

Latest Results from the CUORE

Experiment

Direct Dark Matter Detection

RICAP 2024

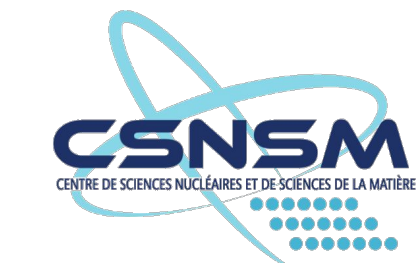


Yale



Samantha Pagan on behalf of the CUORE Collaboration

CUORE collaboration and acknowledgments



The CUORE Collaboration thanks the directors and staff of the Laboratori Nazionali del Gran Sasso and the technical staff of our laboratories. This work was supported by the Istituto Nazionale di Fisica Nucleare (INFN); the National Science Foundation under Grant Nos. NSF-PHY-0605119, NSF-PHY-0500337, NSF-PHY-0855314, NSF-PHY-0902171, NSF-PHY-0969852, NSF-PHY-1307204, NSF-PHY-1314881, NSF-PHY-1401832, and NSF-PHY-1913374; Yale University, Johns Hopkins University, and University of Pittsburgh. This material is also based upon work supported by the US Department of Energy (DOE) Office of Science under Contract Nos. DE-AC02-05CH11231 and DE-AC52-07NA27344; by the DOE Office of Science, Office of Nuclear Physics under Contract Nos. DE-FG02-08ER41551, DE-FG03-00ER41138, DE-SC0012654, DE-SC0020423, DE-SC0019316. This research used resources of the National Energy Research Scientific Computing Center (NERSC) This work makes use of both the DIANA data analysis and APOLLO data acquisition software packages, which were developed by the CUORICINO, CUORE, LUCIFER, and CUPID-o Collaborations. The authors acknowledge Advanced Research Computing at Virginia Tech for providing computational resources and technical support that have contributed to the results reported.

CUORE

The **C**ryogenic **U**nderground **O**bservatory for **R**are **E**vents

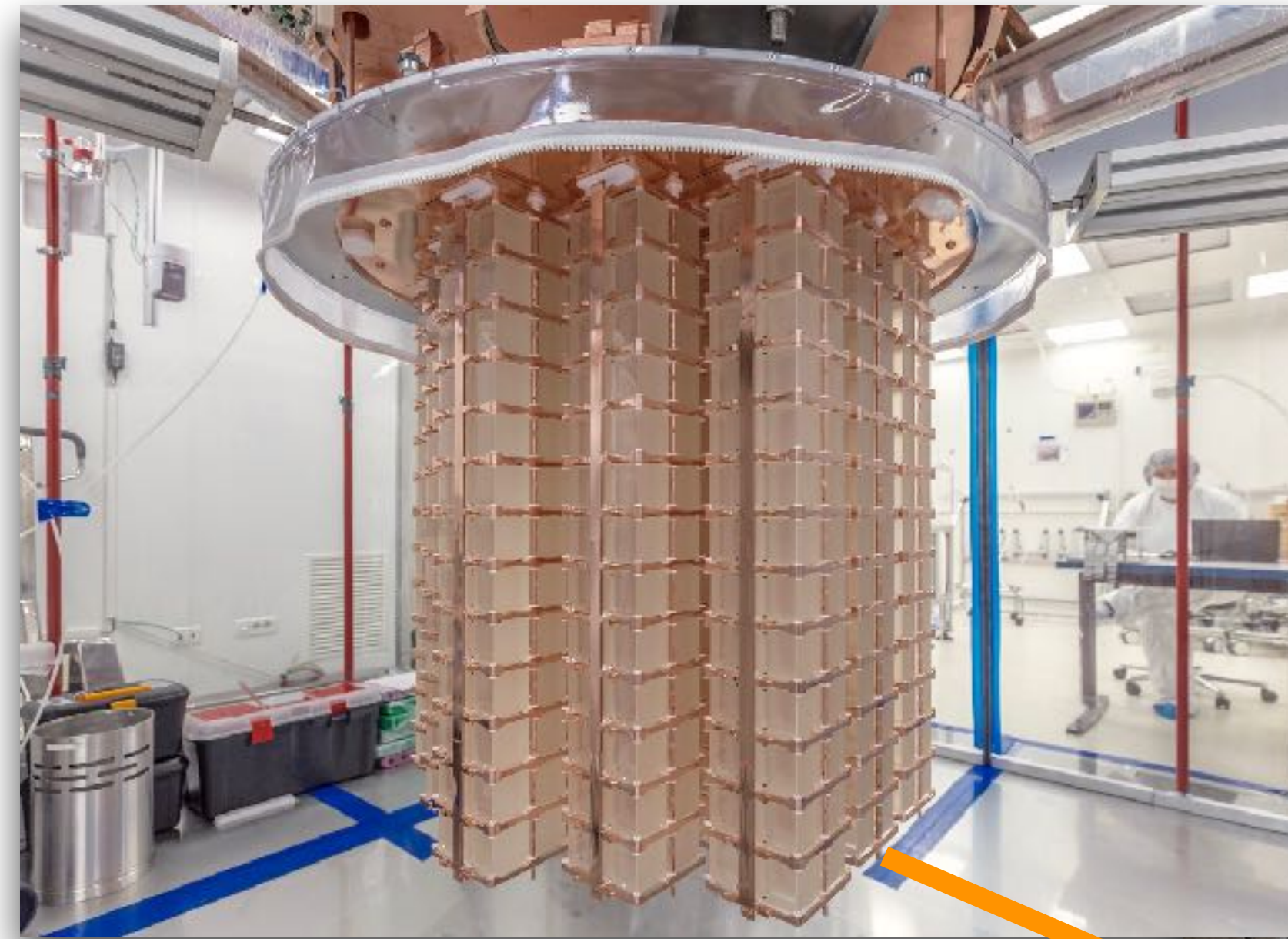


Science Program: Neutrinoless double-beta decay, double-beta decay, decay studies, beyond the standard model searches

Located at LNGS Gran Sasso National Laboratory (~3600 m.w.e.)

988 TeO₂ crystal cryogenic calorimeters

Began collecting data in 2017, stably taking data since 2019



CUORE



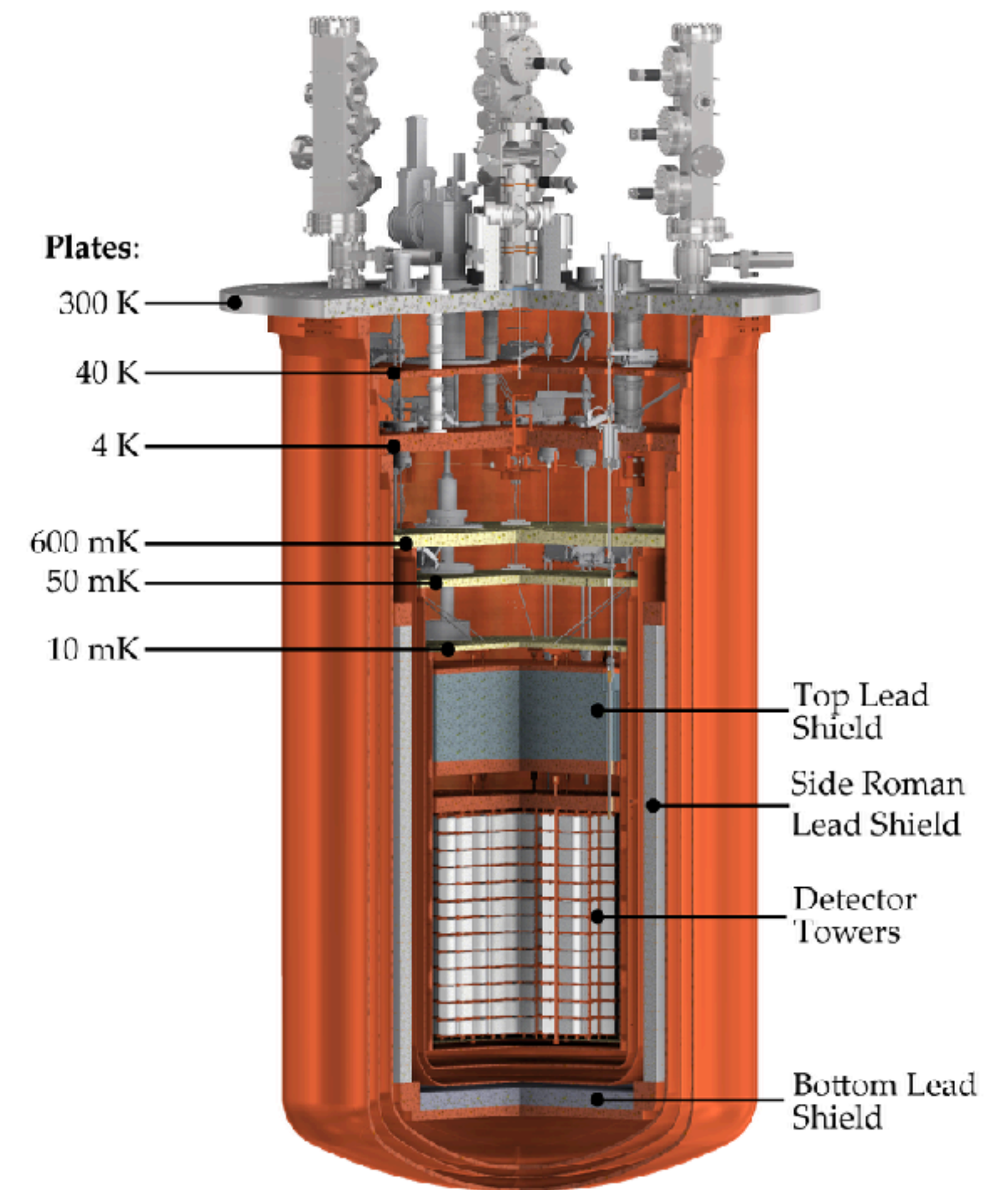
Modern lead on top of detectors and side
Roman lead shielding

Custom dilution refrigerator, nested
vessels, and pulse tubes, with vibration
mitigation

Operated at ~10 mK

19 towers with 13 floors of 4 crystals each

Natural TeO_2 crystals
 $5 \times 5 \times 5 \text{ cm}^3$, 750 g each



CUORE Cryostat

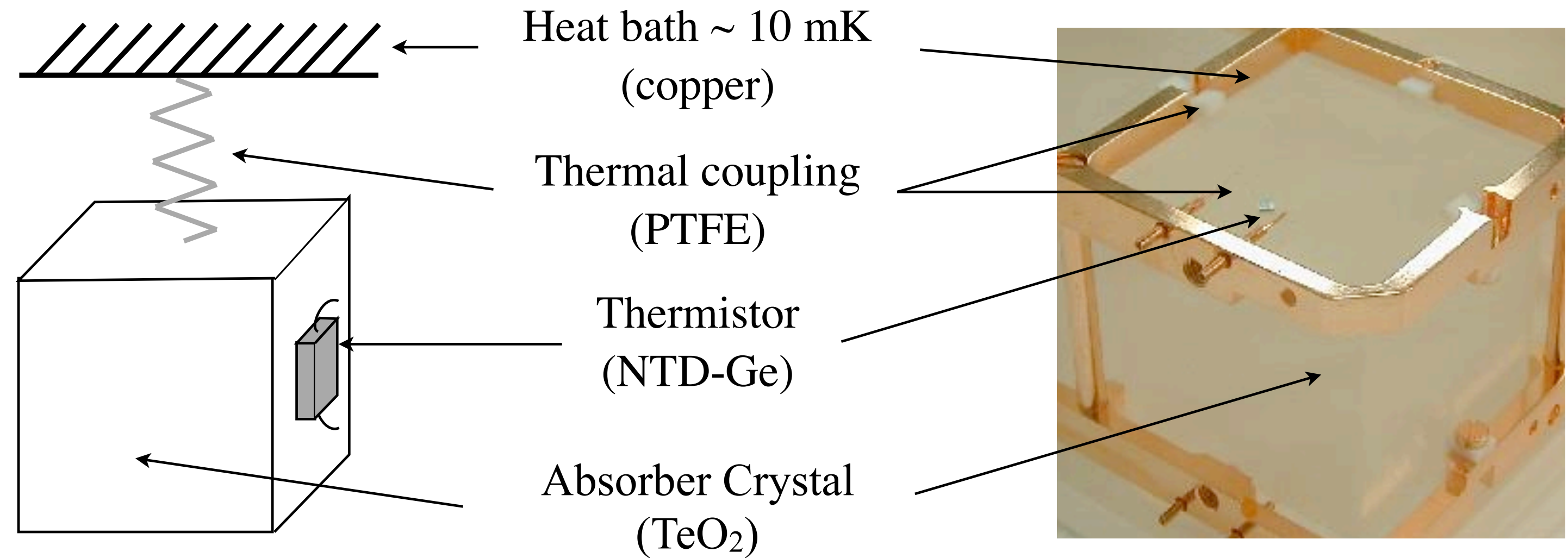
Cryogenic calorimeters: operating principle



Particle interaction
increase crystal
temperature

Thermal coupling to 10
mK heat bath

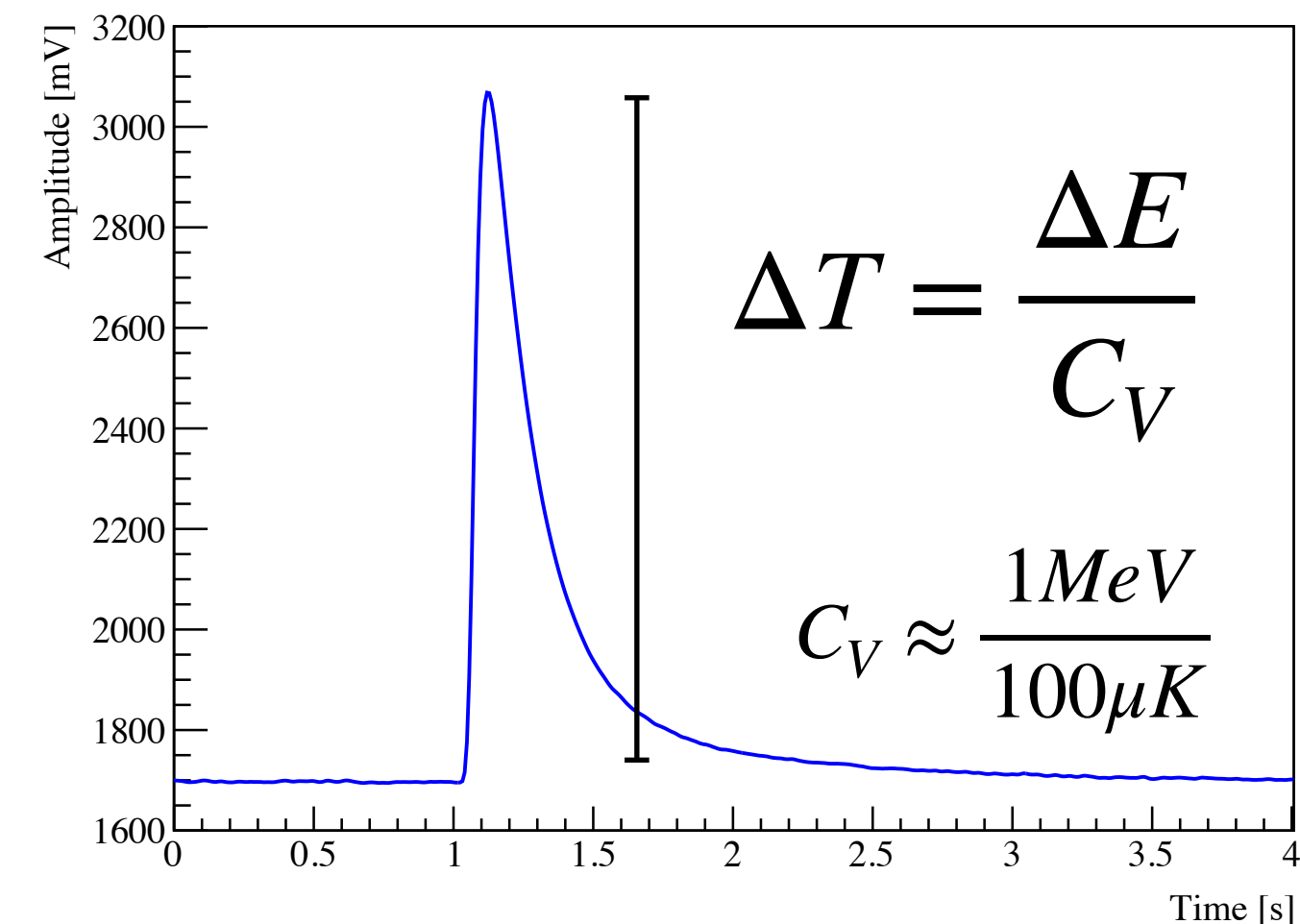
Neutron Transmutation
Doped (NTD) Ge
thermistors



$$C_V(T) \propto T^3$$

$$R(T) = R_0 e^{\sqrt{\frac{T_0}{T}}}$$

E. E. Haller, J. Appl. Phys. 77,
2857-2878 (1995)



CUORE data taking: 2 tonne·yr (2TY)

