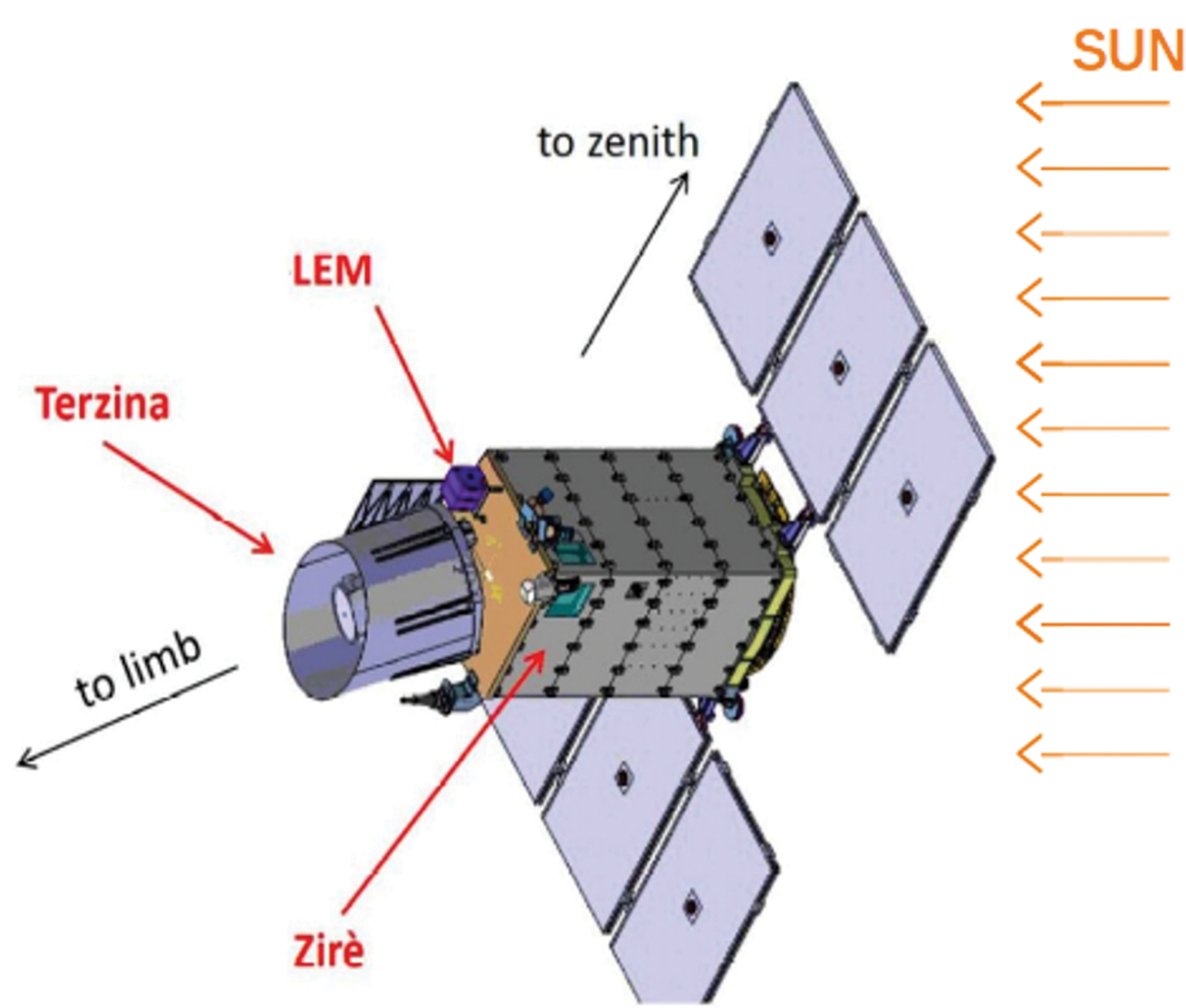


Radiation Hardness and Annealing Strategies for SiPMs of the Terzina Telescope on-board the NUSES space mission

