# Status of the Light Dark Matter eXperiment Lene Kristian Bryngemark, Lund University

TeVPA, Napoli, Italy, September 11-15 2023









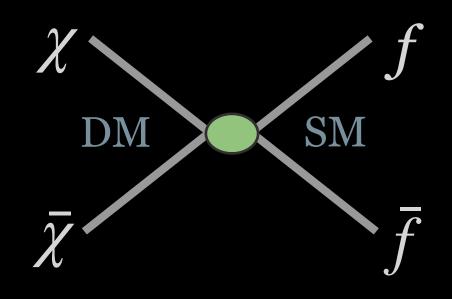
Is present-day dark matter a relic from the hot early universe?

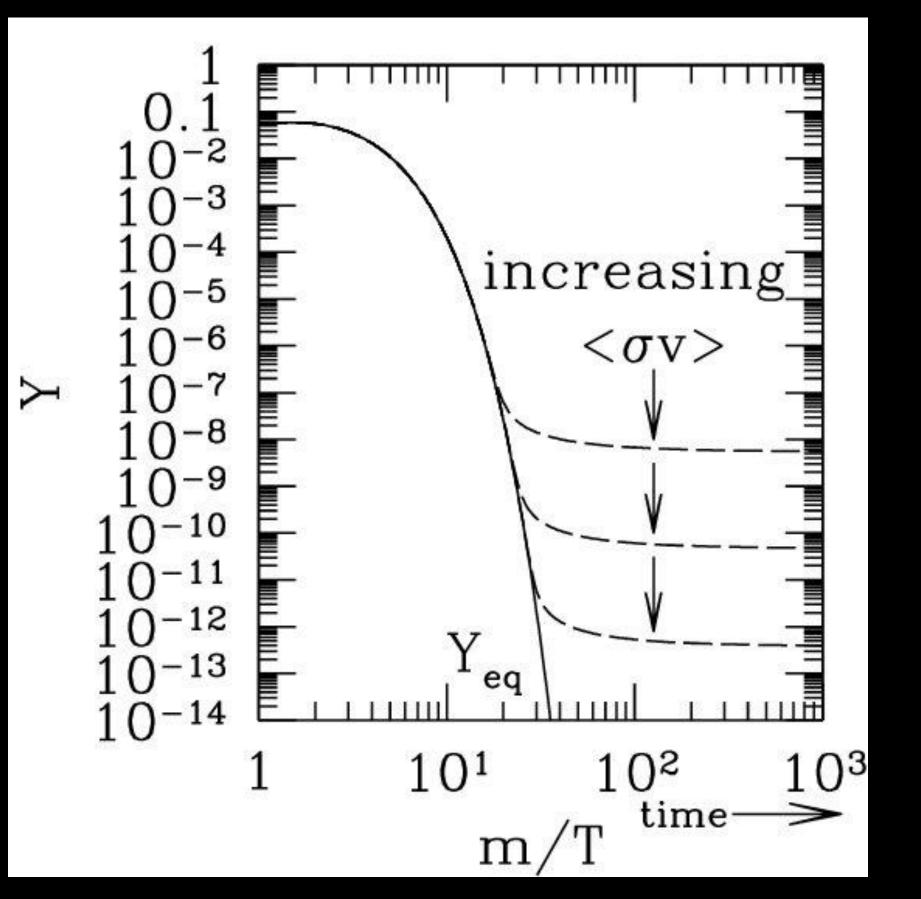








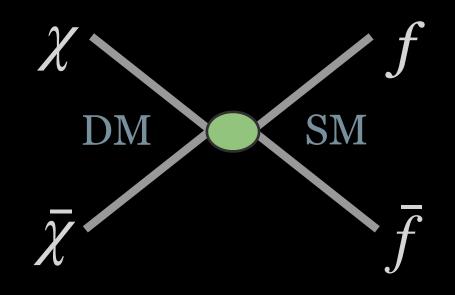


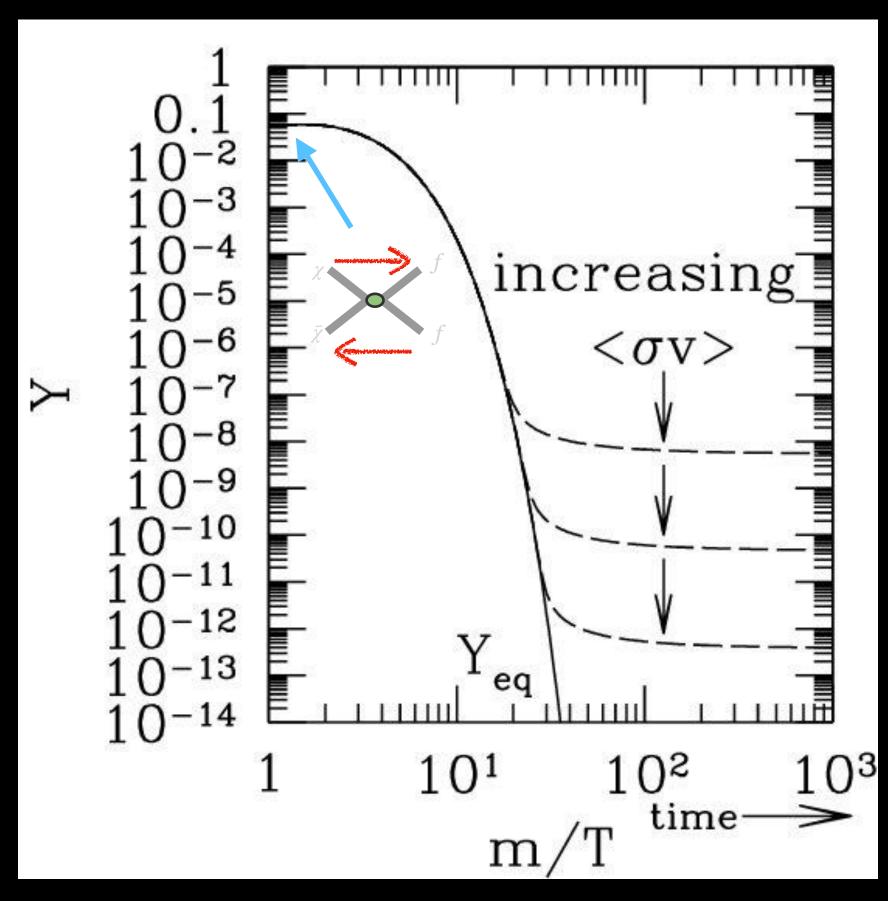


# some minimum non-gravitational interaction → reaches thermal equilibrium at some point in history





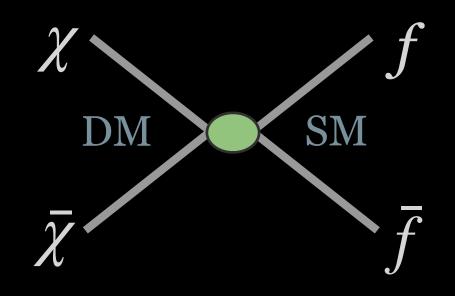


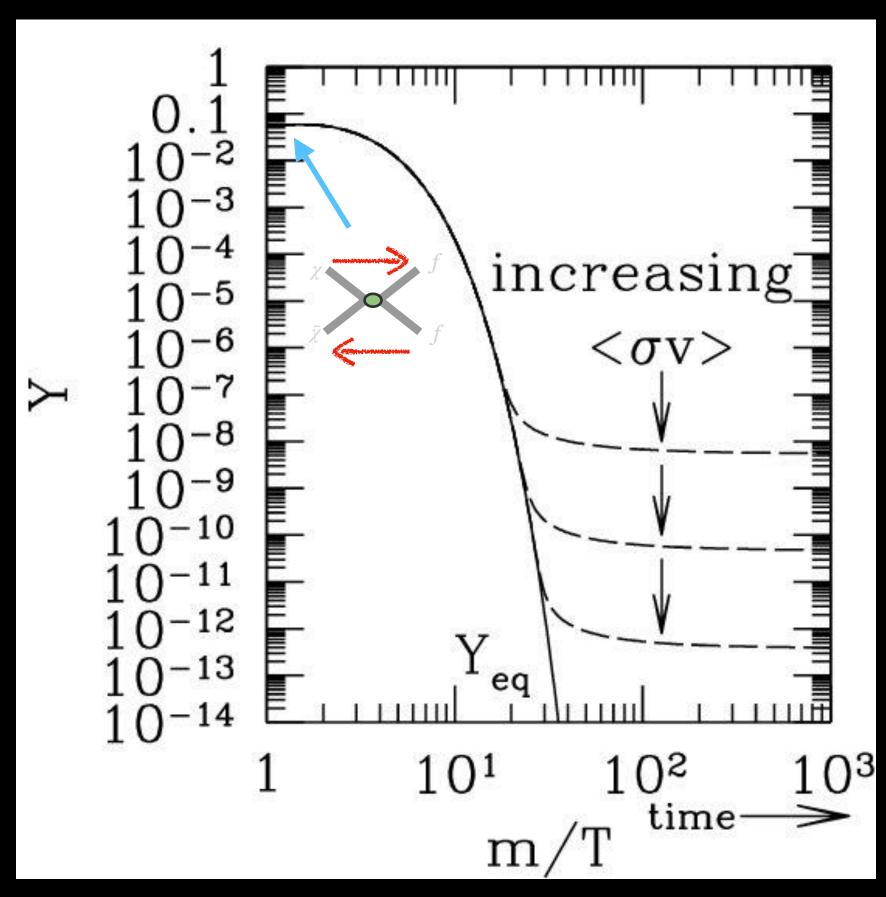


- $\rightarrow$  reaches thermal equilibrium at some point in history
- as the universe expands and cools...





