



Parity-odd terms in the trace anomaly of Weyl fermions: elusive or delusive?

S. Franchino-Viñas

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COST Action CA18108:

Quantum gravity phenomenology in the multi-messenger approach

<https://qg-mm.unizar.es/>

Quantum gravity phenomenology at the dawn of the multi-messenger era—A review, Prog.Part.Nucl.Phys. 125 (2022) 103948,
[arXiv:2111.05659](https://arxiv.org/abs/2111.05659)

Is there an imaginary term in the trace anomaly of Weyl fermions id $D = 4$?

① Anomalies

② Chiral fermions and anomalies

③ The trace anomaly

④ Outlook

① Anomalies

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④ Outlook

Consider a classical field theory. If the system under study possesses a continuous symmetry, then from Noether's theorem (in a classical, i.e. non-quantum theory)

$$\exists j^\mu, \partial_\mu j^\mu \stackrel{*}{=} 0. \quad (1)$$

If the system is quantized, $\partial_\mu \langle \hat{j}^\mu \rangle \stackrel{?}{=} 0$?

If $\neq 0$, the symmetry is said to be anomalous.

- Pion decay (Bell and Jackiw): $\pi \rightarrow \gamma\gamma$ (anomaly in J_A^μ).
- Applications in condensed matter/hydrodynamics (review Chernodub et al., 2021, experiment Gooth et al., 2017): $\vec{J}_\epsilon = (a_{F\tilde{F}} + a_{R\tilde{R}} T^2) \vec{B}$ (anomaly in J_A^μ).
- Kuzmin, Rubakov and Shaposhnikov (1985); still new proposals as QCD baryogenesis [1911.01432v2 [hep-ph]].
- Bouncing universes (instead of Big Bang) from anomalies, Fabris, Pelinson, Shapiro [gr-qc/9810032], Asorey et al. [2202.00154], Camargo and Shapiro [2206.02839] (trace anomaly).
- Axionic dark matter (Basilakos, Mavromatos, Solà [gr-qc/2001.03465]) (anomaly for a Kalb–Ramond field).
- BRST formalism and cohomological aspects.

Anomalies

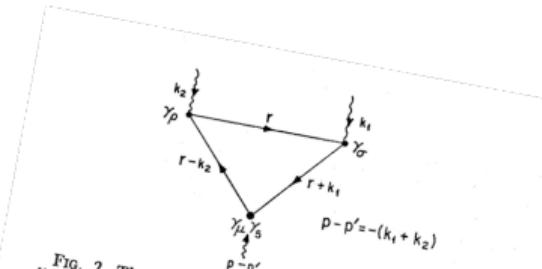
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Consider a massless fermion, coupled to an *external* EM field A_μ

$$\mathcal{L} = i\bar{\psi}\gamma^\mu \partial_\mu \psi + e J_V^\mu A_\mu, \quad J_V^\mu = \bar{\psi}\gamma^\mu \psi, \quad J_A^\mu = \bar{\psi}\gamma^\mu \gamma_5 \psi \quad (2)$$

The system has two global symmetries: $\psi \rightarrow e^{i\alpha}\psi, e^{i\alpha\gamma_5}$, so that classically (on-shell) $\partial_\mu J_{A,V}^\mu = 0$.

$$\begin{aligned} \langle \partial_\mu J_A^\mu(x) \rangle &\sim \int \mathcal{D}\bar{\psi} \mathcal{D}\psi \partial_\mu J_A^\mu e^{i \int i\bar{\psi}\not{\partial}\psi + e J_V^\mu A_\mu} \\ &= -\frac{e^2}{2} \partial_{x^\mu} \iint dy dz \langle J_A^\mu(x) J_V^\nu(y) J_V^\sigma(z) \rangle A_\nu(z) A_\sigma(y) + \dots \quad (3) \\ &= -\frac{e^2}{16\pi^2} \epsilon^{\alpha\beta\sigma\lambda} F_{\alpha\beta} F_{\sigma\lambda}. \end{aligned}$$



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Dimensional dependent properties

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Suppose we define $\gamma_*^{(4)} = -\frac{i}{4!} \epsilon_{\mu\nu\rho\sigma} \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma$ in $n = 4$ dimensions. Then

$$\begin{aligned}\text{tr}(\gamma_*^{(4)} \gamma_\mu \gamma_\nu \gamma_\rho \gamma_\sigma) &= -i \epsilon_{\mu\nu\rho\sigma} \text{tr} \mathbb{1}, \\ \{\gamma_\mu, \gamma_*^{(4)}\} &= 0.\end{aligned}\tag{4}$$

