

The Virtual Observatory in the age of multimessenger Astroparticle Physics

Claudio Galelli - LUTh, Observatoire de Paris
20/06/2023, 13th CRIS-MAC, Trapani



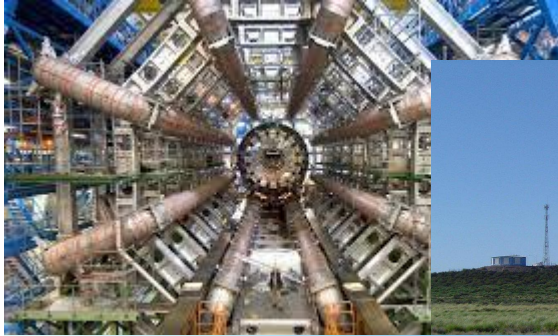
Observatoire
de Paris

| PSL 



The “closed” approach

ATLAS



Auger



HESS

Operation decided by the collaborations
Data released as collaboration decision
Data released fractionally

The “open” approach



Rubin/LSST



Webb

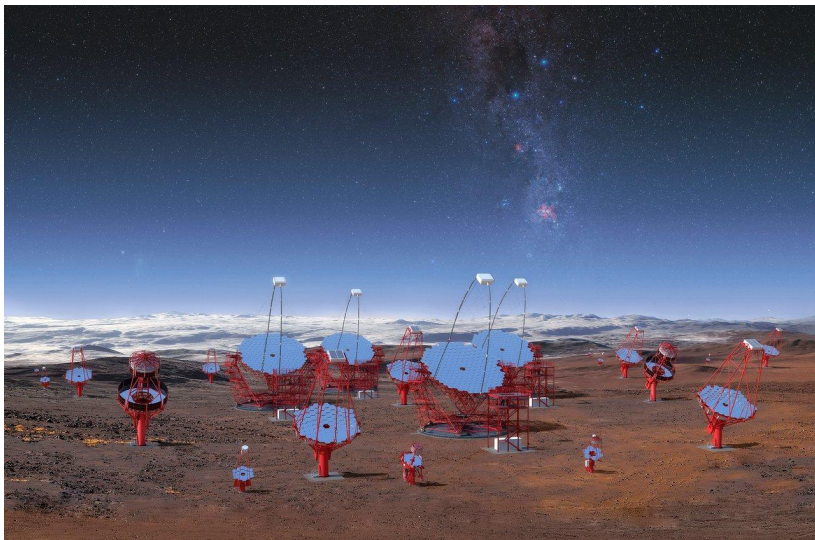


ALMA

Open calls for observation proposals
Data private for a set time (“proprietary”)
All data open and available to the public 2

Why is **CTAO** important for open data?

CTAO will be the first instrument in the VHE environment to operate as an **observatory**



CTA south at Paranal (simulation)

Proposal based - data will be open after a proprietary period

All the **data will be made public** with its **analysis software**

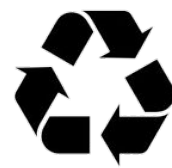
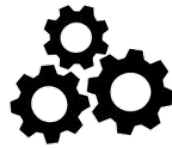


Not as easy as it looks!

Can't we just format our data as .txt files with incomprehensible keywords for column names and put it in an obscure archive page in a website?

Well yes, but it would be very nice if you could be

F_{indable} A_{ccessible} I_{nteroperable} R_{eusable}



Data and metadata should be **easy to find for humans and computers**

Once the data is found, it should be clear how it is **accessed, including authentication**

The data should be easy to **integrate with other data** and analysis and storage workflows

The **reuse of data** should be **optimized** through data and metadata description

FAIR data in astrophysics

Making open data FAIR can be a nightmare - luckily in astrophysics we have a set of guidelines to help us, and we have to thank the **International Virtual Observatory Alliance**



The IVOA is the body that governs the standards on which the **Virtual Observatory (VO)** is built



A collection of data centers containing archives and software tools for astronomical research

A good example are the platforms from the Strasbourg Data Center

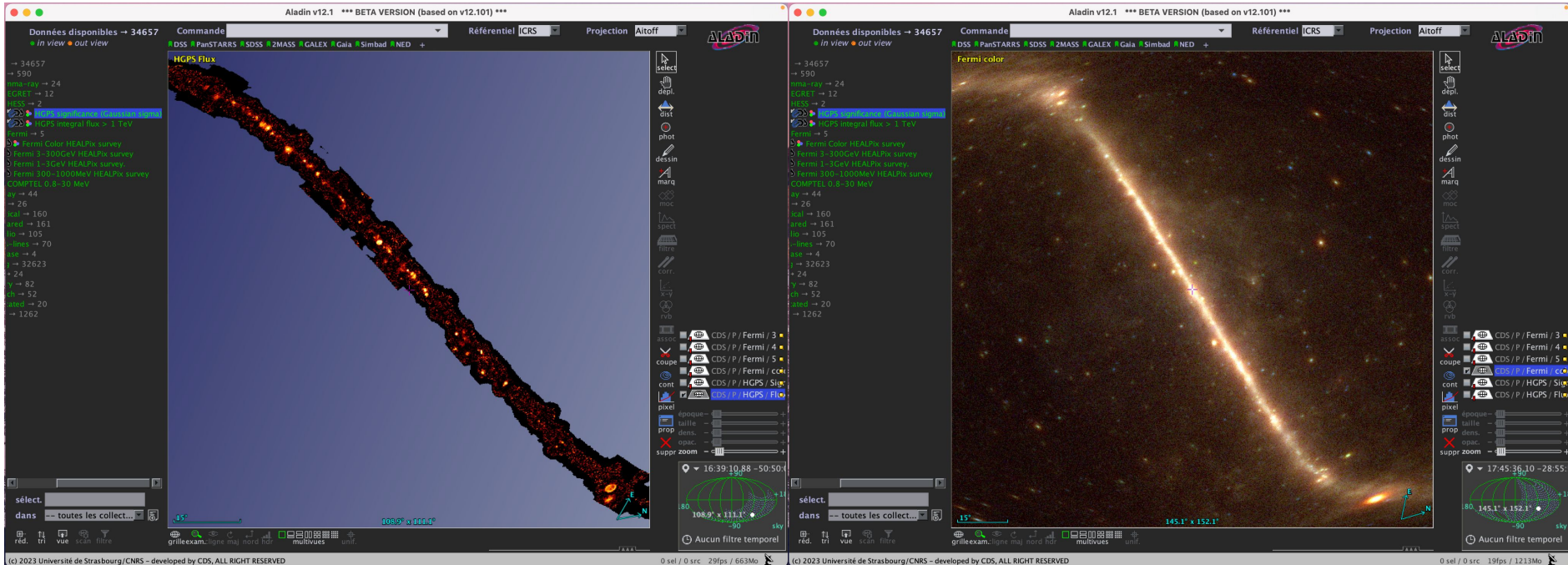


Data is easy to browse, download and plot together in-app



VHE astrophysics and the Virtual Observatory

Some gamma-ray data has already made its way to the VO tools!



