

# Science with Einstein Telescope

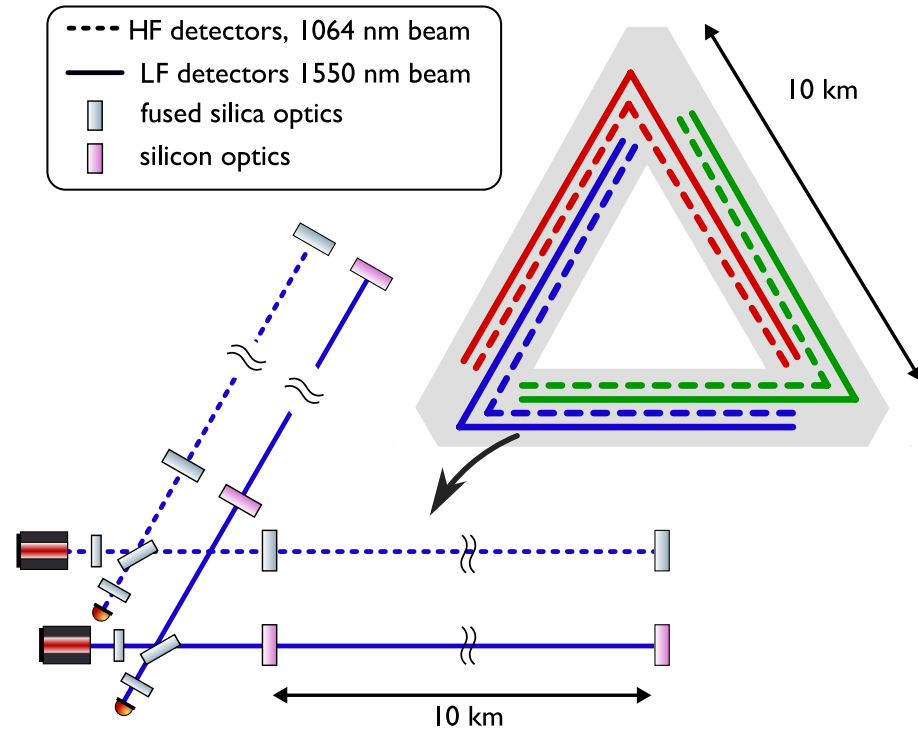
The image is a composite illustration. The top portion shows a cosmic scene with a galaxy, a black hole with an accretion disk, and other celestial bodies. The middle portion shows a landscape with a purple dashed line forming a path that represents a gravitational wave signal. The bottom portion shows a 3D cutaway of the Einstein Telescope detector structure, which is a large underground facility with multiple detector arms and components.

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University of Warsaw

# What is Einstein Telescope?

- 3<sup>rd</sup> generation gravitational wave observatory
- To be built underground
- Triangular configuration\*
- 6 interferometers
- LF and HF

\*) 2L also considered

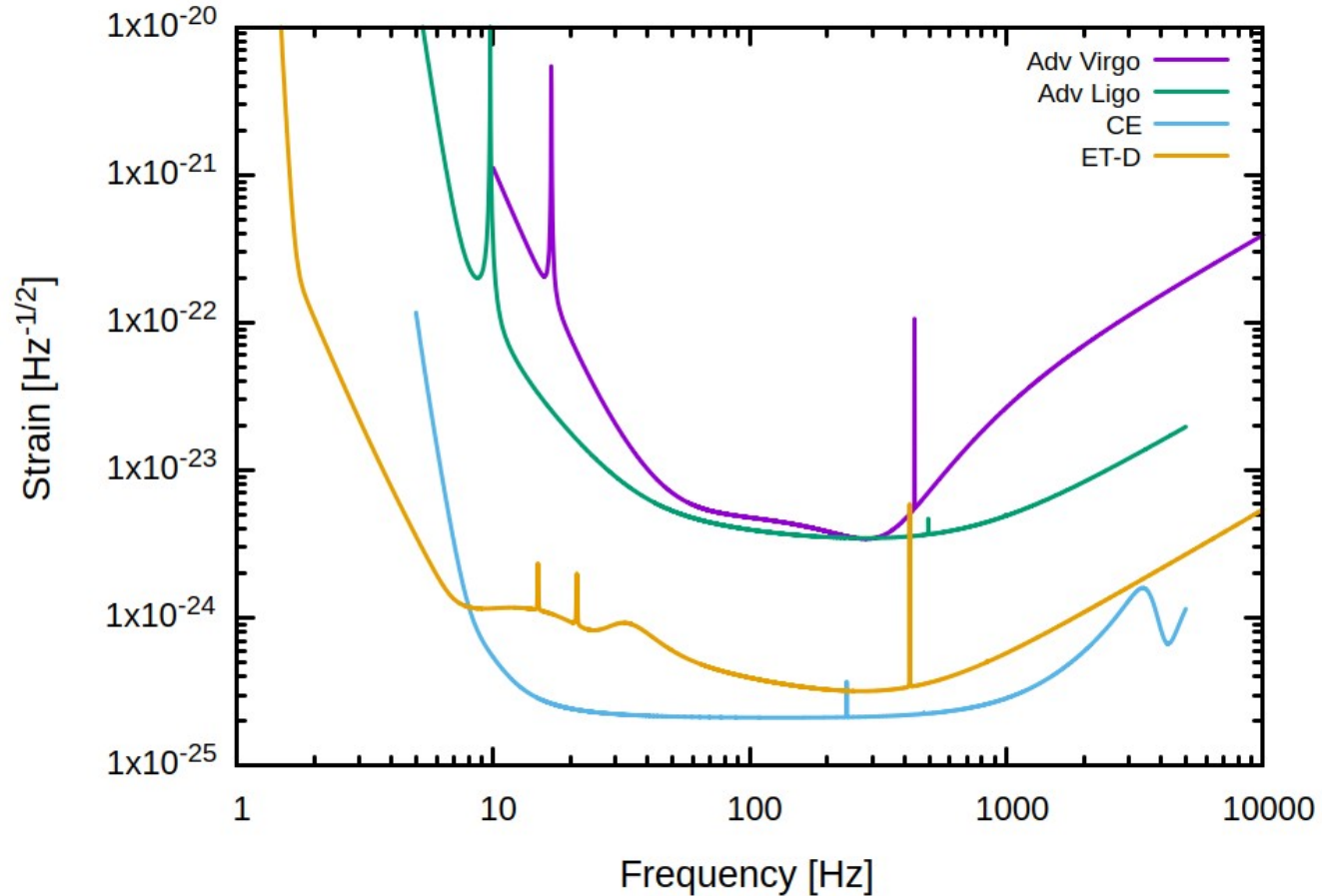


# Technology needed

- Cryogenics
- High power laser
- Heavy mirrors
- New mirror technologies
- Squeezed light
- 



