

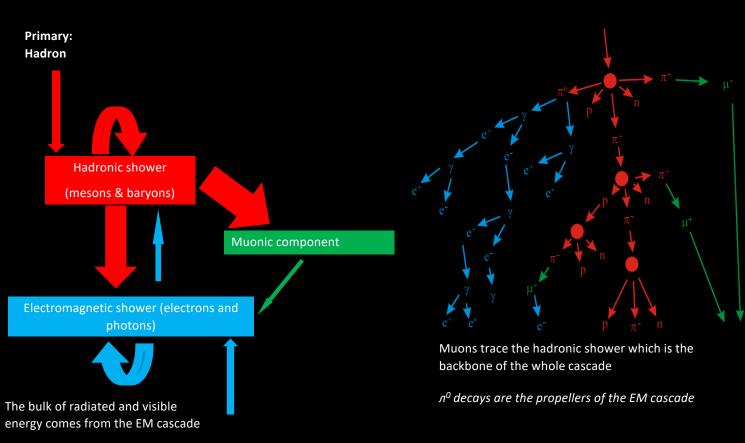
Hadronic and Shower Physics with the Pierre Auger Observatory

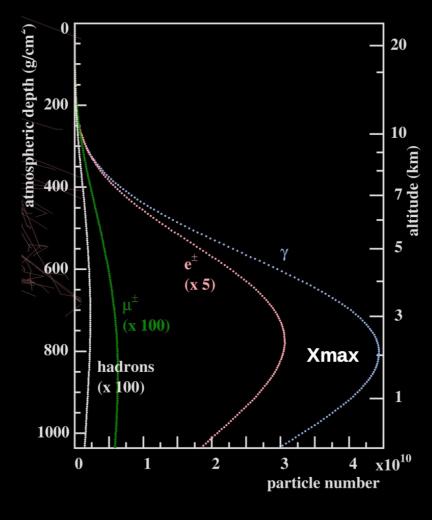
L. Cazon for the Pierre Auger Collaboration



