



Dark Photon physics in SBND

by Gaetano Fricano

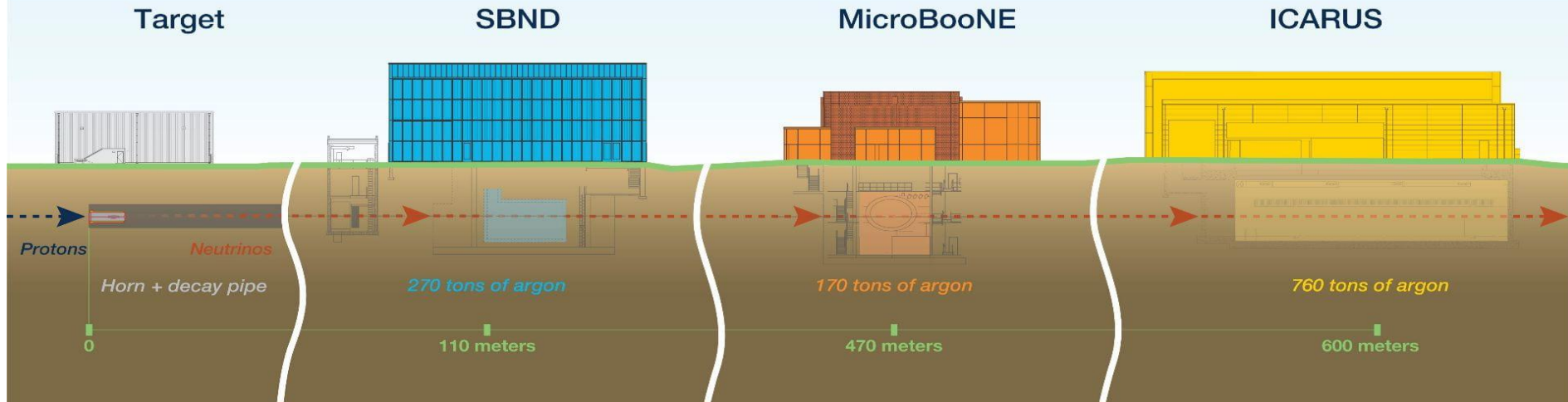
30/11/2023



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Co-supervisor: Vishvas Pandey

Short-Baseline Neutrino Program at Fermilab



SBN program involves three LArTPC neutrino detectors placed on the Booster Neutrino Beam (BNB) axis.

- Neutrino oscillation
- Sterile neutrino hypothesis
- **BSM physics**

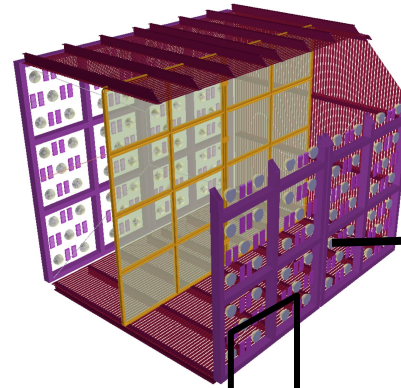


Short Baseline Near Detector

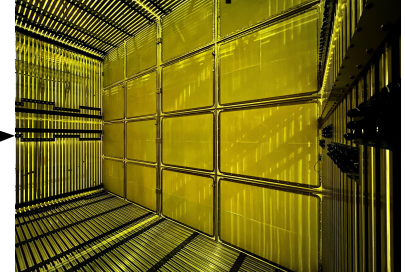
The TPC :

- 4m × 4m in the transverse plane
- 5m along the beam axis
- Two drift volumes
- 3D event reconstruction
- 3mm position resolution
- Low momentum thresholds

Each consists of 3 planes of wires with 3 mm spacing and different angle per plane.

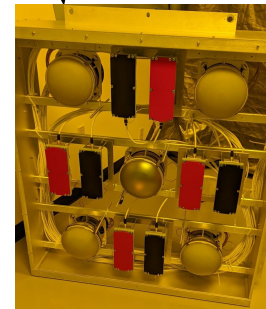
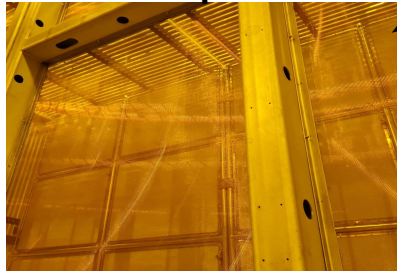


Cathode plane



500 V/cm

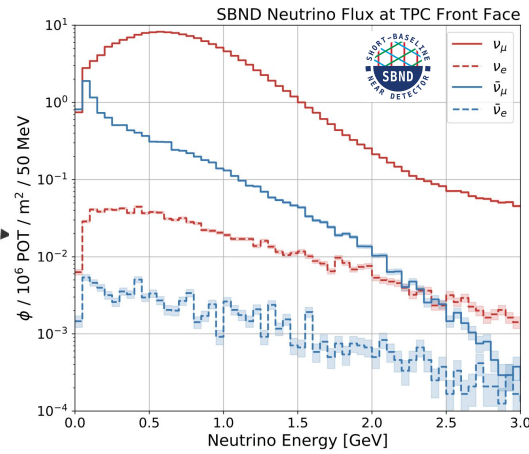
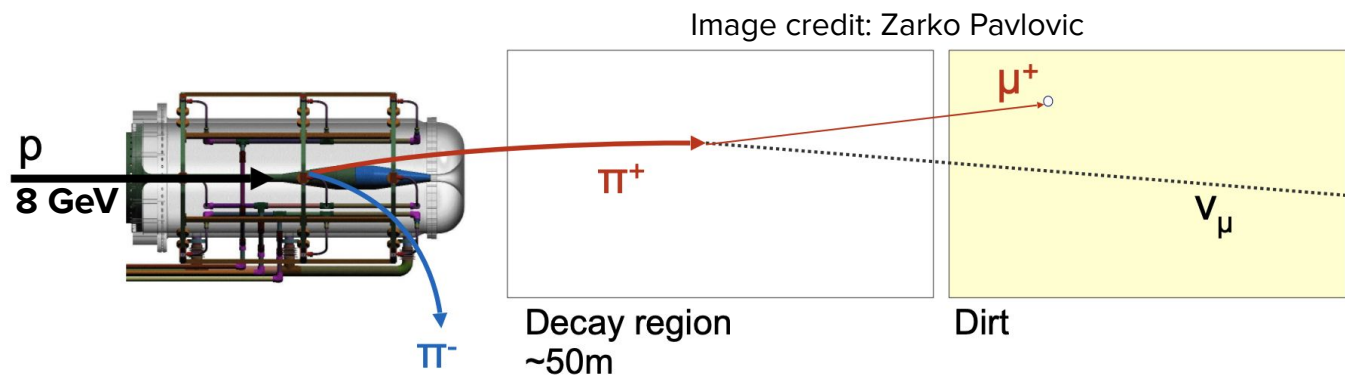
Anode plane



