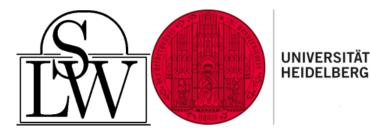
The vanishing of the primary emission region in PKS 1510-089

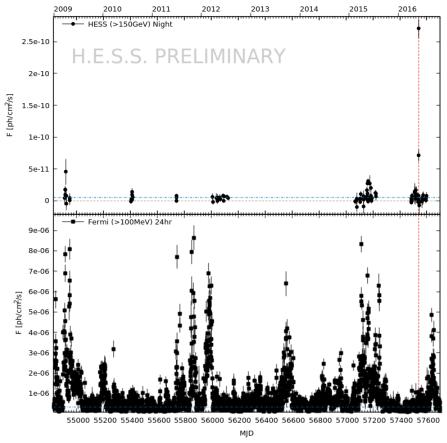
Michael Zacharias, J. Barnard, M. Böttcher, H. Schutte for the H.E.S.S. Collaboration



TeVPA 2023 — Naples



PKS 1510-089 — a brief history

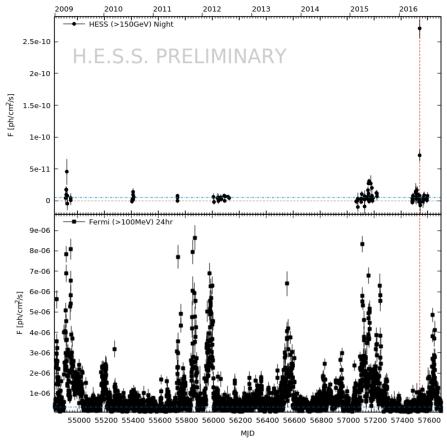


Light curves of PKS 1510-089 (FSRQ) from H.E.S.S. (top) and *Fermi*-LAT (bottom) (preliminary)

- FSRQ at redshift z = 0.361 with a bright BLR
- Expect strong attenuation of VHE γ rays
- Detected with H.E.S.S. in 2009 (HESS+13)
- A deep monitoring campaign revealed...
 - VHE emission is persistent (MAGIC+18)
 - In May 2016, a bright almost orphan VHE flare occurred (HESS+21)
 - Another VHE flare took place in 2019



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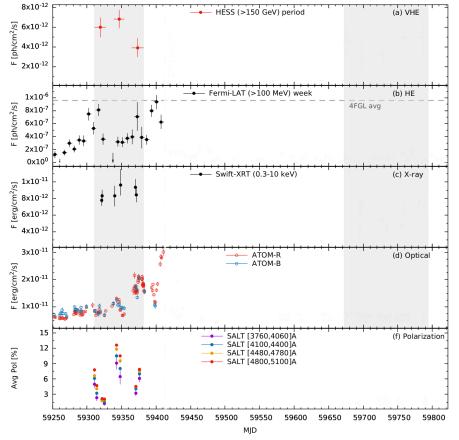


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 - Another VHE flare took place in 2019
 - ... "but we're not here to see any of that!"

(yt:TheTimTraveller)



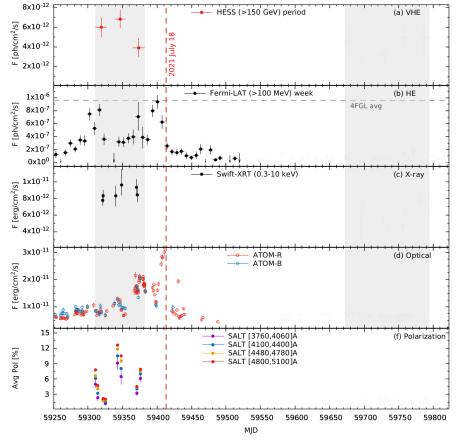


• MWL monitoring program since 2021:

- H.E.S.S., Fermi-LAT, Swift, ATOM
- SALT (spectro-polarimetry)

MWL light curves of PKS 1510-089 (FSRQ) in 2021 from H.E.S.S., Fermi-LAT, Swift, and ATOM, as well as degree of polarization from SALT.

