ARTIFICIAL INTELLIGENCE AND MACHINE LEARNING IN NEUTRINO PHYSICS

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22nd International Workshop on Next Generation Nucleon Decay and Neutrino Detectors (NNN23)

Procida, 13 August 2023

Why do we need machine learning?

- Neutrinos are elusive particles that rarely interact with matter, making their detection and study incredibly challenging.
 - Neutrino experiments need to collect massive amounts of data in order to see a handful of neutrino events.
 - This data is also very complex, as neutrino interactions can produce a variety of different signatures.
- Challenges for standard algorithms:
 - Complex pattern recognition: standard algorithms often struggle to identify intricate patterns in the data.
 - **Potential biases**: human implementation can introduce biases that affect results.
- Machine learning (ML) promise:
 - ML algorithms have the capability to learn from data (simulation) and improve their performance over time.
 - Expect ML to not only meet the challenges but also potentially lead to groundbreaking discoveries in neutrino physics.

But... what is machine learning?

- ML is a subfield of artificial intelligence (AI).
 - AI: branch of computer science that aims to build algorithms capable of performing tasks typically (traditionally) accomplished using human intelligence.



*Source: <u>Stack</u> <u>Exchange</u>

- ML is learning from data.
 - There is no learning without data.
 - ML algorithms only learn from the data.





Neutrinos and images

- Neutrino experiments often involve complex detectors that allow us to collect data in the form of images.
 - Typically serves the purpose of detecting and characterizing events generated by neutrino interactions.
 - Examples: LArTPC detectors (DUNE, MicroBooNE, ICARUS), Cherenkov detectors (Icecube, SuperK/HyperK), scintillators (NOvA, T2K/HyperK)...
 - Require powerful algorithms to deal with such complex data.









Computer vision

- Study of visual data (images and videos).
 - A computer "sees" a grid of numbers.
 - Massive amount of visual data produced every day.
 - Origin in the late 50s.
- Deep learning has been the dominant approach in computer vision research for the past decade.
 - Applications in many areas: automotive, healthcare, robotics, media, agriculture, security, physics...
- What about MLPs?
 - Fully-connected neural networks (FCNNs), also known as dense neural networks or multi-layer perceptrons (MLPs).
 - Type of artificial neural network where each neuron in one layer is connected to every neuron in the next layer.
 - Require a fixed-size one-dimensional input, resulting in an extremely large number of parameters.
 - In the example, the input (flattened) has 3,072 values. If the first layer has 1,000 neurons, that's 3,072,000 parameters (without bias) for only the first layer!
 - They are not translation invariant.



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Hidden layer:

1000 neurons

Input:

32x32x3

[Source: Openframeworks]

Output layer:

Image recognition (intuition)

• Humans break down images into different parts before assembling the information back together.



It is orange with black stripes: it is a tiger.



It is gray with a long trunk: it is an elephant.



It is gray with a long trunk: it is an elephant.

- Convolutional neural networks, or CNNs, are a type of neural network architecture specifically designed for image recognition tasks in computer vision.
 - Convolution: element-wise multiplication and sum of the overlapping elements between the kernel and the input.
 - CNNs use a series of convolutional layers to extract hierarchical features from images.
 - CNNs have achieved state-of-the-art performance in various computer vision tasks.
 - They are (kind of) translation invariant!





CNN example in neutrino physics: the DUNE CVN



Understanding the CVN



Removing the start of the electron shower reduces the CC ν_e score, as expected

The CVN finds the vertex a bit ambiguous, but it uses the end point of the muon to gain a handle on the event type.

Going further: handling sparse data

- In particle physics and astrophysics, data is often sparse due to the nature of the objects being studied or the particles detected.
- This poses a challenge for computer vision, as standard • CNNs are designed to work with dense data. To address this, researchers are developing new algorithms and techniques specifically tailored to sparse data.
 - For example, one approach is to use Submanifold Sparse Convolutional Networks (SSCN), where the convolution operation is performed only on the non-zero elements of the sparse data, resulting in an efficient computation.
 - Example: "Scalable, End-to-End, Deep-Learning-Based Data ٠ Reconstruction Chain for Particle Imaging Detector" (arXiv:2102.01033)



"Dense" image



All pixels might be helpful for the classification.

Ideal for standard **CNNs.**

[Source: britannica.com]

"Sparse" image



Most pixels are background.

A standard CNN would perform loads of useless computations.

Another solution is to use Graph Neural Networks (see Adam's talk!)

Our current work!

- Case study: a detector concept analogous to the SuperFGD detector from the T2K experiment.
 - Part of the upgrade of the near detector (ND280) of the T2K experiment in Japan.
 - Fully-active fine-grained scintillator (FGD) with three views.
 - 2M optically independent cubes, 1 cm³ per cube.
 - Spatial localisation of scintillation light.