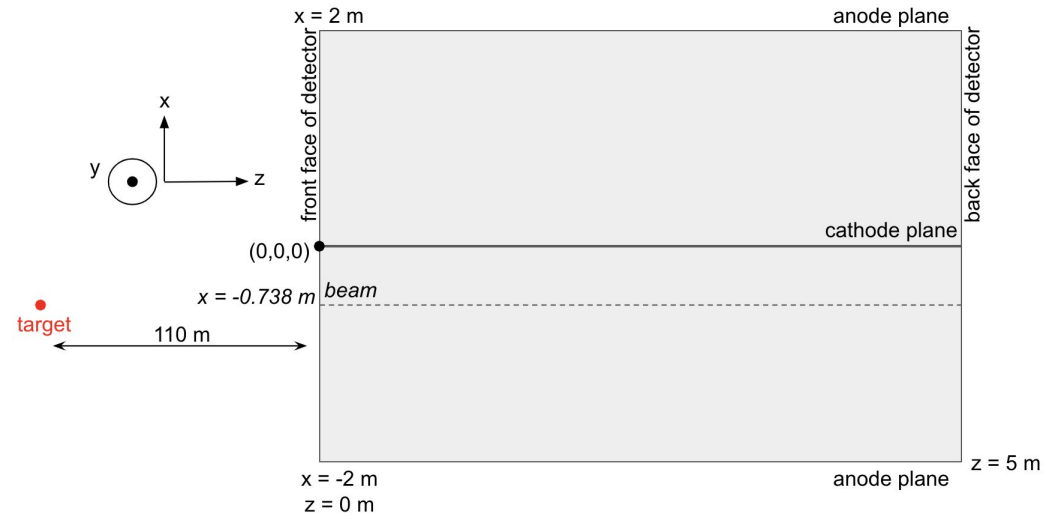
24 boxes for the PMT system



Booster Neutrino Beam

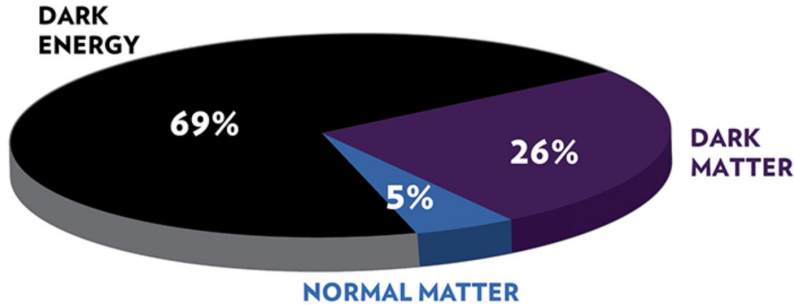


- $\langle E_\nu \rangle \sim 800 \text{ MeV}$.
- $10\text{-}18 \times 10^{20} \text{ POT/yr}$
- BSM phenomena in the beam



Dark Matter

ENERGY DISTRIBUTION
OF THE UNIVERSE



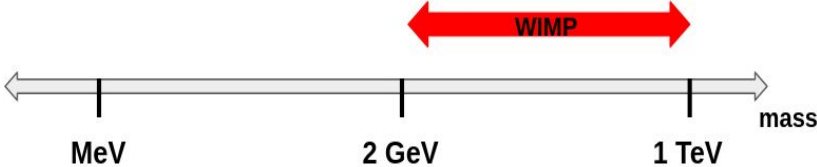
All our evidence for DM comes from its gravitational influence on visible matter

Particle physics approach

- Uncharged
- Non-relativistic and stable
- Very weakly interacting with itself and with visible matter

WIMP's model

The model of Weak Interactive Massive Particle (WIMP) assume dark matter candidates as particles with a mass greater than 2GeV that interact by mean of the Weak force.

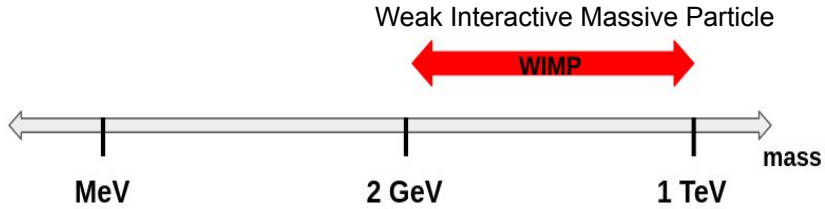


The theoretical model suggests that these particles formed following the Big Bang during the hottest and densest phases in the history of the universe. The high mass range is due to the high energies available during this period.

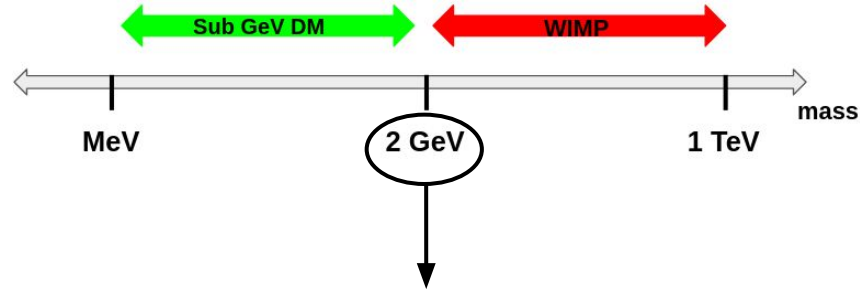
As a consequence of the universe expansion and subsequent cooling, WIMPs would have cooled down, reaching thermal equilibrium with the standard model plasma and interacting with it through the weak interaction.

$$f(t, \vec{x}, \vec{p}) = f_{GCE}(E) = [e^{(E-\mu_i)/T_i} \mp 1]^{-1} \longrightarrow \Omega \sim \frac{1}{\sigma}$$

