## LArTPC TECHNIQUE for NEUTRINO EVENTS and CHARGED PARTICLE RECONSTRUCTION

Visual scanning of neutrino Interactions in LArTPC

Decoding data from FTBF beam counters and conversion into ROOT format

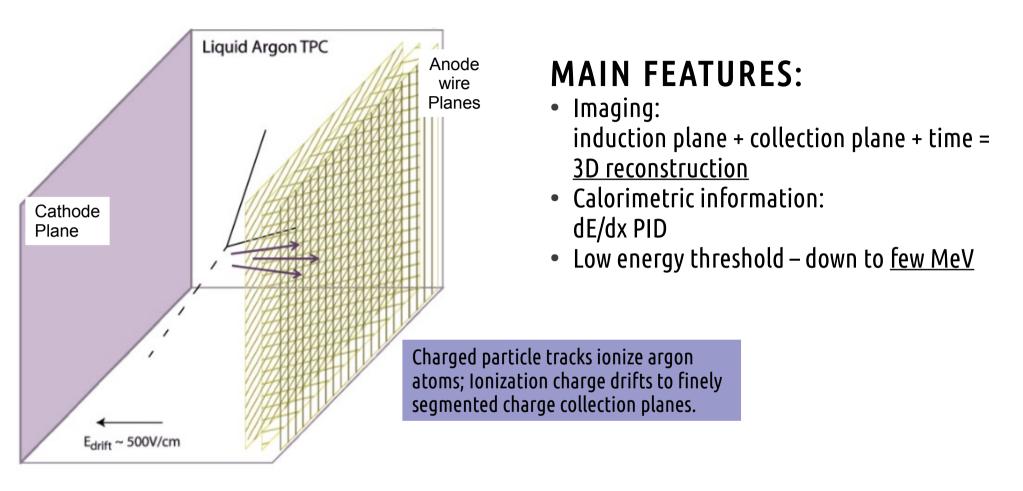
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**Supervisors:** Jennifer Raaf Ornella Palamara Flavio Cavanna

September 26<sup>th</sup> 2013

# LArtpc

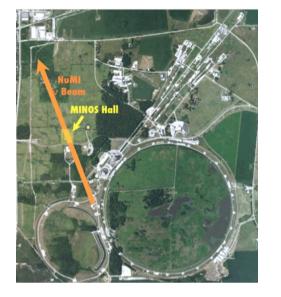


"There are several reasons why pure LAr can be considered as an almost ideal material for a liquid target TPC: it is **dense**, it **does not attach electrons** and hence it permits **long drift-times**, it has a **high electron mobility**, it is **easy to obtain** and to **purify**, it is **inert**, it is **cheap**..."

C. Rubbia, The Liquid-Argon Time Projection Chamber: A New Concept For Neutrino Detector, CERN-EP/77-08 (1977)

# ArgoNeuT





ArgoNeuT was a R&D project at Fermilab (USA) to expose a small Liquid Argon TPC to the NuMI neutrino beam.

ArgoNeuT detector was located between MINERvA and the MINOS near detector (ND) at NuMI Tunnel – 100m underground. Muons escaping the TPC are reconstructed in MINOS ND.

Collecting events in the 0.1 to 10 GeV range, ArgoNeuT was producing the first ever data for low energy neutrino interactions within a LArTPC.

## GOALS:

Demostrate PID capabilities of LArTPC with dE/dx and range measurements.

Study CC and NC neutrino events in the few GeV Range in LAr.

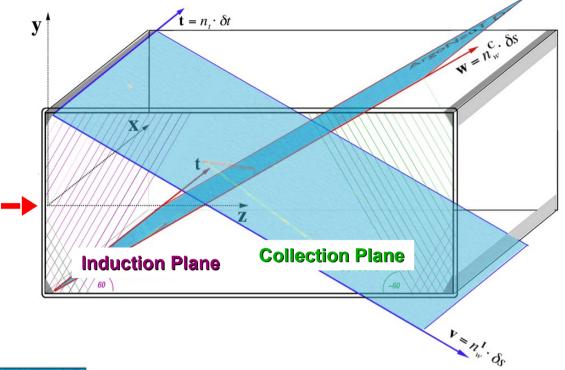
Precise CC QE muon neutrino cross section measurement in Argon.

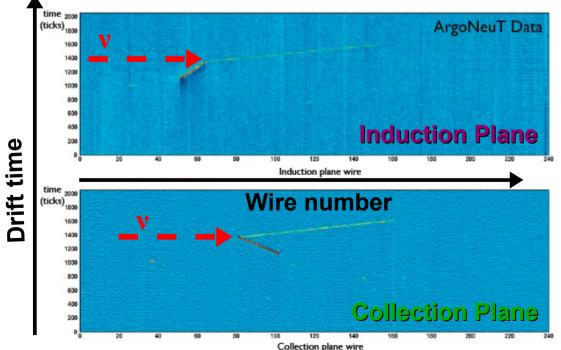


## Neutrino Interactions in ArgoNeuT detector

2 wire planes at the edge of a 170 l TPC 240 wires on each plane 4 mm spacing Plane orientation → 60°

Each of the two instrumented wire-planes provides a <u>2D-image</u> corresponding to the event projection on a plane whose axes are identified as "wire coordinate" and "time coordinate".