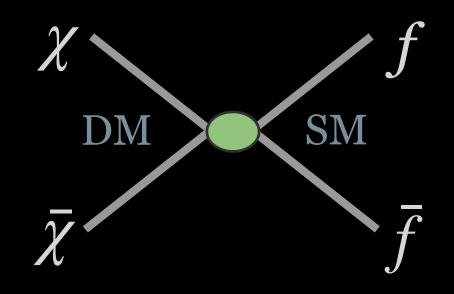


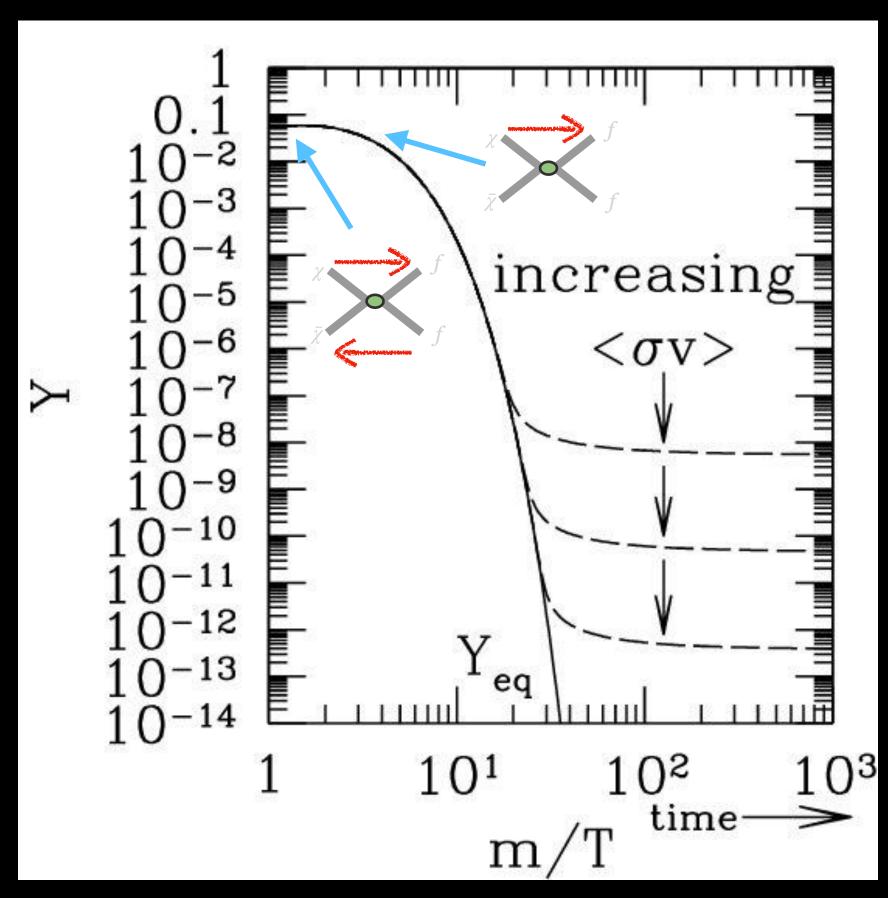


- $\rightarrow$  reaches thermal equilibrium at some point in history
- as the universe expands and cools...
  - process flows in one direction, until





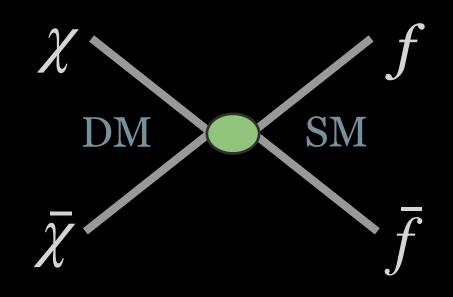


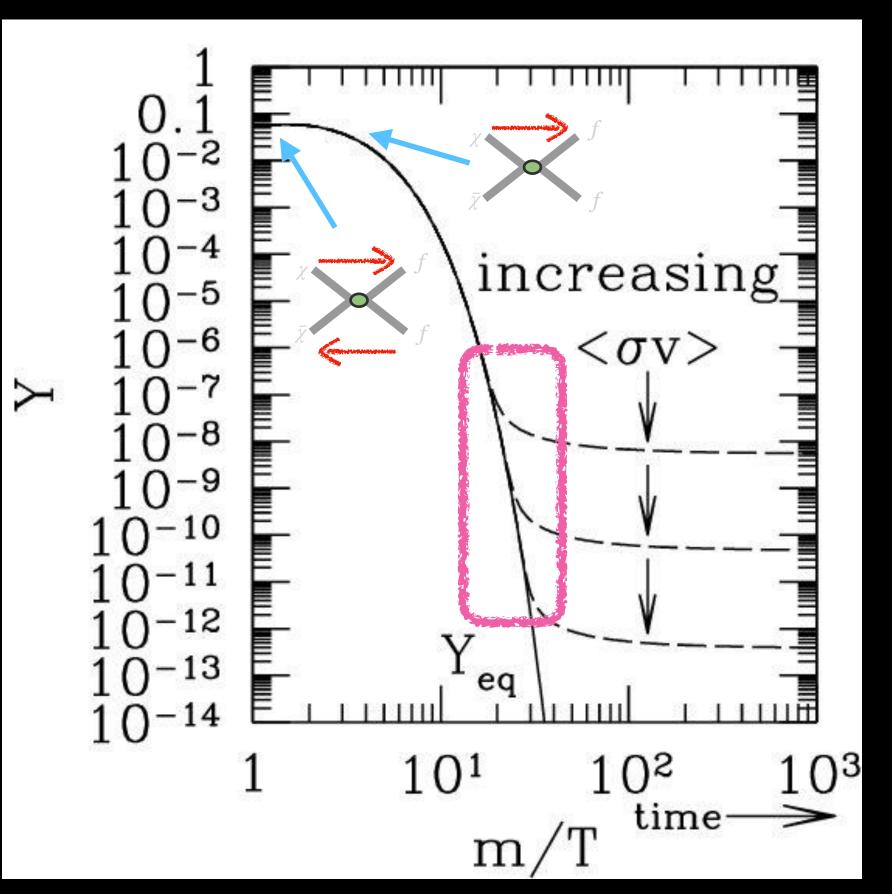


- $\rightarrow$  reaches thermal equilibrium at some point in history
- as the universe expands and cools...
  - process flows in one direction, until
  - annihilation and decays stop  $\rightarrow$  a thermal relic density we see today



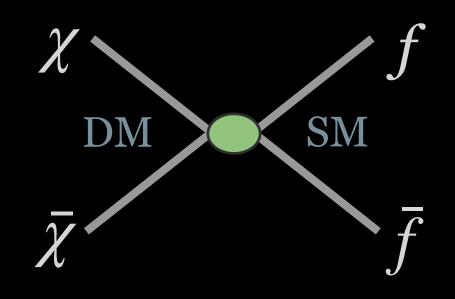


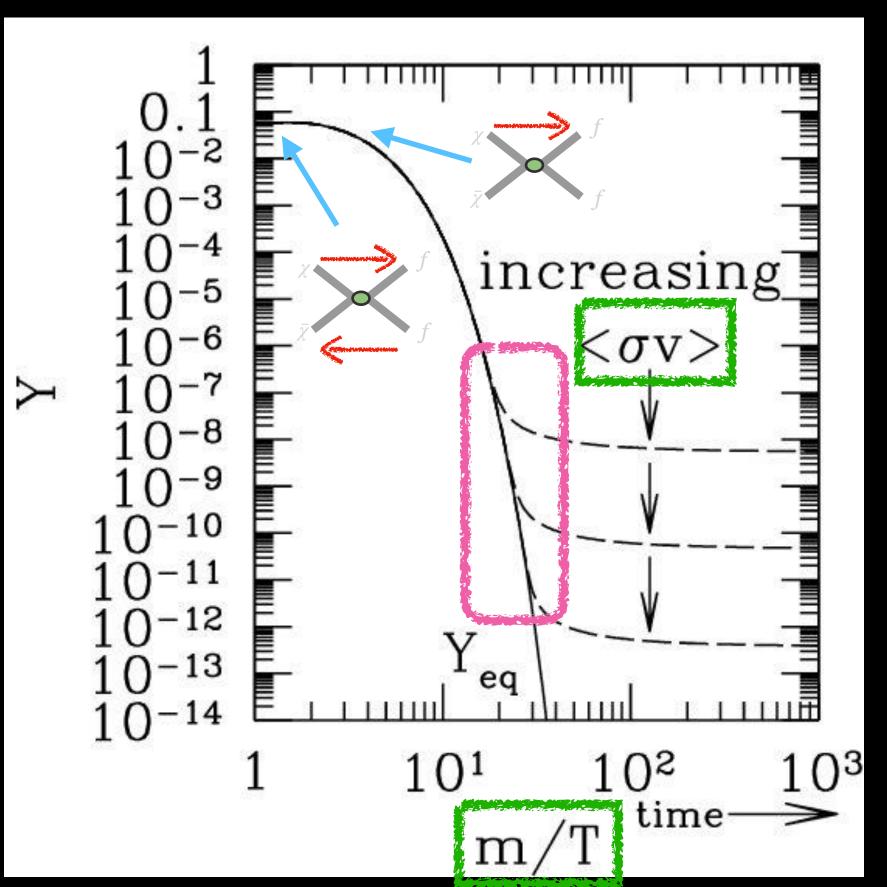




- $\rightarrow$  reaches thermal equilibrium at some point in history
- as the universe expands and cools...
  - process flows in one direction, until
  - annihilation and decays stop  $\rightarrow$  a thermal relic density we see today
- relates particle mass to interaction strength







#### predicts a minimum interaction strength experimental sensitivity target $\bigcirc$





https://arxiv.org/abs/1808.05219

Elastic Scalar

Majorana Fermion

Inelastic Scalar (small splitting)

10

Pseudo-Dirac Fermion V

 $10^{2}$ 

 $m_{\rm DM}$ 

N

(small splitting)

Smmetric Fermion

section Cross scattering DM-electron

 $10^{-35}$ 

 $10^{-37}$ 

10<sup>-39</sup>.

