

Gain Calibration for LiteBIRD detectors

23/05/2023

Workshop LiteBIRD-Italia 2023 @ INFN-LNF
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- In order to have an accurate detection of r we must be able to correctly convert the signal measured by the detectors from μA to K_{CMB}
 - Strict requirements on Gain calibration

Requirements by Ghigna et al.

Band (GHz)	$\Delta_{g,\gamma}$	$\delta_{g,\gamma}$
40	2.5×10^{-3}	2.0×10^{-2}
50	7.5×10^{-3}	6.0×10^{-2}
60	7.5×10^{-3}	6.0×10^{-2}
68	7.5×10^{-3}	10.8×10^{-2}
78	1.0×10^{-2}	14.4×10^{-2}
89	5.0×10^{-3}	7.2×10^{-2}
100	1.0×10^{-3}	2.3×10^{-2}
119	1.0×10^{-3}	2.5×10^{-2}
140	2.5×10^{-3}	5.7×10^{-2}
166	7.5×10^{-4}	1.6×10^{-2}
195	2.5×10^{-4}	0.6×10^{-2}
235	5.0×10^{-4}	0.8×10^{-2}
280	1.0×10^{-3}	1.6×10^{-2}
337	1.0×10^{-4}	0.16×10^{-2}
402	1.0×10^{-4}	0.18×10^{-2}

Revisited by F. Carralot

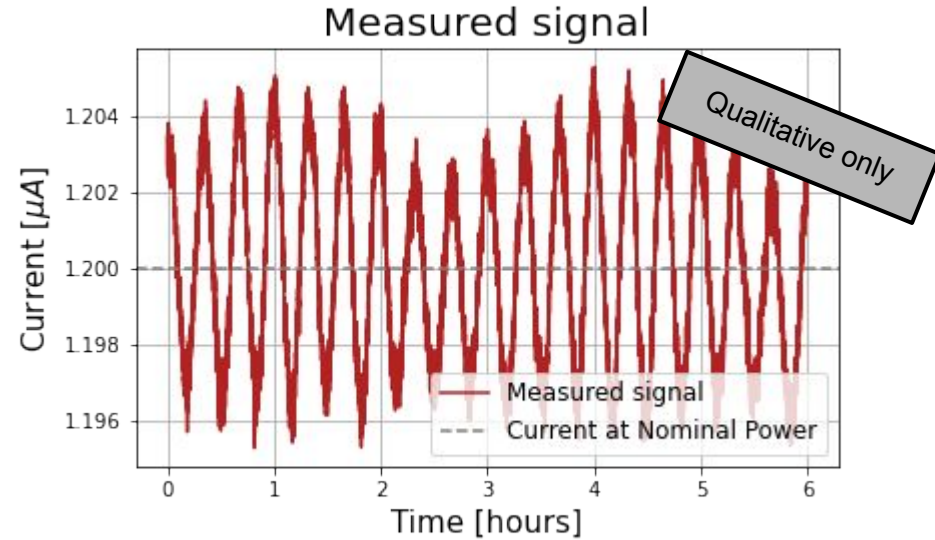
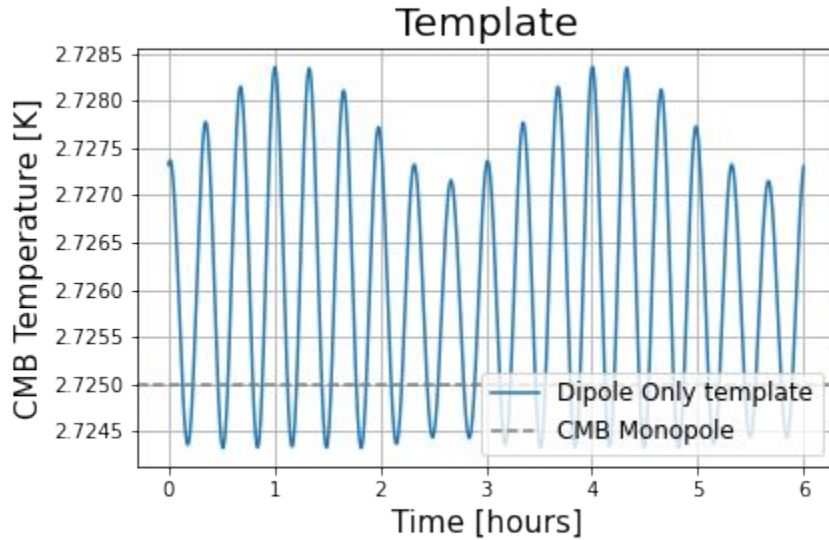
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gain requirements with component separation

