



Dark Matter Searches with COSINE-100, present and perspectives

In-Soo Lee

Center for Underground Physics (IBS)

On behalf of the COSINE-100 collaboration

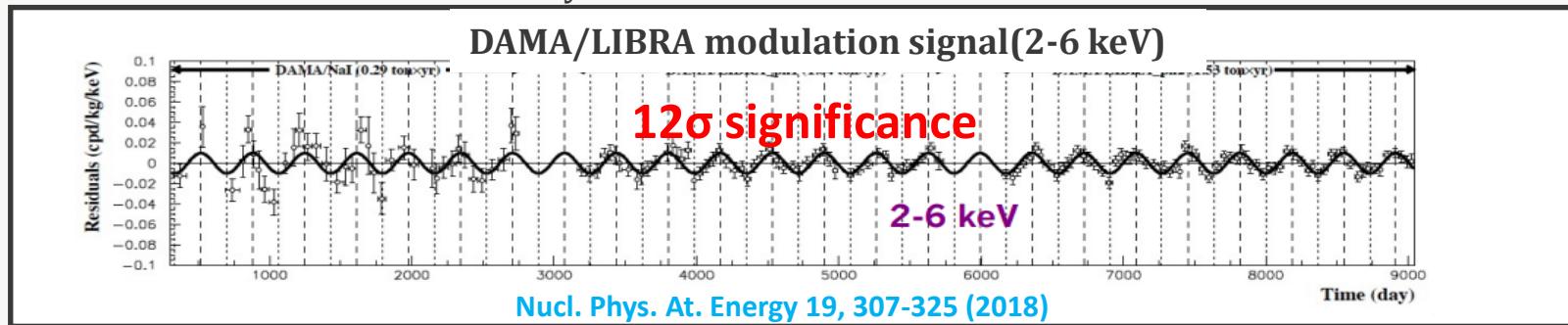
March 7th, 2023



Motivation-DAMA/LIBRA

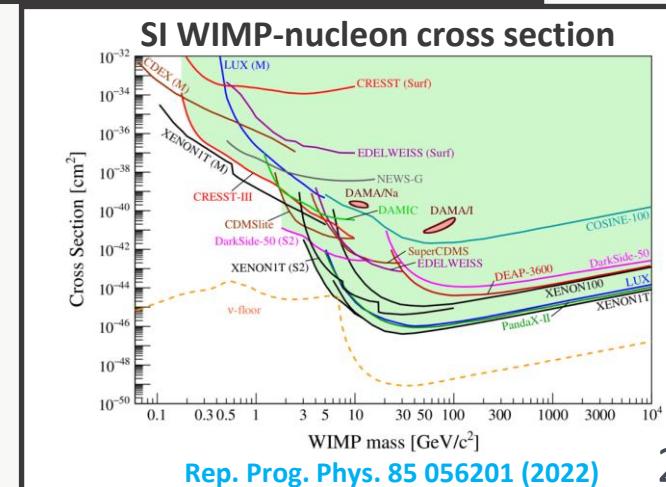
- **DAMA/LIBRA experiment**

- Search for **Dark matter (DM)** annual modulation signature
- Using 25×9.7 NaI(Tl) detectors
- **Claim an observation** of the DM (WIMP modulation signal) at **12σ C.L.** (2-6 keV, 2.17 ton · year)
- **Amplitude** : 0.0096 ± 0.0008 counts/day/kg/keV
 - Phase = 145 ± 5 days
 - Period = 0.9987 ± 0.0008 year

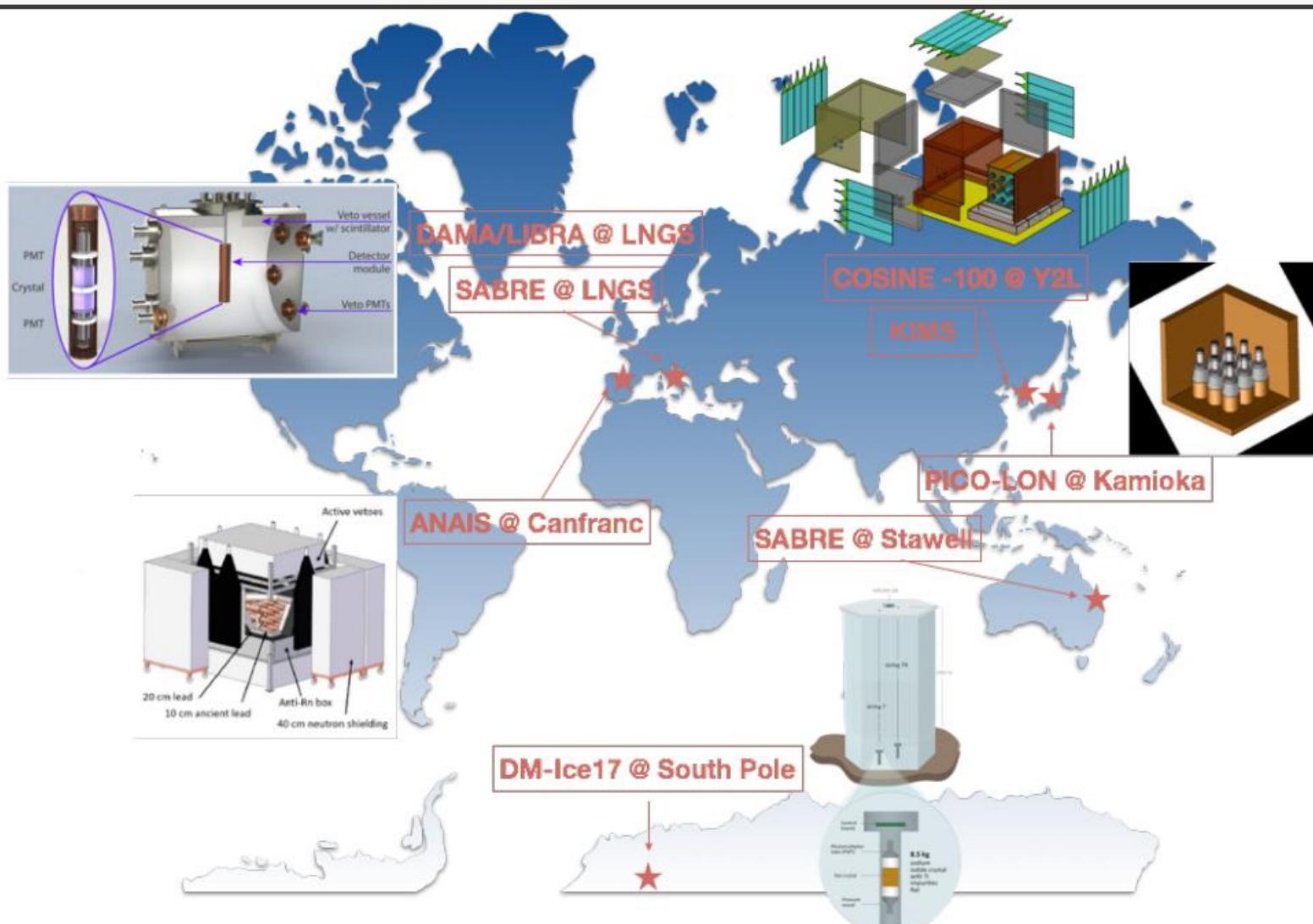


- **No other experiment** has succeeded direct detection of DM except for **DAMA/LIBRA**.
 - Is NaI(Tl) special for DM interaction?

COSINE-100, InSoo Lee



NaI(Tl) Dark matter search

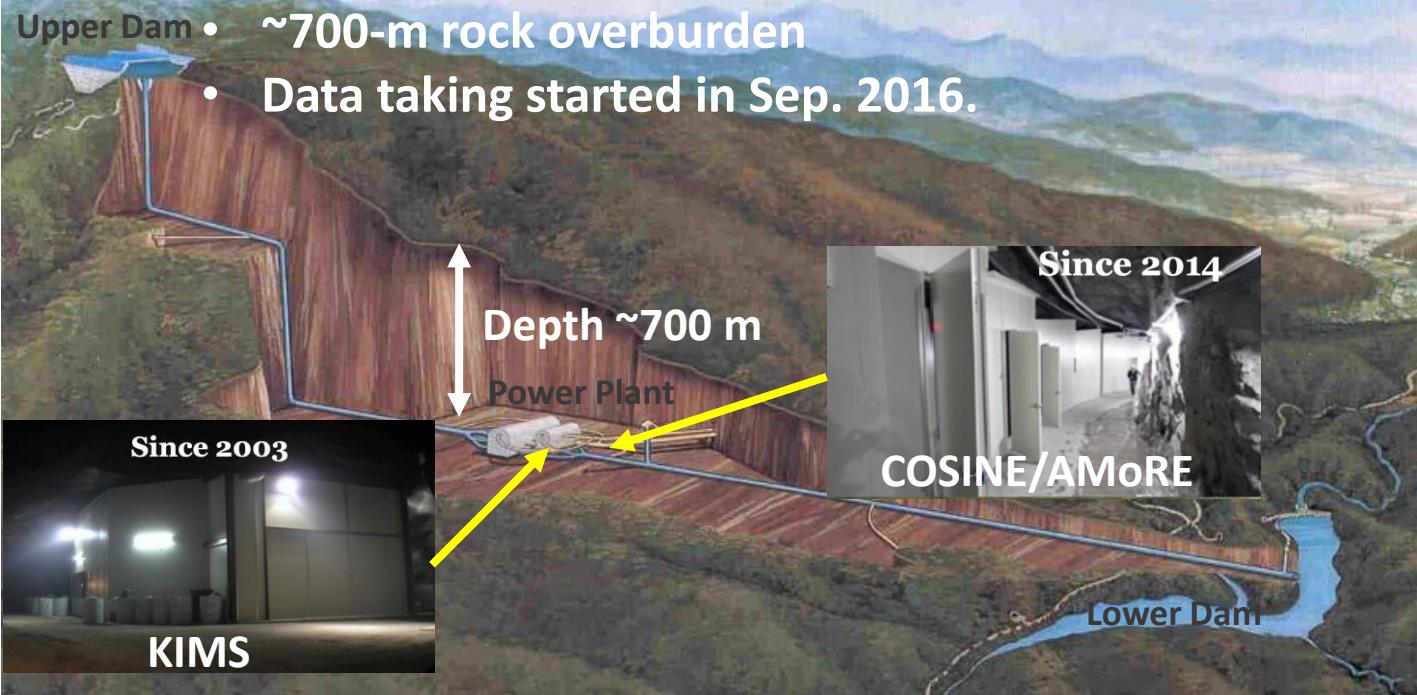


COSINE-100 experiment



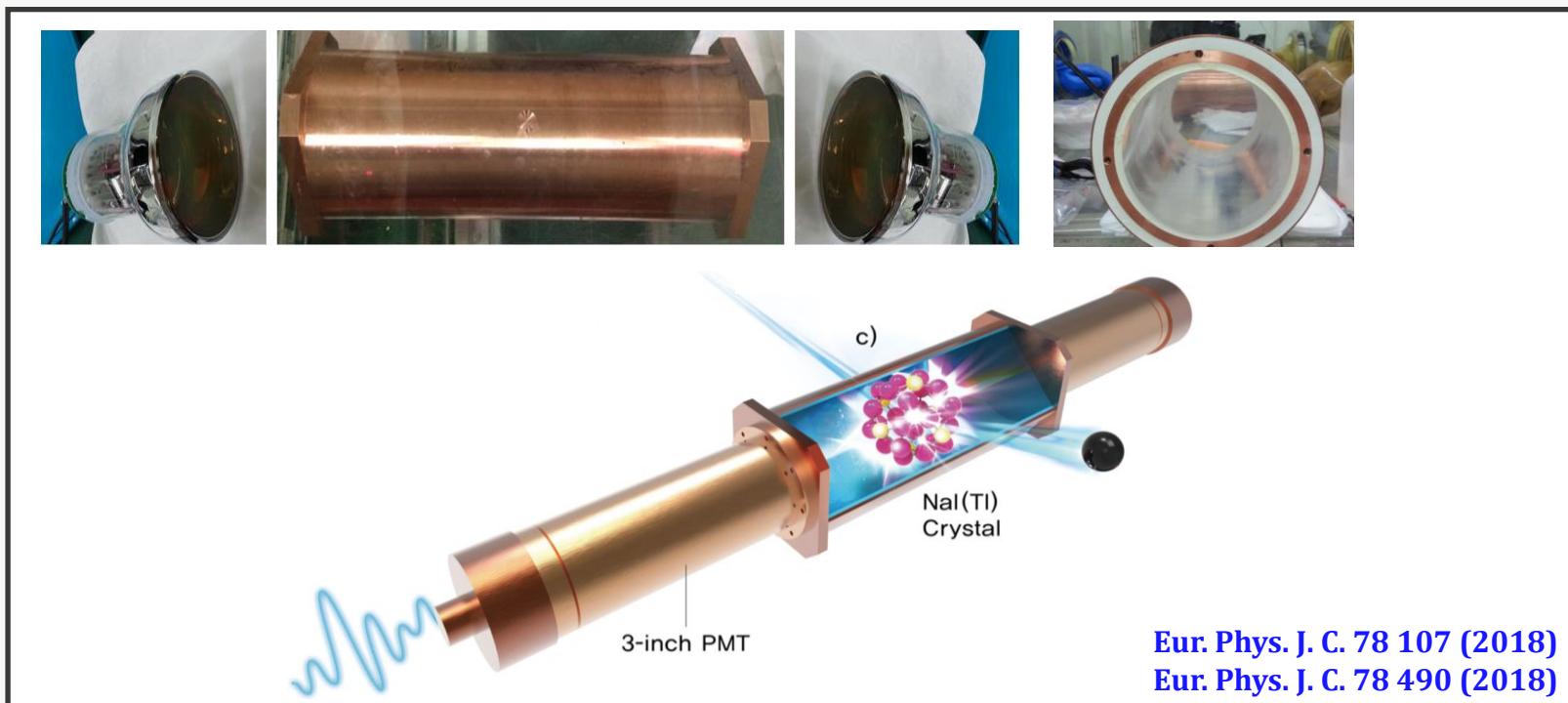
- Joint collaboration of DM-ICE & KIMS
- ~50 collaborators in 17 institutes
- To confirm DAMA/LIBRA's claim using same target material, NaI(Tl)

Yangyang underground laboratory (Y2L), Korea



COSINE-100 experiment-NaI(Tl)