# Planned sensitivity



EMR site



Sos  
Ennatos site

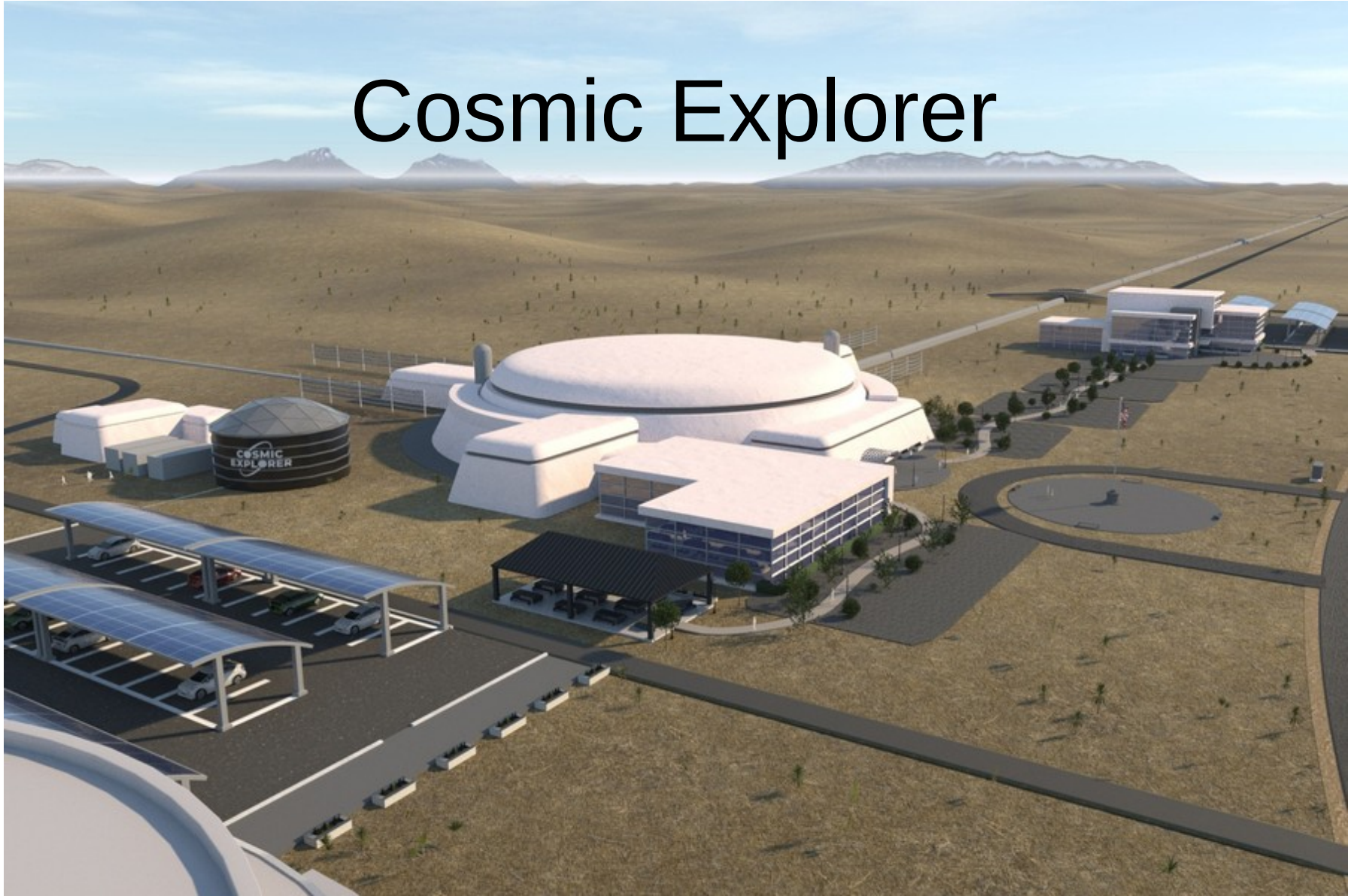


# ET Collaboration



29 countries, 1730 members (yesterday)

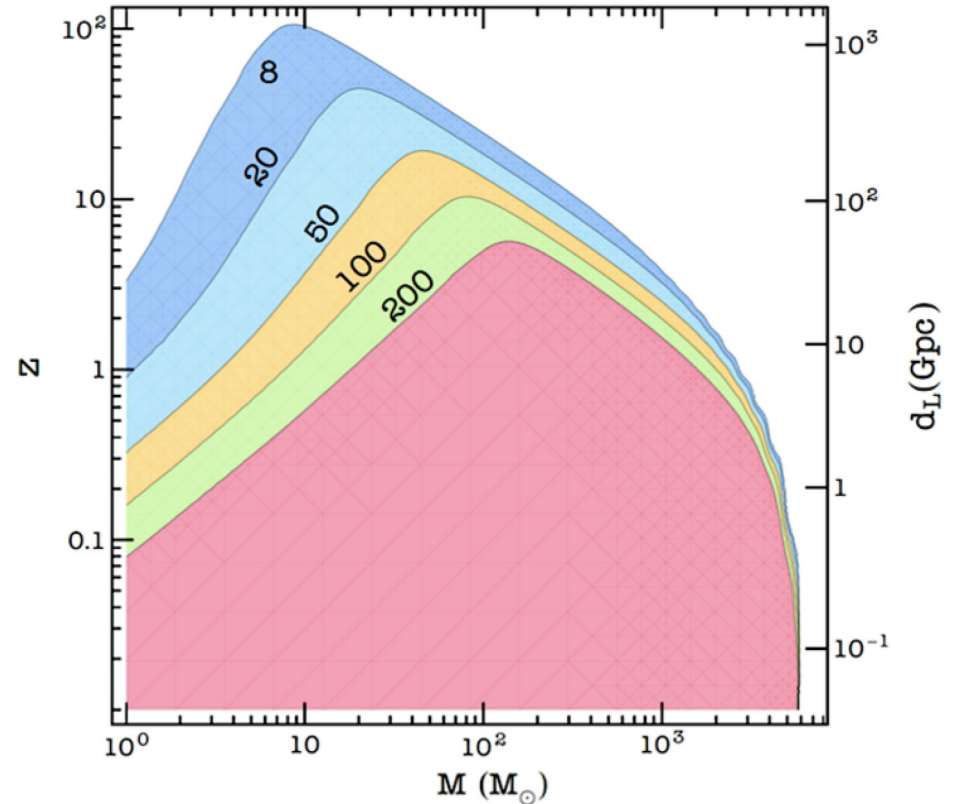
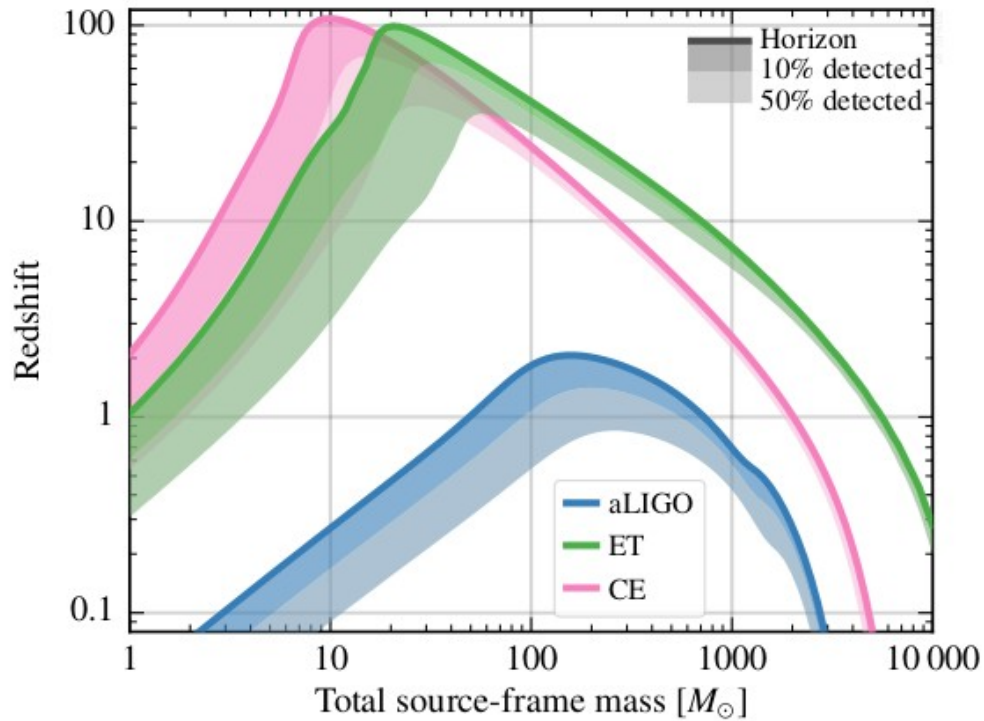
# Cosmic Explorer



What do we want to achieve with  
the Einstein Telescope?



