

# Searching for Neutrinoless Double-Beta Decay with KamLAND-Zen and LEGEND

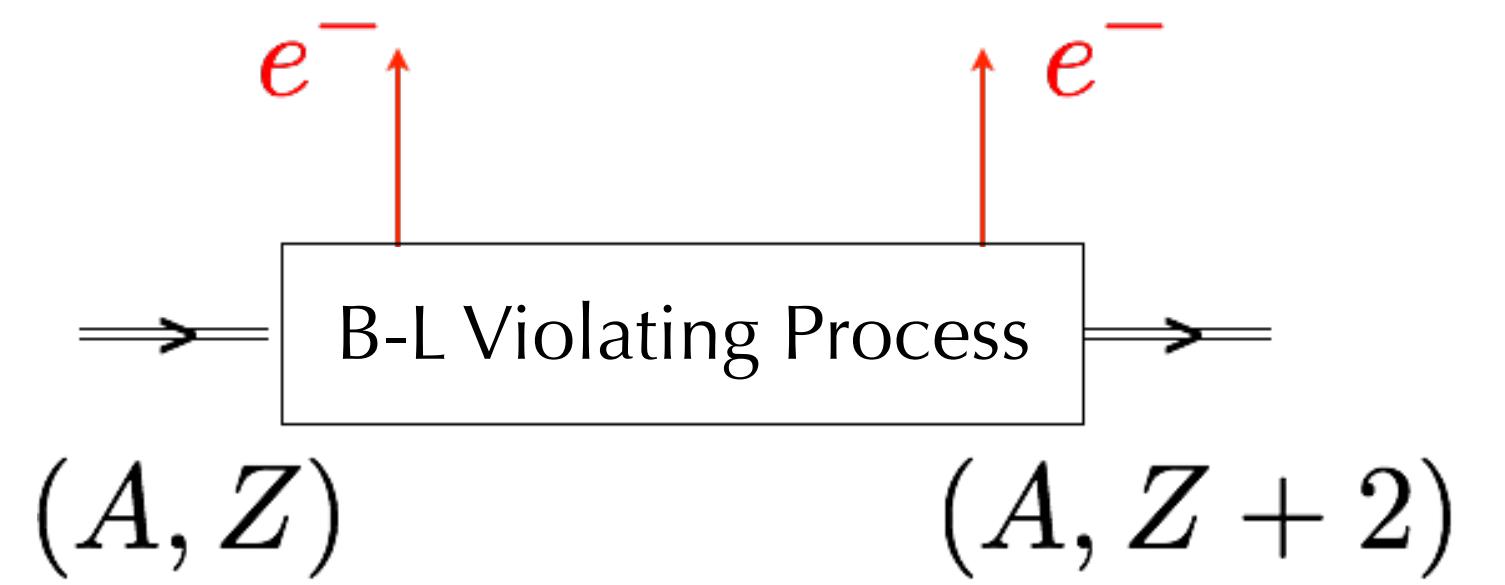
Jason Detwiler, UW  
MAYORANA Workshop  
July 12, 2023

# Outline

- Introduction: strategies for detecting  $0\nu\beta\beta$  decay
- High-exposure: KamLAND-Zen
- Low-background: LEGEND

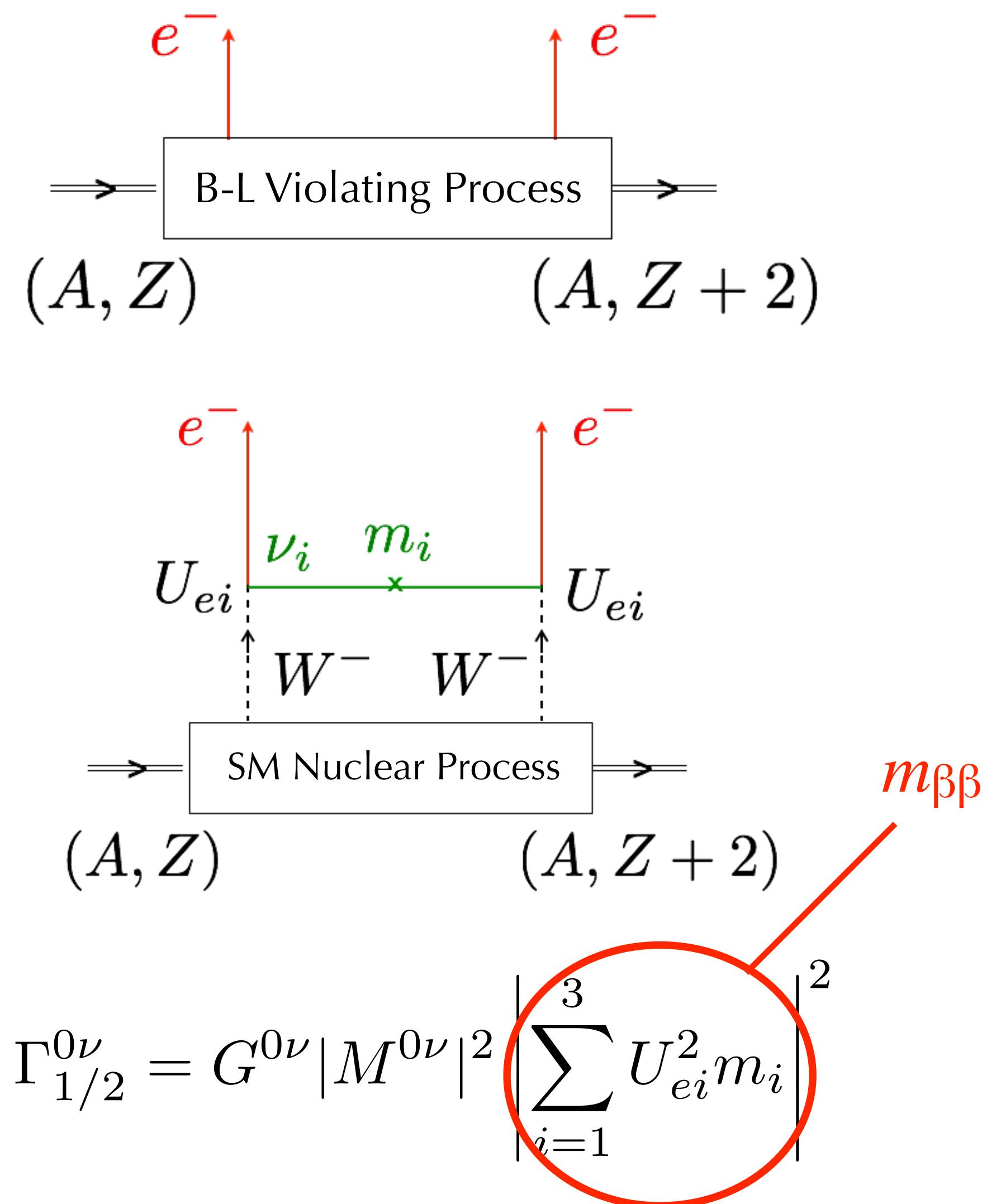
# $0\nu\beta\beta$ Decay

- The creation of matter without antimatter, which has never been observed (since the Big Bang)
- Violates not just  $L$  but  $B-L$ : the last accidentally conserved quantity in the SM
- Generically predicted by GUTs that also explain the cosmic matter asymmetry and the smallness of the neutrino mass



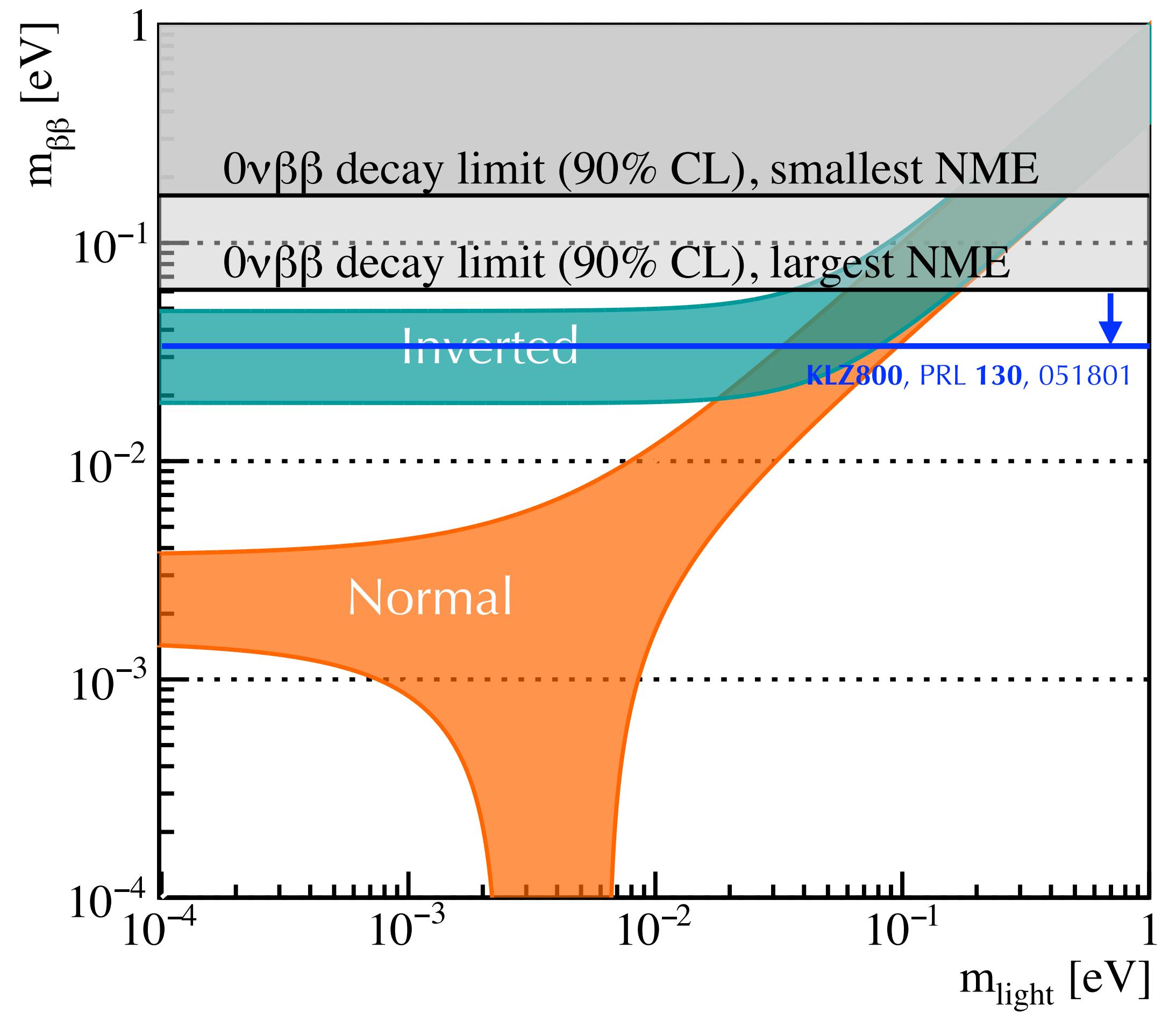
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- Lowest-order LNV operator is at dim 5: a Majorana neutrino mass



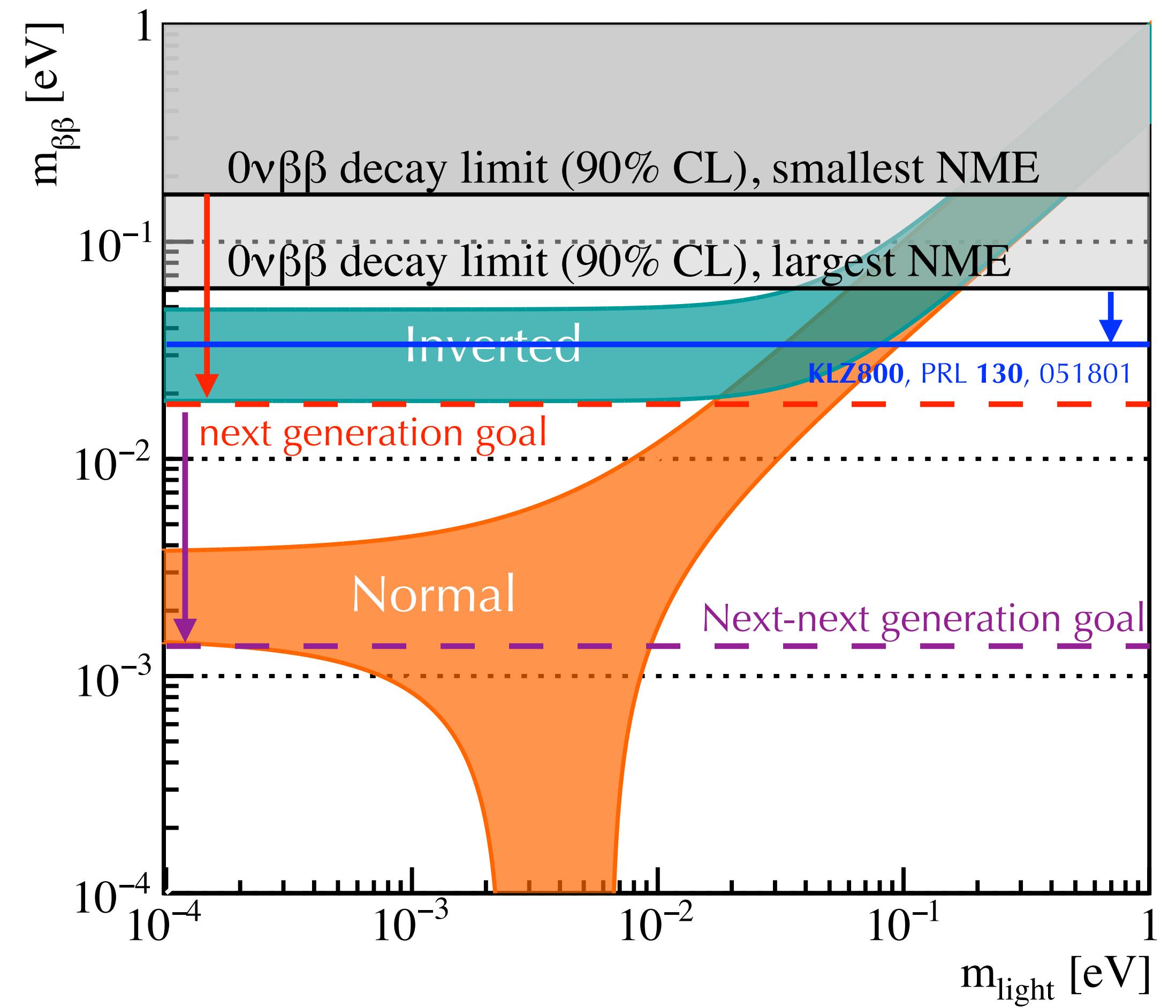
# Light Neutrino Exchange

- Light neutrino exchange is “natural” and “minimalistic”, and sets clear experimental goal posts:
  - IO:  $T_{1/2} \lesssim 10^{28}$  years ( $10^{18}$  times the age of the universe)
  - NO:  $T_{1/2} \lesssim 10^{30}$  years, modulo cancellations, flavor symmetries, etc.
- Other mechanisms are possible: the whole region is “open”!
  - Dim 7 (9) LNV is probed at the PeV (TeV) scale
  - Sterile  $\nu$  make the IO/NO regions cover the entire plane