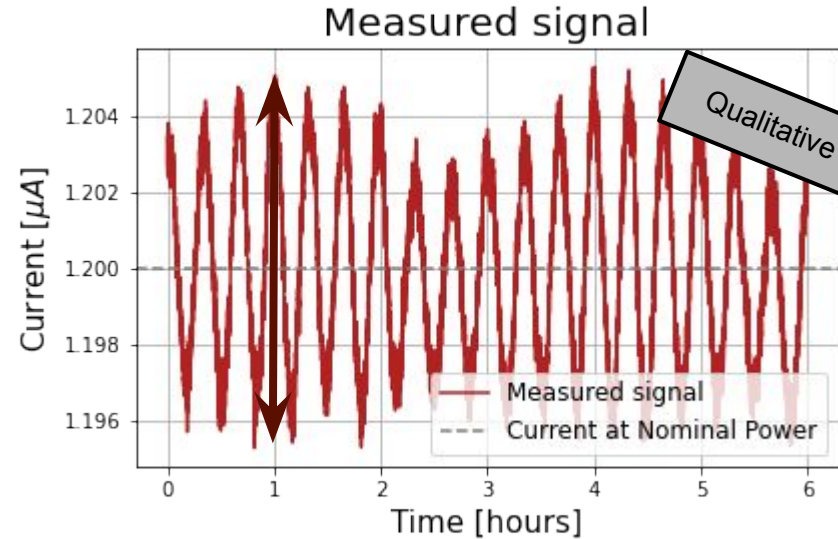
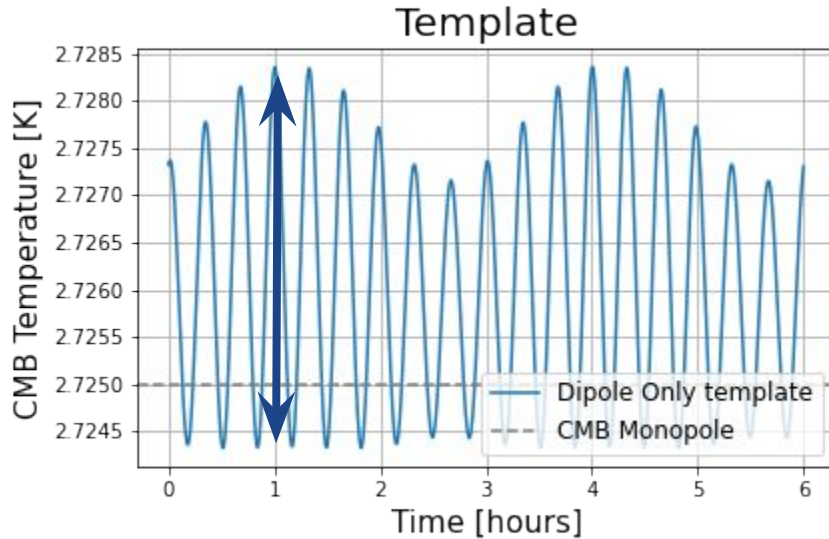
Speaker: Florie Carralot

Table 2: A summary of the requirements in terms of the overall frequency bands ($\Delta_{g,\gamma}$), and per detector ($\delta_{g,\gamma}$) assuming the number of detectors in Table 1.

- We calibrate the gain of the detectors by comparing the measured timelines with a template of the expected signal.