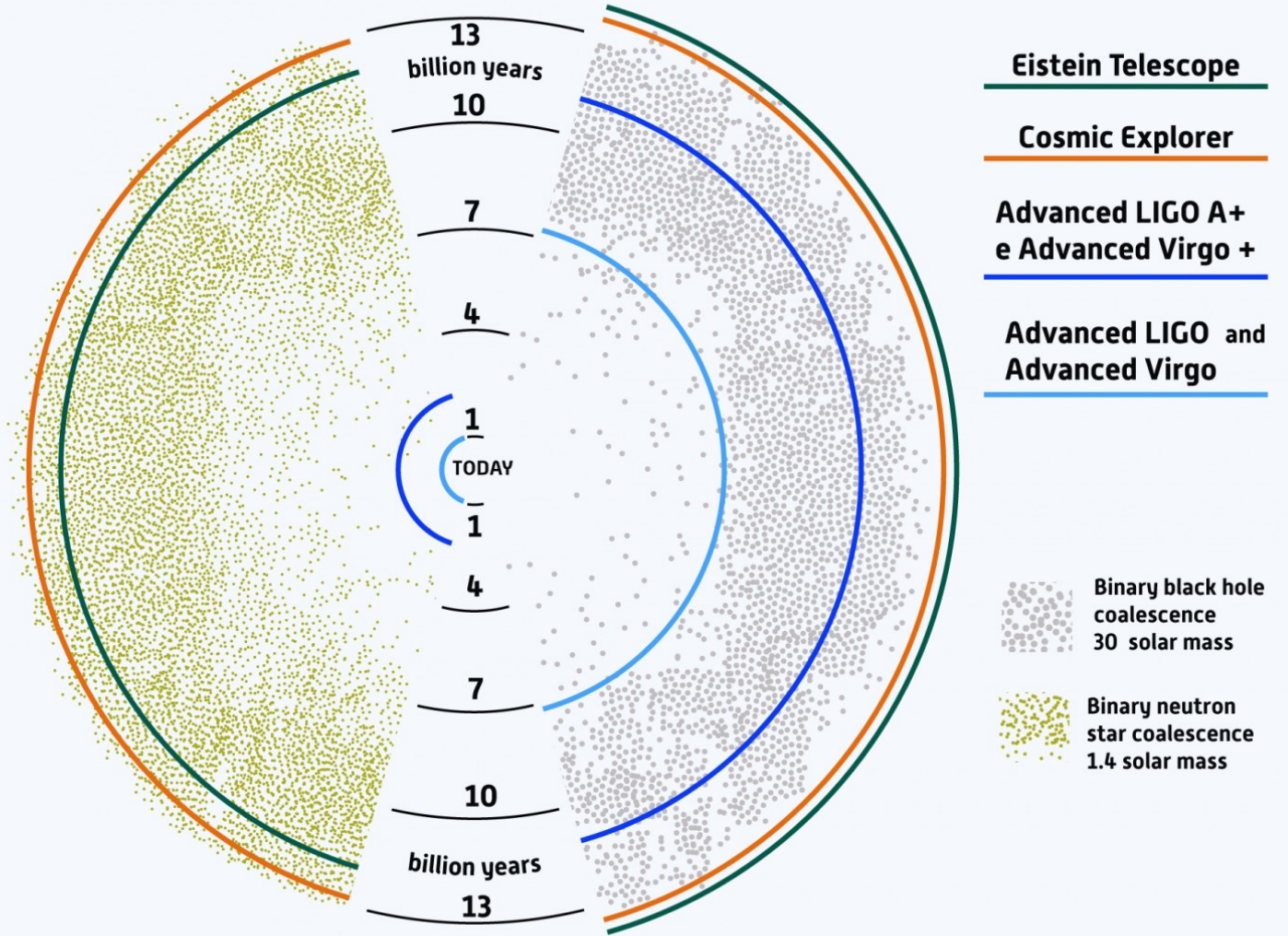
# Compact object merger ranges



NS-NS  
 $10^{4-5} \text{yr}^{-1}$

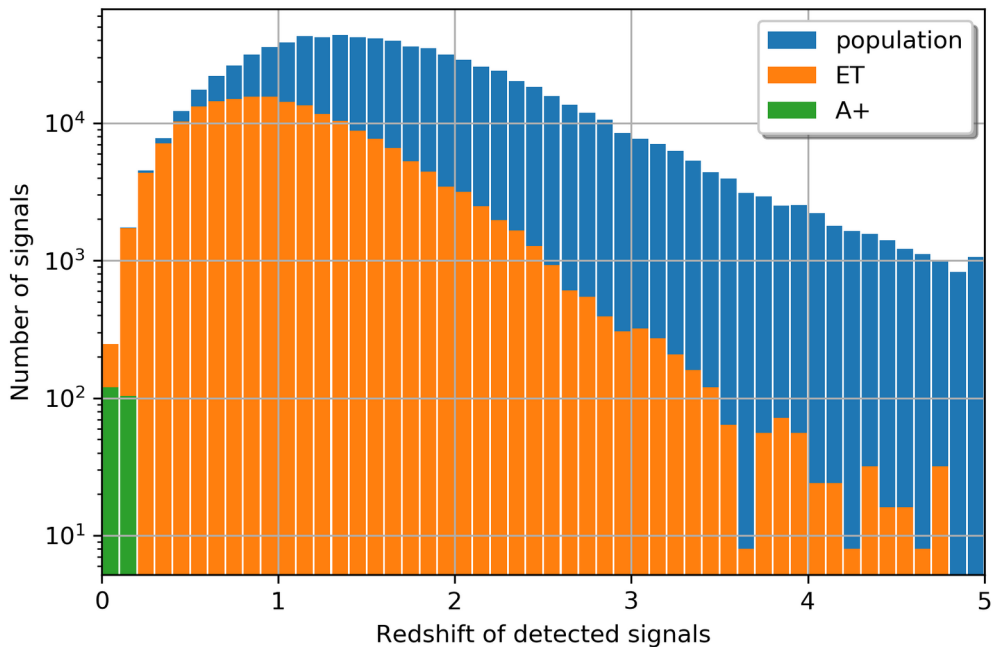
BH-BH  
 $10^{5-6} \text{yr}^{-1}$

REWORKED IMAGE FROM [FONTE ASTRO-PH 2109.09882]

