

Astrofisica, Astronomia e l'esperimento Borexino

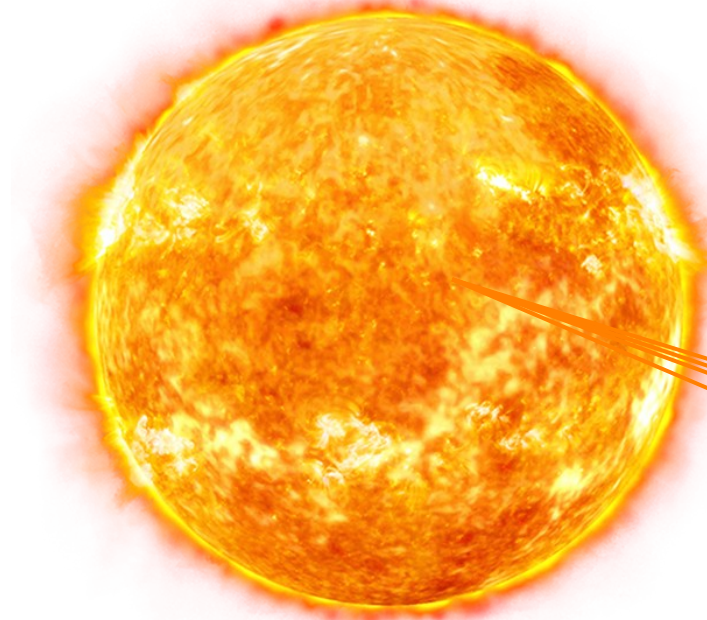


Nicola Rossi

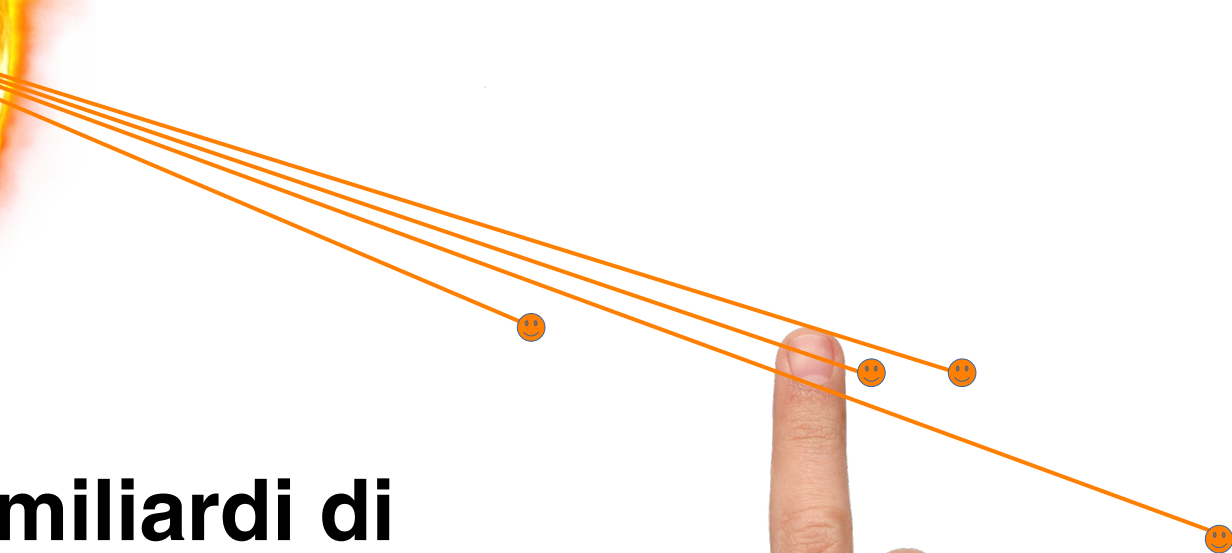
Laboratori Nazionali del Gran Sasso

NEUTRINI SOLARI E MASSIMI SISTEMI

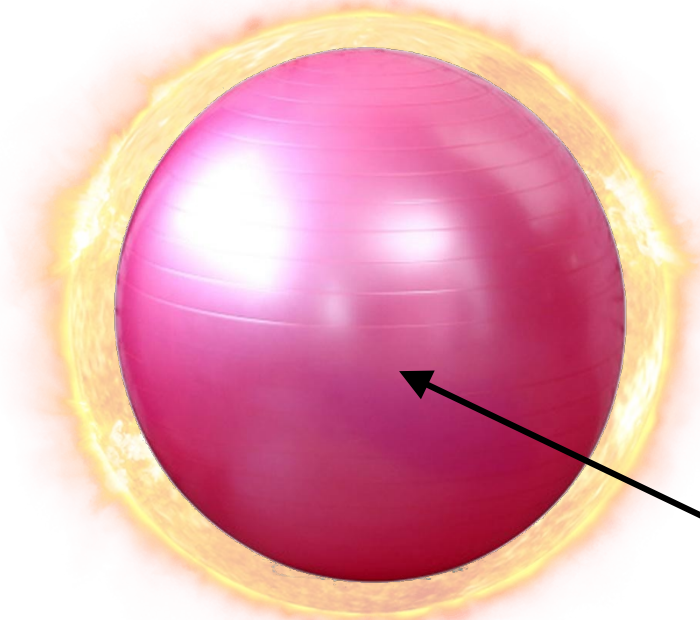
LNGS, 22 febbraio 2023



**Circa 60 miliardi di
neutrini per centimetro
quadro al secondo**



Quando è grande veramente il Sole?



Sole
(1 m)

100 m

Terra
(1 cm)



20 cm

Luna
(2.5 mm)



