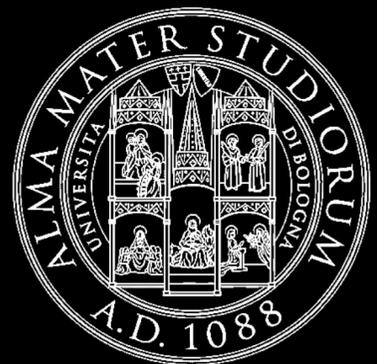
The background of the slide is a vertical strip on the left side showing a cosmological simulation. It features a complex, filamentary structure of matter and dark matter, with a color gradient from dark blue to bright yellow and orange, indicating density variations. The right side of the slide is black, providing a high-contrast background for the white and blue text.

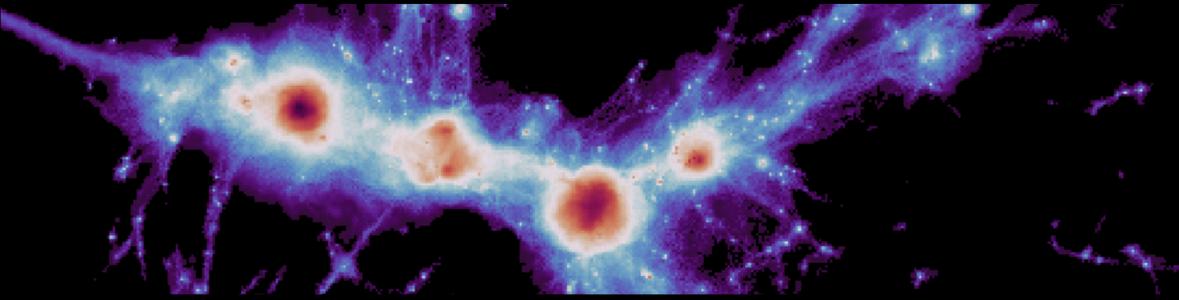
COSMOLOGICAL SIMULATIONS OF SIDM WITH THE TNG MODEL

Giulia Despali

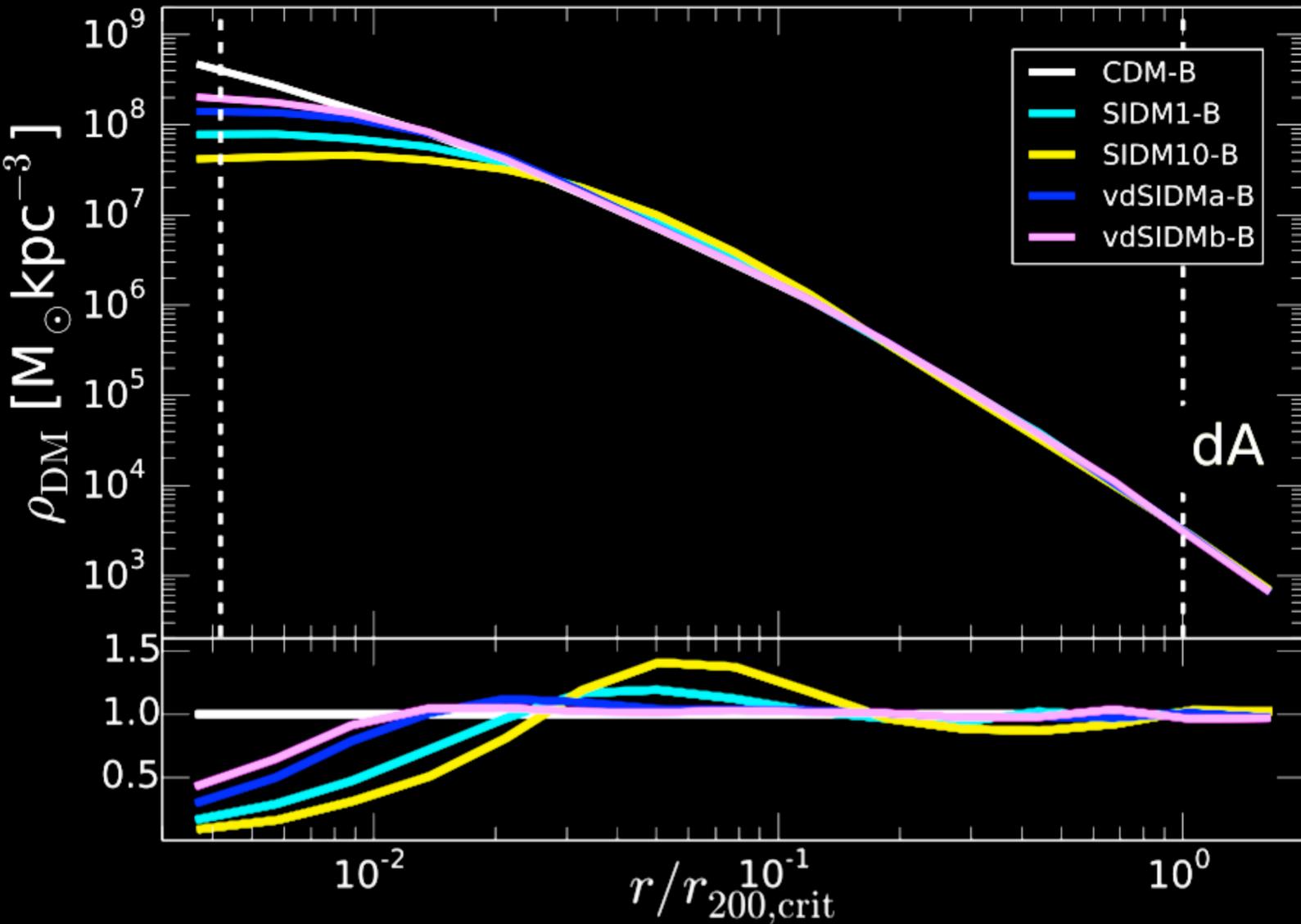
University of Bologna

Simona Vegetti, Mark Vogelsberger,
Lauro Moscardini, Massimo Meneghetti,
Annalisa Pillepich, Dylan Nelson, Claudio
Mastromarino

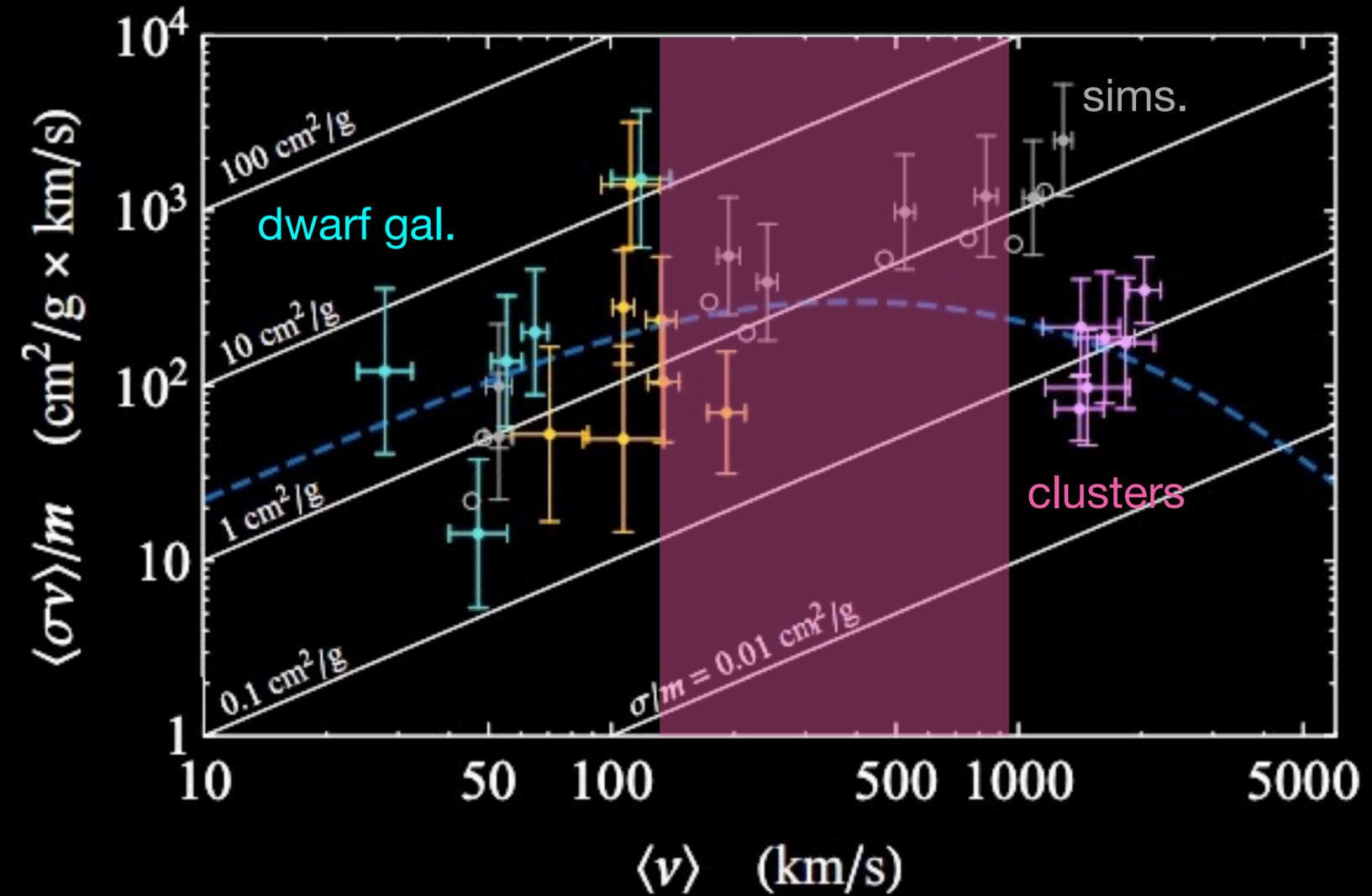




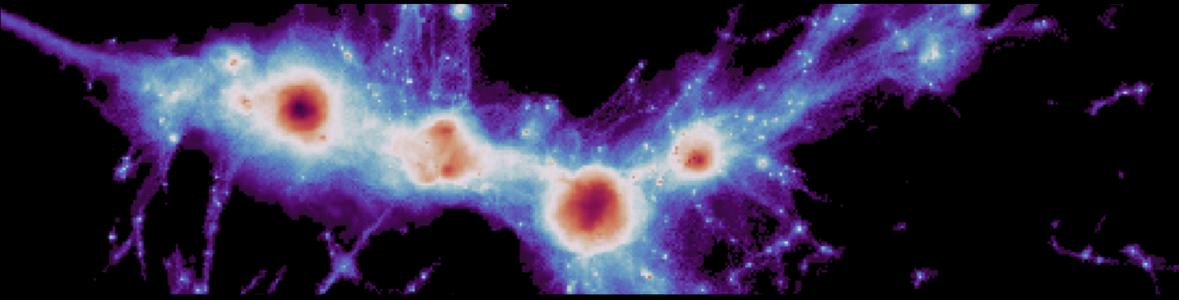
SELF-INTERACTING DARK MATTER



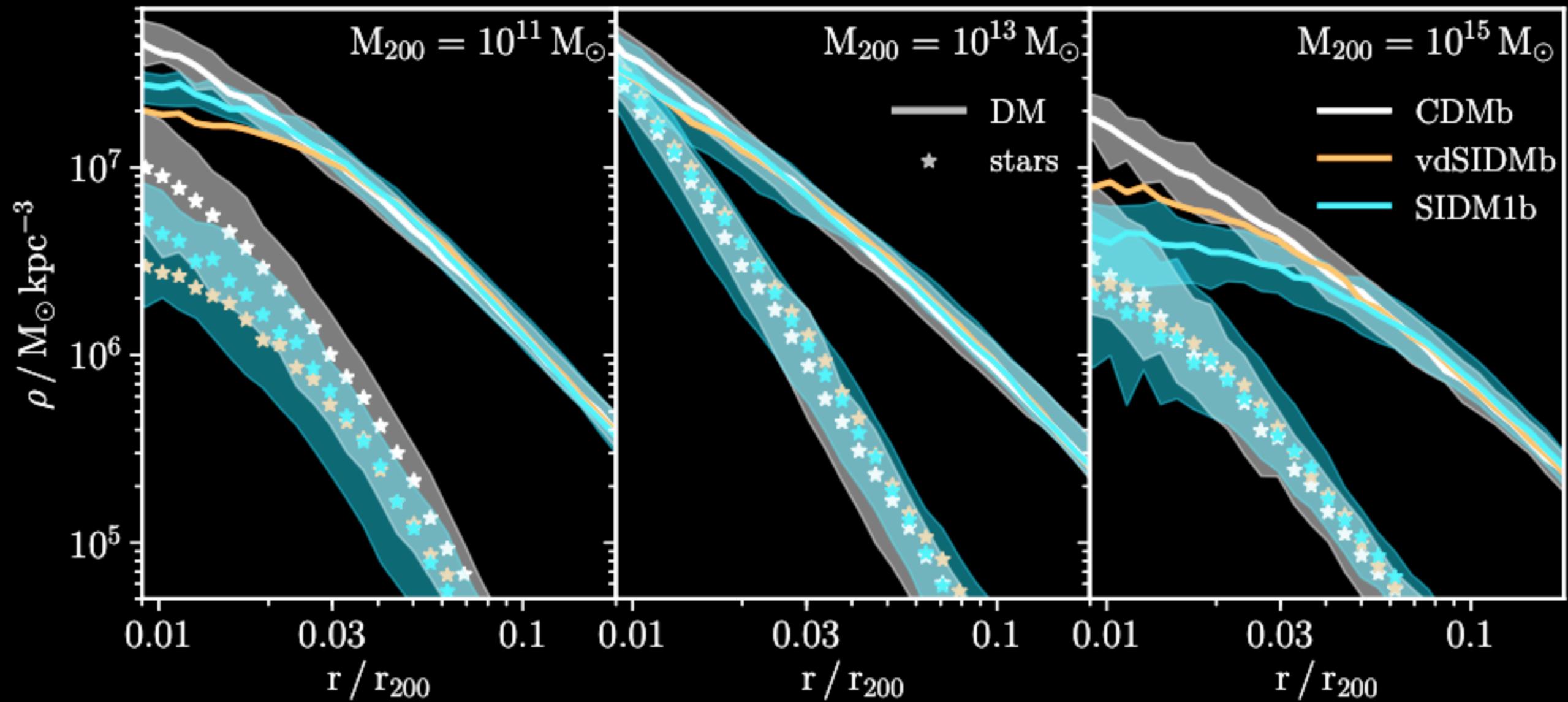
(Vogelsberger et al. 2016)



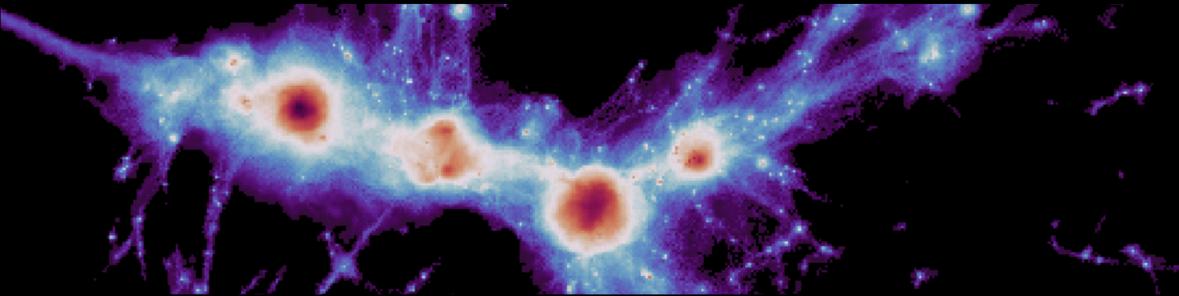
(Kalpinghat et al. 2015)



SIDM HYDRO SIMULATIONS



(Robertson et al. 2021)

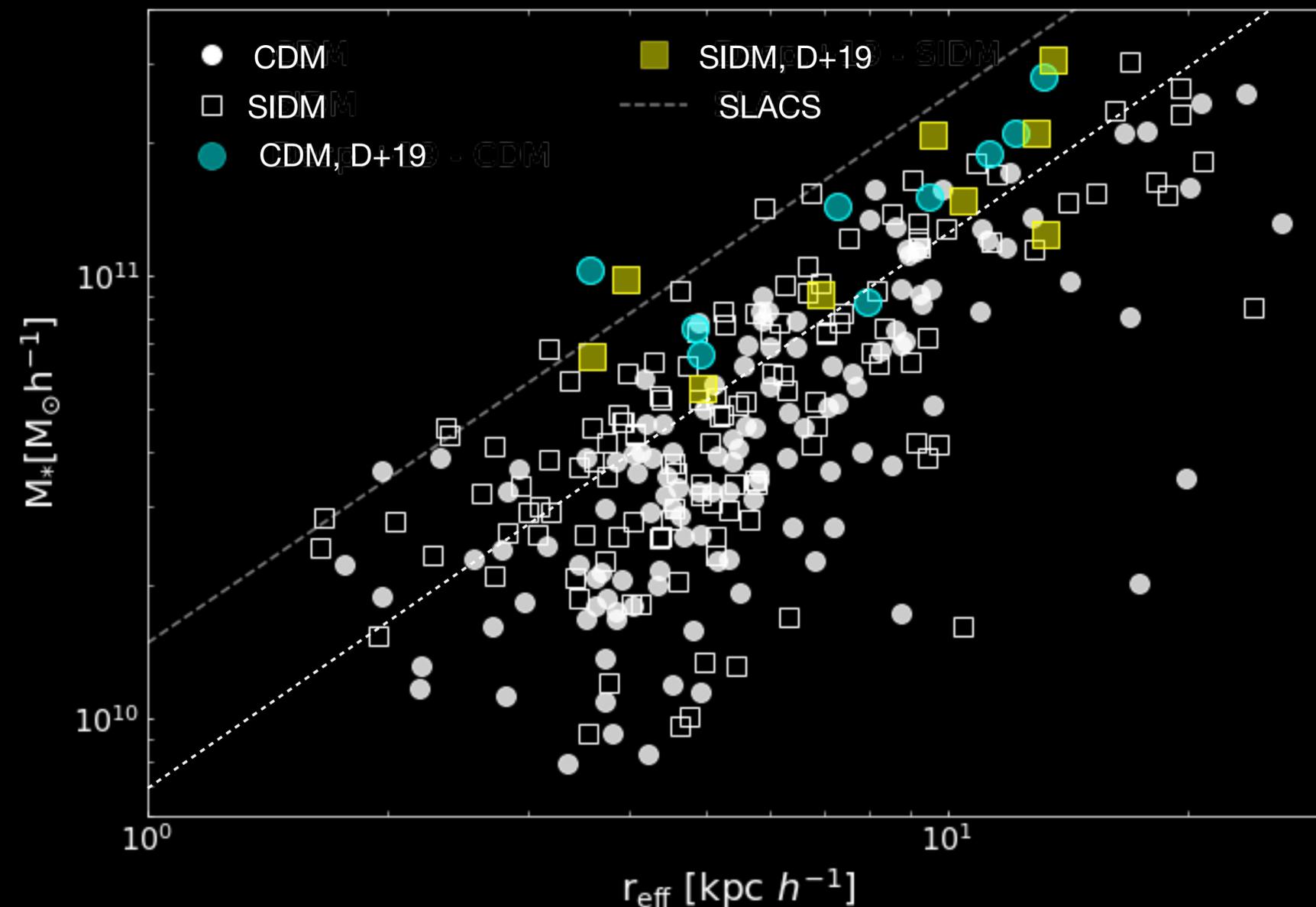


SIDM HYDRO SIMULATIONS

- constant cross section $\sigma=1 \text{ cm}^2/\text{g}$

Selection:

- elliptical galaxies $\log(M) \sim 12.5-13.4 \text{ Msun}$
- analogues of lens galaxies
- redshift $z = 0.2, 0.5, 1$

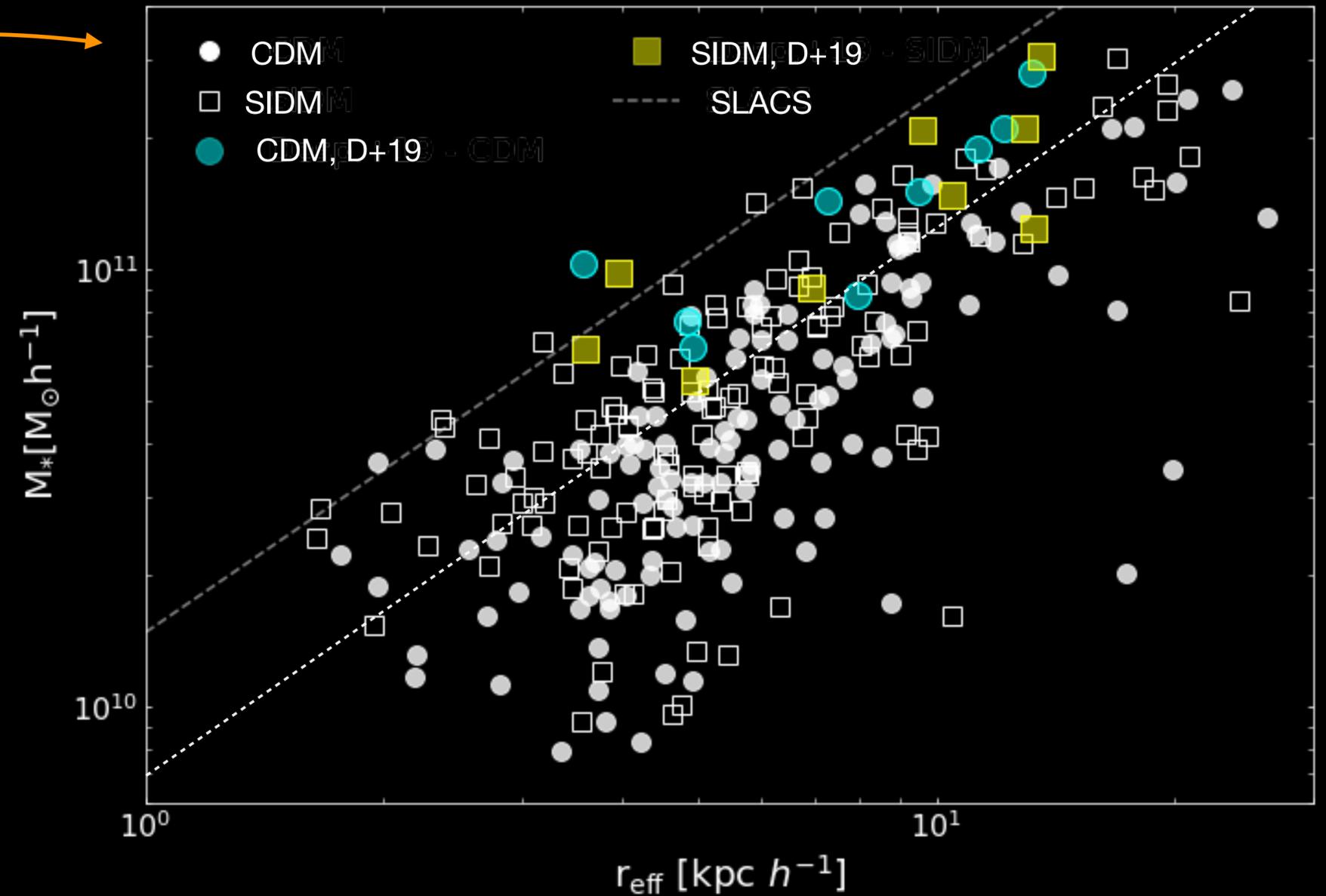
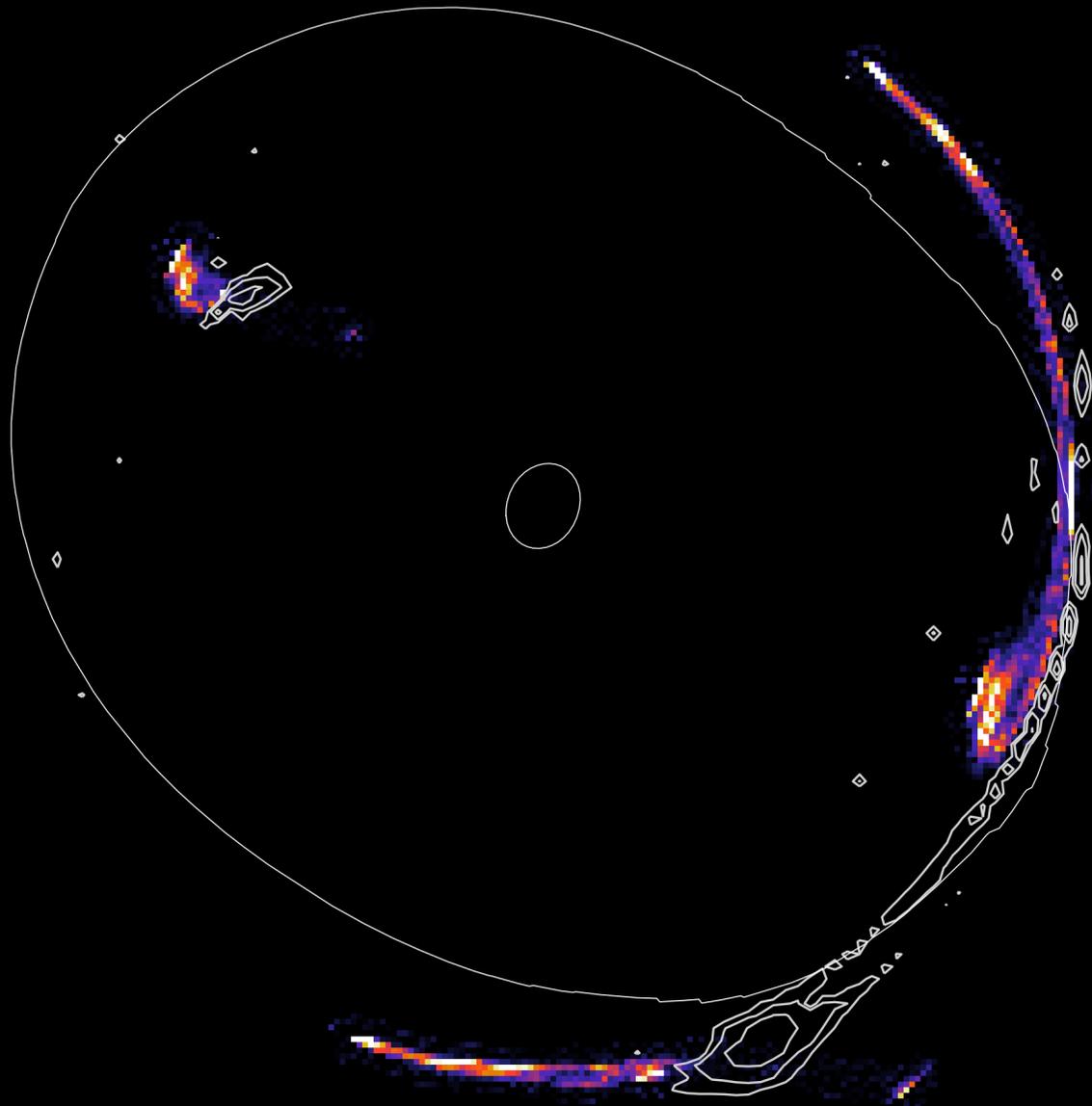


■ Eagle50 SIDM (Robertson+20)

■ ■ Zoom-in TNG+SIDM (Despali+19,22)

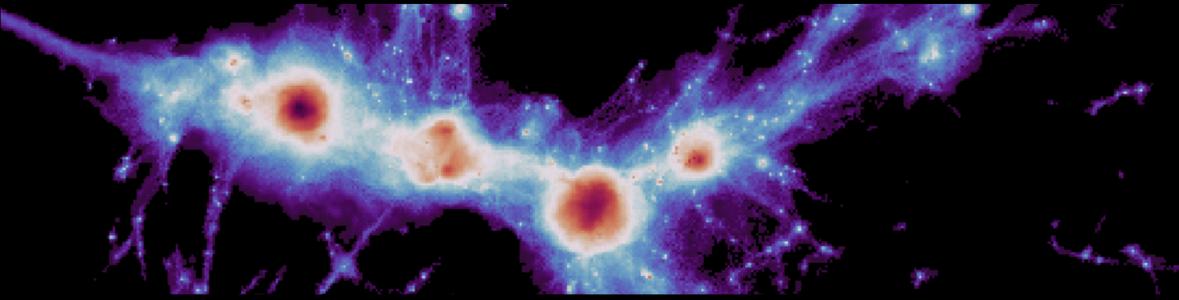
SIDM HYDRO SIMULATIONS

Elliptical (lens) galaxies



■ Eagle50 SIDM (Robertson+20)

■ ■ Zoom-in TNG+SIDM (Despali+19,22)