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Oversimplification

$$g = \frac{\Delta I}{\Delta T}$$

By comparing the amplitude of the oscillations of the measured signal with the expected dipole signal we can obtain the gain of the detector

- More precisely we can model:
 - Data: $d(t) = g(m(t) + n(t))$
 - Template: $m_0(t)$
 - Noise: $n(t)$
- The integral formula used to estimate the gain is:

$$\tilde{g} = \frac{\int d(t) m_0(t) dt}{\int m_0(t)^2 dt} \approx g \left(1 + \frac{\int n(t) m_0(t) dt}{\int m_0(t)^2 dt} \right)$$

- Which means our error on gain calibration is:

$$\delta g = \frac{\int n(t) m(t) dt}{\int m(t)^2 dt} = \frac{\sum_f M(f)N(f)^*}{\sum_f |M(f)|^2}$$

Time Domain
Fourier Domain

- Things to Notice:

★ 1

If our calibration signal goes to 0 the calibration error diverges.

$$\delta g = \frac{\int n(t) m(t) dt}{\int m(t)^2 dt} = \frac{\sum_f M(f) N(f)^*}{\sum_f |M(f)|^2}$$

Time Domain Fourier Domain

It is therefore better to filter out the frequencies where we have low S/N

→ It is possible to find analytically the **optimal filter**:

$$\mathbf{F}^{\text{opt}}_i = \frac{1}{\mathbf{N}_i} \frac{1}{\sqrt{\sum_i \mathbf{M}_i / \mathbf{N}_i}}$$

- Things to Notice:

- ★ 2

- We are assuming a constant gain during the integration time

- This is not true if we calibrate on longer timescales

- Loop gain monitor proposed by Tijmen De Haan will be used to **record fluctuations of the gain**

- We inject a sinusoidal bias into the TES and we use that signal to calibrate the gain

More information on:



TdH_loop_gain_alg.pdf

CMB / ... / Proposal for TOD calibration strategy feb 16, 2023

Loop gain measurement algorithm 9/2/22 Tijmen de Haan 1 **Loop gain** measurement algorithm -- Context • August 2022, I traveled to

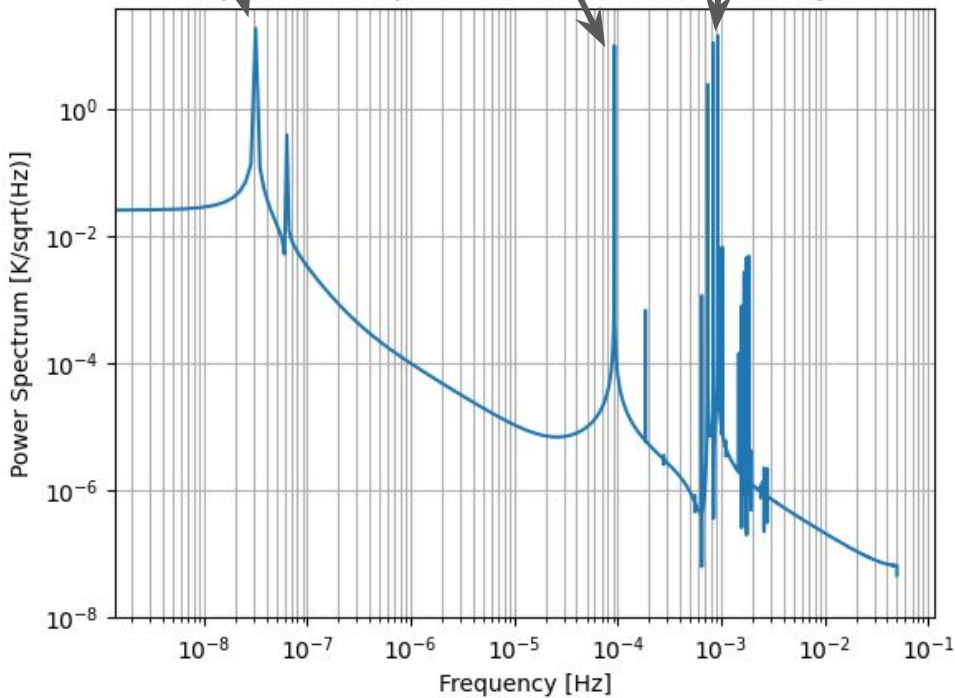
Our main source of calibration: the Dipole signal

orbital dipole
($f=1/\text{year}$)

