CYGNO Analysis Meeting

David Marques

& PMT analysis working group: F. Borra, M. Folcarelli, S. Piacentini, A. Messina, E. Baracchini, E. Kemp

8/06/2023 - Coimbra

PMT group works rundown:

1. <u>On-going:</u>

a. PMT reconstruction <u>TTree</u>

b. (L,x,y) fit

- i. Bayesian fit approach
- ii. Machine learning
- **c. Z diffusion** from Fe spots \Rightarrow See Matteo's discussion
- d. Association of PMT & camera events
 - i. One-to-one (camera & PMT) golden dataset.
 - ii. Closest neighbour approach under study

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- e. Study of the α parameter in L $\propto 1/R^{\alpha}$.
- **f.** Optimize moving average to **filter high frequency noise** ⇒ important to understand constant **sources of noise**.
- **g.** Optimize *find_peaks()* in order to identify better the peaks ⇒ Relevant especially for long tracks.
- **h.** Compute **track angle / \Delta Z \Rightarrow** First steps towards <u>**3D reconstruction**</u>.
 - i. Using **single peaks** approach
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- Future steps:
 - Fully or partial <u>integration of</u>
 x/y/L fit in PMT reconstruction.
 - **<u>Am-Be run</u>** will provide us with a

great new dataset!

1. <u>Goal:</u>

- a. Create a **TTree** with the **<u>PMT reconstructed variables</u>**.
- b. Run **PMT** reconstruction together with **Camera** reconstruction on the cloud.
- c. Create a **framework** that allows the **analysis of PMT events stand-alone and together with camera**.
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2. <u>Steps already taken:</u>

- a. **Retrieve data** from midas data banks through *cygno_libs* -> Both *fast* and *slow* digitizers.
 - i. All four PMTs and all triggers associated with a given picture.
 - ii. Waveforms **automatically corrected**. (maximum digitizer offset of **-0.3V**)
- b. PMT-Reco working for both Run1 and Run2 (which had different midas files)
- c. Tree created with basic variables.
- d. Config files updated.
- e. Bug-testing ongoing:
 - i. Missing tests with pedestals and simulation reconstruction.



→ Mostly completed!

• Further tests will be conducted.

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→ <u>Rundown of the variables:</u>

- pmt_wf_[run/event/trigger/channel/sampling]
- pmt_wf_insideGE
- pmt_[baseline/RMS]
- pmt_tot_[integral/charge]
- pmt_max_ampl
- pmt_nPeaks
- pmt_peak_Number
- pmt_peak_[Position/Height/HalfWidth/FullWidth]

- ⇒ **Basic** info of each waveform
- ⇒ Checks which triggers are inside the camera's *Global Exposure*
- ⇒ Baseline and RMS. Useful to check *quality of data*
- ⇒ Sum and conversion of *all samples* of a waveform
- ⇒ Waveform's max amplitude. Used to *signal selection*
- ⇒ Waveform's number of peaks. Useful to *easily select* ⁵⁵*Fe* vs. others
- ⇒ Useful for *peak coincidence* among PMTs
- \Rightarrow Main analysis point. Basic implementation at the moment.

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The *find_peaks()* is the **most important function** which still needs optimization:

There's space here to some work!

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This was done with 25 calibration runs (5 positions x 5 runs). For higher statistics, it's just a matter of running this on the cluster with many runs...

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Legend: **Pink** = TOT; **Red** and **Green**: Full and Half widths of peaks retrieved with *find_peaks()*

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- 2. Different <u>cuts</u> were studies: **RMS** and **mj2p**



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 \Rightarrow More on the RMS later

\Rightarrow <u>dE/dx on BG signals</u>



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- 2. Eventual **band at low dE/dx** mixed with a diagonal cut \Rightarrow Most long track don't "fit" in the fast digitizer.
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E/A × (keV/n

keV/c

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⇒ The **waveforms' RMSs** show some **weird features** that

could spoil further studies.

One must understand the <u>reason</u> behind this: is it an <u>environment</u> dependency (P,T,n), or does it result from <u>particular particle interaction</u>?

⇒ <u>RMS – isolating the problem:</u>

1. I start with the RMS of 25

calibration runs:



2. Divided by channel:

⇒ The waveforms' **RMS** has a **second main peak** at <u>~4 samples intensity</u>

⇒ I will check with **<u>other factors</u>** like temperature, humidity, light yield, etc.

⇒ We also discussed the possibility of higher RMS coming from hotpots or high energy particles that seem to

destabilize the PMTs.

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3. Divided by run:

PMT- Variables' analysis sum-up

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 - A better **integration** of <u>reco and fit</u> is necessary and is being discussed.
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 - Use **TOT as an approach for 3D** reconstruction of straight(ish) tracks.
- Understand RMS weird behaviour
 - Alphas?
 - Humidity/LY?
- Compare dE/dx with other variables (like number of peaks)
 - Discuss more in detail how to assess TOT once fitted L is calculated.
- Normalize plots to ⁵⁵Fe (eventually LED (?)) to get a better grasp of the units being discussed.
- Understand how to differentiate and integrate the **analysis of the slow digitizer** \Rightarrow Crucial for long waveforms.

→ <u>Cross-check with images:</u>

- At this point is important to <u>look directly at some pictures</u> to understand if there are <u>specific and clear</u>
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→ Eventual plans for the future (to keep in mind):

- PMT R&D:
 - LED calibration at different positions.
 - Multi-particle/energy calibration
- Analysis of PMT simulation:
 - Would be relevant at this point to also start discussing with the PMT simulation group.
 - They can create clear data-set with different particles / energies / directions. etc..

Thank you for

your attention!

The CYGNO Project counts with the collaboration of several international researchers coming from: