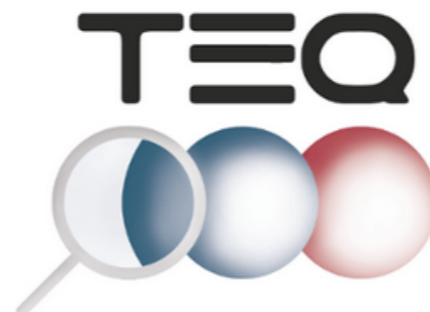


ENHANCING SDD ENERGY RESPONSE WITH ML AND DIFFERENTIAL PROGRAMMING

Fabrizio Napolitano for the VIP Collaboration



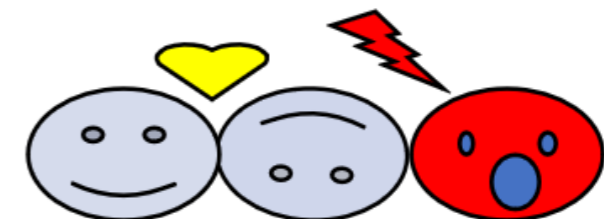
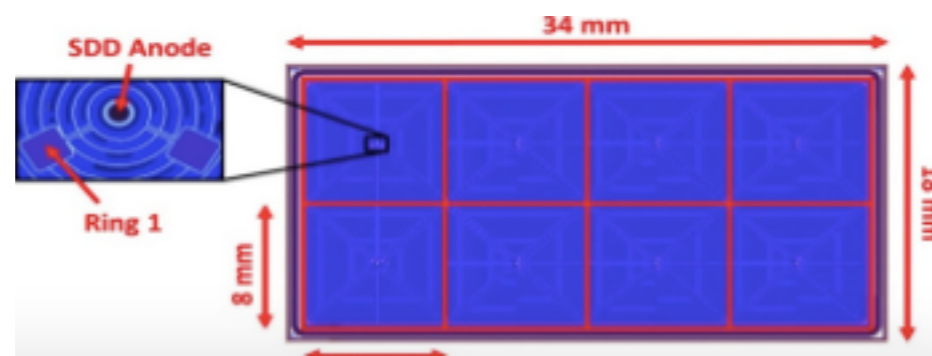
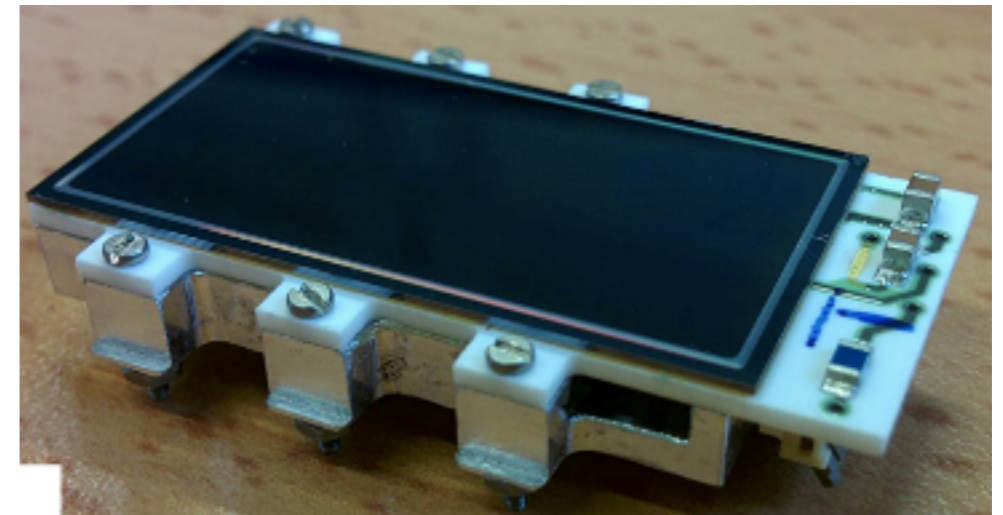
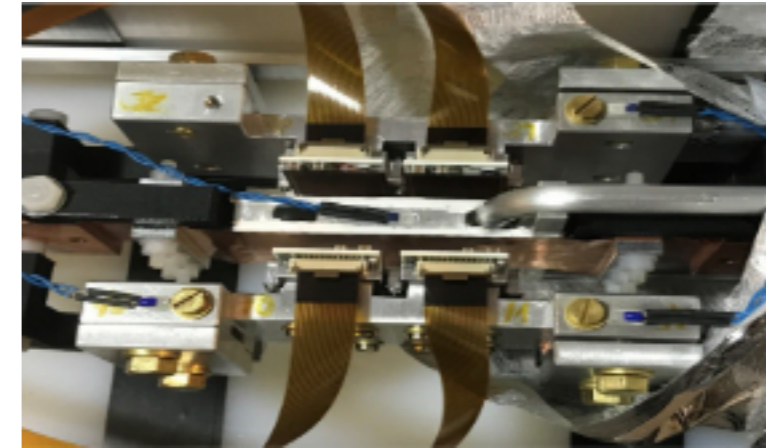
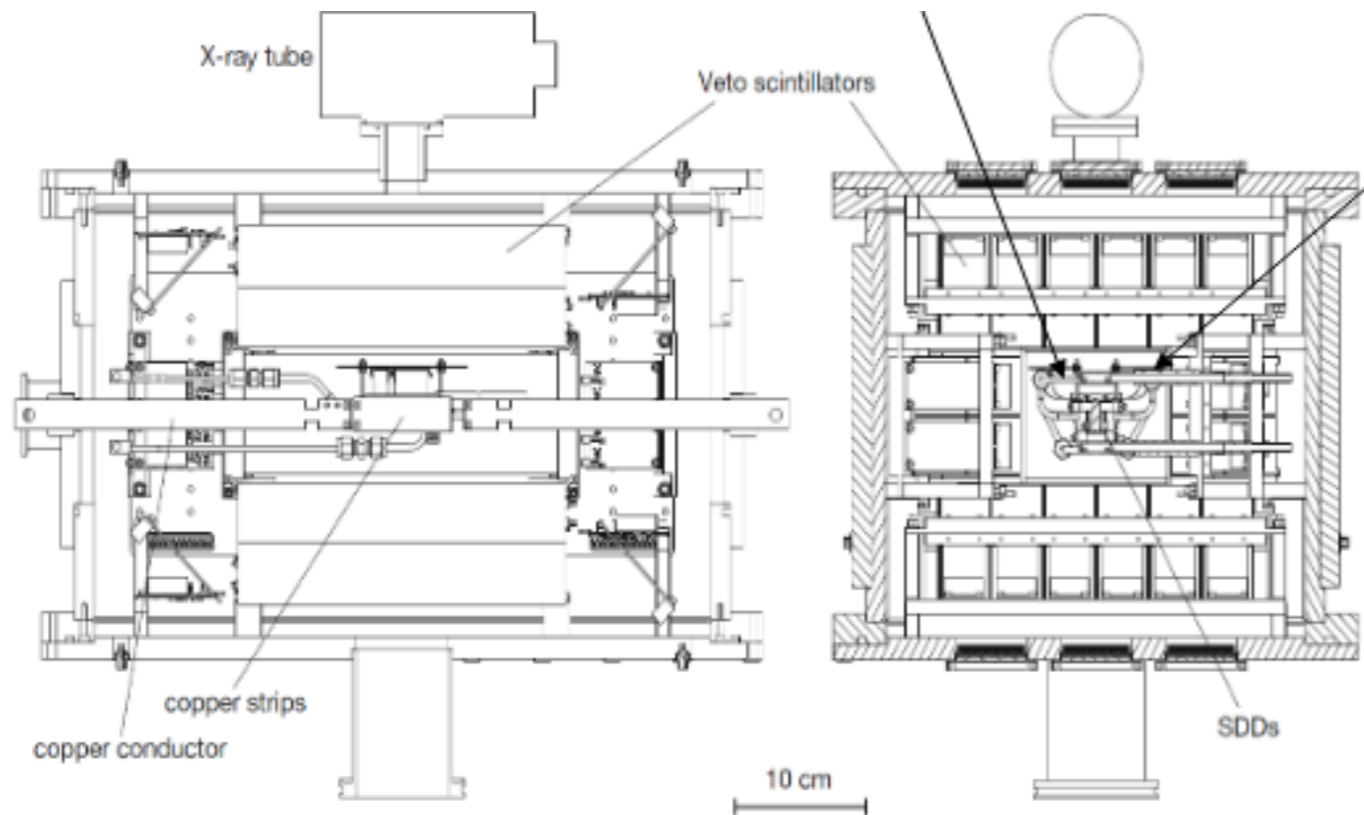
fabrizio.napolitano@lnf.infn.it

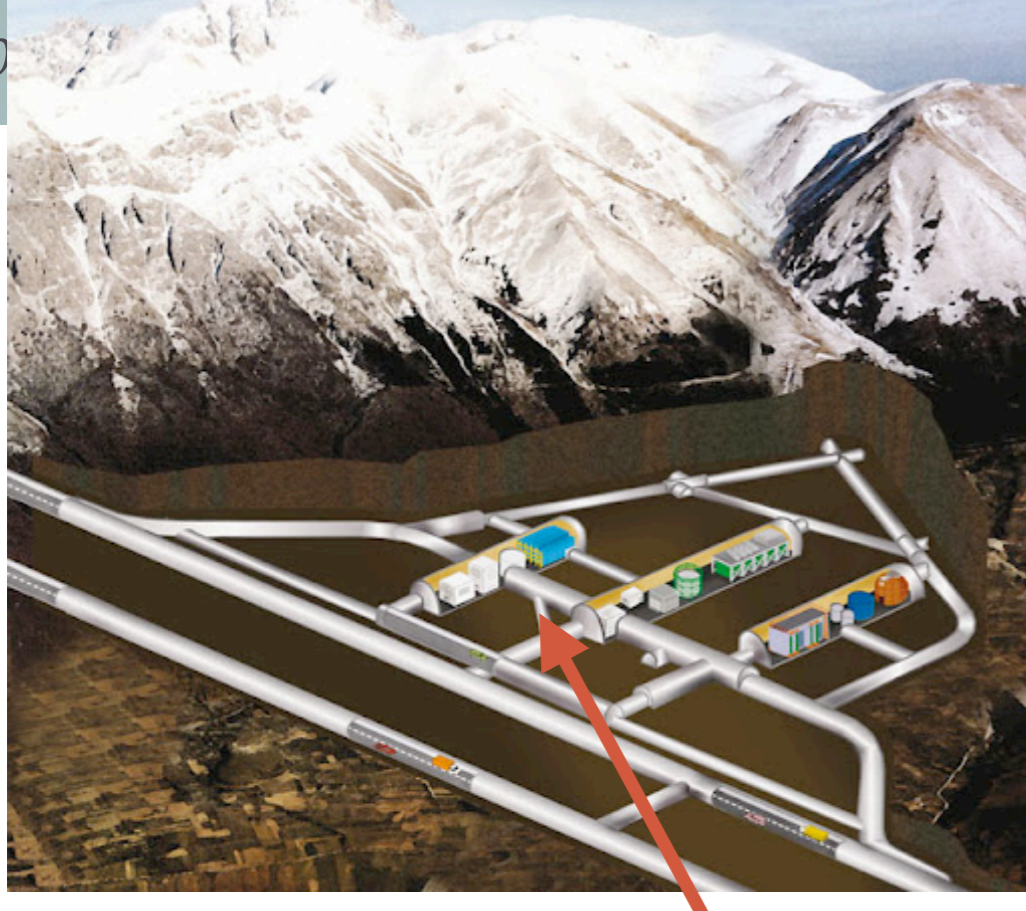
HPXM2023

June 19th to 23rd at Laboratori Nazionali di Frascati (LNF - INFN), Italy

The VIP-2 Experiment

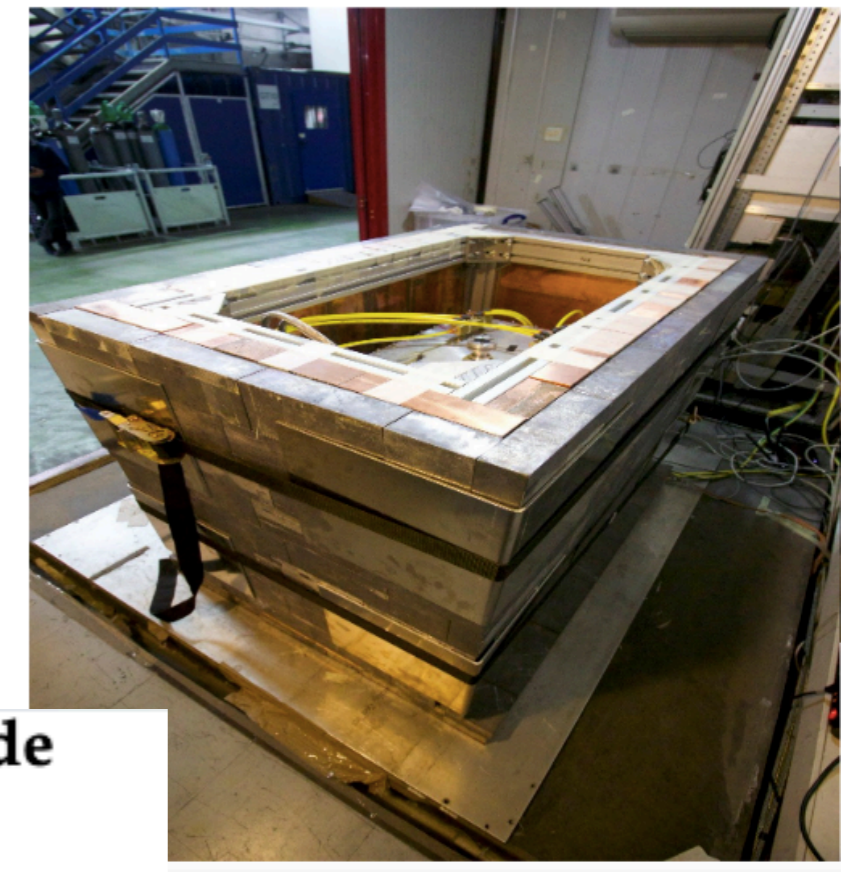
Silicon Drift Detectors (SDDs) higher resolution (190 eV FWHM at 8.0 \rightarrow keV), faster (triggerable) detectors. 4 arrays of 2 x 4 SDDs 8mm x 8mm each, liquid argon closed circuit cooling 170 °C



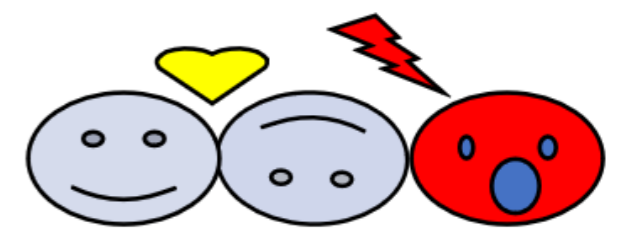
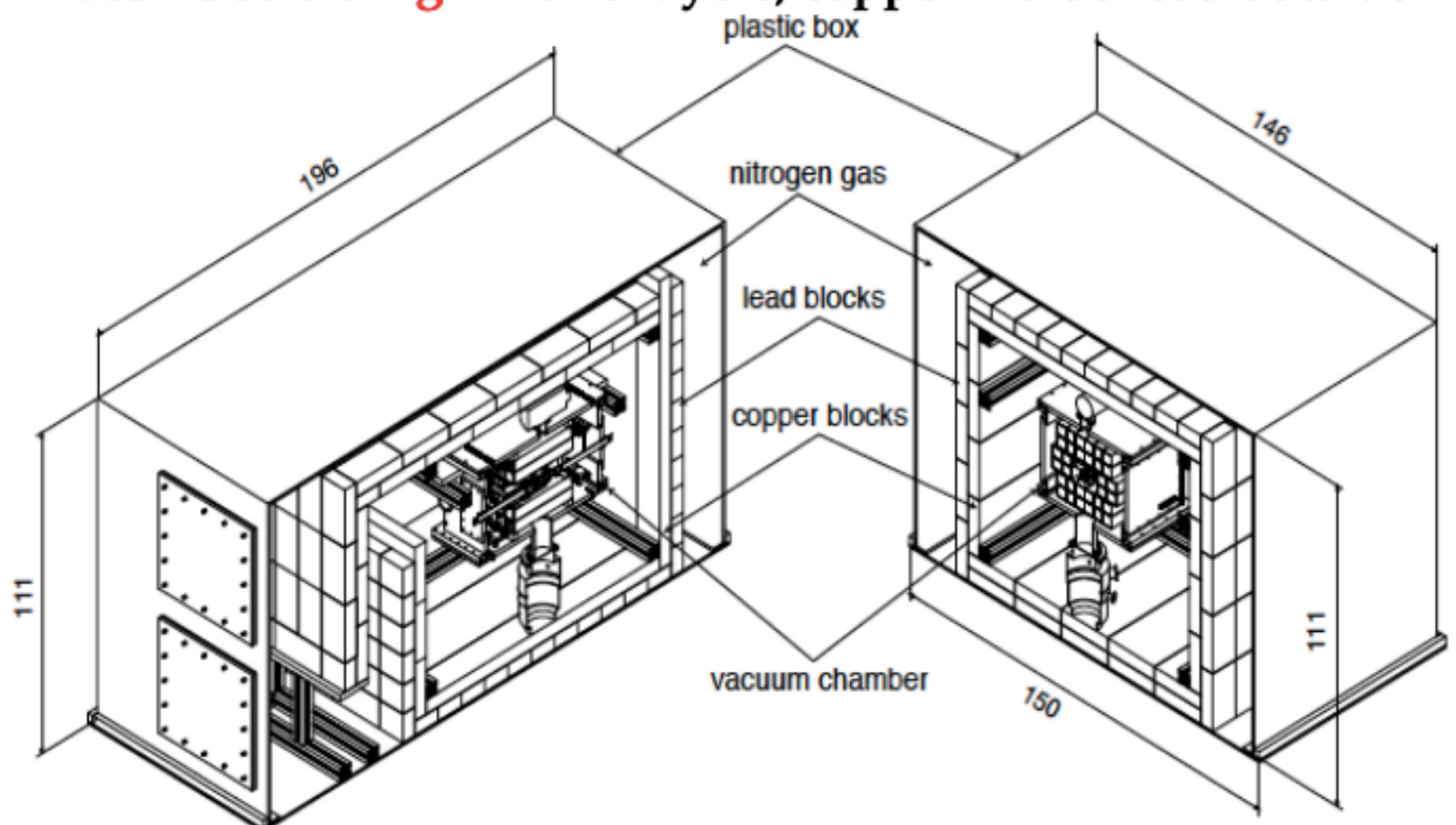


1400 m rock coverage

Upgrade concluded in April 2019:



Passive shielding → two layers, copper inside lead outside





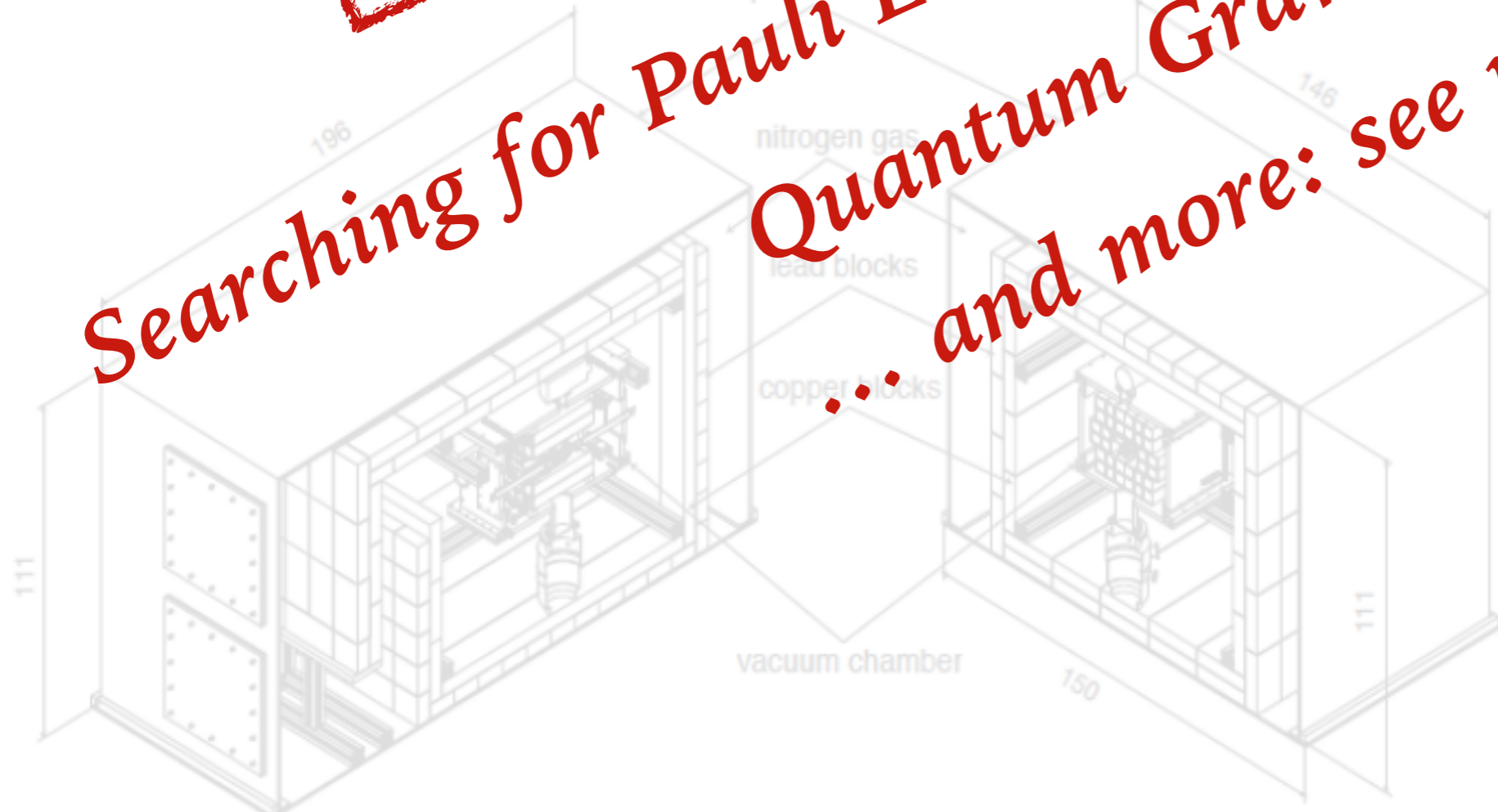
Upgrade concluded in April 2019:

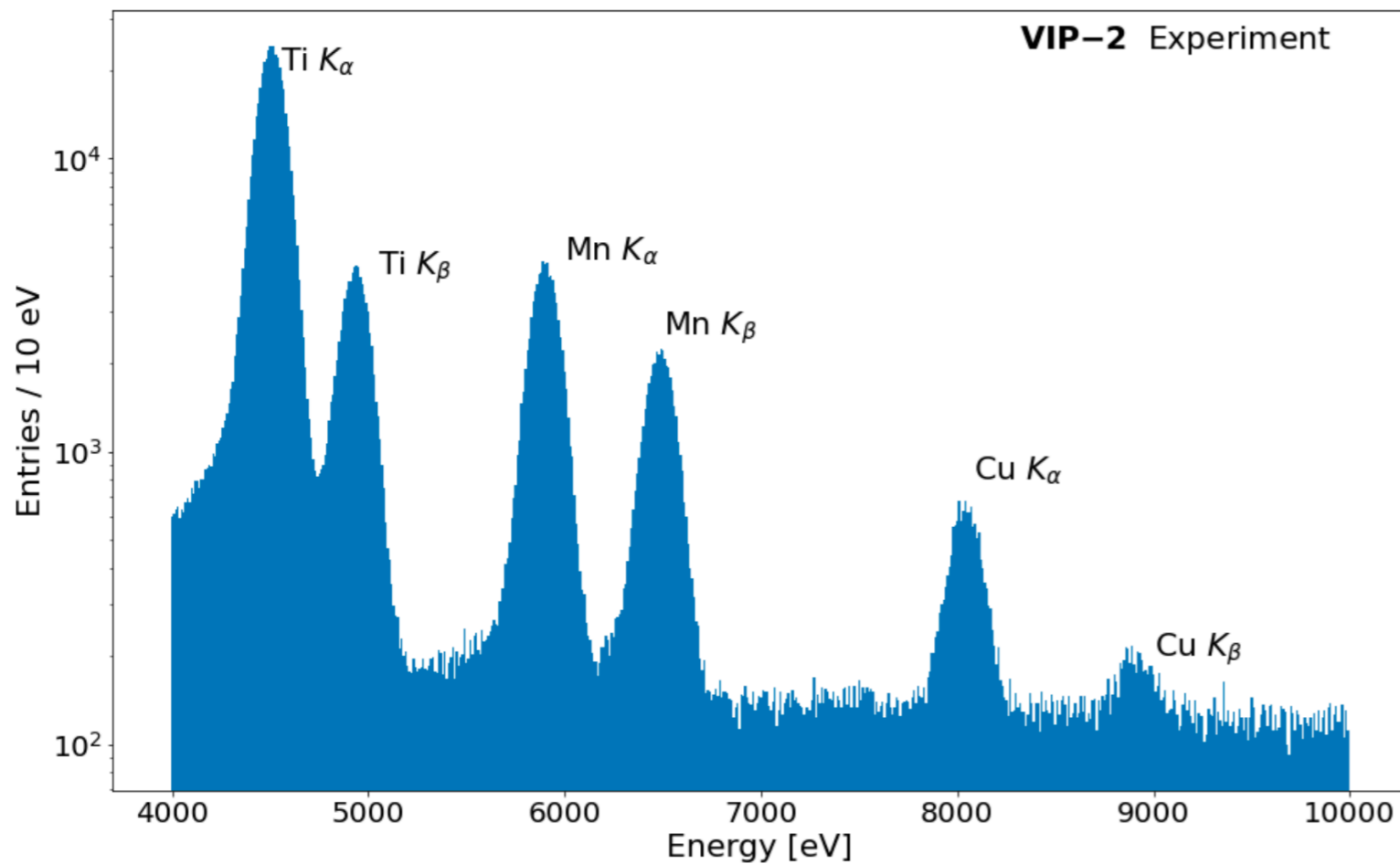


**Ideal place for precision
X-ray spectroscopy**

Passive shielding → two layers, copper inside lead outside

**Searching for Pauli Exclusion Principle
Quantum Gravity
... and more: see next talks!**





Calibrated spectrum of 4 SDD arrays.

