## Research Doctoral programs PON for the academic year 2021/2022 - XXXVII cycle

Theme title	Development of solutions for the generation of architectures	Development of solutions for the generation of efficient FPGA-based computing architectures	
Project title	Integration of Neural Networks on FPGA		
		Giulio Bianchini	
Year	II		
Tutors	Mirko Mariotti, Daniele Spiga		
	5/10/2023 Bertinoro		



5/10/2023



#### **BondMachine**

A framework to build dynamical computer architectures

The BondMachine is an open source (<u>https://github.com/BondMachineHQ</u>) software ecosystem for the dynamical generation of computer architectures that can be synthesized on FPGAs.

- High level programming language (Golang) for both the hardware and software
- Functional style programming
- Architecture generating compiler
- Computational graph and Machine Learning Models

Article published on Parallel Computing,

Elsevier 2022 DOI:10.22323/1.351.0020





#### **Machine Learning Inference on FPGA**

Build firmware for low latency machine learning inference on FPGA

Implement customized and parallel architectures tailored to specific ML models, resulting in faster processing speeds and lower power consumption compared to traditional CPUs or GPUs.

Starting from high-level code and standard ML framework, with HLS tools like HLS4ML or BondMachine, get the firmware implementations of machine learning algorithms





#### Why BondMachine?

The machine learning model is trained with standard frameworks and synthesized in FPGA as a graph of heterogeneous and interconnected processors.

 Optimized resource usage
 Highly customizable
 Available at a high level
 Vendor independent
 User-friendly
 Cloud extension
 We have integrated this solution with Inference-as-a-Service cloud native systems

#### **ML** Inference on cloud

Integrate our system with cloud-native inference as a service system

#### Ease of usage and flexibility

Being able to deploy an inference algorithm on FPGA without caring for "where" the resources are

#### Democratic access and management

← Object Browser

Load model desr.

05-04 11:54:13 \* HLS tool requested bondmachine \*

05-04 12:06:59 \* Going to clean temporary files

05-04 11:54:14 \* Before exec command: make bondmachine S-04 11:54:15 \* Before exec command: make hdl ossibile accedere a '\*.vhd': File o directory non esistente

Request arrived to build firmware

\* Command executed successfully: make hdl

pre exec command: make design bitstream

:54:17 \* Before exec command: make design synthesis

Leveraging cloud/k8s native tools, you can reuse a well established way to orchestrate the bookkeeping and distribution of the payloads

#### **Easy Prototyping**

**OBJECT STORE** 

🚯 \_\_\_\_\_\_\_bject Browser

AGEN & LICENSE

Automation of the build and load process -> the framework take care of vendor specific details



#### Thanks for the attention!

★	Website	http://bondmachine.fisica.unipg.it/
★	Paper	https://www.sciencedirect.com/science/article/pii/S0167819121001150
★	CHEP 2023	https://indico.jlab.org/event/459/contributions/11826/
★	CCR 2023	https://agenda.infn.it/event/34683/contributions/197368/
★	CCR 2022	https://agenda.infn.it/event/30202/contributions/168531/
★	InnovateFPGA 2018 Iron Award	https://github.com/innovatefpga/2018-EM083

★ Main repo

https://github.com/BondMachineHQ

★ ML inference on cloud repo

https://github.com/BondMachineHQ/kserve-bond-extension.git



5/10/2023, Bertinoro



# **CMS ECAL Barrel Upgrade for HL-LHC Lightning Presentation**

XIV edition of ESC School on Efficient Scientific Computing (ESC23)

4-12 October 2023, Bertinoro (Italy)





CMS ECAL Upgrade - ESC23



## **CMS ECAL Barrel Upgrade**



\* Read by a pair of avalanche photo-diodes (APDs)

Cecilia Borca (University and INFN Torino)



HL-LHC upgrade challenges for the ECAL barrel

Increased pileup from 40~60 up to 200

**Increased noise** caused by detector and sensor **ageing** 

Faster front-end (FE) electronics to

**Improve time resolution** for better primary vertex reconstruction -> <30 ps for E>50 GeV

**Discern scintillation signals from APD spikes** (signals from direct ionization of the APDs)









Cecilia Borca (University and INFN Torino)

triggerless: the trigger primitives will be generated in the BCP

## My activities in the CMS ECAL Upgrade



Larger scale ==> efficiency is increasingly important

Cecilia Borca (University and INFN Torino)