Consider now γ_* in an arbitrary n with the same anticommutators + cyclicity of the trace:

$$n \text{tr} \gamma_* = \text{tr}(\gamma_* \gamma^\alpha \gamma_\alpha) = -\text{tr}(\gamma^\alpha \gamma_* \gamma_\alpha) = -\text{tr}(\gamma_* \gamma_\alpha \gamma^\alpha) = -n \text{tr} \gamma_* \implies n \text{tr} \gamma_* = 0.\tag{5}$$

Mutatis mutandis...

$$\begin{aligned}n(n-2) \text{tr}(\gamma_* \gamma_\mu \gamma_\nu) &= 0, \\ n(n-2)(n-4) \text{tr}(\gamma_* \gamma_\mu \gamma_\nu \gamma_\rho \gamma_\sigma) &= 0, \\ &\vdots\end{aligned}\tag{6}$$

which contradicts formula (4)!

with

$$(18) \quad A = -\frac{e^2}{(4\pi)^{n/2}} \cdot m^{v-3} \cdot \Gamma\left(2 - \frac{v}{2}\right) \frac{v-1}{v-3},$$

$$(19) \quad B = -\frac{e^2}{(4\pi)^{n/2}} \cdot m^{v-1} \cdot \Gamma\left(2 - \frac{v}{2}\right) \frac{v-1}{v-3}$$

and

$$(20) \quad \Sigma_r(p, v) = \frac{e^2}{(4\pi)^{n/2}} \cdot m^{v-2} \cdot \Gamma\left(2 - \frac{v}{2}\right),$$

$$\cdot \left[(2-v) \frac{m - i\gamma^\mu p}{m^2 \theta} - 4 \frac{m - i\gamma^\mu p}{m^{2-v}} + \frac{2}{m} \right], \quad 2 - v < 0.$$



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IL NUOVO CIMENTO

Dimensional Renormalization: The Number of Dimensions as a Regularizing Parameter.

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(ricevuto 18 Febbraio 1972)

Summary. — We perform an analytic extension of quantum electrodynamics matrix elements as (analytic) functions of the number of dimensions of space (v). The usual divergences appear as poles for v integer. The renormalization of those matrix elements (for v arbitrary) leads to expressions which are free of ultraviolet divergences for v equal to 4. This shows that v can be used as an analytic regularizing parameter with advantages over the usual analytic regularization method. In particular, gauge invariance is maintained for any v .

1. — Introduction.

In a previous paper (1) we pointed out the possibility of studying the structure of the theory as a function of the number of dimensions. We

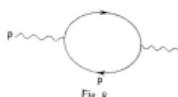


Fig. 8.

$$\int d_\mu p \frac{\text{Tr} [\gamma^\mu (\gamma_\nu (p+k) + m) \gamma^\nu (\gamma_\mu p + m)]}{((p+k)^2 + m^2)(p^2 + m^2)}.$$

The trace may be evaluated using (33), (34) and (35). Taking denominators together one obtains:

$$4 \int_0^1 dx \int d_\mu p \frac{\delta_{\mu\nu}(m^2 + p^2 + pk) - 2p_\mu p_\nu - p_\mu k_\nu - k_\mu p_\nu}{(p^2 + 2pkx + k^2 x + m^2)^2}.$$

Using the equations of appendix A:

$$= \frac{4i\pi^{\frac{1}{2}n}}{\Gamma(2 - \frac{1}{2}n)} \int_0^1 dx \frac{2x(1-x)(k_\mu k_\nu - \delta_{\mu\nu} k^2)}{(m^2 + k^2 x(1-x))^{2-\frac{1}{2}n}}$$

which is manifestly gauge invariant,

6. LIMITATIONS OF THE METHOD

The method fails if in the Ward identities there appear quantities that have the desired properties only in four dimensional space. An example is the completely antisymmetric tensor $\epsilon_{\mu\nu\rho\sigma}$. If the particular properties of this tensor are vital for the Ward identities to hold our method will fail because we cannot generalize $\epsilon_{\mu\nu\rho\sigma}$ to a tensor satisfying the required properties for non-integer n . Similarly for γ^5 . One can write:

$$\gamma^5 = \frac{1}{4!} \epsilon_{\mu\nu\rho\sigma} \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma.$$

Insert this whenever γ^5 occurs and take the ϵ -tensor outside of the expression to be generalized to non-integer n . However, if we are dealing with Ward identities that re-

$$\begin{aligned} \{\gamma^5, \gamma^\alpha\} &= 0 & \text{for} & & \alpha = 1, \dots, n \\ \text{Tr}(\gamma^5 \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma) &= 4\epsilon_{\mu\nu\rho\sigma} \end{aligned}$$



masses of charged massive vector bosons. The derived Feynman rules indicate that particles, and in order to establish unitarity and causality, more terms are needed. The necessary counter

Do we need to consider γ_* ? The standard model is written in terms of chiral fermions.