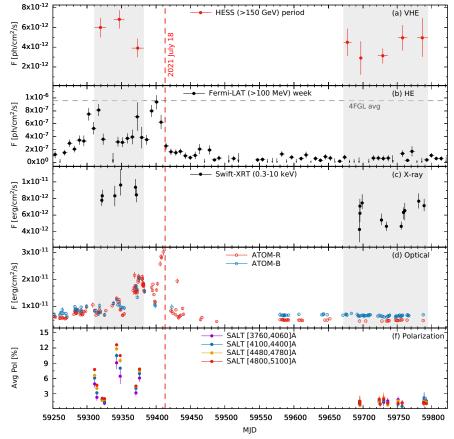




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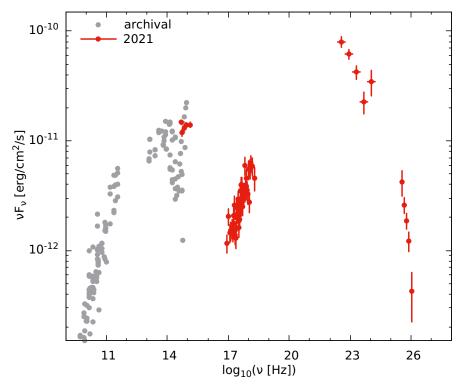
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MWL light curves of PKS 1510-089 (FSRQ) in 2021 & 2022 from H.E.S.S., Fermi-LAT, Swift, and ATOM, as well as degree of polarization from SALT.



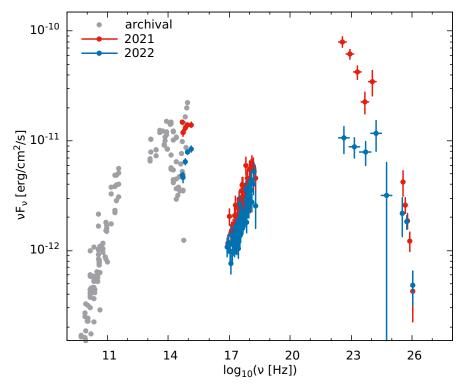
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 - H.E.S.S., Fermi-LAT, Swift, ATOM
 - SALT (spectro-polarimetry)
- For the rest of 2021, HE γ -ray and optical fluxes went down
- And have stayed down since
- Additionally, the degree of polarization has dropped to 0%
- But VHE γ-rays and X-rays have remained almost unchanged



MWL spectrum of 2021 (red) with archival data (gray)

- Spectra averaged during H.E.S.S. observation windows
 - In 2021, this is before the drop!





MWL spectrum of 2021 (red) & 2022 (blue) with archival data (gray)

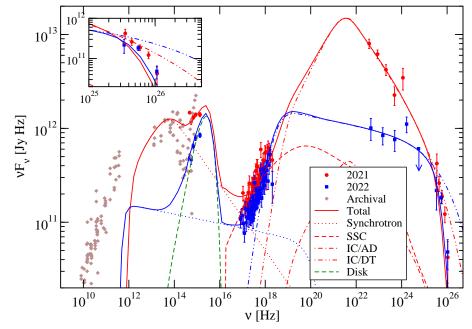
- Spectra averaged during H.E.S.S. observation windows
 - In 2021, this is before the drop!
- Significant change in HE γ -ray spectrum
 - Flux drop and hardening
 - Spectral continuation to VHE domain unchanged
 - Spectral change coincides with flux drop taking place within a few days
- Significant change in optical spectrum

(explored using Schutte+22)

- 2021 is a mix of synchrotron, AD and BLR
- 2022 fully explained by (constant) AD and BLR
- Derived upper limit on synchrotron spectrum



PKS 1510-089 — One-zone modeling

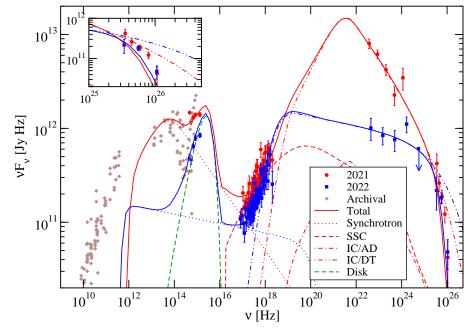


One-zone modeling of the 2021 & 2022 data sets.

