# Systematic treatment of second order NLO QED radiative corrections to exclusive observables

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- Pairs



## **QED Factorization Theorem**

The QCD factorization theorem can be adopted for the QED case (omitting for a while pair corrections) *e.g.* for Bhabha:

$$\begin{split} \mathrm{d}\sigma &= \int_{\bar{z}_1}^1 \!\! \mathrm{d}z_1 \! \int_{\bar{z}_2}^1 \!\! \mathrm{d}z_2 \mathcal{D}_{\mathrm{ee}}^{\mathrm{str}}(z_1) \mathcal{D}_{\mathrm{ee}}^{\mathrm{str}}(z_2) \! \Big( \!\! \mathrm{d}\sigma^{(0)}(z_1,z_2) + \mathrm{d}\bar{\sigma}^{(1)}(z_1,z_2) + \mathcal{O}\left(\alpha^2 L^0\right) \Big) \\ &\times \int_{\bar{y}_1}^1 \!\! \frac{\mathrm{d}y_1}{Y_1} \int_{\bar{y}_2}^1 \!\! \frac{\mathrm{d}y_2}{Y_2} \mathcal{D}_{\mathrm{ee}}^{\mathrm{frg}}(\frac{y_1}{Y_1}) \mathcal{D}_{\mathrm{ee}}^{\mathrm{frg}}(\frac{y_2}{Y_2}), \end{split}$$

where  $\sigma^{(0)}$  is the Born-level cross section,  $\bar{\sigma}^{(1)}$  is the  $\overline{\rm MS}$  subtracted  $\mathcal{O}\left(\alpha\right)$  contribution,

$$\begin{split} \mathcal{D}_{\text{ee}}^{\text{str,frg}}(z) \, &= \, \delta(1-z) + \frac{\alpha}{2\pi} d^{(1)}(z,\mu_0,m_e) + \frac{\alpha}{2\pi} L P^{(0)}(z) \\ &+ \, \left(\frac{\alpha}{2\pi}\right)^2 \left(\frac{1}{2} L^2 P^{(0)} \otimes P^{(0)}(z) + L P^{(0)} \otimes d^{(1)}(z,\mu_0,m_e) \right. \\ &+ \, L P_{\text{ee}}^{(1,\gamma) \text{str,frg}}(z) \right) + \mathcal{O}\left(\alpha^2 L^0, \alpha^3\right) \end{split}$$

## **QED Master Formula Ansatz**

Using slicing in the photon energy, we cast the corrected cross section in the form

$${\rm d}\sigma = {\rm d}\sigma^{(0)} + {\rm d}\sigma^{(1)}_{\rm S+V} + {\rm d}\sigma^{(1)}_{\rm H} + {\rm d}\sigma^{(2)\textit{NLO}}_{\rm S+V} + {\rm d}\sigma^{(3)\textit{LO}}_{\rm H} + {\rm d}\sigma^{(3)\textit{LO}}_{\rm H} + \dots$$

For many observables we need to know the complete kinematics including hard photon angles, which are integrated over in the QCD-like formula.

Let us decompose the  $\mathcal{O}\left(\alpha^2L^{2,1}\right)$  hard radiation contribution

$$\mathrm{d}\sigma_{\mathrm{H}}^{(2)NLO} = \mathrm{d}\sigma_{\mathrm{HH(coll)}}^{(2)} + \mathrm{d}\sigma_{\mathrm{HH(s-coll)}}^{(2)} + \mathrm{d}\sigma_{(\mathrm{S+V)H(n-coll)}}^{(2)} + \mathrm{d}\sigma_{(\mathrm{S+V)H(coll)}}^{(2)}$$

where slicing in the photon emission angle is applied:

- "coll" means collinear photon(s) with  $\vartheta_{\gamma} < \theta_0 \ll 1$ ,
- "n-coll" means non-collinear photon with  $\vartheta_{\gamma} > \theta_{0}$ ,
- "HH(s-coll)" means semi-collinear kinematics, i.e. one collinear photon and one non-collinear



# Particular NLO contributions (1)

The combined effect of virtual corrections and soft photon emission ones within the  $\mathcal{O}\left(\alpha^2L^1\right)$  can be obtained by convolution of the structure functions with the kernel cross section according to the general factorization theorem. Here one requires only one non-trivial convolution

$$\frac{\alpha}{2\pi} L \int_{1-\Delta}^{1} dz \int_{0}^{1} \frac{dx}{x} P^{(0)}\left(\frac{z}{x}\right) d\bar{\sigma}^{(1)}(x)$$

This integral can be found for any relevant process as demonstrated in [A.A., E. Scherbakova, ZhETF Pis'ma 2006] for the large-angle Bhabha case by getting  $d\sigma_{S+V}^{(2)\textit{NLO}}$  in agreement with the complete  $\mathcal{O}\left(\alpha^2\right)$  calculation.

