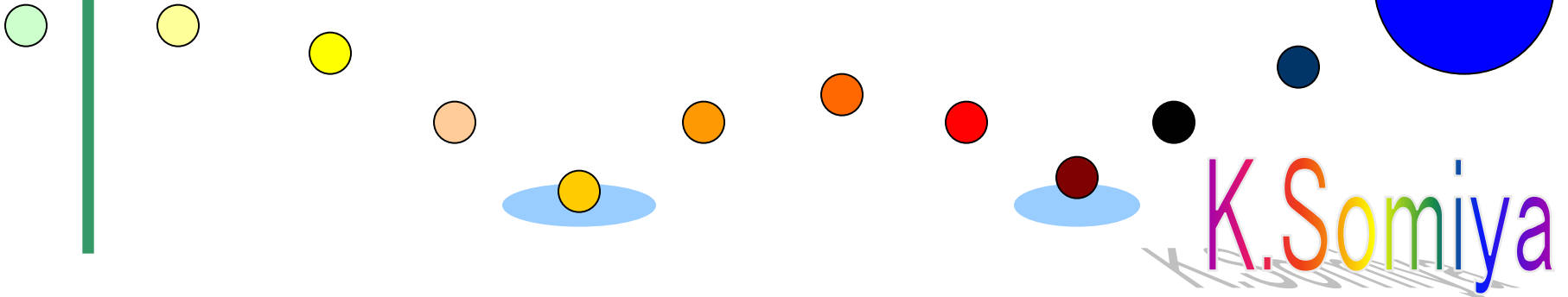


LCGT Detector Configuration

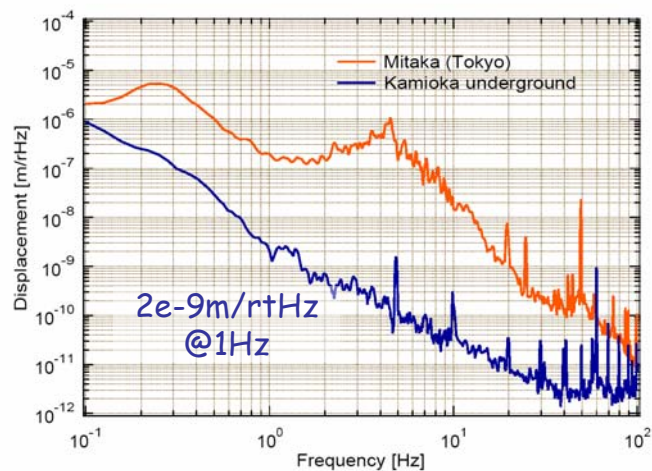
Elba
May 23, 2011

Tokyo Institute of Technology

Kentaro Somiya



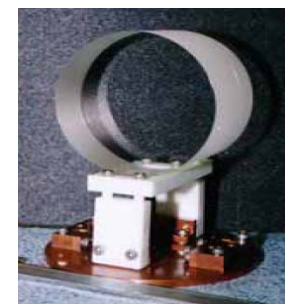
LCGT



Ground motion at Kamioka underground (blue)



SAS

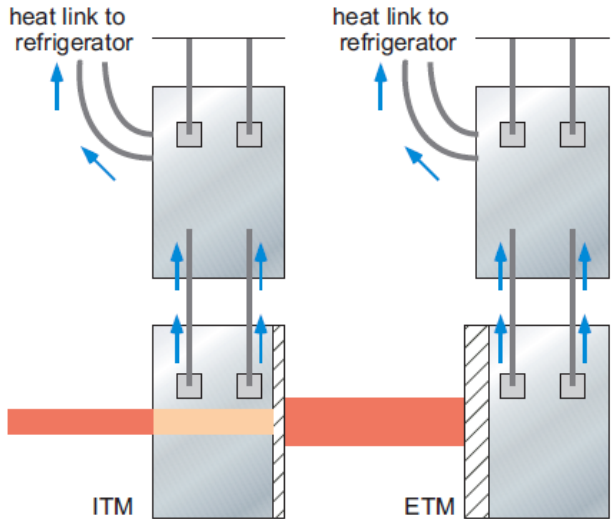


Sapphire mirror

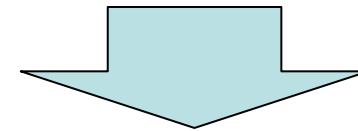
Funded in 2010.6 ☺

- 3km IFO in underground
- 20K Sapphire test masses
- Sub-SQL sensitivity
- Several GW events per year

