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Neutron sd-shell excitations for light nuclei with $N \geq 8$

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In light nuclei, the separation between the p and sd shells decreases with increasing N/Z near $N=8$, and the energy of the $1s_{1/2}$ orbital drops below that of the $0d_{5/2}$ level. This behavior is largely responsible for the halo-like nature of neutron-rich $N=9$ nuclei, which contain a loosely-bound $1s_{1/2}$ neutron. We have used HELIOS to study neutron transfer with the (d,p) reaction using beams of the unstable nuclei ^{12}B , ^{13}B , and ^{15}C , leading to ^{13}B , ^{14}B , and ^{16}C , corresponding to neutron numbers of $N=8$, 9, and 10. Of particular interest are the properties of states with one or two neutrons in the $1s_{0d}$ shell. The exotic beams were produced using the in-flight method. Protons were detected at backward laboratory angles using HELIOS, in coincidence with and forward-recoiling beam-like recoils identified in an array of silicon-detector telescopes. HELIOS is designed specifically to study such reactions in inverse kinematics, and to optimize the excitation-energy resolution in the center-of-mass system. Excitation-energy spectra and angular distributions for these reactions will be presented, as well as relative spectroscopic factors. The data will be compared to the predictions of shell-model calculations for this region. This work was supported by the U.S. Department of Energy, Office of Nuclear Physics under contracts DE-FG02-04ER41320 and DE-AC02-06CH11357, and NSF Grant Nos. PHY-02-16783 and PHY-10-68217.

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