

## **KM3NET STATUS AND RECENT RESULTS**

**ROSA CONIGLIONE ON BEHALF OF THE KM3NET** COLLABORATION INFN - LABORATORI NAZIONALI DEL SUD (ITALY)



### **RICAP-24 Roma International Conference on AstroParticle Physics**

### Frascati 23-27 September 2024





- KM3NeT/ARCA (Astroparticle Research with Cosmics in the Abyss)
  - observation of high energy (GeV ÷ PeV) neutrino sources 👉 a telescope offshore Capo Passero (Sicily-Italy) is under construction at a depth of 3500m
- KM3NeT/ORCA (Oscillation Research with Cosmics in the Abyss)
  - detect neutrinos of tens of GeV is under construction at a depth of 2500m

### 1 collaboration 1 technology - 2 detectors

### KM3NET

KM3NeT is a Mediterranean research infrastructure hosting two neutrino detectors and instrumentations for earth and sea sciences

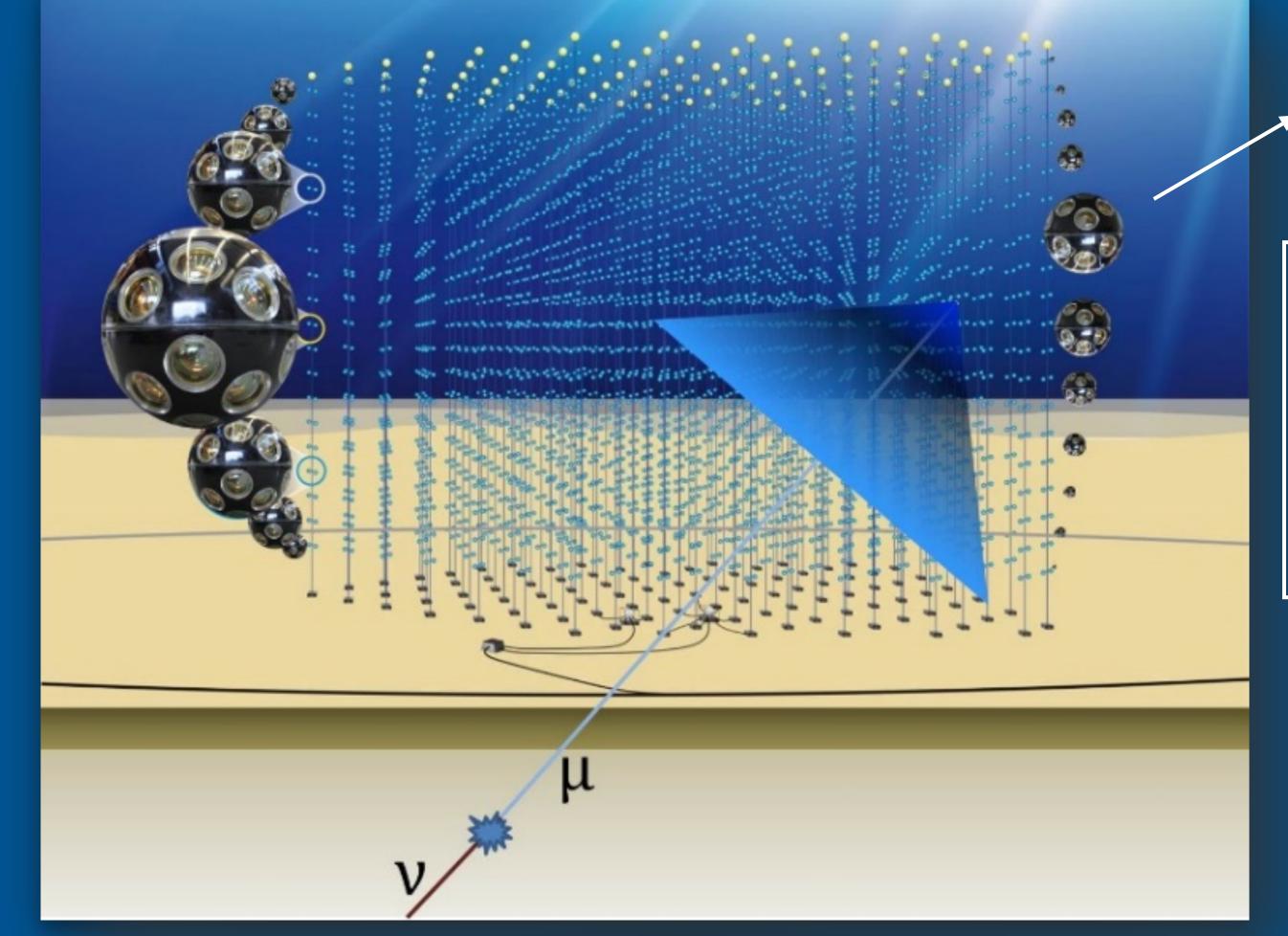
determination of the neutrino mass hierarchy f a detector offshore Toulon (France) able to







## THE KM3NET DETECTORS Same technology for the two detectors



**Detection Unit (DU)** 

Detectors under construction and taking data

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### Digital Optical Module DOM) 31 PMTs of 3 inches

Neutrino direction can be reconstructed from PMT positions and hit times

Photomultipliers to detect the light Acoustic system to measure the PMT position each 10 minutes

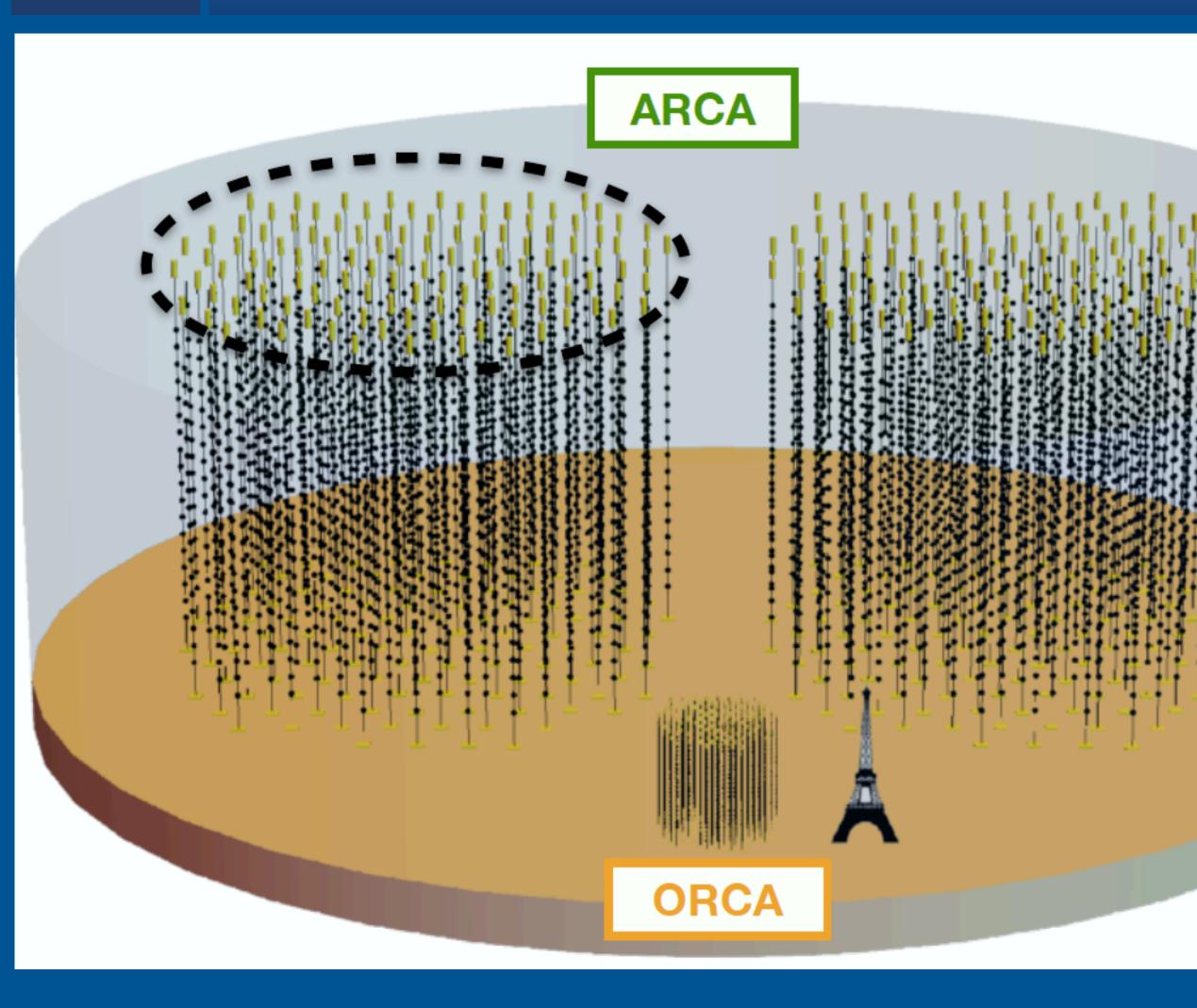
$$\nu_{\mu} \quad \begin{pmatrix} \theta \\ \approx \frac{1.5^{\circ}}{\sqrt{E_{\nu} [TeV]}} \\ N \\ \bullet \\ X \end{pmatrix}$$

For neutrino energies >TeV 🧲 neutrino and muon collinear



## THE KM3NET DETECTORS





Difference in the spatial distance of optical sensors

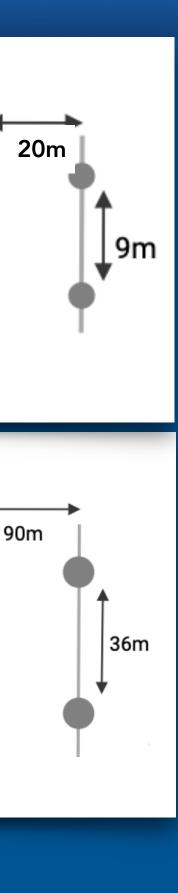
### **ORCA**

- Depth ~2500 m
- One block of 115 DUs
- Distance between DUs ~20 m
- Vertical distance between DOMs ~9 m
- ≈8 Mton

### ARCA

- Depth ~3500 m
- Two blocks of 115 DUs each
- Distance between DUs ~90 m
- Vertical distance between DOMs ~36 m
- Volume (0.5 × 2 ) km<sup>3</sup>

### 1 Building Block (BB) *—*115 Detection Units ARCA 2 BB (230 DUs) ORCA 1BB (115DUs)

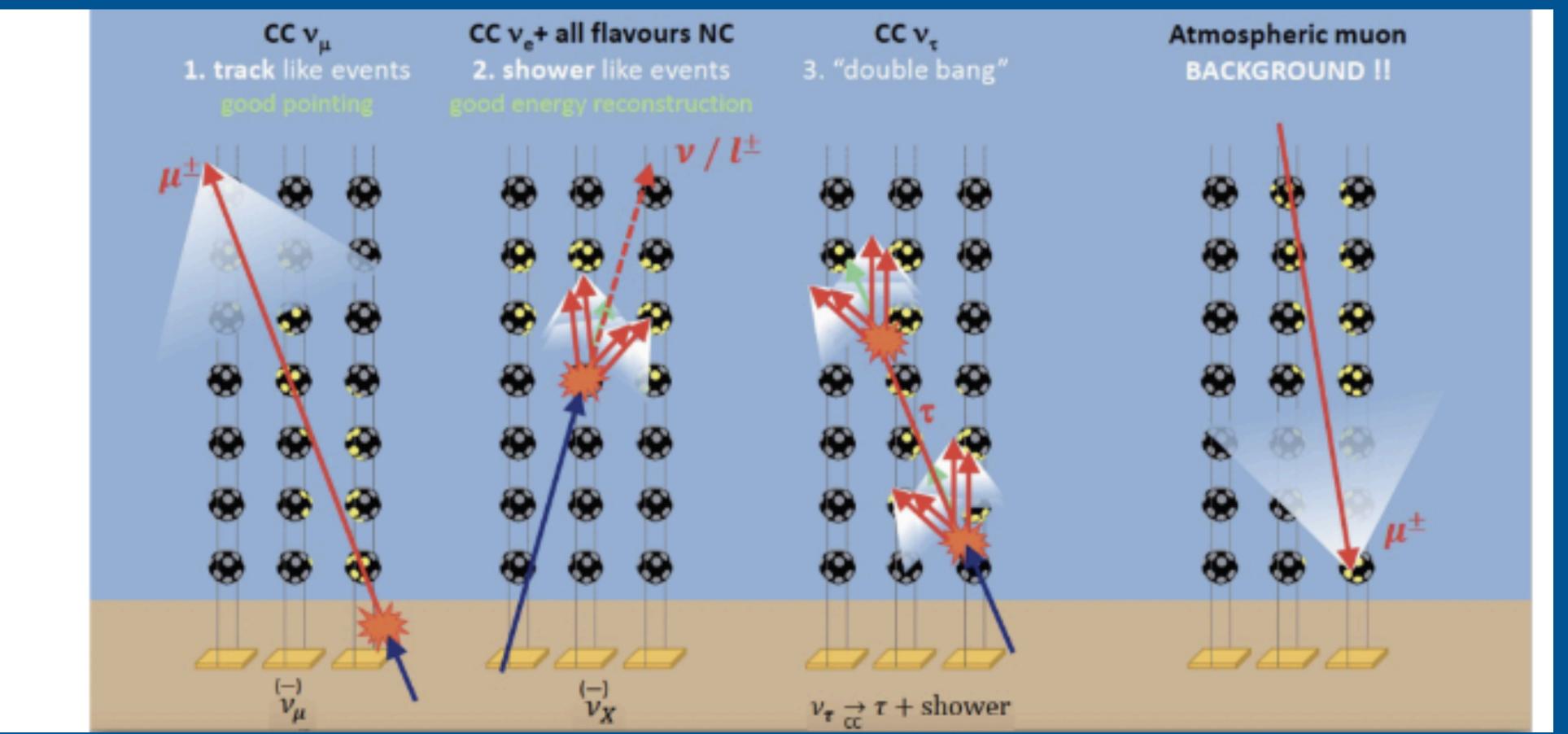




## DETECTION PRINCIPLE

Two kinds of event:

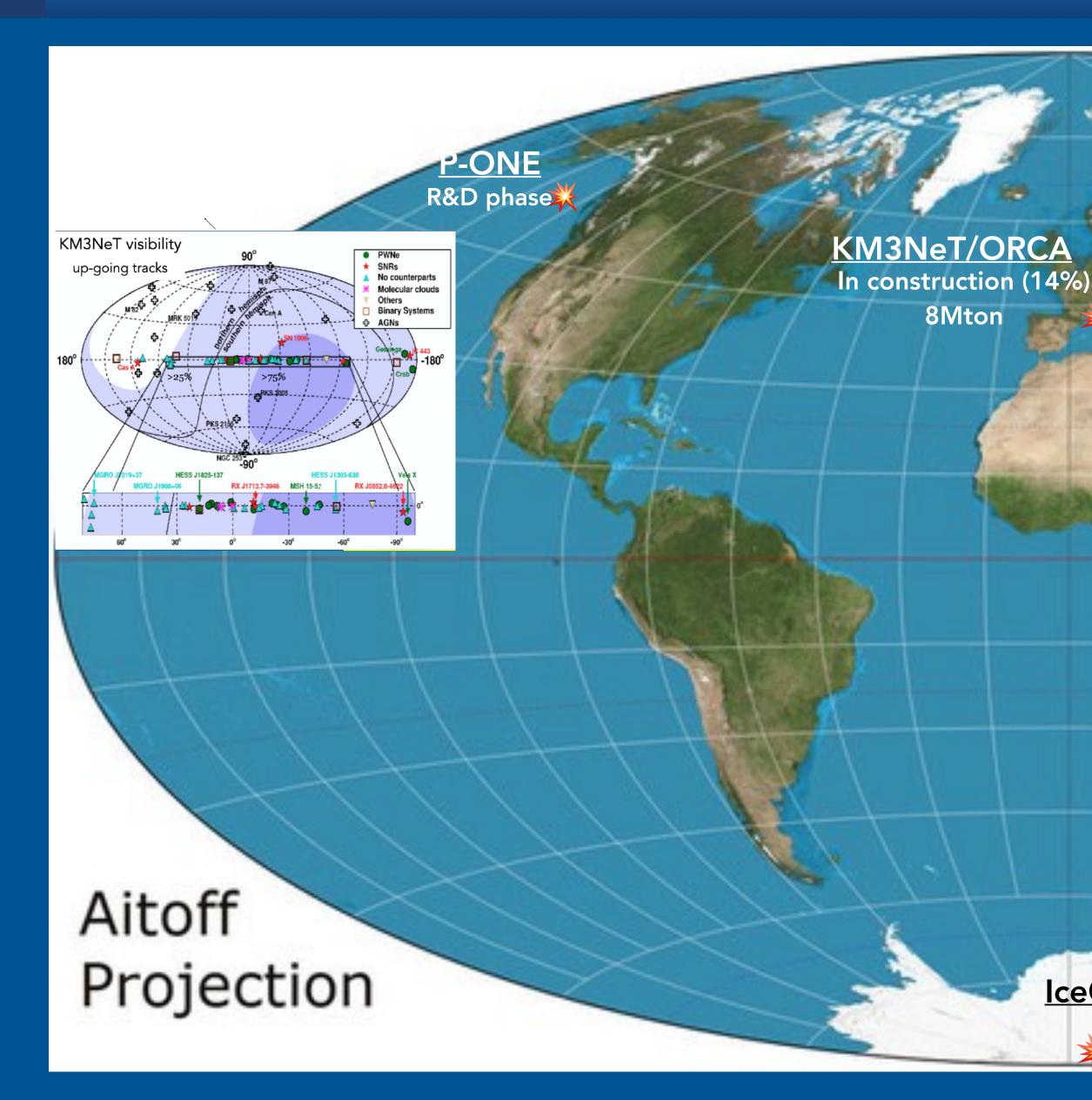
• "Track-like" events mainly from  $v_{\mu}$  CC interactions • "Cascade-like" events mainly from v<sub>e</sub> CC and NC interactions



# Shower ← @E<sub>v</sub>>100 TeV Ang. res. below 2° - Energy res. ~6%

**Tracks**  $\leftarrow @E_v > 100$  TeV Ang. res. below 0.1° - Energy res. ~ factor 2

### THE HIGH ENERGY NEUTRINO DETECTORS



#### In China also:

- HUNT ~30 km<sup>3</sup> in Lake Baikal or the South China Sea <u>http://hunt.ihep.ac.cn/</u>
- NEON ~10 km<sup>3</sup> in the

South China Sea https:// arxiv.org/pdf/2408.05122

R&D phase <u>TRIDENT</u> ~8 km<sup>3</sup>

lacksquare

**Baikal-GVD** 

In construction

1km<sup>3</sup>

IceCube visibility up-going tracks

Northern hemisphere



ANTARES

(dismantled) 0.01 km<sup>3</sup>

constructior

KM3NeT/ARCA

IceCube ~8 km<sup>3</sup>

Southern hemisphere



### The basic elements:

- Optical sensors DOMs

#### The Digital Optical Module

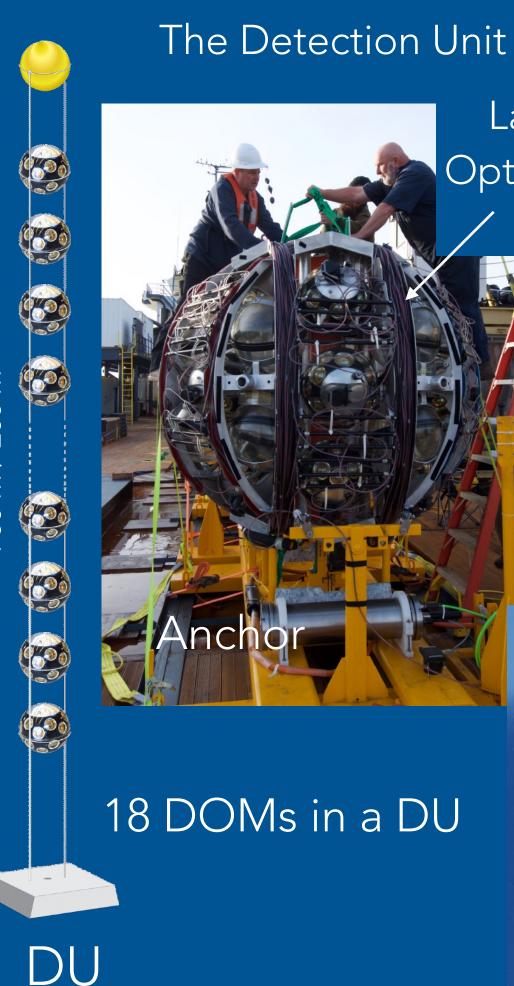


### DOM

It is a 17" glass sphere containing:

- 31 3" PMTs (photocathode aerea  $\simeq 3 \times 10^{\circ}$  PMTs)
- LED and Piezo

Front-end electronics -> FPGA



Video of a DU integration

### THE TECHNOLOGY

#### Sea talks of:

- L. Morales Gallegos (technology and construction)
- F. Benfenati (DAQ)
- D. Real (electronics)
- C. Lastoria (data quality)

Launcher of **Optical Modules** (LOM)

The Sea floor network:

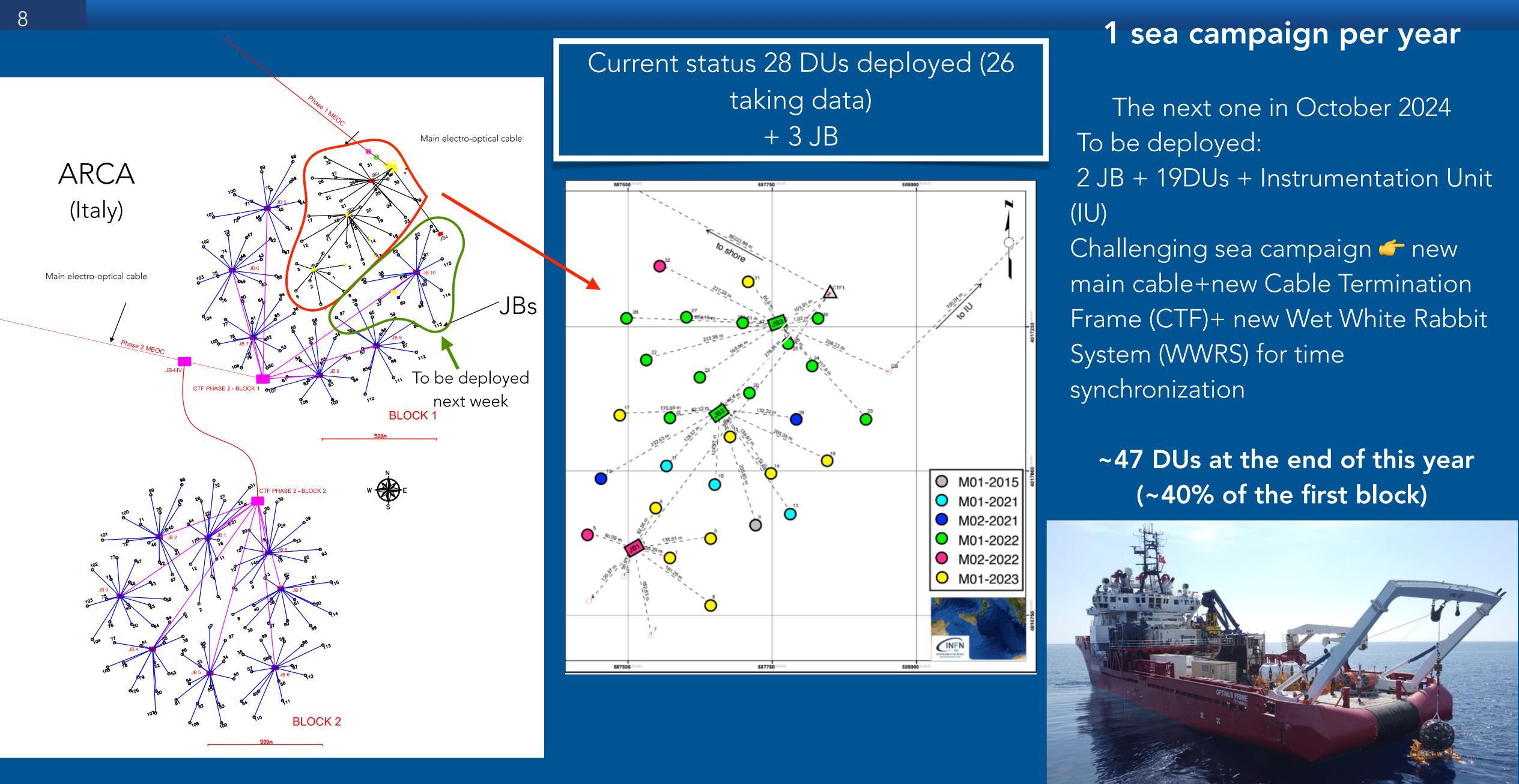
- Electro-optical cables from shore to the deep sea
- Junction boxes (JB)/nodes to distribute power and optical fibers
- In ARCA Cable Termination Frames (2 already installed)
- Interlink cables for connection of DU to JB and JB to the main cable



### ARCA JB

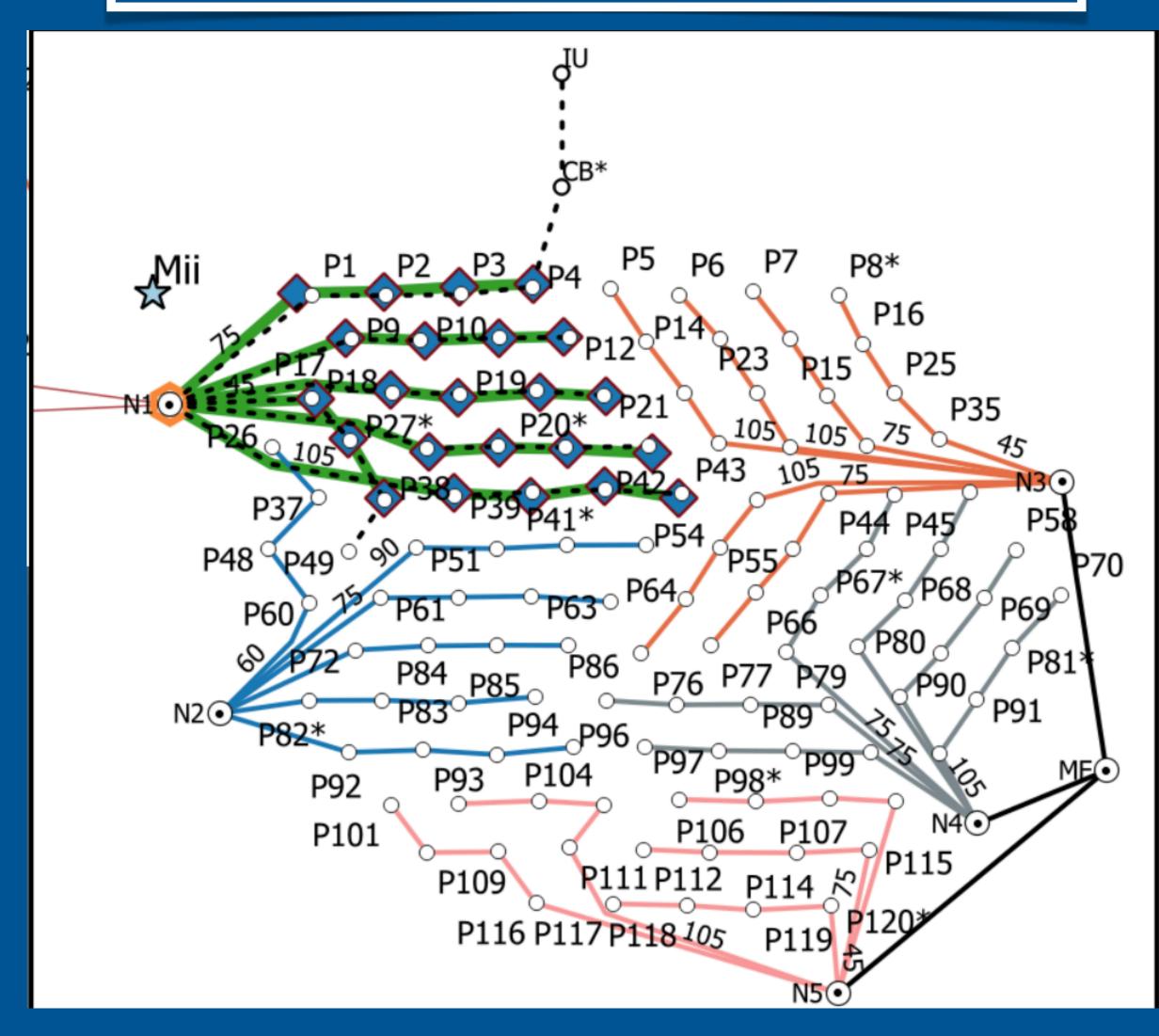


## THE KM3NET/ARCA STATUS



## THE KM3NET/ORCA STATUS

### Current status 23 DUs deployed (22 taking data)



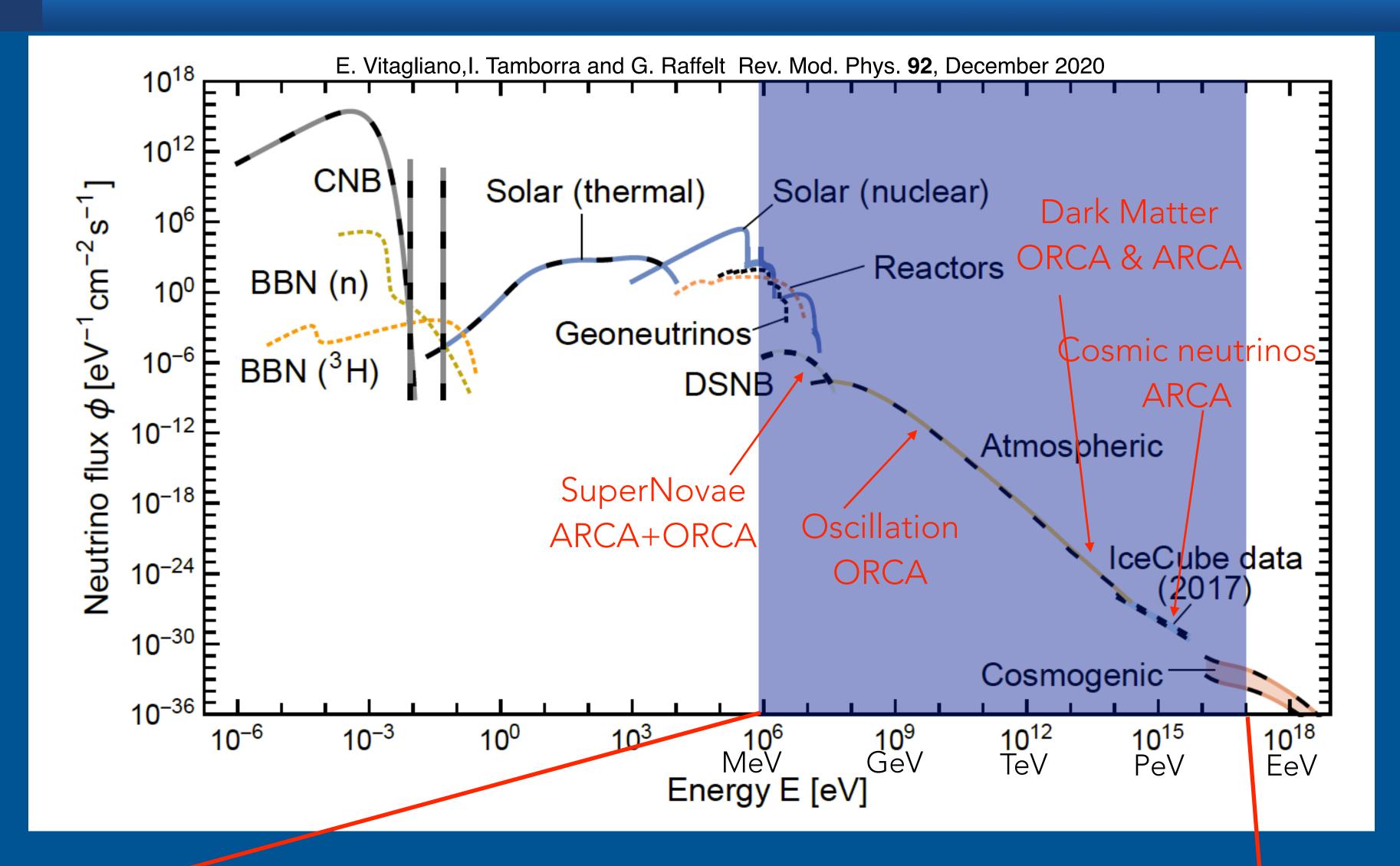
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Many sea campaigns/year

Next one foreseen before the end of the year 👉 complete the DUs of node1 (+2DUs) and deploy 4-5 DUs

> Almost completed the first node 20% of the full detector in water





KM3NeT detects neutrinos from MeV to PeV

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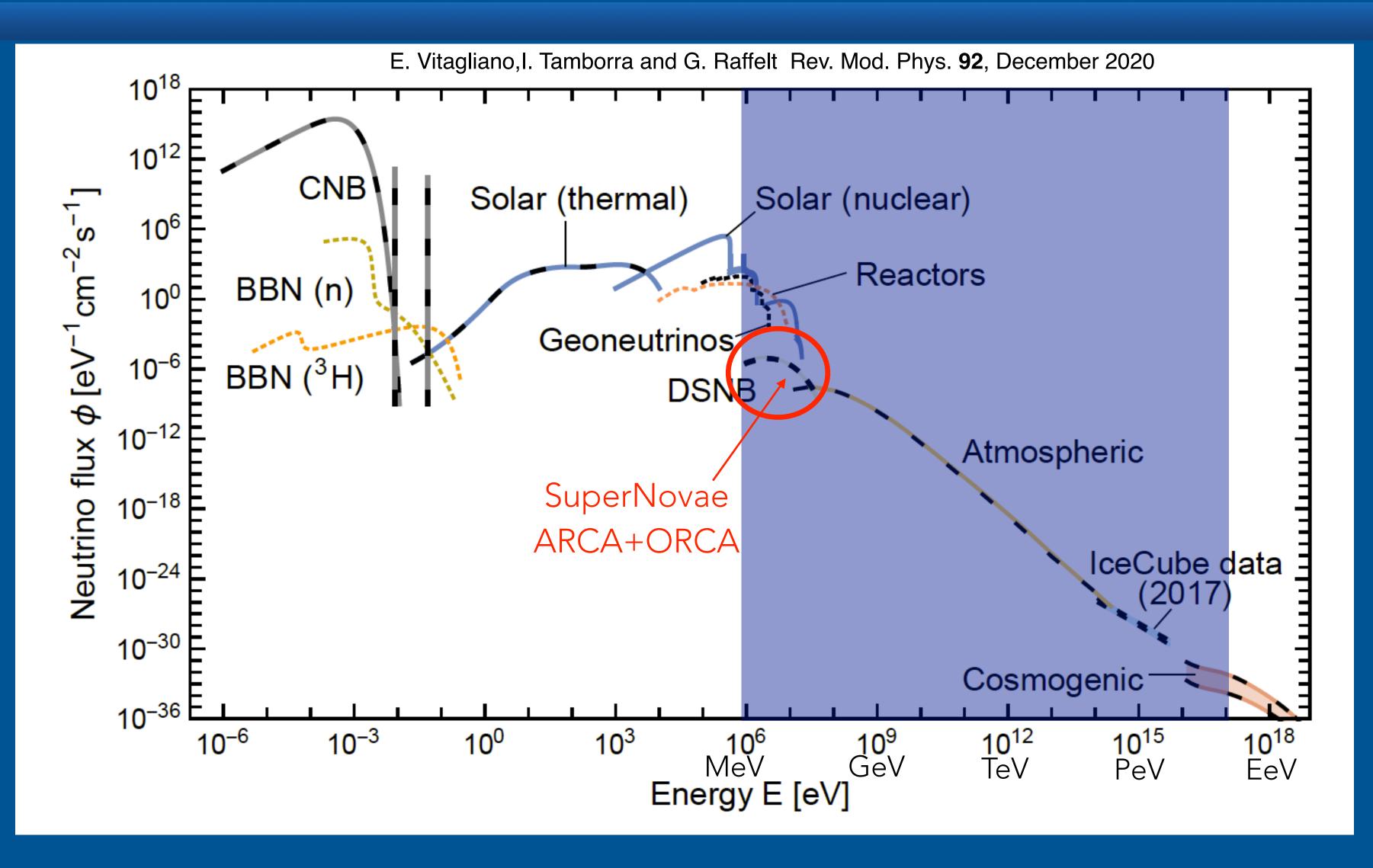
### THE NEUTRINO SPECTRUM AT EARTH

Dark Matter: talk of C. Gatius GeV Neutrinos: talk of J. Mauro

**GRB**: talk of J. Palacios Gomez



## SUPERNOVA EXPLOSION

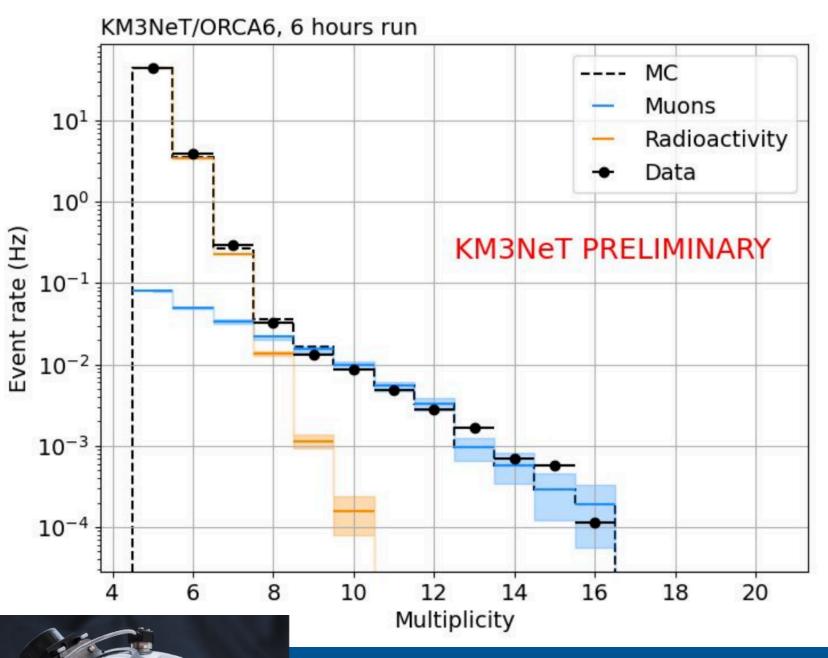


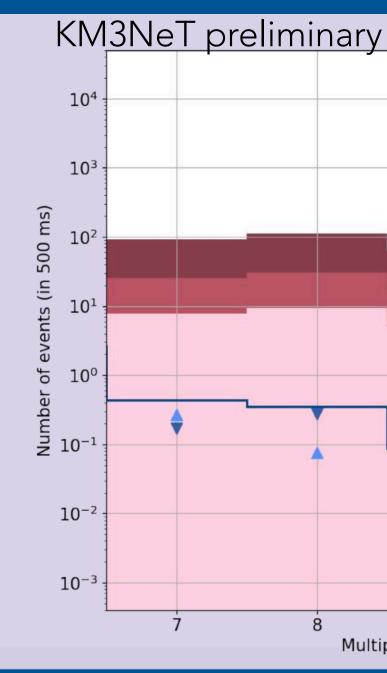
## SUPERNOVA EXPLOSION

## A DOM as a single detector Muon background rejection improved

### PMT multiplicity plot

#### Signal expected above background







#### Sea talk of Godefroy Vannoye

Significance

#### KM3NeT preliminary ORCA115 background 11 M☉, without BDT ARCA115 background 50 Total background 27 Mo, without BD7 40 Signal + background 40 M<sub>☉</sub>, without BDT $40 M_{\odot}$ at 10 kpc 30 11 Mo, with BDT $27~M_{\odot}$ at 10 kpc 27 Mo, with BDT Sensitivity (σ) 00 $11 \text{ M}_{\odot}$ at 10 kpc 40 M<sub>☉</sub>, with BDT Cloud ARCA28+ORCA24 lagellan geu 80 120 20 100 60 40 Distance (kpc) 10 Multiplicity

 $>5\sigma$  for ARCA+ORCA for 27M $_{\odot}$  at a distance <50kpc

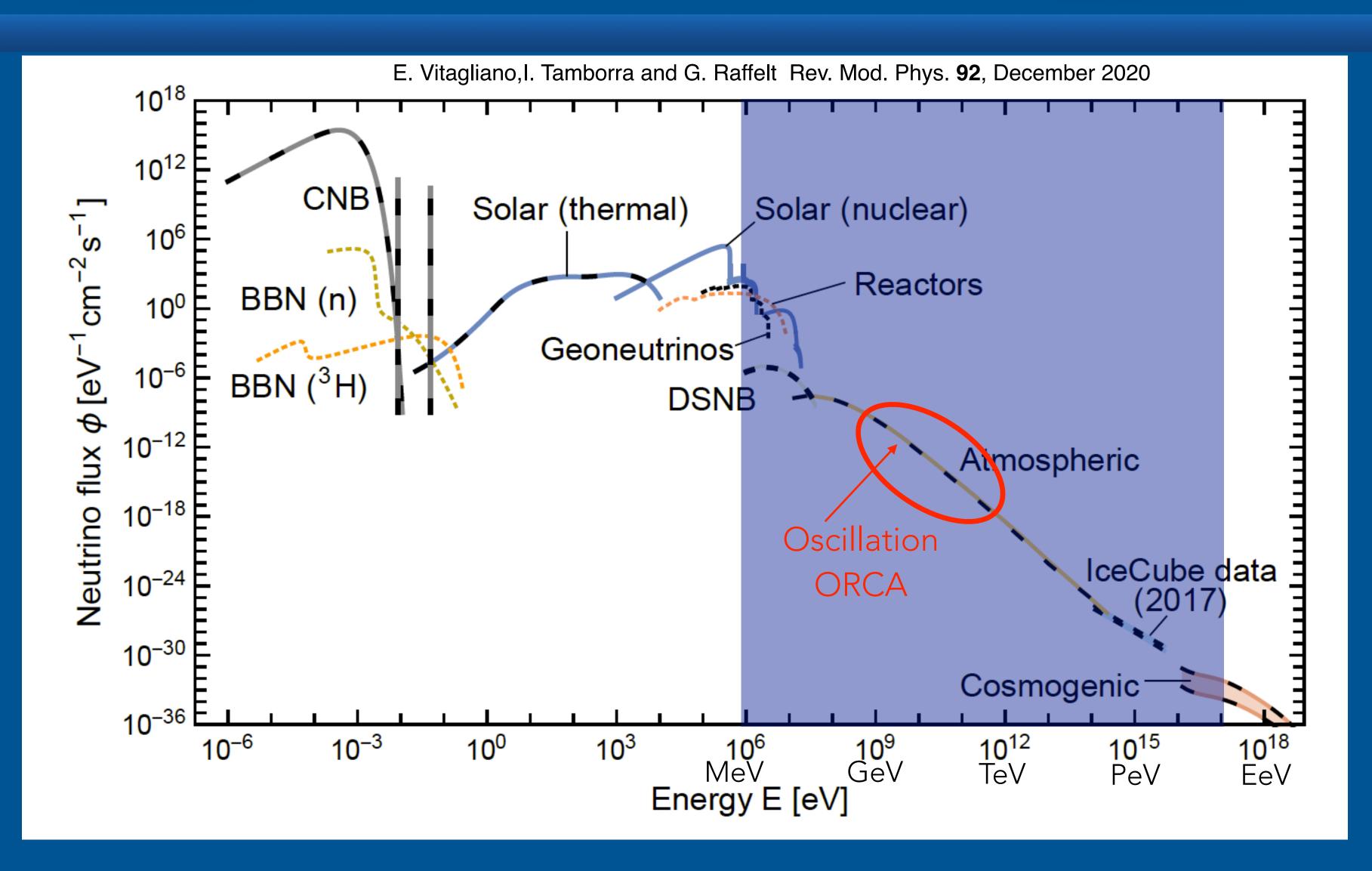
An on-line alert system for CCSN already implemented Integrated in SuperNova Early Warning System (SNEWS)







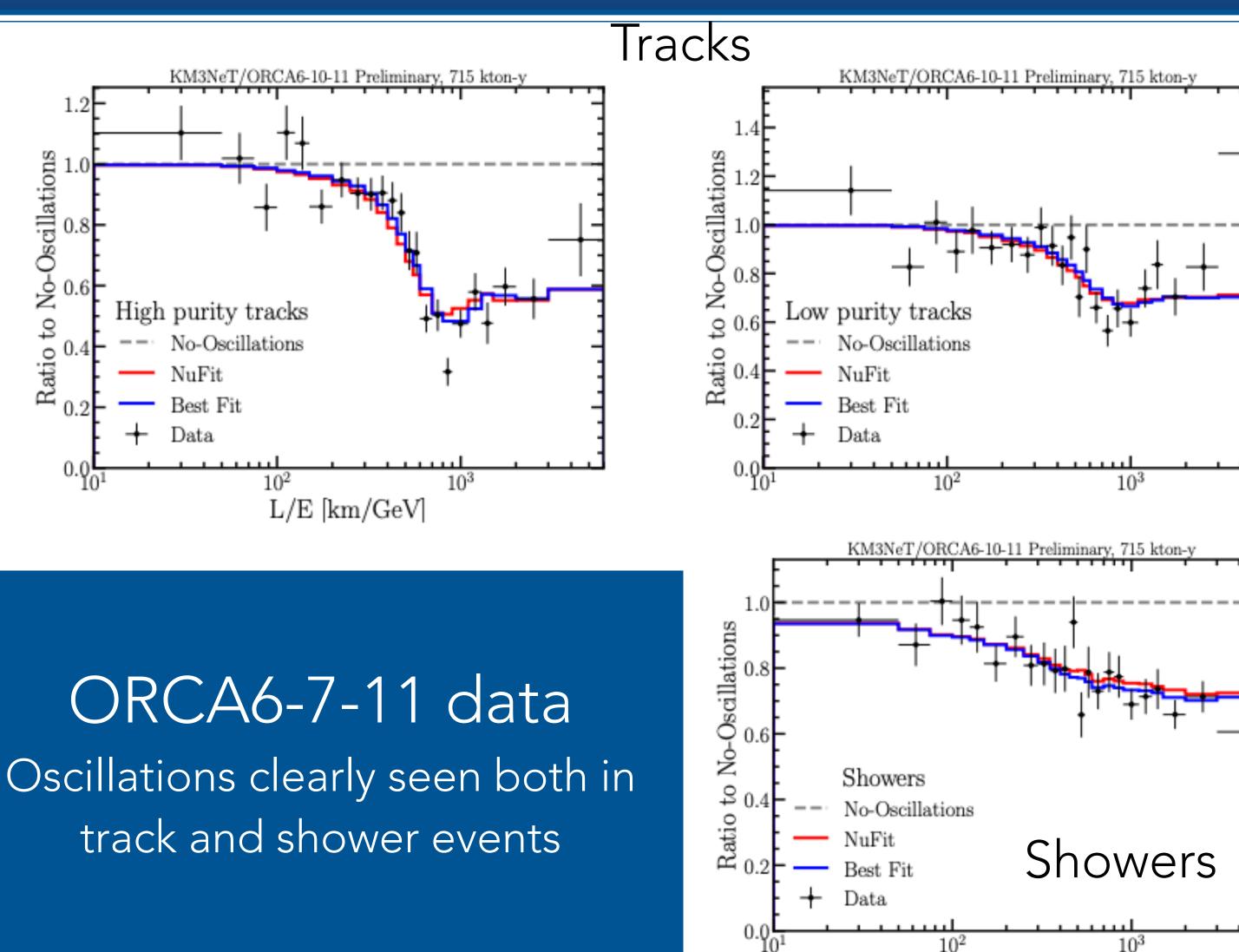
### ATMOSPHERIC NEUTRINOS AND OSCILLATION



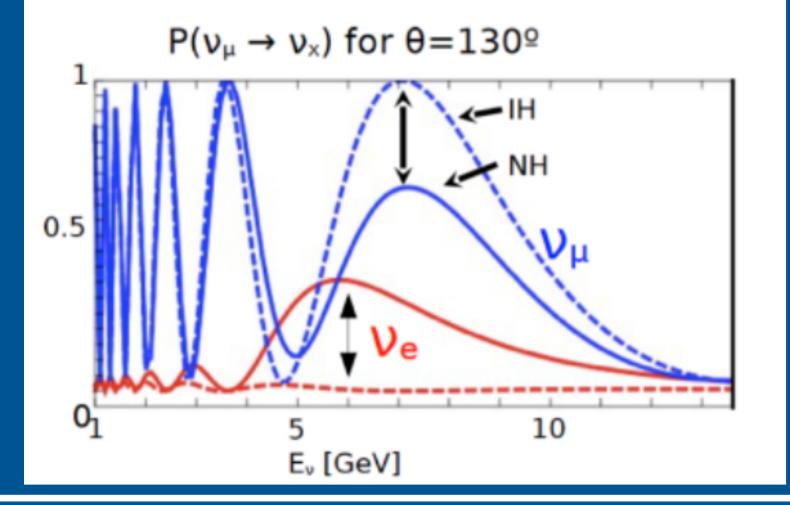
## NEUTRINO OSCILLATION WITH ORCA

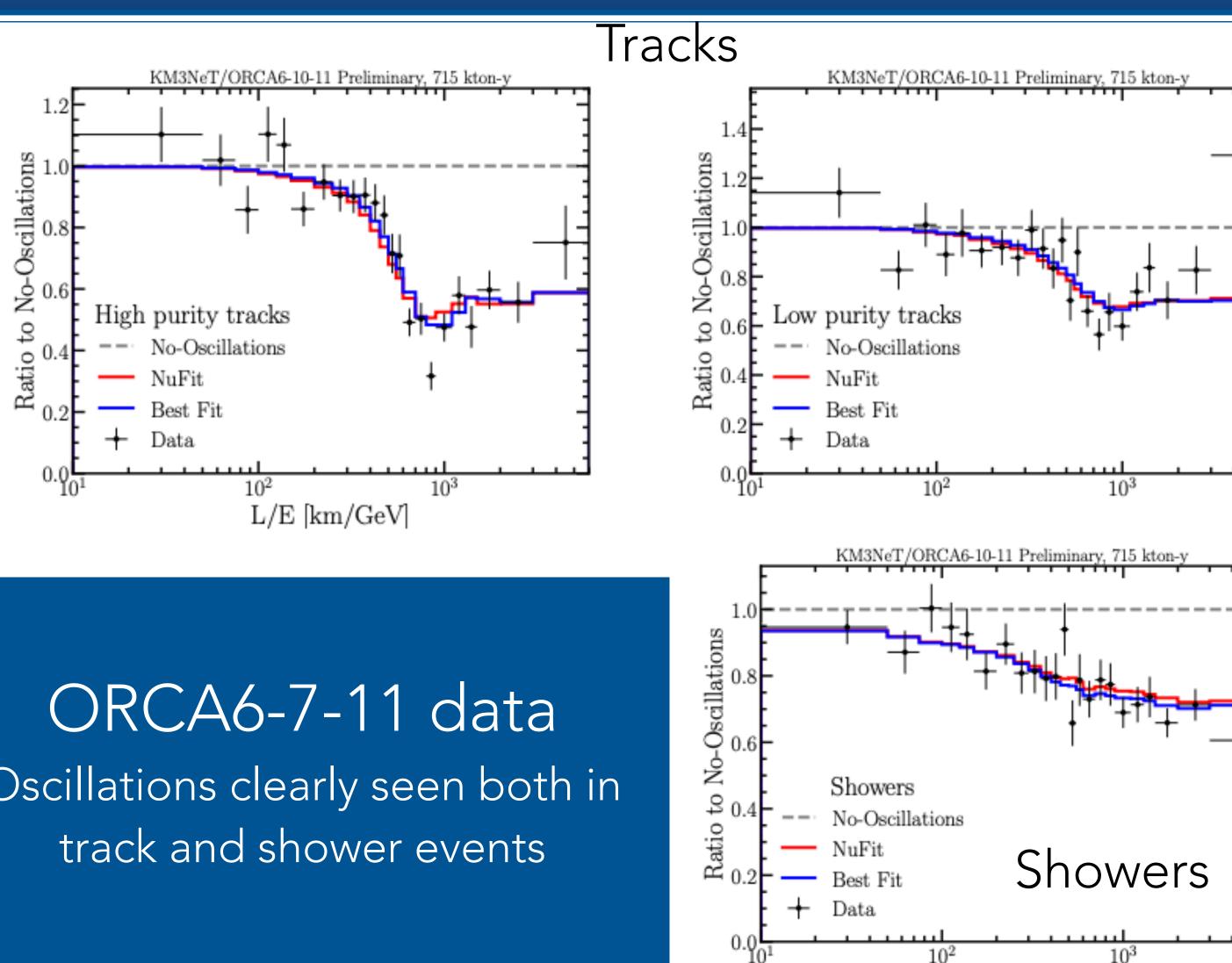
# Baseline from 50 to 12800 km detector

Neutrino Mass Ordering measuring atmospheric neutrinos crossing the Earth



#### Energy range of interest 5-15 GeV





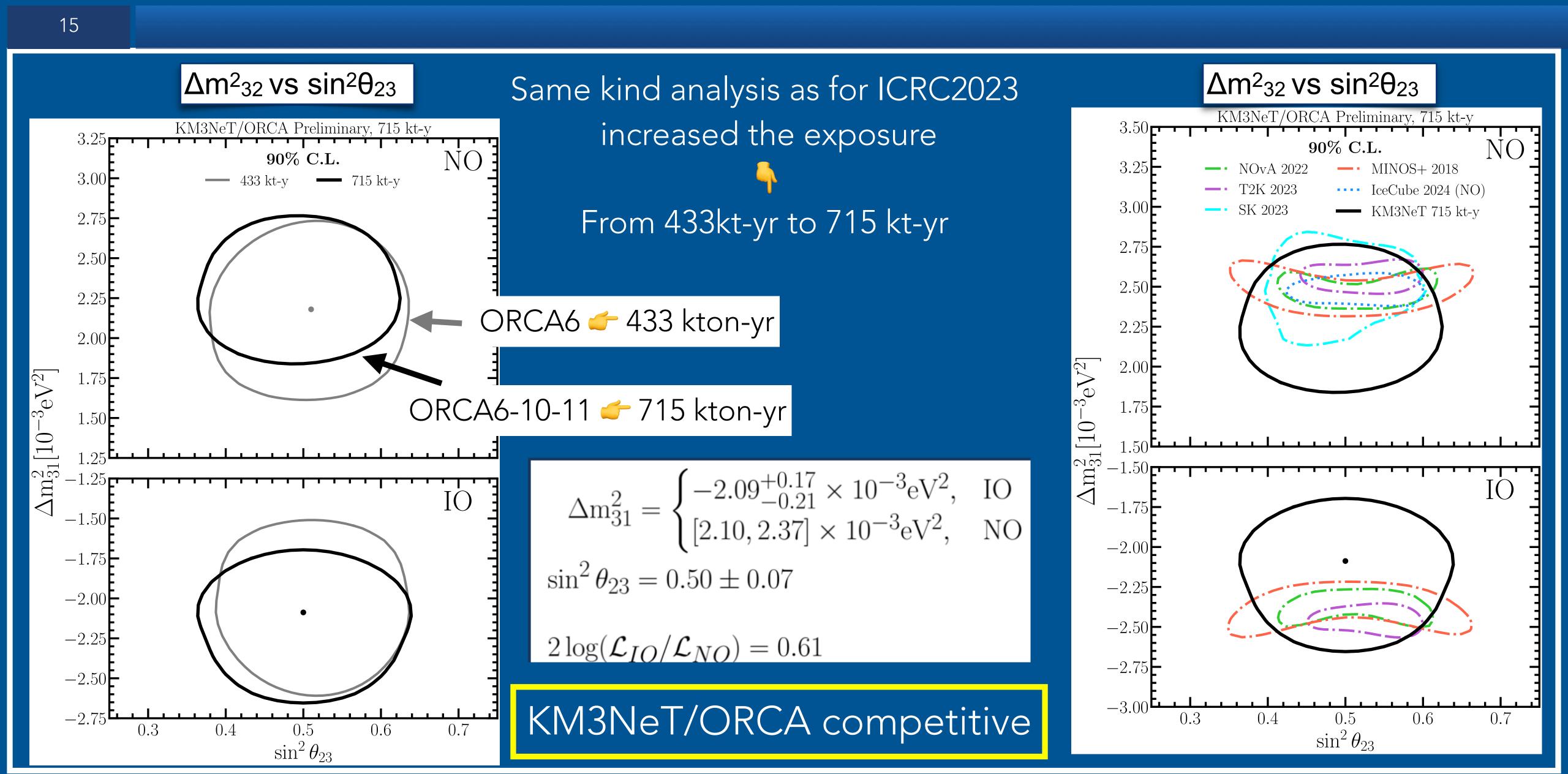




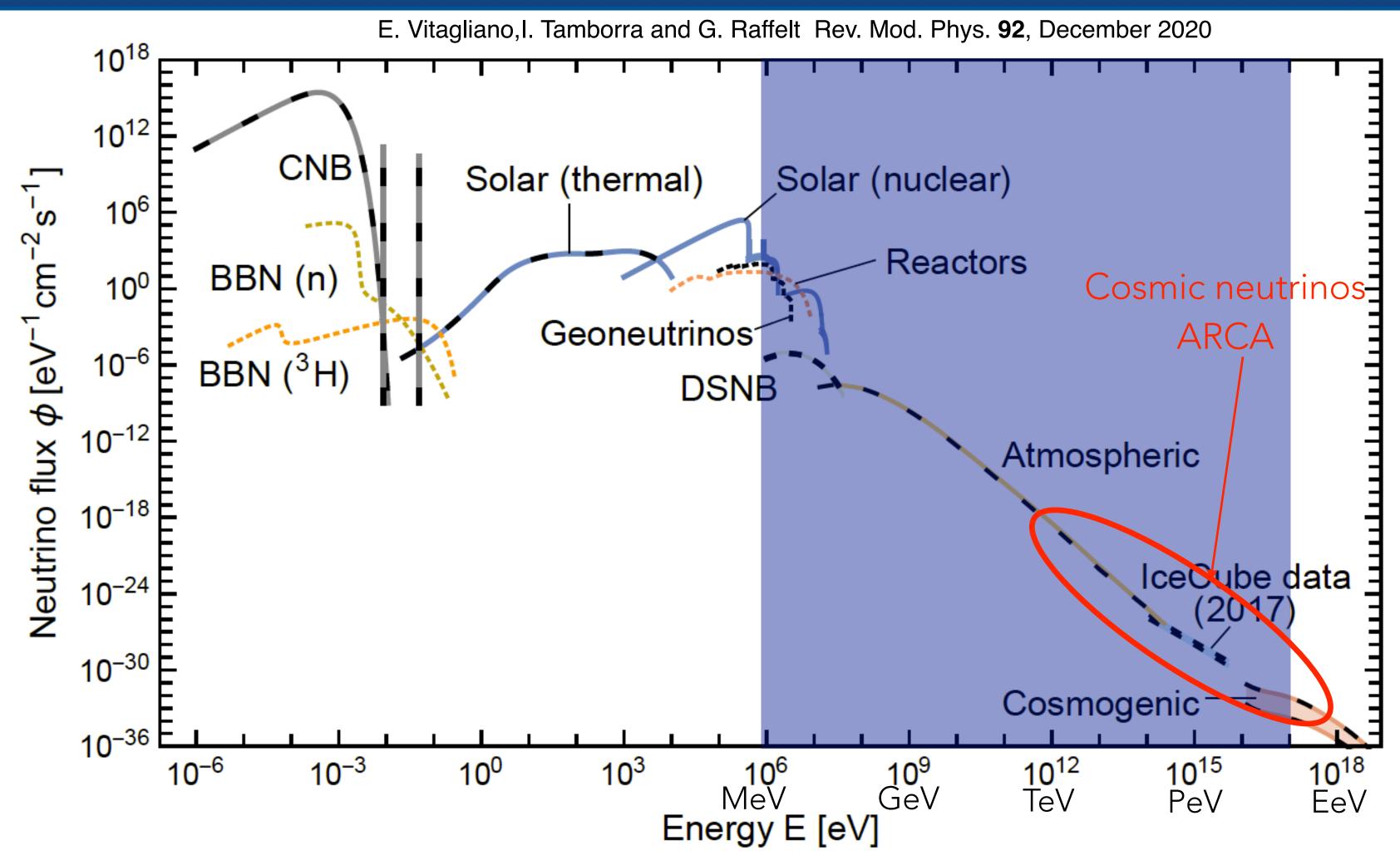
L/E [km/GeV]



## NEUTRINO OSCILLATION WITH ORCA

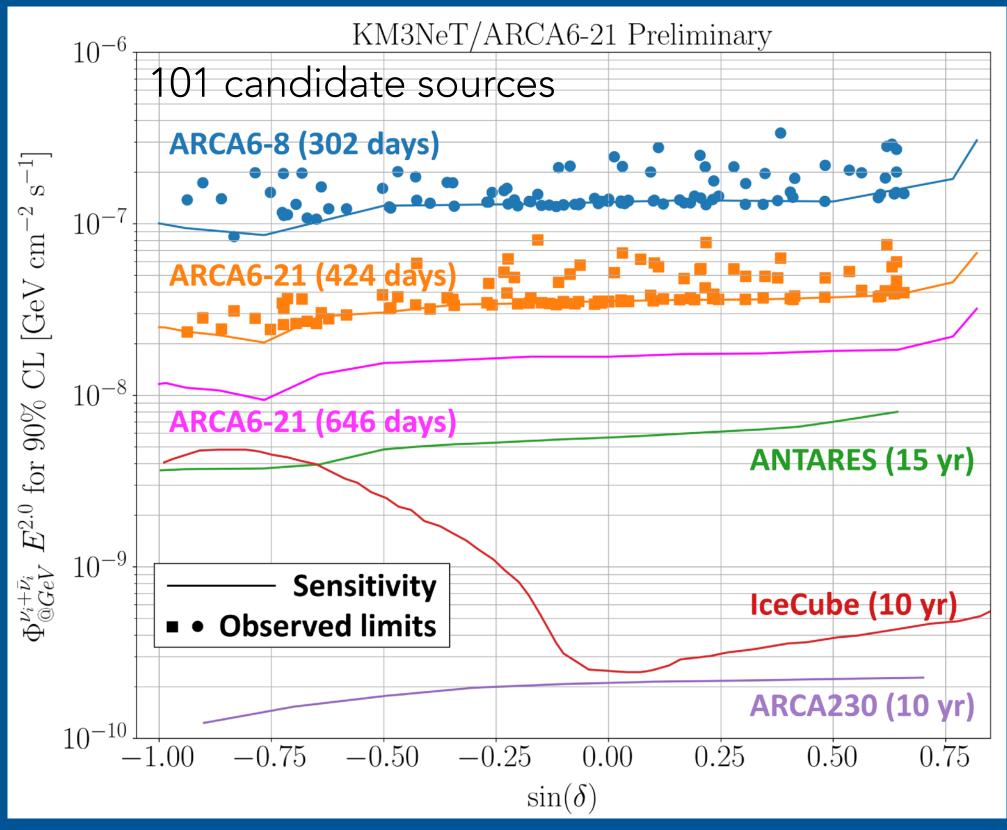


## THE COSMIC HIGH-ENERGY NEUTRINOS



## SEARCH FOR POINT-LIKE SOURCES

ARCA21 unblinded until December 2022 Unblinding of ARCA21 data full period expected very soon

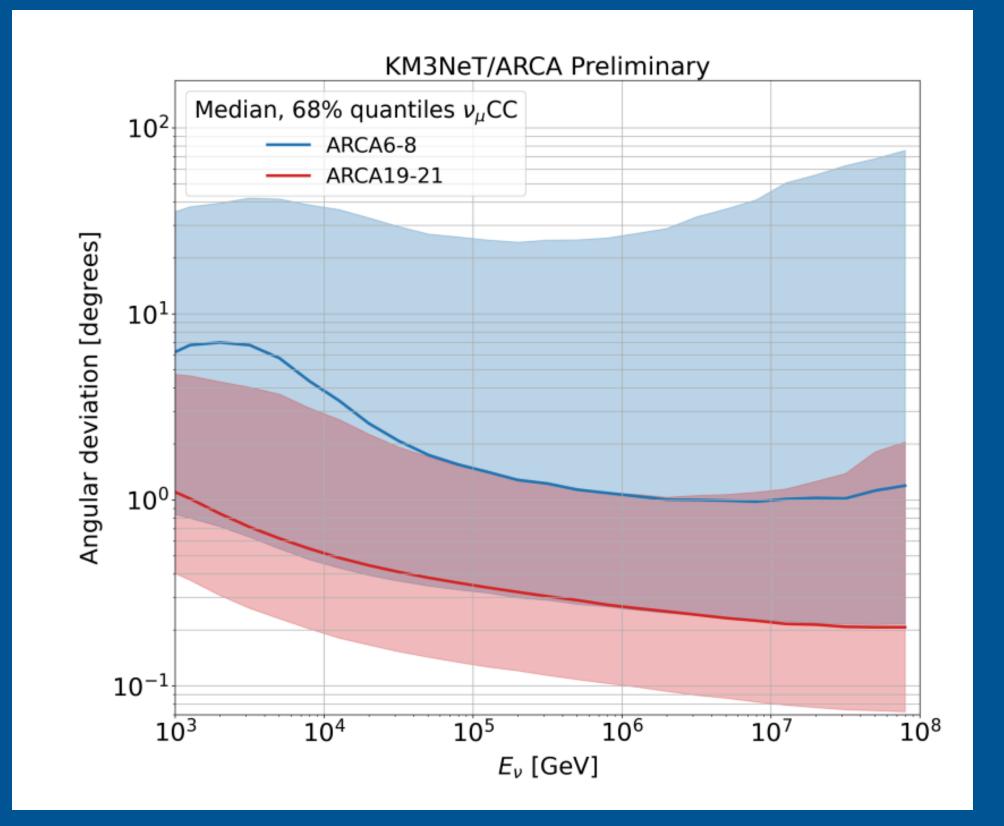


Large improvement in sensitivity is expected in the next year: + ARCA28 from sept 2023 + ARCA47 (?) from November 2024

KM3NeT upper limits are quickly reaching the ANTARES 15yr limits

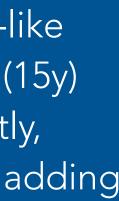
Improvements also in angular resolution

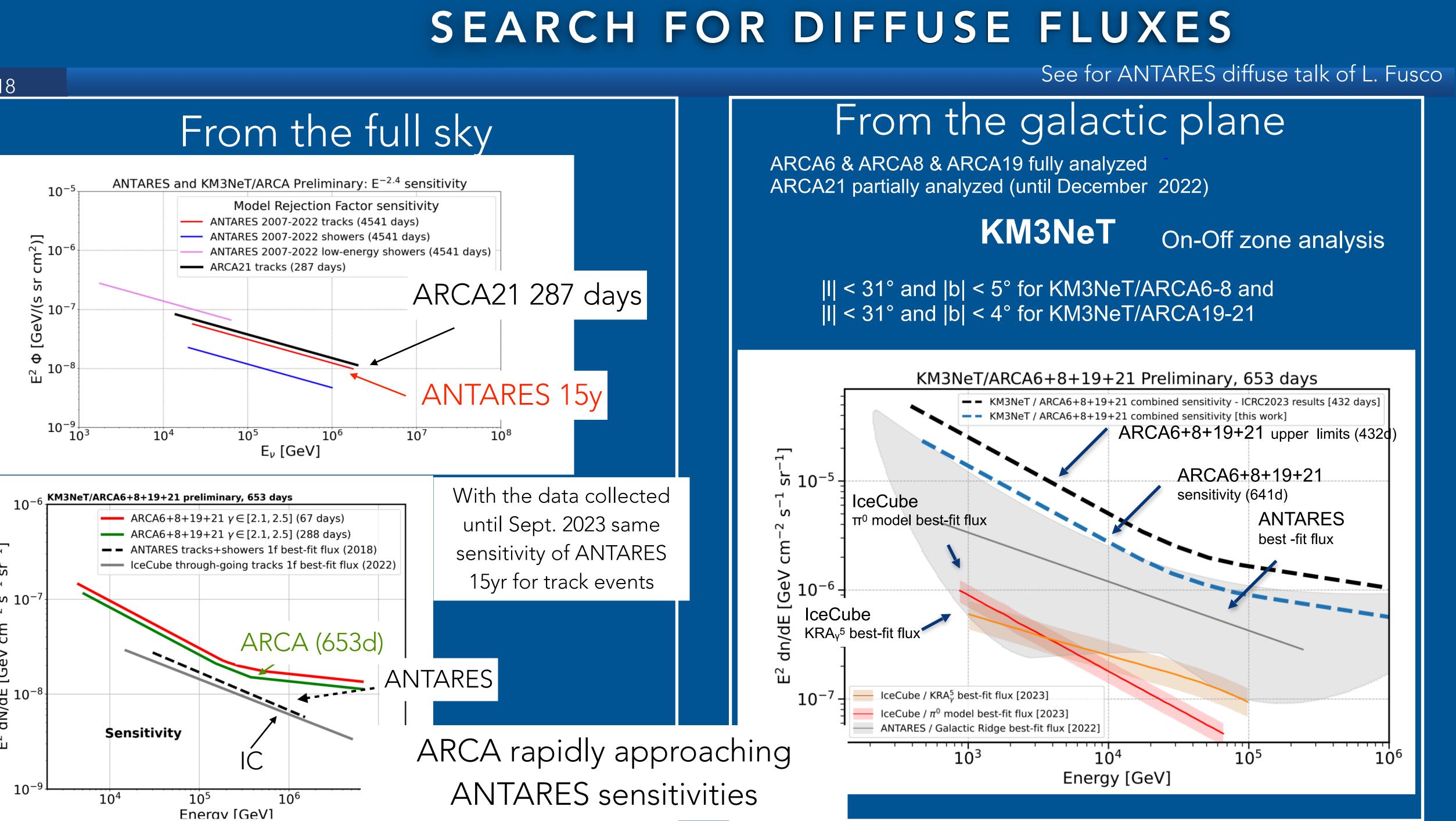
### Angular resolution

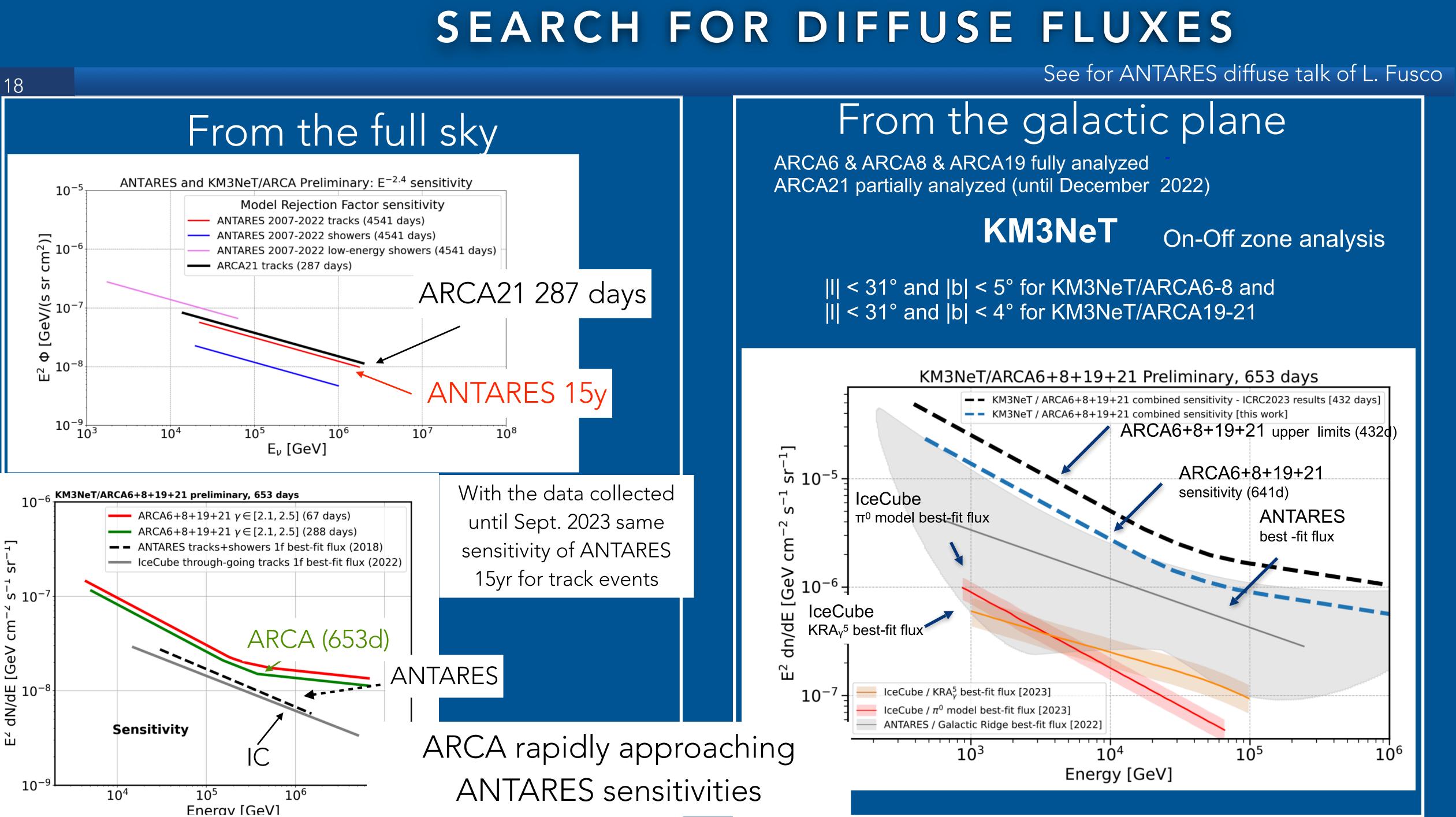


Joint ARCA-ANTARES point-like searches ongoing. ANTARES (15y) contributes most significantly, enhancement by 10% observed adding ARCA data (424d)

See Sandra Zavaterelli talk



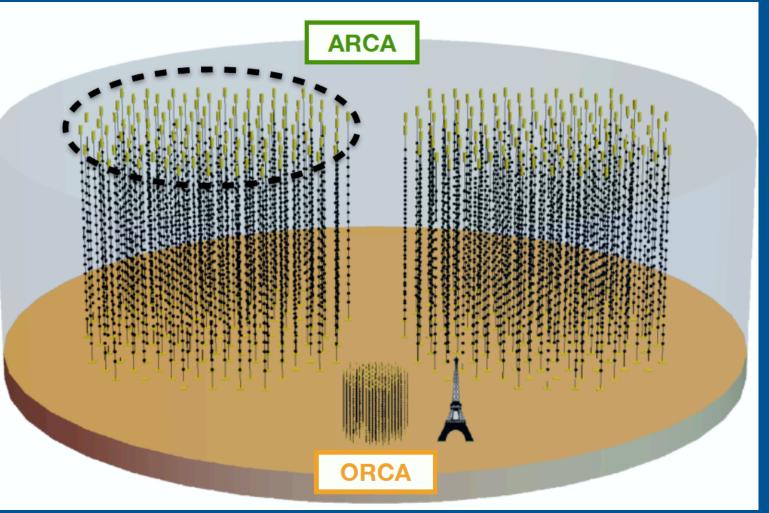




## MULTI-MESSENGER PROGRAM

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A dedicated software is installed at the shore stations for Real-Time Analysis (RTA)



### Receiving alerts

Receive alert from external communities - on-line analysis and follows ups

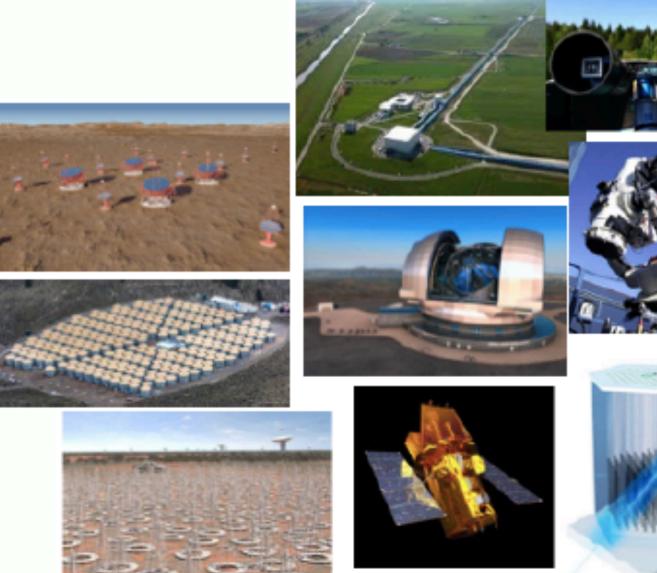
Receiving alert system operative *F* RTA platform already active from November 2022 in ARCA and in ORCA detectors Sending alert system *f* High-energy neutrino alerts will be sent in real-time by end of 2024