 $10^{-41}$ 

 $10^{-43}$ 

 $10^{-45}$ 

 $10^{-47}$ 

 $10^{-49}$ 

 $10^{-51}$ 

10<sup>-53</sup>

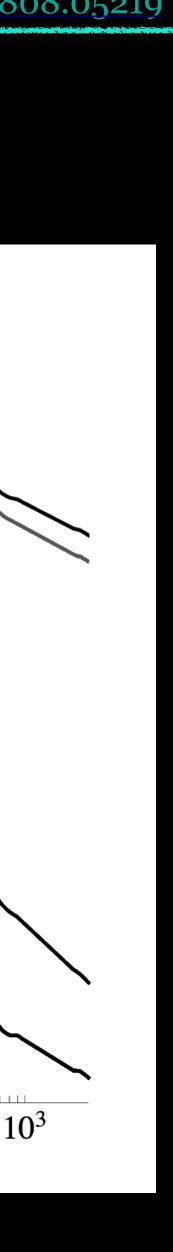
 $10^{-55.}$ 

 $10^{-57.}$ 

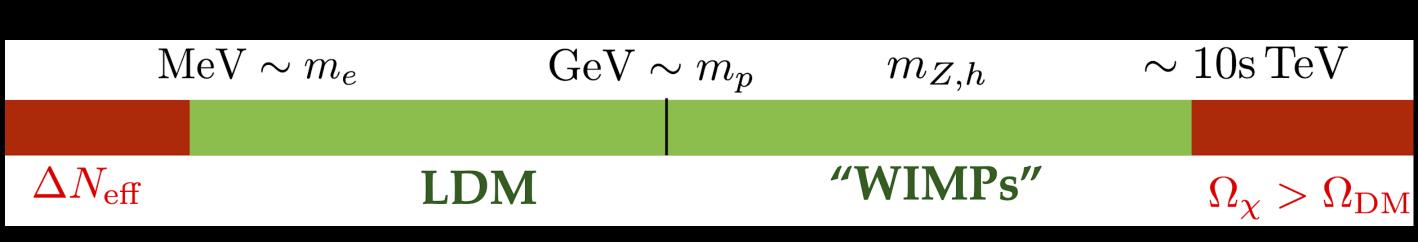
 $10^{-59.}$ 

 $\sigma_e$ 





- predicts a minimum interaction strength experimental sensitivity target  $\bigcirc$
- constrains the ~90 orders of mag. DM mass range

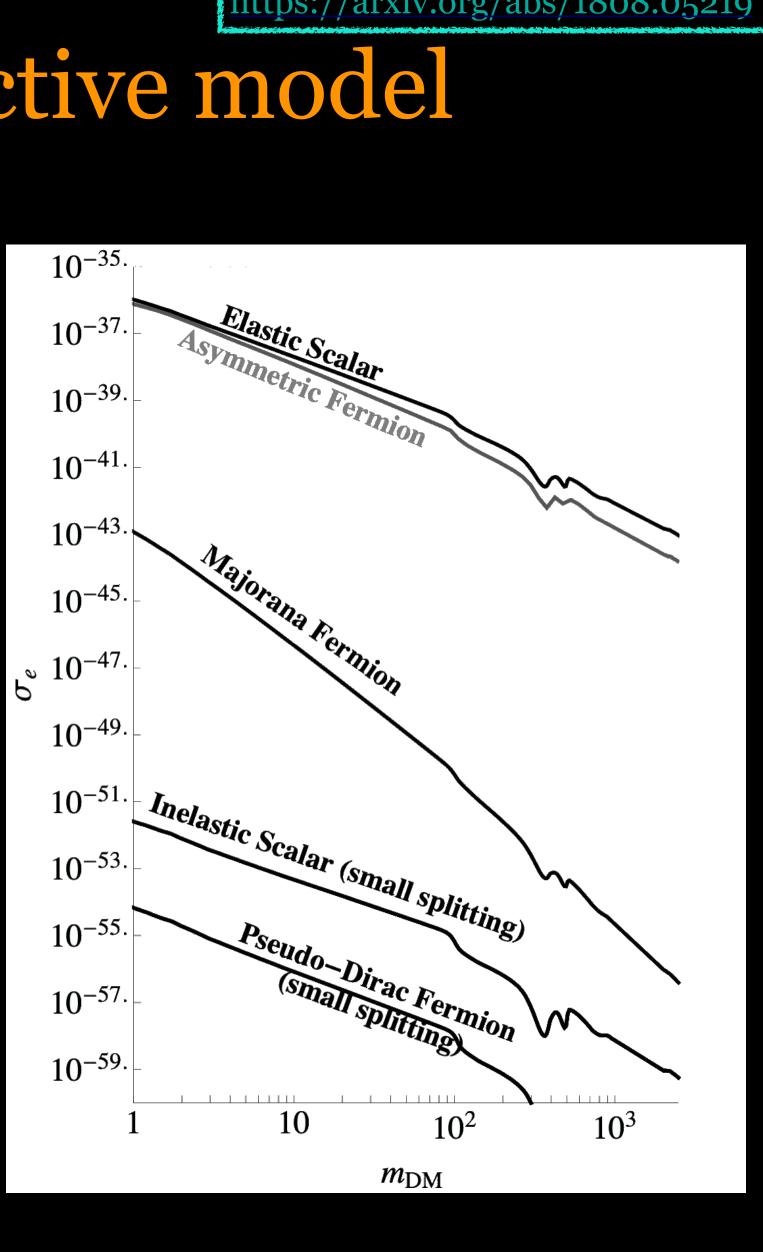




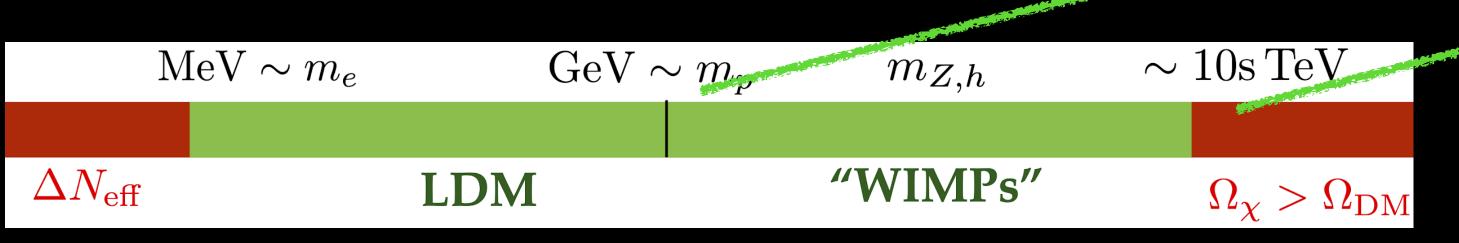


https://arxiv.org/abs/1808.05219

section Cross scattering DM-electron



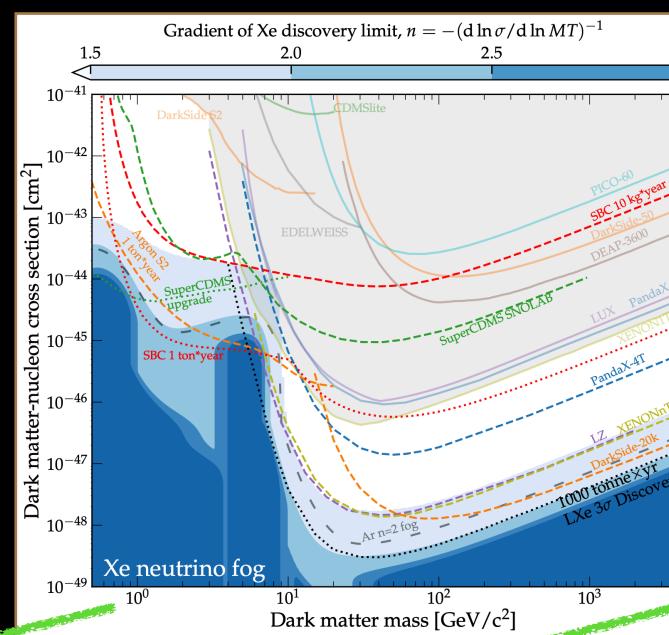
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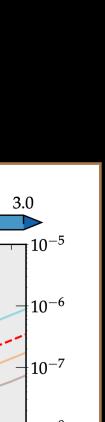


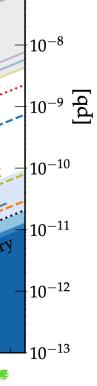
XMQ

https://arxiv.org/abs/1808.05219

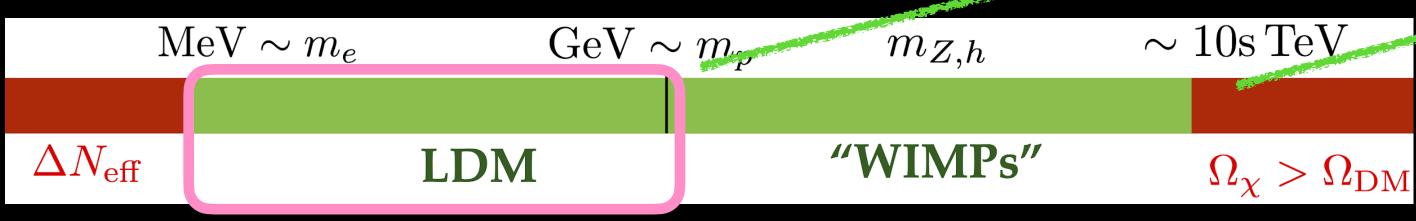


#### extensively explored by e.g. direct detection





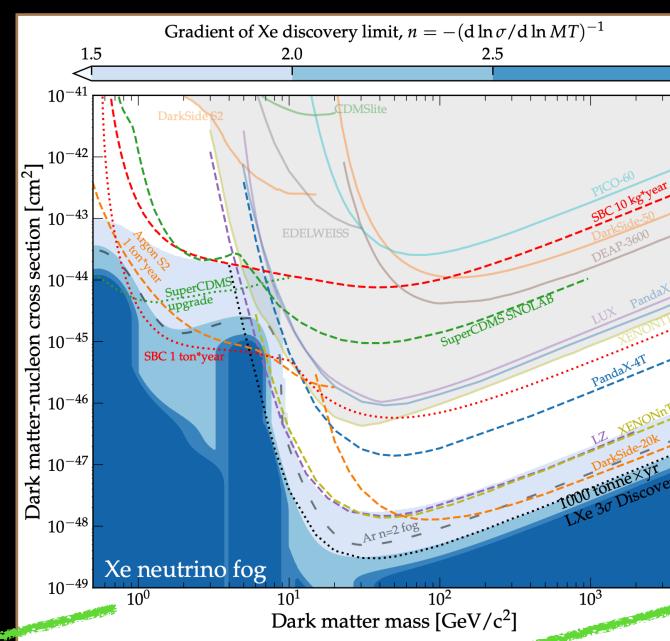
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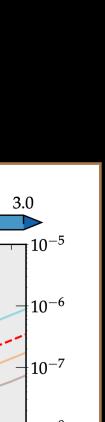