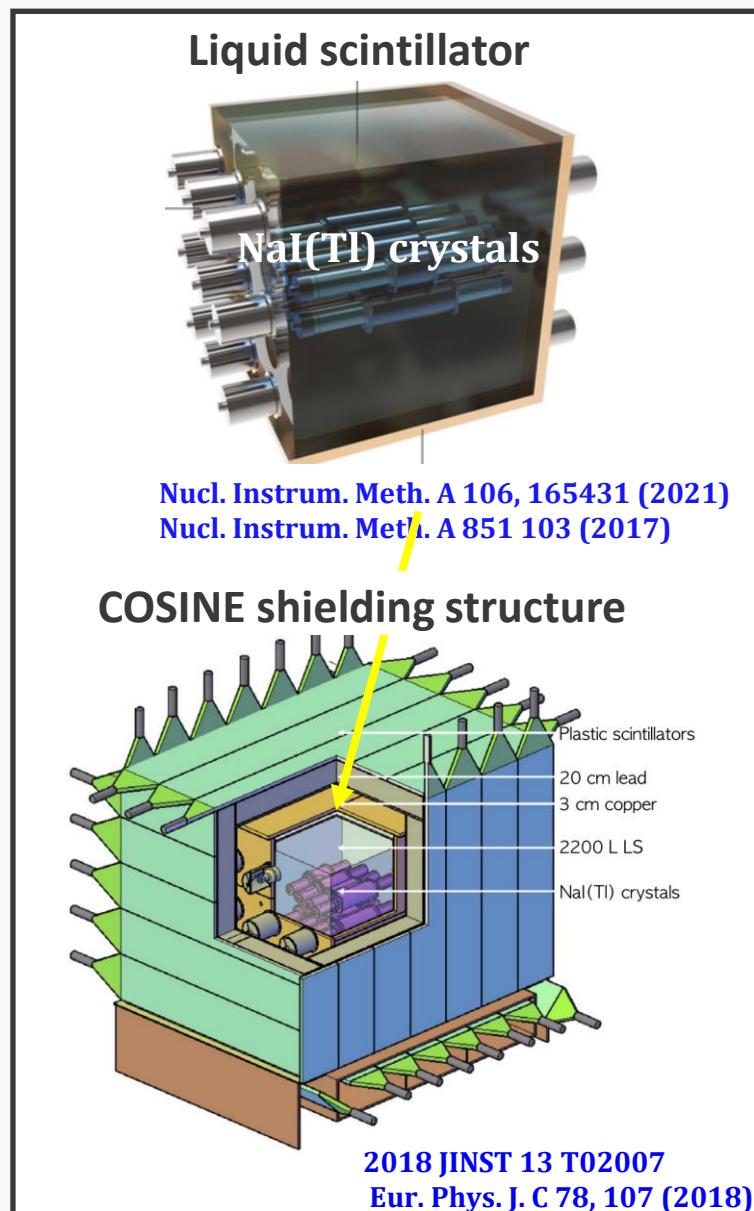
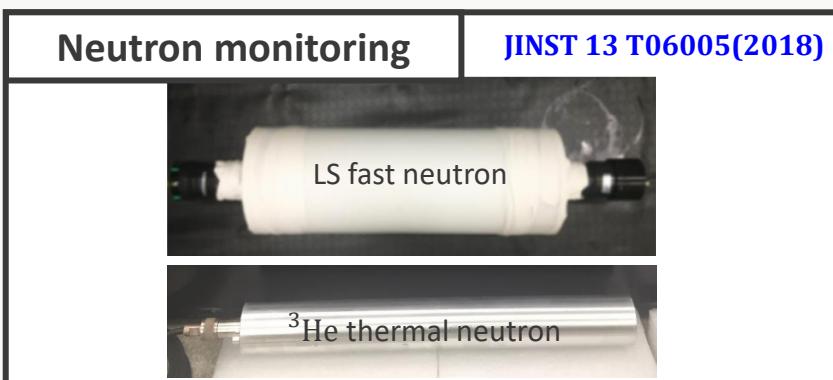
- 8 low-background NaI(Tl) crystals with 106 kg in total
 - U/Th/K level is less than DAMA, but total alphas (^{210}Pb) are higher than DAMA.
 - Total **background** level is 2-3 times that of DAMA/LIBRA.
 - **Higher light yield (15 P.E./keV)** than DAMA/LIBRA (5-10 P.E./keV)
 - Can make the threshold lower easily
 - Each crystal is **encapsulated in copper and quartz windows**.
 - Two 3-inch PMTs (R12669SEL) are attached to each crystal.
 - Quantum efficiency: 35% @ 420 nm



Eur. Phys. J. C. 78 107 (2018)
Eur. Phys. J. C. 78 490 (2018)

COSINE-100 experiment-Shielding

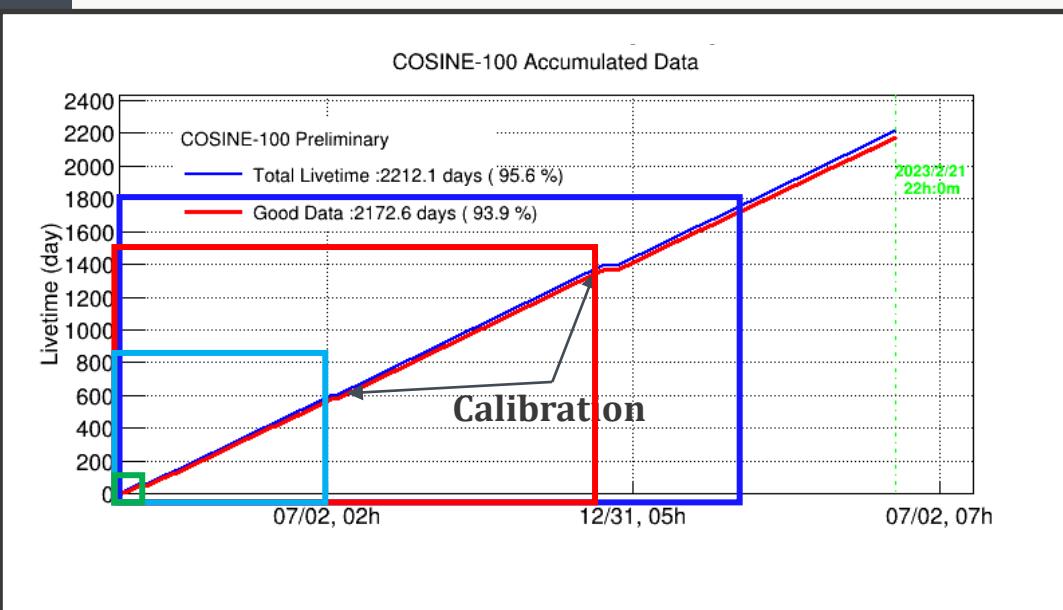
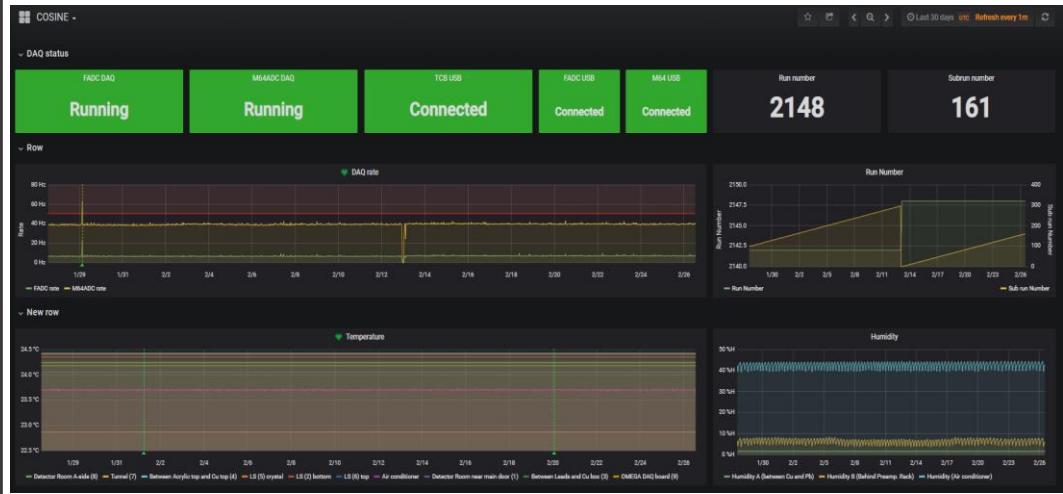
- Active veto
 - Liquid scintillator (LS)
 - 2200-L LAB-based LS
 - 5-inch PMT(R877) for LS detector
 - 4π Muon counter
 - 37 plastic scintillator panels
 - 2-inch PMT(H7195) for muon counter
- Passive veto
 - 3-cm thick copper box
 - 20-cm thick lead castle
- Neutron monitoring
 - Fast neutron detector (Liquid Scintillator)
 - Thermal neutron detector (^3He gas detector)



Operation & slow monitoring

JINST 17 T01001 (2022)

Slow monitoring



• Slow monitoring system

- > 200 parameters
- DAQ system trigger rate
- Electronics status
- Environmental parameters
- Neutron rate

• Stable run from Sep. 2016

- DAQ efficiency > 96% (calibration runs, power outage)
- Exposure time > 6 years

SET 1 : 59.5 days

SET 2 : ~1.7 years

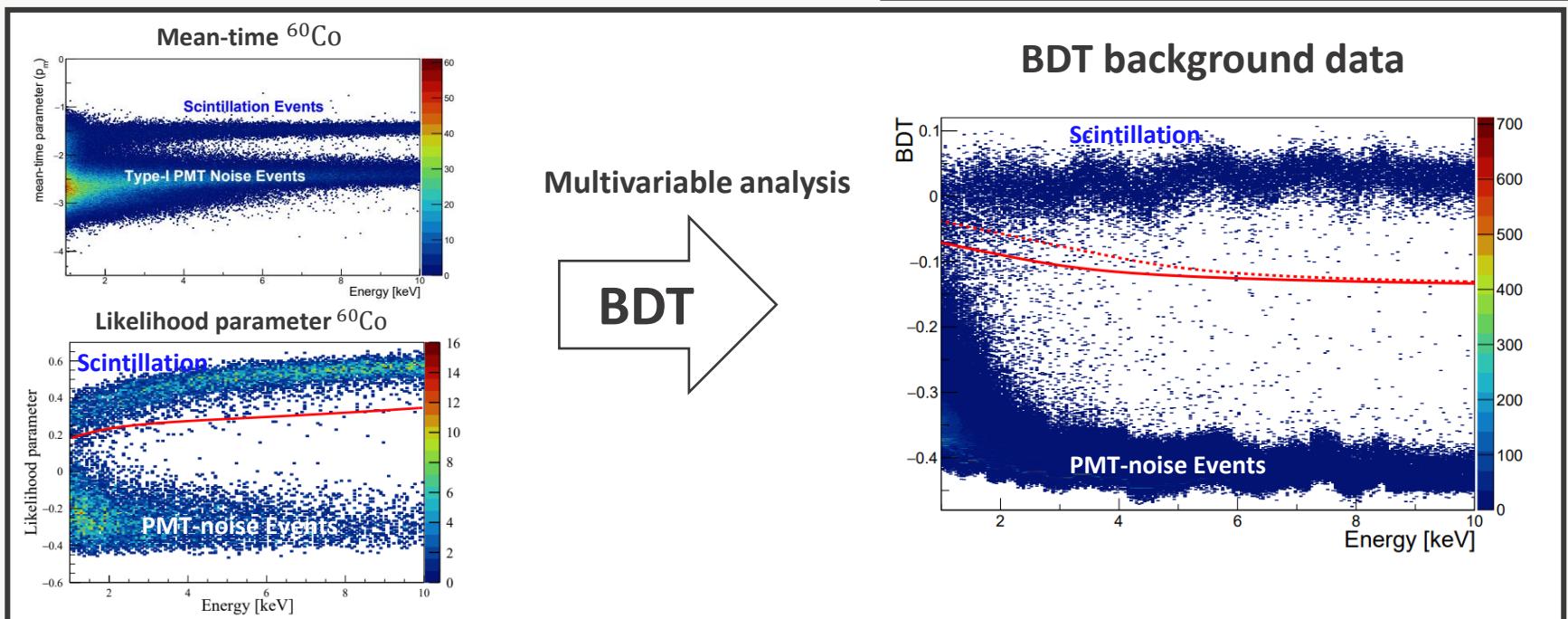
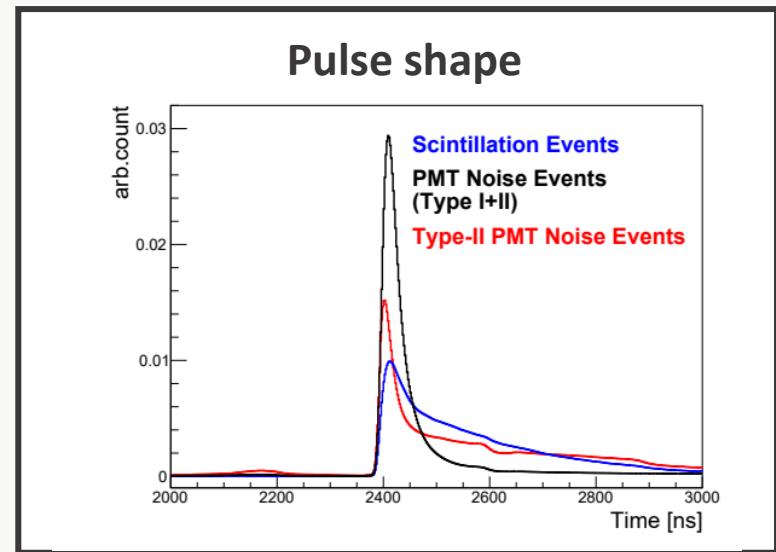
SET 3 : ~3.0 years

SET 4 : ~4.5 years

Event Selection

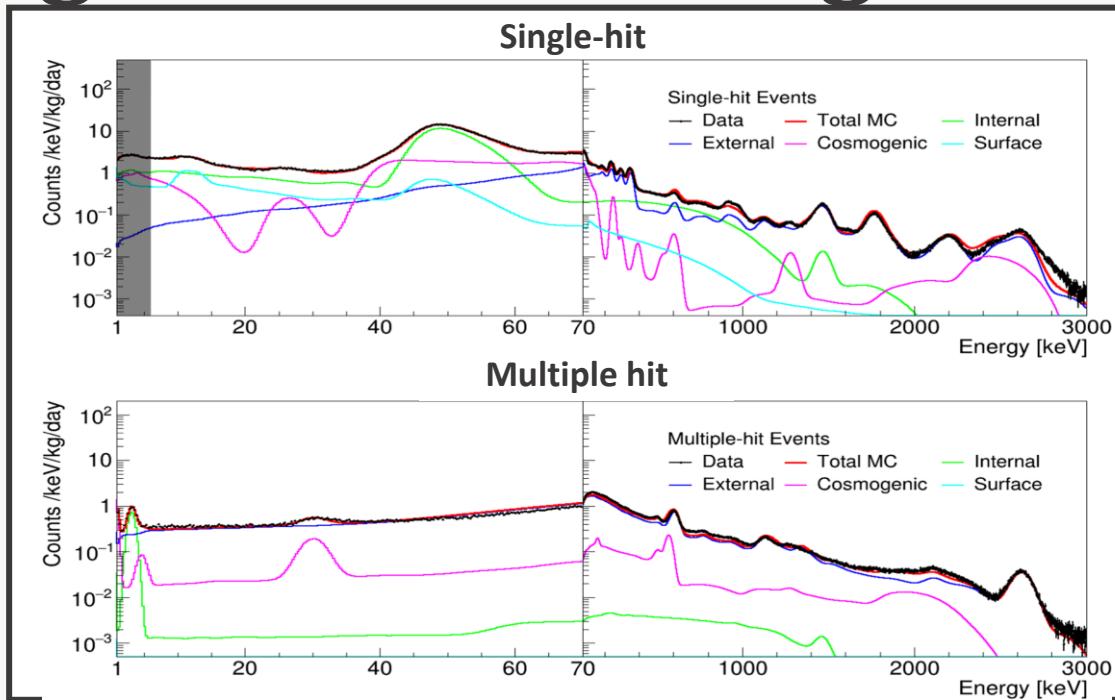
Astropart. Phys 130, 102581 (2021)

- Use Set 2 (1.7 year) data
- Use pulse-shape parameters to select NaI scintillation signals
 - Meantime, likelihood parameters
- Use Boosted decision tree (**BDT**) to separate PMT-noise from scintillation signal
- Achieve 1 keV threshold



Background Modeling

EPJC 81, 837 (2021)
Astropart. Phys. 126, 102528 (2021)



- SET2 1.7-year data with 1 keV Threshold
- Use detector simulation based on **GEANT4**
- Single (Multiple) Hit : Time coincidence w/o (w/) other crystal or LS
- Exclude **WIMP ROI** (<6keV single hit) in the fitting
- **Background** in low energy region
 - Cosmogenic activation (**mostly** ^3H)
 - Contamination of ^{210}Pb
 - Contamination of crystals from K/U/Th