Not easy to calibrate because:

- *Copper line at orders of magnitude smaller than Ti and Mn*
- *Tiny distortions of FEE*

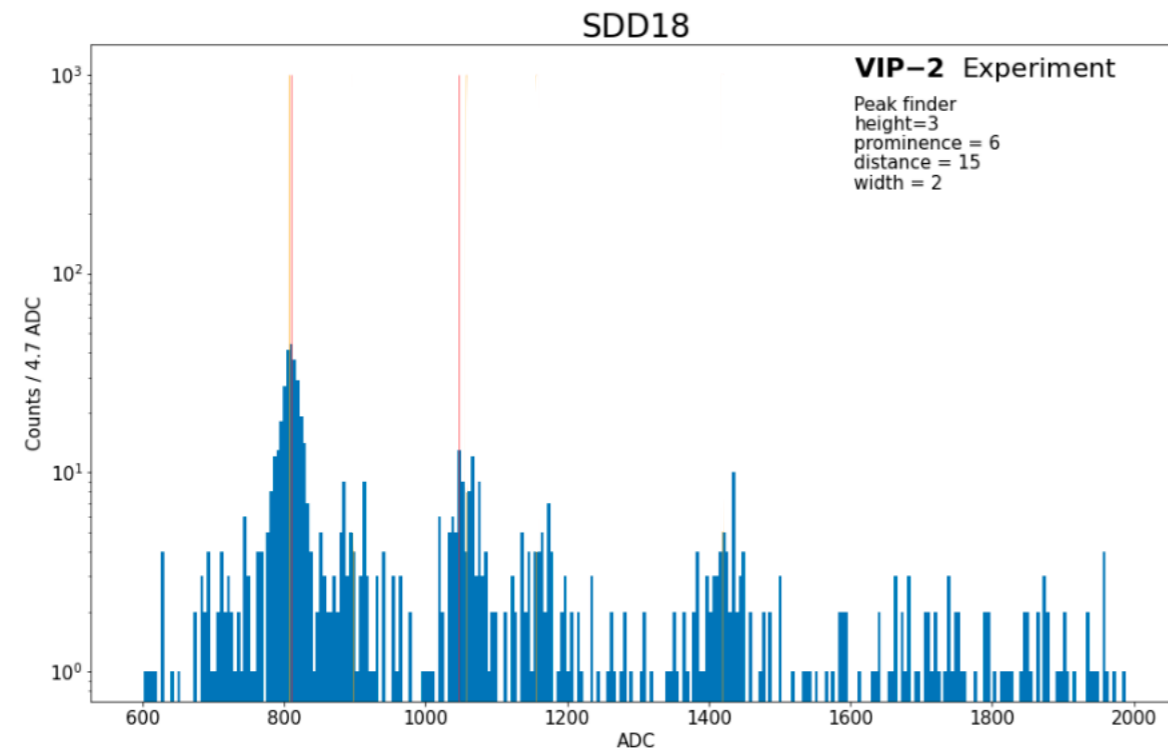
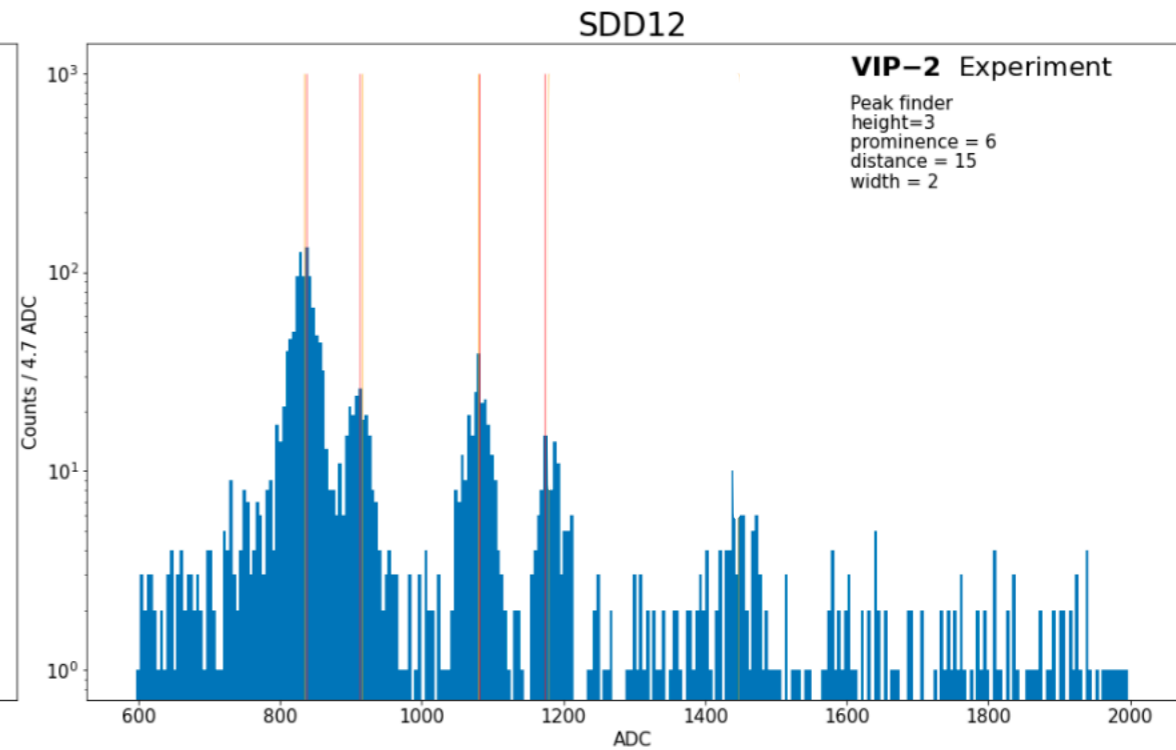
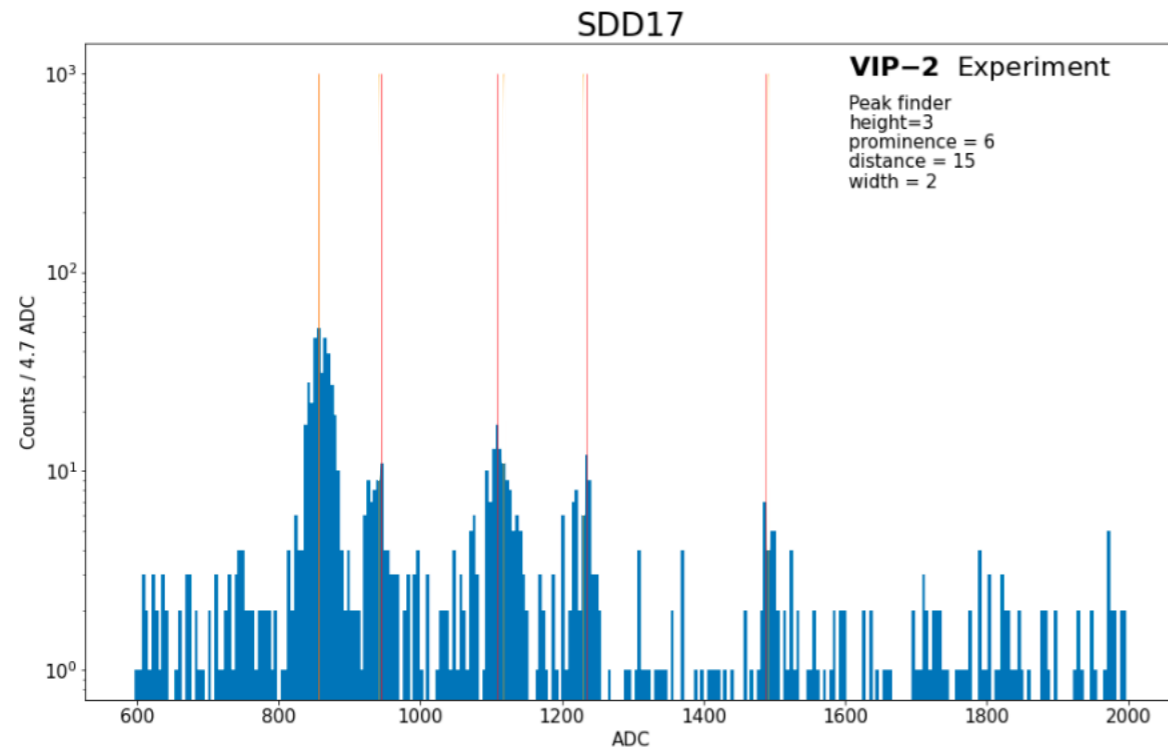
Calibration can be done in big or small batches

Big batches

Can determine better the Copper position but cannot capture fluctuations

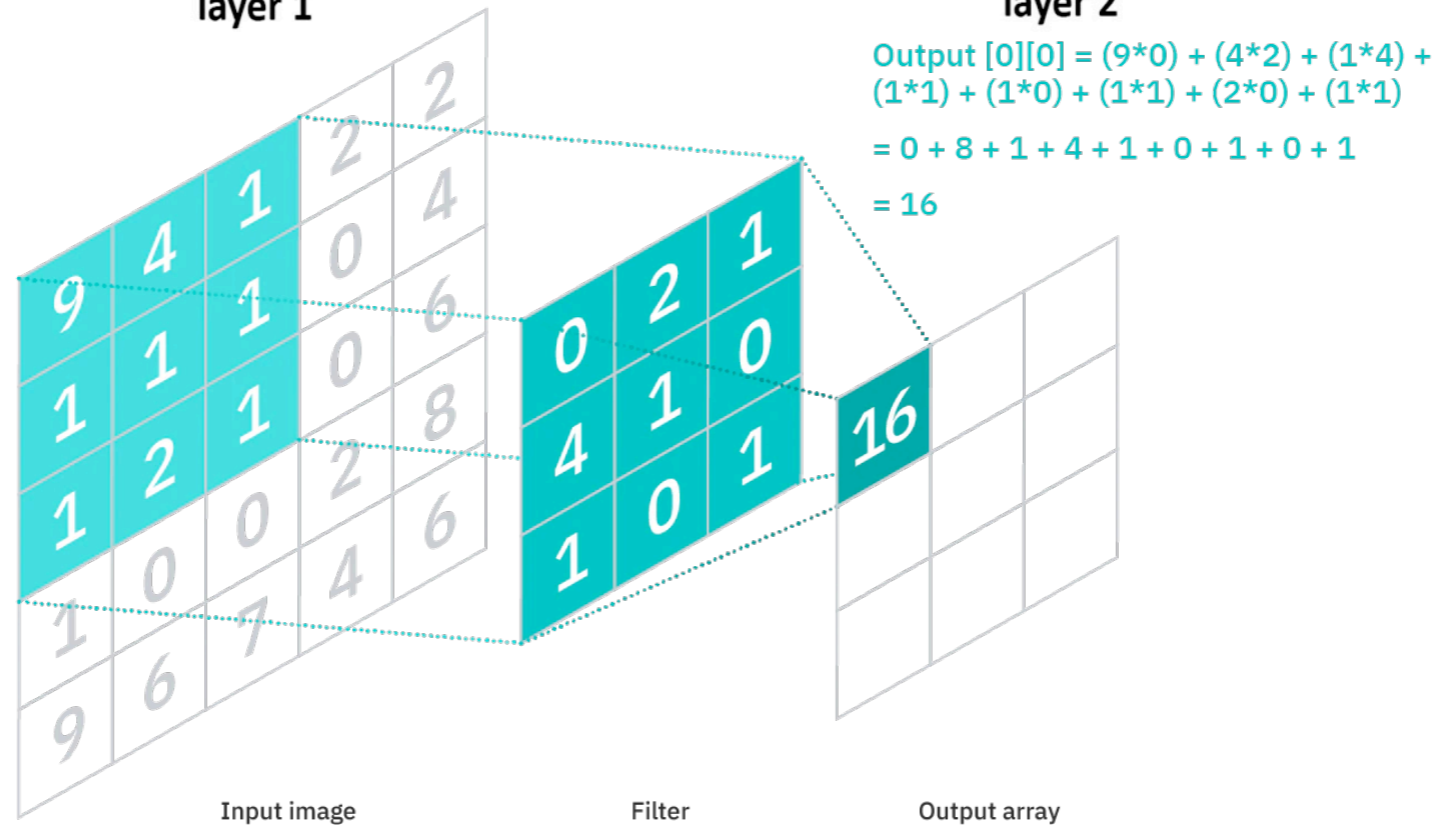
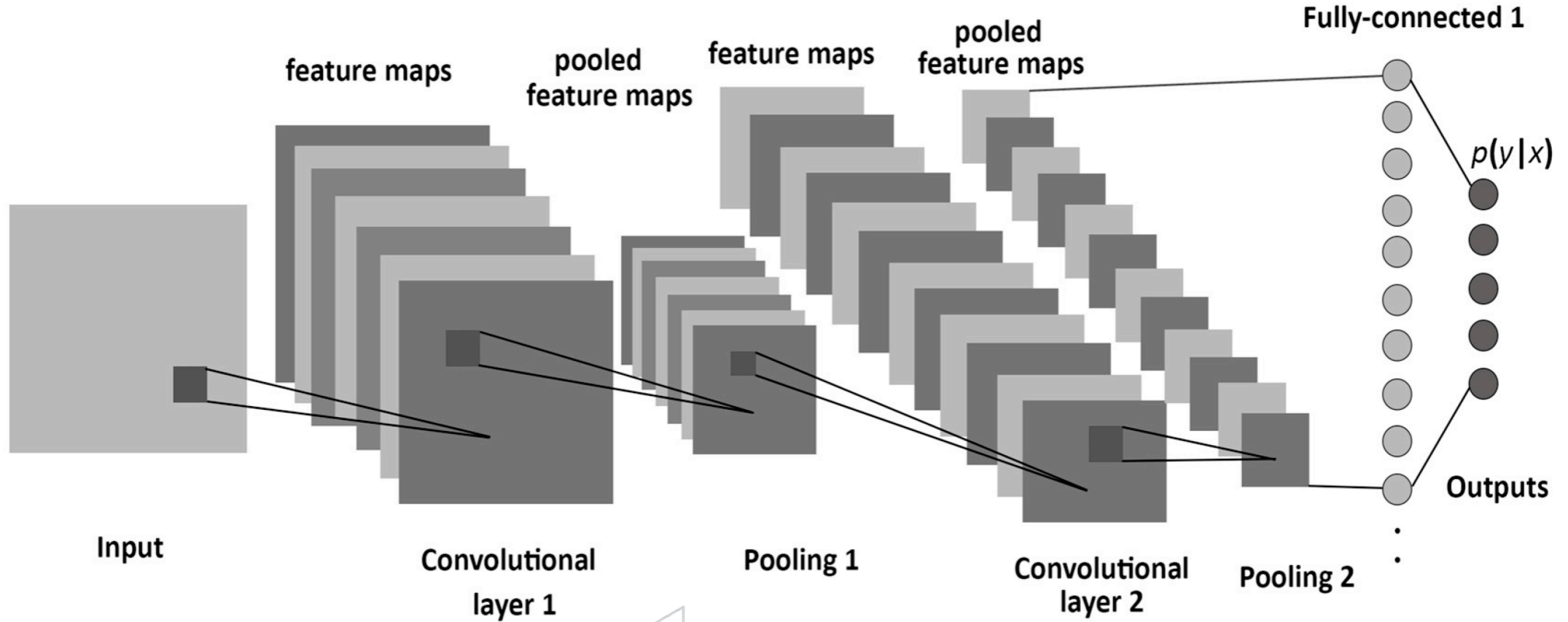
Small batches

Can capture fluctuations but cannot determine the Copper position well



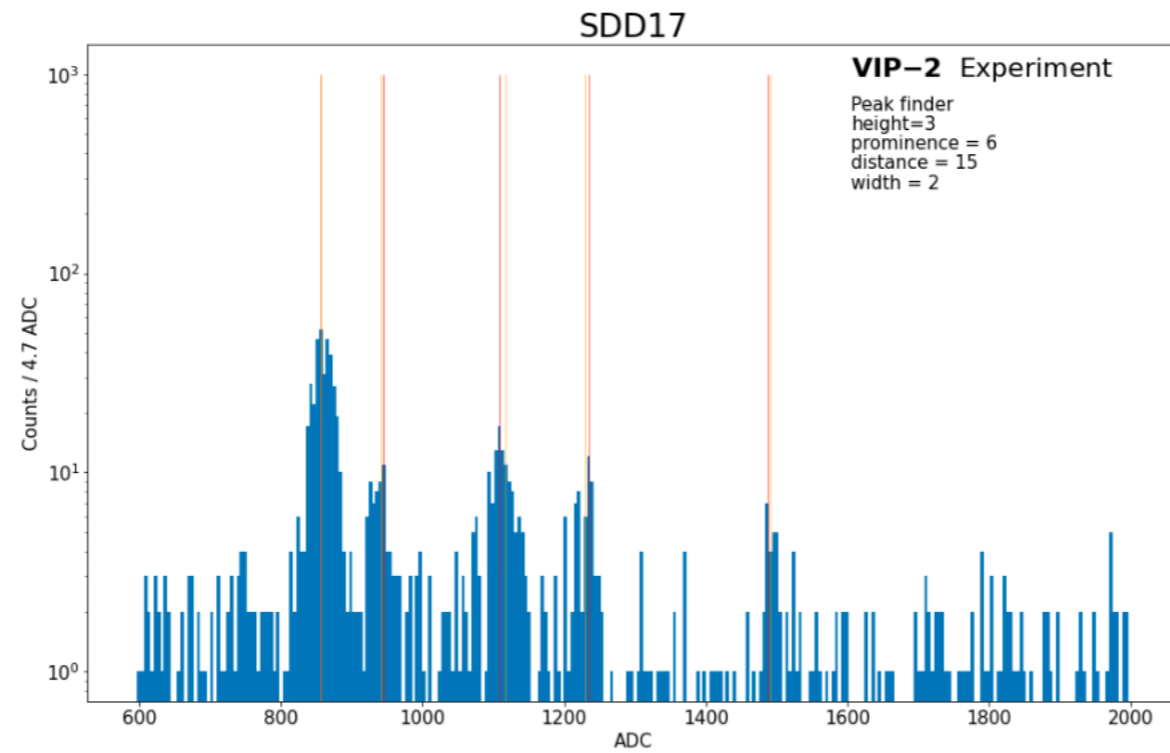
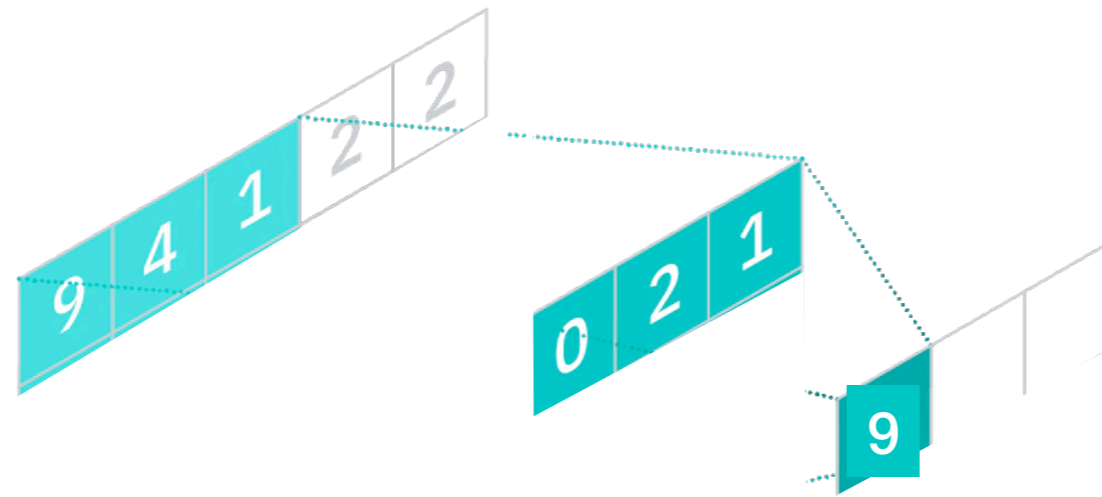
*Statistical fluctuations
 at low counts can make
 the use of peak finder
 algorithm tricky to setup
 needs constant care
 calibrated to be resilient
 algo params need to be
 tuned*

