Supported by ERC through Starting Grant no. 759253 and Synergy Grant no. 101071865



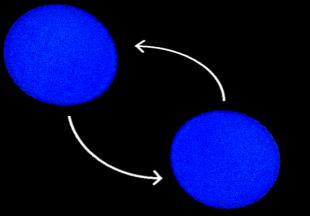
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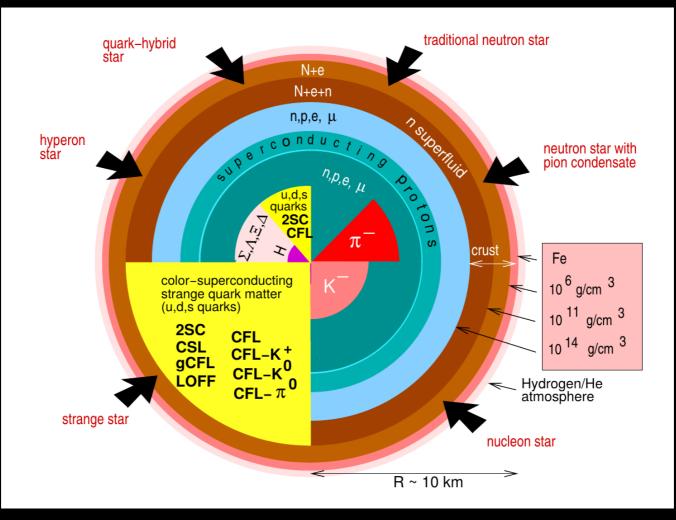






Pions, hyperons and quark matter in neutron star mergers GraSP23, PISA, 16/10/2023 Andreas Bauswein (GSI Darmstadt, HFHF)

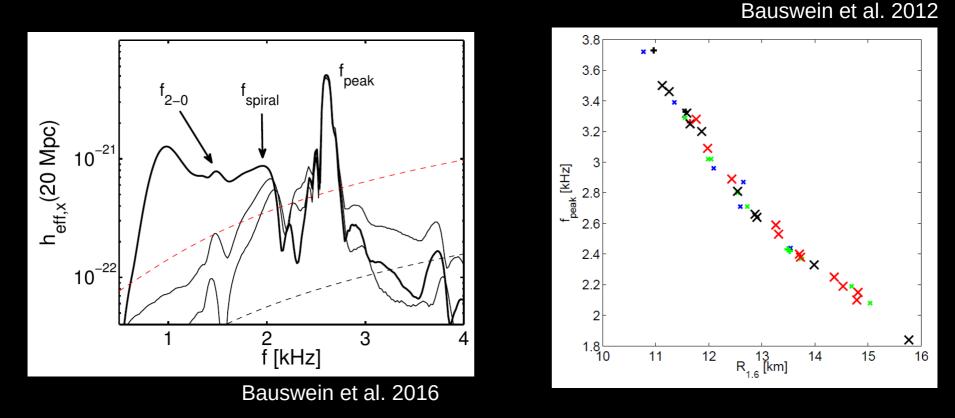




Weber 2004

Empirical relations

- ► To determine NS properties and EoS
- For postmerger GW emission, ejecta properties, threshold mass for black-hole formation etc.



Do non-nucleonic degrees of freedom lead to deviations from empirical relations?

Impact of pions in NS mergers

Vimal Vijayan, Ninoy Rahman, AB, Gabriel Martinez-Pinedo, Ignacio Arbina arXiv: 2302.12055