These are point contributions! Showcase example: HESS public releases on VO

F = Findable

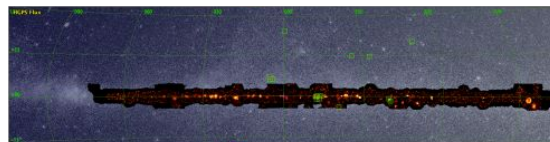


International Virtual Observatory Alliance (IVOA) standards tailored to make data findable:

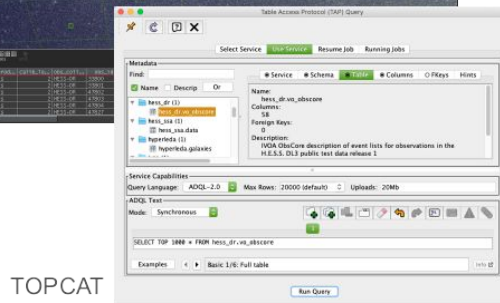
- **IVOA** Observation Data Model Core Components (**ObsCore**)
[\[link to IVOA REC\]](#) → adapted to Cherenkov data
- **IVOA** Table Access Protocol (**TAP**)
[\[link to IVOA REC\]](#)
- **Deployed service** at ObsParis
<https://hess-dr.obspm.fr>
- Registered to the VO **Registry** via PADC (Paris Astronomical Data Centre)

→ **Data widely findable**

e.g. Aladin, TOPCAT, TapHandle, PyVO...
+ dedicated web pages



Aladin



TOPCAT

F = Findable : generation of an ObsCore Table

Module **ivoa.py** now included in **Gammapy**

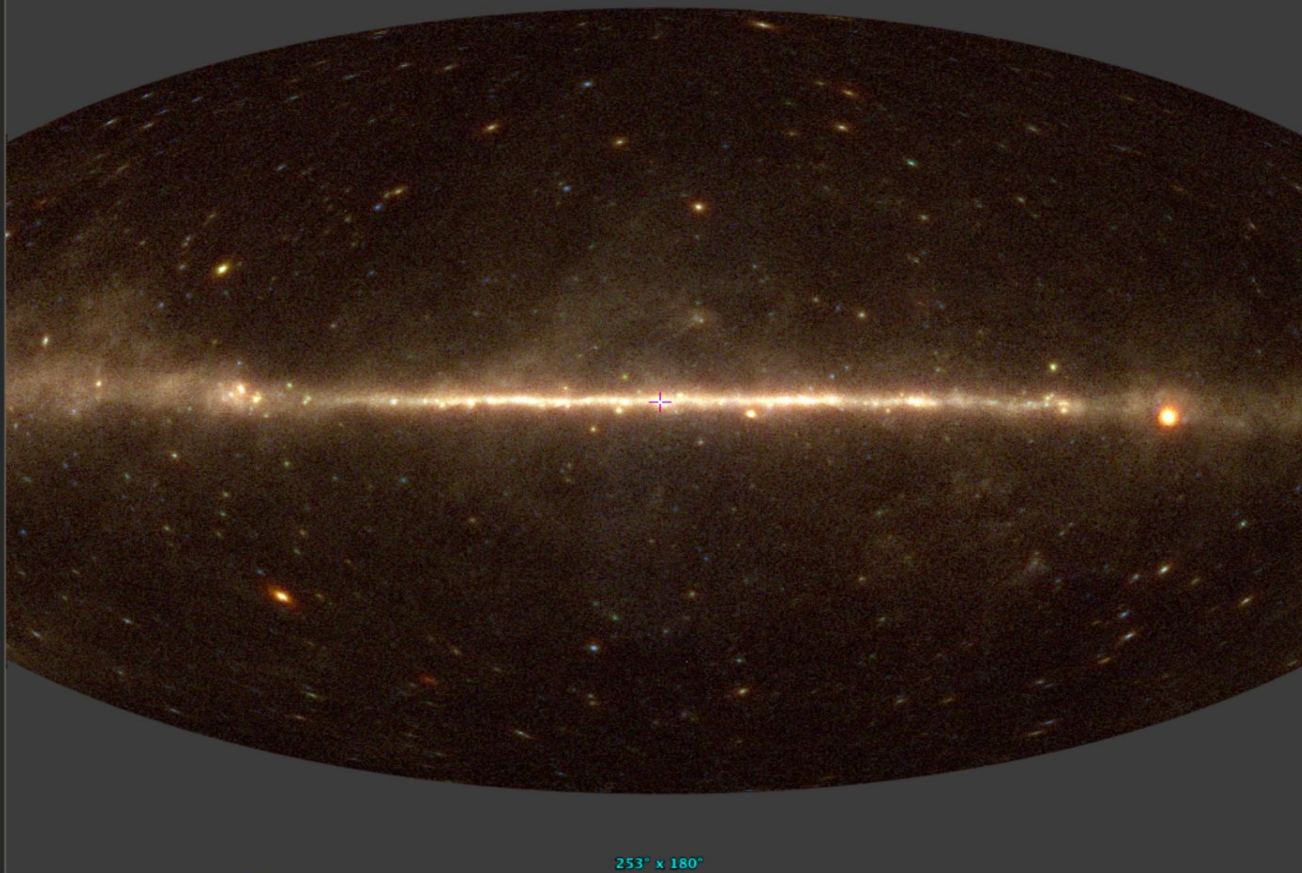
Export of the DataStore into an IVOA ObsCore table with proper metadata to build an IVOA TAP service



| dataproduct_type | calib_level | target_name | obs_id | obs_collection | obs_publisher_id | access_url | access_format | access_estsize | s_ra | s_dec | s_fov |
|-------------------------|--------------------|--------------------|---------------|-----------------------|--------------------------|-------------------|----------------------|-----------------------|----------------|----------------|----------------|
| | | | | | | | | kbyte | deg | deg | deg |
| str10 | int32 | str25 | str10 | str10 | str30 | str30 | str30 | int32 | float64 | float64 | float64 |
| EVENTS | 2 | AGN monitoring | 513837 | DL3 | ivo://ctao#<internal_id> | URL<internal_id> | application/fits | 1797 | 327.5722 | -14.7231 | 10.0 |
| EVENTS | 2 | AGN monitoring | 513839 | DL3 | ivo://ctao#<internal_id> | URL<internal_id> | application/fits | 1785 | 356.2607 | -16.4372 | 10.0 |
| EVENTS | 2 | AGN monitoring | 513833 | DL3 | ivo://ctao#<internal_id> | URL<internal_id> | application/fits | 1664 | 262.7 | -0.2026 | 10.0 |

Collections → 35132

Fermi color


 CDS / P / Fermi

epoch -

size -

dens. -

opac. -

zoom -

 359.94424 -00.04

 253° x 180°

 80

 no time filter sky

253° x 180°

 select
 from -- all collections --

exp. sort view scan filter

grid study wink redo north hdr multiview match

- ▼ Collections → 26 / 35132
 - ▼ Image → 2 / 620
 - ▼ Gamma-ray → 2 / 24
 - ▼ HESS → 2
 - HGPS significance
 - HGPS integral flux**
 - ▼ Catalog → 20 / 33068
 - Others → 4 / 1262
 - SSA (spectrum) → 2 / 159
 - padc.obspm.astro → 1 / 1
 - hess_ssa
 - vopdc.obspm → 1 / 3
 - High Energy Stereosc
 - TAP (table) → 2 / 220
 - nasa.heasarc → 1 / 35
 - RHESSI Gamma-Ray B
 - padc.obspm.he → 1 / 2
 - H.E.S.S. DL3 public te:

HGPS Flux



360° x 180°


 CDS / P / HGPS

 CDS / P / Fermi
epoch +size +dens. +opac. +zoom +

359.94023 -00.04

360° x 180°

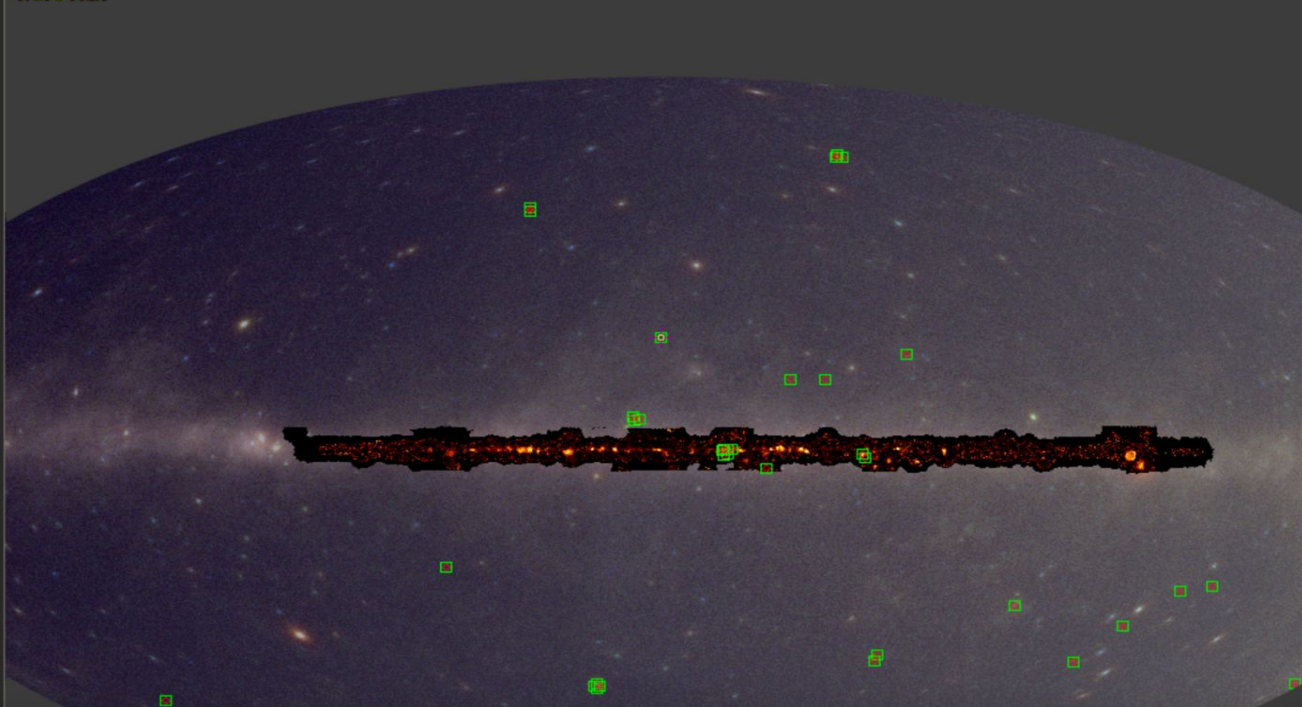
80

no time filter sky

0 sel / 0 src 32fps / 406Mb

- ▼ Collections → 26 / 35132
 - ▼ Image → 2 / 620
 - ▼ Gamma-ray → 2 / 24
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 - H.E.S.S. DL3 public te

HGPS Flux



obscore i

- Field: target_name
- Value: Sco X-1
- UCD: meta.id;src

name of target

select

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padc.obspm.he /

GDS/P / HGPS

Aladin Java measurements frame

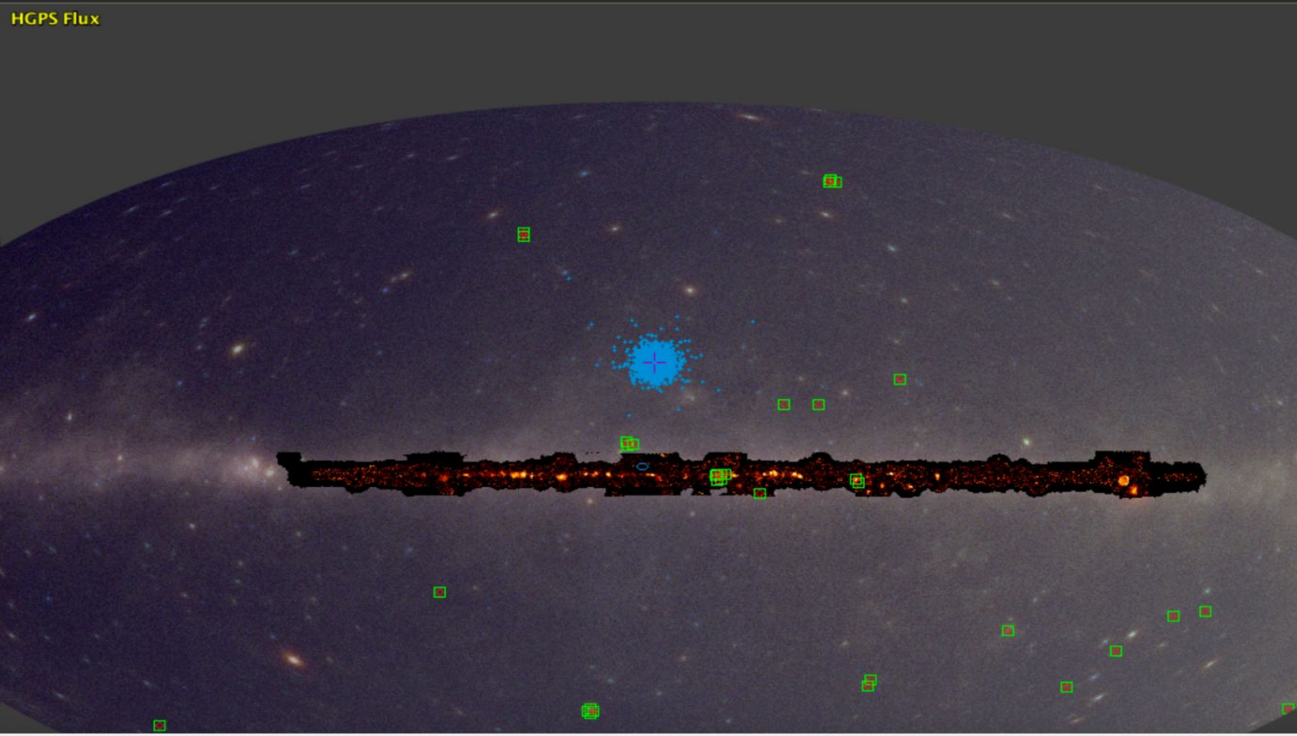
obscore - target_name: name of target

Search

| access_url | dataprodct_type | dataprodct_sub | calib_level | obs_collection | obs_id | obs_publisher_did | access_format | access_estsi | target_name | s_ra |
|---|-----------------|----------------|-------------|----------------|--------|---------------------------|------------------|--------------|-------------|--------------------|
| https://hess-dr.obspm.fr/ | event-list | events | 2 | HESS-DR | 26827 | ivo://padc.obspm/hess#... | application/fits | 241920 | Arp 220 | 234.5016784667... |
| https://hess-dr.obspm.fr/ | event-list | events | 2 | HESS-DR | 26791 | ivo://padc.obspm/hess#... | application/fits | 172800 | Arp 220 | 233.7383728027... |
| https://hess-dr.obspm.fr/ | event-list | events | 2 | HESS-DR | 26077 | ivo://padc.obspm/hess#... | application/fits | 239040 | Sco X-1 | 244.9916687011... |
| https://hess-dr.obspm.fr/ | event-list | events | 2 | HESS-DR | 25511 | ivo://padc.obspm/hess#... | application/fits | 233280 | 3C 273 | 187.77819824218... |
| https://hess-dr.obspm.fr/ | event-list | events | 2 | HESS-DR | 25443 | ivo://padc.obspm/hess#... | application/fits | 325440 | 3C 273 | 187.2778930664... |
| https://hess-dr.obspm.fr/ | event-list | events | 2 | HESS-DR | 25345 | ivo://padc.obspm/hess#... | application/fits | 334080 | 3C 273 | 187.2778930664... |
| https://hess-dr.obspm.fr/ | event-list | events | 2 | HESS-DR | 23736 | ivo://padc.obspm/hess#... | application/fits | 259200 | SN 1987A | 83.86775207519... |



