### **Espresso Seminars**

# Theoretical Modelling of ECR Plasmas for Application to PANDORA

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## Talk Overview

#### Nuclear Astrophysics

**R-** and **S-**processes

Role of stellar  $\beta$  decay rates and heavy element opacities

#### ECR Plasmas as Stellar Testbench

Physics and properties of ECR plasmas

Measurement of in-plasma decay rates using  $\gamma$ -tagging

Plasma diagnostics

#### **Theoretical Modelling**

Interplay between plasma environment and nuclear decay rates

Models as a tool to complement γ-tagging results

Extend theory to stellar interiors

Extract plasma parameters from diagnostics data

Advance fundamental research into ECRIS operation

## **Elements of Nuclear Astrophysics**



Nucleosynthesis chart showing neutron-capture processes [1]

## **Elements of Nuclear Astrophysics**



[3] S. Taioli et al., ApJ 933, 158 (2022)

## **Electron Cyclotron Resonance Plasmas**



## **Properties of ECR Plasmas**

Electrons are heated to high energies ( $k_B T_e$  between 0.1 - 100 keV)



### PANDORA

Plasmas for Astrophysics, Nuclear Decay Observations and Radiation for Archaeometry



Two grand classes of in-plasma phenomena to be studied:

- $\beta$ -decay rates of radioisotopes in stellar atmospheres for s-process branching
- opacity of heavy elements in early-stage KN for r-process nucleosynthesis

## $\gamma$ -Tagging and T<sub>1/2</sub> Measurement

ECR magnetoplasma can be maintained in MHD equilibrium for days or even weeks







Is o-significance contour plots for <sup>176</sup>Lu isotope showing measurement time required to overcome background, as a function of lifetime and significance level [6]

Numerical simulations to determine detection efficiency according to chosen plasma model - 14 HpGe detectors [5]

## **Plasma Diagnostics**

#### Plasma Emitted Radiation



Diagnostic tool	Sensitive Range	Measurement	Resolution & Meas. Error
SDD	1.0 ÷ 30 keV	Volumetric soft X-ray Spectroscopy:	Res. ~ 120 eV
		warm electrons temperature and density	$\epsilon_{n_e} \sim 7\%, \epsilon_{T_e} \sim 5\%$
HpGe	30 ÷ 400 keV	Volumetric hard X-ray Spectroscopy:	Res. ~ 200 eV
		hard electrons temperature and density	$\epsilon_{n_e} \sim 7\%, \epsilon_{T_e} \sim 5\%$
Visible Light Camera	1.0 ÷ 12 eV	Optical Emission Spectroscopy:	$\Delta \lambda = 0.04$ nm
		cold electrons temperature and density	R=12500
Microwave Interferometer	K-band	Interferometric measurement:	$\epsilon_{n_e} \sim 50\%$
	18 ÷ 26.5 GHz	line integrated total density	
Microwave Polarimeter	K-band	Faraday-rotation measurement:	$\epsilon_{n_e} \sim 25\%$
	18 ÷ 26.5 GHz	line integrated total density	
X-ray pin-hole camera	2 ÷ 15 keV	2D Space-resolved spectroscopy	Energy Res. ~ 0.326 keV
		soft X-ray Imaging and plasma structure	Spatial Res. ~ 0.56 mm
Multi-pins RF probe +	10 ÷ 26.5 GHz	Frequency-resolved Spectroscopy	SA Resolution bandwidth:
Spectrum Analyzer (SA)	(probe)	plasma emitted EM wave in GHz range	RBW = 3 MHz
Multi-pins RF probe	10 ÷ 26.5 GHz	Time-resolved X-ray Spectroscopy	80 Gs/s (scope)
+ Scope + HpGe	(probe)		time scales below ns

## Plasma Diagnostics – Warm Electrons



## ECR Plasma Modelling – Why?



Mode-like EM field structure in plasma chamber under vacuum [10]

3.5

4.5

3D structure of electrons from simulations – the spatial anisotropy and non-homogeneity with respect to density and energy are visible [11]

[9] B. Mishra *et al*, Frontiers in Phys., special issue (2022)
[10] G. Torrisi *et al*, JEWA 28, 9 (2014)
[11] R. Racz *et al*, Plasma Sources Sci. Technol. 26, 075011 (2017)

3

1.5

2

2.5

[12] B. Mishra, EPJ Web of Conferences, 2022.