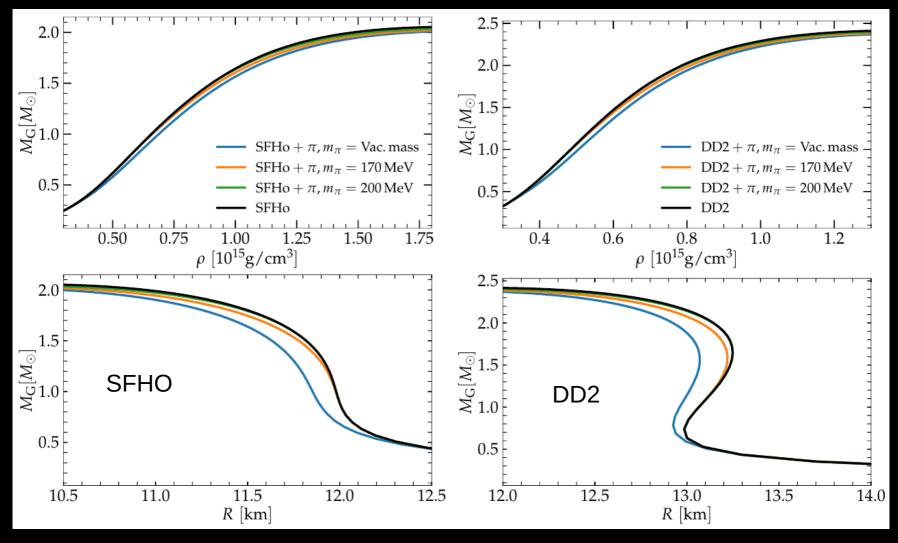
Including pions

- ▶ π^{\pm}, π^{0} mesons with rest mass of about 140 MeV in vacuum
- Impact on NS matter already discussed for decades but neglected in basically all EoS tables used in merger simulations
- ▶ Simple model to include pions as non-interacting Bose gas with chosen effective mass
 → pion condensation (ground state) and thermal pions
- Two base EoSs (DD2 and SFHO) and chosen constant effective mass

Pion condensation discussed since decades, e.g. Sawyer+ 1972, Migdal 1973, Baym & Flowers 1974, ...; more recently thermal pions Oertel+ 2012, Fore & Reddy 2020

Impact on stellar structure

- Stronger impact for smaller pion mass (earlier condensation)
- ► Radius decrease by 200 m; Mmax slightly reduced

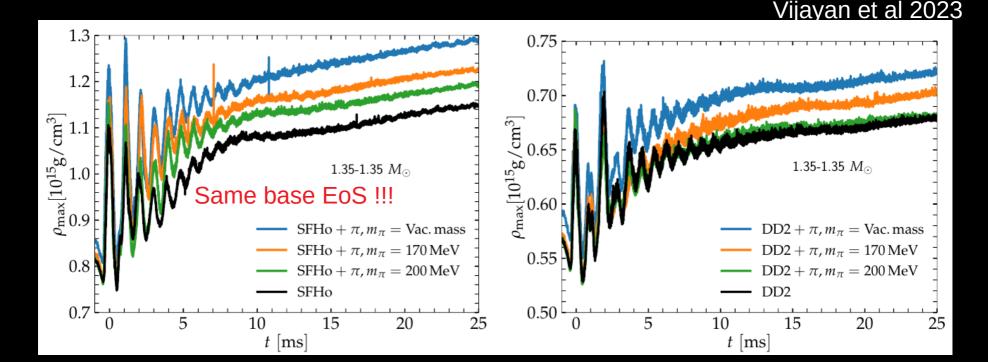


Merger simulations

Motivation:

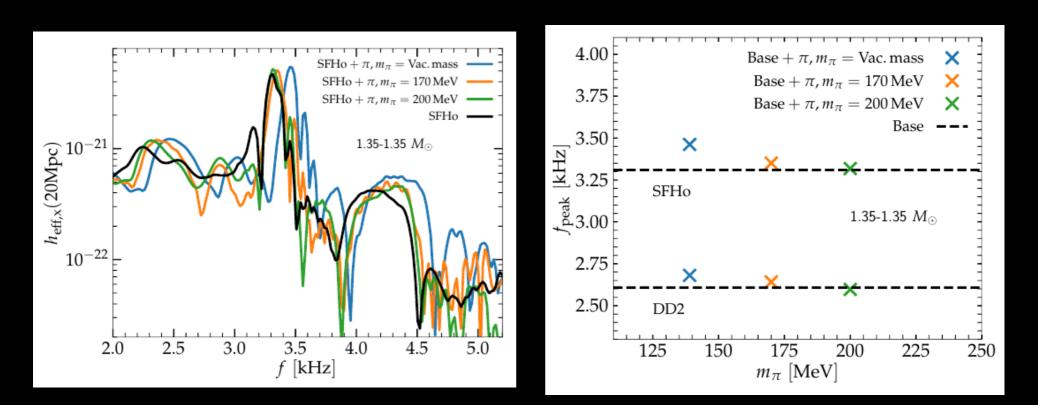
- Possibly combined effects condensate and thermal pions
- Empirical relations of merger observables often expressed by TOV properties

 → to which extend to those relations hold when pions are included ?
- SPH merger simulations in CFC with modified EoS tables compared to originals
- Electron fraction advected (okaish for high-density part where pions occur)



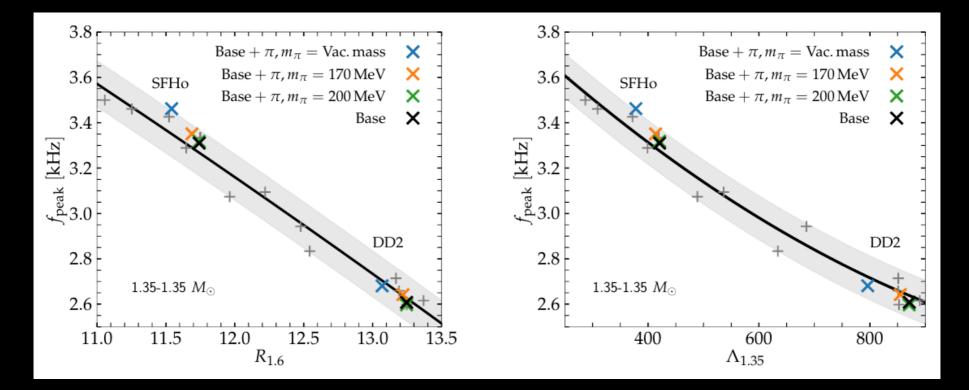
Postmerger GWs

- ► Inspiral: Tidal deformability reduced by 10 per cent (for mpi close to vac mass)
 → potentially problematic if nuclear parameters are deduced and pis neglected
- Postmerger: frequency shifts up to 150 Hz