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 - RHESSI Gamma-Ray B
 - ▼ padc.obspm.he → 1 / 2
 - H.E.S.S. DL3 public te



obscore ⓘ

- Field: access_url
- Value: https://hess-dr.obspm.fr/
- UCD: meta.ref.url

URL used to access dataset

- select
- pan
- dist
- phot
- draw
- tag
- moc
- spect
- filter
- cross
- x-y
- rgb
- assoc
- crop
- cont
- pixel
- prop

https://hess-dr.obspm.fr/

https://hess-dr.obspm.fr/

https://hess-dr.obspm.fr/

padc.obspm.he/

CDS/P/HGPS/

Aladin Java measurements frame

obscore - access_url: URL used to access dataset

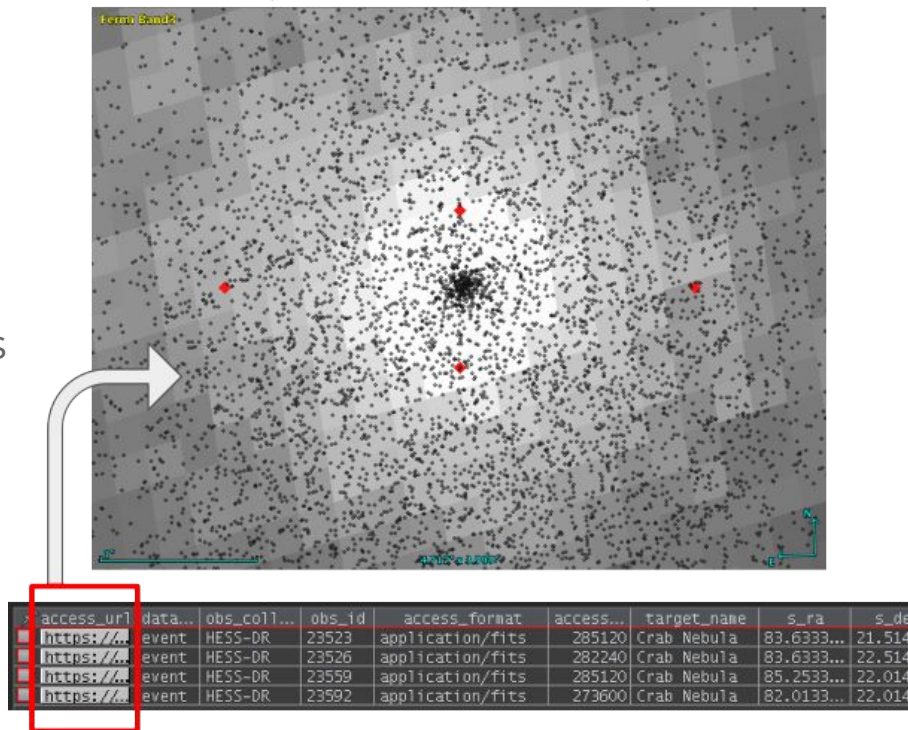
Search

| access_url | dataprodct_type | dataprodct_sub | calib_level | obs_collection | obs_id | obs_publisher_did | access_format | access_estsi | target_name | s_ra |
|--|-----------------|----------------|-------------|----------------|--------|---------------------------|------------------|--------------|-------------|--------------------|
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| https://hess-dr.obspm.fr/retrieve/hess_dl3_dr1_obs_id_026077.fits.gz | event-list | events | 2 | HESS-DR | 26077 | ivo://padc.obspm/hess#... | application/fits | 239040 | Sco X-1 | 244.9916687011... |
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A = Accessible

DL3 event list directly opened in Aladin
(each black dot is an event)

- ObsCore **access_url**
 - Direct download link to the FITS file
- IVOA **DataLink** (to be implemented):
 - Access to different storage services
 - Access to analysis services, previews
- **Access rights**
 - Public data: no restrictions
 - Anticipating need for **permissions**:
 - PI proprietary period
 - Federation authentication



I = Interoperable



A **Python** package for **gamma-ray** astronomy

A community initiative to define **common data formats for gamma-ray astronomy** based on FITS

<https://vodf.readthedocs.io>

<https://gamma-astro-data-formats.readthedocs.io>

- Includes formats for: event lists, effective area, energy resolution, point spread function, instrumental background...
- More and more used by current instruments: Fermi-LAT, HESS, VERITAS, MAGIC, FACT, ...

[A&A 625, A10, 2019] [A&A 632, A72, 2019] [A&A 632, A102, 2019]

Open-source Python package (Astropy affiliated package)

- Core library for the Science Tools of CTA
- Used in the analysis of existing gamma-ray instruments, such as H.E.S.S., MAGIC, VERITAS, HAWC...

FAIR4RS: FAIR Principles for Research Software

→ <https://doi.org/10.15497/RDA00065>

The road to an open format for VHE

The building phase of CTA gave us (quite a lot...) of time to develop its data and analysis tools



GADF

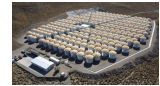
A (short History of GADF)

- 2011** Prototypes for the CTA data format and science tools
- 2016**
 - Establishment of the Gamma-ray Astronomy Data Formats (GADF) initiative
 - First preliminary release version (0.1), mainly focused on IACTs
- 2018**
 - Version 0.2 released
 - Support implemented in the science tools Gammapy and ctools
 - H.E.S.S. releases \approx 50 h of observations of different sources using the format
- 2019**
 - FACT, Fermi-LAT, H.E.S.S. MAGIC and VERITAS observations of the Crab Nebula are used to perform the first multi-instrument analysis
[doi:10.1051/0004-6361/201834938]
<https://github.com/open-gamma-ray-astro/joint-crab>
 - **ctools** based analysis of the H.E.S.S. data release
[doi:10.1051/0004-6361/201936010]
 - Comparison of Gammapy and ctools using the H.E.S.S. data release
[doi:10.1051/0004-6361/201936452]

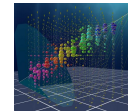
from Maximilian Linhoff

Validation of standardized data formats and tools for ground-level particle-based gamma-ray observatories

[doi:10.1051/0004-6361/202243527](https://doi.org/10.1051/0004-6361/202243527)



GADF @Gammapy



From B. Khelifi



Prospects for combined analyses of hadronic emission from γ -ray sources in the Milky Way with CTA and KM3NeT [arXiv:2309.03007](https://arxiv.org/abs/2309.03007) 15

From GADF to VODF

GADF has been a cornerstone step for interoperability and development, but it has problems...