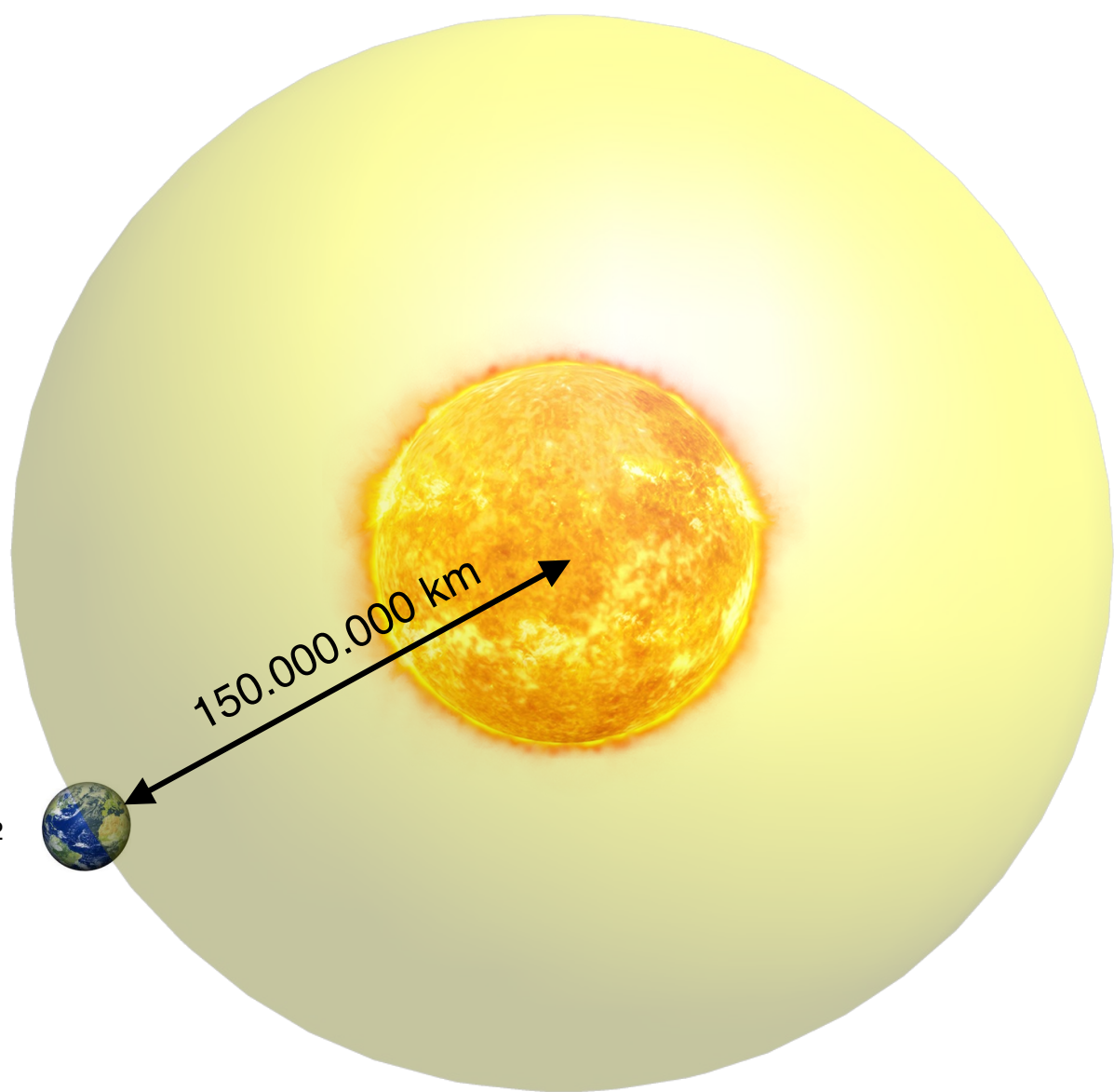
**Il sole irraggia sulla terra
circa 1.3 kW/m^2**



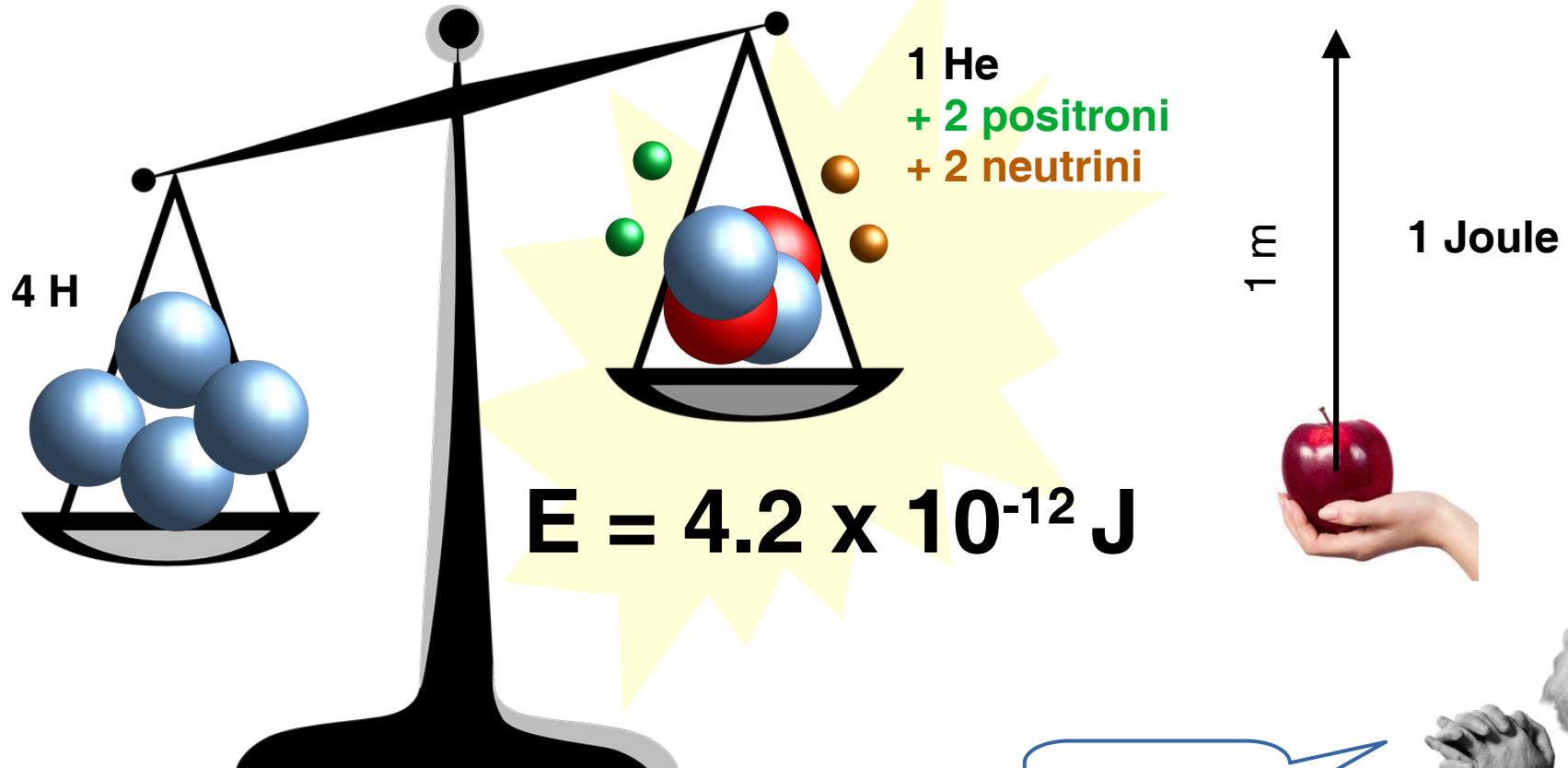
Potenza generata dal Sole:

$$1300 \text{ W} \times [\text{sfera}] = 4 \times 10^{26} \text{ W!}$$

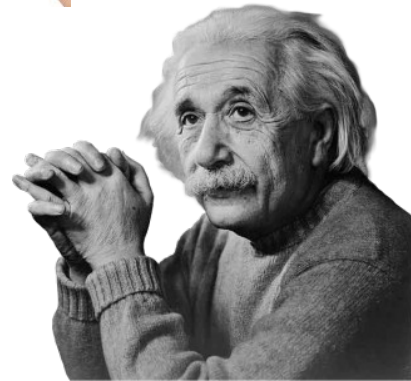
1.3 kW/m²



La fusione dell'idrogeno



$E = mc^2!$



A che velocità brucia il sole?

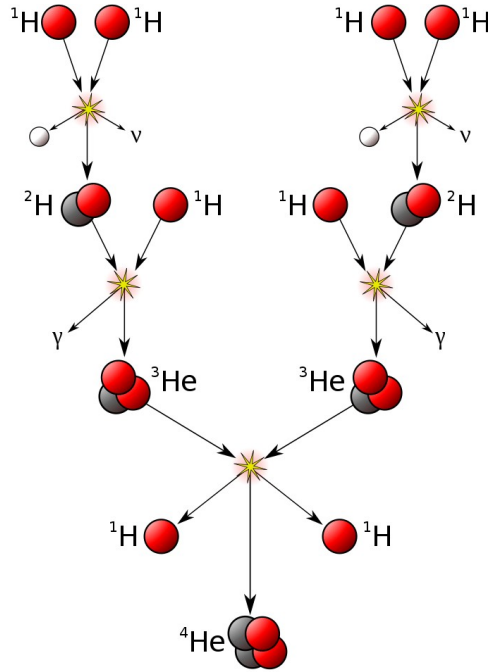
$$\frac{\text{potenza totale}}{\text{energia singola}} = \frac{4 \times 10^{26} \text{ J/s}}{4.2 \times 10^{-12} \text{ J}} \approx 10^{38} \frac{\text{fusioni}}{\text{secondo}}$$

600.000.000 tonnellate/sec di idrogeno

Quanti neutrini arrivano sulla terra?

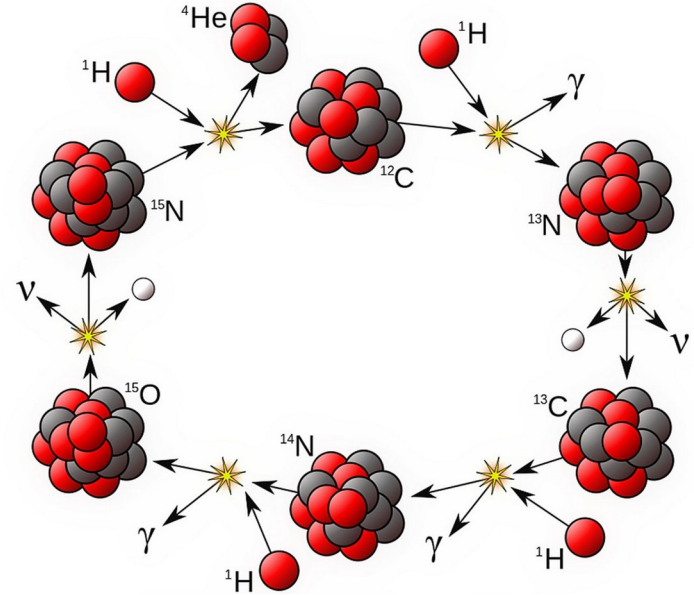
$$\frac{2 \times 10^{38}}{[\text{sfera}]} \approx 60 \times 10^9 \frac{\text{neutrini}}{\text{secondo} \times \text{cm}^2}$$