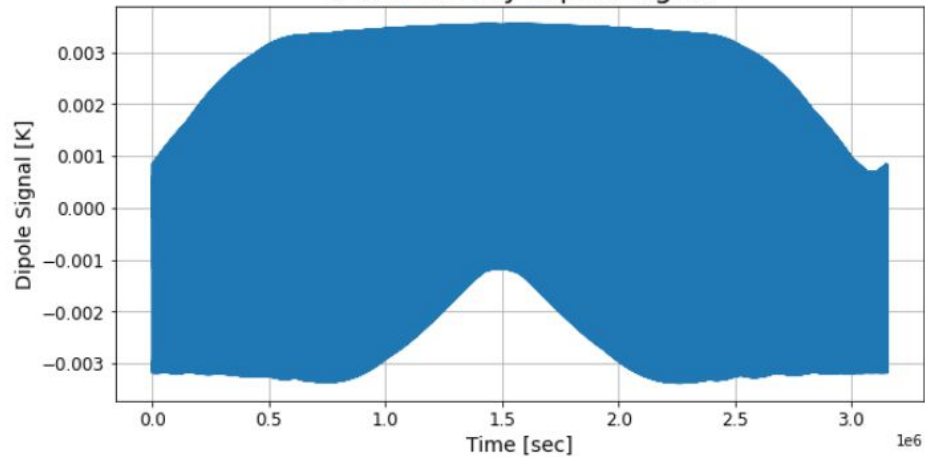
precession
($f=1/3\text{hr}$)

satellite spin
($f=1/20\text{min}$)

Dipole Power spectrum (tot obs time 3650 days)



1 Year of only Dipole signal



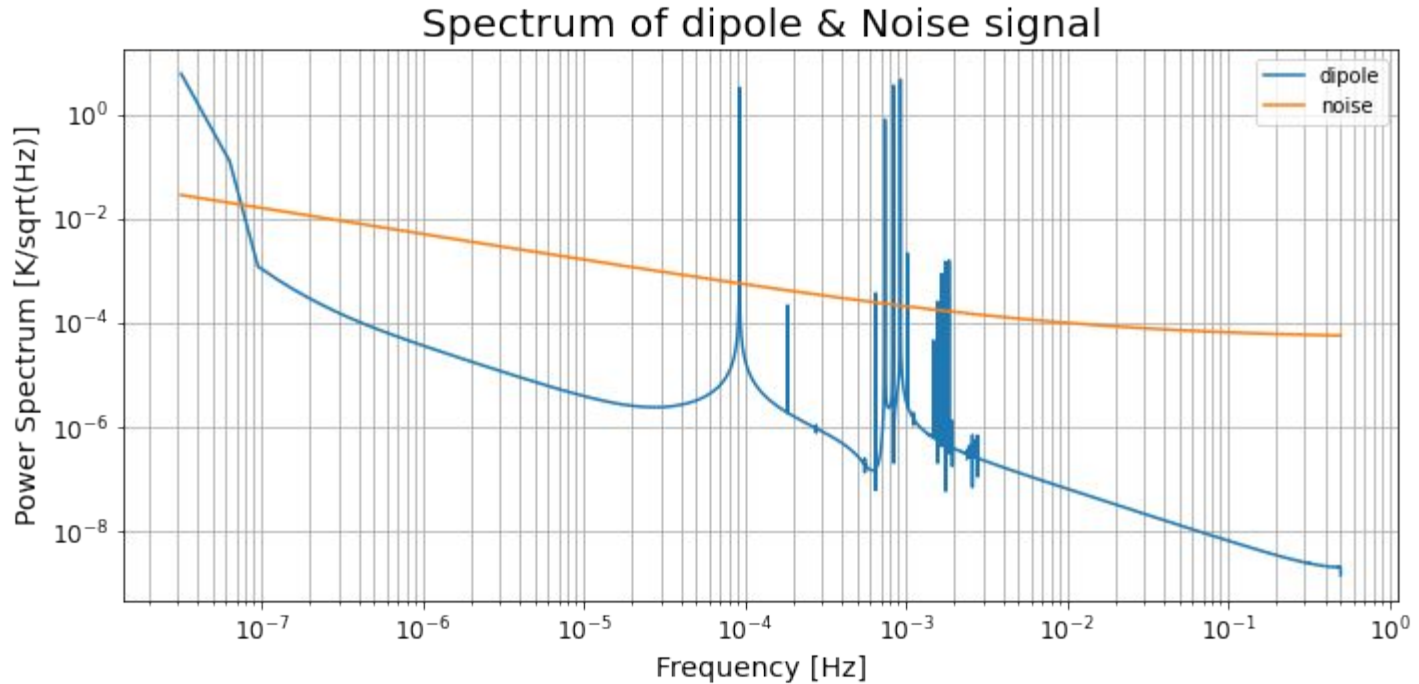
NOTE:
There are several parts of the spectrum where the dipole signal is very faint