Empirical relations for dominant postmerger frequency

- Empirical relations build without pions remain approximately valid
- For softer SFHO model possibly stronger shifts for m_{π} = vac. mass



Grey data points EoS models without pions (black curve fit to those)

Threshold mass for prompt collapse

1.4 \times Base + π , \times DD2 0.6 DD2	$m_{\pi} = \text{Vac.}$ $m_{\pi} = 1701$ $m_{\pi} = 2001$ 2.75 f_{peak} [k	MeV MeV 1.35-1.3	5 M⊙ 3.25	3.50			3.2 [] W 3.0 W 3.0 W 3.0 2.8 2.6 13.0	12.5 _{12.011.5} 11.0 ₁		2.2 (Mo) Mmax (Mo)
Model	$M_{\rm thres}$	$M_{\rm max}$	$R_{1.6}$	$R_{\rm max}$	$\Lambda_{1.4}$	$\tilde{\Lambda}_{\rm thres}$	$M_{\rm thres}^{\rm fit}$	$M_{\rm thres}^{\rm fit}$	$M_{\rm thres}^{\rm fit}$	$M_{\mathrm{thres}}^{\mathrm{fit}}$
(Max. dev./ M_{\odot})			6 1	6 3			$(Y = R_{1.6})$ (0.042)	$(Y = R_{\max})$ (0.059)	$(Y = \Lambda_{1.4})$ (0.056)	$(Y = \tilde{\Lambda}_{\text{thres}})$ (0.085)
	$[M_{\odot}]$	$[M_{\odot}]$	[km]	[km]			$[M_{\odot}]$	$[M_{\odot}]$	$[M_{\odot}]$	$[M_{\odot}]$
$SFHo + \pi$, Vac. mass	2.810	2.017	11.542		296.937		2.806(0.004)	2.804(0.006)	2.784(0.026)	2.796(0.014)
$SFHo + \pi, 170 MeV$	2.845	2.026	11.688	10.212			2.835(0.010)	2.832(0.013)	2.811(0.034)	2.816(0.029)
$SFHo + \pi, 200 MeV$	2.855	2.038	11.741	10.277	332.950	291.953	2.851(0.004)	2.850(0.005)	2.825(0.030)	2.832(0.023)
SFHo Base	2.870	2.056	11.743	10.285	332.970	282.036	2.861(0.009)	2.859(0.011)	2.835(0.035)	2.830(0.040)
$DD2 + \pi$, Vac. mass	3.250	2.381	13.069	11.692	639.278	256.841	3.257(-0.007)	3.271(-0.021)	3.271(-0.021)	3.228(0.022)
$DD2 + \pi, 170 MeV$	3.290	2.390	13.220	11.791	699.649	261.744	3.287(0.003)	3.294(-0.004)	3.325(-0.035)	3.256(0.034)
$DD0 + - 000 M_{\odot} V$	0.010	0.400	19.946	11.865	700 166	256.079	3.298(0.012)	3.314(-0.004)	3.333(-0.023)	2 250(0 051)
$DD2 + \pi, 200 MeV$	3.310	2.403	13.240	11.000	100.100	200.010	5.230(0.012)	0.014(-0.004)	3.333(-0.023)	3.259(0.051)

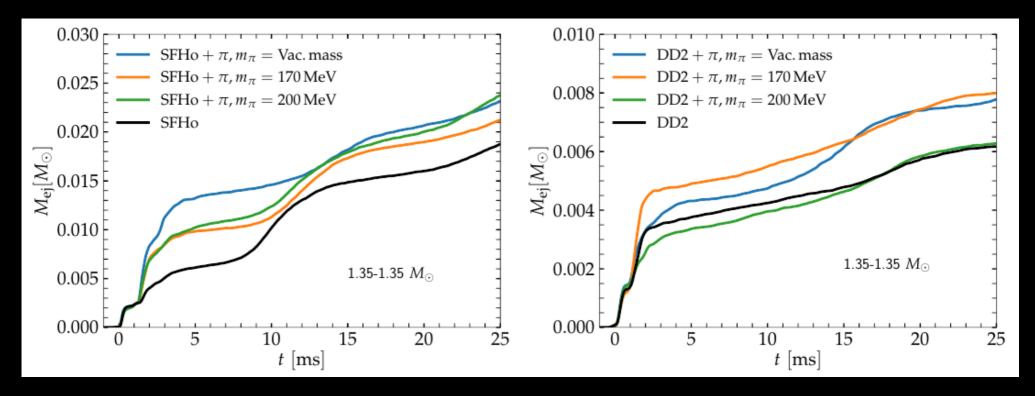
- Mthres reduced by up to 0.08 M_{sun} (for m_{π} ~vac mass)
- But empirical relations remain valid (within scatter) combined effect on M_{thres} and TOV properties