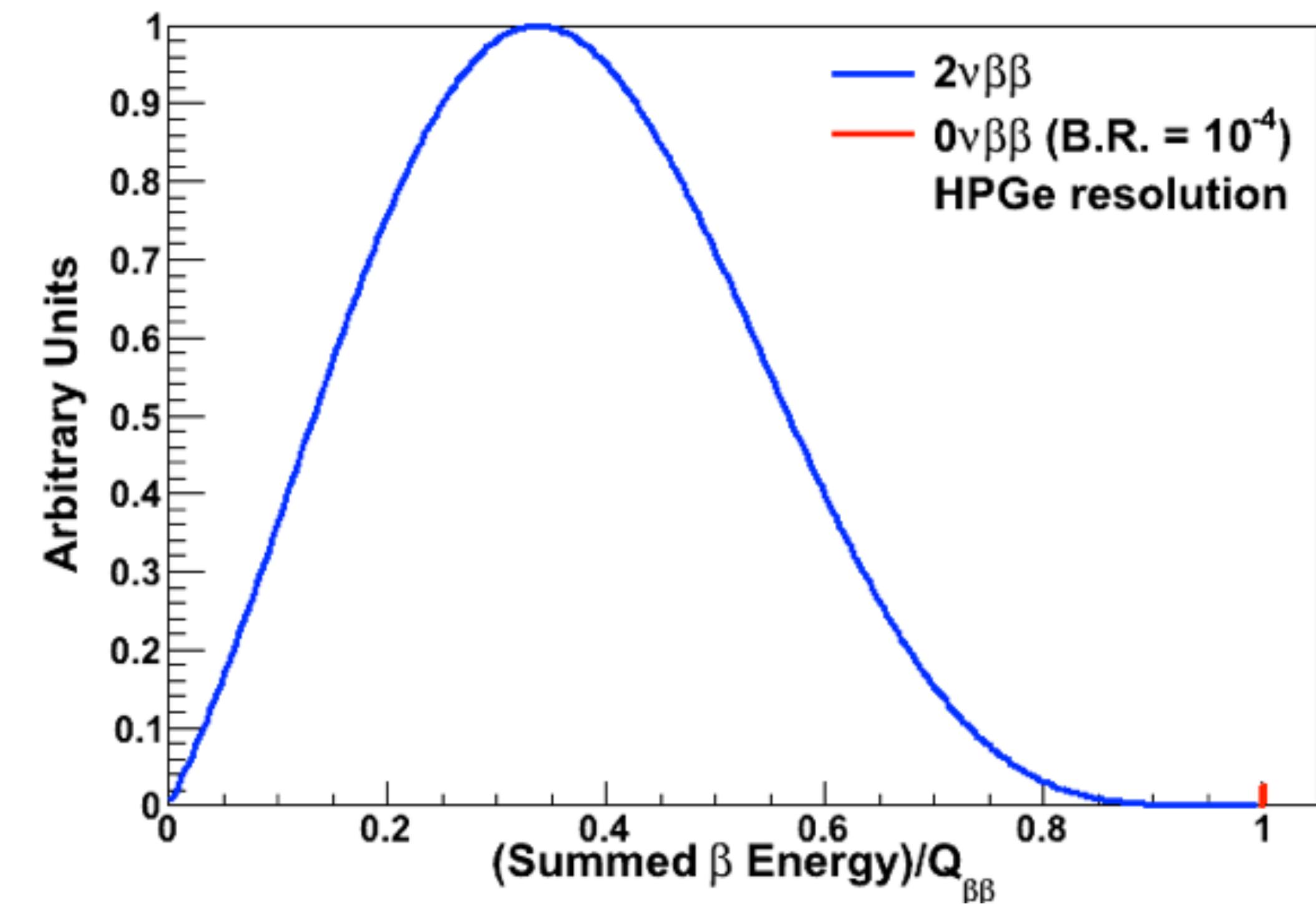
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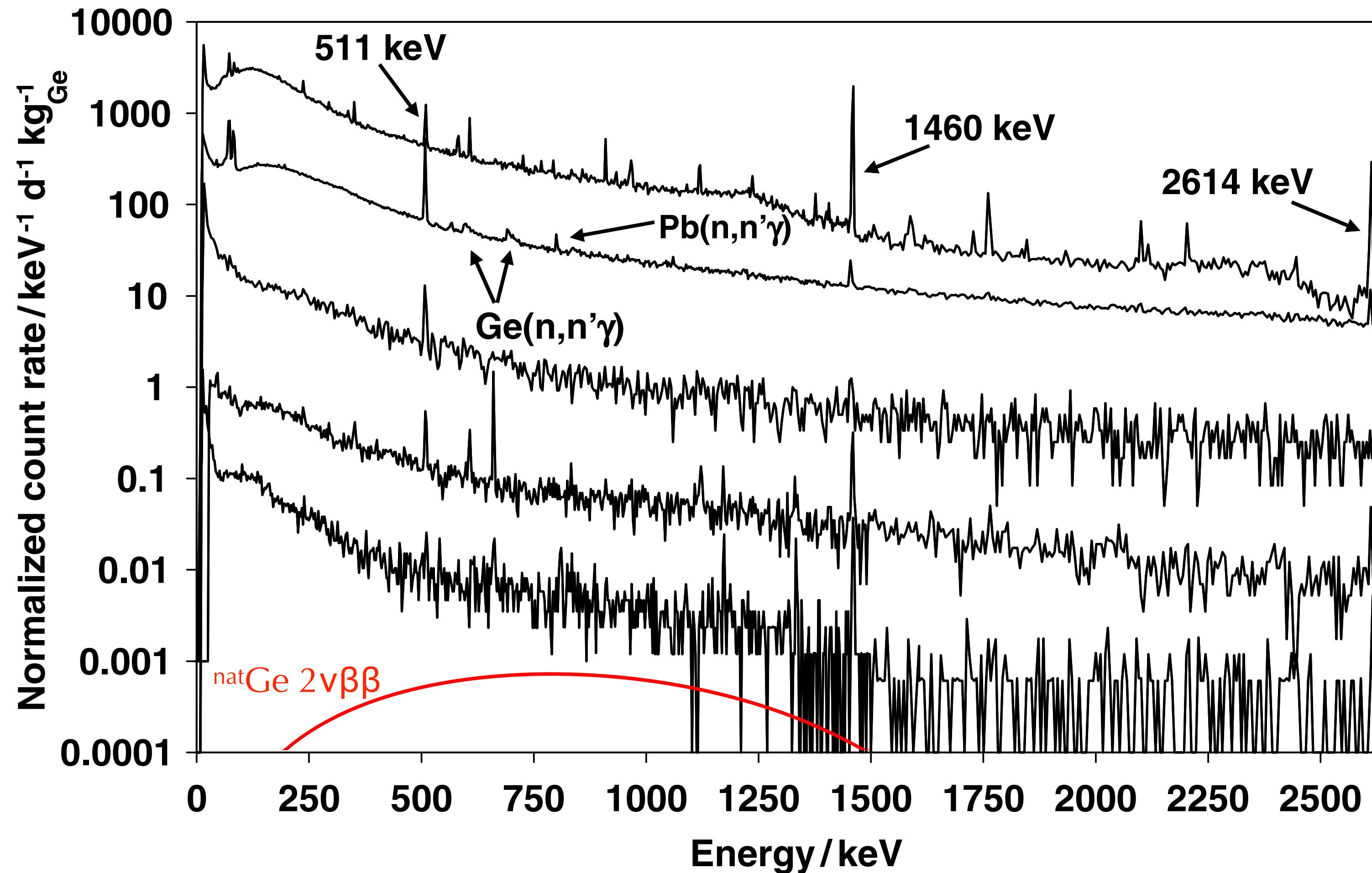


# Neutrinoless Double-Beta Decay

- Must measure summed electron kinetic energy to distinguish  $0\nu\beta\beta$  from the Standard-Model  $2\nu\beta\beta$  process: search for a peak at  $Q_{\beta\beta}$
- The peak in the plot exceeds current limits by  $>1$  order of magnitude



# The Background Problem



- Typical surface detector (HPGe): natural radioactivity dominates
- Low-bg surface detector: muon and primary n cosmic rays
- Low-bg detector, 125 mwe: muons
- Low-bg detector, 500 mwe: muons + natural radioactivity
- Ultra-low-bg detector, 3400 mwe: natural radioactivity

Need an underground detector made of pure materials, and typically need enrichment.

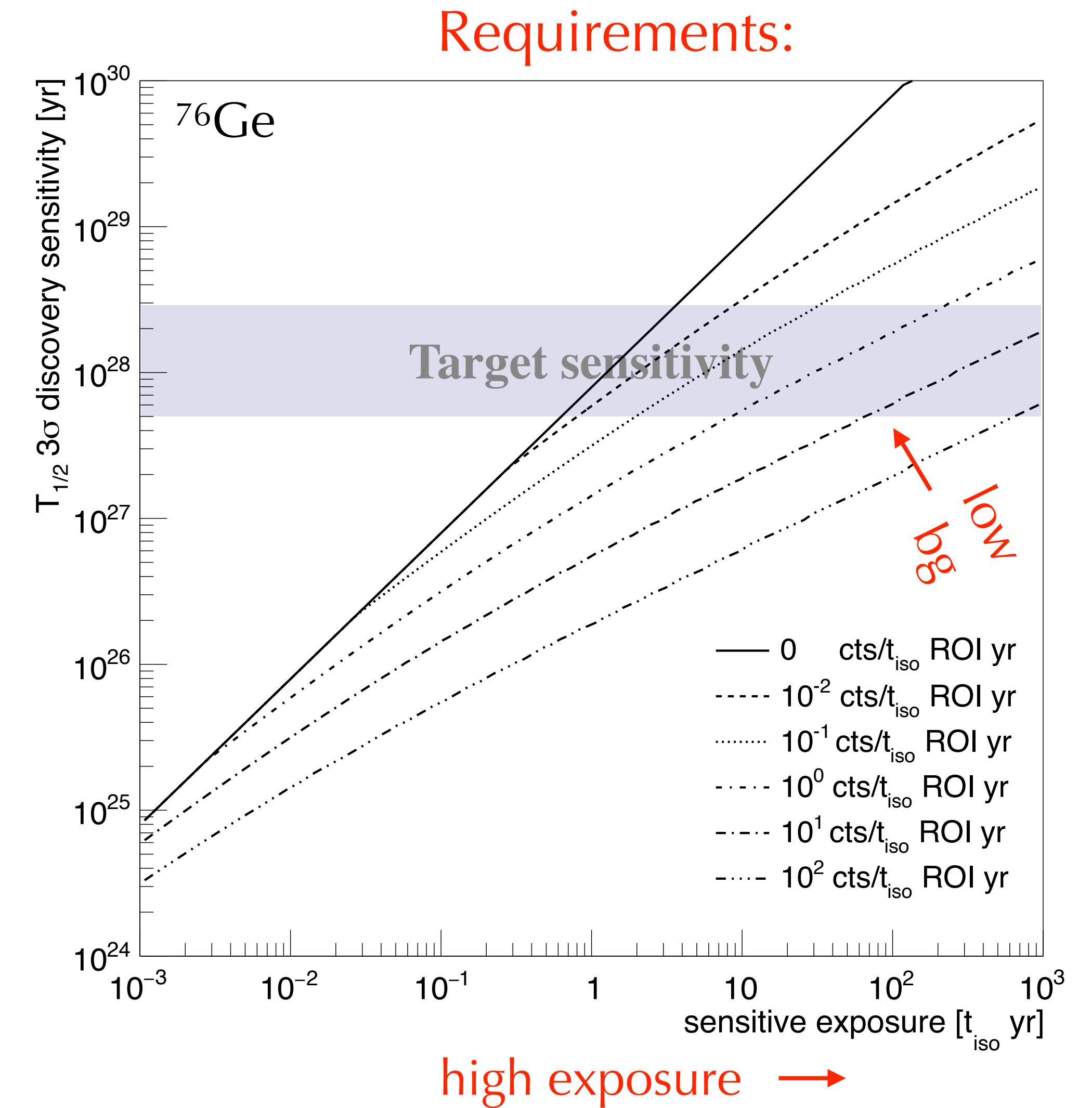
# Experimental Sensitivity

- Effectively a Poisson counting experiment near  $Q_{\beta\beta}$
- Relevant parameters: sensitive exposure and sensitive background

$$\mathcal{E} = \epsilon m_{iso}^{FV} t \quad \mathcal{B} = N_{bg}/\mathcal{E}$$

- Discovery sensitivity: the minimum signal strength for which an experiment has a  $\geq 50\%$  chance to observe a signal above background with significance  $\geq 3\sigma$ :

$$T_{1/2}^{3\sigma} = \ln 2 \frac{N_A \mathcal{E}}{m_a S_{3\sigma}(\mathcal{B}\mathcal{E})}$$



# Tutorial for observing a $10^{28}$ yr half-life

- Get O(tons) of  $\beta\beta$  isotope 😎
- Instrument it so that it can detect  $0\nu\beta\beta$  decay with high efficiency 🧙
- Eliminate ~all random events that can mimic  $0\nu\beta\beta$  🤯
- Wait ~10 years 😐

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- Eliminate ~all random events that can mimic  $0\nu\beta\beta$  😬
- Wait ~10 years 😐
- However: if  $0\nu\beta\beta$  decay is just beyond current limits, these same ton-scale experiments will observe O(100) events!

# Experimental Techniques

- Bolometers (CUORE/CUPID, AMoRE, CANDLES IV)

- Measure  $E$  ( $\sigma \sim 0.1\text{-}0.3\%$ ) from phonons; granularity gives position info
- Instrumenting with photon detectors for background rejection



CANDLES

- External trackers (NEMO3, SuperNEMO)

- Trackers + calorimeters, measure  $E$  ( $\sigma \sim 3\text{-}10\%$ ) + tracks / positions + PID



NEXT-100

- Scintillators (KamLAND-Zen, SNO+, CANDLES-III, Theia, ZICOS)

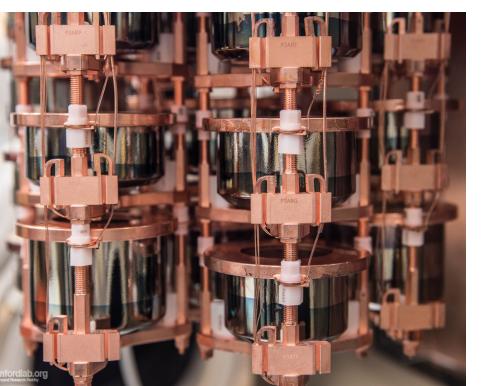
- Measure  $E$  ( $\sigma \sim 3\text{-}10\%$ ) + position from scintillation light; some PID



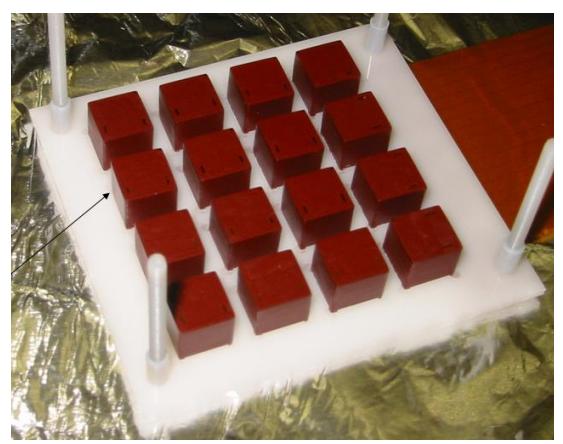
EXO-200

- Semiconductors (COBRA, MAJORANA, GERDA, LEGEND)

- Measure  $E$  ( $\sigma \sim 0.05\text{-}0.3\%$ ) from ionization; some tracking / position sensitivity



MAJORANA



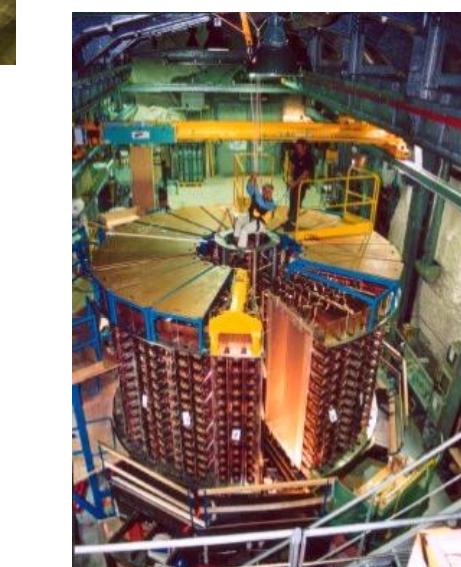
COBRA

- TPCs (nEXO, NEXT, PandaX, AXEL, NvDEx, DARWIN, LZ)

- Collect scintillation + ionization: measure  $E$  ( $\sigma \sim 0.4\text{-}3\%$ ) + tracks / position + PID