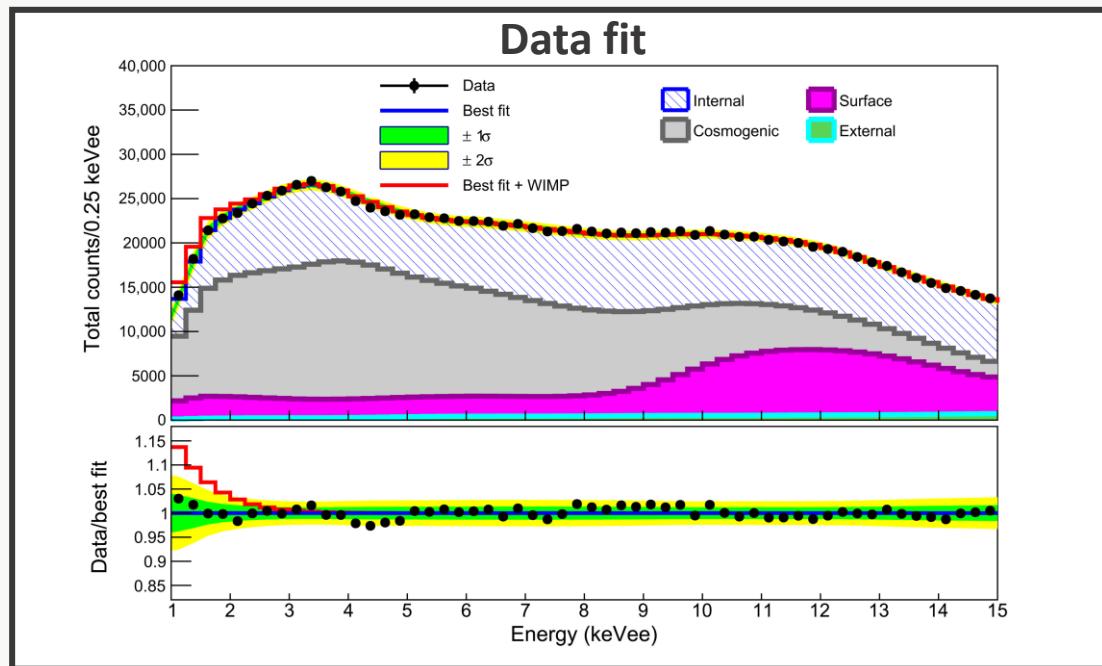


Main background

WIMP extraction

Science Advances. 7, 46 (2021)

- Spin-independent(SI) WIMP Search with SET2 data
 - 1.7 years data with 1 keV threshold
 - Fit background + WIMP signal data in [1, 15] keV

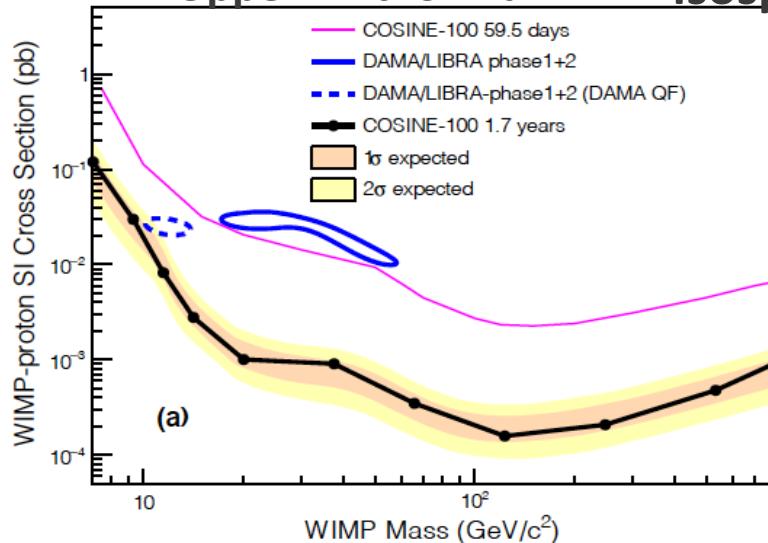


- No WIMP signal in 1 keV threshold

WIMP extraction

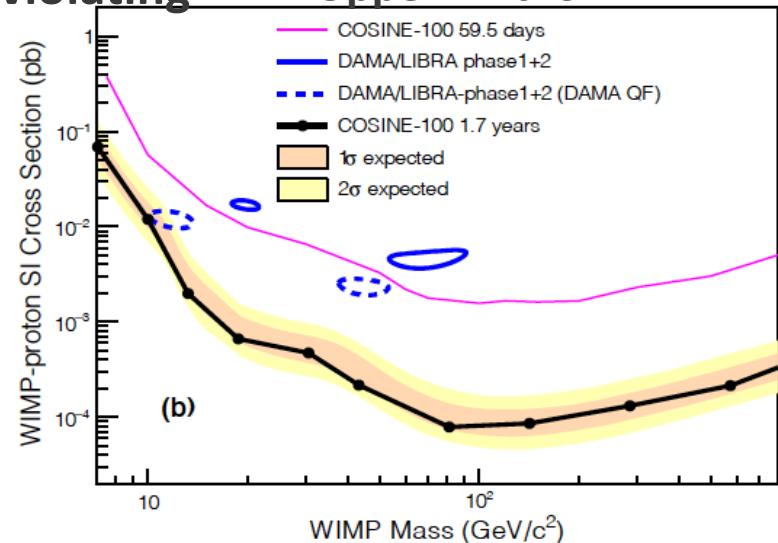
Science Advances. 7, 46 (2021)

Upperlimit for Na

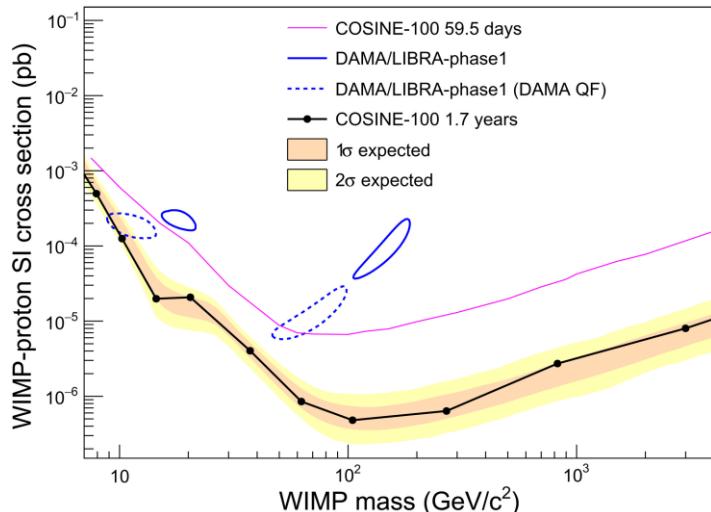


Isospin violating

Upperlimit for I



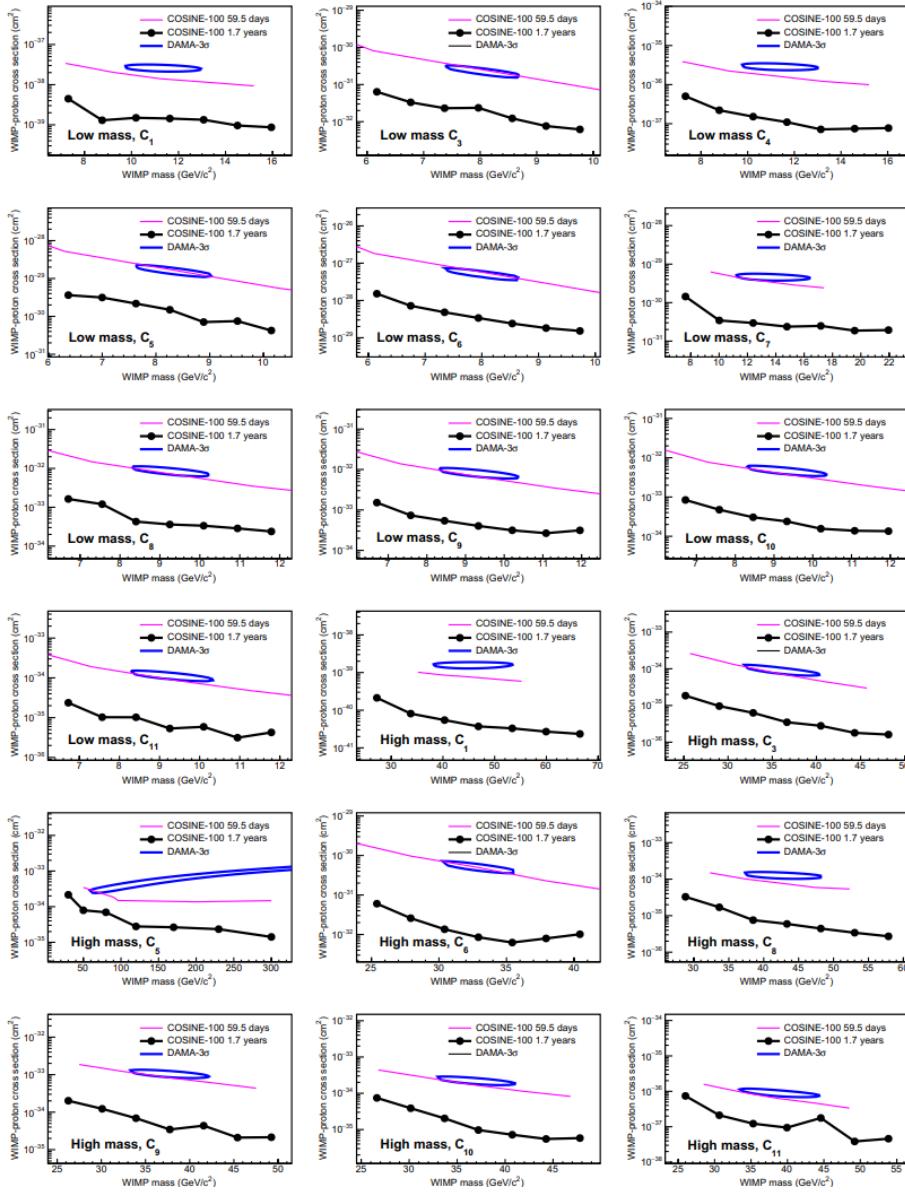
Isospin conserving



- **Exclude DAMA/LIBRA phase 1 considering quenching factor.**

WIMP extraction

Science Advances. 7, 46 (2021)

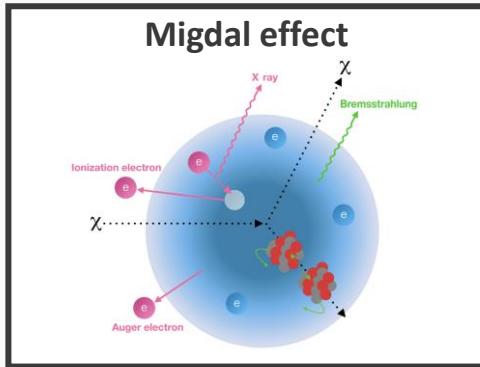


- Exclude possible EFT operators.

WIMP extraction using Migdal effect

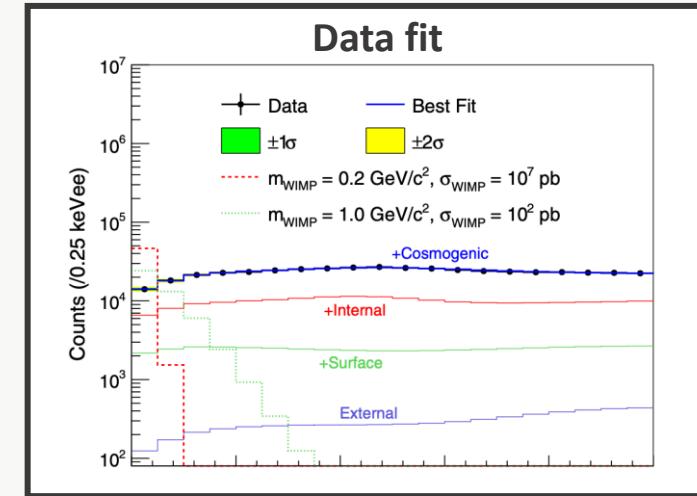
- Nuclear recoil → Boost of electrons → Secondary radiation

Phys. Rev. D 105, 042006

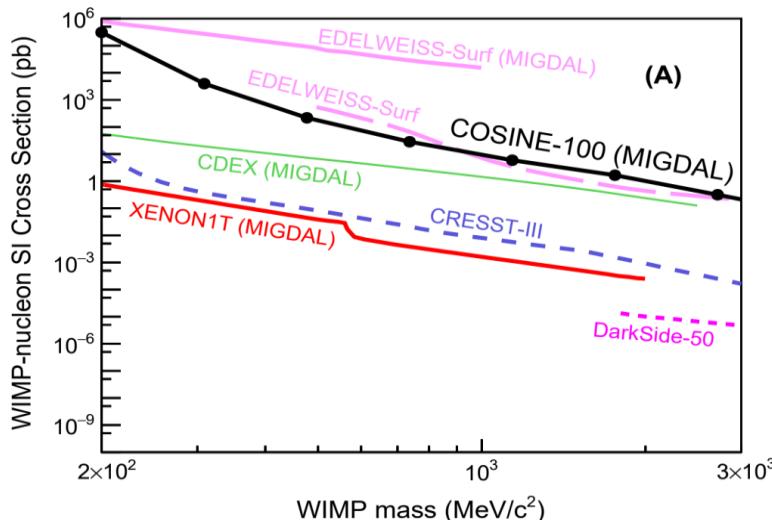


- Large visible energy of electron recoil compared to nuclear recoil.

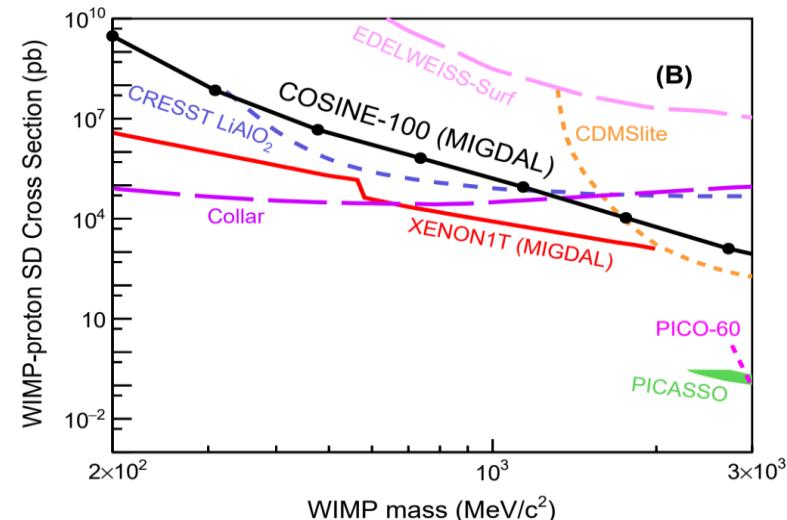
→ Enables to search low mass WIMP (100 MeV Scale)