## ECR Plasma Modelling – Why?

ECR plasmas properties are non-uniform. Models can:

- (1) Help calculate space-resolved decay rates and opacities
- (2) Connect laboratory and stellar plasma theory
- (3) Complement γ-tagging results
- (4) Couple with diagnostics for extracting plasma parameters
- (5) Improve fundamental understanding of ECRIS devices



Mass of neutral atom at ground state

Kinetic energy of continuum electron





The lepton phase volume quantifies the number of ways a decay can occur. The phase volume changes with variations in atomic configuration, depending on type of decay

Information needed: level probability distribution (LPD), orbital occupancy, orbital electron wavefunction, decay energy and shape factor



## In-Plasma $\beta$ -Decay Rates $Q = Q_0 + (E_{X,K}^* - E_{Y,K'}^*) + (\epsilon^{i,j} - \epsilon^{i',j'}) + (\Delta_X - \Delta_Y)$

The decay energy depends on not just the difference in nuclear masses, but on the overall system energy which includes atomic/ionic energy

Information needed: energy of different atomic configurations of parent system and coupling with daughter system



All configurations should have at No selection on daughter configuration least one K-shell electron K-shell vacant states autoionising

#### **K-SHELL CAPTURE**

All configurations should have at least one L-shell electron

No selection on daughter configuration L-shell vacant states autoionising

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<sup>7</sup>Be decay Q-value as a function of charge state and level for neutral and 1+ (left) and 2+ and 3+ (right)

Conservation of total angular momentum implies that only certain electron orbitals can interact with the nucleus, depending on the spin and parity of the decay

Information needed: spin-parity of electron orbitals and decay transition

$$S_{(m)x} = \begin{cases} 1 & \text{for } m = a, nu \text{ and } x = ns_{1/2}, np_{1/2} \\ q^2 & \text{for } m = u \text{ and } x = ns_{1/2}, np_{1/2} \\ 9/R^2 & \text{for } m = u \text{ and } x = np_{3/2}, nd_{3/2} \\ 0 & \text{otherwise.} \end{cases}$$

The probability of electron capture from bound states depends on the square of the radial component of the orbital wavefunction evaluated on the nuclear surface

Information needed: formalism for radial wavefunctions of different orbitals





Larger of  $f_x^2$  or  $g_x^2$  as calculated for <sup>7</sup>Be taking  $R = R_0 A^{1/3}$ 



The ion CSD and LPD strongly depends on electron density and temperature

Information needed: CSD and LPD of <sup>7</sup>Be for various  $n_e$  and  $T_e$ (calculated using FLYCHK)





LPD of <sup>7</sup>Be<sup>0+</sup> and <sup>7</sup>Be<sup>3+</sup> for different temperatures as calculated by FLYCHK (no effect of density)

CSD and LPD calculated using grid of density and temperature values in FLYCHK under LTE approximation



CSD of <sup>7</sup>Be for different temperatures as calculated by FLYCHK (no effect of density)





#### CALCULATING THESE IN ECR PLASMA IS THE ULTIMATE GOAL!

### Self-Consistent Steady-State Electron Simulations



[13] D. Mascali *et al*, EPJ D, 69:27 (2015)
[14] A. Galatà *et al*, Plasma Sources Sci. Tech. 25, 045007 (2016)



 $W_3 = N_{L3}/N_{tot}$ 







 $N_{1} = K_{3}(1 - k_{1 \to 2} + 2k_{1 \to 2}k_{2 \to 1} + k_{1 \to 2}k_{2 \to 3}k_{3 \to 2})n_{1}$   $N_{2} = K_{3}k_{1 \to 2}(1 - k_{2 \to 3} - k_{2 \to 1} + k_{2 \to 3}k_{3 \to 2})n_{2}$   $N_{3} = K_{3}k_{1 \to 2}k_{2 \to 3}(1 - k_{3 \to 4} - k_{3 \to 2})n_{3}$ 

$$\rho_{\Delta} = e[n_e - (N_1 + 2N_2 + 3N_3)]$$



[9] B. Mishra *et al*, Frontiers in Phys., special issue (2022) [12] B. Mishra, EPJ Web of Conferences, 2022.





ECR plasmas models can advance fundamental research in ECRIS operation

### **Isotope-Plasma Simulations**



### Soft X-Ray Diagnostics Models



[20] B. Mishra *et al*, Phys. Plasmas, (2021)
[21] B. Mishra *et al*, Condens. Matter, (2021)
[22] S. Biri *et al*, JINST 16, P03003 (2021)

### Soft X-Ray Diagnostics Models



### **Conclusions and Future Perspectives**

ECR plasmas can be reliable systems for emulating stellar environments and reactions thereof

Experimental activities can be complemented by ECR plasma models for optimised analysis

Self-consistent steady state routines developed to describe electrons, ions and radio-isotopes in plasma

In-plasma decay rate formalism already developed

Plasma models (soft X-ray) can also be used to extract electron and ion parameters from spectra

Models useful not only for application like PANDORA but also for fundamental ECRIS research



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# THANK YOU FOR YOUR ATTENTION!

...the PANDORA collaboration, and everyone here whose work I have used in some way!

