

# The Gamma-400 space mission

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(on behalf of the Firenze, Pisa/Siena, Roma2, Trieste collaboration)

RD11 – 10<sup>o</sup> International Conference on Large Scale Application and  
Radiation Hardness of Semiconductor Detectors (6-8 Jul 2011)

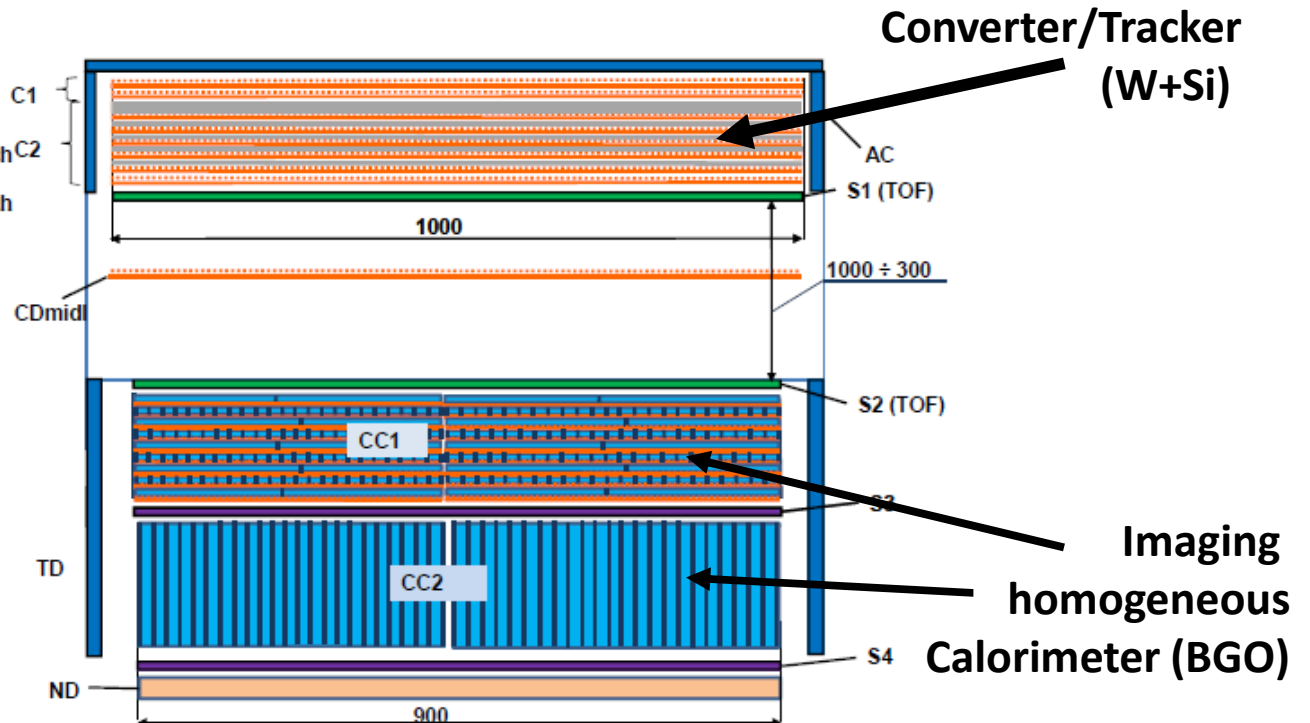
# Gamma-400

- “ Russian mission originally devoted to the study of high-energy (30 GeV – 3 TeV) gamma rays and high-energy electrons and positrons (Ginzburg).
- “ Italian PAMELA groups officially asked by the Russians to consider participation in the project (several meetings already).
- “ The characteristics of the satellite (scientific payload 2600 kg, power budget 2 kW, expected lifetime 10 years) provide excellent opportunities to configure the apparatus for accomplishing extremely important physics tasks:
  - **Study of the p and He spectra up to the “knee” region ( $10^{14}$  –  $10^{15}$  eV)**
  - **Extension of the gamma capability in the 50 – 300 MeV region**
- “ The Italian contribution to the project (with key responsibilities in hardware and analysis) would concentrate on these two items
- “ Preliminary simulation work already well established in Trieste, Florence, Pisa/Siena and Roma 2 (several possible configurations under study)