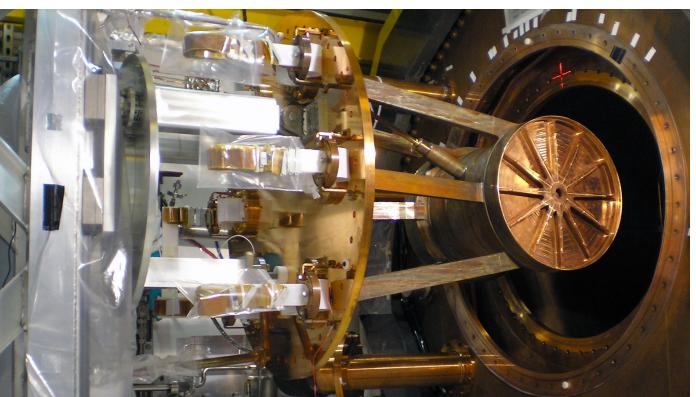
SuperNEMO



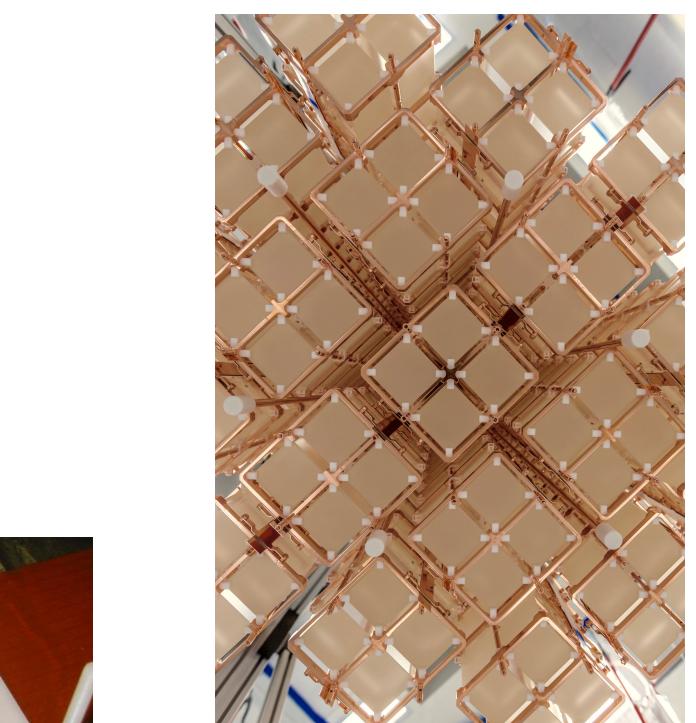
NEMO3



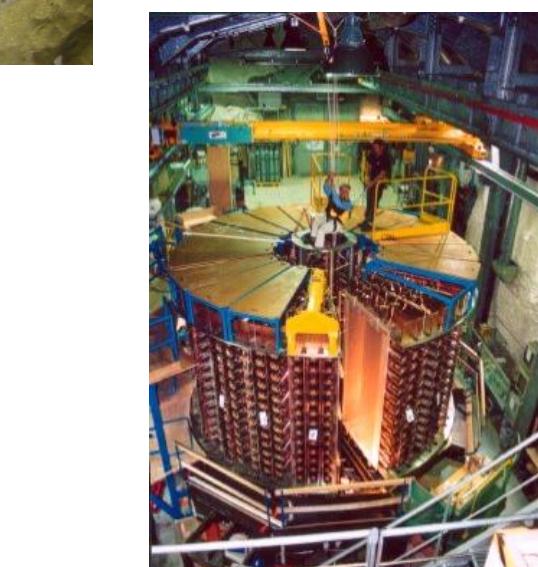
KamLAND-Zen



CANDLES



EXO-200



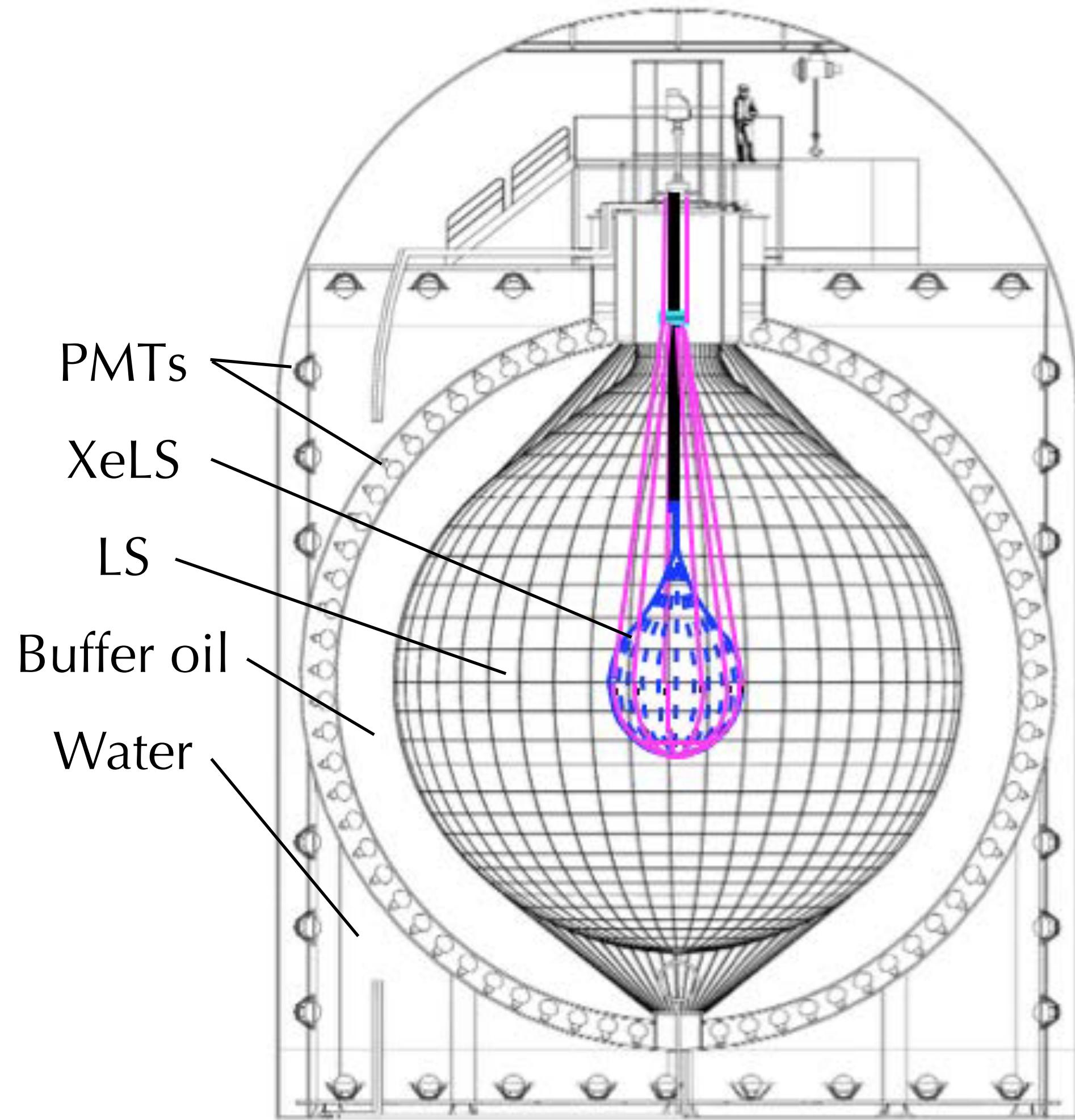
KamLAND-Zen

# Outline

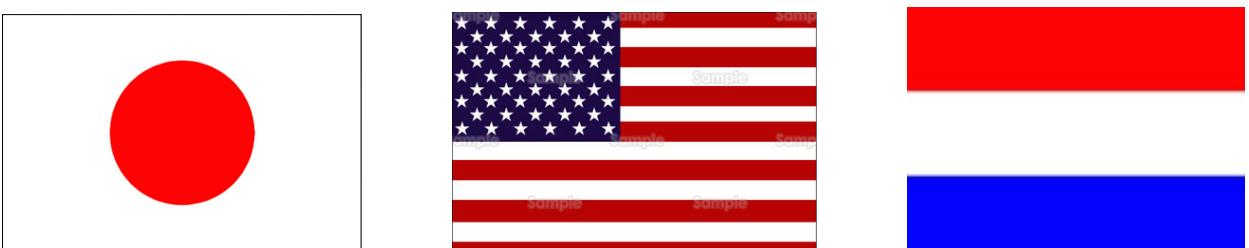
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# KamLAND-Zen



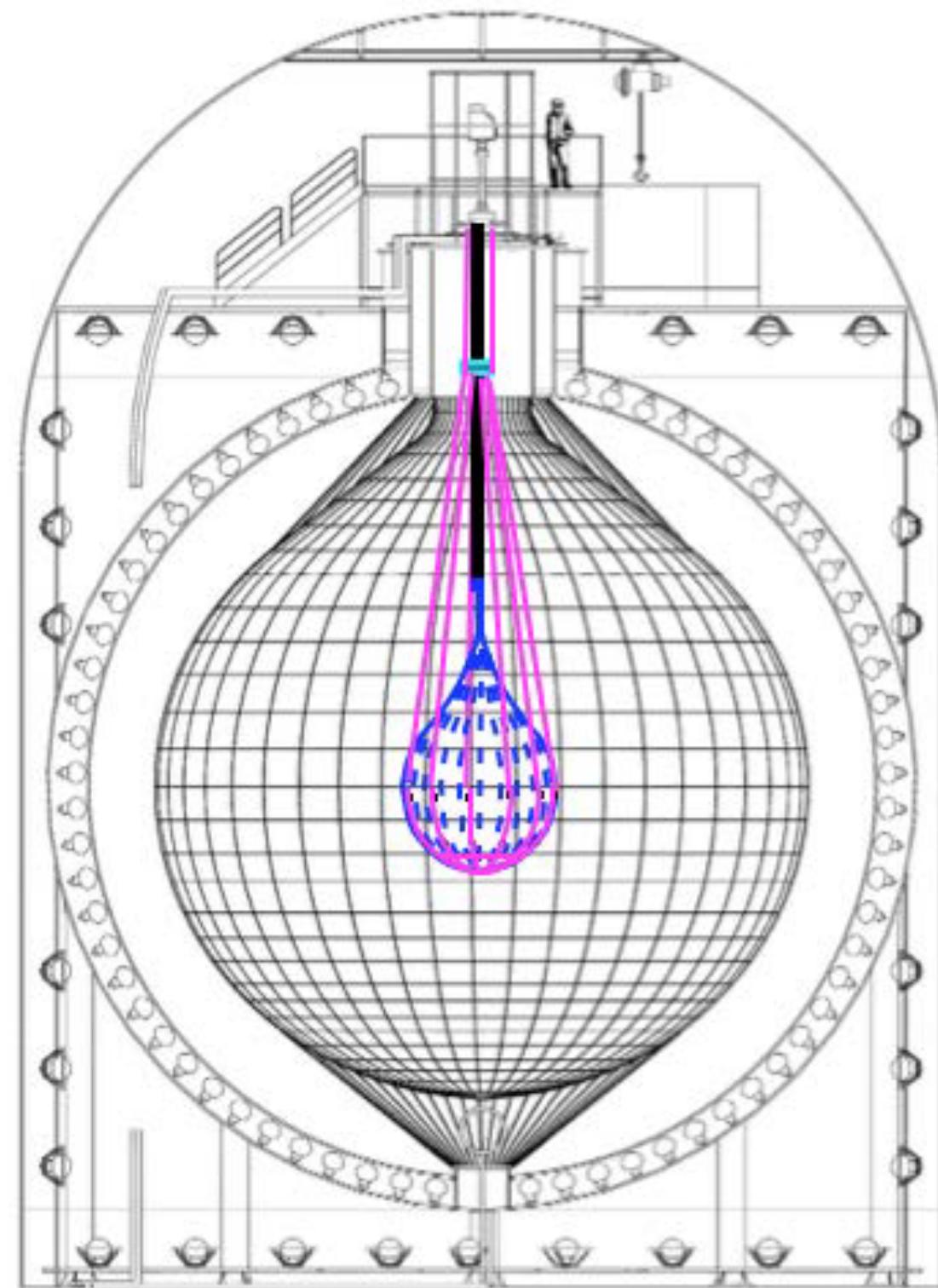
>50 researchers from Japan, the US, and the Netherlands



Measures light:  $E$ ,  $r$ , and  $t$  (+  $\mu$  veto)

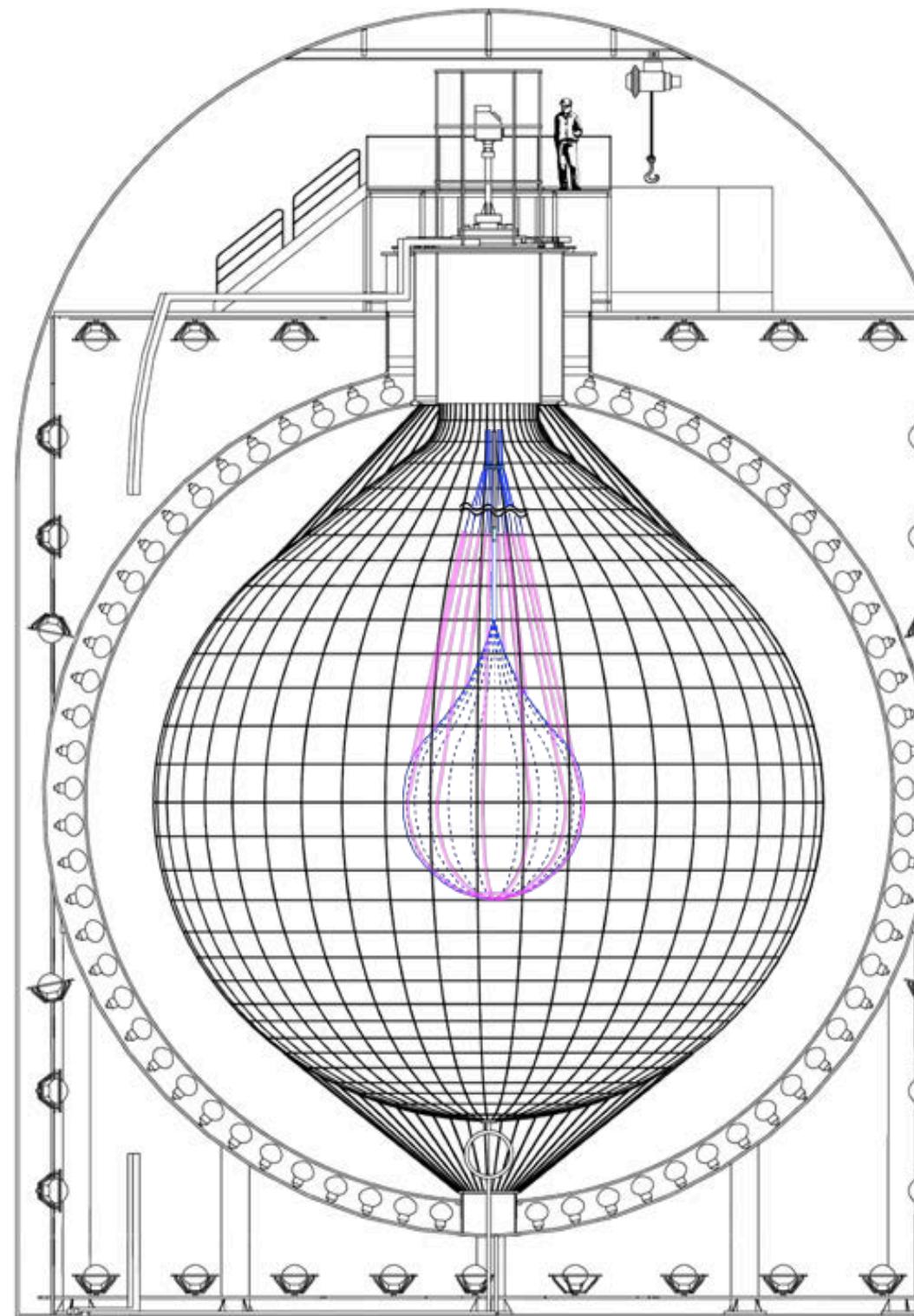
# KamLAND-Zen Timeline

2011-2015: KLZ-400



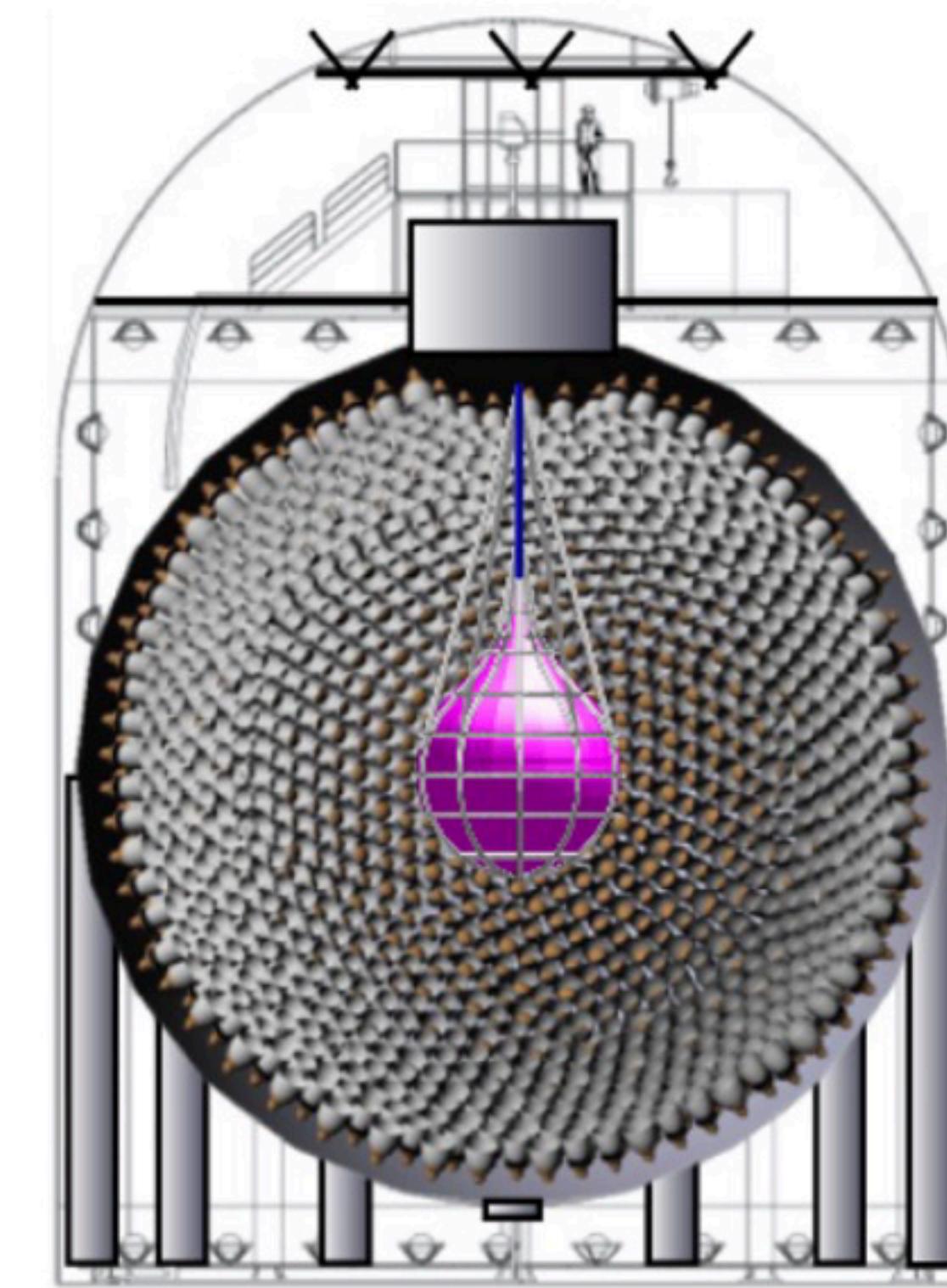
380 kg Xe  
 $T_{1/2} > 1.07 \times 10^{26}$  yr  
PRL 117, 082503 (2016)