Cryogenic system

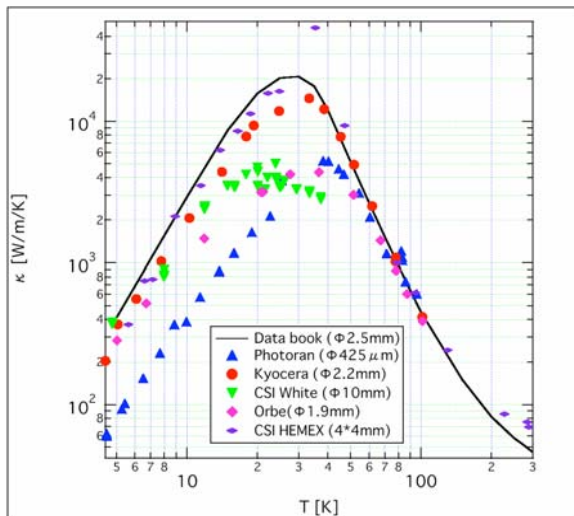


- (1) Sapphire is good with 1064nm laser, and has high thermal conductivity
- (2) 20K fiber can transfer $\sim 1W$ heat
- (3) Absorption of Sapphire substrate is not small



High-finesse RSE

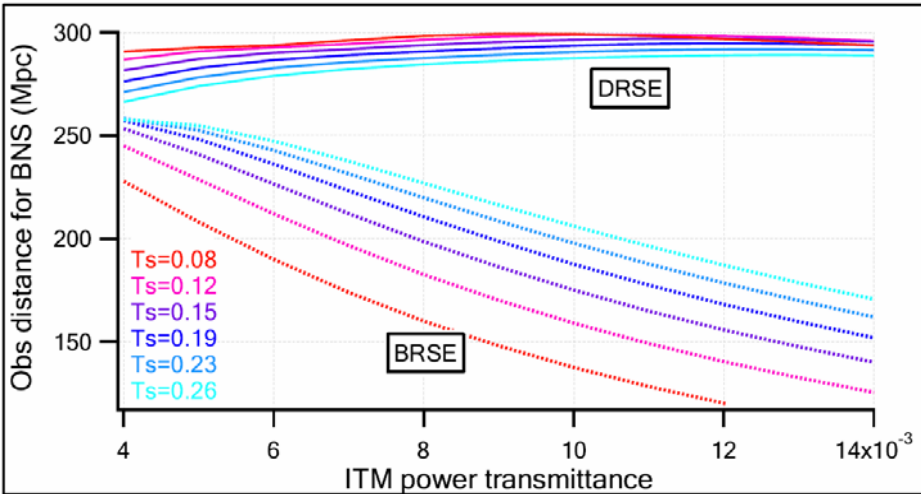
- decent power in PRC
- high power in the arm
- less absorption in ITM substrate