L. Burmistrov, S. Davarpanah*, M. Heller, T. Montaruli, C. Tönnis, C. Trimarelli

1. NUSES mission

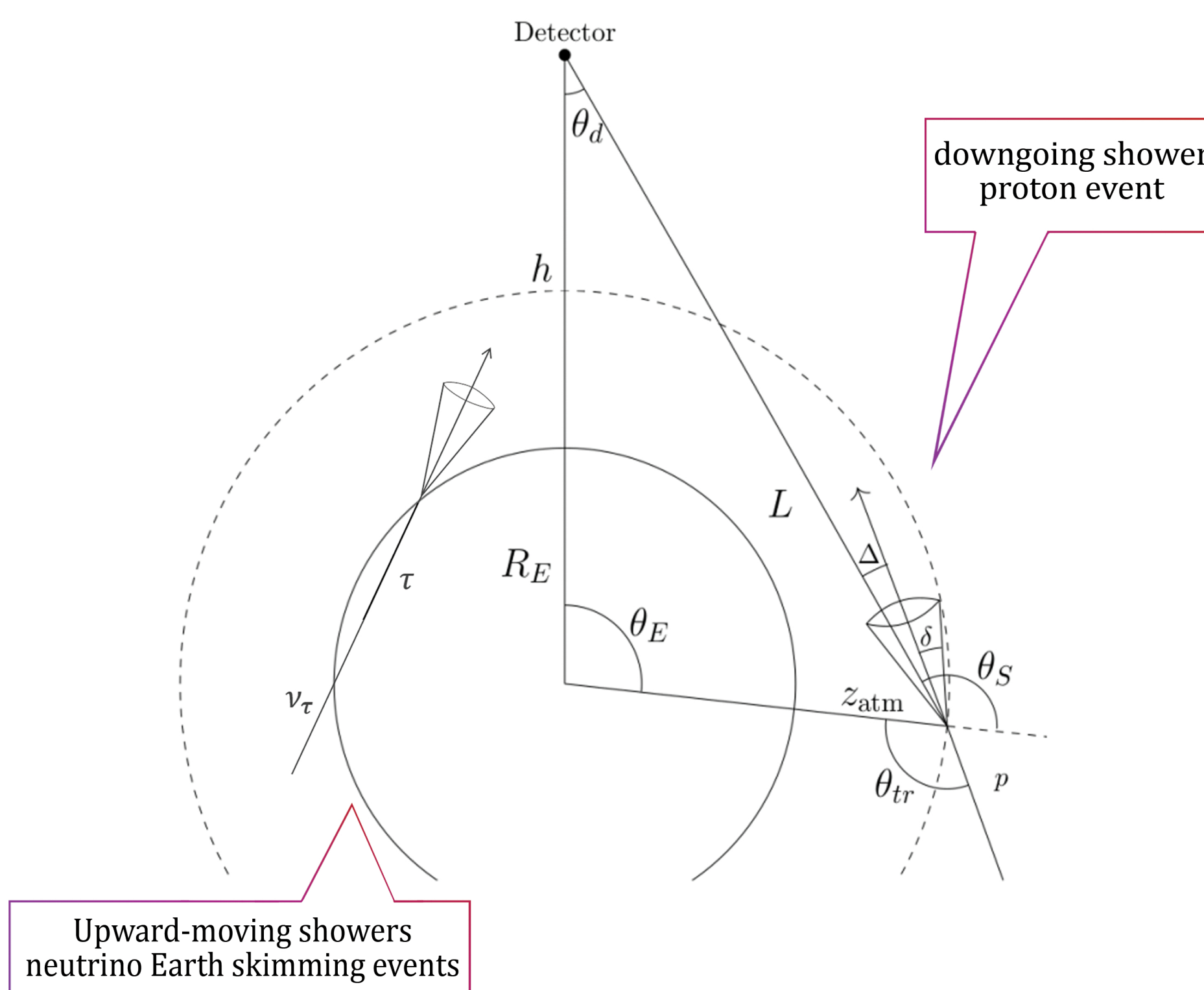
- Lifetime 3 years
- Sun-synchronous orbit
- Low Earth orbit, initial altitude 535 km
- Two payloads: ZIRE and TERZINA