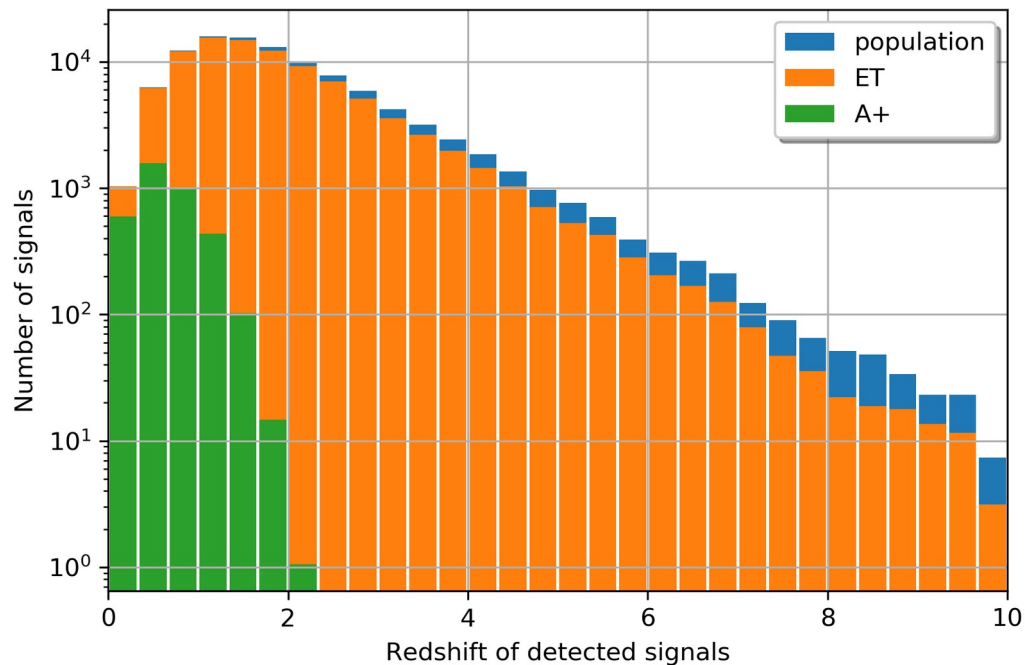


# Completeness

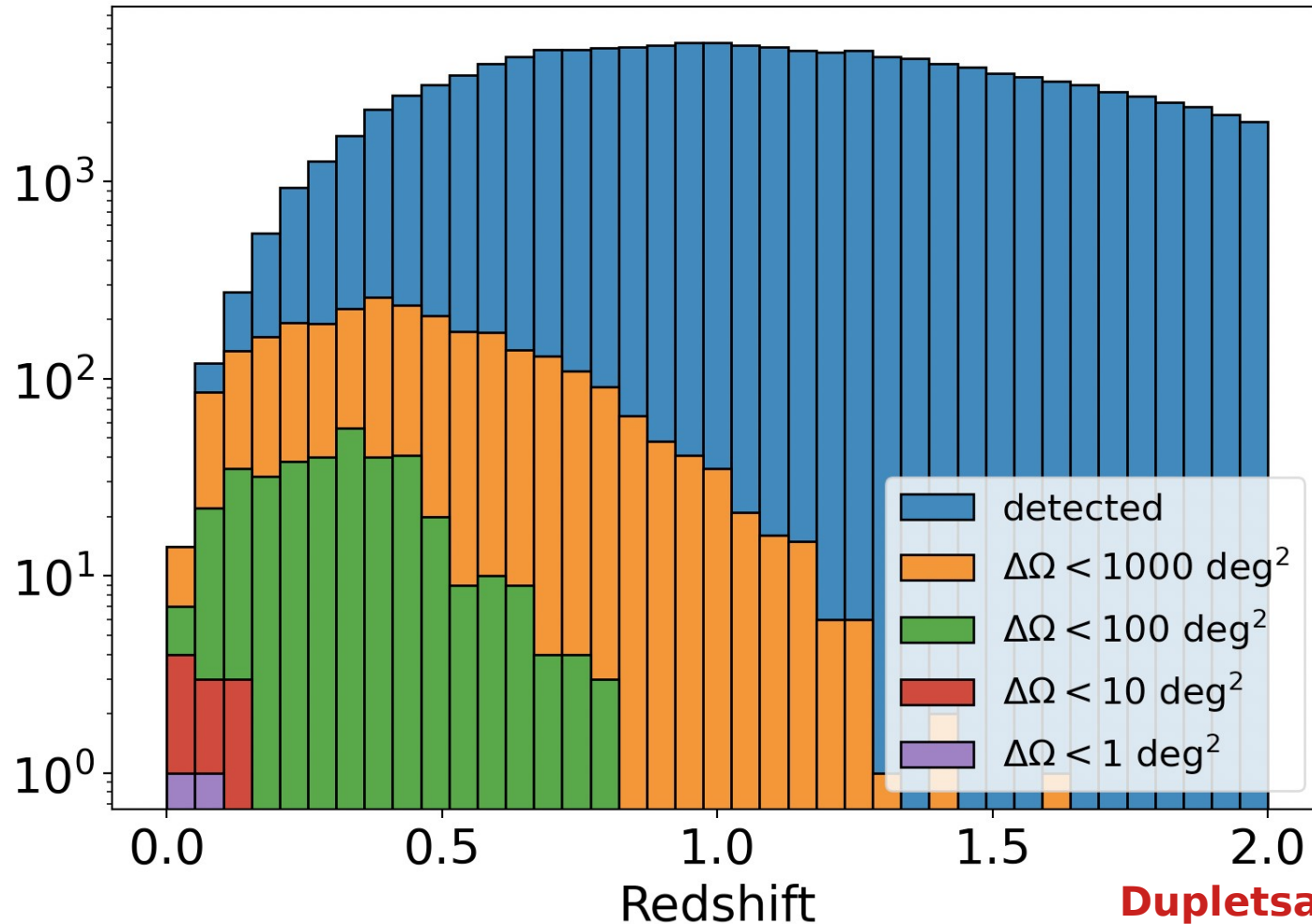
BH BH



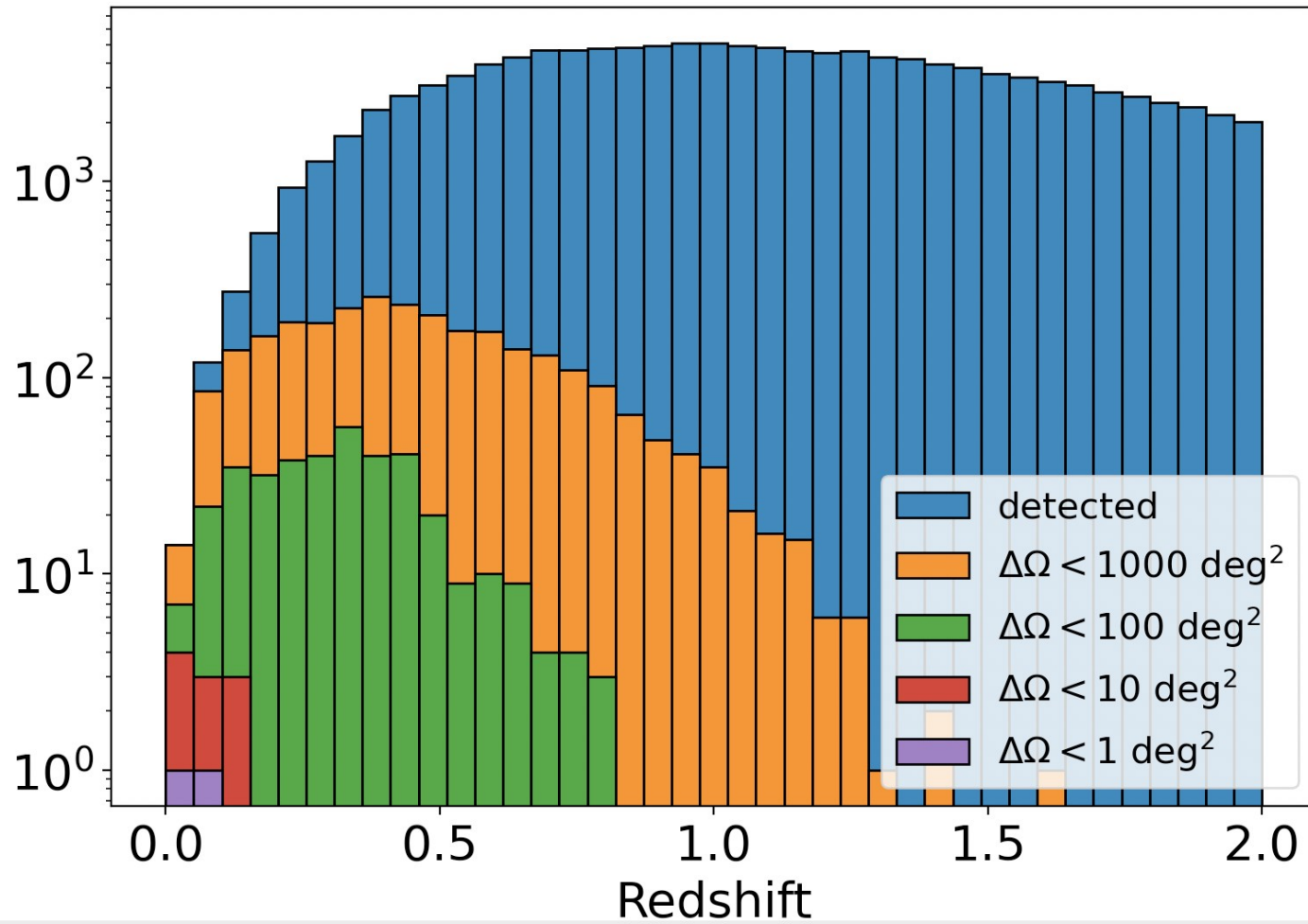
NS NS



# Sky localization with ET



# Network with CE

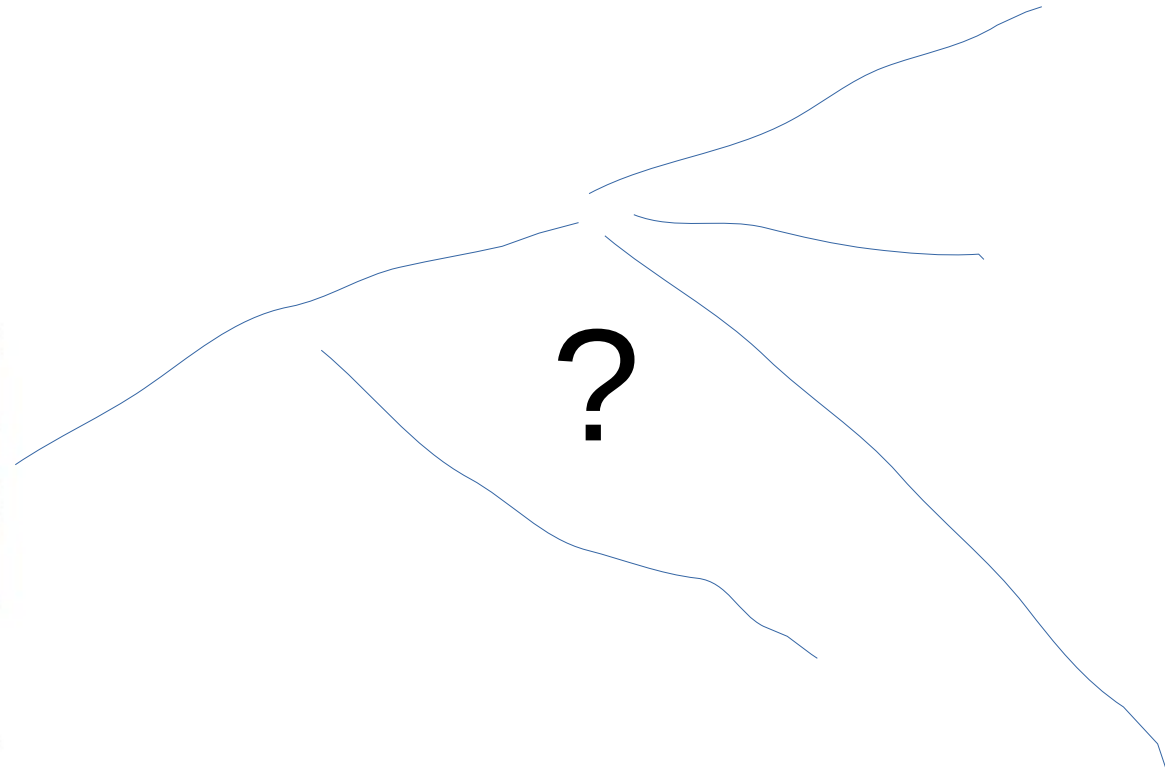
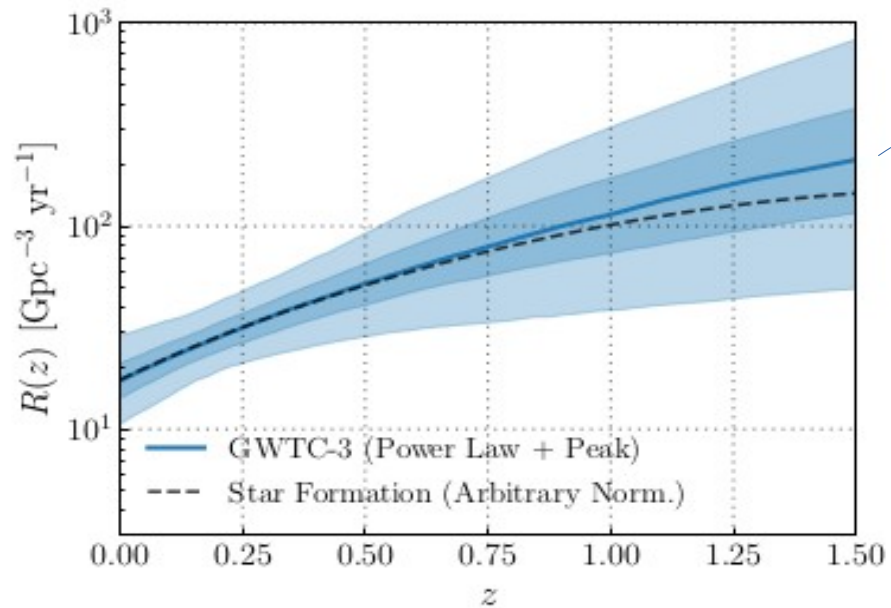


A few detections per day with localization better than 1 deg square!