### Additional Content 1 – Soft X-Ray Spectroscopy

 $T = 300\rho_e \rho_i \omega_{K\alpha} \varepsilon_g \varepsilon_q \int_{-\infty}^{\infty} \sigma_{K,ion}(E) v(E) f(E) dE$ 

Evaluating 3D space-resolved electron energy distribution function (EEDF) and analysing behaviour of K-shell ionisation cross-section [6]









Evaluating 3D space-resolved geometrical efficiency and quantum efficiency of simulation domain [7]



### Additional Content 1 – Soft X-Ray Spectroscopy







#### Line Emissivity Density (Ar ions)

$$J_{theo,2.96} = rac{h 
u_{2.96}}{\Delta E} 
ho_e 
ho_i \omega_{2.96} \int_{3.205}^\infty \sigma_{K,ion}(E) v_e(E) f(E) \mathrm{d}E$$

$$J_{theo,3.19} = \frac{h\nu_{3.19}}{\Delta E} \rho_e \rho_i \omega_{3.19} \int_{3.205}^{\infty} \sigma_{K,ion}(E) v_e(E) f(E) dE$$

$$\sigma_{K,ion} = a_K q_K rac{\ln arepsilon/I}{arepsilon I} \{1 - b_K ext{exp}[-c_K(arepsilon/I - 1)]\}$$

Pseudo-Voigt profile for line broadening

$$D_{PV}(x-x_0,f) = \eta L(x-x_0, au_L) + (1-\eta)G(x-x_0,\sigma_G)$$

$$oldsymbol{V} oldsymbol{J}_{theo,line,plasma}(h
u) = [J_{theo,2.96}D_{PV}(h
u-2.96,f_{2.96})+ \ J_{theo,3.19}D_{PV}(h
u-3.19,f_{3.19})]\Delta E$$

4 +



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Plasmas for Astrophysics

rchaeometry

Nuclear Decay Observation and Radiation for

Lotz cross-section for Ar superposed on Maxwell EEDFs at different temperatures [6]

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#### **Electron Diagnostics – X-Ray Spectroscopy - Volumetric**



[6] B. Mishra et al, Phys. Plasmas, (2021)

Objective: Launch electron simulations for higher energy ranges and compare model-generated maps with experiment again

#### SECOND ATTEMPT DATA: Debrecen 2018







#### Qualitative EEDF fit of ROI 2 – best function Maxwellian+Druyvesteyn



#### **Electron Diagnostics – X-Ray Spectroscopy - Fluorescence**



Schematic of multi-collimator CCD-pinhole setup used in Debrecen 2018 [9]

collimator (top) and with multi-collimator (bottom) - enhanced screening and resolution
 3D space-resolved charge density and EEDF
 K-shell ionisation cross-section

Local geometrical efficiency calculated using the ray-tracing Monte Carlo method without multi-



Local geometrical efficiency

#### **Electron Diagnostics – X-Ray Spectroscopy - Fluorescence**



Grid of 1024 x 1024 points Map photon to nearest neighbour grid points (1st order approximation)



#### **Electron Diagnostics – X-Ray Spectroscopy - Fluorescence**





0<sup>th</sup> order projection

- 50



#### **Conclusions:**

New method for resolving plasma non-homogeneity established, allowing quantitative study of EEDF on a ROI-by-ROI basis

Match between experiment and model much better for Debrecen 2018 as compared to Debrecen 2014 – electron density on correct order of magnitude, general features reproduced

Some aspects not reproduced – notably the hole in the centre – can be attributed to incompleteness of electron simulations

Only 3 magnetic branches visible in experiment – may be attributed to ion density distribution, incomplete implementation of clusterisation



[9] S. Biri et al, JINST 16, P03003 (2021)

### Additional Content 2 – Level Grouping



### Additional Content 2 – Level Grouping

<sup>7</sup> Be <sup>0+</sup>	<sup>7</sup> Be <sup>1+</sup>	<sup>7</sup> Be <sup>1+</sup>	<sup>7</sup> Be <sup>1+</sup>	
Continuum	Continuum	Continuum	Continuum	
04g10	li10	he22p	hy25	
				$\ln 2$ $\ln 2$
				$\lambda_{tot} = \lfloor f_{IF(m)} t_{1/2} \rfloor \sum f_{IF(m)}^*$
<u>04g05</u>	li3p	he2pt	hy4	Nuclear Matrix
04g04	li3s	he2ss	hy3	Element (NME) from
04g03_	li2p	he2st	hy2	uccuy tubics
04g02	li2s	hels	hy1	

- Use grouped levels as super-groups for calculating reaction rates
- Variations within each super-group do not affect decay rates significantly only ratio of population in each does
  - Reaction rate matrix to be solved till steady-state final population represent plasma equilibrium

Only parent system can be studied