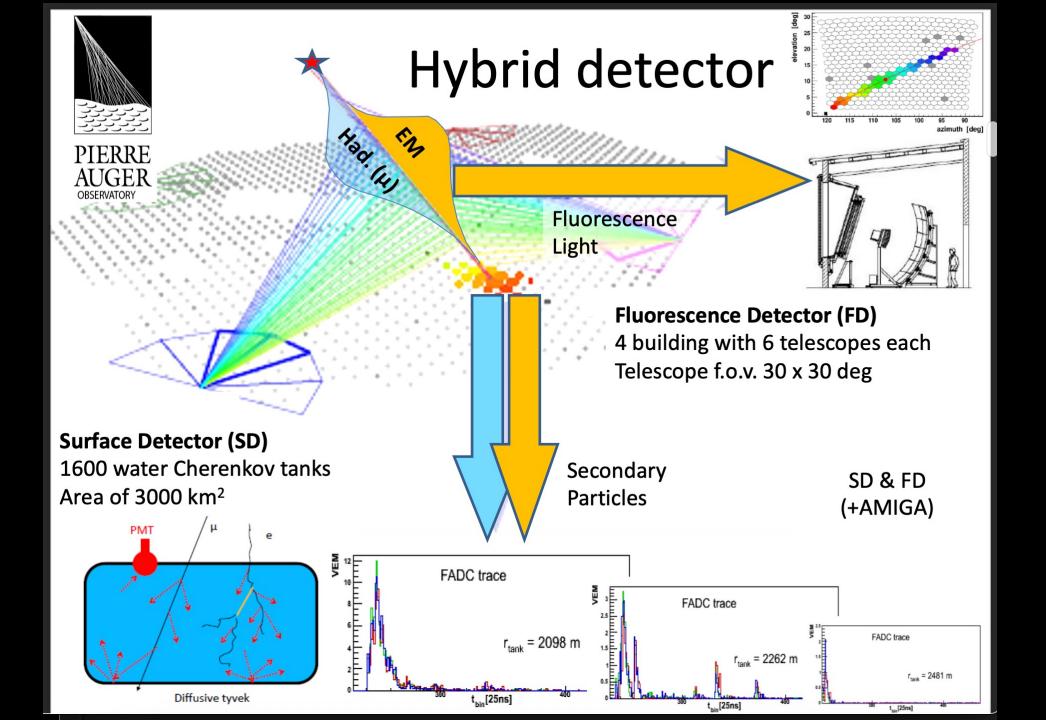


Air Shower:

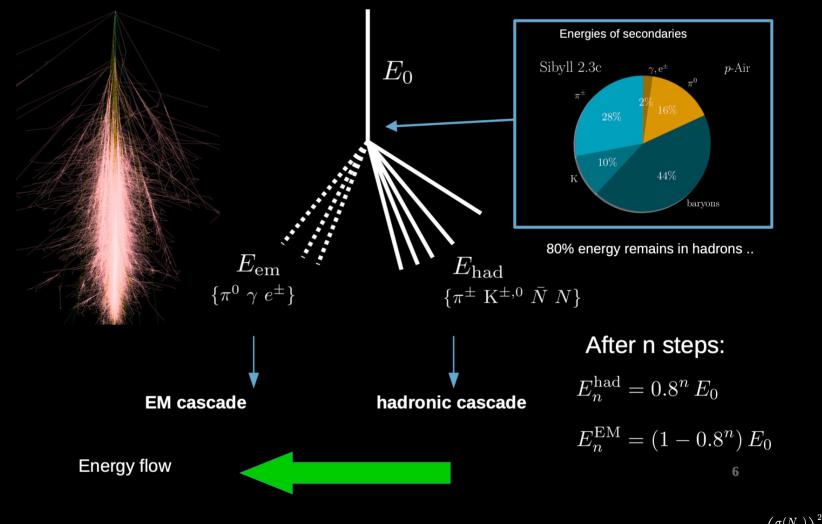


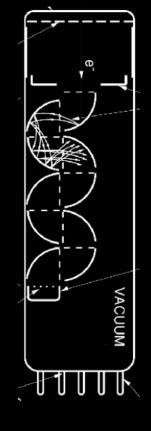


Primary:Photon



Energy flow, averages and fluctuations





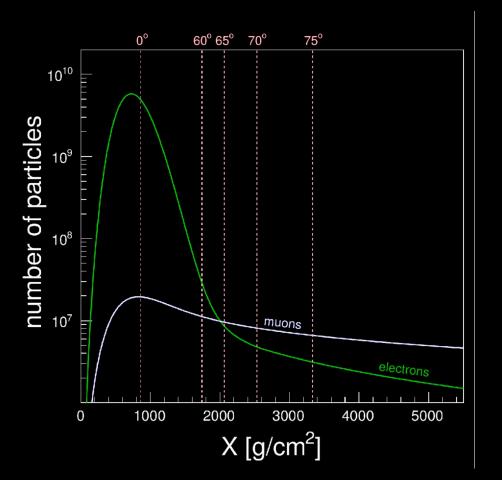
PMT analogy to shower

70% of fluctuations from first interaction

 $\left(\frac{\sigma(N_{\mu})}{N_{\mu}}\right)^2 \simeq \left(\frac{\sigma(\alpha_1)}{\alpha_1}\right)^2 + \left(\frac{\sigma(\alpha_2)}{\alpha_2}\right)^2 + \ldots + \left(\frac{\sigma(\alpha_c)}{\alpha_c}\right)^2$