# Black hole science

- Populations
  - PBHs, PopIII BHs,
  - Merger rate density as a function of redshift
  - Cosmology
- Precise waveform measurements
  - GR tests
  - Formation scenarios

# Rates vs redshift



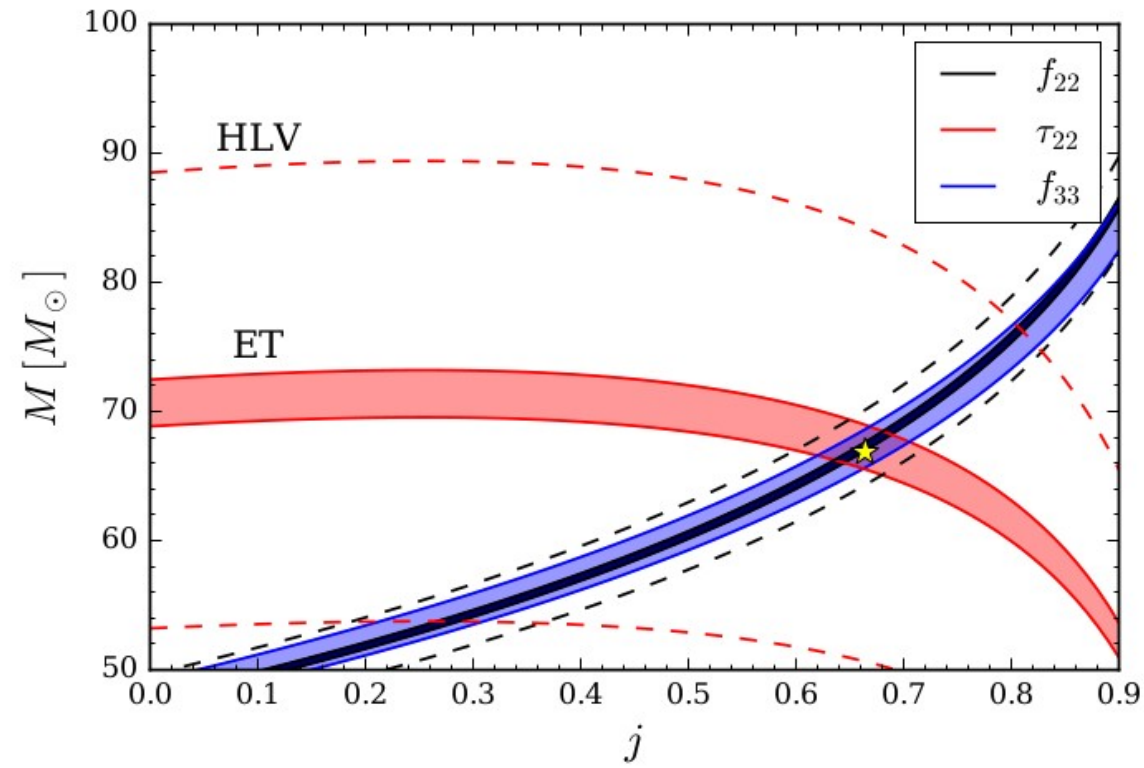
# GR tests

QNM analysis with very high accuracy

No hair theorem

Alternative theories

Exotic objects





# Cosmology with standard sirens

- Energy emitted  $E \sim M$
- Timescale of merger  $t \sim M$
- Luminosity  $L = E/t \sim \text{const}$
- In practice using the measured  $M_z$  and amplitude we obtain  $d_L$
- Need a separate measurement of  $z$ :
  - from known(?) mass distribution
  - from EM observation
  - From statistical considerations

# Cosmology with GW - now

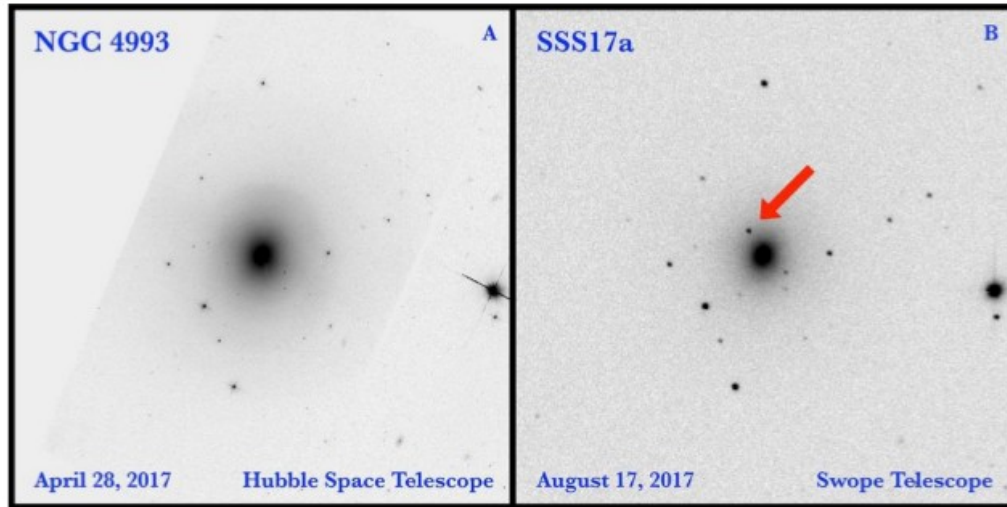
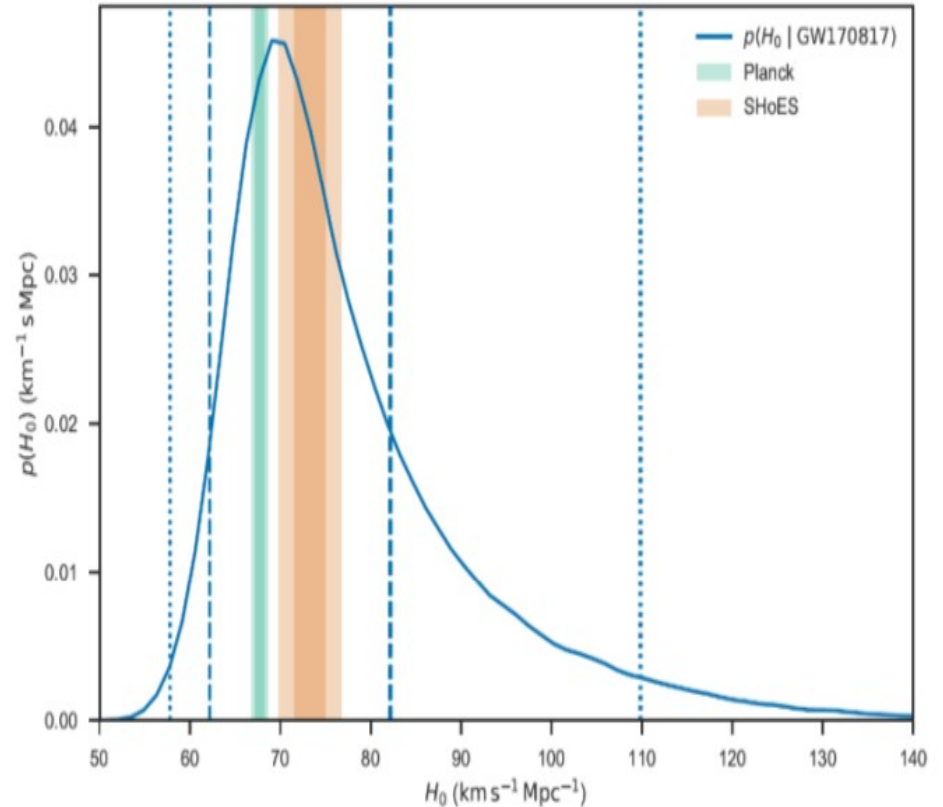