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- Goal: develop a deep-learning-based analysis strategy that does not depend on the neutrino interaction model:
 - a) Distinguish between single and multi-primaryparticle hits.
 - b) Fit the trajectory of single-particle objects.
 - c) Understand the activity at the vertex of neutrino interactions.

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Approach

a) Hit identification.

- b) Particle trajectory fitting.
- c) Vertex activity fitting.



Hit identification: single vs multi-particle hits

- Classify each individual hit as:
 - **Single-particle hit**: only one particle passes through the hit cube.
 - Multiple-particle hit: at least two different particles pass through the hit cube.
 - Other: crosstalk or ghost.
- Using a submanifold sparse U-Net-based neural network architecture (<u>https://arxiv.org/abs/1706.01307</u>).
 - More computationally efficient than standard CNNs.
- Efficiencies:

	True multi-particle	True single-particle	True other
Pred. multi-particle	0.7777	0.1511	0.0711
Pred. single-particle	0.0055	0.9654	0.0291
Pred. other	0.0079	0.0479	0.9442

• Excellent single-particle isolation accuracy allows running a further NN-based track trajectory fitting on single particles, relying on detailed MC simulations of single particles.





Approach

- a) Hit identification.
- b) Particle trajectory fitting.
- c) Vertex activity fitting.



Fitting particle trajectories

• Each algorithm outputs the fitted 3D trajectory point for each input hit.

•Sampling Importance Resampling Particle Filter (SIR-PF): first reconstructed hit used as prior. The likelihood relies on a precomputed 5-dimensional histogram.

Recurrent neural network (RNN): five bi-directional GRU layers, 50 hidden units.

Transformer: 5 encoder layers, 8 heads, hidden size of 64.

• Main results (<u>Commun. Phys 6, 119 (2023)</u>):



>30% better transformer resolution compared to the SIR-PF

• The improved trajectory fitting significantly improves the charge reconstruction, PID by range, and momentum by curvature.



 $y_1 = (x_1, y_1, z_1, E_1)$

 $\vec{S}_{in} = (x_i, y_i, z_i, E_i)_{\times \# hit}$

 $\vec{S}_{in,N} = (x_N, y_N, z_N, E_N)$

 $\vec{S}_{in,2} = (x_2, y_2, z_2, E_2)$

Approach

- a) Hit identification.
- b) Particle trajectory fitting.
- c) Vertex activity fitting.



Vertex activity: standard fitting method

- Vertex activity (VA): particles releasing their energy in the proximity of the neutrino interaction vertex but that do not show visible tracks from where the kinematics can be reconstructed.
 - A "blob" of scintillation light is observed.



• Standard VA fitting method:

•Goal: build the neutrino VA in forward folding from the sum of single particle reconstructed objects. Particle information to reconstruct: # of particles (mostly protons), energy, direction, vertex position.

- Method: likelihood fitting.
 - 1. Simulating any possible combination of the VA parameters and build VA.
 - 2. Finding the VA 3D image (e.g. SFGD hits) that "best fits" the data and find the "best-fit" parameters.

- The fitting method is highly computationally expensive.
 - Requires a large number of combinations of parameters to be simulated.
 - Unfeasible in practice.

Alternative: deep-learning approach



• Event config: 1 muon, 1-5 protons.

- 70% accuracy in reconstructing the correct number of particles.
 - >98% assuming a ± 1 error.
- ~2 mm vertex resolution.
- Good reconstruction of kinematics.



• Also tested configurations with nuclear clusters (deuterium, tritium)!

Alternative: deep-learning approach



- The generator (GAN) is fully differentiable.
 - Run a gradientdescent minimiser to optimise the particle kinematics further.



 $⁽KE_{true} - KE_{reco})/KE_{true}$

- \circ GAN-BF = Brute force.
- GAN-GD (avg) = Gradient-descent (average result).
- GAN-GD (best) = Gradient-descent (best result).

Summary

- Machine learning could be a key tool for neutrino experiments.
 - In particular, for detectors that provide fine interaction details but are hard to analyse using traditional methods.
- Strongly linked with computer vision.
 - Most of our data is in the form of images.
- Successful application to different problems, such as:
 - Flavour identification, final-state particle counting, track fitting, etc.
- Current and future work requires an extensive validation of the methods and application to experimental data.

Bonus

• Asked the Bing AI model (<u>https://www.bing.com/images/create</u>) to generate images from the following prompt:

"Neutrino interaction powered by AI happening at Procida Island in Italy".





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