Spin independent interaction



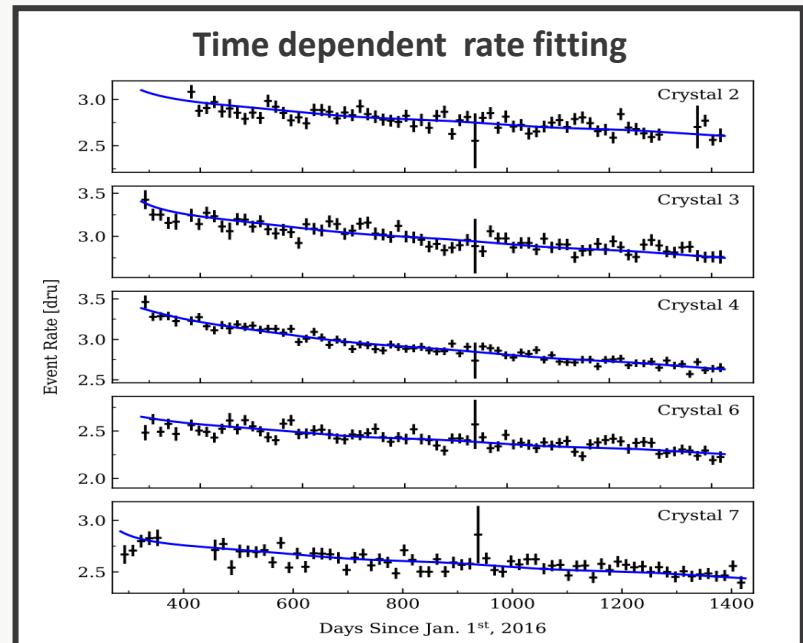
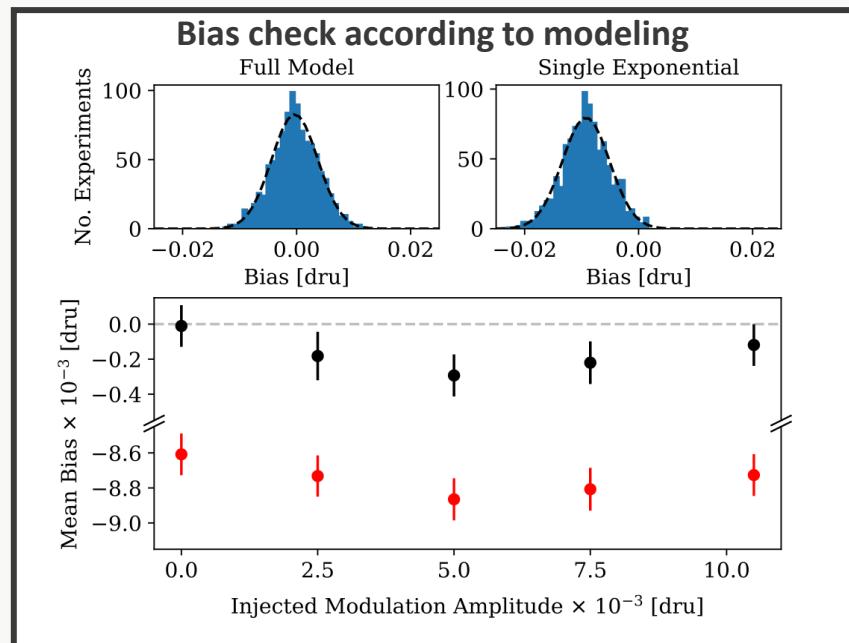
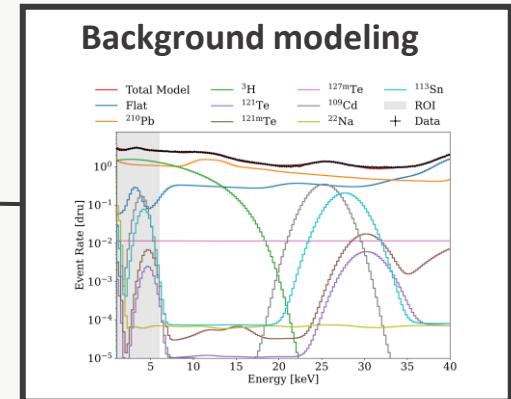
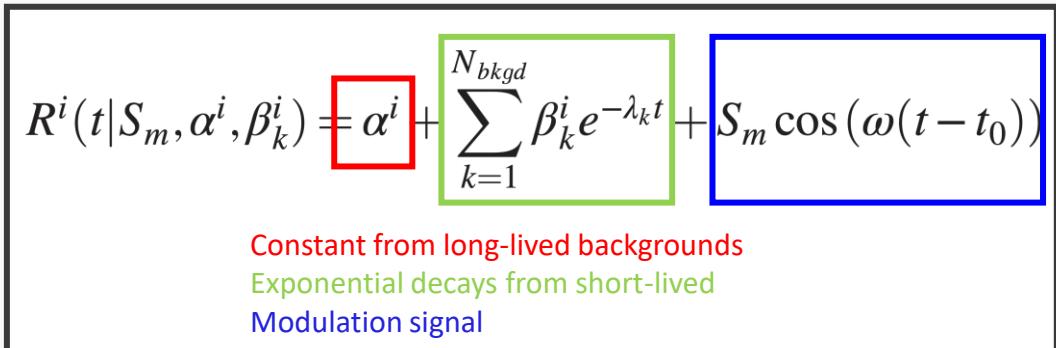
Spin dependent interaction (proton)



Annual modulation analysis (3 years data)

- Full model of time dependent background components

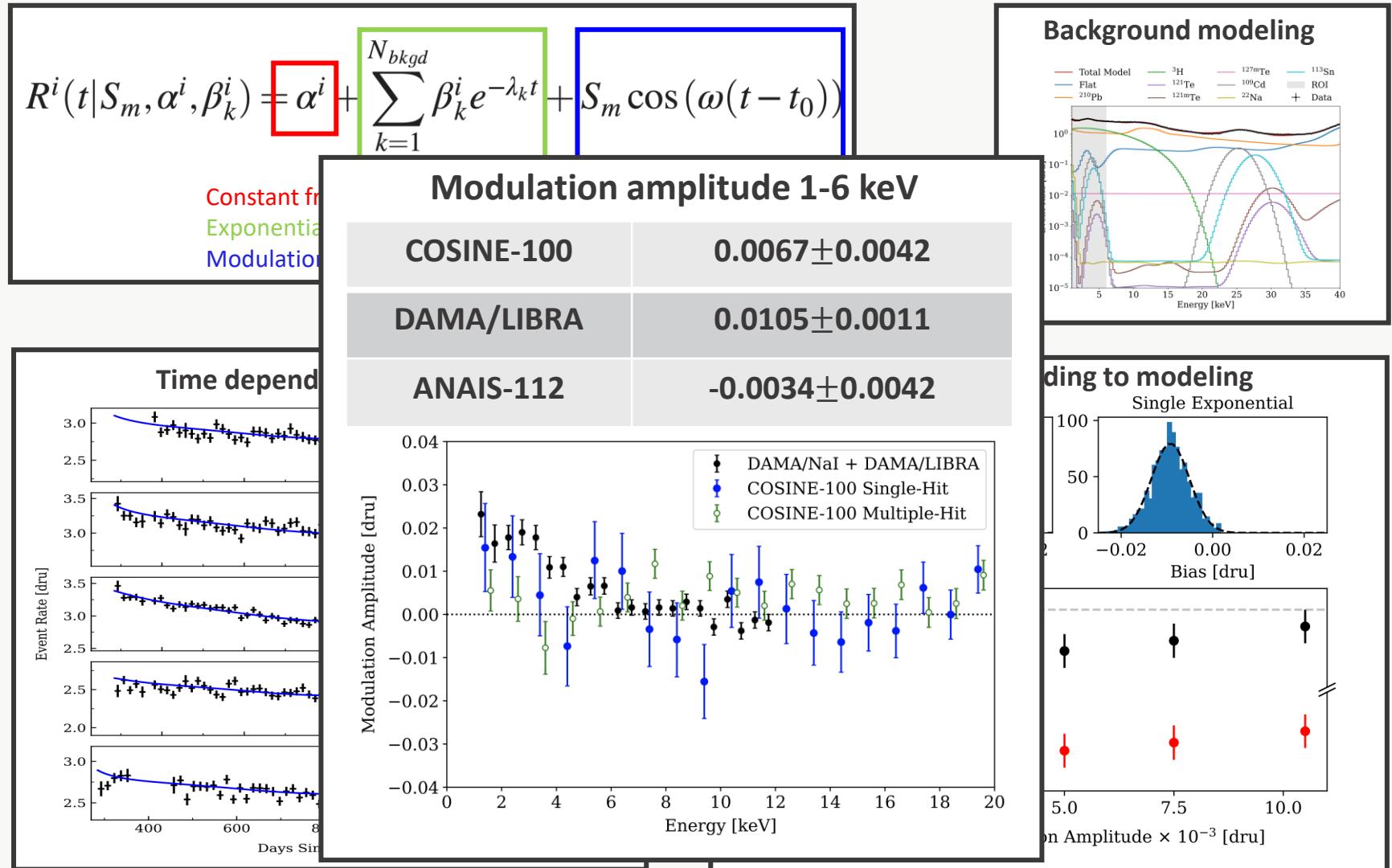
Phys. Rev. D 106, 052005 (2022)



Annual modulation analysis (3 years data)

- Full model time dependent rate fitting

Phys. Rev. D 106, 052005 (2022)



Annual modulation w/ DAMA/LIBRA method

arXiv:2208.05158

- Test DAMA/LIBRA method in COSINE-100 data

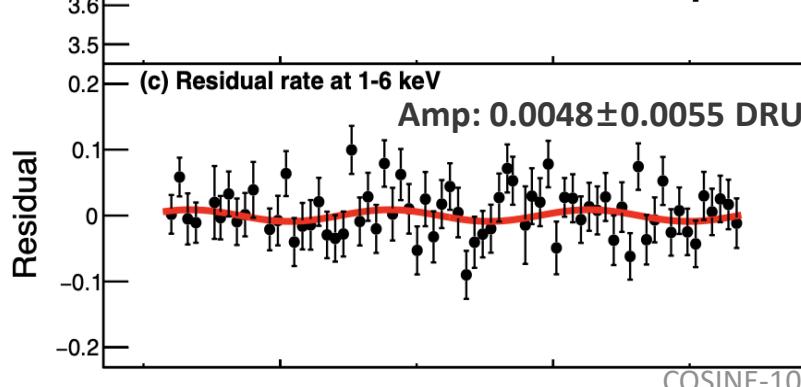
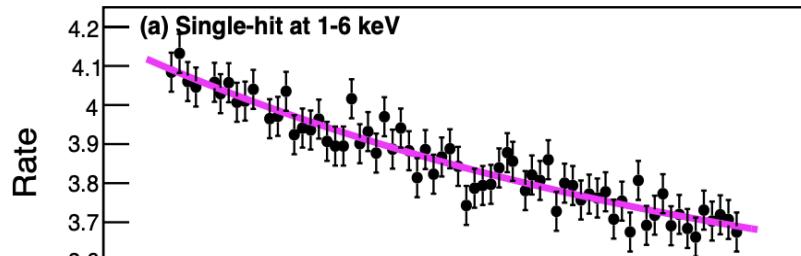
- Apply similar DAMA/LIBRA event selection

- No LS Veto
 - No Muon Veto
 - DAMA parameter for event selection
 - 600 ns waveform integration window.

- Test DAMA/LIBRA background modeling

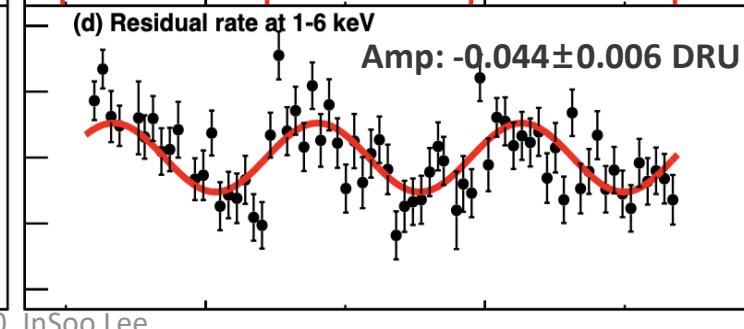
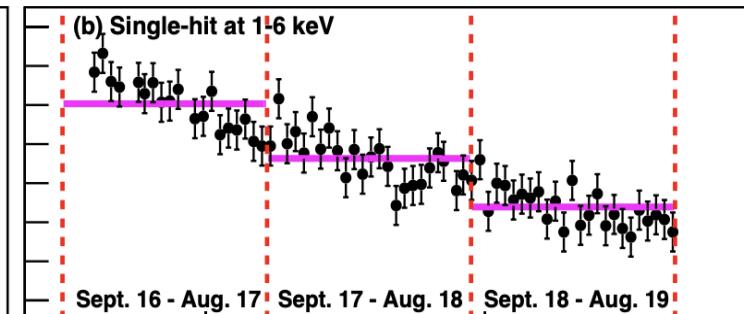
COSINE : $C + p_0 e^{-\frac{\ln 2 t}{p_1}} + A \cos \frac{2\pi(t-t_0)}{T}$

Single exponential



DAMA/LIBRA : $C + A \cos \frac{2\pi(t-t_0)}{T}$

Yearly average



Annual modulation w/ DAMA/LIBRA method

arXiv:2208.05158

- Test DAMA/LIBRA method in COSINE-100 data

- Apply DAMA/LIBRA event selection

- No LS Veto
 - No Muon Veto
 - DAMA parameter for event selection
 - 600 ns waveform integration window.

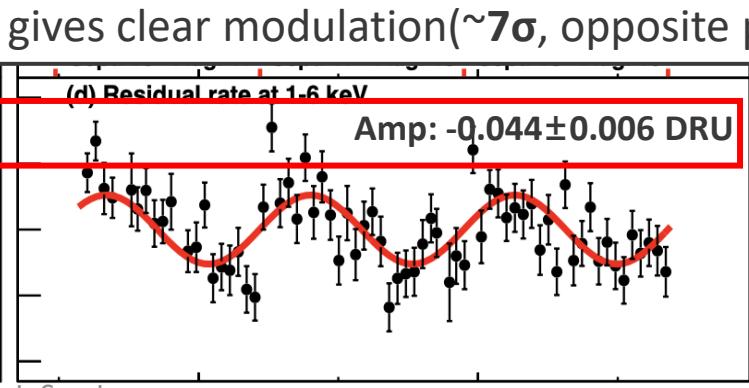
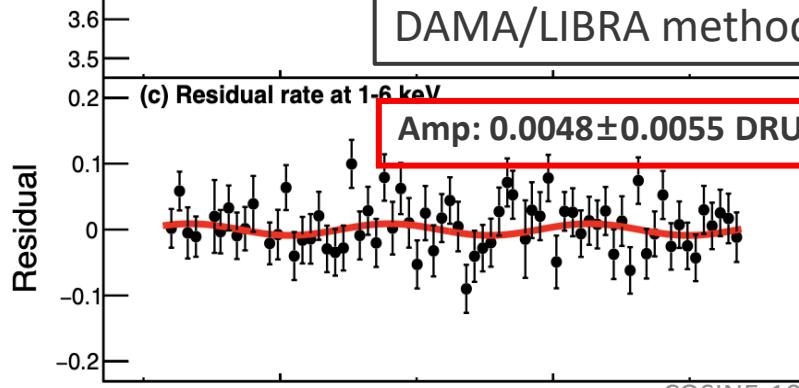
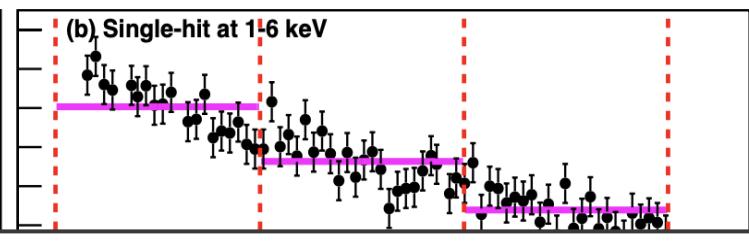
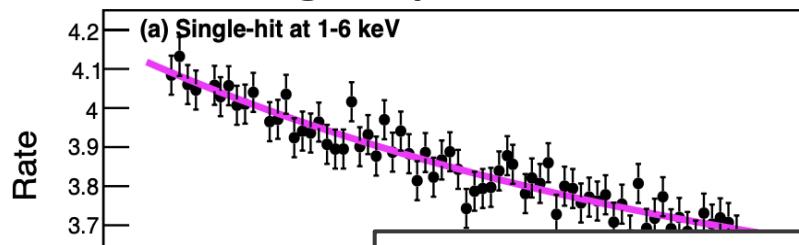
- Test DAMA/LIBRA background modeling

COSINE : $C + p_0 e^{-\frac{\ln 2 t}{p_1}} + A \cos \frac{2\pi(t-t_0)}{T}$

Single exponential

DAMA/LIBRA : $C + A \cos \frac{2\pi(t-t_0)}{T}$

Yearly average



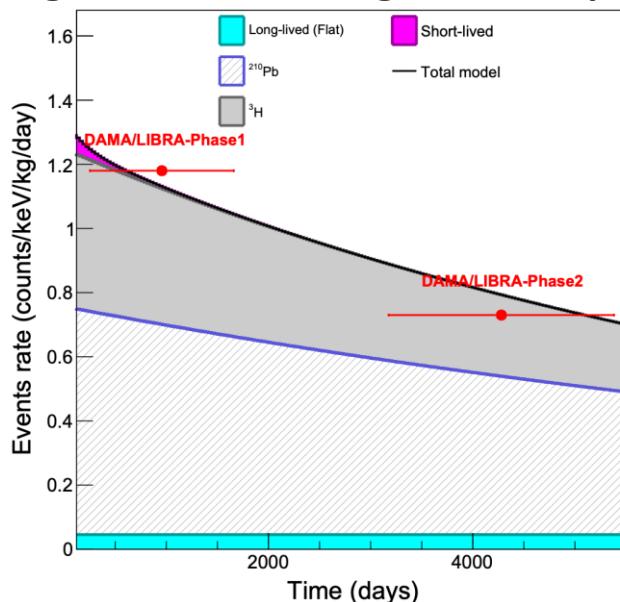
DAMA/LIBRA method gives clear modulation ($\sim 7\sigma$, opposite phase)