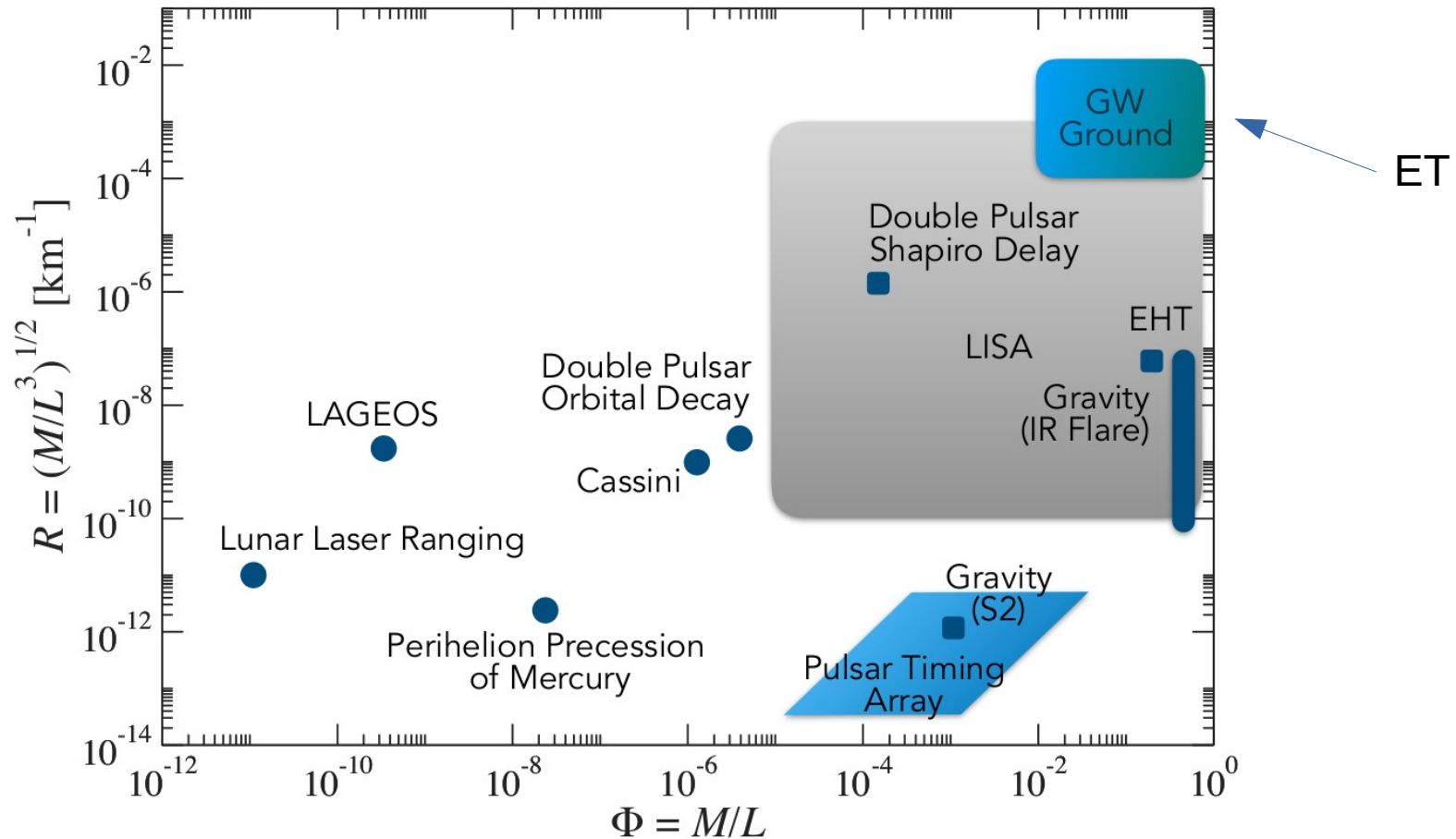
Fig. 4.  $3 \times 3$  arcminute images centered on NGC 4993 with North up and East left. (A) *Hubble Space Telescope* F606W-band (broad V) image from 4 months before the GW trigger (25, 35). (B) Swope image of SSS17a. The *i*-band image was obtained on 2017 August 17 at 23:33 UT by the Swope telescope at Las Campanas Observatory. SSS17a is marked with the red arrow. No object is present in the *Hubble* image at the position of SSS17a (25, 35).

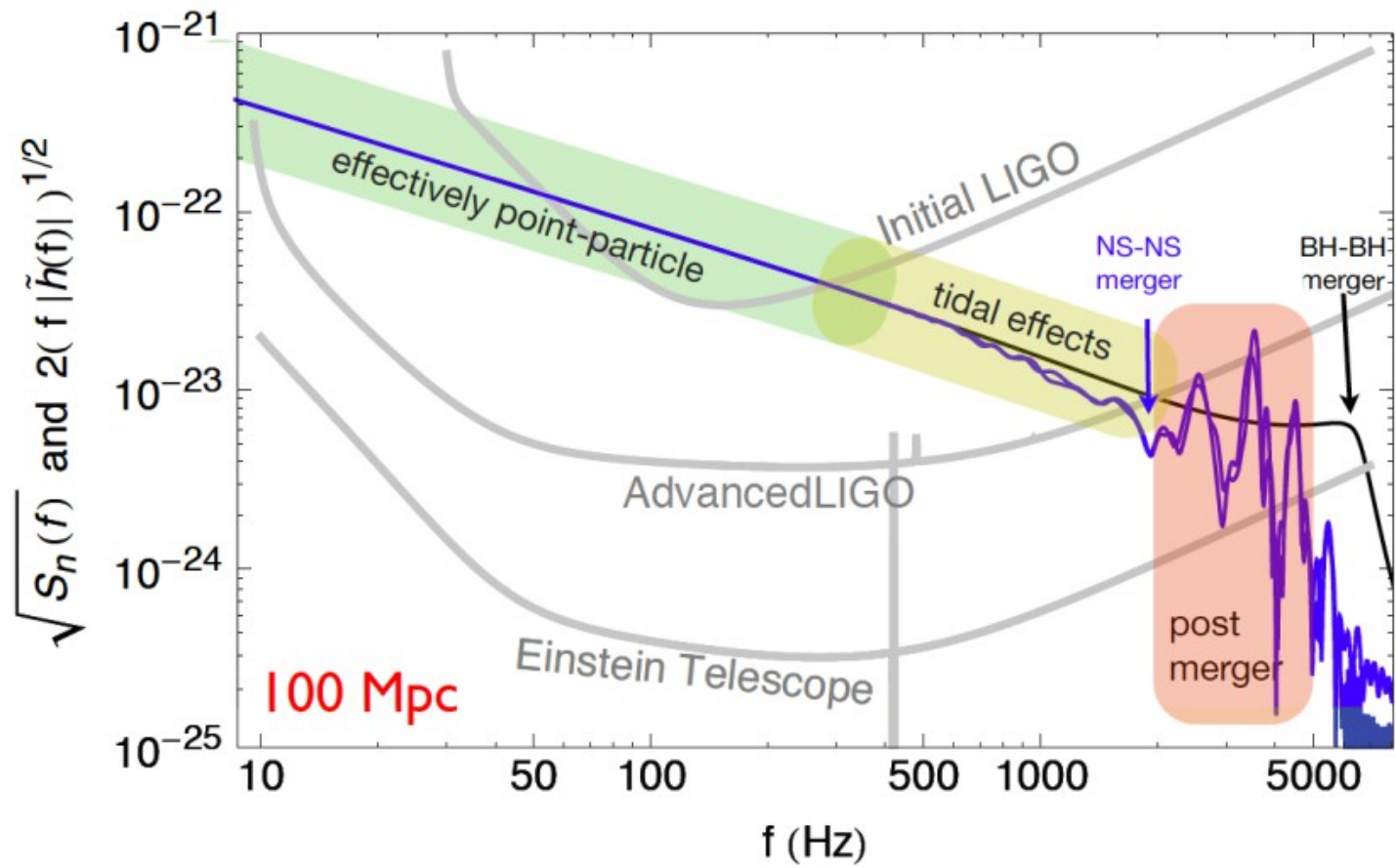


# Neutron Star Mergers

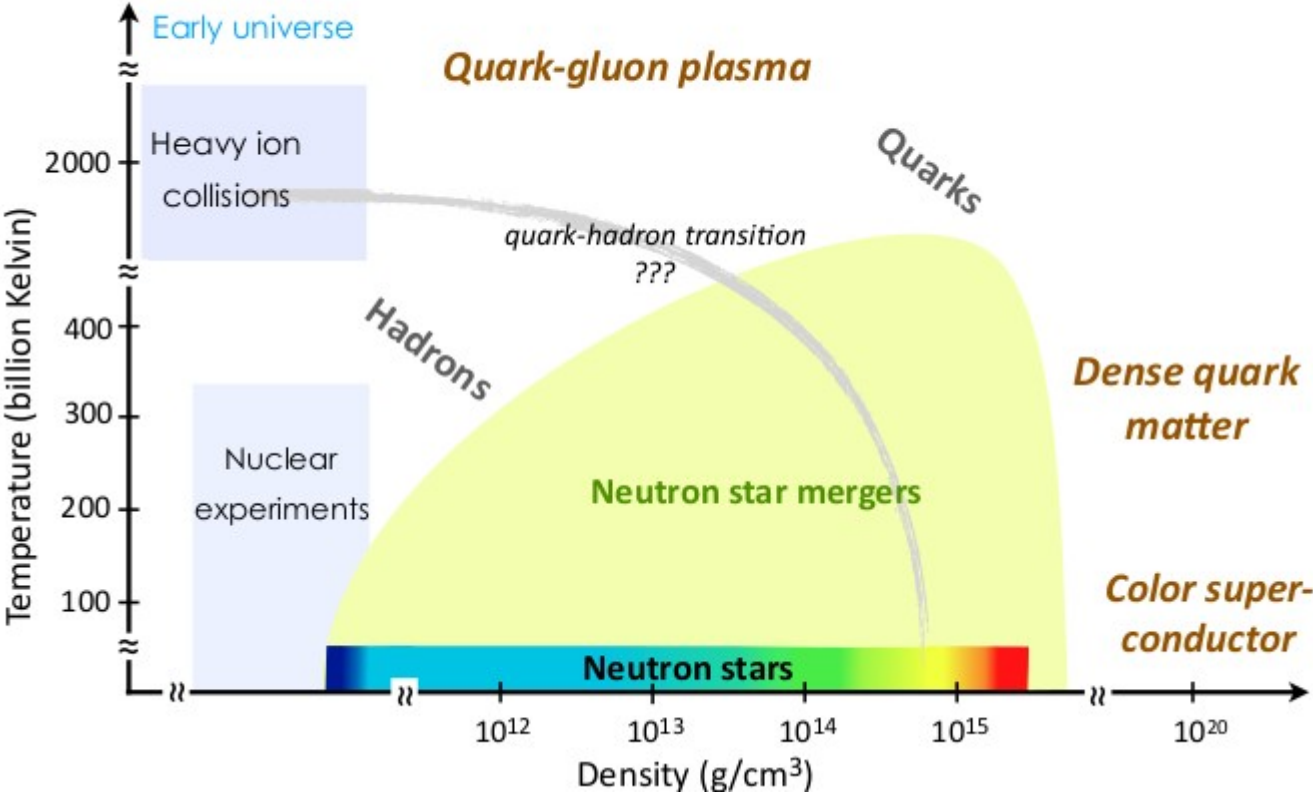
- Equation of state
- Multi-messenger studies
- Early warning on mergers
- GRB connection