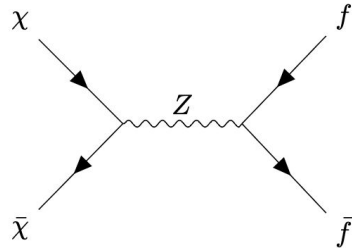




- Thermal relic model
- WIMP's miracle $\Omega \sim 1/\sigma$
(Ω = abundance)



Lee-Weinberg bound



$$L \sim G_F (\bar{\chi} \gamma^\mu \chi) (\bar{f} \gamma^\mu f)$$

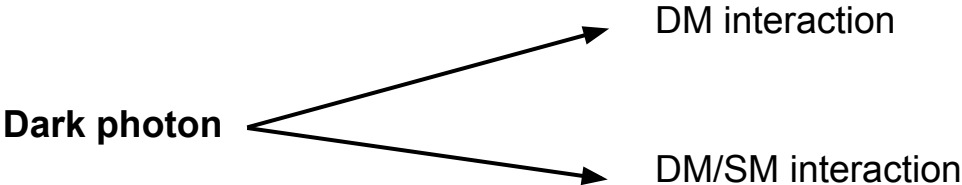
$$\sigma \sim \frac{G_F^2 m_\chi^2}{\pi} \sim 1 \text{pb} \left(\frac{m_\chi}{5 \text{GeV}} \right)^2$$

SubGeV Dark Matter

According to thermal relic model a new interaction is needed in order to have a SubGeV DM candidate.



New gauge boson !



- Model parameters
- m_ν
 - $m_\chi < 2 \text{ GeV}$
 - ε , kinetically mixing parameter.



Dark photon: Vector Portal

The gauge symmetry employed is $U(1)'$, and we are able to apply the Higgs mechanism, resulting in a non-zero mass associated with the dark photon

In order to allow the dark photon to transform into a standard model photon we introduce the concept of a **kinetically mixing parameter ϵ**

$$U(1)_{\gamma} \sim U(1)_{em}$$



$$L = \frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \frac{1}{4} F'_{\mu\nu} F'^{\mu\nu} + \frac{\epsilon}{2} F_{\mu\nu} F'^{\mu\nu} + eA_{\mu} J_{em}^{\mu} + d_D V_{\mu} J_D^{\mu} + \frac{1}{2} m_V^2 V_{\mu} V^{\mu}$$

$$V \longrightarrow V$$

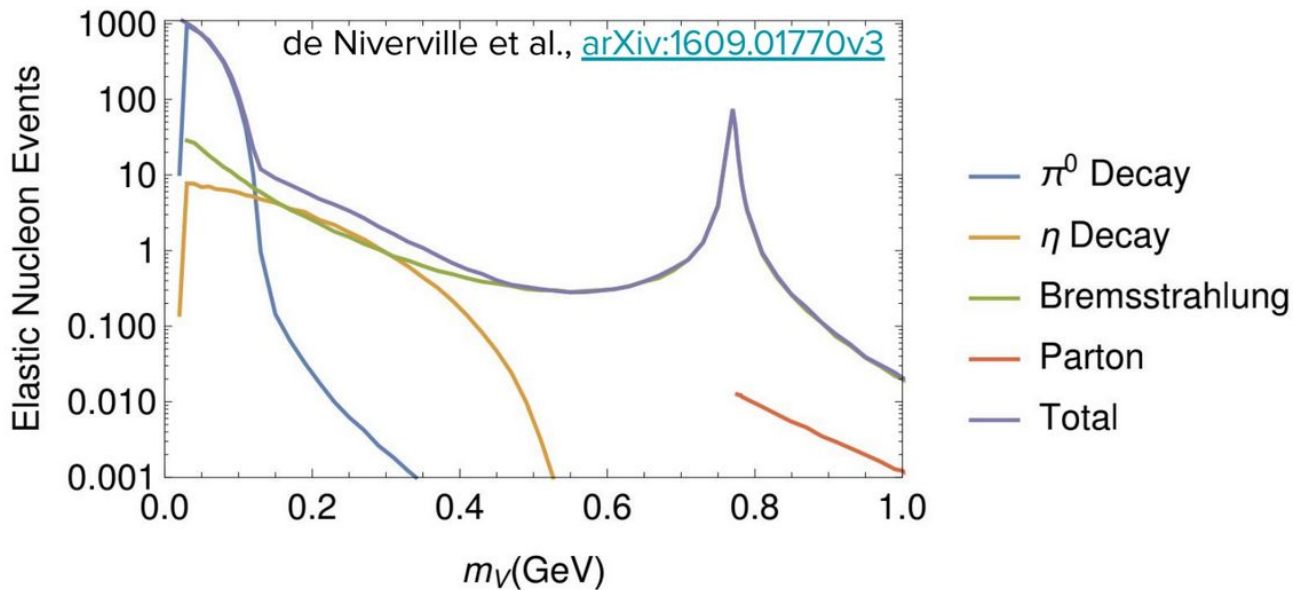
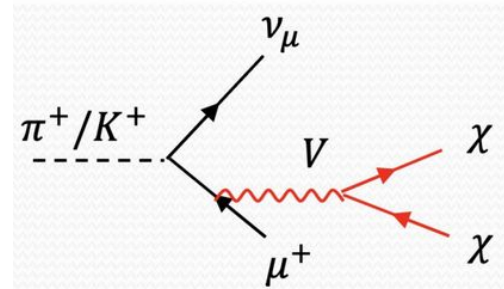
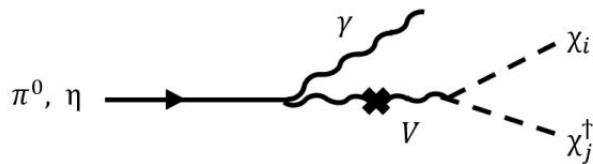
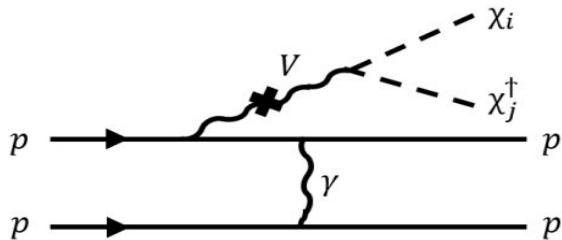
$$A \longrightarrow A + \epsilon V$$