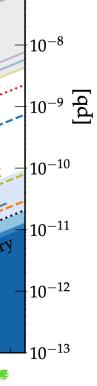


XMO

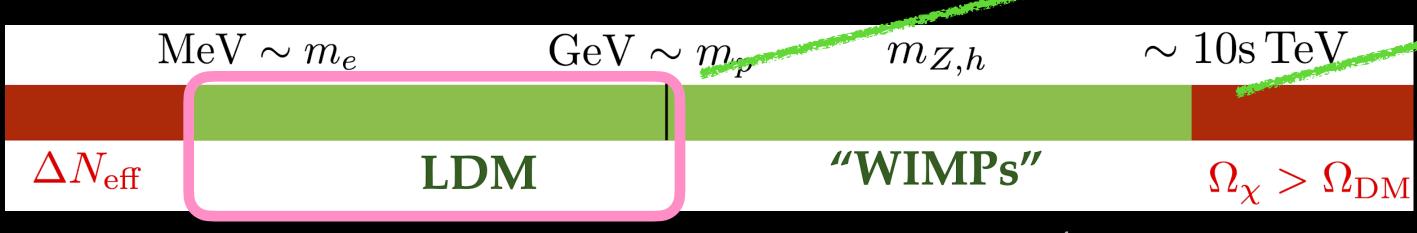
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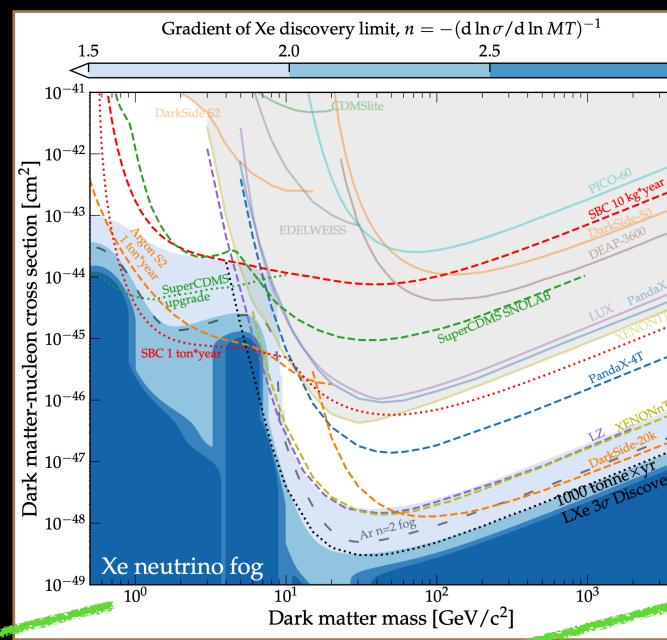
- predicts a minimum interaction strength experimental sensitivity target  $\bigcirc$
- constrains the ~90 orders of mag. DM mass range
  - light mediator: requires non-SM force/dark sector  $\bigcirc$

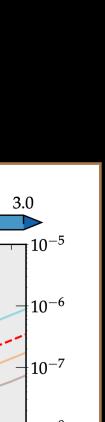


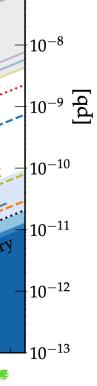


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- predicts a minimum interaction strength experimental sensitivity target  $\bigcirc$
- constrains the ~90 orders of mag. DM mass range
  - light mediator: requires non-SM force/dark sector  $\bigcirc$
  - not very constrained by experiments  $\bigcirc$

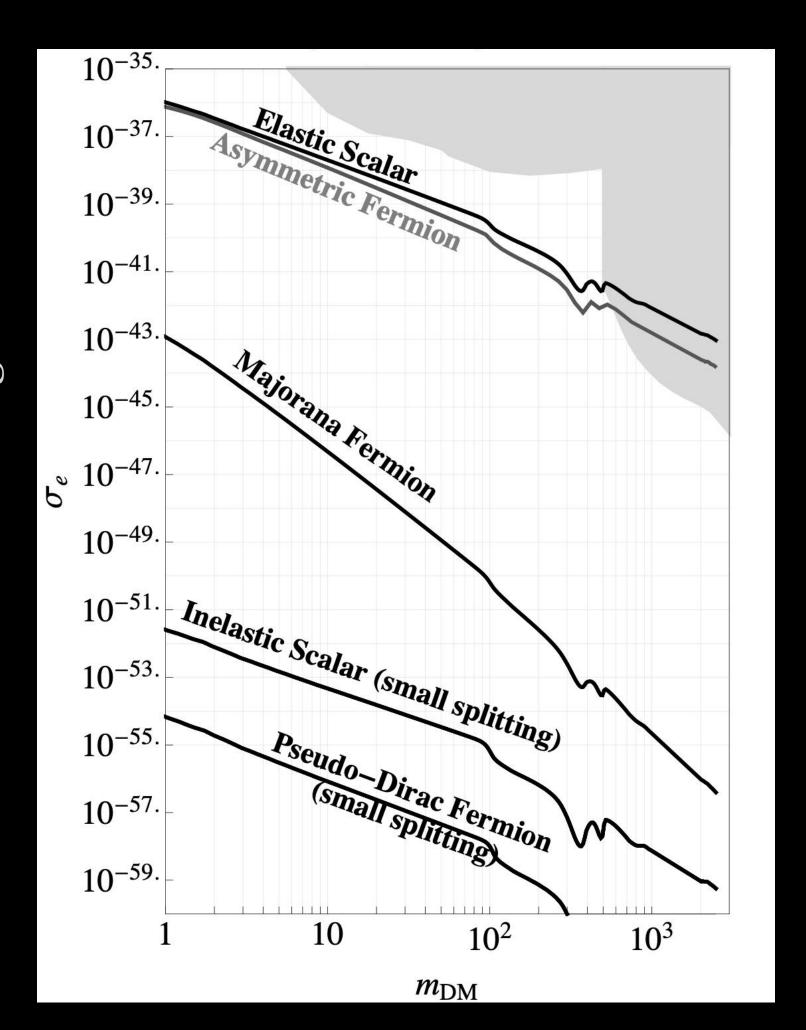




XMQT

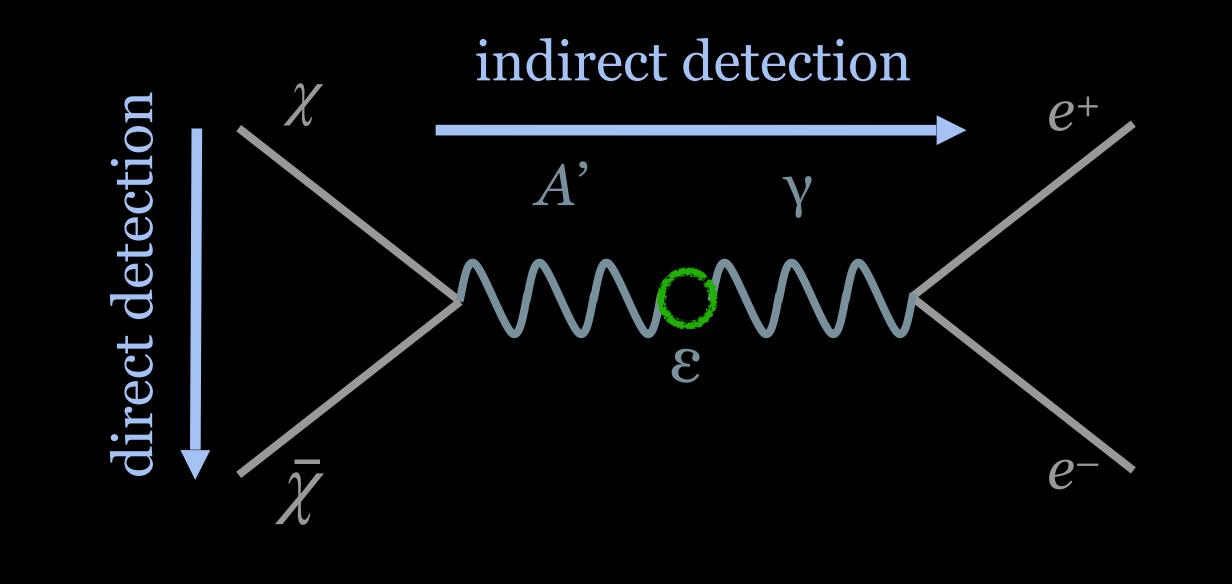
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 $\sim 10 \mathrm{s} \,\mathrm{TeV}$  $\Omega_{\chi} > \Omega_{\rm DM}$  section **CTOSS** scattering DM-electron



How can we probe this open parameter space?

"Simplest" possible dark sector extension is U(1) dark QED + kinetic mixing ( $\varepsilon$ ) = a feeble interaction with SM matter 



