



Self-Interacting Dark Matter:

Bound State within SIMPs

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I. from SIDM to SIMP

We only see dark matter from the sky.

Almost everything fits at large-scales, while several puzzles at **small-scale** remain.

 DM cores preferred by observations in many (sub-)halos [Moore 1994; Burkert 1995, Newman et al 2013, ...] (core/cusp problem)

2. Non-observation of massive sub-halos which should host brightest dwarfs [M.Boylan-Kolchin et al. 2011, 2012, Ferrero et al. 2011]

(too-big-to-fail problem)

3. Some globular/star clusters expected to be destroyed, or sink to halo centre for cuspy profile [J. Binney & S.Tremaine 2008, F. Contenta et al. 2017, P. Boldrini et al. 2018, ...]

(GC timing problem)



4. Diversity of galaxies,

Possible explanations:

Systematic uncertainties?

Baryonic effects (by bursty star formation)?

 $10^{53} - 10^{55}$ erg

Self-interacting dark matter (SIDM)?

Observational evidence for self-interacting cold dark matter

D.N. Spergel and P J. Steinhardt [astro-ph/9909386]

Infalling dark matter is scattered before reaching the center of the galaxy so that the orbit distribution is isotropic rather than radial. These collisions increase the entropy of the dark matter phase space distribution and lead to a dark matter halo profile with a shallower density profile.



Strong DM self-scattering at dwarfs



inner halo DM self-thermalization (heating up the halo center)

O(1) scatters per DM particle

Popular models of SIDM:

1. SIDM via a light mediator [D.N. Spergel & P. J. Steinhardt 1999, J. Feng, M. Kaplinghat & H.-B. Yu 2009, ...]



t-channel elastic scattering & light mediator enhances cross sections

2. SIDM via Breit-Wigner resonance [M. Ibe & H.-B. Yu 2009, E. Braaten & E.W.Hammer 2013, XC, C. Garcia-Cely, H.



3. N-to-2 strongly-interacting-massive-particles (SIMP) [E. Carlson, M. Machacek & L. Hall,

1992, A. de Laix, R.Scherrer & R.Schaefer 1995, Hochberg, Kuflik, Volansky & Wacker 2014, ...]



kinetic equilibrium with radiation & relic abundance requires strong interaction.

Zoom-in strongly-interacting-massive-particles (SIMP)



- 1. Very-constrained sub-GeV mass range;
- 2. Only constant self-scattering.

$$\frac{\sigma_{\rm SI}}{m_{\rm DM}} \simeq \frac{1}{m_{\rm DM}^3} \simeq \frac{1}{(60\,{\rm MeV})^3}$$

Meanwhile, order-one interactions indicate potentially on-shell bound state formation;



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II. The framework: SIMP + X



step-1: formation of bound state, X

* Use Bethe-Sapleter equation at non-relativistic limit [e.g. K.Petraki, M.Postma&J.de Vries 2016, ...]:

Set all initial states to be at rest, one obtains a t-channel enhancement:

$$\frac{s}{t - m_{\pi}^2} \propto \frac{m_{\pi}^2}{m_X^2 - 4m_{\pi}^2} \propto \frac{m_{\pi}}{E_B} \gg 1$$



step-1: formation of bound state, X

* Use Bethe-Sapleter equation at non-relativistic limit [e.g. K.Petraki, M.Postma&J.de Vries 2016, ...]



- 1. Only **even-number** vertices involved, so far;
- 2. The rate of $3\pi \rightarrow \pi X$ can be much **larger** than the (v-suppressed) rate of $3\pi \rightarrow 2\pi$.

step-2: mass reduction with bound state, X



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One-step freeze-out: numerical solutions

If consider the Wess-Zumino-Witten (WZW) term with simple parametrization:



1. When the $3\pi \rightarrow \pi X$ is larger than H, chemical equilibrium lasts;

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- 1. When the $3\pi \rightarrow \pi X$ is larger than H, chemical equilibrium lasts;
- 2. At $\Gamma_{3\pi \to \pi X}$ = H: free- π abundance freezes,

X abundance keeps decreasing (EQ, then quasi-EQ).

2-body processes **generally** sufficient before BBN/CMB. **But not in this case** with even-number vertex, when

 $Y_X \ll Y_\pi$

here the bottleneck for DM freeze-out is more complicated:



Nevertheless, it is appealing: odd-number vertex is not necessary for SIMP.

It thus applies to fermionic DM, allowing much larger masses than direct $4 \rightarrow 2$ SIMP [N.Bernal & XC, 2015].

If only consider even-vertex interactions (neither Yukawa nor WZW):



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 - 2. Fast $3\pi \leftrightarrow \pi X$ leads to **non-zero** chemical potential;

Boltzmann-suppression only from **binding energy**: $Y_{\pi}^2 = Y_X \frac{(Y_{\pi}^{EQ})^2}{Y_Y^{EQ}}$

 $\sim Y_X e^{|E_B|/T}$

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- 1. First decoupling happens when $XX \rightarrow \pi\pi$ can not change free- π abundance;
- 2. Fast $3\pi \leftrightarrow \pi X$ leads to **non-zero** chemical potential;
- 3. Second decoupling from insufficient $3\pi \leftrightarrow \pi X$ (while X re-starts to decrease via XX $\rightarrow \pi\pi$).