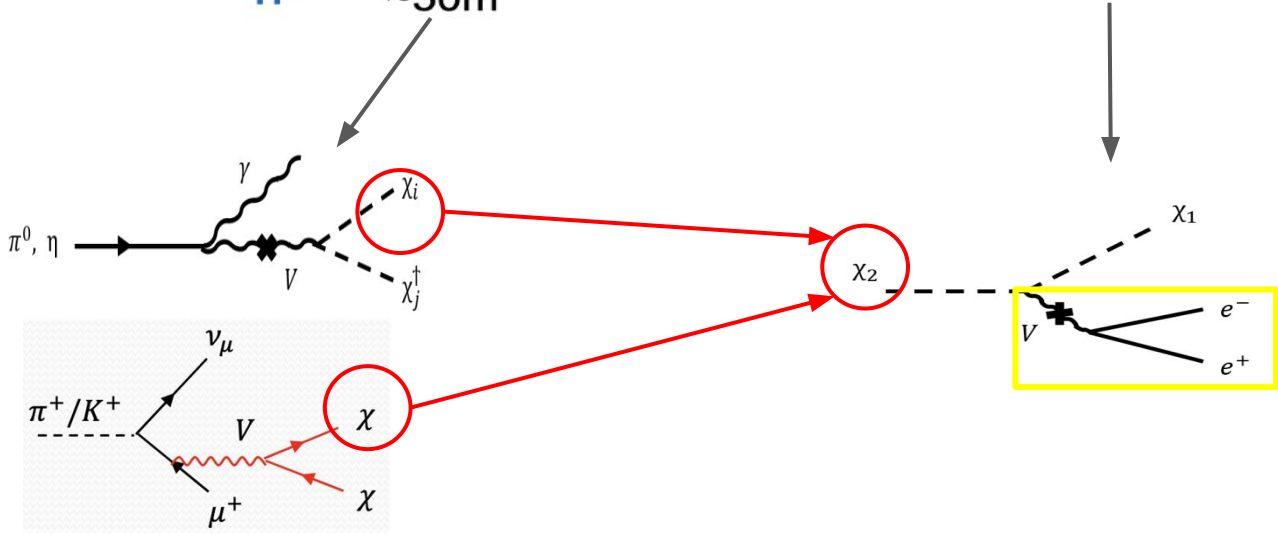
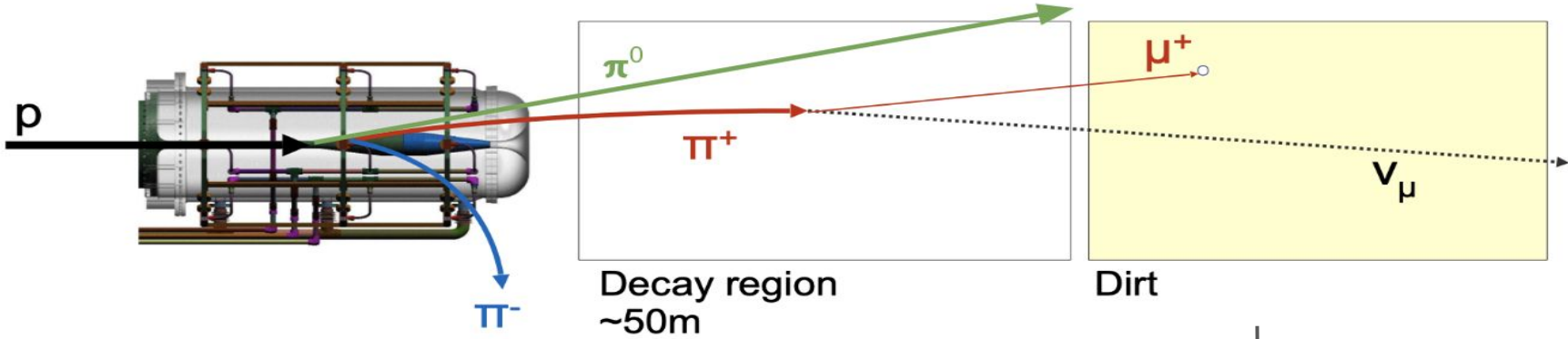
$$eA_{\mu} J_{em}^{\mu} \longrightarrow eA_{\mu} J_{em}^{\mu} + \boxed{\epsilon e V_{\mu} J_{em}^{\mu}}$$