# Probing gravity at extremes



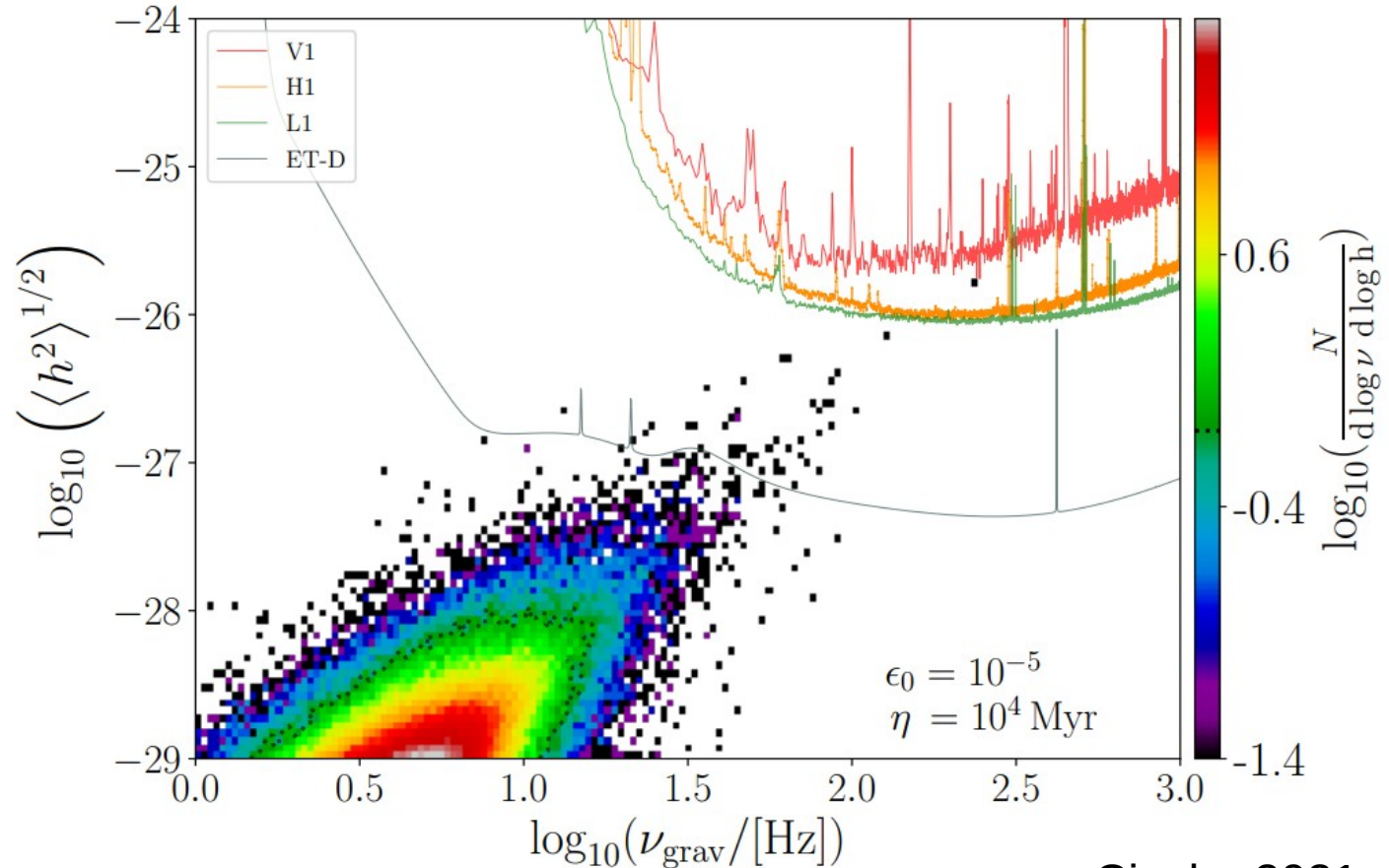


# Dense matter parameter space



# Pulsar population

Sensitivity in one year observation.



