



# Latest Results from the LZ Dark Matter Experiment

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## LZ in the News

ARTICLE • MYSTERIES OF MATTER

#### LZ Experiment Sets New Record in Search for Dark Matter

LUX-ZEPLIN Experiment Narrows The EDUCATION Search For Elusive Dark Matter By Lauren Biron OUICK READS August 26, 2024 August 26, 2024 | U.S. Department of Energy Scientists at underground lab announce progress in hunt for dark matter 08-26-2024 Unending search for elusive dark matter GHT STAFF - AUGUST 26, 2024 3:48 PM ScienceNews sets a new record NEWS PARTICLE PHYSICS **By Eric Ralls** The possibilities for dark Earth.com staff writer matter have just shrunk — by NEWS | PHYSICS Hunt for dark matter particles bags nothing—again a lot The massive LUX-ZEPLIN detector leaves few hiding places for hypothetical particles called WIMPs The LZ experiment reports no signs of dark matter in their latest search **NewScientist By Emily Conover** 26 AUG 2024 · 6:00 PM ET · BY ADRIAN CHO AUGUST 26, 2024 AT 1:00 PM **Physics** 

#### Another blow for dark matter as biggest hunt yet finds nothing

The hunt for particles of dark matter has been stymied once again, with physicists placing constraints on this mysterious substance that are 5 times tighter than the previous best

# LZ @ Sanford Underground Research Facility



SURF in Lead, South Dakota is the deepest underground lab in the U.S.

Davis Campus

as double-beta d

Ross Cam

- LZ is located on the 4850 level ~1.5 km underground
- ~10<sup>6</sup> reduction in cosmic muon flux
- Primary goal is to detect WIMPs



Ray Davis in the Homestake mine, 1971

#### Dual Phase Xenon Time Projection Chamber (TPC)



- Signal vs. background discrimination
  - Charge (S2)/ light (S1) ratio is different between electron recoil (ER) and nuclear recoil (NR)



- Electrons and gammas interact with atomic electrons, produce ER
- WIMPs (and neutrons) interact with Xe nuclei, produce NR

#### Grid Voltage Changes since LZ's First WIMP Search

- Anode voltage lowered to reduce spurious light and charge emission
- Cathode voltage lowered to mitigate light emission from the Skin detector observed after the first WIMP search
- ER/NR discrimination is not impacted

	anode (A) gate (G)	Run	C/G/A Voltage [kV]	Drift Field [V/cm]	Livetime [d]
		WS2022*	-32/-4/+4	193	60
	cathode (C) bottom	WS2024**	-18/-4/+3.5	97	220
Bottom PMTs		*Dec. 2021	– May 2022		

\*\*March 2023 - March 2024

## Calibration Data in WS2024

- Dark blue points: Tritium beta data (ER)\* (continuum betas up to 18.6 keV)
- Orange points: DD neutron data (NR)\* (2.45 MeV neutrons produced through Deuterium-Deuterium fusion)
  - ER/NR discrimination: 0.2% ER leakage past the median of the NR population
    - Light gain g1= 0.112 ± 0.002 phd/photon
    - Charge gain  $g_2 = 34.0 \pm 0.9$  phd/electron



\*Details about calibration source deployment: J. Aalbers et al 2024 JINST 19 P08027

# Bias Mitigation in WS2024

- Bias mitigation via 'salting': fake WIMP signals injected randomly during data taking (rate bounded by WS2022 cross section)
- Salt events are revealed only after all analysis inputs are finalized



#### LXe Flow Control for Background Reduction in WS2024



- Data collected in two flow states through circulation and cooling systems control
- High Mixing (40.9 live days)
  - Turbulent flow
  - Uniform distribution of injected calibration sources
  - Low Mixing (179.1 live days)
    - Laminar-like flow
- Liquid flow mapping using <sup>222</sup>Rn-<sup>218</sup>Po coincidences