TeO₂ exposure

~86 kg·yr

Alduino, C et al. (CUORE Collaboration), *Phys. Rev. Lett.* **120**, 132501 (2018)

~300 kg·yr

Adams, D.Q. et al. (CUORE Collaboration), *Phys. Rev. Lett.* **124**, 122501 (2020)

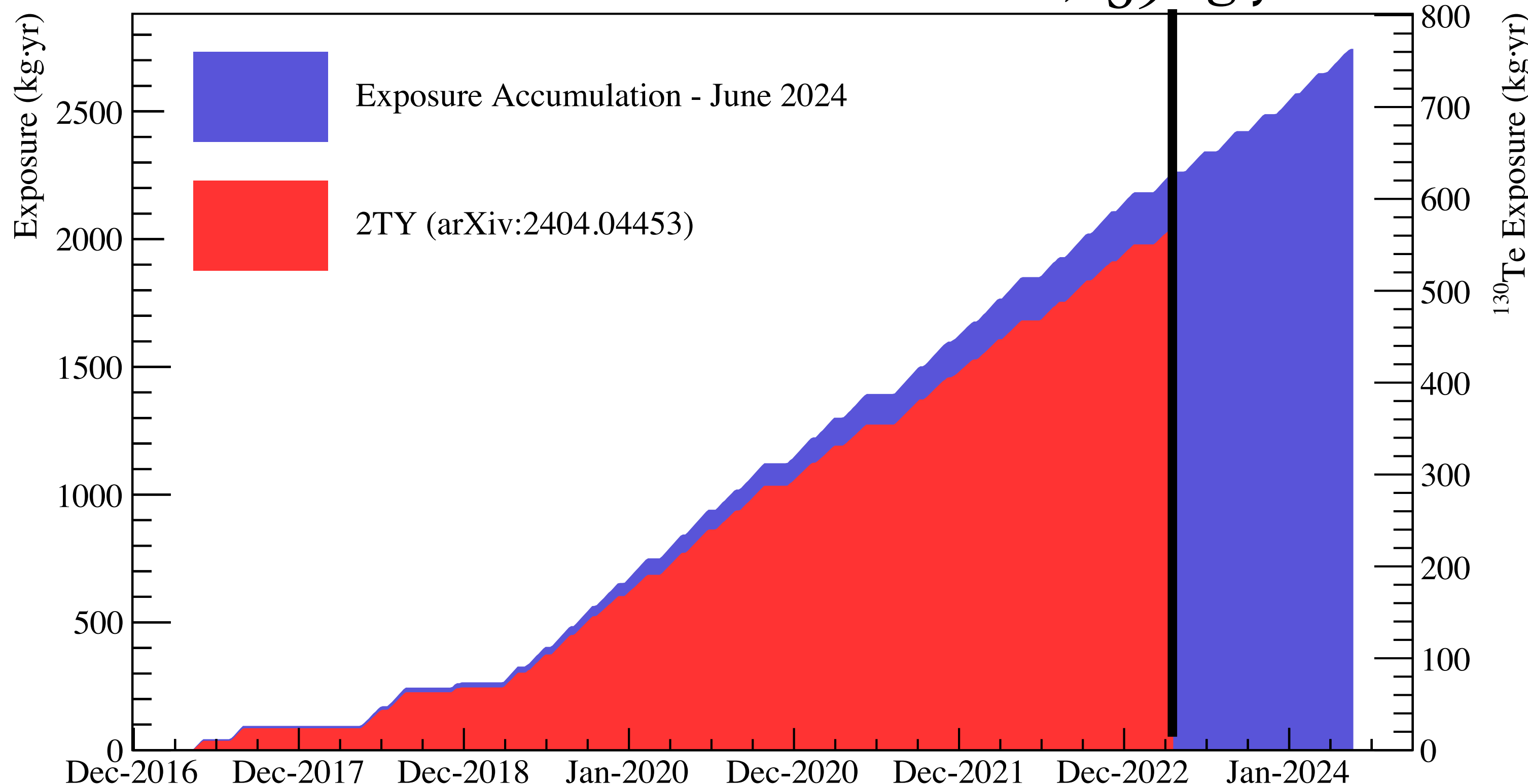
~1,000 kg·yr

Adams, D.Q. et al. (CUORE Collaboration), *Nature* **604**, 53-58 (2022)

~2,000 kg·yr

Adams, D.Q. et al. (CUORE Collaboration), *arXiv:2404.04453* (2024)

2,039 kg·yr

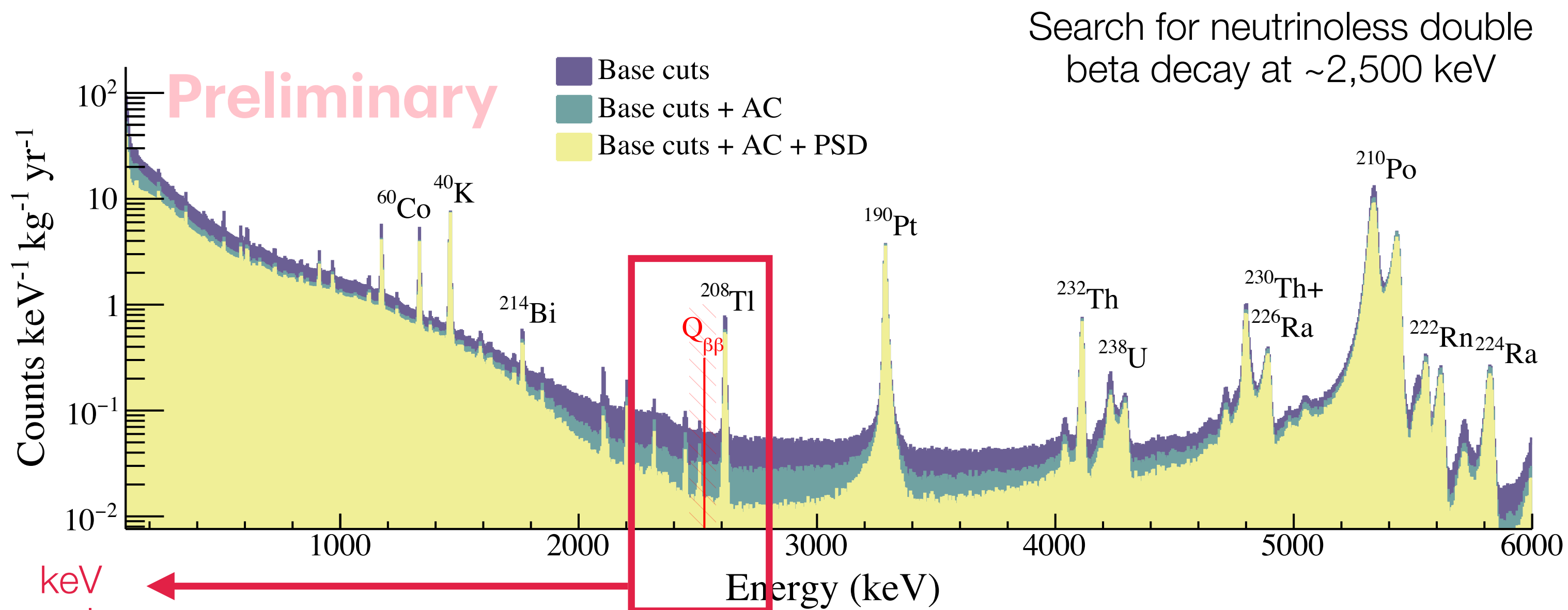


Data taking is divided into ~1.5 months long datasets

CUORE 2TY energy spectrum



Adams, D.Q. et al. (CUORE Collaboration),
arXiv:2404.04453 (2024)

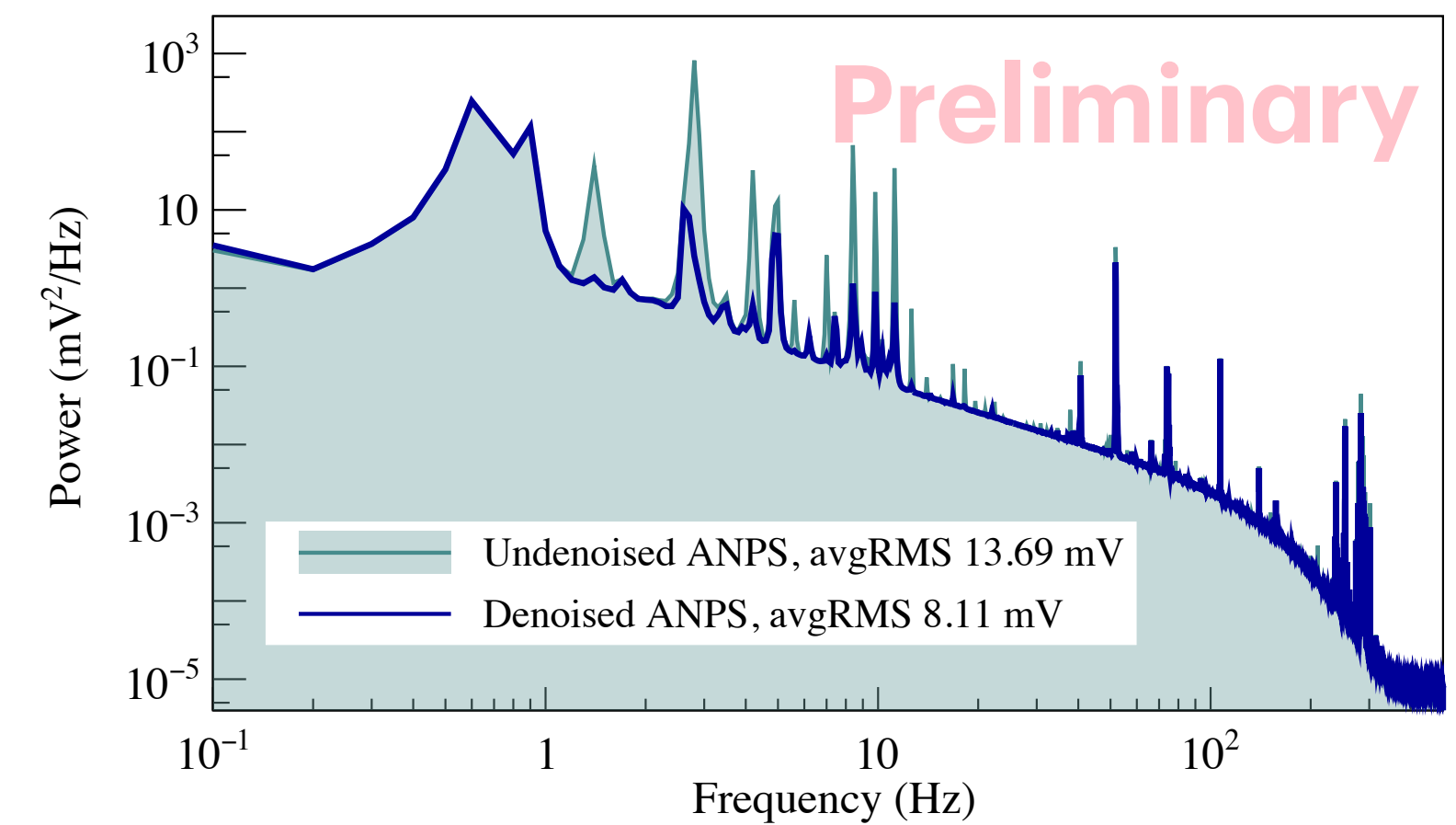


2039.0 kg·yr TeO_2 (567.0 kg·yr ^{130}Te)

2 TY Improvements

De-noising using auxiliary devices

Vetter, K. J. et al. arXiv:2311.01131 (2024)



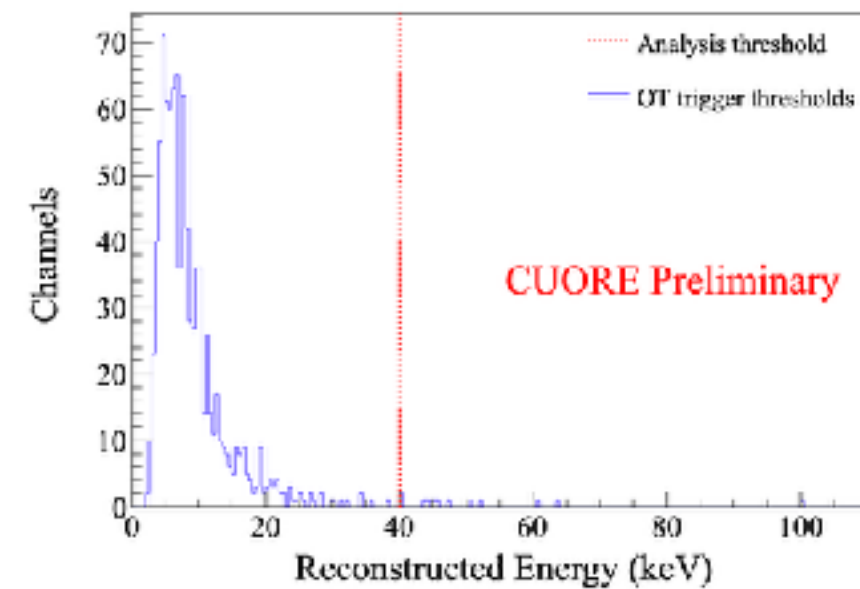
Updates to multiple analysis methods and addition of low energy variables