Catena protone-protone



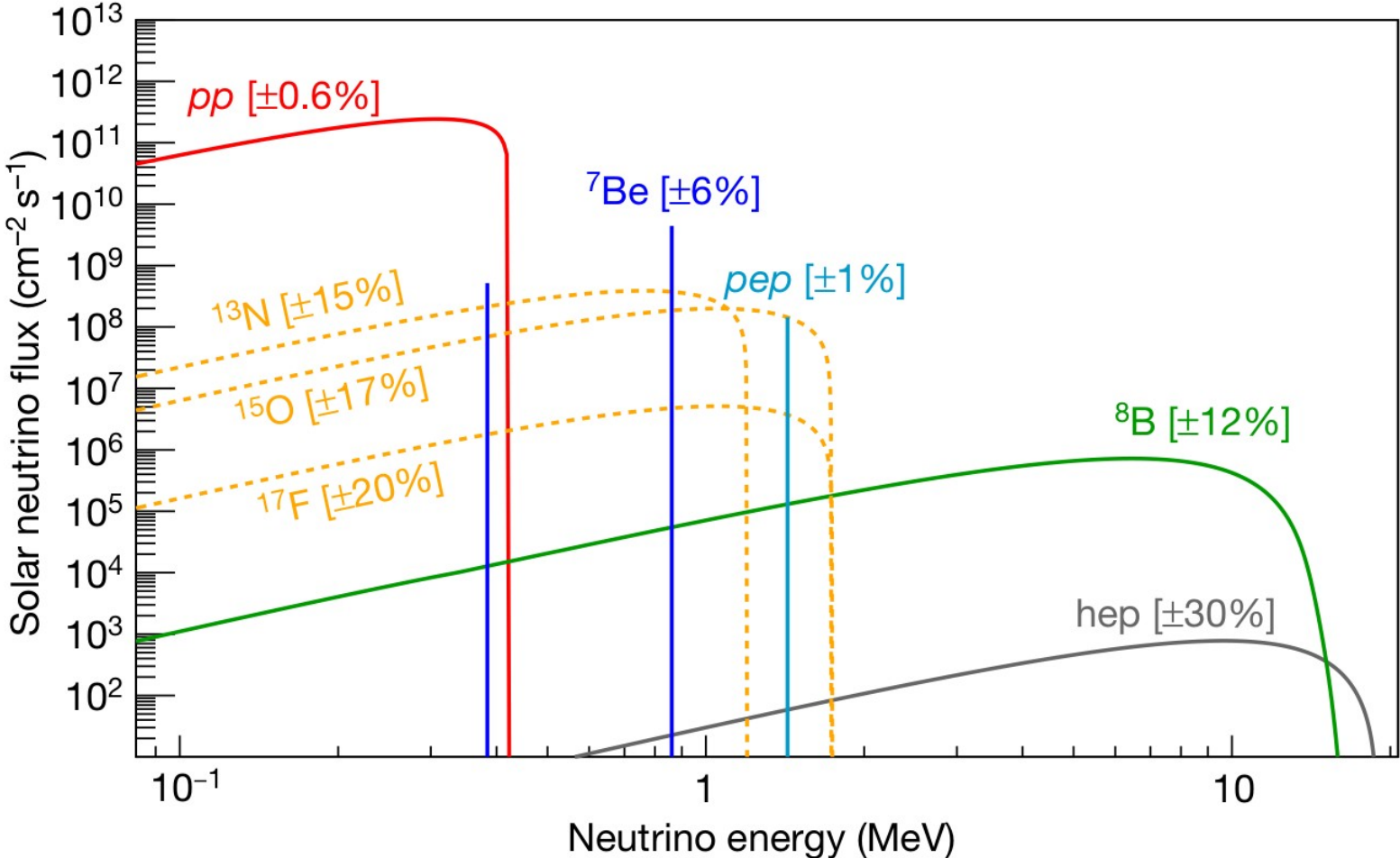
99%

Ciclo CNO

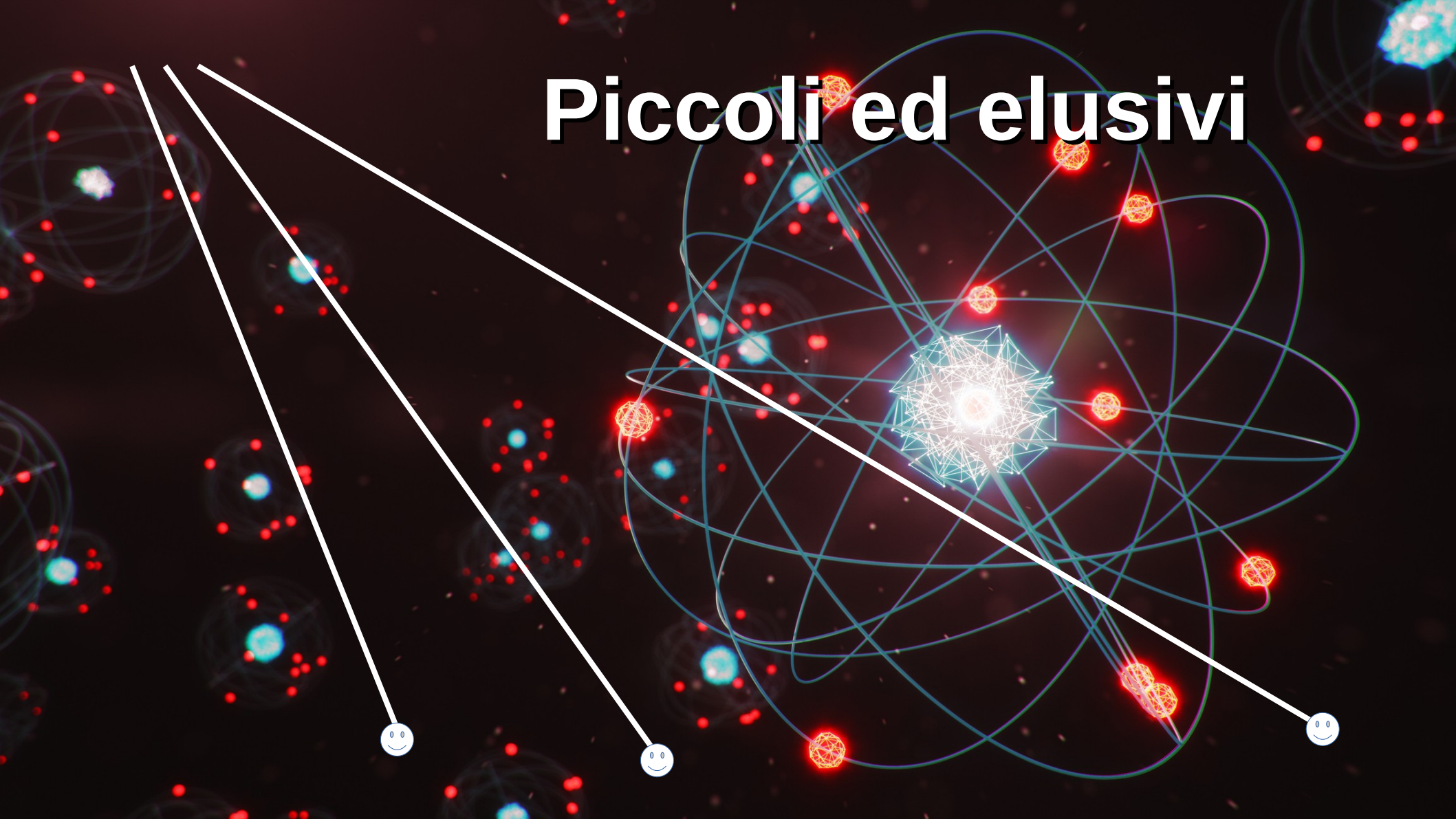


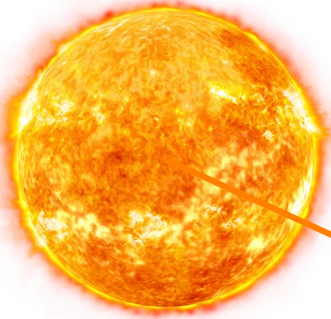
1%

Neutrini solari



Piccoli ed elusivi





**Che probabilità ha
un neutrino di
interagire in un
litro di acqua?**

$$60 \cdot 10^9 \frac{\nu}{\text{cm}^2} \times 3 \cdot 10^{26} \times 10^{-45} \text{cm}^2 \times 3 \cdot 10^7 \text{s} \approx 1 \frac{\nu}{\text{y}}$$

Flusso
di neutrini
sulla Terra

Quanti elettroni
in un litro
di acqua?