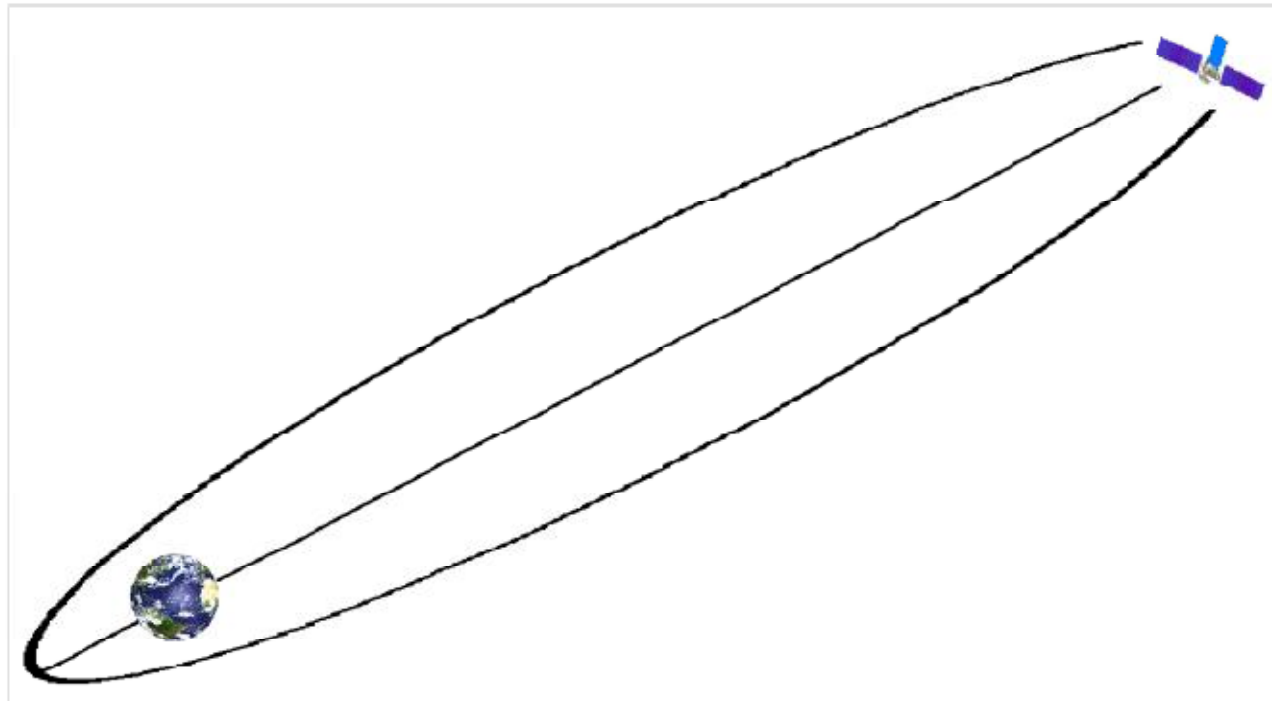
# GAMMA-400

- ❑ Russian Project for the detection of High Energy Gamma's and Electrons (RAS + MEPhI + ...)
- ❑ Angular resolution better than  $0.01^\circ$  (@100GeV) and energetic resolution  $\sim 1\%$  (>10GeV) (Fermi-GLAST ( $0.05^\circ$ ,  $\sim 10\%$ ))
- ❑ Very high apogee (300,000 km) elliptical orbit
- ❑ Approved by *ROSCOSMOS* ( 2.600kg, 2kW power, > 8 m<sup>3</sup> for the instrument)
- ❑ Availability for a revision of the project that does not alter the original objectives

