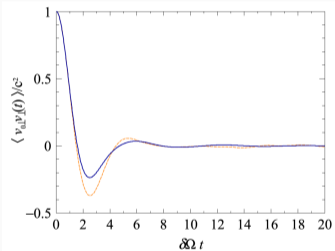
- Second iteration:

$$\begin{aligned} W_z(s) = & \frac{1}{s} - \frac{2\delta\Omega^2}{3} \frac{W_z(s)}{s} \mathcal{L} [\varphi(x) \langle w_z(x) \rangle_0 \cos \Omega_0 x] (s) \\ & + 2 \left(\frac{\delta\Omega^2}{3} \right)^2 \frac{W_z(s)}{s} \mathcal{L} \left[\varphi(x_1 + x_2) \varphi(x_2 + x_3) \right. \\ & \left. \langle w_z(x_1) \rangle_0 \langle w_z(x_2) \rangle_0 \langle w_z(x_3) \rangle_0 \cos \Omega_0 (x_1 - x_3) \right] (s). \end{aligned}$$

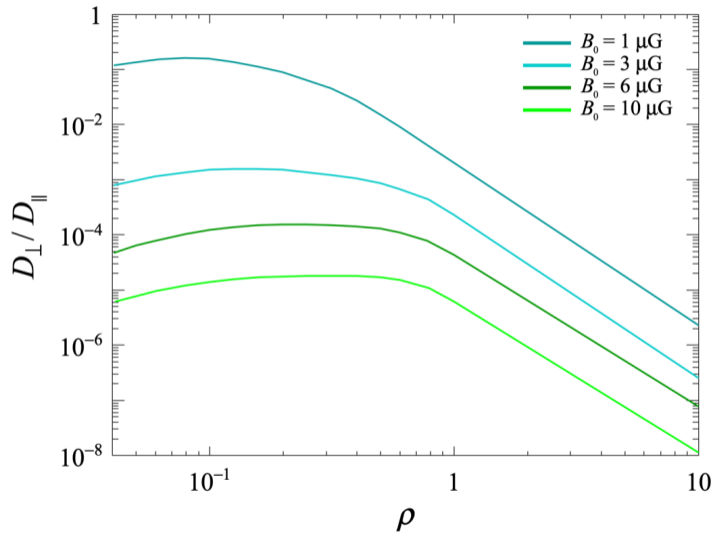
Parallel diffusion



Perpendicular diffusion



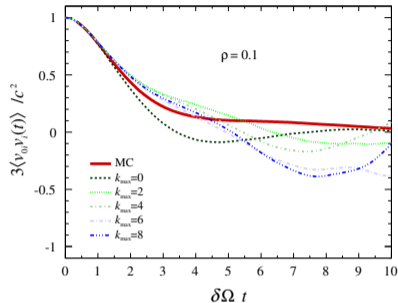
“Benchmark plot”



- Physical solution provided by partial summation schemes
- Convenient way to get dependencies in E , δB^2 , B_0^2 , etc. compared to heavy MC campaigns
- Main features captured in the gyro-resonant and high-rigidity regimes
- Details in *Astrophys.J.* 920 (2021) 2, 87 and arXiv:2406.00361 (to appear in *Phys. Rev. D*)

Backup

Truncation of the series?



- Summation of all terms up to some order k_{\max} ?
- Series absolutely convergent for all t , very many terms required for $t > 3/(2\delta\Omega^2\tau)$ [e.g. Kraichnan, JMP 2 (1961) 124]
- → combinatorics that determines the number of occurrence of each diagram rapidly non-tractable

Partial summation

- Simplest scheme: Bouret propagator [Bouret, NCim 26 (1962) 3833]
- → substitute the “mass operator” in the Dyson equation for unconnected contributions:

$$\equiv \approx \text{---} + \text{---} \text{---} \text{---}$$

- → sum of unconnected diagrams
- NB: exact solution in the case of white-noise process (cancellation of nested/crossed diagrams) [Plotnikov, Pelletier, & Lemoine, A&A 532 (2011) A68]
- Decoupling in Laplace-transform space (change of variables $t = x + x_1 + x_2$, $t_1 = x_1 + x_2$, $t_2 = x_2$): $U^{(1)}(p) = p - \frac{2\delta\Omega^2}{3} \frac{U^{(1)}(p)}{p+\tau}$
- Inverse transform: $u(t) = \frac{A}{2B} e^{-At/6\tau} (e^{Ct/\tau} - 1)$, with $A = 3 + \sqrt{9 - 24\tau^2\delta\Omega^2}$, $B = A - 3$, and $C = \sqrt{1 - 8\tau^2\delta\Omega^2/3}$
- → nonphysical oscillations around 0 for $t > 3/(2\delta\Omega^2\tau)$