

Astrofisica, Astronomia e l'esperimento Borexino

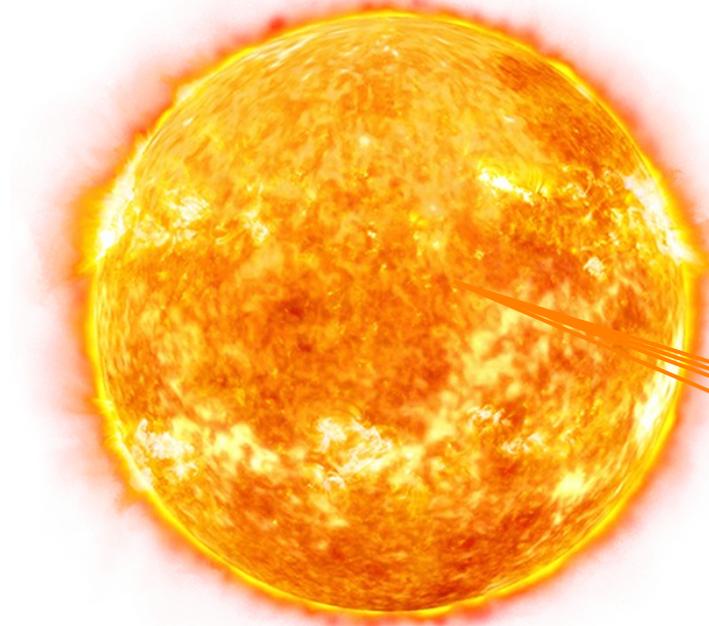


Nicola Rossi

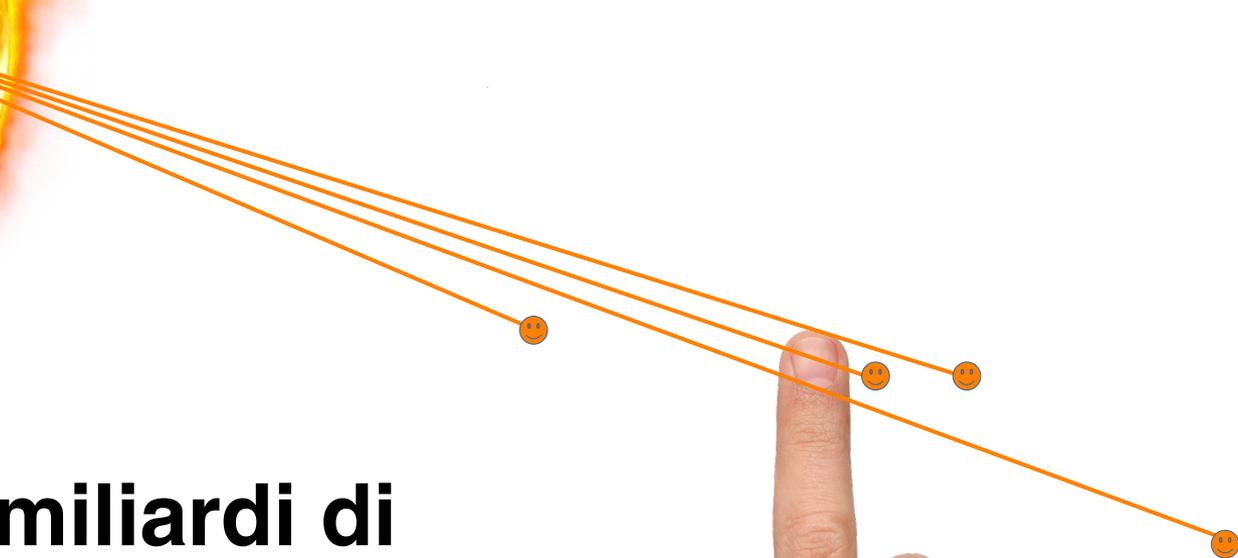
Laboratori Nazionali del Gran Sasso

NEUTRINI SOLARI E MASSIMI SISTEMI

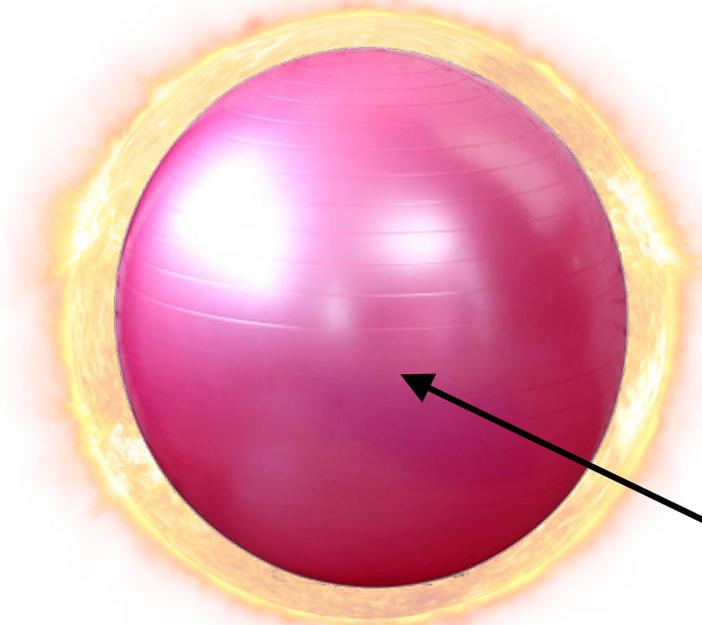
LNGS, 22 febbraio 2023



**Circa 60 miliardi di
neutrini per centimetro
quadro al secondo**



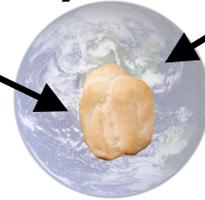
Quando è grande veramente il Sole?