*shideh.davarpanah@unige.ch
on behalf the NUSES collaboration

2. Terzina target events

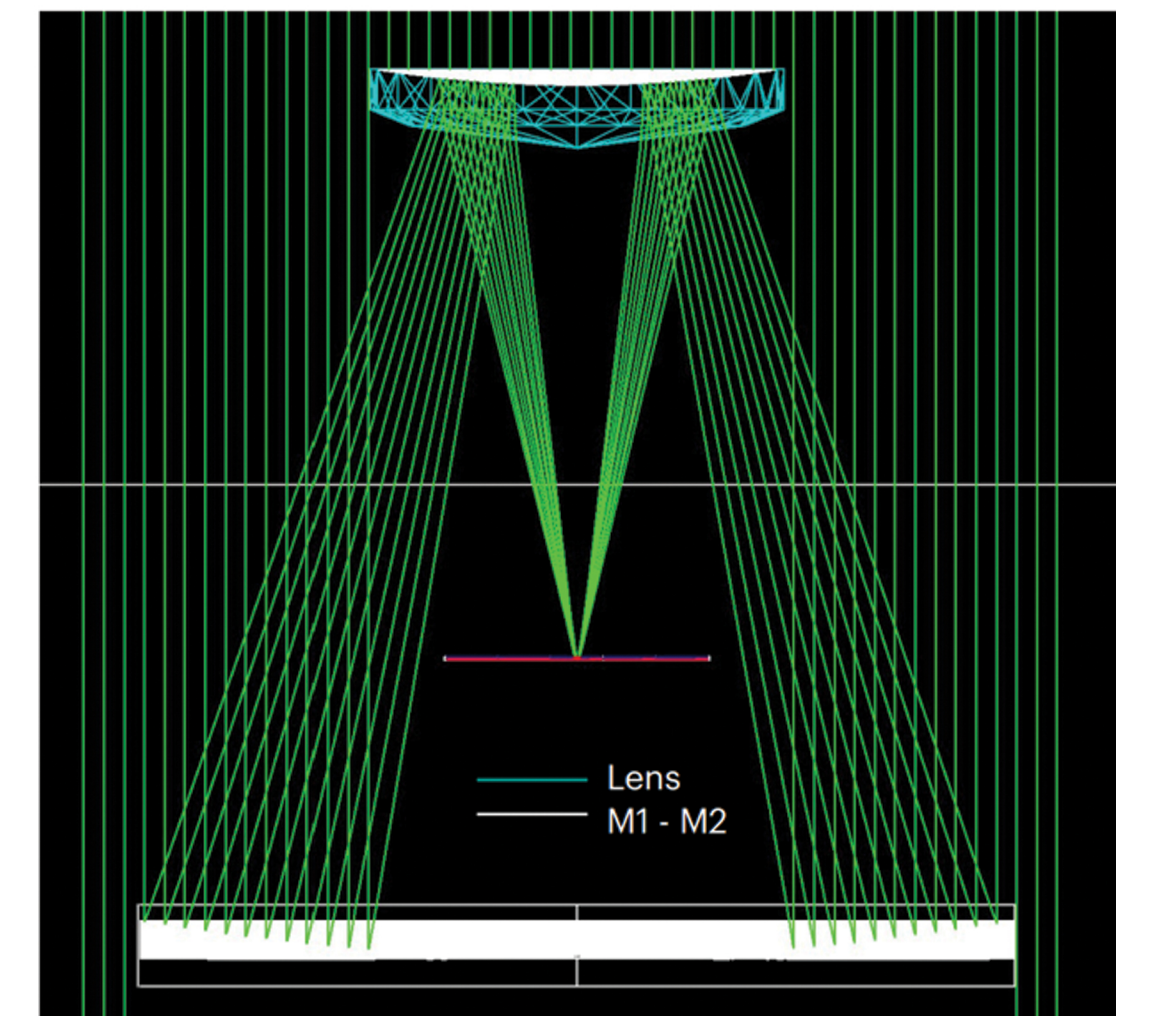
- 1st observation of Cherenkov light emitted by atmospheric showers from space!
- upgoing and downgoing Extensive Air Shower