SIDM HYDRO SIMULATIONS

$\log(M) \sim 12-12.5$

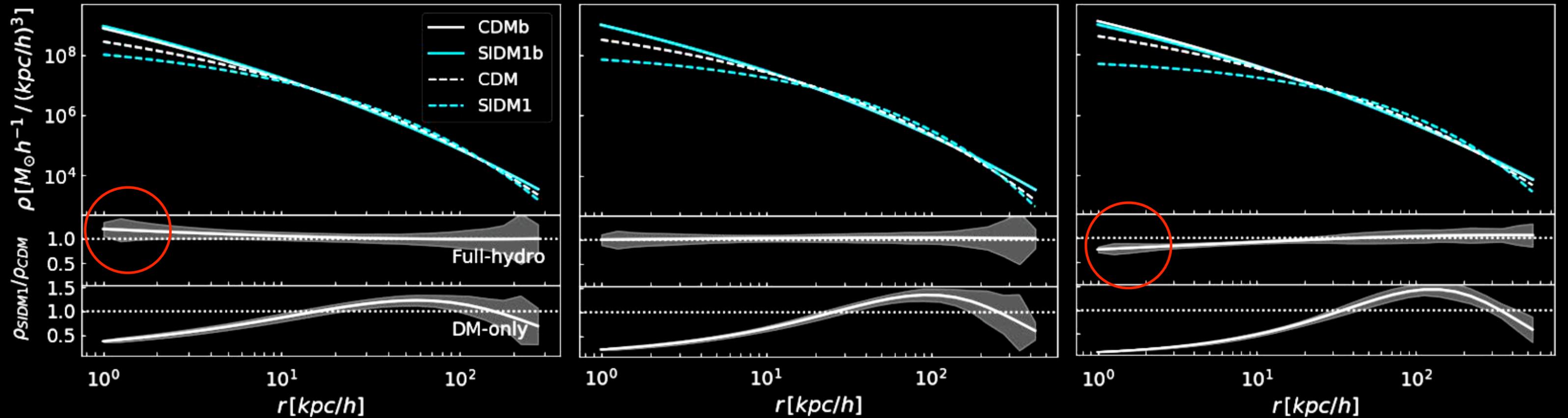
$\log(M) \sim 12.5-13$

$\log(M) \sim 13-13.5$

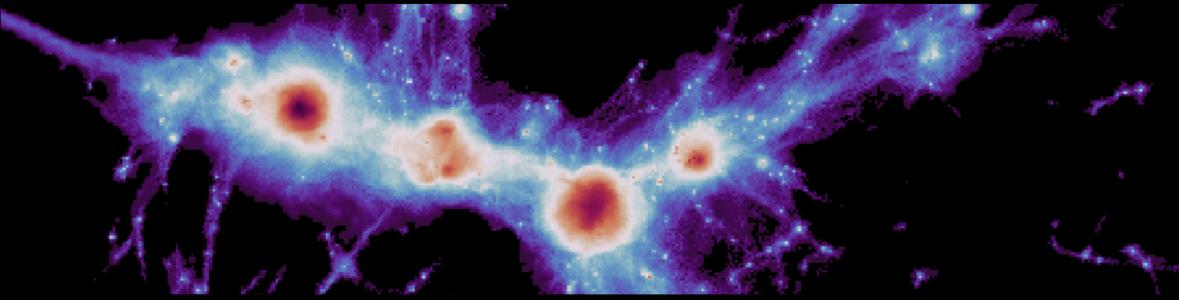
Group 1

Group 2

Group 3

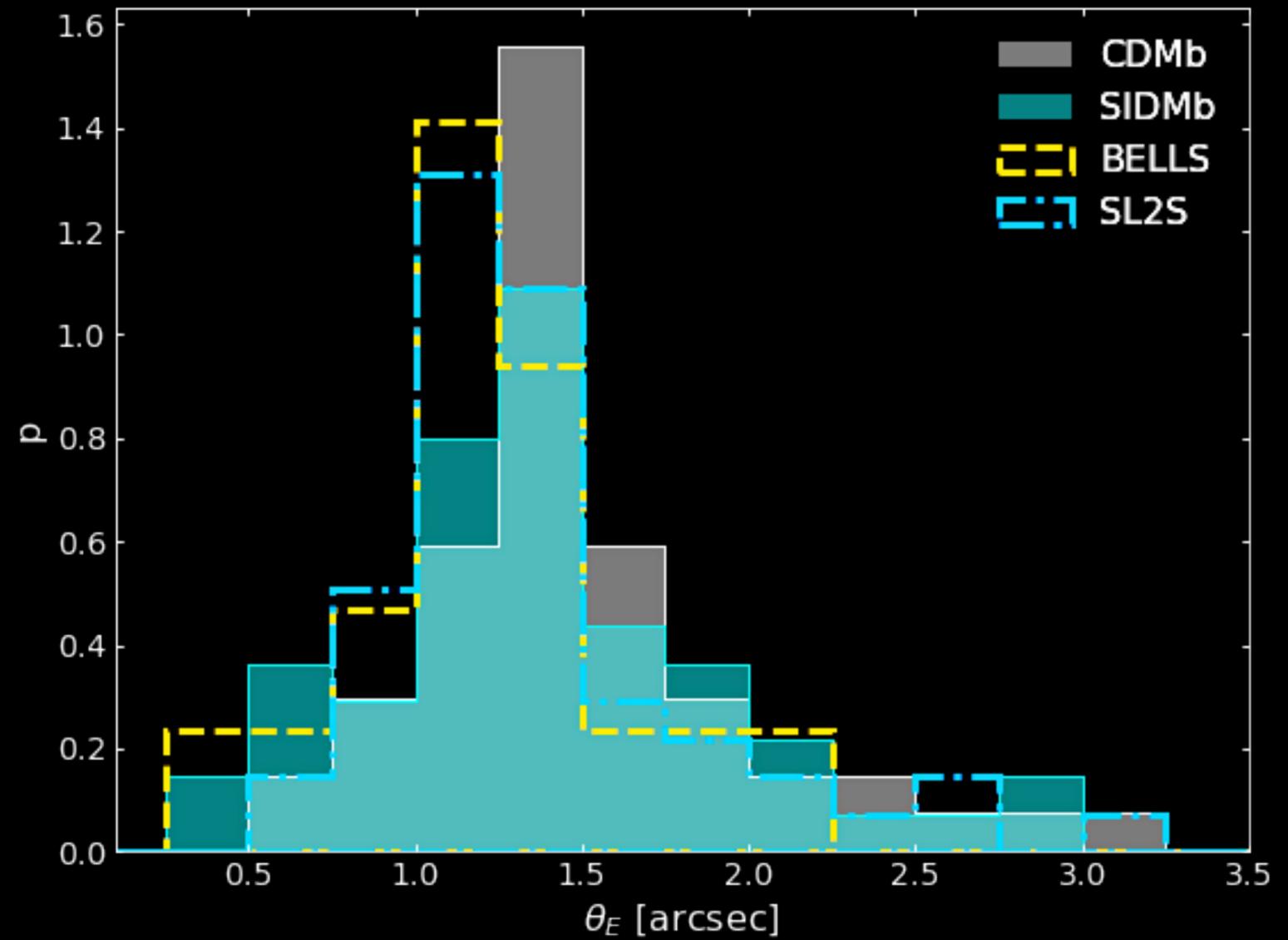
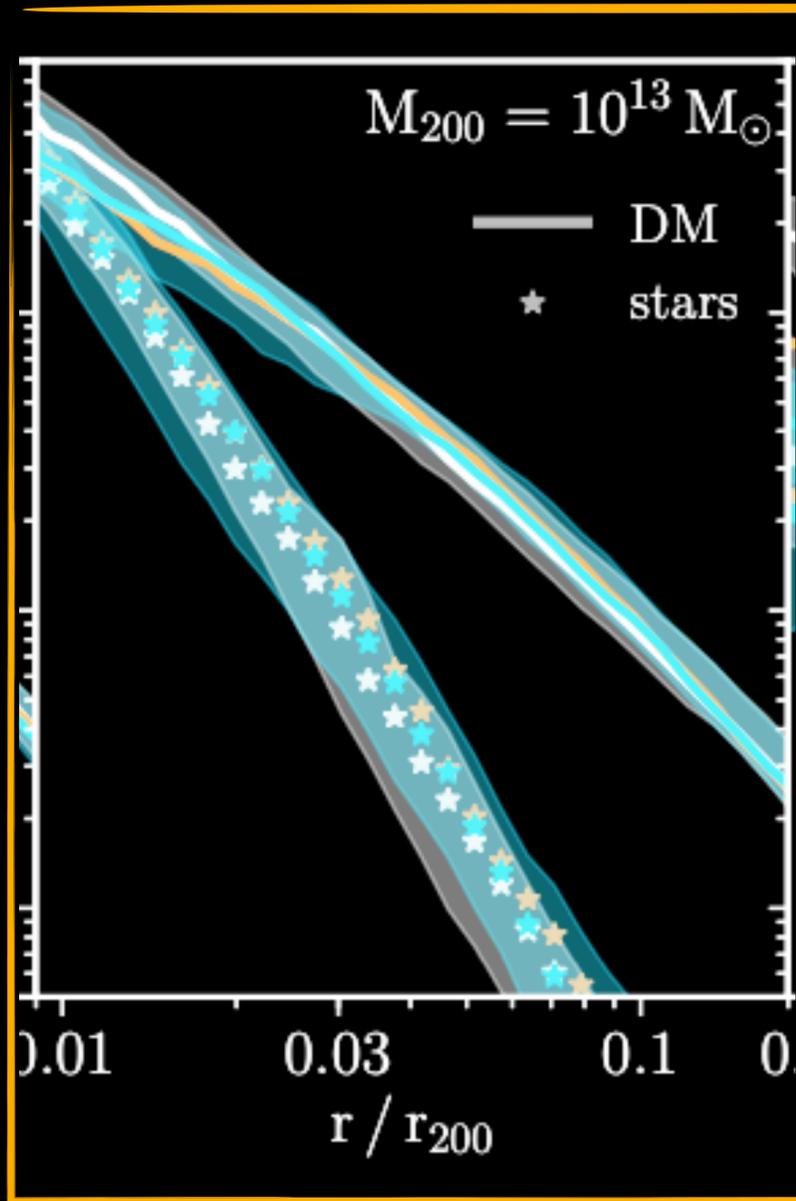


(Mastromarino, Despali et al. 2023)

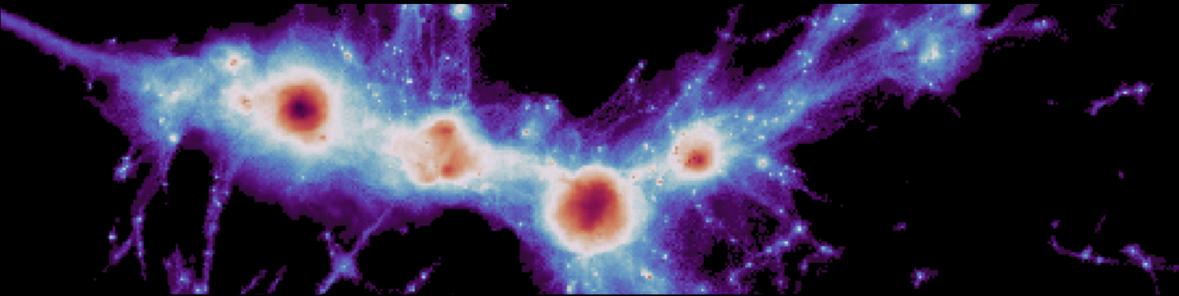


SIDM + BARYONS: LENSING EFFECT

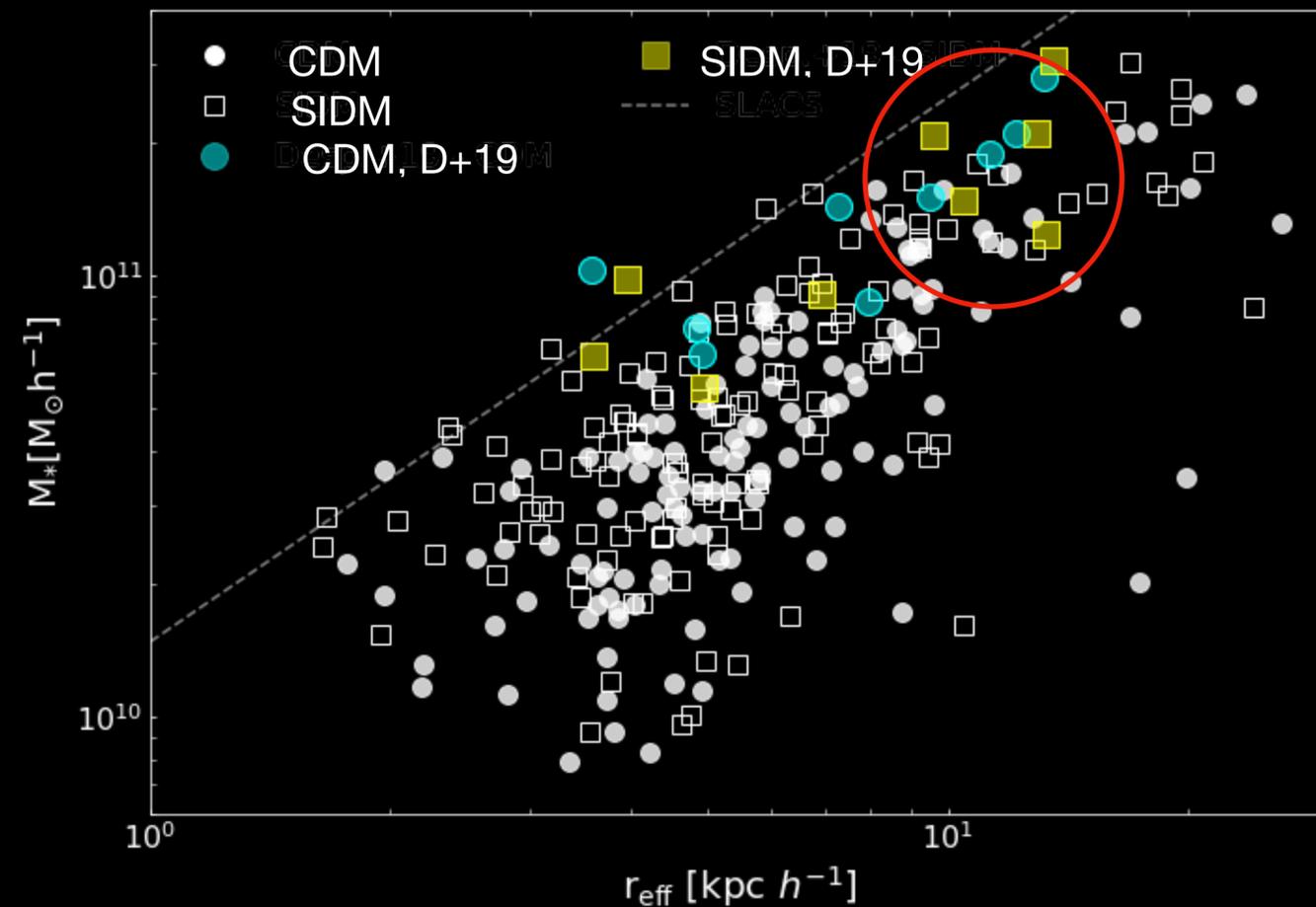
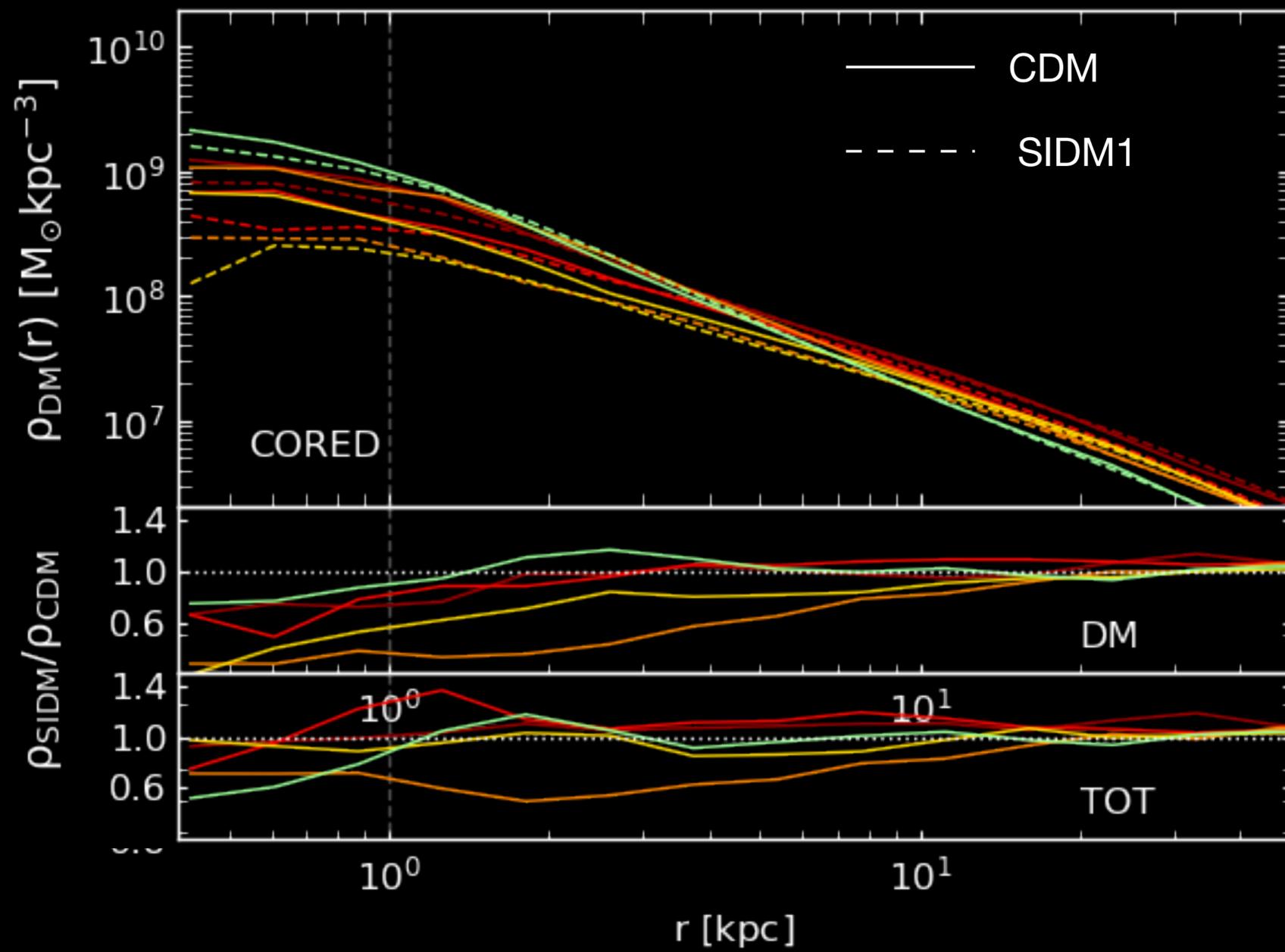
similar distribution of Einstein radii



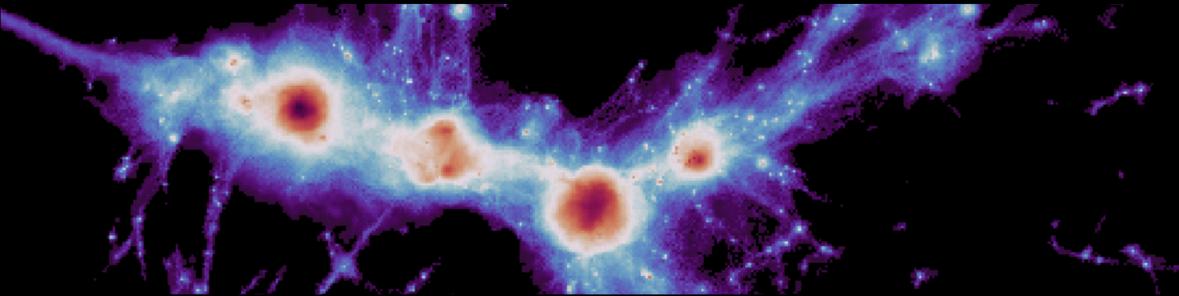
(Mastromarino, Despali et al. 2023)



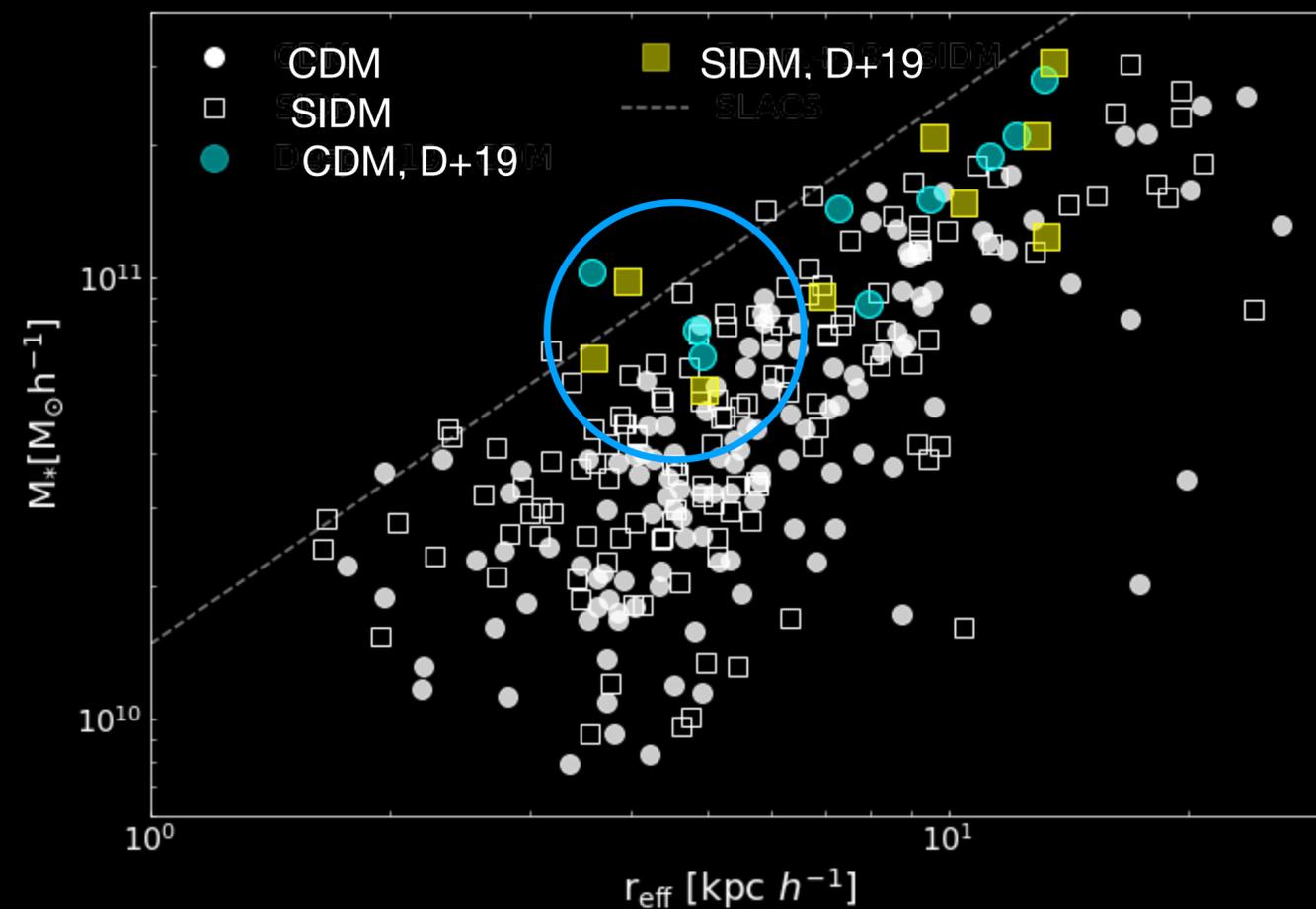
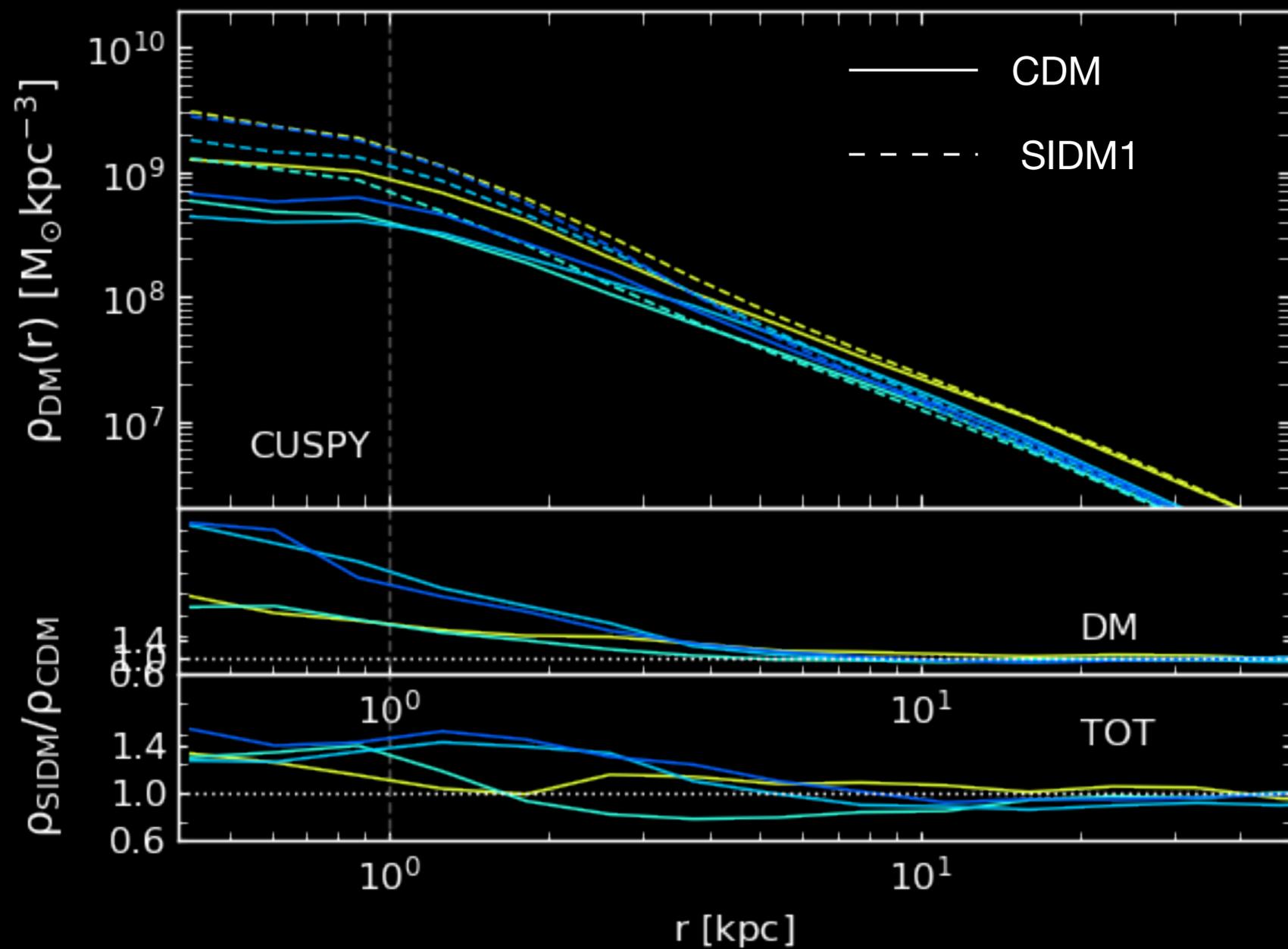
SIDM HYDRO SIMULATIONS



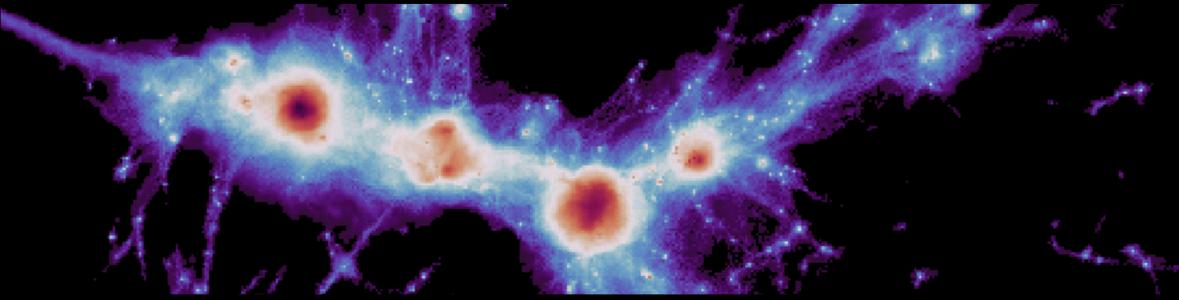
- Eagle50 SIDM (Robertson+20)
- Zoom-in TNG+SIDM (Despali+19,22)



