

Recent IceCube results on the origin of cosmic neutrinos

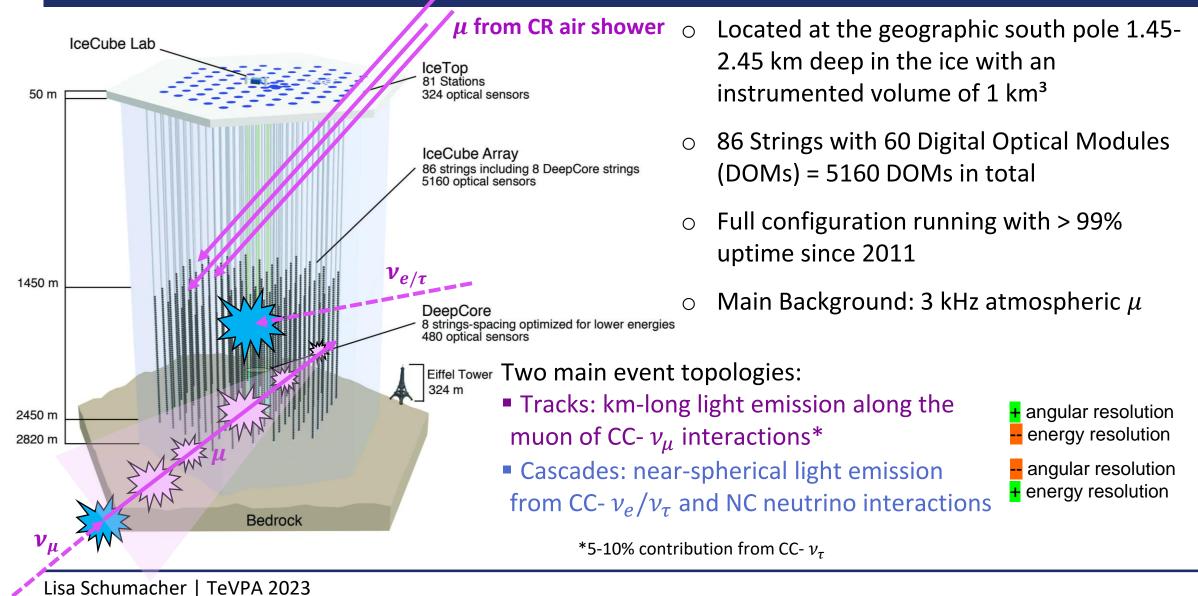
Lisa Schumacher

for the IceCube Collaboration



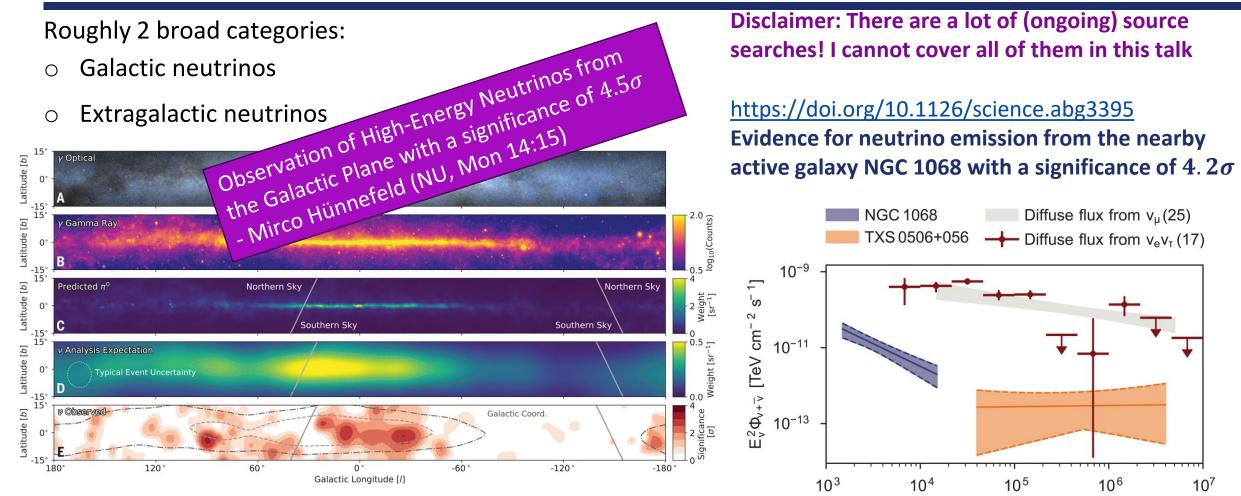
The IceCube Neutrino Observatory





Overview – our recent highlights





https://doi.org/10.1126/science.adc9818 Major contributions by Mirco Hünnefeld and Steve Sclafani