Use two step approach - 1st: convolutional neural network as peak finder

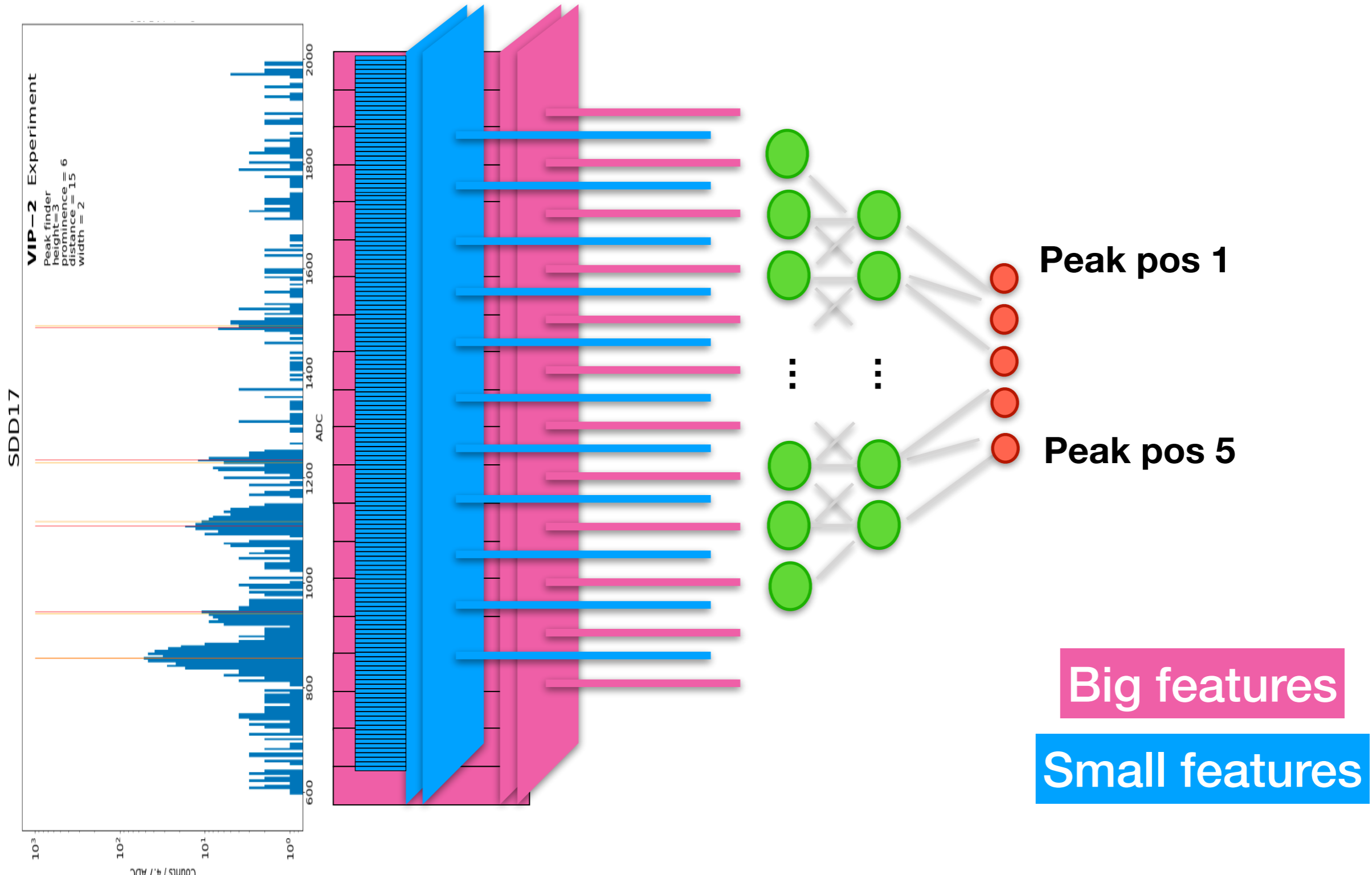


https://www.mdpi.com/1099-4300/19/6/242

Use two step approach - 1st: convolutional neural network as peak finder

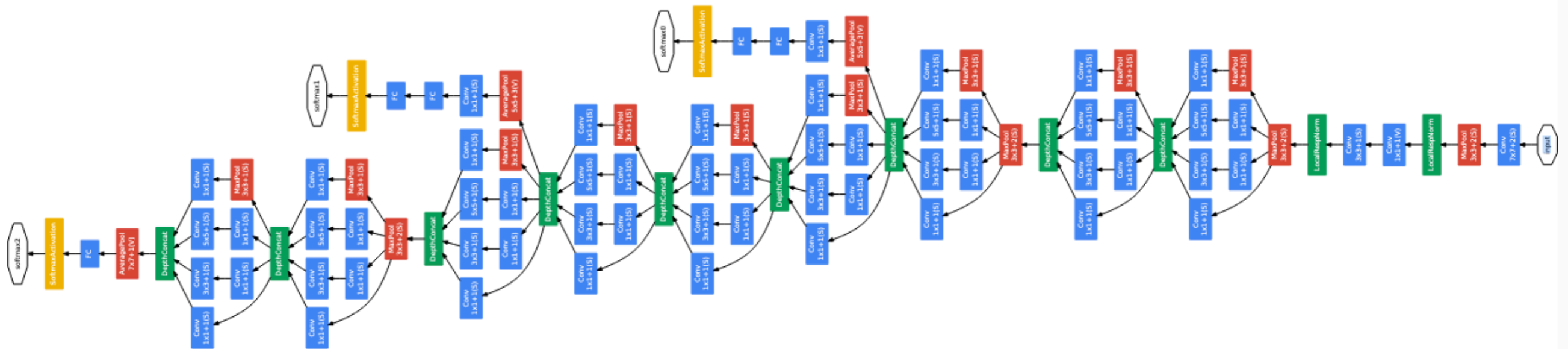
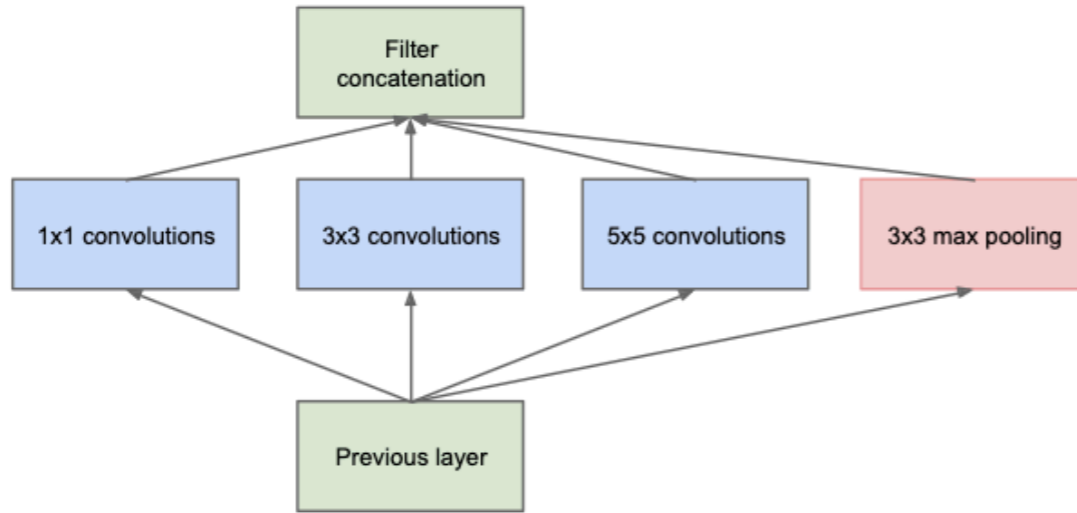


Use two step approach - 1st: convolutional neural network as peak finder



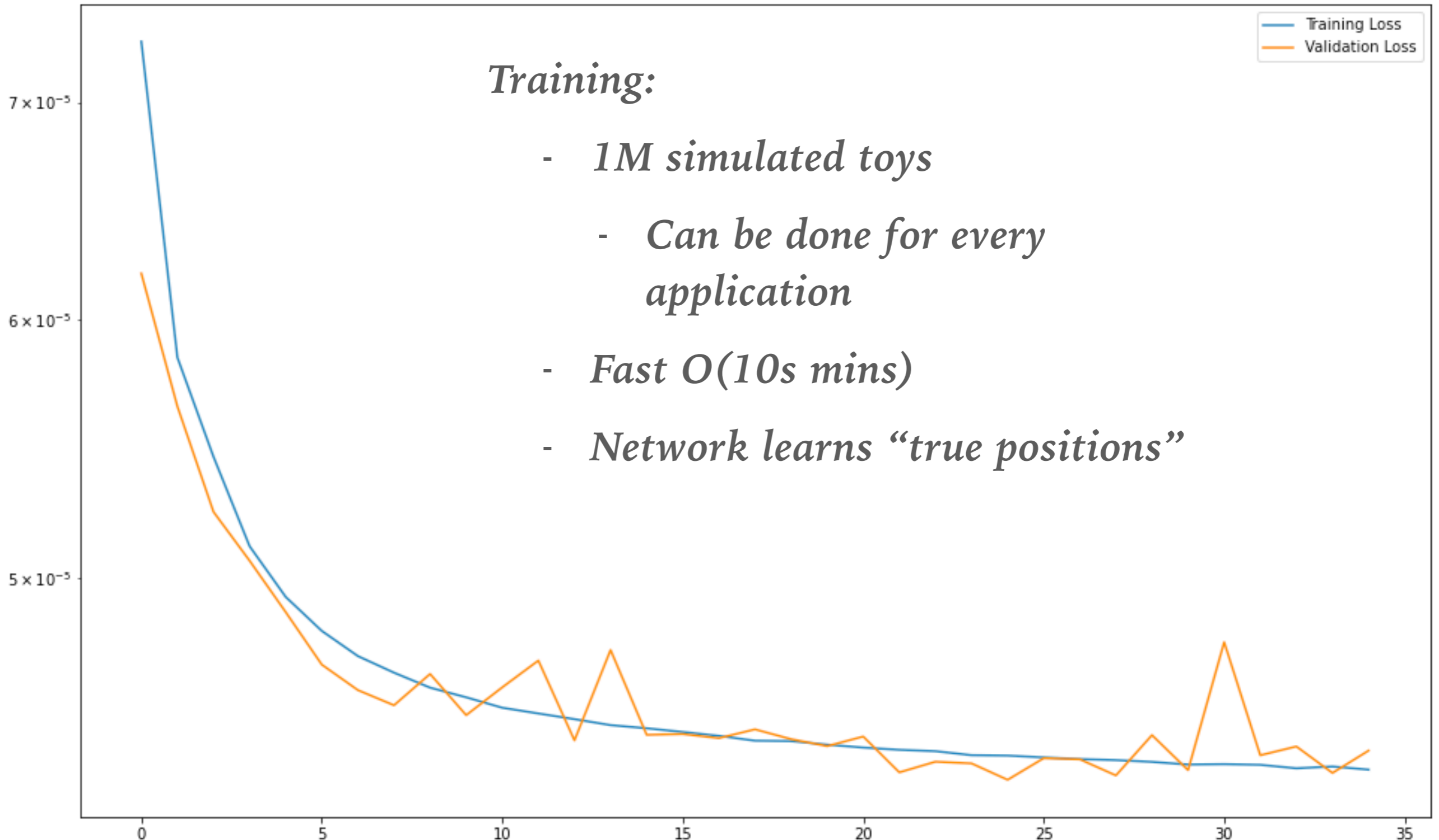
Use two step approach - 1st: convolutional neural network as peak finder

The "Inception" architecture

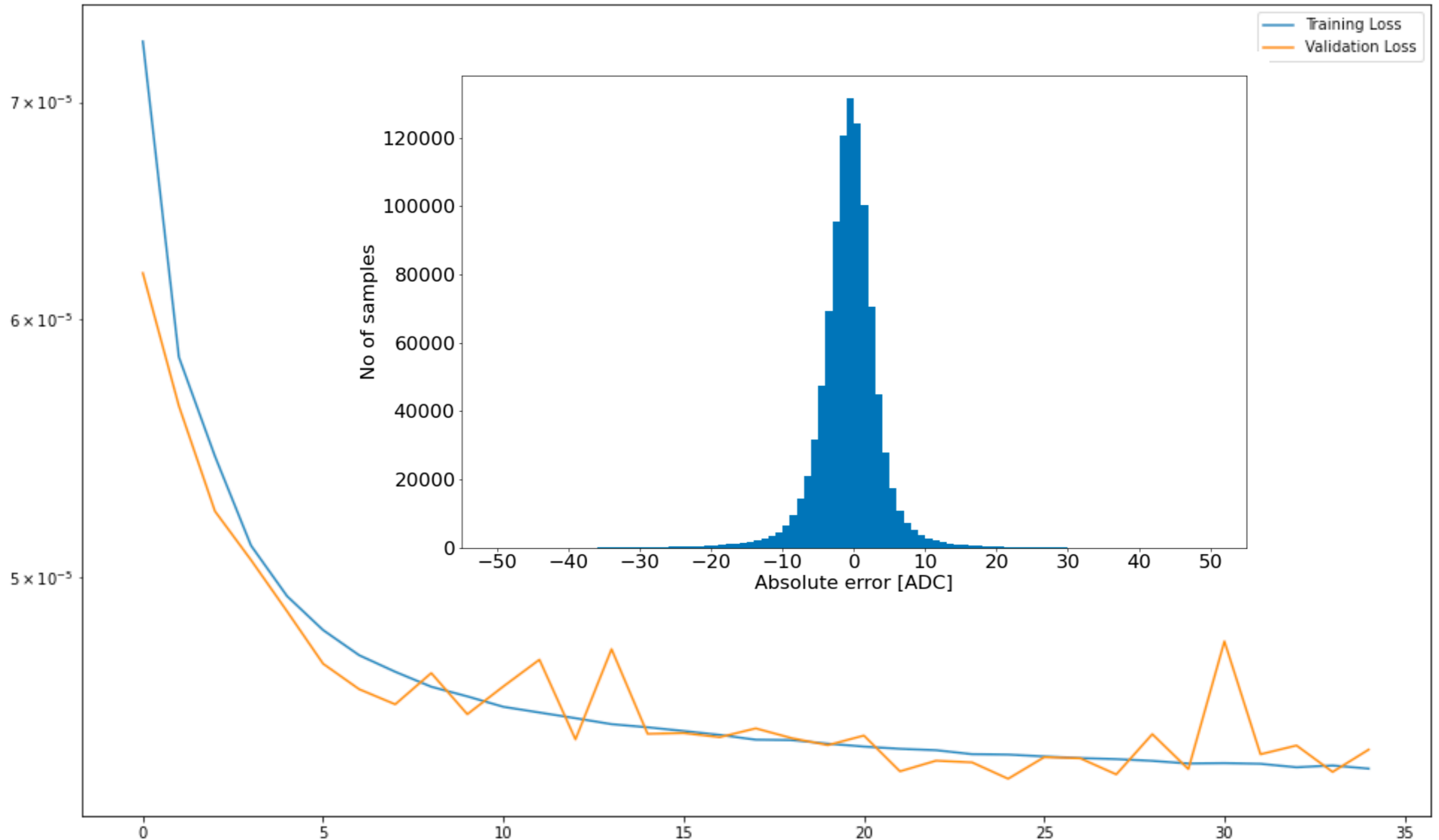


GoogLeNet network

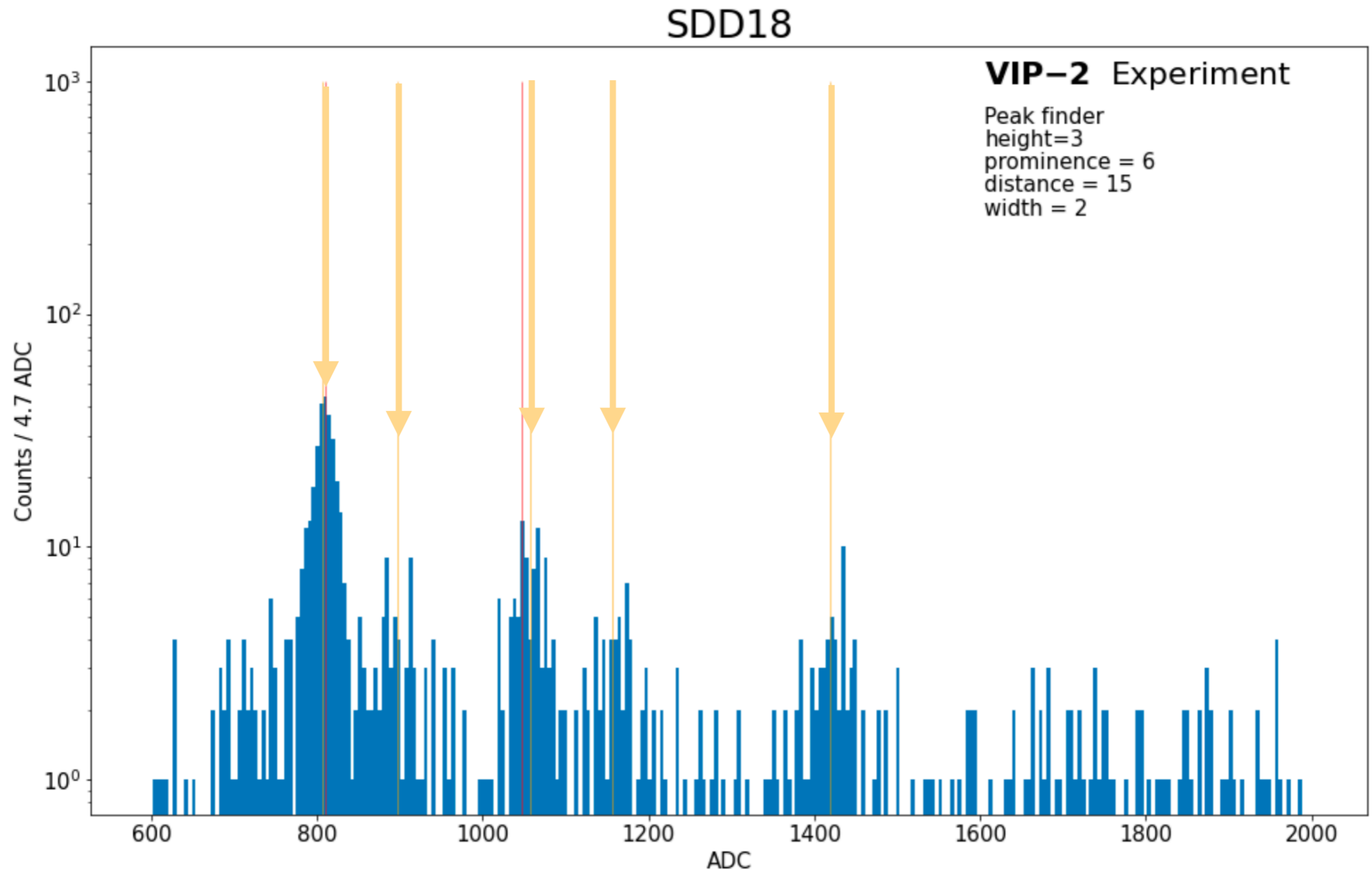
Use two step approach - 1st: convolutional neural network as peak finder

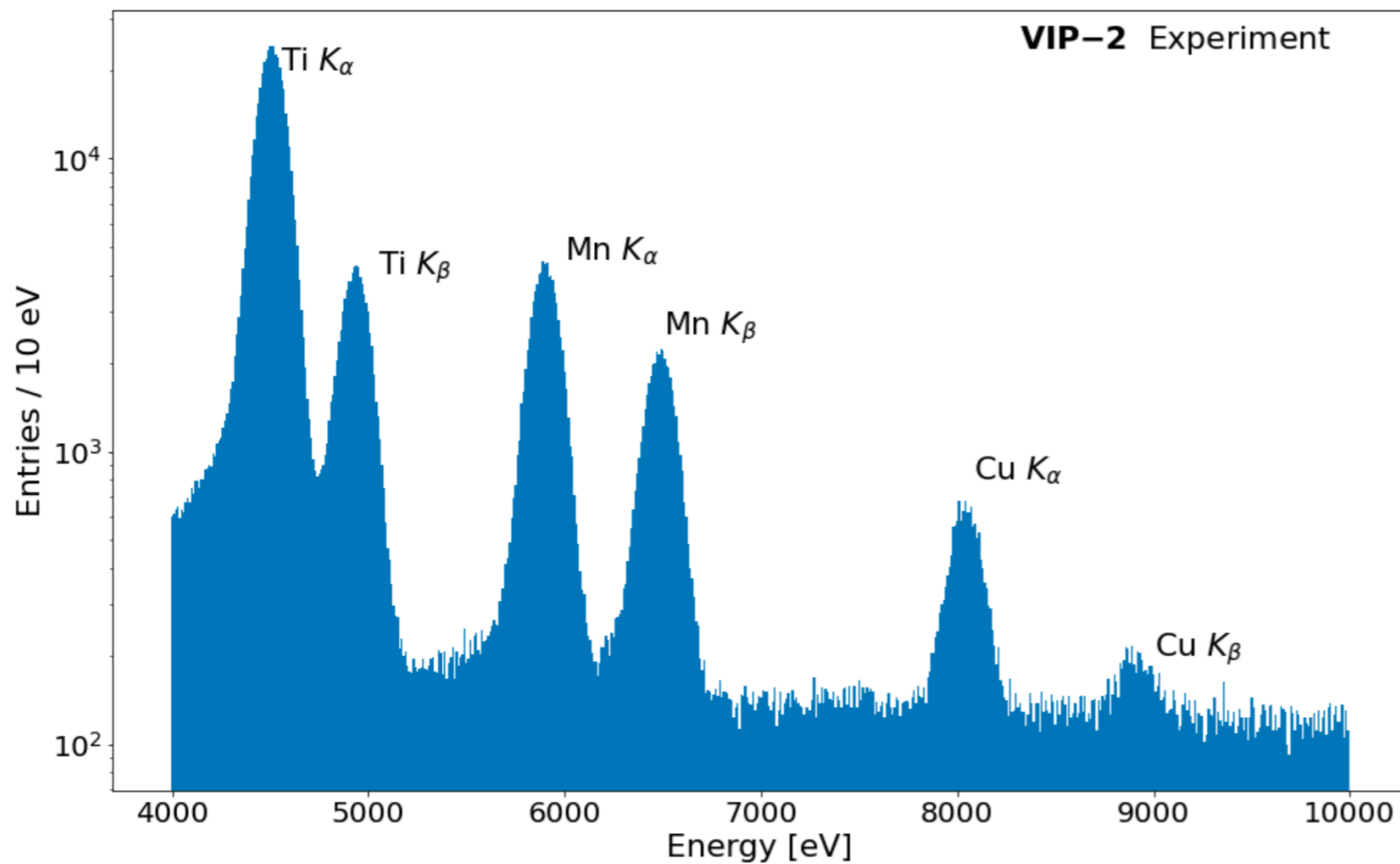


Use two step approach - 1st: convolutional neural network as peak finder



Use two step approach - 1st: convolutional neural network as peak finder





*Calibrated spectrum of
4 SDD arrays.*

Not easy to calibrate because:

- *Copper line at orders of magnitude smaller than Ti and Mn*
- *Tiny distortions of FEE*

Calibration can be done in big or small batches

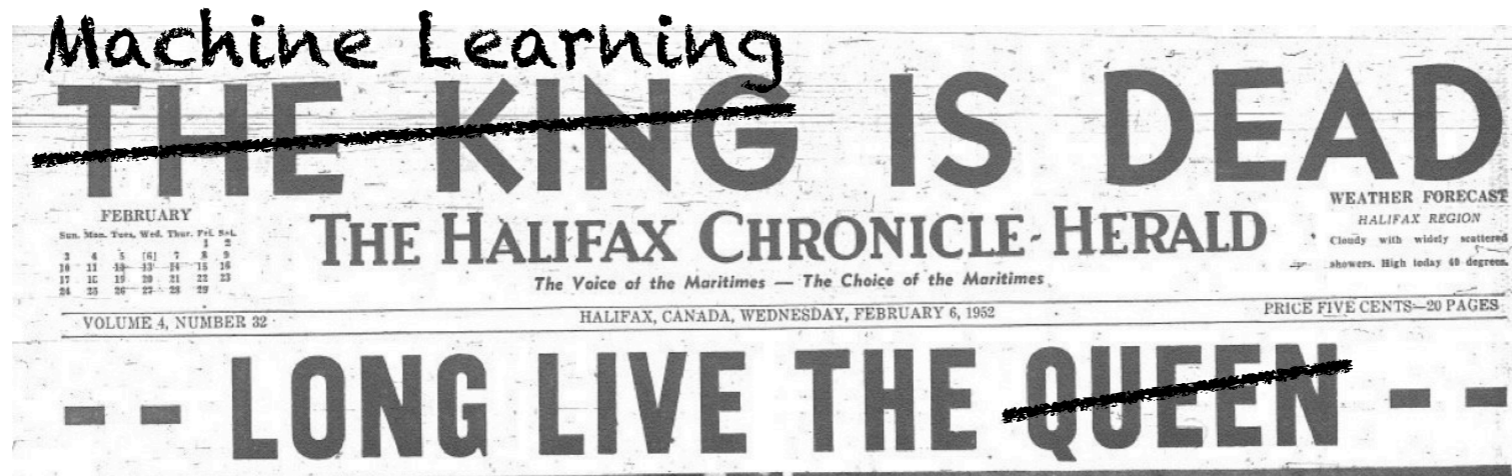
Big batches

*Can determine better the
Copper position but
cannot capture fluctuations*

Small batches

*Can capture fluctuations but
cannot determine the Copper
position well*

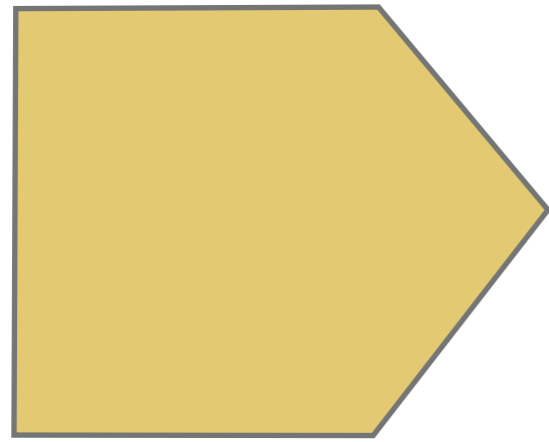
Use two step approach - 2nd: to the limit with differentiable programming



