

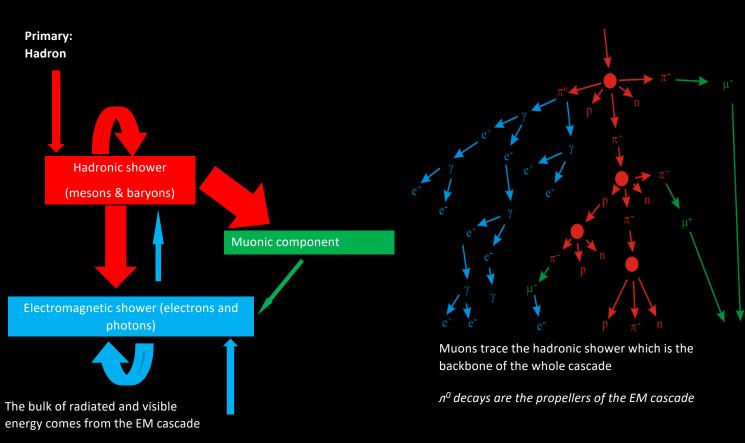
Hadronic and Shower Physics with the Pierre Auger Observatory

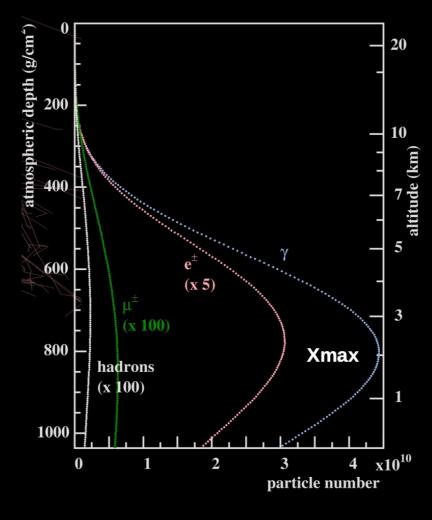
L. Cazon for the Pierre Auger Collaboration



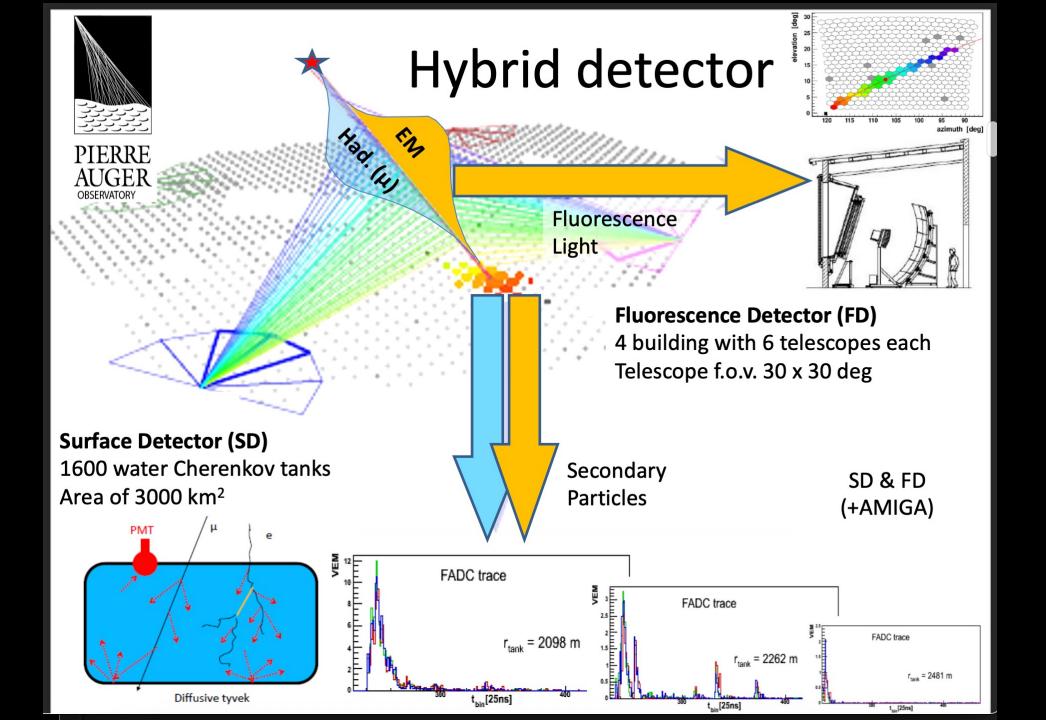


Air Shower:

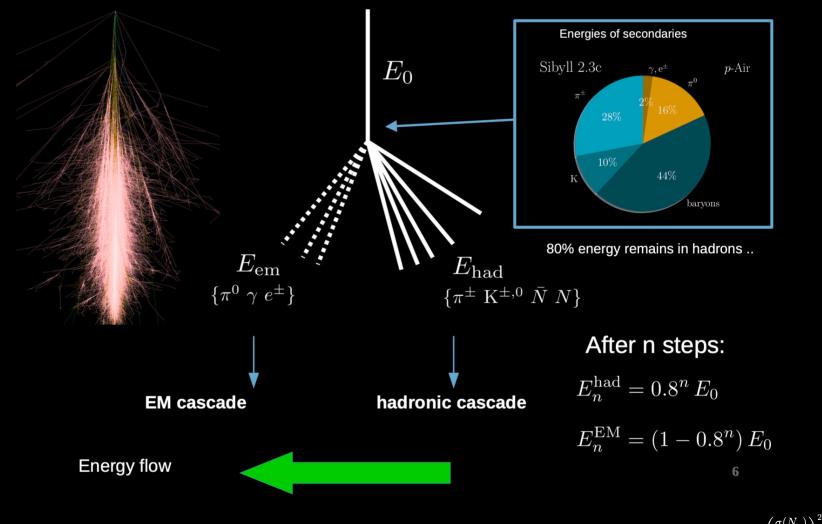


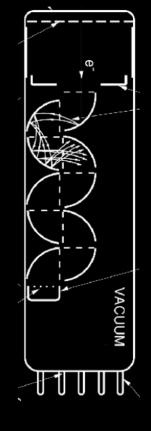


Primary:Photon



Energy flow, averages and fluctuations





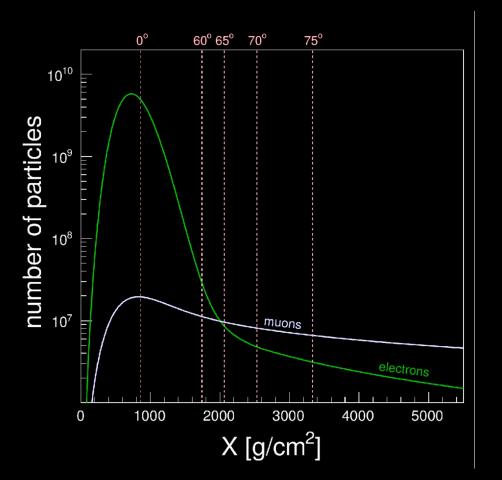
PMT analogy to shower

70% of fluctuations from first interaction

 $\left(\frac{\sigma(N_{\mu})}{N_{\mu}}\right)^2 \simeq \left(\frac{\sigma(\alpha_1)}{\alpha_1}\right)^2 + \left(\frac{\sigma(\alpha_2)}{\alpha_2}\right)^2 + \ldots + \left(\frac{\sigma(\alpha_c)}{\alpha_c}\right)^2$

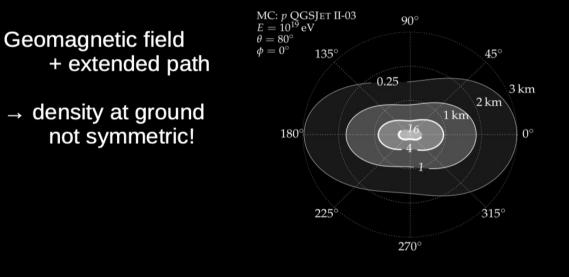
Muon content

Muon content in inclined showers



EM component attenuates after 2000 gr/cm2

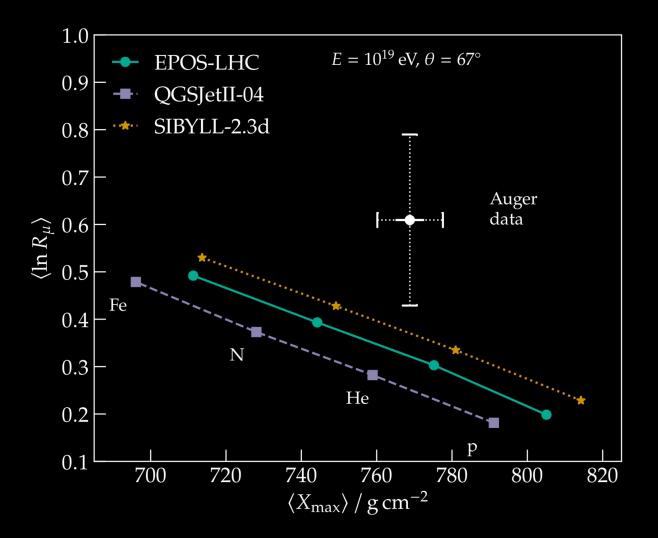
(PRD 91(2015) 032003, PRL 126 (2021) 152002)