Annual modulation w/ DAMA/LIBRA method

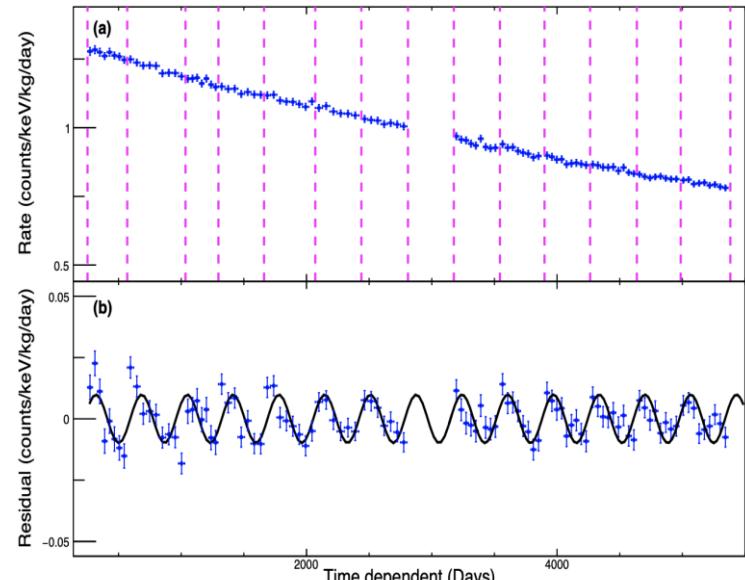
arXiv:2208.05158

- Background model for DAMA/LIBRA using the COSINE-100 background compositions

Composing DAMA/LIBRA Background,
using COSINE-100 background component



13 years pseudo-data

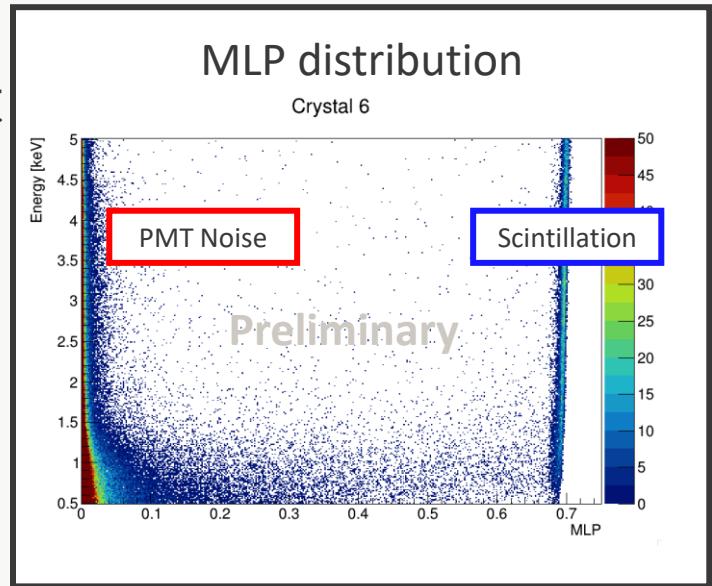


- Similar modulation amplitude in 1000 pseudo-experiment (Opposite phase)

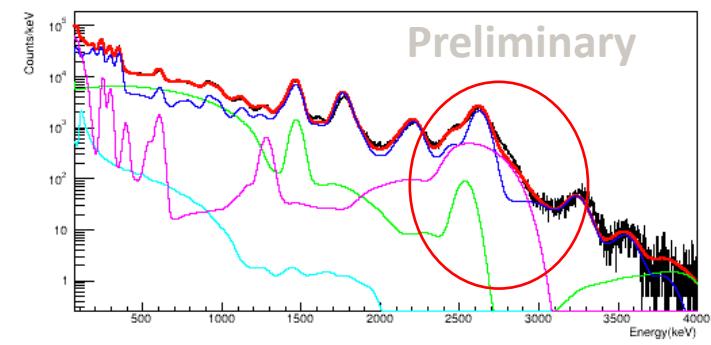
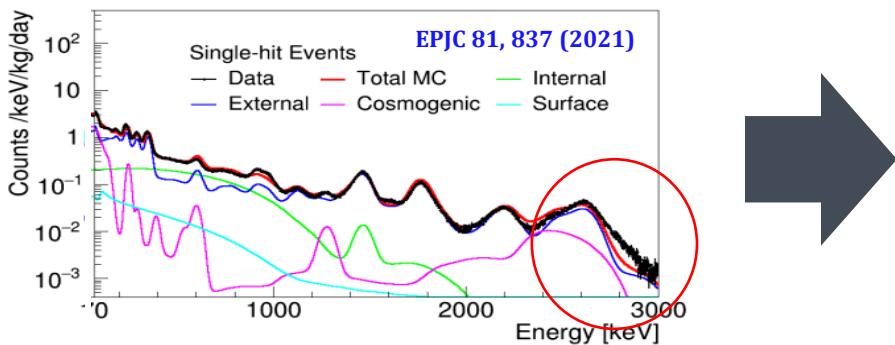
Pseudo data	-0.0098 ± 0.0008
DAMA/LIBRA	0.0105 ± 0.0011

On going works

- Event selection improvement
 - Aim to 0.5 keV threshold
 - Deep learning (MLP) method
- Background modeling
 - Non-Proportionality update
 - Extended modeling range 6-3000keV → 6-4000 keV

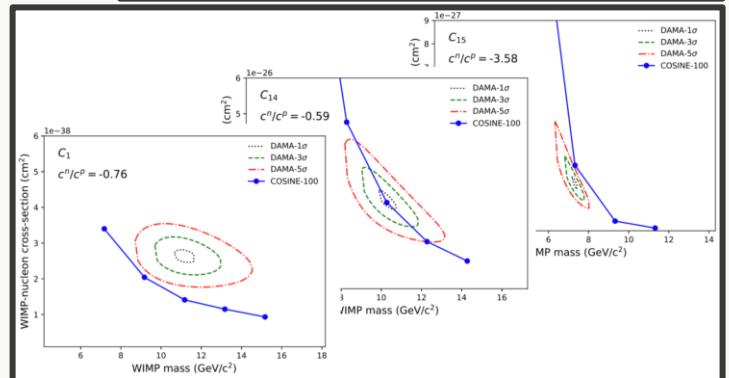
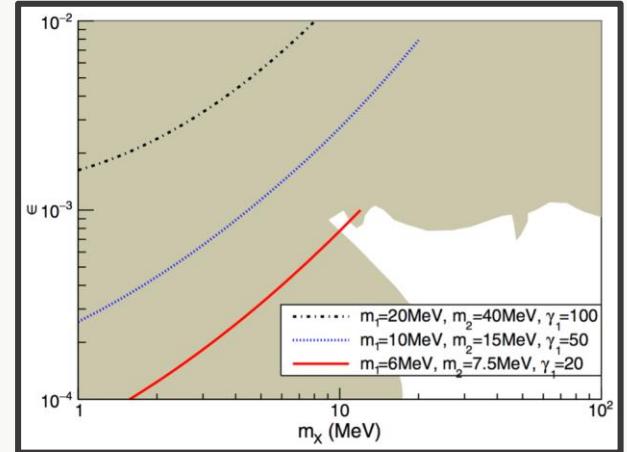
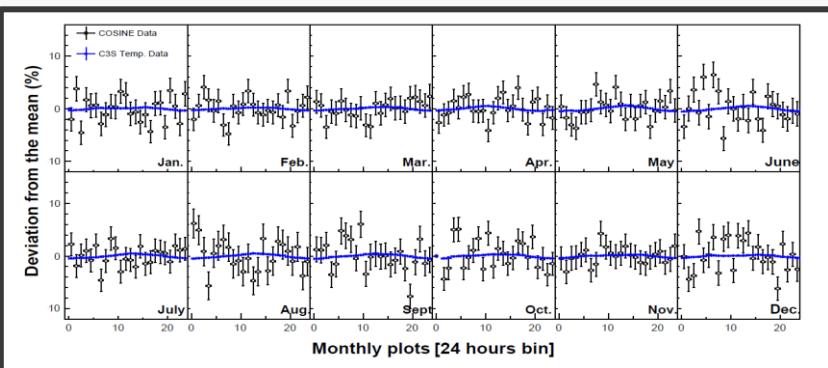
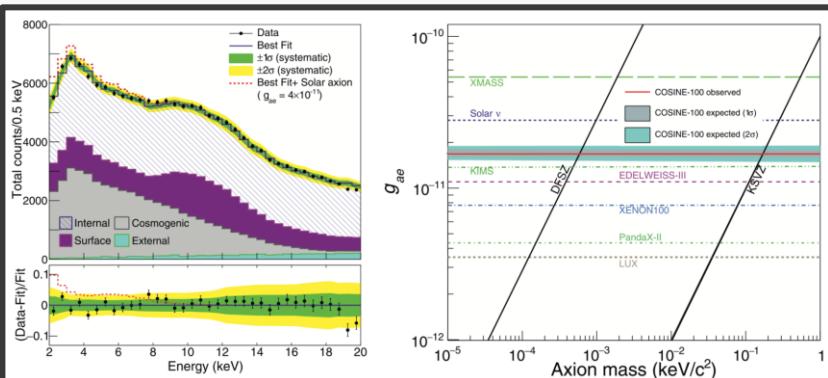


1.7-year background modeling



Exotic DM interaction

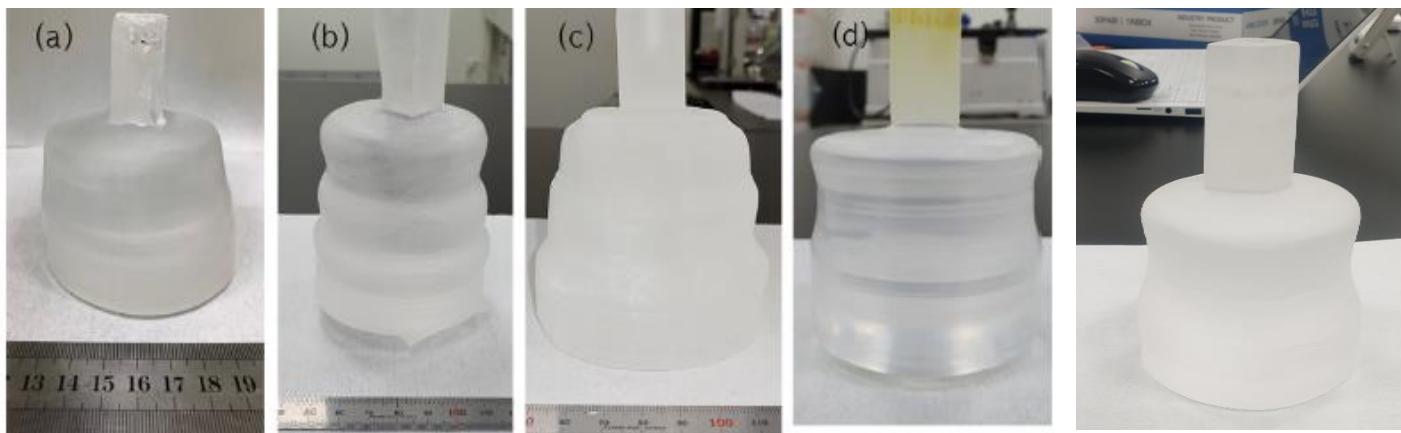
- Inelastic Boosted Dark Matter
- Solar Axions
- Several effective operators
- Inelastic WIMP- ^{127}I interaction
- Cosmic-ray Boosted Dark matter



JCAP 06, 048 (2019)
Astropart. Phys. 114 101 (2019)
Phys. Rev. Lett. 122 131802 (2019)

Effort for next phase

- Crystal development for **COSINE-200**
 - In-house development of **entire process** (at IBS , Korea)
 - NaI powder purification
 - Crystal growing
 - Detector assembly
 - Prepare ~0.7 kg R&D crystals
 - Expect ~0.2 DRU background level for COSINE-200 crystal
 - On going development for full size crystal growing (~100 kg)
 - Prepare COSINE-200 crystal by end of 2024



Effort for next phase

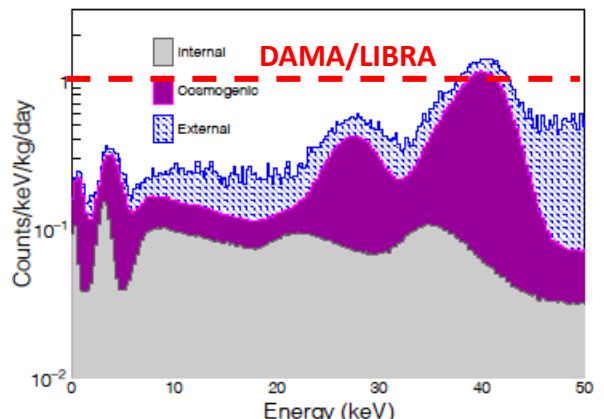
- Crystal development for **COSINE-200**
 - In-house development of **entire process** (at IBS , Korea)
 - NaI powder purification
 - Crystal growing
 - Detector assembly
 - Prepare ~0.7 kg R&D crystal
 - Expect ~0.2 DRU background level for COSINE-200 crystal
 - On going development for full size crystal growing (~100 kg)
 - Prepare COSINE-200 crystal by end of 2024



Background level in R&D crystals

Eur. Phys. J.C. 80 814 (2020)
Frontiers in Physics 11 (2023)

Expected background level
in COSINE-200



	K (ppb)	^{210}Pb (mBq/kg)	^{238}U (ppt)	^{232}Th (ppt)
COSINE	16.8 ± 2.5	1.87 ± 0.09	<0.02	0.7 ± 0.2
Powder	5	-	<20	<20
Aug.2018	<53	0.01 ± 0.02	0.9 ± 0.3	1.7 ± 0.5
Sep.2019	<42	0.42 ± 0.27	36.5 ± 3.9	<4.9
Feb.2021	8.3 ± 4.6	0.38 ± 0.10	<2.0	<0.8
DAMA	<20	$0.01 \sim 0.03$	0.7-10	0.5-7.5

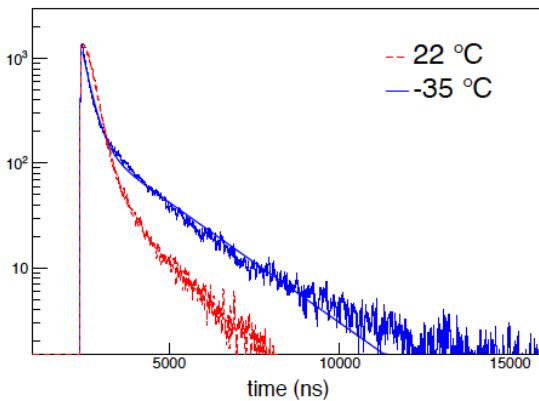
Effort for next phase

JINST 17 P02027 (2022)

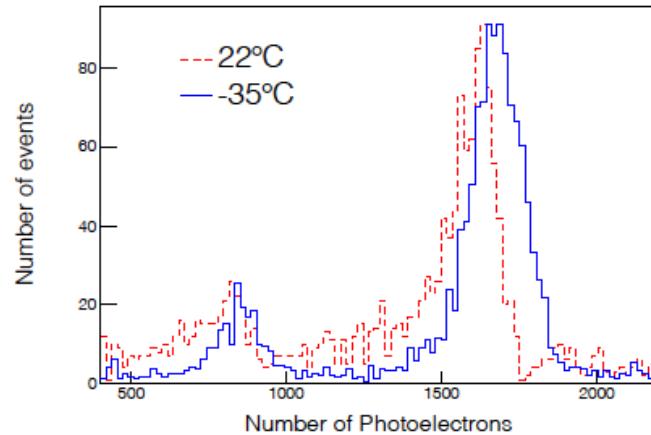
- Low temperature operation (at -35°C)
 - ~5 % increased Light yields
 - ~9% increased alpha quenching