We need a solution to this γ_* problem:

- give up the cyclicity of the trace (Kreimer+).
- give up $\{\gamma_\mu, \gamma_*\} = 0$
 - give up all (Thompson-Yu);
 - keep the first four (Breitenlohner-Maison);

BM scheme

- n -dimensional usual metric $\eta_{\mu\nu}, \gamma_\mu, p_\mu, \dots$
- decompose them into a four-dimensional part (\bar{X}) and an $(n - 4)$ -dimensional part (\hat{X}):
$$\begin{aligned}\eta_{\mu\nu} &= \bar{\eta}_{\mu\nu} + \hat{\eta}_{\mu\nu}, & \gamma_\mu &= \bar{\gamma}_\mu + \hat{\gamma}_\mu, & p_\mu &= \bar{p}_\mu + \hat{p}_\mu, \dots \\ \eta_\mu{}^\nu \hat{\eta}_{\nu\rho} &= \hat{\eta}_{\mu\nu} \hat{\eta}^\nu{}_\rho = \hat{\eta}_{\mu\rho}, & \hat{\eta}_{\mu\nu} &= \hat{\eta}_{\nu\mu}, \\ \bar{\eta}^{\mu\nu} \hat{\eta}_{\nu\rho} &= 0, & \bar{\eta}_\mu{}^\nu p_\nu &= \bar{p}_\mu, & \hat{\eta}_\mu{}^\nu \gamma_\nu &= \hat{\gamma}_\mu, \dots\end{aligned}$$
- $\epsilon_{\mu\nu\rho\sigma}$ is a purely four-dimensional object:
 $\epsilon_{\mu\nu\rho\sigma} = \bar{\epsilon}_{\mu\nu\rho\sigma}, \hat{\eta}^{\alpha\mu} \epsilon_{\mu\nu\rho\sigma} = 0.$
- γ_* the same as in four dimensions,

$$\gamma_* \equiv -\frac{i}{4!} \epsilon_{\mu\nu\rho\sigma} \gamma^\mu \gamma^\nu \gamma^\rho \gamma^\sigma.$$

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The trace anomaly

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Consider a field theory in curved space, and analyze the (local) symmetry under changes of scales

$$g_{\mu\nu} \rightarrow e^{-2\sigma} g_{\mu\nu}, \Phi_i \rightarrow e^{k_i \sigma} \Phi_i, \quad (7)$$

where the ($n = 4$) Weyl weights for matter fields are $k_i = (-1, -3/2, 0)$ for scalar, fermionic and vector fields.

$$\begin{aligned} \frac{\delta S}{\delta \sigma(x)} &= \int d^4y \frac{\delta g_{\mu\nu}(y)}{\delta \sigma(x)} \frac{\delta S}{\delta g_{\mu\nu}(y)} + \# \text{EOM} \\ &= -2g_{\mu\nu} \frac{\delta S}{\delta g_{\mu\nu}}(x) + \# \text{EOM} \\ &= -\sqrt{-g} g_{\mu\nu} T^{\mu\nu} + \# \text{EOM} \end{aligned} \implies \text{if scale invariant, } T^{\mu}_{\mu} \stackrel{*}{=} 0.$$

Does it survive the quantization, i.e. $g_{\mu\nu} \langle T^{\mu\nu} \rangle = 0$?

Trace Anomalies in Dimensional Regularization.

D. M. CAPPEI

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M. J. DUFF

Department of Astrophysics, University of Oxford - Oxford

(ricevuto il 22 Gennaio 1974)

Summary. — An unusual perturbation theory anomaly is pointed out. If there exists a trace identity valid in an arbitrary number of dimensions, then employing dimensional regularization can result in an amplitude satisfying the identity in an arbitrary number of dimensions, but the finite part of the amplitude violating it in four dimensions. An example given here is the one-loop neutrino contribution to the graviton propagator. Anomalous behaviour, of a different origin, also occurs in the one-loop photon contribution. Both kinds of anomaly can be removed at the expense of introducing a-dimensional, rather than 4-dimensional, counterterms.

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Dimensional regularization⁽¹⁻³⁾, the scheme whereby the dimension of spacetime is used as a regulating parameter in dealing with the divergences of quantum field theory, has gained a good deal of popularity. In gauge theories especially, where it is essential that the regularization scheme respect the Ward identities, this method seems particularly appealing.