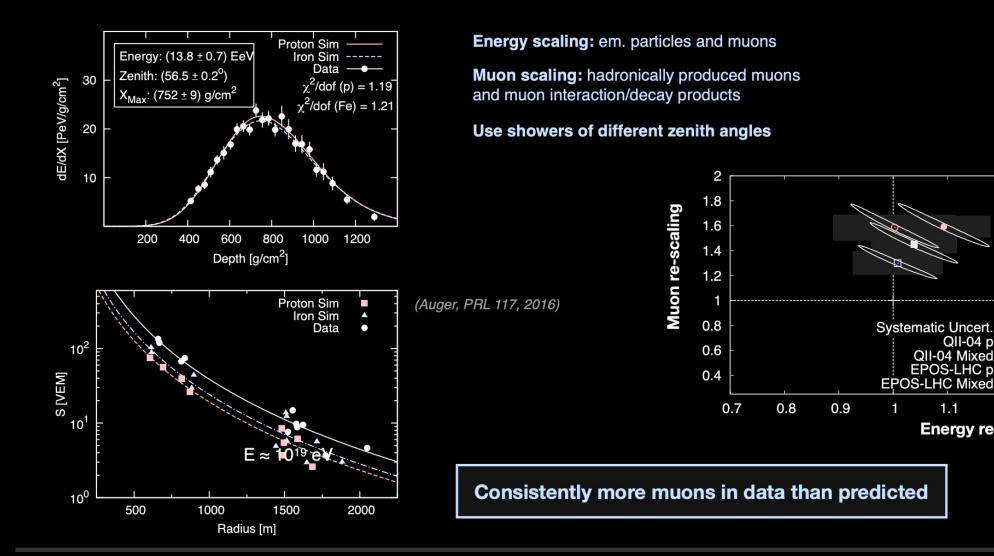
$$ho_\mu(E, heta,x,y) \;=\; R_\mu(E)
ho_\mu^{
m ref}(heta,x,y)$$

Measure scale relative to reference model

Muon content in inclined showers



Muon content in vertical showers



5

QII-04 Mixed

EPOS-LHC p

1.1

 \odot

1.3

1.2

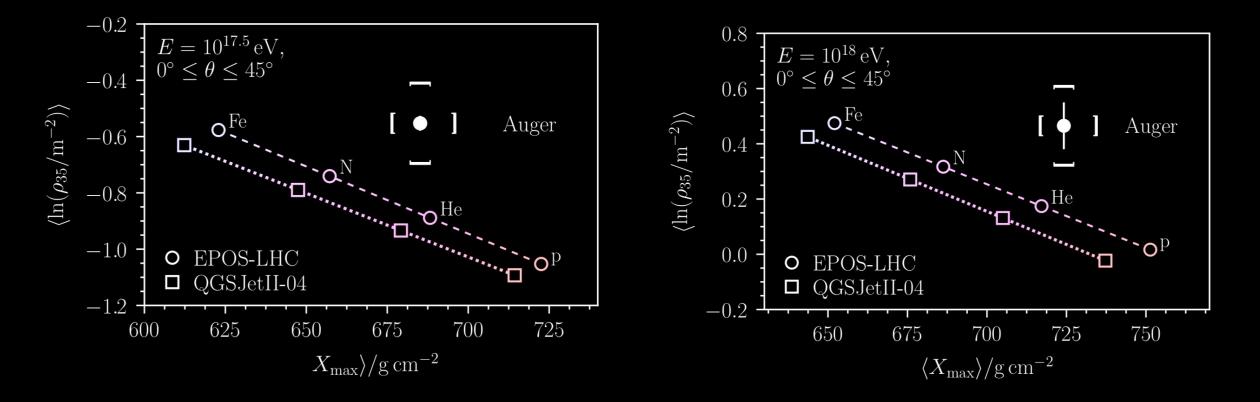
Energy re-scaling

Direct muon content in vertical showers

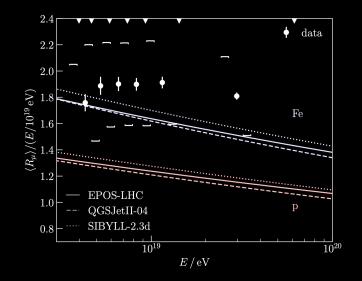


Underground Scintilators (UMD)