- ole
- critical points
- the production

(3) Beam Tests Preparation and Data Analysis

- Last successful test beam campaign at \* CERN H4 facility in July 2023 with the <u>full system</u> (previous in 2022, 2021) Tested 200 channels of a supermodule \* read by the full readout chain (frontend to back-end)
- I developed a data readout application to integrate the controller of the BCP (written in Python) with the existing C++ DAQ software



... for the future => integration into the CMS certral DAQ => large scale commissioning tests

CMS ECAL Upgrade - ESC23

4





## A uniform and heterogeneous global event reconstruction at the CMS experiment

### Alessandro Brusamolino, Karlsruhe Institute of Technology



## Who am I?



- Master's degree in Physics at the University of Milano-Bicocca
  - Data reduction techniques for CMS Run 3
  - Started working on event reconstruction
- Currently PhD student at the Karlsruhe Institute of Technology
  - Working on the **CMS global event reconstruction** for Phase 2

## The CMS Experiment for HL-LHC



- General purpose detector
  - **Tracker**  $\rightarrow$  track charged particles
  - **MTD**  $\rightarrow$  time of arrival of charged particles
  - ECAL, HCAL, HGCAL  $\rightarrow$  energy measurement
  - Superconducting magnet → 3.8 T magnetic field
  - **Muon chambers**  $\rightarrow$  identify muons
- ~ 40 MHz collision rate
  - Level-1 Trigger  $\rightarrow$  750 kHz
  - **HLT** → 7.5 kHz
- Up to 200 simultaneous collisions (pileup)





## **Global event reconstruction**

- Combine information from the different parts of the sub-detectors
  - Identify final state particles
  - Reconstruct particle properties
- The current global reconstruction is Particle Flow
- High pileup scenario during Phase 2
  - Additional contributions from particles coming from pileup interactions
    - $\rightarrow$  performance are degraded
  - Novel algorithms developed for Phase 2
    - TICL (The Iterative CLustering) developed for HGCAL
      - Start from hits in the calorimeter and return particle properties and identification probabilities
      - Developed with <u>heterogeneous computing in mind</u>





Algorithms

Data structures

## **Extending TICL to the Barrel region**



- Goal is to have a uniform event reconstruction across the entire detector
- TICL is easily generalizable to be run on different detectors
  - $\rightarrow\,$  extension to ECAL and HCAL
- Reconstruction would benefit from such an approach
  - Exploiting other steps of the reconstruction offloaded to GPUs
  - Use same data structures and algorithms
  - Good performances in critical region of the detector





### Alessandra Casale

Tecnologist

INFN LNGS Computing and Networking Systems Services



Istituto Nazionale di Fisica Nucleare Laboratori Nazionali del Gran Sasso

Education and career history Master's degree in Nuclear Physics (University of Genova, 2011) Ph.D. in Physics Underwater Acoustics (University of Genova, 2017)

Patent submission, optical hydrophone (granted in 2020) Software developer (supply chain, machine learning) +8 years

> User Experience Design course (POLI.design Politecnico Milano, 2021)

Tecnologist at INFN Laboratorio Nazionale del Gran Sasso since July 2023

#### HPC4DR project

Started in 2020, involving the universities and research centers of the Abruzzo, Marche and Molise regions, hard hit by earthquakes and other catastrophic events in 2016/2017.



Reduction of risks connected to disasters due to natural and human-made phenomena Highperformance computing (HPC) technological infrastructure

My current activity

#### HPC infrastructure @LNGS

On January 27, 2022, the agreement was signed between INFN and CINECA for the free transfer of:

- n. 5 GALILEO racks each containing 72 nodes
- n. 1 GALILEO rack containing 36 calculation nodes
- n. 1 GALILEO rack containing spare parts

The computing infrastructure is Lenovo NextScale Each rack holds up to 6 Lenovo NeXtScale n1200 Enclosures (6RU) Each enclosure contains 12 NeXtScale nx360 M5 Compute Nodes

- 2\*Intel Xeon E5-2697 v4 @ 2.30GHz 18-core each (Broadwell)
- 128GB RAM/node, 3.5GB RAM/core