Mass ejection

- Inclusion of pions leads to (tentatively) more ejecta
- Increase stronger than expected from TOV properties (employing common fit formulae, see e.g. Henkel et al 2022)

 \rightarrow potentially problematic of EoS inference

Torus mass similar effects

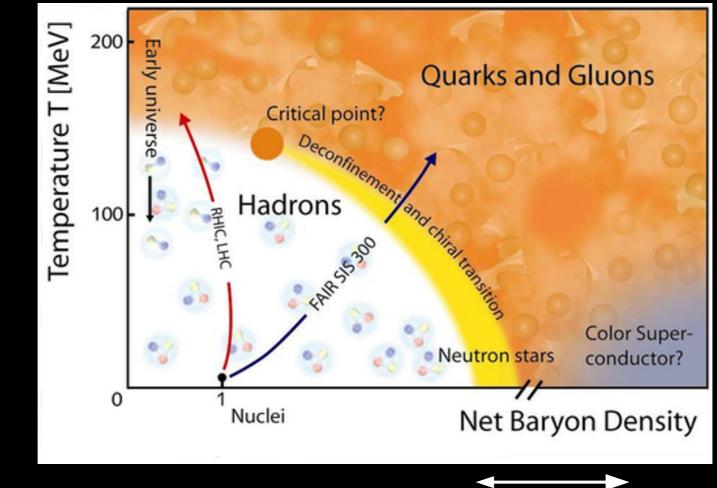


Quark matter in NS mergers

Bauswein et al 2019, Bauswein & Blacker 2020, Bauswein et al 2020, Blacker et al. 2020, Blacker et al. 2023

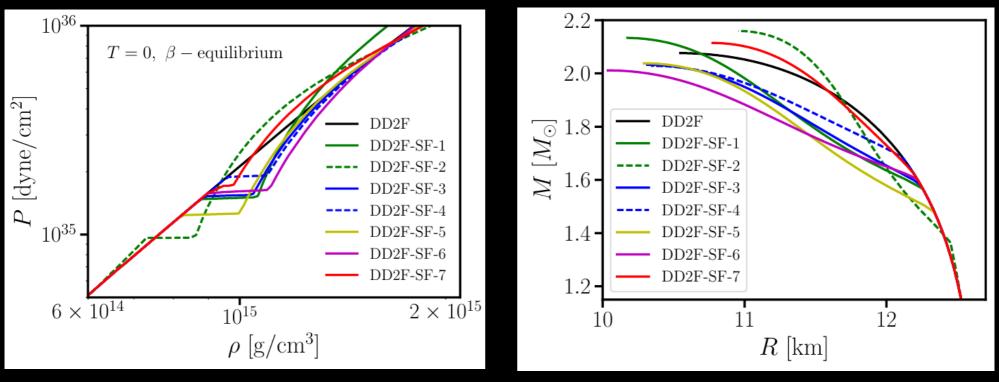
Phase diagram of matter of strongly interacting matter