Sole
(1 m)

100 m

Terra
(1 cm)



20 cm

Luna
(2.5 mm)



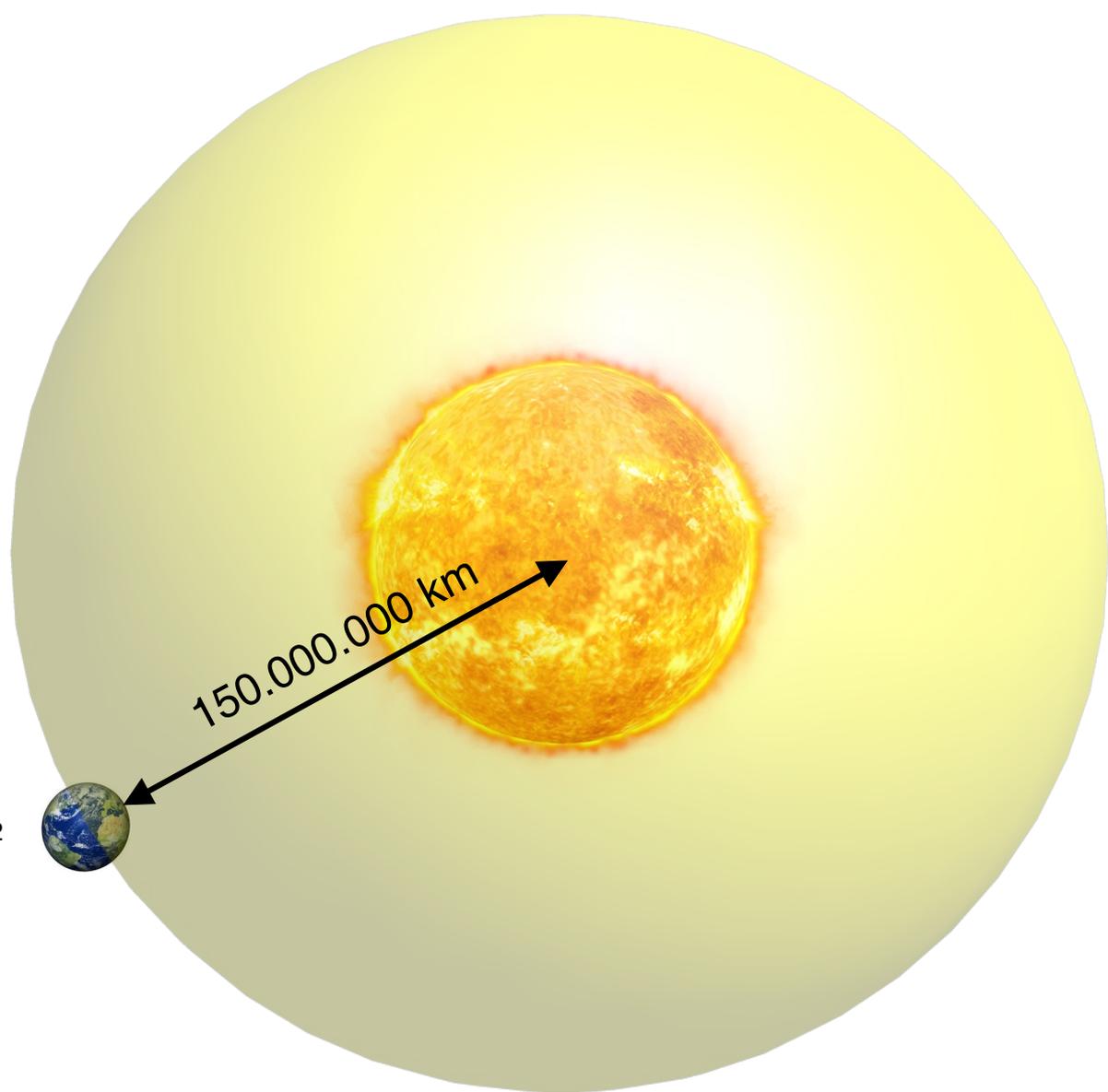
**Il sole irraggia sulla terra
circa 1.3 kW/m^2**



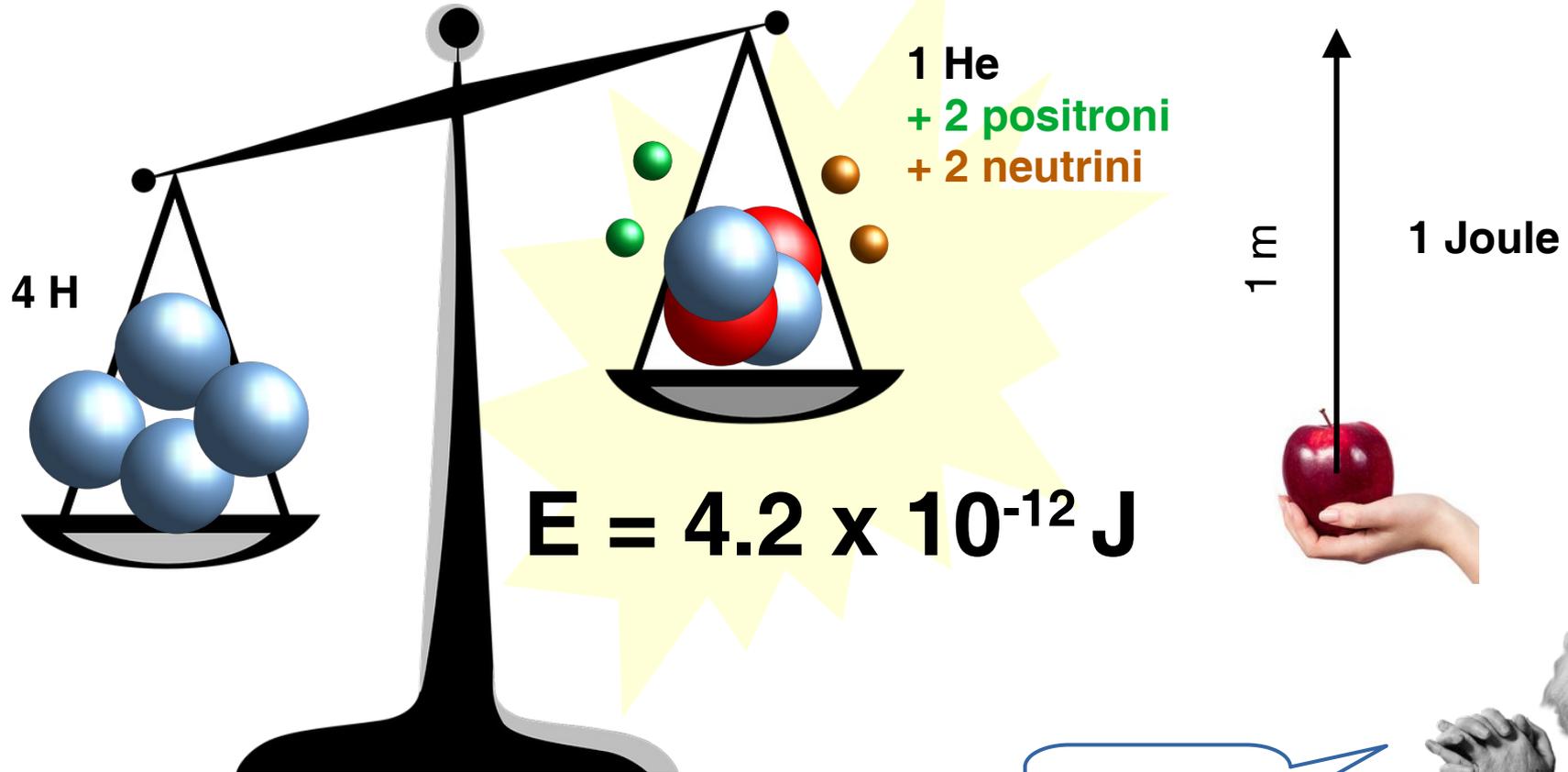
Potenza generata dal Sole:

$$1300 \text{ W} \times [\text{sfera}] = 4 \times 10^{26} \text{ W!}$$

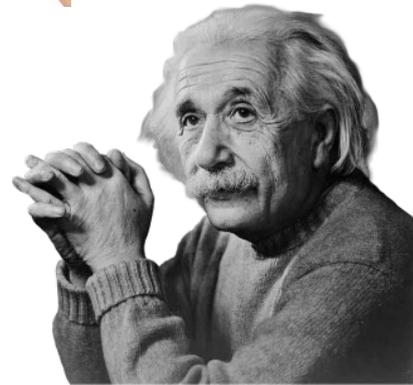
1.3 kW/m²



La fusione dell'idrogeno



$$E = mc^2!$$



A che velocità brucia il sole?

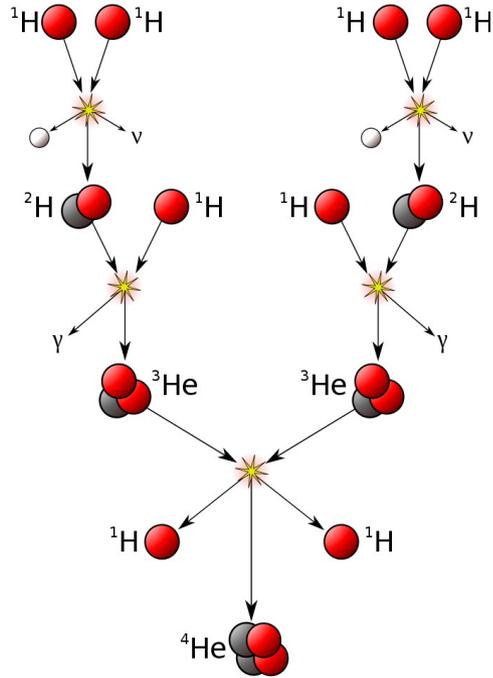
$$\frac{\text{potenza totale}}{\text{energia singola}} = \frac{4 \times 10^{26} \text{ J/s}}{4.2 \times 10^{-12} \text{ J}} \approx 10^{38} \frac{\text{fusioni}}{\text{secondo}}$$

600.000.000 tonnellate/sec di idrogeno

Quanti neutrini arrivano sulla terra?

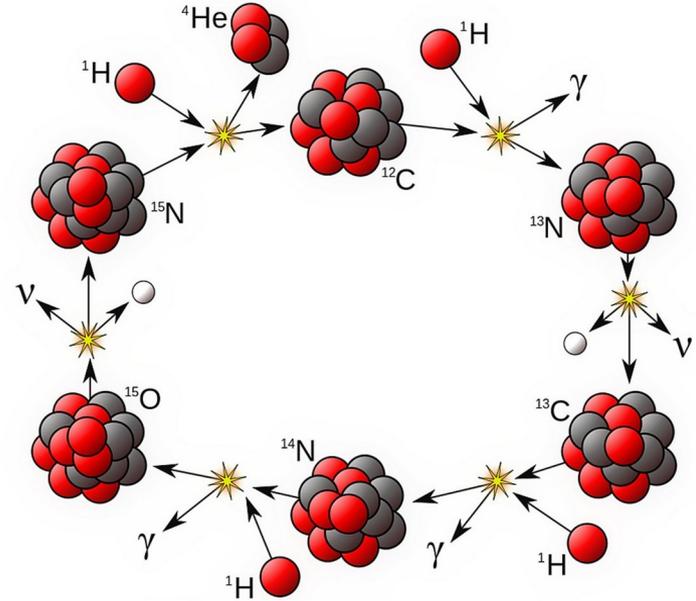
$$\frac{2 \times 10^{38}}{[\text{sfera}]} \approx 60 \times 10^9 \frac{\text{neutrini}}{\text{secondo} \times \text{cm}^2}$$