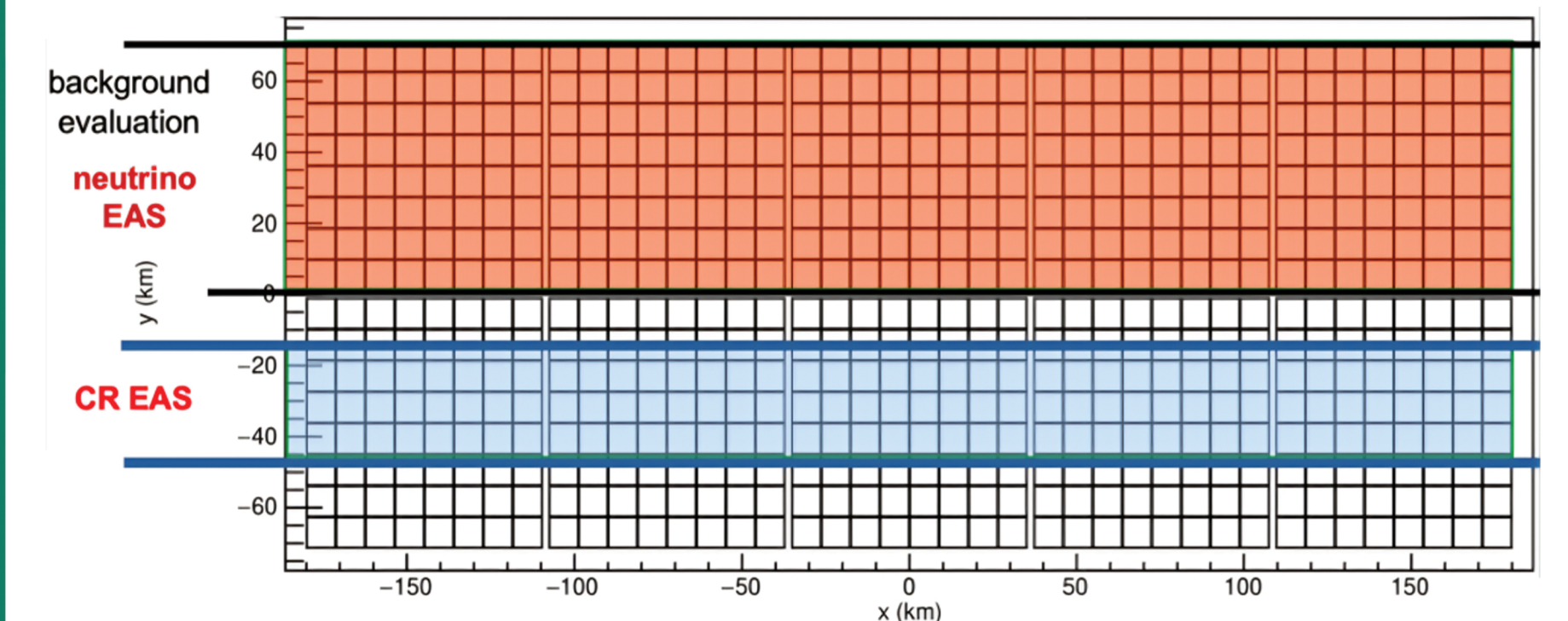
3. Terzina detector

- Schmidt-Cassegrain optical system
- Camera Field of View ~ 7.2° x 2.88° (40 x 16 pixels)

- 2 x 5 = 10 tiles (24.9 x 24.9 mm²)
- each 8 x 8 pixels (3 x 3 mm²)