GSI/FAIR

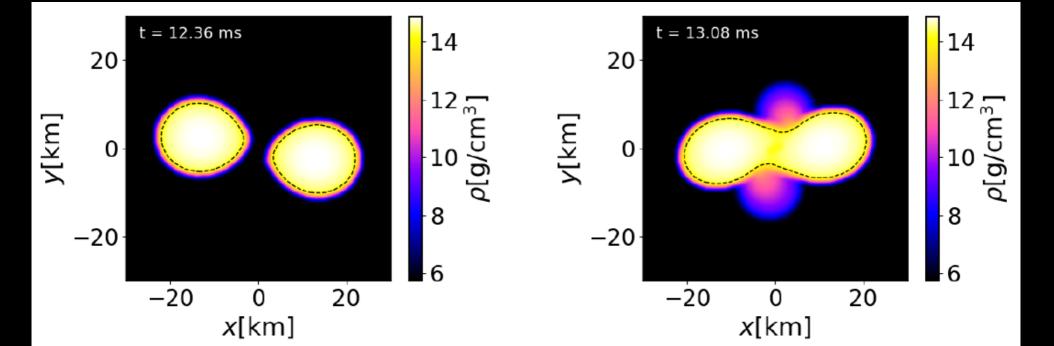


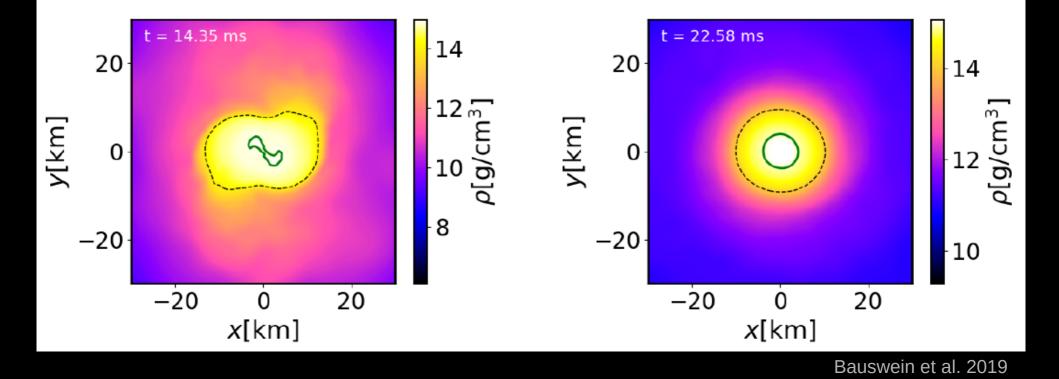
Does the phase transition to quark-gluon plasma occur (already) in neutron stars or only at higher densities? (low T, high rho not accessible by experiments or ab-initio models)

High T, low μ: experiments and lattice QCD 7 different models for quark matter: different onset density, different density jump, different stiffness of quark matter phase

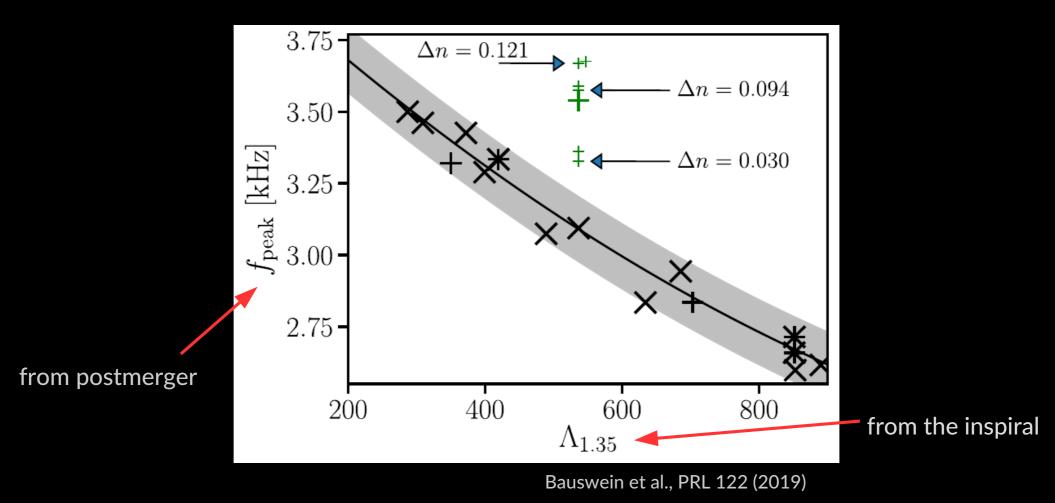


Bauswein et al. 2019, Bastian 2020





Signature of phase transition

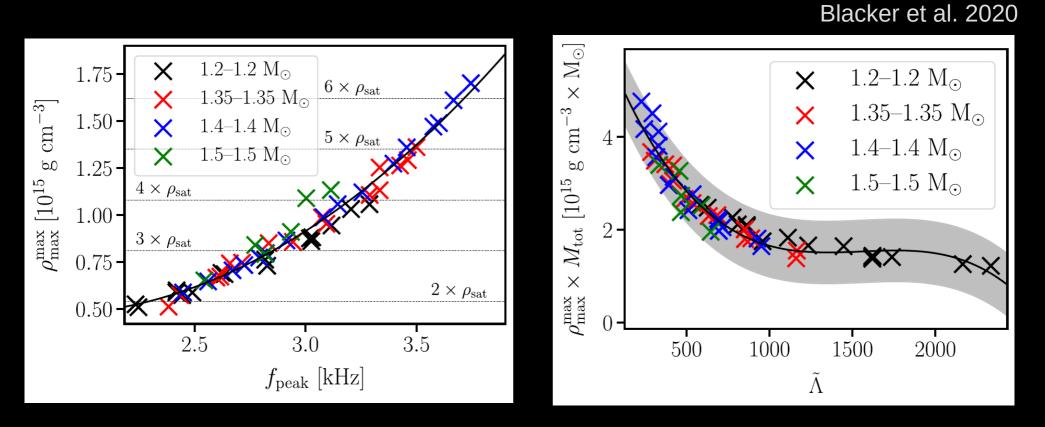


- Tidal deformability from inspiral
- Postmerger frequency measurable to within a few 10 Hz @ a few 10 Mpc (eg by Adv. Ligo upgrade, e.g. Chatziioannou+2017, Torres-Riva+2019, Wijngaarden+ 2022)