- Additional parametrization is needed for IRFs (Event types, simulated datasets)
- Standard is vague instead of well described
- Interoperability is not complete or immediate
- There is no governing structure
- Lack of standardized metadata - the **format is not FAIR or VO compliant**



The GADF team decided to create a new more structured initiative with the intent of **linking the VHE community to VO standards**

VODF: a new structure for open science

<https://vodf.readthedocs.io>

Steering committee (11)
Lead editors (3)
Conveners (2)
Since 2022

GADF
Data Formats for
Gamma-Ray Astronomy

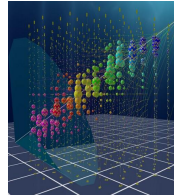
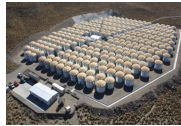
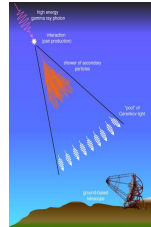


VODF
very-high-energy open data format



 A Python package for
gamma-ray astronomy

 **cherenkov
telescope
array**



VODF Steering committee

Steering Committee

| Facility | Category | Representative |
|-----------|-----------------------------------|--------------------|
| ASTRI | Pointing γ -ray instrument | Fabio Pintore |
| CTAO | Pointing γ -ray instrument | Roberta Zanin |
| FACT | Pointing γ -ray instrument | Maximilian Linhoff |
| Fermi-LAT | Slewing γ -ray instrument | Nicola Omodei |
| HAWC | Slewing γ -ray instrument | Xiaojie Wang |
| H.E.S.S. | Pointing γ -ray instrument | Bruno Khélifi |
| IceCube | Neutrino detector | Marcos Santander |
| KM3Net | Neutrino detector | Kay Graf |
| MAGIC | Pointing γ -ray instrument | Cosimo Nigro |
| SWGO | Slewing γ -ray instrument | Andrew Smith |
| VERITAS | Pointing γ -ray instrument | Amanda Weinstein |

Lead editors:

Karl Kosack (IACTs)

Laura Olivera-Nieto (Slewing instruments)

Jutta Schnabel (Neutrino instruments)

Conveners:

Bruno Khélifi

Roberta Zanin

What's next for **VODF**?

very-high-energy open data format

Now that the VODF structure is in place, it's time to take concrete steps towards the creation of the new standards for VHE

- Allow for well-defined usage of multiple IRFs in the same data processing
- Introduction of Event Types
- Define VO keywords with VHE specificities - strong base in X-Ray astronomy
- Set standards for **Metadata**, not just data
- Improve discoverability using VO recommendations - Observatory Core Component Data Model (**ObsCore**)
- Integrate completely into IVOA interest groups (**HEIG**)

 π A Python package for **gamma-ray** astronomy



From F-A-I to FAIR

F-A-I are very technical points.

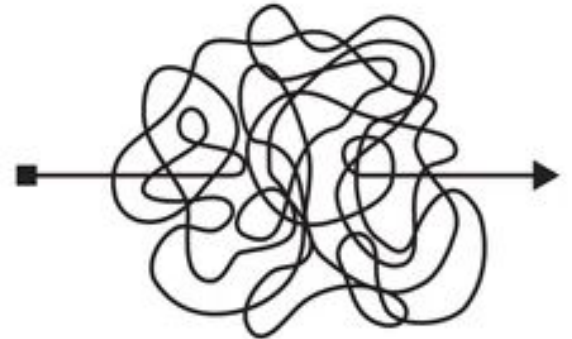
Reusability?

- Based on the **quality / reliability / trustworthiness** of the products
- What calibration was applied? What tools were used and how?
What assumptions were made during the data preparation?
- **Sustainability**: with time, key information may disappear...

Provenance information as an answer to reusability

- Need for the **origin, trace**, and detailed manipulations
- Need to **structure, store** it and **link** it to the data

→ **IVOA Provenance standard data model!**



Summary and conclusion

The VHE-multimessenger community strongly values **open, accessible and interoperable data and analysis software** and is taking priority steps towards it



- Built on the **big success of GADF** as an interoperable standard
- Will be **VO compliant**
- Has a defined governing structure

June 2023 in Paris

IVOA standards for High Energy Astrophysics
IVOA Note to justify a HE Interest Group

<https://github.com/ivoa/HighEnergyDataNote>

<https://indico.obspm.fr/event/1963>



BACKUP

Current High Energy data in the VO

- Images
 - Fermi Full Sky, eRosita, XMM-Newton, H.E.S.S. galactic plane
- Catalogs
 - Vizier dedicated section
 - NASA HEASARC (many legacy archives)
 - Generally catalogs of sources
- SIA/TAP services
 - High level data, catalogs, proposals, ...



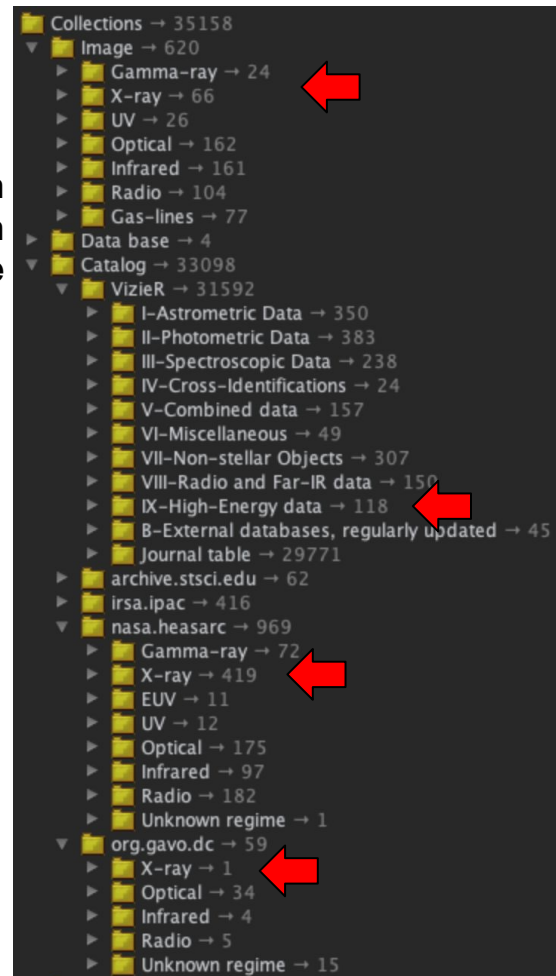
XMM-Newton Science Archive



TAP queries to the XSA database

NASA's HEASARC
High Energy Astrophysics Science Archive Research Center

Aladin
Search
Tree



A long way

It's a Long Way... from Private Ground-based Gamma-ray Data to Public Release: Open-data, Open-source Tools, First Real TeV Data Release from H.E.S.S.

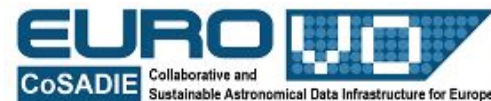
[C. Boisson et al. 2020, ADASS XXVII Santiago, ASPC 522 497B]

- Imaging Atmospheric Cherenkov Telescopes
- Gamma **data format** initiative
- **Open tools** for analysis

→ A story of the relations between Cherenkov observatories and the IVOA



Cherenkov
telescope
array



2012-2014

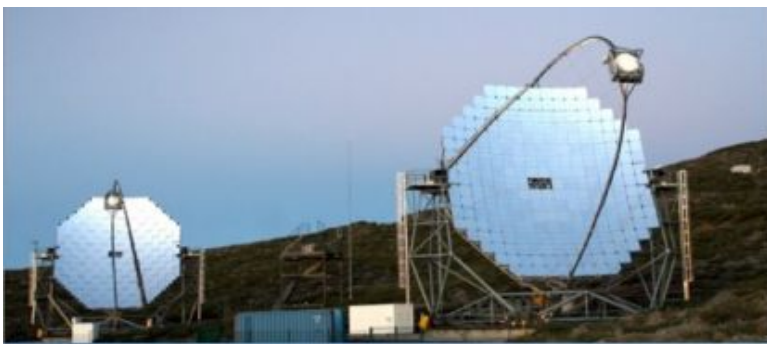


2015-2019



2019-2023

Cherenkov Astronomy



MAGIC: located in La Palma, Spain
Since 2004: single 17m telescope
Since 2009: system of two 17m telescopes