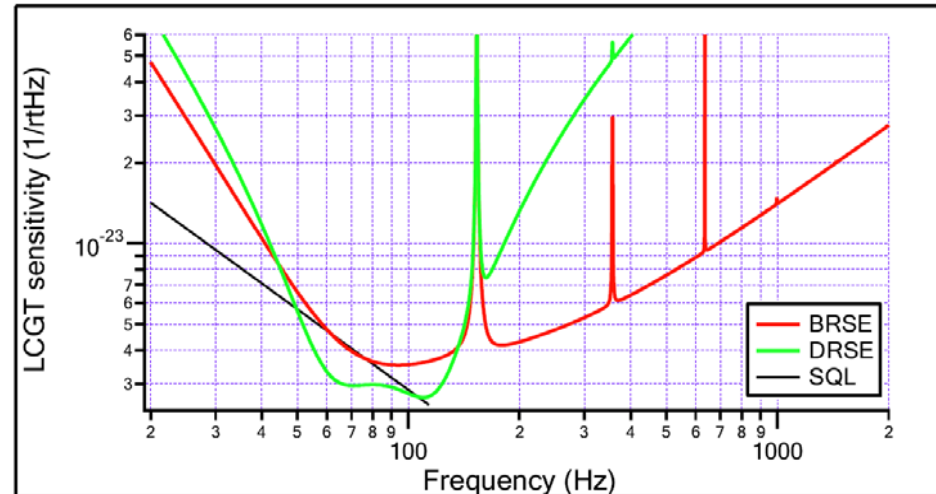
Thermal conductivity of Sapphire fiber

($P_{prc}=825W$, $F=1550$, $W_{sub}=0.24W$, $W_{coa}=0.20W$)

BRSE or DRSE

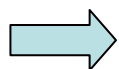


* GW from the optimal direction



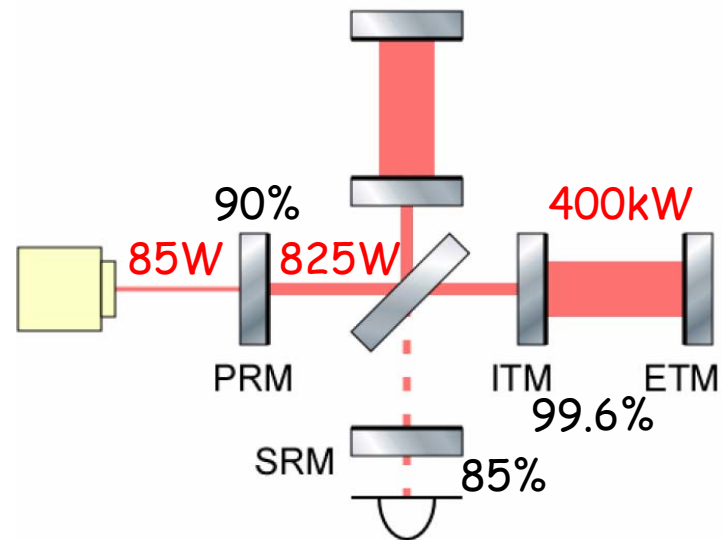
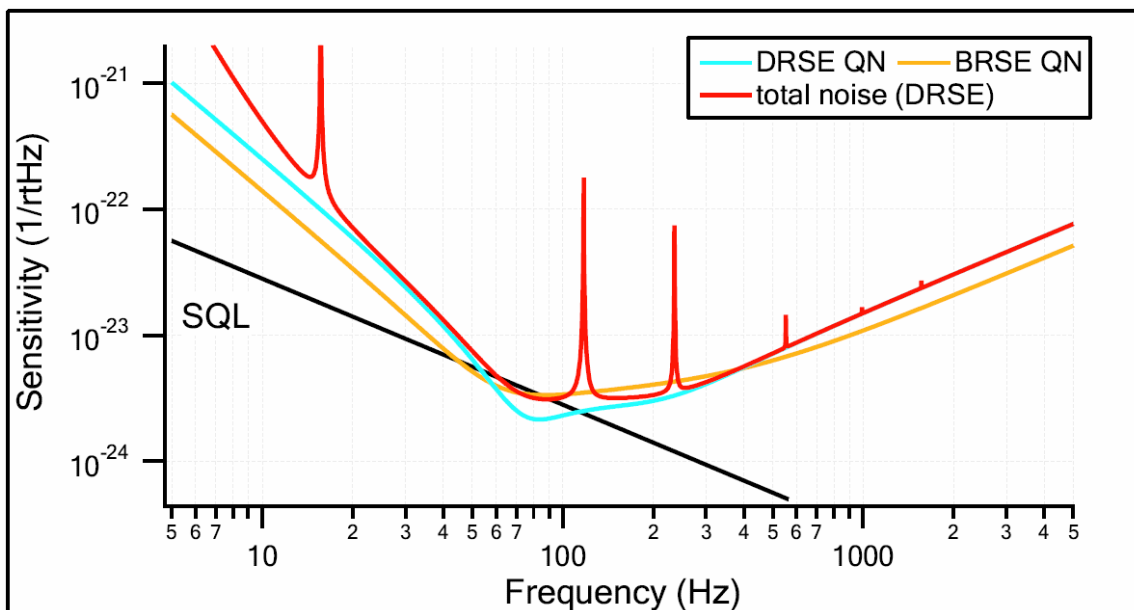
* These curves are old ones.