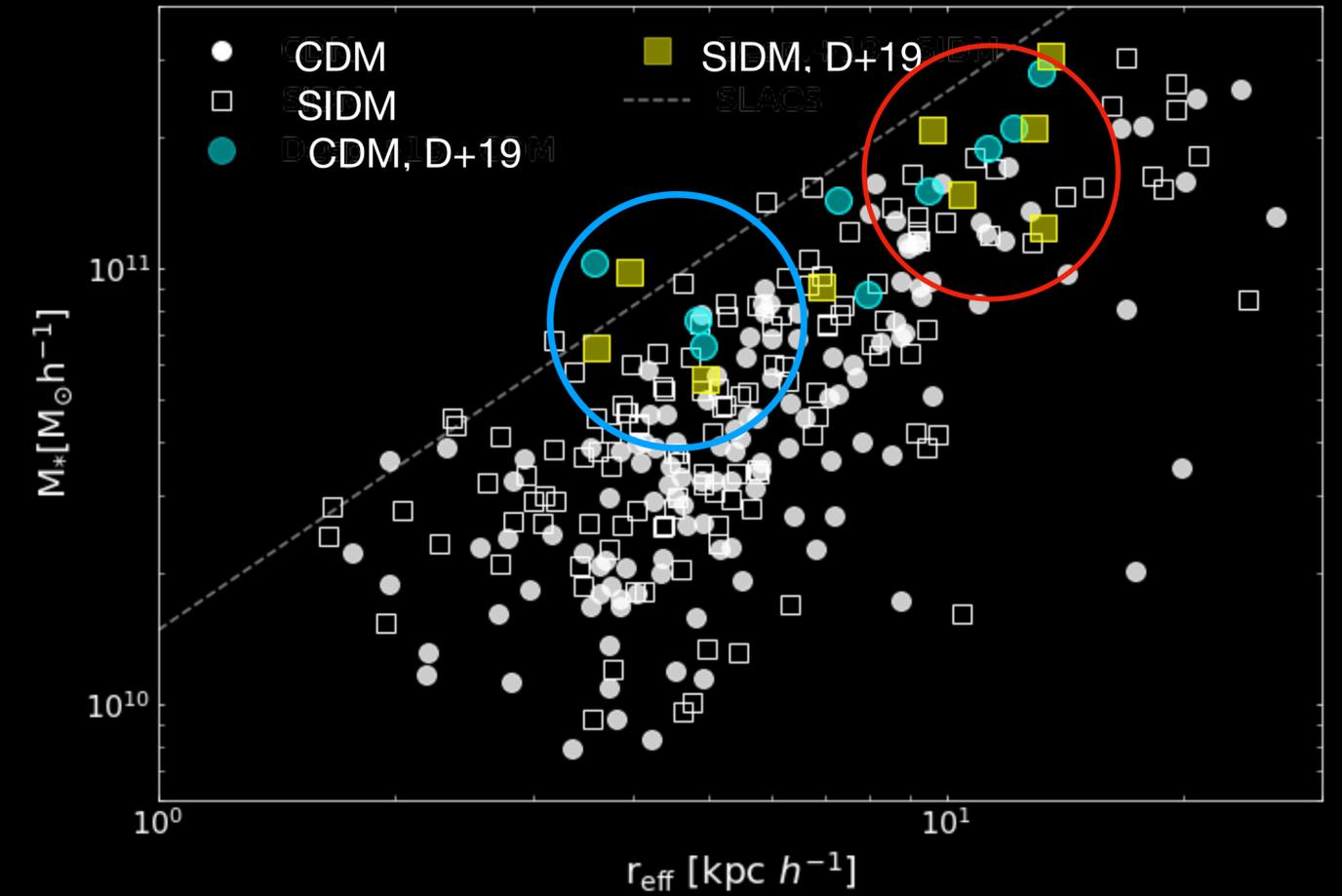
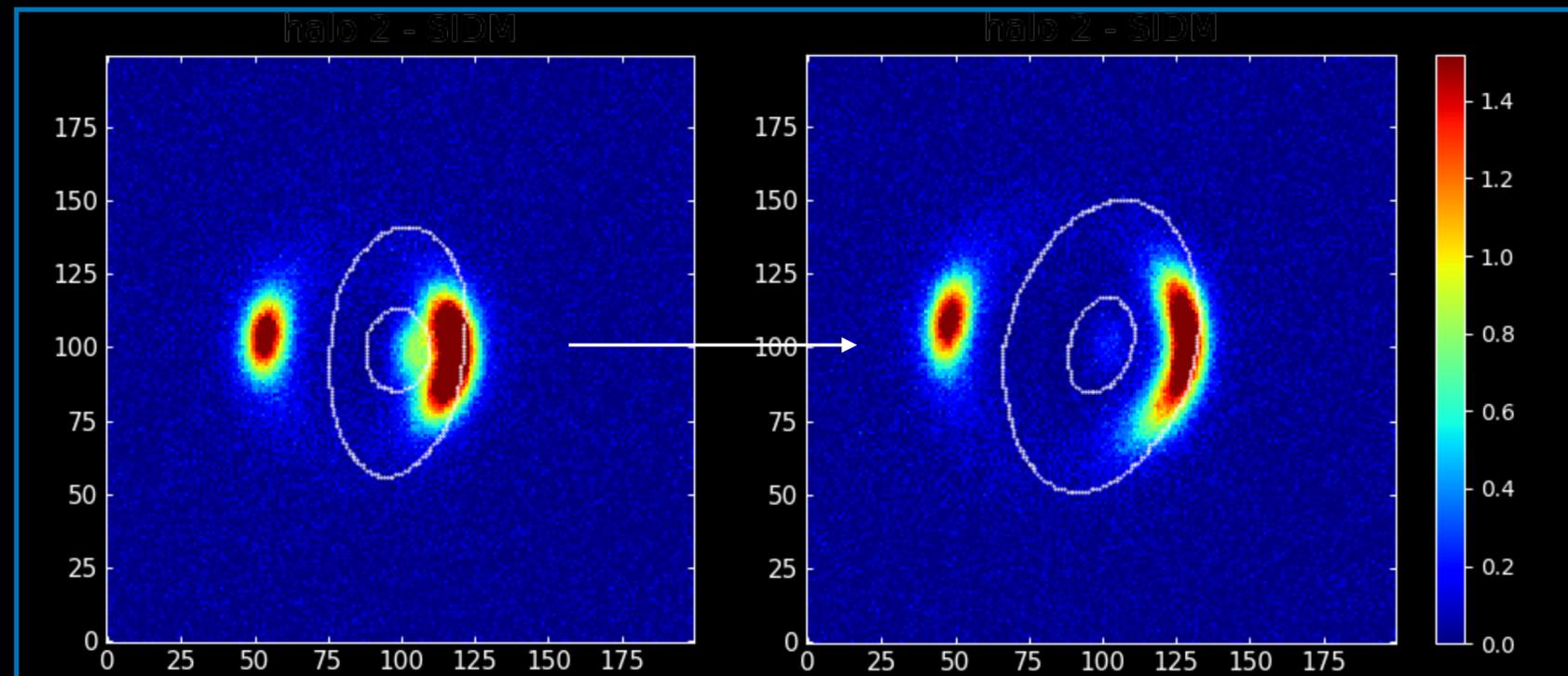
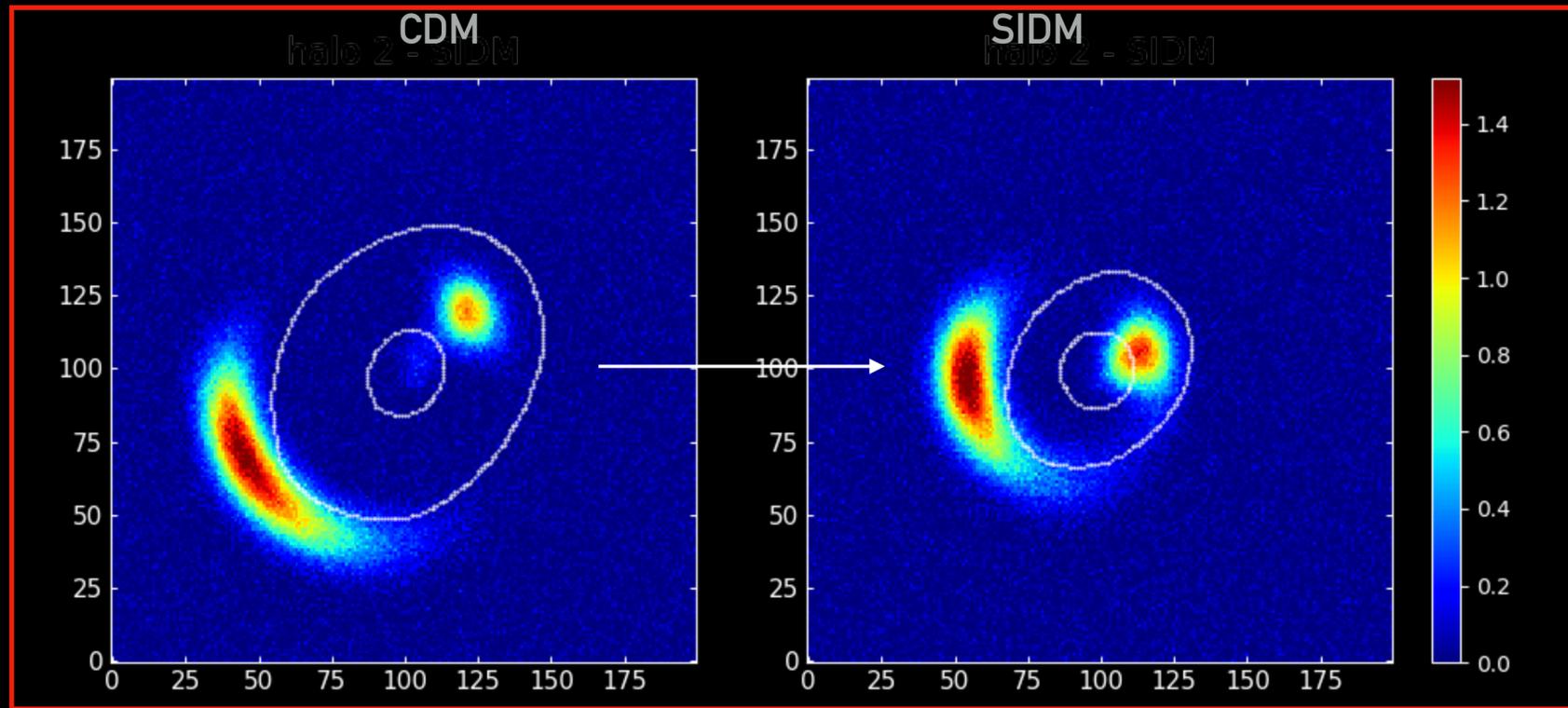
SIDM HYDRO SIMULATIONS



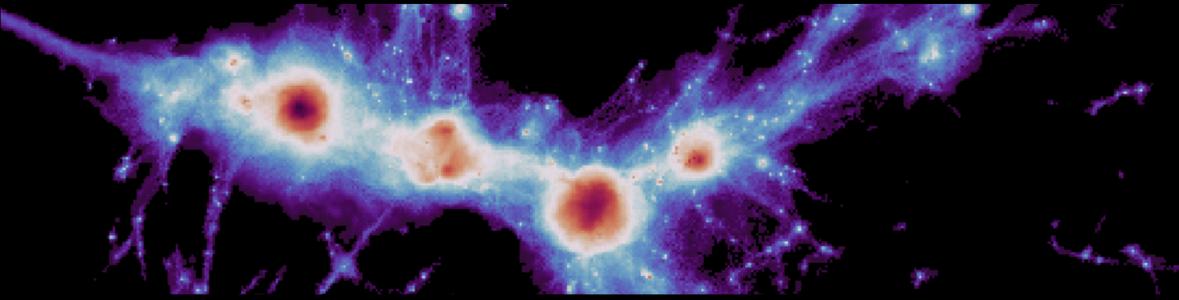
- Eagle50 SIDM (Robertson+20)
- Zoom-in TNG+SIDM (Despali+19,22)



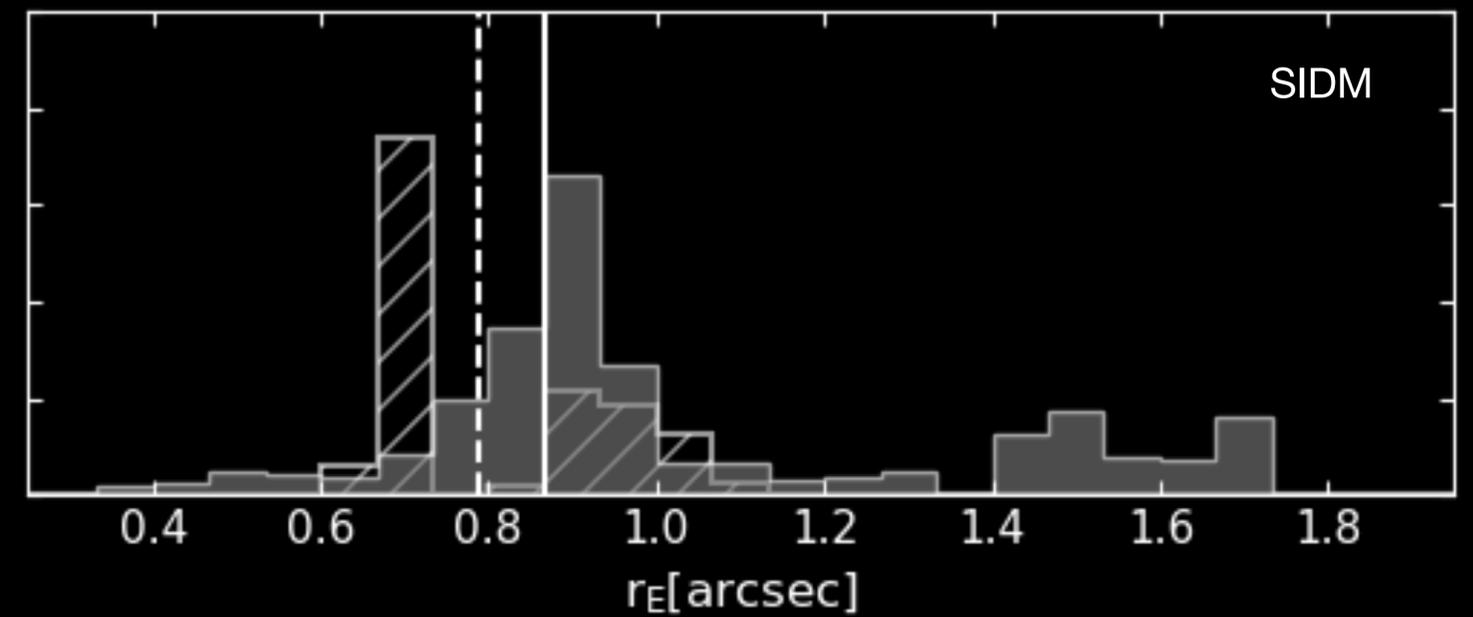
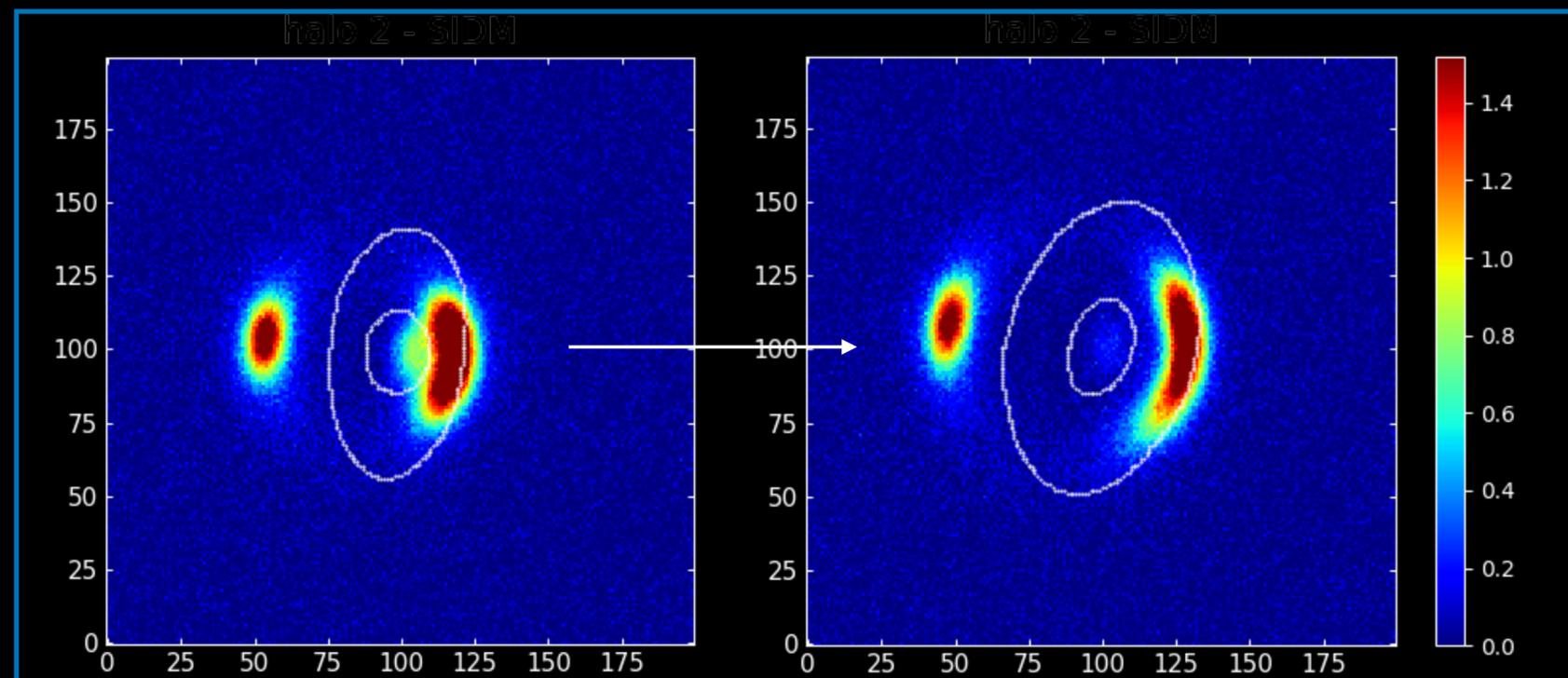
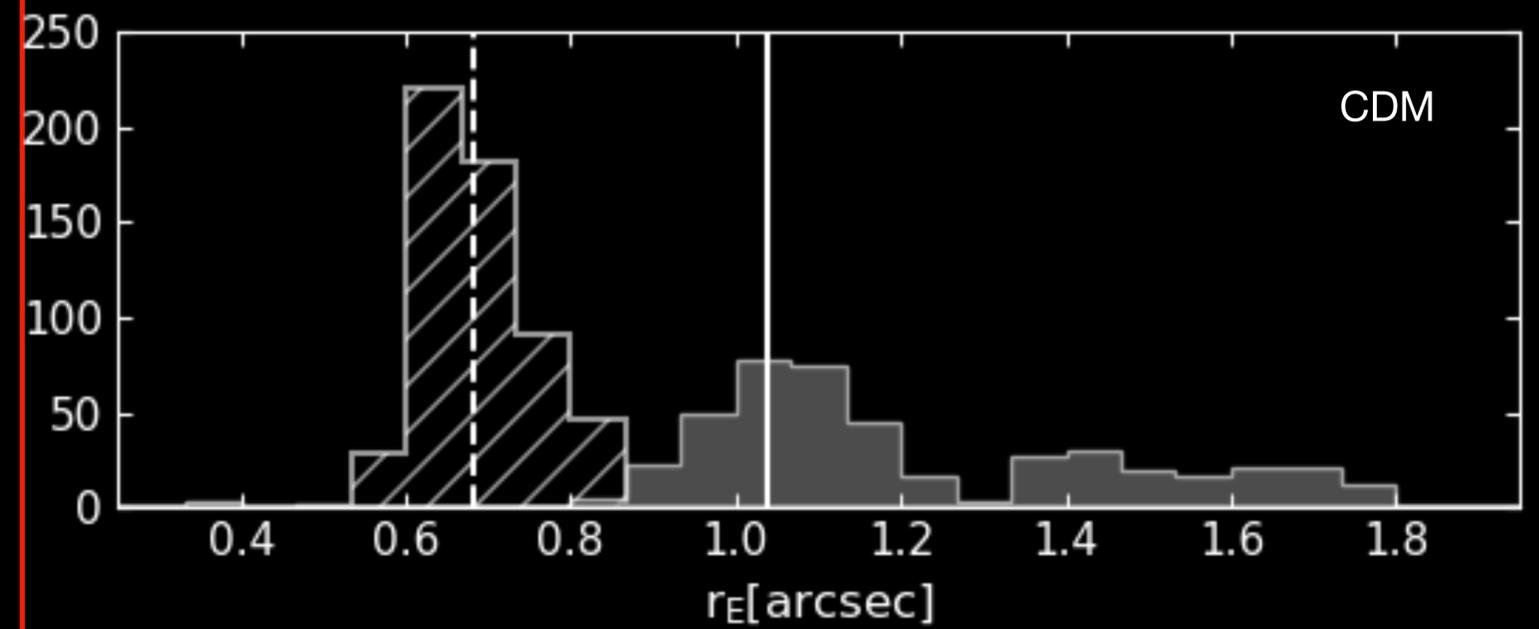
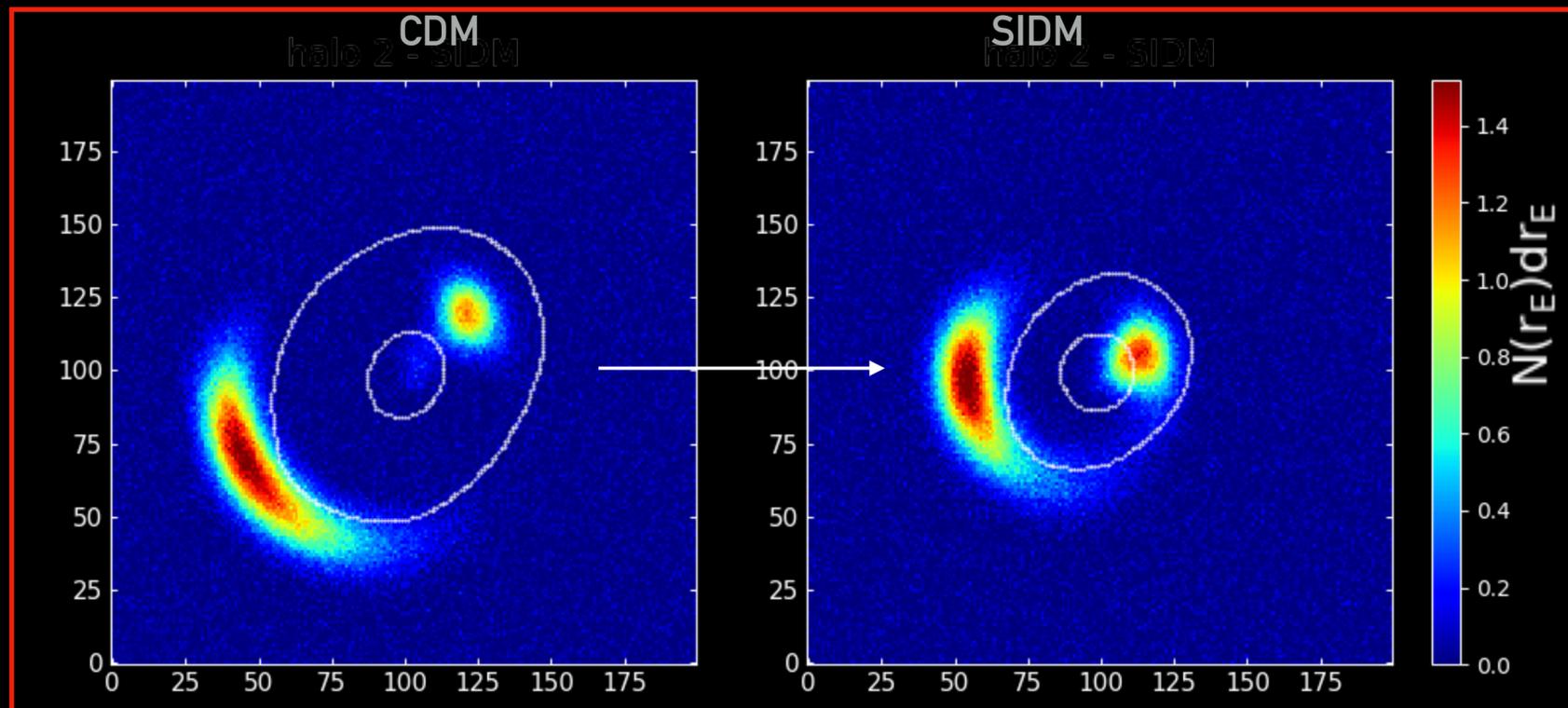
LENSING EFFECT



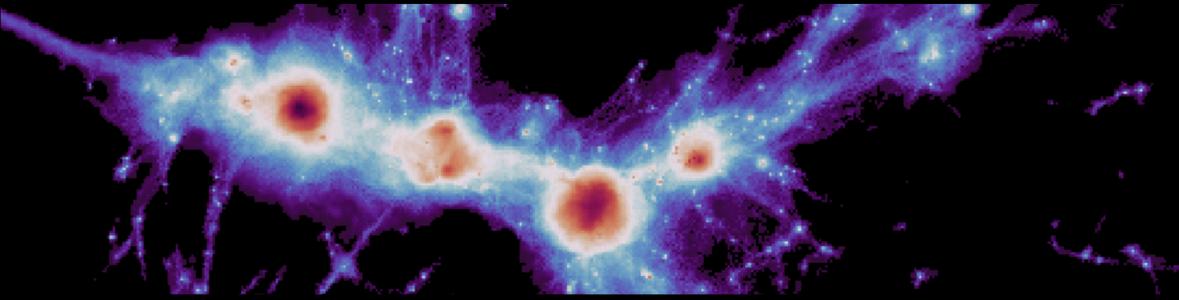
Zoom-in TNG+SIDM (Despali+19,22)



