

ACTIVE-STERILE NEUTRINO OSCILLATIONS IN DENSE ENVIRONMENTS

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Proposta di Use Case per il WP 1 del CN-HPC del PNRR

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SCIENTIFIC CASE

Neutrino flavor conversion in dense environments (e.g. **early Universe, Supernovae...**)

- Theoretically interesting
- Phenomenological implications:
 - Primordial nucleosynthesis and light elements yields
 - Cosmological density of relic neutrinos, both active and sterile (neutrino masses, effective number of species)
 - Nucleosynthesis of heavy nuclei in stars
 - Explosion of massive stars

FORMALISM

2+1 density matrix

$$\rho(x, y) = \begin{pmatrix} \rho_{ee} & \rho_{e\mu} & \rho_{es} \\ \rho_{\mu e} & \rho_{\mu\mu} & \rho_{\mu s} \\ \rho_{se} & \rho_{s\mu} & \rho_{ss} \end{pmatrix}$$

Equation of motion:

$$i \frac{d\rho}{dx} = + \frac{x^2}{2m^2 y \overline{H}} [M^2, \rho] + \frac{\sqrt{2}G_F m^2}{x^2 \overline{H}} \left[\left(-\frac{8 y m^2}{3 x^2 m_W^2} E_\ell - \frac{8 y m^2}{3 x^2 m_Z^2} E_\nu + N_\nu \right), \rho \right] + \frac{x \widehat{C}[\rho]}{m \overline{H}}$$

Vacuum oscillations

Matter effects
(e+ and e-)

Matter effects
(neutrino self
interactions)

Collisions

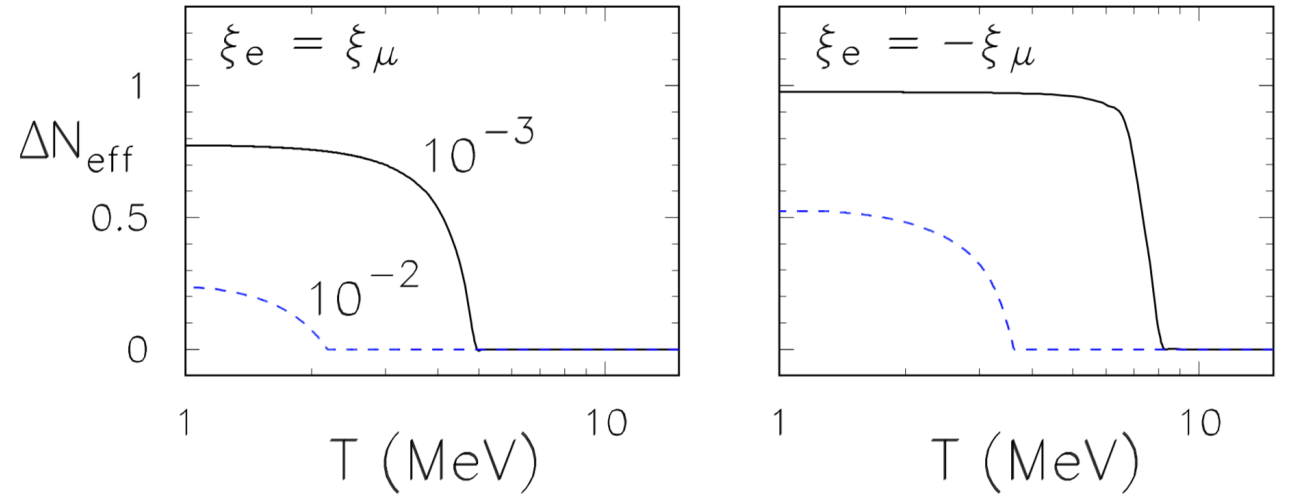
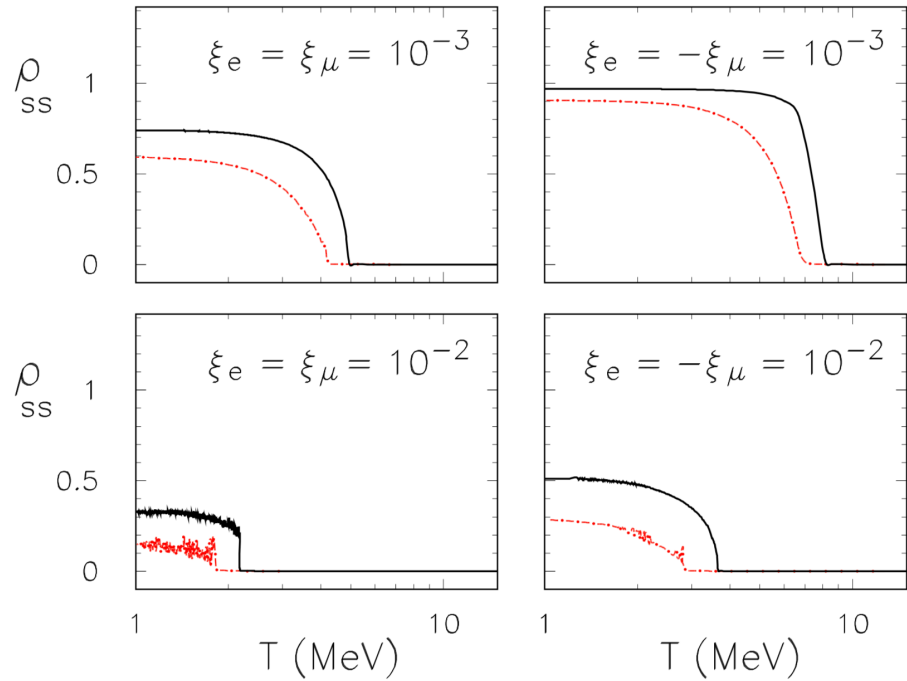
+ similar equation for the density matrix of antineutrinos
+ continuity equation

OBJECTIVE AND BENCHMARKS

We aim at being able to solve the equation of motion for ρ :

- Using ~ 10 corehrs
- With a 10^{-2} absolute precision on N_{eff}
- Multimomentum
- Multiflavour
- Accounting for lepton asymmetries

... with the final objective of performing a scan in parameter space (asymmetries+mixing angles+masses).



N. Saviano et al., PRD 2013

Possible strategies (not mutually exclusive...) to bring down the time requested to solve the EoM at fixed parameter values:

- Optimization of the solver
- Choice of the momentum discretization scheme
- Use emulators for the computation of collision terms

+ use emulators to solve the EoM when performing the MC ?

PERSONNEL AND REQUESTS

Personnel

- M. Lattanzi (INFN Ferrara, staff)
- N. Saviano (INFN Napoli, staff)
- + assegnista/RTD to be recruited with open call at SSM (?)

Requests

- Access to a small cluster for development purposes (200 kch)
- Access to an HPC cluster for test and production purposes (1Mch)

Possible extensions

Neutrino oscillations in SN

Axion-photon conversion in cosmological/astrophysical magnetic fields