- AC - anticoincidence detectors
- C1 - 2 layers x 2 single side Si 0.3mm (pitch 0.5 mm)
- C2 - Converter - total 1 Xo
  - 1 layer W 0.7 Xo (2.45 mm) +Si (x,y) 0.5mm (pitch 0.1mm)
  - 2 layers W 0.1 Xo (0.35 mm) +Si (x,y) 0.5 mm (pitch 0.1 mm)
  - 2 layers W 0.05 Xo (0.175 mm) +Si (x,y)0.5 mm (pitch 0.1 mm)
  - 1 layer Si (x,y) 0.5 mm (pitch 0.1 mm)
- CDmidl - Si (x,y) detectors (pitch 0.1 mm)
- S1, S2 - TOF detectors
- S3, S4 trigger detectors
- CC1 - imaging calorimeter (10Xo)
  - 10 layers BGO 1.12x2.24x23 + 1.12x2.24x22 cm (20+20)x4x10 crystals (WLS 1mm 2 layers x, 2layers y) Si (x,y)(pitch 1cm)
- CC2 - electromagnetic calorimeter BGO (14Xo)
  - 2.24x2.24x15.7cm - 400x4 crystals
- ND - neutron detector
- TD - 4 trigger detectors



## GAMMA-400 ORBIT



GAMMA-400 space observatory with the Navigator service module will be launched by Zenit-2SB rocket into a high-apogee orbit (apogee 300000 km, perigee 500 km, inclination  $51.8^\circ$ ).

## Gamma-400 main parameters

Gamma-ray energy range	0.1-3000 GeV
Calorimeter	90 x 90 cm <sup>2</sup> ~ 30 X <sub>0</sub>
Angular resolution (E <sub>γ</sub> ≥ 100 GeV)	~ 0.01°
Energy resolution (E <sub>γ</sub> ≥ 10 GeV)	~ 1%
Proton rejection	10 <sup>6</sup>
Telemetry downlink	100 GB/day
Power consumption	2000 W
Max. dimensions	2x2x3 m <sup>3</sup>
Total mass	2600 kg

# PHYSICS of COSMIC RAYS from SPACE

Search for signals from  
annihilation of dark matter



Differential fluxes of antimatter  
(antiprotons and positrons)

Search for anomalies in  
the electron spectrum

by

Magnetic spectrometers  
(**PAMELA** and **AMS-2**)

Study of origin and propagation  
of cosmic rays in the Galaxy



Differential fluxes of proton and nuclei

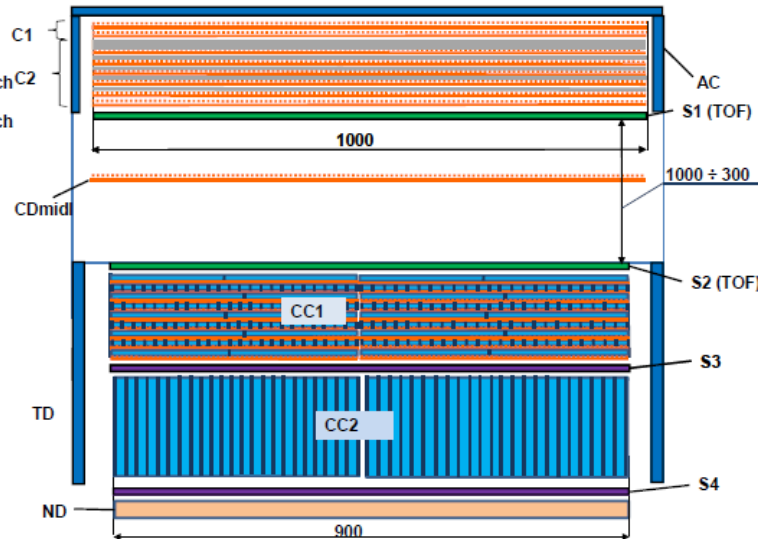
Measurement of the ratio between  
Primary and secondary cosmic rays

by

Magnetic spectrometers  
(**PAMELA** and **AMS-2**)

Calorimeters  
(**ATIC** and **CREAM** on balloons)