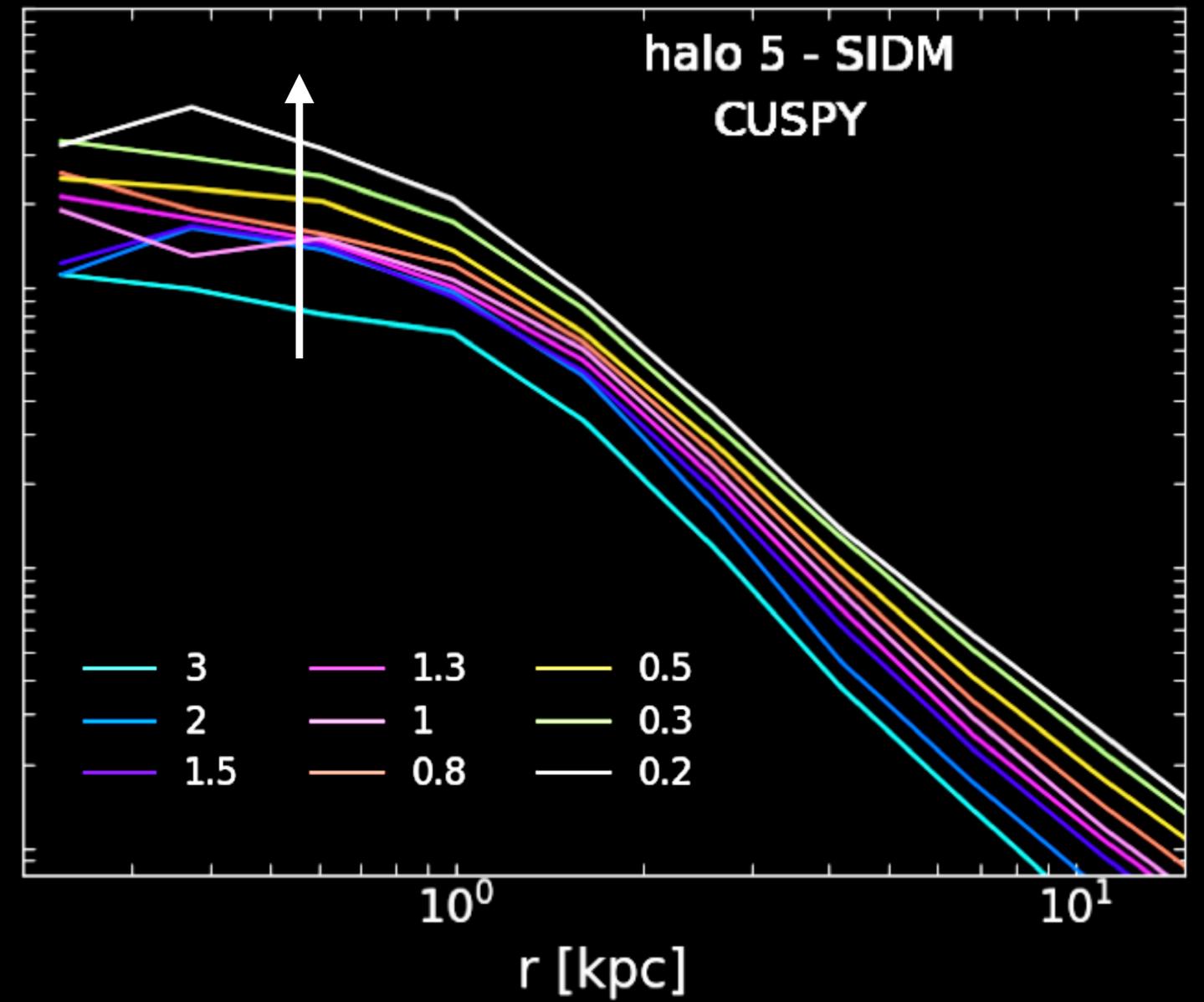
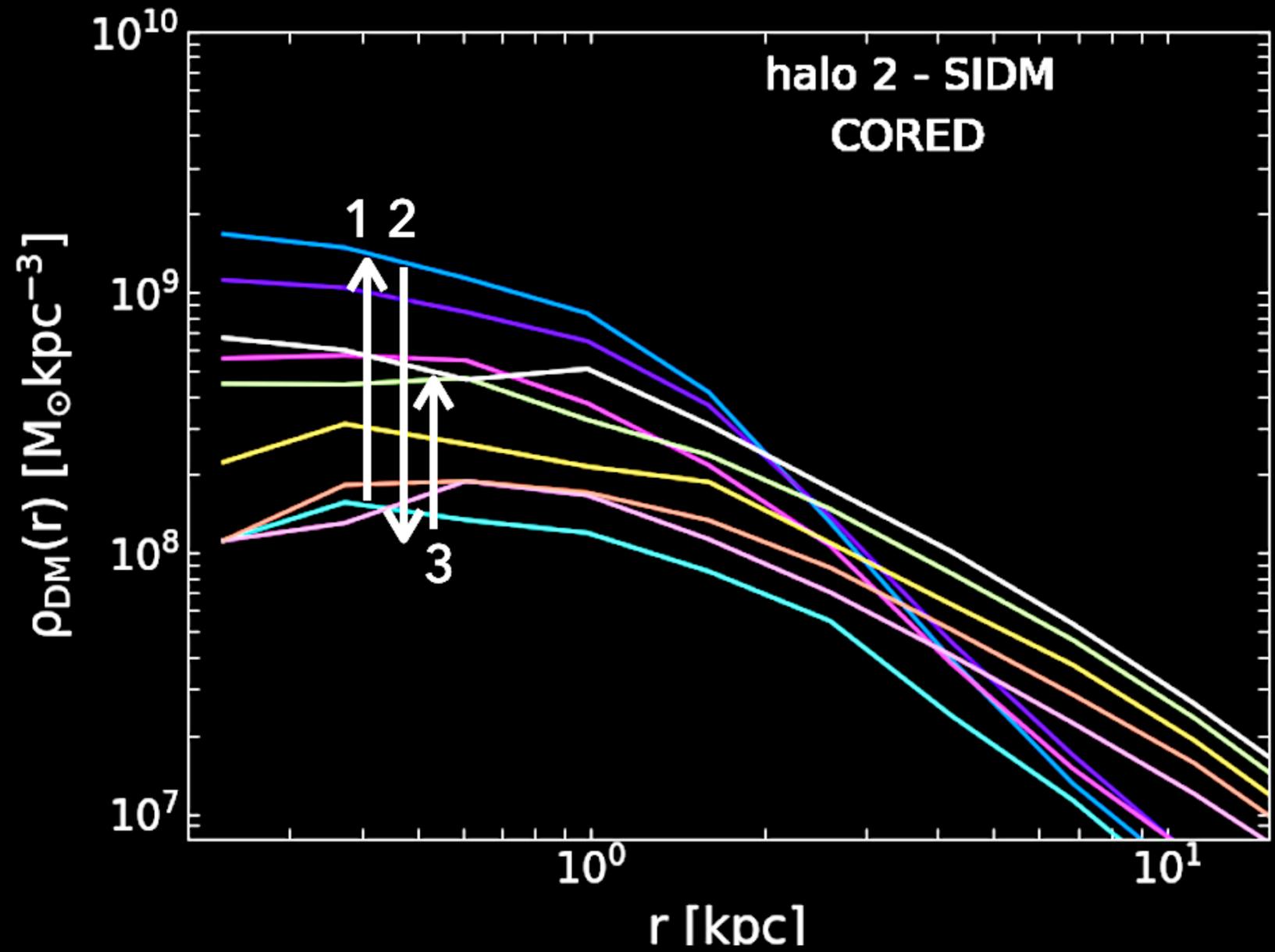
LENSING EFFECT

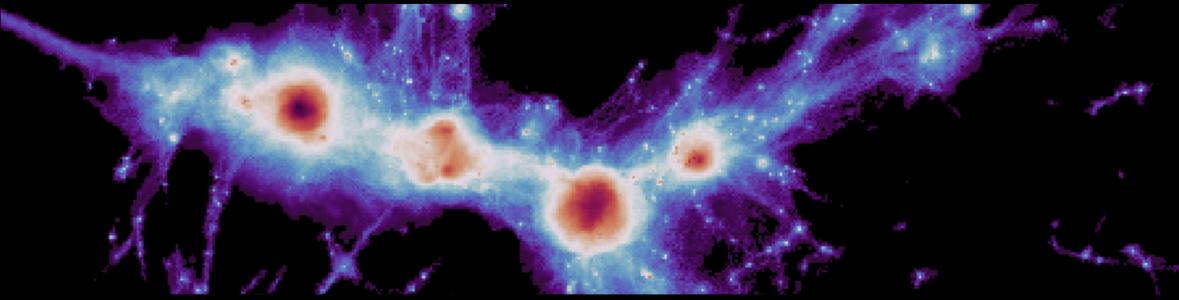


Zoom-in TNG+SIDM (Despali+19,22)



SIDM HYDRO SIMULATIONS

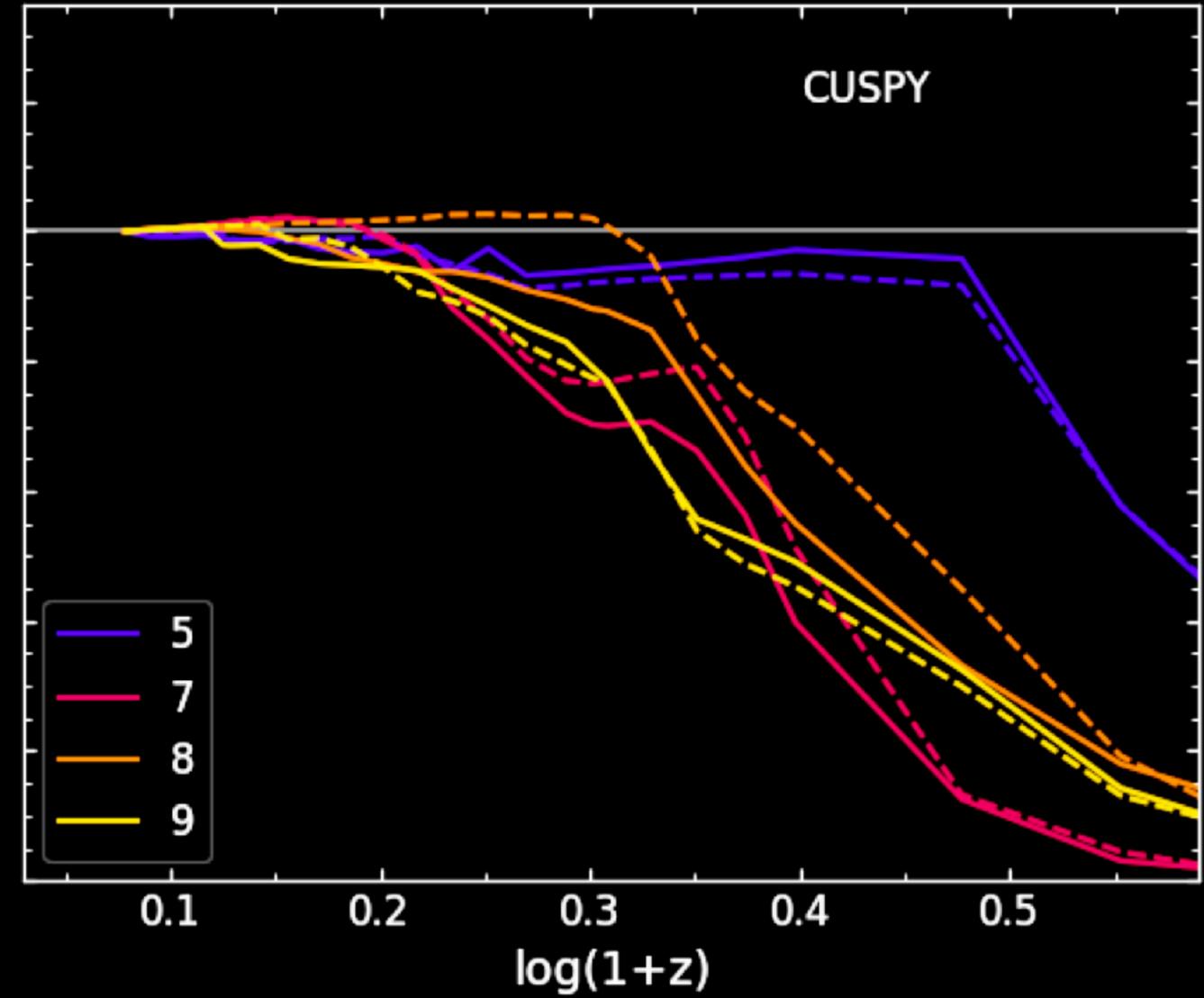
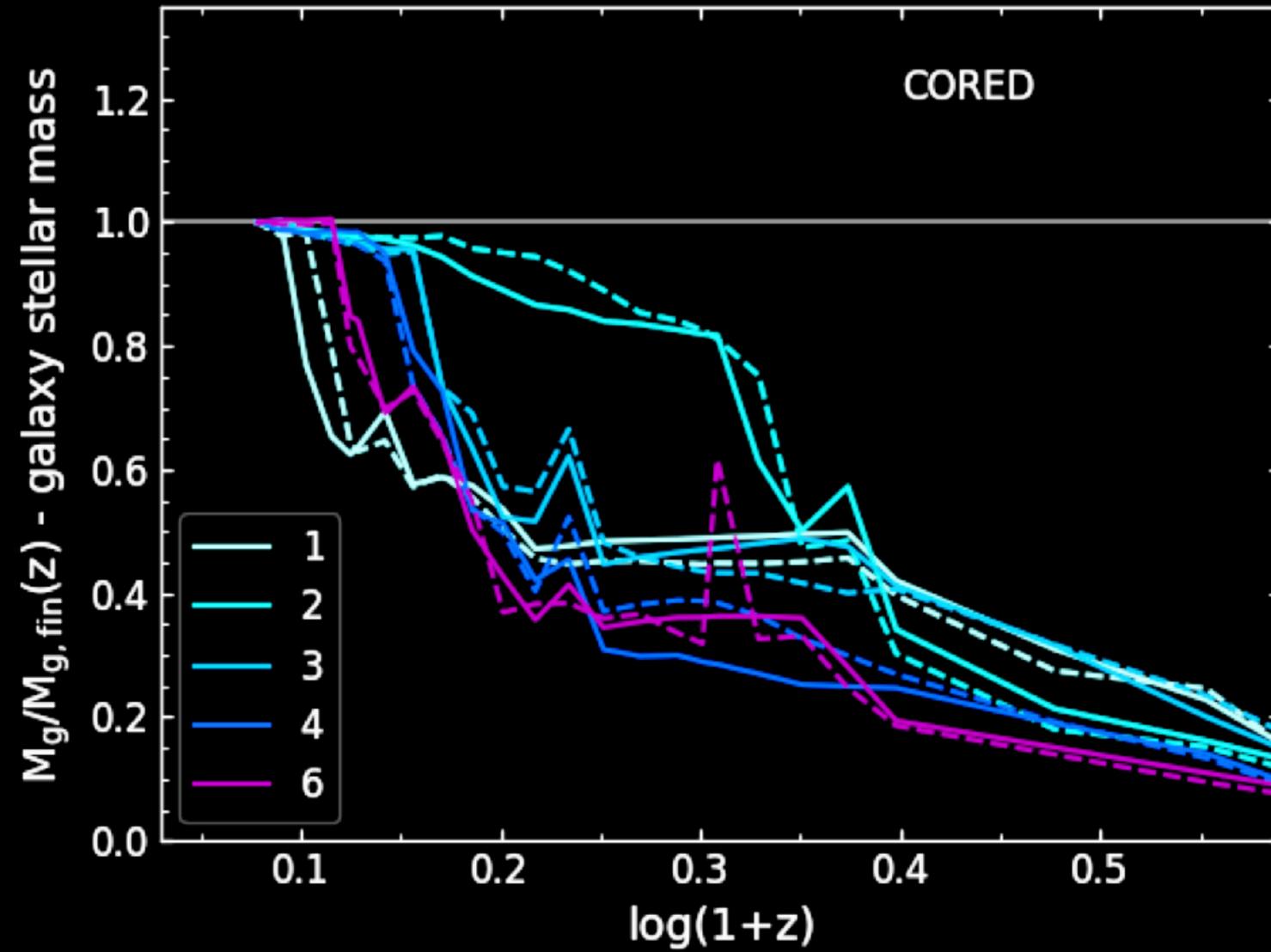


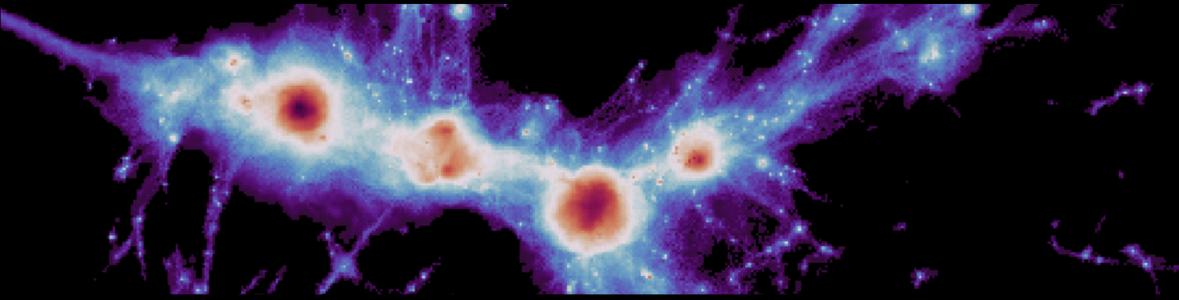


SIDM HYDRO SIMULATIONS

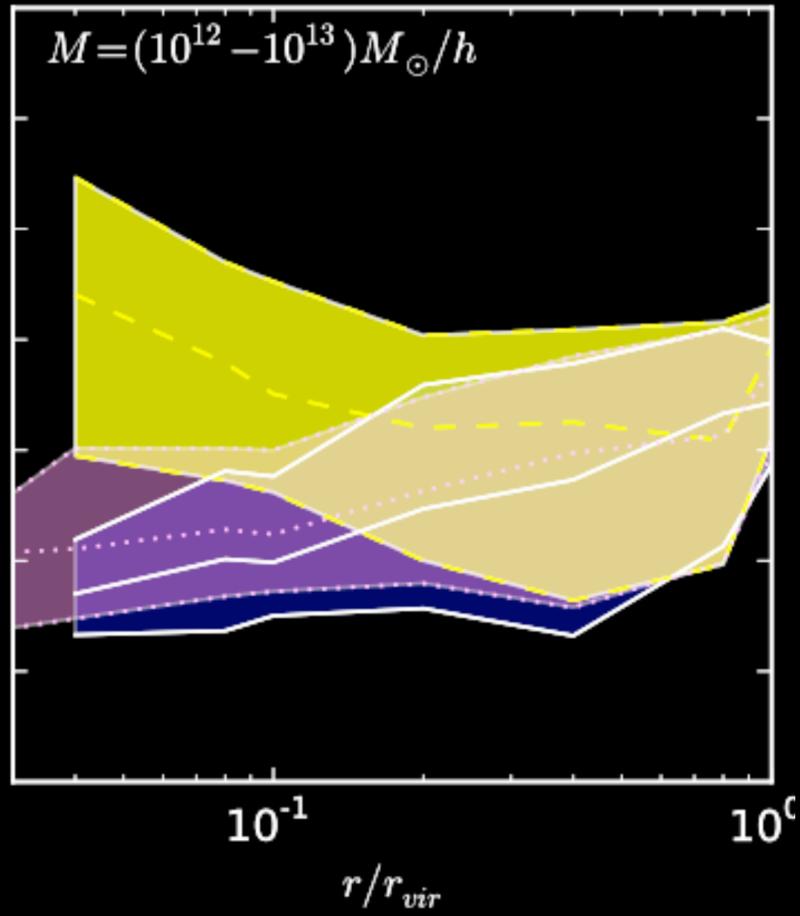
$z_f < 1$

$z_f > 1$



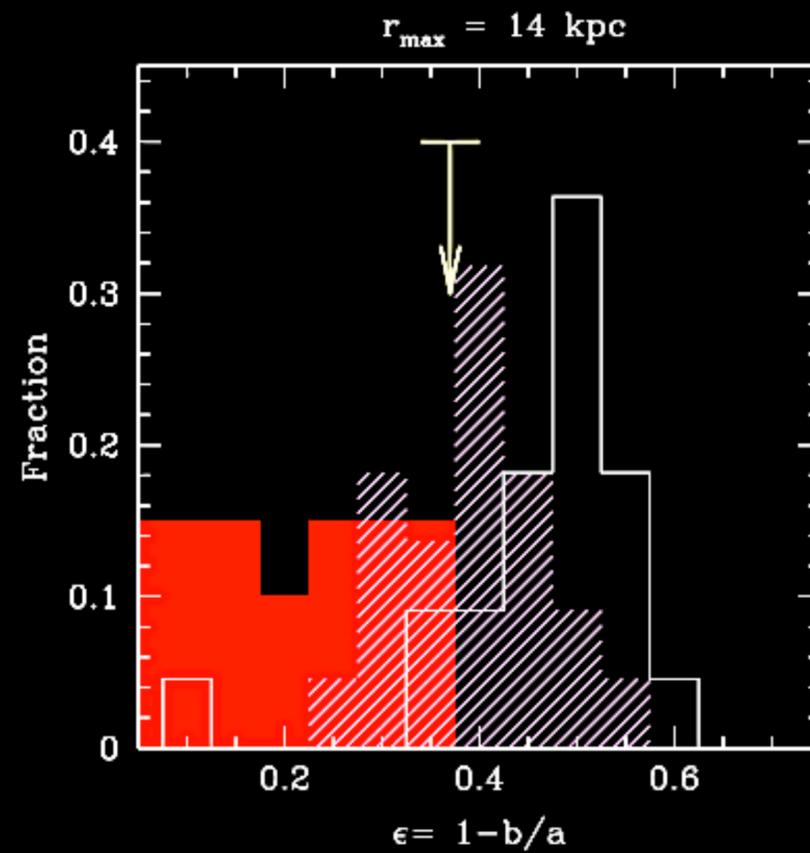


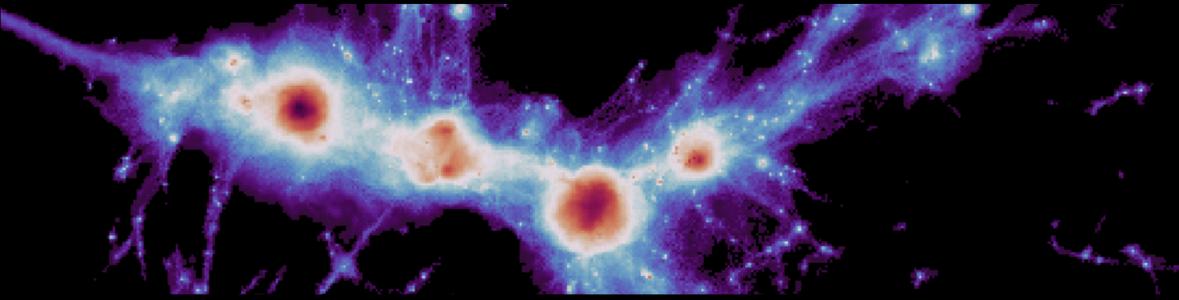
SHAPES



SIDM produces
rounder haloes

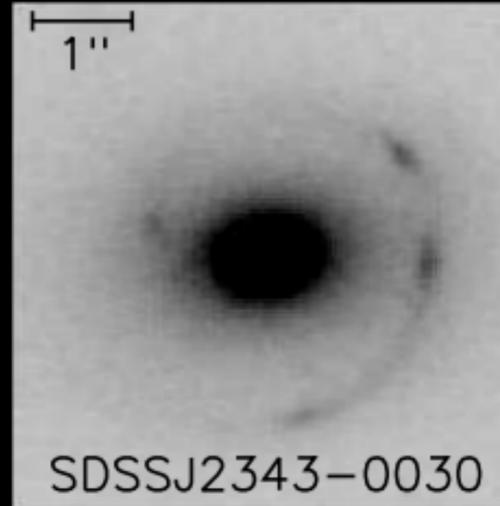
(Peter et al. 2013)





TOTAL MASS

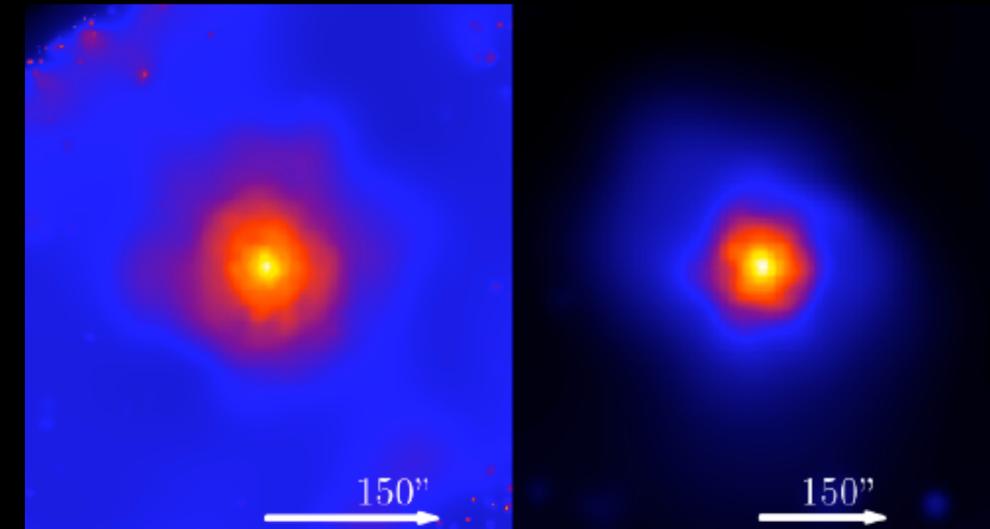
Grav. Lensing - galaxies modelled as SIE



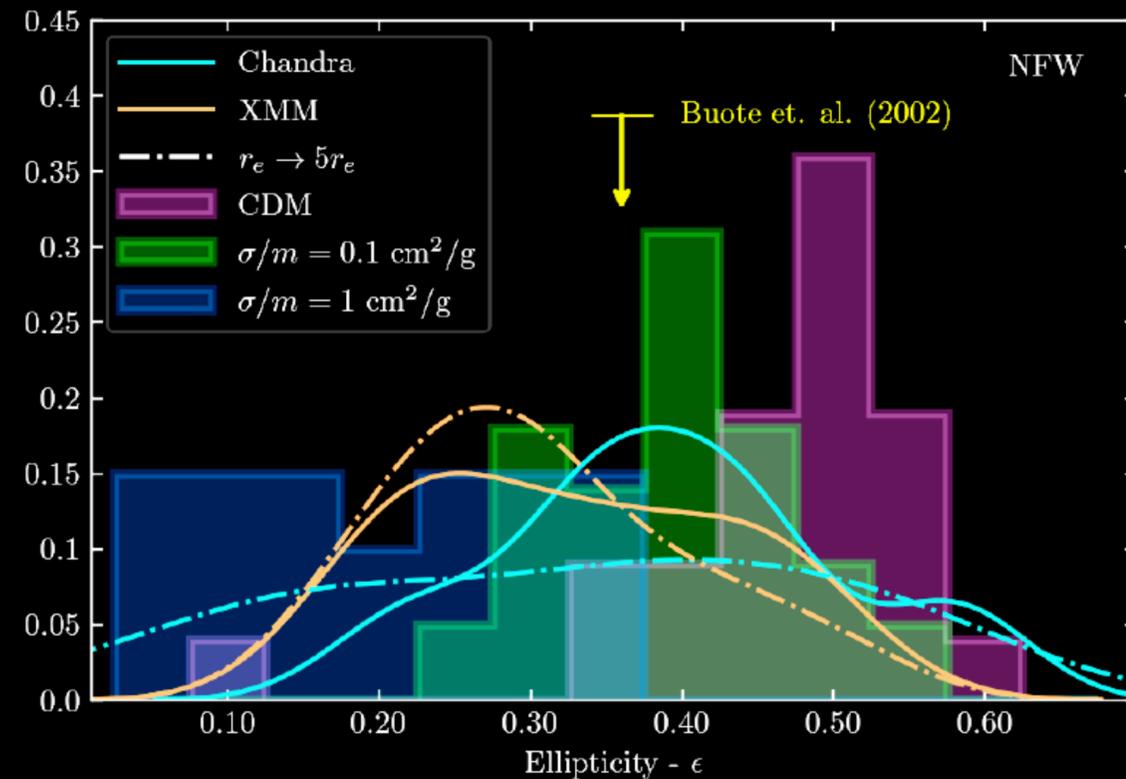
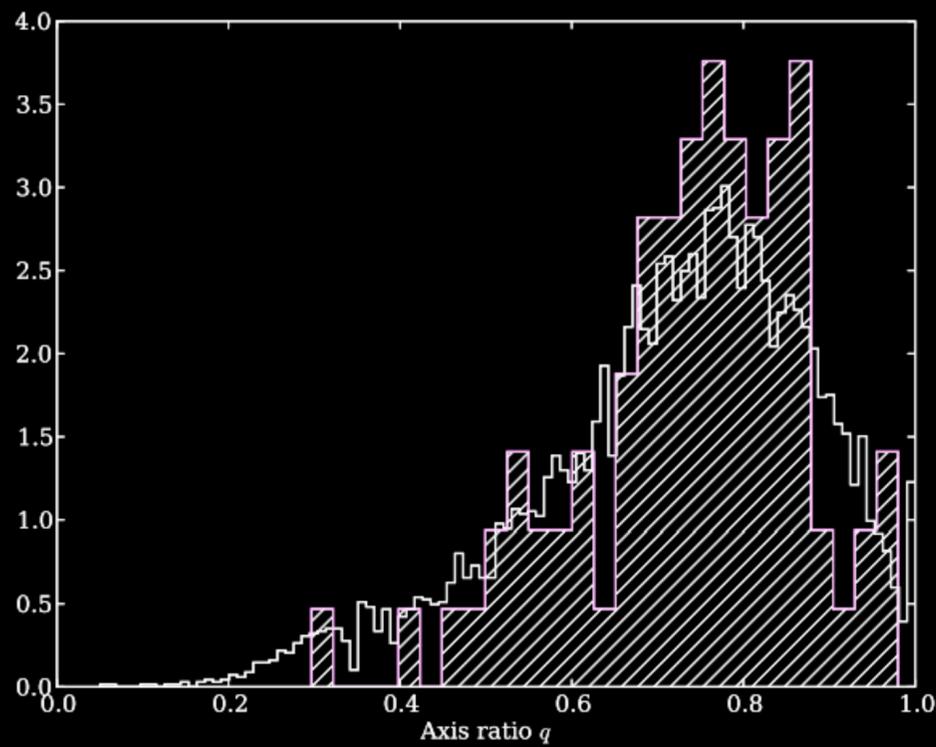
SHAPES