2019-present: KLZ-800



750 kg Xe  
Cleaner, larger balloon

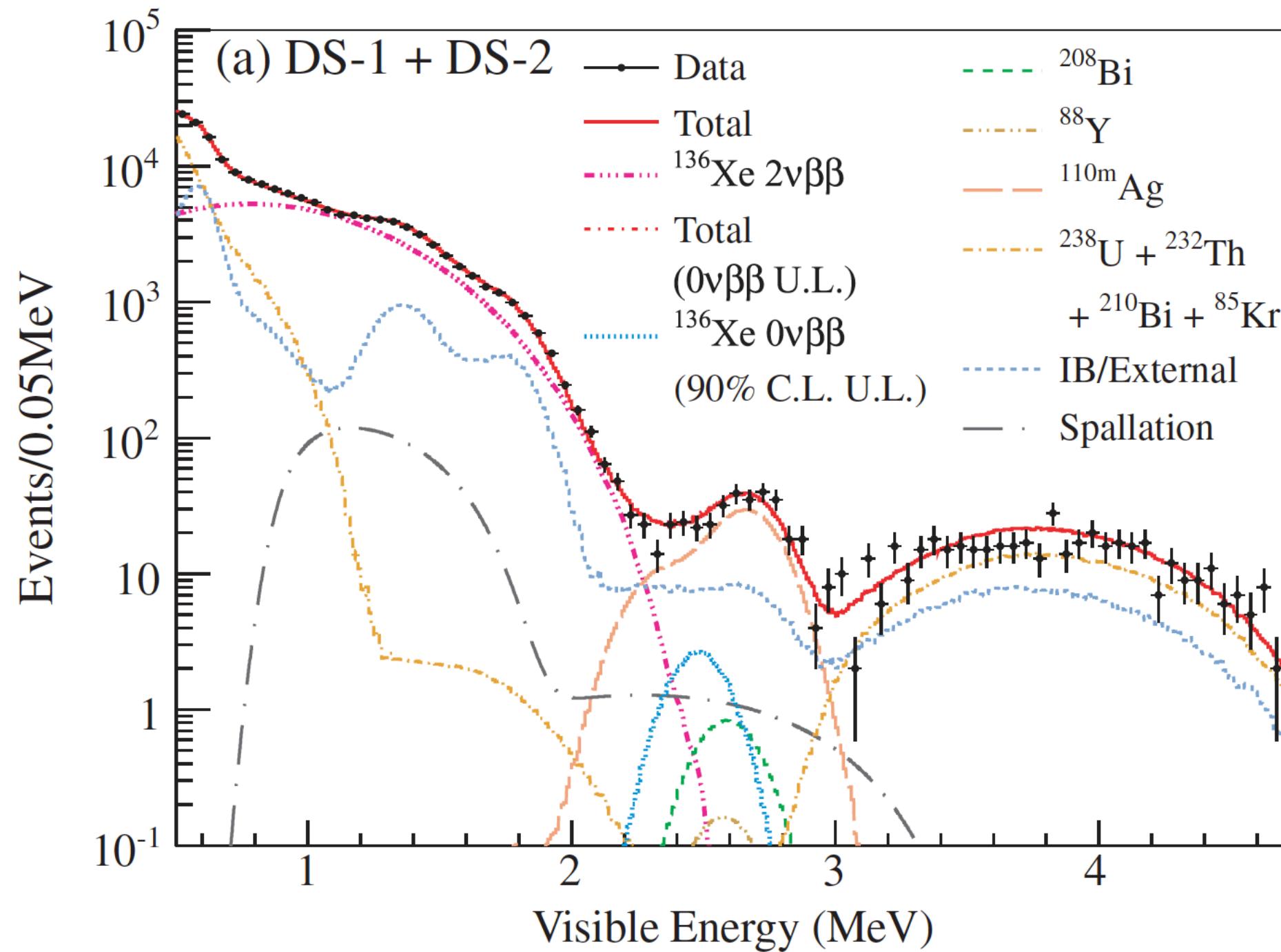
Future: KamLAND2-Zen



1 t Xe  
Improved light collection  
and background rejection

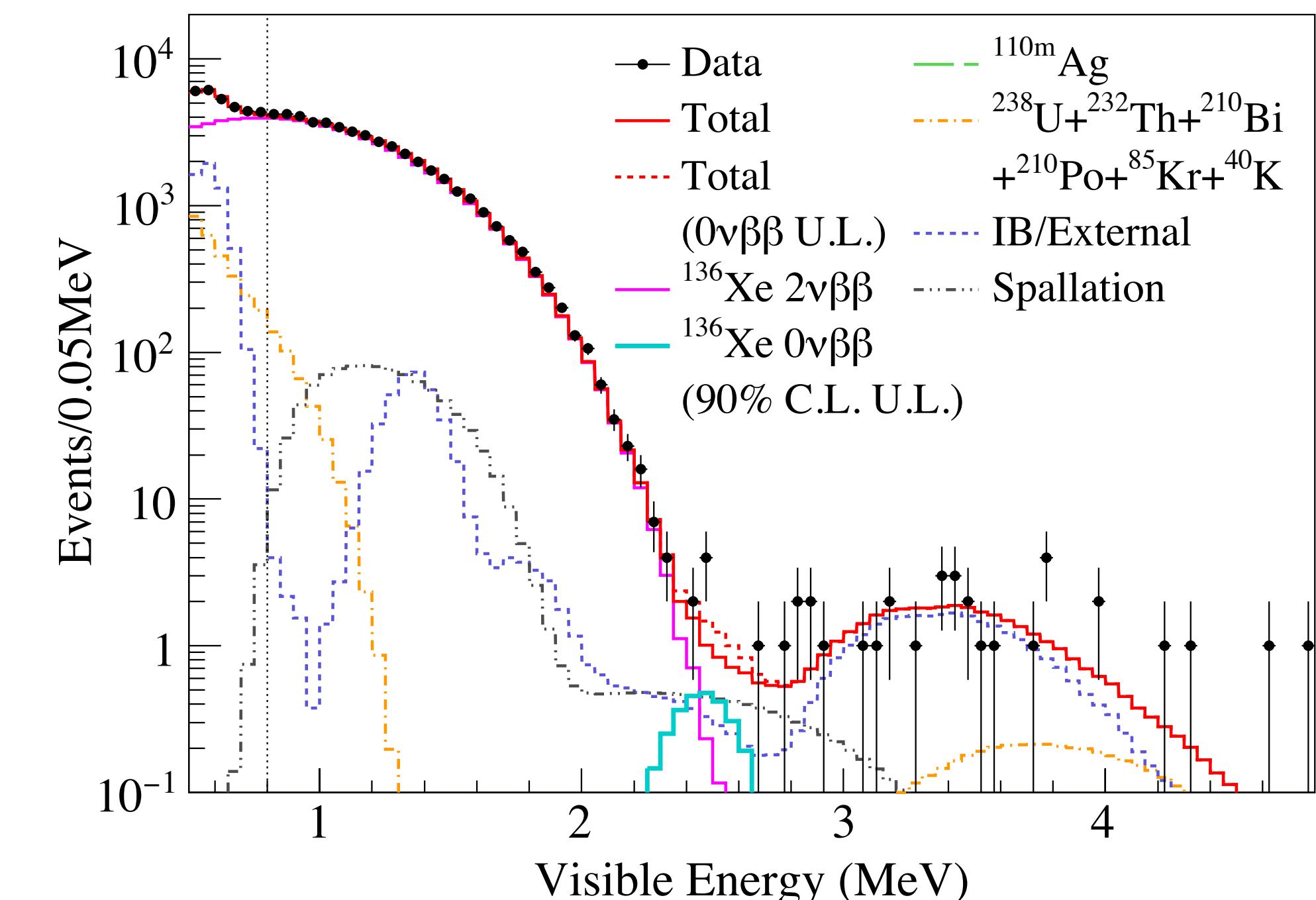
# KamLAND-Zen 400 Backgrounds

Phase I spectrum: Fukushima Fallout

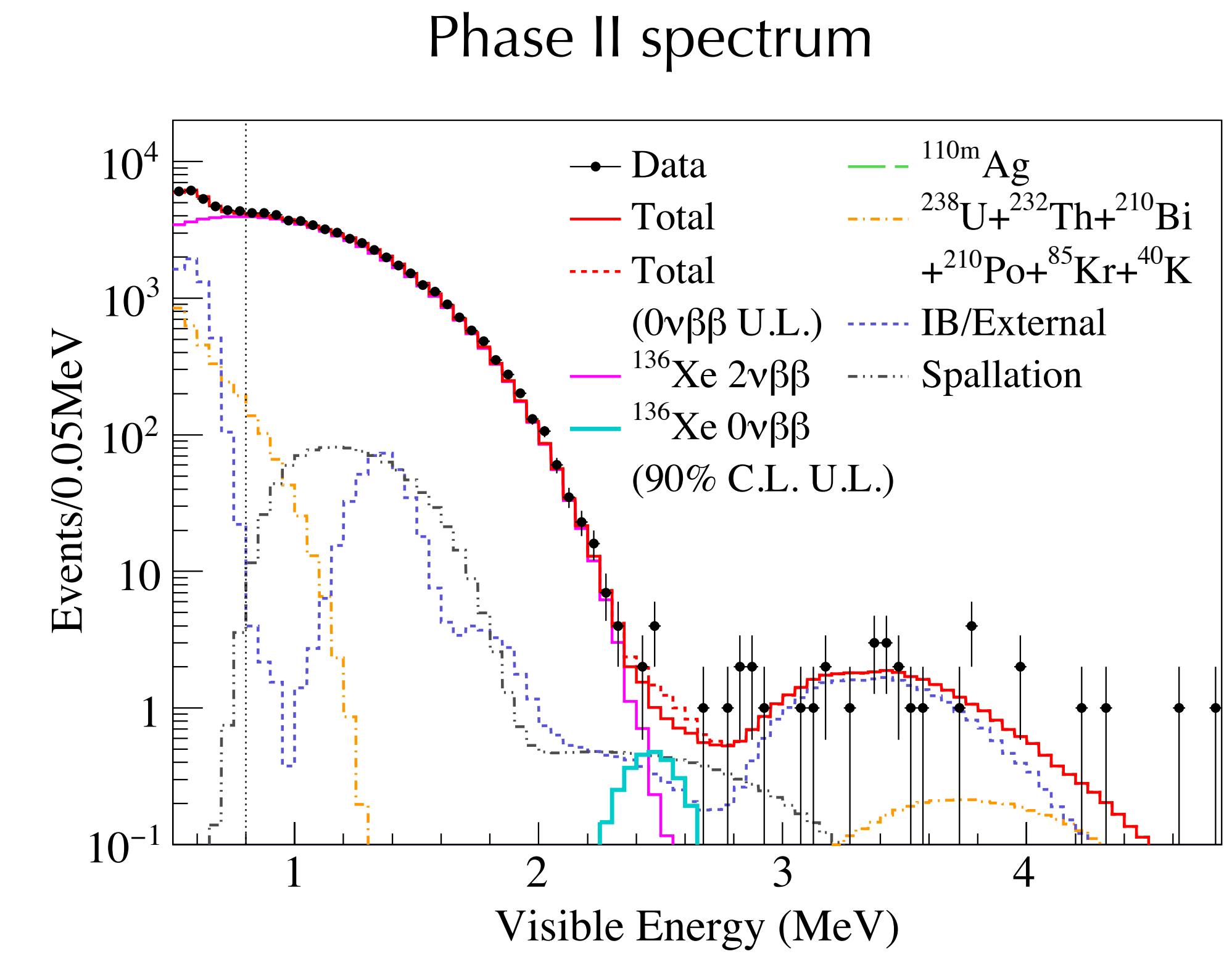
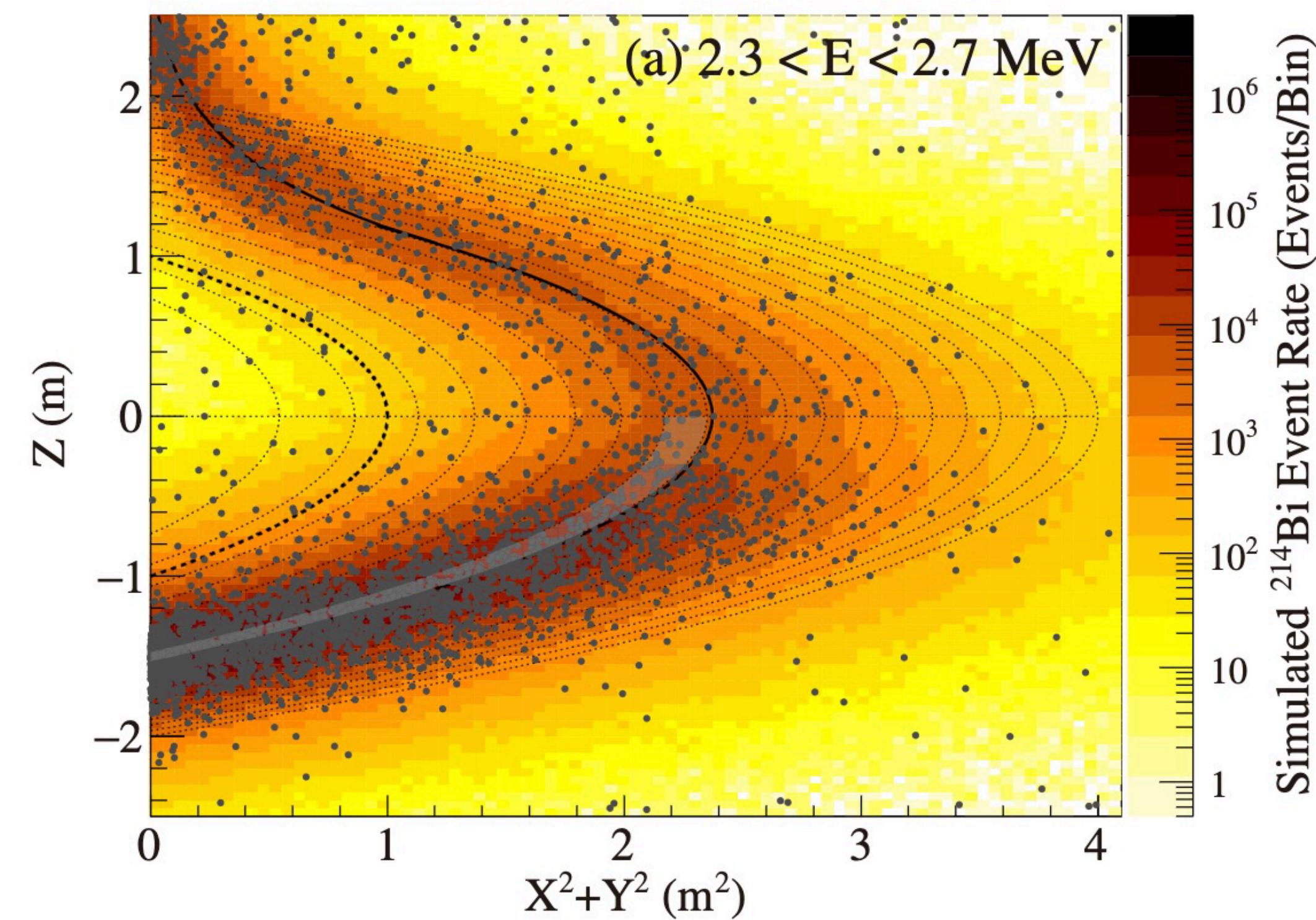


purification

Phase II spectrum



# KamLAND-Zen 400 Backgrounds



Fiducial volume limited by balloon backgrounds

# New Balloon Fabrication

- Performed in a class 1 cleanroom
- Full body covering, laundered after each use
- All materials/tools cleaned with ethanol and pure H<sub>2</sub>O
- 1.5 years, 20+ researchers

1. Washing



2. Cutting



3. Welding



4. Leak tests / repairs

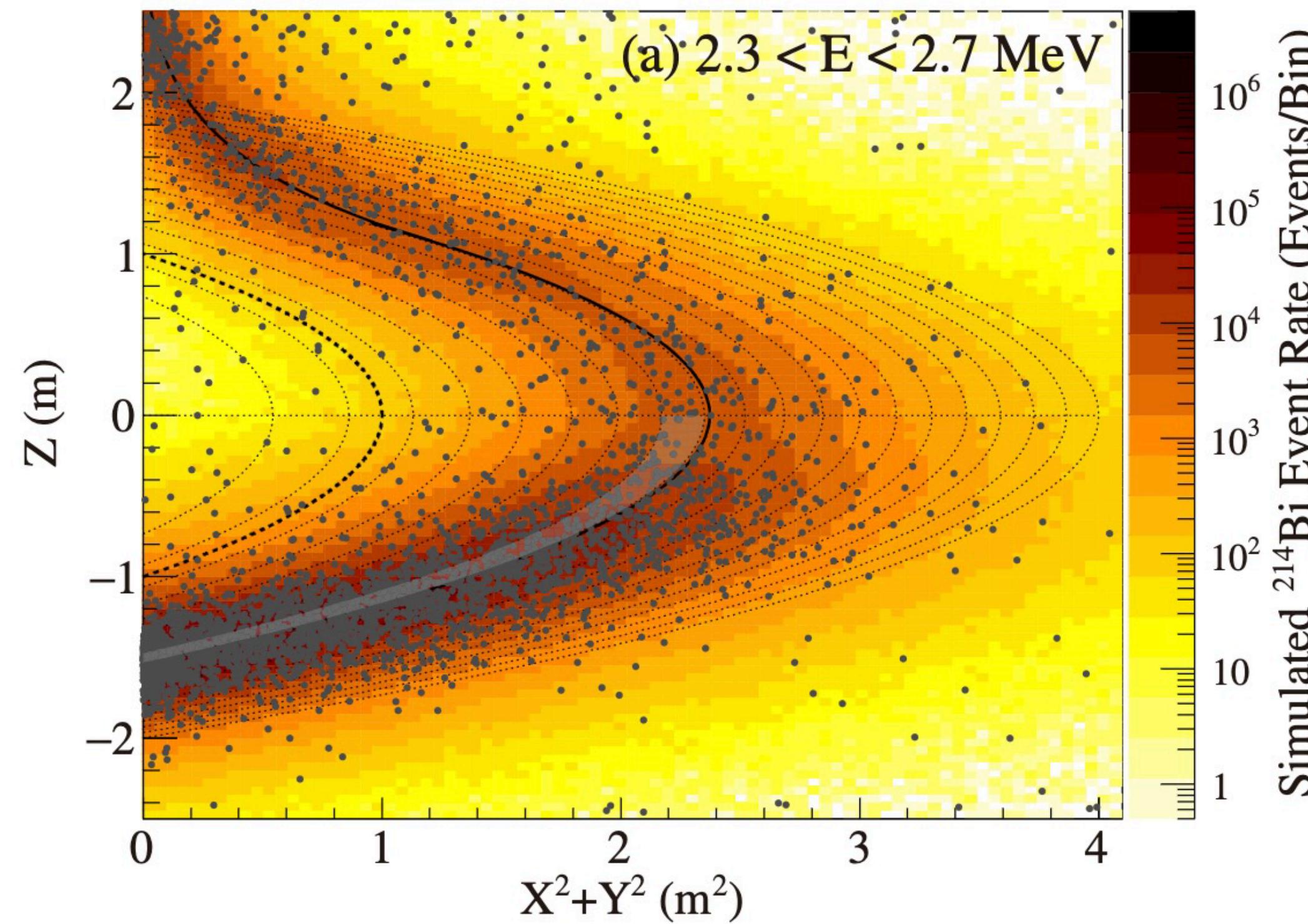


5. Deployment

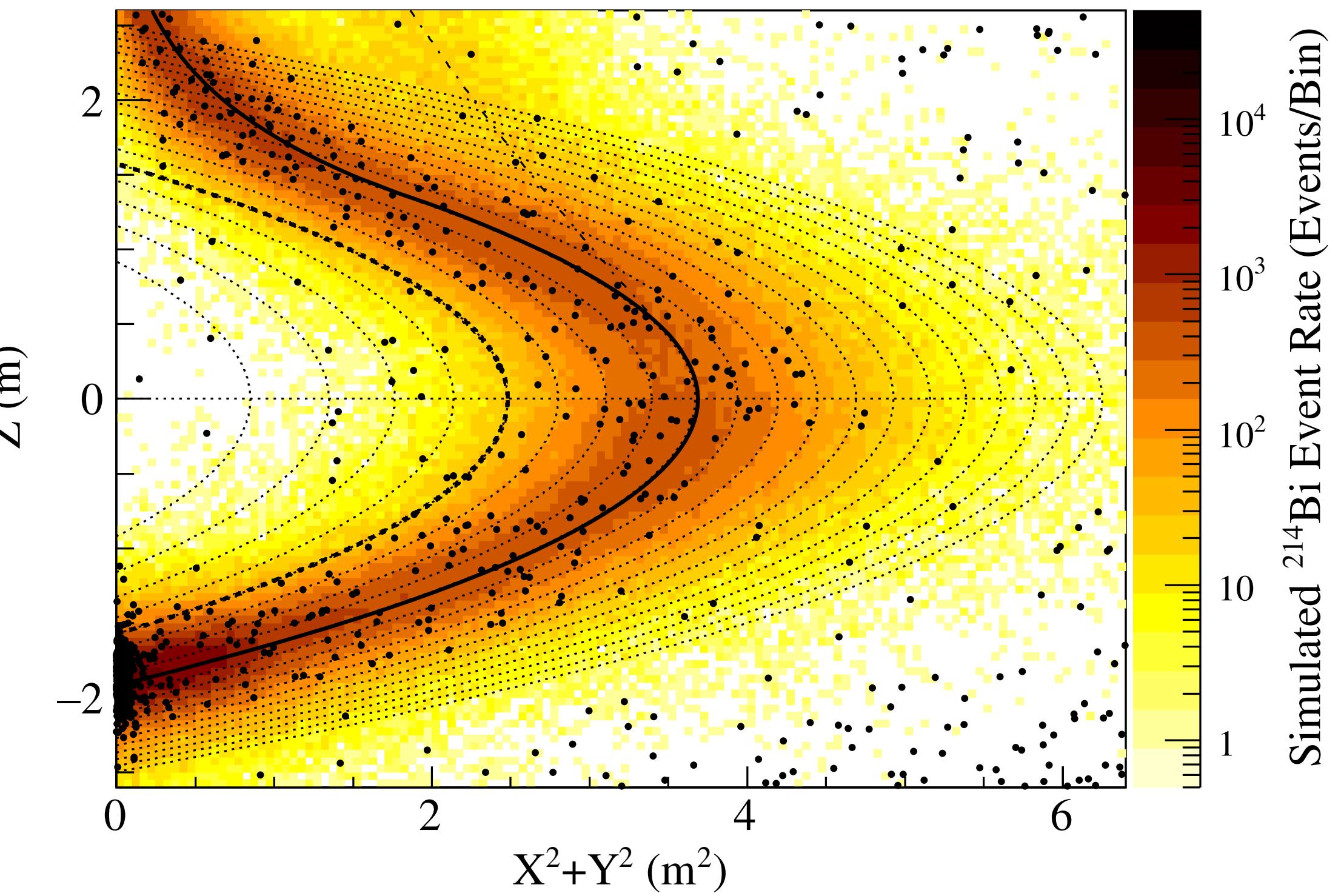


# Balloon Backgrounds

KamLAND-Zen 400 (530 days)