Accumulated waveform

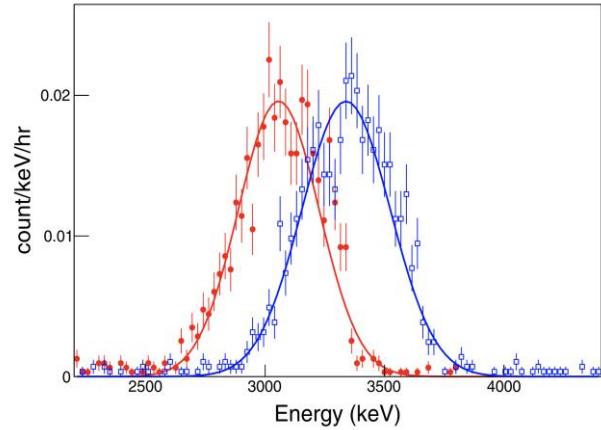
Accumulated waveform of ^{241}Am Events



NPE Measurement

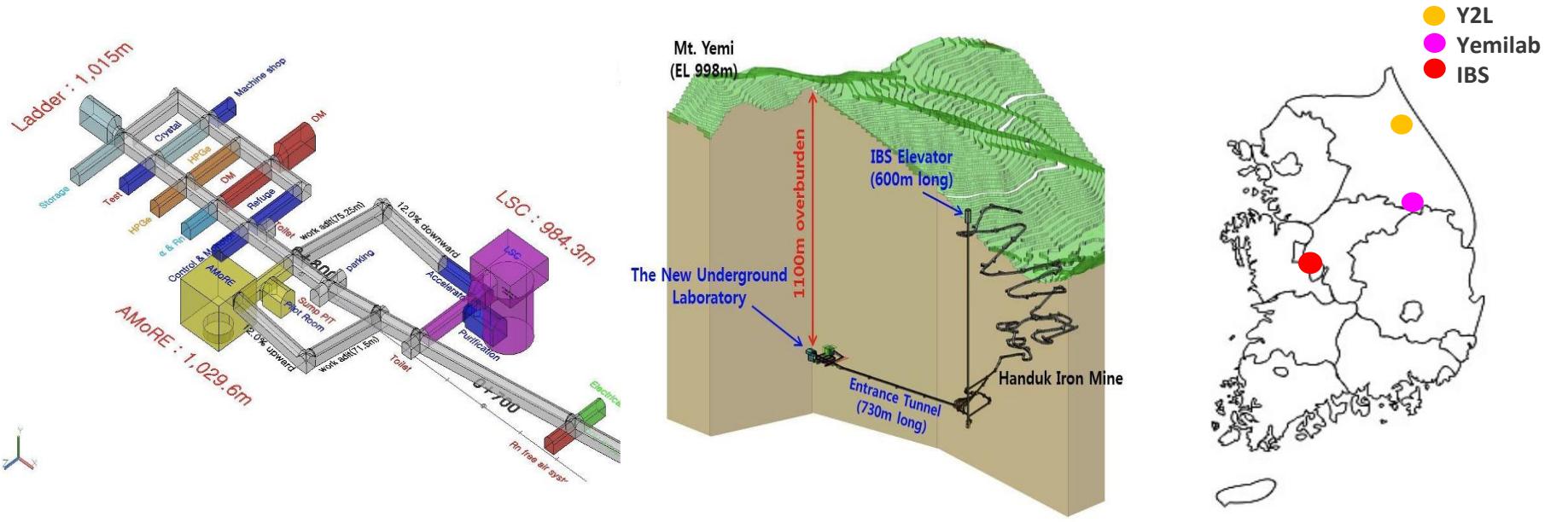


Alpha response using ^{210}Po

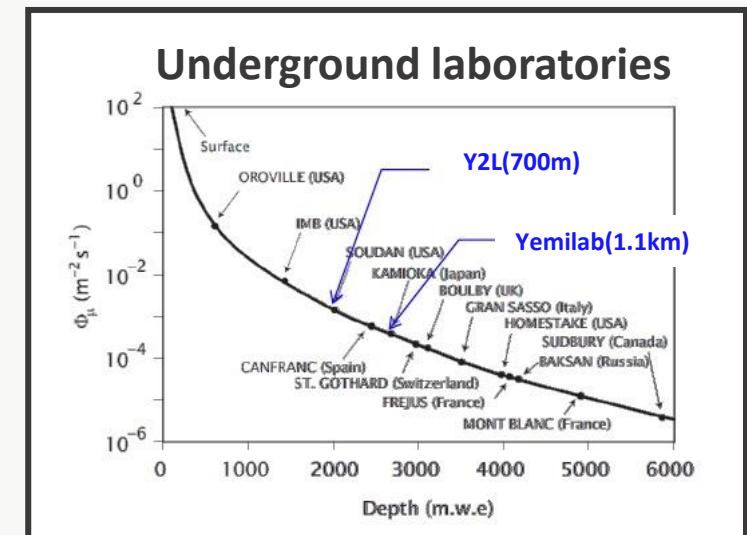


- COSINE-200 will operate at -35°C

Yemilab @ Jeonseon

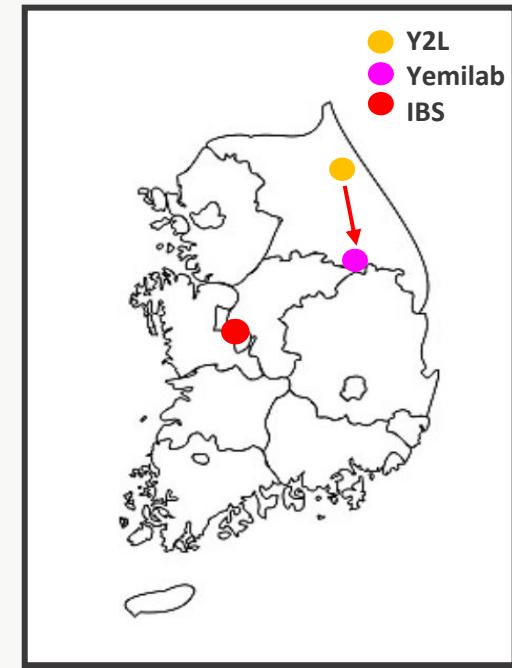


- Newly constructed underground laboratory at **Jeongseon** in 2022
 - **1.1 km** rock overburden
- Move to Yemilab in 2023



Upgrade COSINE-100

- Move COSINE-100 to Yemilab in 2023
- As an **intermediate experiment** of the **COSINE-200 @ Yemilab**
 - Use **COSINE-100** crystal
 - Use new encapsulation method
 - Improve light yields >20 P.E./keV
 - Operate at -35°C environment
- Operate **upgrade COSINE-100** until COSINE-200 crystal preparation



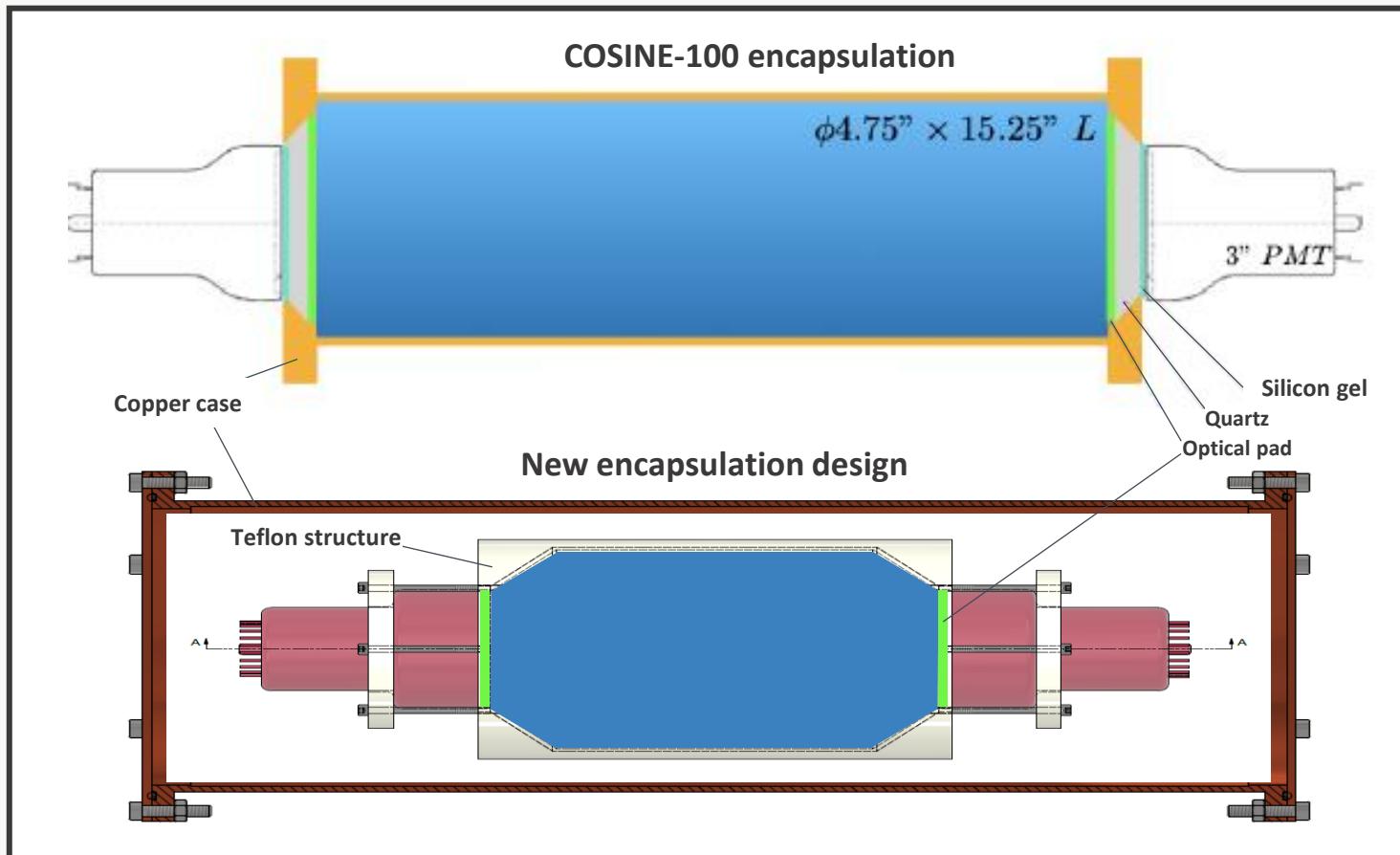
COSINE detector fridge room @Yemilab



Effort for next phase

- New encapsulation technique
 - Direct attachment of PMT to crystal
 - ~50% increased light Yield
 - 15 P.E. → 22 P.E.

Nucl. Instrum. Meth. A 981, 164556 (2020)

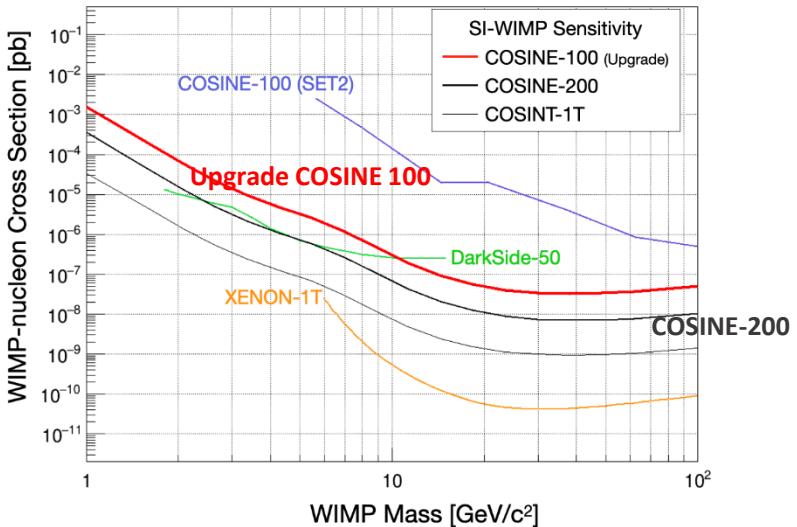


Sensitivity estimation

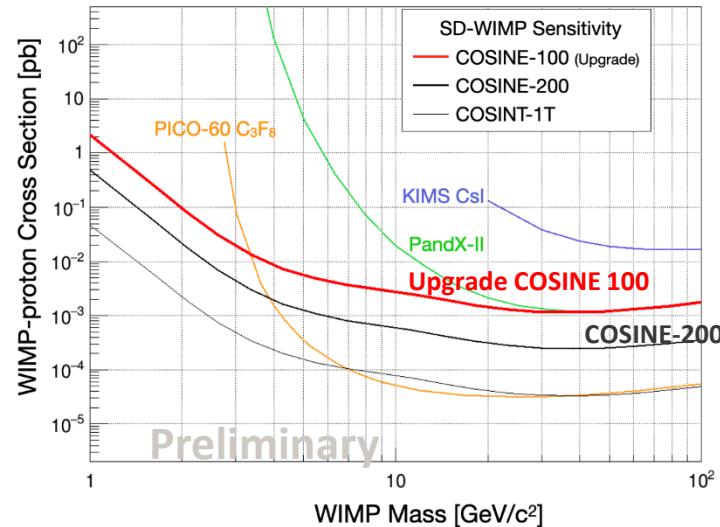
- Condition for upgrade COSINE-100 (COSINE-200)

- 1 year
- Light yields 22 P.E./keV
- COSINE-100 (R&D crystal) background
- No systematics

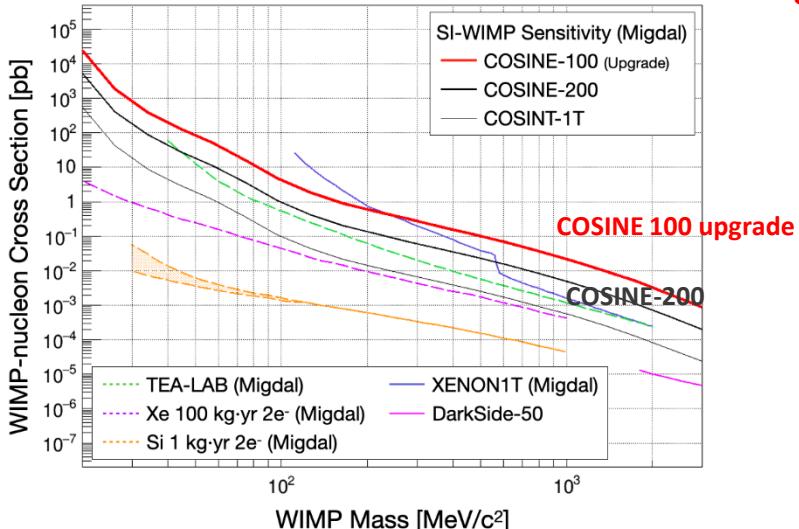
SI WIMP interaction



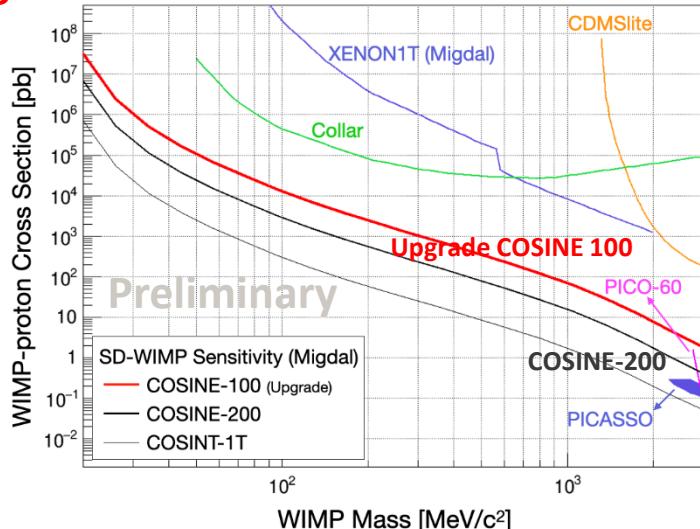
SD WIMP interaction



With Migdal



No systematics



Schedule for COSINE-100&200

	2023				2024				2025		
DAQ run		DAQ stop			Upgrade	COSINE-100				COSINE-200	
Encapsulation upgrade											
Y2L→Yemilab											
COSINE-200 development					COISINE-200 R&D						