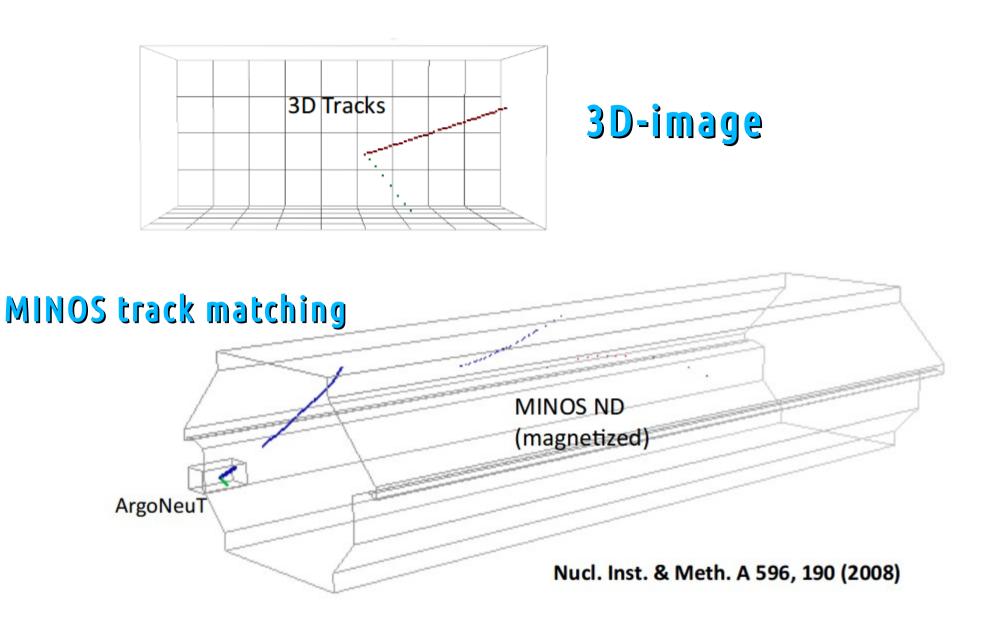




 $\frac{2\text{D views}}{(w, t) + (v, t)} \rightarrow \frac{3\text{D-image}}{(x, y, z)}$ 

Color is representative of the amount of charge detected by the wires.

Low charge



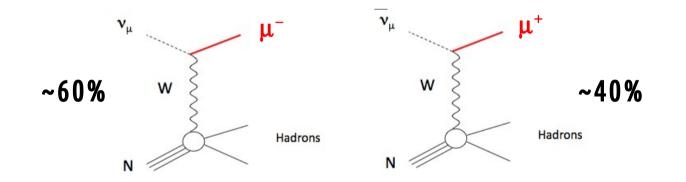
The presence of the **MINOS ND** allows for <u>energy reconstruction</u> and <u>charge identification</u> of escaping muons.

#### CURRENT ANALYSIS TOPIC Muon AntiNeutrino Inclusive Charged Current Cross-Section

Measurements of the muon neutrino and antineutrino CC-inclusive cross-section using the <u>Antineutrino running data</u>

→ NuMI beam: % neutrinos > % antienutrinos

 $\rightarrow$  use the magnetized MINOS ND to distinguish between neutrino and antineutrino.



#### **Automatic Event Selection**

(frame recorded at every spill – most frames are empty or with crossing particles generated by neutrino interactions in the upstrem material)

Visual Scanning inefficiencies of the automated selection (and validate final CC-inclusive sample)

Muon kinematics and Cross-section measurement

## Visual scanning of CC-inclusive sample

**TASK:** Remove from the CC inclusive sample those events that were not rejected by reconstruction.

Neutrino events characterization takes place with the LARSOFT automated reconstruction software: hit finding  $\rightarrow$  hit clustering  $\rightarrow$  cluster fitting as linelike objects  $\rightarrow$  <u>TRACK</u> identification.

# DATA:

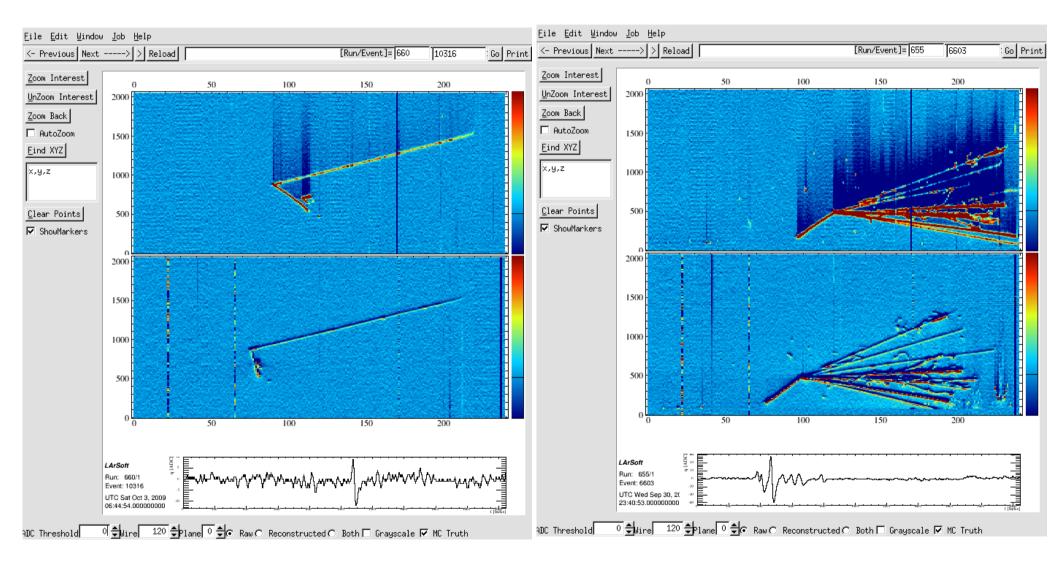
<u>Neutrino-mode</u> (2 weeks  $\rightarrow$  8.5x10<sup>18</sup> POT) 414 events

<u>Antineutrino-mode</u> (6 months  $\rightarrow$  1.2x10<sup>20</sup> POT) ~2000 events

Selection criteria on the events:

- MINOS matching
- Fiducial volume

## $\boldsymbol{\nu}_{\!\mu}$ interaction to keep in the CC-inclusive sample



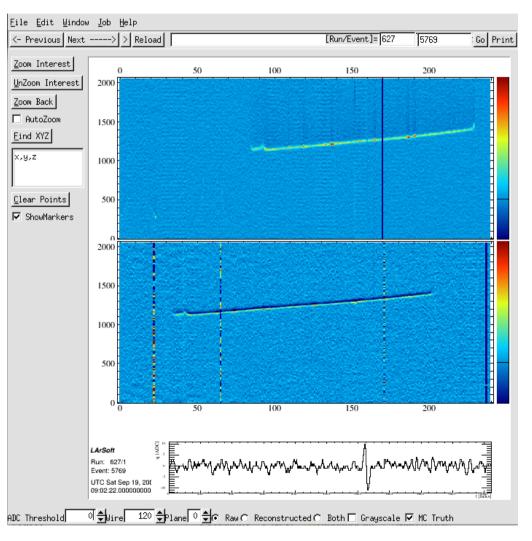
# **EVENT CLASSIFICATION**

- upstream events: upstream events produced by interactions outside the detector.
- heavily ionizing: events that probably would be rejected if we checked the dE/dx info.
- piddly tracks: very short tracks (contained or not).
- **garbage:** events where erroneously a track was reconstructed.
- the unknown: events that we don't understand how to interpret.