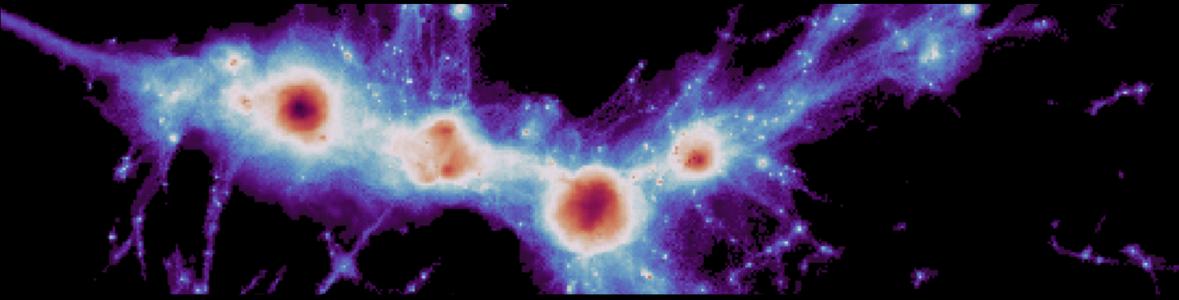
GAS: X-RAY

the X-ray emitting gas traces the gravitational potential



(McDaniel+21) - 11





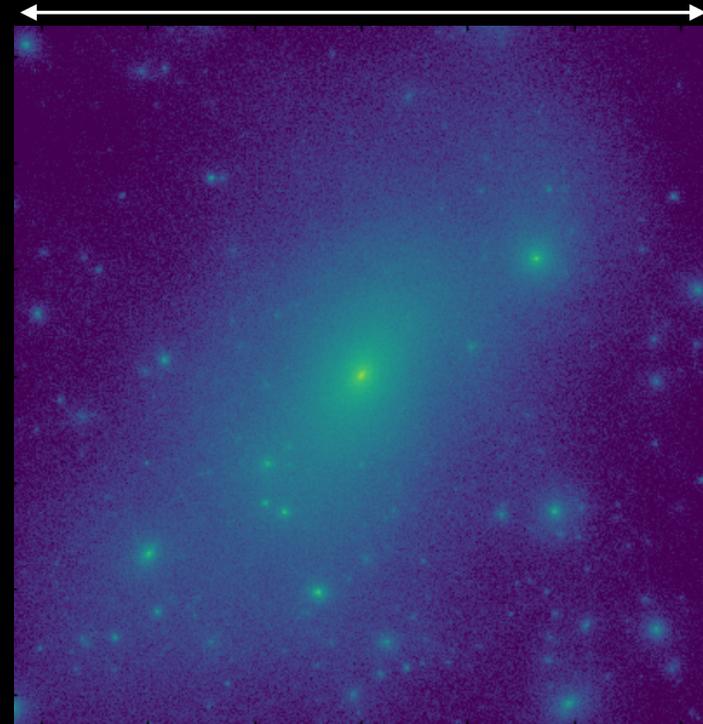
SIDM + BARYONS: SHAPES

Chandra

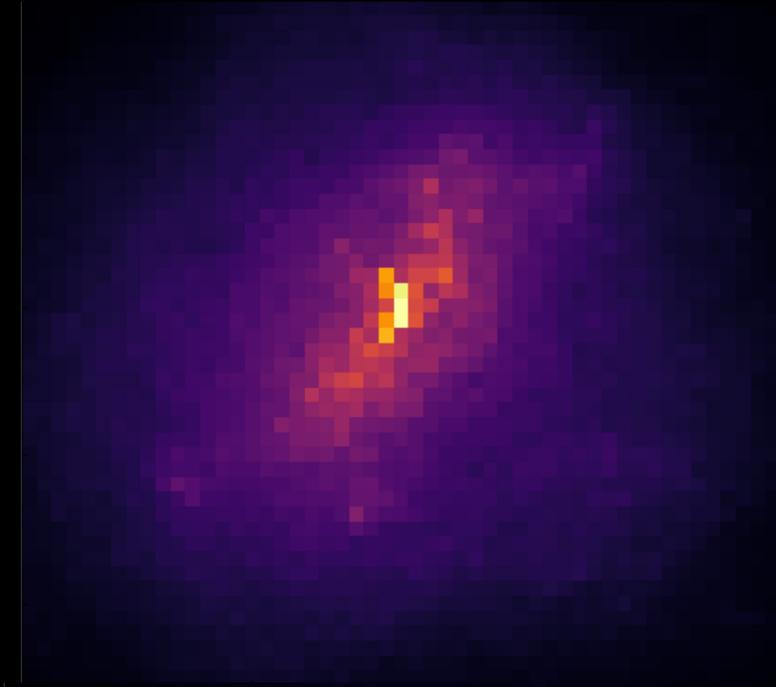
XMM

0.5 R_{vir}

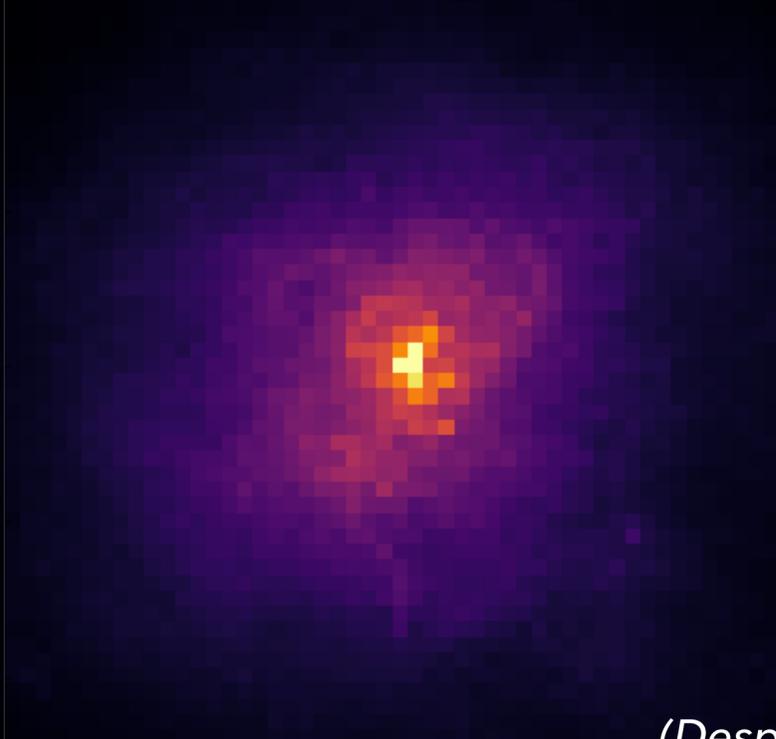
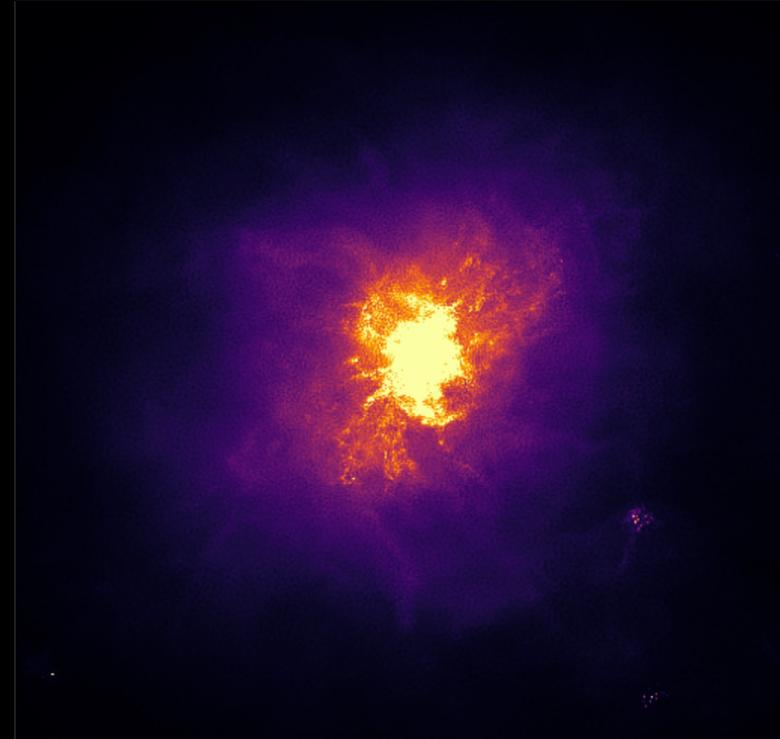
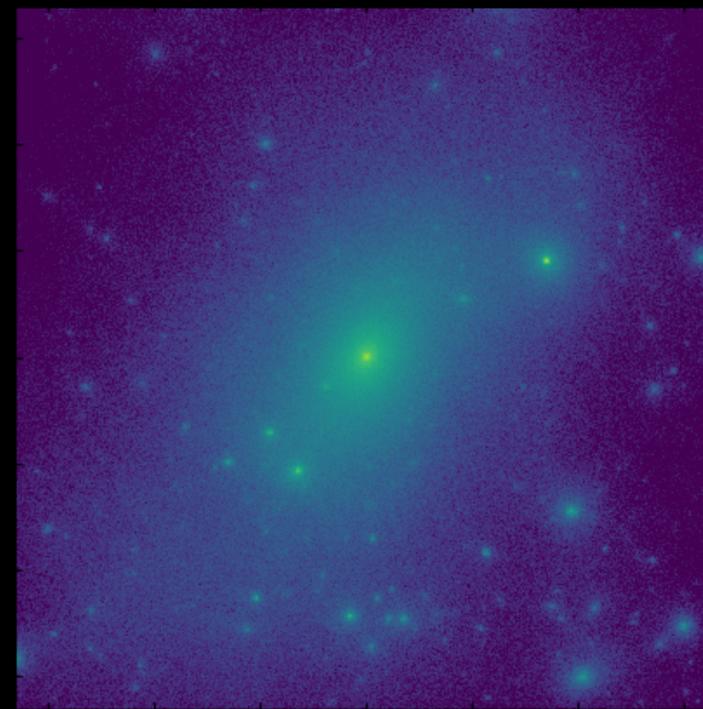
CDM



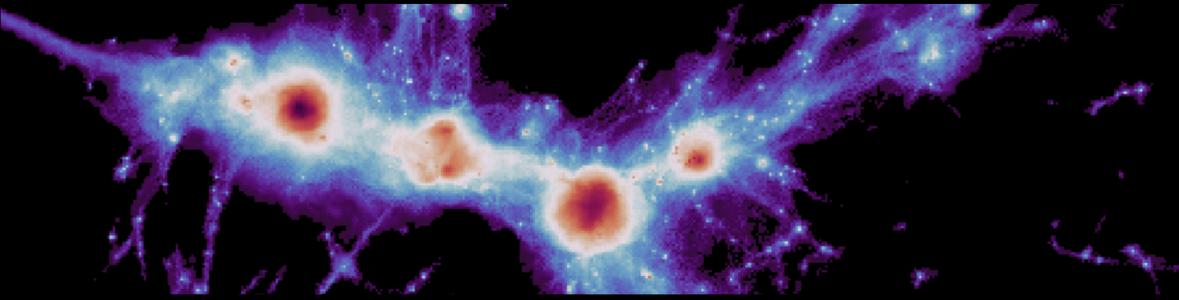
50
0
-50
-100



SIDM

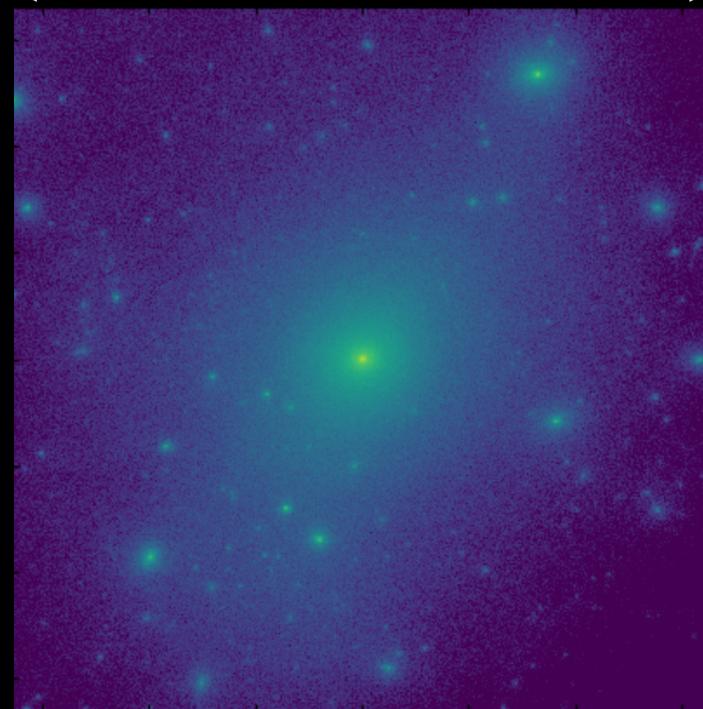


(Despali et al. 2022)

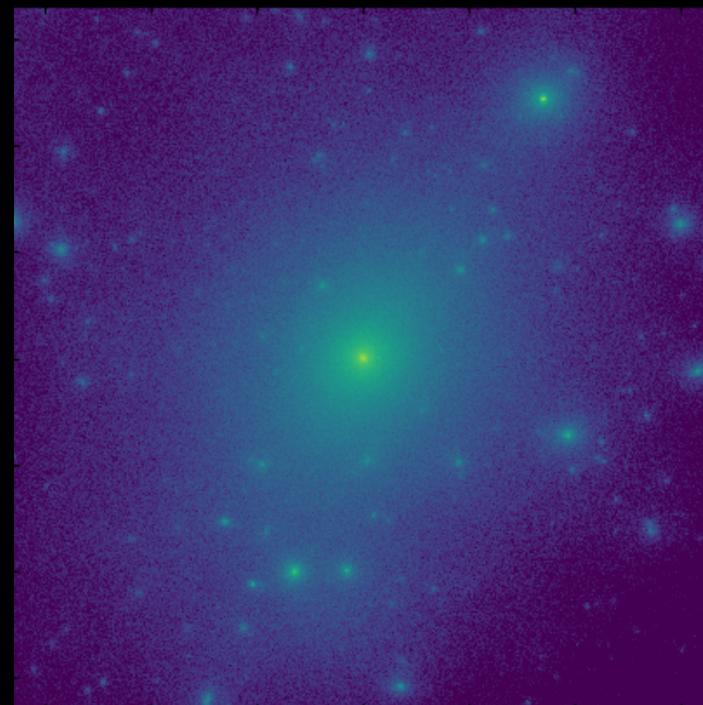


SIDM + BARYONS: SHAPES

CDM



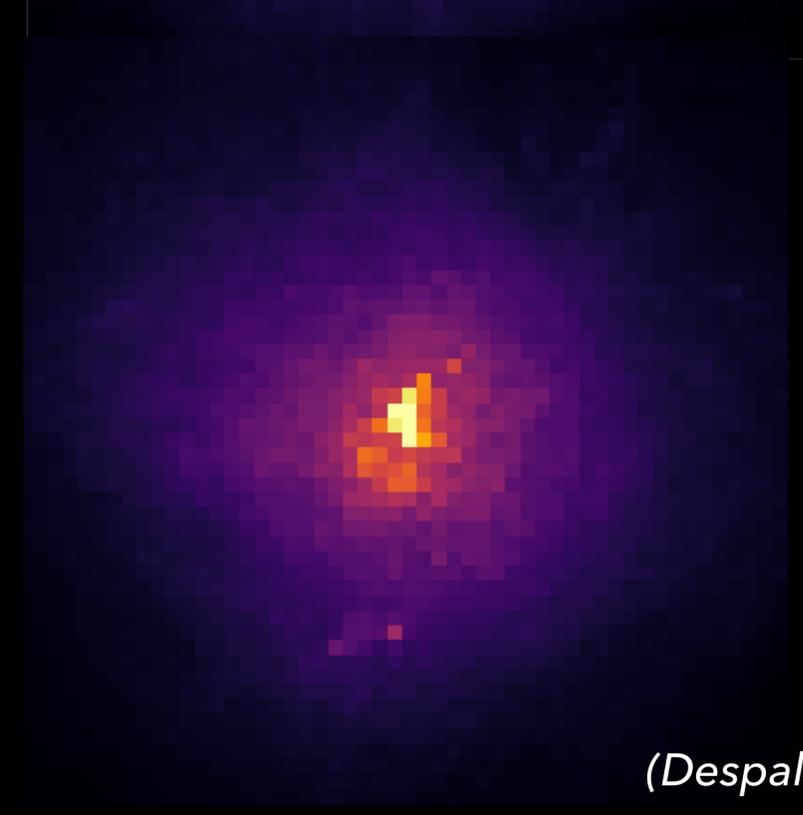
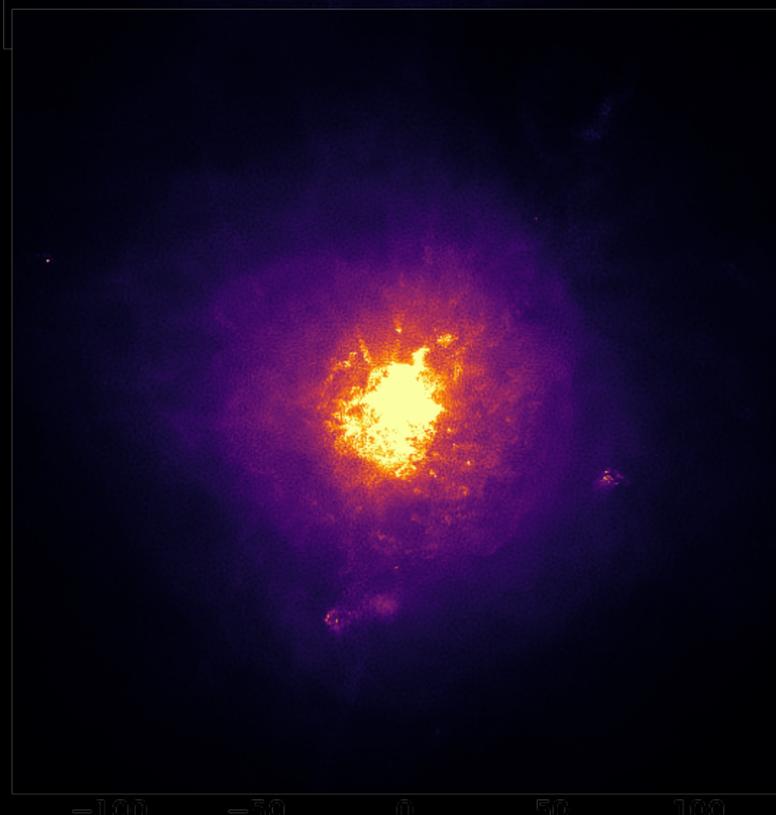
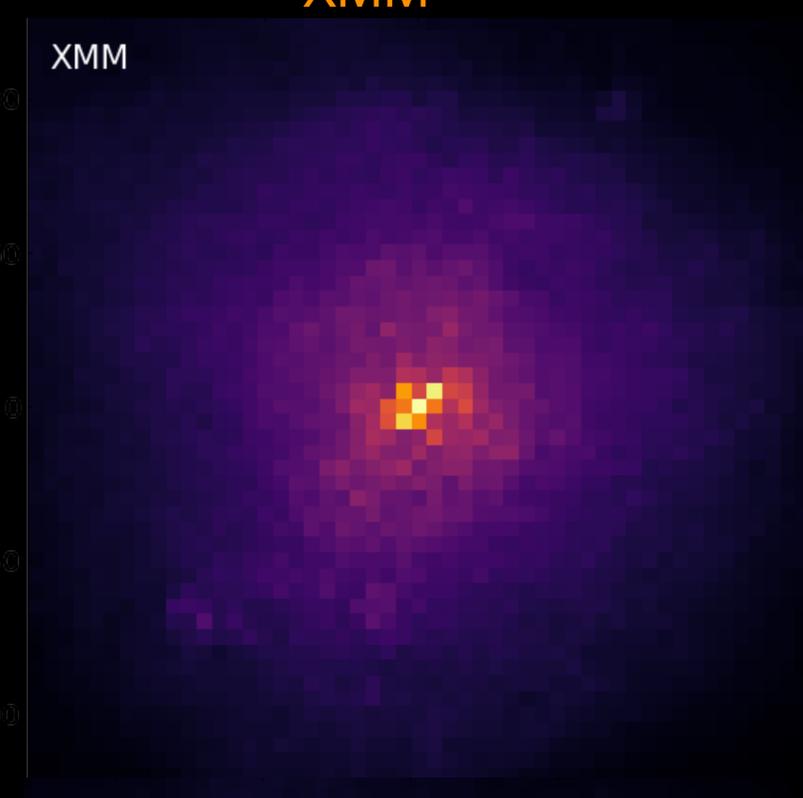
SIDM

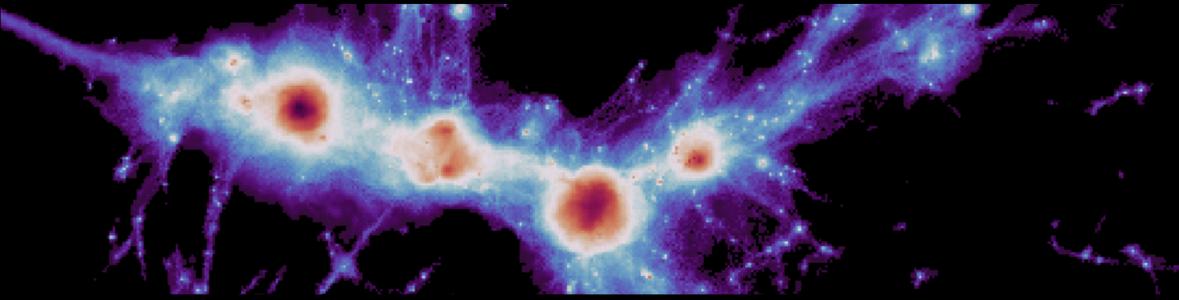


Chandra



XMM





- CDM haloes are more elongated in the inner parts

- SIDM makes haloes rounder and reverses the trend

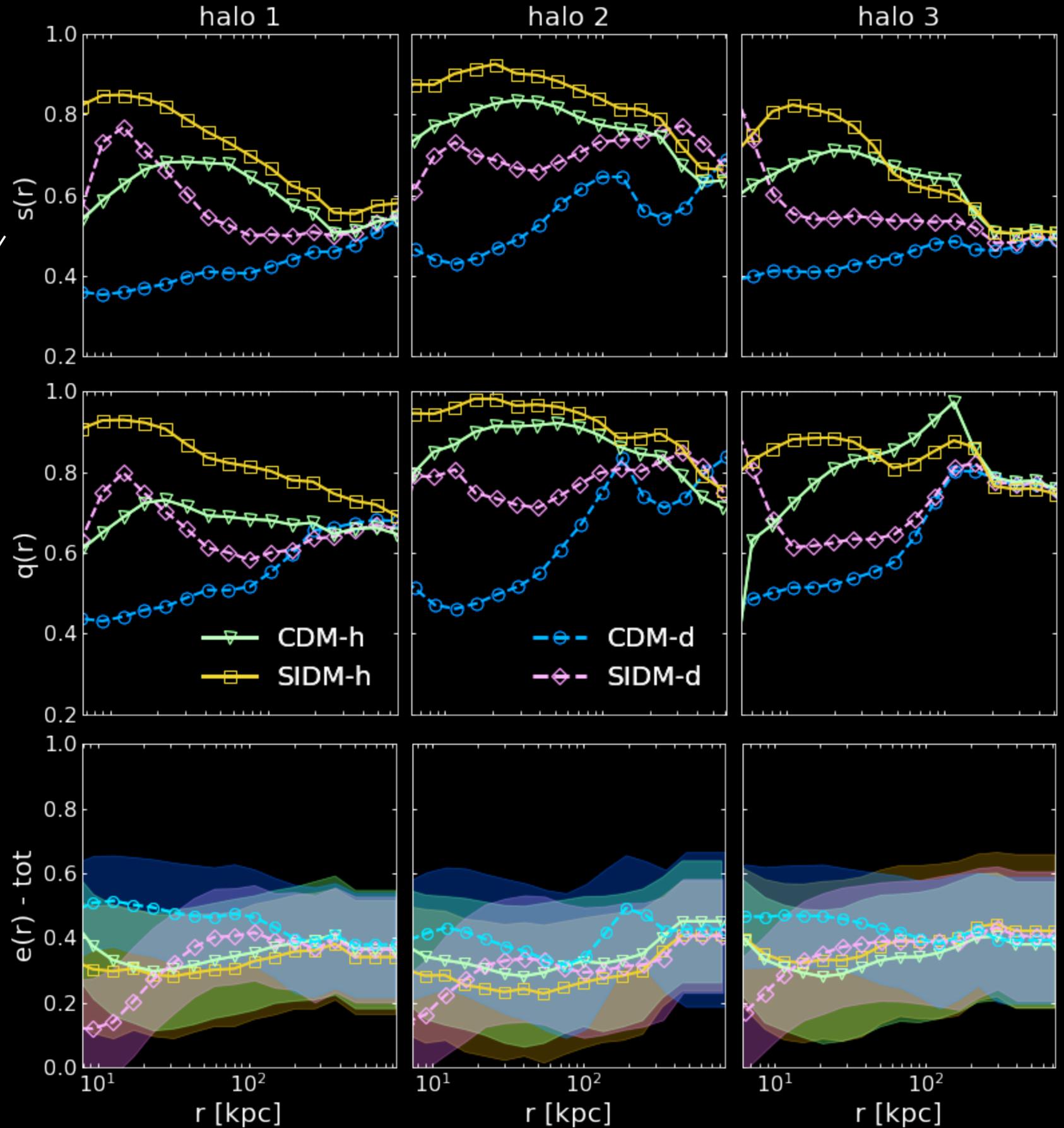
- baryons also make haloes rounder

- both only influence the shapes in the inner regions

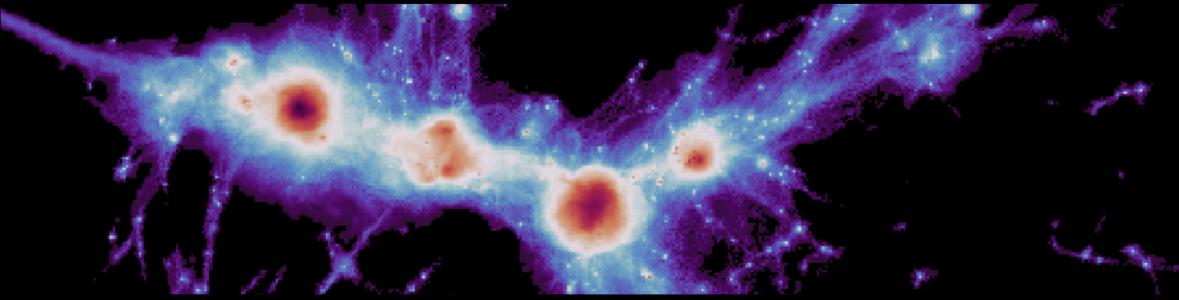
- hydro runs are more similar to each other than dm-only

- in projection most of the difference is lost

3D



2D
projected
mass



- CDM haloes are more elongated in the inner parts

- **SIDM makes haloes rounder and reverses the trend**

- **baryons also make haloes rounder**

- both only influence the shapes in the inner regions

- **hydro runs are more similar to each other than dm-only**

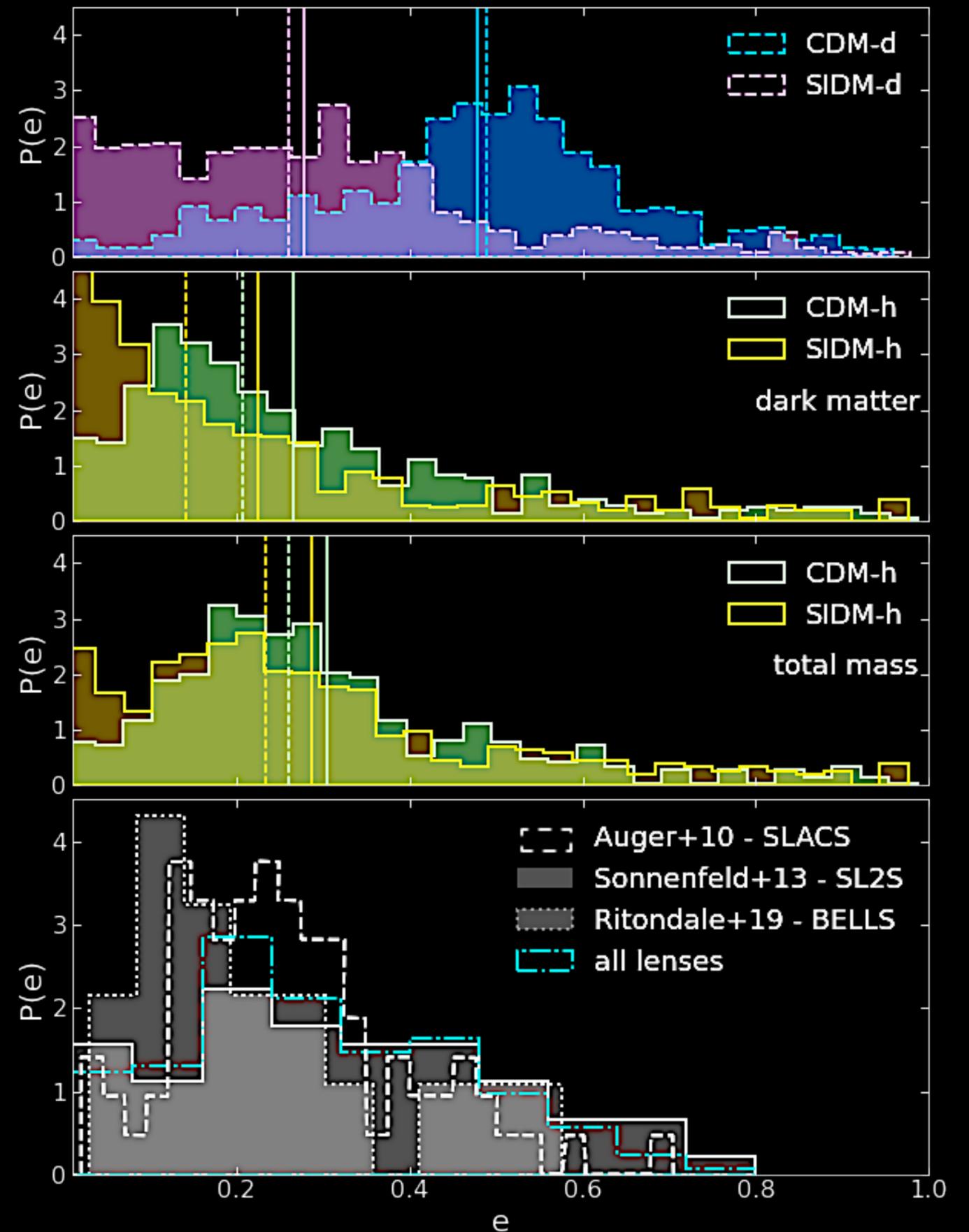
- **in projection most of the difference is lost**