See also Most+ 2019, Blacker+ 2020, Weih +2020, Bauswein+2020, Prakash+ 2021, Liebling+ 2021, Hanauske+ 2021, Fujimoto+2022, Tootle+ 2022, Huang+ 2022, ...

► GWs inform about highest density in the remnant

 \rightarrow constraint on onset density (if PT is identified/excluded)

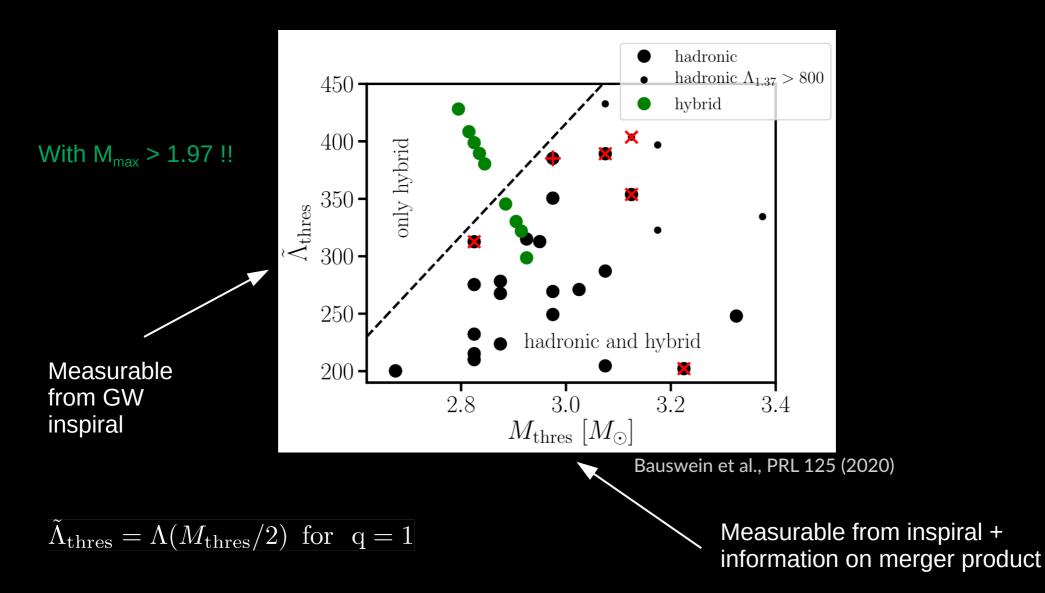


Postmerger frequency fpeak

tidal deformability from inspiral

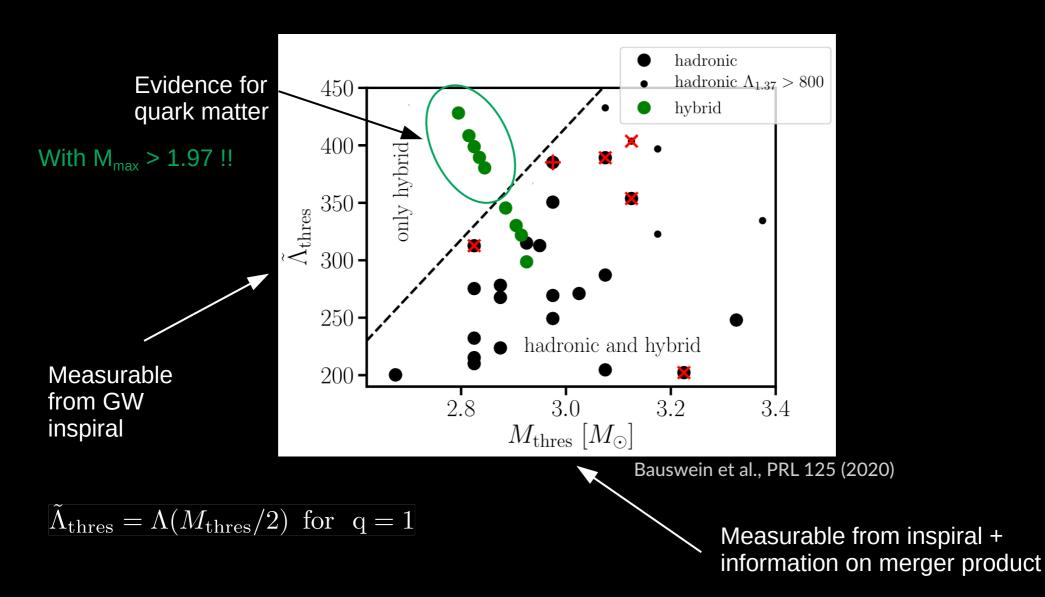
QCD phase transition from collapse behavior