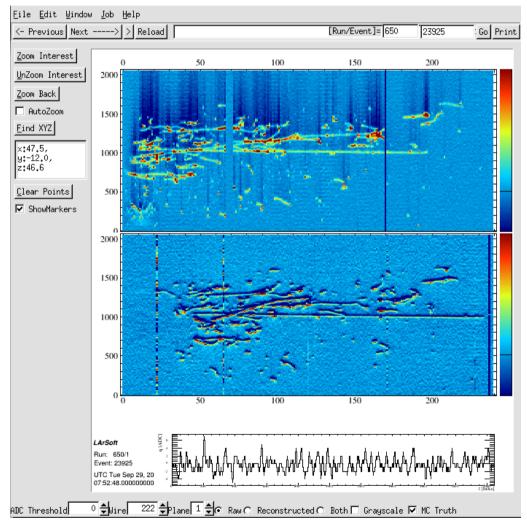
### • "v's" : events that look like electron neutrino interactions.

# **UPSTREAM EVENTS**

#### **CROSSING MUONS**

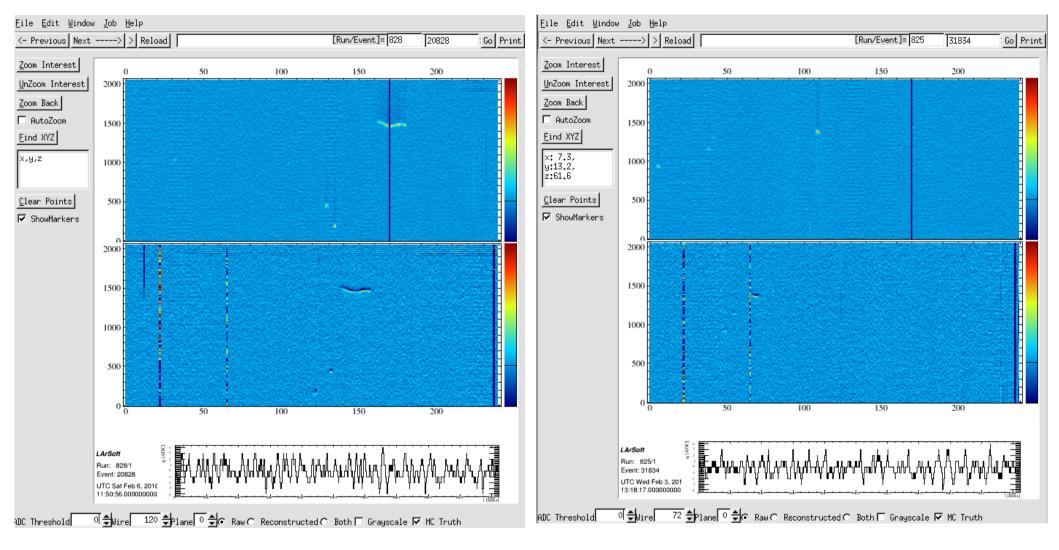


#### **NEARBY INTERACTIONS**



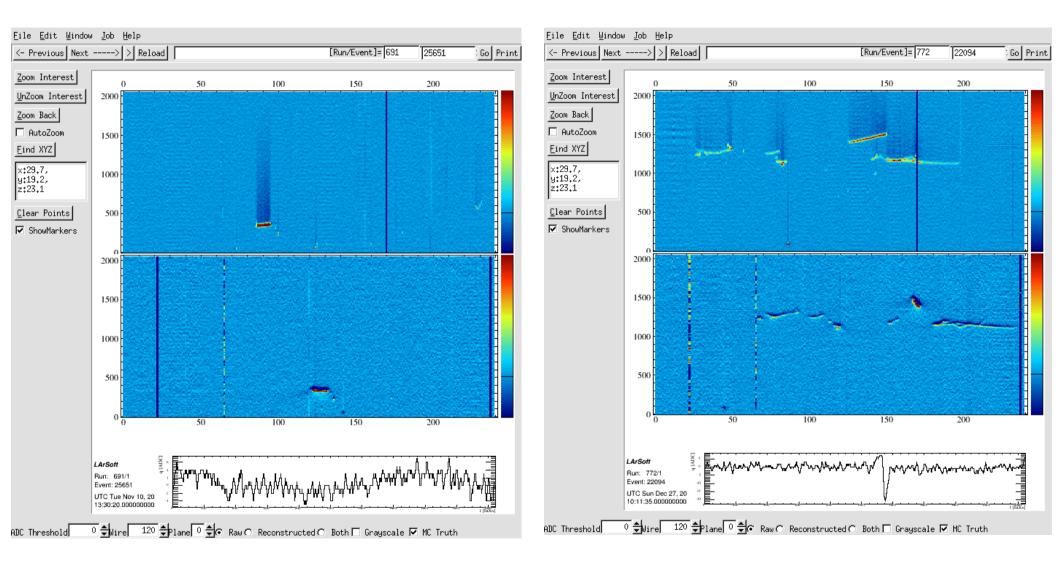
#### **PIDDLY TRACK**

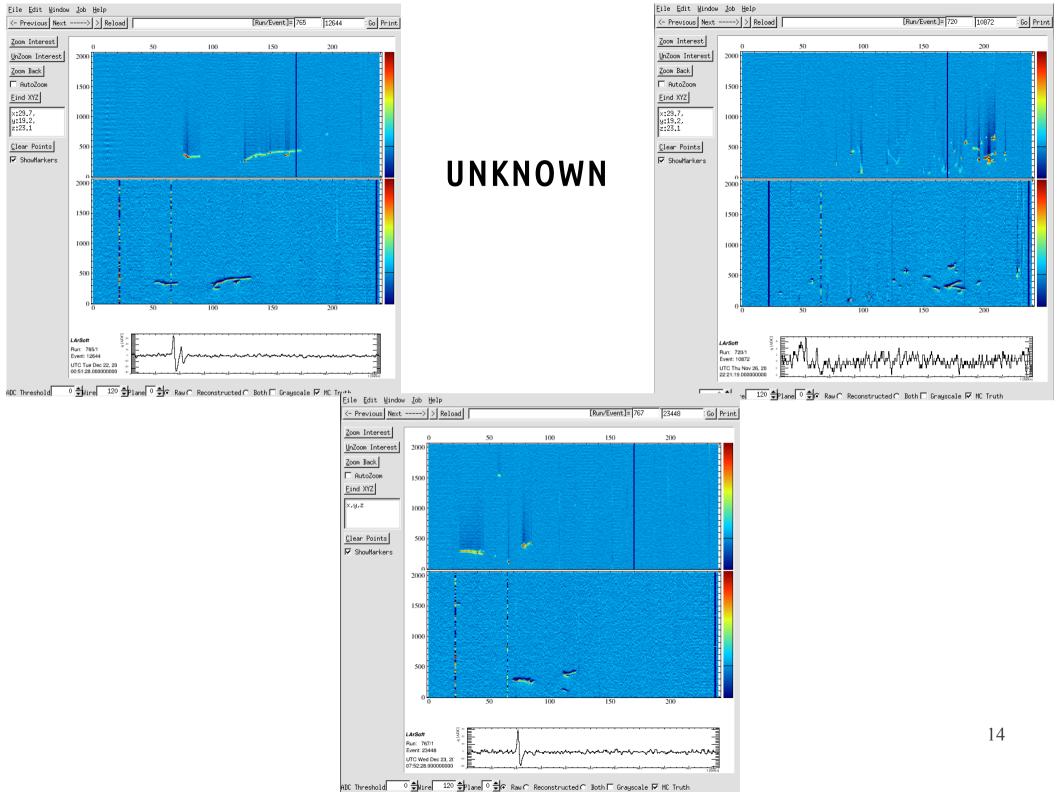
#### GARBAGE



#### **HEAVLY IONIZING**

#### UNKNOWN





# CONCLUSIONS

- EVENTS TO BE REJECTED:
- 54 CROSSING MUONS
- 5 NEARBY INTERACTIONS
- 10 PIDDLY TRACKS
- 3 HEAVLY IONIZING
- 3 GARBAGE
- 3 ELECTRON NEUTRINO LIKE
- 5 EVENTS TO BE CHECKED (UNKNOWN)

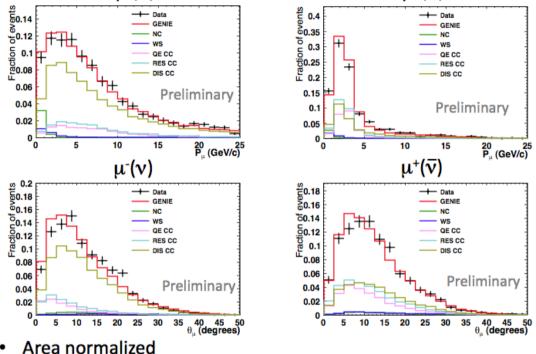
## ~15% of events potentially to be removed.

Antineutrino Mode

Veutrino Mod

- EVENTS TO BE REJECTED:
- 360 CROSSING MUONS
- 36 NEARBY INTERACTIONS
- 45 PIDDLY TRACKS
- 9 HEAVLY IONIZING
- 29 GARBAGE
- 45 EVENTS TO BE CHECKED (UNKNOWN)
  - ~26% of events potentially to be removed.

#### Muon kinematics analysis and Cross-section measurement - IN PROGRESS -Anti-neutrino Mode (1.2E20) $\mu^{-}(\bar{v})$



- Need to improve flux prediction
- Paper in preparation

Data from a subsample (~1/4) from which the bad reconstructed events were rejected thanks to visual scanning analysis.



Now the whole data file is ready to be analyzed.

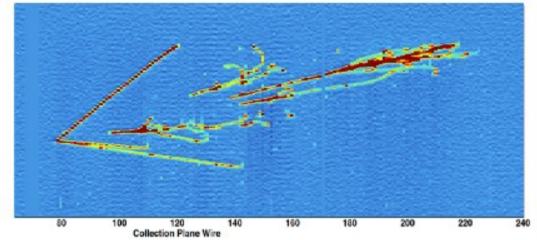
## current analysis topic Study EM shower events

EM showers sources:

1) electron neutrino contamination in muon neutrino beam (CC reaction) 2) neutral pion production (Delta resonance channel/DIS channel)  $\rightarrow 2 \gamma \rightarrow 2$  EM Showers.

3)  $\nu_{\mu} \rightarrow \nu_{e}$  oscillations !!!

(excluded in ArgoNeuT because at near location from the neutrino beam source)



Neutrino interaction with 4 photon conversions

#### Automated EM showers event selection

Visual scanning to remove inefficiencies of the automated selection (and validate final EM shower event sample)

## Visual scanning of shower events

**TASK** Classification of events from a data sample obtained after a shower filtering selection was applied  $\rightarrow$  improvement of the Filter performace.

# DATA:

**<u>Neutrino-mode</u>** (2 weeks  $\rightarrow$  8.5x10<sup>18</sup> POT) 611 events

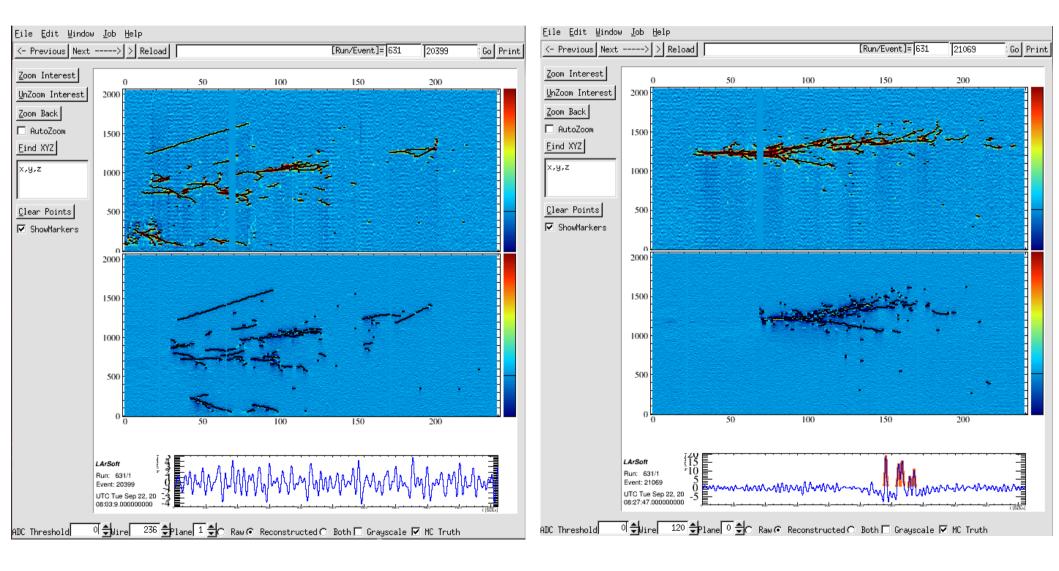
Selection criteria on the events: - Shower Filter

# **EVENT CLASSIFICATION**

- shower with gap: showers probably due to  $\pi^0 \rightarrow \gamma \gamma$
- **shower NO gap:** showers connected to the primary vertex or isolated.
- **shower muon**: showers with a well defined muon track.
- **single track + gamma rays:** single tracks with spots induced by photon interactions inside the TPC.
- single track + delta rays: single tracks with low energy secondary electrons close to the main track.
- many tracks: muon neutrino interactions with more than one particle at the vertex.
- **busy:** events with high multiplicity (contained or not).
- other: all other kind of events.

