



Anharmonic Effects on the Squeezing of Axion Perturbations

Supervisors:

Takeshi Kobayashi

Sabino Matarrese

Collaborators:

Nicola Bartolo

Matteo Viel

Introduction

- ✓ The history of the Universe undergoes a period of exponential expansion, **inflation**.
- ✓ Quantum fluctuations provide the seeds for structure formation.
- ✓ The CMB sky we see today is classical.

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Quantum to classical transition

- ✓ Inflation itself provides an explanation due to **squeezing**.
- ✓ Further source of classicalization: **reheating**.

Framework

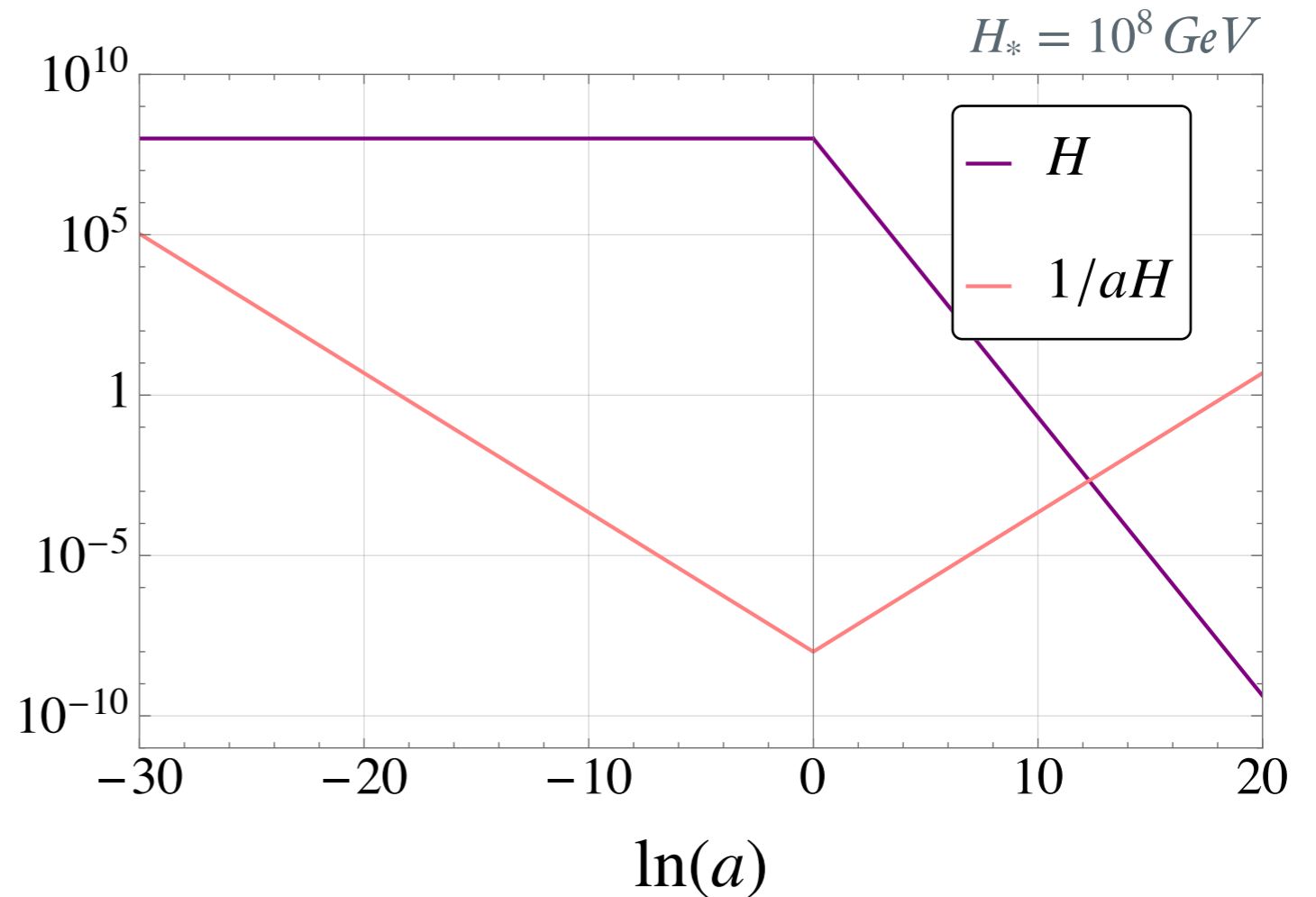
✓ de Sitter (DS) inflation followed by a Radiation Domination (RD) phase