II. Realization of SIMP + X

Requirements on model realization



*** X** needs to be long-lived

1. <u>suppress $X \rightarrow \pi\pi$ </u>: bound states appear in QCD,

but they are usually unstable [e.g. glueball-onium, tetraquark].

Easy to achieve with $m_X \leq 2m_{\pi}$, induced by dark-

QCD with heavy quarks, or other short-range forces [e.g. G.Kribs&E.Neil 2016, Y.Tsai, R.McGehee&H.Murayama 2020, R. Mahbubania, M.Redic&A.Tesi 2020,].

Suppress X → radiation: number-changing
 processes generally leave X-stability symmetry unprotected. Thus, kinetic equilibrium of DM and
 (dark) radiation may force X quickly decay to radiation.

Not easy to achieve [\mathbb{Z}_4 DM is safe, where DM-R scattering is tricky].

Requirements on model realization



* X needs to be long-lived

But, there are several ways out:

a) introduce heavier intermediate sates, \mathbf{K}^{\pm}



Extra mass of K may come from heavier dark-quark, additional-charge corrections, or K is X ...

b) impose quantum number for X (e.g. dark pion case)

 $X_{ij} = [\pi_i \pi_j]$

Energy leak via $[\pi_i \pi_i]$ could last:

improve by enhancing off-diagonal potential?

Requirements on model realization



Boost the formation of **X**

 $3\pi \rightarrow \pi X$ needs three initial states:



- 1. Introducing **nearly-massless gauge boson** strong enough to make DM in **kinetic equilibrium** with (dark) radiation [e.g. N.Bernal, C.Garcia-Cely&R.Rosenfeld 2015, ...];
- 2. Additional channel $\pi\pi \rightarrow VV$ and/or long-liveness of X: $g_V \sim 10^{-7} 10^{-5}$



$$n_V \gg n_\pi$$



After taking a short-range potential:

In practice, all information is embedded in the **wave-functions**, or for ground state, the effective **volume of the potential**:

$$\psi_X^\star(\vec{r}=0) \sim rac{1}{\sqrt{V_{
m pot.}}}$$

Relevant Processes with bound states:

1. As mentioned already one takes Peskin's notation (or Bethe-Sapleter equation):

$$\pi \longrightarrow M(p_1, p_2, p_3 \to k, Q)_{3\pi \to \pi X}$$

$$\approx \frac{1}{\sqrt{\mu_{re.}}} \int \frac{d^3q}{(2\pi)^3} \tilde{\psi}_X^{\star}(\vec{q}) \int \frac{dq_0}{2\pi} \frac{S(q;Q)}{S_0(\vec{q};\vec{Q})} \times M_{(p_1,p_2,p_3 \to k,Q/2+q,Q/2-q)}^{\text{free}}$$

$$\pi \longrightarrow \frac{\sqrt{2m_X}}{2m_\pi} \int \frac{d^3q}{(2\pi)^3} \int d^3r \, \psi_X^{\star}(\vec{r}) \, e^{-i\vec{q}\vec{r}} \times M_{(p_1,p_2,p_3 \to k,Q/2+q,Q/2-q)}^{\text{free}}$$
• For constant $M: \, \psi_X^{\star}(\vec{r}=0) \sim \frac{1}{\sqrt{V_{pot.}}}$
• For $M \propto p_i p_j: \, \nabla_i \nabla_j \psi_X^{\star}(\vec{r}=0)$ divergent?
non-vanish for ground-state wave-function
w.r.t. a finite-range spherical potential.
• For $M \propto p_i p_j: \, \nabla_i \nabla_j \psi_X^{\star}(\vec{r}=0)$ divergent?
to be **regularized** to $\sim |m_\pi E_{\mathbf{B}}| \psi(0) | e.g. \, N.Brambilla, D.Eiras et al, 2002, S. Biondinia&V.Shtabovenko 2020, ..].$

After taking a short-range potential:



3. All free-states processes, e.g. WZW-like interaction



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IV. Conclusions

Conclusions

- Halo mass deficit may be hint of DM self-interaction, such as SIMP
- SIMP involves N-body processes, where **bound state X** could play a role
 - automatically induce velocity-dependence;
 - modify $N \rightarrow 2$ freeze-out significantly;
- Freeze-out enhanced by on-shell X and t-channel resonance
 - w/o odd-number vertex: one-step/two-step;
 - bottleneck is **three initial-states**: better with extensions;
- Concrete model has to satisfy several conditions, leading to constraints
- X can be a **fundamental** particle, sitting near 2π -resonance by accident

Thanks!