Lisa Schumacher | TeVPA 2023

Ev [GeV]

Galactic plane – diffuse emission



- Observation of High-Energy Neutrinos from the Galactic Plane with a novel machine-learning reconstruction of Cascades (all sky)
 Significance: 4.5σ
- 2) Results from searches for astrophysical neutrino sources in the southern sky and galactic plane using IceCube starting Track events Significance: $< 2\sigma$
- Galactic and Extragalactic Analysis of the Astrophysical Muon Neutrino Flux with 12.3 years of IceCube Track Data (Northern Sky)
 Significance: 2.7σ
 - Three independent results by now using different detection channels
 - Results using tracks are not yet significant, but are compatible with cascades
 - Work ongoing on combining these independent data sets

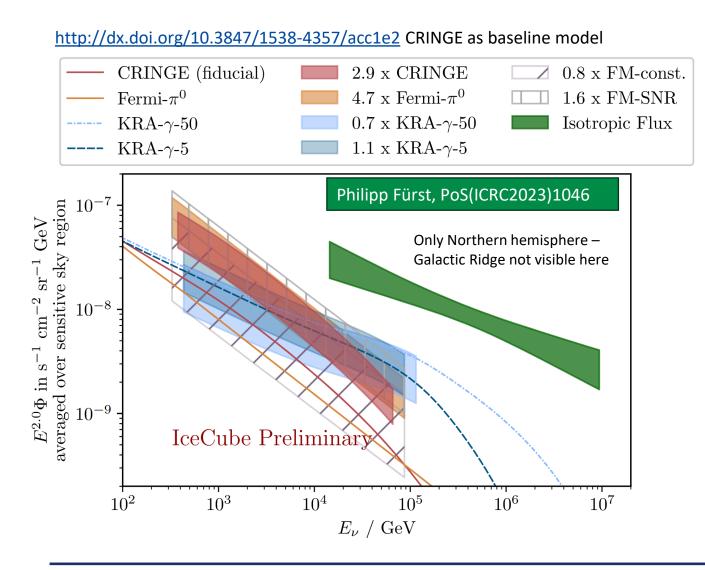
Mirco Hünnefeld (NU, Mon 14:15)

Sarah Mancina (NU, Thu 16:30)

Philipp Fürst, PoS(ICRC2023)1046

What we (don't) know





- ✓ We see the galactic plane in neutrinos!
- View of our Galaxy at highest energies & complementary to Fermi and LHAASO gamma rays
- ✓ About 6-13% of total measured (diffuse) neutrino flux comes from GP
- Multiple diffuse neutrino emission models were tested, but no clear preference yet
- Normalization larger than expected could hint at sources in addition to diffuse emission, but this is still speculative

Galactic sources

- Starting tracks: no galactic sources identified
- Enhanced track selection at the Galactic Center region: no time-dependent emission found at the location of Sgr A*
- $\circ~$ Through-going tracks: no significant evidence for extended sources in the Galactic Plane, but 2.6σ significance at region of unidentified TeV gamma-ray source 3HWC J1951+266
- Cascades: $> 3\sigma$ significance for correlation with TeV-gamma ray sources in GP, cannot be disentangled from the diffuse GP emission due to large angular uncertainty of cascades

We start to see the (diffuse) neutrino emission of the plane, but individual sources cannot be resolved yet ICECUBE