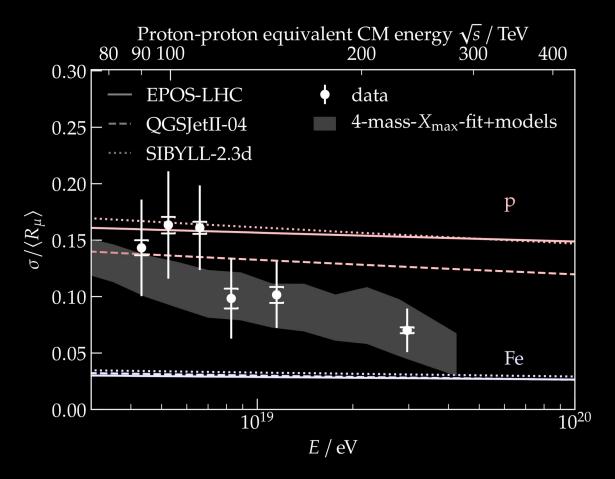




Muon fluctuations



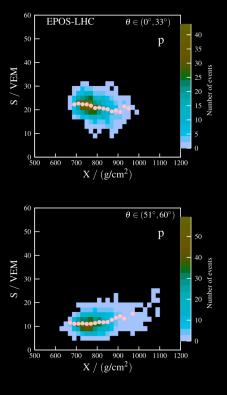
1st interaction of air shower is not exotic in terms of fluctuations

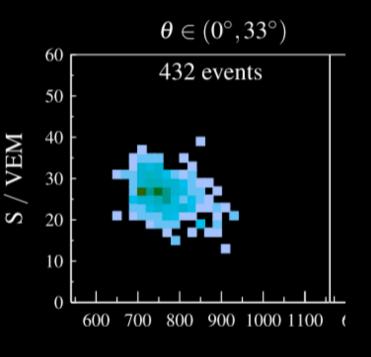


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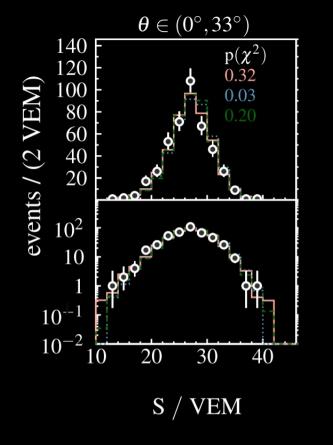
Longitudinal developement

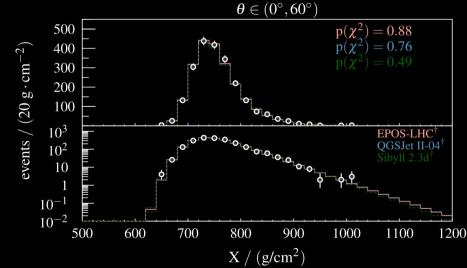
Fitting the Xmax - S1000 joint distribution