In this paper, however, we wish to issue a word of warning. Dimensional regularization, at least in its conventional form, can and does give rise to anomalies. That is to say, symmetries present in the original Lagrangian are

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$$S_g[g, \psi, \bar{\psi}] = \int dx \left[\frac{i}{4} \sqrt{g} \bar{\psi} D^\mu \gamma_\mu (1 + \gamma^5) \bar{\psi} \right],$$

the covariant derivative and $D^\mu(x)$ is the vierbein field

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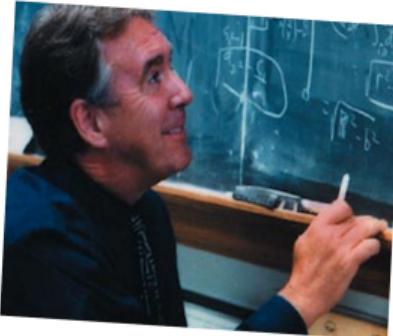
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$$Q_{\mu\nu}^{\text{trans}} \neq 0$$

with the requirements of conformal invariance. In the photon loop example, although we can patch up objection B) \rightarrow Introduction, dimensional regularization inevitably « fails » (1). In our second example, the converse would via objection A). In our second example, the converse

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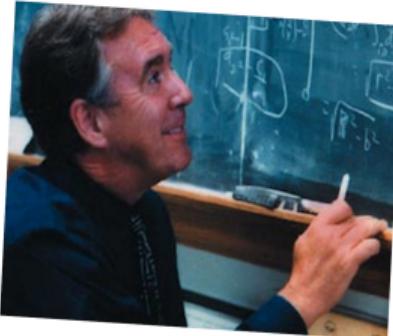
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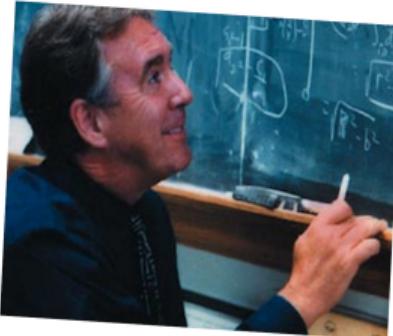
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M. CAPPEI and M. J. DUFF: Imperial College preprint, ICTP/T3/12.

[1803:09764 [hep-th] Bruque, Cherchiglia and Pérez-Victoria]: consider a two-dimensional space in which the following integral appears in a perturbative expansion.

$$I_{\mu\nu} = \int d^2k \frac{k_\mu k_\nu}{(k^2 + m^2)^2}. \quad (8)$$

We need to make sense of it by applying some regularization $R : I_{\mu\nu} \rightarrow [I_{\mu\nu}]^R$.

$$\begin{aligned} [I_{\mu\nu}]^R &= \frac{1}{2} \left[\int d^2k \frac{\delta_{\mu\nu}}{k^2 + m^2} - \int d^2k \partial_{k^\mu} \left(\frac{k_\nu}{k^2 + m^2} \right) \right]^R \\ &= \frac{1}{2} \left[\int d^2k \frac{\delta_{\mu\nu}}{k^2 + m^2} \right]^R \qquad \qquad \qquad \Rightarrow \qquad \qquad \delta^{\mu\nu} [I_{\mu\nu}]^R = [\delta^{\mu\nu} I_{\mu\nu}]^R + \pi. \\ &= \frac{\delta_{\mu\nu}}{2} \left[\int d^2k \frac{k^2}{(k^2 + m^2)^2} + \frac{m^2}{(k^2 + m^2)^2} \right]^R \\ &= \frac{\delta_{\mu\nu}}{2} [I_{\alpha\alpha} + \pi]^R, \end{aligned}$$

Anomaly induced action

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In $n = 4$ spacetime dimensions, on dimensional grounds

$$g^{\mu\nu}\langle T_{\mu\nu} \rangle = wC^2 + bE_4 + c\square R + \alpha R^2 + \beta F^2,$$

E_4 : Euler characteristic,

C : Weyl tensor,

w, b, c, α, β : depend on the theory.

Consider semi-classical gravity: integrate out the (free) matter fields.



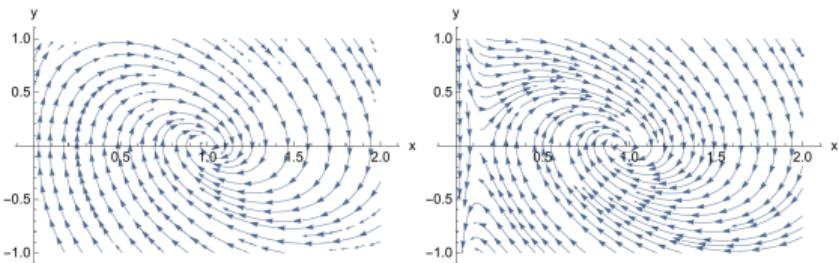
If the matter action is classically conformal invariant, the induced action for the (classical) gravitational field should satisfy the variational equation

$$\frac{\delta\Gamma_{\text{ind}}}{\delta\sigma(x)} = -\sqrt{-g}g_{\mu\nu}\langle T^{\mu\nu} \rangle.$$

We can integrate the anomaly to obtain Γ_{ind} !