CUORE analysis chain



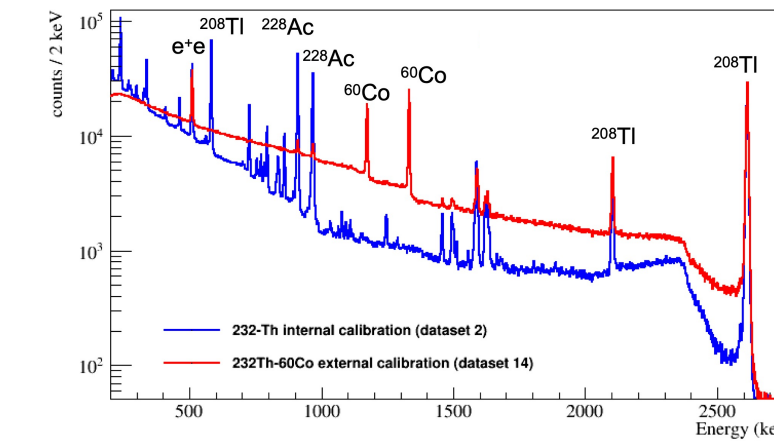
1. Triggering

Data is continuously saved and can be re-triggered offline



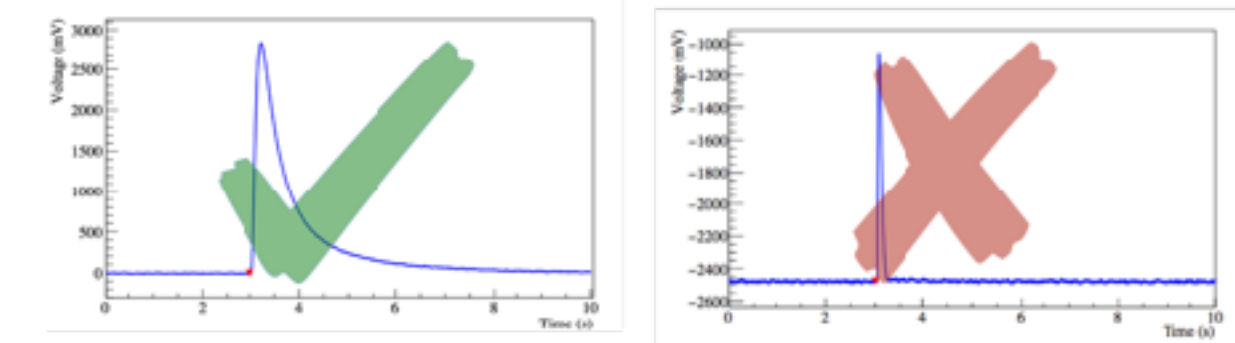
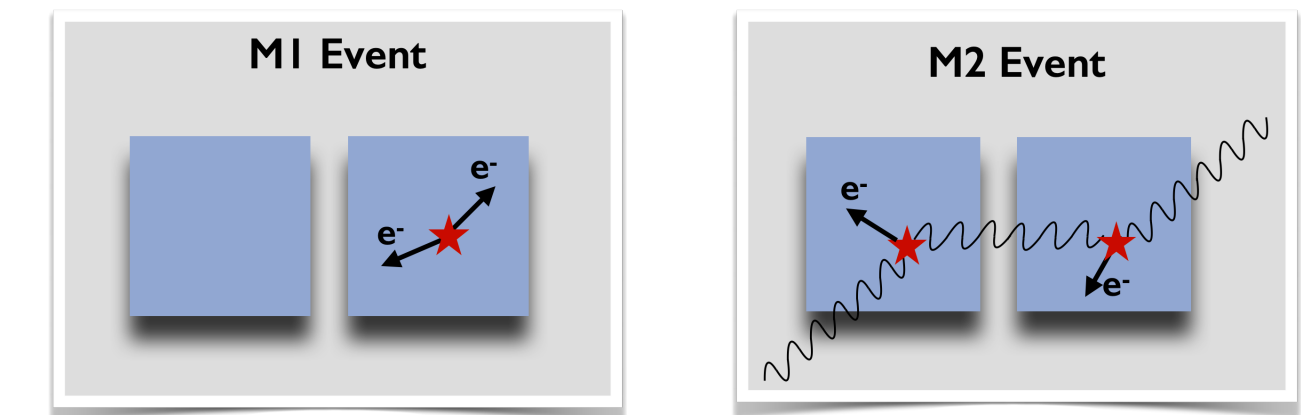
4. Energy Calibration

Calibrate calorimeters using known peaks



5. Building coincidence events

Differentiate events that occur in one or multiple crystals

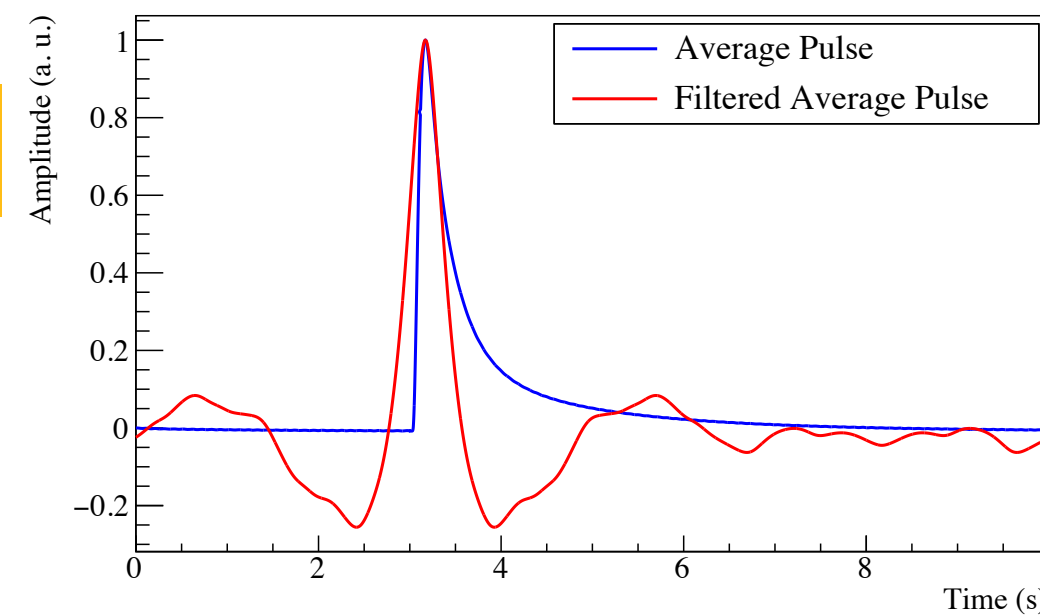


6. Pulse Shape Analysis

Save pulses with physical shapes

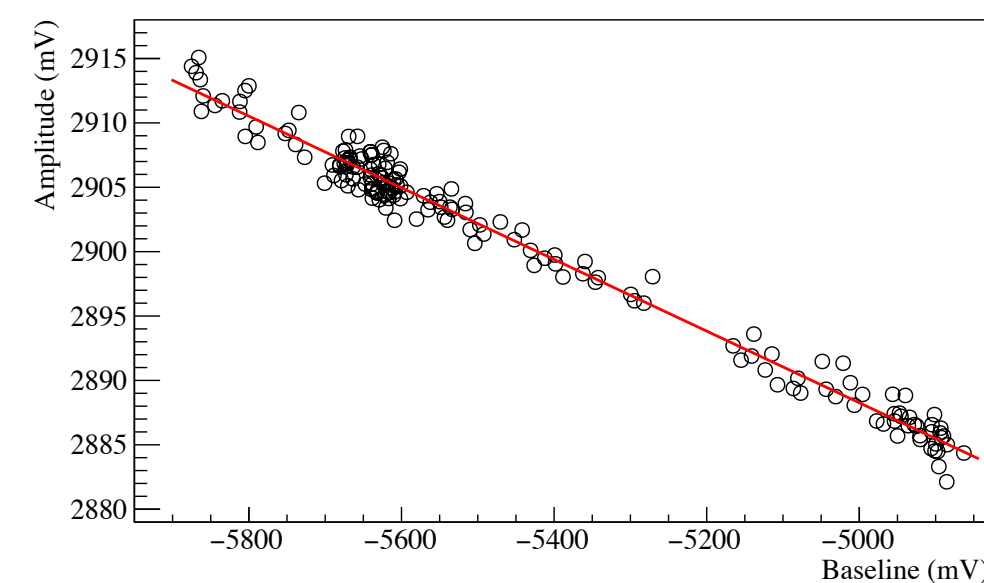
2. Amplitude Evaluation

Calculate the amplitude of pulses using an optimum filter, which maximizes signal to noise ratio



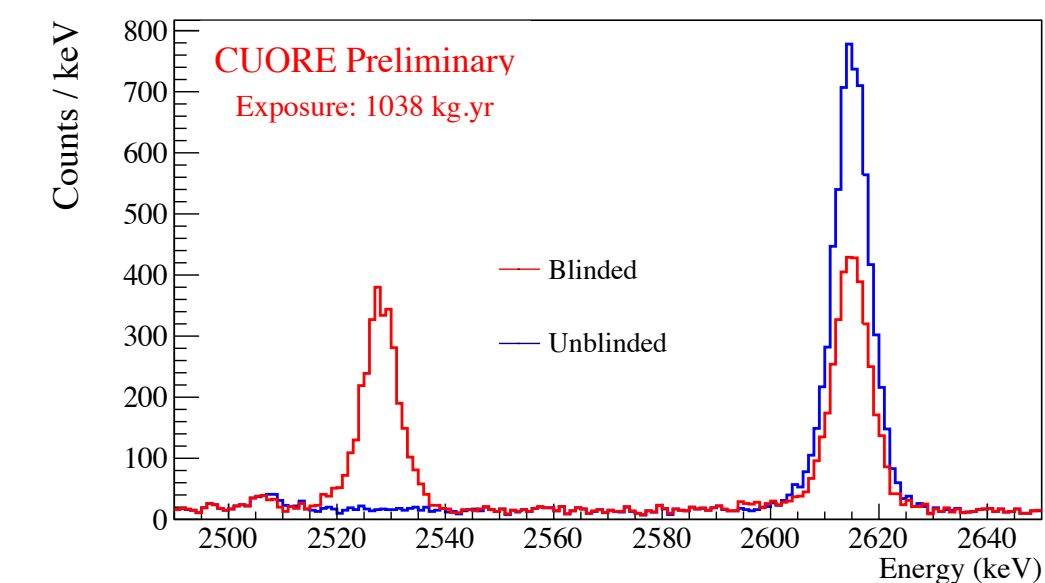
3. Gain Stabilization

Correct for calorimeter gain dependence on drifts in temperature



7. Peak Salting

Data blinding technique, that moves data from a nearby peak



CUORE analysis chain



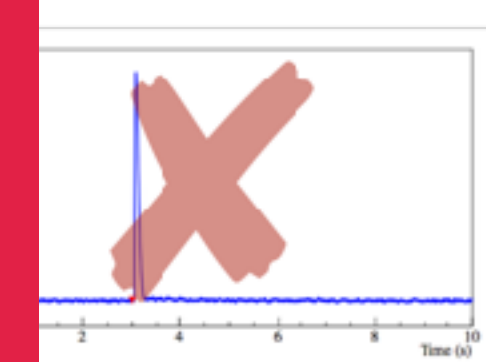
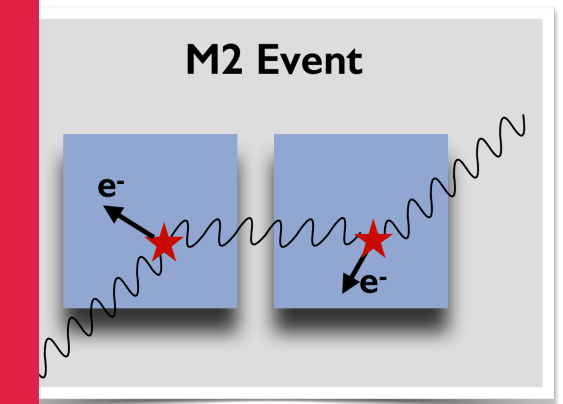
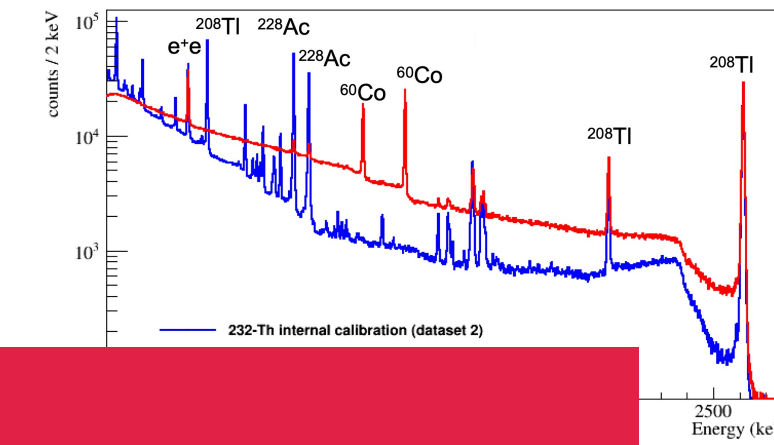
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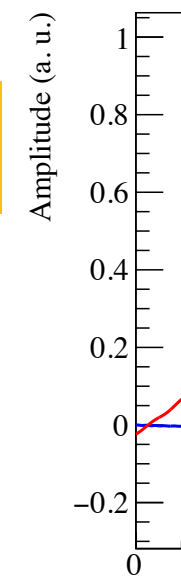
Calibrate calorimeters using known peaks



In the 2TY dataset, we built upon these techniques for keV-scale searches such as direct dark matter detection