- Directly measurable from events around M_{thres}
- Already single events yielding constraints may indicate presence of quark matter



QCD phase transition from collapse behavior

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- Already single events yielding constraints may indicate presence of quark matter



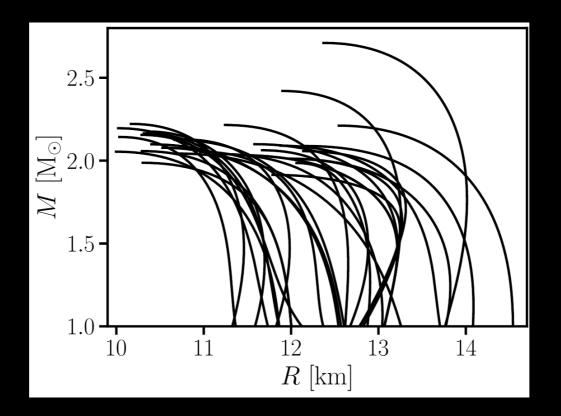
Hyperons in NS mergers

Blacker, Kochankovski, Bauswein, Ramos, Tolos, arXiv:2307.03710

See Sekiguchi et al. 2011, Raice et al 2017 for early studies of individual EoS models

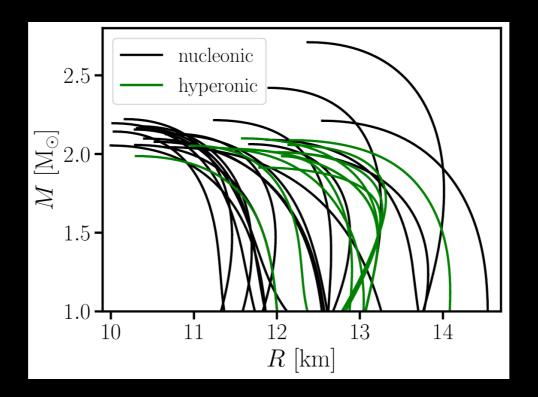
Hyperon puzzle

- Natural to expect that nucleons are converted to hyperons once chemical potential reaches hyperon mass
- ► Hyperon puzzle: Hyperons would soften the EoS which is in tension (?) with 2 Msun NSs
- ► Several modern hyperonic EoS fulfill the 2 Msun constraint
 - \rightarrow hyperon puzzle unsolved interacting Fermi gas with unknown interactions
 - \rightarrow generally hyperons leave weak impact on NS structure indistinuishable MR



Hyperon puzzle

- Natural to expect that nucleons are converted to hyperons once chemical potential reaches hyperon rest mass
- ► Hyperon puzzle: Hyperons would soften the EoS which is in tension (?) with 2 Msun NSs
- Several modern hyperonic EoS fulfill the 2 Msun constraint
 - \rightarrow hyperon puzzle unsolved interacting Fermi gas with unknown interactions
 - \rightarrow generally hyperons leave weak impact on NS structure indistinuishable MR



► A nucleonic EoS could mimic T=0 behavior of any hyperonic EoS !

 \rightarrow Comprehensive study of hyperonic EoSs in NS mergers

Isolate thermal behavior of hyperons

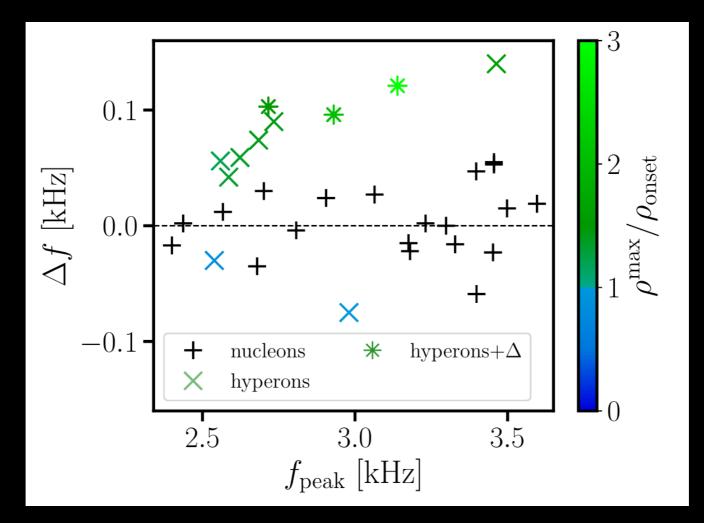
- Idea: assume T=0 EoS do not contain any information and adopt hyperonic EoS to be purely nucleonic (obviously incorrect assumption but necessary)