 $\langle \sigma v \rangle \propto \varepsilon^2 \alpha_D$ 



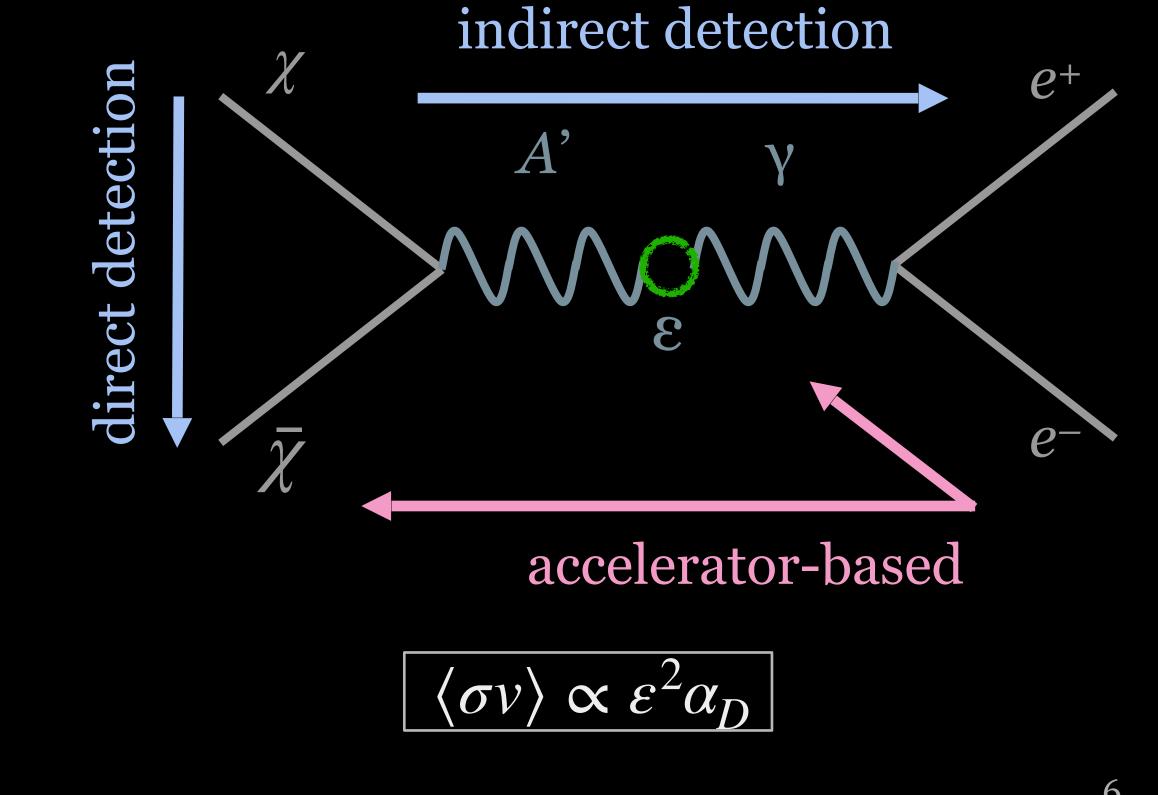
XMC

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LK

XMC

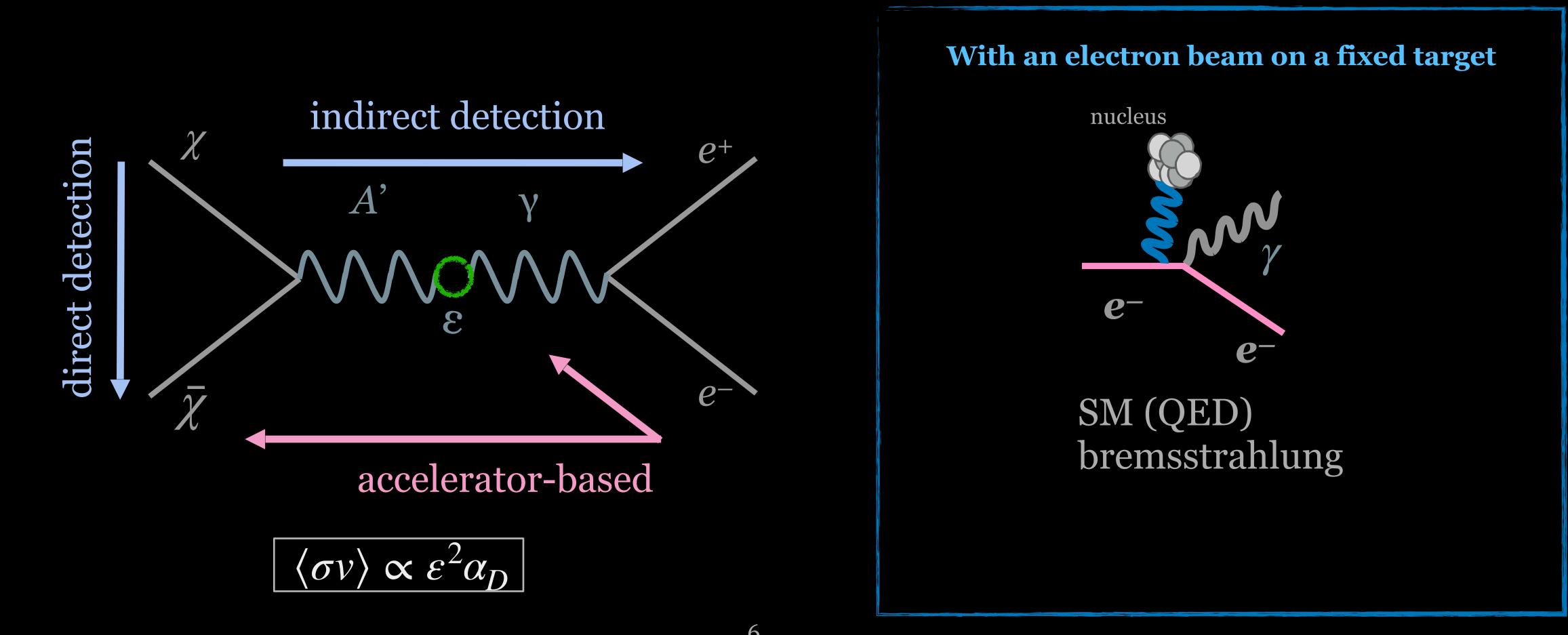


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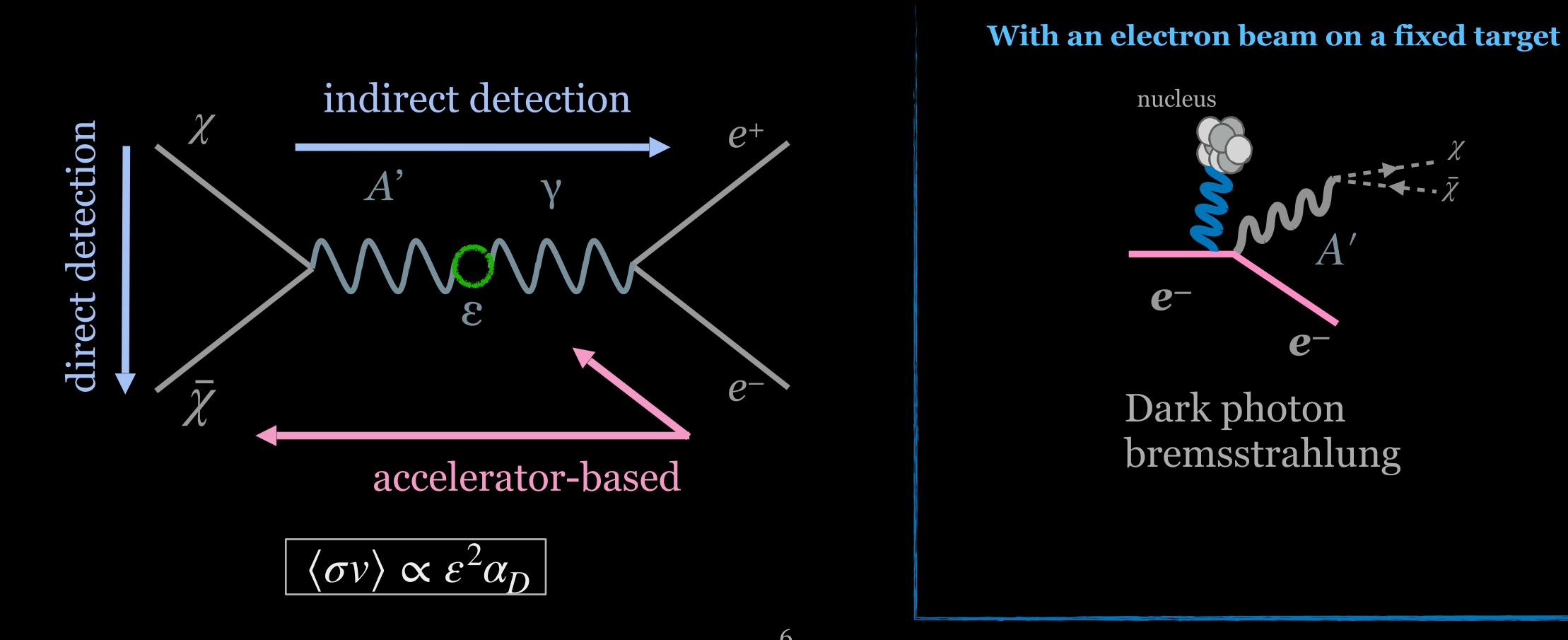


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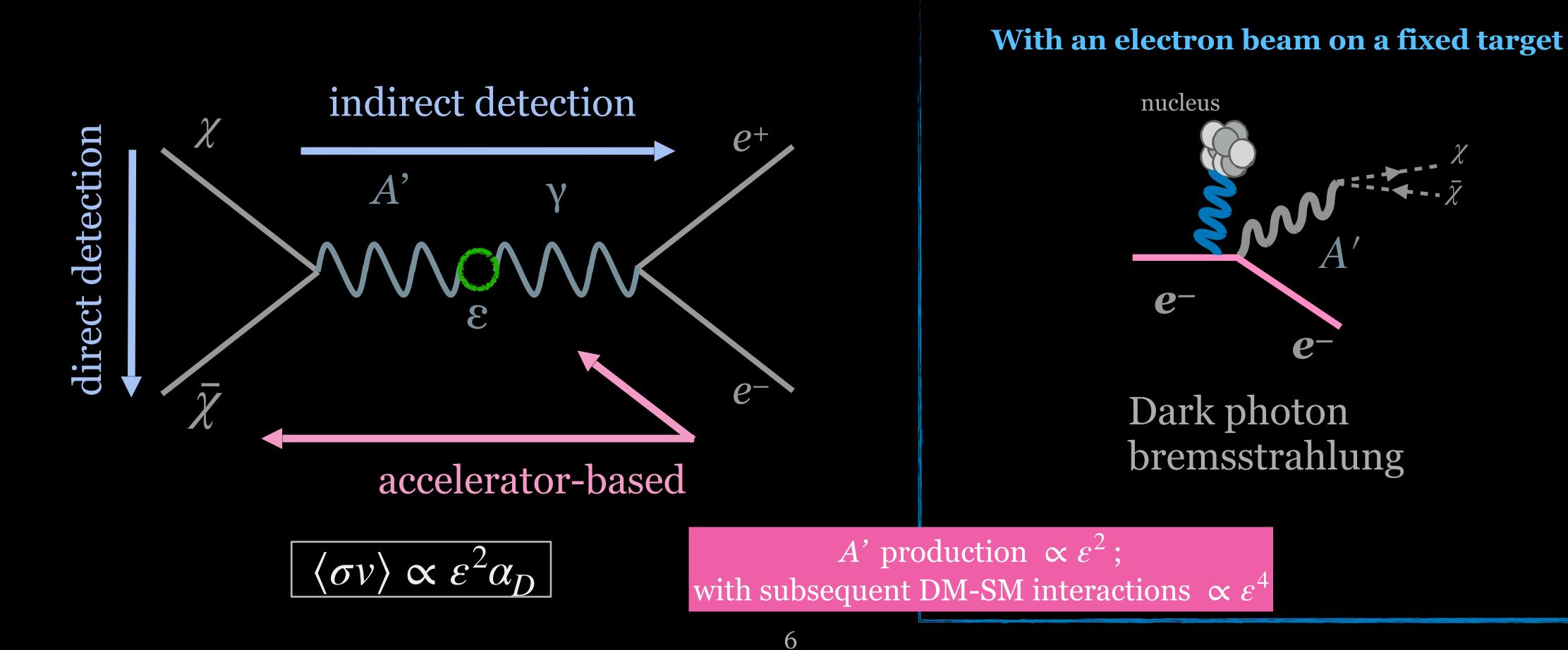