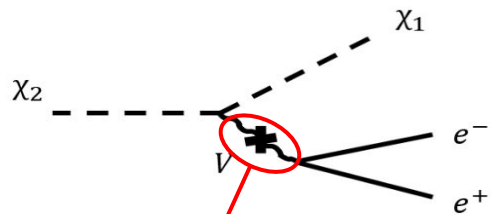
Dark millicharge



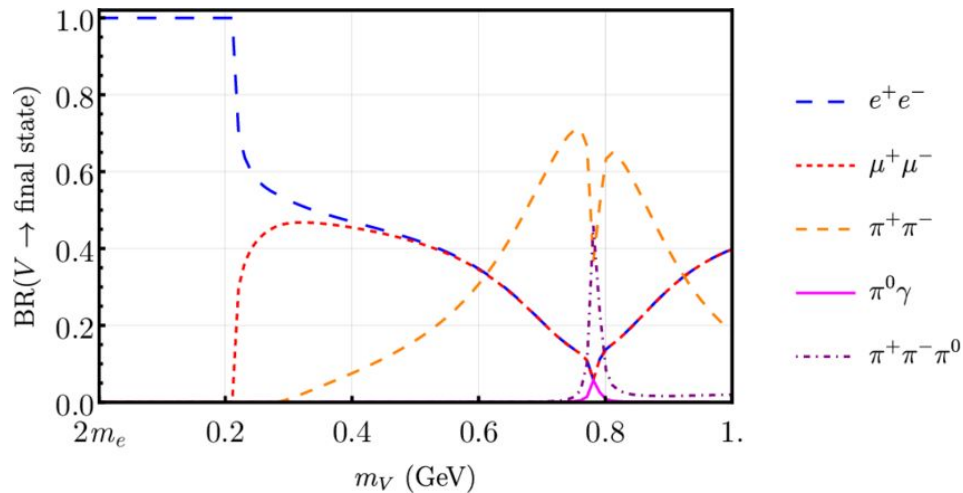


DM decay





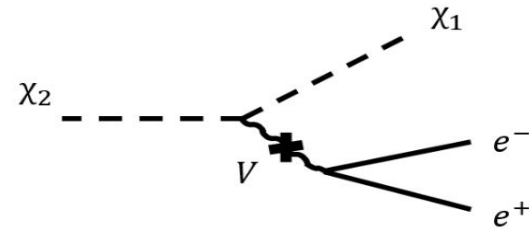
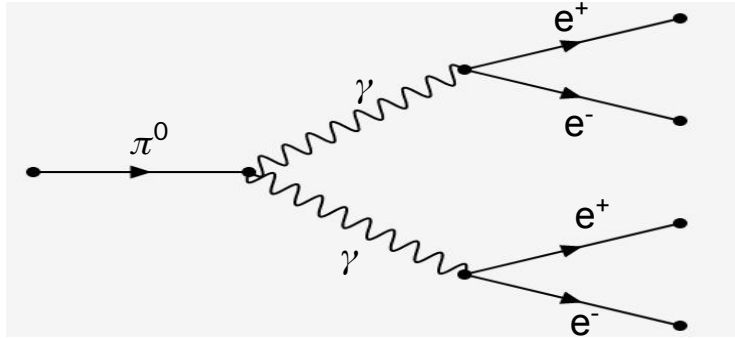
- 0.2 GeV
- 0.3 GeV
- 0.4 GeV



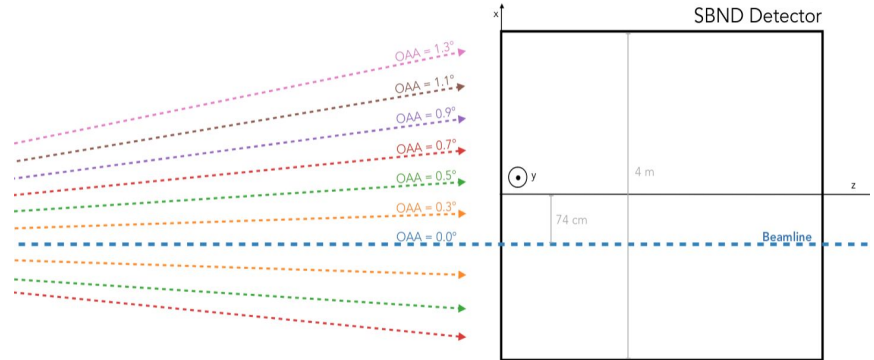
<https://arxiv.org/abs/2206.02802>

Background \implies NC π^0

Neutrino interaction with the Argon nuclei



Event selection needed \implies



About the e^+e^- selection

VertX1 - 1 cm < VertX2 < VertX1 + 1 cm

VertY1 - 1 cm < VertY2 < VertY1 + 1 cm

VertZ1 - 1 cm < VertZ2 < VertZ1 + 1 cm

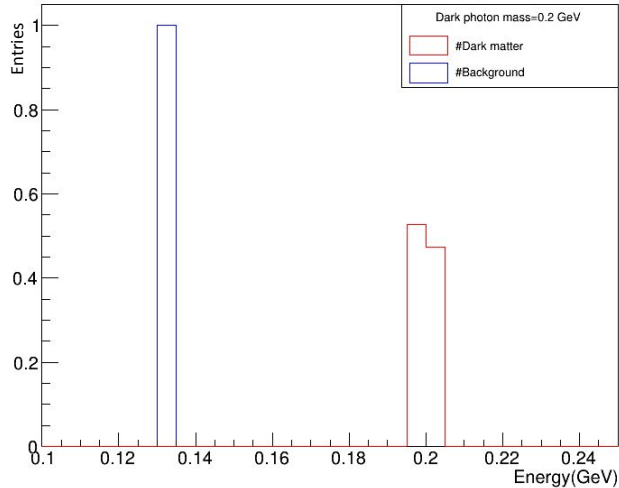


Fiducial Volume



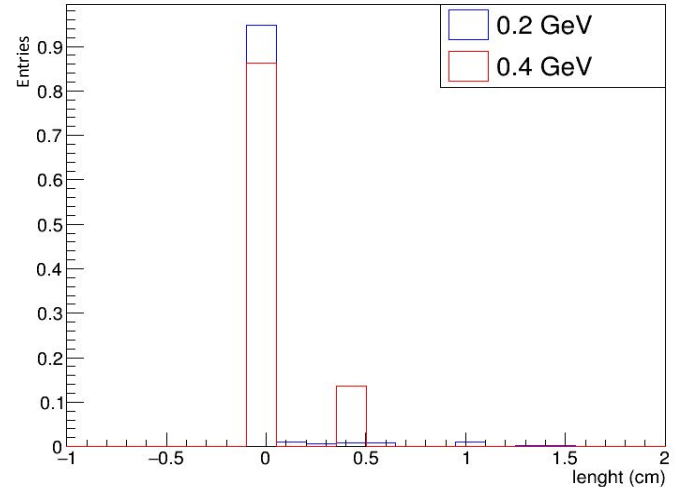
The two particles are
an e^+e^- pair !