# Particular NLO contributions (2)

Emission of two hard photons, HH, can be considered in three regions:

- 1. non-collinear:  $\theta_{1,2} > \theta_0$  suited for Monte Carlo simulation
- 2. semi-collinear:  $\theta_1 > \theta_0$  and  $\theta_2 < \theta_0$  in  $\mathcal{O}(\alpha^2 L)$  has factorized form  $d\sigma_{\mathrm{H}}^{(1)} \otimes R_{\mathrm{H}}^{\mathrm{ISR,FSR}}(z)$
- 3. collinear:  $\theta_{1,2} < \vartheta_0$  is described by the HH radiation factor convoluted with the Born

Emission of one hard photon in  $\mathcal{O}\left(\alpha^2L\right)$  can be sliced into two domains:

- 1. non-collinear:  $\theta_{\gamma} > \vartheta_0$  as a product of two factors  $\mathrm{d}\sigma_{\mathrm{H}}^{(1)} \times \delta_{\mathrm{Soft+Virt}}^{\mathrm{LO}}$
- 2. collinear:  $\theta_{\gamma} < \vartheta_0$  is described by the collinear NLO H radiation factor (see below)



## QED Collinear Radiation Factors in NLO (1)

A. Arbuzov, E. Scherbakova,

Phys. Lett. B 660 (2008) 37 [arXiv:0706.2984]

$$\mathrm{d}\sigma[a(p_1) + b(p_2) \to c(q_1) + d(q_2) + \gamma(k \sim (1-z)p_1)] = \mathrm{d}\hat{\sigma}[a(zp_1) + b(p_2) \to c(q_1) + d(q_2)] \otimes R_\mathrm{H}^\mathrm{ISR}(z)$$

Emission of collinear photons in FSR and ISR with conditions

$$\vartheta_{\gamma} < \vartheta_{0}, \qquad \frac{m}{E} \ll \vartheta_{0} \ll 1, \qquad I_{0} = \ln \frac{\vartheta_{0}^{2}}{4}, \quad \frac{E_{\gamma}}{E} > \Delta \ll 1$$

In  $\mathcal{O}(\alpha)$  the result is well known:

$$R_{\mathrm{H}}^{\mathrm{ISR}}(z) = \frac{\alpha}{2\pi} \left[ \frac{1+z^2}{1-z} \left( \ln \frac{E^2}{m^2} - 1 + l_0 \right) + 1 - z + \mathcal{O}\left(\frac{m^2}{E^2}\right) + \mathcal{O}\left(\vartheta_0^2\right) \right]$$



## QED Collinear Radiation Factors in NLO (2)

Emission of two collinear photons (HH) in the same direction is described by a one-fold integral of results from [A.A. et al., Nucl. Phys. B 483 (1997) 83]:

$$R_{\rm HH}^{\rm ISR}(z) = \left(\frac{\alpha}{2\pi}\right)^2 L \left\{ (L+2l_0) \left(\frac{1+z^2}{1-z} (2\ln(1-z) - 2\ln\Delta - \ln z) + \frac{1+z}{2} \ln z - 1 + z \right) + \frac{1+z^2}{1-z} \left(\ln^2 z + 2\ln z - 4\ln(1-z) + 4\ln\Delta\right) + (1-z) \left(2\ln(1-z) - 2\ln\Delta - \ln z + 3\right) + \frac{1+z}{2} \ln^2 z \right\}$$

FSR factor is restored with help of the Gribov-Lipatov relation generalized for the collinear emission case:

$$R_{
m HH}^{
m FSR}(z) = \left. -z R_{
m HH}^{
m ISR} \left(rac{1}{z}
ight)
ight|_{\ln \Delta 
ightarrow \ln z; \ l_0 
ightarrow l_0 + 2 \ln z}$$



# QED Collinear Radiation Factors in NLO (3)

Emission of one collinear hard photon accompanied by one-loop soft and virtual correction (H(S+V)) is received using the NLO QED splitting functions

$$\begin{split} R_{\mathrm{H(S+V)}}^{\mathrm{ISR}}(z) \otimes \mathrm{d}\hat{\sigma}(z) &= \delta_{\mathrm{(S+V)}}^{(1)} R_{\mathrm{H}}^{\mathrm{ISR}}(z) \otimes \mathrm{d}\sigma^{(0)}(z) \\ &+ \left(\frac{\alpha}{2\pi}\right)^2 L \bigg[ 2\frac{1+z^2}{1-z} \bigg( \mathrm{Li}_2 \left(1-z\right) - \ln(1-z) \ln z \bigg) \\ &- (1+z) \ln^2 z + (1-z) \ln z + z \bigg] \otimes \mathrm{d}\sigma^{(0)}(z), \\ \delta_{\mathrm{(S+V)}}^{(1)} &= \frac{\mathrm{d}\sigma_{\mathrm{Soft}}^{(1)} + \mathrm{d}\sigma_{\mathrm{Virt}}^{(1)}}{\mathrm{d}\sigma^{(0)}} \,, \end{split}$$

where  $\sigma^{(0)}(z)$  is the boosted Born cross section, and  $\delta^{(1)}_{(\mathrm{S+V})}$  is the relative  $\mathcal{O}\left(\alpha\right)$  Soft + Virtual radiative correction with  $E_{\gamma}^{\mathrm{Soft}}<\Delta E$ . The corresponding FSR factor is received again using the Gribov-Lipatov relation.

#### **Pair Corrections**

Leptonic and hadronic pair corrections are important for a number of precision observables. Exclusive treatment here is of ultimate importance. Monte Carlo has to be used for real or for hard pairs, then soft and virtual ones can be treated analytically (semi-analytically for the hadronic case).

Singlet and non-singlet NLO pair contributions in  $\mathcal{O}\left(\alpha^2L\right)$  to inclusive observables can be described within the QCD-like factorization approach.

But if we have a MC for hard pairs, we can extract analytically the soft+virtual part, so that

$$\mathrm{d}\sigma_\mathrm{pair}^{(2)} = \mathrm{d}\sigma_\mathrm{H~pair}^{(2)MC} \ + \ \mathrm{d}\sigma^{(0)} \times \delta_\mathrm{S+V~pair}^{(2)}$$

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- Negatively weighted events within this approach are possible but not numerous



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- MC integrator and generator for Bhabha scattering is under development (upgrade of SAMBHA MC)