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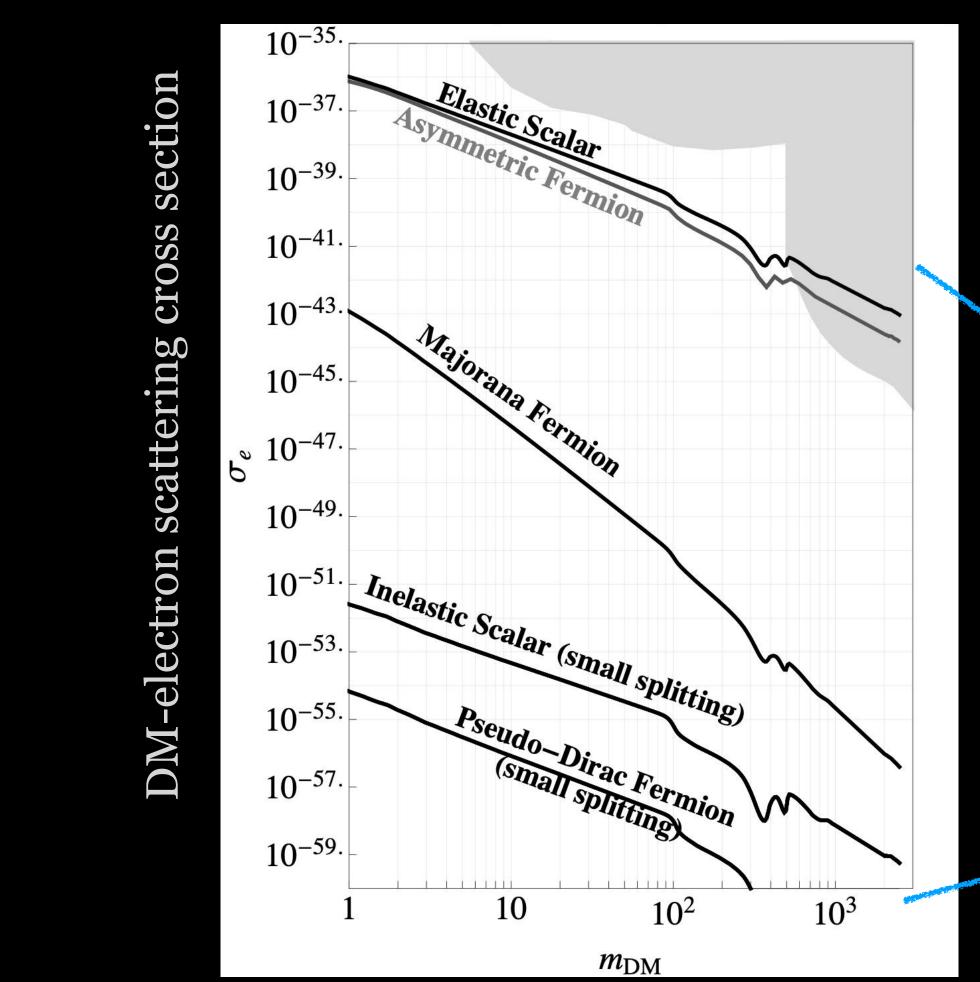


# Accelerator $\rightarrow$ relativistic DM production

7

Reduces sensitivity to spin and velocity suppression effects

sub-GeV thermal prediction targets line up within reach at accelerators 

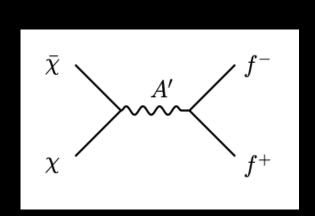


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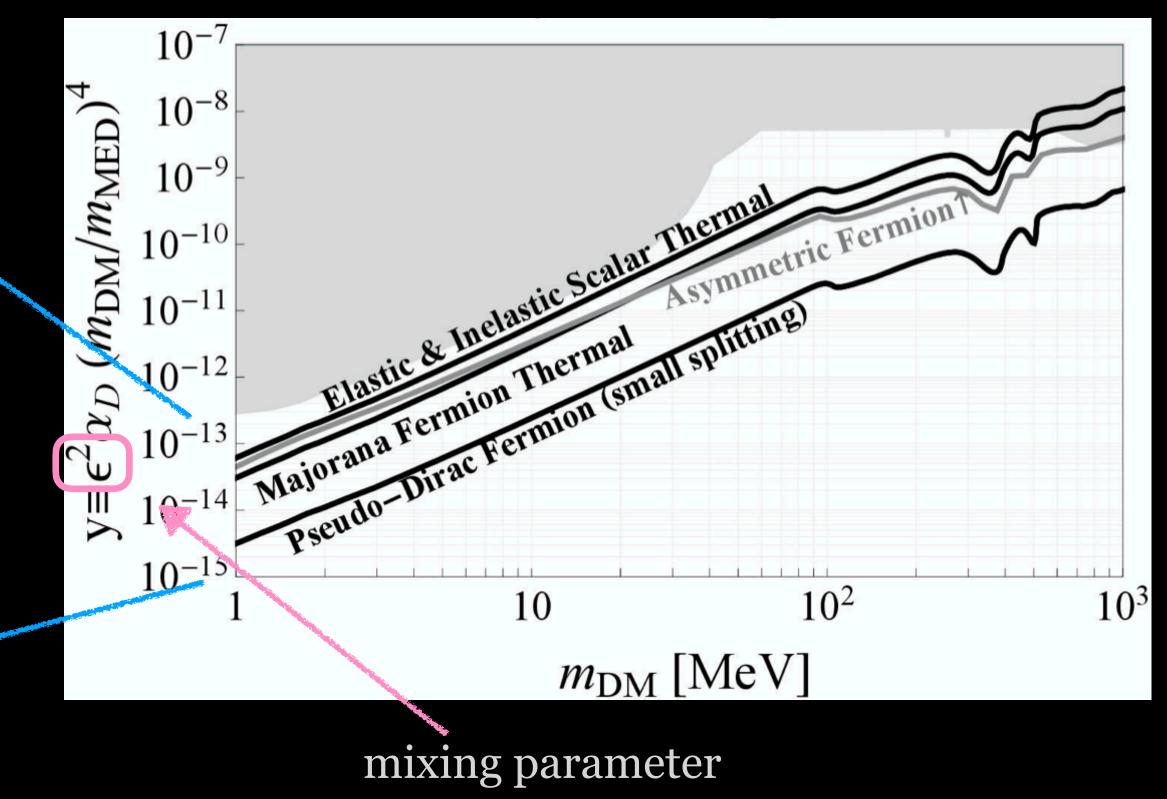
LK

XMQT

https://arxiv.org/abs/1808.05219



 $\sigma v(\chi \chi \to A' \to ff) \propto \varepsilon^2 \alpha_D \frac{m_{\chi}^2}{m_{\Lambda'}^4} = \frac{y}{m_{\Lambda'}^4}$ 



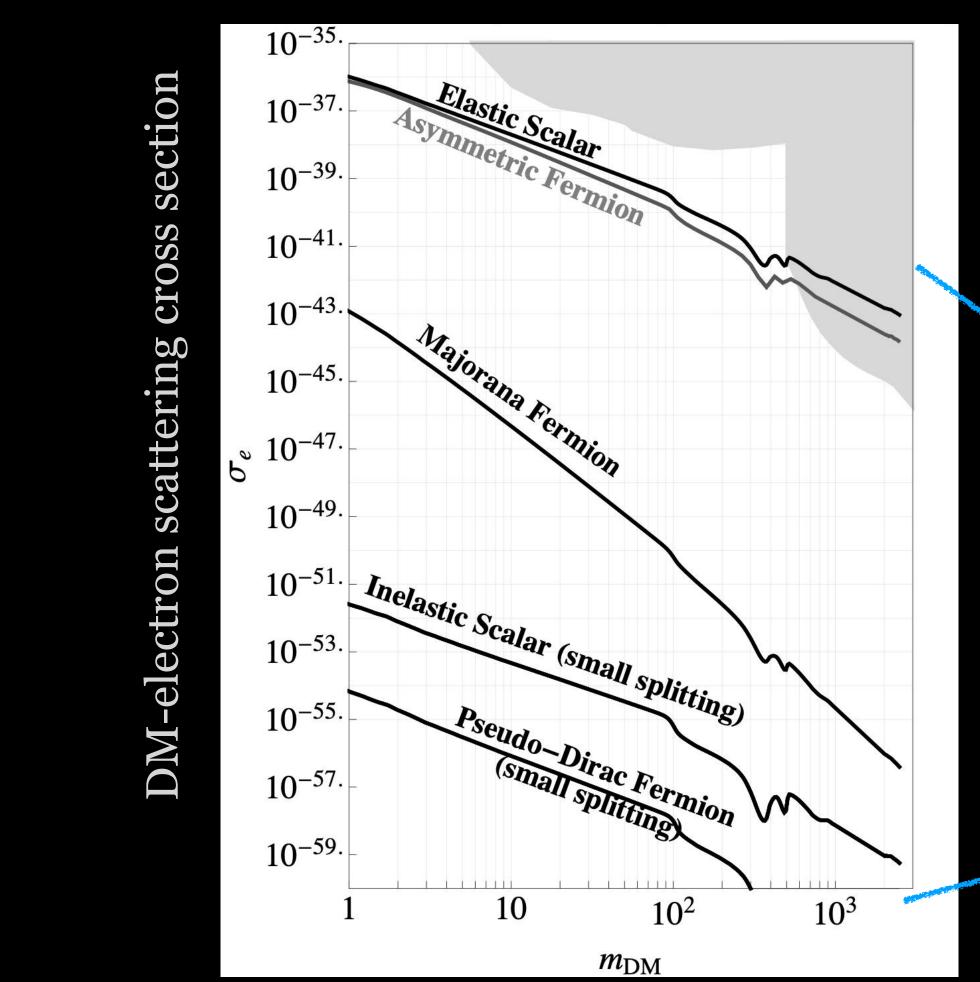


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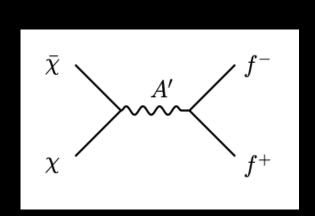