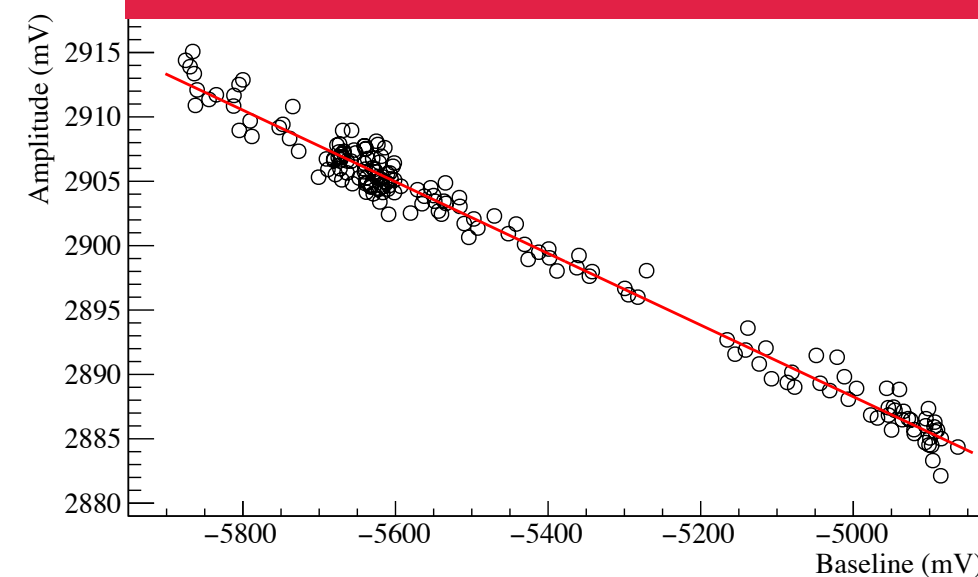
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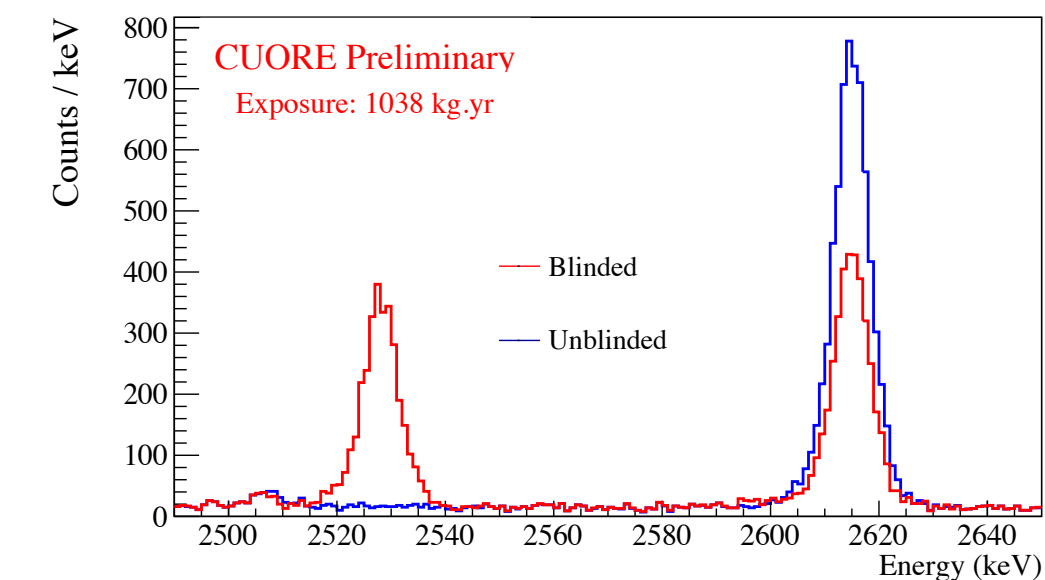
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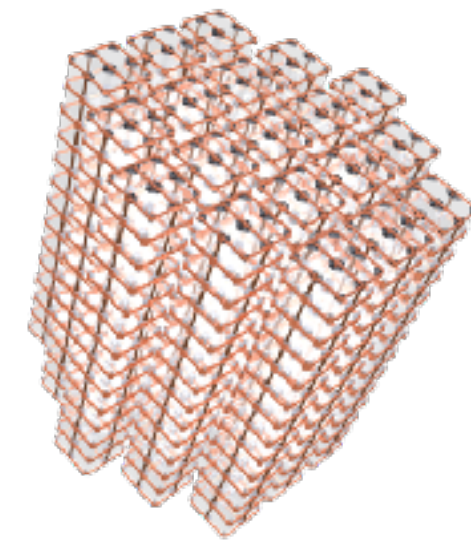


Tools for keV-scale searches

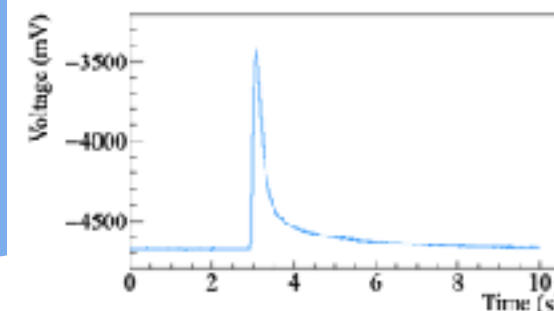


CUORE Data

Optimized detector selections



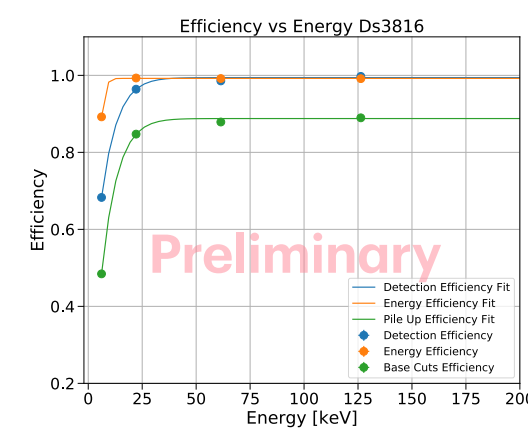
Event level cuts



New filters, de-noising



Calibrations, efficiencies evaluations



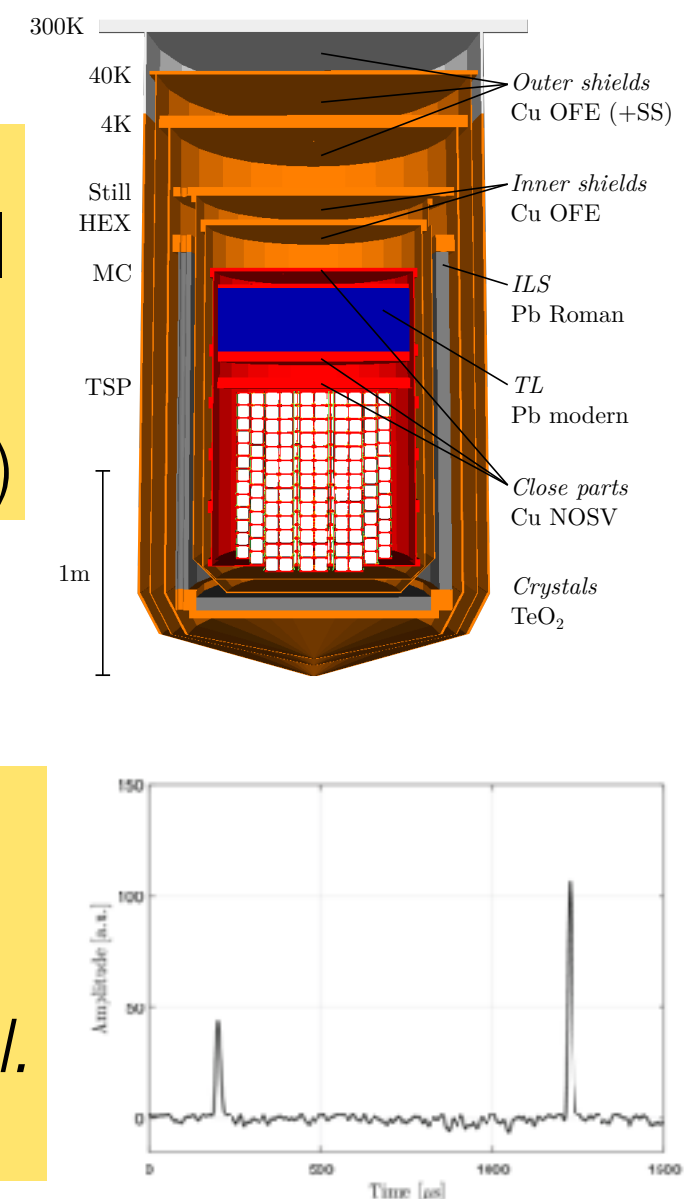
Simulations

Monte Carlo Background Model

Adams, D.Q. et al. (CUORE Collaboration), *arXiv:2405.17937* (2024)

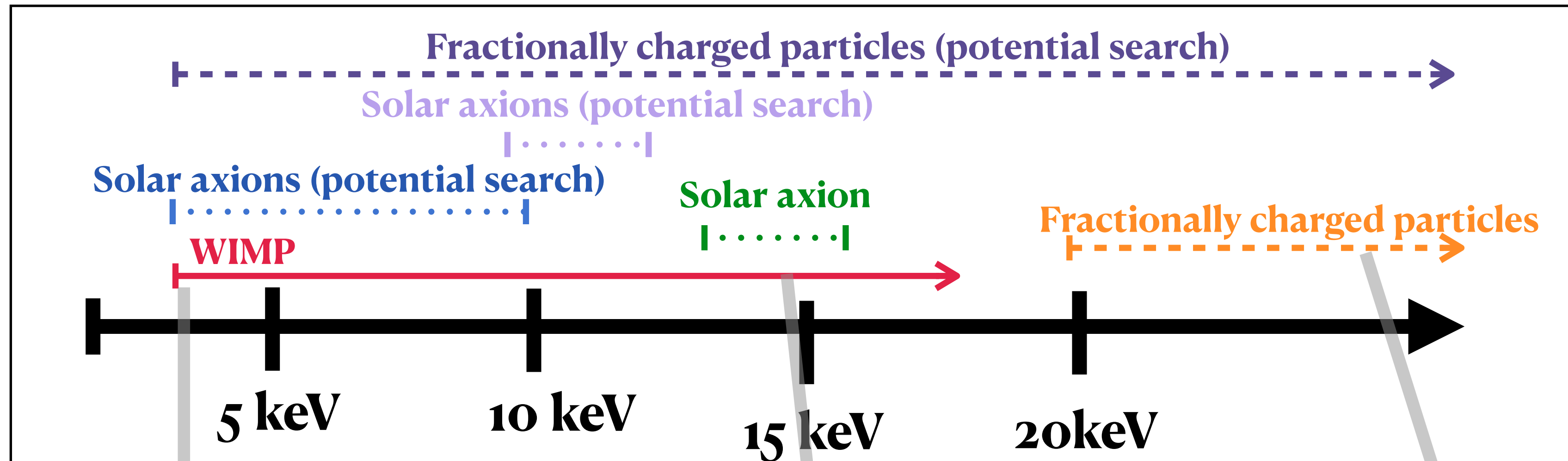
Mock Data Stream

Beretta, M., Biassoni, M., Gironi, L. *et al.* *Eur. Phys. J. Plus* **136**, 89 (2021)



Building upon: Low energy analysis techniques for CUORE. *Eur. Phys. J. C* **77**, 857 (2017).

Status of low energy analyses



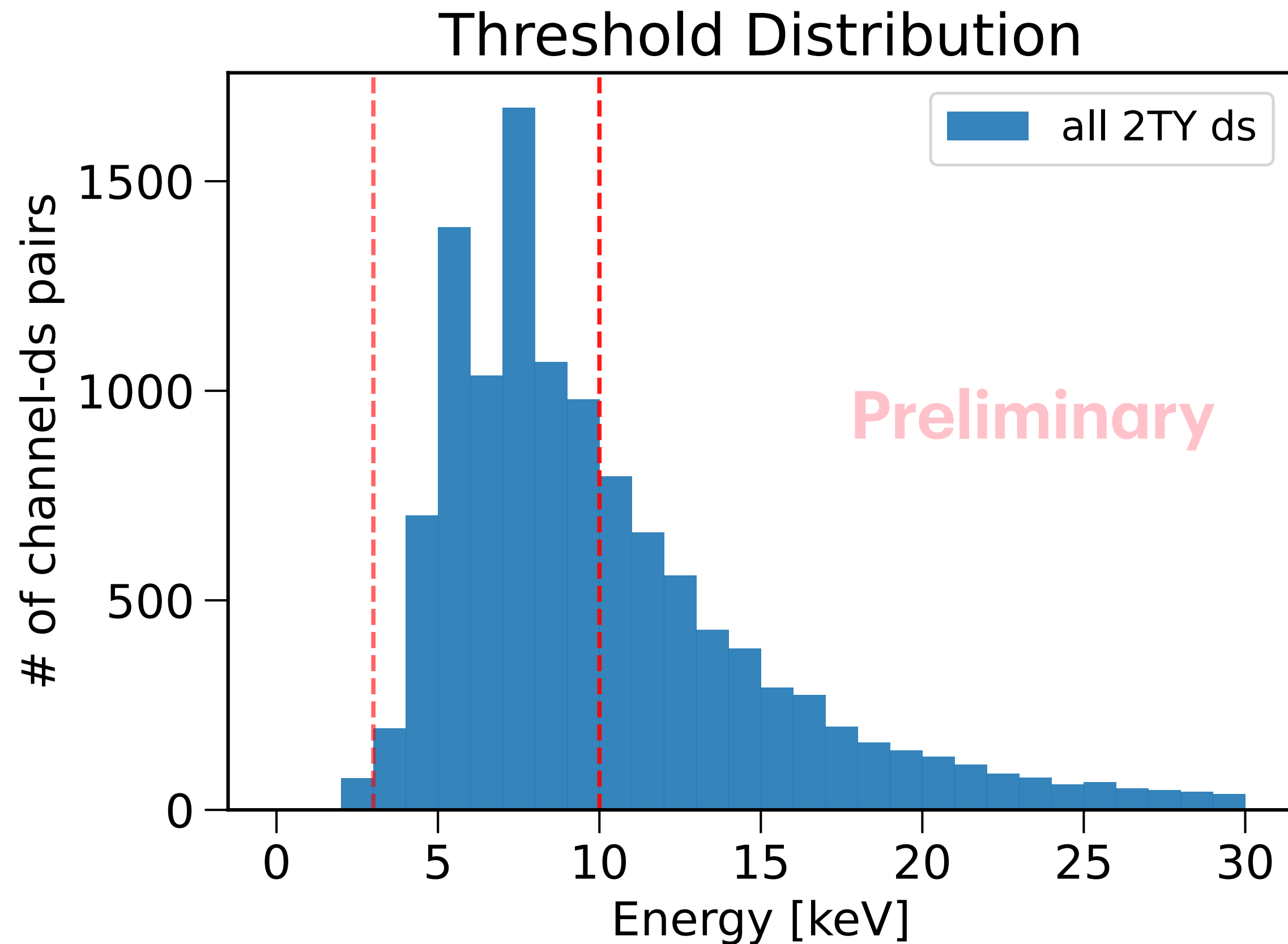
On 2TY data
In progress

On 2TY data
In progress



On 1TY data
Search for fractionally charged particles with CUORE. arXiv:2406.12380

Results of 2TY keV-scale processing



Thresholds of bolometer-dataset pairs in the 2TY dataset, method from [1]. Vertical lines at 3 keV and 10 keV.