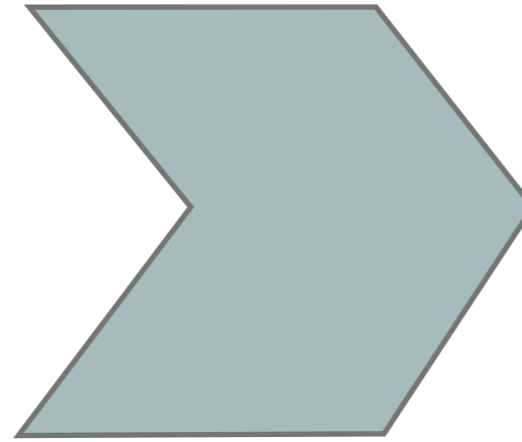
Differentiable Programming



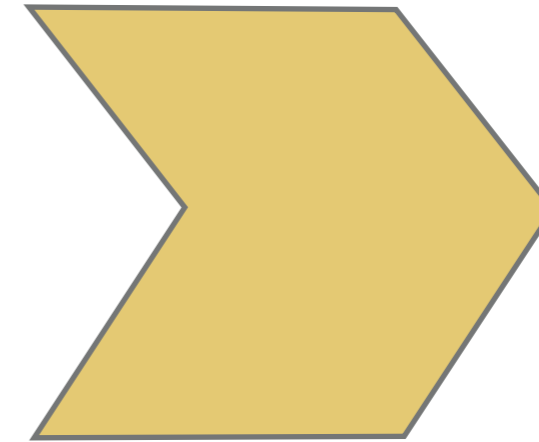
Use two step approach - 2nd: to the limit with differentiable programming



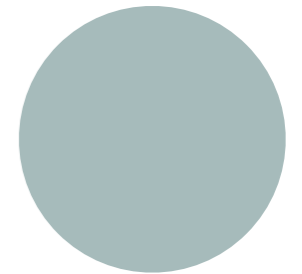
Starting
Conditions,
with e.g. many parameters



Complex manipulation.
Differential equations
Simulations

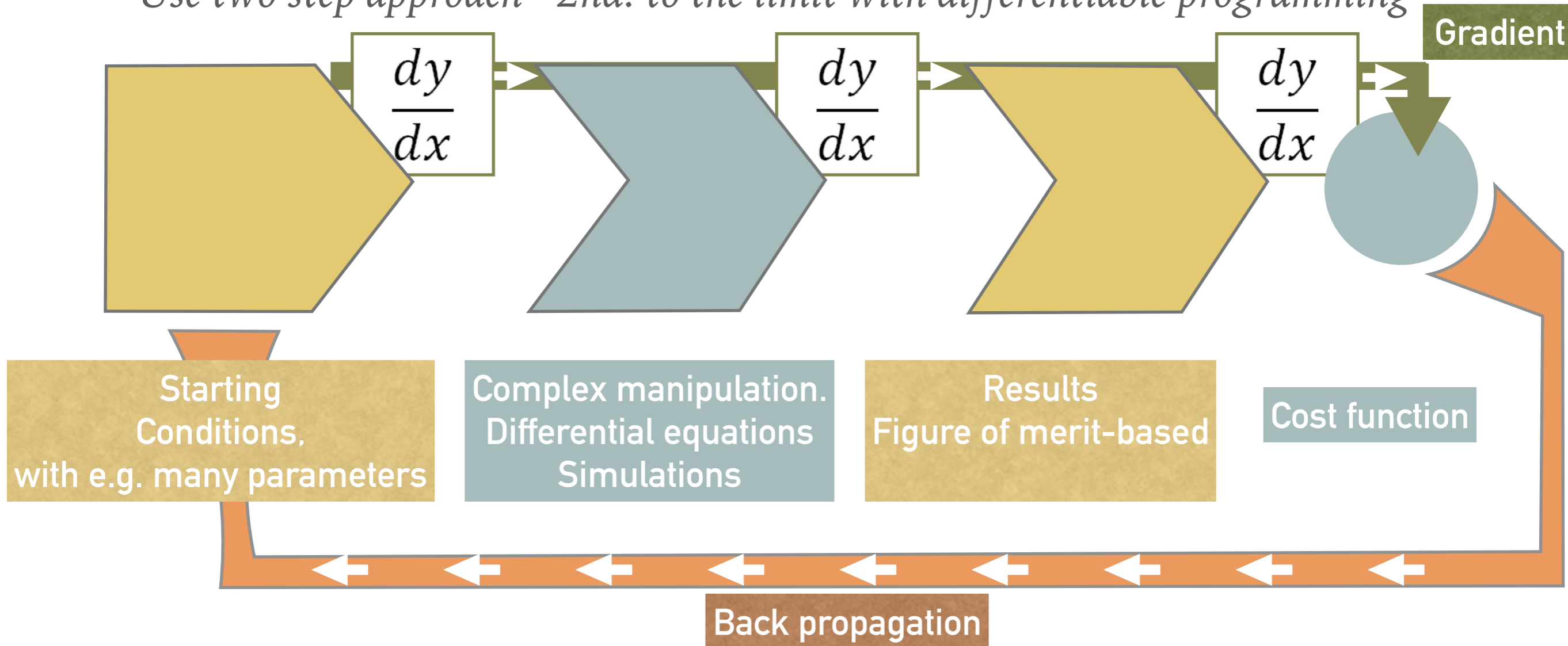


Results
Figure of merit-based

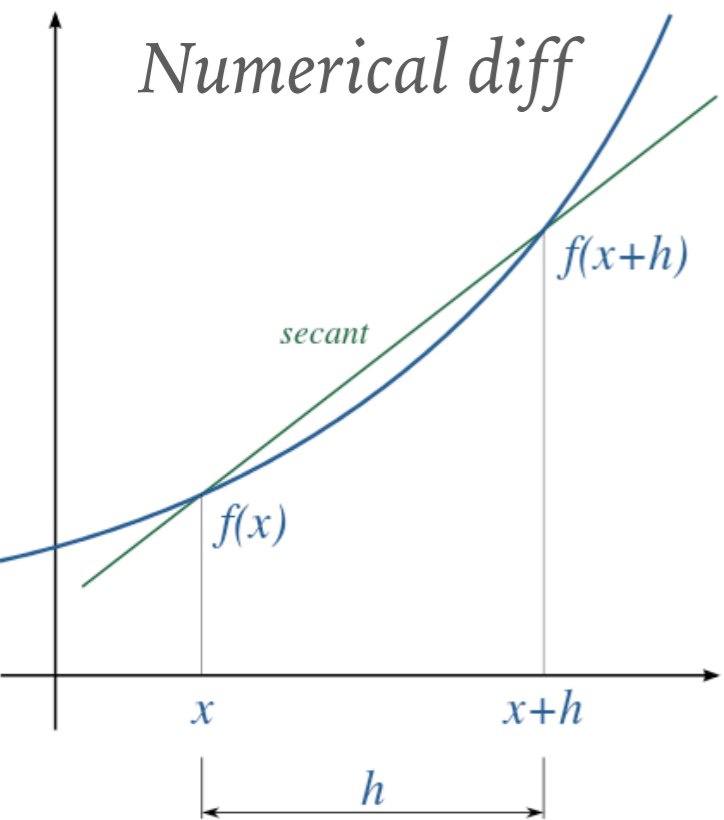
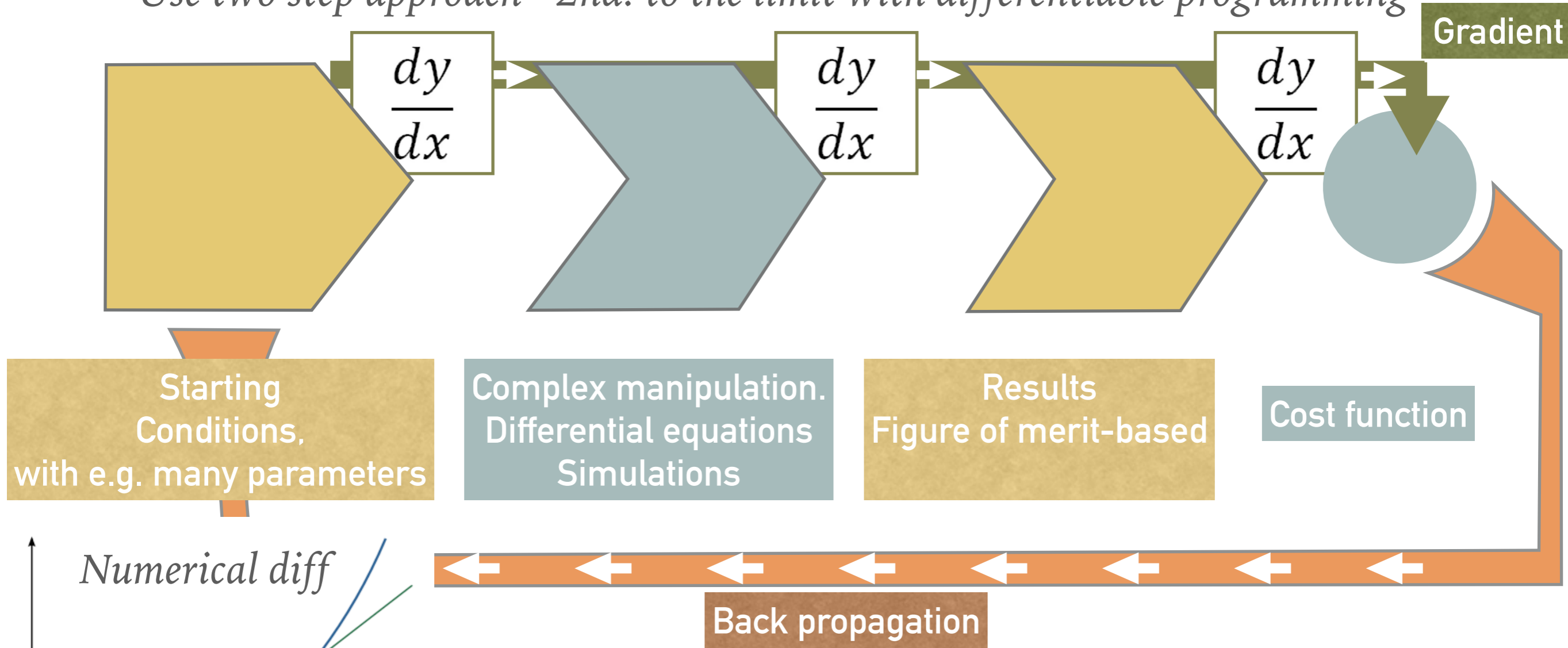


Cost function

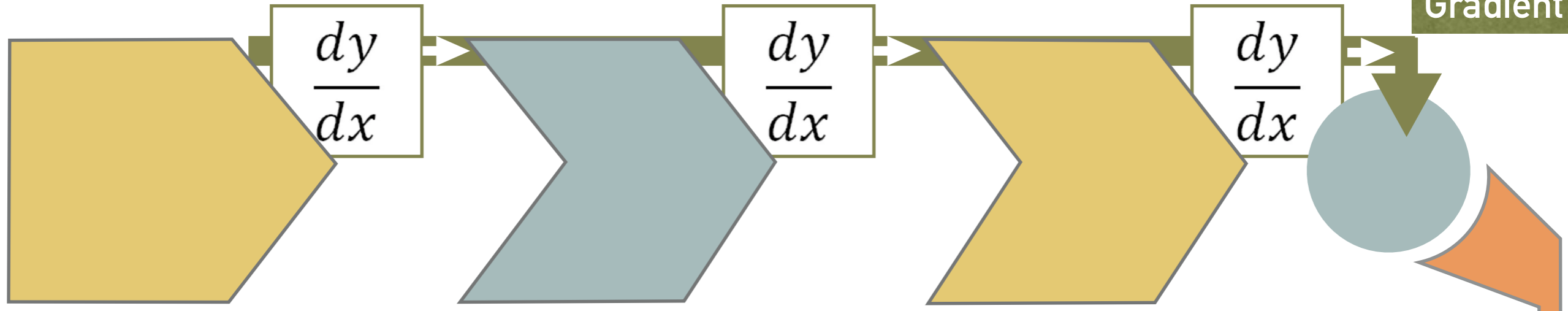
Use two step approach - 2nd: to the limit with differentiable programming



Use two step approach - 2nd: to the limit with differentiable programming



Use two step approach - 2nd: to the limit with differentiable programming

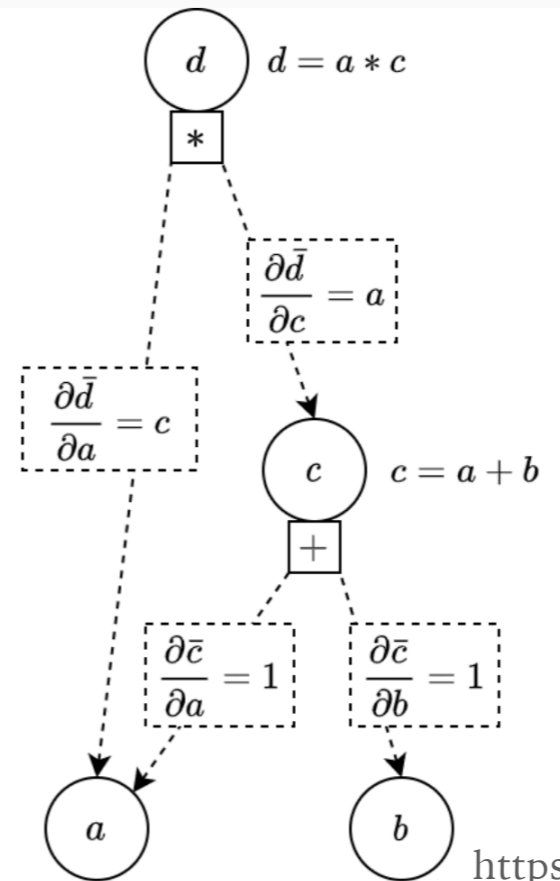
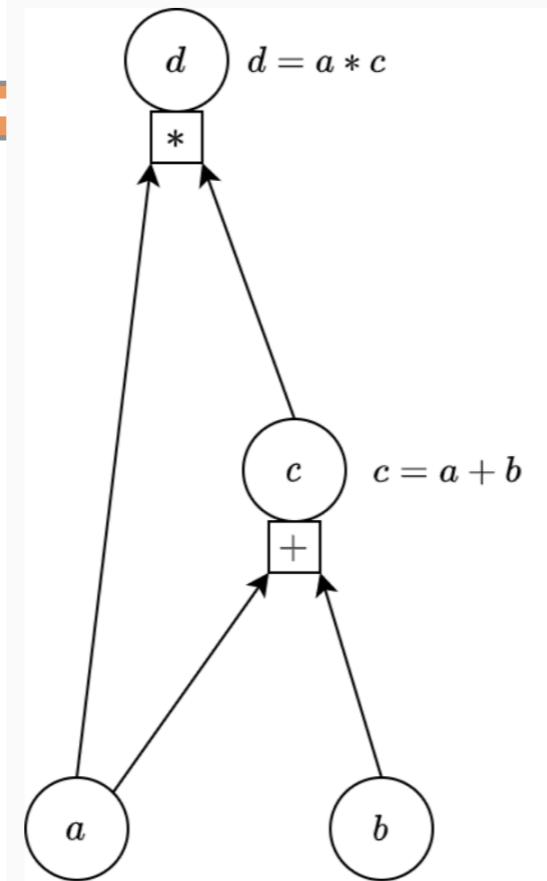
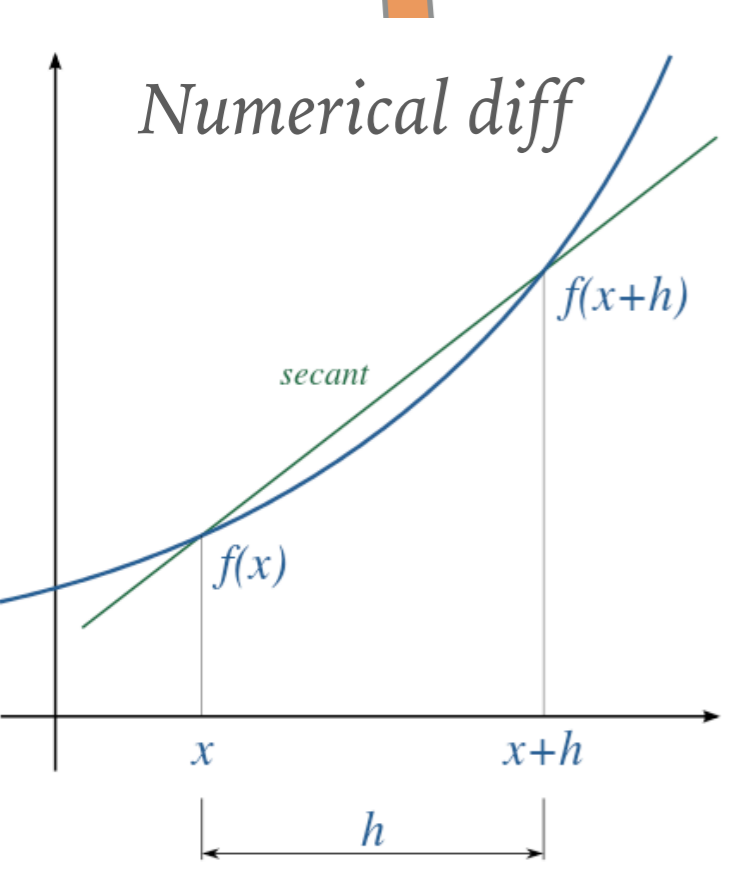


Starting Conditions, with e.g. many parameters

Complex manipulation. Differential equations Simulations

Results Figure of merit-based

Cost function



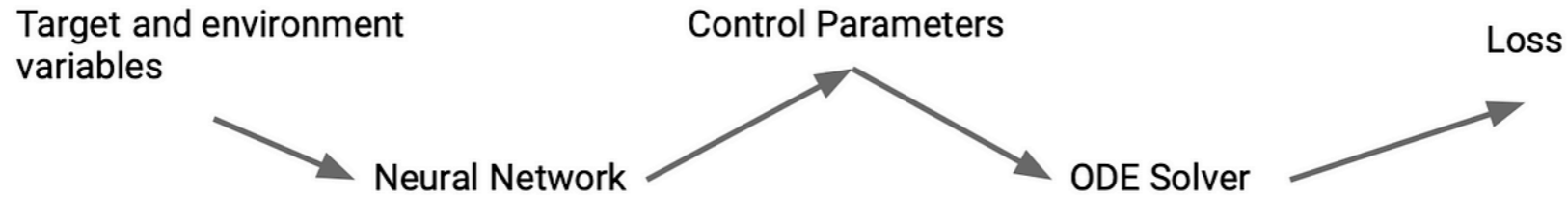
$$\frac{\partial d}{\partial a} = \frac{\partial \bar{d}}{\partial a} + \frac{\partial \bar{d}}{\partial c} * \frac{\partial \bar{c}}{\partial a}$$

$$\frac{\partial d}{\partial a} = c + a * 1$$

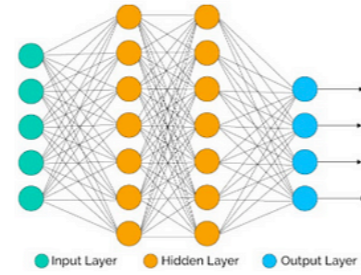
$$\frac{\partial d}{\partial a} = a + b + a$$

$$\frac{\partial d}{\partial a} = 2a + b$$

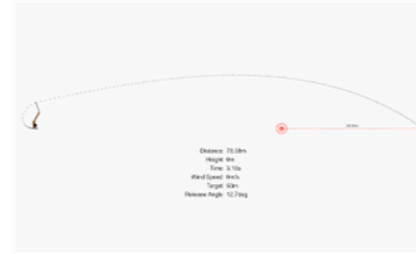
Use two step approach - 2nd: to the limit with differentiable programming