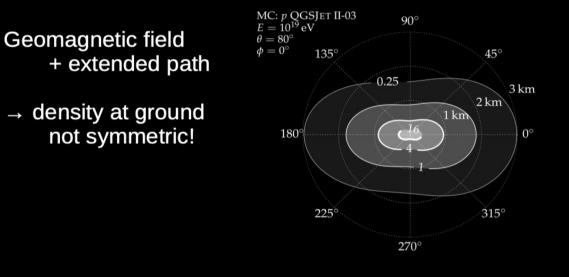
Muon content

Muon content in inclined showers



EM component attenuates after 2000 gr/cm2

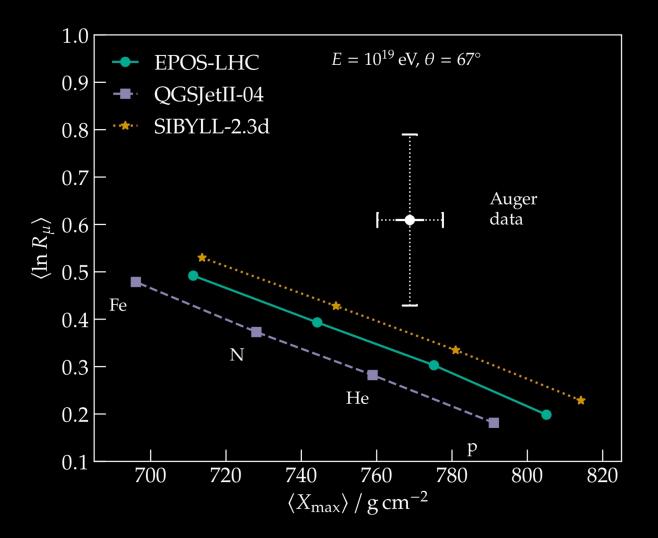
(PRD 91(2015) 032003, PRL 126 (2021) 152002)



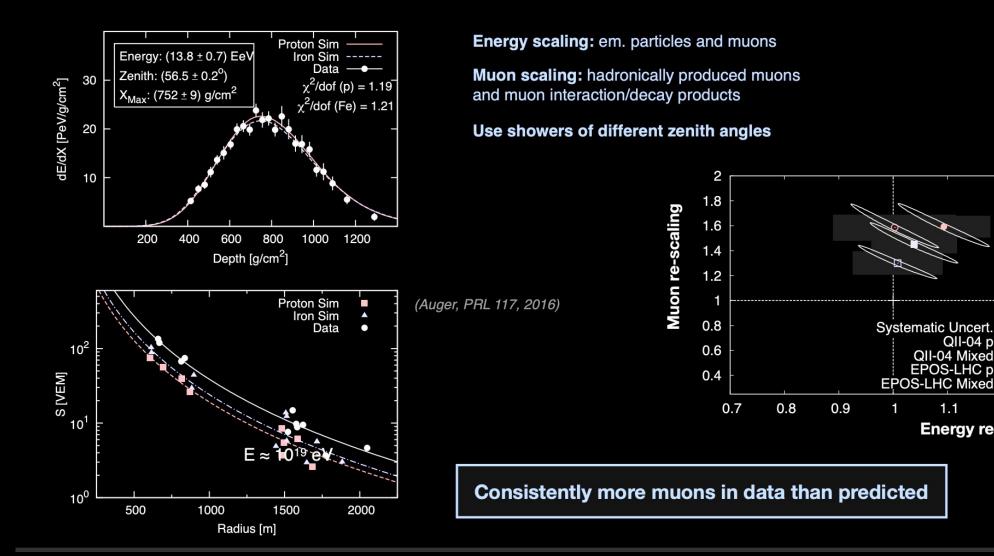
$$ho_\mu(E, heta,x,y) \;=\; R_\mu(E)
ho_\mu^{
m ref}(heta,x,y)$$