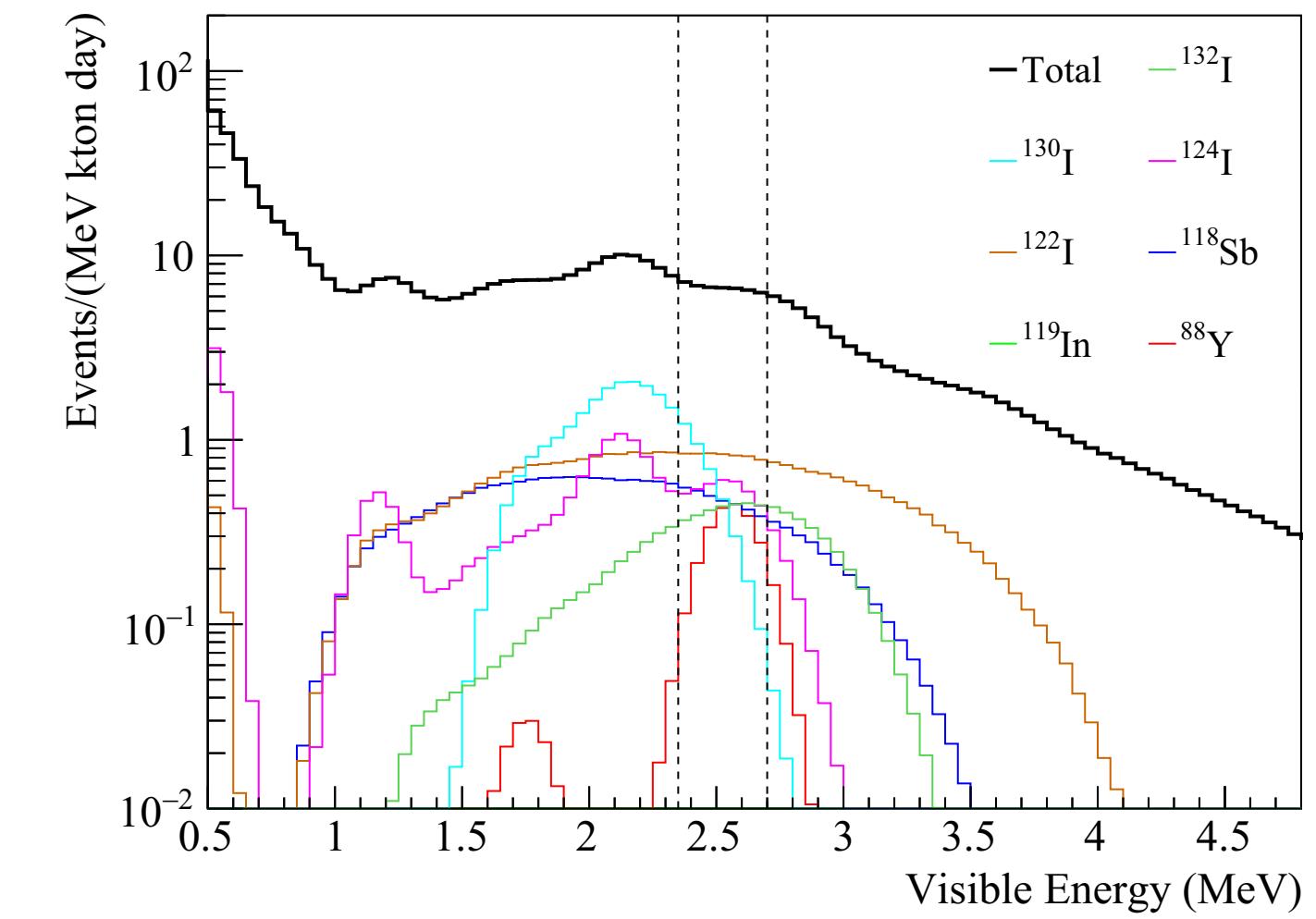
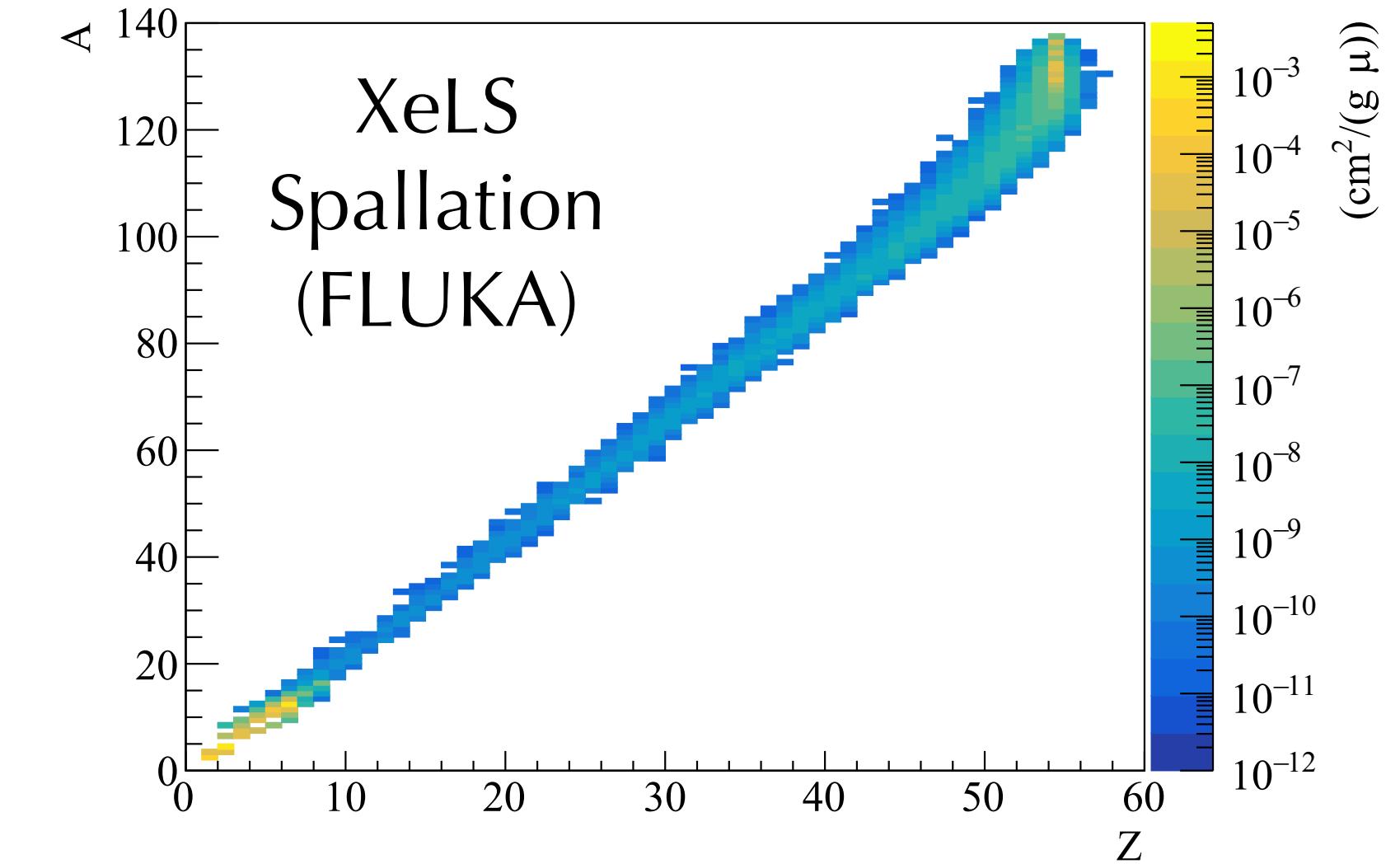
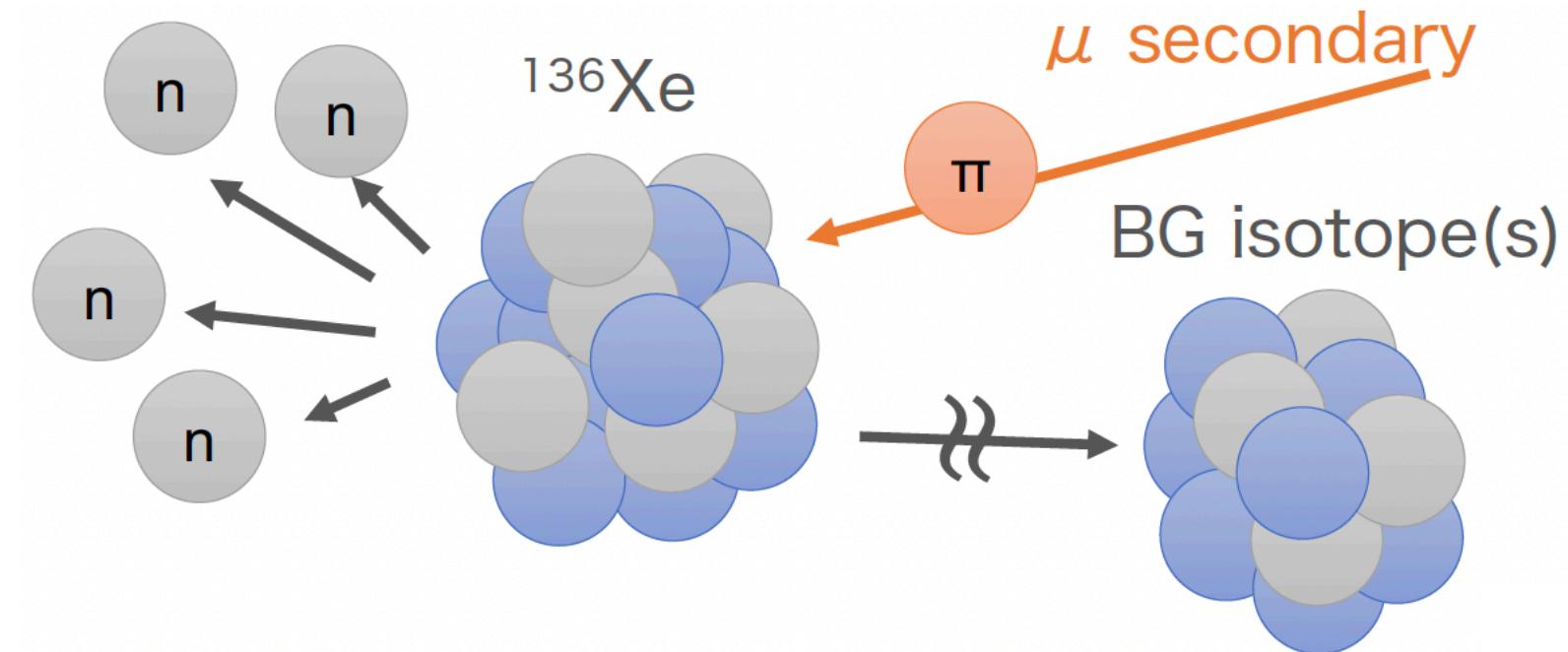
KamLAND-Zen 800 (520 days)



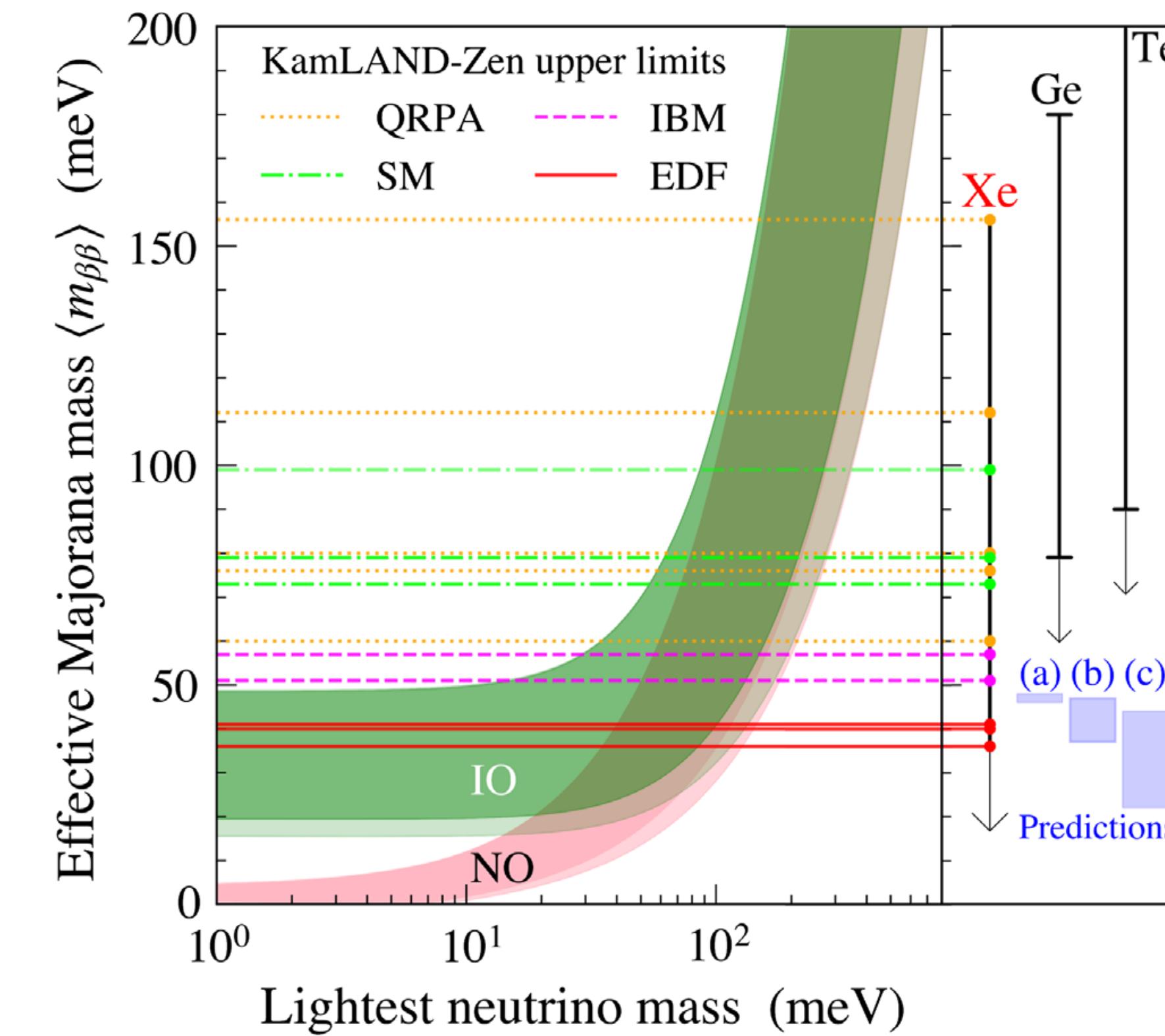
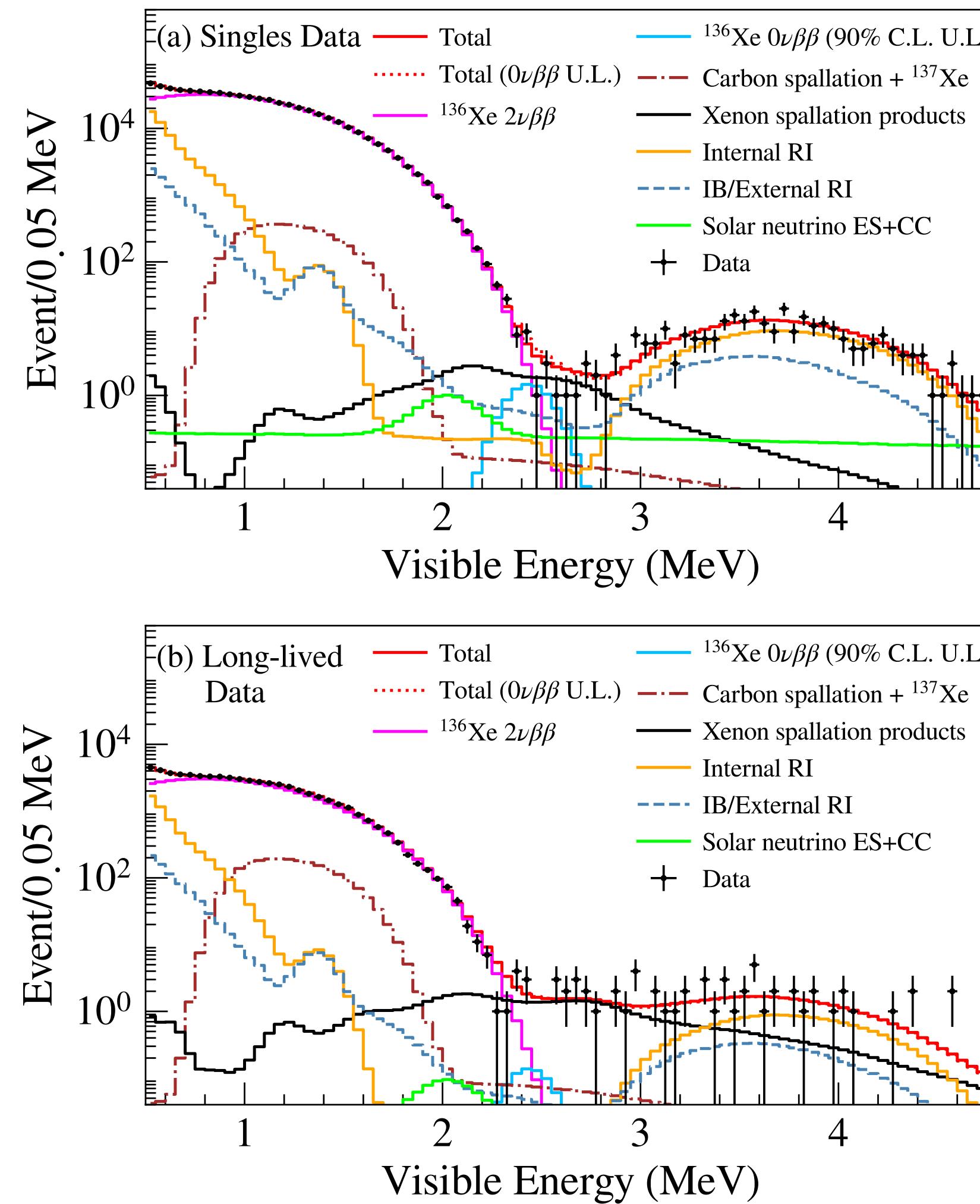
- U/Th reduced by factor of 10
- Fiducial volume increased by factor of 2.5!