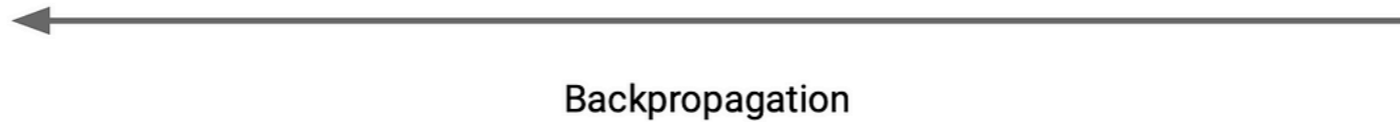
wind = -10m/s
target = 50m



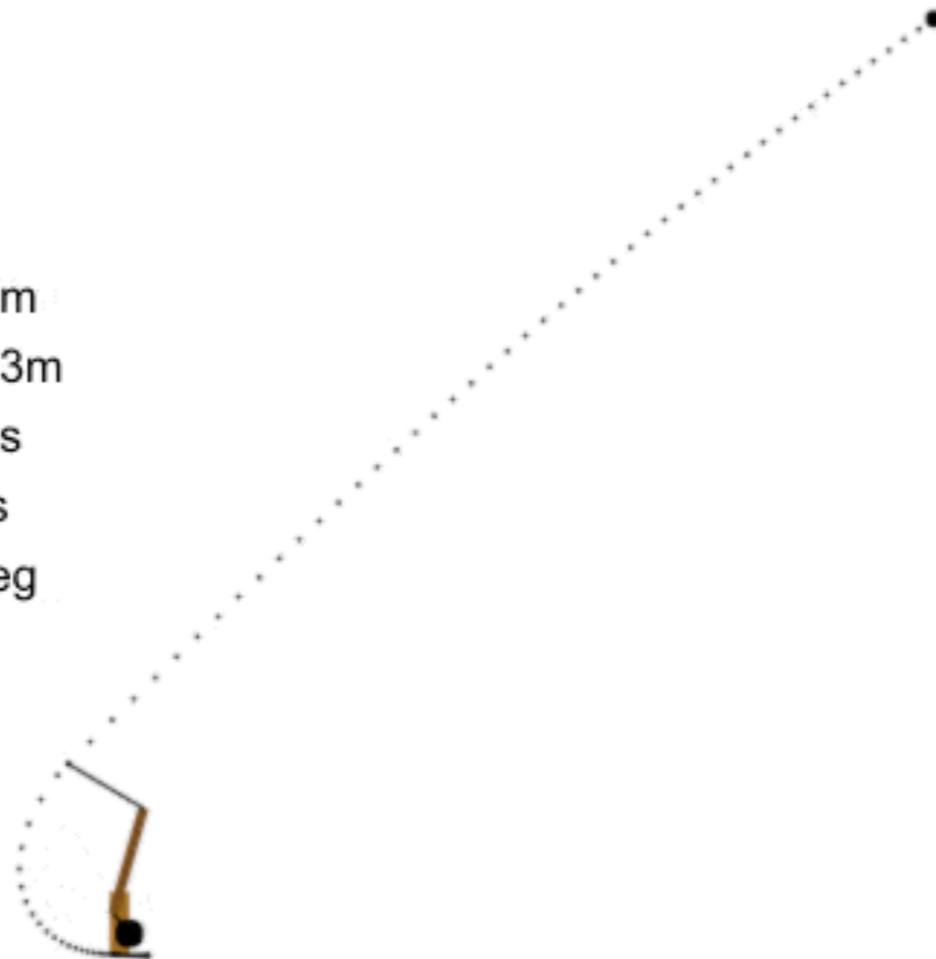
angle = 25°
weight = 200kg



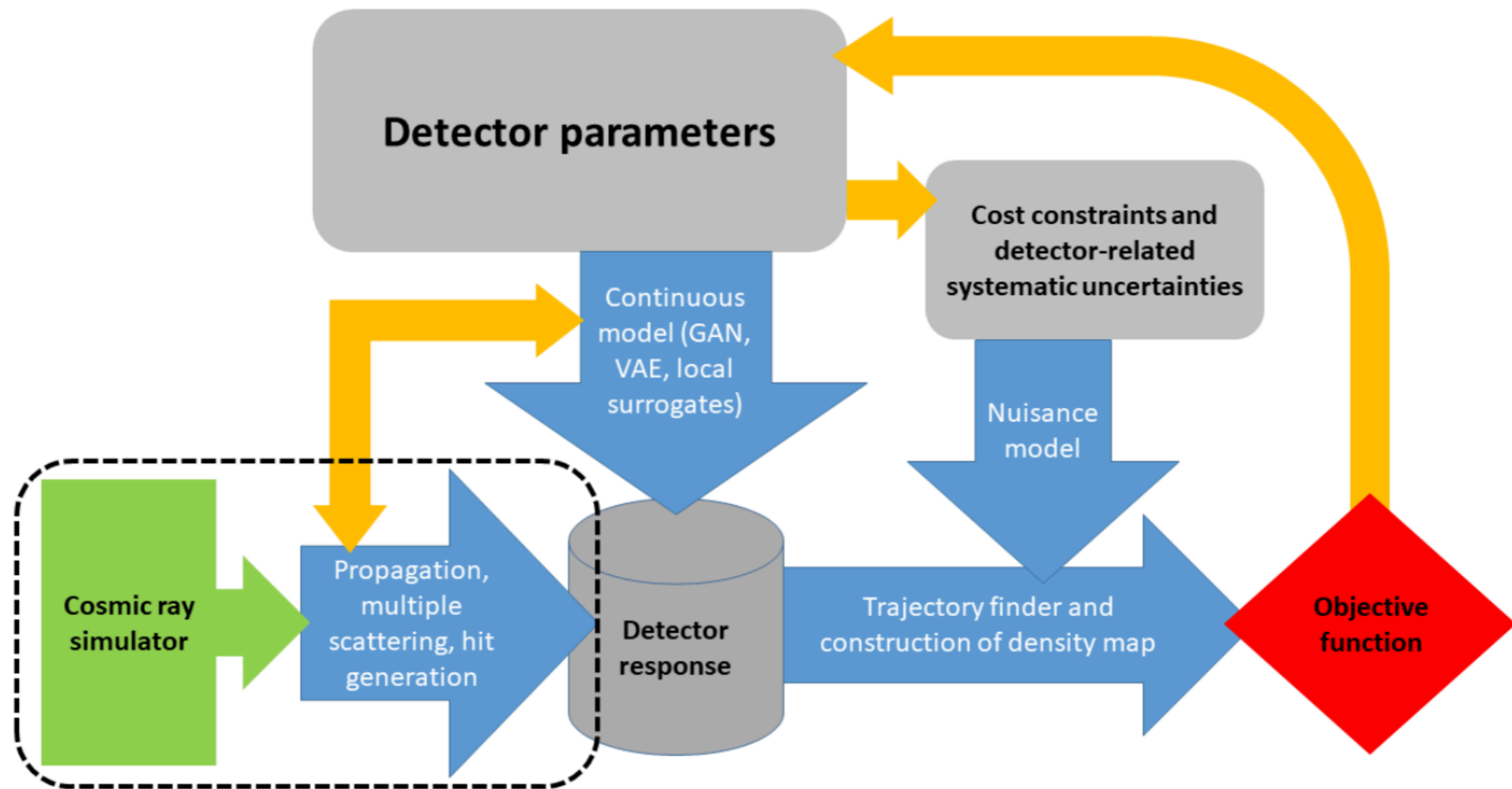
$(target_distance - actual_distance)^2$



Distance 19.4m
 Height 22.33m
 Time 1.46s
 Wind Speed 1m/s
 Release Angle 45deg



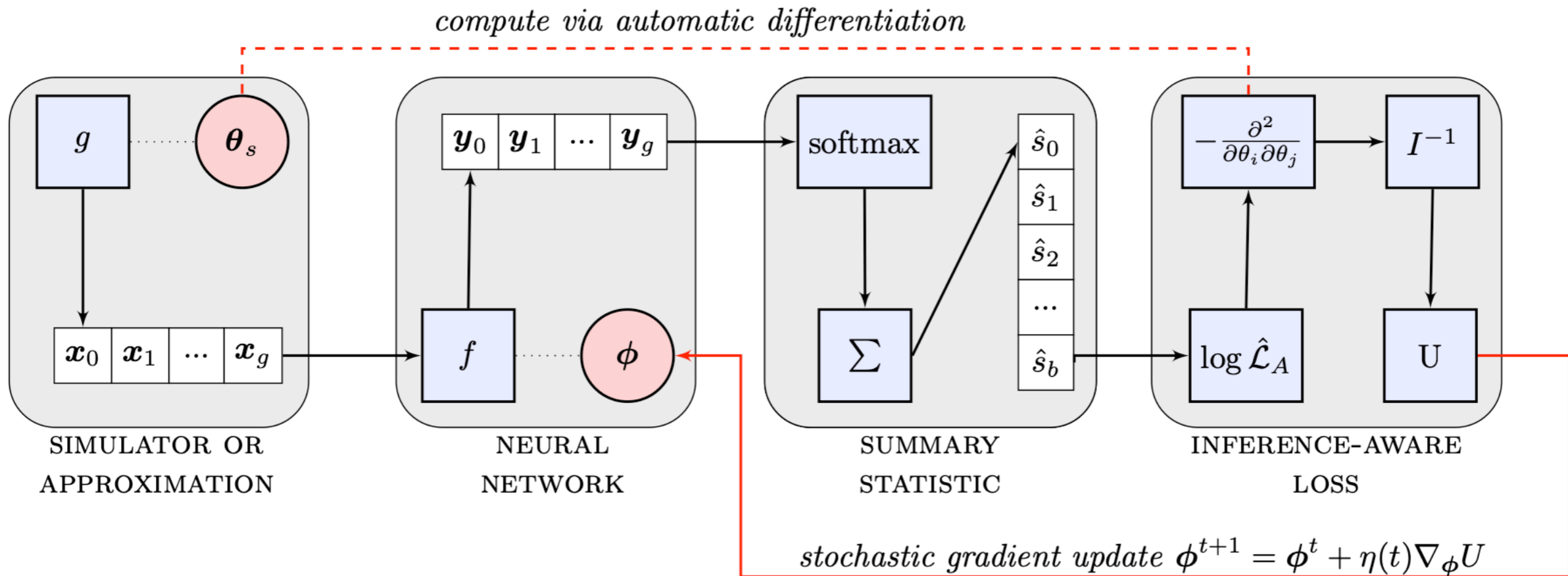
Use two step approach - 2nd: to the limit with differentiable programming



Optimization of detector design and operation



Use two step approach - 2nd: to the limit with differentiable programming

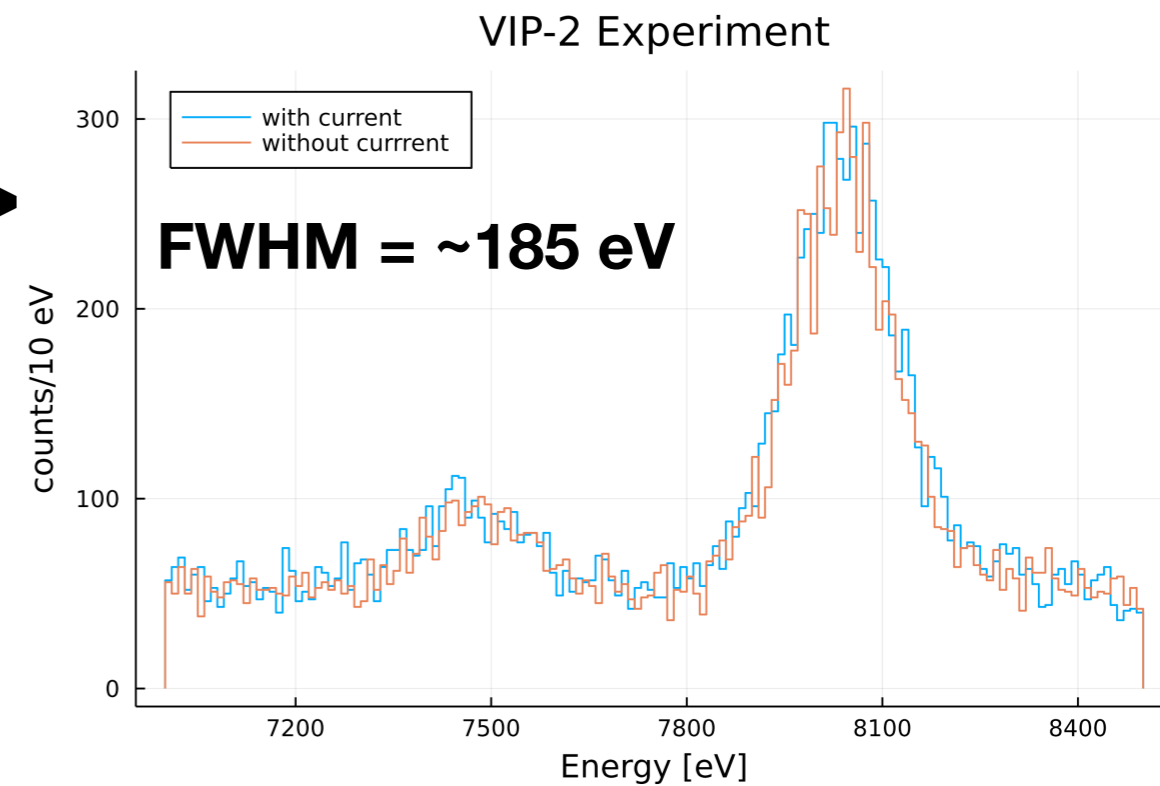
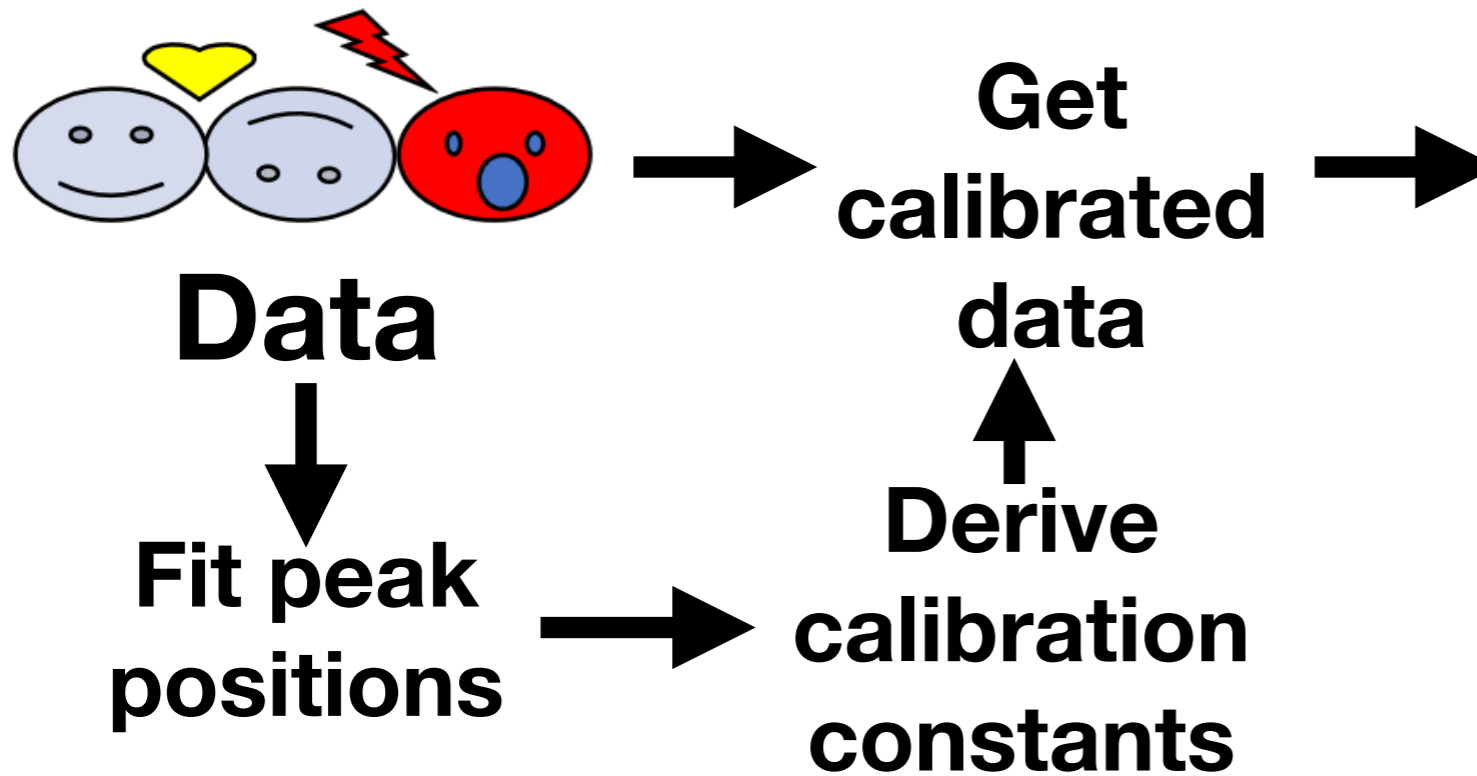


Sketch of the INFERNO algorithm. Batches from a simulator are passed through a neural network and a differentiable summary statistic is constructed that allows to calculate the variance of the POI. The parameters of the network are then updated by stochastic gradient descent.

Use two step approach - 2nd: to the limit with differentiable programming

Idea: use automatic differentiation to compute gradients of functions

Use gradients to find global optima

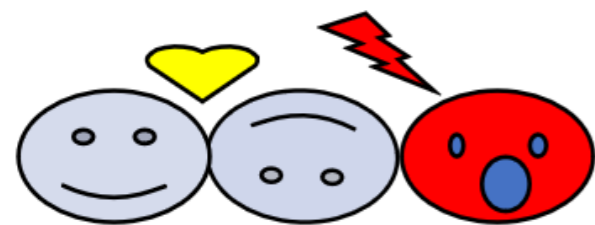


Our Calibration Flow

Use two step approach - 2nd: to the limit with differentiable programming

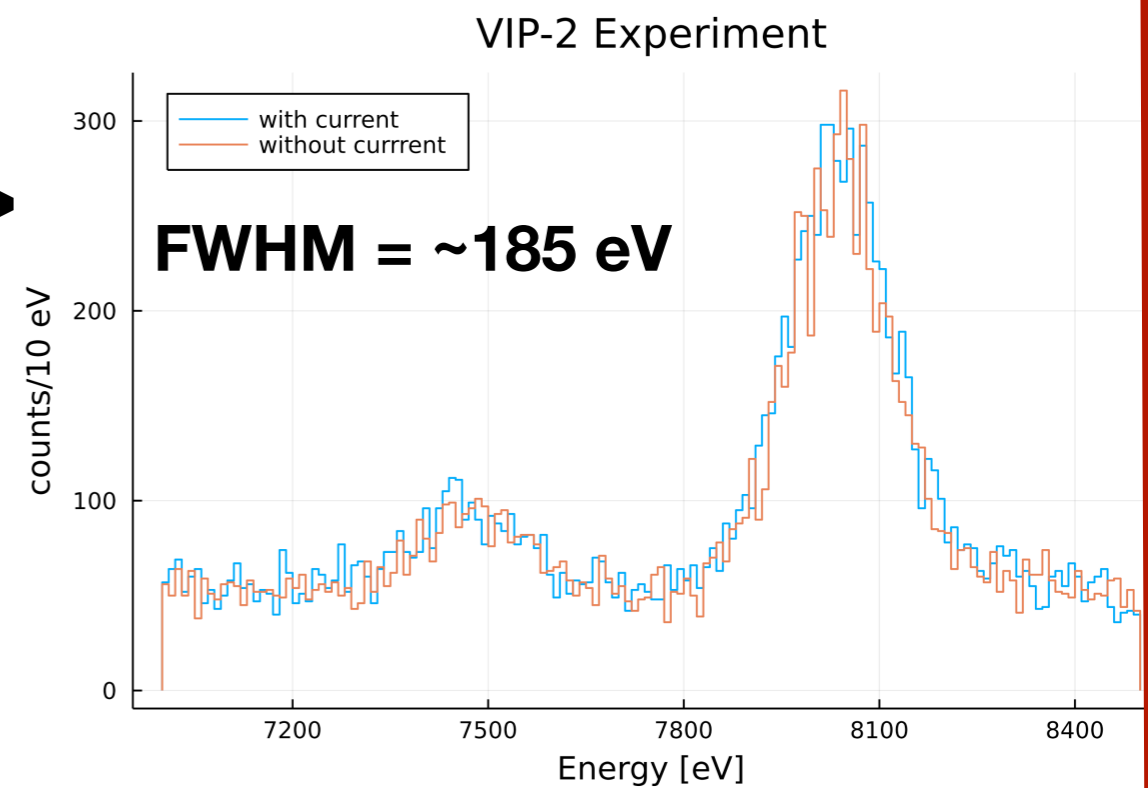
Idea: use automatic differentiation to compute gradients of functions

Use gradients to find global optima



Data
↓
Fit peak positions

Get calibrated data
↑
Derive calibration constants



Following the gradient, change the constants to enhance FWHM

Use two step approach - 2nd: to the limit with differentiable programming

Loss function for the gradient: log likelihood

Derive calibrated

dataset

**Take new
calibration
constants**

$$\mathbf{E} = \mathcal{C}(\mathbf{ADC} | \mathbf{P})$$

$$\mathbf{P} = \begin{bmatrix} p_{0,1} & p_{1,1} & p_{2,1} \\ \vdots & \vdots & \vdots \\ p_{0,i} & p_{1,i} & p_{2,i} \\ \vdots & \vdots & \vdots \\ p_{0,N} & p_{1,N} & p_{2,N} \end{bmatrix}$$

$$\mathcal{L}(\mathbf{E})$$

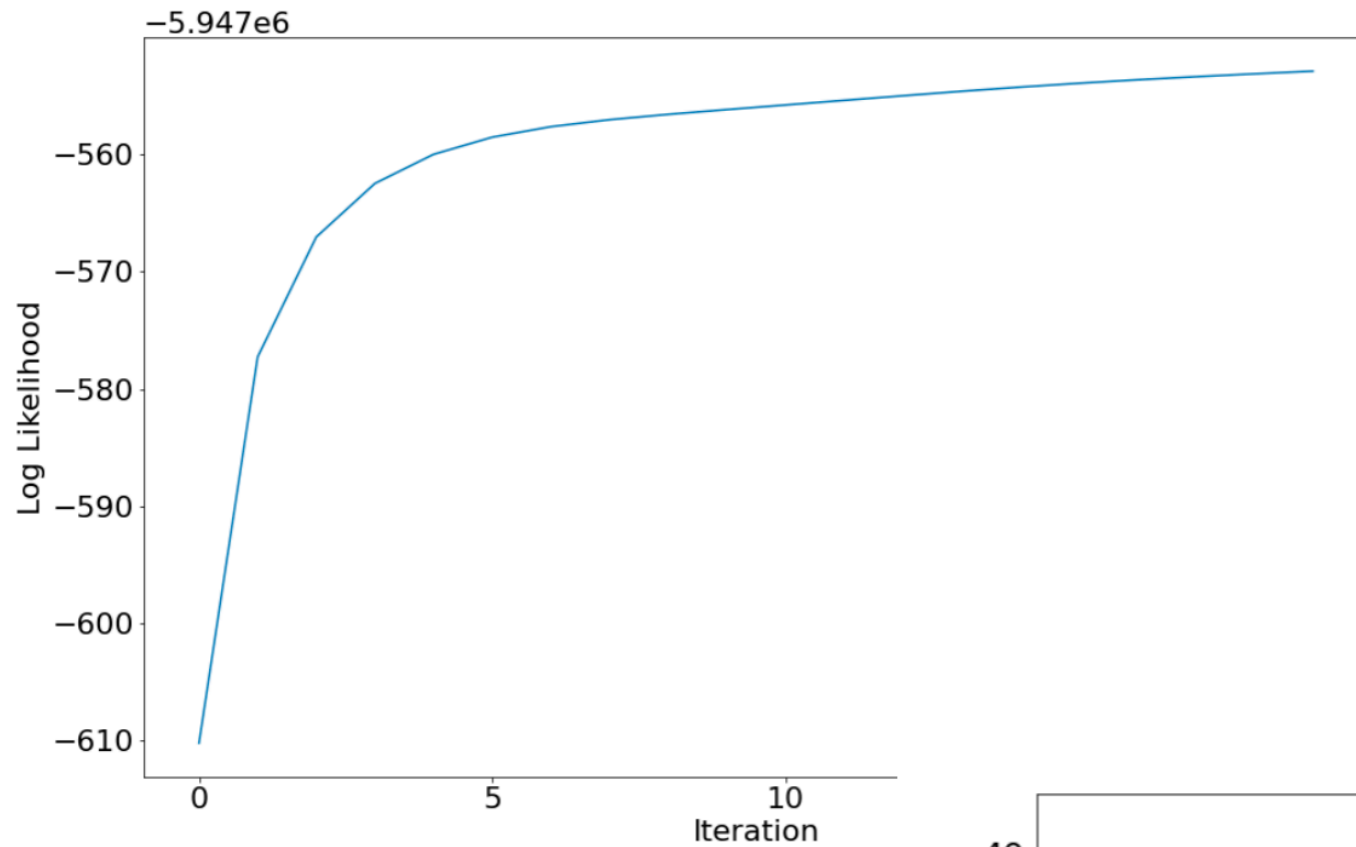
Compute likelihood

**Update \mathbf{P} to minimize
the likelihood**

$$\mathcal{L}(\mathbf{ADC}, \mathbf{P}) = \prod_i \prod_{j \in i} (\text{Gauss}(\mathcal{C}(\text{ADC}_{i,j}, P_i) - \mu_{Cu_{K_{\alpha 1}}}, \sigma))$$

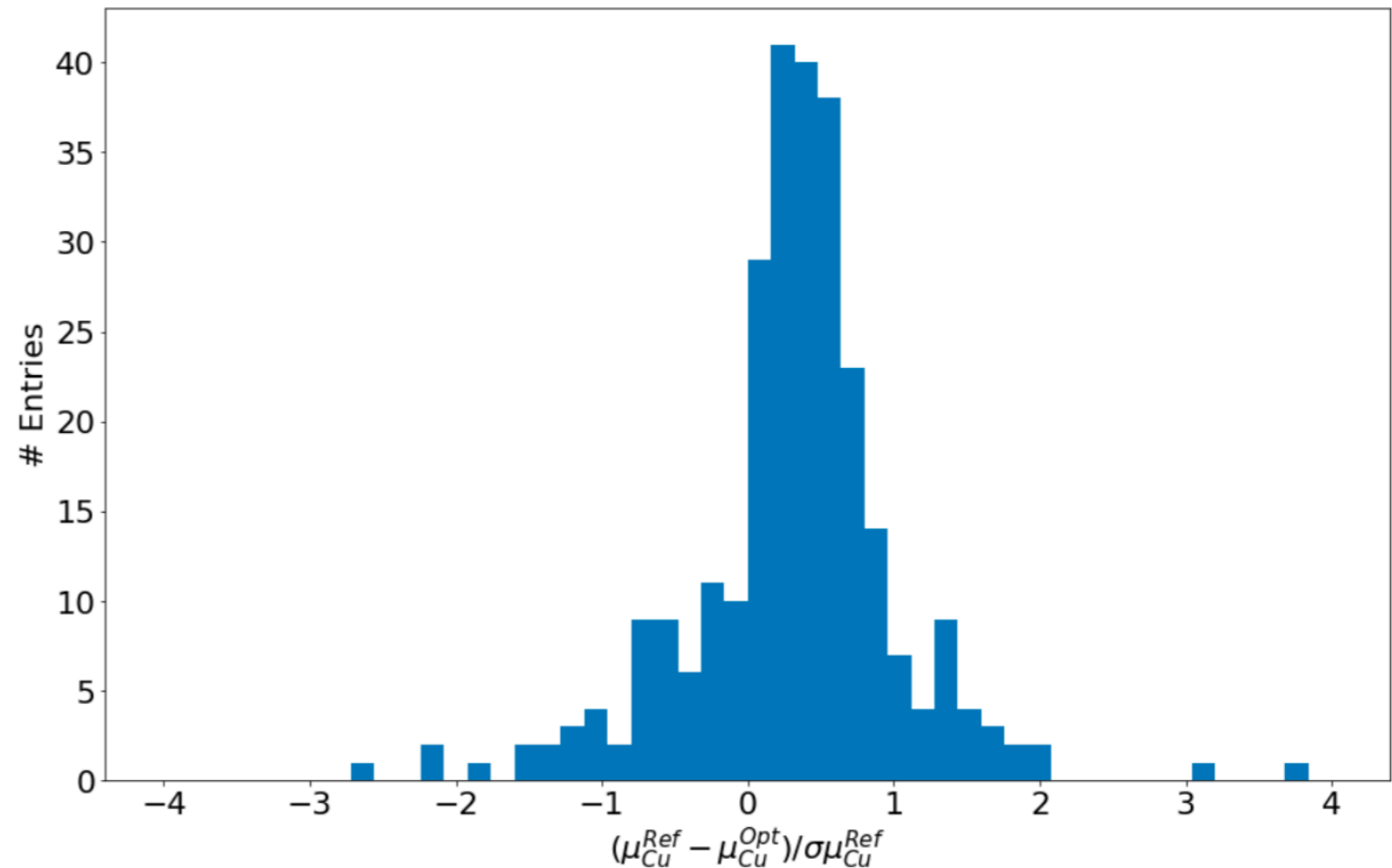
$$+ \text{Gauss}(\mathcal{C}(\text{ADC}_{i,j}, P_i) - \mu_{Cu_{K_{\alpha 2}}}, \sigma))$$