Objective:

- I use TOAST / LB_SIM to simulate a detector's TOD
- I generate multiple realizations of $1/f$ noise and find the f_k that matches the requirements by Tommaso

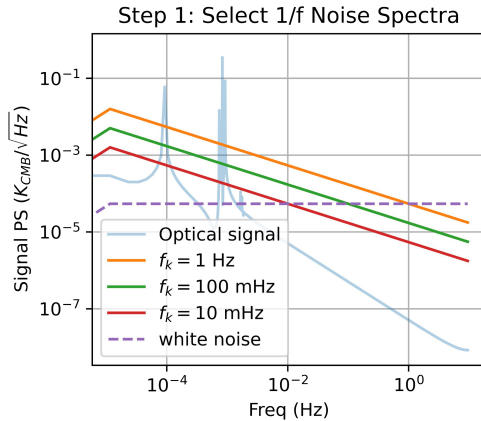
1/f Noise:

- NET 50 $\mu\text{K}/\sqrt{\text{Hz}}$
- $f_k = 10 \text{ mHz}$

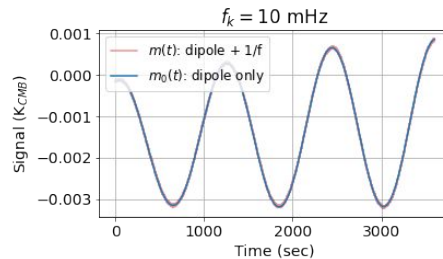
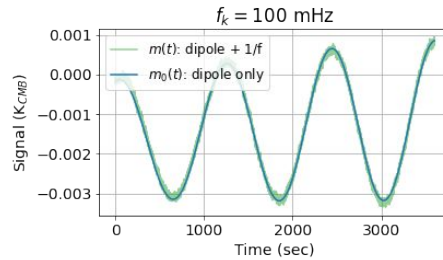
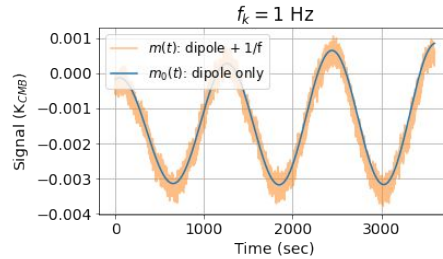


1/f Noise - Monte Carlo Simulations

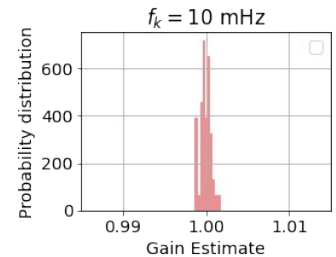
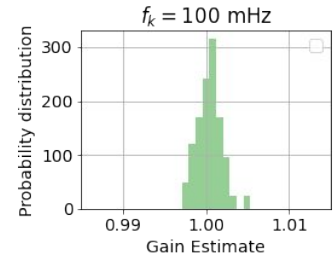
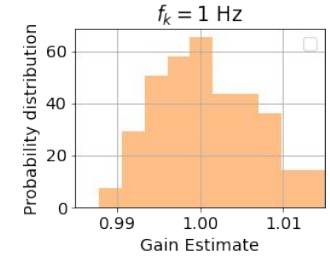
- Noise with different values of f_k