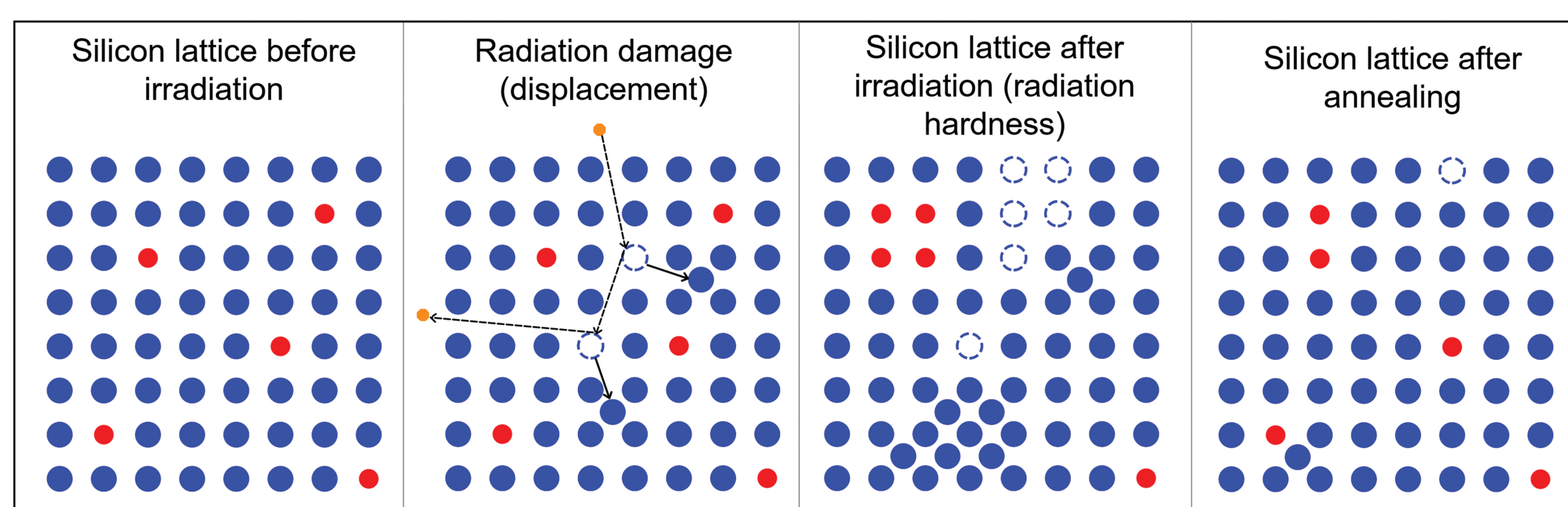
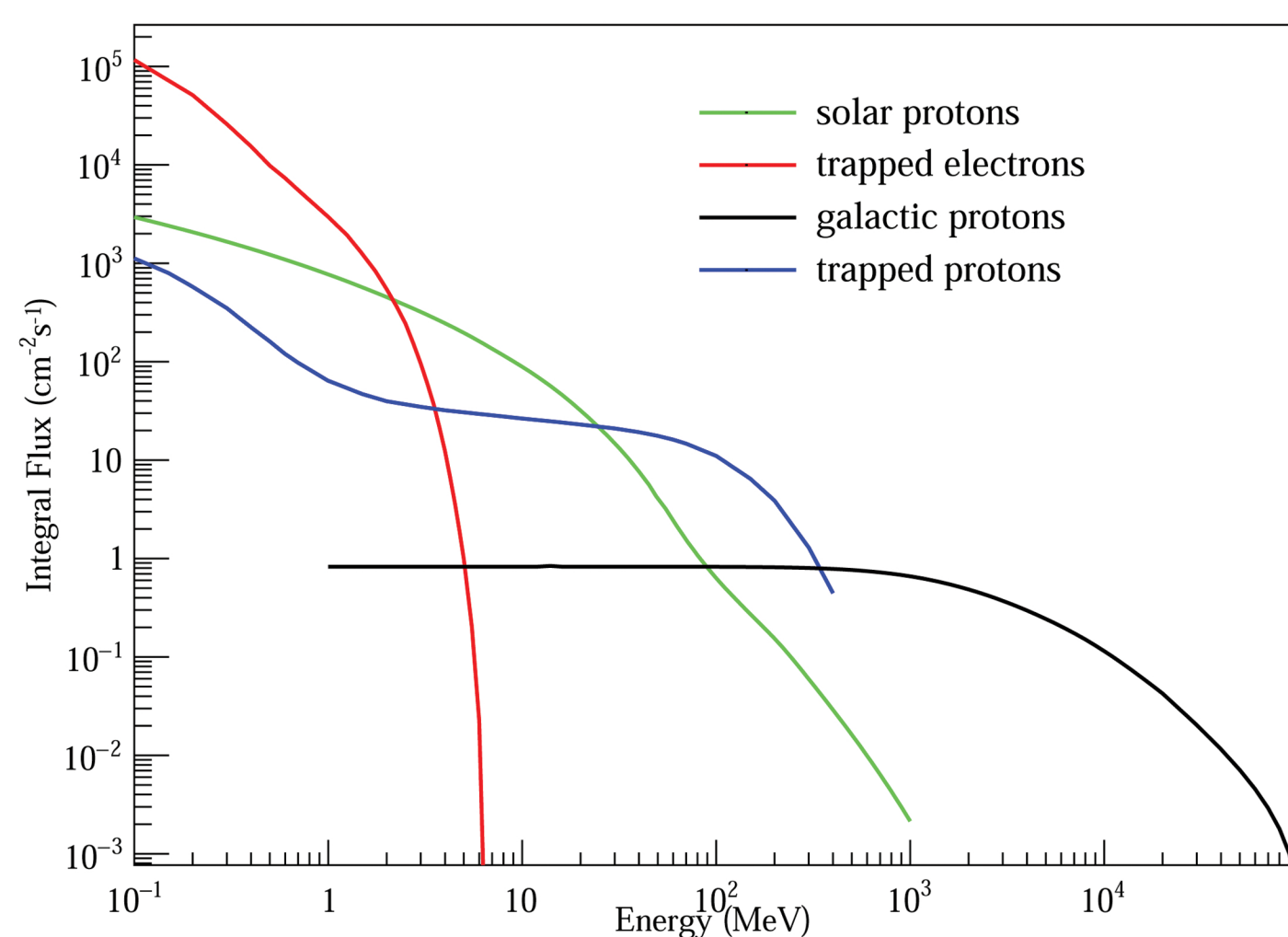