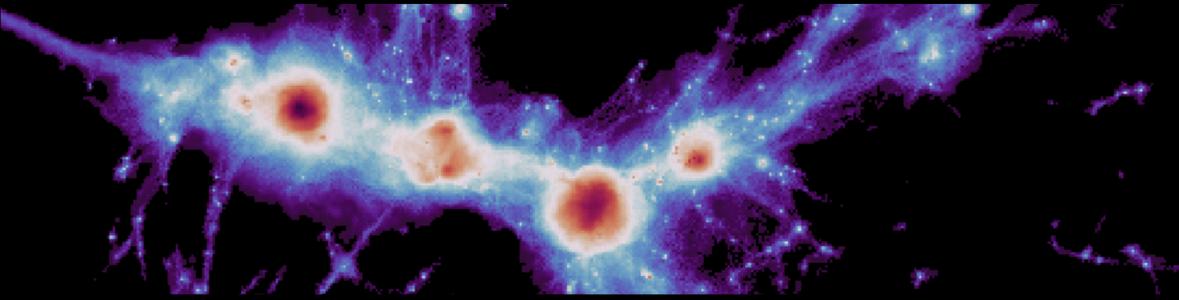
DM-only

DM in hydro runs

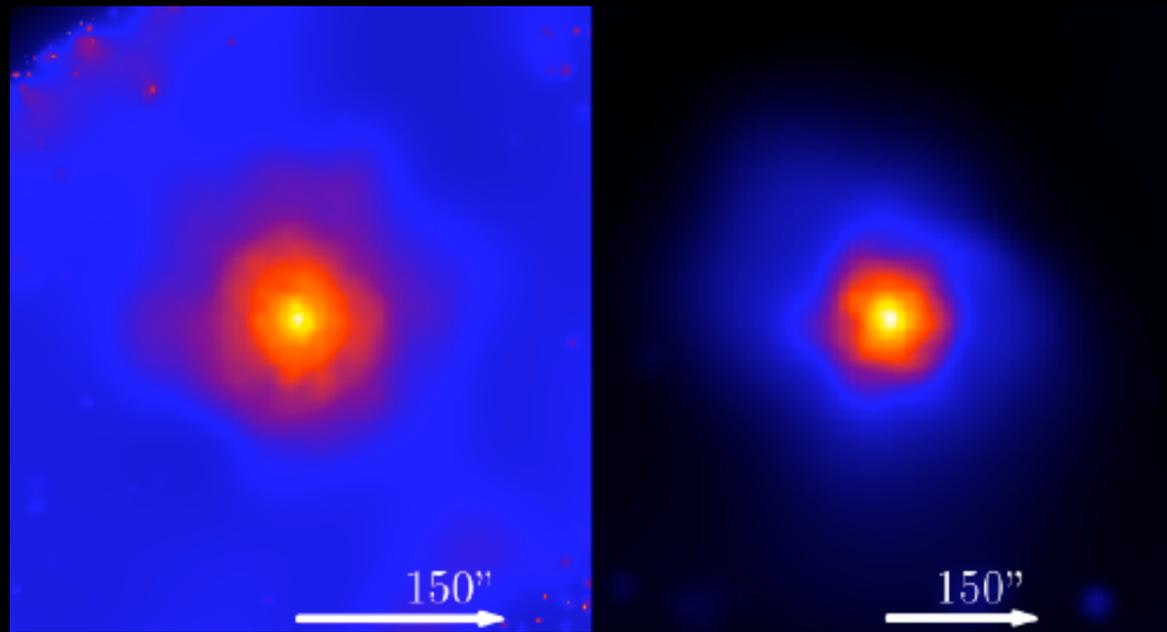
total mass in hydro runs

lensing observations

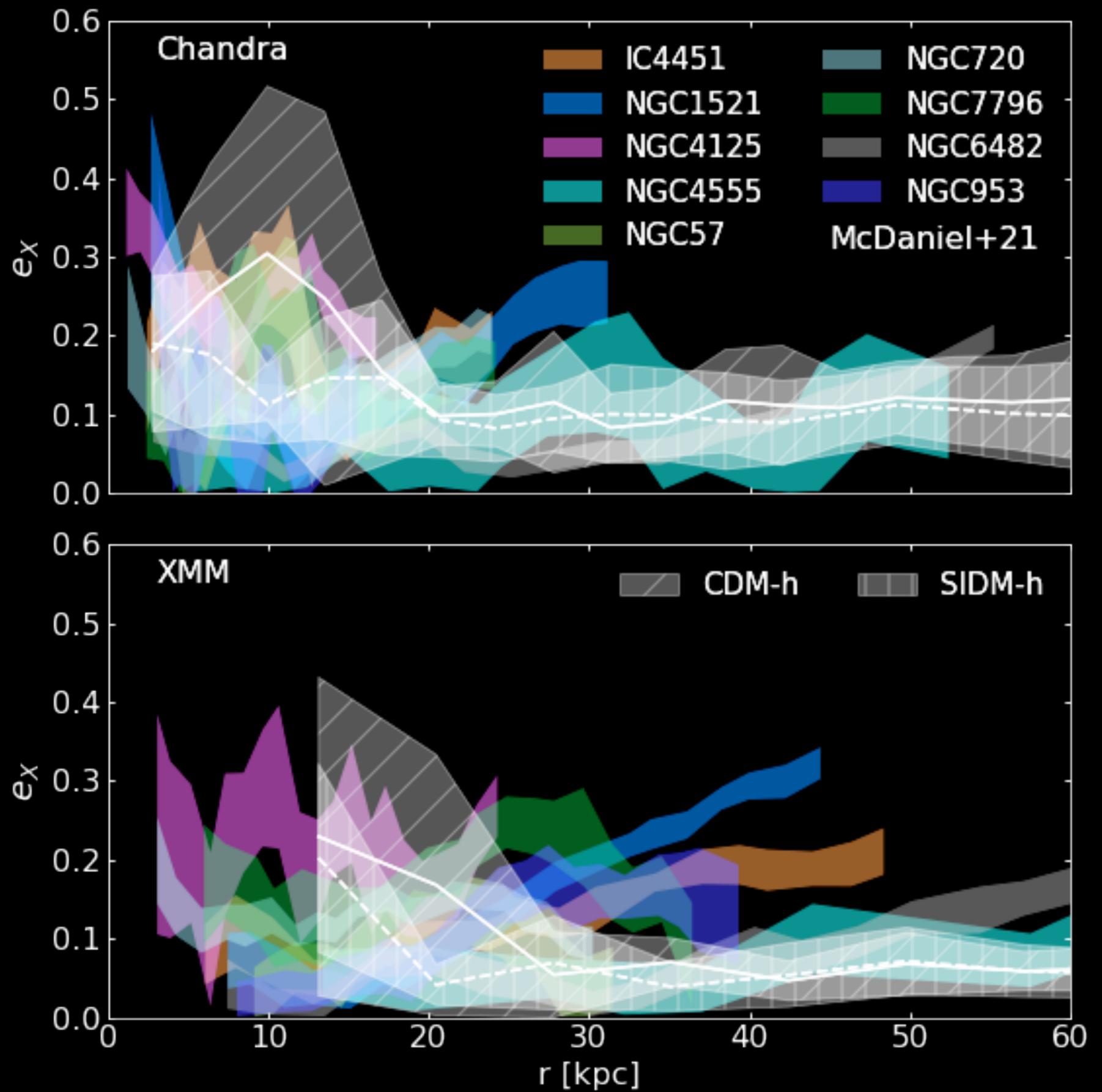


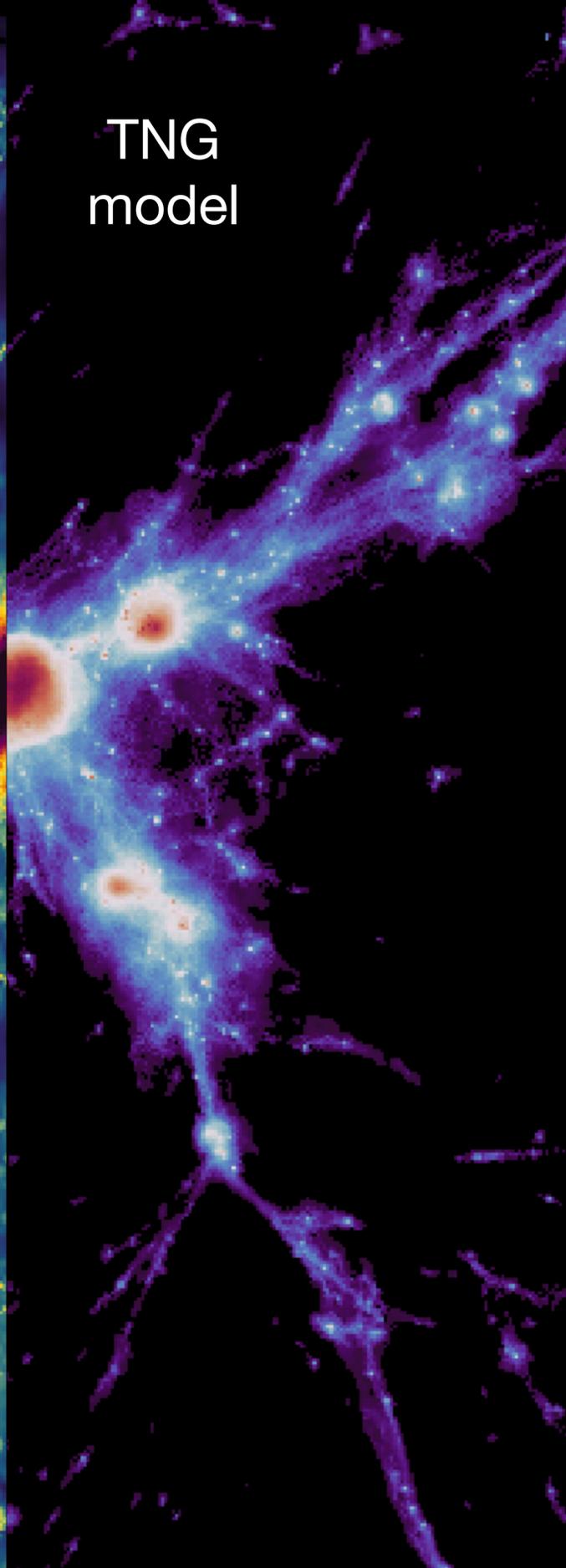
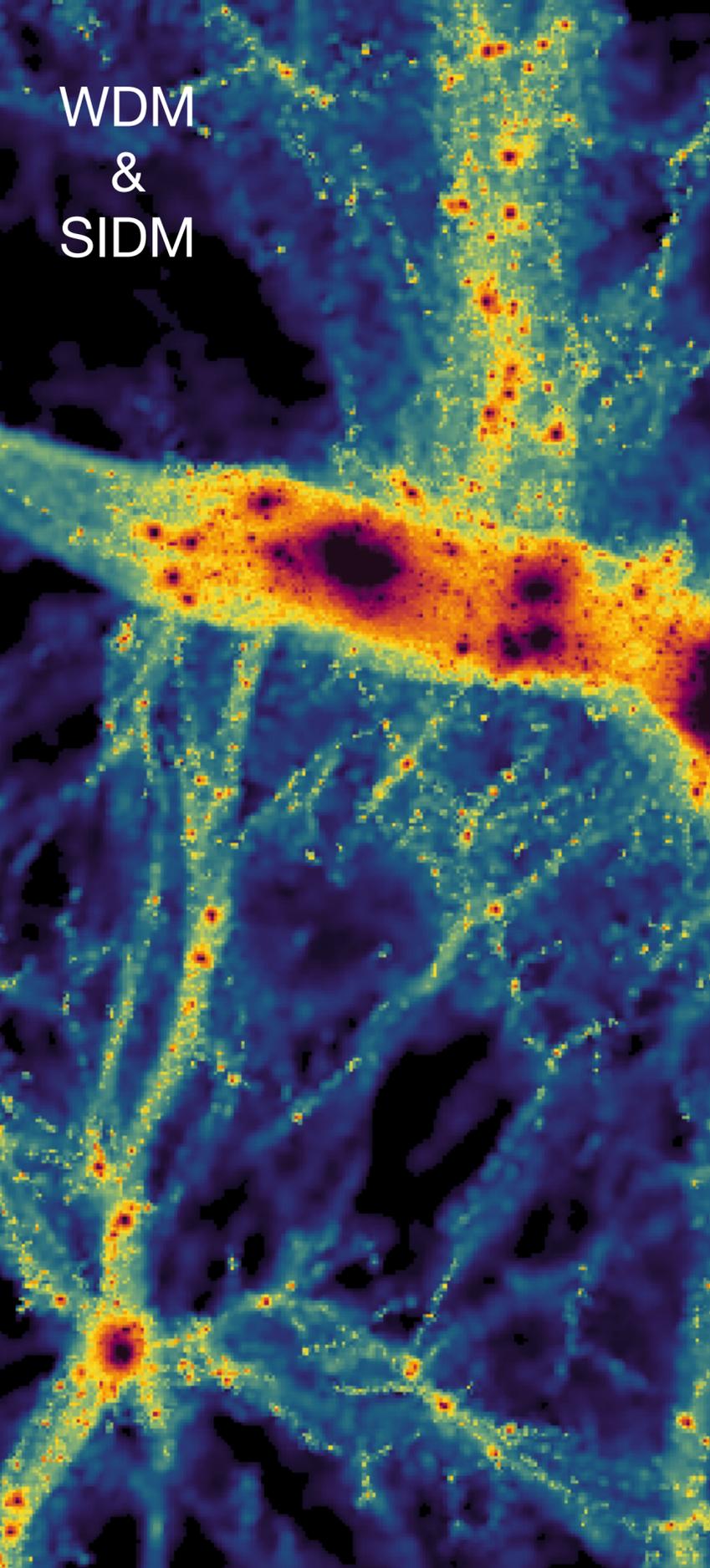


(McDaniel+21) -
11 elliptical galaxies in Xray

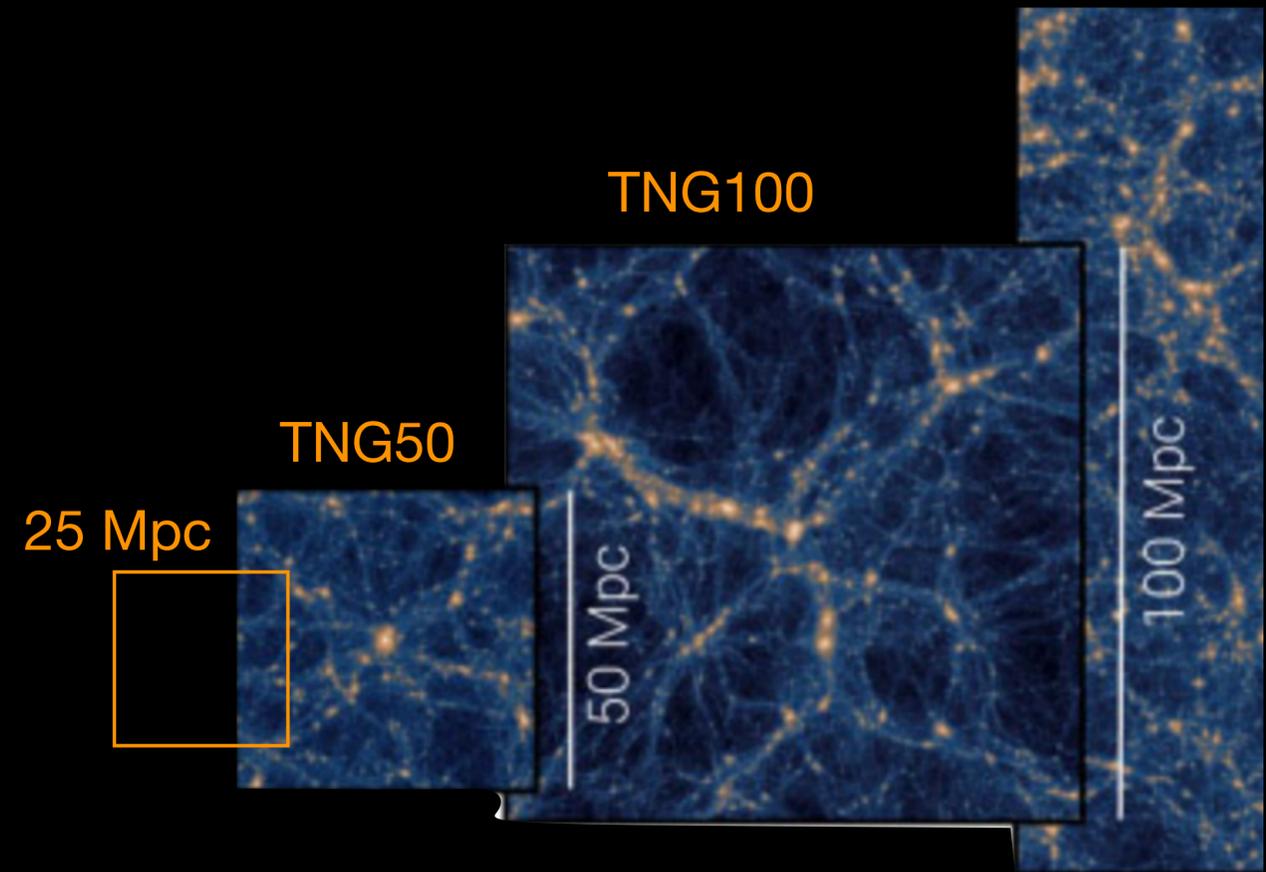


(Despali et al. 2022)





NEW SIMULATIONS

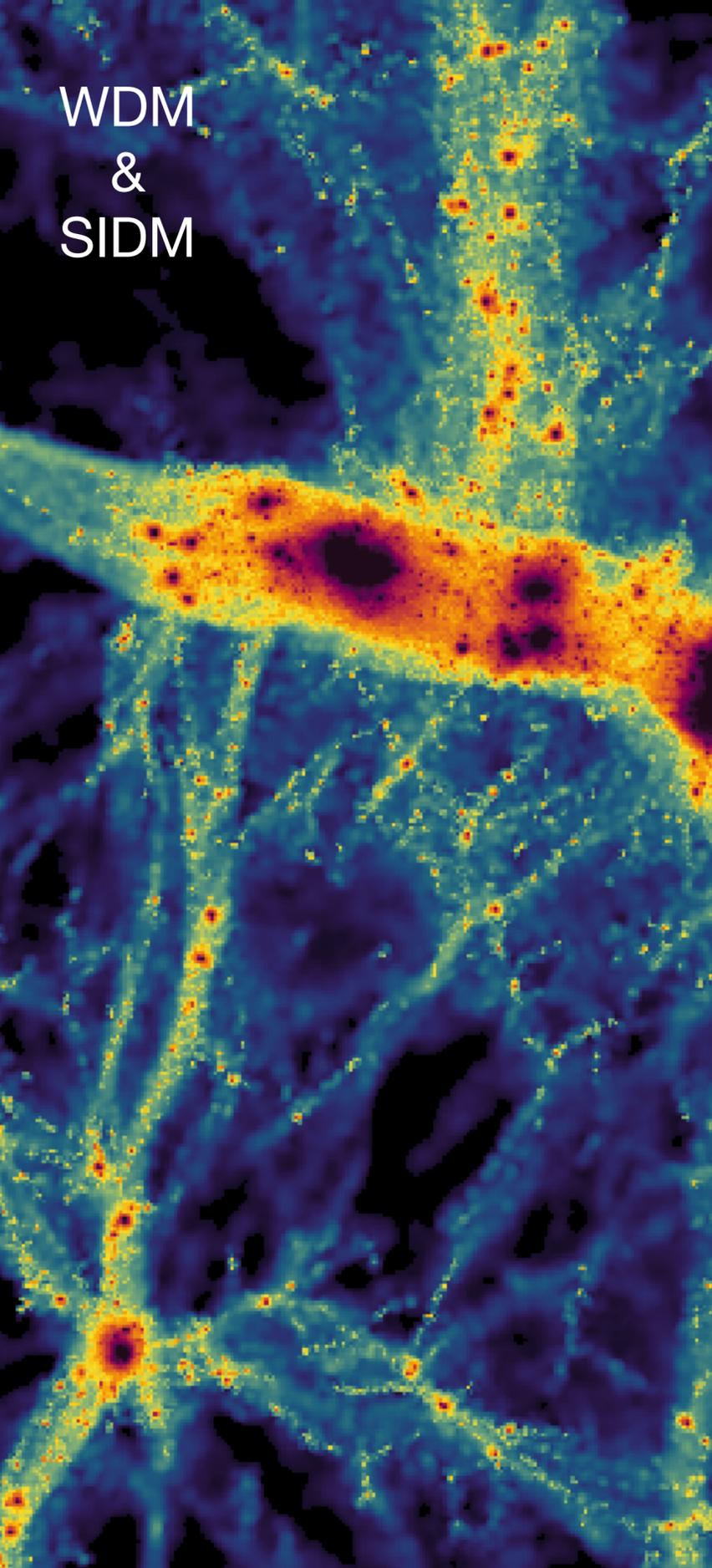


max resolution: $4 \times 10^5 M_{\odot}$, 280 pc

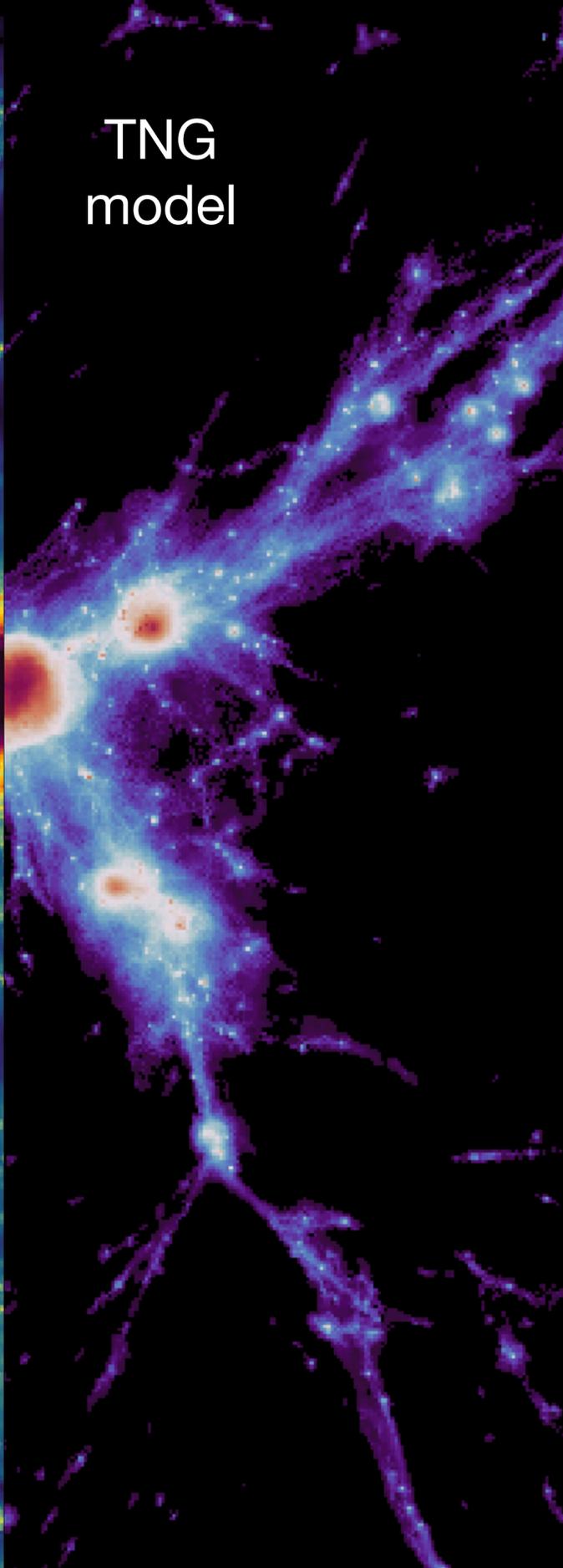
WDM, SIDM (constant, velocity-dependent & inelastic)

Despali et al. in prep: Annalisa Pillepich, Dylan Nelson, Volker Springel, Lauro Moscardini