#### Talks of M. Mastrodicasa and and V. Cecchini

#### EM/MM external communities











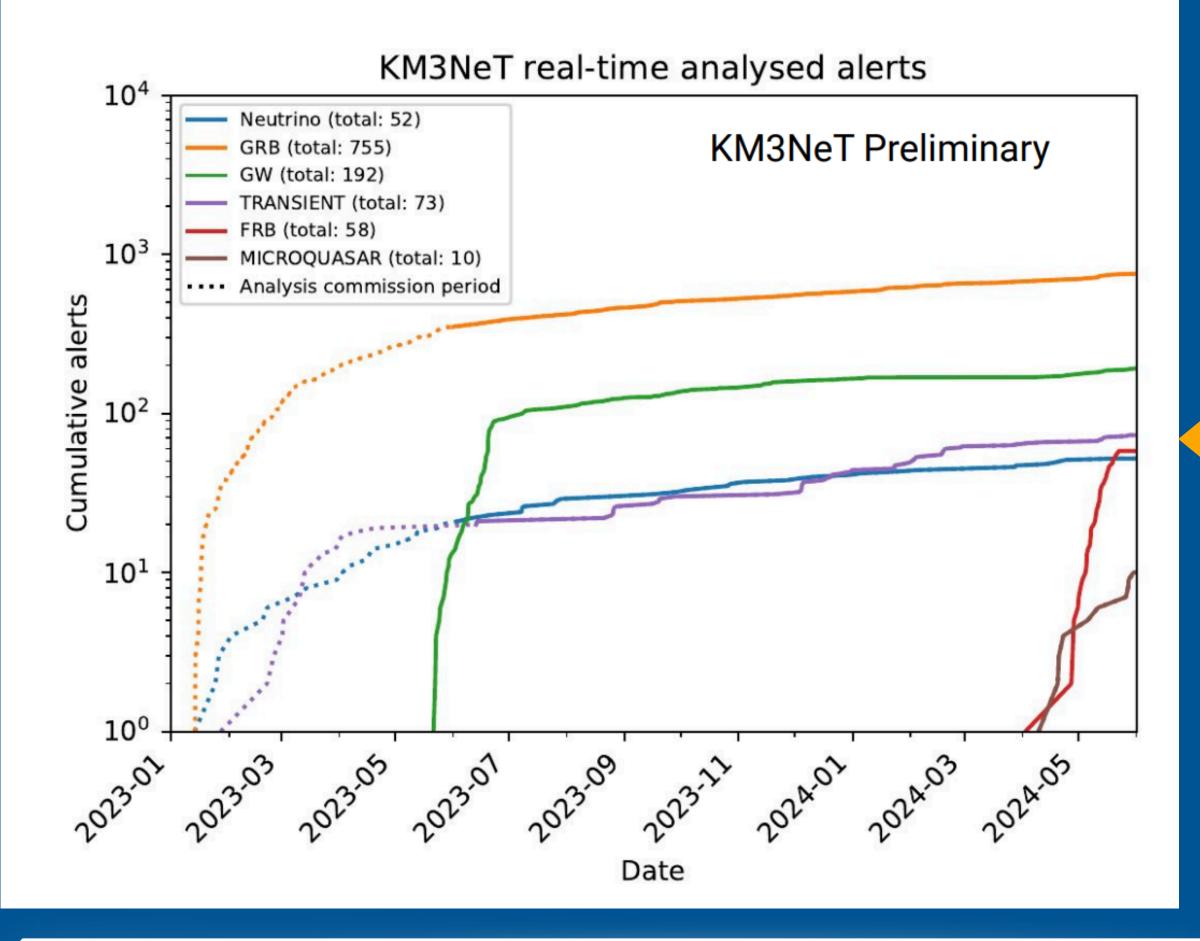






## MULTI-MESSENGER PROGRAM

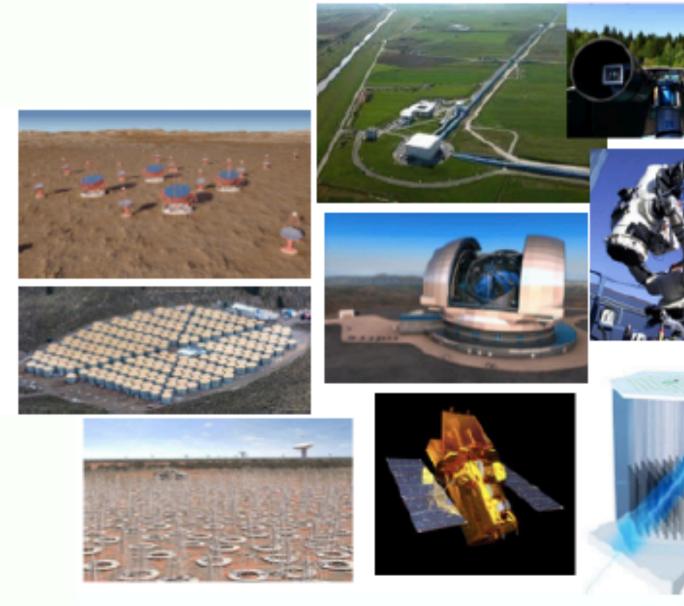
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Several thousands of alert received and analyzed in real time *f* so far no significant excess found in any of the observed alerts

Talks of M. Mastrodicasa and and V. Cecchini

#### EM/MM external communities







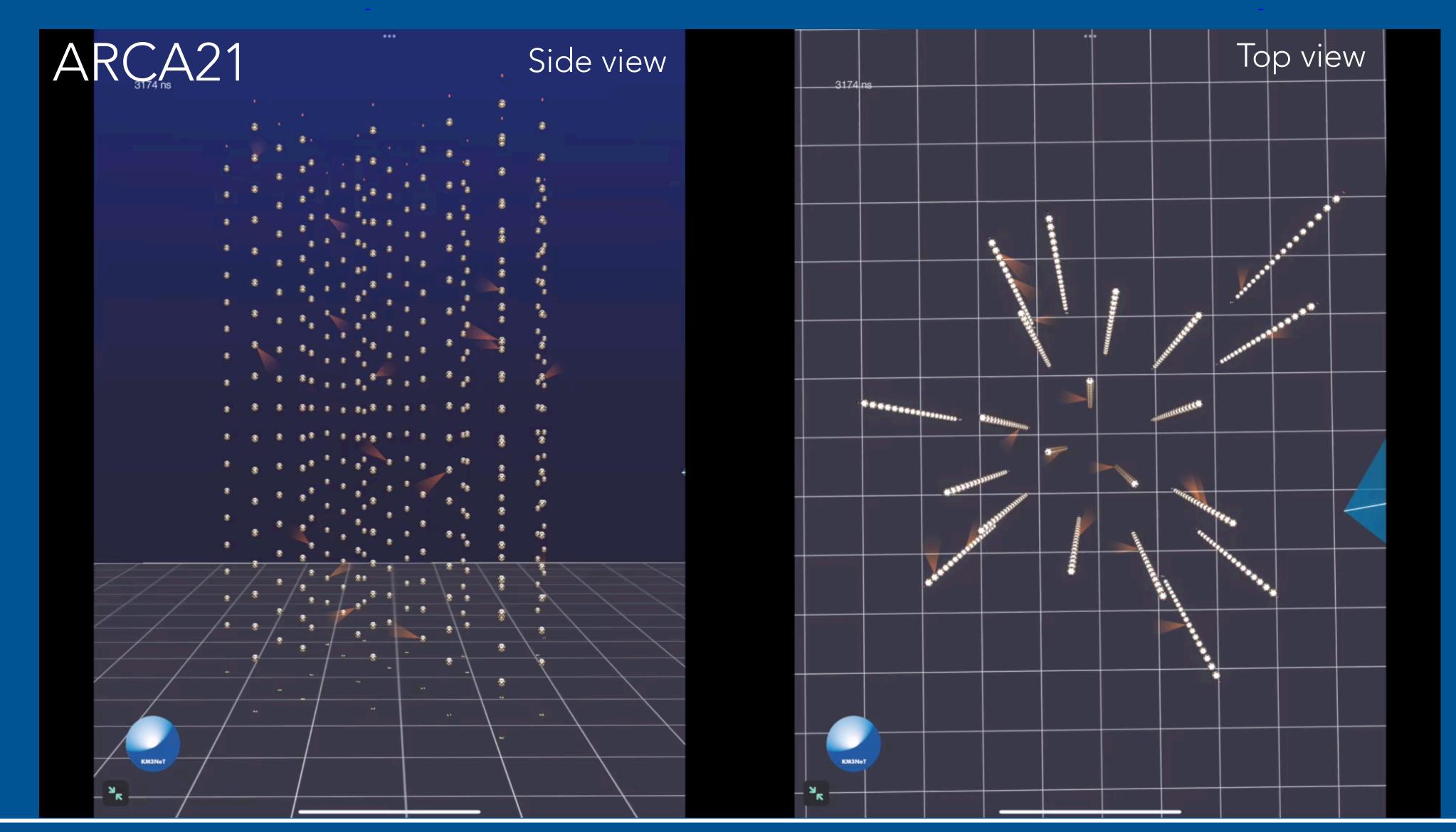




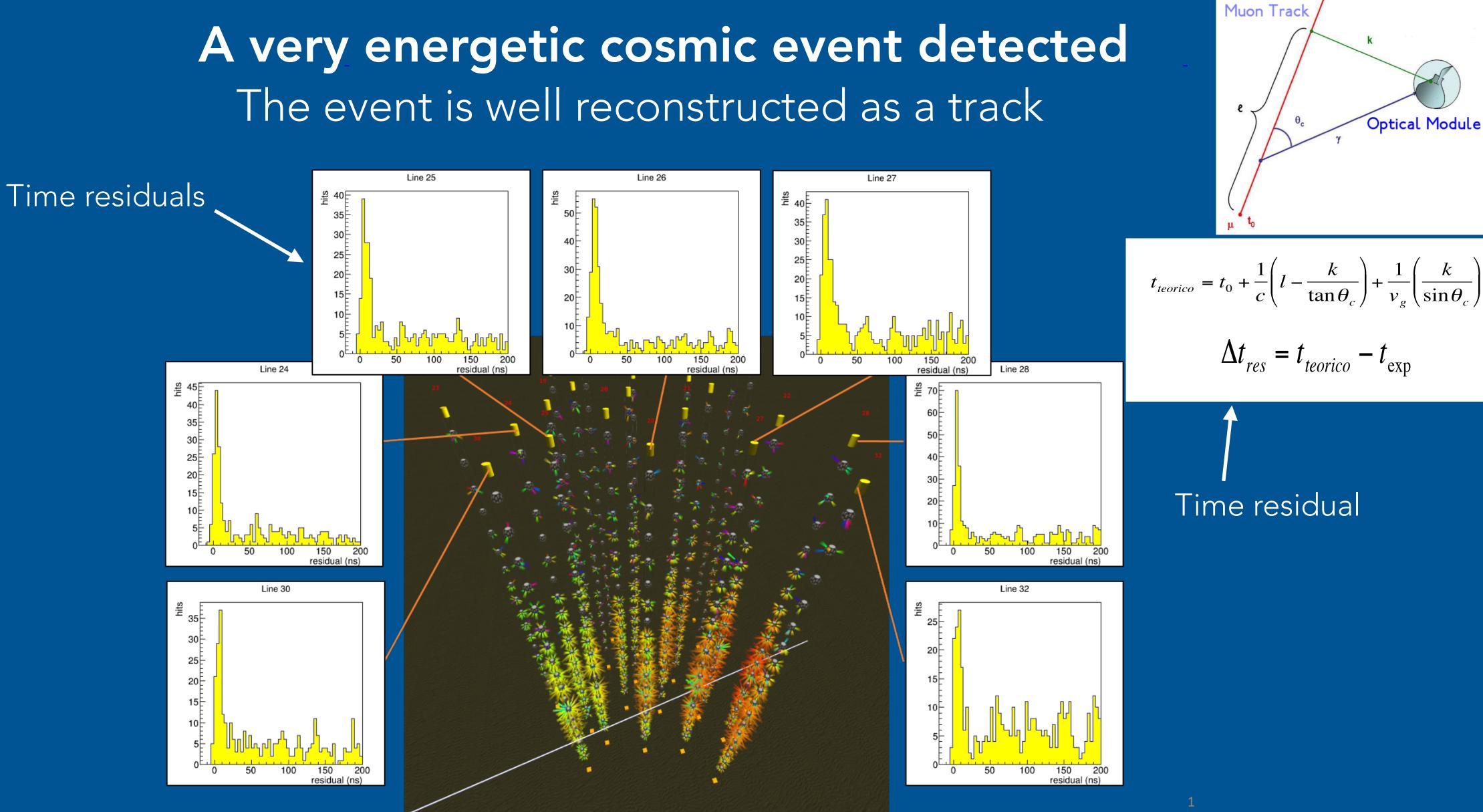


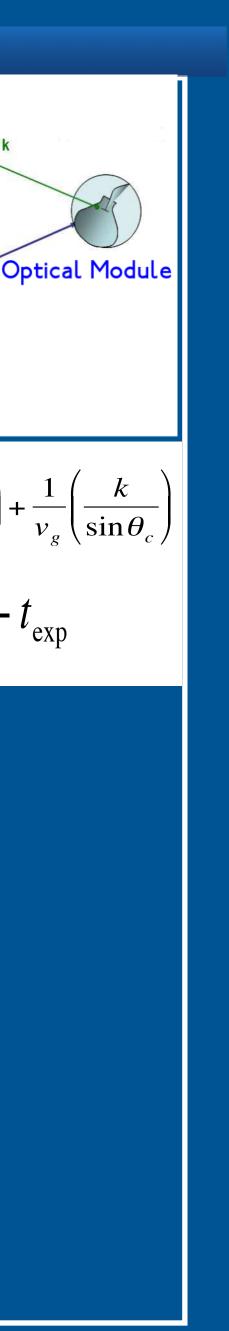
## A very energetic cosmic event detected