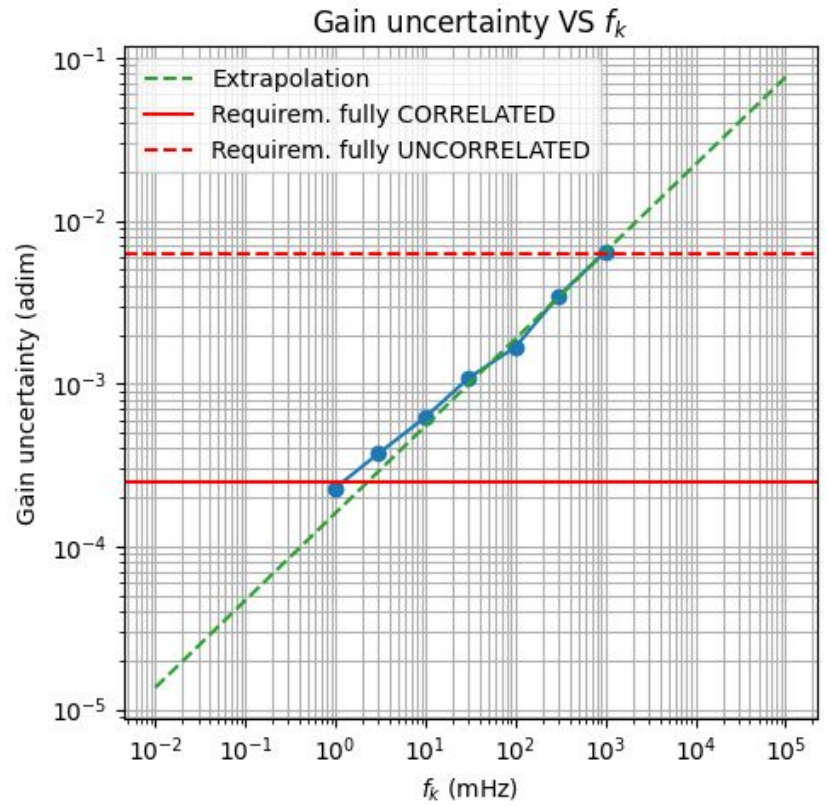
- Simulate TOD



- Calculate Uncertainty on our gain estimate



How to extract a requirement for calibration every 24 hrs



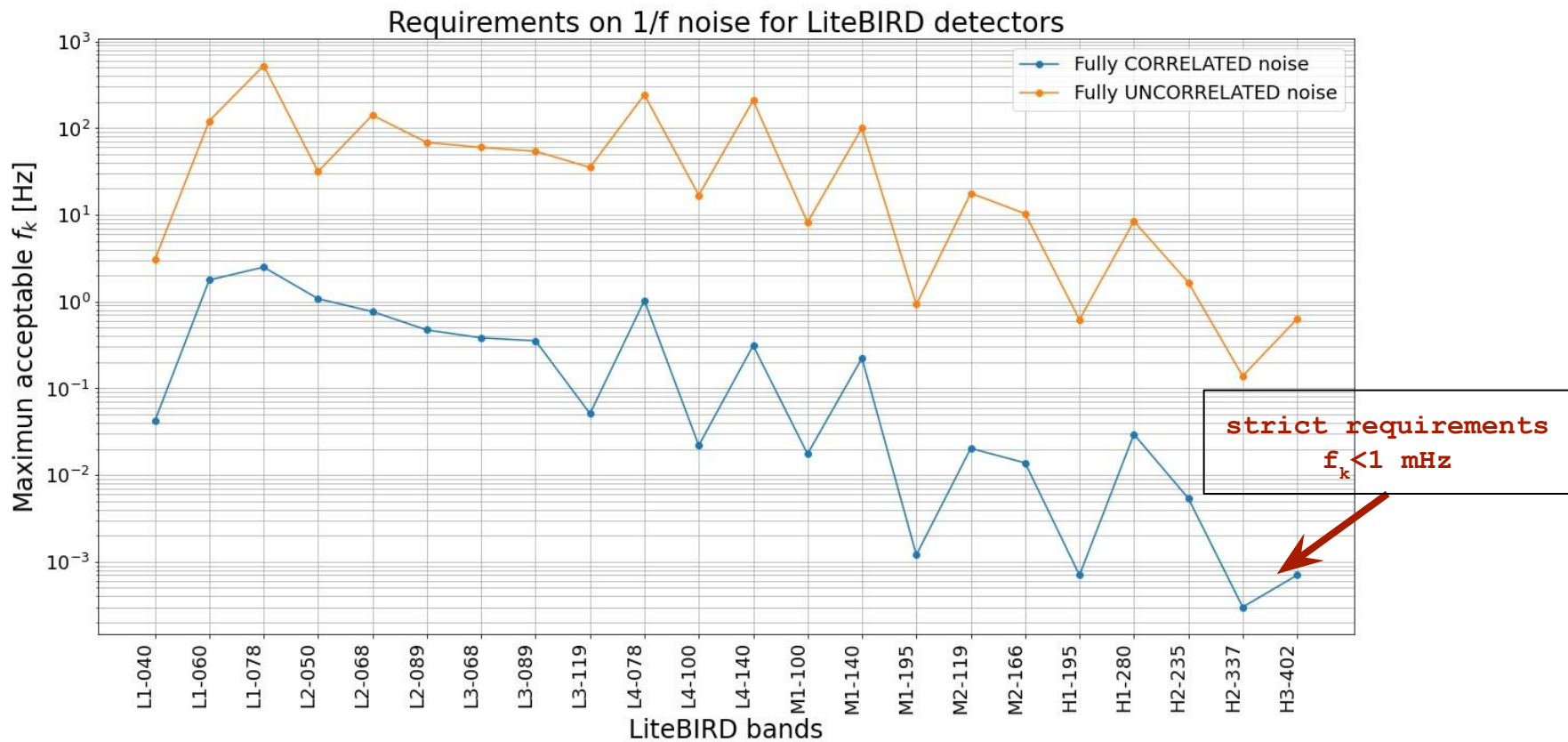
Maximum gain uncertainty if the noise is UNCORRELATED between detectors