- Detuned RSE = narrow band, high inspiral range ($\sim 300\text{Mpc}$)
low-finesse \times high SR is the best
- Broadband RSE = broad band, low inspiral range ($\sim 250\text{Mpc}$)
high-finesse \times low SR is the best



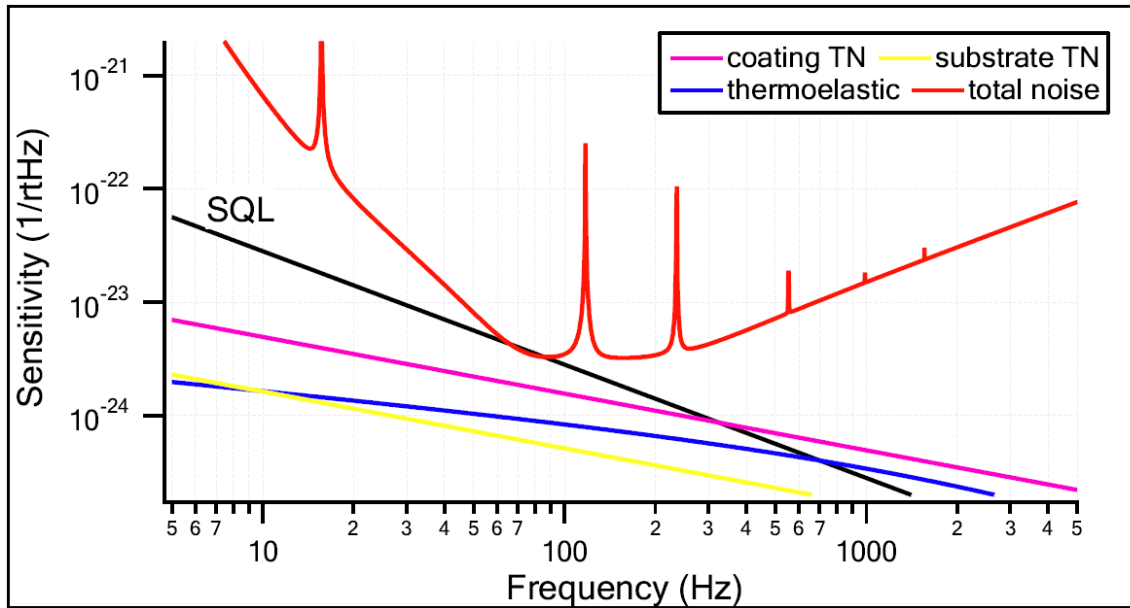
We decided to choose an intermediate one
(273Mpc and quite broad)

Quantum noise



- For DRSE, $\phi=86.5$ deg, $\zeta=134.2$ deg
- For BRSE, $\zeta=119.3$ deg
- The best sensitivity is better with DRSE
- Bandwidth is broader with BRSE
- QN exceeds the SQL at around a certain frequency