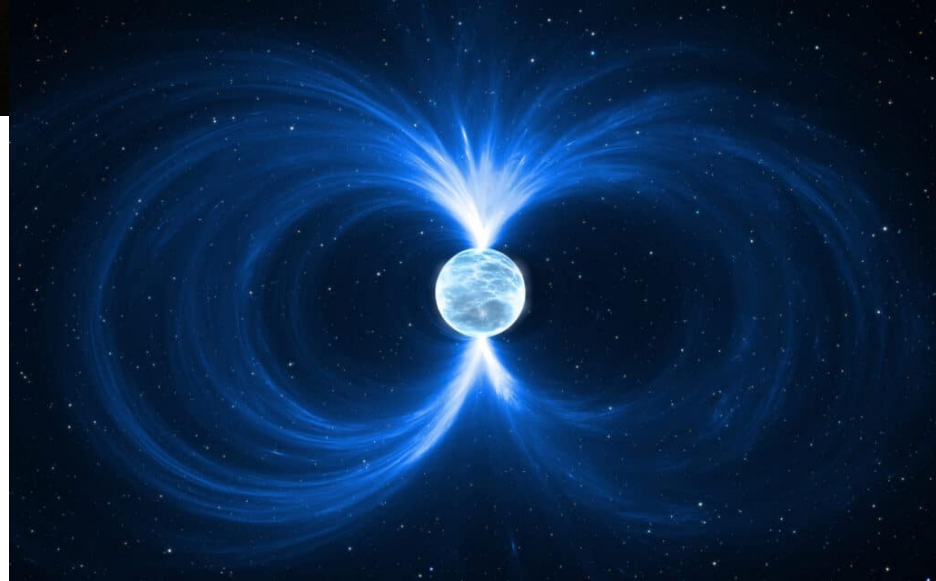
# Bursts

Supernovae

Fast radio  
bursts

Magnetars

Many other



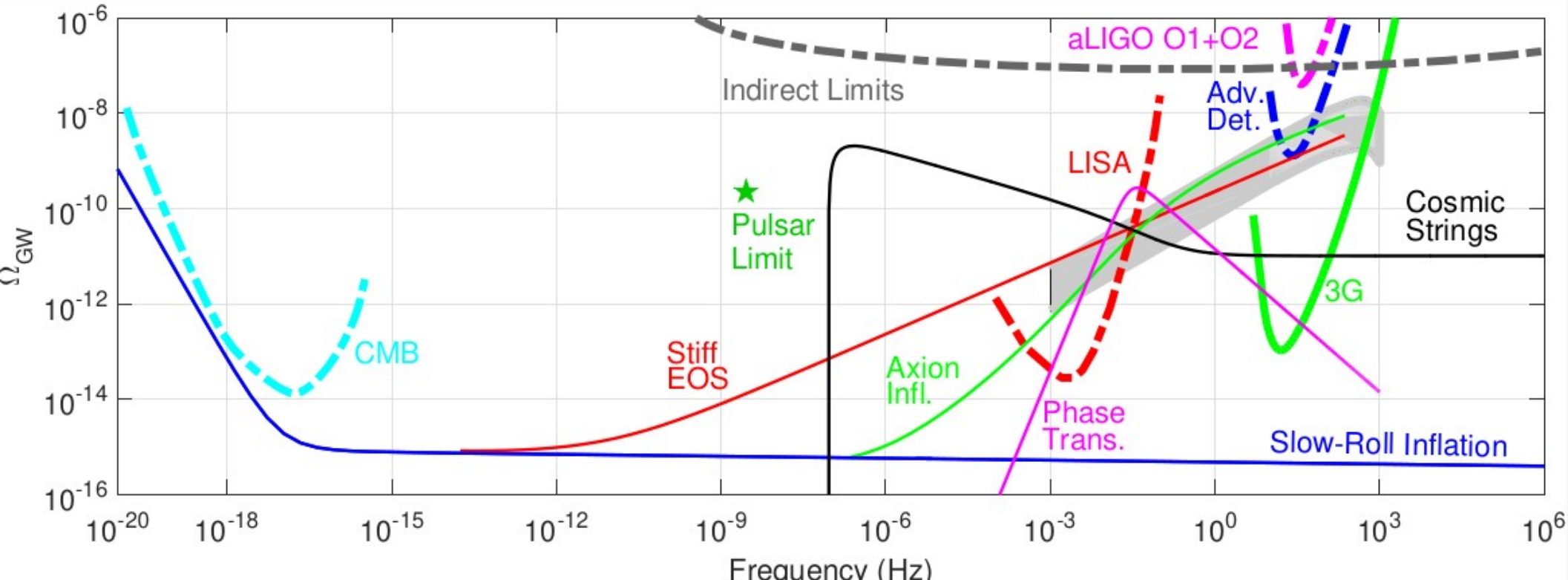


Ability to subtract  
binary foreground

# Backgrounds



Three (x2)  
interferometers,  
independent noise.



# ET Science in a nutshell

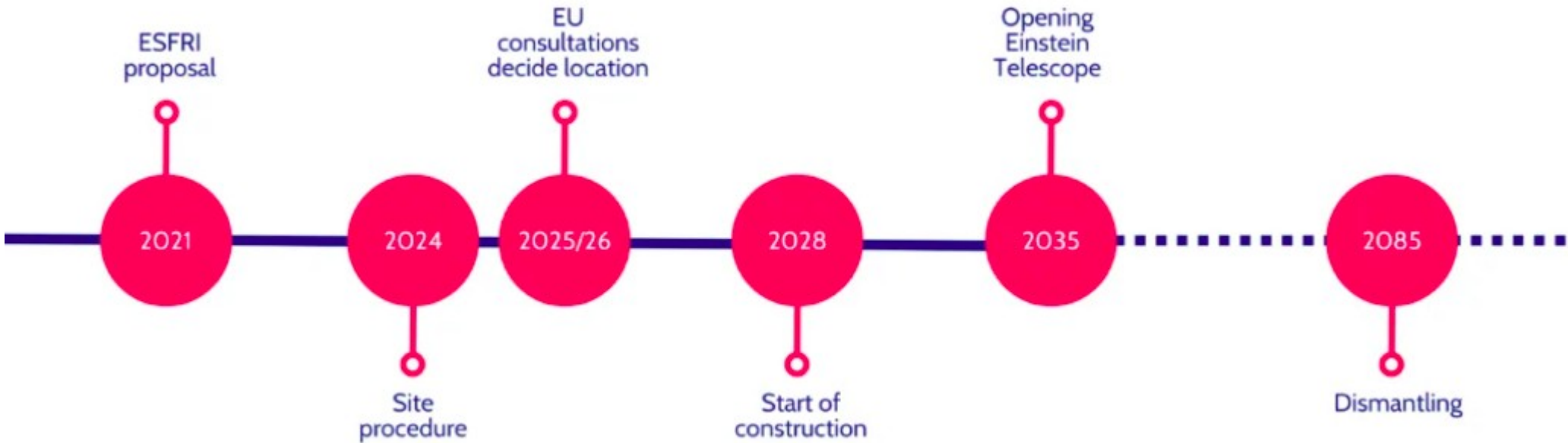
## ASTROPHYSICS

- **Black hole properties**
  - origin (stellar vs. primordial)
  - evolution, demography
- **Neutron star properties**
  - interior structure (QCD at ultra-high densities, exotic states of matter)
  - demography
- **Multi-band and -messenger astronomy**
  - joint GW/EM observations (GRB, kilonova,...)
  - multiband GW detection (LISA)
  - neutrinos
- **Detection of new astrophysical sources**
  - core collapse supernovae
  - isolated neutron stars
  - stochastic background of astrophysical origin

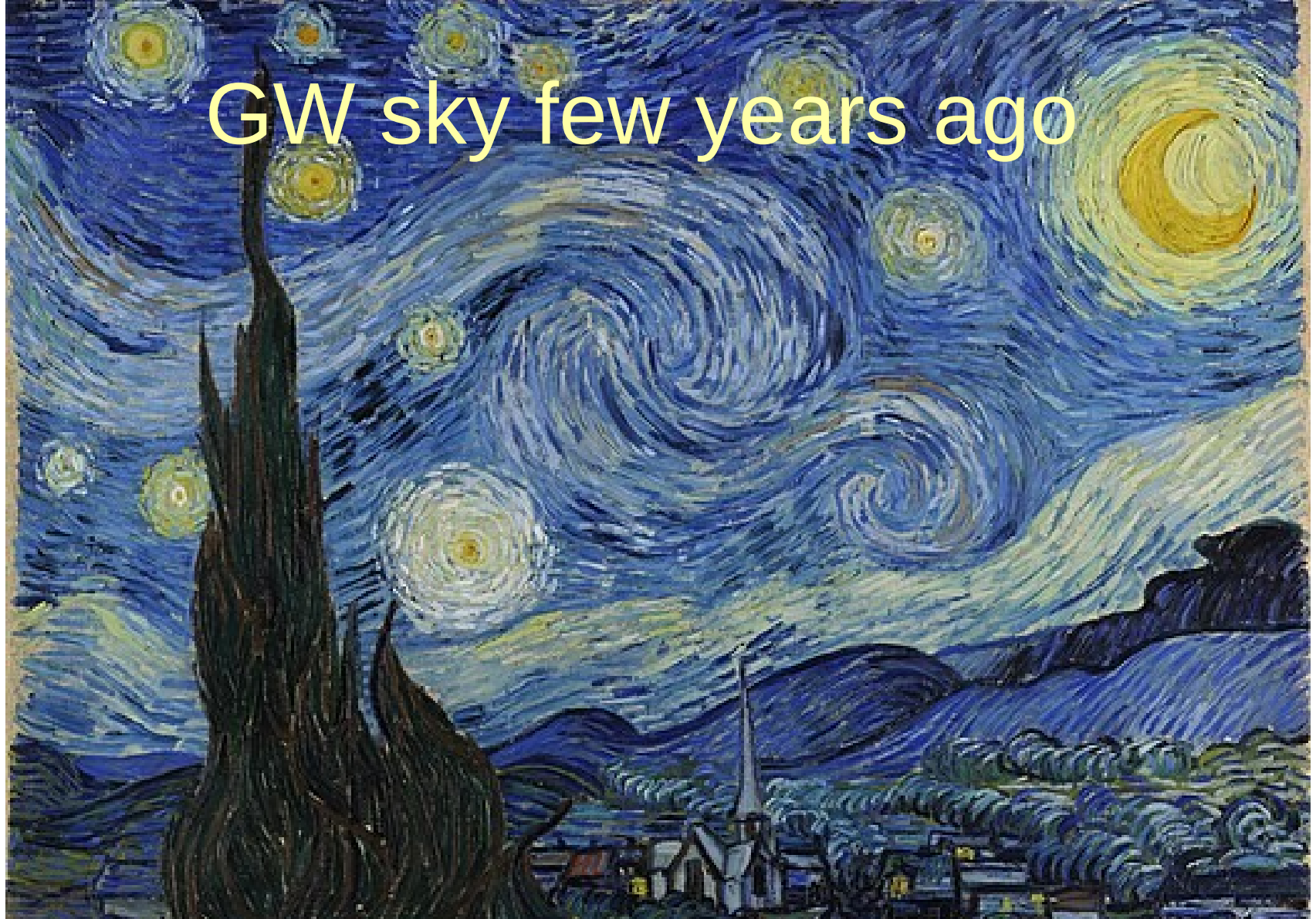
## FUNDAMENTAL PHYSICS AND COSMOLOGY

- **The nature of compact objects**
    - near-horizon physics
    - tests of no-hair theorem
    - exotic compact objects
  - **Tests of General Relativity**
    - post-Newtonian expansion
    - strong field regime
  - **Dark matter**
    - primordial BHs
    - axion clouds, dark matter accreting on compact objects
  - **Dark energy and modifications of gravity on cosmological scales**
    - dark energy equation of state
    - modified GW propagation
- Stochastic cosmological backgrounds**
- inflation, phase transitions, cosmic strings

# Timeline



GW sky few years ago



# GW sky with Advanced detectors



GW sky with ET

The image features a dark, star-filled night sky. A prominent, bright band of the Milky Way galaxy stretches diagonally across the frame, from the upper right towards the lower left. The stars are numerous and vary in brightness, creating a dense field of light points. The overall color palette is dominated by deep blues and blacks, with the white and yellowish light of the stars providing contrast.



# Present and future GW astronomy