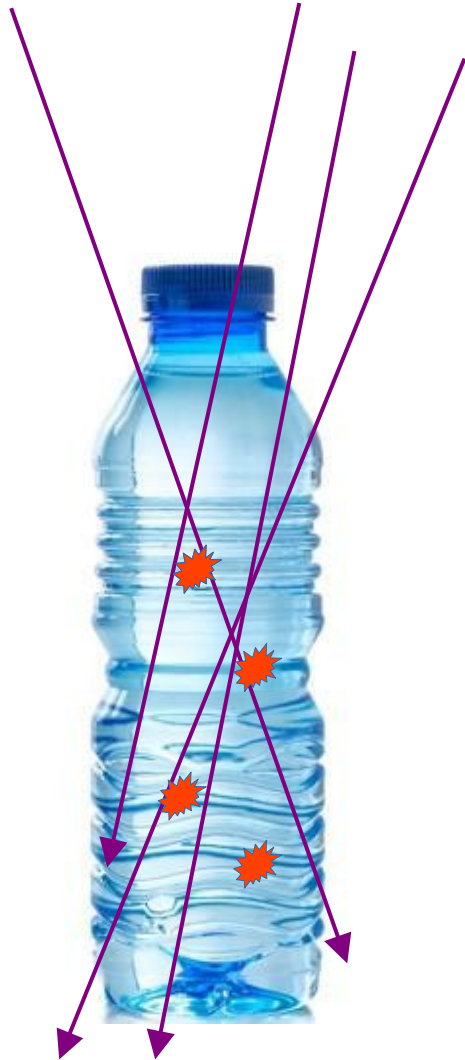
“dimensione”
dei neutrini

Secondi
In un anno

**Solo un neutrino
all'anno fermato
da un litro di
acqua!**

**Pioggia incessante
di raggi cosmici**

**Un litro di acqua
minerale contiene parti
per miliardo di
impurezze
(una disintegrazione al
secondo)**



**Un neutrino su
100.000.000
eventi di
fondo:
Cercare l'ago
nel pagliaio!**

Un rivelatore di neutrini solari deve essere

1. Protetto dalla roccia
2. Un miliardo di volte più puro di una buona acqua potabile
3. Di grandissimo volume

→ **BOREXINO**



**x100000
centinaia di neutrini
al giorno**

Anni '90



BOREXINO
Un'avventura
lunga pù di 30
anni

Gianpaolo Bellini
Frank Calaprice
Raju Ragavan

Sardegna 2015





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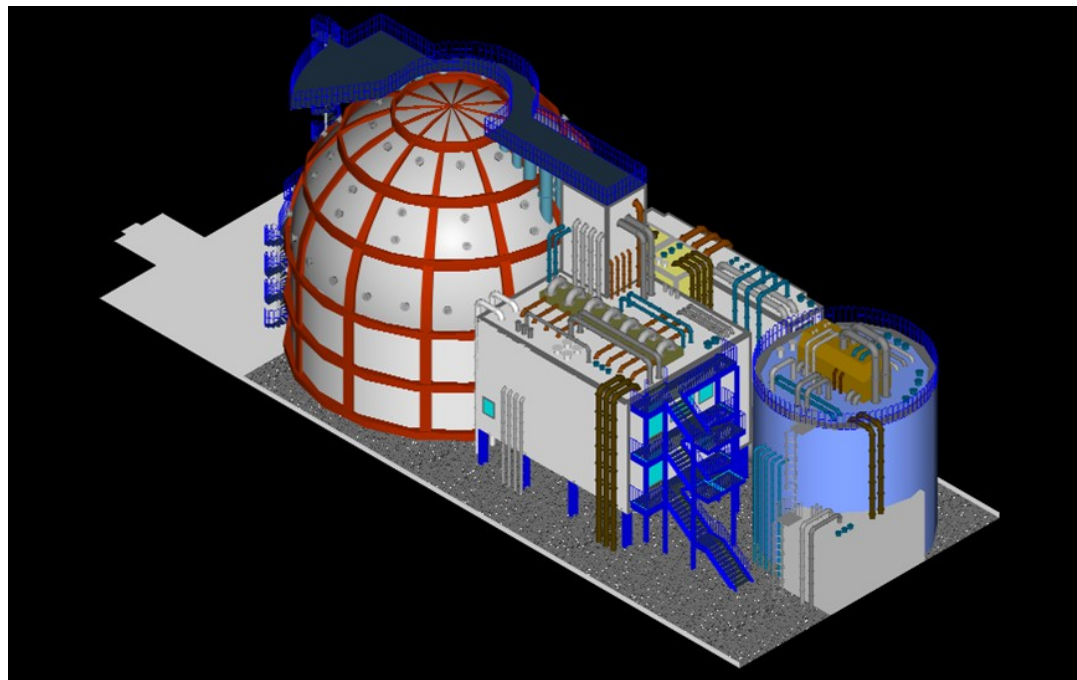
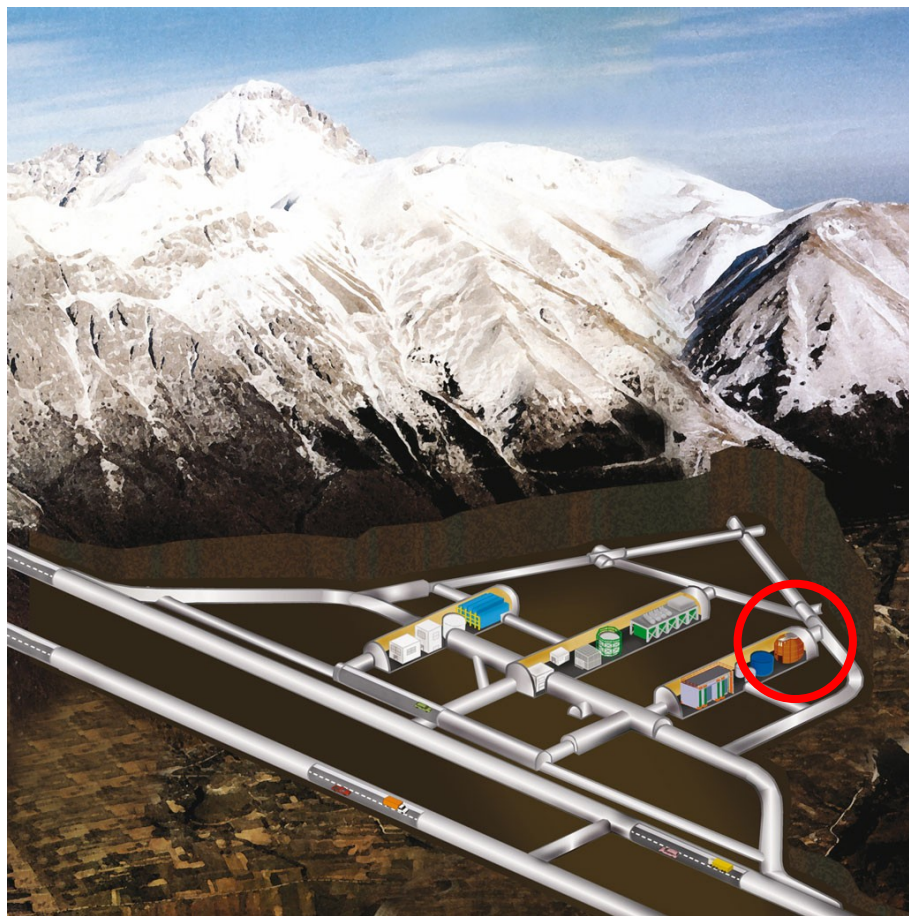