Huge amount of light detected *4* 35% of the total number of PMTs were triggered



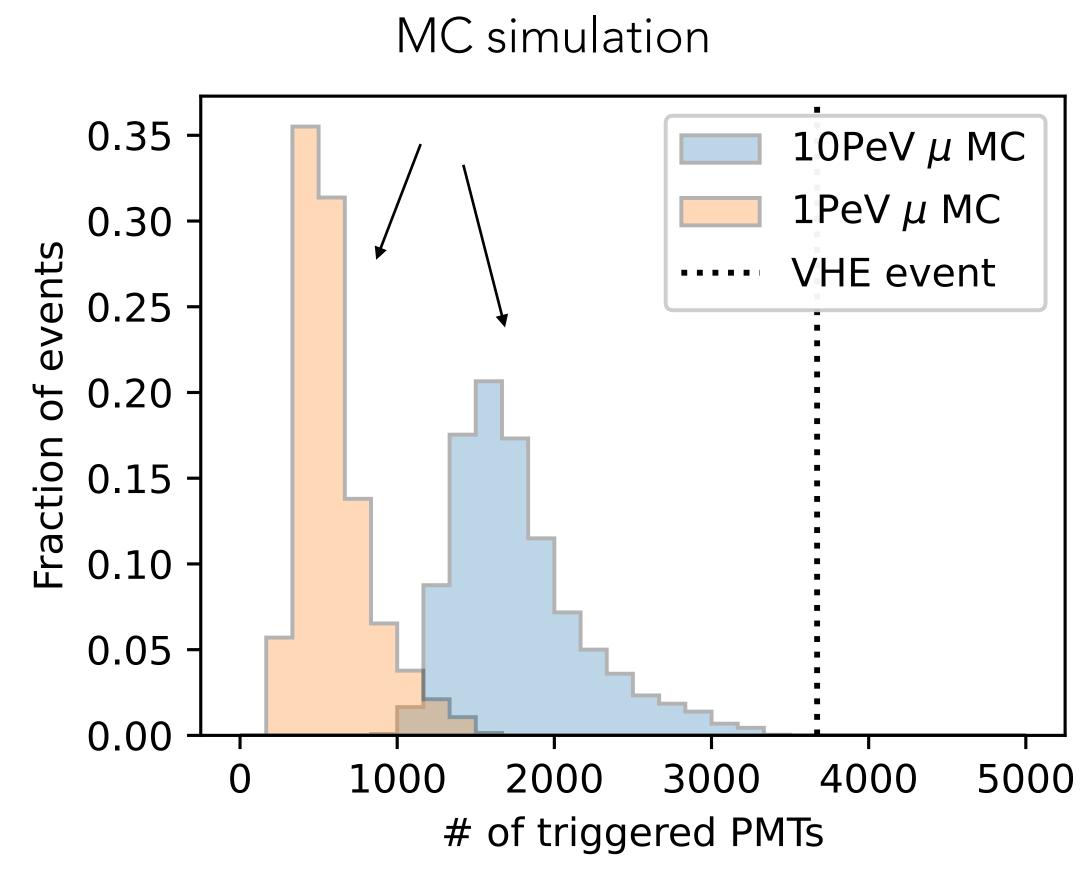


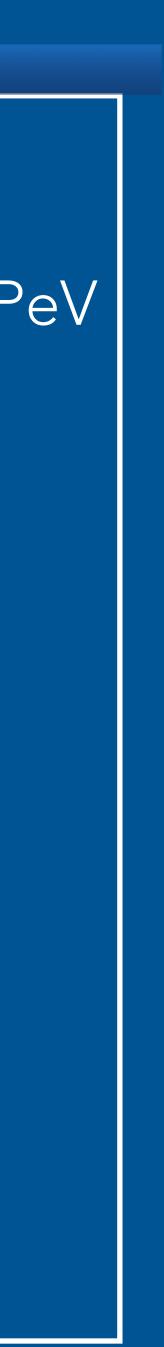




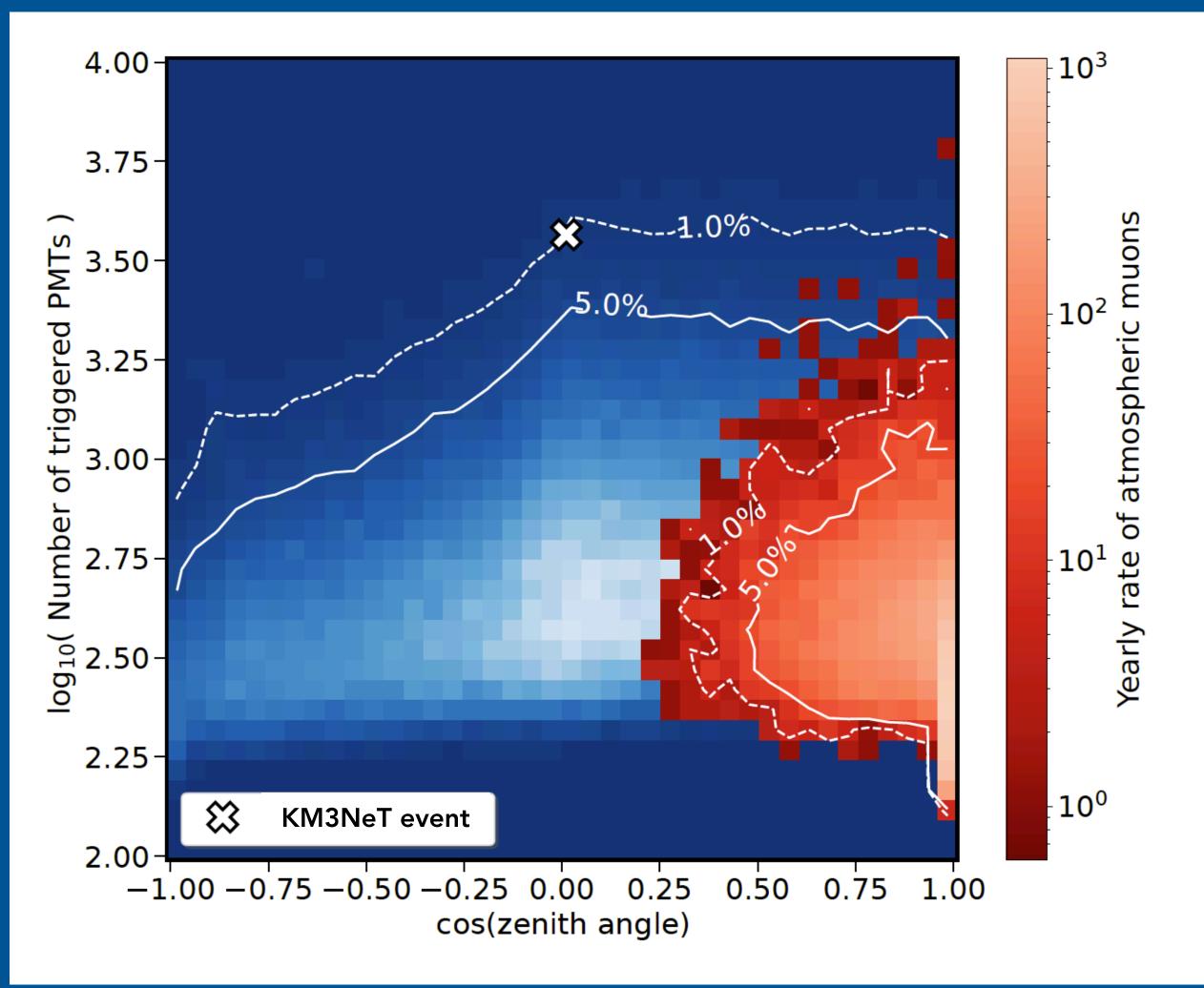
## A very energetic cosmic event detected

### The event is a horizontal event (~0.6° above the horizon) with energy well above 10 PeV





## Is it background?



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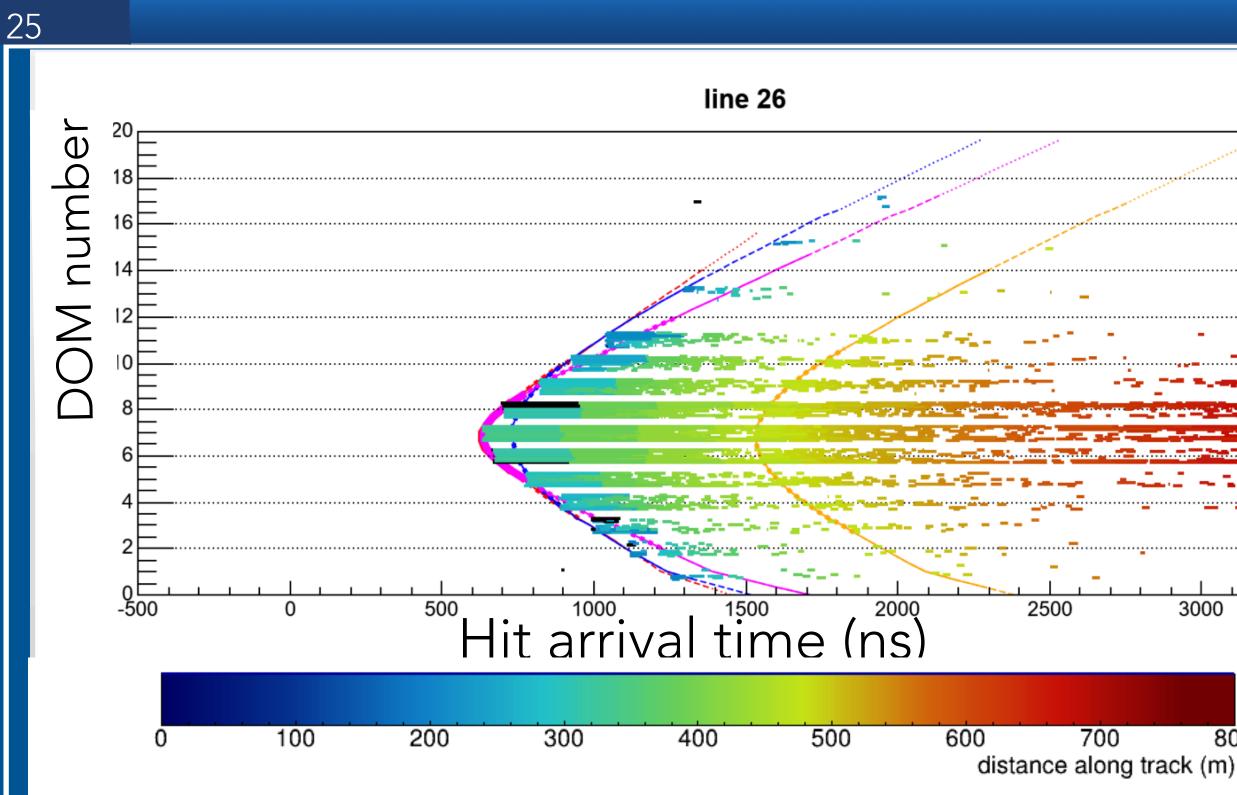
0.012 0.010 uentrinos 0.006 ō 0.0 7early rate 0.002 L0.000

Given the detected energy and direction the expected rate of atmospheric muon reaching the detector is very low

• The amount of material traversed along the event direction is >100km of water equivalent (uncertainties included)

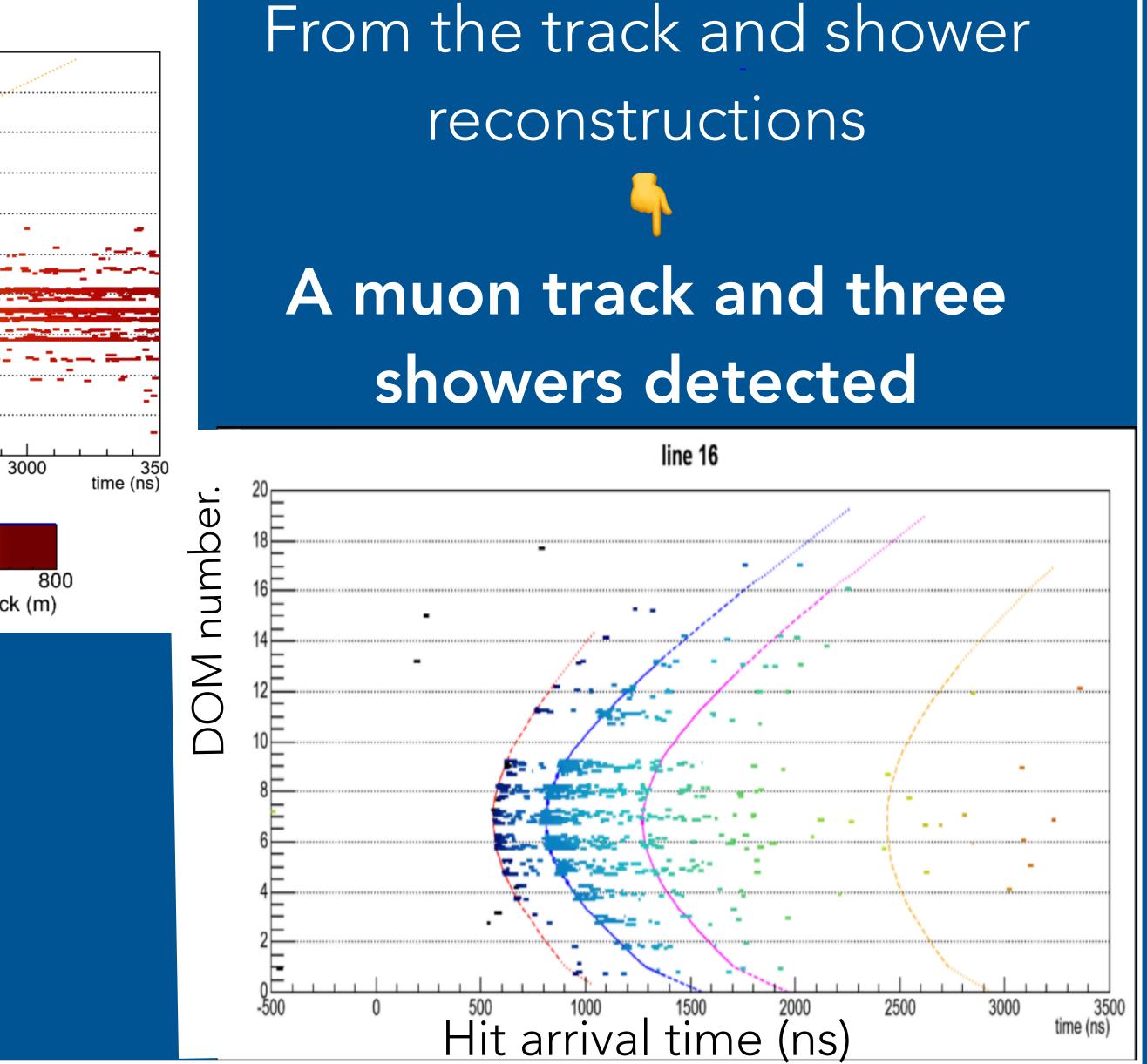
At this energies the expected rate of an atmospheric neutrino (prompt component) is about 10<sup>-5</sup> events per year





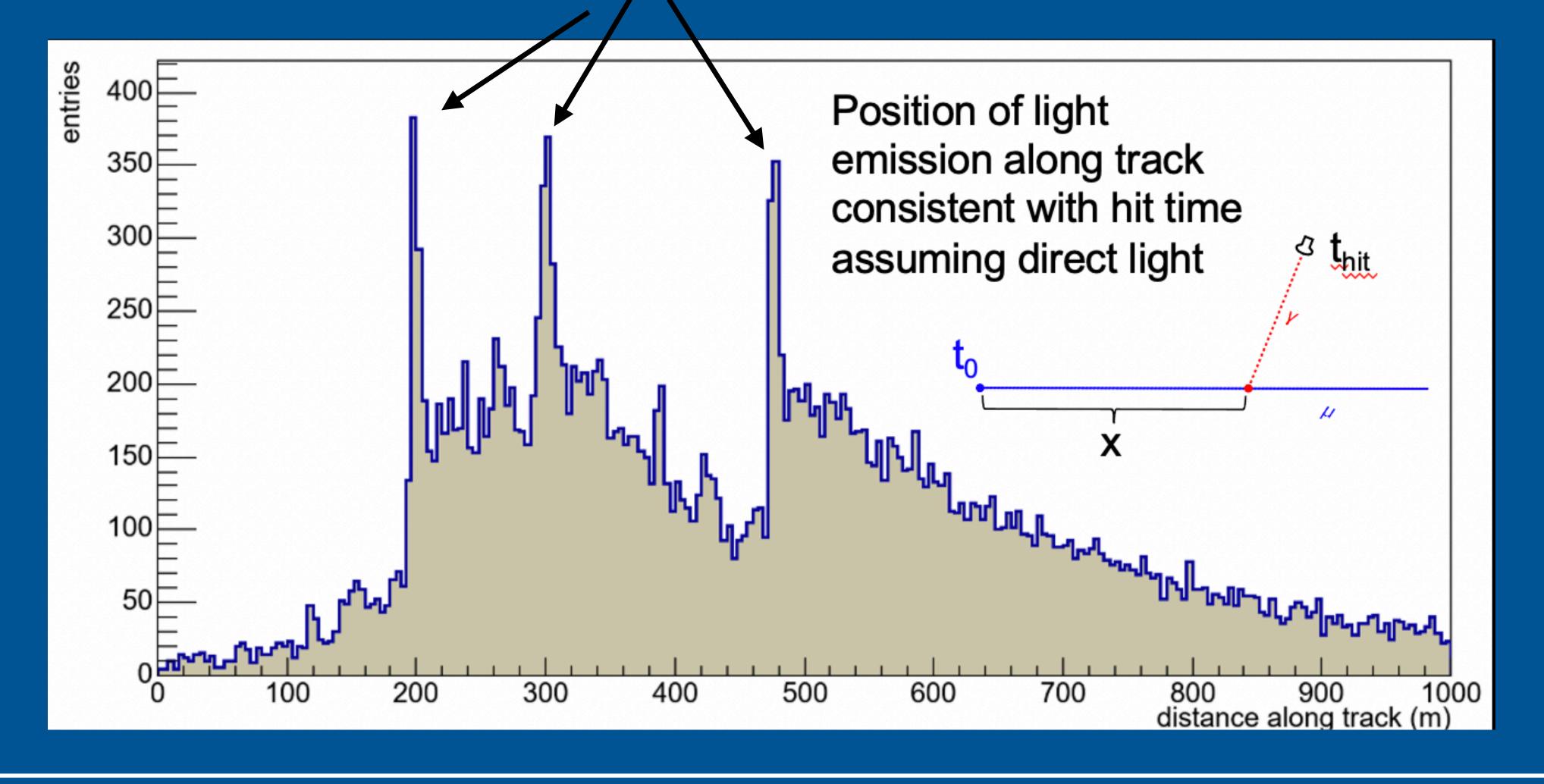
Hit times are fully consistent with photons from Cherenkov emission