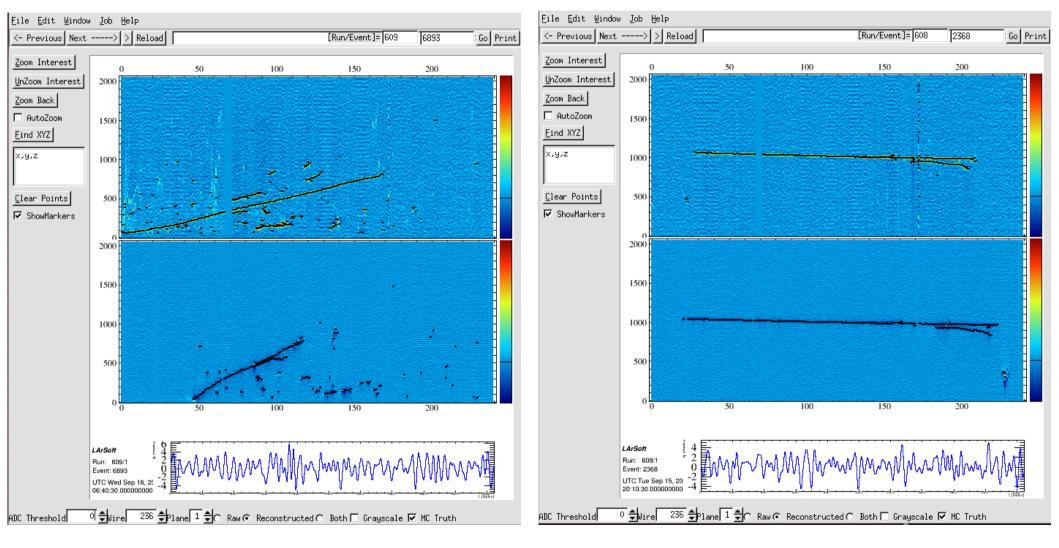
#### SHOWER WITH GAP

#### SHOWER NO GAP



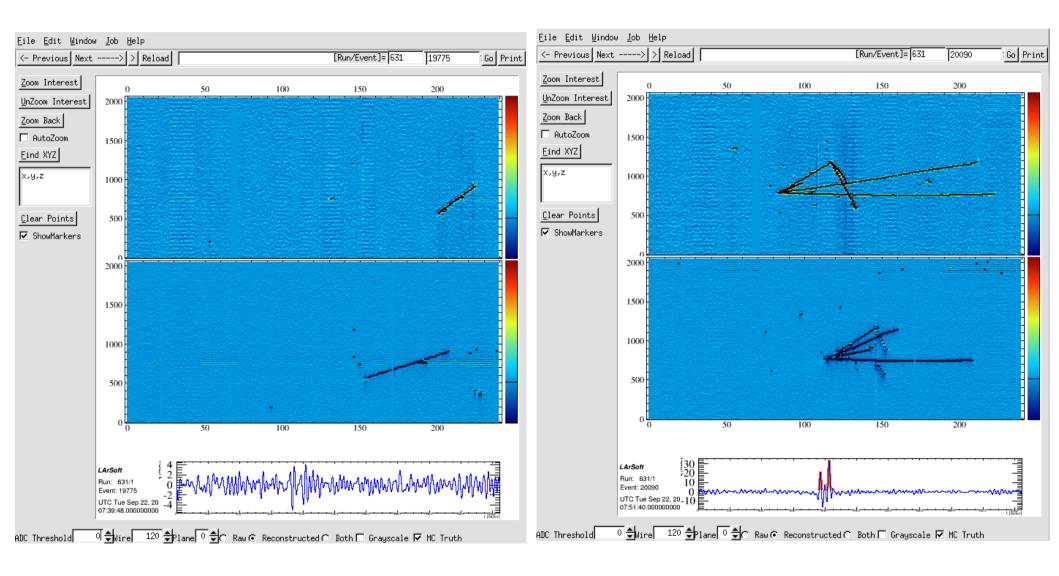
#### SHOWER + MUON

#### TRACK + DELTA RAYS



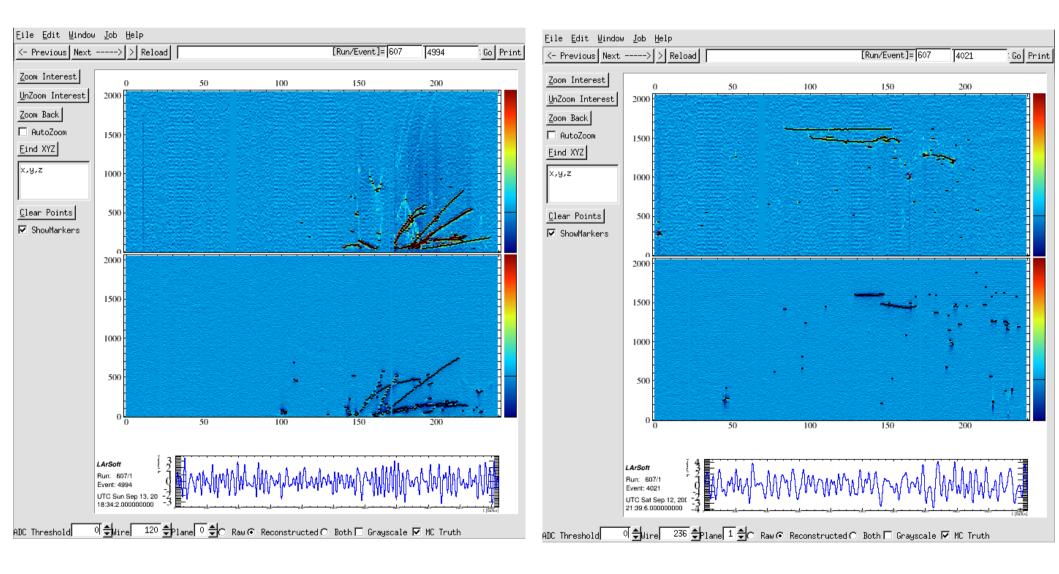
#### **TRACK + GAMMA RAYS**

#### MANY TRACKS



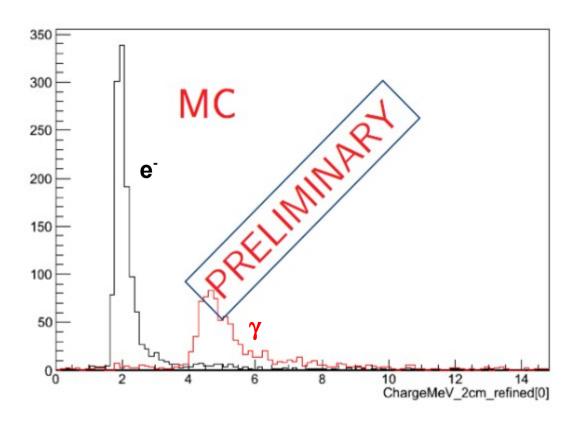
#### BUSY

#### OTHER



## THE NUMBERS:

- 108 SHOWERS WITH GAP
- 57 SHOWERS NO GAP
- 33 SHOWERS +  $\mu$
- 74 SINGLE TRACK +  $\gamma$
- 82 SINGLE TRACK +  $\delta$
- 63 MANY TRACKS
- 37 BUSY
- 157 OTHER

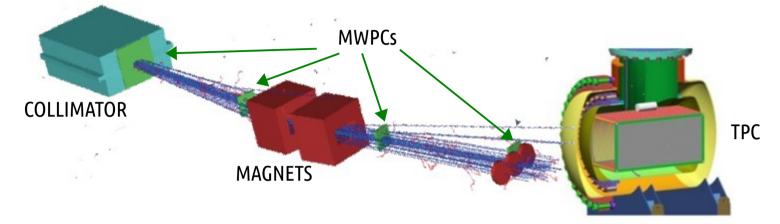


#### <u>165</u> shower events ←

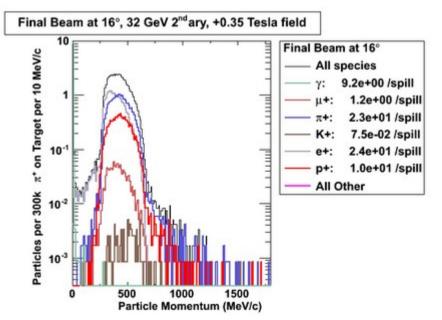
Preliminary studies of dE/dx for electron-gamma separation are feasible!

# LAriat

Experiment set to calibrate LarTPC technology by palcing the detector on a beam of charged particles of kown type an momentum at FTBF.