✓ Axions produced via misalignment mechanism with $f > \max(T_{rh}, H_I)$

✓ Axion is a spectator field

✓ Instantaneous reheating

$$H = \begin{cases} H_* & a < 0 \\ H_* a^{-2} & a > 0 \end{cases}$$



Framework

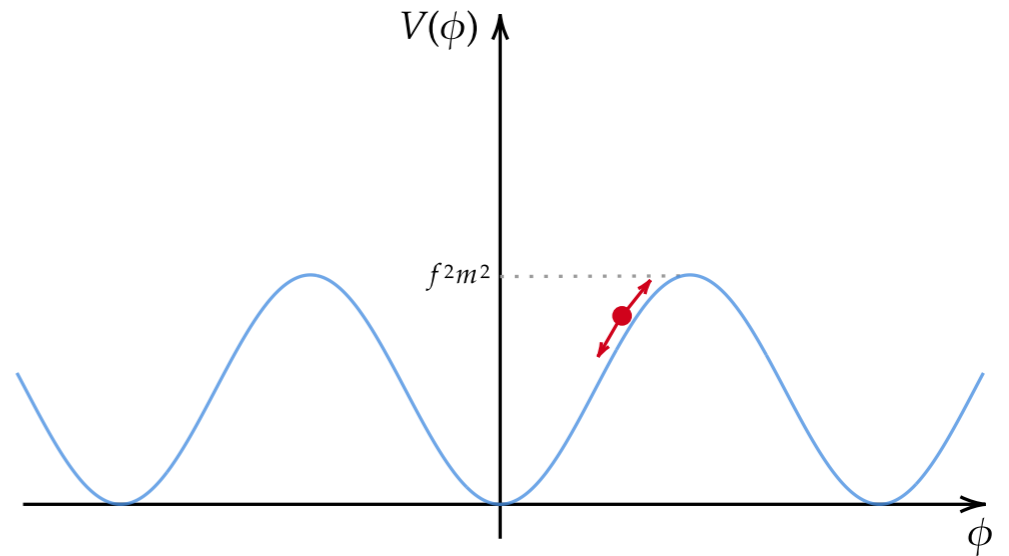
What about the axion potential?

$$V(\phi) = f^2 m_\phi^2 \left[1 - \cos \left(\frac{\phi}{f} \right) \right]$$



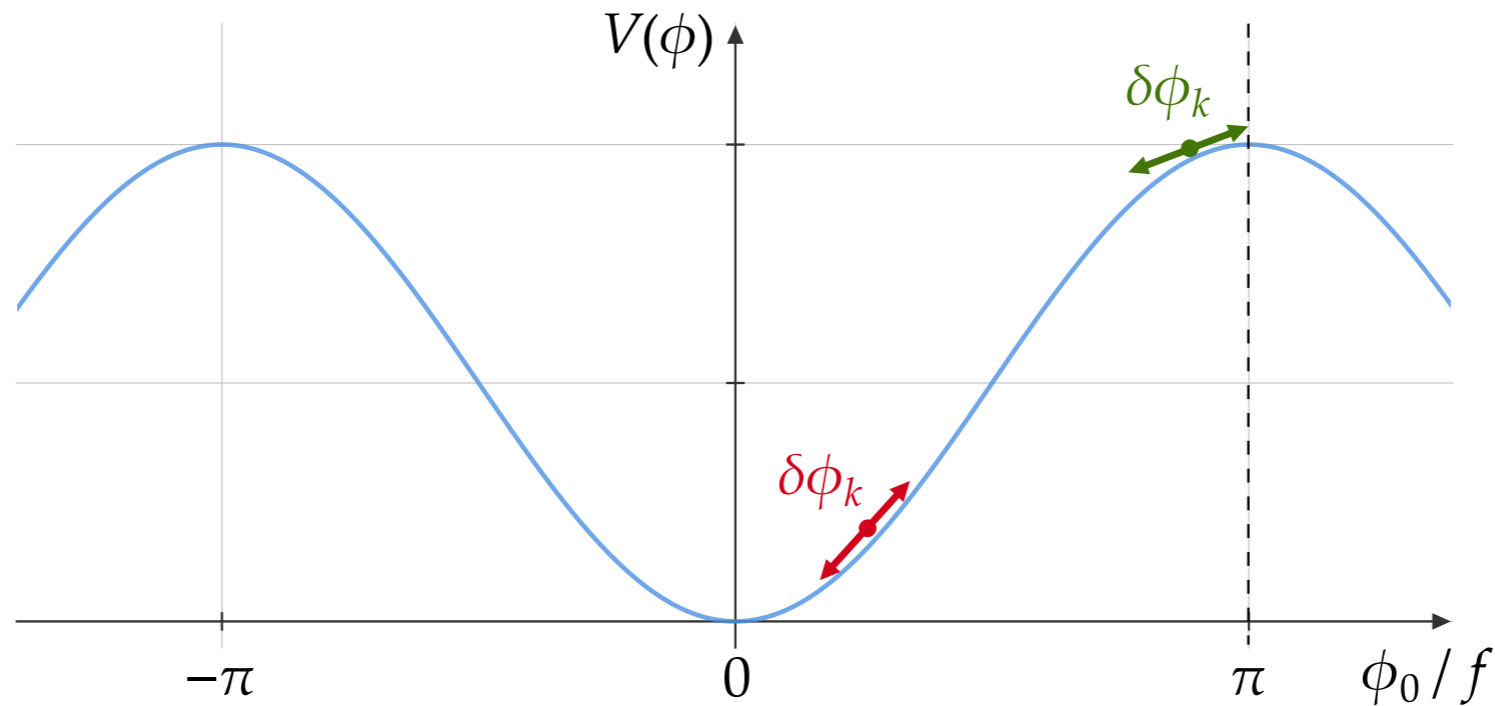
$$\ddot{\phi}_0 + 3H\dot{\phi}_0 + fm_\phi^2 \sin \left(\frac{\phi_0}{f} \right) = 0$$