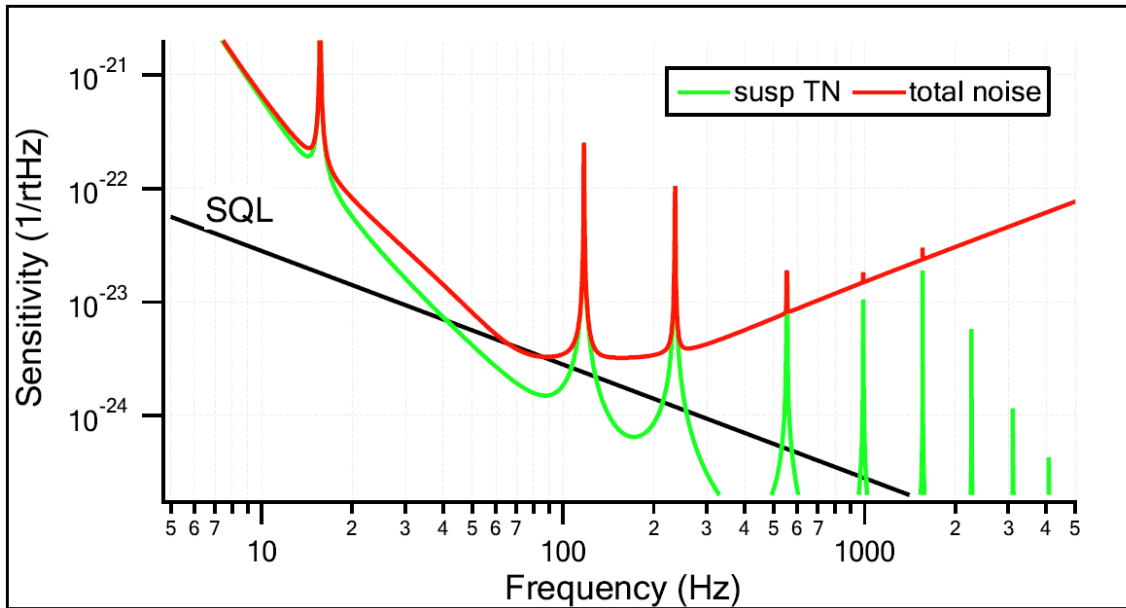
Mirror thermal noise



Mirror temperature 20K
Substrate $Q=1e8$
Tantala coating $\phi=5e-4$
Silica coating $\phi=3e-4$
ITM:9 layer, ETM:18 layer
Beam radii are 3.4/4.5cm
(flat-concave)

- Averaged coating loss requirement = $3.9e-4$
- Measured value in Glasgow = $6.7e-4$ ($8e-4/5e-4$)
- Measured value in U of Tokyo = $5.0e-4$
- A new coating experiment at NAO

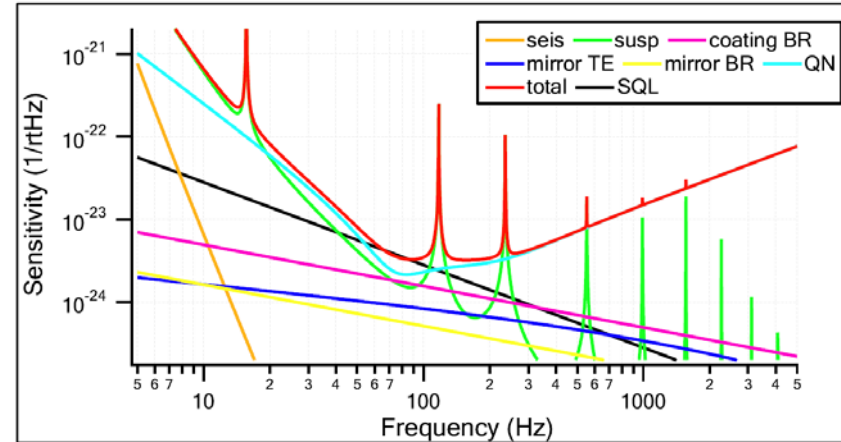
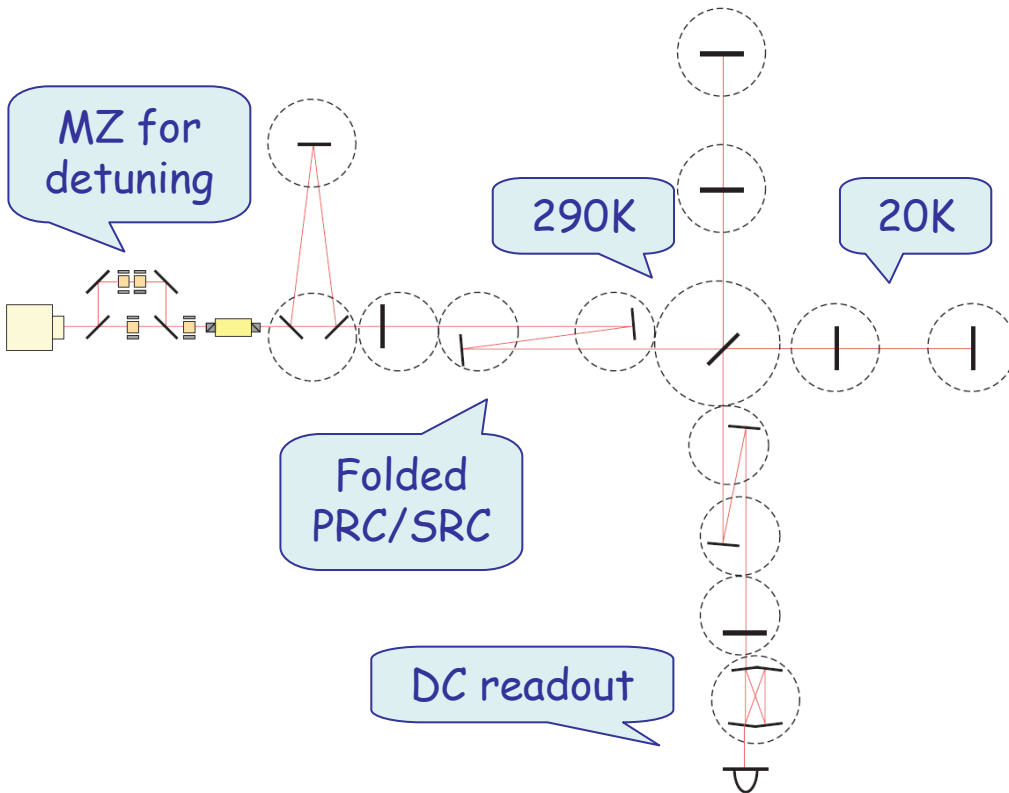
Suspension thermal noise