5 rack \* 72 nodes, 1 rack \* 36 nodes -> 396 nodes, 14256 cores

- Each server has a peak computing power of 1.3Tflop/s.
- The total computing power is approximately 0.5PFlops
- The compute nodes are interconnected by a 100Gb/s Intel Omnipath network and a 1Gb/s Ethernet network
- n. 1 node with **8 GPU** from GSSI (Gran Sasso Science Institute)

It's time to deepen my knowledge in parallel computing for testing and user support <sup>(3)</sup>



My current activity

#### **INFN CLOUD project**

The INFN Cloud is an internal project which aims to

- manage a (large) fraction of the INFN resources in a sustainable and optimized way;
- make different INFN communities able to access resources, regardless of the availability of local and dedicated hardware



My current activity



The federation of the LNGS infrastructure will allow:

- To expand the INFN Cloud backbone (BARI, CNAF) with a scalable set of satellite sites,
- To share with the INFN community a portion of LNGS-HPC resources







## Study of matter-antimatter asymmetry and strangeness equilibration at the LHC with ALICE

#### Mario CIACCO PhD Student at Politecnico and INFN, Torino

ESC 2023 - Bertinoro (FC)

#### Matter-antimatter asymmetry at the LHC

#### • LHC $\rightarrow$ antimatter factory

- Matter created in Pb–Pb collisions at TeV scale → (almost) exact equilibration of matter and antimatter
- Similar to the conditions of the early Universe according to standard cosmology
- How close to perfect symmetry?
  - Measure the antimatter-matter balance
  - $\circ \quad \textbf{ALICE} \rightarrow outstanding \\ particle identification \\ performance \\ \end{cases}$



#### Matter-antimatter asymmetry at the LHC

- Antimatter/matter yield ratios measured for different identified species
  - $\circ$  Protons and light nuclei  $\rightarrow$  sensitive to the baryonic asymmetry of the system
  - Baryon chemical potential  $\mu_{\rm B} \rightarrow$  global variable providing the net-baryon content



fits based on statistical hadronization model

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#### Strangeness production at the LHC

- Enhanced strangeness production in nuclear collision  $\rightarrow$  bulk of thermalised partons
- Similar effects observed also in hadronic collisions
  - Does strangeness equilibrate also in small collision systems? 0



Lund string fragmentation (PYTHIA)

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#### Strangeness conservation in (non-)thermal models

Canonical statistical model

- Charge conservation over a volume
- Symmetric correlation of same- and opposite-charge hadron pairs
- Large rapidity correlation for strangeness conservation

$$\mathcal{Z}(B, Q, S) = \int_{-\pi}^{\pi} \frac{d\phi_B}{2\pi} \int_{-\pi}^{\pi} \frac{d\phi_Q}{2\pi} \int_{-\pi}^{\pi} \frac{d\phi_S}{2\pi} e^{-i(B\phi_B + Q\phi_Q + S\phi_S)}$$

$$\times \exp\left[\sum_j z_j^1 e^{i(B_j\phi_B + Q_j\phi_Q + S_j\phi_S)}\right].$$
Conserved exactly over correlation volume  $V_c$ 

V. Vovchenko et al. Phys. Lett. B 785, 171 (2018)

Lund string fragmentation (PYTHIA)

- Local conservation of charges
- Correlated production of **opposite** charges
- Small rapidity correlation for strangeness conservation



→ Can be explored via event-by-event measurements! ESC 2023 - Bertinoro (FC)

#### Net- $\Xi$ net-kaon number correlation

- Same- and opposite-sign correlations
  - $\rightarrow$  2 different species  $\rightarrow$  charged **kaons** and  $\Xi$ 
    - Negligible effect of heavy Ο resonance decays
- Correlation of net-particle number
  - Contains effects of both charge 0 pairings
  - "How much does  $(\Xi^+ \Xi^-)$  number 0 difference influence  $(K^+ - K^-)$ number difference due to strangeness conservation?"
- The results obtained can be described by thermal models
- **String fragmentation**, despite matching the measured yields, fails to describe higher-order moments