Inflation, bouncing driven by quantum effects:

- Fabris, Pelinson, Shapiro [gr-qc/9810032]
- Hawking, Hertog, Reall [hep-th/0010232]
- ...



(©Shapiro+ [hep-th/2012.10554], $y \sim \dot{H}$, $x \sim H$)

- Nakayama [1201.3428]: what if a Pontryagin term ($P = \tilde{R}R$) ...
- Bonora, Giaccari, Lima de Souza, +++, [arXiv:1403.2606, 1503.03326, 1703.10473, ..., 2207.03279]: there is a P contribution to the chiral trace anomaly, which is purely imaginary ($\frac{i}{180(4\pi)^2} \frac{15}{4} P$). Diagrammatic, heat-kernel (axial gravity).
- Bastianelli, Martelli, Broccoli [arXiv:1610.02304, 1911.02271, 2203.11668]: Pauli-Villars regularization, no P contribution.
- Abdallah, SF, Fröb [arXiv:2304.08939, arXiv:2101.11382]: diagrammatic, no P contribution.
- Duff [arXiv:2003.02688]: no computation.
- Discussion also on $F\tilde{F}$ contributions to trace anomaly.

Chiral trace anomaly

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We consider a fermion in curved space,

$$S = - \int \bar{\psi} \gamma^\mu \nabla_\mu \psi \sqrt{-g} d^n x, \quad (9)$$

where as usual we introduce the vielbein

$$\gamma^\mu \equiv e^\mu{}_b \gamma^b,$$

and the covariant derivative involves the spin connection

$$\omega_{\mu\rho\sigma} = \eta_{ab} \left(e_\sigma{}^a \partial_{[\mu} e_{\rho]}{}^b - e_\rho{}^a \partial_{[\mu} e_{\sigma]}{}^b + e_\mu{}^a \partial_{[\sigma} e_{\rho]}{}^b \right).$$

The fermion is chiral:

$$\psi = \mathcal{P}_+ \psi \equiv \frac{1}{2} (\mathbb{1} + \gamma_*) \psi, \quad \bar{\psi} = \bar{\psi} \mathcal{P}_- \equiv \frac{1}{2} \bar{\psi} (\mathbb{1} - \gamma_*) .$$

The EM tensor is given by

$$T^{\mu\nu} = \frac{1}{2} \bar{\psi} \mathcal{P}_- \gamma^{(\mu} \overleftrightarrow{\nabla}^{\nu)} \mathcal{P}_+ \psi + \frac{1}{2} g^{\mu\nu} \bar{\psi} \mathcal{P}_- \gamma^\rho \overleftrightarrow{\nabla}_\rho \mathcal{P}_+ \psi,$$

which is classically traceless (on-shell), since the action is scale invariant.

How can we compute $\langle T^{\mu\nu} \rangle$?

Chiral trace anomaly

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The simplest thing we know to do is to perform an expansion around a given metric

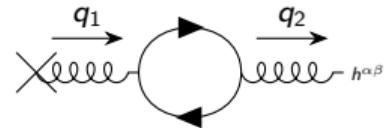
$$g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu}$$

↓

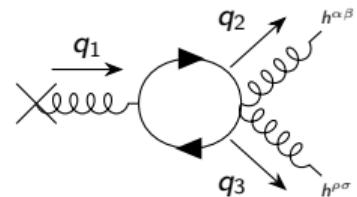
$$e^\mu{}_a = e^{(0)\rho}{}_a \left(\eta^\mu{}_\rho - \frac{1}{2} \kappa h^\mu{}_\rho + \frac{3}{8} \kappa^2 h^{\mu\sigma} h_{\sigma\rho} \right) + \mathcal{O}(\kappa^3),$$

$$g^{\mu\nu} = \dots,$$

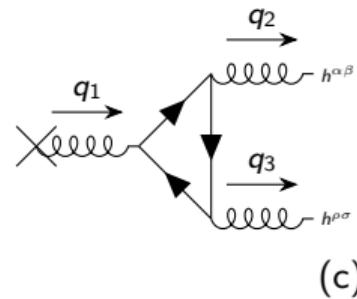
$$\begin{aligned} \langle T^{\mu\nu}(x) \rangle_g &= \frac{\langle T^{\mu\nu}(x) \exp [i(\kappa S_{(1)} + \kappa^2 S_{(2)})] \rangle}{\langle \exp [i(\kappa S_{(1)} + \kappa^2 S_{(2)})] \rangle} + \mathcal{O}(\kappa^3) \\ &= \langle T^{\mu\nu}(x) \rangle_{(0)} + \kappa \langle T^{\mu\nu}(x) \rangle_{(1)} \\ &\quad + \kappa^2 \langle T^{\mu\nu}(x) \rangle_{(2)} + \mathcal{O}(\kappa^3). \end{aligned}$$