Sarah Mancina, NU, Thu 16:30

Xinyue Kang PoS(ICRC2023)1051

Upcoming publication arXiv:2307.07576

Same cascade data as used for GP analysis

NGC 1068



Seyfert II galaxy at a distance of 14.4 Mpc (very Ο Evidence for neutrino emission from Ο close by!) with Compton-thick AGN NGC 1068 with significance of 4.2σ using tracks Significance based on pre-defined source list Ο X-Ray corona around accretion disk may enable Ο A-posteriori analysis: NGC 4151 might be Ο neutrino production & gamma-ray absorption associated with another neutrino hotspot IceCube (this work) Electromagnetic observations (26) Theoretical v model (52,55) 0.1 to 100 GeV gamma-rays (40,41) +75° Theoretical v model (53) > 200 GeV gamma-rays (42) 10^{-9} **10**⁻¹⁰ 's PKS 1424+240 X-ravs $E^2 \Phi$ [TeV cm⁻² 10 - 11 TXS 0506+056 NGC 1068 10 ⁻¹² Oh **10**⁻¹³ 10^{-14} 3 5 **10**⁻¹⁵ **10**⁻¹² 10⁻⁶ 10³ $-LOG_{10}(p_{IOCAI})$ 10^{-9} 10^{-3} 10^{0} 10^{6} https://doi.org/10.1126/science.abg3395 Energy [GeV]

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 10^{1}

ິທີ

20

 10^{0}

 10^{-9}

intrinsic flux (2-10 keV) [erg $\ cm^2$ 10_{-13} [erg $\ cm^2$ $\ cm^2$

21

Seyfert Galaxies & X-ray bright AGN

Excess of neutrinos associated with two sources, Ο NGC 4151 and CGCG 420-015 @ 2.7σ significance

23

24

NGC 4151

data from Ricci et al. (2017)

 10^{4}

BAT AGN Spectroscopic Survey

 10^{3}

Results constrain the collective neutrino emission \cap from chosen source catalogue

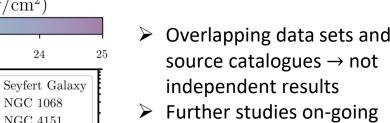
 $\log_{10}(\text{column density/cm}^2)$

22

 10^{2}

Shiqi Yu, (MM, Thu 14:40)

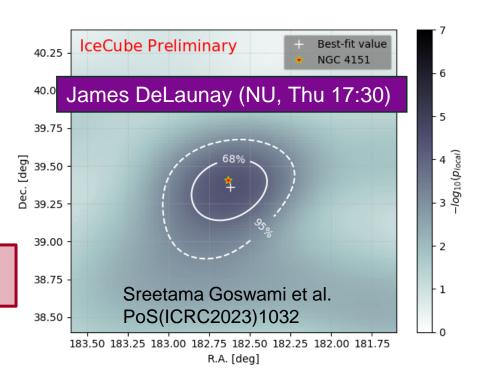
distance [Mpc]



Open questions remain about \geq neutrino production mechanism in source candidates

Intriguing results, but no rejection of isotropic neutrino flux yet!

- Search for high-energy neutrino emission Ο from hard X-ray AGN
- Confirmed emission of NGC 1068 and found Ο NGC 4151 @ 2.9 σ significance