($c=0$)

Maximum gain uncertainty if the noise is FULLY CORRELATED between detectors

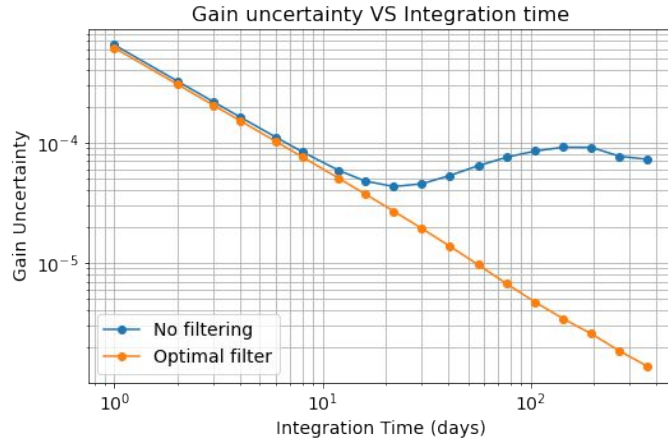
($c=1$)

Requirements calibrating in 24 hours



Extrapolating requirements for a calibration every 6 months

We therefore have the following proportionalities:



Using the optimal filter:
(T: integration time)

$$\delta g \propto 1/\sqrt{T}$$

$$\delta g = \frac{\int n(t) m(t) dt}{\int m(t)^2 dt} = \frac{\sum_f M(f) N(f)^*}{\sum_f |M(f)|^2}$$

Time Domain
Fourier Domain

For every frequency:
(ASD_N: Amplitude Spectral Density uK/sqrt(t))

$$\delta g \propto ASD_N$$

Too keep the same gain uncertainty calibrating in 6 months:

- $APS_{1D/6M}$: acceptable Amplitude Spectral Density if we calibrate in 1 day / 6 months
- We are **assuming $\alpha=1$** in the $1/f$

$$APS_{1D} \frac{1}{\sqrt{1 \text{ day}}} = APS_{6M} \frac{1}{\sqrt{6 \text{ months}}}$$
$$A \left(\frac{f_k^{1D}}{f} \right)^{\frac{1}{2}} \frac{1}{\sqrt{1 \text{ day}}} = A \left(\frac{f_k^{6M}}{f} \right)^{\frac{1}{2}} \frac{1}{\sqrt{6 \text{ months}}}$$
$$f_k^{1D} 365/2 = f_k^{6M}$$

We can rescale the requirements on f_k by the number of days we use to calibrate the gain

Requirements calibrating in 6 months