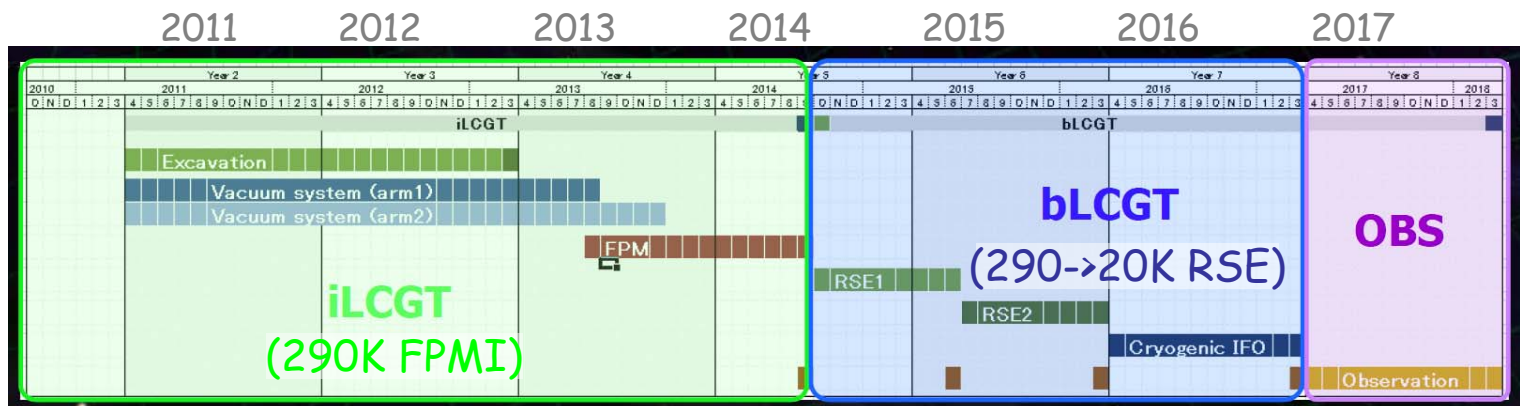
Values are for TM/IM/RM fiber
(test mass/intermediate mass/recoil mass)
Material=Sapphire/Tungsten/BeCu
Structure loss= $5e-8/1e-4/5e-6$
Fiber length= $30\text{cm}/50\text{cm}/30\text{cm}$
Fiber $d=1.6\text{mm}/0.6\text{mm}/0.4\text{mm}$
Clamp loss= $0/1e-3/0$
Ave Temperature= $16\text{K}/10\text{K}/16\text{K}$
Mini GAS freq= 0.4Hz
HV coupling= $1/200$
IM/RM mass= $60\text{kg}/30\text{kg}$