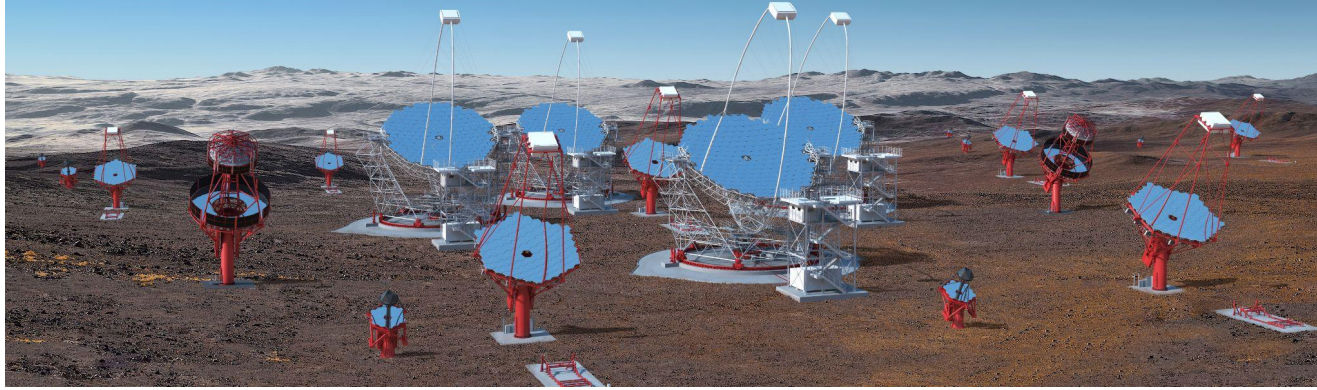


VERITAS: located in Mt Hopkins, Arizona
Since 2007: four 12m telescopes
Since 2012: upgraded PMTs



H.E.S.S.: located in Khomas Highlands, Namibia
Since 2002: four 12m telescopes
Since 2012: added 32m by 24m telescope
Since 2015: camera upgrades on 12m telescopes

Atacama, Chile



La Palma, Spain



Alpha configuration

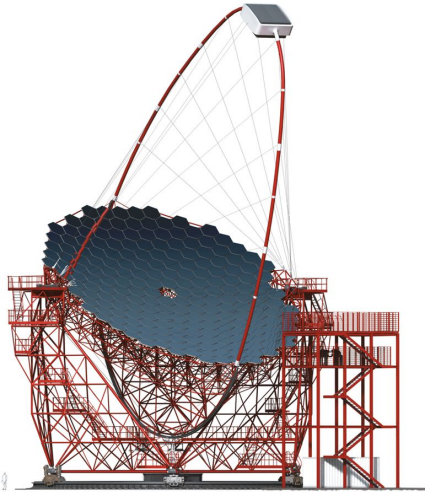
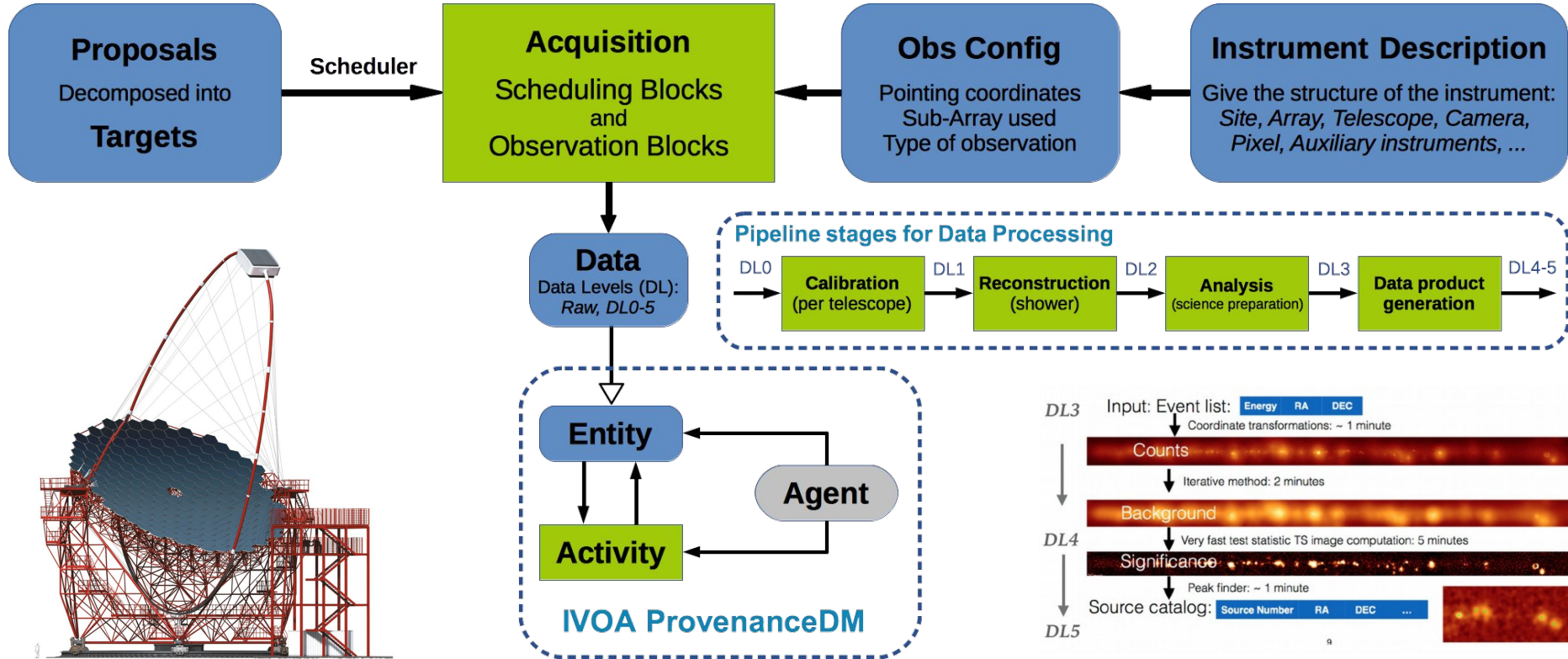
CTAO Southern array
51 telescopes over a ~ 3 km² area

14 **Medium-Sized Telescopes (MST)**
37 **Small-Sized Telescopes (SST)**

CTAO Northern array
13 telescopes distributed over an
area of about 0.5 km²

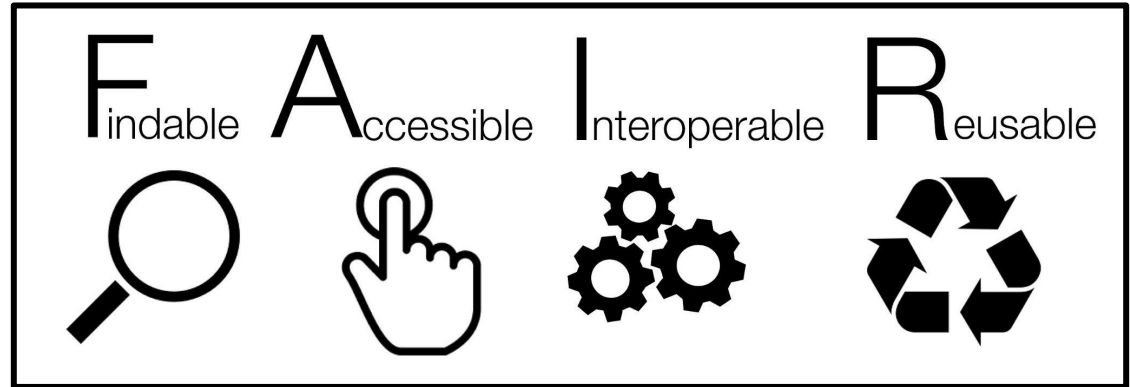
4 **Large-Sized Telescopes (LSTs)**
9 **Medium-Sized Telescopes (MSTs)**