Use two step approach - 2nd: to the limit with differentiable programming



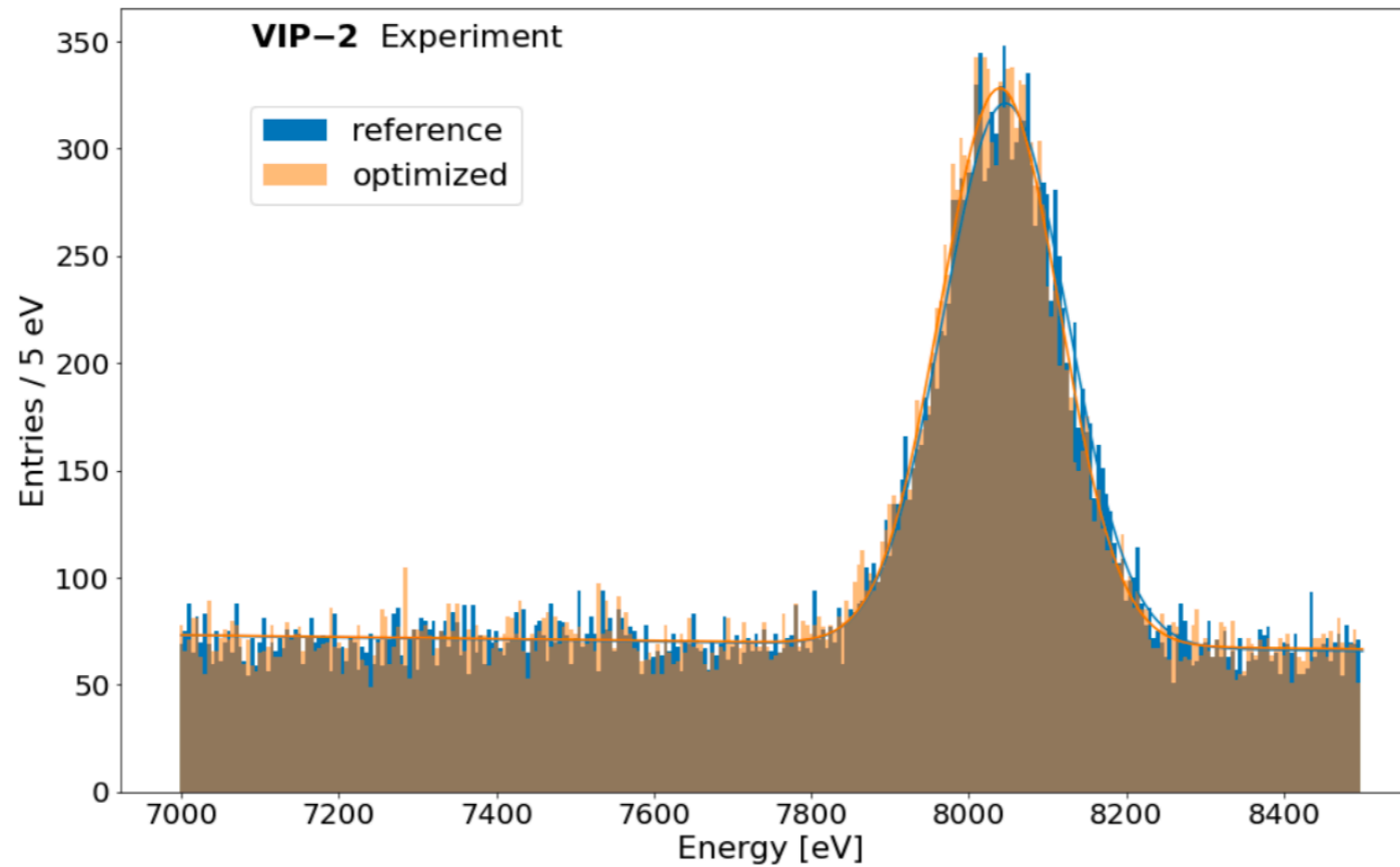
*Gradient descended
into plateau*

*Most of the calibration
parameters do not change
much*



Results

<https://arxiv.org/abs/2305.17153>
Submitted to Meas. Sci. Tech.



	Position [eV]	FWHM [eV]	χ^2/ndf
Reference	8050 ± 1	185 ± 2	1.64
Optimized	8048 ± 1	176 ± 2	1.25

$$f(x, A, \mu, \sigma) = A \times \frac{51}{100} \times \text{Gauss}(x - \mu - 20, \sigma) + T_2(x) + A \times \text{Gauss}(x - \mu, \sigma) + T_1(x) + m \times x + C$$

$$T_i(x) = \frac{A_i}{2\beta\sigma} \times e^{\frac{x-\nu}{\beta\sigma} \frac{1}{2\beta^2}} \times \text{erfc} \left(\frac{x-\nu}{\sqrt{2}\pi} + \frac{1}{\sqrt{2}\beta} \right)$$

Conclusions

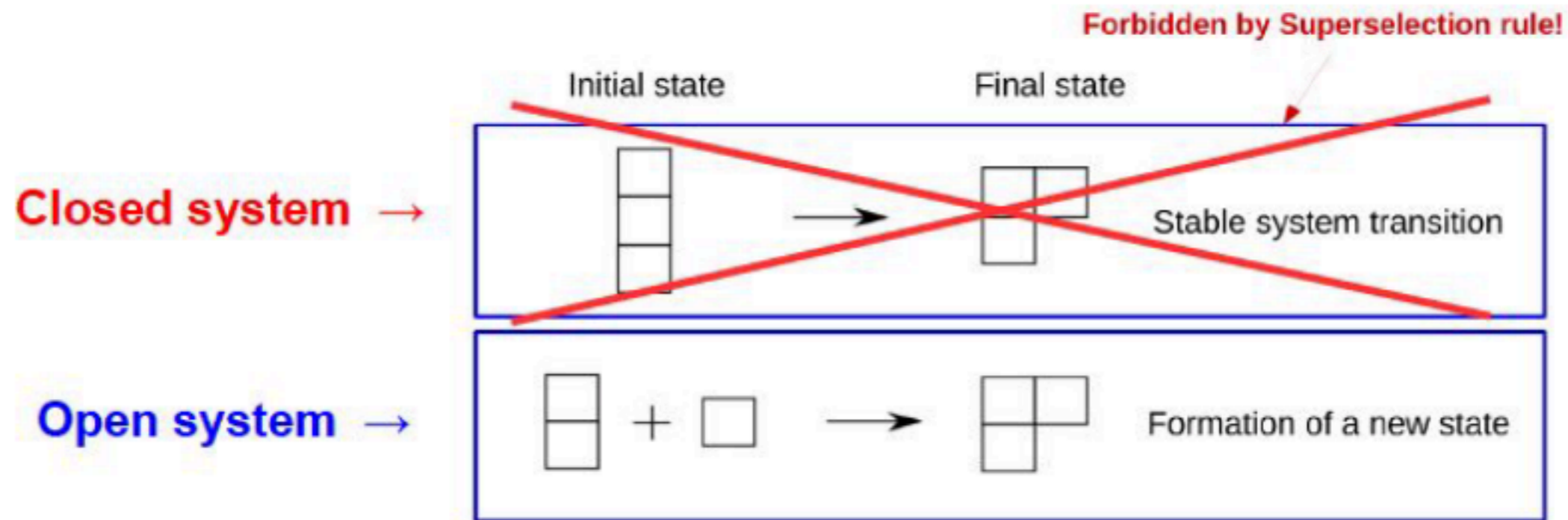
- *VIP-2 experiment at LNGS uses X-ray spectroscopy as tool for precision tests QM and Standard Model*
- *Showed a ML and DP way to increase the detector resolution*
 - *Below 180 eV FWHM at Cu for the first time in VIP-2*
 - *The method can correct for miscalibration*
 - *Closer compatibility with the model*
 - *Higher discovery significance of forbidden transitions*

Thank you for your attention!
Questions?

Messiah-Greenberg super-selection rule:

Superposition of states with different symmetry are not allowed \rightarrow

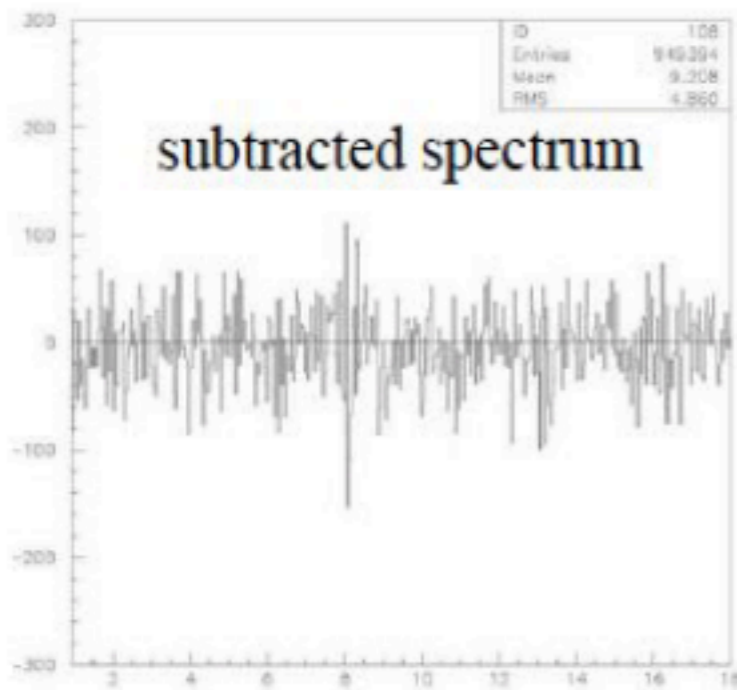
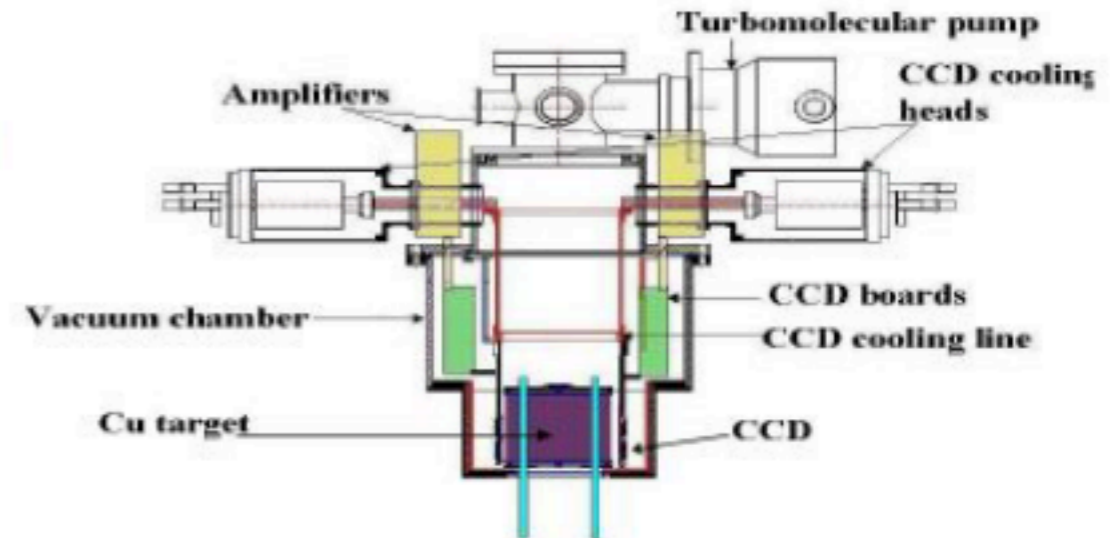
Transition probability between two symmetry states is ZERO



VIP-2 Experiment: best limits on PEP violation of an elementary particle respecting the Messiah-Greenberg super-selection rule

From VIP to VIP-2

- copper ultrapure cylindrical foil
- surrounded by 16 Charge Coupled Devices (CCD) res. at 8 keV 320 eV (FWHM)
- inside a vacuum chamber: CCDs cooled to 168K by a cryogenic system
- amplifiers + read out ADC boards.



$$\beta^2/2 \leq 4.7 \times 10^{-29}$$

improved the limit obtained by Ramberg & Snow by a factor ~ 400

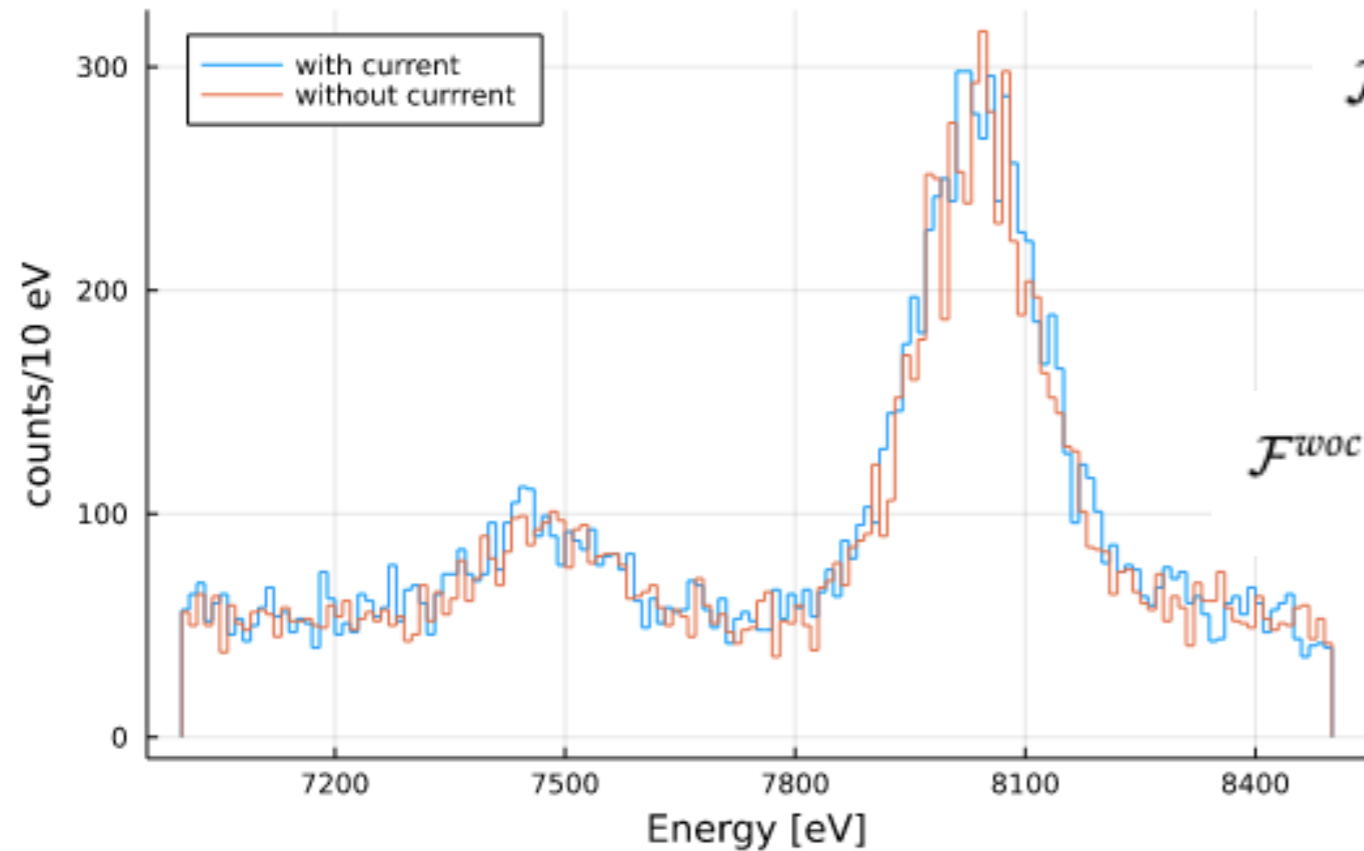
(Foundation of Physics 41 (2011) 282+ other papers)

GOAL OF VIP-2 : improve the VIP result of 2 orders of magnitude

Symmetry **2022**, 14(5), 893;
<https://doi.org/10.3390/sym14050893>

Six months of data taking

VIP-2 Experiment



Description spectrum with current

$$\mathcal{F}^{wc}(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S}) = y_1 \times Ni(\theta_1, \theta_2) + y_2 \times Cu(\theta_3, \theta_4) + y_3 \times pol_1(\theta_5) + \mathcal{S} \times PEPV(\theta_4)$$

Description spectrum without current

$$\mathcal{F}^{woc}(\boldsymbol{\theta}, \mathbf{y}) = y_1 \times Ni(\theta_1, \theta_2) + y_2 \times Cu(\theta_3, \theta_4) + y_3 \times pol_1(\theta_5)$$

Likelihood

$$\mathcal{L}(\mathcal{D}^{wc}, \mathcal{D}^{woc} | \boldsymbol{\theta}, \mathbf{y}, \mathcal{S}) = \text{Poiss}(\mathcal{D}^{wc} | \mathcal{F}^{wc}(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S})) \times \text{Poiss}(\mathcal{D}^{woc} | \mathcal{F}^{woc}(\boldsymbol{\theta}, \mathbf{y} \times \mathcal{R}))$$

Bayesian

$$p(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S} | \mathcal{D}^{wc}, \mathcal{D}^{woc}) =$$

$$= \frac{\mathcal{L}(\mathcal{D}^{wc}, \mathcal{D}^{woc} | \boldsymbol{\theta}, \mathbf{y}, \mathcal{S}) p(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S})}{\int d\boldsymbol{\theta} d\mathbf{y} \mathcal{L}(\mathcal{D}^{wc}, \mathcal{D}^{woc} | \boldsymbol{\theta}, \mathbf{y}, \mathcal{S}) p(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S})}$$

Frequentist

$$t_{\mathcal{S}} = -2 \ln \Lambda(\mathcal{S}) = -2 \ln \frac{\mathcal{L}(\hat{\boldsymbol{\theta}}, \hat{\mathbf{y}}, \mathcal{S})}{\mathcal{L}(\hat{\boldsymbol{\theta}}, \hat{\mathbf{y}}, \hat{\mathcal{S}})}$$

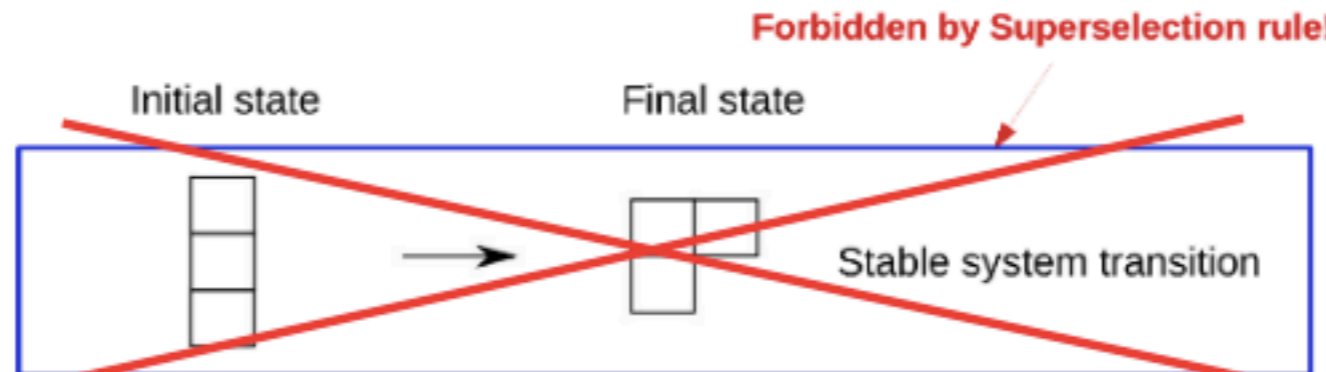
$$p_{\mathcal{S}} = \int_{t_{obs}}^{\infty} f(t_{\mathcal{S}} | \mathcal{S}) dt_{\mathcal{S}}$$

Messiah - Greenberg superselection rule

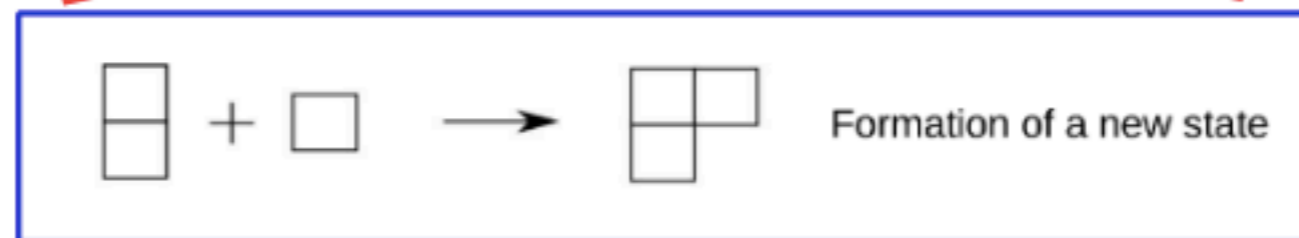
Superpositions of states with different symmetry are not allowed →
 transition probability between two symmetry states is ZERO

Messiah-Greenberg superselection rule :

Closed system →



Open system →



**VIP-open systems sets the best limit on PEP violation
 for an elementary particle
 respecting the M-G superselection rule**

*VIP-2 experiment goal**(Upper limit not using Close Encounters (CE) treatment)**As reference for past experiments*

Experiment	Target	Upper limit of $\beta^2/2$	reference
Ramberg-Snow	Copper	1.7×10^{-26}	[5]
S.R. Elliott et al.	Lead	1.5×10^{-27}	[14]
VIP(2006)	Copper	4.5×10^{-28}	[12]
VIP(2012)	Copper	4.7×10^{-29}	[13]
VIP2(goal)	Copper	$\times 10^{-31}$	[15]

New paradigm for VIP-2

Quantum gravity models can embed PEP violating transitions!

PEP is a consequence of the spin statistics theorem based on: Lorentz/Poincaré and CPT symmetries; locality; unitarity and causality. Deeply related to the very same nature of space and time



most effective theories of QG foresee the non-commutativity of the space-time quantum operators (e.g. k -Poincaré, θ -Poincaré)



non-commutativity induces a deformation of the Lorentz symmetry and of the locality \rightarrow naturally encodes the violation of PEP

S. Majid, Hopf algebras for physics at the Planck scale, *Class. Quantum Grav.* 5 (1988) 1587.

S. Majid and H. Ruegg, Bicrossproduct structure of Kappa Poincare group and noncommutative geometry, *Phys. Lett. B* 334 (1994) 348, hep-th/9405107.

M. Arzano and A. Marciano, *Phys. Rev. D* 76, 125005 (2007) [arXiv:0707.1329].

G. Amelino-Camelia, G. Gubitosi, A. Marciano, P. Martinetti and F. Mercati, *Phys. Lett. B* 671, 298 (2009) [arXiv:0707.1863].

A. Addazi, A. Marcianò *International Journal of Modern Physics A* Vol. 35, No. 32, 2042003 (2020)



PEP violation is suppressed with $(E/\Lambda)^n$, n depends on the specific model, E is the energy of the PEP violating transition, Λ is the scale of the space-time non-commutativity emergence.