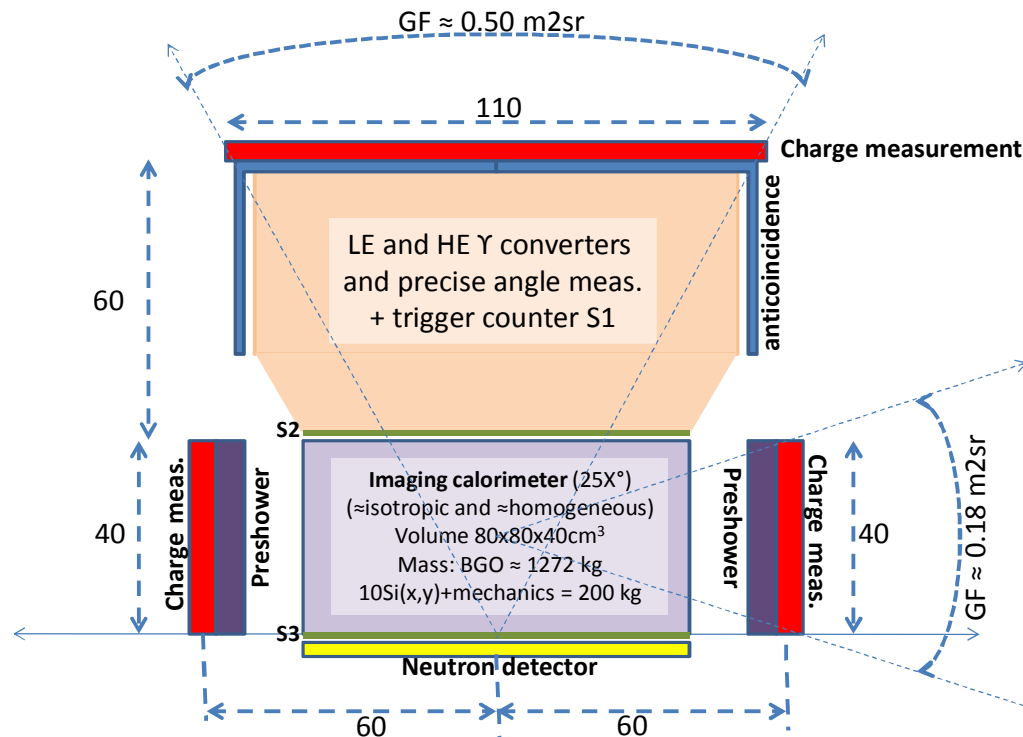
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## GAMMA 400 R&D

INFN teams from  
PAMELA, AGILE, FERMI-GLAST

Two years of R&D  
(project of the new instrument,  
simulations, prototypes)



+ Low energy gamma's  
( $E > 10 \text{ MeV}$ )

+ High energy nuclei up  
to the knee region with  
high geometry factor  
( $\sim 1 \text{ m}^2 \text{sr}$ )

## **Converter+tracker for good angular resolution**

1X° (5 layers: 0.7+0.1+0.1+0.05+0.05) of W  
interleaving 6 microstrip Si(x,y) layers (thick=0.5mm, pitch=0.1mm)  
+ 1 middle microstrip Si(x,y) layer (thick=0.3mm, pitch=0.1mm)

## **Calorimeter for good energy resolution**

Deep imaging calorimeter (>24X°)  
large area (>0.6m<sup>2</sup>) for pointing

## **Extention to low energy gamma's**

it must be added a number of very thin W layers on top,  
interleaved by Si(x,y) layers (thick=0.5mm, pitch=0.5mm)



## STUDY OF THE ORIGIN AND PROPAGATION OF COSMIC RAYS

@ knee until now only indirect measurements on ground:  
very difficult to understand the chemical composition



