A Statistical Formulation of

Conjunction Assessment

Soumaya Elkantassi¹ and Anthony Davison²

¹Department of Operations, Université de Lausanne (UNIL)

²Institute of Mathematics, École polytechnique fédèrale de Lausanne (EPFL)

soumaya.elkantassi@unil.ch

Inil

UNIL | Université de Lausanne

Motivation

With the new space race

- Hundreds of satellites are launched into Earth's orbit every year for different purposes.
- In total, there are nearly 6,000 satellites circling our planet and this number keeps growing.
- About 60% of these satellites are defunct, contributing to the growing problem of space debris or "space

2) $P\{\widehat{p}_c < p_c(\xi)\} > 1/2$, i.e., \widehat{p}_c is too small more often than not. -og10 probability

2.2 Higher-order likelihood theory

For continuous responses, third-order inference can be based on limiting N(0,1) distribution of the modified likelihood root

$$r^*(\psi) = r(\psi) + \frac{1}{r(\psi)} \log\left\{\frac{q(\psi)}{r(\psi)}\right\},\,$$

where $q(\psi)$ depends on the model, and for which

$$P\{r^*(\psi_0) \le r^*_{obs}\} = \Phi(r^*_{obs}) + O(n^{-3/2}).$$

junk".

- Managing the increasing number of space objects poses a significant challenge in collision avoidance.

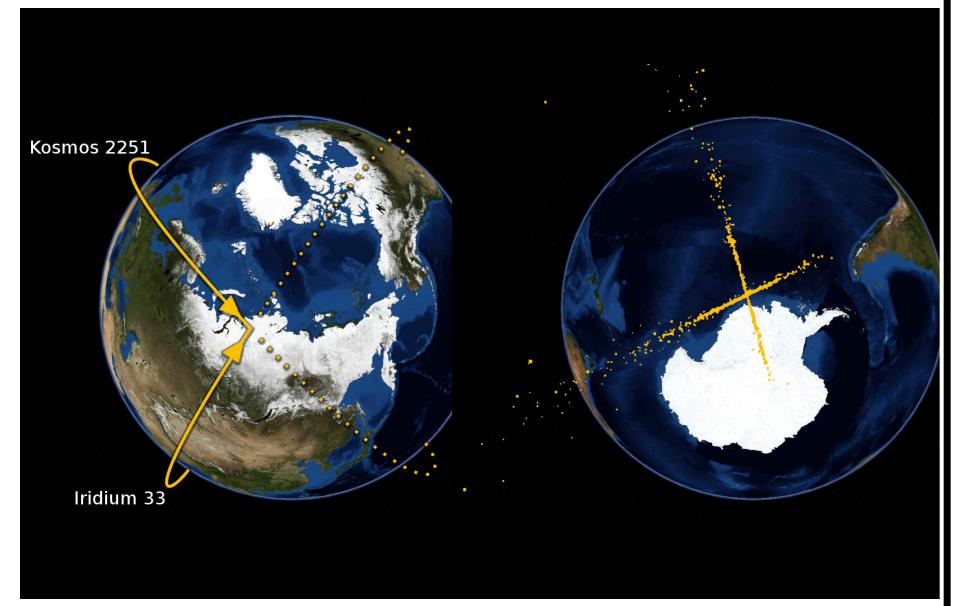


Figure 1: Collision of Kosmos 2251 and Iridium 33 on February 10, 2009.

In this work

- → We point out that the usual collision probability estimate can be badly biased.
- → We formulate an approach to satellite conjunction assessment based on a statistical model.
- → We discuss inference on the miss distance between