ALI-PREL-556590

ESC 2023 - Bertinoro (FC)







dall'Unione europea NextGenerationEU







ESC23 Scuola Internazionale INFN

### Architetture, strumenti e metodologie per lo sviluppo di applicazioni di calcolo scientifico su larga scala

Solar Flare Forecasting based on Deep Learning Models

Elizabeth Doria Rosales

Supervisor: Prof. Vincenzo Carbone Co-Supervisor: Prof. Fabio Lepreti







Finanziato dall'Unione europea NextGenerationEU



### Impact of Space Weather on Earth and Human Technologies

Coronal Mass Ejections (CMEs) Large portrions of the corona, or outer atmosphere of the Sun, can be explosively blown into space, sending billions of tons of plasma, or superheated gas Earth's direction. These CMEs have their own magnetic field, resulting in geomagnetic storms. The fastest of these CMEs can reach Earth in under a day, with the slowest taking a for 5 days to reach Earth.

#### Solar Wind

The solar wind is a constant outflow of electrons and protons from the Sun, always present and buffeting Earth's magnetic field. The background solar wind flows at approximately one million miles per hour!

Fart

Earth's Magnetic Field

Earth's magnetic field, largely like that of a bar magnet, gives the

etic field can become extreme. In addition to some buffer

Earth some protection from the effects of the Sun. Earth's

During geomagnetic storms, the disturbances to Earth's

the charged particles of a radiation storm

magnetic field is constantly compressed on the day side and stretched on the night side by the ever-present solar wind.

ing by the atmosphere, this field also offers some shielding from



Space weather refers to the variable conditions on the Sun and in the space environment that can influence the performance and reliability of space-based and ground-based technological systems, as well as endanger life or health. Just like weather on Earth, space weather has its seasons, with solar activity rising and falling over an approximate 11 year cycle.

#### Sun's Magnetic Field

Strong and ever-changing magnetic fields drive the life of the Sun and underlie sunspots. These strong magnetic fields are the energy source for space weather and their twisting, shearing, and reconnection lead to solar flares.

#### Solar Radiation Storms

Charged particles indufing electrons and protons, can be accelerated by coronal mass ejections and solar flares. These particles bounce and gyrate their way through space cupilly following the magnetic field lines and utilinately bombarding Earth from every direction. The fastest of these particles can affect Earth tens of mixtus after a solar flare.



#### Geomagnetic Storms

A geomagnetic storm is a temporary disturbance of Earth's magnetic field typical associated with enhancements in the solar wind. Thuses storms are created when the solar wind and its magnetic field interacts with Earth's magnetic field. The primary source of geomagnetic storms is OAE's which stretch the magnetosphere on the nightsde causing it to relaxe energy through magnetic reconnection. Disturbances in the ionosphere (a region of Earth's upper atomosphere) are usually associated with geomagnetic storms.





Technological and infrastructure affected by space weather events. *Credits: NASA* 

NOAA space weather infographic.

Solar Flares

iust 8 minutes

Reconnection of the magnetic fields on the surface of the Sun drive the biggest explosions in our solar

tem. These solar flares release immense amount

nma rays to radio

ergy and result in electromagnetic emis

ves. Traveling at the speed of light, these

emissions make the 93 million mile trip to Earth in

Sunspots

space weather

Sunspots are comparatively cool areas at up to 7,700° F and show the location of

through what we would see as the Sun's

surface. Large, complex sunspot group are generally the source of significant

strong magnetic fields protruding











## Heliophysics Missions and Programs



SDO launched on February 11, 2010, 10:23 am EST on an Atlas V from SLC 41 from Cape Canaveral.

After launch, the spacecraft was placed into an orbit around the Earth with an initial perigee of about 2,500 km (1,600 mi). SDO then underwent a series of orbit-raising maneuvers which adjusted its orbit until the spacecraft reached its planned circular. geosynchronous orbit at an altitude of 35,789 km (22,238 mi), at 102° West longitude, inclined at 28.5°.

Credit: NASA's Goddard Space Flight Center







dall'Unione europea NextGenerationEU



### How SDO Sees the Sun



Each of the wavelengths observed by NASA's Solar Dynamics Observatory (SDO) was chosen to emphasize a specific aspect of the Sun's surface or atmosphere.

This image shows imagery both from the Advanced Imaging Assembly (AIA), which helps scientists observe how solar material moves around the Sun's atmosphere, and the Helioseismic and Magnetic Imager (HMI), which focuses on the movement and magnetic properties of the Sun's surface.