- Using a leptonic steady-state one-zone code (Böttcher+13)
- One-zone modeling of both states possible
 - Requires significant changes to electron distribution function (normalization, index)
 - Must be located relatively far away from the black hole (\gtrsim 10pc)
 - Changes not likely to happen on such a short time scale
 - Fit to VHE data not so good



PKS 1510-089 — One-zone modeling

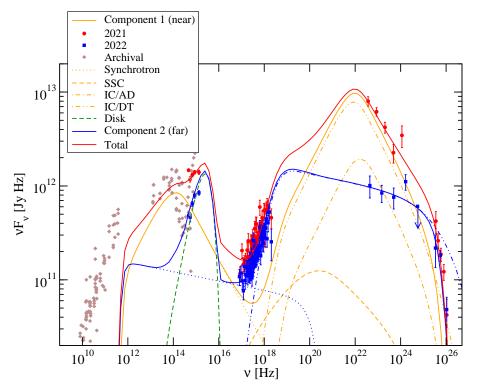


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- What if it were two zones?



PKS 1510-089 — Two-zone modeling

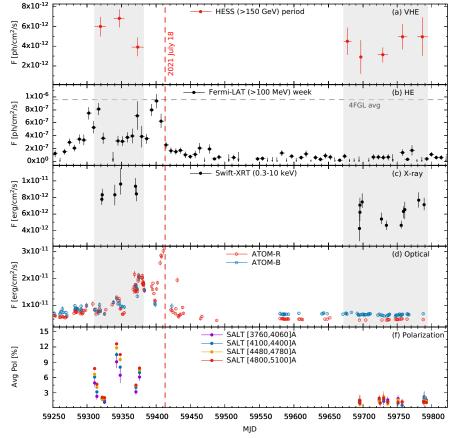


Two-zone interpretation of the data with a near zone (yellow), a far zone (blue) and the total (red).

- 2022 "one-zone model" fits X-rays and VHE γ rays in a "far zone" (\gtrsim 10pc)
- Another "near zone" (close to the BH) added for 2021 data explaining optical and HE γ -ray data
- "Near zone" disappeared in July 2021
 - Inner jet became inefficient
 - Inner jet has swung away from the line-of-sight
- It would explain
 - ... rapid change of state
 - ... persistent VHE flux (MAGIC+18)
 - ... unclear correlation patterns (Zacharias+19)
 - ... most of "ancient" variability in near zone



PKS 1510-089 — Summary



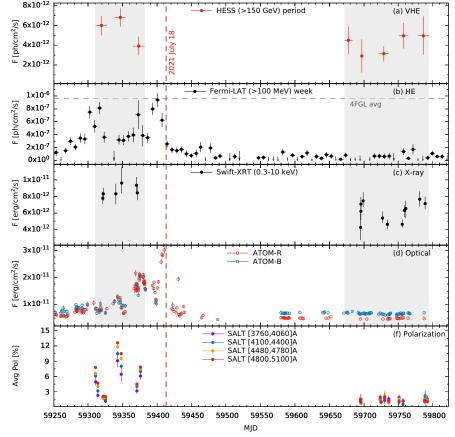
MWL light curves of PKS 1510-089 (FSRQ) in 2021 & 2022 from H.E.S.S., Fermi-LAT, Swift, and ATOM, as well as degree of polarization from SALT.



- Historically, 2 zones were active at the same time
- In July 2021, the near-zone emission region disappeared
- Only the far-zone emission region is still active
- Radio (and far-IR) data may shed light on the cause of this identity crisis

All details: HESS+23, ApJL, 952, L38

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Thank you!