Direct measurements possible  
up to  $10^{15}$ - $10^{16}$  eV

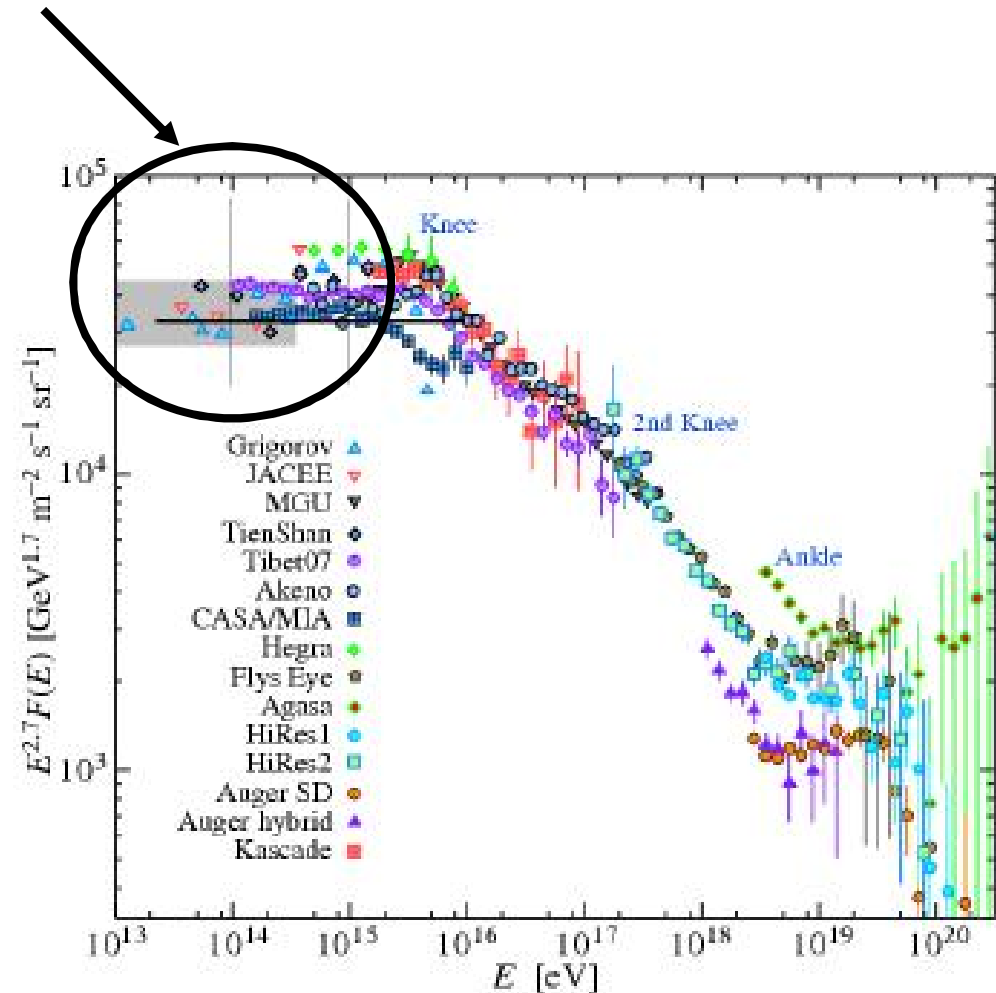
Deep calorimeter needed  
( $>25 X^\circ$ ,  $1.5\lambda$ )

Large acceptance x time needed  
( $GF \geq 1 \text{ m}^2 \text{ sr}$  for 5 years on satellite  
 $>10^3$  events in the  $10^{14}$ - $10^{15}$  eV)

Present situation:

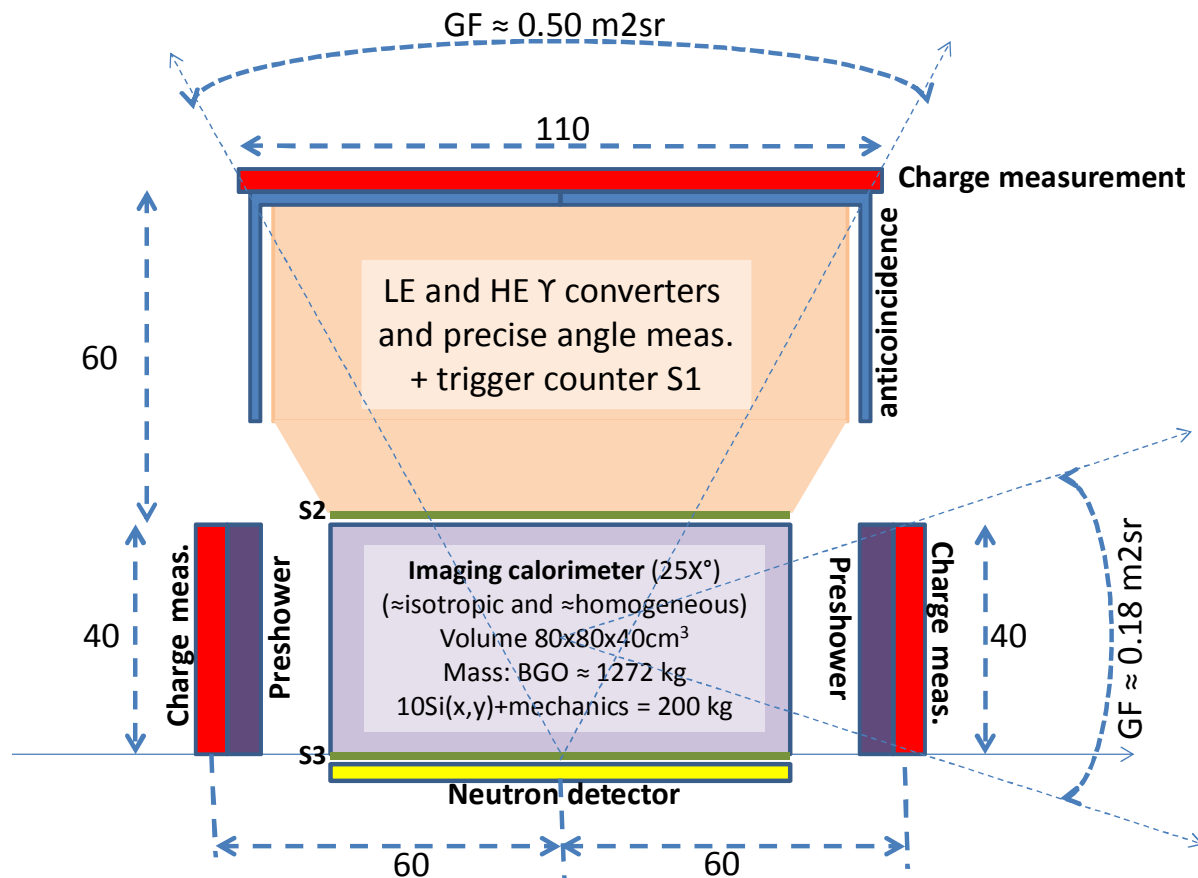
**AMS-2** calorimeter  $GF = 0.04 \text{ m}^2 \text{ sr}$

**CALET** experiment (2014-2018)  $GF = 0.12 \text{ m}^2 \text{ sr}$

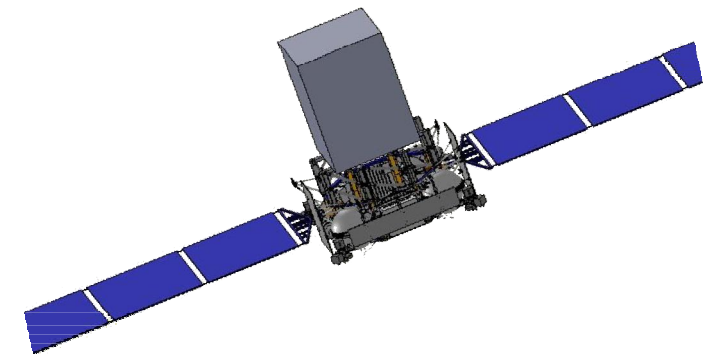


## Modifications for high energy protons and nuclei:

- Charge measurement: resolution and dynamics.
- Make use of the sides of the calorimeter for increasing the geometry factor



It must be developed an ISOTROPIC configuration of the instrument and of the calorimeter