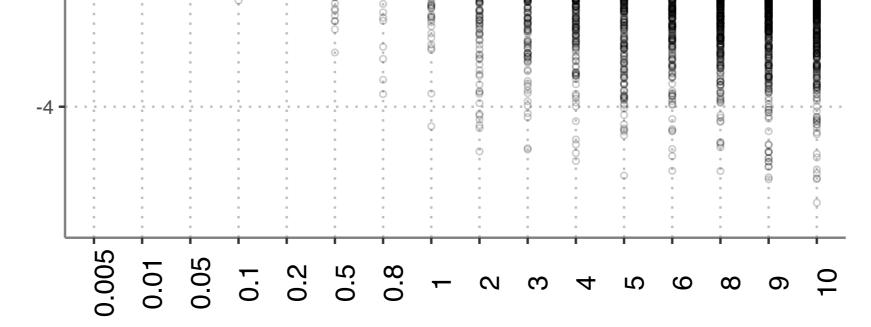
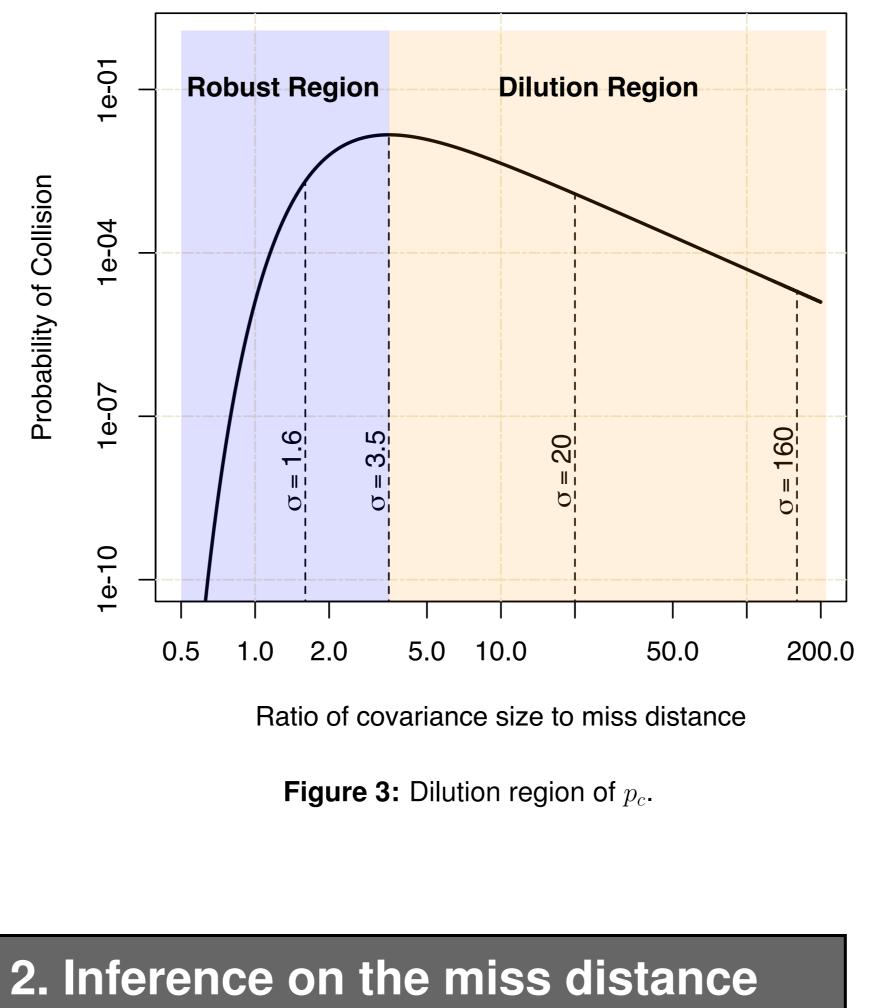


Figure 2: Simulated \hat{p}_c when $\psi \approx 12$, HBR=20m. The blue lines show $p_c(\xi)$, the red the averaged \hat{p}_c .

3) Probability of collision has a dilution region, seen in the literature as a 'paradox' in need of clarification, see [1] on false confidence.



3. Decision making

Aim for interpretation of significance probabilities to be like \widehat{p}_c , we take

$$H_0: \psi = \psi_0 \quad H_A: \psi > \psi_0.$$

We hope that the 'null' will be rejected so we don't have to take evasive action. Under H_0 , we expect

small $1 - \Phi \{r(\psi_0)\} \equiv \text{ small } \widehat{p}_c$.

Proposed procedure:

- \rightarrow choose ψ_0 to represent a 'safe' minimum distance,
- \rightarrow choose ε (typically 10^{-4} but it could depend on potential losses),
- \rightarrow if $1 \Phi \{r^*(\psi_0)\} > \varepsilon$, then plot $\Phi \{r^*(\psi)\}$ as measure of evidence about ψ relative to ψ_0 ,

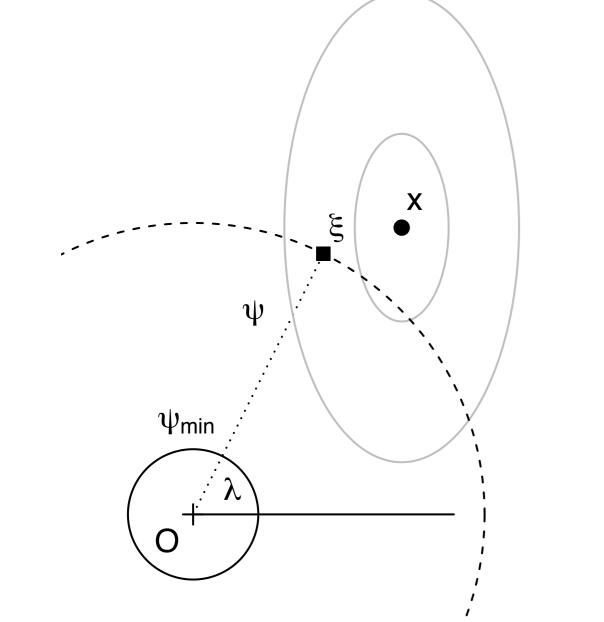
 \rightarrow decide whether to take evasive action or not.