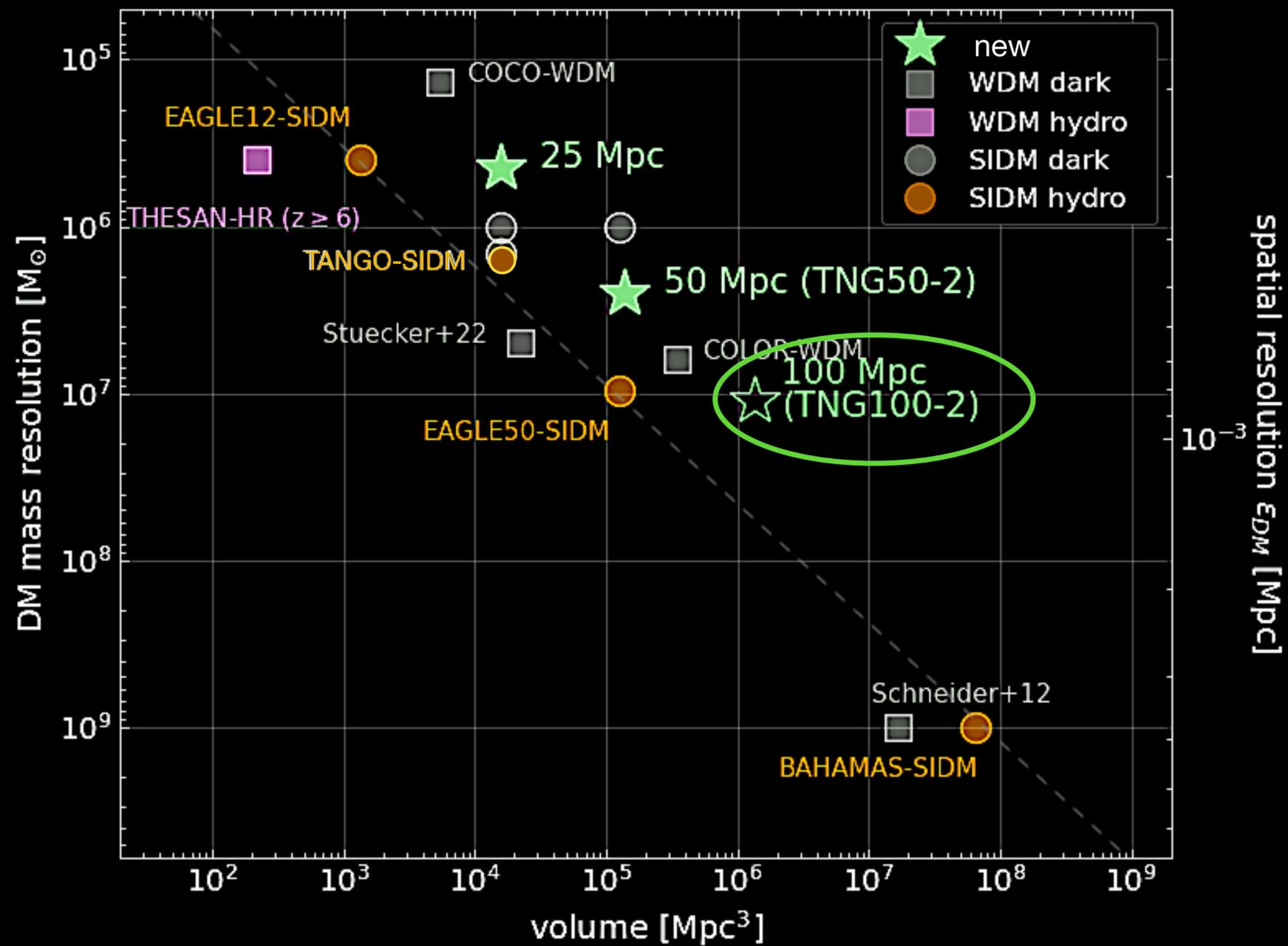
WDM
&
SIDM



TNG
model

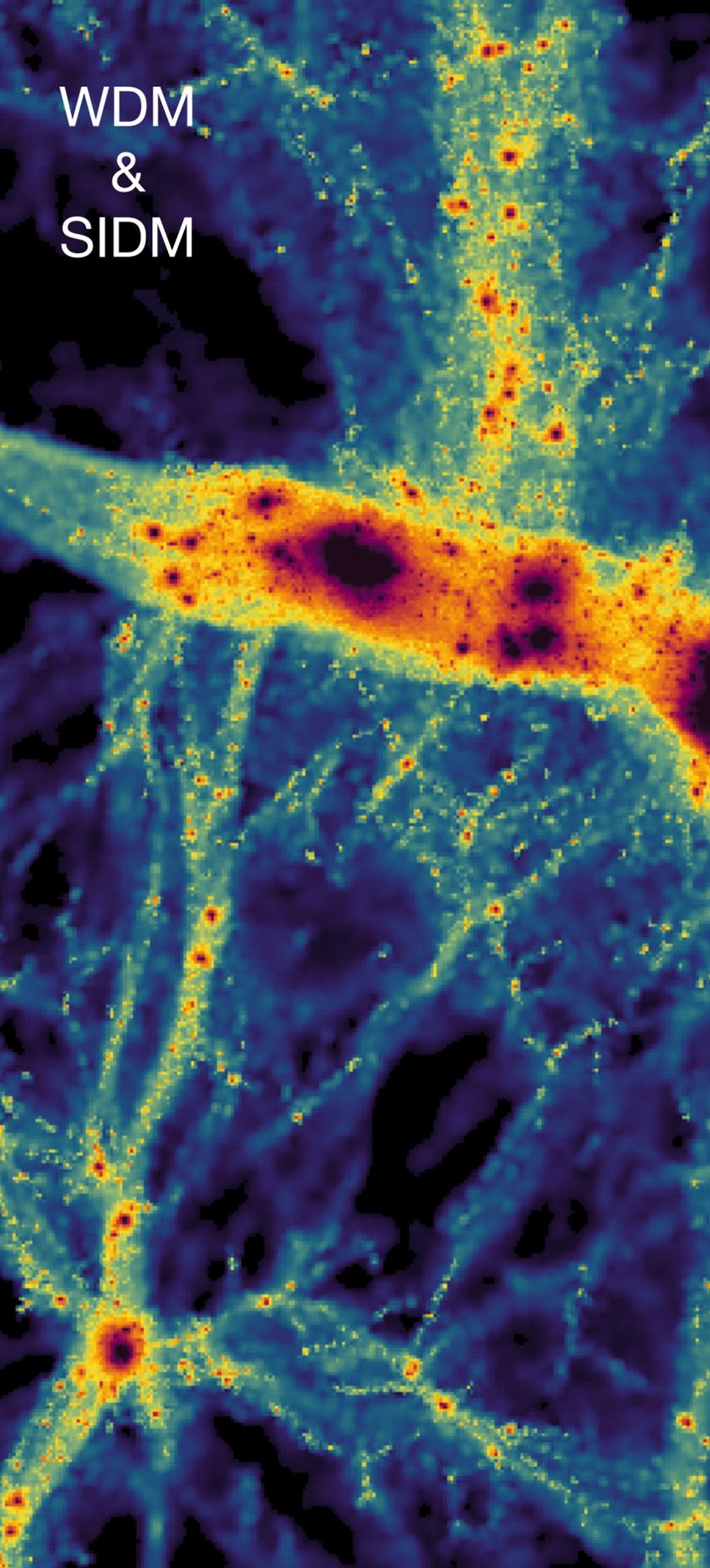


NEW SIMULATIONS

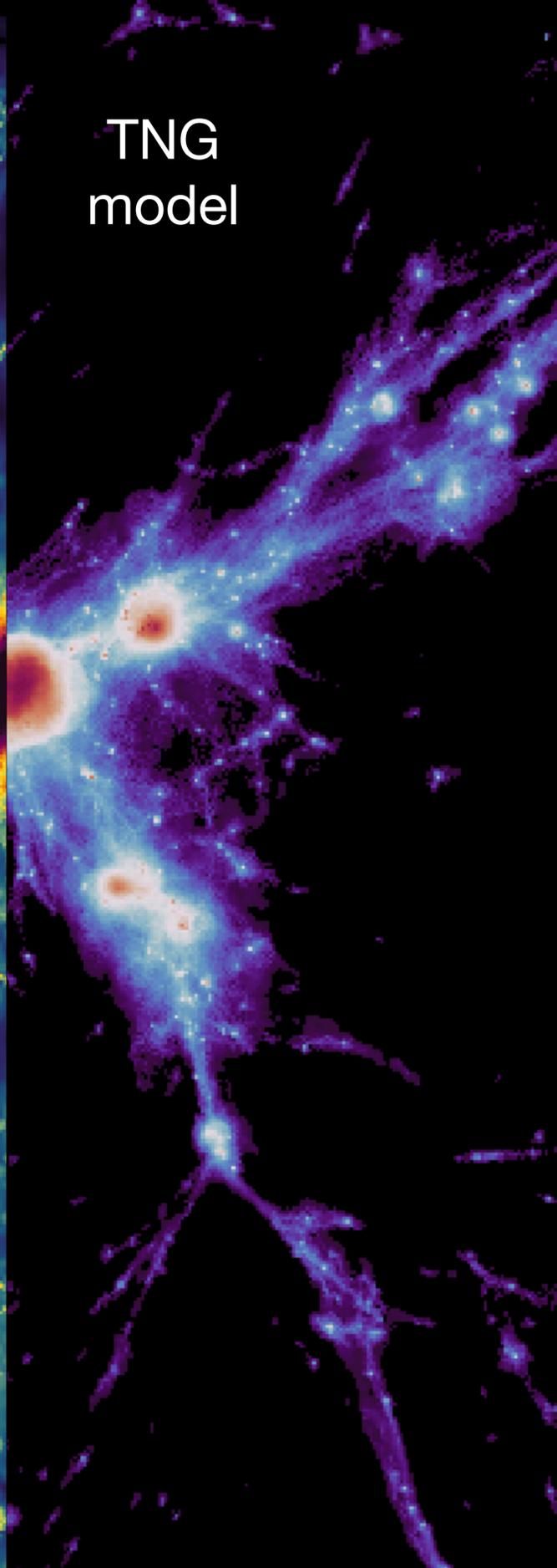


Despali et al. in prep: Annalisa Pillepich, Dylan Nelson, Volker Springel, Lauro Moscardini

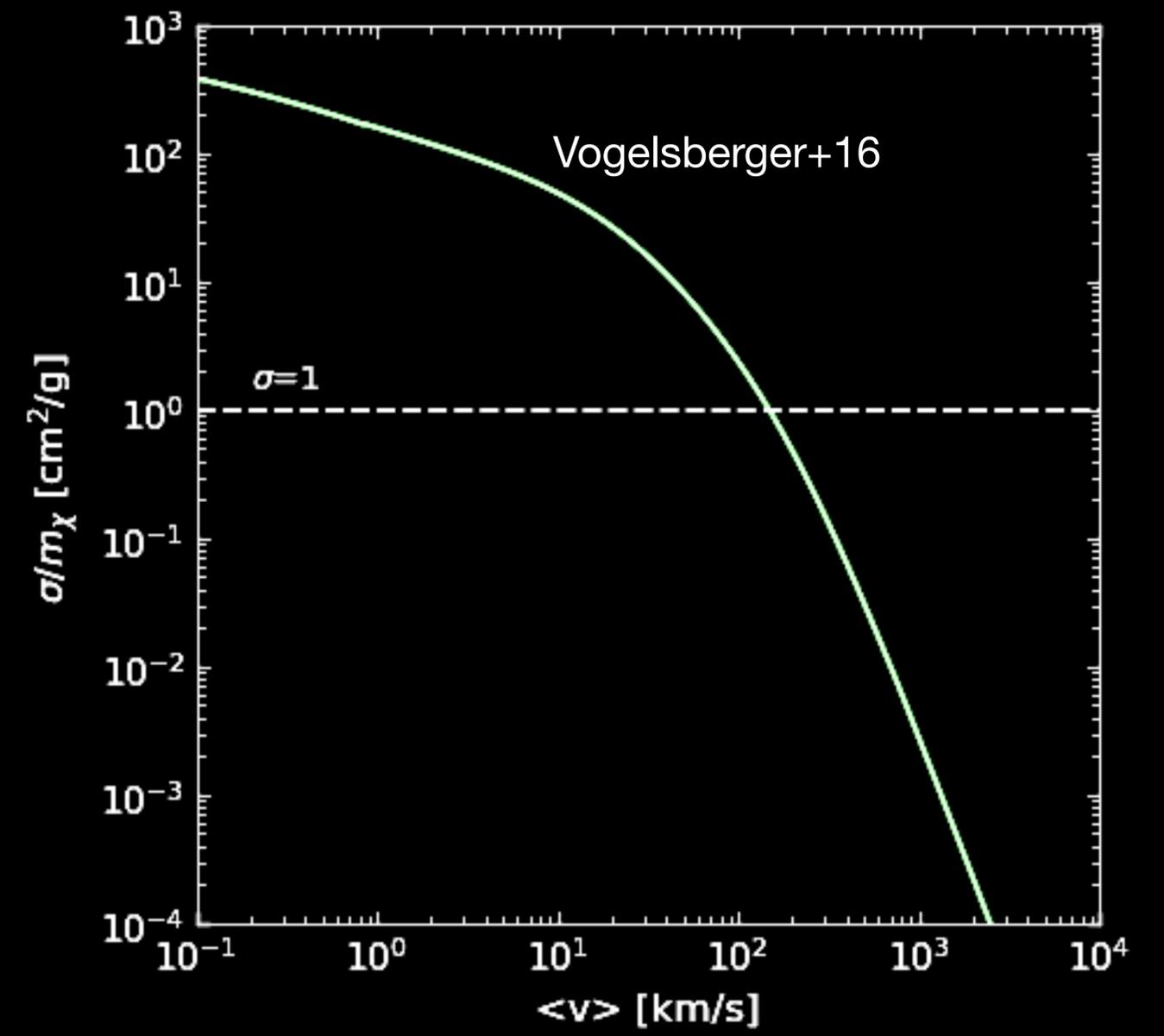
WDM
&
SIDM



TNG
model

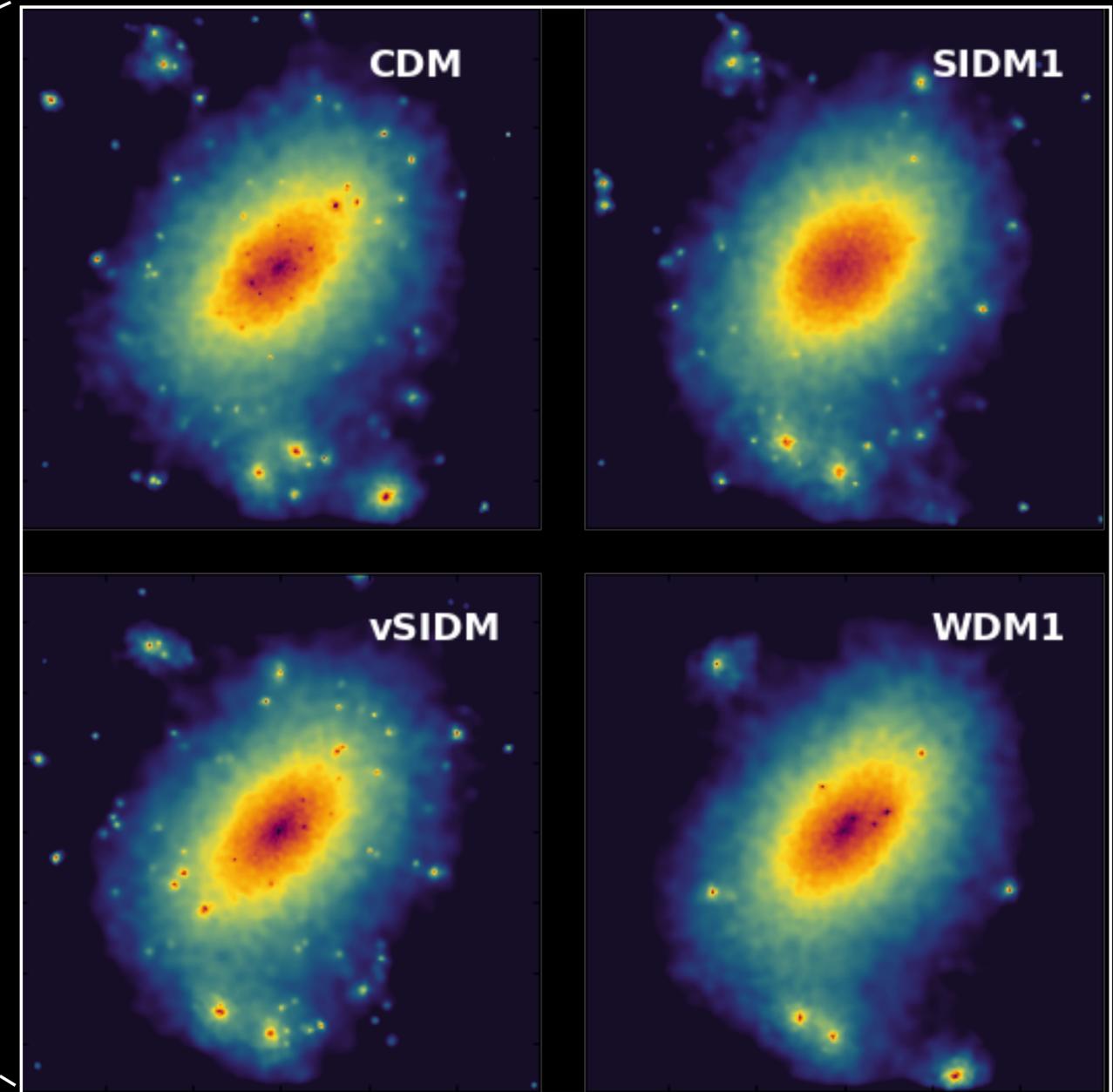
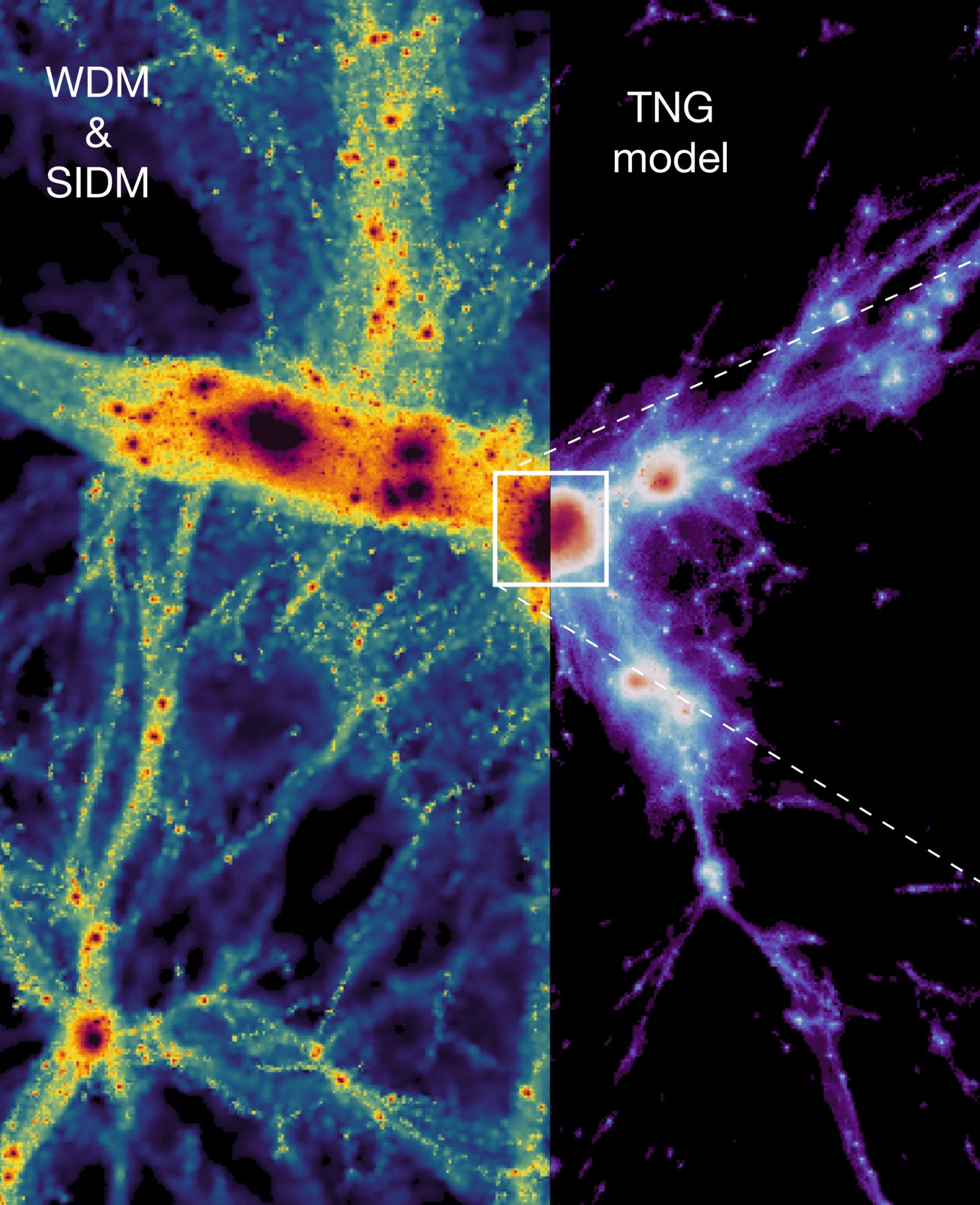


NEW SIMULATIONS



Despali et al. in prep: Annalisa Pillepich, Dylan Nelson, Volker Springel, Lauro Moscardini

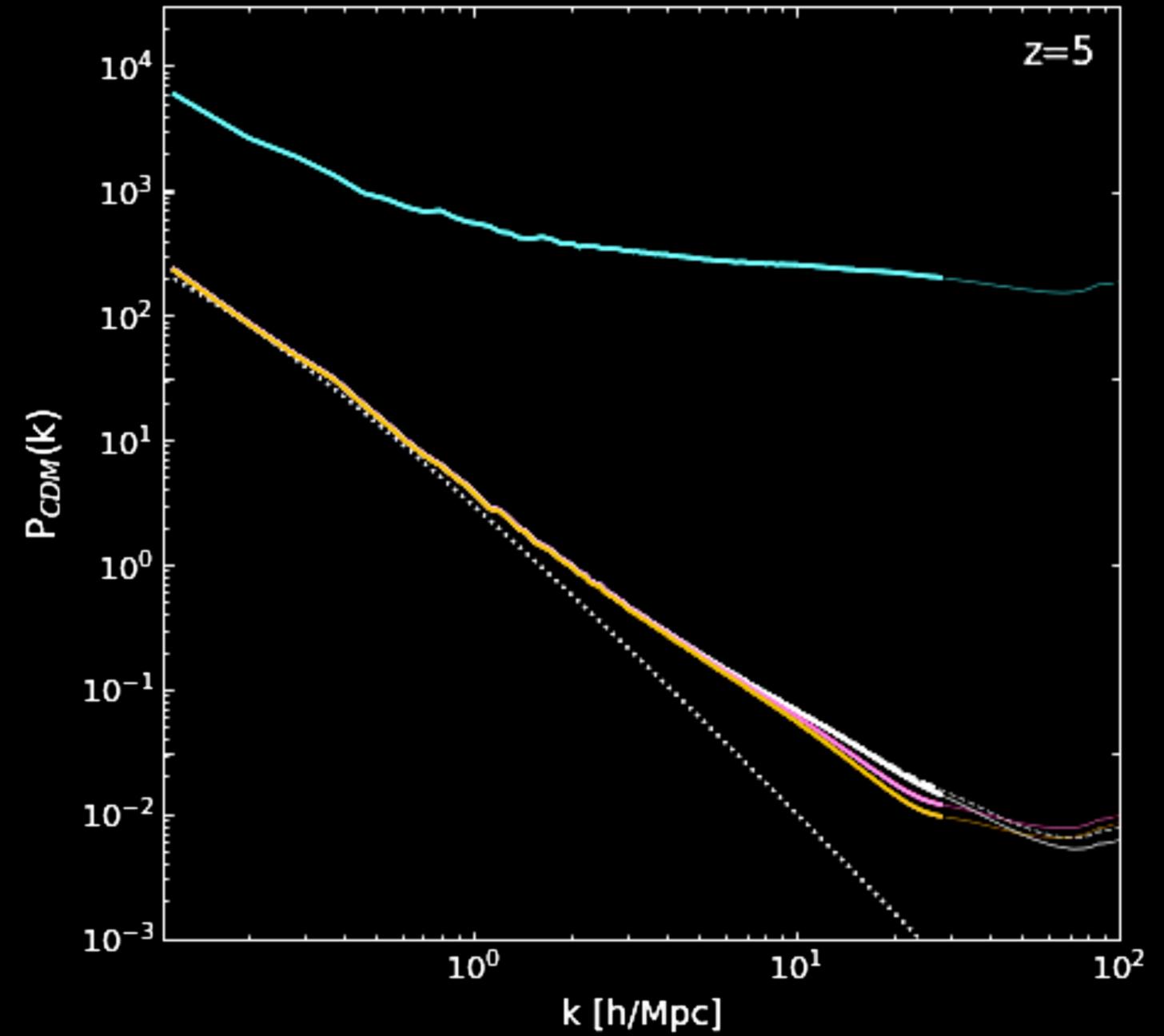
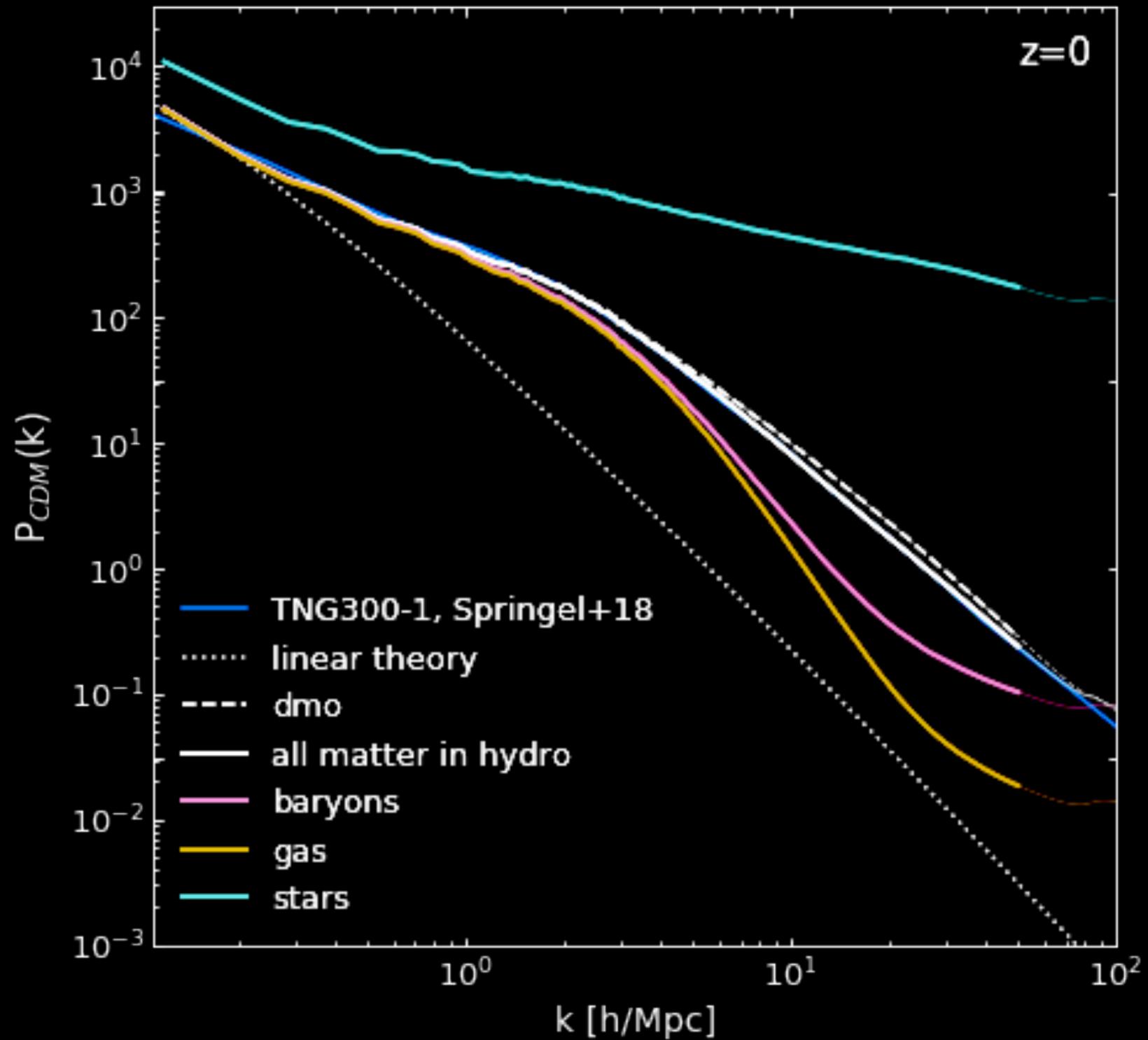
NEW SIMULATIONS