Catena protone-protone



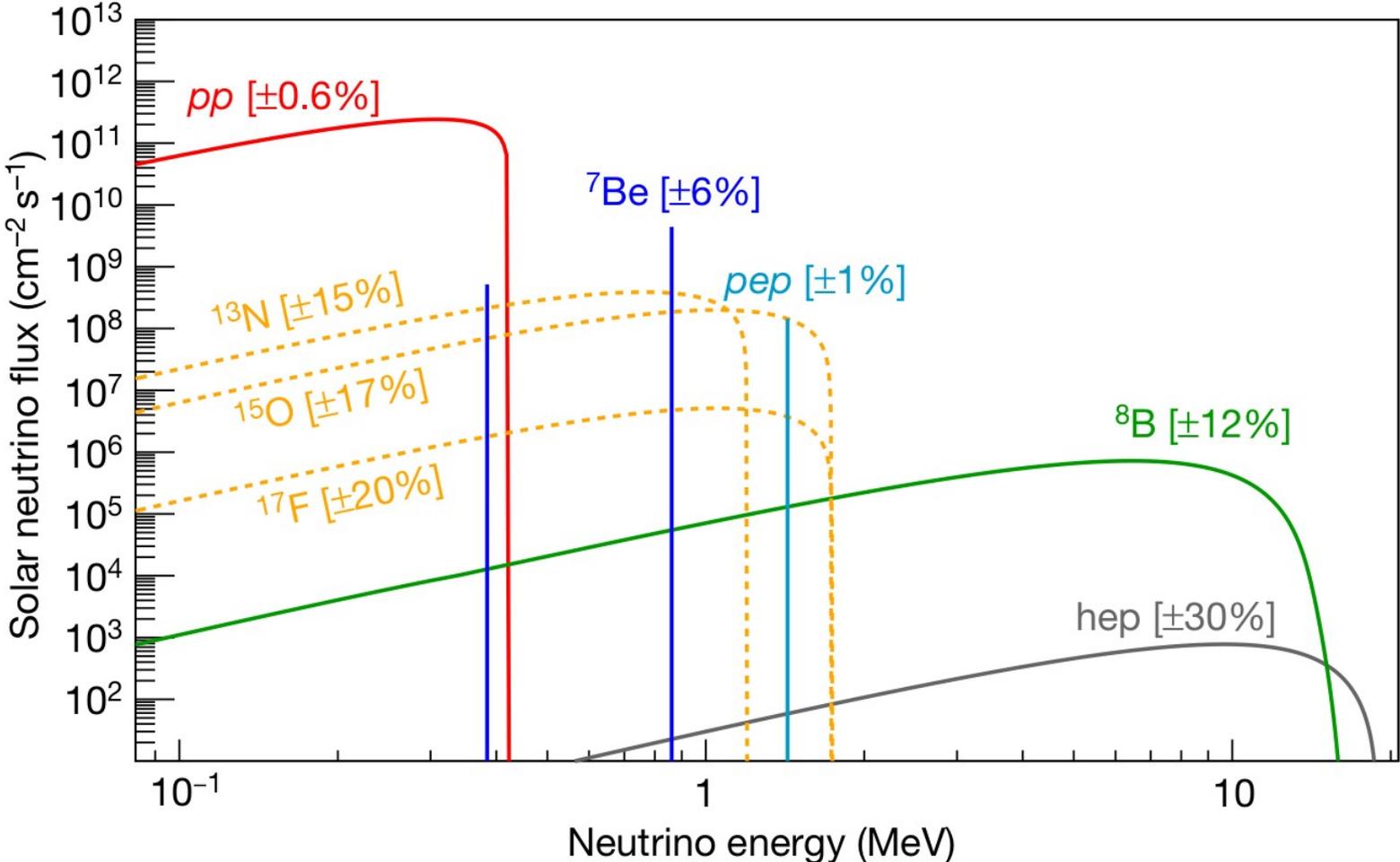
99%

Ciclo CNO

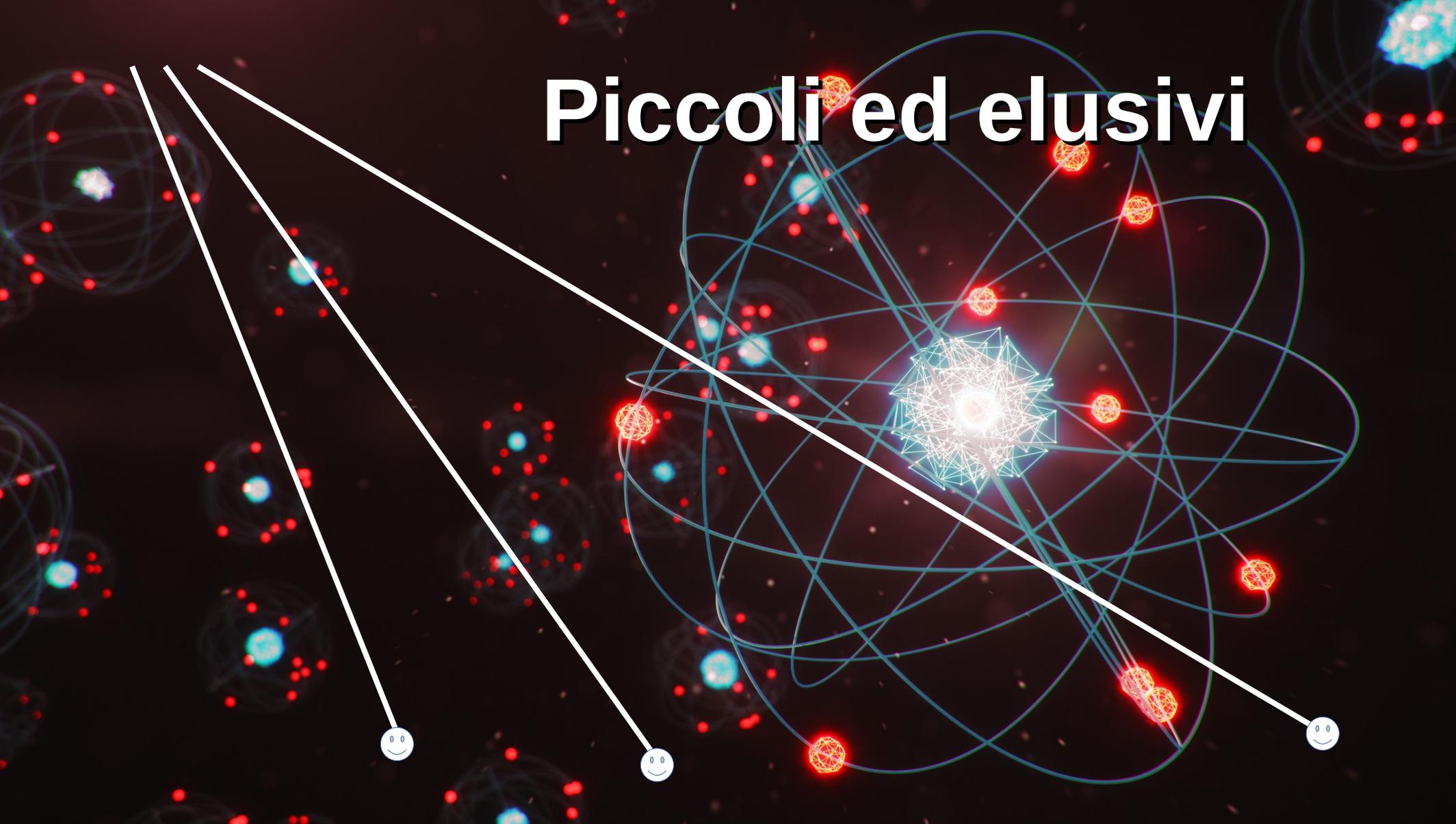


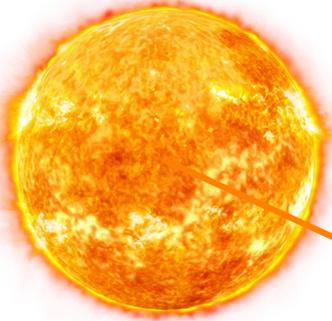
1%

Neutrini solari



Piccoli ed elusivi





**Che probabilità ha
un neutrino di
interagire in un
litro di acqua?**

$$60 \cdot 10^9 \frac{\nu}{\text{cm}^2} \times 3 \cdot 10^{26} \times 10^{-45} \text{cm}^2 \times 3 \cdot 10^7 \text{s} \approx 1 \frac{\nu}{\text{y}}$$

Flusso
di neutrini
sulla Terra

Quanti elettroni
in un litro
di acqua?

