To detect dark matter with gravitational wave interferometers



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Cosmic WISPers WG4 meeting

Mar 29, 2023



EUROPEAN UNION European Structural and Investment Funds Operational Programme Research, Development and Education



The spin-2 field is special:

$$S_{
m int}[g,M_{ij},\Psi]\sim lpha\int\!{
m d}^4x\,\sqrt{-g}\,M_{ij}\,T_{\Psi}^{ij}$$

It is in fact an extension of general relativity
<> Bimetric gravity <<

S =

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 \Re R(f) is the Xicci for the metric $f_{\mu\nu}$, with strength αm_g

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This action contains **no ghosts**! ~ ~80 yrs to get right

The dark matter field is described by

$$M_{ij}(t) = \frac{\sqrt{2\rho_{\rm DM}}}{m}\cos{(mt + \Upsilon)\varepsilon_{ij}(r)}$$

where

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:: Perturb ::

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This looks like a continuous gravitational wave



Persistent

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- Continuous waves can be detected at much smaller sentitivity Thanks to a longer integration time and $h_0 \propto T_{obs}^{-1/2} \sim T_{obs}^{-1/4} T_{chunk}^{-1/4}$
- The overall magnitude scales as 1/m thanks to α C.f. the $1/m^2$ for spin-0 and spin-1 (without fifth forces)



:: references ::

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