Measure scale relative to reference model

Muon content in inclined showers



Muon content in vertical showers



5

QII-04 Mixed

EPOS-LHC p

1.1

 \odot

1.3

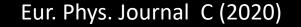
1.2

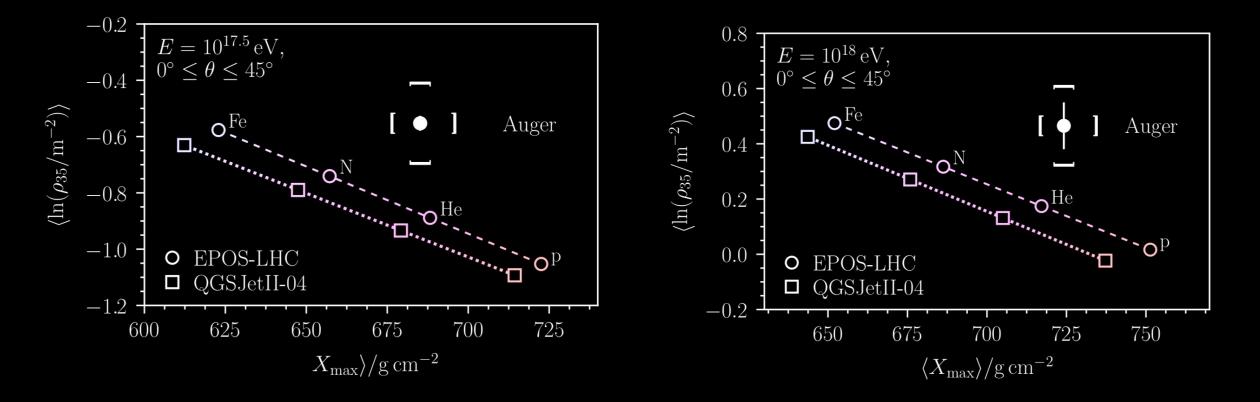
Energy re-scaling

Direct muon content in vertical showers

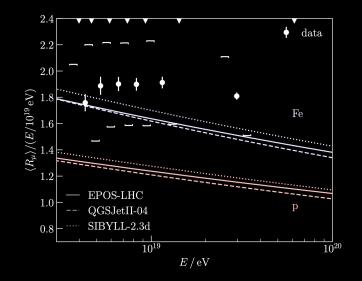


Underground Scintilators (UMD)

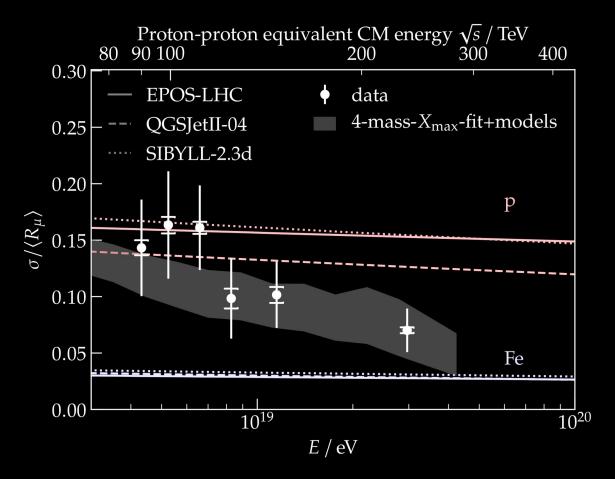




Muon fluctuations



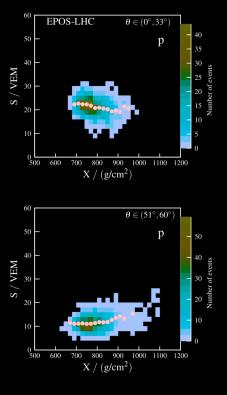
1st interaction of air shower is not exotic in terms of fluctuations