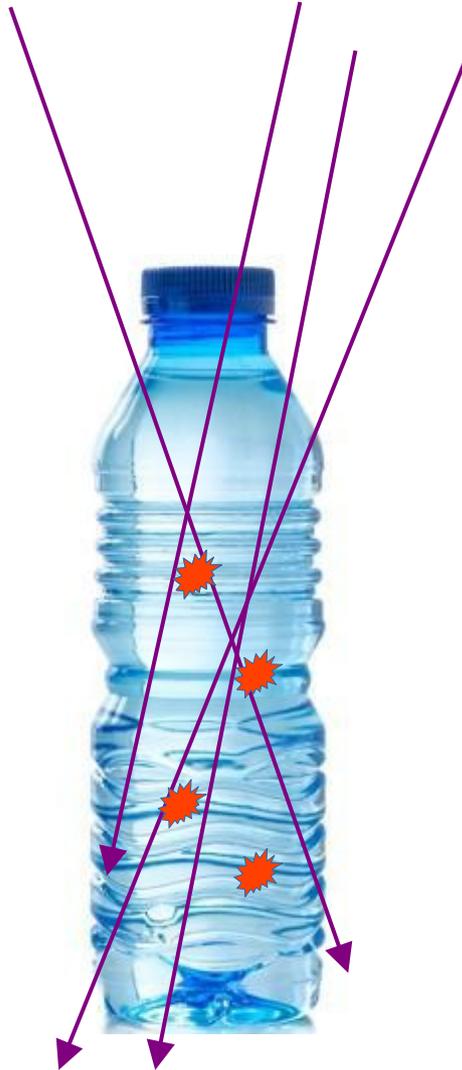
“dimensione”
dei neutrini

Secondi
In un anno

**Solo un neutrino
all'anno fermato
da un litro di
acqua!**

**Pioggia incessante
di raggi cosmici**

**Un litro di acqua
minerale contiene parti
per miliardo di
impurezze
(una disintegrazione al
secondo)**



**Un neutrino su
100.000.000
eventi di
fondo:
Cercare l'ago
nel pagliaio!**

Un rivelatore di neutrini solari deve essere

1. Protetto dalla roccia
2. Un miliardo di volte più puro di una buona acqua potabile
3. Di grandissimo volume

→ **BOREXINO**



**x100000
centinaia di neutrini
al giorno**

Anni '90



BOREXINO
Un'avventura
lunga pù di 30
anni

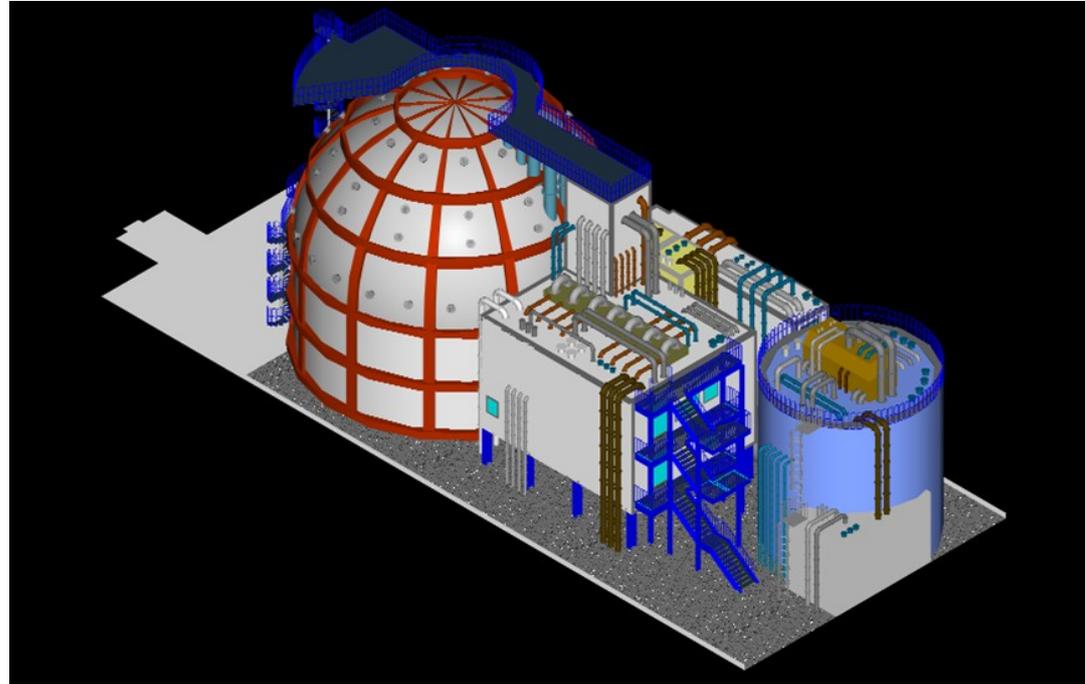
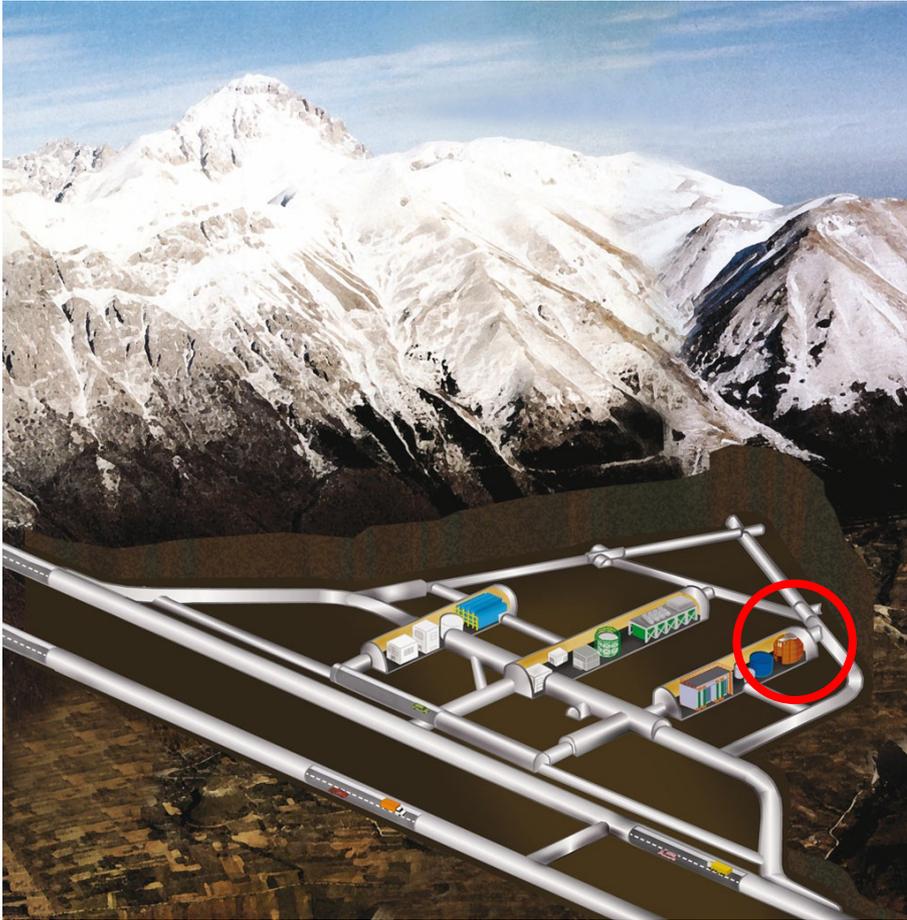
Gianpaolo Bellini
Frank Calaprice
Raju Ragavan