CTA Master Configuration Data Model

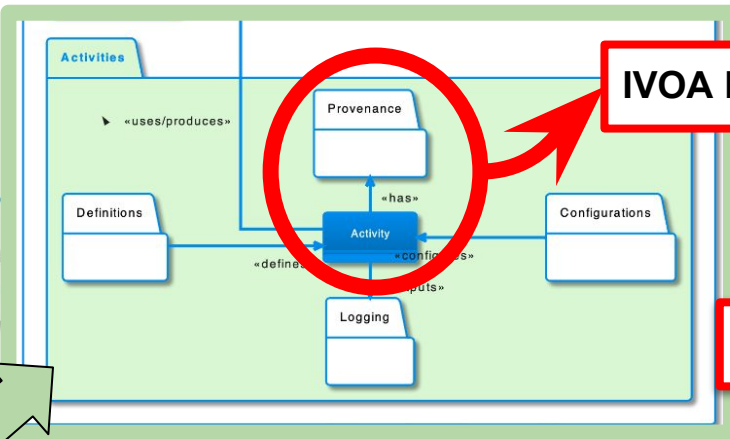
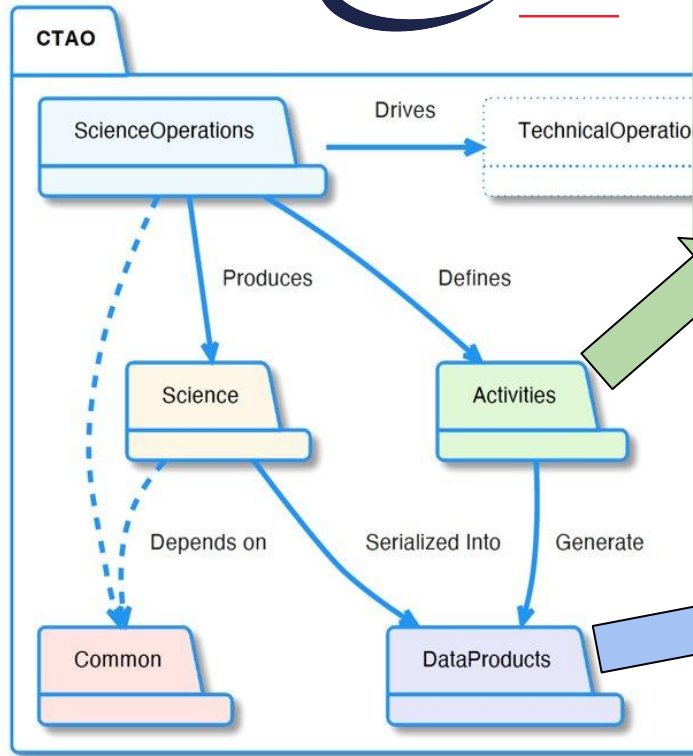


Open Observatory and Open Science

CTA will operate as an **open observatory** and will provide data to the scientific community. In the context of **Open Science**, the data provided by CTA must follow the **FAIR Guiding Principles** for scientific data management:

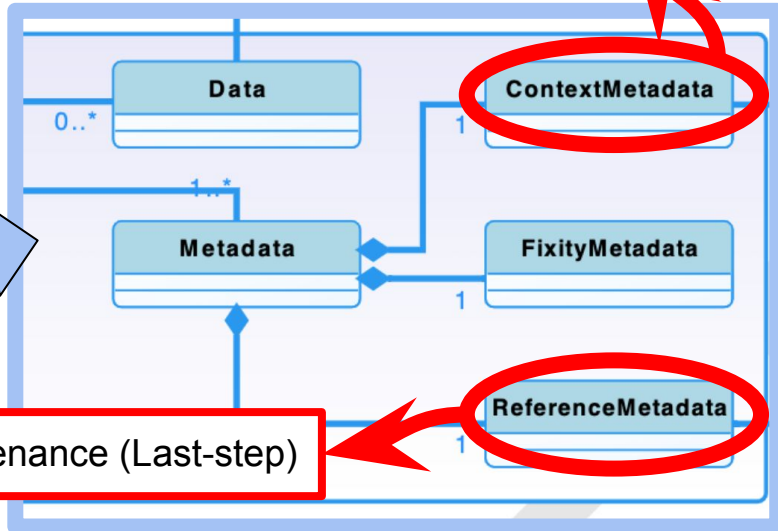


See e.g. Servillat et al. 2022, [ADASS XXXI, ASP Conference Series](#)
"FAIR high level data for Cherenkov astronomy"
<https://hal-obspm.ccsd.cnrs.fr/obspm-03516688>



IVOA Provenance (Processing)

IVOA ObsCore metadata

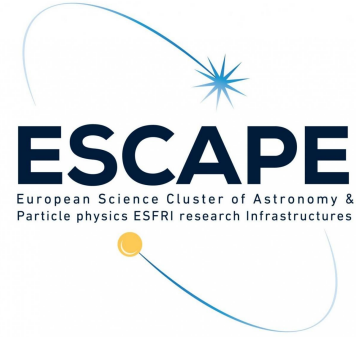


IVOA Provenance (Last-step)

CTAO Top-level Data Model

A long way...

- Interactions between observatories and the IVOA
 - Astronomers and engineers need to be part of IVOA
 - They then need to convince their organisation!
 - And maintain the momentum!
- Early engagement of observatories/facilities
 - Construction and operations = different priorities
 - Seeds in the CTA requirements and data models
 - CTAO involved in ESCAPE open collab., OSTrails...
 - Role of European Projects
 - Role of **OV-France**



HE dedicated workshop at OV-France

- **October 2022 in Strasbourg**
 - <https://indico.obspm.fr/event/1489>
- Continue activities of the **ESCAPE European project** that focused on High Energy Facilities (**2019-2023** H2020 project).
- Bring together representatives of high energy observatories (VHE, HE, GW, neutrino)
- Presentations of HE observatory operations and data:
 - CTA (Mathieu Servillat)
 - Ligo Virgo Kagra (Pierre Chaniel)
 - Neutrino (Damien Dornic)
 - XMM & SVOM (Laurent Michel)
 - GADF/VODF (Bruno Khelifi)



HE meeting at IVOA Interop

- **May 2023 at the IVOA Bologna**

- Dedicated talk at the DM session by M. Servillat:
 - <https://wiki.ivoa.net/internal/IVOA/IntropMay30>
- Fruitful splinter IVOA meeting



- **Creation of HE “Club”**

- IVOA mailing list and wiki page
 - <https://wiki.ivoa.net/twiki/bin/view/IVOA/HEGroup>
- Several online meetings
 - IVOA data models (cube, dataset)
 - Detailed of HE data, with a focus on Instrument Response Functions
- Next call: **April 3rd 16h30**

Second OV-France workshop enlarged to IVOA

- **June 2023 in Paris**
 - *IVOA standards for High Energy Astrophysics*
 - <https://indico.obspm.fr/event/1963>
- Review of previous documents since 2021
- Focus on **user scenarios** in HE
 - Access and Analysis of HE data
 - Used IVOA standards
 - What specific developments are needed
- Prepare an IVOA Note to justify a HE Interest Group
 - <https://github.com/ivoa/HighEnergyDataNote>