How to model PEP violations

- *Ignatiev & Kuzmin model: Fermi oscillator with a third state*

(Ignatiev, A.Y., Kuzmin, V. , Quarks '86: Proceedings of the 229 Seminar, Tbilisi, USSR, 1517 April 1986)

$$\begin{array}{ll}
 a^+|0\rangle = |1\rangle & a|0\rangle = 0 \\
 a^+|1\rangle = \beta|2\rangle & a|1\rangle = |0\rangle \\
 a^+|2\rangle = 0 & a|2\rangle = \beta|1\rangle
 \end{array}$$

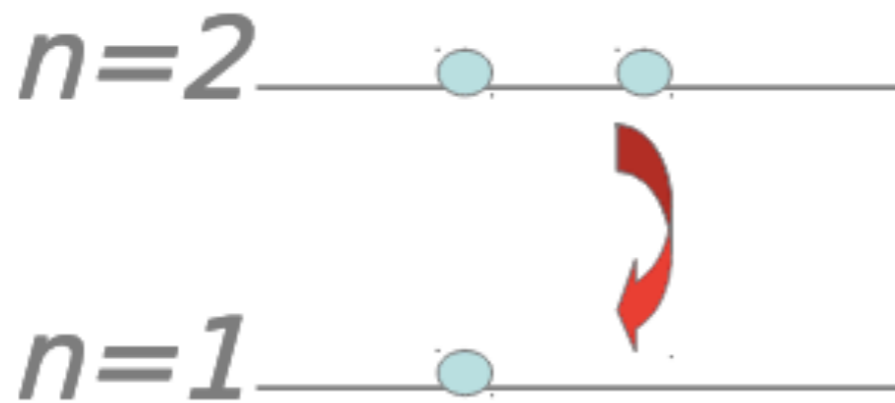
β quantifies the degree of violation in the transition

- *Greenberg & Mohapatra: Local Quantum Field Theory, q parameter deforms anticommutators [Phys. Rev. Lett. 1987,59,2507]:*

$$a_k a^+ l - q a^+ l a_k = \delta_{k,l}$$

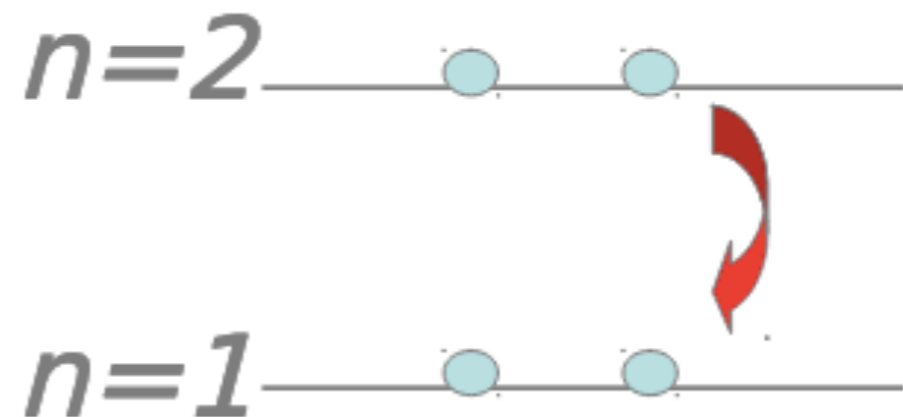
- *Rahal & Campa: global wave function of the electrons not exactly antisymmetric, PEP holds as long as the number of wrongly entangled pairs is small*

Search for anomalous X-ray transitions performed by electrons introduced in a target *trough a DC current (open system)*



Normal $2p \rightarrow 1s$ transition

~ 8.05 keV in Cu



$2p \rightarrow 1s$ transition violating Pauli principle

~ 7.7 keV in Cu

Paul Indelicato (Ecole Normale Supérieure et Université Pierre et Marie Curie)

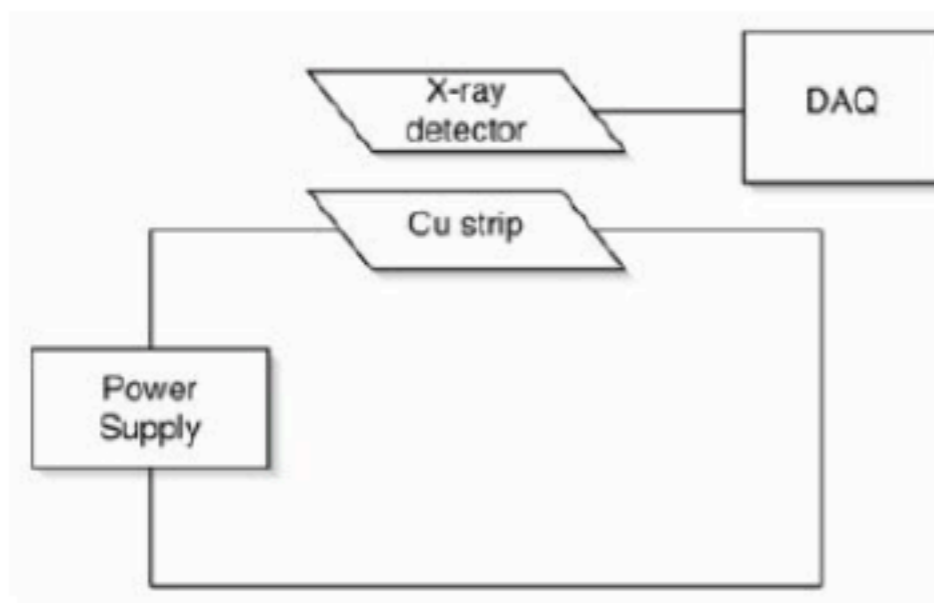
Multiconfiguration Dirac-Fock approach

Accounts for the shielding of the two inner electrons

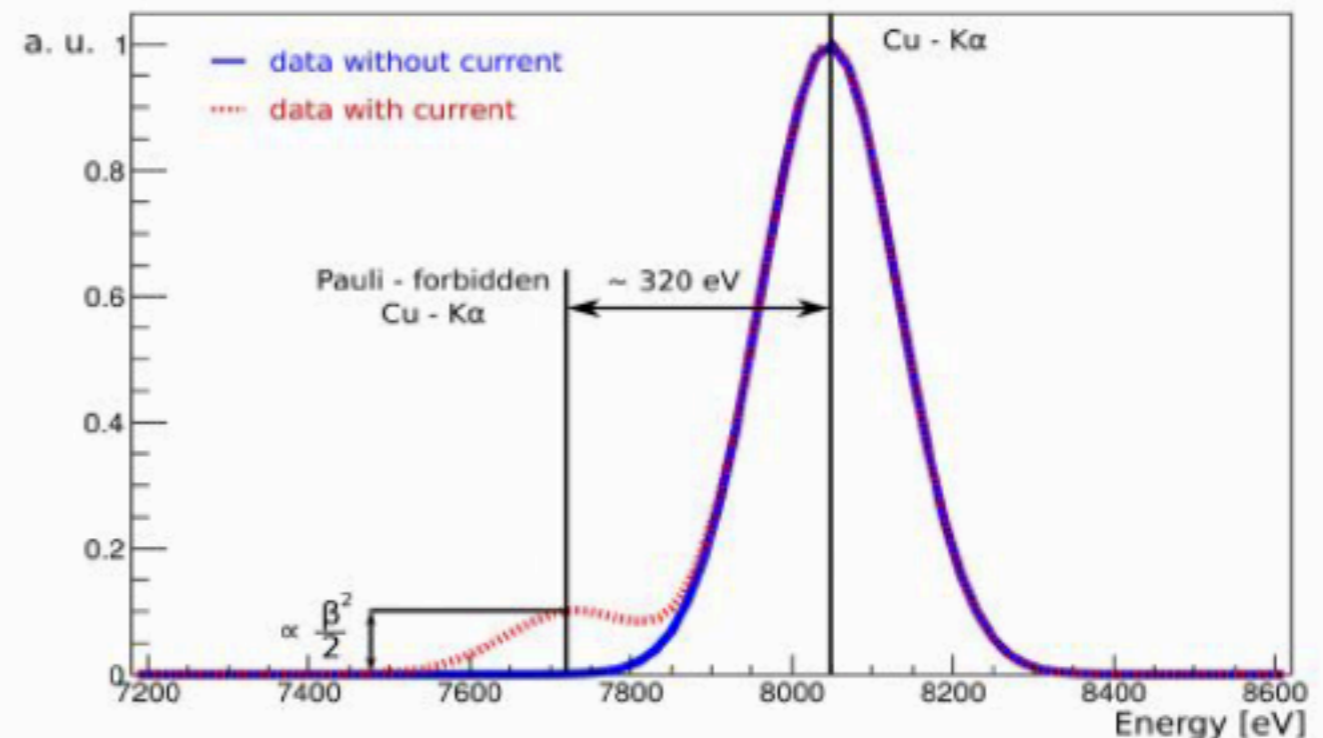
**Greenberg, O. W. & Mohapatra, R. N., Phys Rev Lett 59, (1987).
E. Ramberg and G. A. Snow, Phys Lett B 238, 438-441(1990)**

**Search for anomalous electronic transitions in Cu
induced by a circulating current**

**introduced electrons interact with the valence electrons
search transition from 2p to 1s already filled by 2 electrons
alternated to X-ray background measurements without current**

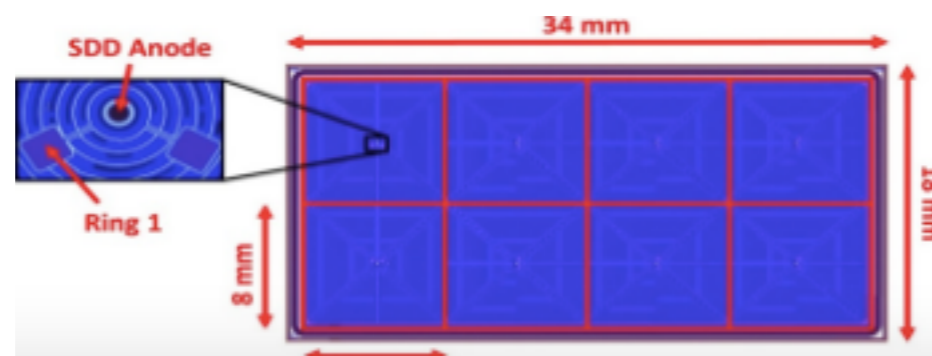
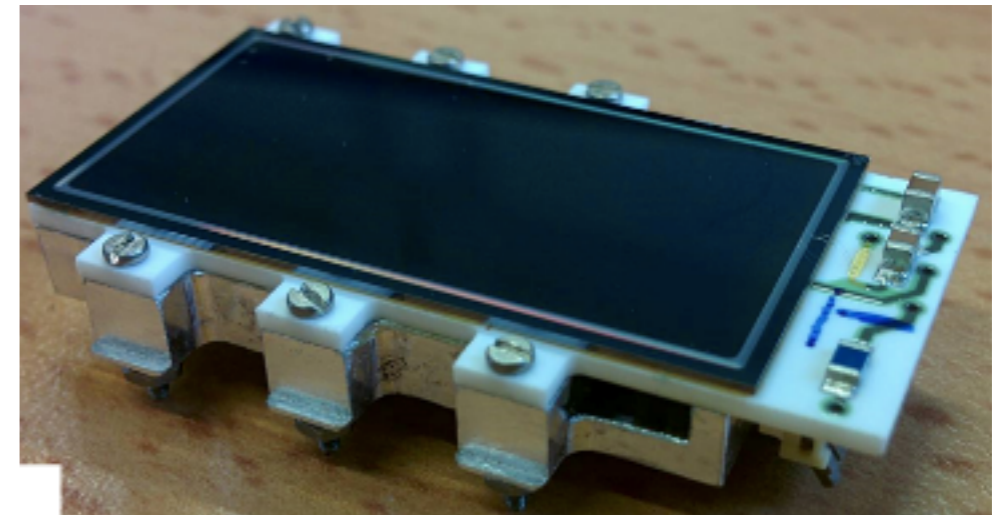
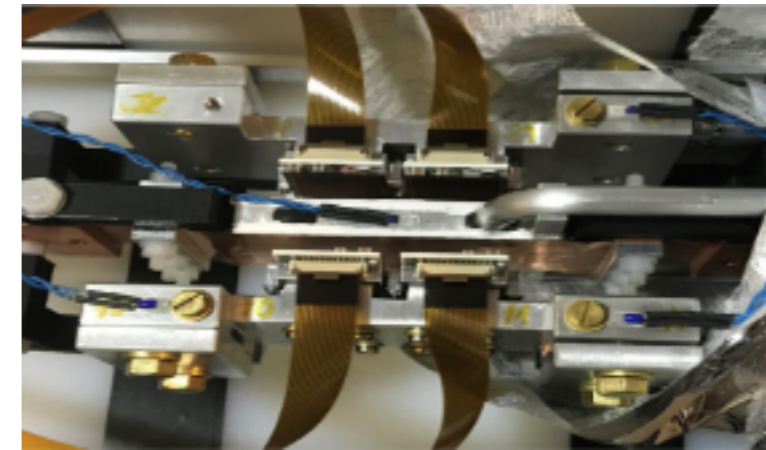
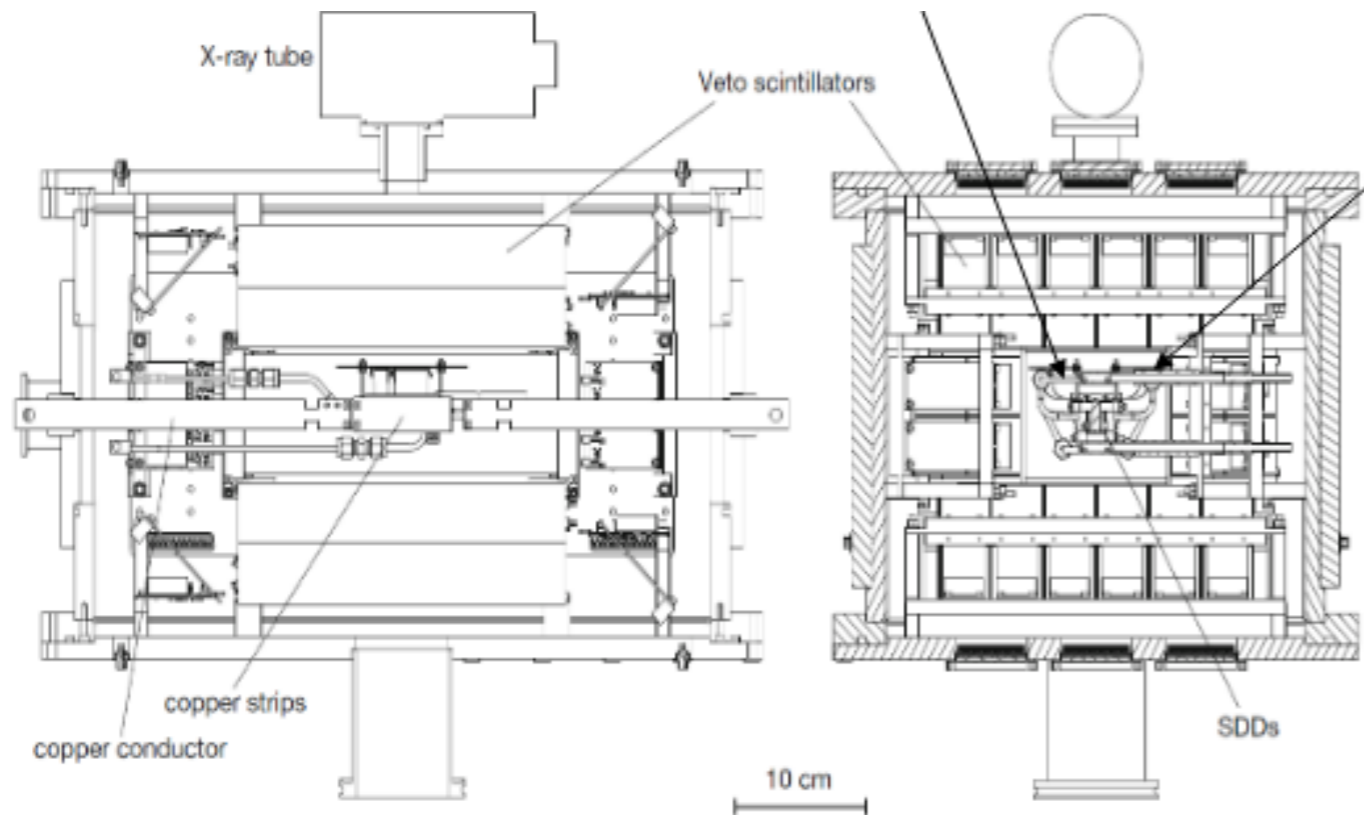


PEP Violation Signal



The VIP-2 Experiment

Silicon Drift Detectors (SDDs) higher resolution (190 eV FWHM at 8.0 \rightarrow keV), faster (triggerable) detectors. 4 arrays of 2 x 4 SDDs 8mm x 8mm each, liquid argon closed circuit cooling 170 °C

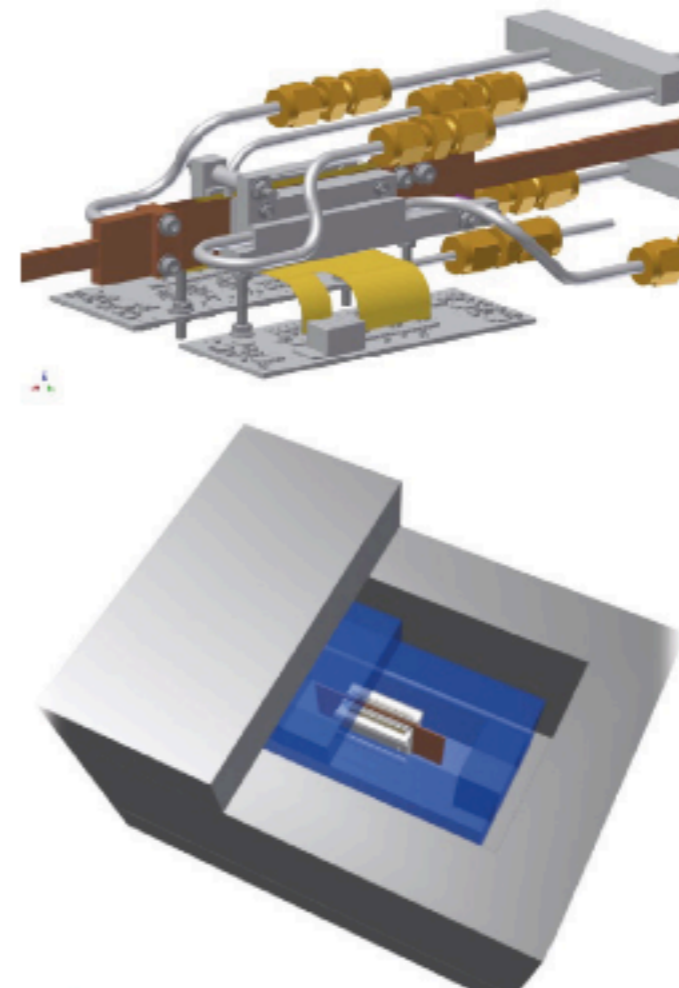
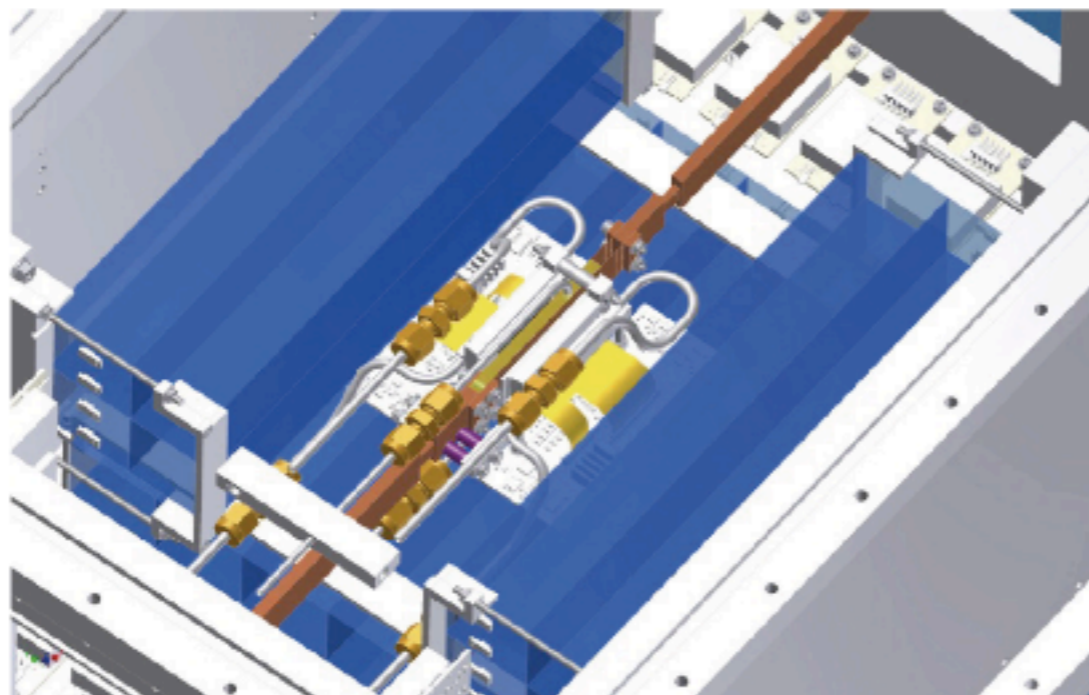


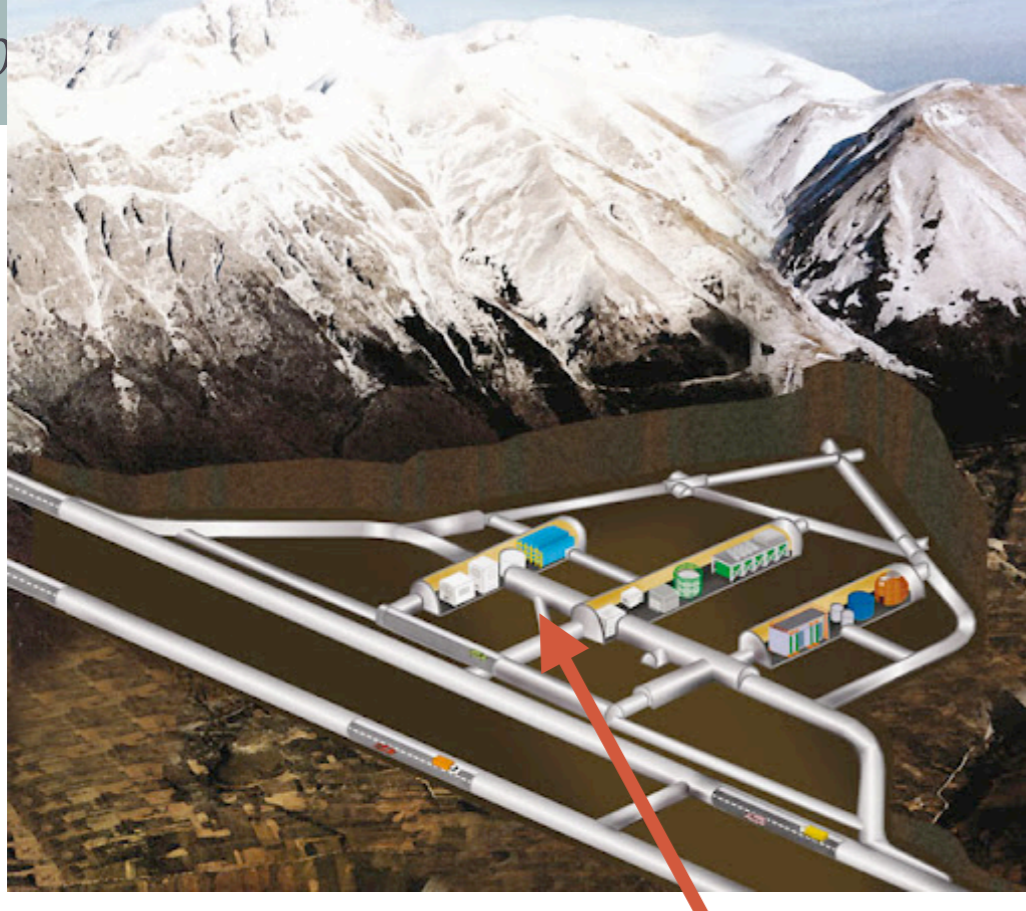
The VIP-2 Experiment

2 strip shaped Cu targets (25 μm x 7 cm x 2 cm) more compact target \rightarrow higher acceptance, thinner \rightarrow higher efficiency
DC current supply to Cu bars

Cu strips cooled by a closed Fryka chiller circuit \rightarrow higher current (100 A) @ 20 °C of Cu target implies 1 °K heating in SDDs