Sardegna 2015





Il rivelatore Borexino ai LNGS

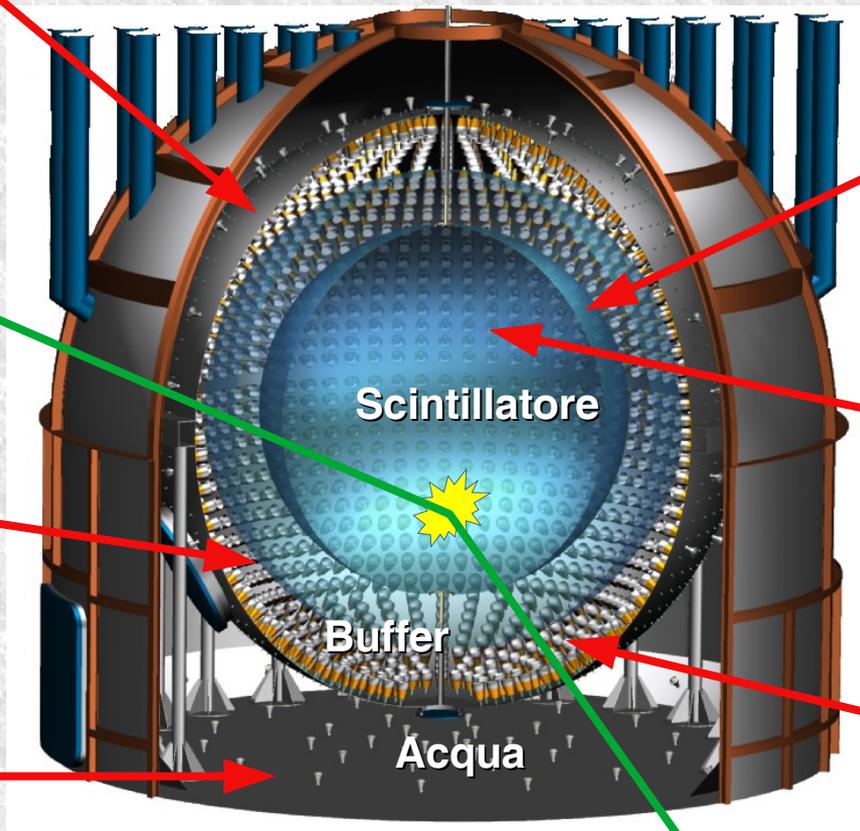


Sfera di acciaio



Pallone esterno di nylon

Acqua ultra-pura



Pallone interno di nylon

Scintillatore

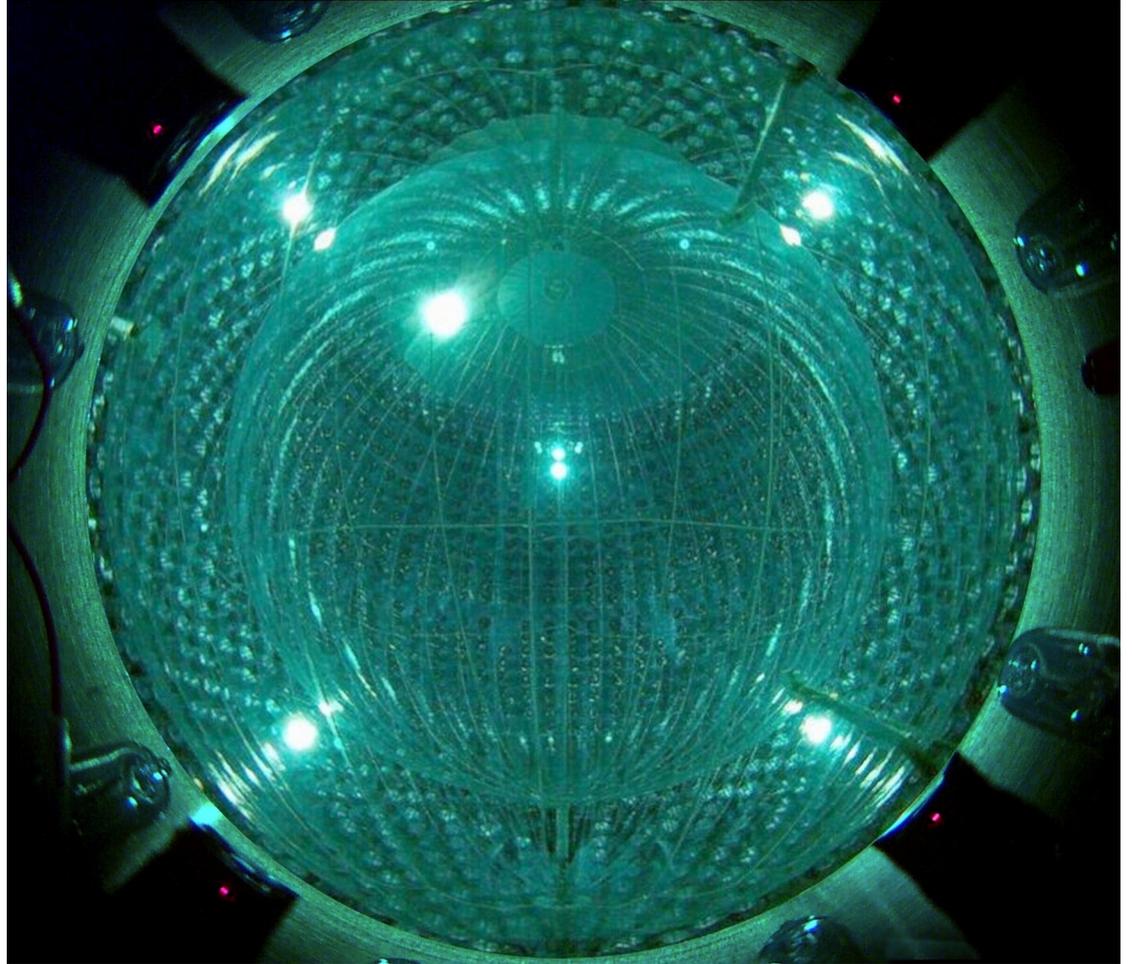
Tubi fotomoltiplicatori



Prima del riempimento

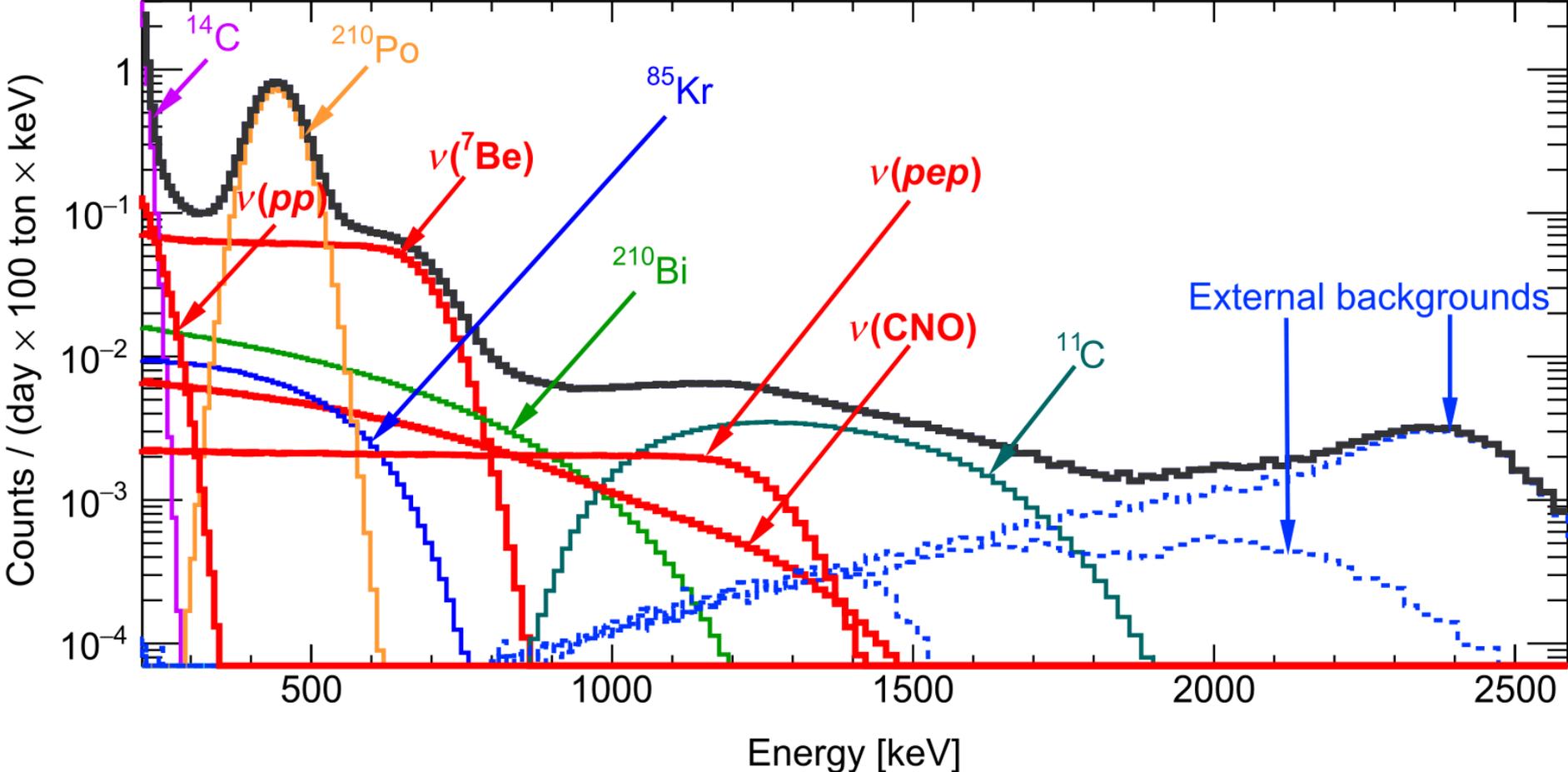


Nella Sala C



Dalla telecamera

Distribuzione energetica degli eventi rivelati da BOREXINO



Risultati di BOREXINO

nature

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Article | [Published: 24 October 2018](#)

Comprehensive measurement of *pp*-chain solar neutrinos

[The Borexino Collaboration](#)

[Nature](#) **562**, 505–510 (2018) | [Cite this article](#)

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Abstract

About 99 per cent of solar energy is produced through sequences of nuclear reactions that convert hydrogen into helium, starting from the fusion of two protons (the *pp* chain). The neutrinos emitted by five of these reactions represent a unique probe of the Sun's internal working and, at the same time, offer an intense natural neutrino beam for fundamental physics. Here we report a complete study of the *pp* chain. We measure the neutrino–electron elastic scattering rates for neutrinos produced by four reactions of the chain: the initial proton–proton fusion, the electron-capture decay of beryllium-7, the three-body proton–electron–proton (*pep*) fusion, here measured with the highest precision so far achieved, and the boron-8 beta decay, measured with the lowest energy threshold. We also set a limit on the

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Article | [Published: 25 November 2020](#)

Experimental evidence of neutrinos produced in the CNO fusion cycle in the Sun

[The Borexino Collaboration](#)

[Nature](#) **587**, 577–582 (2020) | [Cite this article](#)