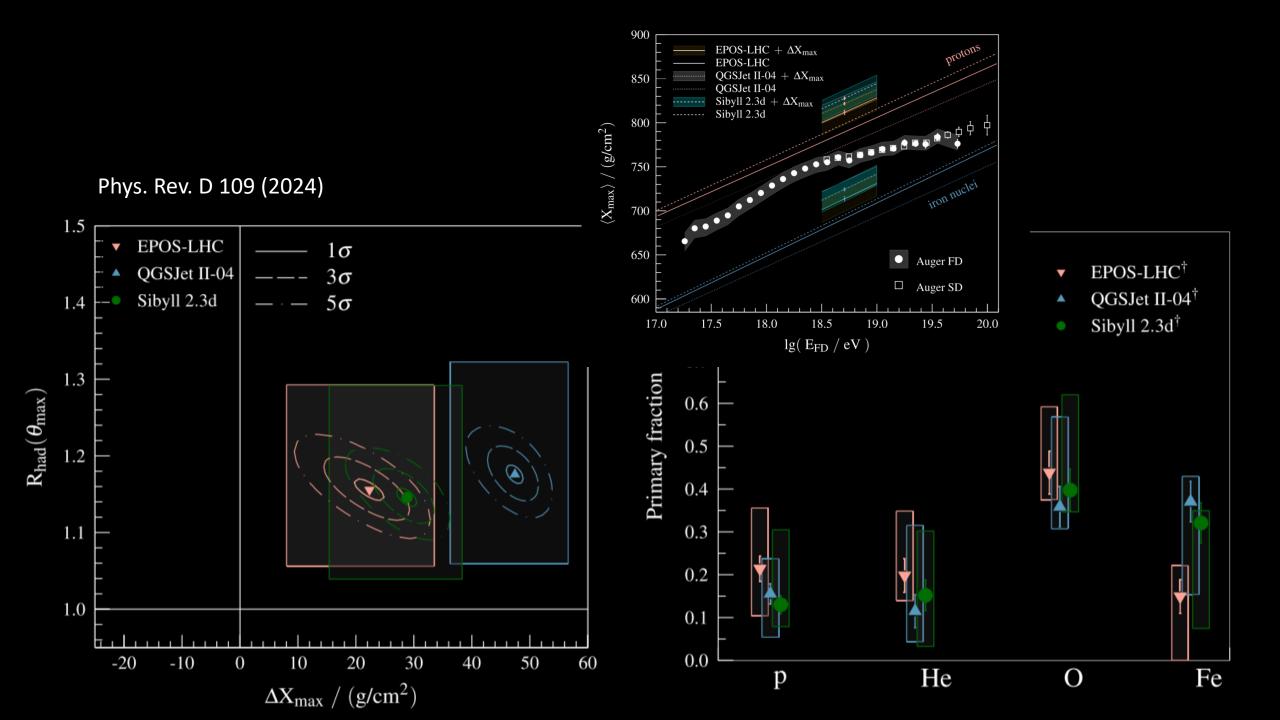




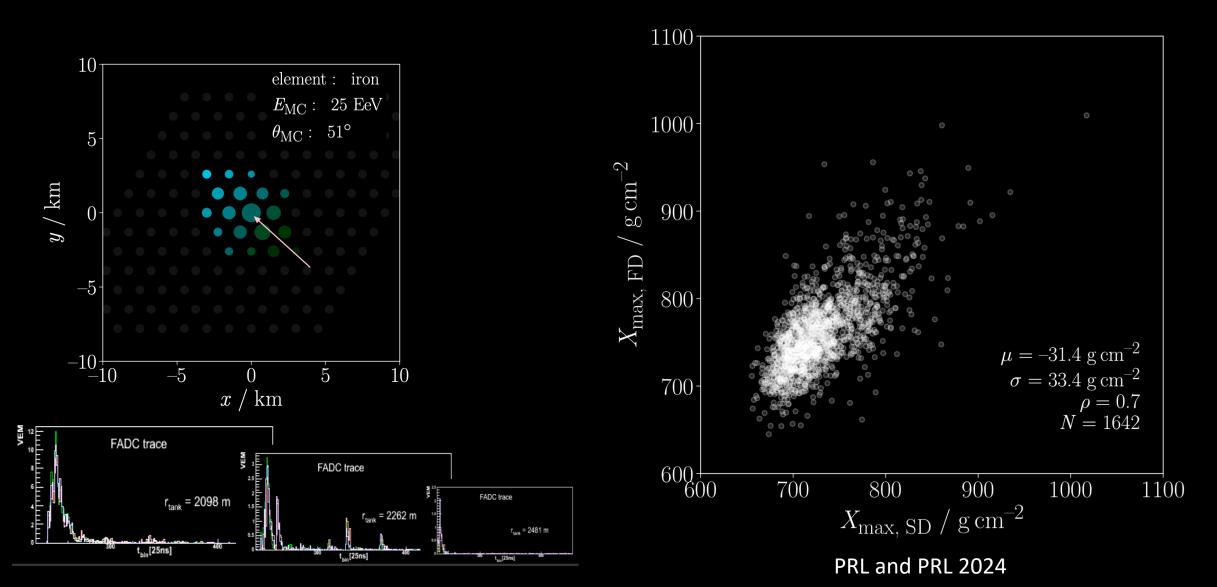




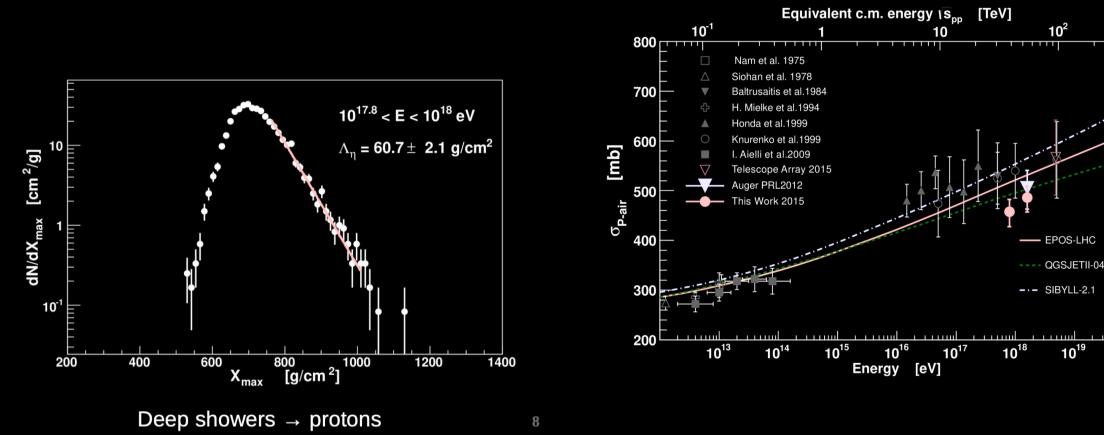




Xmax from DNN on SD traces

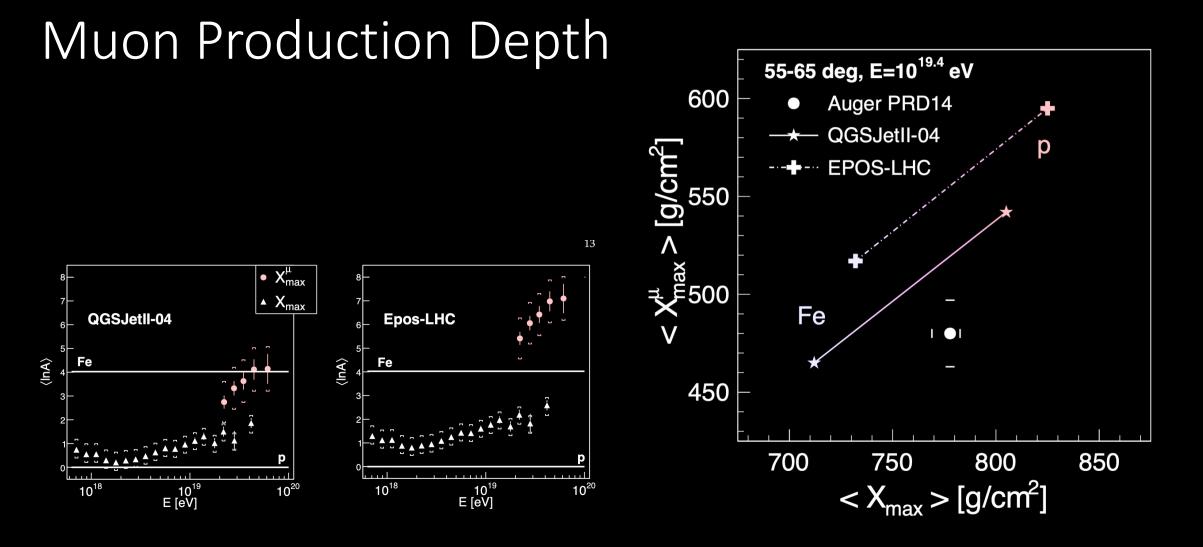


p-Air cross section



PRD 2012

10²⁰



PRD 2015

Conclusions

MUON PUZZLE

- Auger finds a deficit of muons in sims. (All HI models) Results confirmed by independent measurements and different energies
 - Experimental confirmation that energy scale is not the problem
- Muon fluctuations compatible with sims
 - A large deviation from expectations in the 1st interaction it is unfavoured
 - Most likely scenario is accumulation of small deviations along generations

MUON PRODUCTION DEPTH

- Off in models.
 - Related to pion-Air difraction cross section.
 - Fine tunning feedback with Xmax through Hadronic-EM flow in the cascade
- Xmax SCALE
 - not well explained by models by 20-50 gr/cm2

More to come

New publications coming soon

- LIV limits based on muon fluctuations (C. Trimarelli, ICRC2023)
- Muon measurements based on **Shower Universality** (M. Stadlemeier ICRC2023)
- Muon SD and Radio (AERA) (M. Gottowik ICRC2023)
- p-Air cross sections & composition fit (O. Tkatchenko ICRC23)

New detectors, new data: AugerPrime

- Radio
- Surface Scintillators
- Underground Scintillators
- Larger dynamic range for Water Cherenkovs
- New and faster electronics

Back up