(a)



(b)



(c)

Chiral trace anomaly

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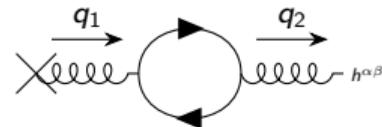
It is easy to write down the computations that should be done...

$$= \frac{i}{8} \mathcal{P}_- [(p_1 + p_2)_\mu \gamma_\nu + (p_1 + p_2)_\nu \gamma_\mu] \mathcal{P}_+,$$

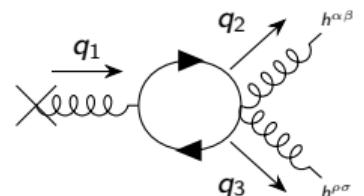
$$\mathcal{I}^{\mu_1 \dots \mu_m}(p) \equiv \int \frac{q^{\mu_1} \dots q^{\mu_m}}{(q^2 - i0)[(q+p)^2 - i0]} d^n q,$$

$$\text{tr} [\gamma^\mu \mathcal{P}_+ \gamma^\tau \mathcal{P}_- \gamma^\sigma \mathcal{P}_+ \gamma^\delta \mathcal{P}_- \gamma^\alpha \mathcal{P}_+ \gamma^\lambda \mathcal{P}_-],$$

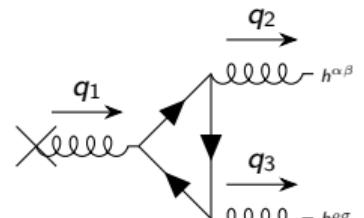
$$\mathcal{I}^{\mu_1 \dots \mu_m}(p, k) \equiv \int \frac{q^{\mu_1} \dots q^{\mu_m}}{(q^2 - i0)[(q-k)^2 - i0][(q-p)^2 - i0]} d^n q.$$



(a)



(b)



(c)

$$\begin{aligned}
I^{\mu\nu\rho\sigma\alpha\beta}(p, k) = & \frac{1}{192} \frac{i}{(4\pi)^2} \eta^{(\mu\nu} \eta^{\rho\sigma} \eta^{\alpha\beta)} \left[3(k^2)^2 - 6k^2(pk) + 4(pk)^2 + 5k^2 p^2 - 6(pk)p^2 + 3(p^2)^2 \right] \left(\mathcal{D} + \frac{3}{2} \right) \\
& - \frac{1}{16} \frac{i}{(4\pi)^2} \eta^{(\mu\nu} \eta^{\rho\sigma} \left[p^\alpha p^\beta \left[3p^2 + 2k^2 - 3(pk) \right] + k^\alpha k^\beta \left[2p^2 + 3k^2 - 3(pk) \right] + p^\alpha k^\beta [3p^2 + 3k^2 - 4(pk)] \right] (\mathcal{D} + 1) \\
& + \frac{1}{4} \frac{i}{(4\pi)^2} \eta^{(\mu\nu} \left(p^\alpha p^\beta p^\rho p^\sigma + p^\alpha p^\beta p^\rho k^\sigma + p^\alpha p^\beta k^\rho k^\sigma + p^\alpha k^\beta k^\rho k^\sigma + k^\alpha k^\beta k^\rho k^\sigma \right) \mathcal{D} \\
& - \frac{15}{16} \frac{i}{(4\pi)^2} \eta^{(\mu\nu} \eta^{\rho\sigma} \eta^{\alpha\beta)} \left[(p^2)^2 (G_{04}(p, k) + G_{02}(p, k) - 2G_{03}(p, k)) + (k^2)^2 (G_{40}(p, k) + G_{20}(p, k) - 2G_{30}(p, k)) + 4p^2(pk) (G_{13}(p, k) \right. \\
& \quad \left. - G_{12}(p, k)) + 4k^2(pk)(G_{31}(p, k) - G_{21}(p, k)) + 2p^2 k^2 (G_{11}(p, k) - G_{12}(p, k) + G_{22}(p, k) - G_{21}(p, k)) \right] \\
& + \frac{45}{2} \frac{i}{(4\pi)^2} \eta^{(\mu\nu} \eta^{\rho\sigma} p^\alpha k^\beta \left[p^2 (G_{12}(p, k) - G_{13}(p, k)) + k^2 (G_{21}(p, k) - G_{31}(p, k)) - 2(pk)G_{22}(p, k) \right] \\
& + \frac{45}{4} \frac{i}{(4\pi)^2} \eta^{(\mu\nu} \eta^{\rho\sigma} k^\alpha k^\beta \left[p^2 (G_{21}(p, k) - G_{22}(p, k)) + k^2 (G_{30}(p, k) - G_{40}(p, k)) - 2(pk)G_{31}(p, k) \right] \\
& + \frac{45}{4} \frac{i}{(4\pi)^2} \eta^{(\mu\nu} \eta^{\rho\sigma} p^\alpha p^\beta \left[p^2 (G_{03}(p, k) - G_{04}(p, k)) + k^2 (G_{12}(p, k) - G_{22}(p, k)) - 2(pk)G_{13}(p, k) \right] \\
& - \frac{15}{2} \frac{i}{(4\pi)^2} \eta^{(\mu\nu} \left[p^\rho p^\sigma p^\alpha p^\beta G_{04}(p, k) + 4p^\rho p^\sigma p^\alpha k^\beta G_{13}(p, k) + 6p^\rho p^\sigma k^\alpha k^\beta G_{22}(p, k) + 4p^\rho k^\sigma k^\alpha k^\beta G_{31}(p, k) + k^\rho k^\sigma k^\alpha k^\beta G_{40}(p, k) \right] \\
& + \frac{i}{(4\pi)^2} \left[p^\mu p^\nu p^\rho p^\sigma p^\alpha p^\beta F_{06}(p, k) + 6p^{(\mu} p^\nu p^\rho p^\sigma p^\alpha k^\beta) F_{15}(p, k) + 15p^{(\mu} p^\nu p^\rho p^\sigma k^\alpha k^\beta) F_{24}(p, k) + 20p^{(\mu} p^\nu p^\rho k^\sigma k^\alpha k^\beta) F_{33}(p, k) \right. \\
& \quad \left. + 15p^{(\mu} p^\nu k^\rho k^\sigma k^\alpha k^\beta) F_{42}(p, k) + 6p^{(\mu} k^\nu k^\rho k^\sigma k^\alpha k^\beta) F_{51}(p, k) + k^\mu k^\nu k^\rho k^\sigma k^\alpha k^\beta F_{60}(p, k) \right],
\end{aligned}$$