- Stop **COSINE-100** DAQ in Mar. 2023
- Start DAQ run after **6 months** encapsulation upgrade at Yemilab (~Oct. 2023)
- Movement **Y2L→Yemilab** in Aug. 2023
- Operate **upgrade COSINE-100** until COSINE-200 crystal preparation
- Start **COSINE-200** in middle of 2025

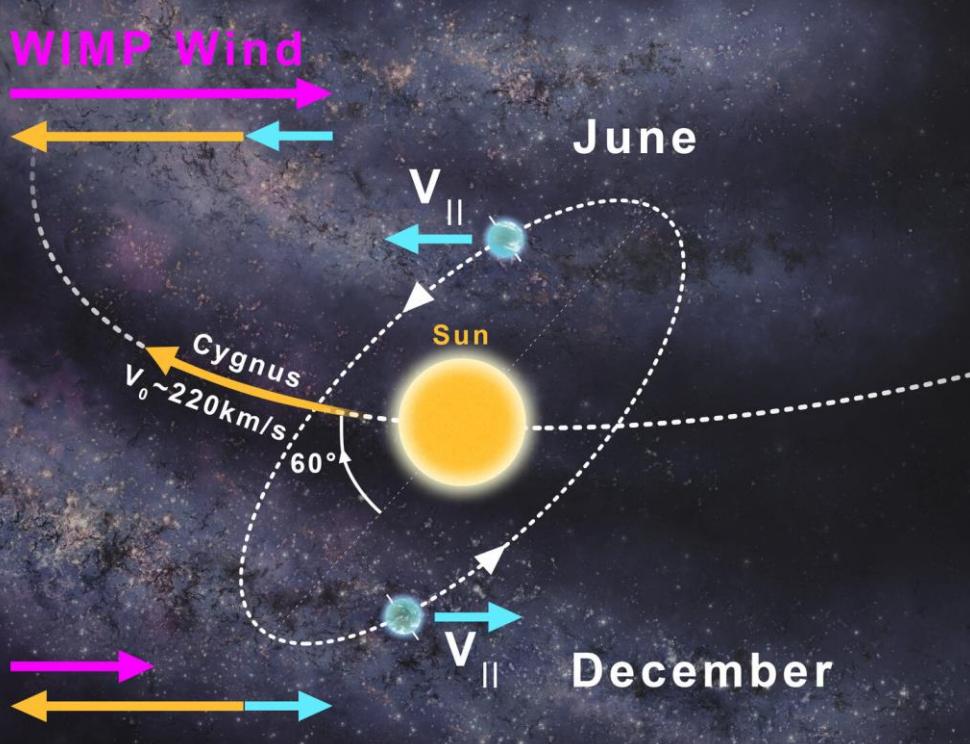
Summary

- COSINE-100 is an experiment to reproduce the DAMA/LIBRA signal via the same target material, NaI(Tl).
- Spectral Analysis
 - **Exclude** DAMA/LIBRA **signal** using 1.7 data
 - **Low mass WIMP** region was checked considering **Migdal Effect**.
- Modulation analysis
 - Could **not make clear decision** about DAMA/LIBRA modulation signal using **3 years data**
 - **Negative amplitude** using DAMA/LIBRA modeling method
- Preparations are ongoing for next phase, **COSINE-200**.
 - In-house development of **entire crystal growing process** at IBS , Korea
 - R&D NaI(Tl) crystals with **~0.2 count/day/kg/keV** of background level.
 - **50% light yields** improvement with **new encapsulation technique**.
 - Tested the properties of NaI(Tl) **at -35°C**
- We are planning an **intermediate experiment** of the **COSINE-200**.
 - Upgrade COSINE-100 **encapsulation** for **higher light yield**
 - Start upgrade COSINE-100 **DAQ** after 6-month **upgrade period (~Oct.2023)**

Back up



Annual modulation signal



- Spherical halo of dark matter in our galaxy is expected to make dark matter flux to Solar system (WIMP Wind)
- Speed of WIMP wind changes according to seasonal rotation of the Earth.
- Changing WIMP speed induces changing interaction rate.

Rate of WIMP elastic scattering

$$R \propto N_T \cdot \sigma_{\chi^N} \cdot \frac{\rho_\chi}{m_\chi} \int_{v_{min}} \frac{f(v)}{v} dv$$

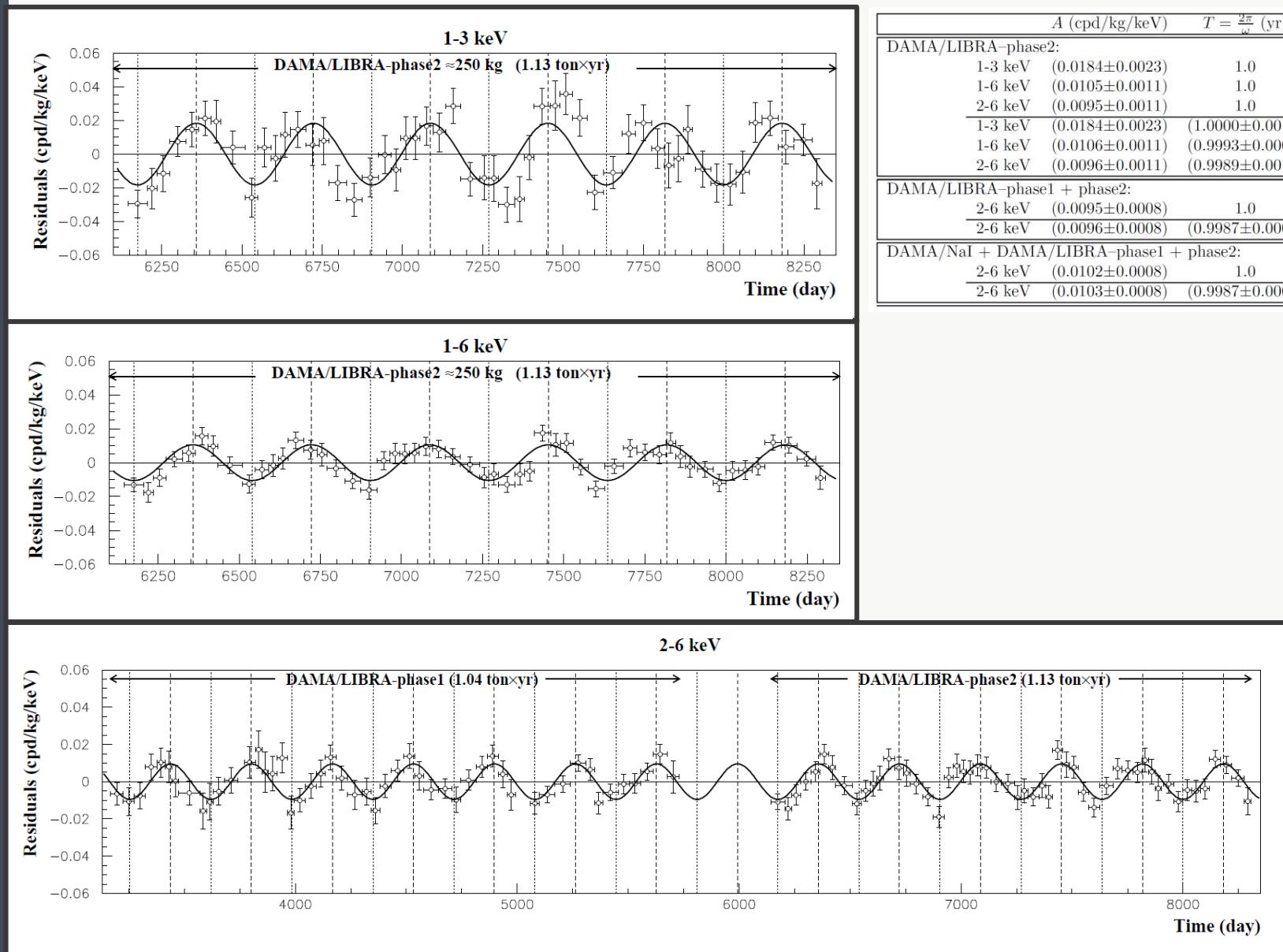
$f(v)$: WIMP velocity distribution

v_{min} : minimum velocity for interaction

ρ_χ : local WIMP density

m_χ : WIMP mass

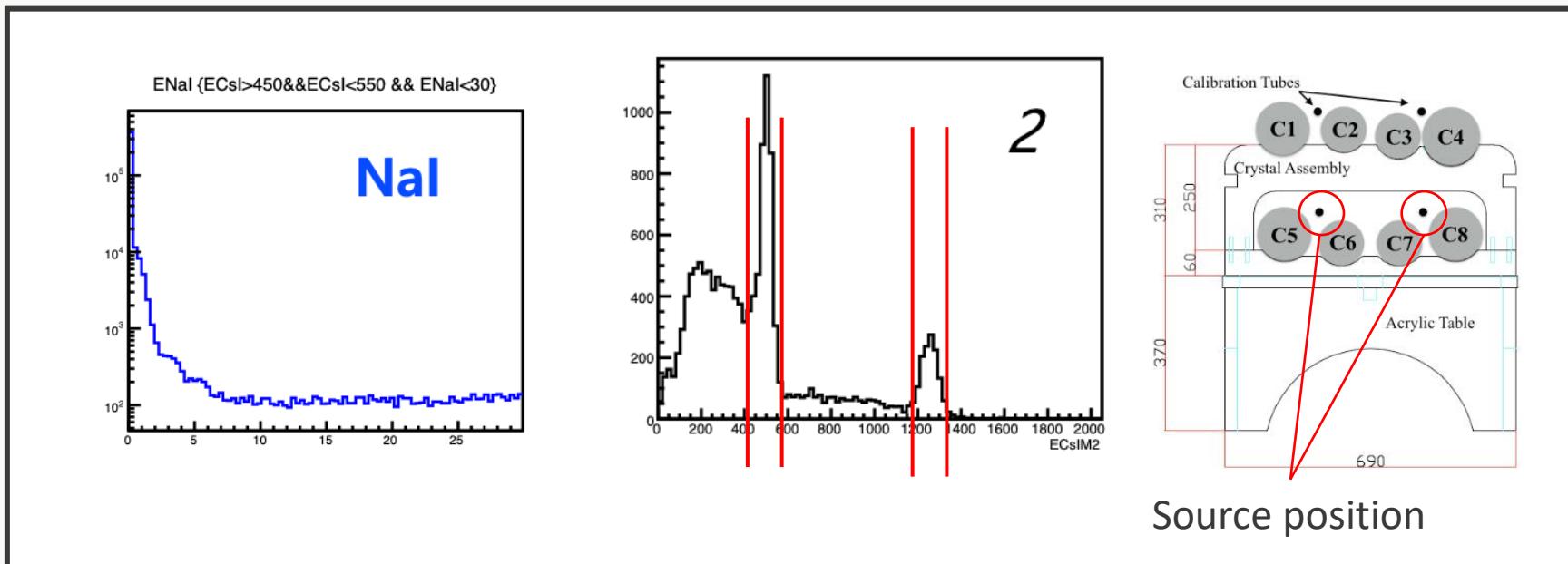
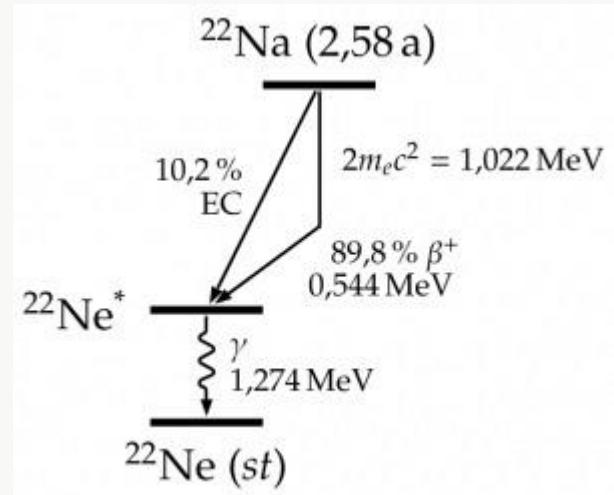
DAMA/LIBRA result



	A (cpd/kg/keV)	$T = \frac{2\pi}{\omega}$ (yr)	t_0 (days)	C.L.
DAMA/LIBRA-phase2:				
1-3 keV	(0.0184±0.0023)	1.0	152.5	8.0 σ
1-6 keV	(0.0105±0.0011)	1.0	152.5	9.5 σ
2-6 keV	(0.0095±0.0011)	1.0	152.5	8.6 σ
1-3 keV	(0.0184±0.0023)	(1.0000±0.0010)	153±7	8.0 σ
1-6 keV	(0.0106±0.0011)	(0.9993±0.0008)	148±6	9.6 σ
2-6 keV	(0.0096±0.0011)	(0.9989±0.0010)	145±7	8.7 σ
DAMA/LIBRA-phase1 + phase2:				
2-6 keV	(0.0095±0.0008)	1.0	152.5	11.9 σ
2-6 keV	(0.0096±0.0008)	(0.9987±0.0008)	145±5	12.0 σ
DAMA/NaI + DAMA/LIBRA-phase1 + phase2:				
2-6 keV	(0.0102±0.0008)	1.0	152.5	12.8 σ
2-6 keV	(0.0103±0.0008)	(0.9987±0.0008)	145±5	12.9 σ

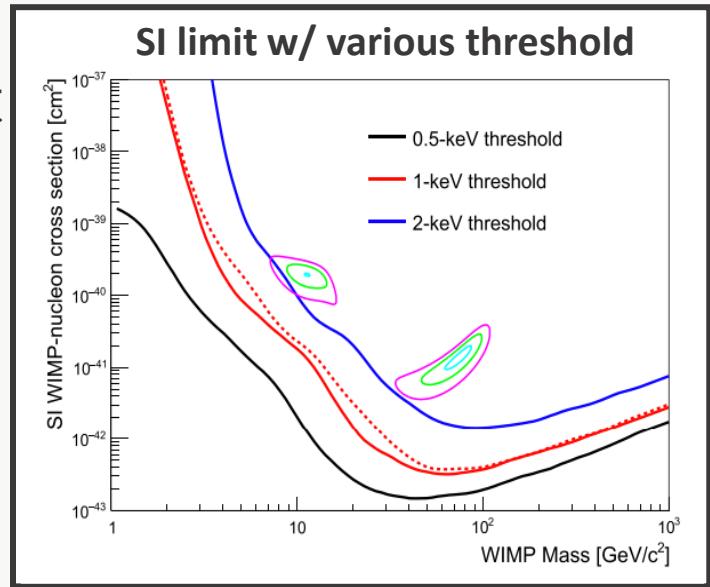
Why ^{22}Na calibration?