- sensitive area for each pixel = 6.58 mm²



4. Radiation background, radiation damage and annealing

- No atmospheric shield
- Radiation background on the telescope



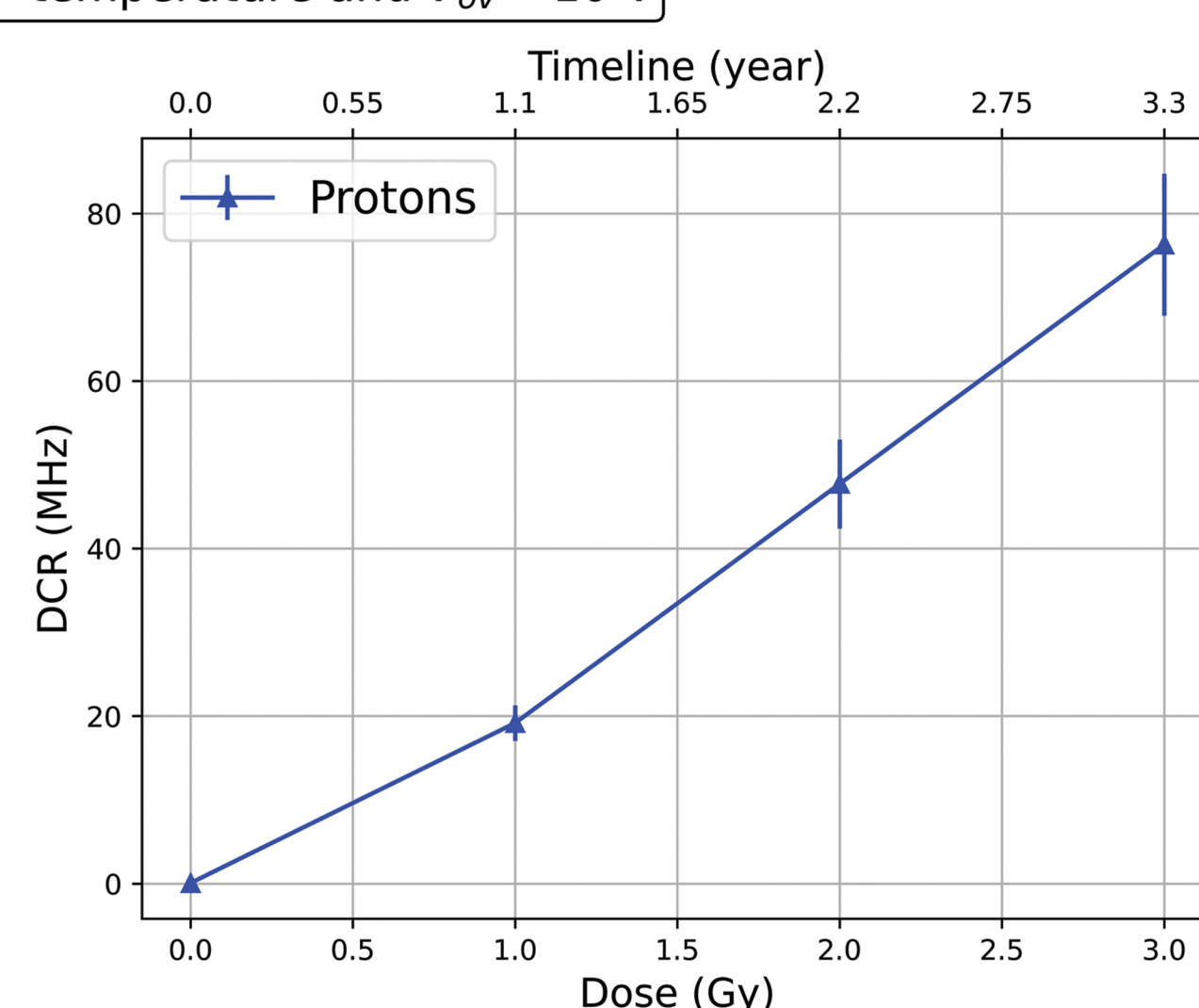