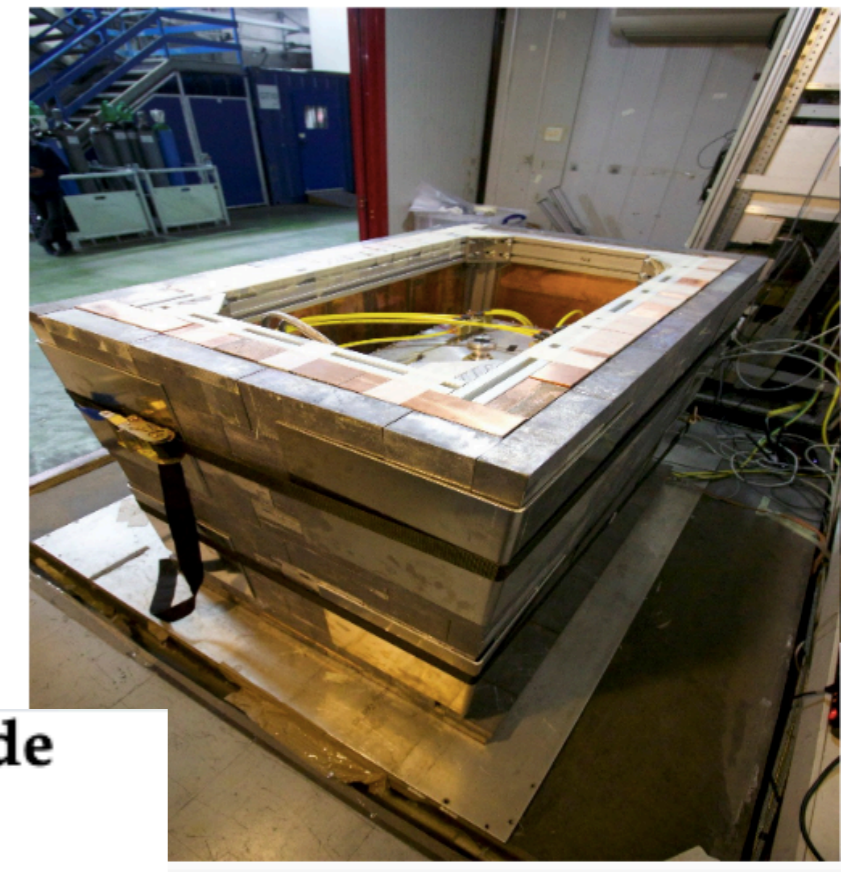
Sketch of the VIP2 Setup:



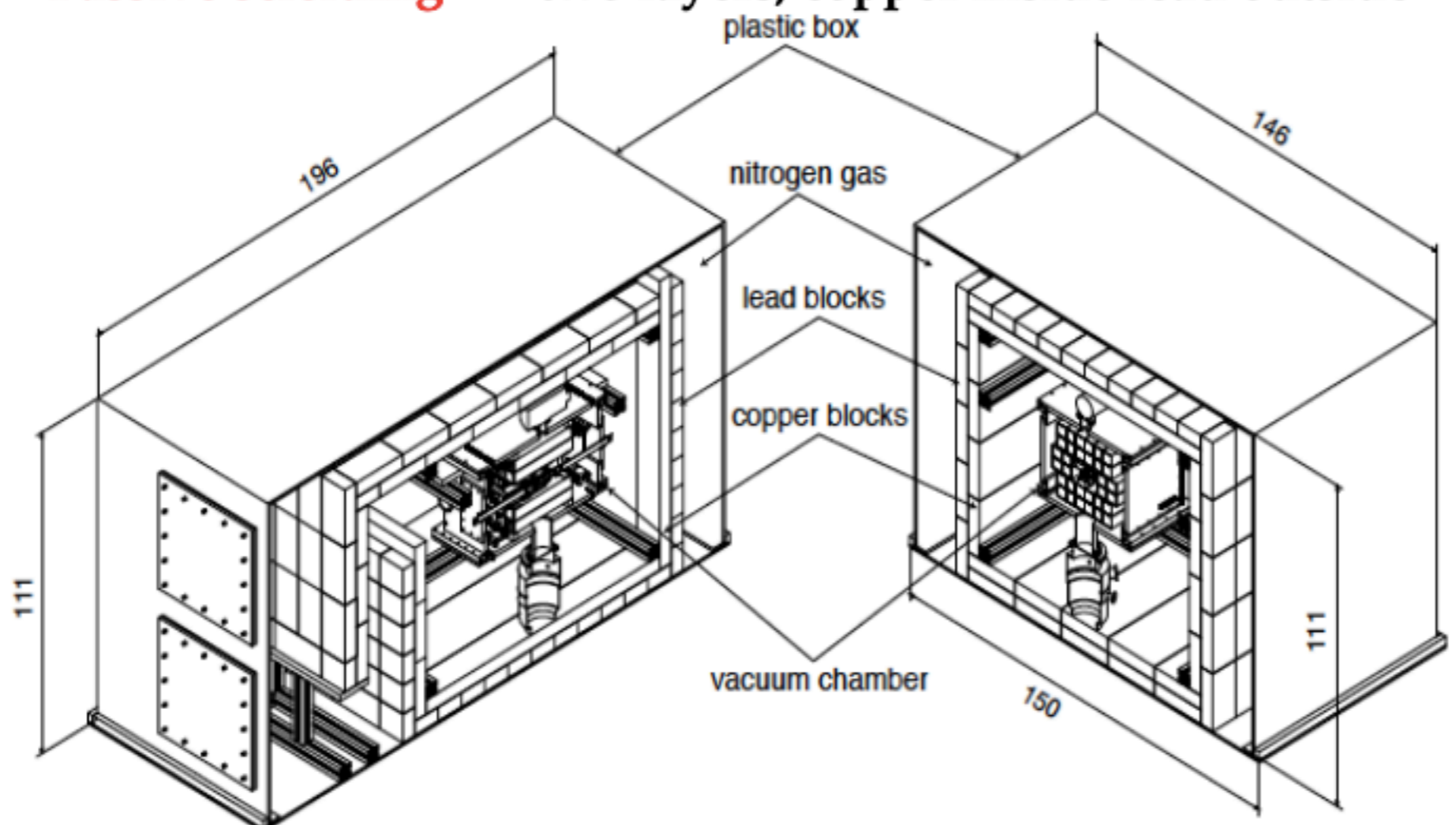


1400 m rock coverage

Upgrade concluded in April 2019:



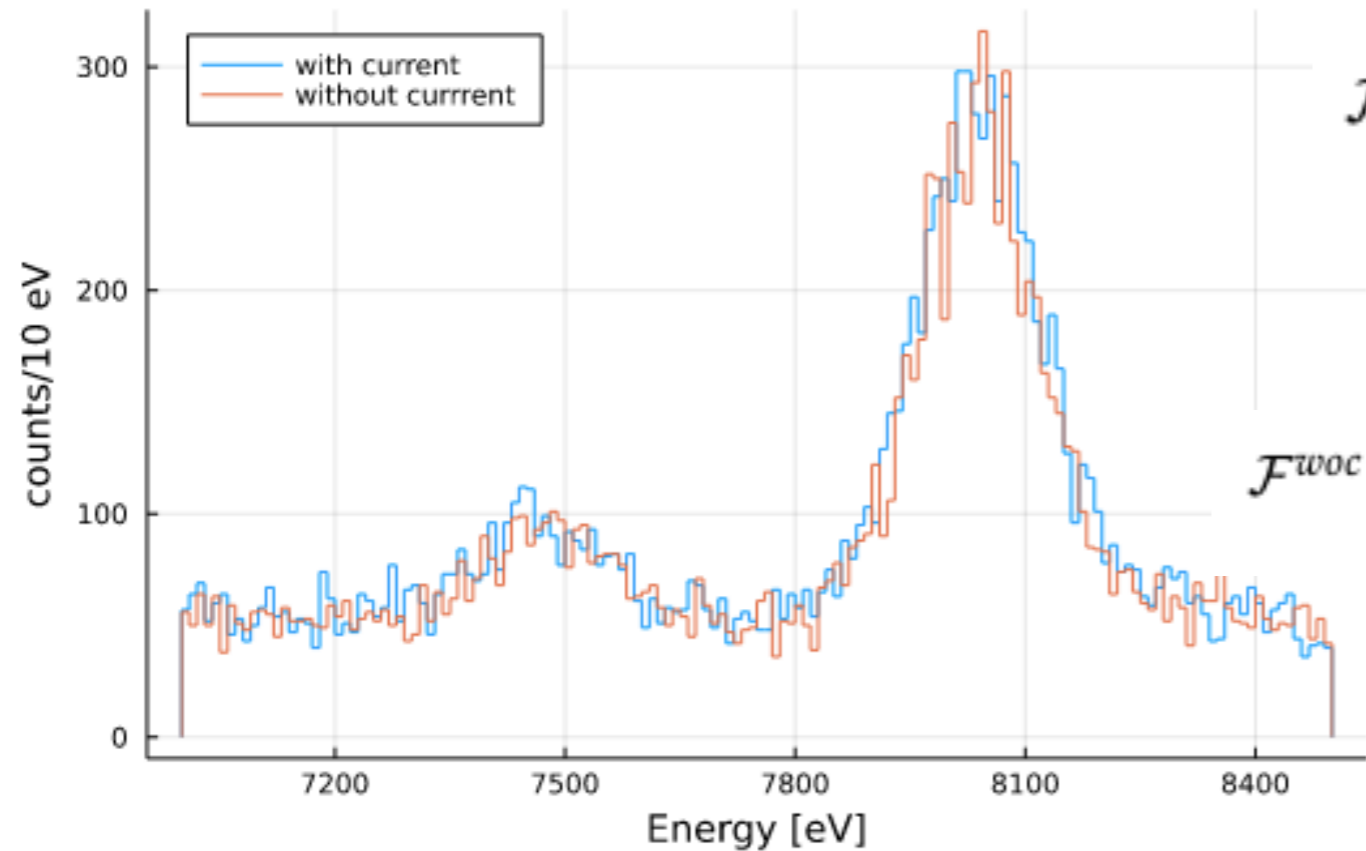
Passive shielding → two layers, copper inside lead outside



Symmetry **2022**, 14(5), 893;
<https://doi.org/10.3390/sym14050893>

Six months of data taking

VIP-2 Experiment



Description spectrum with current

$$\mathcal{F}^{wc}(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S}) = y_1 \times Ni(\theta_1, \theta_2) + y_2 \times Cu(\theta_3, \theta_4) + y_3 \times pol_1(\theta_5) + \mathcal{S} \times PEPV(\theta_4)$$

Description spectrum without current

$$\mathcal{F}^{woc}(\boldsymbol{\theta}, \mathbf{y}) = y_1 \times Ni(\theta_1, \theta_2) + y_2 \times Cu(\theta_3, \theta_4) + y_3 \times pol_1(\theta_5)$$

Likelihood

$$\mathcal{L}(\mathcal{D}^{wc}, \mathcal{D}^{woc} | \boldsymbol{\theta}, \mathbf{y}, \mathcal{S}) = \text{Poiss}(\mathcal{D}^{wc} | \mathcal{F}^{wc}(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S})) \times \text{Poiss}(\mathcal{D}^{woc} | \mathcal{F}^{woc}(\boldsymbol{\theta}, \mathbf{y} \times \mathcal{R}))$$

Bayesian

Frequentist

$$p(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S} | \mathcal{D}^{wc}, \mathcal{D}^{woc}) =$$

$$= \frac{\mathcal{L}(\mathcal{D}^{wc}, \mathcal{D}^{woc} | \boldsymbol{\theta}, \mathbf{y}, \mathcal{S}) p(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S})}{\int d\boldsymbol{\theta} d\mathbf{y} \mathcal{L}(\mathcal{D}^{wc}, \mathcal{D}^{woc} | \boldsymbol{\theta}, \mathbf{y}, \mathcal{S}) p(\boldsymbol{\theta}, \mathbf{y}, \mathcal{S})}$$

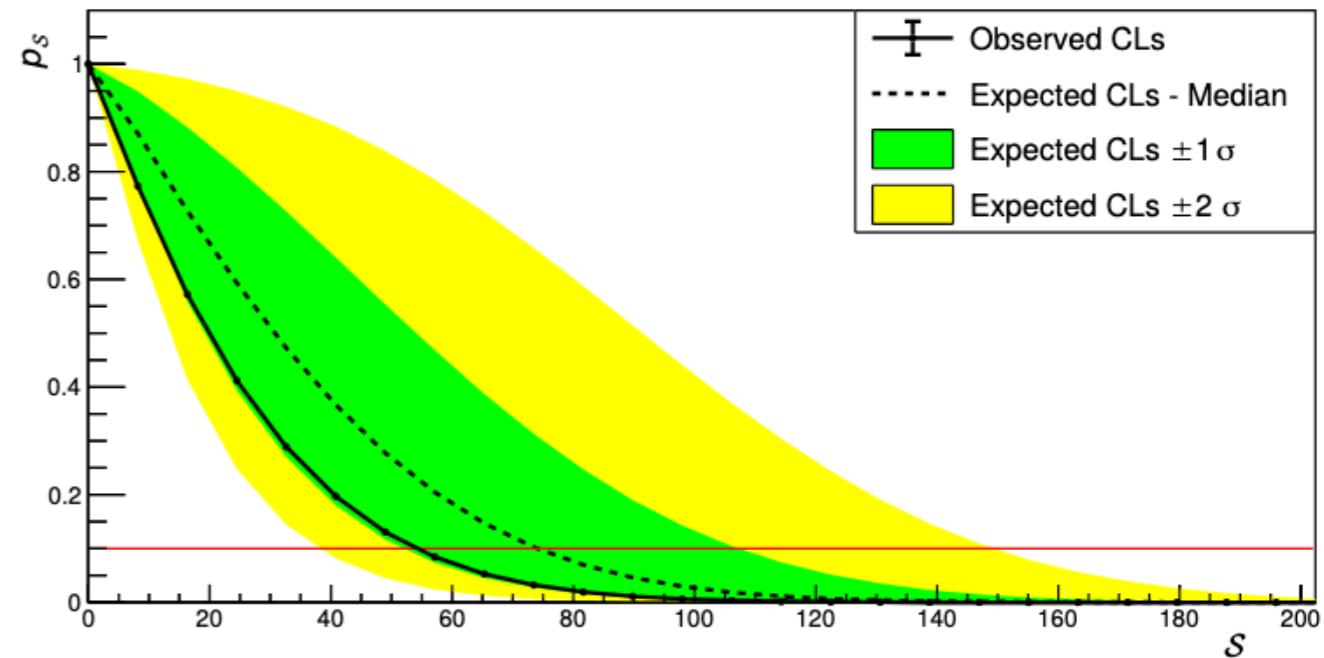
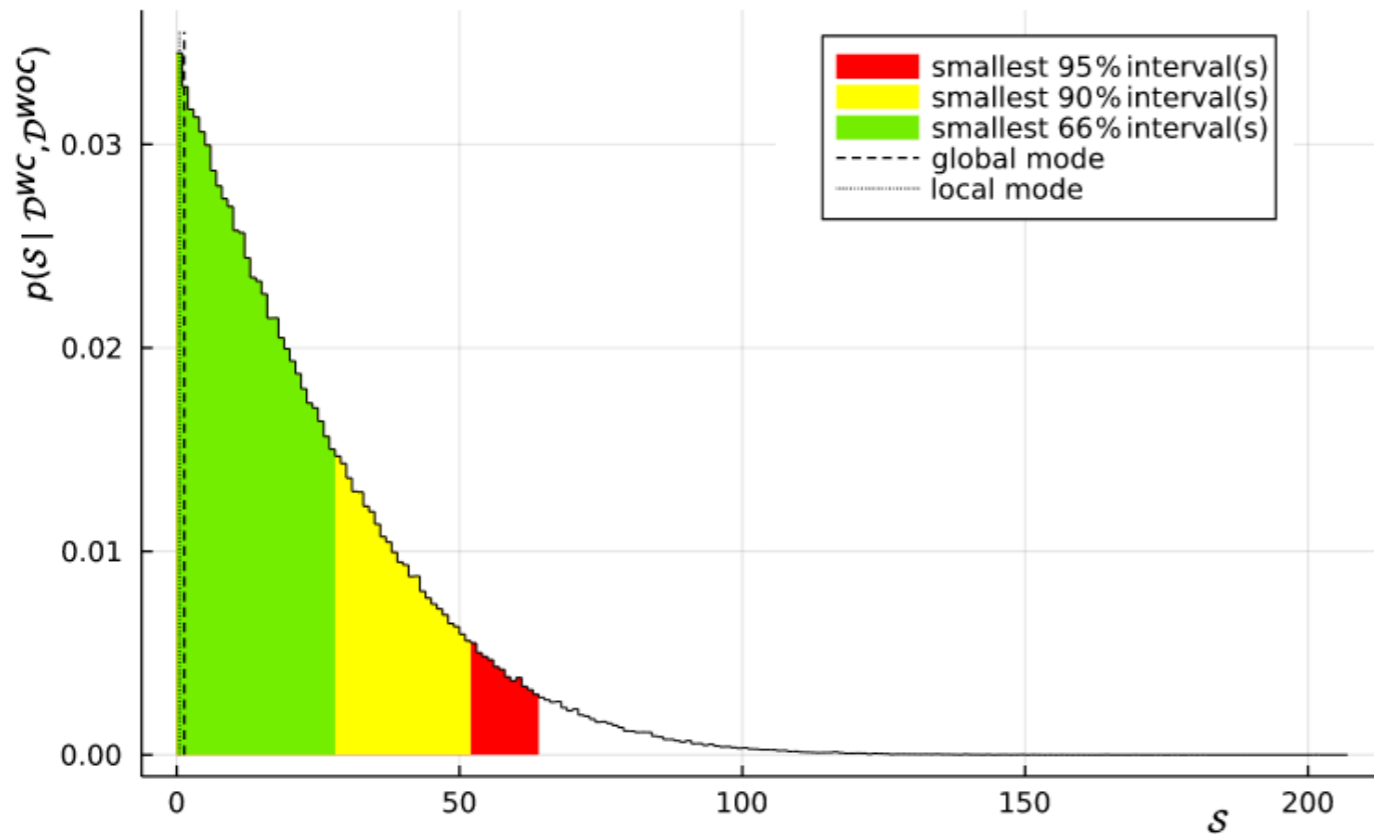
$$t_S = -2 \ln \Lambda(\mathcal{S}) = -2 \ln \frac{\mathcal{L}(\hat{\boldsymbol{\theta}}, \hat{\mathbf{y}}, \mathcal{S})}{\mathcal{L}(\hat{\boldsymbol{\theta}}, \hat{\mathbf{y}}, \hat{\mathcal{S}})}$$

$$p_S = \int_{t_{obs}}^{\infty} f(t_S | \mathcal{S}) dt_S$$

Symmetry **2022**, *14*(5), 893;
<https://doi.org/10.3390/sym14050893>

Bayesian

Frequentist

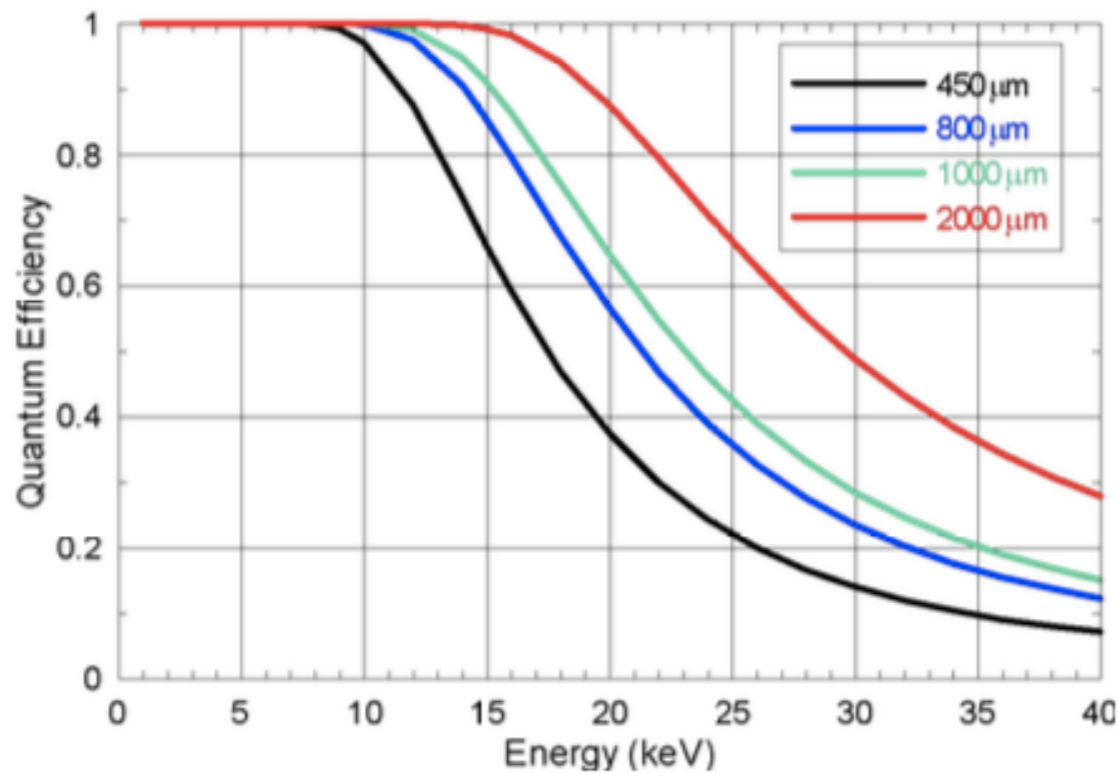
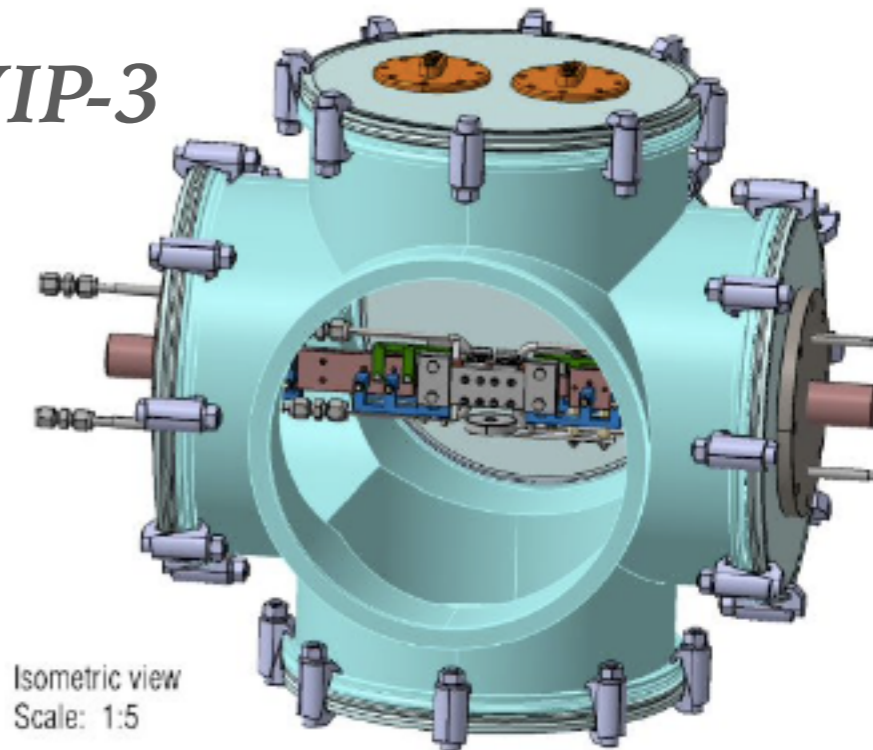


$$\beta^2 / 2 \leq 8.6 \times 10^{-31} \text{ (Bayesian)}, \quad \beta^2 / 2 \leq 8.9 \times 10^{-31} \text{ (CL}_s\text{)}$$

New article in preparation with all the available statistics!

VIP-2 experimental upgrade: VIP-3

- new vacuum chamber, increase the number of SDD detectors, increase the geometrical efficiency, higher current up to 400 A
- New thermal contact between cold finger and SDDs
- New target cooling system
- Higher quantum efficiency needed for the SDDs at higher Z: use 1 mm thick SDDs, allowing to scan e.g. Ag, Sn and Pd



- 2x4 SDDs, 8x8 mm² each, in production with FBK & politecnico di Milano