- supplement with approximate thermal pressure treatment to mimic "nucleonic" thermal behavior

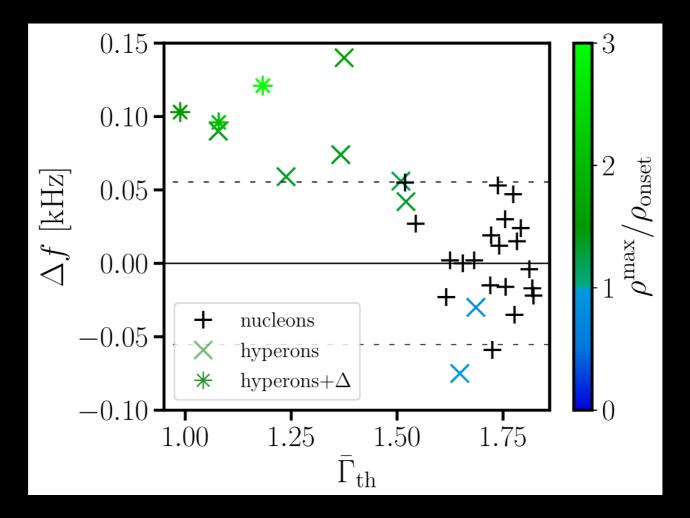
 $P_{th} = (\Gamma_{th} - 1)\epsilon\rho$ $\Gamma_{th} = 1.75$ found to reproduce nucleonic EoSs

Compare $\Gamma_{th} = 1.75$ runs vs. full T-dependent simulations \rightarrow thermal behavior of hyperons

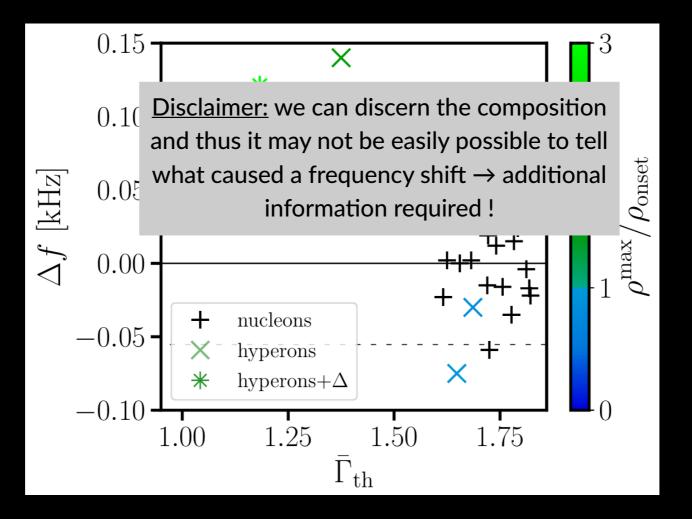
▶ Delta f describes impact of hyperons on thermal behavior \rightarrow in principle measurable !!



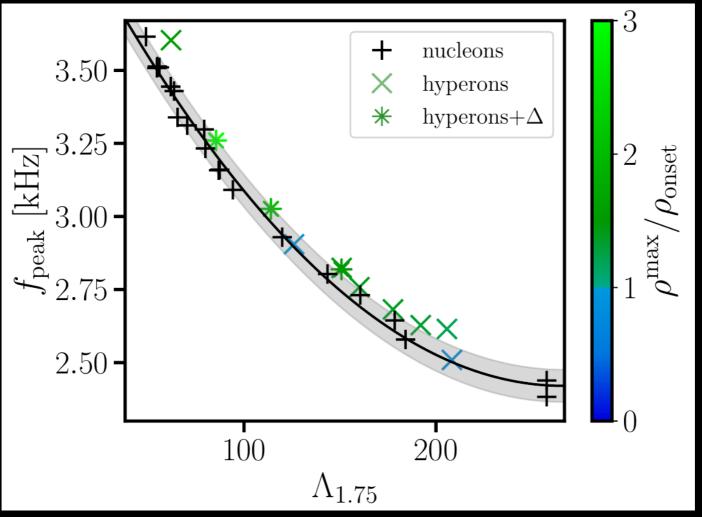
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► Concrete scenario: hyperonic models have tendency to yield increased fpeak



Blacker et al. 2023

Summary

- ► Pions may affect stellar structure and merger dynamics
 - \rightarrow empirical relations still hold (cancelation effect)
 - but should be considered for certain applications (systematic bias)
- Quark matter can lead to a characteristic shift of postmerger frequency
 - by compactification of remnant
 - also threshold mass affected
- Hyperons modify thermal behavior of $EoS \rightarrow only minor$ frequency shift