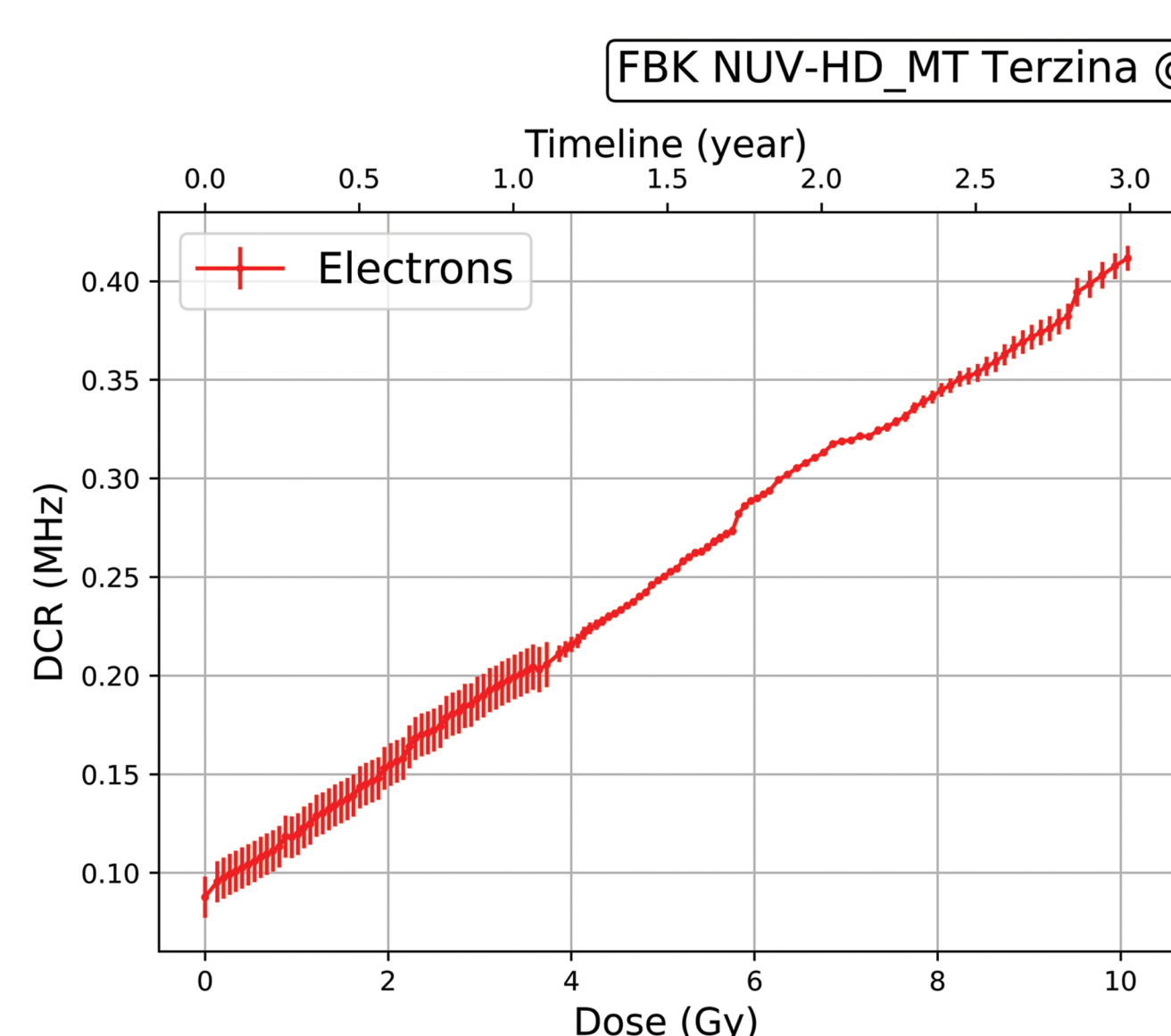
- Radiation damage increase the Dark Count Rate
- Mitigation: by heating them; measuring efficiency of treatment at temperature comparable to the power consumption