where F_{ij} and G_{ij} are functions of k and p which satisfy several nontrivial relations.

Chiral trace anomaly

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You'd better use tensorial simplifications:

$$\mathcal{I}^{\mu_1 \dots \mu_m}(p, k) = 2 \int_0^1 \int_0^{1-y} \int \frac{(q + yp + xk)^{\mu_1} \dots}{(q^2 + M_E - i0)^3} d^n q dx dy.$$

$$M_E \equiv y(1-y)p^2 + x(1-x)k^2 - 2xy(p \cdot k).$$

Because of symmetry, we can substitute

$$\begin{aligned} q^\mu q^\nu &\rightarrow \frac{1}{n} \eta^{\mu\nu} q^2, \\ q^\mu q^\nu q^\rho q^\sigma &\rightarrow \frac{3}{n(n+2)} \eta^{(\mu\nu} \eta^{\rho\sigma)} q^4, \\ &\vdots \end{aligned} \tag{10}$$

⇓ the traceless part of $\eta^{(\mu\nu} \dots$ vanishes

$$[\mathcal{I}^{\mu \dots}(p, k)]_{\text{trless}} = 2 \int_0^1 \int_0^{1-y} \int \frac{[(yp + xk)^\mu \dots]_{\text{trless}}}{(q^2 + M_E - i0)^3} d^n q dx dy$$

$$\begin{aligned}
\mathcal{I}^{\mu\nu\rho\sigma\alpha\beta}(p, k) &= \left[\mathcal{I}^{\mu\nu\rho\sigma\alpha\beta}(p, k) \right]_{\text{trless}} \\
&\quad + \frac{15\mathcal{I}(p - k)}{16(-1+n)(1+n)(8+n)} \eta^{(\rho\sigma} \\
&\quad \left[(2+n)(4+n)k^\alpha k^\beta k^\mu k^\nu + 4(-4+n^2)k^\alpha k^\beta k^\mu p^\nu + 6(-2+n)nk^\alpha k^\beta p^\mu p^\nu \right. \\
&\quad \left. + 4(-4+n^2)k^\alpha p^\beta p^\mu p^\nu + (2+n)(4+n)p^\alpha p^\beta p^\mu p^\nu \right] \\
&\quad - \frac{45\mathcal{I}(p - k)}{8(-1+n)(1+n)(6+n)(8+n)} \eta^{(\mu\nu} \eta^{\rho\sigma)} \\
&\quad \left[2k^\alpha p^\beta \left((-2+n(4+n))k^2 - 10n(p \cdot k) + (-2+n(4+n))p^2 \right) \right. \\
&\quad \left. + k^\alpha k^\beta \left((2+n)((6+n)k^2 - 10(p \cdot k)) + (8+n(4+n))p^2 \right) \right. \\
&\quad \left. + p^\alpha p^\beta \left((8+n(4+n))k^2 + (2+n)(-10(p \cdot k) + (6+n)p^2) \right) \right] \\
&\quad + \frac{15\mathcal{I}(p - k)}{16(-1+n)(1+n)(4+n)(6+n)(8+n)} \eta^{(\mu\nu} \eta^{\rho\sigma} \eta^{\alpha\beta)} \\
&\quad \left[3(4+n)(6+n)k^4 + 4(48+n(4+n))(p \cdot k)^2 - 60(4+n)(p \cdot k)p^2 \right. \\
&\quad \left. + 3(4+n)(6+n)p^4 + 2k^2 \left(-30(4+n)(p \cdot k) + (64+n(22+3n))p^2 \right) \right]
\end{aligned}$$

A couple of contributions of this type...