#### Credits: NASA/SDO/GSFC









### Spaceweather HMI Active Region Patch (SHARP)



Joint Science Operations Center (JSOC) (http://jsoc.stanford.edu/)

#### HARP 401 (NOAA AR 11166) on 9 March 2011 at 23:24:00 TAI



Geostationary Operational Environmental Satellite (GOES) (https://www.swpc.noaa.gov/)

Universal Time







dall'Unione europea



### First Part: Feature extraction







### Second Part: Time series classification



(dimension reduction)

## Lightning Students Talk

4-12 October, Bertinoro

Sabrina Giorgetti

# **XIV INFN International School on Efficient Scientific Computing (ESC23)**

14 0 8 7 6 5



## **Getting to Know Me Academic Path**

- Master's degree in Physics of Data at the University of Padua (2020-2022)
- Technical student at CERN (2022)
- Member of the CMS collaboration (since March 2022)
- PhD student in Physics at the University of Padua (since February 2023) sabrina.giorgetti@studenti.unipd.it
- INFN Padova associate (since February 2023)

## Why I'm interested in the school?

- Learn tools and techniques to improve my programming skills
  - Memory usage and parallel programming
- Boost my current work, and potentially stimulate new projects
  - GP-GPU programming

## The CMS experiment

## Large Hadron Collider (LHC)

- LHC delivers proton-proton collisions at ECM = 13.6 TeV since the start of Run 3 (July 2022)
- LHC bunches collide every 25 ns with a resulting event's rate of 40MHz

## CMS

• General-purpose detector

Jura

 Measure the energy and momentum of collisions products

CMS





CMS Experiment at the LHC, CERN Data recorded: 2023-Apr-21 17:14:17.991232 GMT Run / Event / LS: 366403 / 108625285 / 113

# FIRST COLLISIONS OF 2023 IN CMS


# The CMS data flow



# The Level-1 Data Scouting system The Run 3 demonstrator



- Novel data acquisition system that stores data before the Level-1 (L1) accept
- The Level-1 Data Scouting (L1DS) system is designed for the CMS Phase-2 at HL-LHC
- A demonstrator, parallel and independent from the standard trigger chain, is currently being developed in Run 3
- Gathers L1 primitives at the full bunch crossing frequency from different sources:
  - Global Muon Trigger (GMT) : up to 8 muons
  - Barrel Muon Track Finder (BMTF): stubs primitives
  - Calorimeter Layer-2: up to 12 each of  $e/\gamma$ , jets,  $\tau$ missing transverse energy and energy sums
  - Global Trigger (GT): trigger decision algorithms bits

x7







# The Level-1 Data Scouting system **Analysis and ML applications** BX

## L1DS analysis

- Characterisation of data collected since the start of Run 3
  - Monitor the rates, occupancy over bunch crossings
  - Histograms of objects' parameters, study di-muon events \_
- Trigger-less data at 40MHz
  - 1 min of L1DS data taking corresponds to ~2.5GB of data formatted for analysis
  - Need to use distributed framework for processing (e.g. -Spark)



The distribution of η, φ, p<sub>T</sub> differences between the neural network (Micron DLA) prediction (or GMT) values, and the offline reconstructed muon tracks, for matched muons. A significant improvement in track parameter resolution is observed in the MDLA result (blue) when compared to the GMT output (red) [CMS-DP-2022-066].



Global Muon Trigger (GMT) muons occupancy per bunch crossing (BX) within an LHC orbit for muons reconstructed by the Barrel Muon Track Finder (BMTF) [CMS-DP-2023-025]

## **ML** applications

- Recalibration of L1 trigger primitives to improve resolution
- Implementation of algorithms in hardware (e.g hl4ml)







## Presentation of Research Activity







Jekatěrina Jaroslavceva<sup>1,2</sup>

## Supervisor: Felice Pantaleo<sup>1</sup>

<sup>1</sup> CERN, European Organization for Nuclear Research, Meyrin, Switzerland <sup>2</sup> Czech Technical University in Prague, Prague, Czech Republic

## Who am I?



- First-year PhD student in Computer Science, since September 2023, at the Czech Technical University in Prague.
- Master thesis title: "A New Trackster Linking Algorithm Based on Graph Neural Networks for the CMS Experiment at the Large Hadron Collider at CERN".
- Currently, a DOCT student @ CERN working on improving reconstruction in HGCAL (new CMS sub-detector) using Graph Neural Networks.