Side views are free

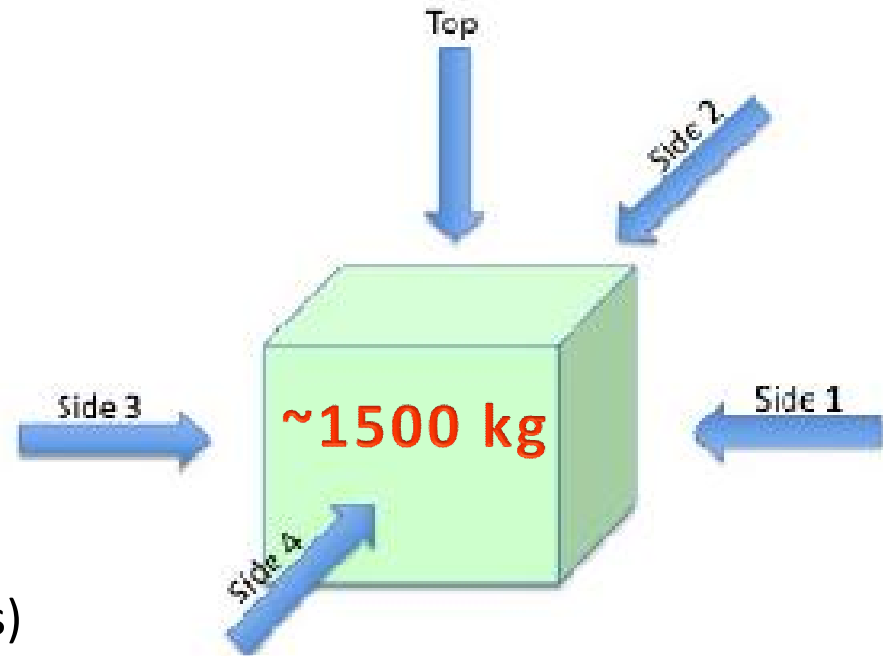
# Calorimeter for Gamma-400 - 1

“Simulation of different structures of the detector and their characterization for the measurement of nuclei and gamma’s

“Study and design of the calorimeter

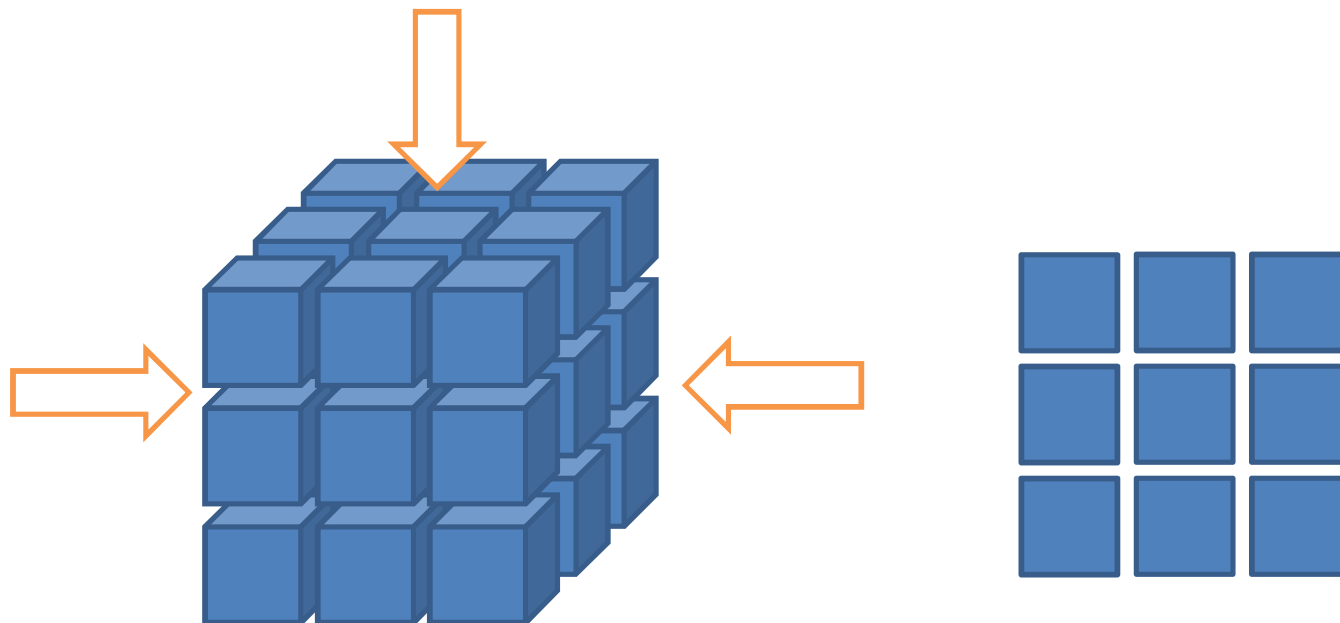
“Features :