# Spallation Backgrounds

- Long-lived Xe spallation products are the dominant background in KLZ-800!
- Tag with 3-fold spatial / temporal coincidence
- 42% rejection with 8.6% sacrifice

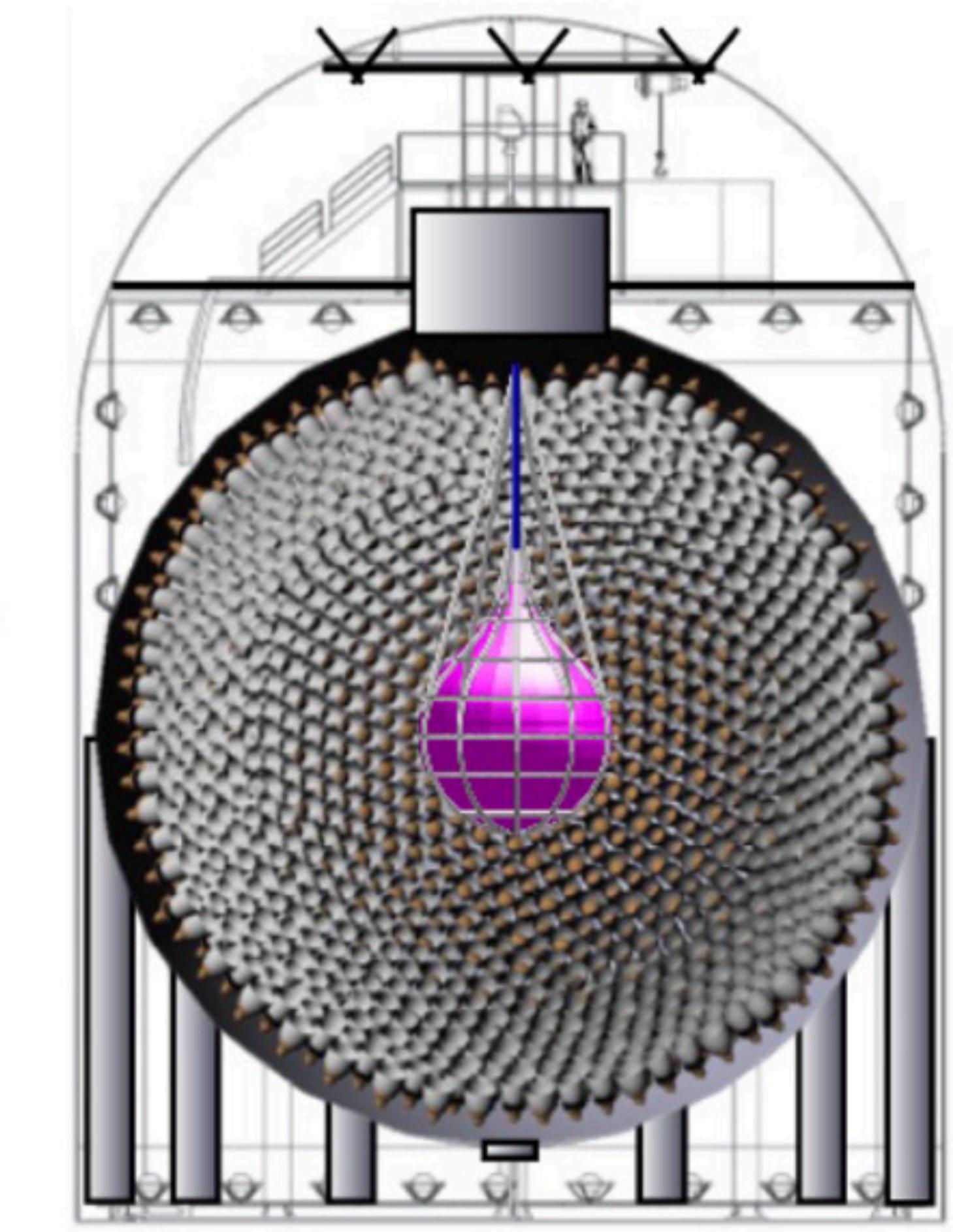


# KamLAND-Zen 800 Results



# Future: KamLAND2-Zen

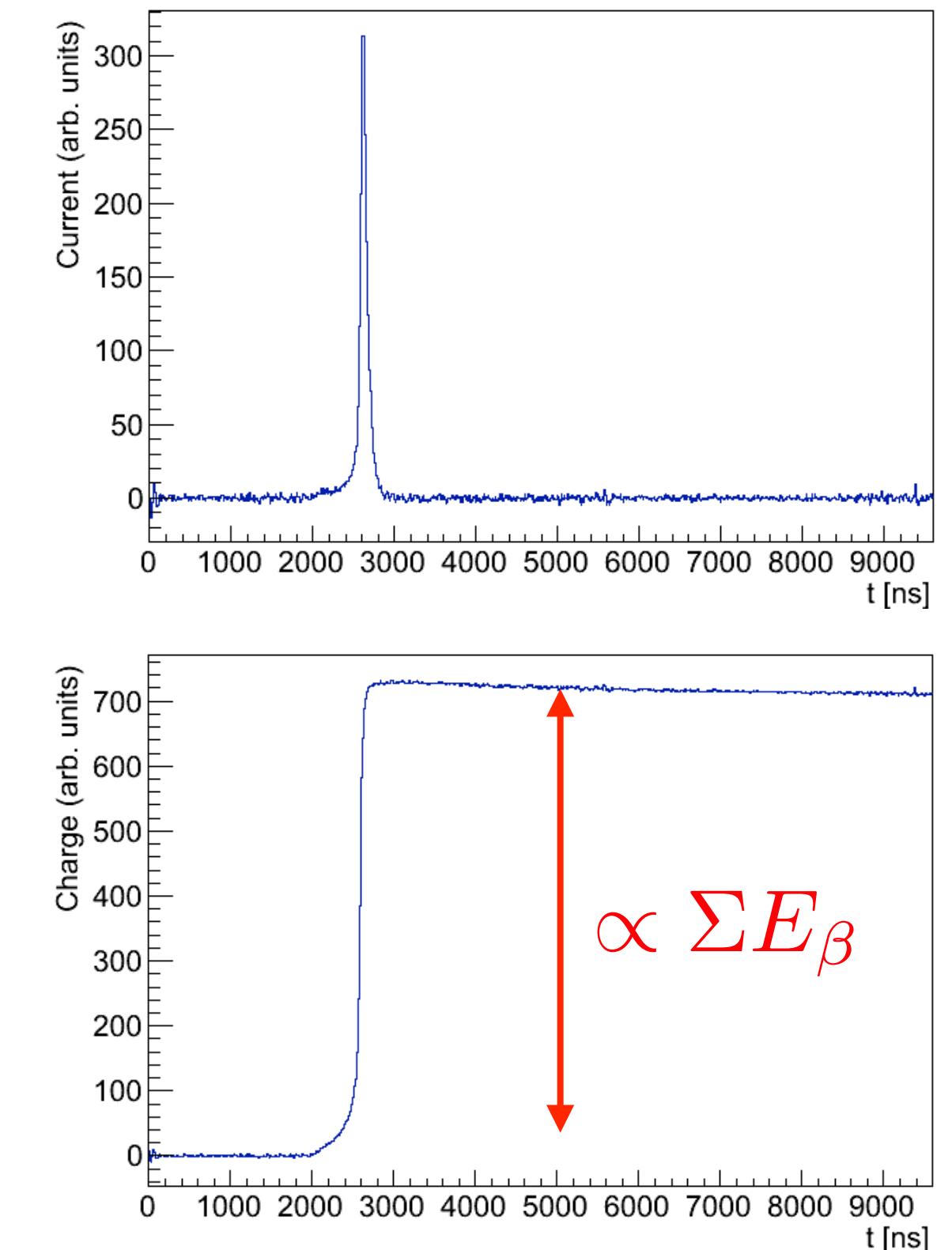
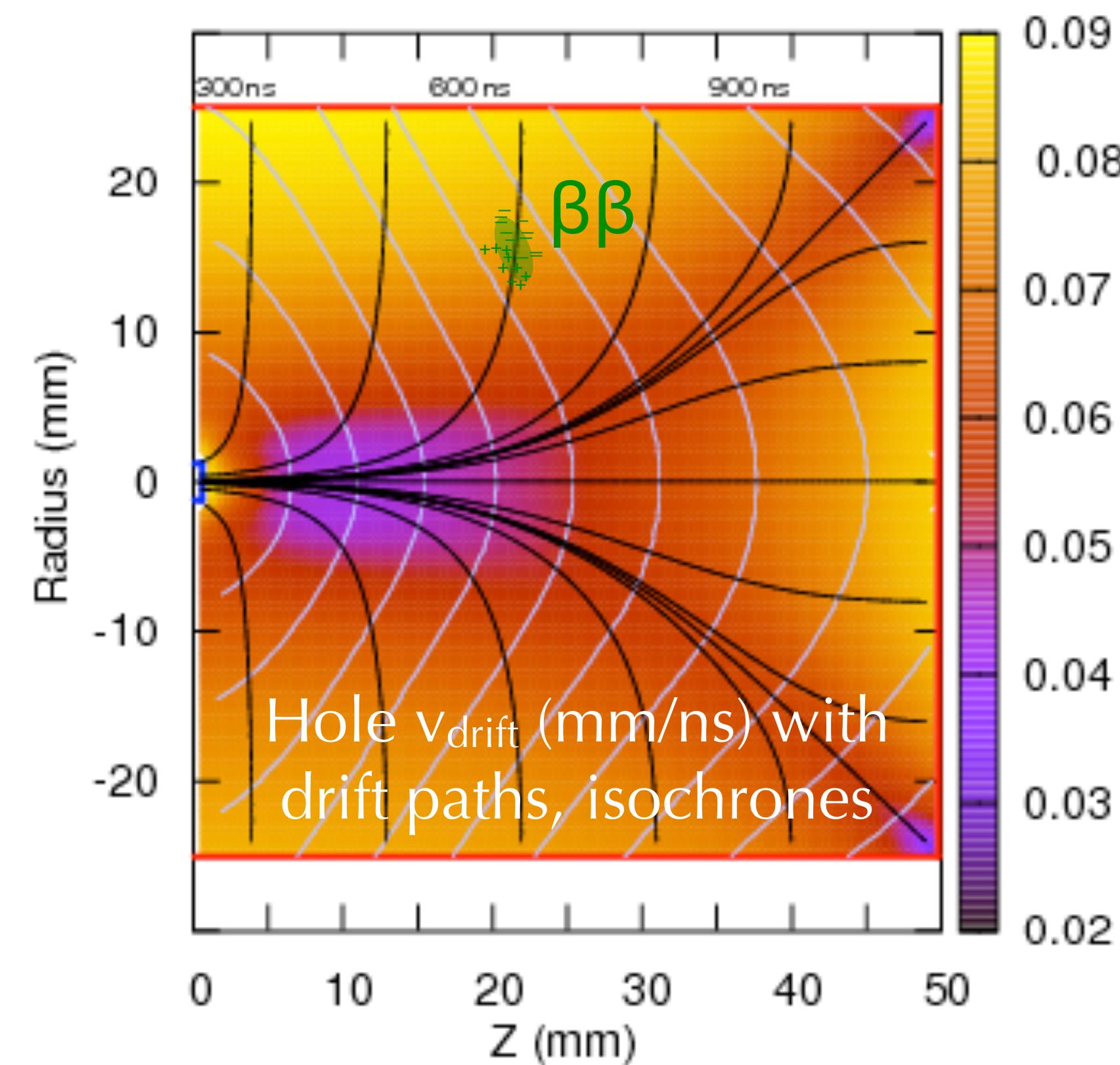
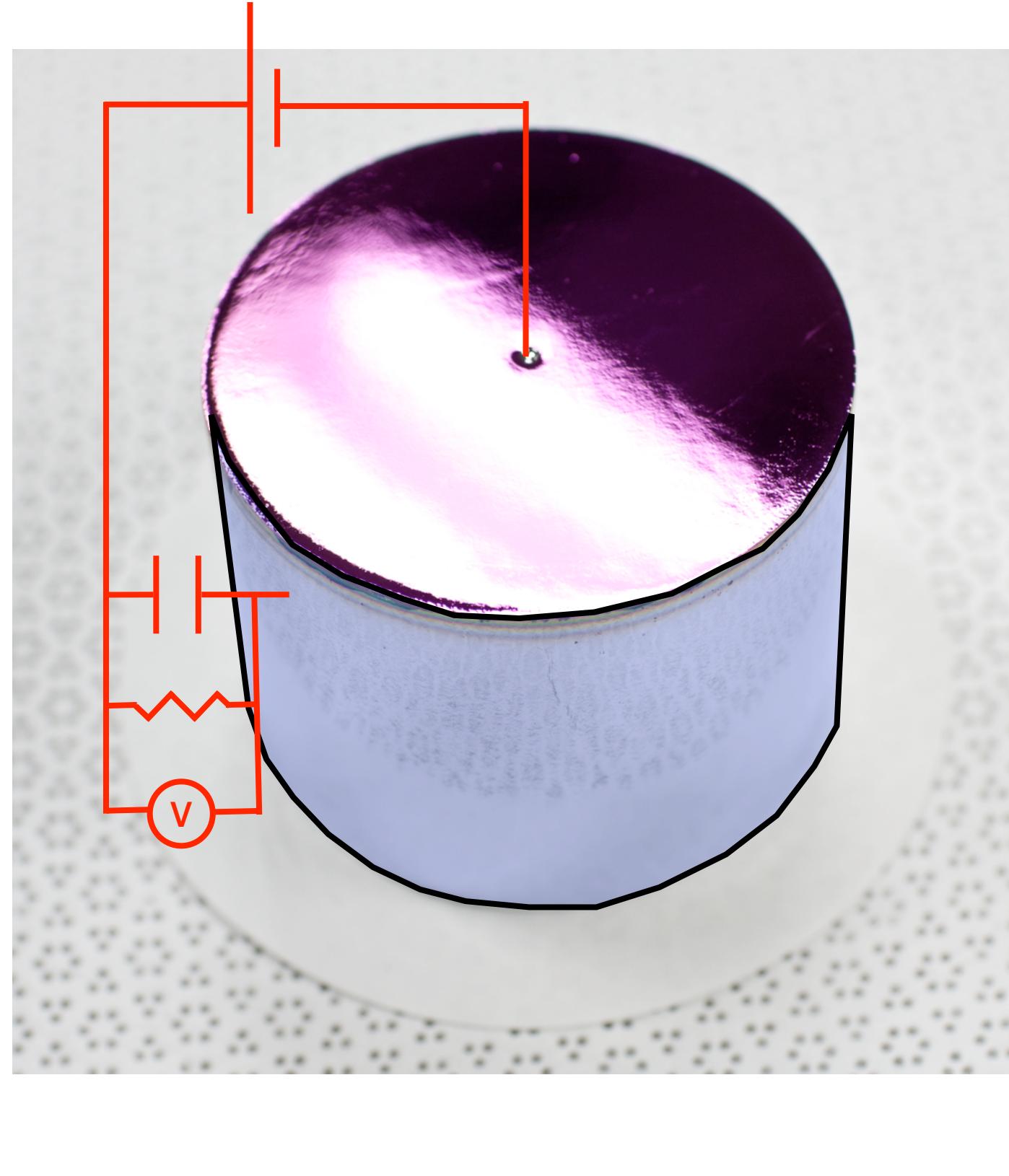
- Improved energy resolution  
(x5 increased light collection)
  - Higher light yield LS (x1.4)
  - New high-quantum-efficiency PMTs (x1.9)
  - Light-collecting Winston cones (x1.8)
- Improved background rejection
  - New scintillating mini-balloon
  - Upgraded electronics (~100% spallation  $n$  detection)
- More xenon ( $\rightarrow$  1 ton)
- Target half-life sensitivity:  $2 \times 10^{27}$  yr ( $m_{\beta\beta} \sim 20$  meV)