Invariant mass SBND simulation in progress



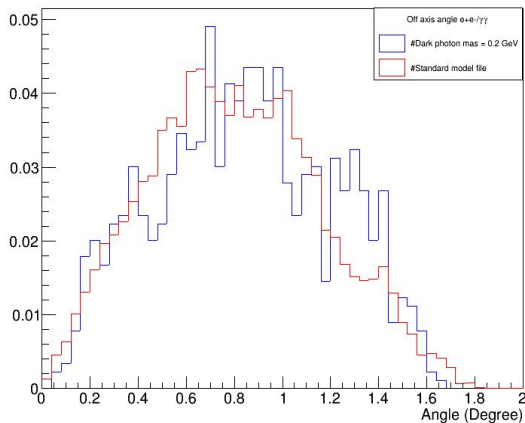
Norm of vectorial
difference

SBND simulation in-progress

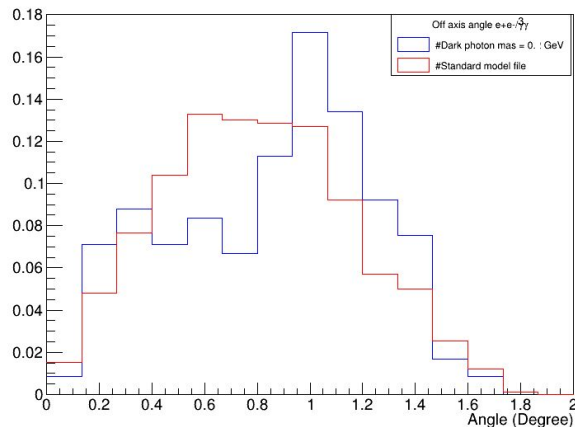


PRISM measures \longrightarrow Off-Axis angle

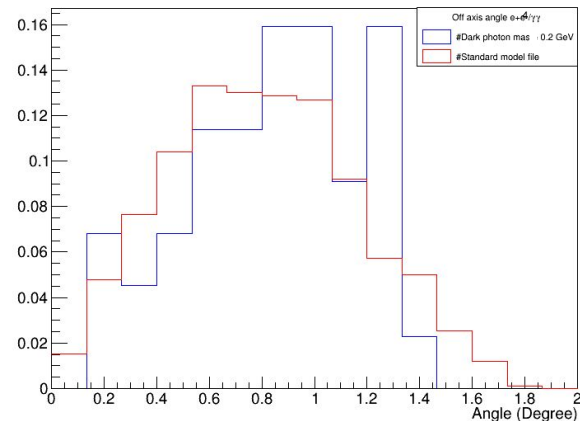
Off-axis angle SBND simulation in progress



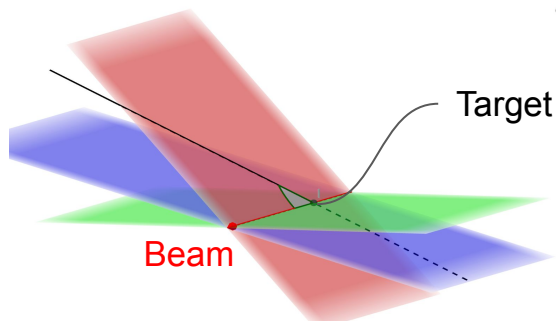
Off-axis angle SBND simulation in progress