## **RESEARCH ACTIVITIES:**

>> Enhancing the CMS Event Reconstruction using Artificial Intelligence: efficient techniques for linking granular information, such as calorimetric tracksters, tracks, and MTD timing, in the Phase 2 CMS detector.

>> Designing a **heterogeneous** inference approach within the CMSSW framework, enabling seamless resource management and execution on GPUs when available, while gracefully falling back to CPU execution.

## **The CMS High Granularity Calorimeter**

- HGCAL operation starts in 2029.
- Measures the energy of traversing particles like large 3D camera.
- Particles interact with the detector material and produce particle showers.
- Interactions captured in sensors as hits.



## **Reconstruction in HGCAL**





**Hits** (*x*, *y*, *z*, *t*, *E*): Energy deposits in the detector sensors Layer-cluster (LC): cluster of hits from the same particle on a detector layer (2D) **Trackster**: collection of LCs from the same particle shower (3D)

## **Trackster Linking**

- Ideally, individual particle showers = individual tracksters.
- A clustering algorithm produces trackster fragments due to the detector's irregularity, distant and non-aligned energy deposits, tuning for high purity.
- An additional trackster linking step is needed.



Reduced example of a small window event graph of 2 particles.

**Project goal**: recover complete particle showers.

**How:** Develop a **Graph Neural Network** (GNN) approach for learning structures on a graph of nearby tracksters.



Jekatěrina Jaroslavceva



### \* ONNX-Exportable for inference in production environment



ESC23

Jekatěrina Jaroslavceva

## **Performance Evaluation**





## CERN-THESIS-2023-110 - J. Jaroslavceva

## **Double Pion dataset evaluation\***



- CLUE3D Trackster can be used as nodes in a Graph, which GNN takes as input
- Local information aggregation.
- graph the TICLGraph

ESC23

- Nodes Feature: Trackster properties
- Edge Features: Geometrical, Energy and Time compatibility
- Fast CPU inference (ONNX)



#### Trigger performance monitor at the LHCb detector

#### Bogdan Kutsenko, Dorothea vom Bruch, Anton Poluektov

Aix Marseille Univ, CNRS/IN2P3, CPPM, Marseille, France

ESC23 - October 6, 2023







1/6

#### LHCb detector

Single-arm forward spectrometer:

- Designed for the heavy flavor physics with  $2<\eta<5$
- Coverage is complementary to ATLAS and CMS
- Excellent performance of tracking, powerful particle identification and the fact that B meson produced in the forward region allow to perform high precision measurement of semileptonic B decays



HLT1 RTA

#### LHCb Dataflow Run 3



Bogdan Kutsenko (AMU/CNRS/CPPM)

Trigger performance monitor

Allen - high-level trigger (HLT) application on GPUs. The HLT in the LHCb experiment at CERN's Large Hadron Collider (LHC) plays a crucial role in efficiently filtering data from the collisions that occur within the detector. It runs with the upgraded LHCb detector in the Run 3. Two new algorithms for the Allen project were developed on GPUs:

- Splitting of tracks into two different categories based on the defined selection criteria (right/left side selection, random selection, etc. )
- Two different sets of tracks are used to reconstruct two sets of primary vertexes (PVs)

Motivation:

- Real-time reconstruction needs detailed monitoring of the physics performance of the algorithms
- Opportunity to split tracks, in general, can be useful for resolution estimation on data



#### Allen project

• The performance of the PV reconstruction algorithm heavily relies on good alignment. Therefore an additional PV performance monitor is required



• Resolution monitor is currently work in progress

#### Status and outlook

The performance of the PV reconstruction algorithm was studied:

- Motivation for an additional PV performance monitor become apparent
- To create monitor new algorithms for the Allen project were developed on GPUs and tested

Plans:

 Using new Allen algorithms make an additional performance monitor of the PV resolution, which will be shown in the control room during the data-taking period

Interests in the field of scientific computing:

- Programming with Threading Building Blocks
- GP-GPU Programming with CUDA

Thank you!