## <sup>214</sup>Pb tagging in the Low-mixing State in WS2024

- Liquid flow map enables tagging of <sup>214</sup>Pb background in the low-mixing state
- Both tagged and untagged populations are used in the final analysis
  - <sup>214</sup>Pb reduced to 1.8  $\pm$  0.3 µBq/kg in the untagged population (3.9  $\pm$  0.6 µBq/kg in the total exposure)



#### <sup>214</sup>Pb tagging in the Low-mixing State in WS2024 Liquid flow map enables tagging of <sup>214</sup>Pb Neutral background in the low-mixing state /olume Both tagged and untagged populations are used in the final analysis <sup>214</sup>Pb reduced to 1.8 $\pm$ 0.3 $\mu$ Bq/kg in the untagged t=5 mi 214Pb+ population (3.9 $\pm$ 0.6 $\mu$ Bq/kg in the total exposure) <sup>214</sup>Pb fraction Exposure (low-mixing state) t=15 min Charged Tagged 60±4 % 0.3 tonne-yr t=25 min α <sup>222</sup>Rn <sup>214</sup>Bi 218Po Untagged 40±4 % 1.8 tonne-yr <sup>214</sup>Pb 3.8 d 26.8 mir 19.9 min 3.1 min Dominant ER background



## Electron Capture Decay Backgrounds

Isotope	Decay mode	Primary energy in WIMP search region	Half-life
<sup>127,125</sup> Xe (neutron source activation)	Single EC	5.2 keV (L-shell)	36.3d ( <sup>127</sup> Xe) 16.9h ( <sup>125</sup> Xe)
<sup>124</sup> Xe (0.095% nat. abundance)	Double EC	5.98 keV (LM shell) & 10 keV (LL shell)	$(1.09 \pm 0.14_{stat} \pm 0.05_{sys}) \times 10^{22}$ year (LZ preliminary *)

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 Single L-shell EC charge yield is suppressed compared to β's of the same energy

- Measured in XELDA\*
- Preliminary result from WS2024:  $Q_{EC}/Q_{\beta} = 0.86 \pm 0.01$

\*Temples et al, Phys. Rev. D 104, 112001 (2021)

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 <sup>124</sup>Xe LL captures is expected to display further charge yield suppression compared to single-L capture

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# Backgrounds in WS2024

ER backgrounds:

![](_page_17_Figure_2.jpeg)

#### NR backgrounds

- Detector material neutrons
- Solar + Atmospheric neutrinos
- Accidental coincidences background

## WS2024 WIMP Search Data (salted)

- 220 live days x 5.5 t = 3.3 tonne-yr
- 7 events pass all analysis cuts out of 8 total injected (consistent with evaluated efficiency)
- 1220 events after unsalting

![](_page_18_Figure_4.jpeg)

#### WS2024 Final WIMP Search Data (salt removed)

All data (3.3 tonne-yr):

Radon-tagged set (0.3 tonne-yr):

![](_page_19_Figure_3.jpeg)

Radon tagged (<sup>214</sup>Pb rich) sample does not contain leakage from <sup>124</sup>Xe!

#### WS2024-only Limit on Spin-independent WIMP-nucleon Cross Section

- No evidence of WIMPs at any mass
- WS2024 min cross section:  $\sigma_{SI}$ = 2.3 x 10<sup>-48</sup> cm<sup>2</sup> @ 43 GeV/c<sup>2</sup>
- Upper limit (solid black) is power constrained @  $-1\sigma$  sensitivity band

![](_page_20_Figure_4.jpeg)

#### WS2022+WS2024 Combined Limit Results

- No evidence of WIMPs at any mass
- Combined min cross section:  $\sigma_{SI}$  = 2.2 x 10<sup>-48</sup> cm<sup>2</sup> @ 43 GeV/c<sup>2</sup>
- Upper limit (solid black) is power constrained @  $-1\sigma$  sensitivity band

![](_page_21_Figure_4.jpeg)

## Summary and Outlook

- LZ has set new record on the WIMP-nucleon cross section limit with a total exposure of 4.2 tonne-year
  - More than x4 improvement compared to WS2022
  - First use of flow map-based <sup>214</sup>Pb tagging technique leads to 60% reduction of primary ER background
  - First observation of charge-suppressed <sup>124</sup>Xe double electron capture
  - LZ will continue taking data until 2028 (1000 live days) with 'salt' events injected for bias mitigation
- Future physics searches: <sup>8</sup>B CEvNS, low-mass WIMPs, neutrinoless double beta decay, supernova neutrinos, etc.