#### LArIAT dedicated <u>Tertiary Beam</u> at FTBF (MCentral)



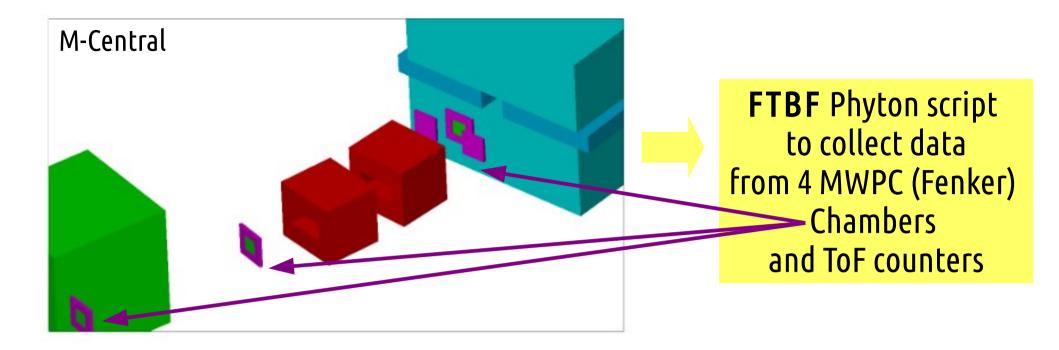
## GOALS:

Cryogenic/purification facility at FNAL designed to allow future test of LAr detectors.

Optimize the particle ID:

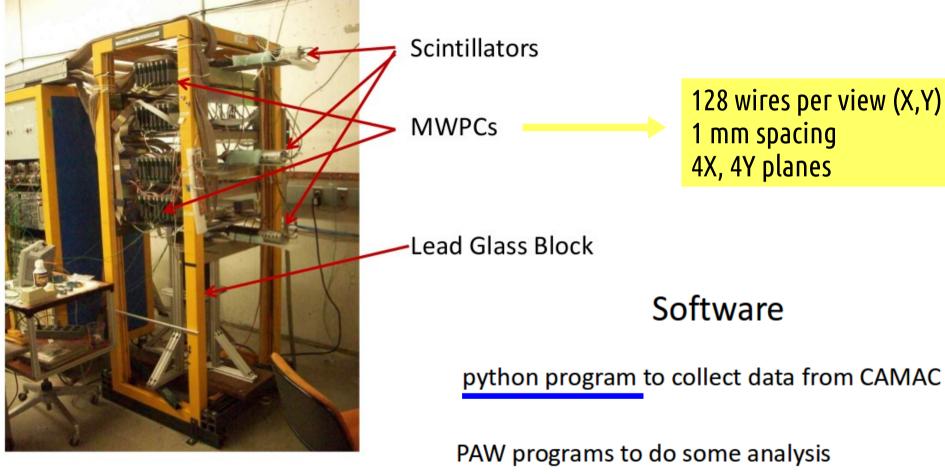
- For protons, kaons and pions measuring the recombination factors.
- For electrons and gammas measuring the dE/dx separation, crucial for studing  $v_{\mu} \rightarrow v_{e}$  oscillations.

Non-magnetic muon sign determination. 25



# **TASK:** We want to define an online routine to convert the output from the data acquisition into a ROOT file in order to define a standard format output to be used for further analysis.

#### Cosmic Ray Telescope at FTBF



128 wires per view (X,Y) 1 mm spacing 4X, 4Y planes

FTBF counters initially arranged in a Cosmic Ray Telescope geometry for test purposes.