- To get pure low energy gamma spectrum,
 - Tagging coincidence event with other crystals' 511 keV or 1274 keV event

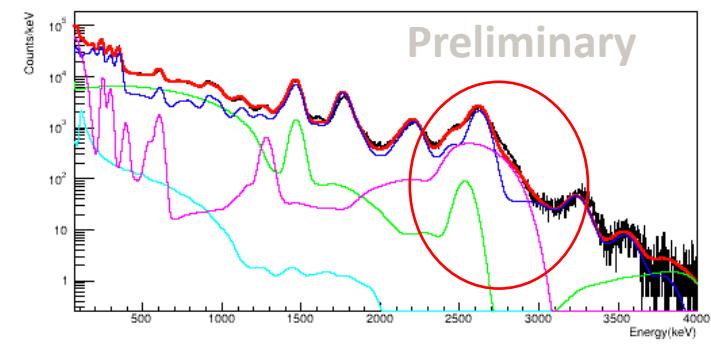
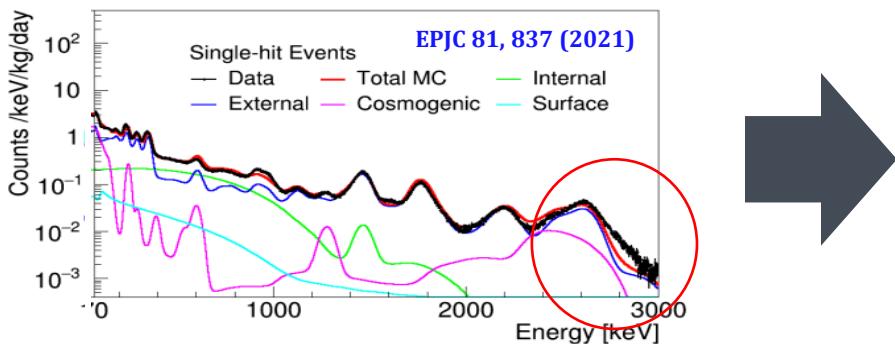


On going works

- Event selection improvement
 - Aim to 0.5 keV threshold
 - Deep learning (MLP) method
- Background modeling
 - Non-Proportionality update
 - Extended modeling range 6-3000keV → 6-4000 keV

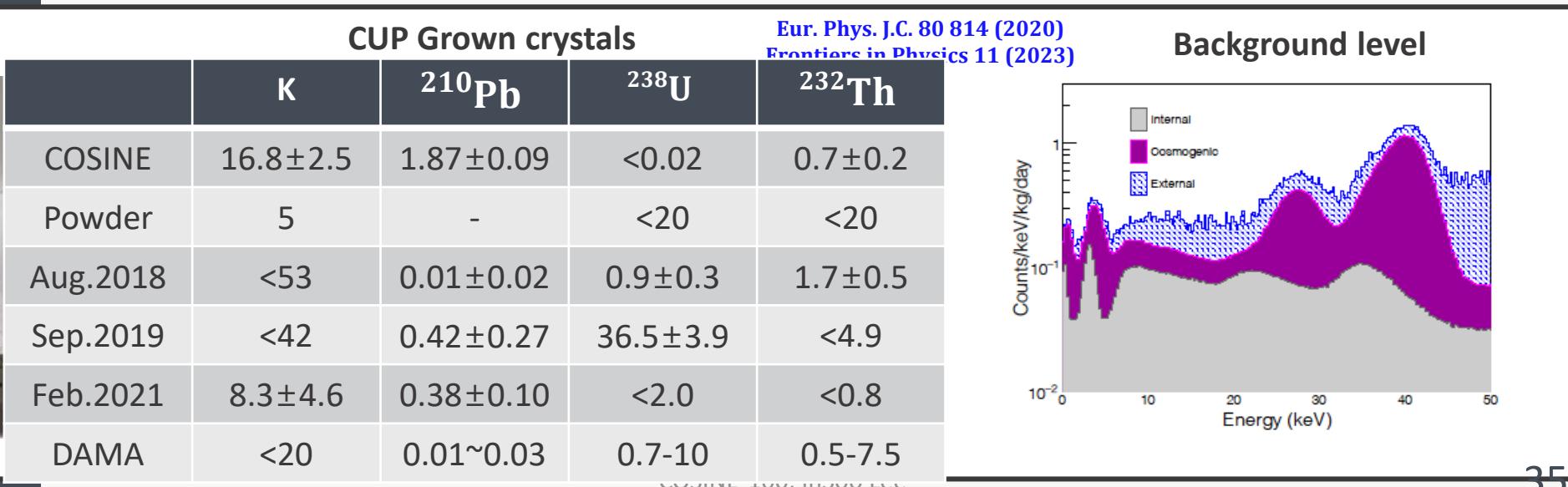


1.7-year background modeling



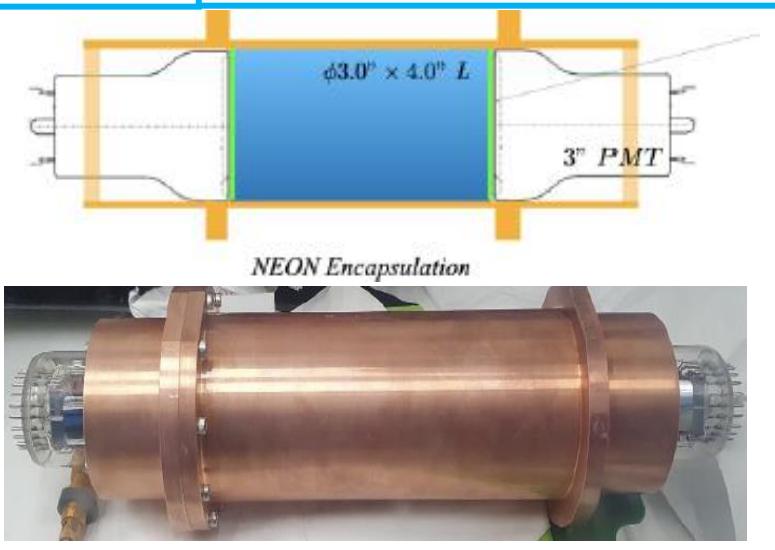
Effort for next phase

- Crystal development for COSINE-200
 - In-house development of entire process (at IBS , Korea)
 - NaI powder purification
 - Crystal growing
 - Detector assembly
 - $\sim 0.7 \text{ kg}$ crystal w/ $\sim 0.2 \text{ count/day/kg/keV}$
 - On going development for full size crystal ($\sim 100 \text{ kg}$)

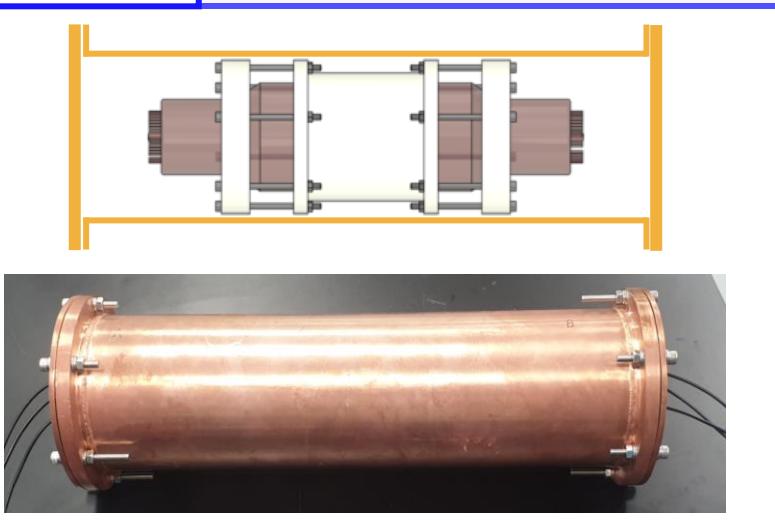


NEON-Phase 2 upgrade

Phase 1



Phase 2

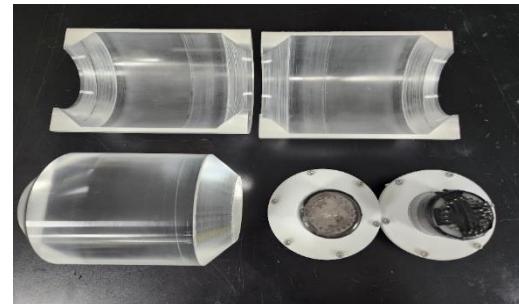
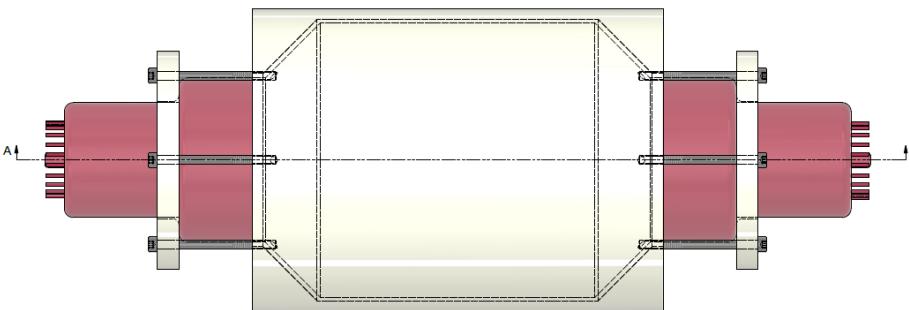


- Upgrade encapsulation (**Apr. 2022**)
 - Fix mechanical instability (Air or LS leak)
 - Achieve **higher light yields**
 - Add more mass : **13.6 → 16.3 kg**
- Exclude two high background crystals
- Include two **8" length crystals**

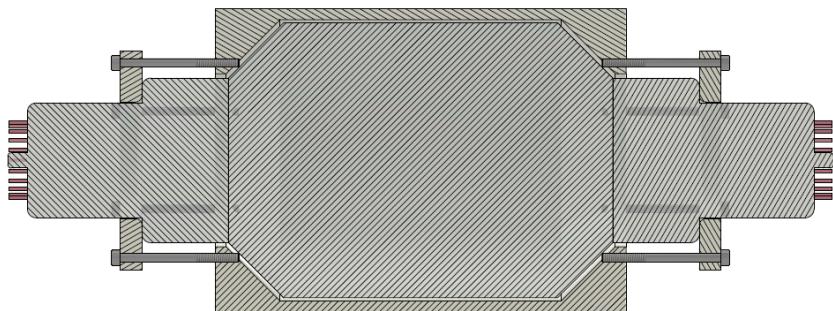
Crystal #	Size (dia x length)	Light yields(PE/keV)	
		Phase 1	Phase 2
1	3" x 4"	20.5 ± 0.9	Excluded
2	3" x 4"	19.3 ± 0.9	Excluded
3	3" x 4"	21.8 ± 0.9	23.9 ± 0.8
4	3" x 8"	22.4 ± 1.0	24.5 ± 0.7
5	3" x 8"	21.8 ± 0.9	22.9 ± 0.4
6	3" x 4"	21.7 ± 1.0	20.2 ± 0.4
7	3" x 8"	-	22.0 ± 0.5
8	3" x 8"	-	26.7 ± 0.8

Encapsulation

- Inner structure for PMT-crystal connection
 - Direct contact PMT and crystal with optical pad
 - PTFE body, Brass bolt
 - PMT base shield (PTFE)

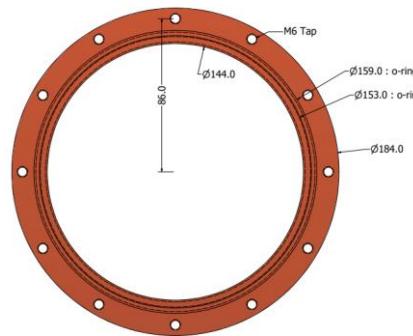
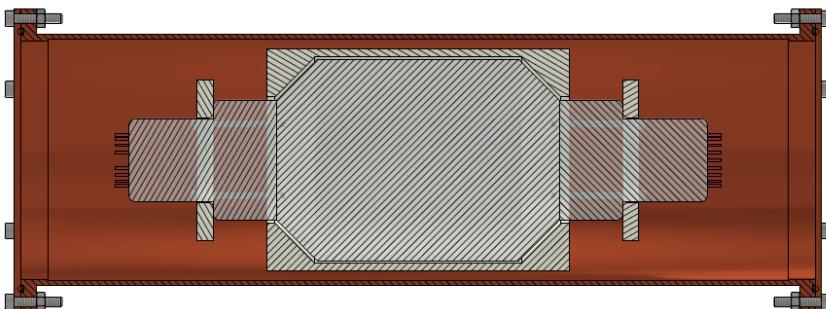
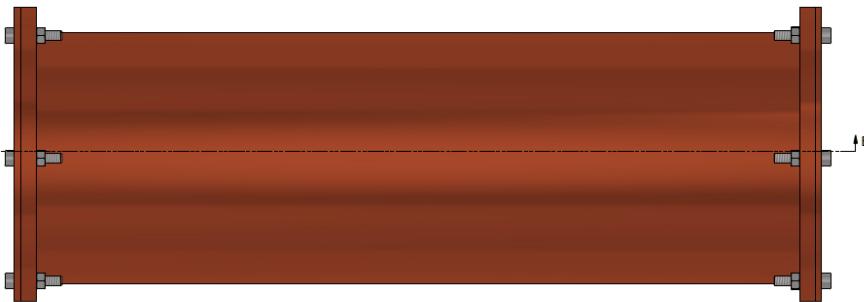


PMT base shield

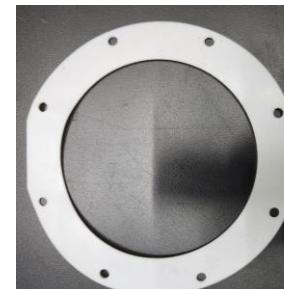


Encapsulation cont'd

- Copper case
 - Same as NEON experiment
 - To prevent LS & air leak
 - PTFE Gasket
 - Cable gland



PTFE gasket



cable gland



NaI(Tl) Crystal

- 8 ultra low-background NaI(Tl) crystal with 106 kg in total
- U/Th/K level is less than DAMA, but total alphas ($\sim 210\text{Pb}$) are higher than DAMA.
- Higher light yield (15 p.e./keV) than DAMA
 - Easier to reach lower threshold
- Total background is 2-3 times that of DAMA

Crystal	Mass (kg)	Powder	Alpha rate (mBq/kg)	^{40}K (ppb)	^{238}U (ppt)	^{232}Th (ppt)	Light yield (p.e./keV)
Crystal 1	8.3	AS-B	3.20 ± 0.08	43.4 ± 13.7	< 0.02	1.31 ± 0.35	14.88 ± 1.49
Crystal 2	9.2	AS-C	2.06 ± 0.06	82.7 ± 12.7	< 0.12	< 0.63	14.61 ± 1.45
Crystal 3	9.2	AS-WS II	0.76 ± 0.02	41.1 ± 6.8	< 0.04	0.44 ± 0.19	15.50 ± 1.64
Crystal 4	18.0	AS-WS II	0.74 ± 0.02	39.5 ± 8.3		< 0.3	14.86 ± 1.50
Crystal 5	18.0	AS-C	2.06 ± 0.05	86.8 ± 10.8		2.35 ± 0.31	7.33 ± 0.70
Crystal 6	12.5	AS-WS III	1.52 ± 0.04	12.2 ± 4.5	< 0.018	0.56 ± 0.19	14.56 ± 1.45
Crystal 7	12.5	AS-WS III	1.54 ± 0.04	18.8 ± 5.3		< 0.6	13.97 ± 1.41
Crystal 8	18.3	AS-C	2.05 ± 0.05	56.15 ± 8.1		< 1.4	3.50 ± 0.33
DAMA			< 0.5	< 20	0.7 - 10	0.5 - 7.5	5.5 - 7.5

Eur. Phys. J. C. 78 107 (2018)

Upgrade plan for COSINE-100

Review

- COSINE-100 @ Yemilab
 - Crystals used in COSINE-100
 - Recover un-used crystals (# 1, 5, 8)
 - With **new** technique of **encapsulation** (Light yields >20 PEs/keV)
 - At -35°C environment

Nucl. Instrum. Meth. A 981 (2020) 164556

JINST 17 P02027 (2022)

COSINE-100 crystals

Crystal #	Size (diameter x length)	Light yield(PEs/keV)
1	5.0" x 7.0"	14.9 ±1.5
2	4.2" x 11.0"	14.6 ±1.5
3	4.2" x 11.0"	15.5 ±1.5
4	5.0" x 15.3"	14.9 ±1.5
5	5.0" x 15.5"	7.3 ±0.7
6	4.8" x 11.8"	14.6 ±1.5
7	4.8" x 11.8"	14.0 ±1.4
8	5.0" x 15.5"	3.5 ±0.3