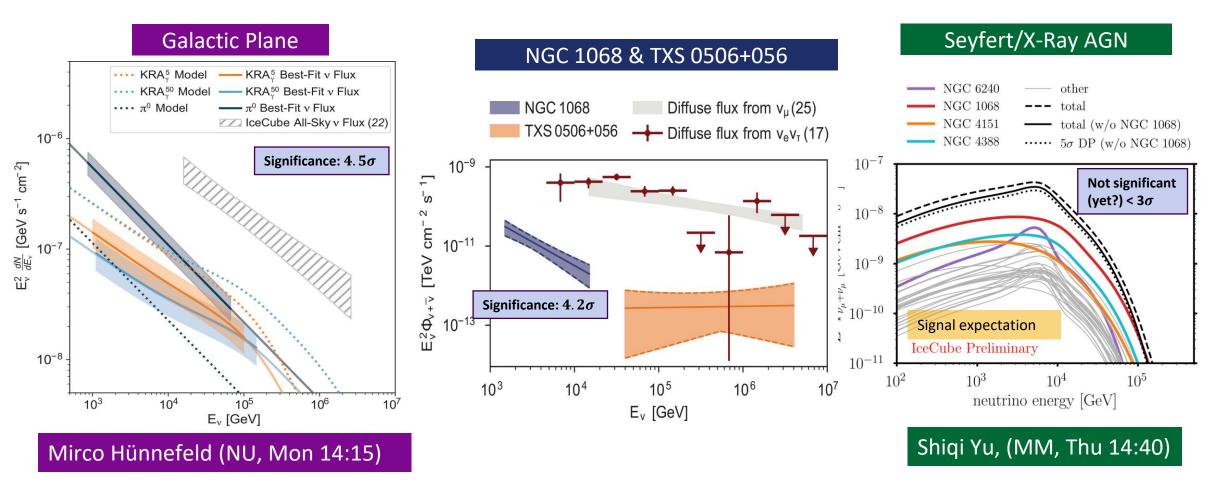




The bigger picture



- 1) Galactic & extragalactic neutrino associations (start to) appear
- 2) Significant proportion of overall neutrino flux is still not accounted for



Overview – IceCube at this conference



Contributions regarding the origin of astrophysical neutrinos:

- Observation of High-Energy Neutrinos from the Galactic Plane Mirco Hünnefeld (NU, Mon 14:15)
- Investigating Millimeter-Bright AGN as IceCube's Astrophysical Neutrino Sources - Alina Kochocki (NU, Tue 15:00)
- The IceCube Realtime Program Giacomo Sommani (MM, Wed 17:10)
- Development of a new IceCube realtime alert using multiplet signal for optical follow-up Nobuhiro Shimizu (MM, Wed 17:25)
- IceCube Search for High Energy Neutrino Emission from X-ray Bright Seyfert Galaxies - Shiqi Yu (MM, Thu 14:40)
- Results from searches for astrophysical neutrino sources in the southern sky and galactic plane using IceCube starting track events -Sarah Mancina (NU, Thu 16:30)
- A search for neutrinos from dark matter in the Galactic Centre with IceCube Nhan Chau (IDM, Thu 17:00)
- Searching for high-energy neutrino emission from hard X-ray AGN with IceCube - James DeLaunay (NU, Thu 17:30)

Other contributions:

- A combined Flavour composition measurement of astrophysical neutrinos using multi-sample IceCube data -Neha Lad (NU, Thu 16:45)
- Inelasticity studies using TeV-scale starting track neutrino events in IceCube - Marjon Moulai (NU, Thu 17:00)
- IceCube Starting Events For Diffuse Astrophysical Neutrino Measurements - Aswathi Balagopal V. (NU, Thu 17:45)
- New Results for eV-scale Sterile Neutrino Searches with IceCube - Alfonso Andres Garcia Soto (PL, Thu 9:00)
- Reconstruction and identification methods of sub-PeV gamma rays at the IceCube Neutrino Observatory -Federico Bontempo (GRA, Thu 17:15)

... + external contributions using IceCube results and/or public data

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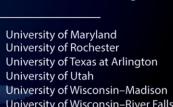
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Fonds de la Recherche Scientifique (FRS-FNRS) Fonds Wetenschappelijk Onderzoek-Vlaanderen C h (FWAVAanheen)

FUNDING AGENCIES

German Research Foundation (DFG) eVPA 202 Beutsches Elektronen-Synchrotron (DESY)

Federal Ministry of Education and Research (BMBF) Japan Society for the Promotion of Science (JSPS) The Swedish Research Council (VR) Knut and Alice Wallenberg Foundation Swedish Polar Research Secretariat

University of Wisconsin Alumni Research Foundation (WARF) US National Science Foundation (NSF)



Back up

IceCube data sets – an incomplete overview



- \circ Northern-Sky Tracks (NT), ν_{μ} -CC, only Northern Sky, high neutrino purity
- Point-Source Tracks (PS), v_{μ} -CC, full Sky, high neutrino purity only in Northern Sky,
 Southern Sky is CR-muon dominated
- \circ DNN Cascades: $v_{e/\tau}$ & NC all flavor, full sky, competitive with PS in Southern Sky
- $\,\circ\,$ ESTES: Enhanced starting track event selection, ν_{μ} , full sky, competitive with PS in Southern Sky
- Realtime alerts: Please see Giacomo Sommani's talk (MM, Wed 17:10) ☺
- \circ Cascades $v_{e/\tau}$ & NC, currently only used for diffuse analyses
- HESE & MESE: High-energy & Medium-energy starting events, All flavors, "traditionally" used for diffuse analyses

IceCube events in numbers





= 3000 atmospheric μ per second