PAW programs to do some analysis ana4r.for - quick monitor while running anatk.for - 4 chamber reconstruction >> ntuple (paw, or 😓 root) graphic tracks

## Data file from the Cosmic Ray Telescope: apr19.dat

Wait for ev				Module	б		
event s	seen at Fr	i Apr 19	09:55:12 2013	Module	7		
EVENT	1		Fri Apr 19 09:55:12 2013	Module	8		
Crate	CC0215			Channel	28	284	142.0 ns
Module	1			Module	9		
scaler	0	1		Module	10		
scaler	1	2		Channel	20	179	89.5 ns
scaler	2	Θ		Module	11		
scaler	3	Θ		Module	12		
Module	2			Module	13		
tdc	Θ	4095		Module	14		
tdc	1	3444		Module	15		
tdc	2	3532		Module	16		
tdc	3	892		Channel	25	269	134.5 ns
tdc	4	Θ		Crate	CC0208		
tdc	5	Θ		Module	1		
tdc	6	Θ		Module	2		
tdc	7	Θ		Channel	12	284	142.0 ns
tdc	8	Θ		Module	3		
tdc	9	Θ		Module	4		
tdc	10	Θ		Module	5		
tdc	11	Θ		Module	б		
tdc	12	Θ		Module	7		
tdc	13	Θ		Module	8		
tdc	14	Θ		Channel	21	295	147.5 ns
tdc	15	Θ		Module	9		
Module	3			Module	10		
adc	Θ	227		Channel	4	288	144.0 ns
adc	1	220		Module	11		
adc	2	160		Module	12		
adc	3	438		Module	13		
Crate	CC0211			Module	14		
Module	1			Module	15		
Module	2			Module	16		
Channel	29	291	145.5 ns	Channel	19	288	144.0 ns
Module	3						
Module	4			Wait for e			
Module	5					Арг 19	09:55:46 2013
				EVENT	2		Fri Apr 19

09:55:46 2013





event number



event number – crate – module – scalerID – scaler value



event number – crate – module – tdcID – tdc value



event number – crate – module – adcID – adc value



event number – crate – module – channel – wireID – time

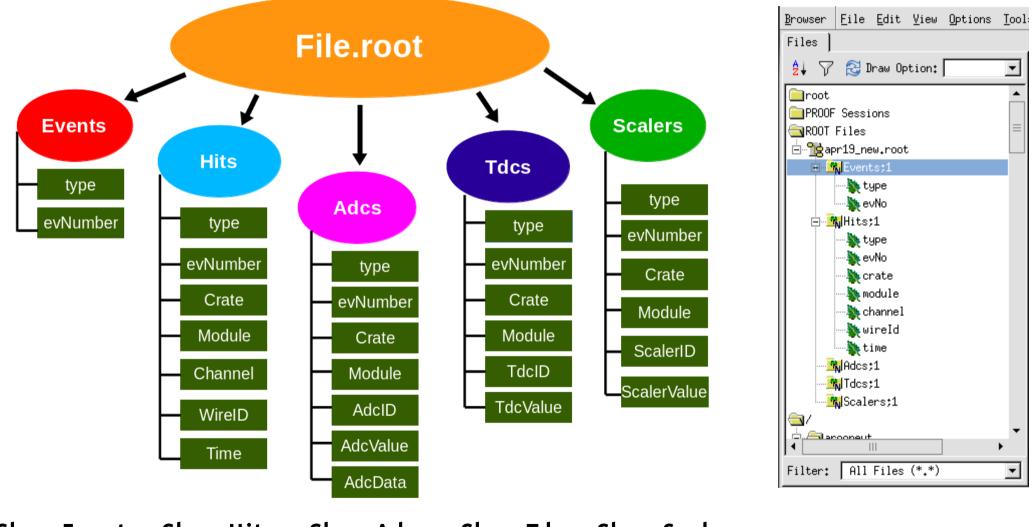
## Simplified .txt Data file

apr19.txt

19880301 1		EVENT
20070103 1 215	5 1 0 1	
20070103 1 215		
20070103 1 215	5 1 2 0	SCALER
20070103 1 215	5 1 3 0	
19882007 1 215	5 2 0 4095	
19882007 1 215	5 2 1 3444	
19882007 1 215	5 2 2 3532	
19882007 1 215	5 2 3 892	
19882007 1 215	5 2 4 0	TDC
19882007 1 215	5 2 5 0	
19882007 1 215	5 2 6 0	
19882007 1 215	5 2 7 0	
19882007 1 215	5 2 8 0	
19882007 1 215	5 2 9 0	
19882007 1 215	5 2 10 0	
19882007 1 215	5 2 11 0	
19882007 1 215	5 2 12 0	
19882007 1 215	5 2 13 0	
19882007 1 215	5 2 14 0	
19882007 1 215	5 2 15 0	
19880103 1 215	5 3 0 227	ADC
19880103 1 215	5 3 1 220	
19880103 1 215	5 3 2 160	
19880103 1 215	5 3 3 438	1.1177
19880720 1 211	1 2 29 291 145.5	HIT
19880720 1 211	1 8 28 284 142	
19880720 1 211	1 10 20 179 89.5	
19880720 1 211	1 16 25 269 134.5	
19880720 1 208	8 2 12 284 142	
19880720 1 208	8 8 21 295 147.5	
19880720 1 208		
	8 16 19 288 144	
19880301 2		

## Final .root Data file

#### apr19.root



Class Event Class Hit Class Adc Class Tdc Class Scaler

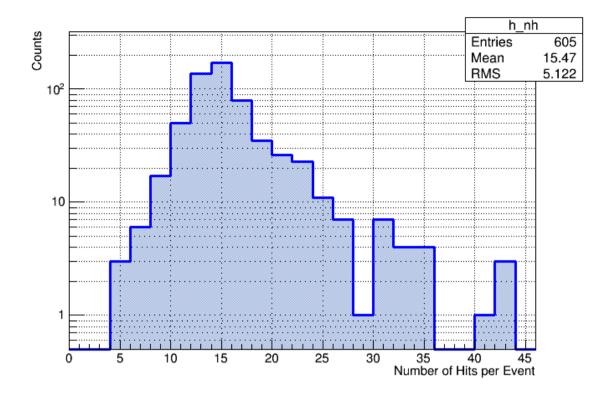
## SUMMARY

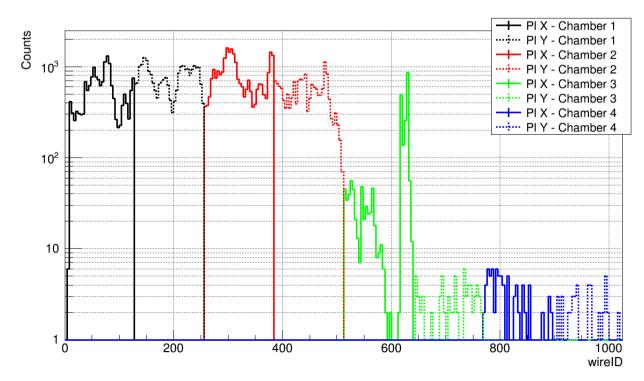
dataformat.h	Defines the classes needed to fill the root file with different kind of entries.
makeTxT.C	Takes the original data file as input and converts it into a simplified txt file.
makeRoot.C	Takes the .txt file as input and converts it into a root file with all the information from the data acquisition.