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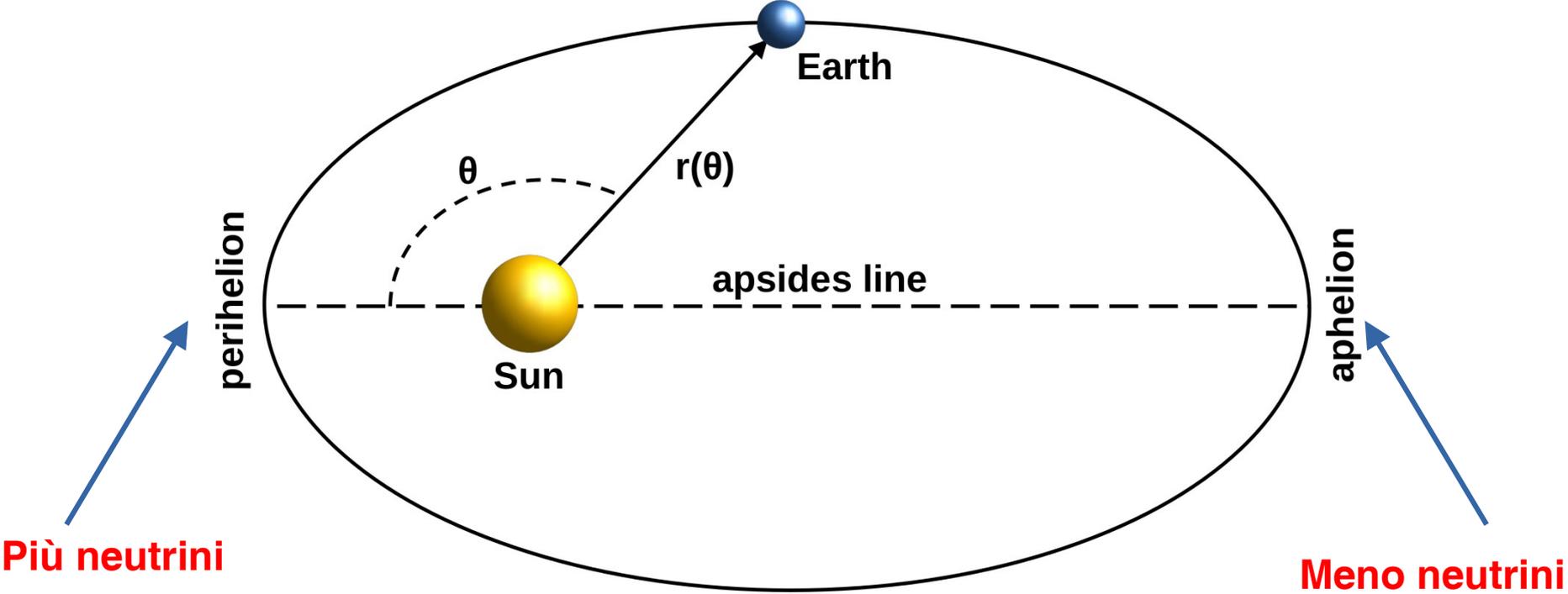
Abstract

For most of their existence, stars are fuelled by the fusion of hydrogen into helium. Fusion proceeds via two processes that are well understood theoretically: the proton–proton (*pp*) chain and the carbon–nitrogen–oxygen (CNO) cycle^{1,2}. Neutrinos that are emitted along such fusion processes in the solar core are the only direct probe of the deep interior of the Sun. A complete spectroscopic study of neutrinos from the *pp* chain, which produces about 99 per cent of the solar energy, has been performed previously³; however, there has been no reported experimental evidence of the CNO cycle. Here we report the direct observation, with a high statistical significance, of neutrinos produced in the CNO cycle in the Sun. This experimental evidence was obtained using the highly radiopure, large-volume, liquid-scintillator detector of Borexino, an experiment located at the underground Laboratori

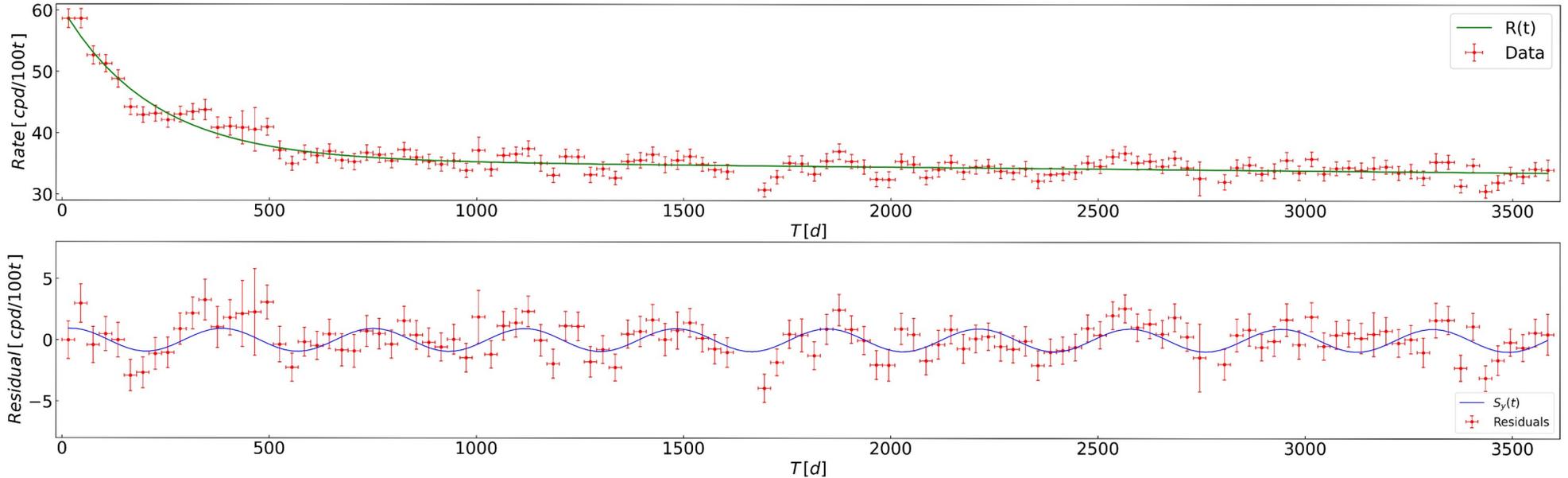


**Cosa c'entra BOREXINO
con l'astronomia?**

Variazione annuale del flusso di neutrini rivelati da BOREXINO



Prima misura dei parametri orbitali terrestri fatta con i neutrini!



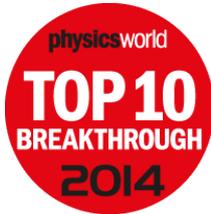
Eccentricità = 0.0184 ± 0.0032 (0.0167)

→ **Presentazione di Riccardo Biondi**

Grazie!



Premio
Cocconi
2021 - EPS



Premio Pontecorvo
2015 G. Bellini



Premio Fermi
2017 G. Bellini



Ministerstwo
Edukacji i Nauki

Award of
Polish Science
Minister 2022



Bethe Prize
2023 F. Calaprice