[1] S Di Domizio et al 2011 JINST 6 P02007

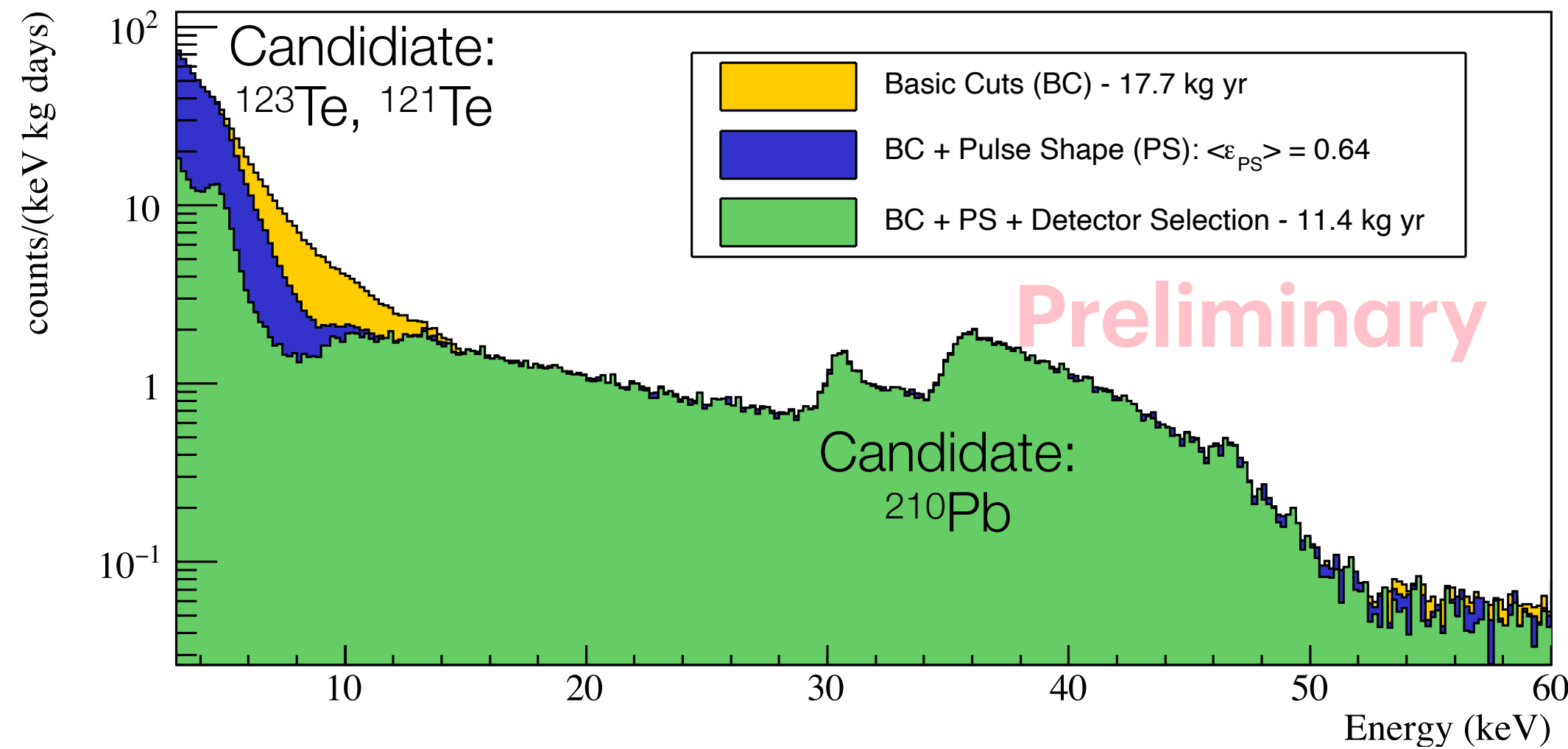
Preliminary

Optimized region of interest	[3,10] keV channels	[10,20] keV channels
Exposure of TeO ₂	11.4 kg·yr	691.0 kg·yr
FWHM Baseline (noise resolution)	1.2 keV	2.5 keV
FWHM Energy resolution at 30 keV from Te x-rays	1.6 keV	2.7 keV

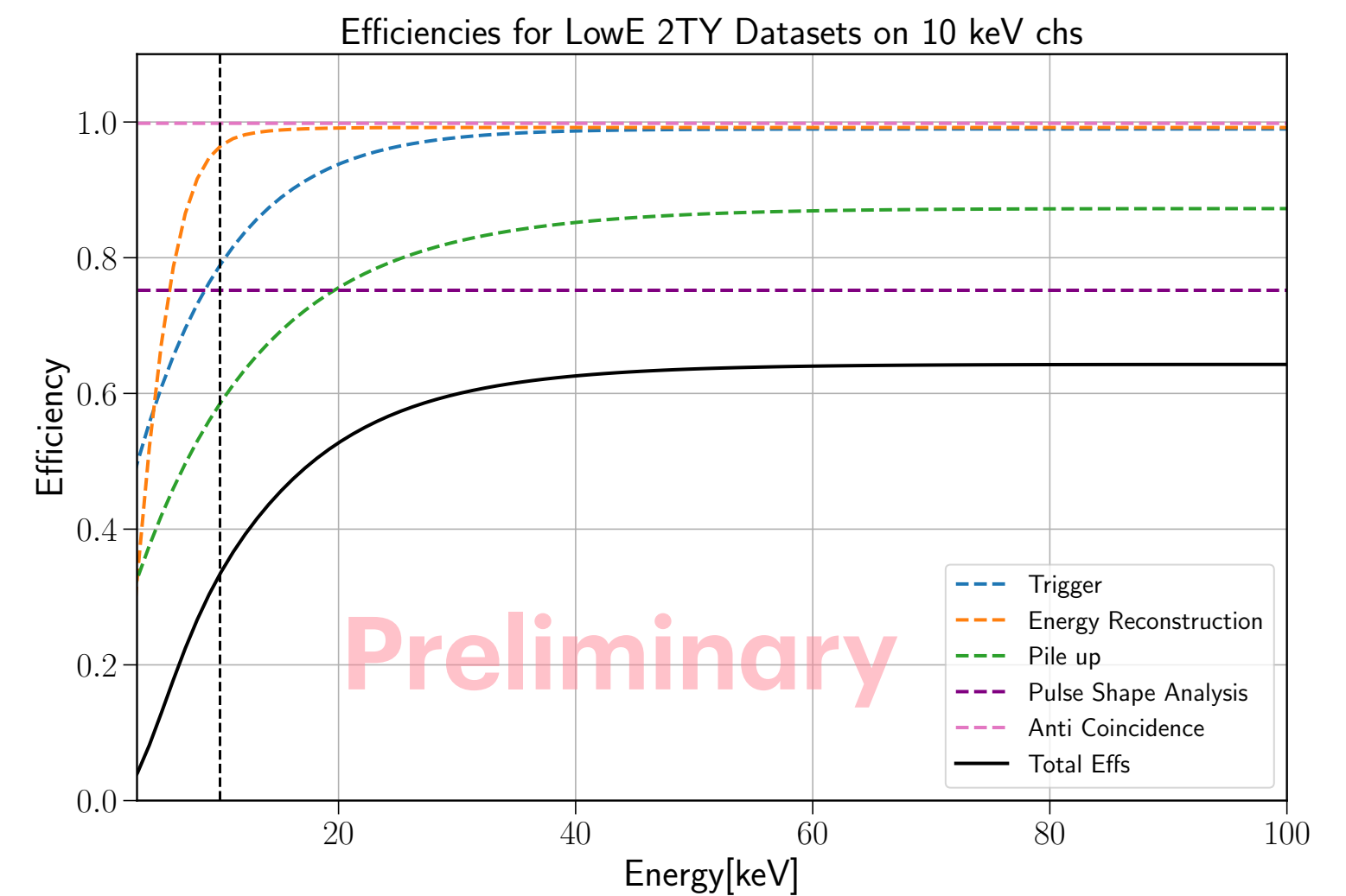
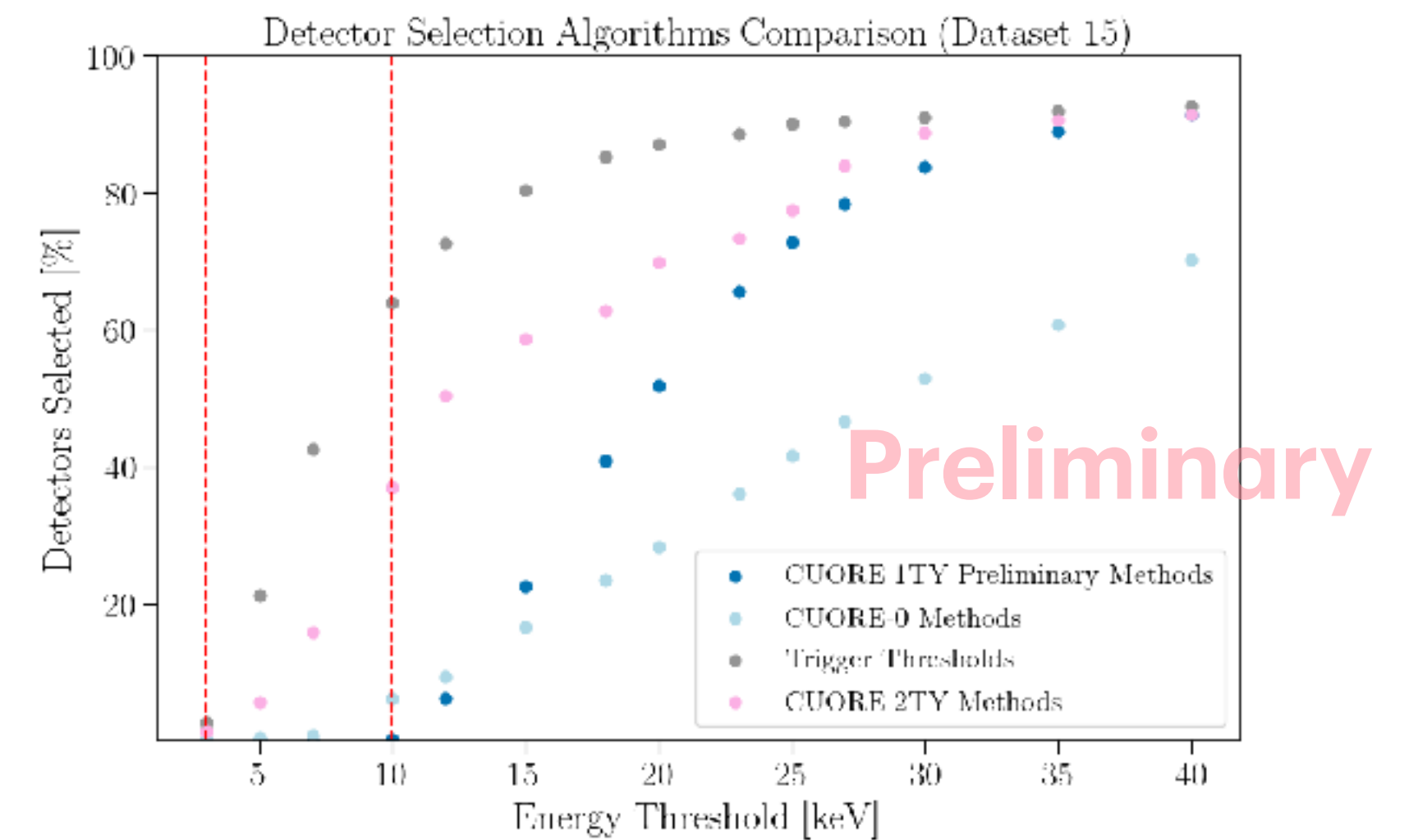
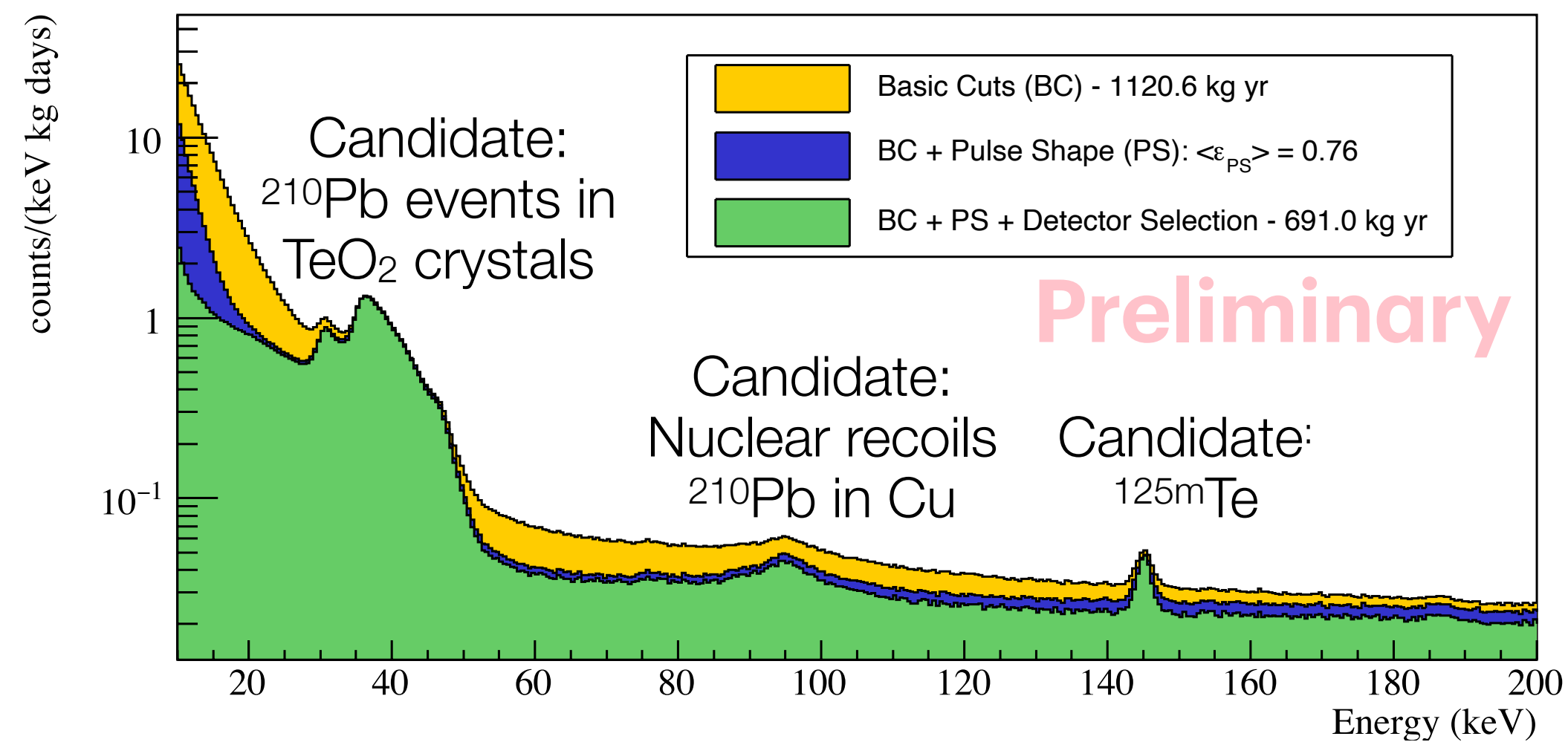
CUORE 2TY low energy spectrum