4. Example: The 2009 event

two space objects.

1. Probability of collision

We consider two space objects in close conjunction described relative to an origin 'O' at the first satel-We assume that the velocity vector is known, lite. and we consider x and ξ as the projections of the observed position and the true position into the plane perpendicular to the velocity vector (encounter plane), i.e., $x = (x_1, x_2)$ and $\xi = (\xi_1, \xi_2) = (\psi \cos \lambda, \psi \sin \lambda)$.



2.1 First-order likelihood theory

The measured data in the encounter plane is assumed to be normally distributed $x \sim \mathcal{N}_2(\xi, D)$, where D is a known diagonal matrix. The minimum distance between the two objects is

 $\psi = \|\xi\| \cos\beta,$

where $\beta \in (\pi/2, 3\pi/2)$. For $\vartheta = (\psi, \lambda)$, the log likelihood is

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• $\hat{p}_c = 1.14 \times 10^{-5}$ (no action)

• Significance probabilities for $\psi_0 = 20 \text{ m}$ are 1.2×10^{-3} (Wald and r) and $7.2 \times 10^{-3} (r^*)$ (take action).

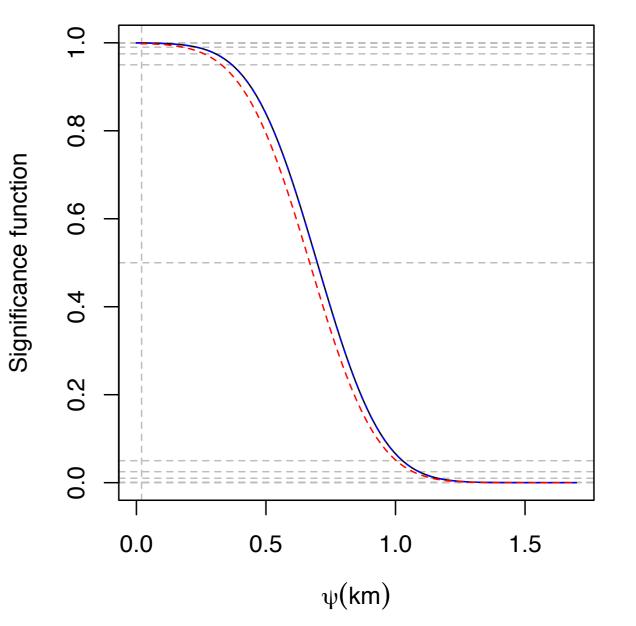


Figure 4: Significance function for the US and Russian satellite collision event.

✤ Promising results + can be used in parallel with conventional tools, see [3].

Most existing literature focuses on the probability of collision p_c , i.e., the integral of the density of x over the Hard body Ratio (HBR or ψ_{min}) disk, i.e.,

$$\widehat{p}_c = p_c(x),$$
 where $p_c(\xi) = \int_{\{x': \|x'\| \le \psi_{\min}\}} f(x';\xi) dx'$

which can be seen as a plug-in estimate of $p_c(\xi)$. What is the problem with \hat{p}_c ?

1) $p_c(x)$ is the probability that the observation x lies inside the HBR, not the probability that the second object will hit the first object.

 $\ell(\vartheta) = -\frac{1}{2}(x-\xi)^{\mathrm{T}}D(x-\xi).$

Computations are more complicated in the general 6D case.

If ψ scalar, first order inference based on limiting N(0,1) laws of

*likelihood root, $r(\psi) = \operatorname{sign}(\widehat{\psi} - \psi)[2\{\ell(\widehat{\psi}) - \ell(\psi)\}]^{1/2}$, *Wald statistic, $t(\psi) = j(\widehat{\psi})^{1/2}(\widehat{\psi} - \psi).$

as $n \to \infty$ have, for instance,

$$\mathrm{P}\left\{r(\psi) \leq r_{\mathsf{obs}}\right\} = \Phi\left(r_{\mathsf{obs}}\right) + O\left(n^{-1/2}\right).$$

which yields tests and confidence sets for ψ_0 based on an observed $r_{obs} = r(\psi_0)$.

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