Dic. 2021



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Il rivelatore Borexino ai LNGS

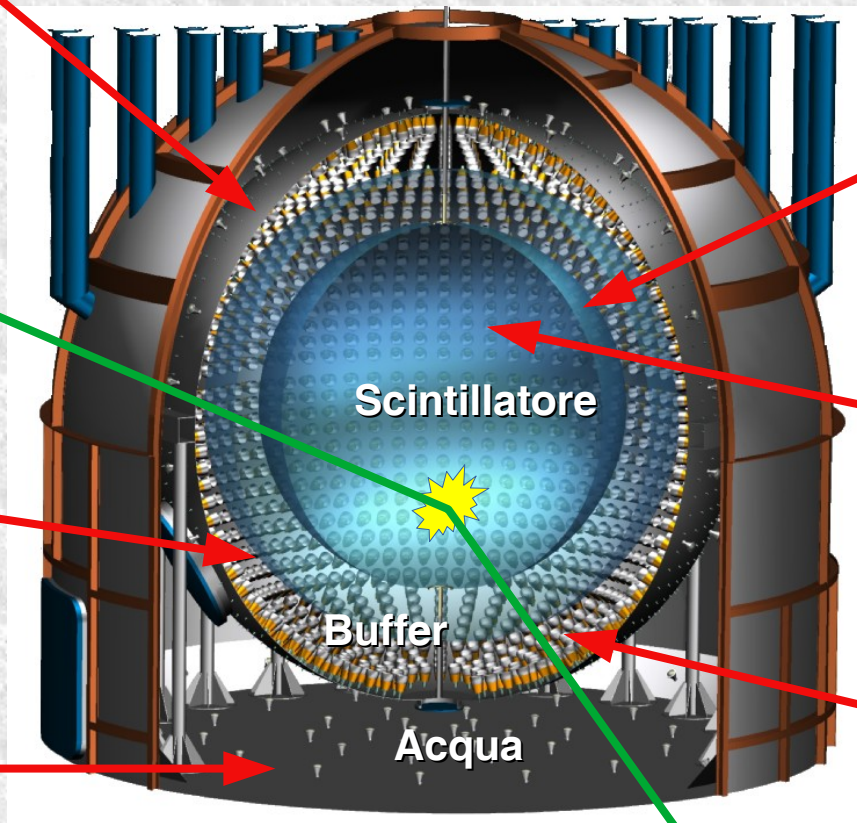


Sfera di acciaio



Pallone esterno di nylon

Acqua ultra-pura



Pallone interno di nylon

Scintillatore

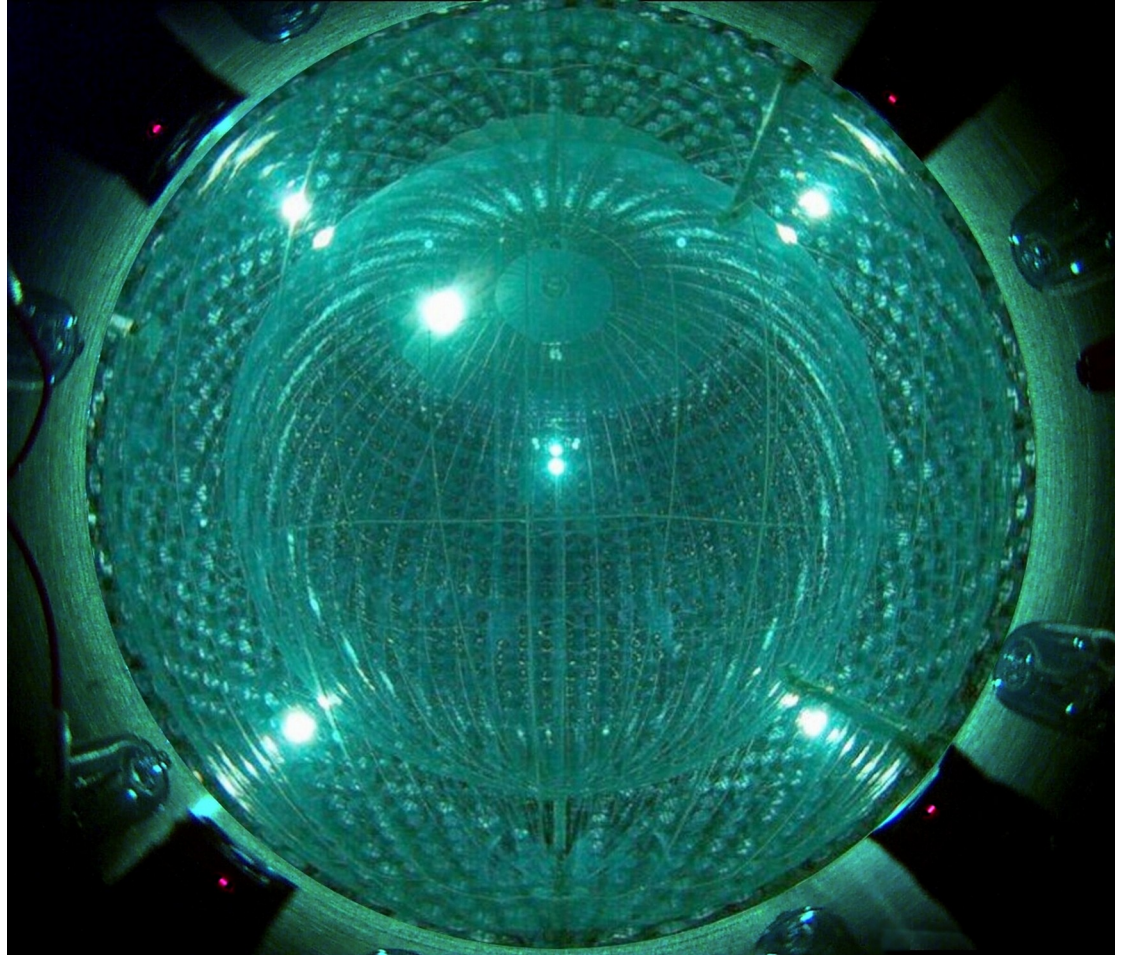
Tubi fotomoltiplicatori



Prima del riempimento

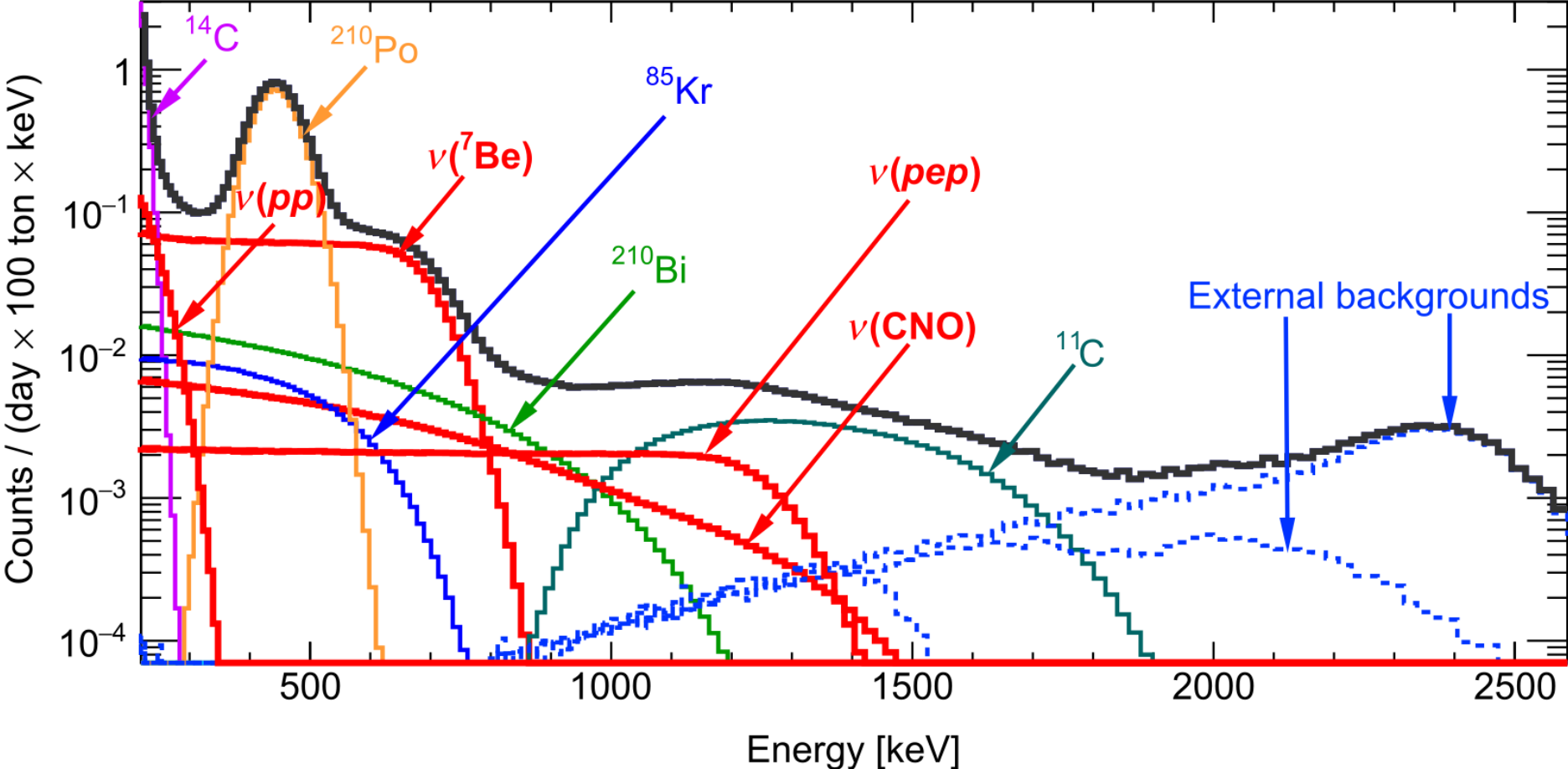


Nella Sala C



Dalla telecamera

Distribuzione energetica degli eventi rivelati da BOREXINO



Risultati di BOREXINO

nature

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Article | [Published: 24 October 2018](#)

Comprehensive measurement of *pp*-chain solar neutrinos

[The Borexino Collaboration](#)

[Nature](#) **562**, 505–510 (2018) | [Cite this article](#)

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Abstract

About 99 per cent of solar energy is produced through sequences of nuclear reactions that convert hydrogen into helium, starting from the fusion of two protons (the *pp* chain). The neutrinos emitted by five of these reactions represent a unique probe of the Sun's internal working and, at the same time, offer an intense natural neutrino beam for fundamental physics. Here we report a complete study of the *pp* chain. We measure the neutrino–electron elastic scattering rates for neutrinos produced by four reactions of the chain: the initial proton–proton fusion, the electron-capture decay of beryllium-7, the three-body proton–electron–proton (*pep*) fusion, here measured with the highest precision so far achieved, and the boron-8 beta decay, measured with the lowest energy threshold. We also set a limit on the

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Experimental evidence of neutrinos produced in the CNO fusion cycle in the Sun

[The Borexino Collaboration](#)