- . Depth` ( $\sim 1.5 \lambda$  for protons)  
→ shower containment
- . high granularity ( $\sim 1 X_0$ )  
→ e/h separation (by shower topology)
- . Homogeneity  
→ Energy resolution (reduced fluctuations)
- . **Isotropy**  
→ **uniform response for the top and side particles**



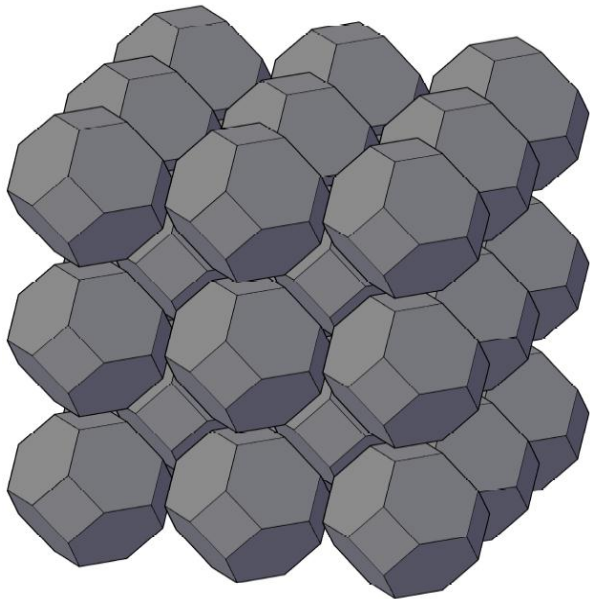
## Calorimeter for Gamma-400 - 2

- “ Design of a calorimeter composed by cubic ( $\sim 2 \times 2 \times 2 \text{ cm}^3$ ) BGO crystals
- . Study by simulations of the impact on the performance of the experiment.
  - . Design – Feasibility study (signal output) and realization and test of a prototype ( $\sim 4 \times 4 \times 8$  elements)-

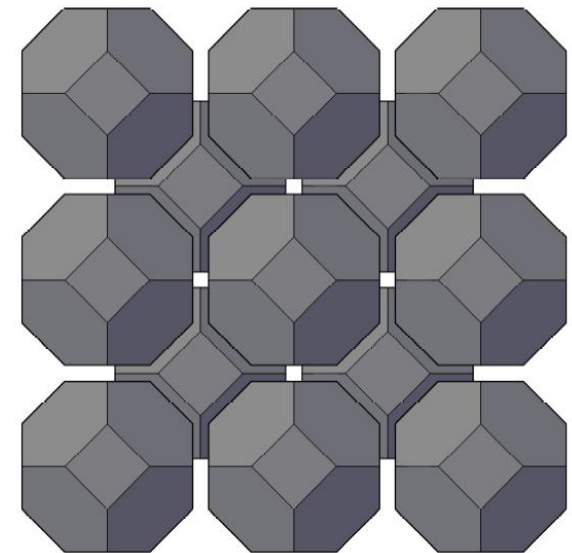


## Calorimeter for Gamma-400 - 3

- “ Feasibility study for a calorimeter composed by **truncated octahedron** shaped crystals.
  - . Study by simulation of the performance and comparison with the cubic lattice calorimeter
  - . Realization of crystal prototypes and test of single elements to verify the influence of the manufacturing on their properties.



“Uniform and cubic symmetric tessellation of the space.  
“Reduced dead zone in comparison to cubic lattice.  
“Possibility of increasing the GF spacing the crystals, for equal mass.



## Aim of the R&D activities:

### Design of a multipurpose instrument for the Gamma-400 mission:

High precision measurements of high energy gamma's  
(source position, E and E variation)

Extention of gamma's measurements down to <30MeV  
(source position, E and E variation)

High precision measurement of electron spectrum (up to > 3 TeV)

Measurement of the energy spectrum of protons beyond 2 PeV

Measurement of the energy spectrum of helium beyond 1 PeV/nucleon

Measurement of the spectra of nuclei up to actinides

### With the following main parameters:

Mass=2600 Kg

Power=2kW

# of channels  $\approx 5 \cdot 10^5$

Thank you for your attention