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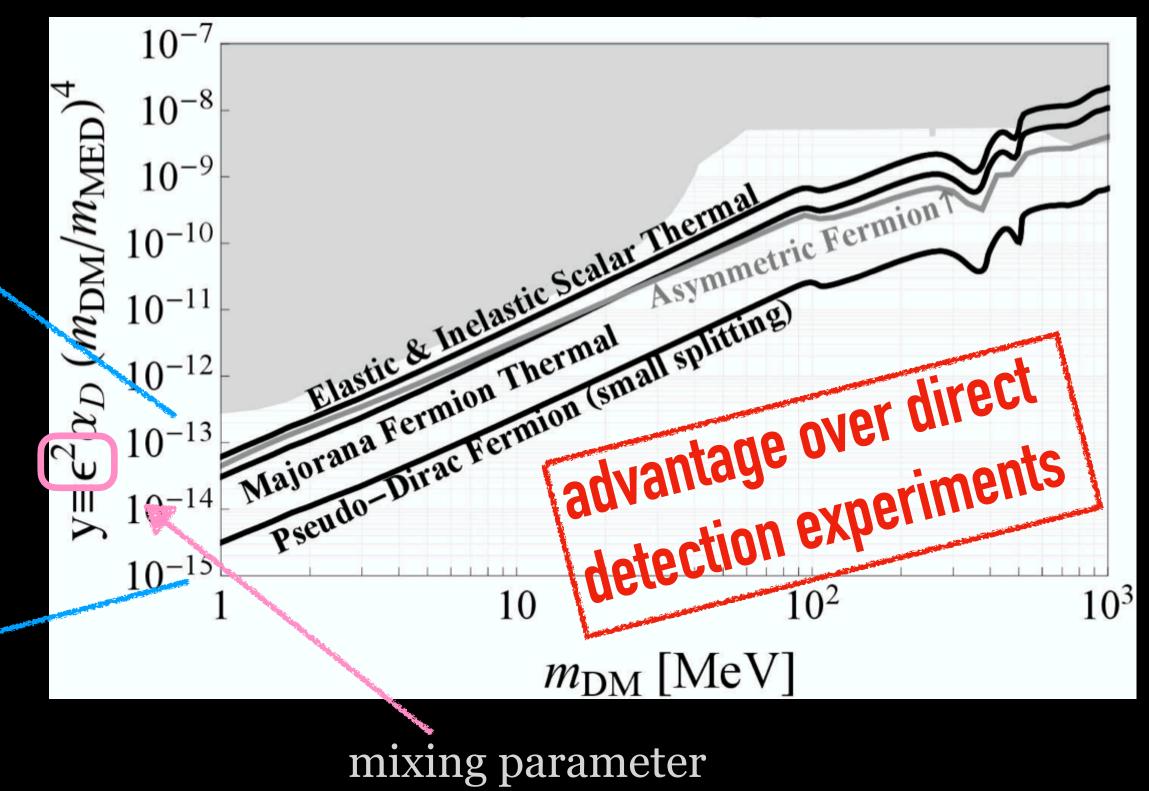
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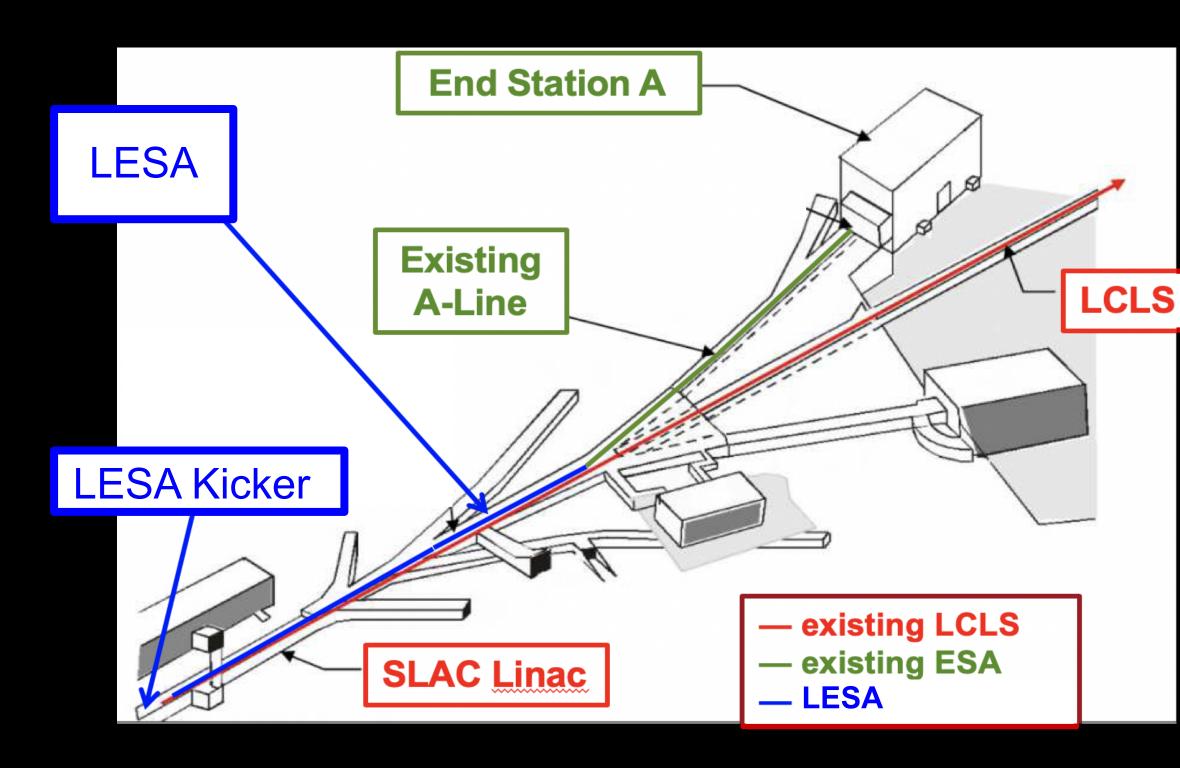


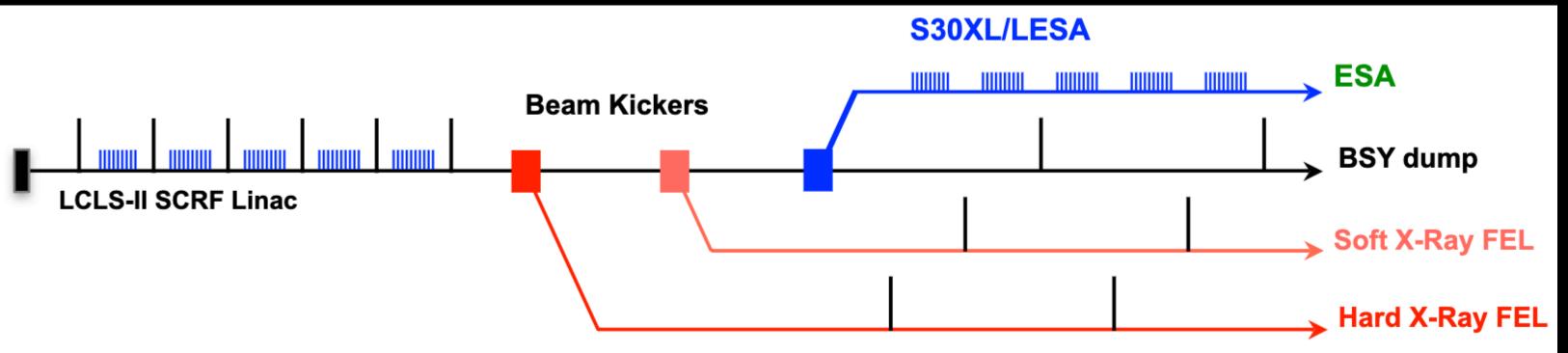
$$\sigma v(\chi \chi \to A' \to ff) \propto \varepsilon^2 \alpha_D \frac{m_{\chi}^2}{m_{A'}^4} =$$