[3,10] keV dataset



[10,20] keV dataset

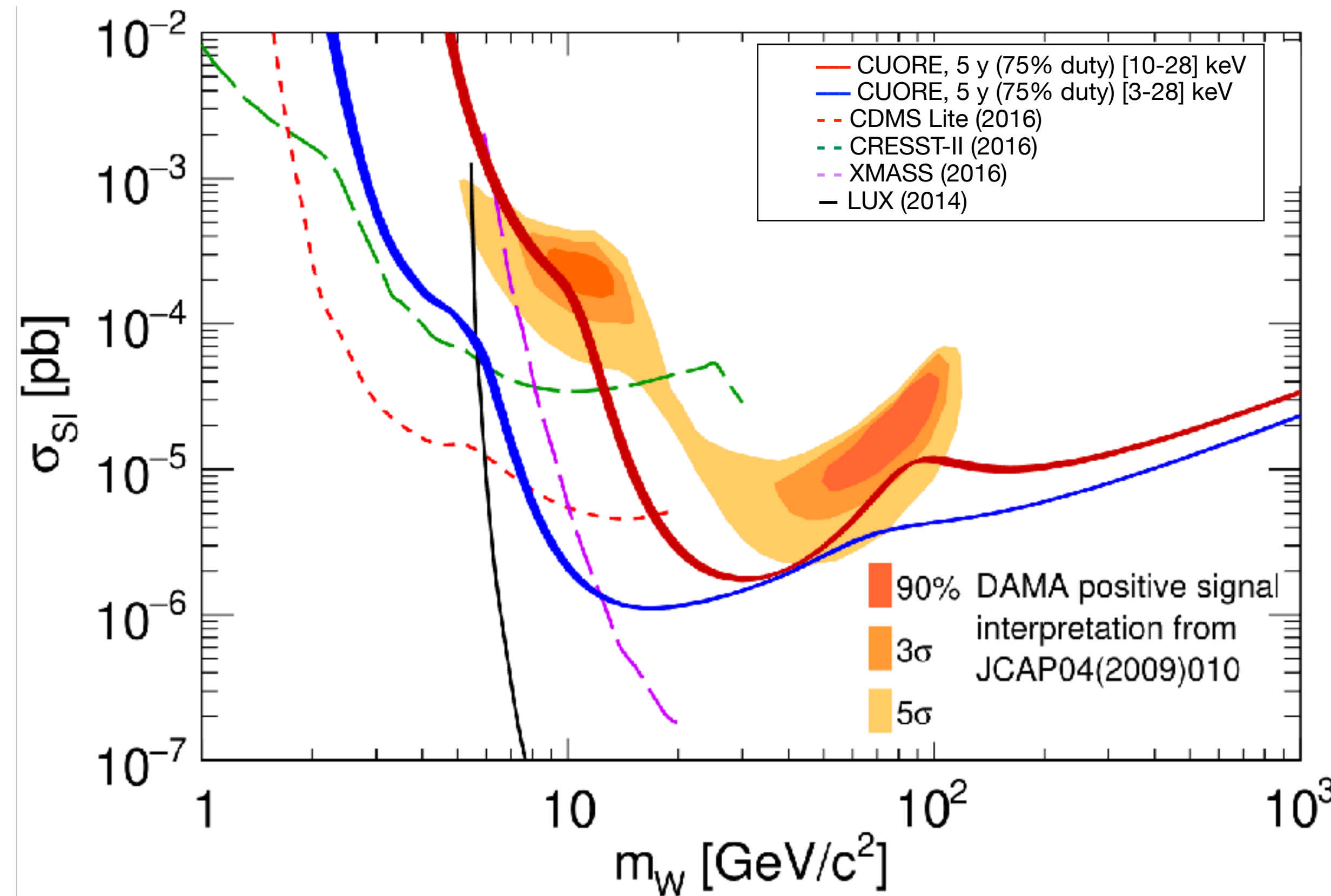


Low mass dark matter

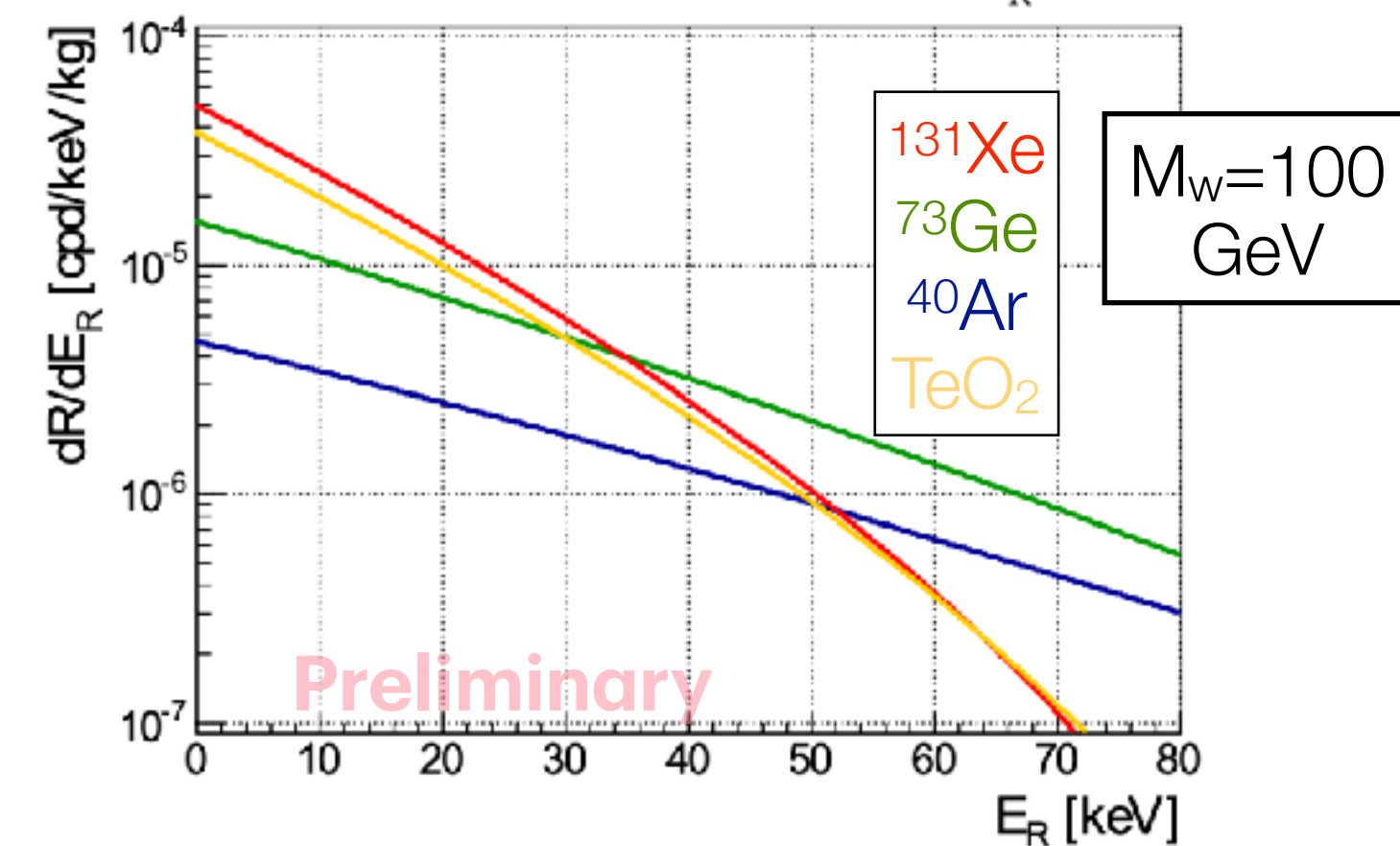
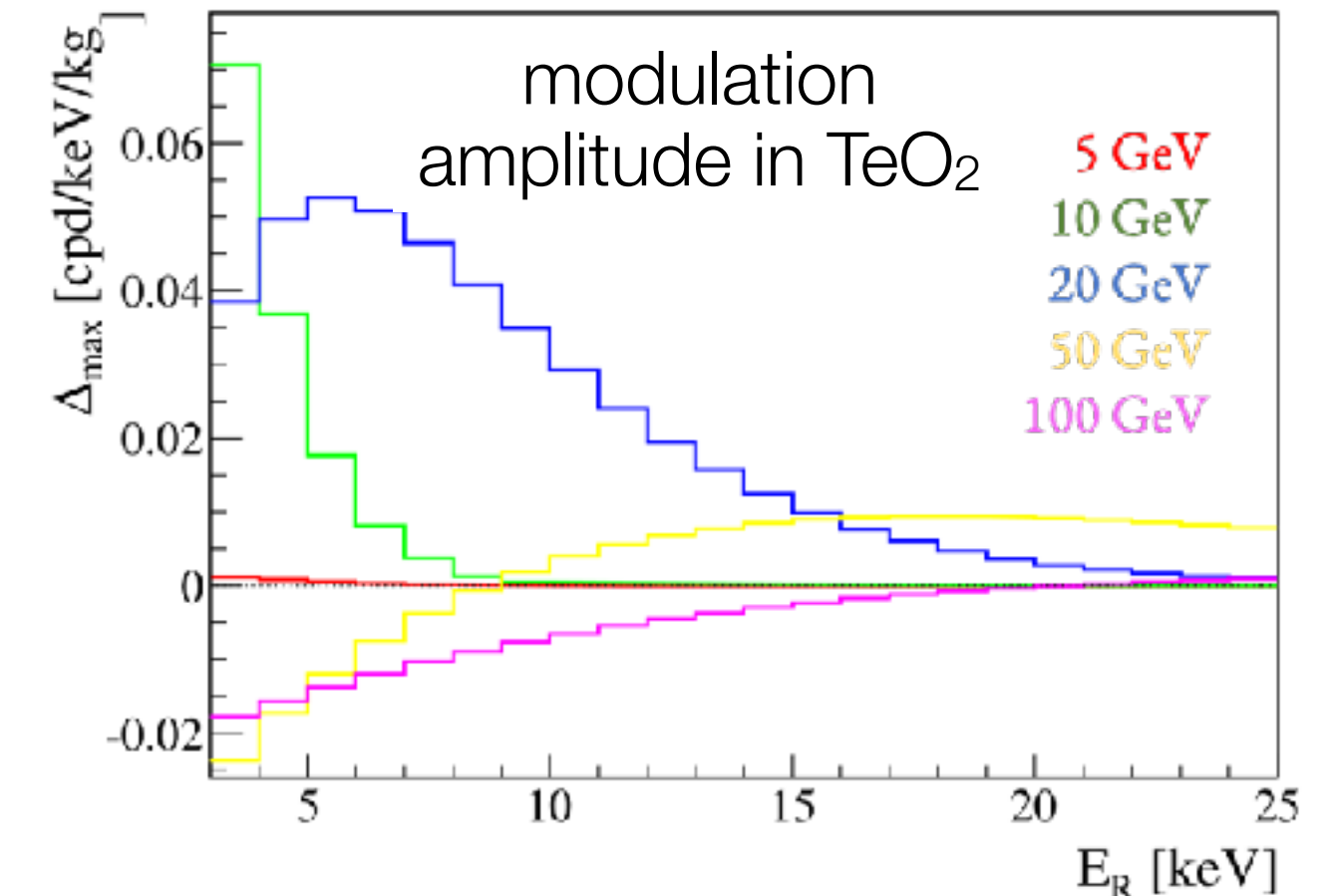
Weakly Interacting Massive Particles (WIMPs) Annual Modulation Search