## Thank you!

![](_page_23_Picture_1.jpeg)

@lzdarkmatter

https://lz.lbl.gov

![](_page_23_Picture_4.jpeg)

#### 250 scientists, engineers, & technical staff 38 institutions

**Black Hills State University Brookhaven National Laboratory Brown University** Center for Underground Physics Edinburah University Fermi National Accelerator Lab. Imperial College London King's College London Lawrence Berkeley National Lab. Lawrence Livermore National Lab. LIP Coimbra Northwestern University Pennsylvania State University Royal Holloway University of London **SLAC National Accelerator Lab.** South Dakota School of Mines & Tech South Dakota Science & Technology Authorit STFC Rutherford Appleton Lab. Texas A&M University University of Albany, SUNY University of Alabama **University of Bristol** University College London University of California Berkeley University of California Davis University of California Los Angeles University of California Santa Barbara **University of Liverpool University of Maryland** University of Massachusetts, Amherst University of Michigan University of Oxford **University of Rochester University of Sheffield** University of Sydney University of Texas at Austin University of Wisconsin, Madison University of Zürich

![](_page_23_Picture_7.jpeg)

Science and Technology **Facilities** Council

Swiss National **Science Foundation** 

![](_page_23_Picture_10.jpeg)

![](_page_23_Picture_11.jpeg)

US

Europe

b Institute for Basic Science

Asia

Oceania

# **Backup Slides**

## WS2024-only Fit Results

Profile likelihood analysis performed in S1 vs. logS2 space

 Best background-only fit projected onto 1D reconstructed energy yields p-value of 0.28

Source	Pre-fit Constraint	Fit Result	• Data — Other FR — Accidentals
$^{214}$ Pb $\beta$ s	$743\pm88$	$733\pm34$	$10^{3} \qquad \text{Total Bkg} \qquad \frac{124}{124} \text{Xe DEC} \qquad \frac{8}{8} \text{B}$
$^{85}$ Kr + $^{39}$ Ar $\beta$ s + det. $\gamma$ s	$162 \pm 22$	$161\pm21$	$10 = \frac{214}{\text{Pb}} \beta_{\text{s}} = \frac{127}{\text{Xe}} + \frac{125}{\text{Xe}} \text{EC}$
Solar $\nu$ ER	$102\pm 6$	$102\pm 6$	
$^{212}\text{Pb} + ^{218}\text{Po} \ \beta \text{s}$	$62.7\pm7.5$	$63.7\pm7.4$	
Tritium+ <sup>14</sup> C $\beta$ s	$58.3\pm3.3$	$59.7\pm3.3$	
$^{136}$ Xe $2 uetaeta$	$55.6\pm8.3$	$55.8\pm8.2$	
$^{124}$ Xe DEC	$19.4\pm3.9$	$21.4\pm3.6$	
$^{127}$ Xe + $^{125}$ Xe EC	$3.2\pm0.6$	$2.7\pm0.6$	
Accidental coincidences	$2.8\pm0.6$	$2.6\pm0.6$	
Atm. $\nu$ NR	$0.12\pm0.02$	$0.12\pm0.02$	
$^8\mathrm{B}{+}hep~ u~\mathrm{NR}$	$0.06\pm0.01$	$0.06\pm0.01$	$10^{-1}$
Detector neutrons	_	$0.0^{+0.2}$	
$40 \text{ GeV}/c^2 \text{ WIMP}$	_	$0.0^{+0.6}$	
Total	$1210 \pm 91$	$1203 \pm 42$	0 2 4 6 8 10 12 14 16 18
			Reconstructed Energy [keV <sub>ee</sub> ]