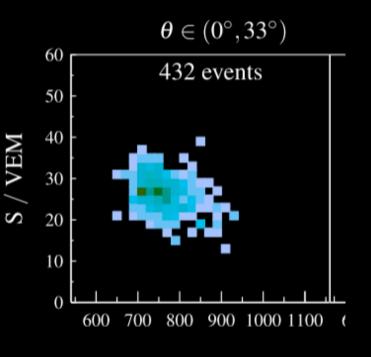


PRL (2021) 126

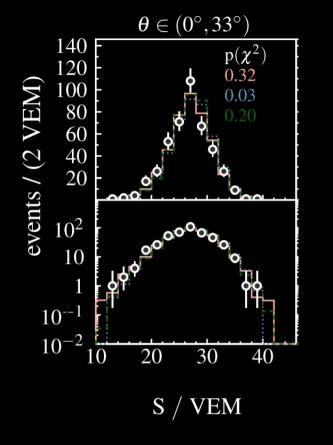
Longitudinal developement

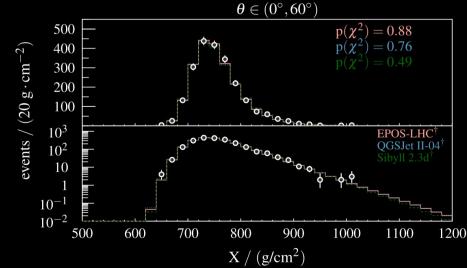
Fitting the Xmax - S1000 joint distribution