readRoot.C

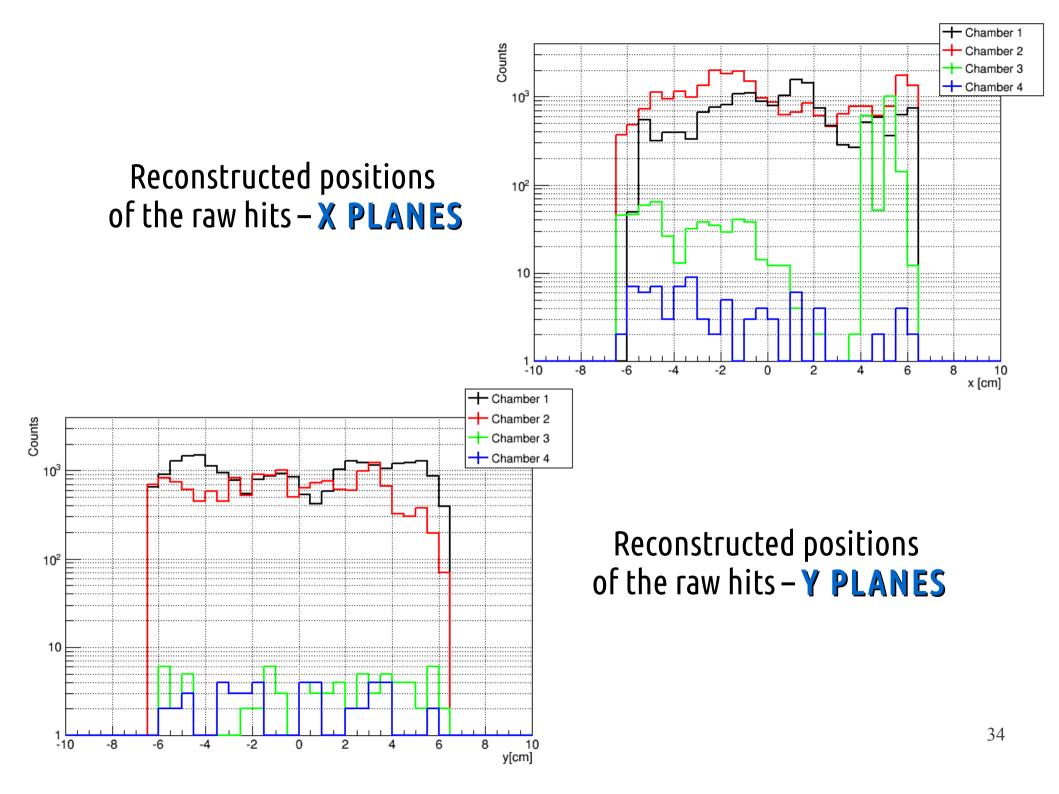
Retrieves the information from the root file.



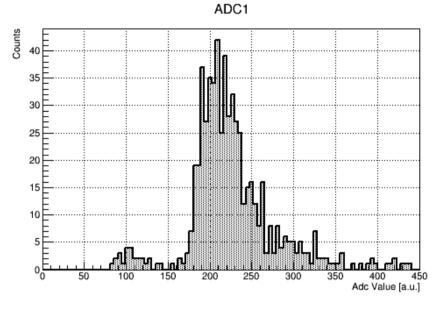




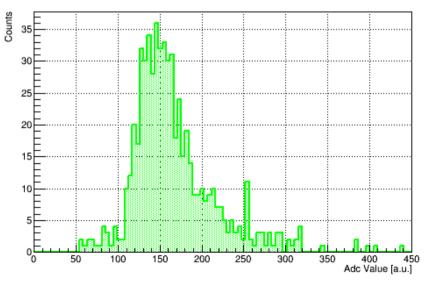
WireID for X and Y PLANES

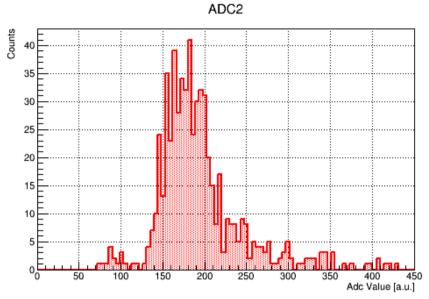


# ADCs

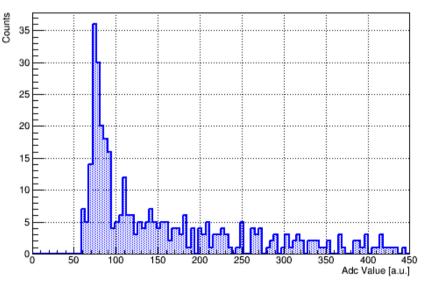












# CONCLUSIONS

We implemented a code to convert the output data file into a ROOT tree.

- + .dat file  $\rightarrow$  .txt file  $\rightarrow$  .root file
- ★ The test data files we used as input are the ones from the Cosmic Ray Telescope → new data files from the FTBF (M-test) should be avaiable in next future output data structure remains almost the same as for the Cosmic Ray Telescope.
  - \* The FTBF counters are now positioned along the M-Test beam line for beam particle definition/characterization.

## Thank you!

Many thanks to Flavio Cavanna, Ornella Palamara, Jennifer Raff, Douglas Jensen, Eric Church, Mitch Soderberg, Andrzej Szelc and Roberto Acciari for their precious help!

Many thanks to Bonnie Fleming, Sam Zeller e Gina Rameika for support for our stay at FNAL!

... and Many thanks to all the LArGroup!