$$\delta\ddot{\phi} + 3H\delta\dot{\phi} + \left[\frac{k^2}{a^2} + m_\phi^2 \cos \left(\frac{\phi_0}{f} \right) \right] \delta\phi = 0$$



Conclusions

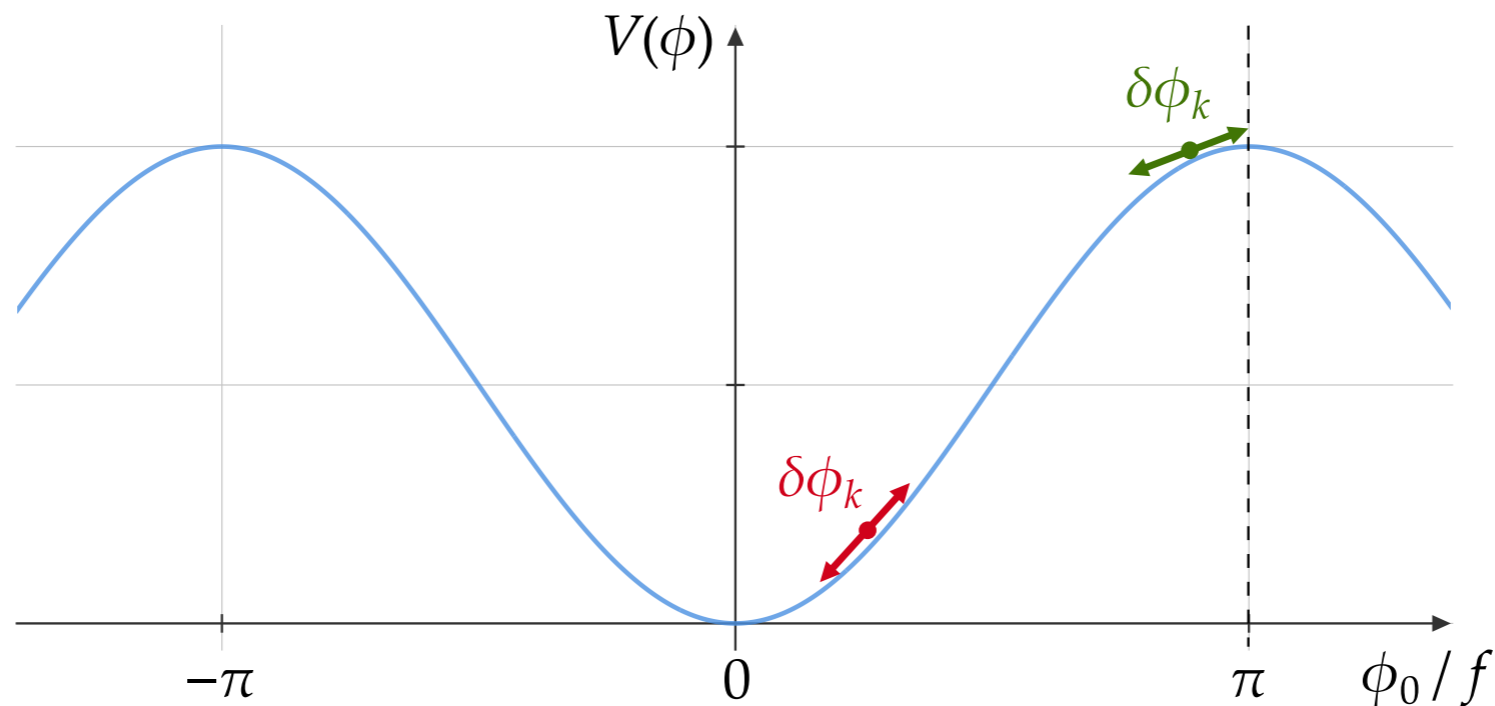
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- ✓ The number of particles and the energy density increase exponentially when approaching the hilltop of the potential
- ✓ Anharmonic effects increase also the amount of squeezing of the perturbations

Conclusions

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THANKS FOR THE ATTENTION