$\sim 10^8 M_{\odot}$, 2.8 kpc, 100 Mpc box

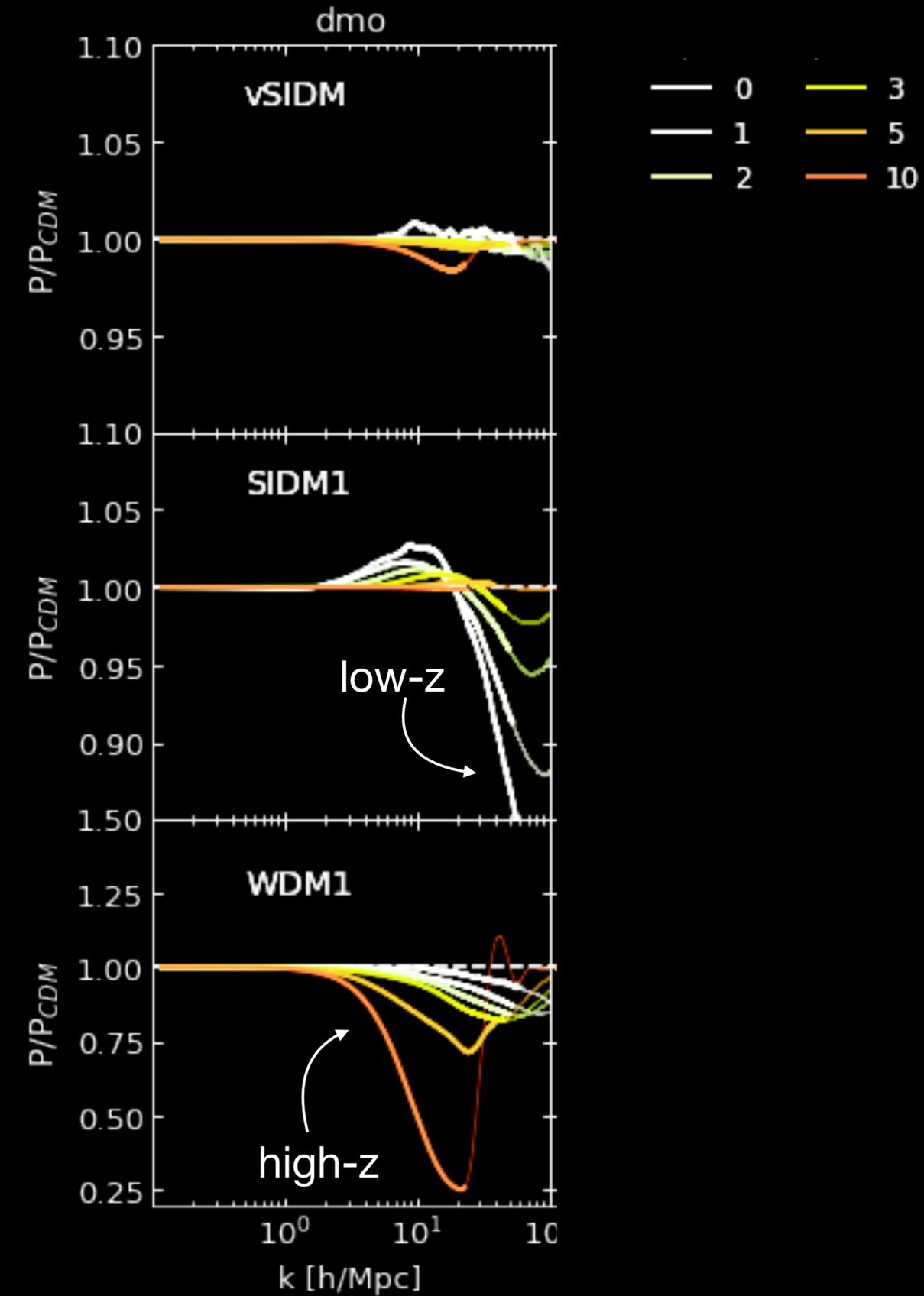
NEW SIMULATIONS

CDM

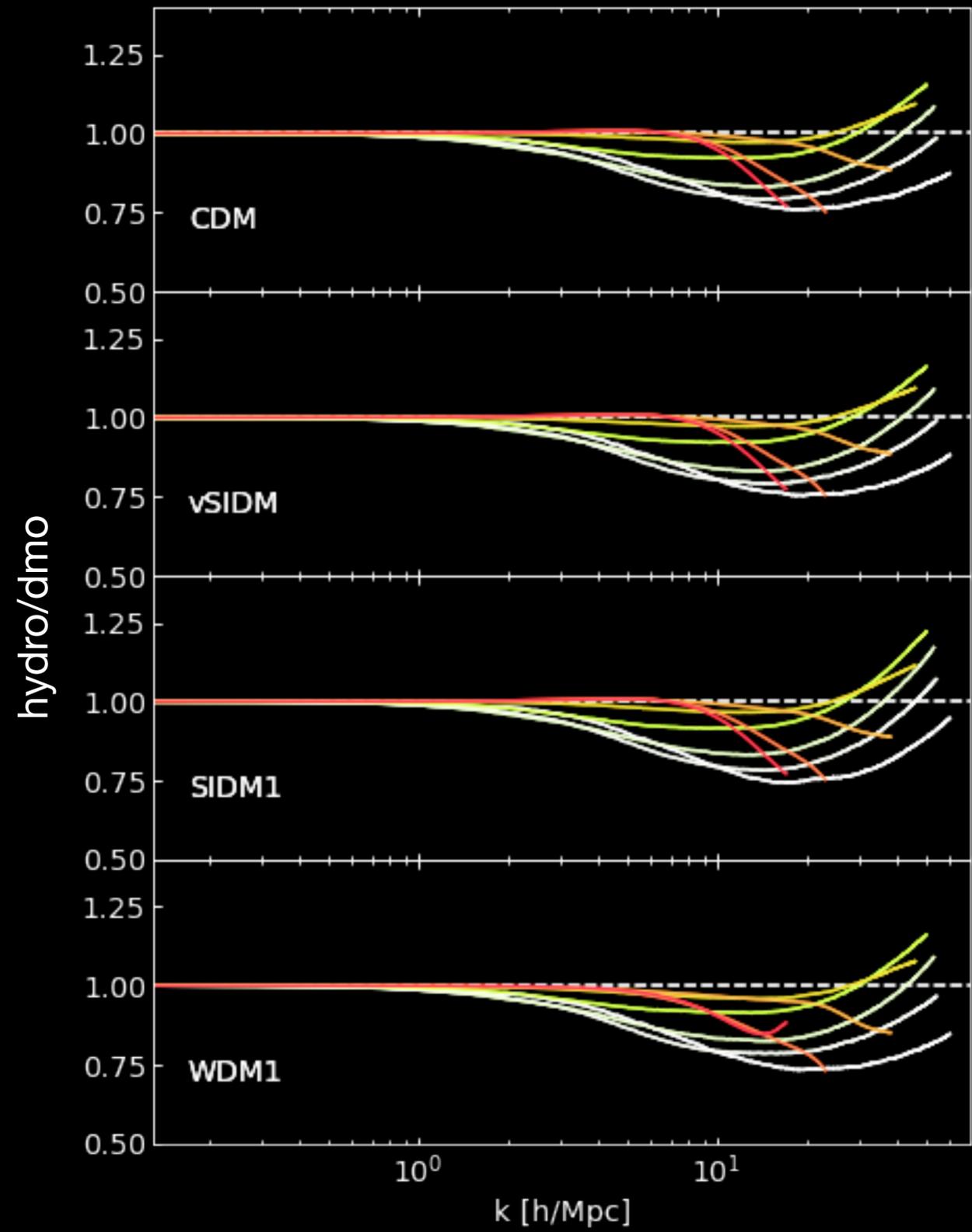
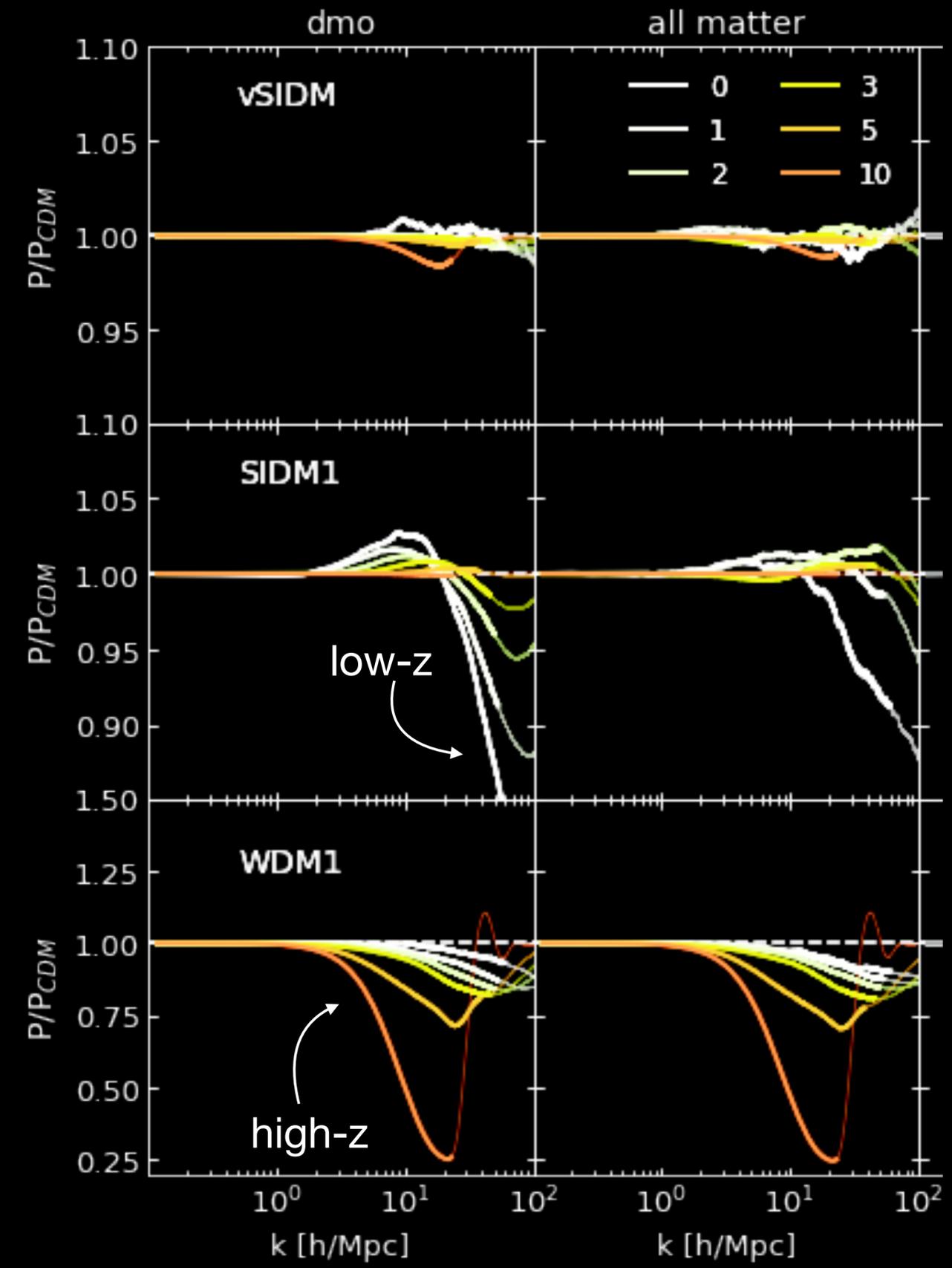


Despali et al. in prep

NEW SIMULATIONS

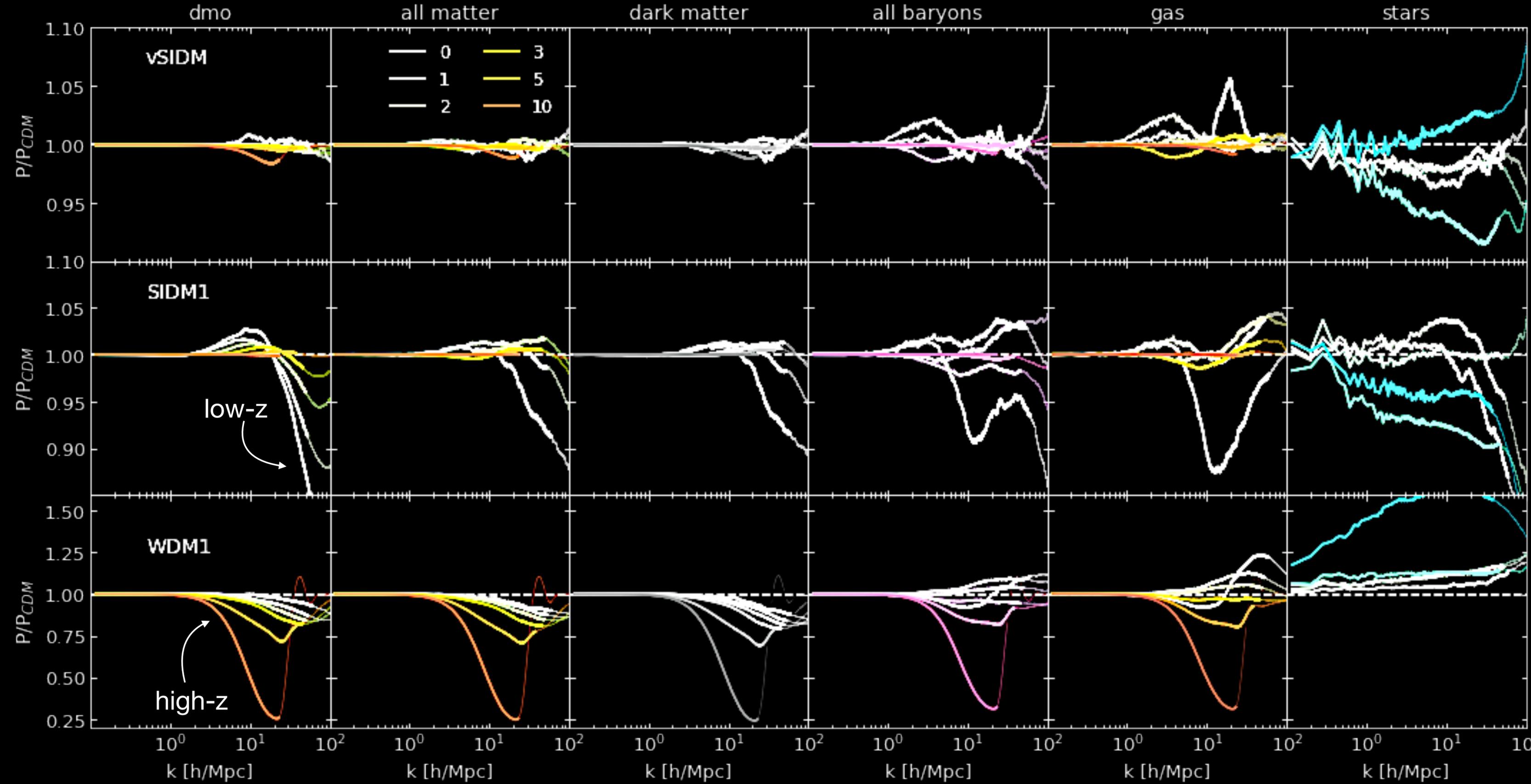


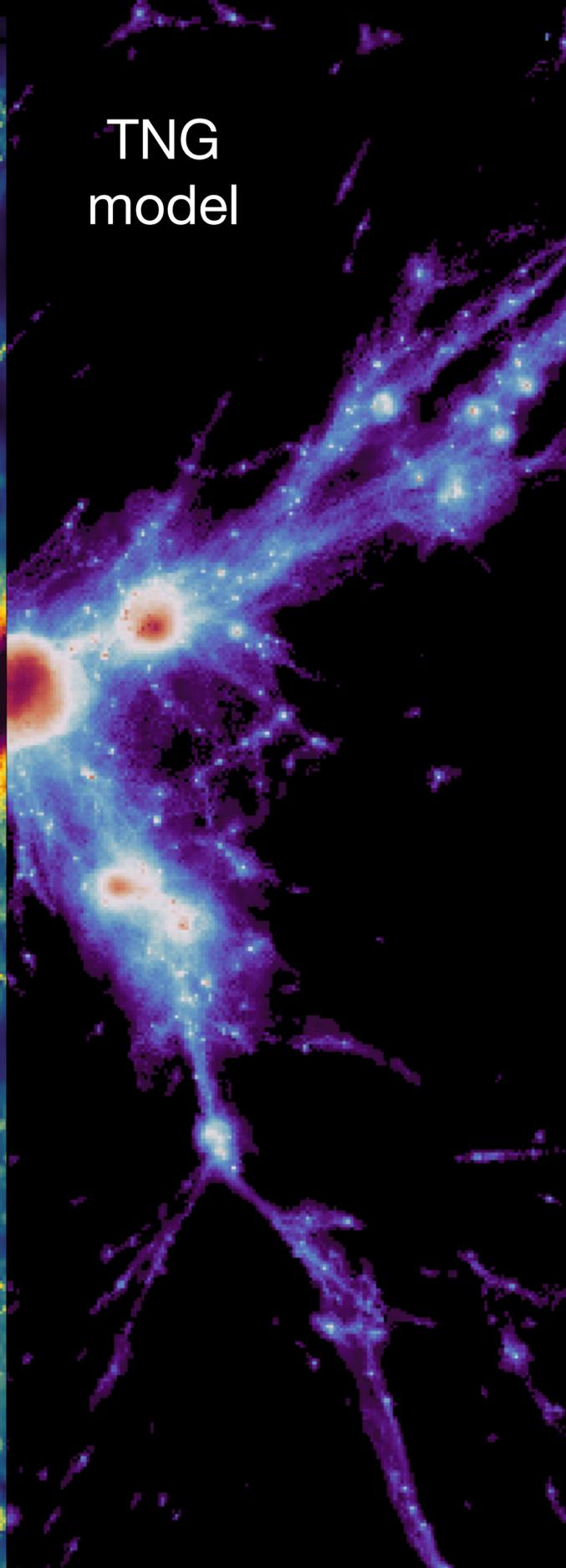
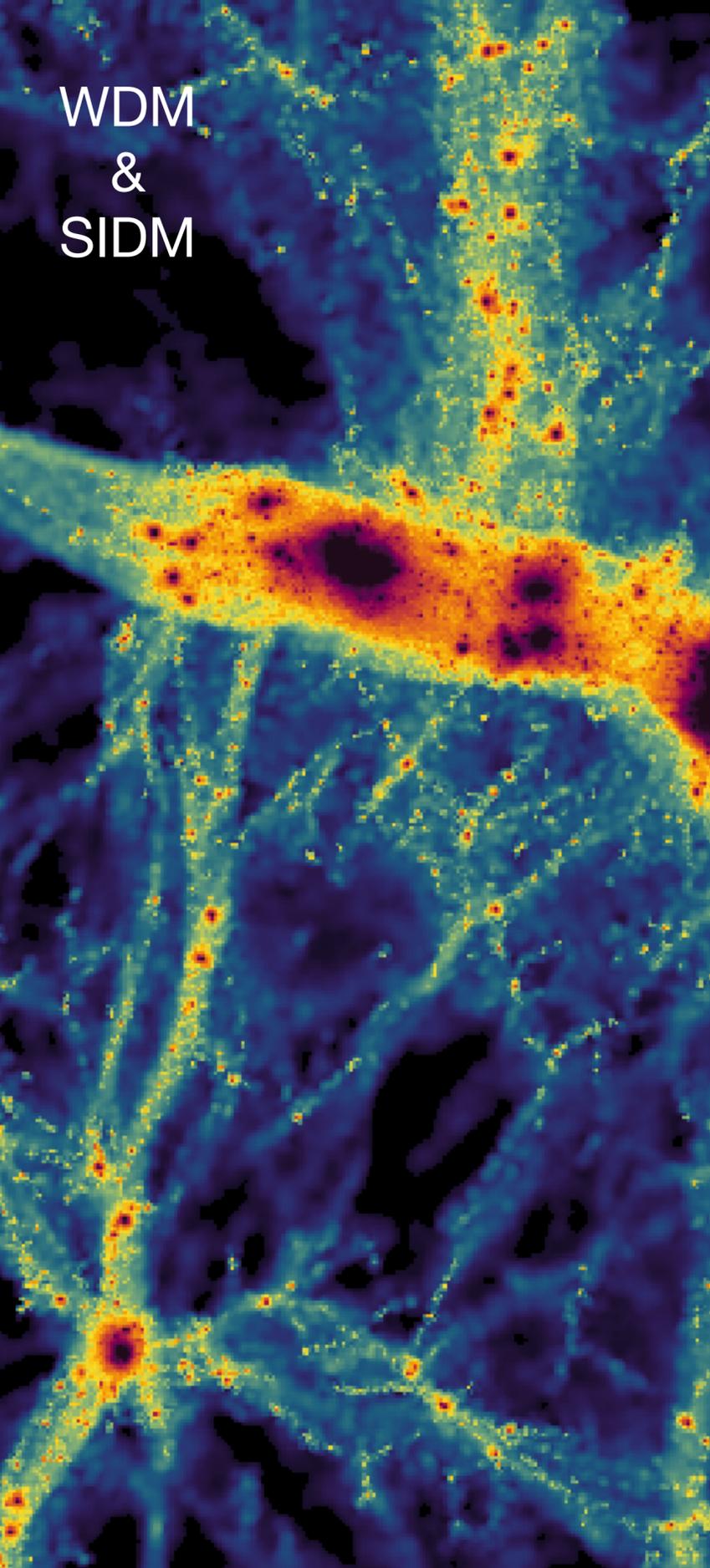
NEW SIMULATIONS



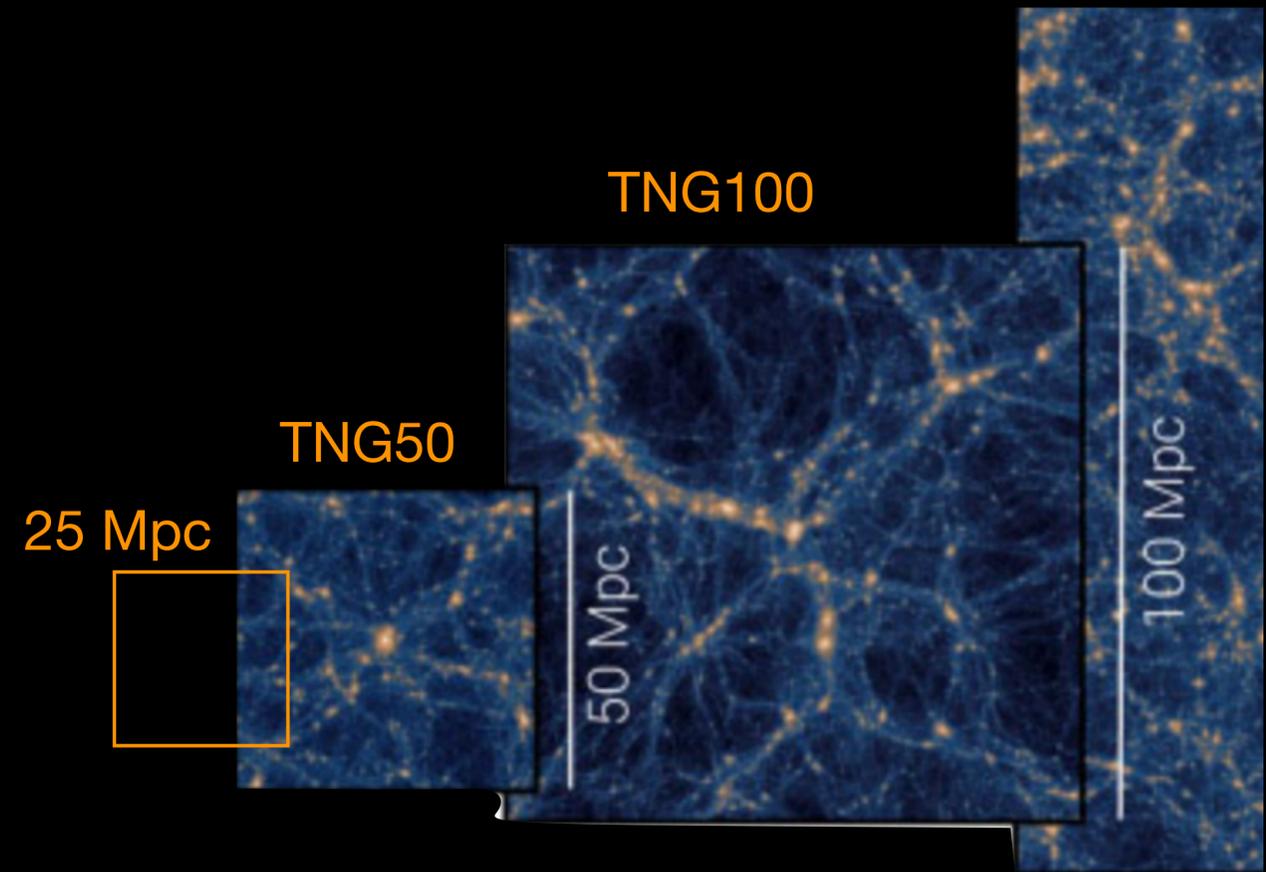
baryonic effect
very similar
in all models

NEW SIMULATIONS





NEW SIMULATIONS

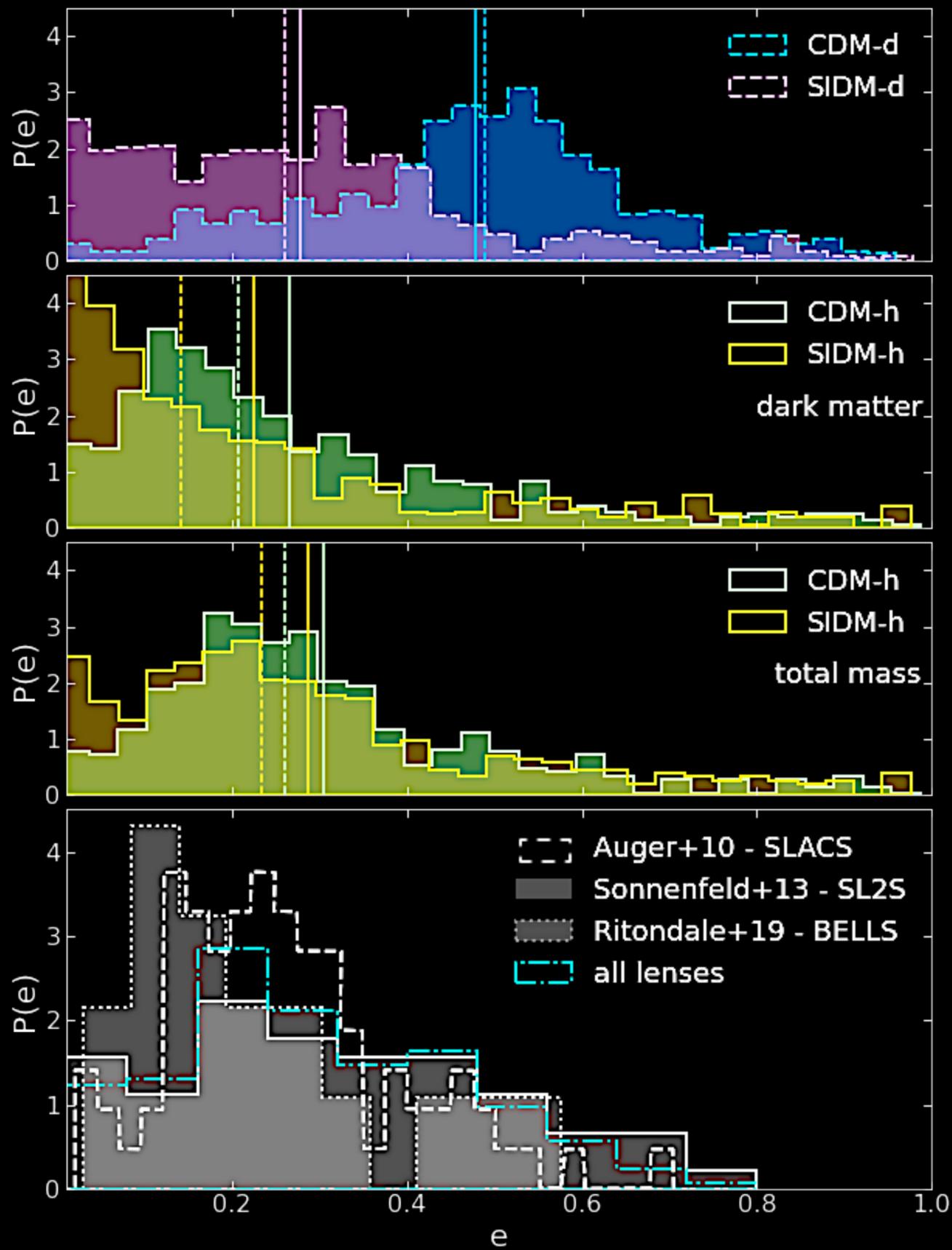


max resolution: $4 \times 10^5 M_{\odot}$, 280 pc

WDM, SIDM (constant, velocity-dependent & inelastic)

Despali et al. in prep: Annalisa Pillepich, Dylan Nelson, Volker Springel, Lauro Moscardini

PROJECTED SHAPES VS LENSING



DM-only

DM in hydro runs

total mass in hydro runs

lensing observations

p -value

SL2S

BELLS

SLACS

ALL

CDM-d

SIDM-d

CDM-h

SIDM-h

9.94e-09

0.23

0.17

0.11

7.90e-07

0.25

0.12

0.51

1.67e-15

2.81e-03

0.45

0.014

5.78e-25

9.00e-03

0.67

0.071

(Despali et al. 2022)

HALO PROFILES VS LENSING

log(M)~12-12.5

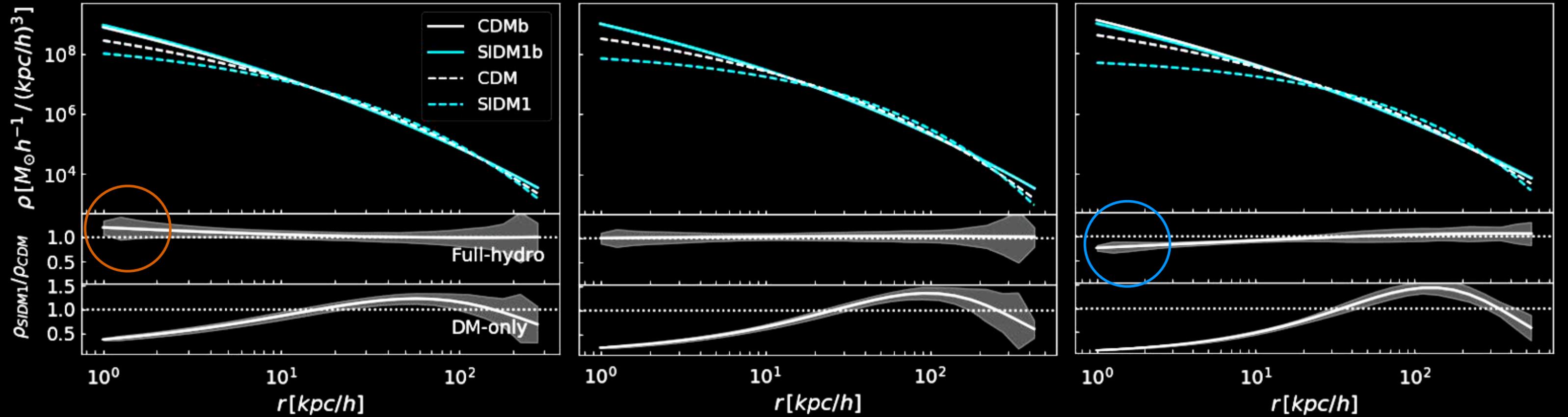
log(M)~12.5-13

log(M)~13-13.5

Group 1

Group 2

Group 3



(Mastromarino, Despali et al. 2023)