Content of the Note

- Common practices and concept of event-list
 - **Lower level** dataset, used to generate images, lightcurves, spectra
 - Generally **reprocessed** from **event lists** for a dedicated analysis
 - Calibrated data, but instrument signature not totally removed
 - **Instrument Response Functions (IRFs)** are tightly connected
- Data Discovery
 - **ObsCore** for a HE event list?
 - Possible extension for HE
- An event-list context data model
 - **Relations** to IRF and Instrument Configuration
- Modelling the content of an event-list
 - Cube and Dataset Data Model



*International
Virtual
Observatory
Alliance*

**Virtual Observatory and High Energy
Astrophysics**

Version 0.3

IVOA Note 2024-01-31

HE “event” in the VO



a HE event is **not** a VOEvent

<https://www.ivoa.net/documents/ObsCore>

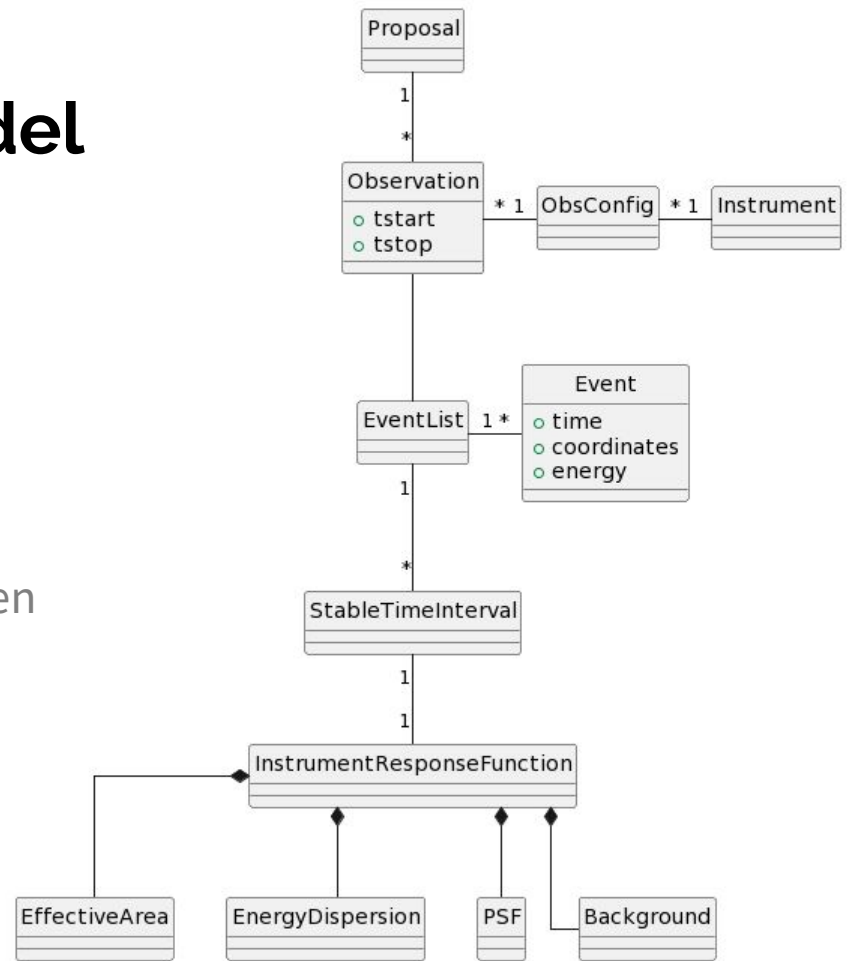
event: An event-counting (e.g. X-ray or other high energy) dataset *of some sort*. Typically this is instrumental data, i.e., "event data". An event dataset is *often a complex object* containing multiple files or other substructures. An event dataset *may contain* data with spatial, spectral, and time information for each measured event, although the spectral resolution (energy) is sometimes limited. Event data may be used to produce higher level data products such as images or spectra.

<http://www.ivoa.net/rdf/product-type> (Preliminary)

event-list: A collection of observed events, such as incoming high-energy particles. A row in an event list is typically characterised by a spatial position, a time and an energy.

Event-list Context Data Model

- Issue
 - What is really in the event-list dataset?
 - Does it include IRFs? only an event-list?
 - Where can one find the corresponding IRFs?
 - Need a way to link an event-list to its IRFs
- A proper **data model** with relations between those elements would help
- Possible **ObsCore extension** fields would appear in this data model

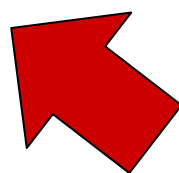


Summary and conclusions

- **HE data available** through the VO
- HE domain have **specificities**
 - In particular the concept of **event-list**
 - and **Instrument Response Functions** (IRFs)
- Very limited number of VO services giving access to event-lists
- On the path to build an **HE Interest Group** at IVOA
 - **Session at IVOA Sydney** in May (plenary or with data model)
 - Next call: April 3rd 16h30



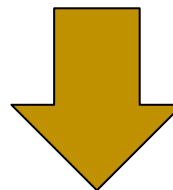
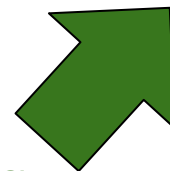
Cherenkov
telescope
array



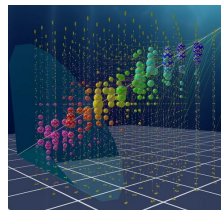
Pointing
Gamma-Ray
instruments

Slewing
Gamma-Ray
instruments

Neutrino instruments



KM3NeT



ICECUBE
SOUTH POLE NEUTRINO OBSERVATORY

Science Analysis: DL3-DL5

Happens at CTA data centers (automatic) + by users on user's laptops or e.g. ESCAPE science platform



Search for data covering region of interest in space, time, maybe other parameters

VO tools?

DL3

Science-Ready

Retrieve Event Lists and IRFs that cover region of interest

obs_id = 84753

| event_id | Time | Energy | RA | Dec |
|----------|----------|--------|------|-------|
| 12351 | 55933.30 | 0.03 | 53.6 | -29.2 |

obs_id = 99584

| event_id | Time | Energy | RA | Dec |
|----------|----------|--------|------|-------|
| 12351 | 55933.30 | 0.03 | 53.6 | -29.2 |
| 12356 | 55933.31 | 0.1 | 53.2 | -29.3 |
| 12378 | 55933.36 | 12.0 | 53.8 | -29.5 |
| 12389 | 55933.45 | 0.2 | 53.7 | -28.7 |

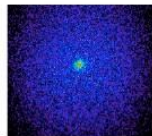
obs_id = 100207

| event_id | Time | Energy | RA | Dec |
|----------|----------|--------|------|-------|
| 12351 | 55933.30 | 0.03 | 53.6 | -29.2 |
| 12356 | 55933.31 | 0.1 | 53.2 | -29.3 |
| 12378 | 55933.36 | 12.0 | 53.8 | -29.5 |
| 12389 | 55933.45 | 0.2 | 53.7 | -28.7 |

DL4

Science-Binned

Use Science Tools to make intermediate Binned Data Products

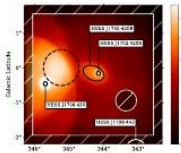


Exposure Cube
Counts Cube
Exclusion Map
...

DL5

Science-Quicklook

Use Science Tools to Fit models to binned data (forward-folding) and make Flux cubes



Sky model
Light-Curve
Spectra
Flux Map/Cube

DL6
Science-Observatory
Observatory or CTAC produces some legacy products

Catalogs,
Survey Products
Karl Kosack



A Python package for gamma-ray astronomy