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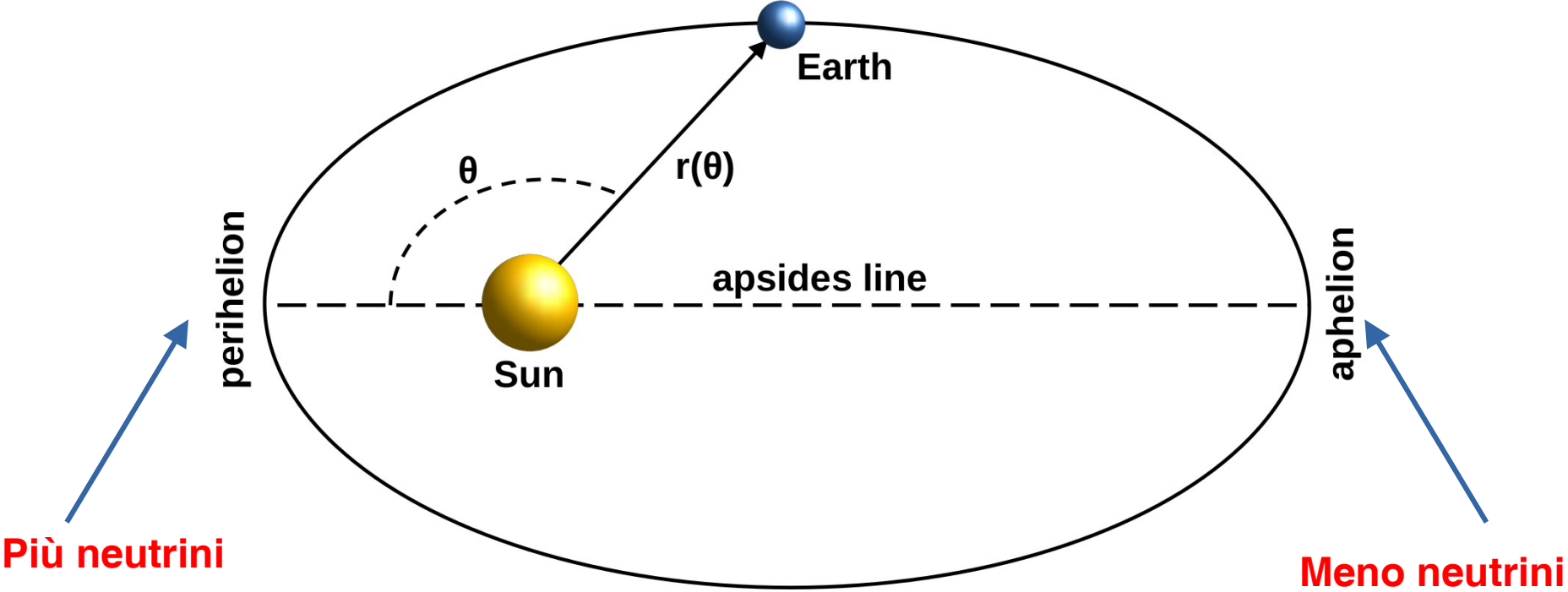
Abstract

For most of their existence, stars are fuelled by the fusion of hydrogen into helium. Fusion proceeds via two processes that are well understood theoretically: the proton–proton (*pp*) chain and the carbon–nitrogen–oxygen (CNO) cycle^{1,2}. Neutrinos that are emitted along such fusion processes in the solar core are the only direct probe of the deep interior of the Sun. A complete spectroscopic study of neutrinos from the *pp* chain, which produces about 99 per cent of the solar energy, has been performed previously³; however, there has been no reported experimental evidence of the CNO cycle. Here we report the direct observation, with a high statistical significance, of neutrinos produced in the CNO cycle in the Sun. This experimental evidence was obtained using the highly radiopure, large-volume, liquid-scintillator detector of Borexino, an experiment located at the underground Laboratori

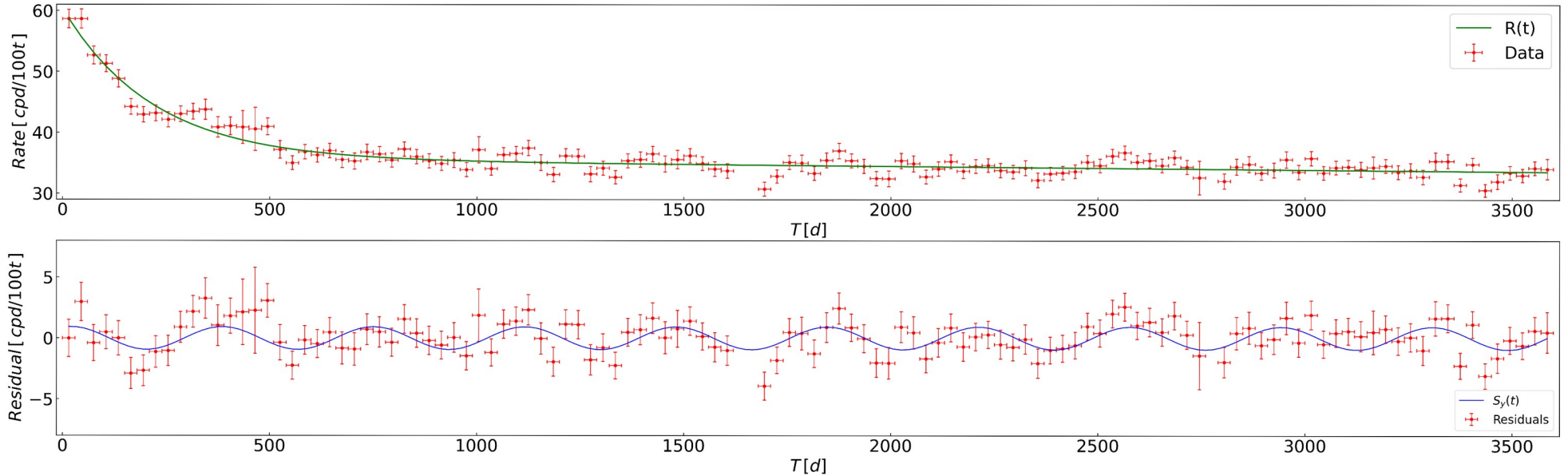


**Cosa c'entra BOREXINO
con l'astronomia?**

Variazione annuale del flusso di neutrini rivelati da BOREXINO



Prima misura dei parametri orbitali terrestri fatta con i neutrini!



Eccentricità = 0.0184 ± 0.0032 (0.0167)

→ **Presentazione di Riccardo Biondi**

Grazie!



Premio
Cocconi
2021 - EPS



Premio Pontecorvo
2015 G. Bellini



Premio Fermi
2017 G. Bellini



Ministerstwo
Edukacji i Nauki

Award of
Polish Science
Minister 2022



Bethe Prize
2023 F. Calaprice