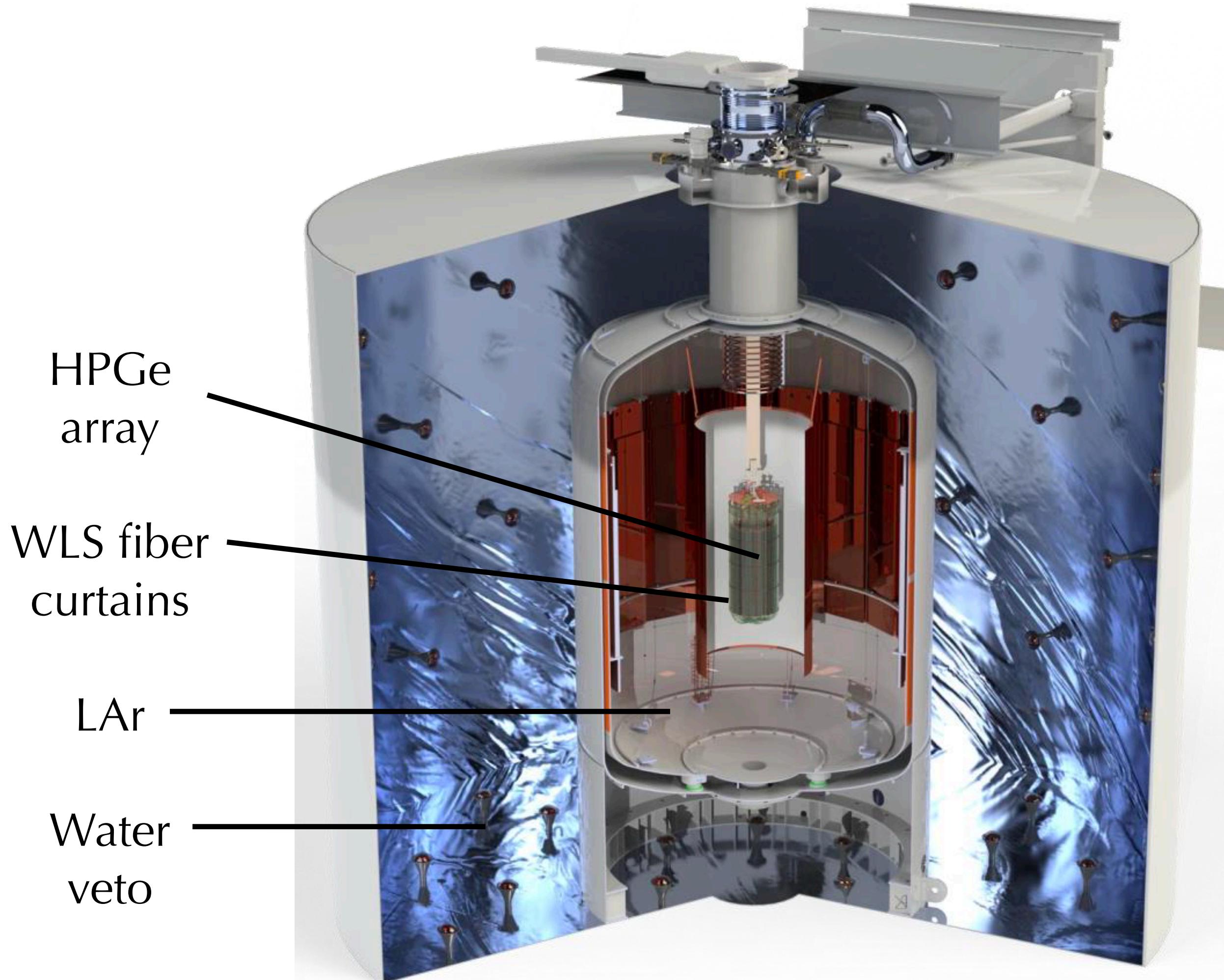
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# Germanium Detectors



# LEGEND

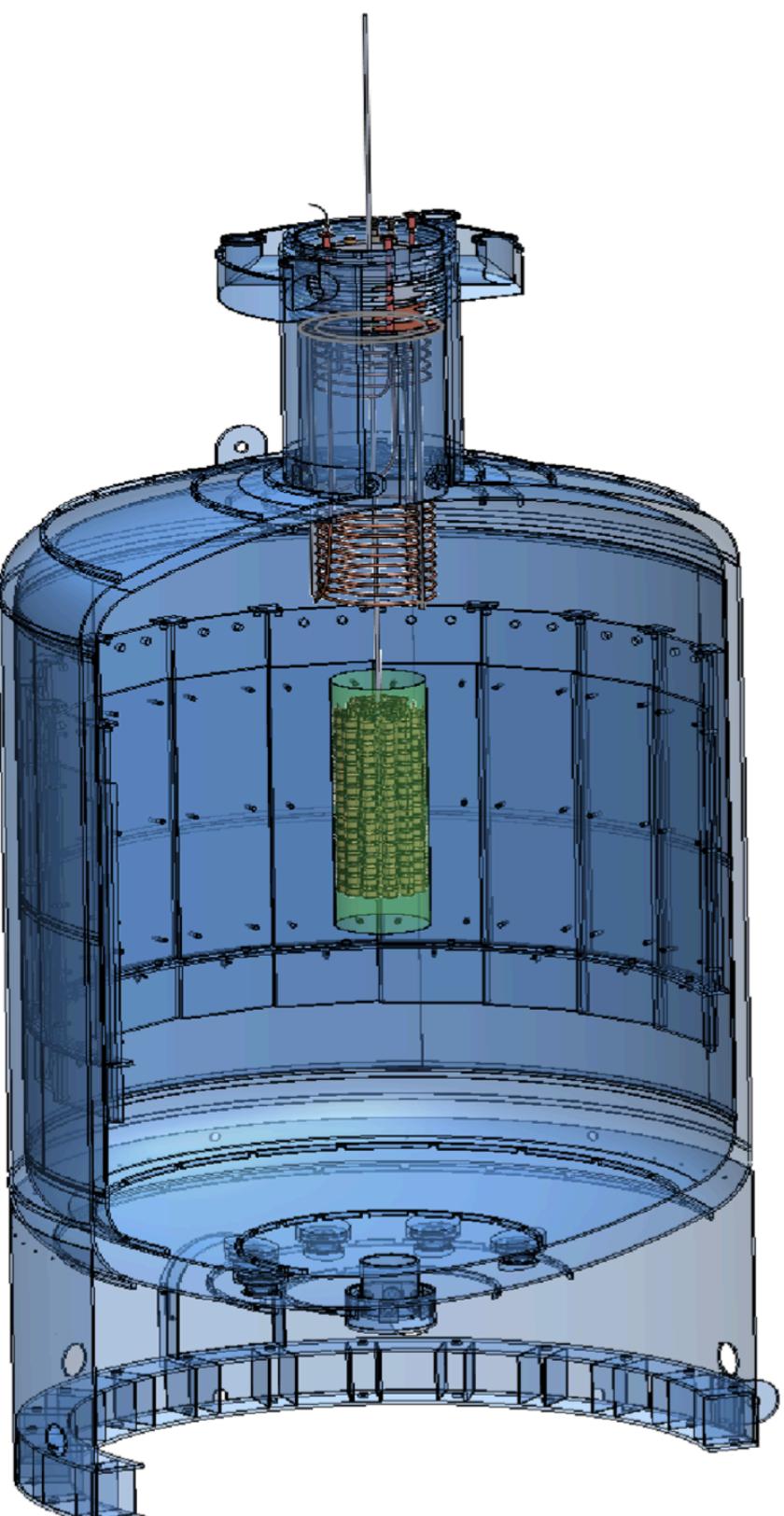


~60 Institutions, ~300 Scientists

# LEGEND Stages

- LEGEND 200

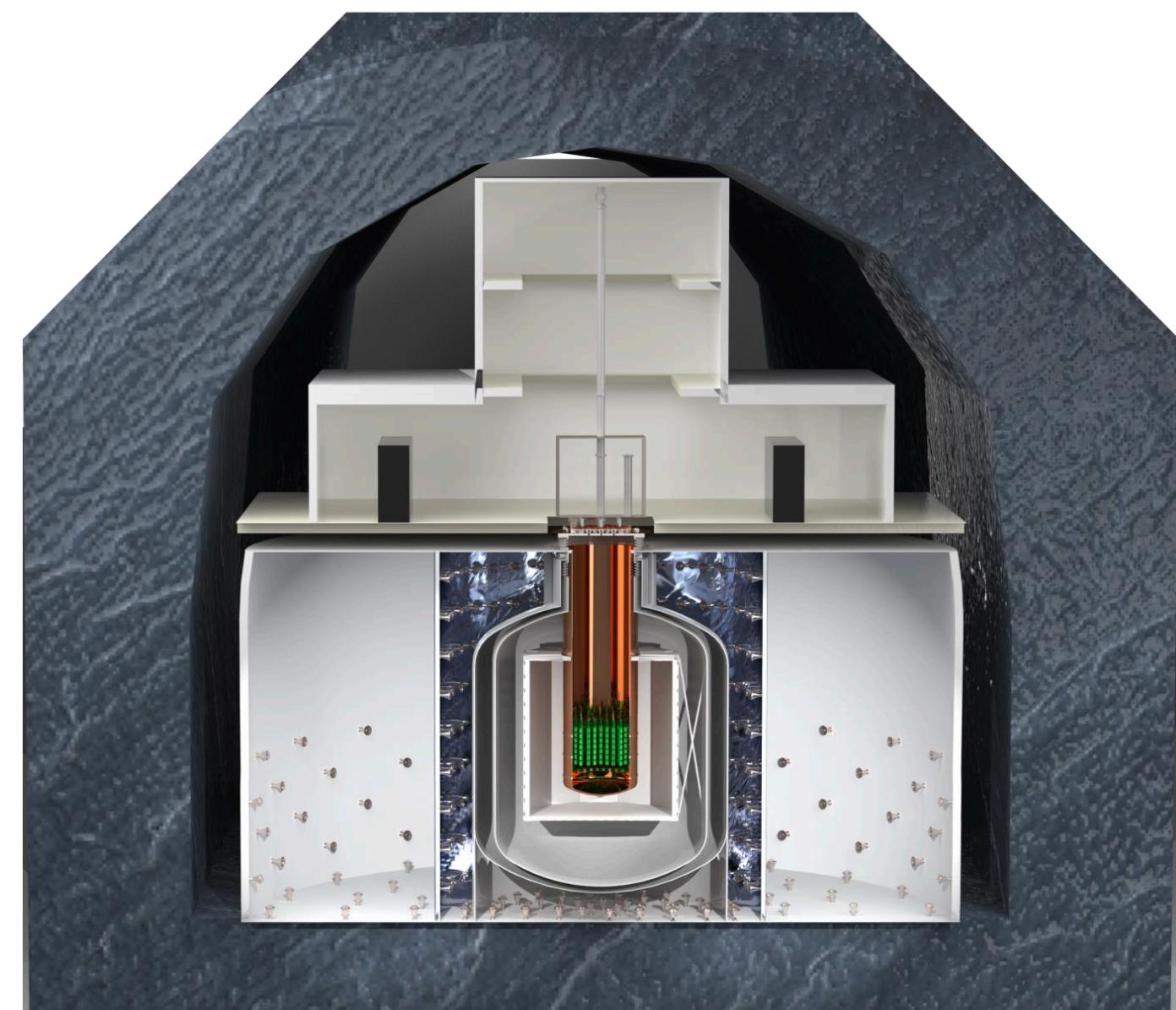
- 200 kg in upgrade of existing GERDA infrastructure at LNGS
- Reuses GERDA and MAJORANA enriched detectors (~60 kg) and adds new detectors (~140 kg)
- Background goal: 0.6 cts/(FWHM t yr)
- Data taking is underway



LEGEND 200

- LEGEND 1000

- 1000 kg (staged in ~250 kg payloads)
- Uses underground argon (eliminate  $^{42}\text{Ar}$  and  $^{39}\text{Ar}$ )
- Background goal: <0.1 cts/(FWHM t yr)
- Location: SNOLAB or LNGS