Off-axis angle SBND simulation in progress



$$\epsilon = 10^{-7}$$

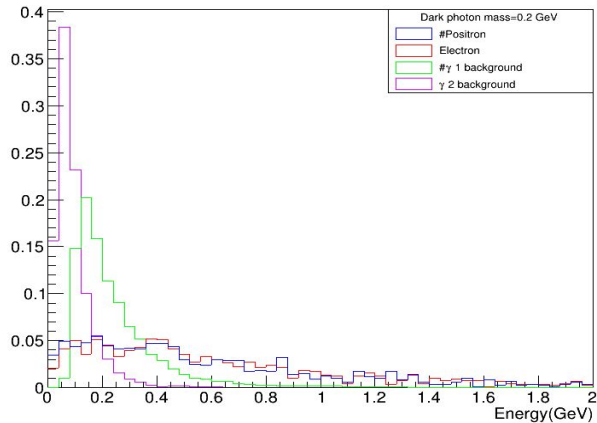


$$\text{Efficiency} = \frac{\text{Events selected}}{\text{Total events}}$$

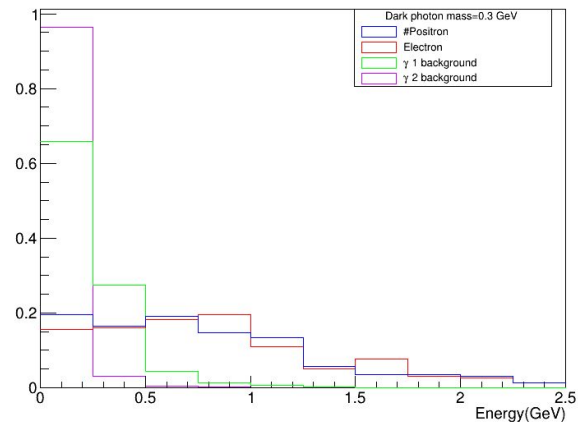


Energy distribution

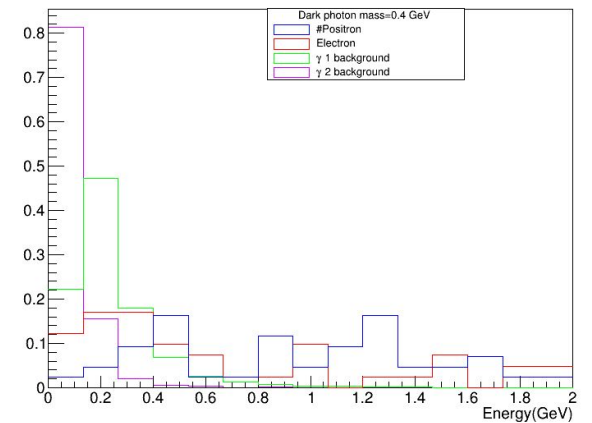
Energy SBND simulation in progress



Energy SBND simulation in progress

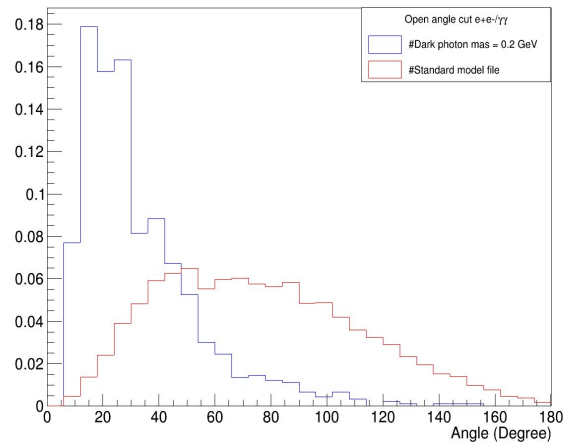


Energy SBND simulation in progress

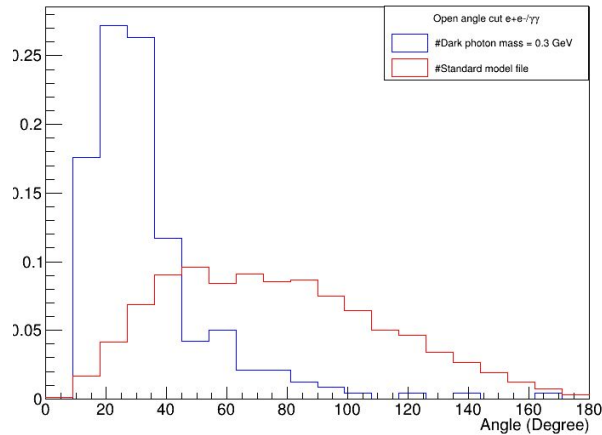


Open angle distribution

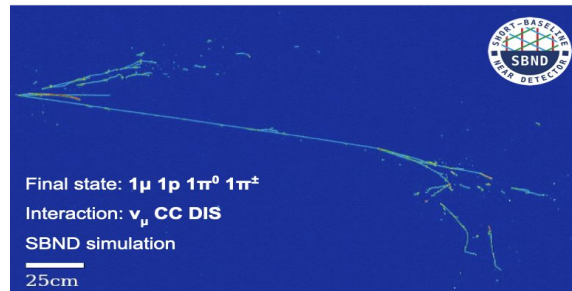
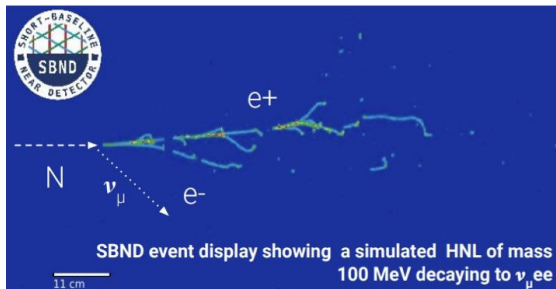
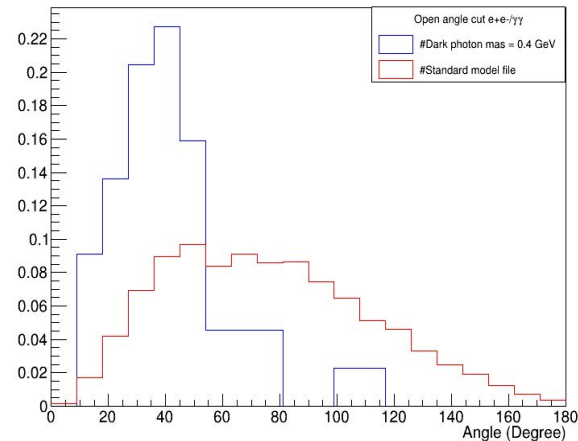
Open angle SBND simulation in progress



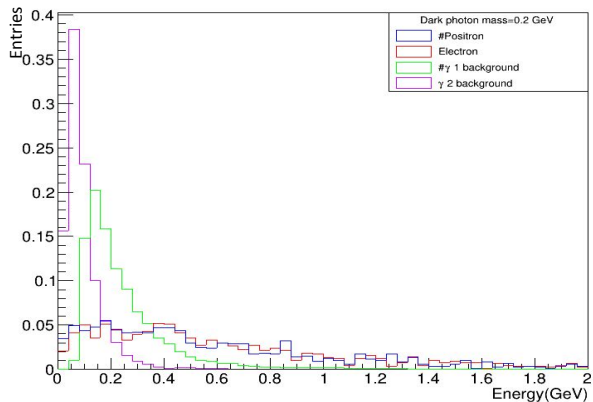
Open angle SBND simulation in progress



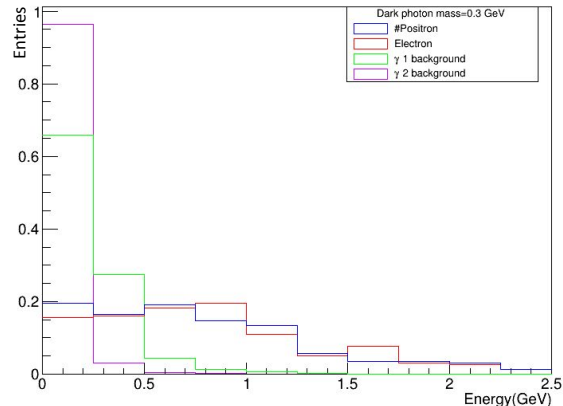
Open angle SBND simulation in progress



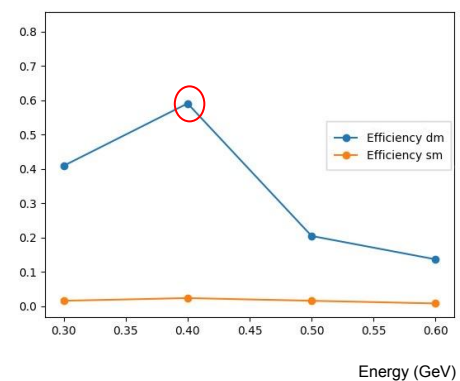
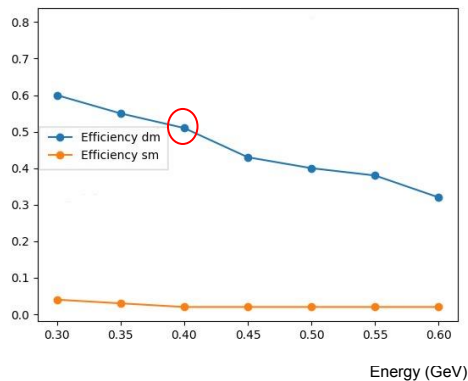
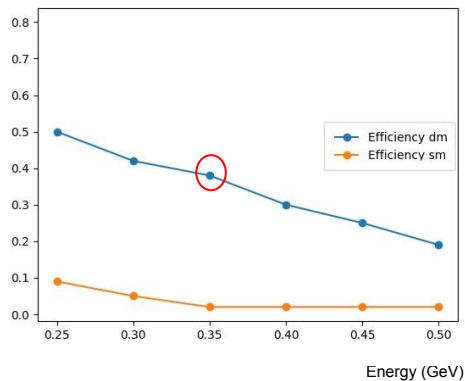
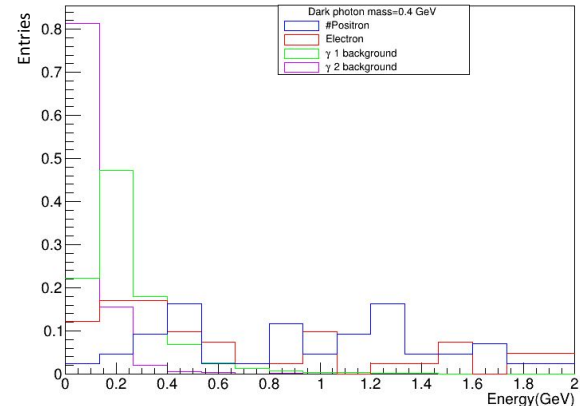
Energy SBND simulation in progress