#### Accidental Coincidence Background

![](_page_26_Figure_1.jpeg)

## WS2024 Data Selection

![](_page_27_Figure_1.jpeg)

## Likelihood breakdown

![](_page_28_Figure_1.jpeg)

WS2024 Skin/OD-tagged events provide direct constraint of neutron background rate (neutron tagging efficiency: 92 ± 1%)

## Statistical Analysis

#### Total log likelihood :

 $\mathcal{L}_{\text{Combined}} = \mathcal{L}_{\text{WS2022}} \times \mathcal{L}_{\text{High mix}} \times \mathcal{L}_{\text{Rn veto inactive}} \times \mathcal{L}_{\text{Rn veto inactive}} \times \mathcal{L}_{\text{Rn tagged}} \times \mathcal{L}_{\text{Not Rn tagged}} \times \mathcal{L}_{\text{Skin+OD tagged}}$ 

models+data from 1st LZ result [PRL 131, 041002 (2023)]

events in high mixing circulation state, contains residual ER calibration events

events in times when Rn-Po flow mapping not reliable (circ. stoppages, etc)

events in Rn veto periods/regions - rich in 214Pb!

complement of above - depleted in <sup>214</sup>Pb & rich in signal

events w/coincident activity in Skin & OD vetoes provides direct constraint on neutron background rate via 92 ± 4% neutron efficiency

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#### WS2024 Event Position distribution

#### LZ Preliminary

![](_page_31_Figure_2.jpeg)

# Livetime removal for data quality

LZ Preliminary

![](_page_32_Figure_2.jpeg)

![](_page_33_Figure_0.jpeg)

#### WS2024 Background Events Pie Chart

![](_page_34_Figure_1.jpeg)

## WS2024 2D Goodness of Fit

![](_page_35_Figure_1.jpeg)

# LZ Skin and Outer Detector Vetoes

![](_page_36_Picture_1.jpeg)

- Delayed veto cut extends to 600 µs w/ 200 & 300 keV OD & skin thresholds to include ncapture on Gd & H
  - Capture on Gd gives 8 MeV gamma
  - Capture on H gives 2.2 MeV gamma
- Measured tagging efficiency for AmLi neutrons: 89 ± 3%
- Predicted tagging efficiency from tuned simulation of background (SF & (α,n)) neutrons: 92 ± 1%
- Accidental tag rate of 3%

# WS2024 Accidental Coincidence: model & unphysical drift sideband comparisons

![](_page_37_Figure_1.jpeg)

## WS2024 Accidental Coincidence Background Impact on Final Limit

- 1. Remove accidental rate constraint: best fit drops  $2.6 \rightarrow 1.4$
- 2. Remove constraint & outlier event: best fit drops  $1.4 \rightarrow 0$ Outlier event holds model up, over subtracting in the WIMP region
- 3. Adding fake events props limit back up
  - $\rightarrow$  under-fluctuation of accidental events in the WIMP region

![](_page_38_Figure_5.jpeg)

#### WS2022+WS2024 Combined Spindependent WIMP-nucleon Cross Section

![](_page_39_Figure_1.jpeg)

- Grey bands : theoretical uncertainties on SD form factors
- Solid black line: power constrained limits

#### WS2022 WIMP Search Result

![](_page_40_Figure_1.jpeg)

## ER leakage vs Drift Field

![](_page_41_Figure_1.jpeg)

## <sup>124</sup>Xe LL-shell Compared to Dark Matter Spectra

- WIMP spectra normalized to LZ's 4.2 tonne-yr median  $3\sigma$  discovery potential:
  - 9 evts @ 40 GeV
  - 11 evts @ 1000 GeV

![](_page_42_Figure_4.jpeg)

## WS2024 Exposure

 Bias mitigation ("salting") began July 3<sup>rd</sup>, 2023; circulation state change July 12th

![](_page_43_Figure_2.jpeg)

# Detector stability monitored by <sup>131m</sup>Xe

![](_page_44_Figure_1.jpeg)

## WS2024 Electron Lifetime

![](_page_45_Figure_1.jpeg)