● silicon ○ vacancies
● doping / impurities ● radiation

5. Measurements and results

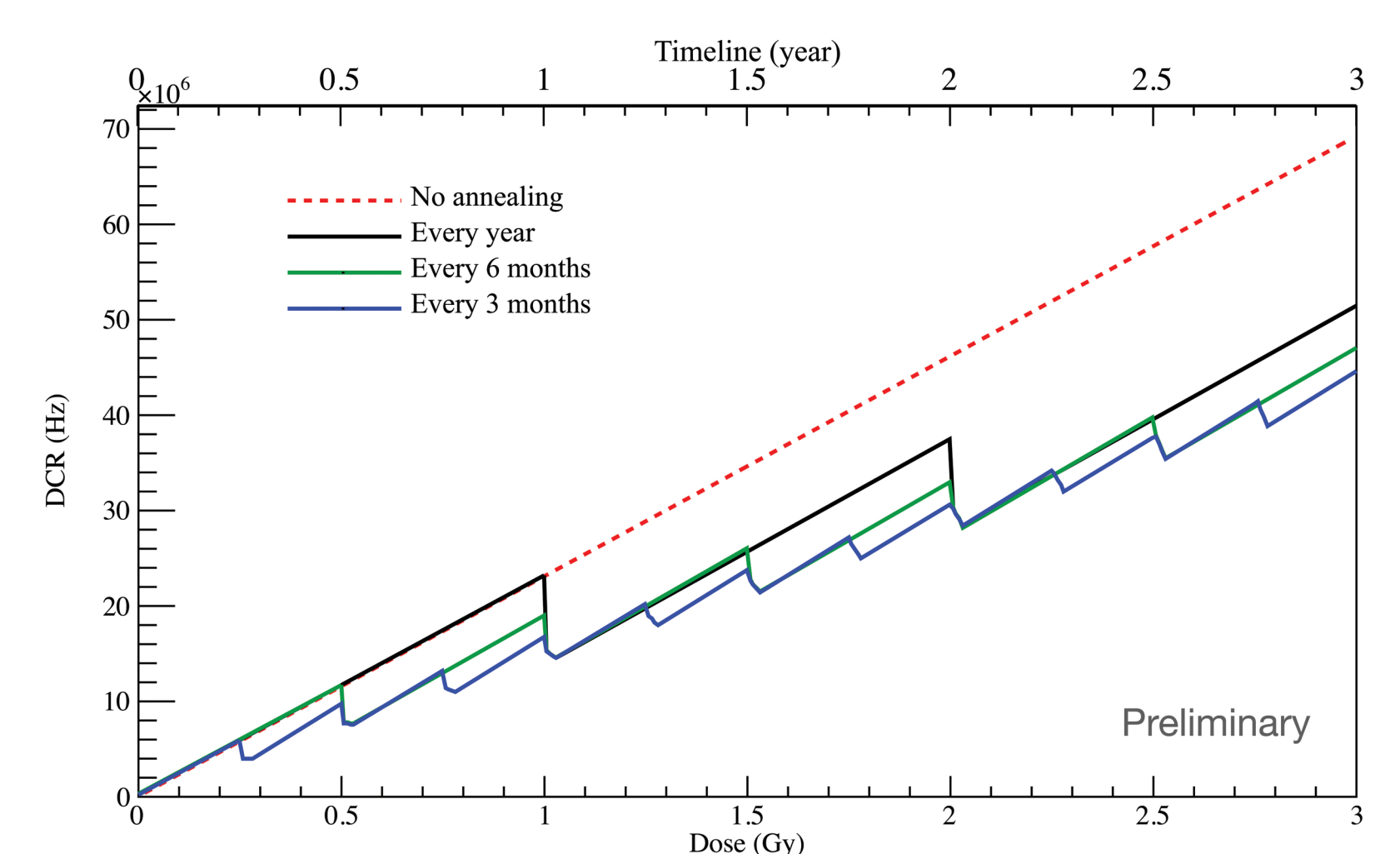
IV measurement done at temperature of 22.0°C for two irradiation cases:

- Proton irradiation: 50MeV proton beam (3cm diameter) with ~3% uniformity
- Electron irradiation: β-source strontium-90, up to 2.2MeV



Inferred DCR per pixel for Terzina SiPMs sensitive area from IV measurement during the irradiation

$$\text{DCR}(10 \text{ ov, dose}) \propto \frac{I(10 \text{ ov, dose})}{\text{Gain} \cdot q \cdot S^*} \cdot \frac{\text{Sensitive area}}{T_{\text{factor}}}$$



Different annealing cycles for the total DCR change during satellite lifetime



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