LEGEND 1000

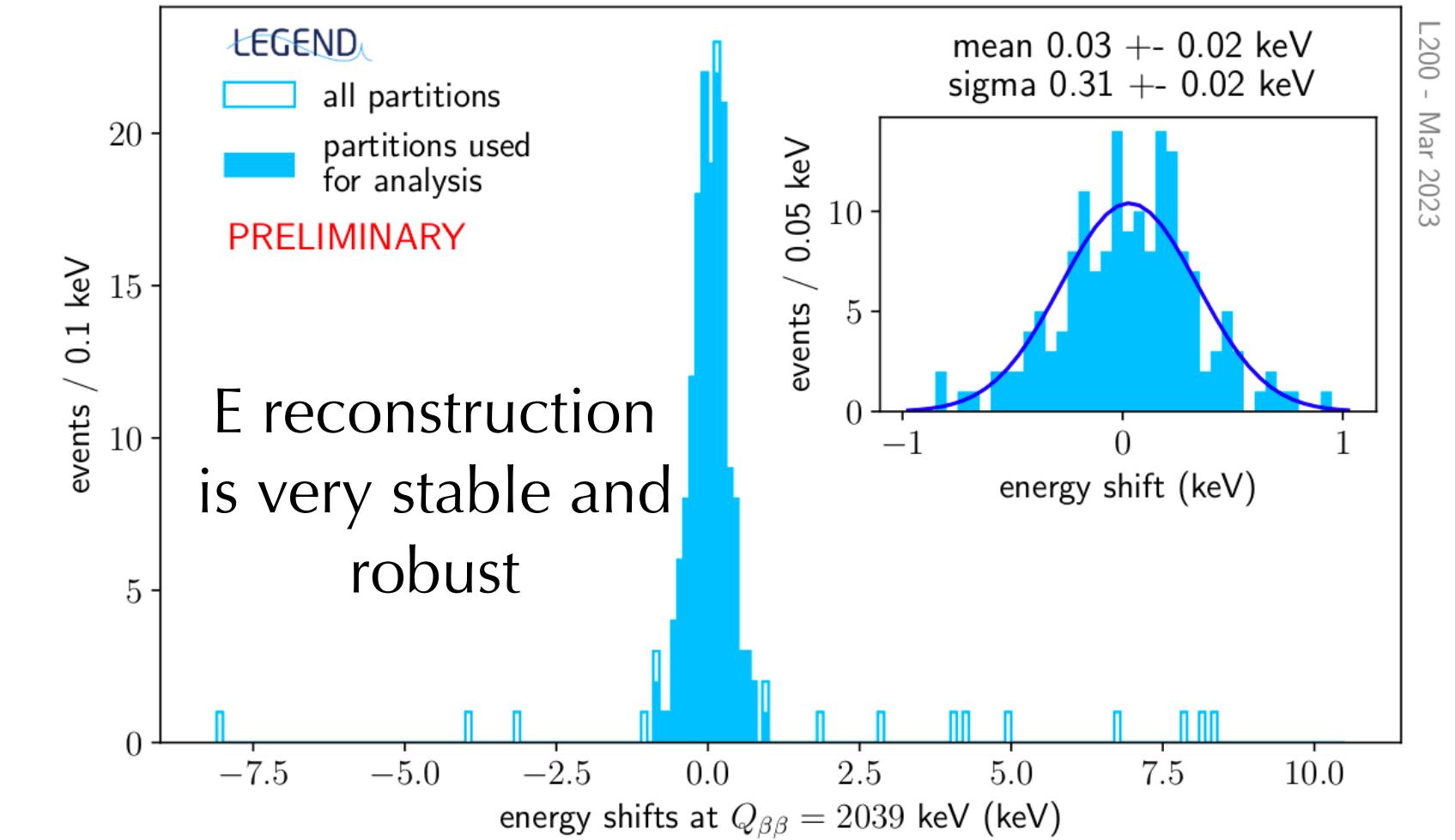
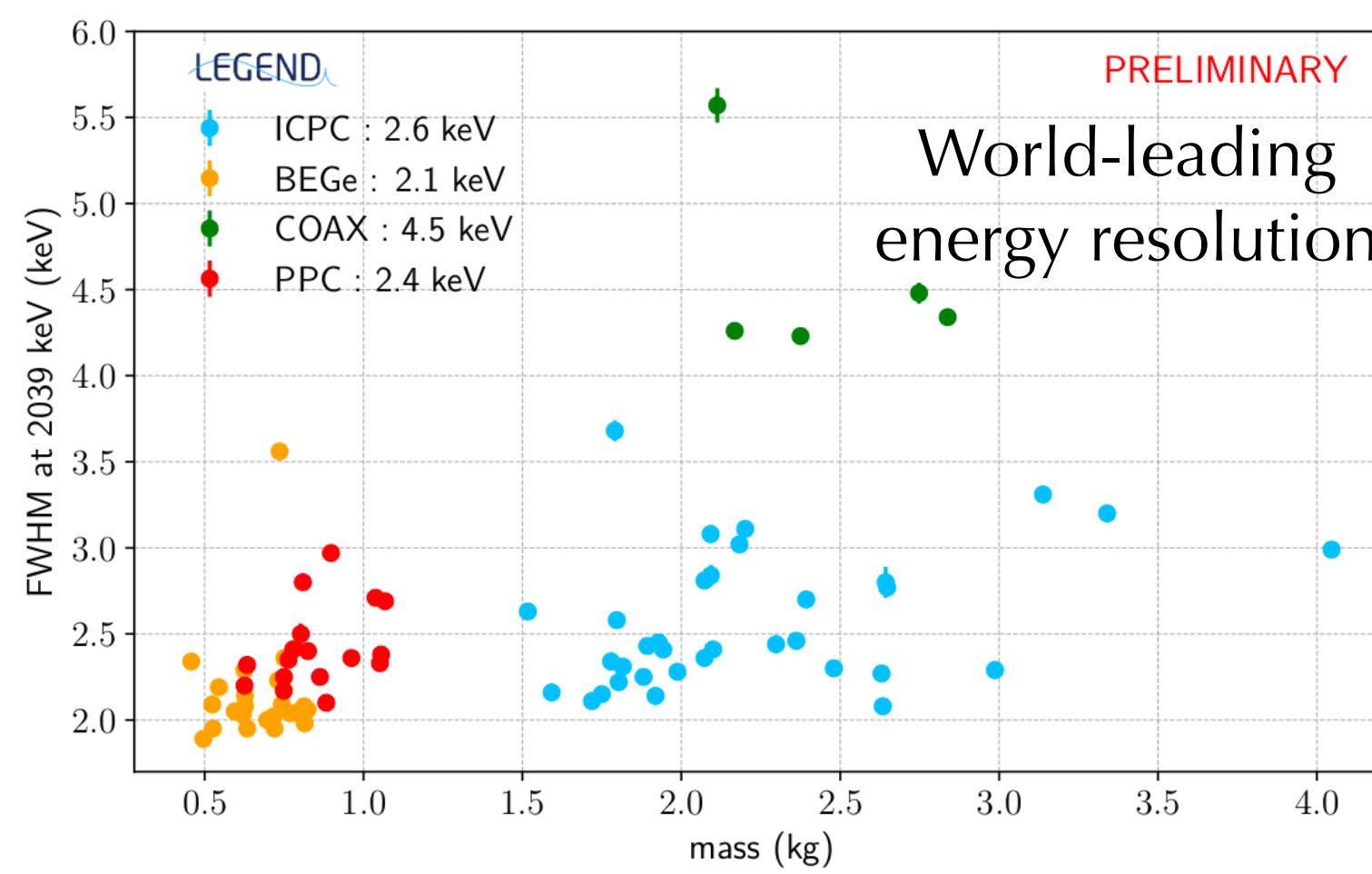
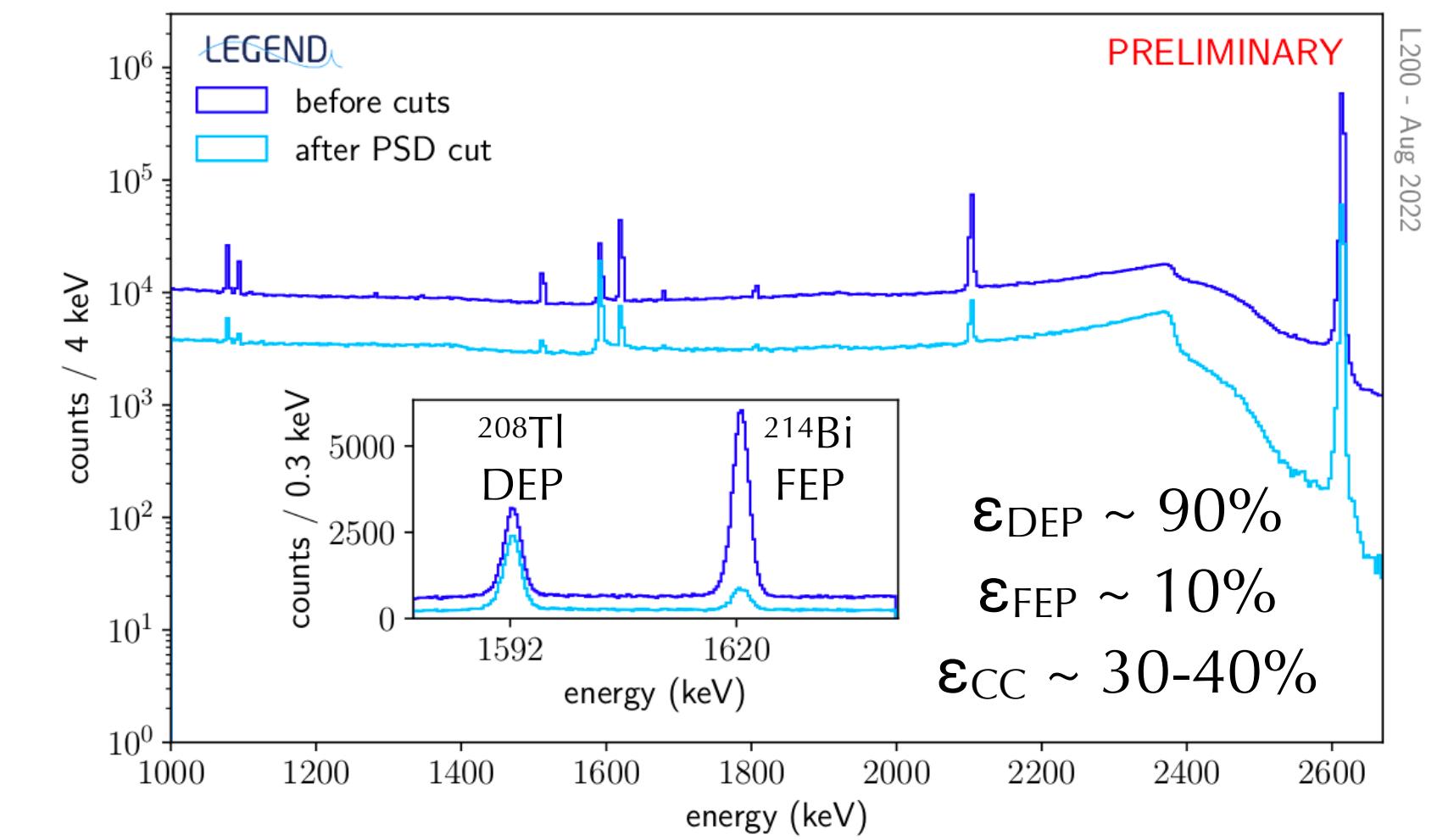
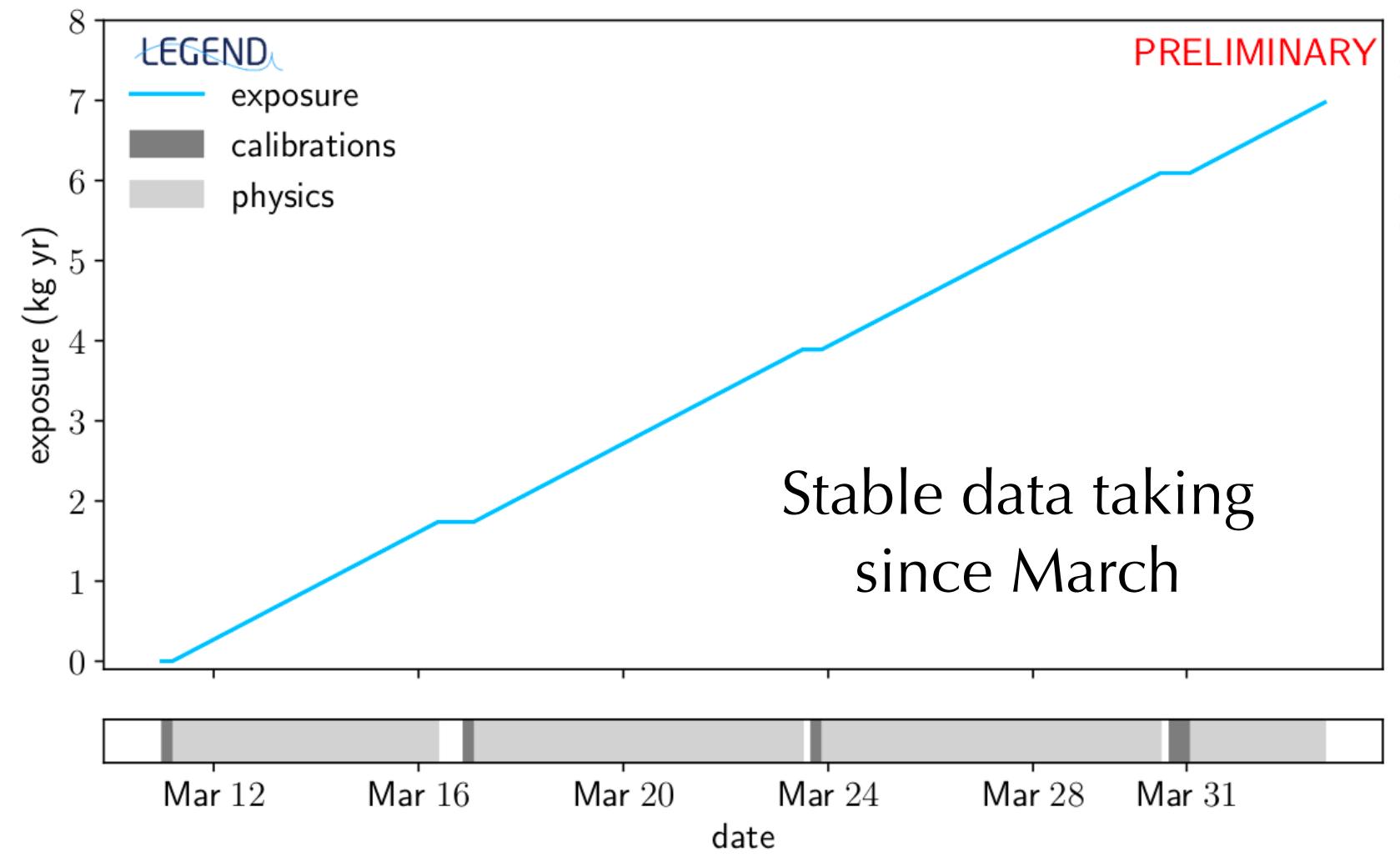
# LEGEND-200 Status

- Cryostat upgrade completed, LAr filling in 2021
- LAr instrumentation, Ge readout implementation, deployment of “the lock” in early 2022
- “L60” commissioning runs taken over summer 2022
- First 140 kg of L200 installed last fall
- Physics data taking in progress!



Photo: Enrico Sacchetti

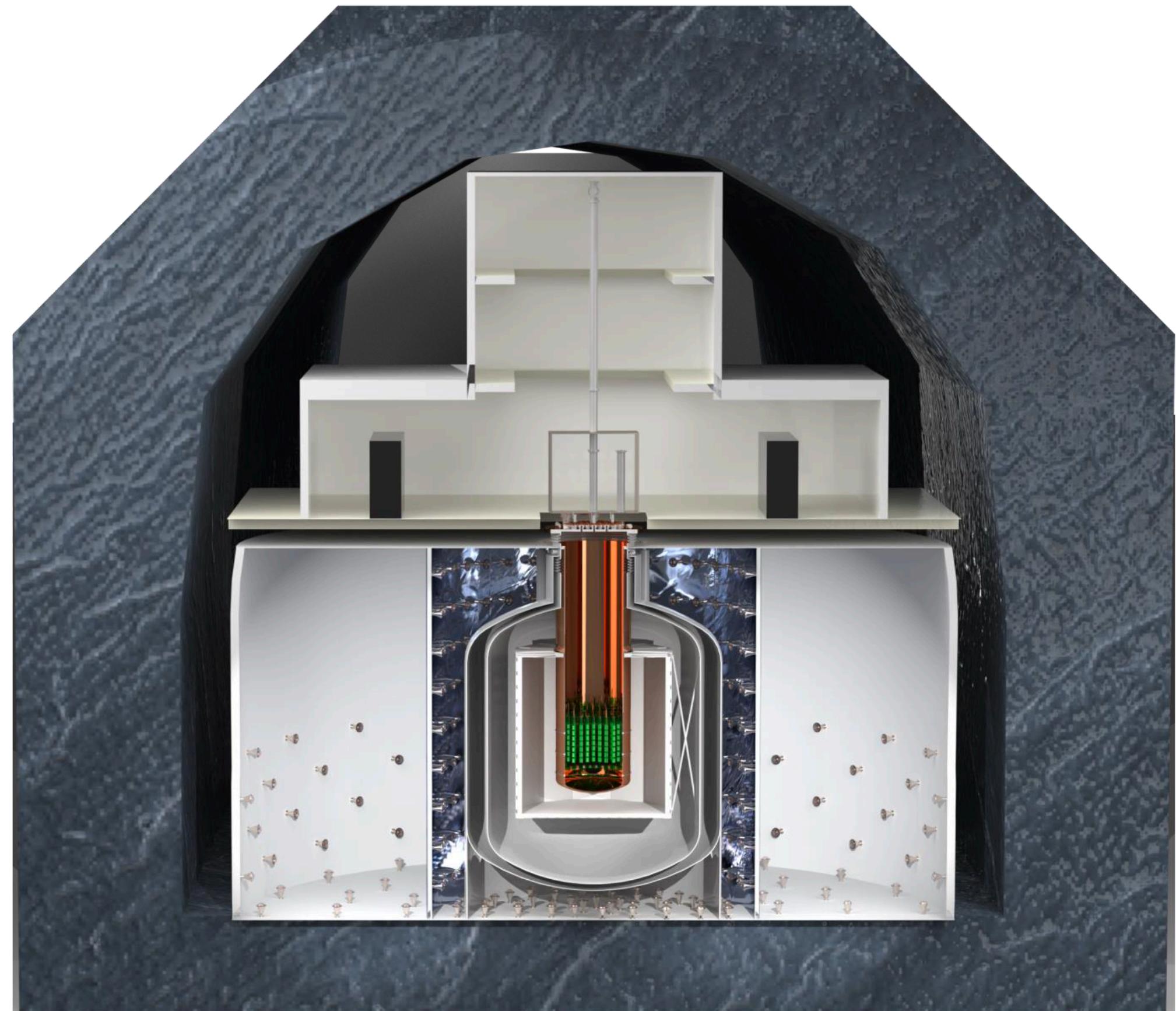
# L200 Preliminary Data



Stay tuned!

# LEGEND-1000 Status

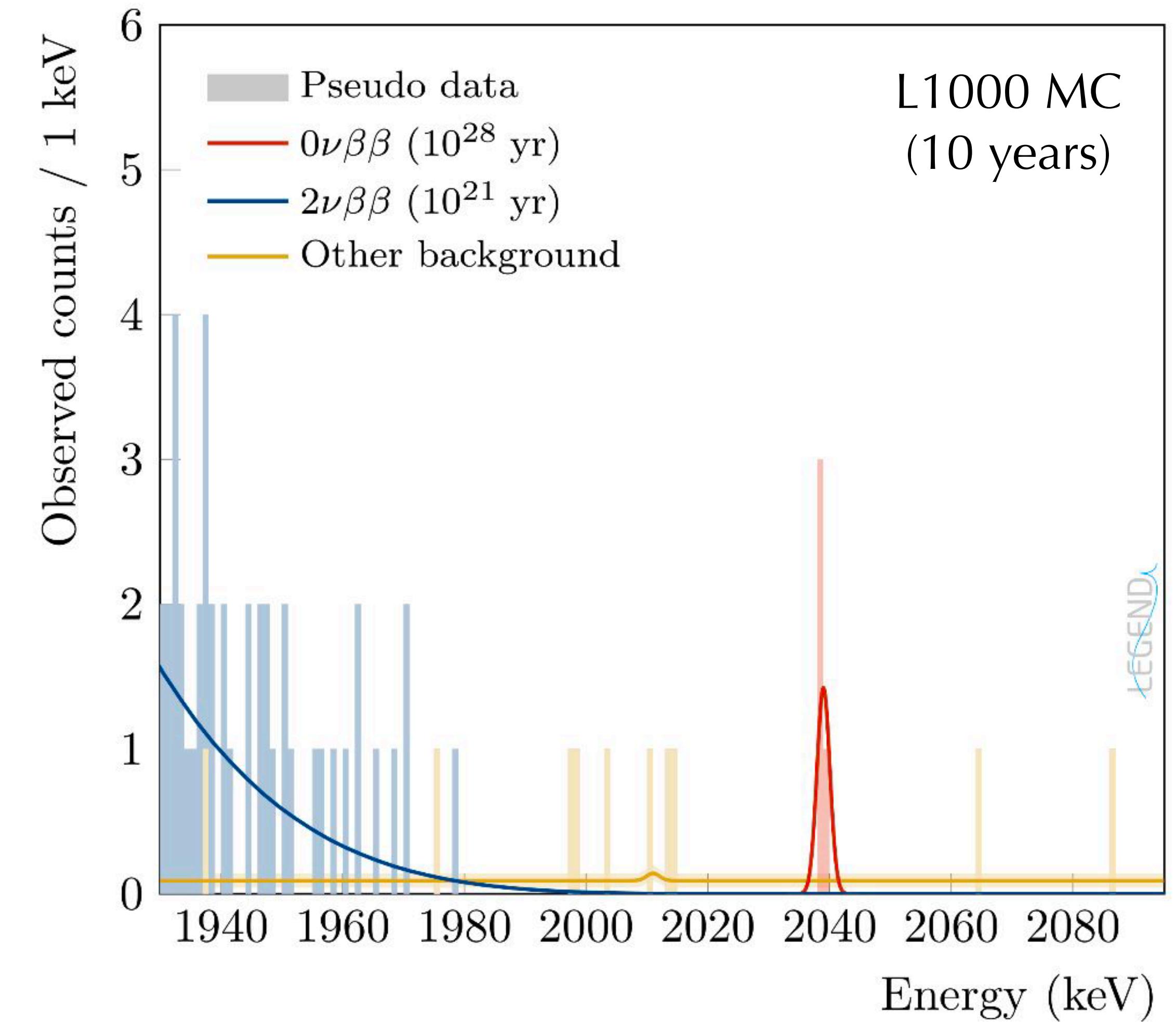
- Preliminary conceptual design:  
arXiv:2107.11462
- Successful performance at 2021 DOE  
“Portfolio Review”
- Site selection underway (in conjunction with  
DOE “Analysis of Alternatives” process)
- Preparing for “CD1” Conceptual Design  
Review this Winter



# LEGEND Sensitivity



- Expect a clear peak over a quasi-background-free continuum near  $Q_{\beta\beta}$
- LEGEND 200:  $T_{1/2} > 10^{27}$  yr
- LEGEND 1000:  $T_{1/2} > 10^{28}$  yr
  - $3\sigma$  discovery sensitivity covers the inverted ordering



# Summary

- Discovery of neutrinoless double beta decay would be the first observation of matter creation (without antimatter) by humans, and is deeply important.
- KamLAND-Zen's high-exposure strategy has produced leading half-life limits; KamLAND2-Zen will push this technology further.
- LEGEND's low-background approach is poised to lead the field in the coming years. LEGEND-200 is taking data, and LEGEND-1000 is in preparation.
- The international experimental program in both experiment and theory is robust and aggressive. A steady march in sensitivity improvement is expected for at least a decade. Discovery could come at any time!