TeO₂ target:
combines high
and low mass
nuclei

Sensitivity
estimated from
CUORE's expected
exposures using
results from
CUORE-o



Eur. Phys. J. C (2017) 77: 857



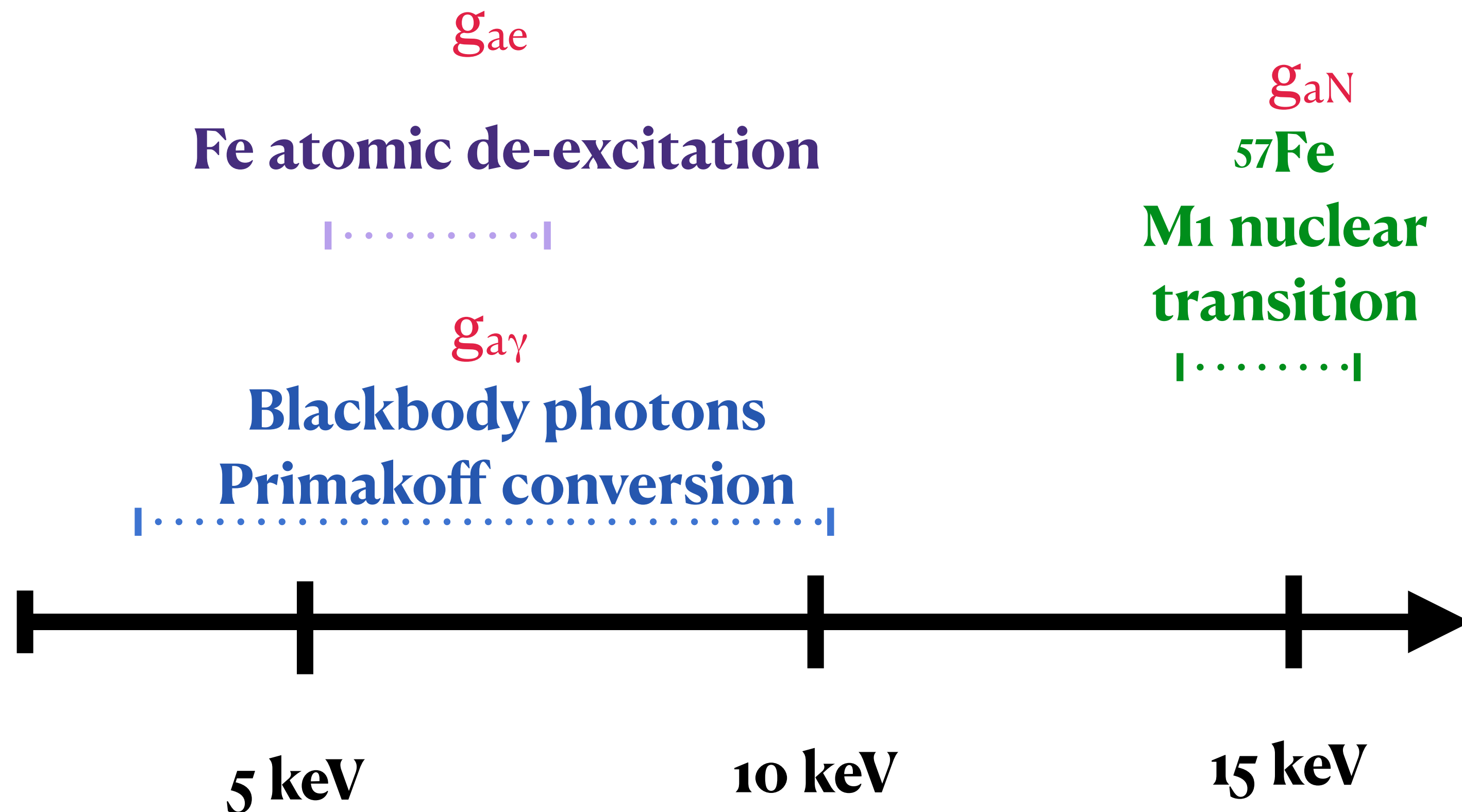
F. Alessandria et al JCAP01(2013)038



Solar axion-like particle searches

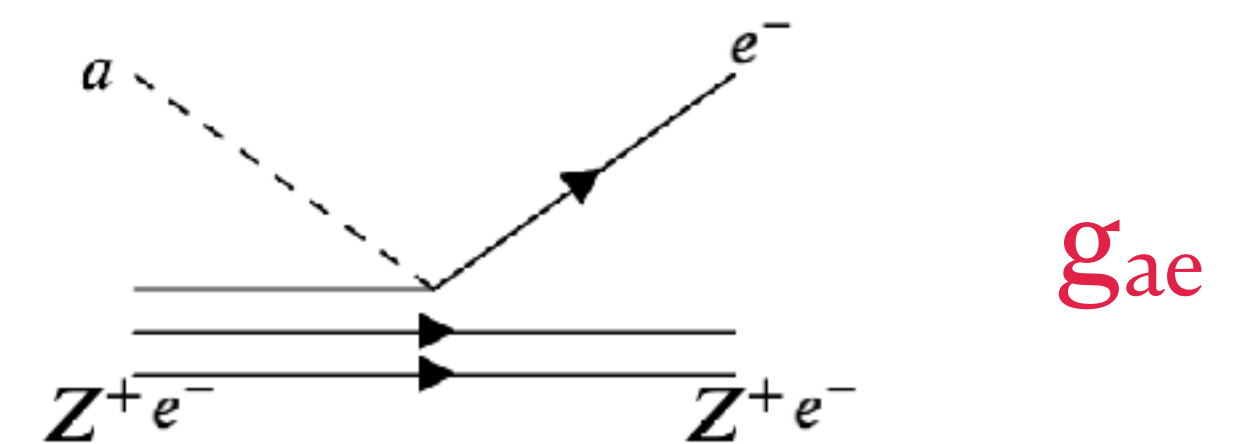


Search region of interest and production mechanism

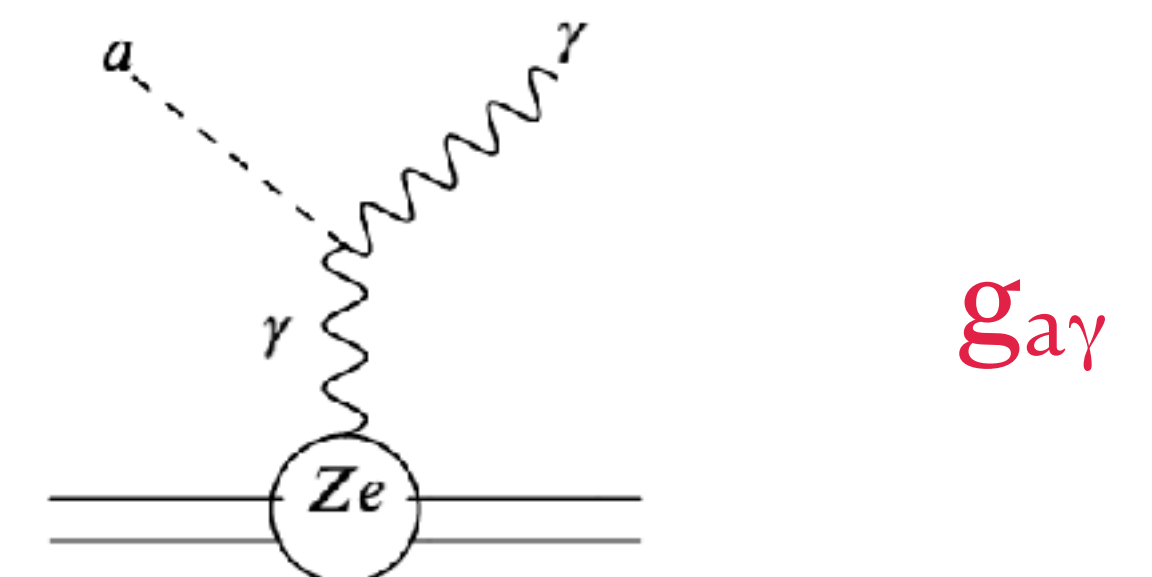


Detection in CUORE

Absorption
Axio-electric effect



Time dependent conversion
Inverse Coherent Bragg-Primakoff conversion

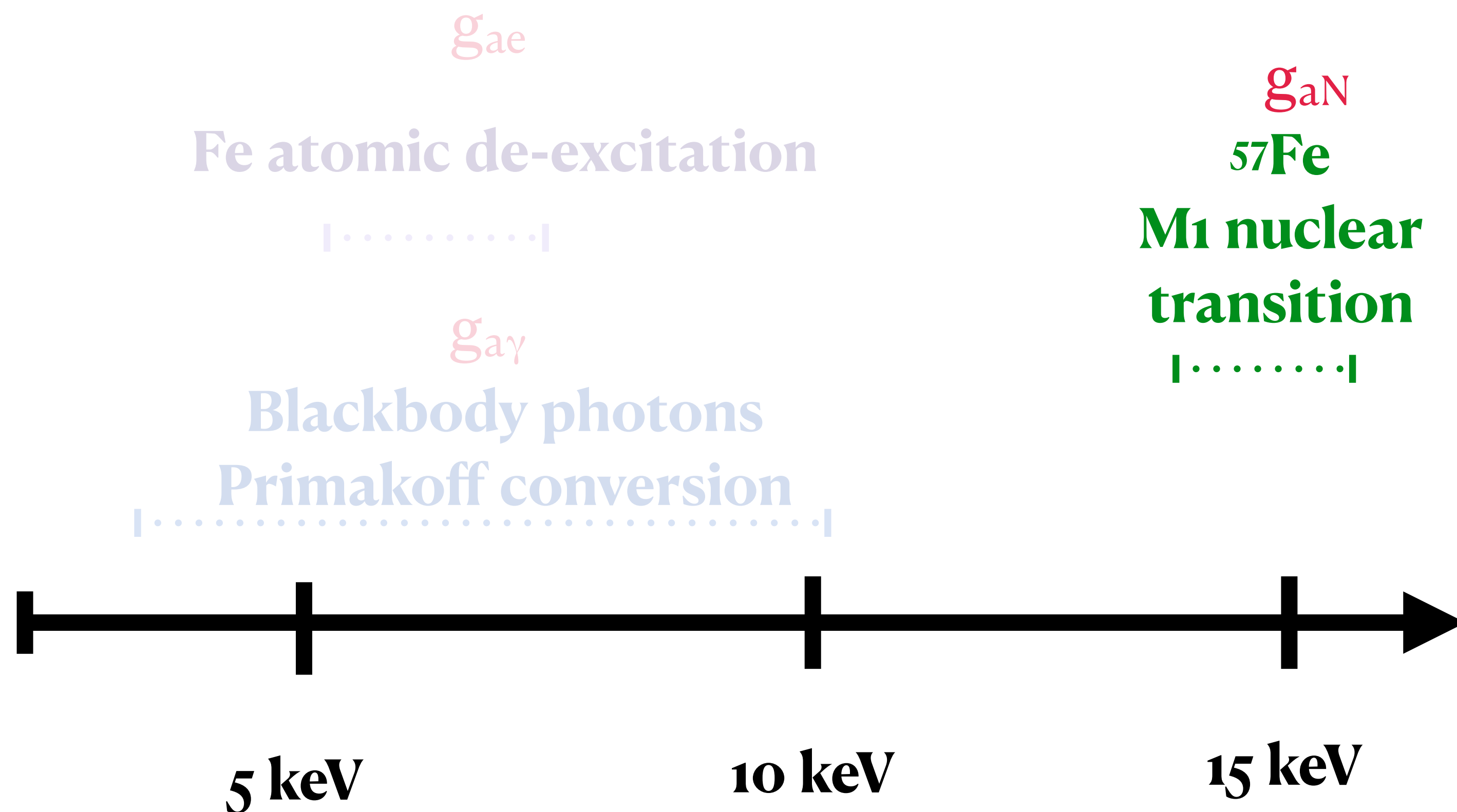




Solar axion-like particle searches

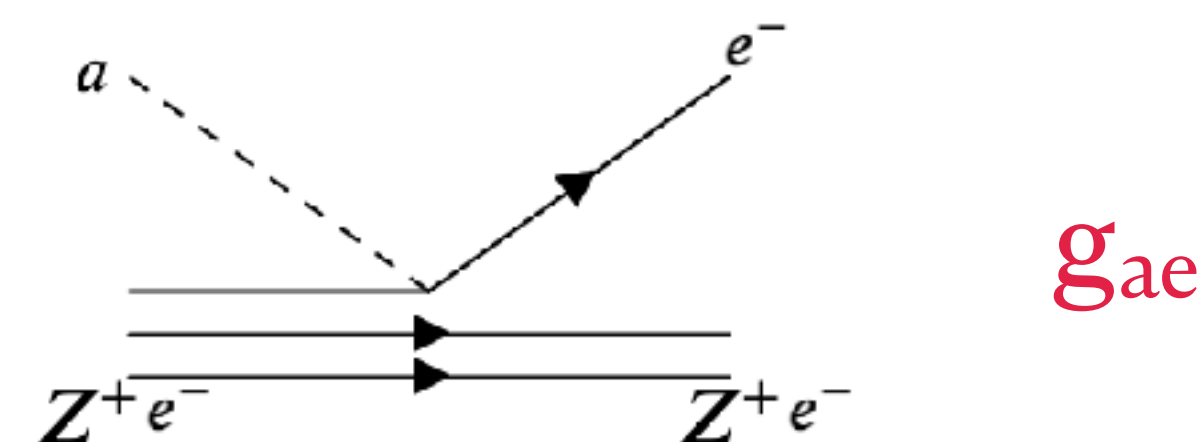


Search region of interest and production mechanism



Detection in CUORE

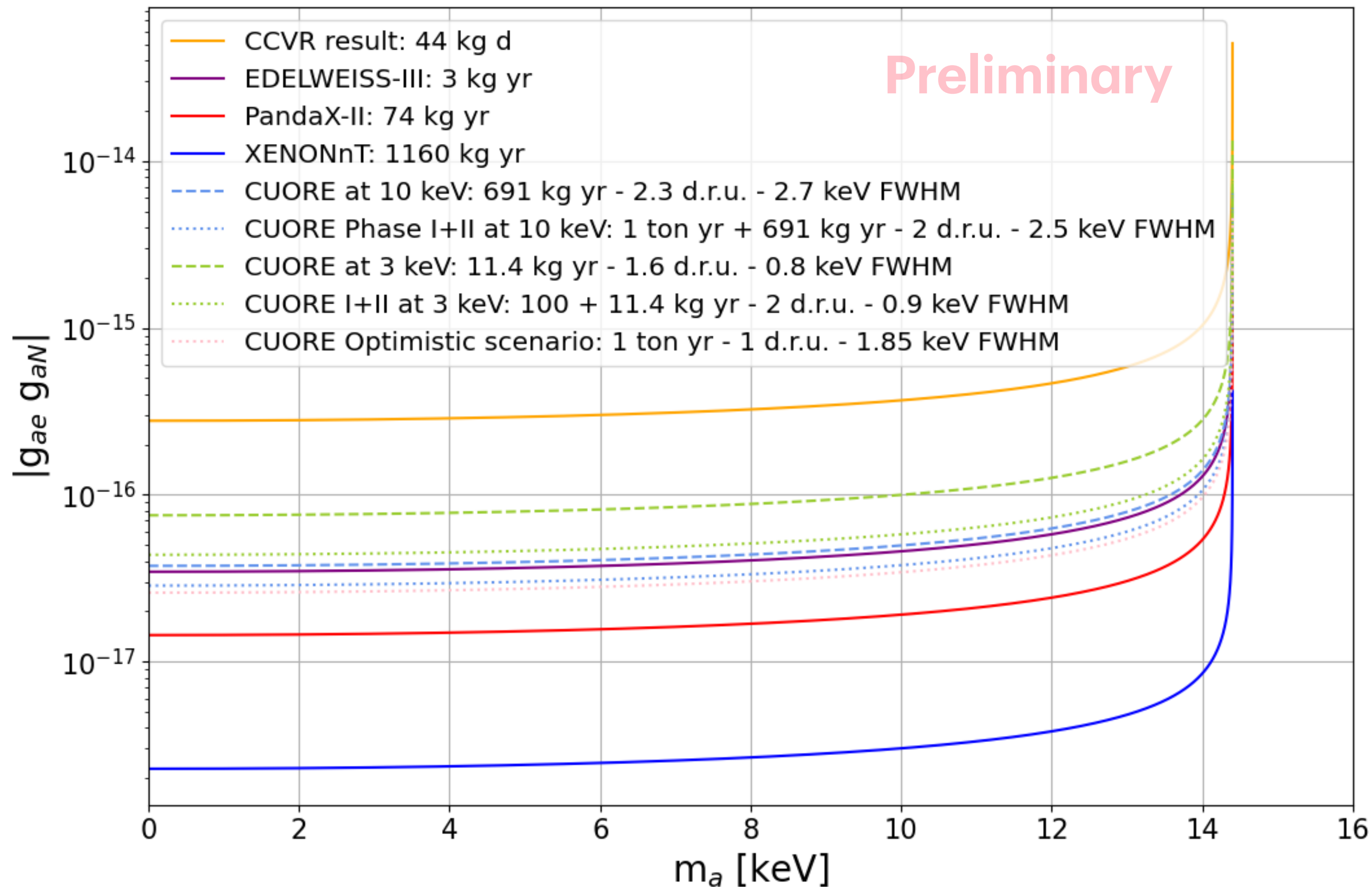
Absorption
Axio-electric effect



Time dependent conversion
Inverse Coherent Bragg-Primakoff conversion



M1 solar axion search sensitivity



CUORE's sensitivity to the M1 solar axion search based on 2TY processing and findings