- compute spinorial factors ($\gamma^\mu \dots$);
- renormalize;
- check divergence (diffeo anomaly);
- compute the trace;
- identify geometrical invariants (which are written as expansions in $h^{\mu\nu}$).

Our results are exactly half of the trace anomaly for the Dirac spinor

$$(g_{\mu\nu} \langle T^{\mu\nu} \rangle)_{(2)}^{\text{ren}}(x) = \frac{1}{16 \cdot 45(4\pi)^2} (-11\mathcal{E}_4 + 18C^{\mu\nu\rho\sigma}C_{\mu\nu\rho\sigma} + 12\nabla^2 R)_{(2)},$$
$$(\nabla_\mu \langle T^{\mu\nu} \rangle)_{(2)}^{\text{ren}}(x) = 0,$$

in terms of the Weyl tensor $C_{\mu\nu\rho\sigma}$ and the four-dimensional Euler density E_4 , which in four dimensions satisfy

$$C^{\mu\nu\rho\sigma}C_{\mu\nu\rho\sigma} = R^{\mu\nu\rho\sigma}R_{\mu\nu\rho\sigma} - 2R^{\mu\nu}R_{\mu\nu} + \frac{1}{3}R^2,$$
$$E_4 = R^{\mu\nu\rho\sigma}R_{\mu\nu\rho\sigma} - 4R^{\mu\nu}R_{\mu\nu} + R^2.$$

What about the parity-odd contribution?

The parity-odd contribution

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Two-point contributions, $\langle T_*^2 \rangle$:

$$\epsilon^{\delta\mu\sigma\tau} \times \text{symmetric in external momenta} = 0 \quad (11)$$

The $\langle T_* J \rangle$ contribution:

$$(2\epsilon^{\alpha\beta\gamma(\tau}\hat{\eta}^{\lambda)\mu} - \epsilon^{\alpha\beta\gamma\mu}\hat{\eta}^{\lambda\tau}) \times \delta^\nu_{[\tau} p_{\lambda]} p^2 = 0 \quad (12)$$

The three-point contribution, $\langle T_*^3 \rangle$:

$$\epsilon^{[\delta\mu\sigma\tau}\hat{\eta}^{\alpha]\lambda} \times \text{something} = 0 \quad (13)$$

Dimensionally dependent identities [arXiv:gr-qc/0105066v1, Edgar & Höglund]. The Cayley-Hamilton theorem for an $n \times n$ matrix ("every square matrix over a commutative ring satisfies its own characteristic equation"):

$$M^{c_1}_{[c_1} M^{c_2}_{c_2} \cdots M^{c_n}_{c_n} \delta^b_{a]} = 0 \quad (14)$$

- The folklore says that the definition of the conformal anomaly for nonconformal theories is

$$\mathcal{A} = g^{\mu\nu} \langle T_{\mu\nu} \rangle - \langle g^{\mu\nu} T_{\mu\nu} \rangle. \quad (15)$$

Why should we employ it for a conformal theory?

- For a CFT in $n = 4$, the correlator of three stress-energy tensors is necessarily parity-even (Stanev, arXiv: 1206.5639)

$$\langle T_{\mu\nu}(x_1) T_{\rho\sigma}(x_2) T_{\alpha\beta}(x_3) \rangle = (\text{parity-even term}) + 0 \times \epsilon_{\bullet\bullet\bullet\bullet}, \quad (16)$$

but this is for the “regular” contribution. One may still have contributions at coincident points (that should be seen in our computation).

① Anomalies

② Chiral fermions and anomalies

③ The trace anomaly

④ Outlook

- We have obtained no parity-odd contribution to the chiral trace anomaly (BM scheme, Feynman diagrams).
- Cosmology (inflation, baryogenesis, ...)
- SMEFT and anomaly cancellations [arXiv:2011.09976, Cata, Kilian and Kreher], [arXiv:2205.10381, 2012.13989, Feruglio], [arXiv:2104.13569, Passarino], [2205.02248 Quevillon et al.].
- Supersymmetric anomalies [arXiv:2104.13391, Minasian, Papadimitriou, Yi], [arXiv:2103.10048, Nakagawa and Nakayama].