## From the track and shower reconstructions A muon track and three showers detected





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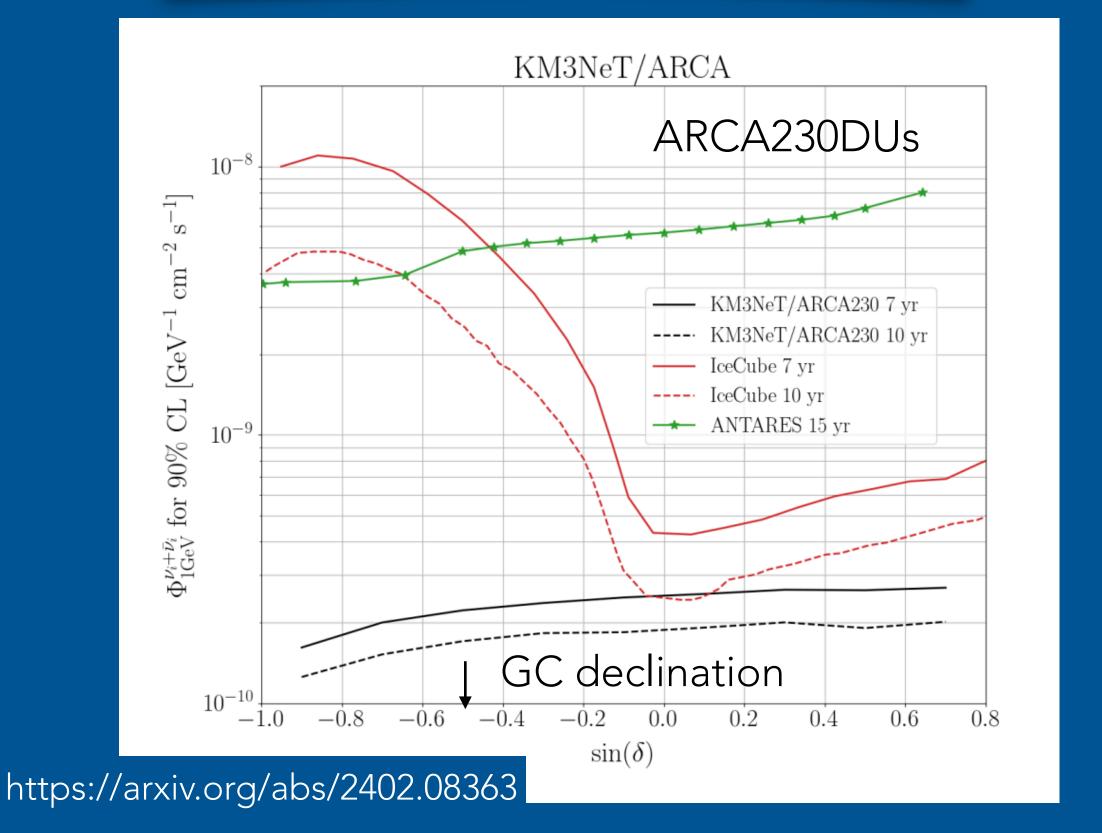


Hit times consistent with the emission from three points along the track 👉 stochastic light emission



### ARCA - Sensitivity for point-like searches

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Big expectations for the southern sky

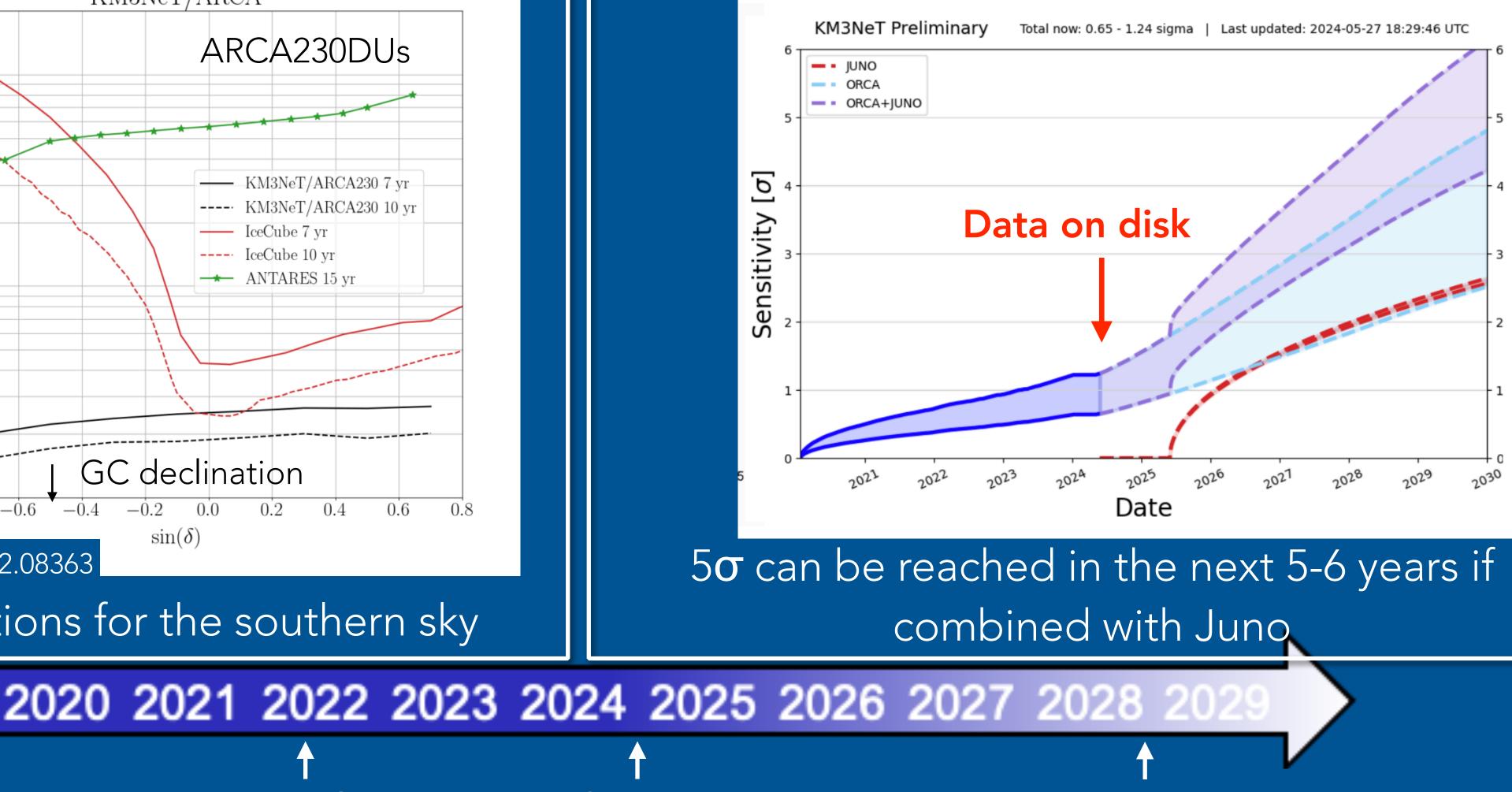
ANTARES decommissioning

ARCA 47DUs ORCA 24 DUs

## **KM3NET PERSPECTIVES**

ORCA - Neutrino mass ordering

#### Predictions based on the current construction plan.



**ARCA & ORCA** completion



## SUMMARY

- ARCA 28 DUs (12% of the full detector) 

  rext month + 19DUs 
  ARCA47 (40% of the full detector) • ORCA 23 DUs (20% of the full detector) 👉 end of the year +2DU 👉 ORCA25 (22% of the full detector)
  - Detectors in data taking from the first strings deployed
  - ARCA sensitivities quickly approaching the ANTARES ones Exciting results expected in a few years especially in the exploration of the southern sky
- Promising results also in the neutrino oscillation 👉 KM3NeT/ORCA in the race for mass hierarchy

- KM3NeT under construction
  - present status:

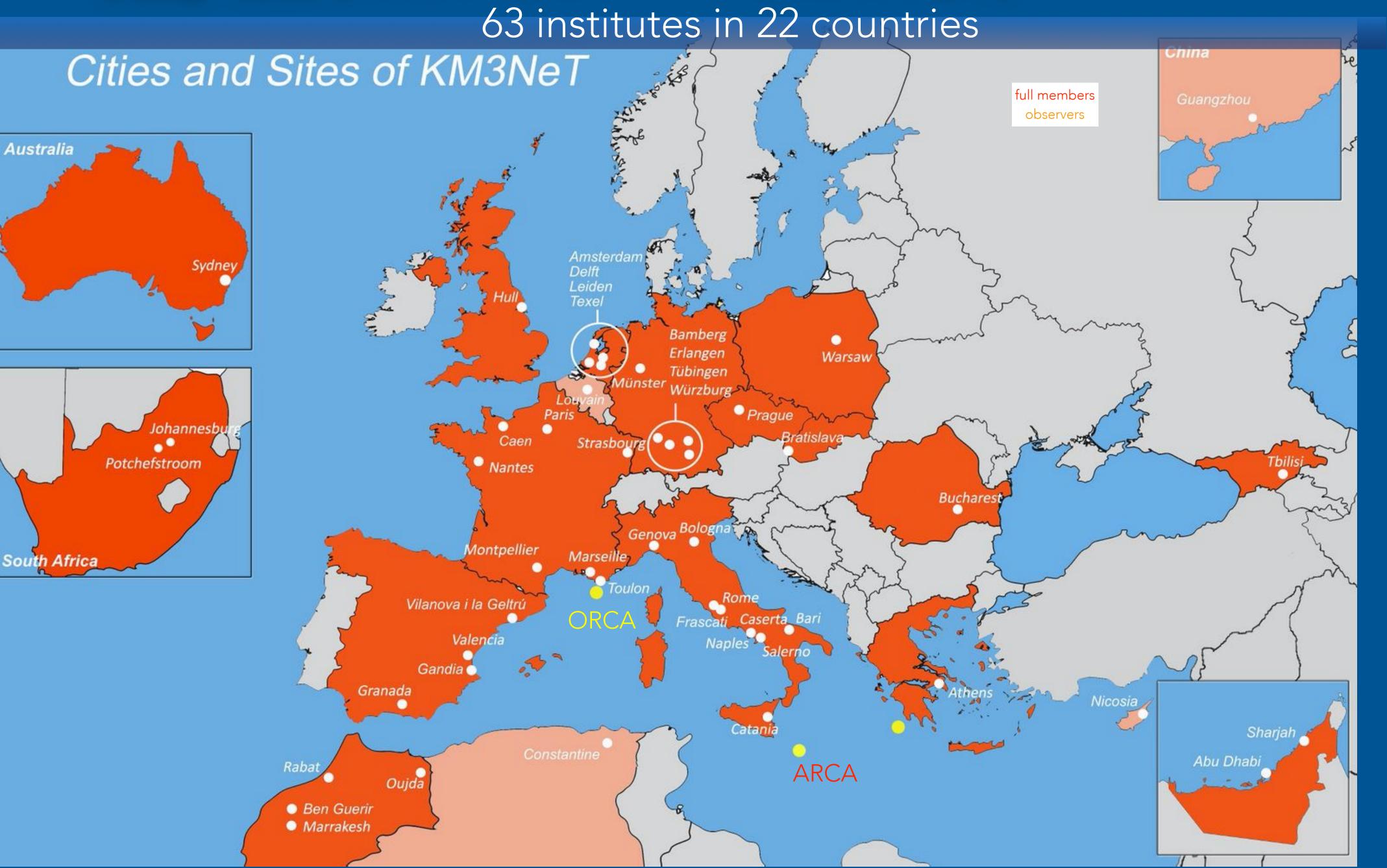
- An exceptional high energy track event detected *—* a horizontal event with energy well above 10 PeV 👉 detailed event description
  - More information will follow soon (paper submitted)



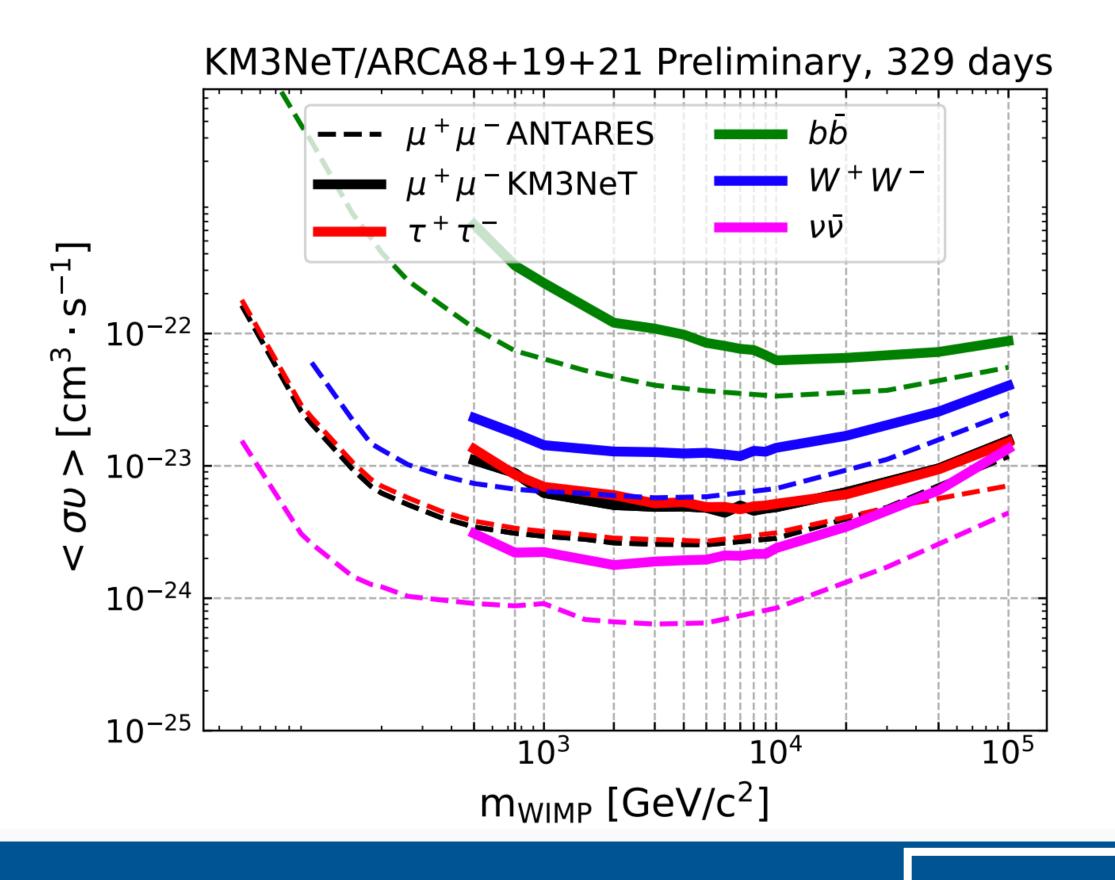


## THE KM3NET COLLABORATION 63 institutes in 22 countries

# + Harvard University (USA) South Africa



## DARK MATTER From the Galactic Center



KM3NeT quickly reaching the ANTARES limits