- The peak at 117Hz: vertical resonance
- The peak at 235Hz: first violin
- HV coupling is bad due to the tilted floor ($1/300$) for water drainage

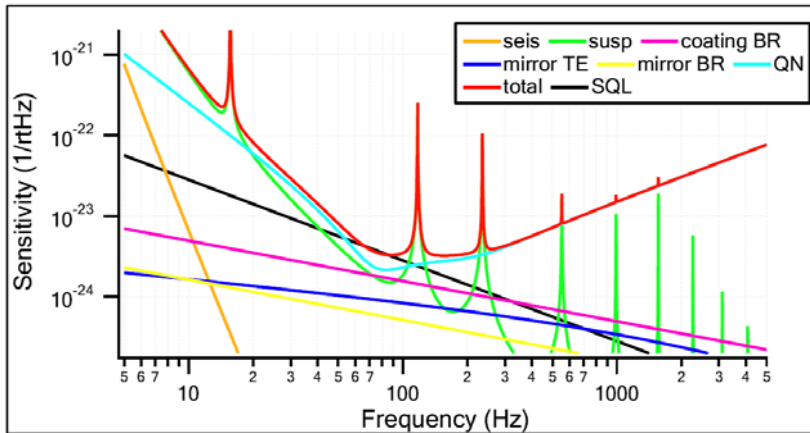
Optical setup and schedule



- BNS IR is 273Mpc
- BRSE Compatible (245Mpc)
- Cryogenic operation 2016~

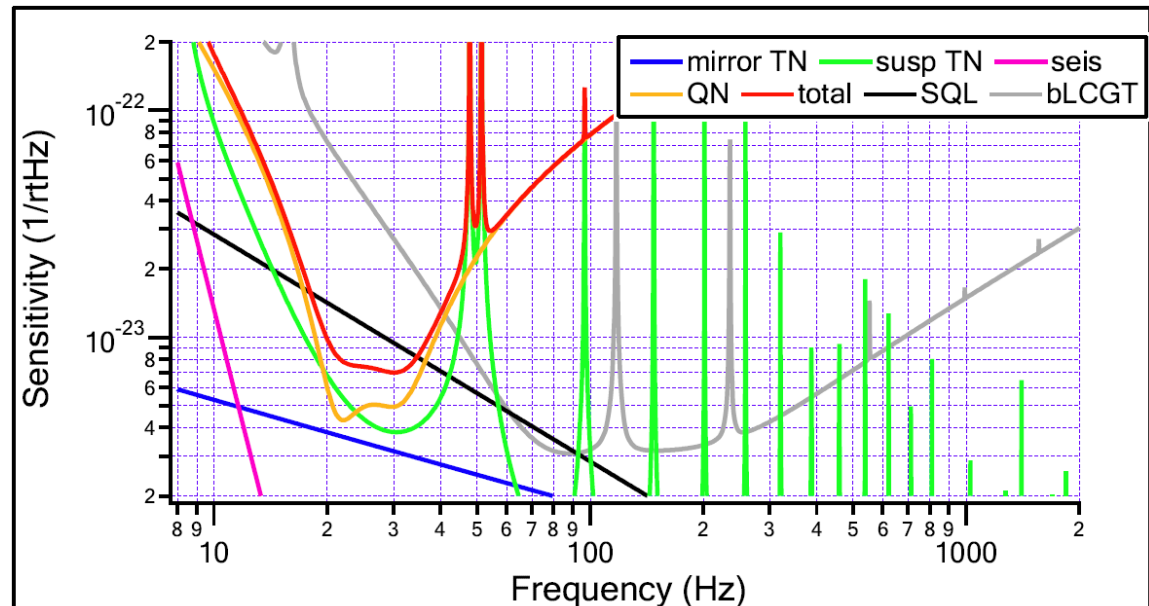


LCGT-LF discussion



- Seismic noise in underground is low
- Low-power operation may shift the curve to a lower frequency band
- Xylophone with aLIGO/AdVirgo