Included scenarios of upgrades to CUORE for future dedicated low energy runs

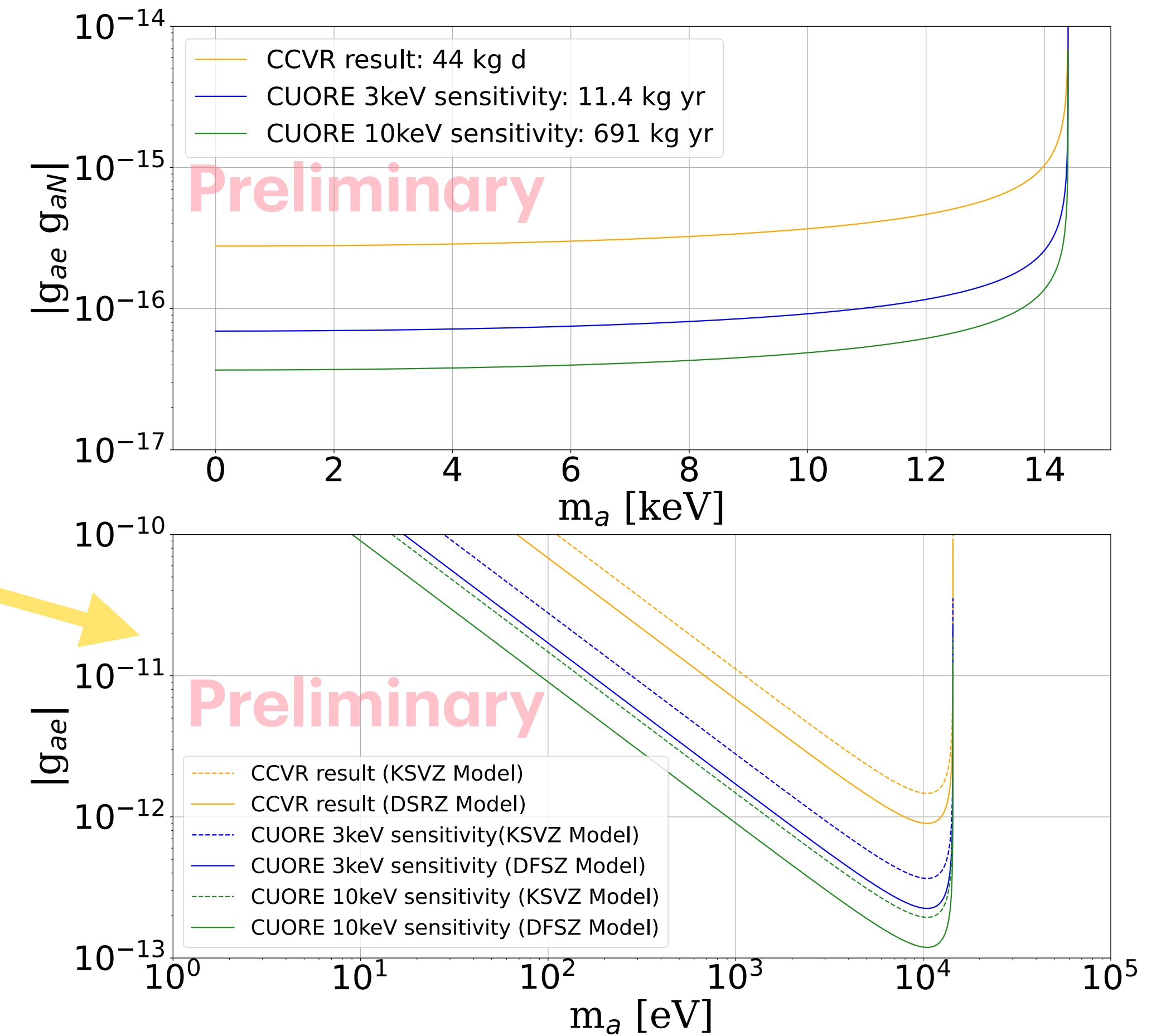
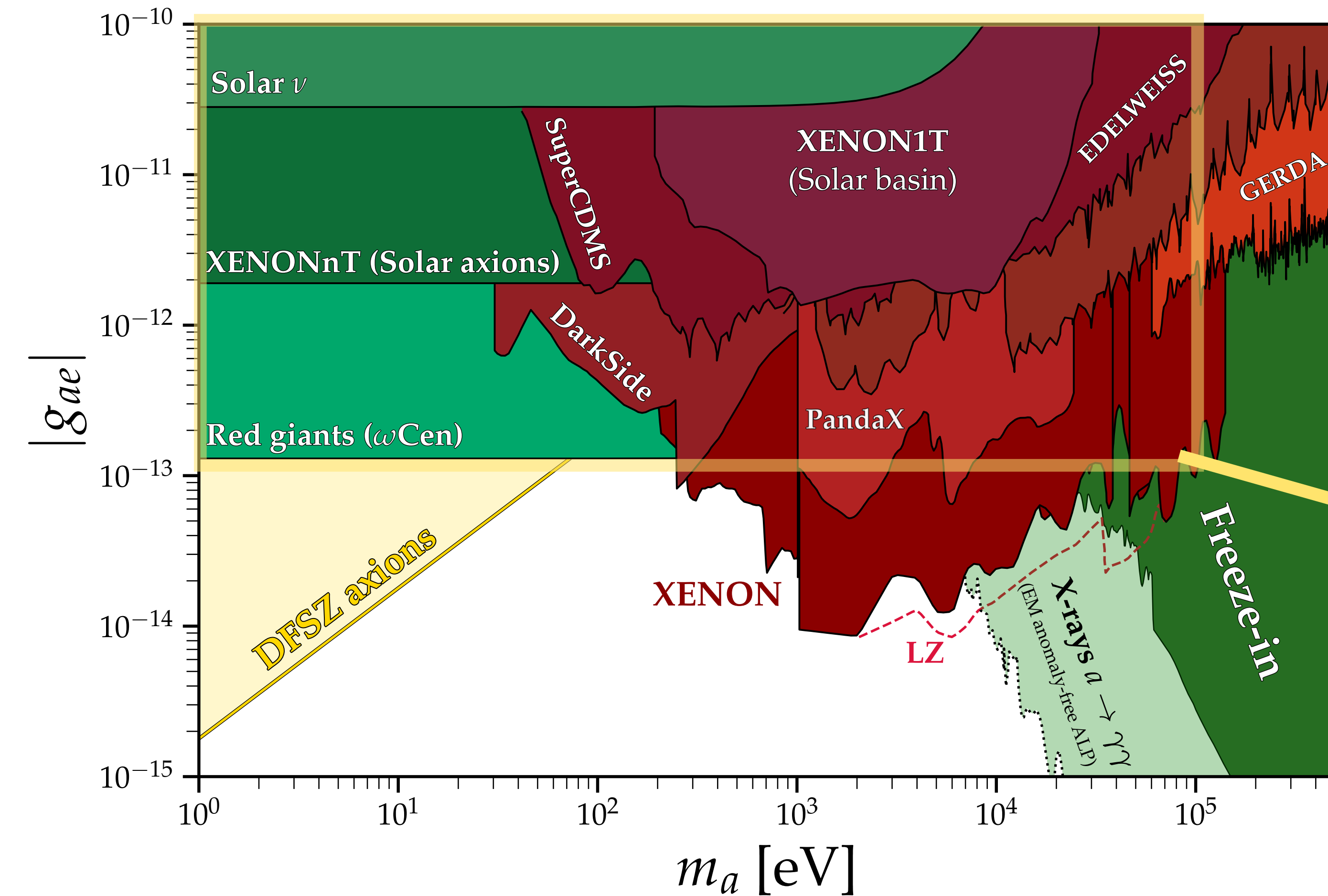
CCVR: JCAP05(2013)007

EDELWEISS-III: Phys. Rev. D 98, 082004

PandaX-II: Phys. Rev. Lett. 119, 181806

XENONnT: Phys. Rev. Lett. 129, 161805

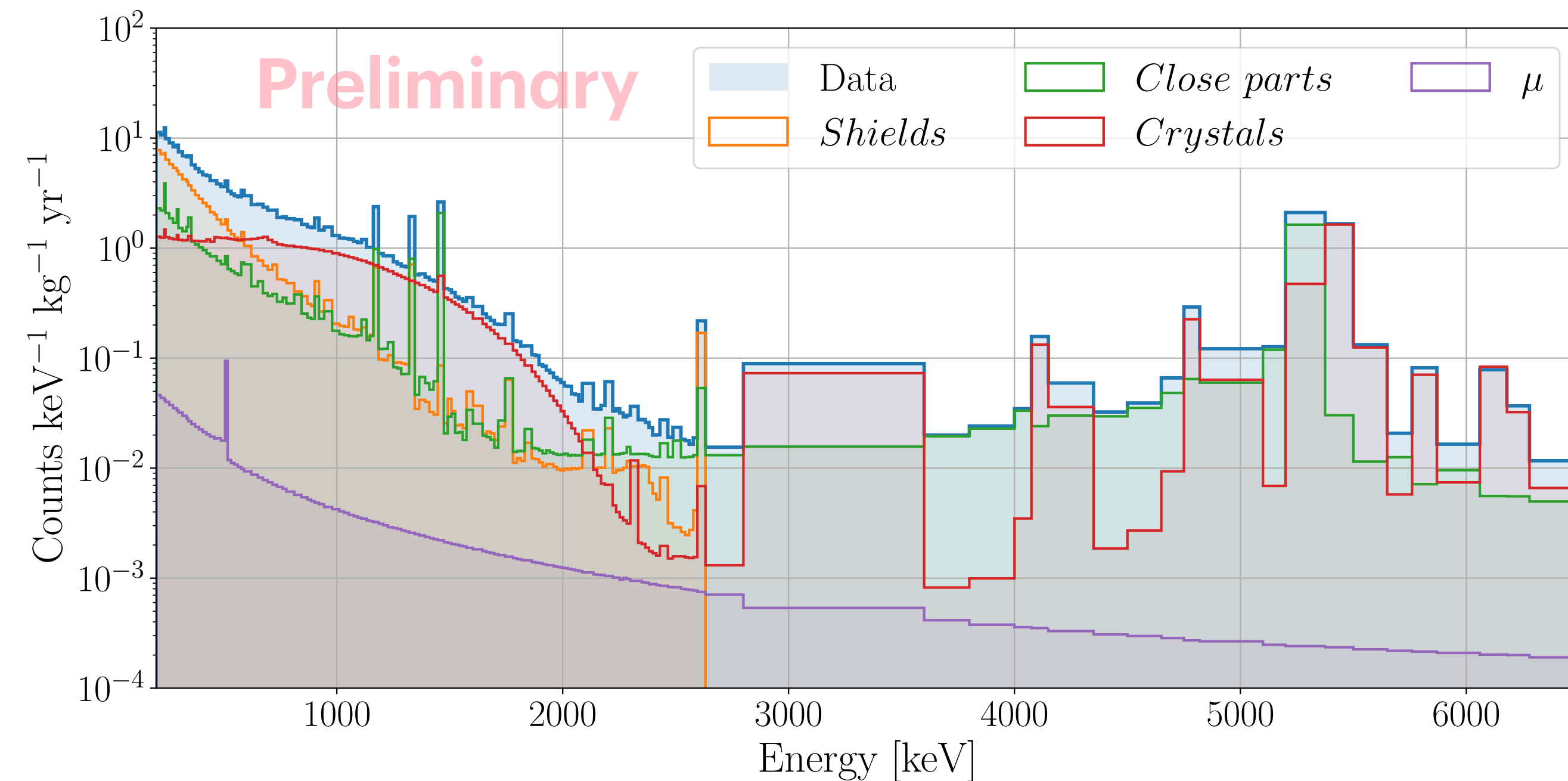
M1 solar axion search sensitivity



Ciaran O'Hare, <https://cajohare.github.io/AxionLimits/>

CCVR: JCAP05(2013)007

CUORE keV-scale background model



Data-driven background model for the CUORE Experiment
arXiv:2405.17937

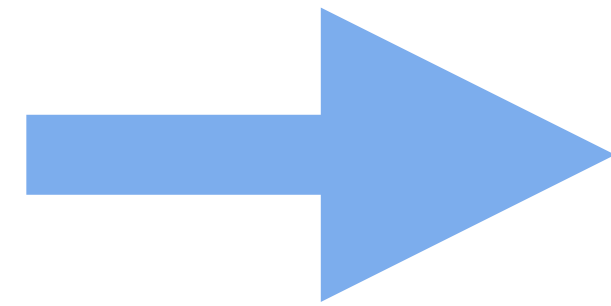
Geant4 simulation combined with CUORE detector response

Over 60 simulated sources: Including $2\nu\beta\beta$ decay in the crystals and contaminants

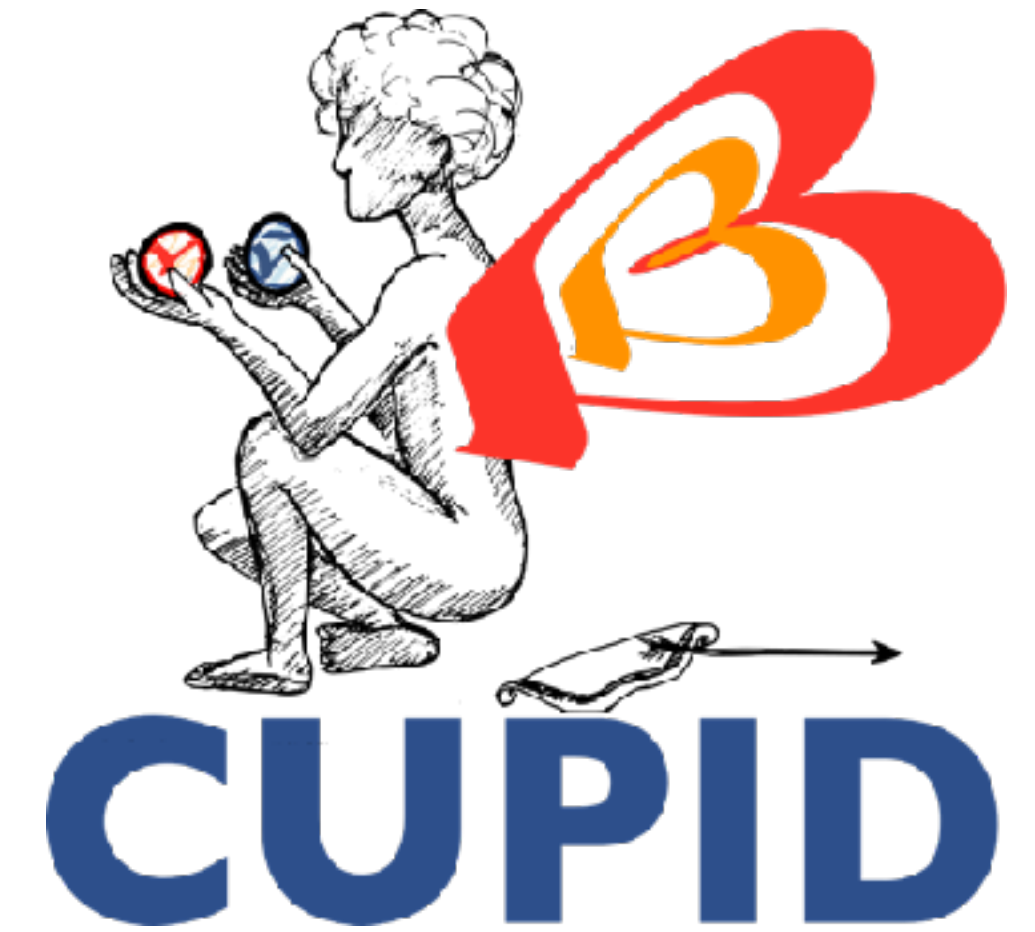
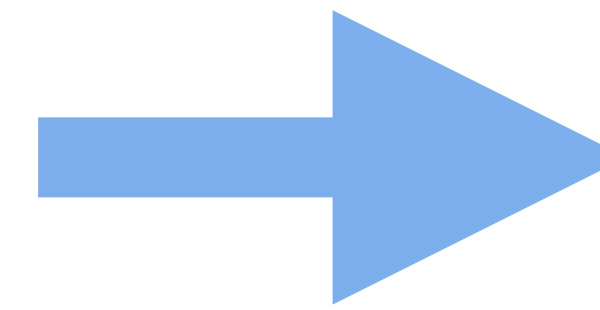
Produces multiplicity, time resolution, energy dependent trigger efficiencies

Work is in progress to expand to the keV-scale

What comes after CUORE?



CUORE Low Energy Runs



Continue successful data taking campaign to reach 3 tonne·yr of TeO_2 exposure

Proposed low-energy focused improvements
Added noise mitigation and detector upgrades

Upgraded Technology/Systems
 $\text{Li}_2^{100}\text{MoO}_4$ scintillating cryogenic calorimeters
Particle identification
Muon veto
Noise mitigation

Conclusions

- CUORE has been stably taking data since 2019, and has collected over 2 tonne·yr of TeO_2 exposure.
- CUORE has produced several results and world-leading limits in the search for neutrinoless double beta decay. We are now pursuing work to extend analyses to lower energies.
- The 2 tonne·yr reprocessing included new variables and methods developed for keV-scale searches.
- We've completed this processing and are studying our energy spectrum below 100 keV down to thresholds as low as 3 keV.
- Dark matter searches are in progress for solar axions and WIMPS using this dataset!
- Dedicated low-energy focused runs are planned for after the completion of CUORE's data taking with enhanced vibration mitigation and other possible improvements.