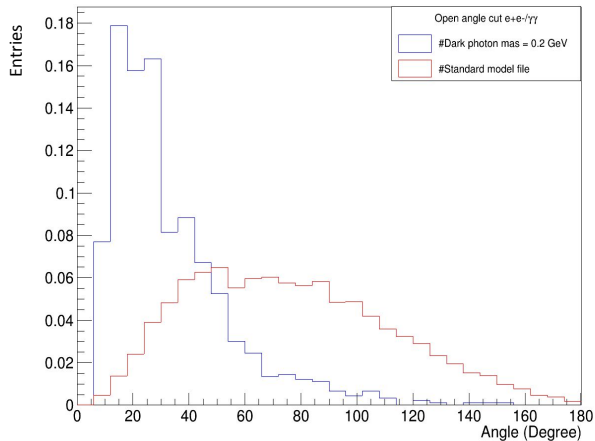
Energy SBND simulation in progress



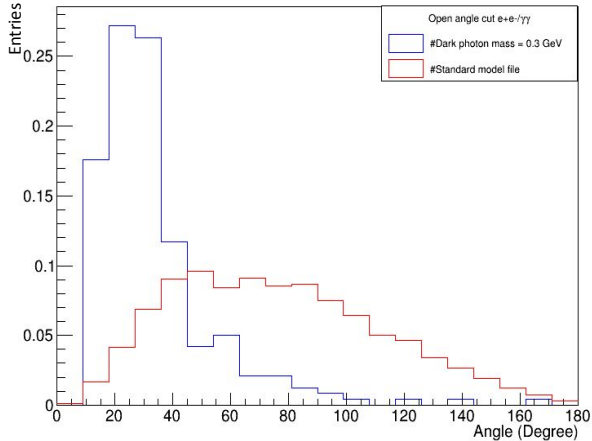
Energy SBND simulation in progress



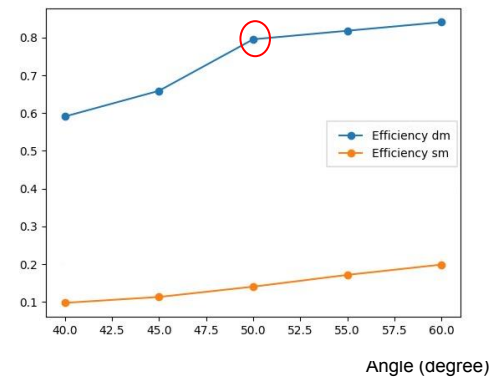
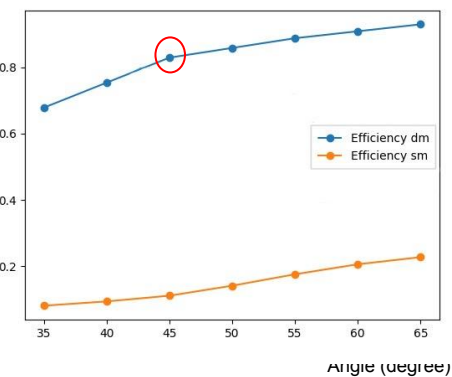
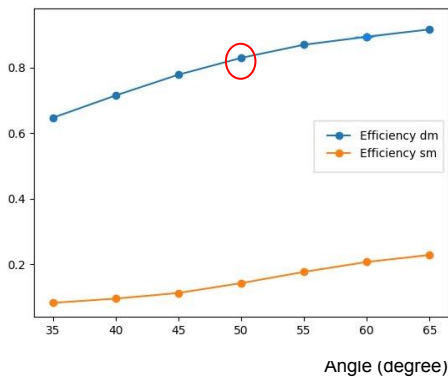
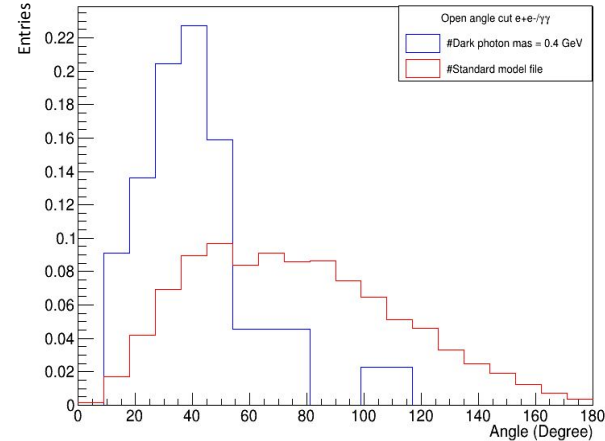
Open angle SBND simulation in progress



Open angle SBND simulation in progress

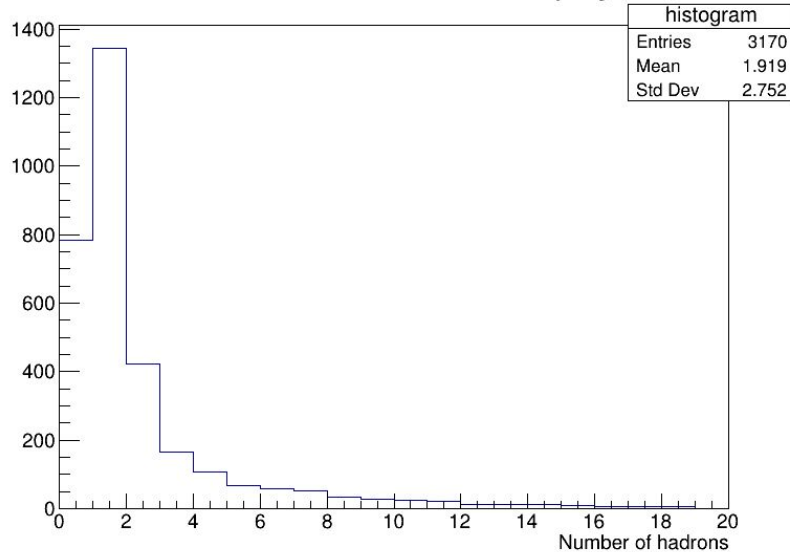


Open angle SBND simulation in progress



Work in progress

Hadron search SBND simulation in progress



- Cuts combination
- More statistics
- dE/dx
- Invariant mass



Thank you for your attention!