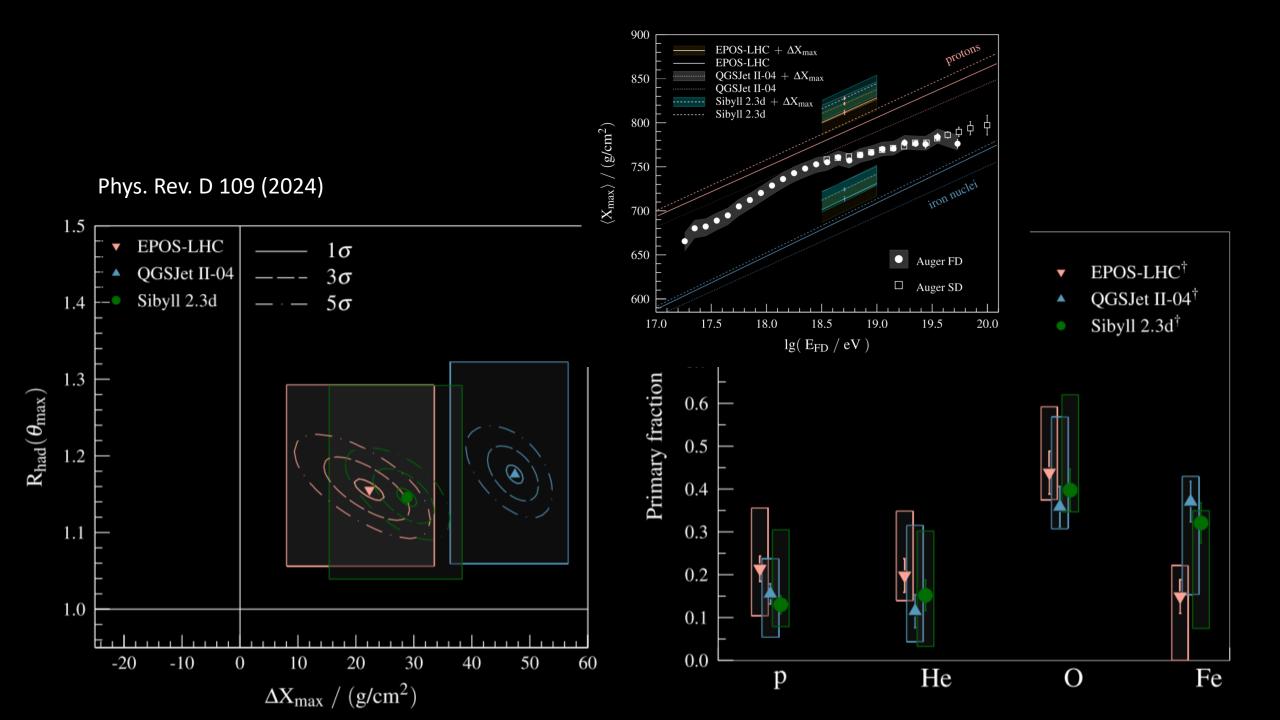




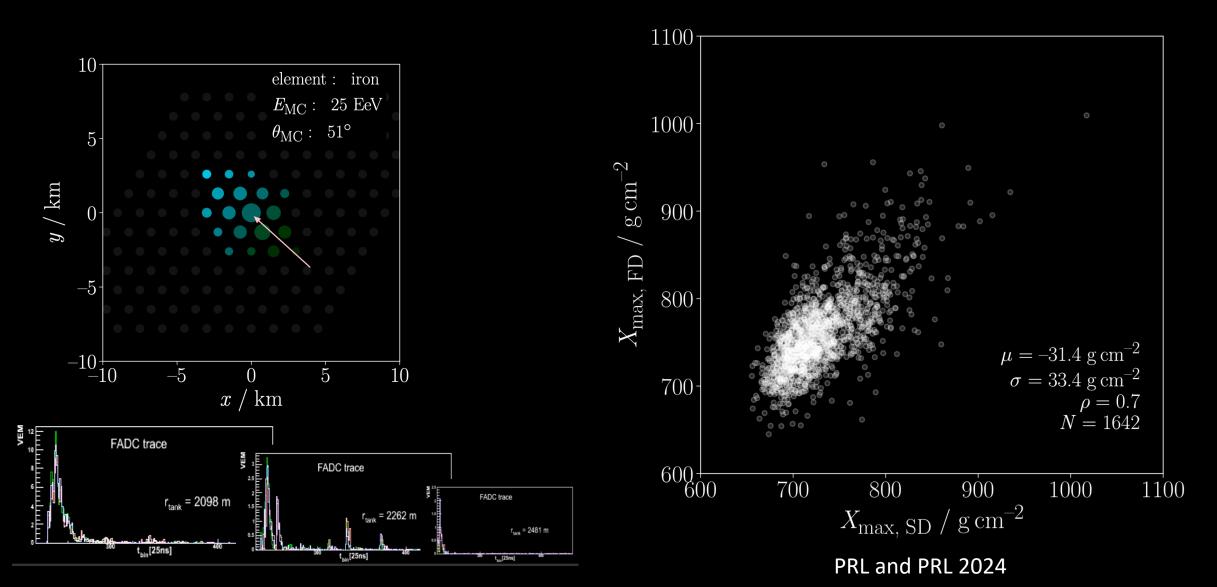




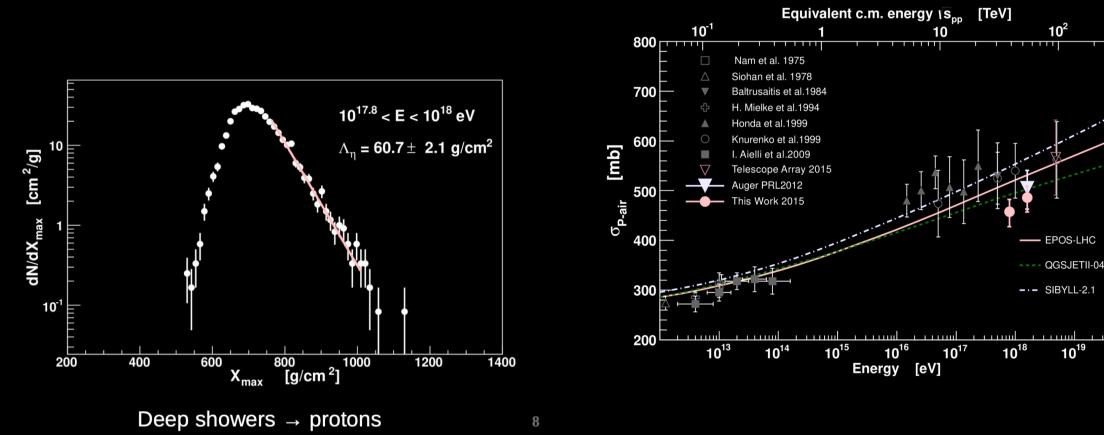




Xmax from DNN on SD traces

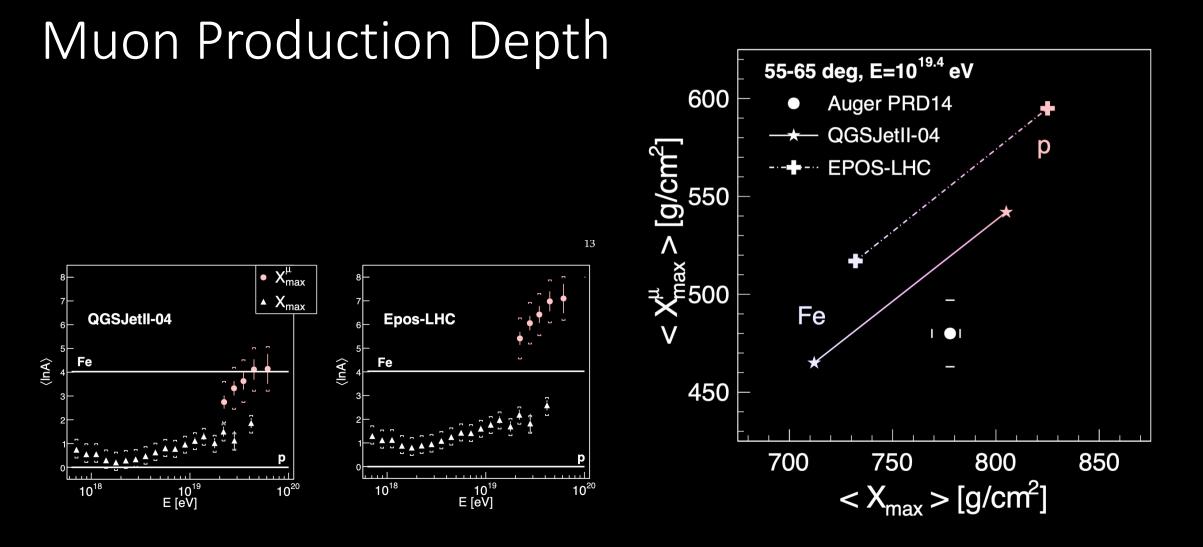


p-Air cross section



PRD 2012

10²⁰



PRD 2015

Conclusions

MUON PUZZLE

- Auger finds a deficit of muons in sims. (All HI models) Results confirmed by independent measurements and different energies
 - Experimental confirmation that energy scale is not the problem
- Muon fluctuations compatible with sims
 - A large deviation from expectations in the 1st interaction it is unfavoured
 - Most likely scenario is accumulation of small deviations along generations

MUON PRODUCTION DEPTH

- Off in models.
 - Related to pion-Air difraction cross section.
 - Fine tunning feedback with Xmax through Hadronic-EM flow in the cascade
- Xmax SCALE
 - not well explained by models by 20-50 gr/cm2

More to come

New publications coming soon

- LIV limits based on muon fluctuations (C. Trimarelli, ICRC2023)
- Muon measurements based on **Shower Universality** (M. Stadlemeier ICRC2023)
- Muon SD and Radio (AERA) (M. Gottowik ICRC2023)
- p-Air cross sections & composition fit (O. Tkatchenko ICRC23)

New detectors, new data: AugerPrime

- Radio
- Surface Scintillators
- Underground Scintillators
- Larger dynamic range for Water Cherenkovs
- New and faster electronics

Back up

