The Quiet Sun with Fermi LAT

Elena Orlando

RICAP 2024 – Sep 2024



2) Inverse Compton (IC) Spatially Extended Component

First theory: <u>Orlando & Strong (</u>2006) arXiv:astro-ph/0607563; (2007) Ap&SS, 309, 59

CR e^{\pm} + eV photon $\rightarrow \gamma$



Inverse Compton Modeling Updates

Orlando & Strong (2021) JCAP 04, 004



IC from Luminous Stars & Associations

Original Article Published: 20 April 2007

Gamma rays from halos around stars and the Sun

E. Orlando 🖂 & A. W. Strong

Astrophysics and Space Science 309, 359–363 (2007) Cite this article

accounted in Cygnus OB2 Association: Ackermann et al. (2011) Science 334, 1103



Fermi upper limits: De Menezes, Orlando et al. (2021) MNRAS 507, 680 – single stars

First Detection of the Quiet Sun in Gamma Rays

Orlando & Strong (2008) A&A 480, 847



The two components are detected and distinguished!

Fermi LAT – daily detections

Fermi LAT Coll. ApJ. (2011) 734, 116



FERMI

Higher significance than before

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A New Component from the Quiet Sun from Radio to Gamma Rays: Synchrotron Radiation by Galactic Cosmic-Ray Electrons

Elena Orlando^{1,2}, Vahe' Petrosian², and Andrew Strong³

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from MHz to UV

many orders lower than present data (e.g. LOFAR and ALMA) and of the solar thermal emission

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a few orders lower than upper limits -> promising in future!



Emission from Spiraling Electrons

THE ASTROPHYSICAL JOURNAL, 962:52 (16pp), 2024 February 10

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Yet Another Sunshine Mystery: Unexpected Asymmetry in GeV Emission from the Solar Disk

Bruno Arsioli^{1,2,3,4}⁽¹⁾ and Elena Orlando^{3,4,5}



Analysis

Morphology over time and deviations from isotropic emission as assumed by current models

- 14 years of data
- Solar Disk Photons > 5 GeV in Helioprojection system
- Careful data cleaning
- Optimization of spatial resolution and photon counts



Non-uniformity of the Disk ~2014



Significant deviations from isotropic emission for ~2014

Asymmetry in Energy ~2014



~2014: Sun's Polar Magnetic Field Flip



Data from the Wilcox Solar Observatory (WSO)

~2014: Sun's Polar Magnetic Field Flip



Crucial Role of the Magnetic Field, but still unexplained, it opens up a link among Elena Orlando plasma physics, astronomy, and particle physics

~2014: Sun's Polar Magnetic Field Flip



Crucial Role of the Magnetic Field, but still unexplained, it opens up a link among Elena Orlando plasma physics, astronomy, and particle physics

AAS Press Conference



Current Mysteries of the Gamma-Ray Sun

• Observed emission higher and harder than expected (Orlando & Strong 2008, Abdo et al. 2011, Ng et al. 2016)

- A possibly new equatorial component? (Linden 2018)
 - Unexpected dip in the spectum (Tang et al. 2018)
- Unexplained high energy emission (Albert et al. 2023, HAWC)
- Synchrotron Radiation on the solar B-field as a new component (Orlando et al. 2023)
 - Unexpected asymetry in the disk emission (Arsioli & Orlando 2024)

References to our works

Arsioli & Orlando (2024) ApJ 962, 52 - Unexpected asymmetry of the solar emission during magnetic field reversal

Orlando, Petrosian, Strong (2023) ApJ 943, 173 - A New Component from the Quiet Sun from Radio to Gamma Rays: Synchrotron Radiation by Galactic Cosmic-Ray Electrons

De Menezes, Orlando et al. (2021) MNRAS 507, 680 - Gamma-rays from stars with Fermi LAT

Orlando & Strong (2021) JCAP 04, 004 - Refining model of the solar emission

Abdo et al (2011) ApJ 734, 116 – Observations of the Sun with Fermi LAT data

Orlando & Strong (2008) A&A 480, 847 - First detection of the quiet Sun in gamma rays

Orlando & Strong (2007) Ap&SS 309, 359 - First theory of the extended inverse compton emission from the Sun and stars

Thank you for your attention

Results: Asymmetry



Kolmogorov Smirnov and Rayleigh statistical tests

Synchrotron Emission Modeling



Orlando, Petrosian, Strong 2023

Solar B-field 10¹ **GY11** Patzold 10⁰ B-field Intensity [G 10^{-1} 10⁻² $10^{-3} \begin{bmatrix} B(r) = \begin{cases} 1.0 \ (r/R_{sun})^{-1.9}, & \text{Parker Solar Probe for } 0.1 < r/AU < 1.0. \\ 0.31 \ (r/R_{sun})^{-1.5}, & \text{GY11, for } r/AU < 0.1. \\ 8.4 \ (r/R_{sun})^{-2.6}, & \text{Patzold, for } r/AU < 0.1. \end{cases}$ Blue line: based on Patzold et al. (1987) 10^{-4} Red line: based on Gopalswamy & Yashiro (2011) Also on Parker Solar Probe measurements from Badman et al. (2020, 2021). 10⁻⁵ $10^{\overline{0}}$ 10¹ 10^{2}

Elena Orlando

Distance from the Sun [Rsun]