WISPs Dark Matter and Cosmology pans and organisation

WWII

Javier Redondo COST21106 kick off meeting LNF 23-24/02/2023





Dep. Theoretical Physics Universidad de Zaragoza

COST : big picture

Bring together European WISP researchers to :





COST : big picture

In detail: 4 years of funding (~500,000 total), for :















- WISP (axion) dark matter and cosmology

- Working Group ``leaders" (we'd prefer coordinators)



Nicholas Rodd CERN Dark matter indirect detection. EFTs, axions, statistics, and collider physics. https://nickrodd.com



Javier Redondo Universidad de Zaragoza, Spain Max Planck for Physics, Munich BSM @ Low energy physics, Axions, Dark matter, other WISPs

WG 2: WISPs Dark Matter and Cosmology. This WG aims to study the multi-faceted and diverse cosmology of WISPs. Many WISPs are cold DM candidates whose production mechanisms may be non-thermal and depend on the dynamics of phase transitions or topological defects (as is the case of the QCD axion) [Sikivie (2008)]. This WG aims to identify and study the production mechanisms for WISPs DM, in order to reliably compute the DM relic density. In the case of the QCD axion, the relic density depends on the dynamics of two phase transitions, the second being the QCD Phase Transition, when the axion mass turns on. The WG would explore what lattice gauge theory can tell about this temperature-dependent axion mass [Borsanyi (2016)]. If the Peccei-Quinn symmetry is restored after inflation, the DM axion production from cosmic string decays has a dominant role [Gorghetto (2018)]. The WG will study the axion relic abundance from this mechanism and the production of miniclusters. It will study how WISP DM grows Large Scale Structures, and explore whether it generates distinct features that could allow to distinguish WISPs DM from WIMPs. The WG will also explore the cosmological signatures of WISPs as subleading hot DM. In this context, it has been recently pointed out [Di Luzio (2021)] that the current computations of the mass bounds on axion hot dark matter are not reliable since they are based on an extrapolation of the axion-pion interaction in a range where chiral perturbation theory breaks down. The WG will find a strategy to derive a reliable hot dark matter axion mass bound using lattice QCD techniques.

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WG 2: WISPs Dark Matter and Cosmology.

WISPs as dark matter candidates Production mechanisms Axion Thermal mass (lattice, ChiPT) Cosmic string decays Effects on Large and small-scale structure hot DM (lattice, ChiPT)

In practice

- Bring together scientists around WISPy Dark Matter & Cosmology
- Accelerate CURRENT research directions
- Create NEW research directions
- Identify bottlenecks (for CURRENT and NEW directions)
- widen the necks
- Identify key expertise, invite to the network

legacy, scientific outreach, accountability

- Produce reports on state of the art WISP dark matter and cosmology
 - Draft (year 1)
 - interim (year 2)
 - final (year 4)

The aim is NOT to produce another snowmass-type document We would like it to be new, exciting, new perspectives, new connections

Bold idea: perhaps connect it to the schools?



Snowmass 2021 White Paper Axion Dark Matter

J. Jaeckel¹, G. Rybka², L. Winslow³, and the Wave-like Dark Matter Community ⁴

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- Plenty of information available, executive summary of axion cosmology

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		4.1.2 The Post-Inflation Scenario	15
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	4.2	Late-Time Structure Formation and the Local Properties	16

Initial plans

- Establish a discussion forum
- Organize focussed meetings on specific topics
- Widen the community

We do not plan (yet) a structure for the reports, let it emerge from the discussions and interest of the community

Plans 1/3



Online Discussion Forum

👬 slack



will it work? let us try for a couple of months

Plans 2/3

- Focused workshops on very concrete topics:

- input from lattice QCD into axion (WISP) cosmology?

QCD potential away from ~ 0, axion thermal width, couplings,

- Important steps towards an accurate computation of m_a for post-inflation PQ

Numerical simulations, HPC, AMR, theoretical modeling, etc...

- What needs to be done to identifying WISP signals in LSS?

Birefringence, PS, Ly-alpha, strings, ...

- Community decides using the forum

Plans 3/3

- currently we are ~100 people in the WG2 mailing list
- Identify new plausible collaborators

- We may have ~exhausted the EURO axion ``field", but not the

- lattice QCD community
- Cosmology community
- Early Universe Phase transitions
- Computational cosmology/structure formation community
- Theoretical cosmology
- Solid-state?
- Scan by institution, intuitions, search for breakthroughs in fields

- Discuss most urgent/needed fields, recruit new scientists?

1st WG2 meeting

Nice attendance

~ D'Oro, Rodd, Diego, Pierobon, Lella, Terças, Lombardo, Sala, Sigl, Mirizzi, Schwetz, Tammaro, Lucente, Demir Pulice, Bernal, Masrh, Mevemsek, Cogollos, Kaltschmidt, O'Hare, Redondo, Gerbino, Lattanzi, Gorghetto, Cembranos

Very heterogenous community

CTA, warm DM, astro-signatures, CMB birefringence, axion DM simulations, Miniclusters, ALP pheno, SN, ALPs in plasmas lattice QCD, ALP star pheno, HE astro, dark matter, a-gamma mixing in B-fields, neutrino astro, neutrino cosmo, machine-learning WISPs in astro/cosmo, DM from Higgs stability, vector fields, WIMPs, phase transitions, haloscopes, small scale distribution of DM, CMB on ALPs, neutrinos, precision QCD, ultralight dark matter ...

- some students, young postdocs, but not nearly enough :-)

Some suggestions

- common calendar, compatibility of events

- common topics, coordination (miniclusters?)

Conclusions



- Widen the network
- Forum
- Topical workshops
- Doing relevant science
- Plant the seeds for even more relevant science