- Input power 1.5W
- PRG=11, Rsr=88%
- Finesse 1050
- Fiber length 120cm
- Fiber thickness 1.4mm
- Safety factor in cooling



Summary

- LCGT is now under construction
- 290K obs (2014), 20K obs (2017~)
- Low seismic noise [SAS design -> DeSalvo's talk]
- Cryogenic operation
[Mirror Cryostat design -> Kimura & Sakakibara's talk]
- Sub-SQL
- LCGT-LF discussions

-
- Risk management
 - Possible future extension
- } Nishida's talk

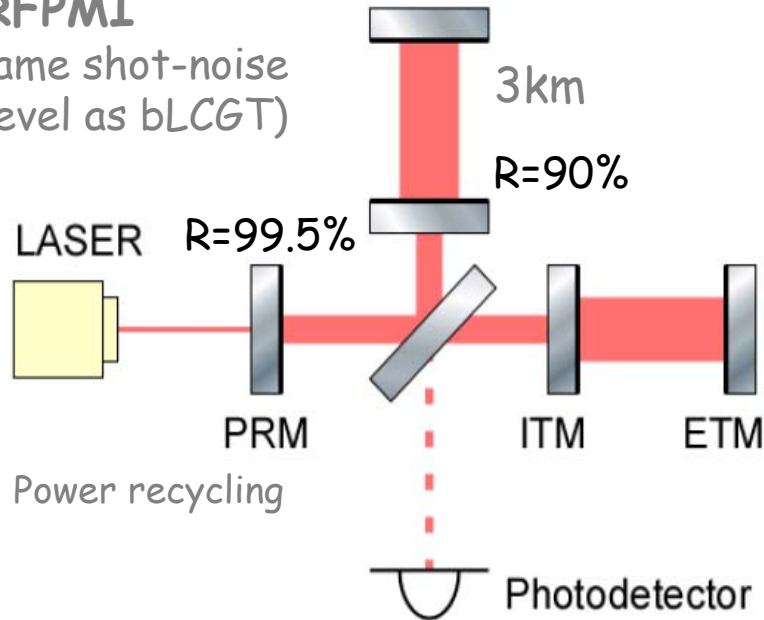
Supplementary slides

RSE

RSE=Resonant Sideband Extraction

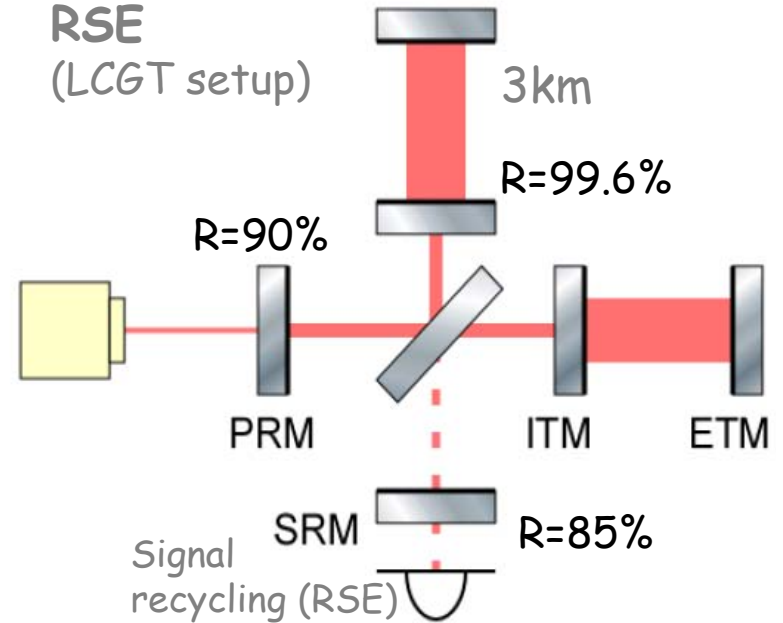
PRFPMI

(same shot-noise level as bLCGT)



RSE

(LCGT setup)



Laser power transmitting ITM would be 25 times higher w/o RSE



RSE is good for LCGT

825W in PRC, Arm power=400kW
Absorption in ITM substrate=0.24W
Absorption in coatings=0.20W