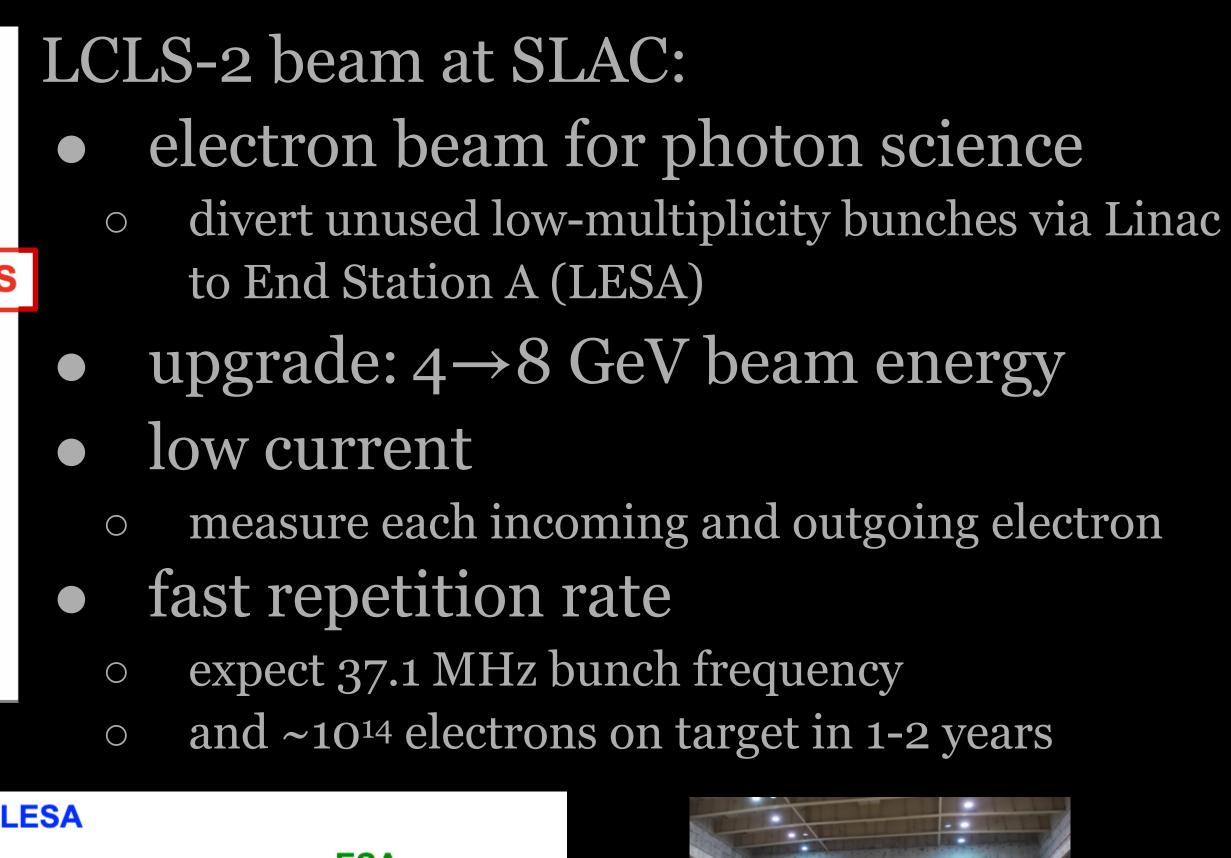
### GeV-scale electron beam at SLAC







XMQ





upstream

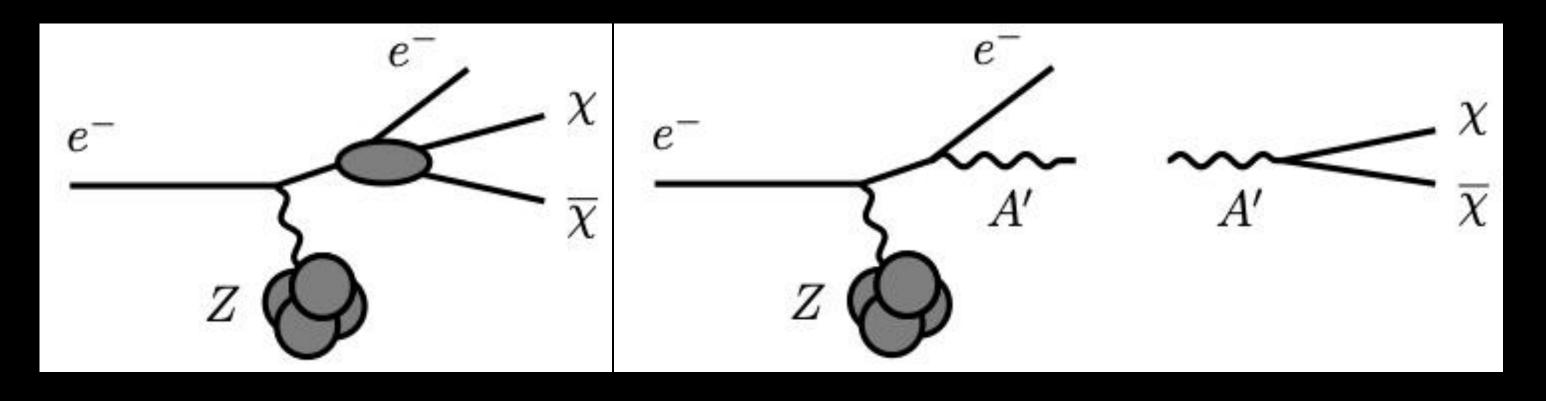






# LDMX: a fixed-target missing momentum experiment

#### LDMX focuses on *escaping* dark matter:



- massive dark photon (A') bremsstrahlung in thin target
- strategy: make all SM backgrounds appear in detector
  - veto everything but low-activity events  $\bigcirc$
- agnostic when it comes to (invisible) fate of the A'
  - notice that initial electron energy goes missing  $\bigcirc$

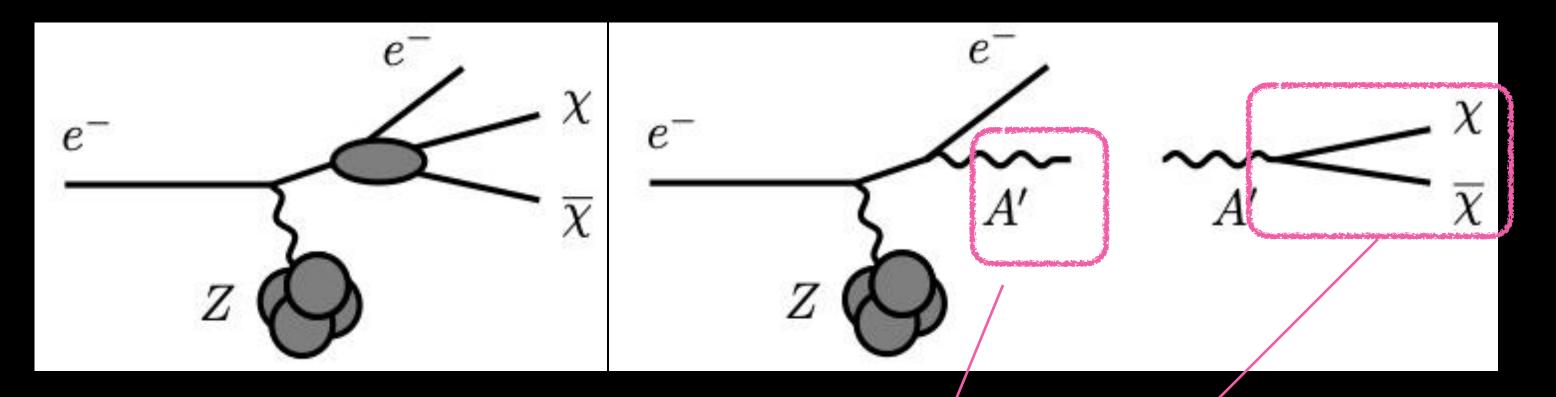
More detail: A high efficiency photon veto for the Light Dark Matter eXperiment, JHEP04 (2020) 003





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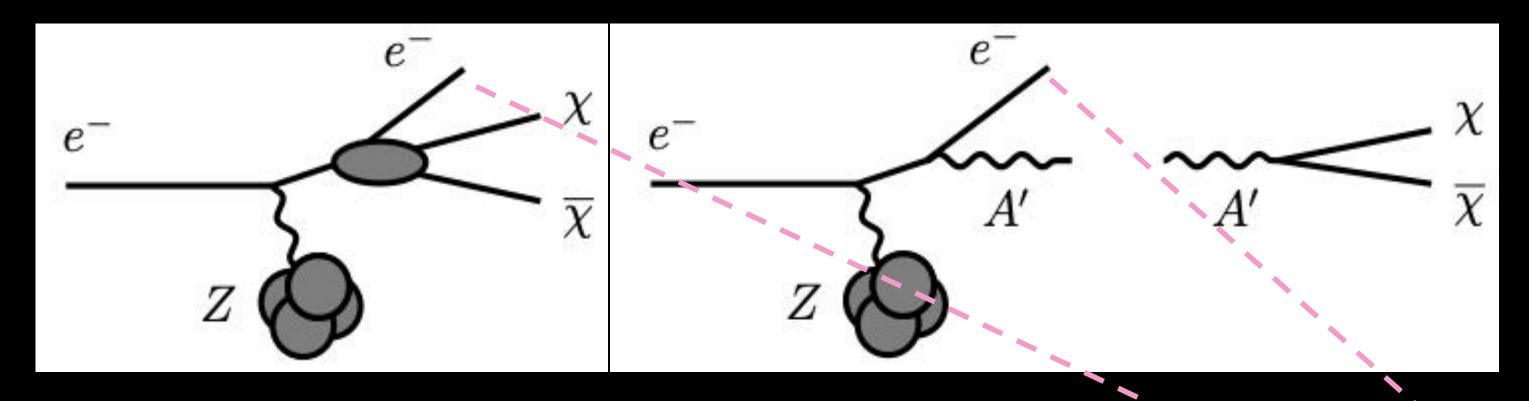
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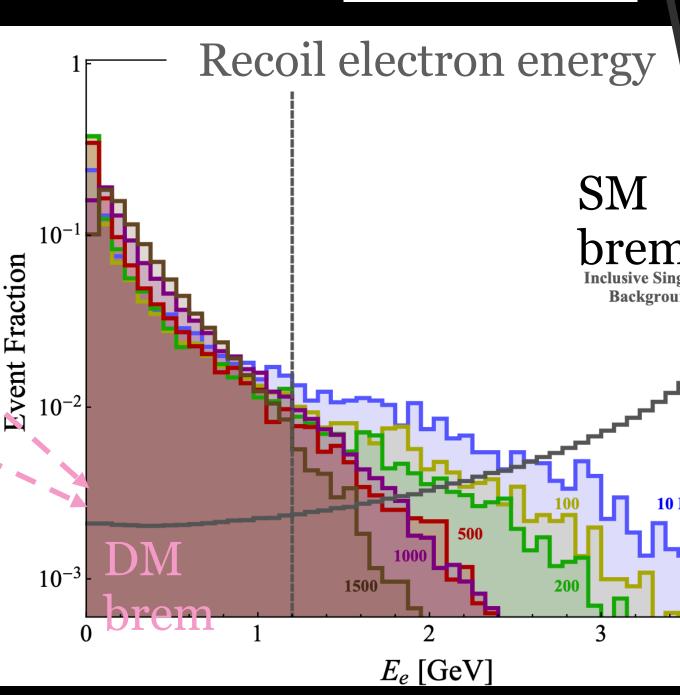
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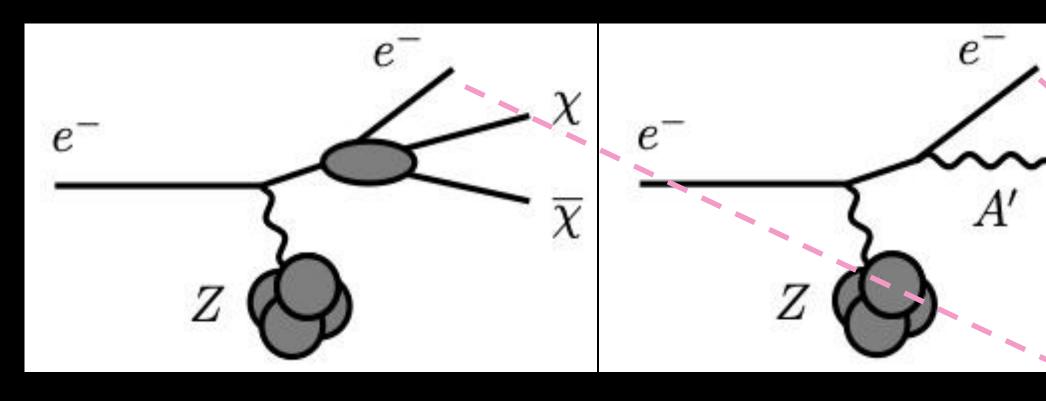






# LDMX: a fixed-target missing momentum experiment vs e<sup>4</sup> advantage

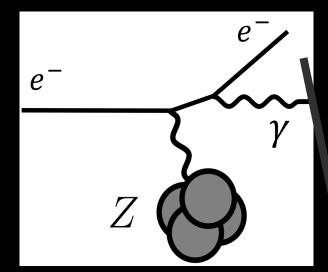
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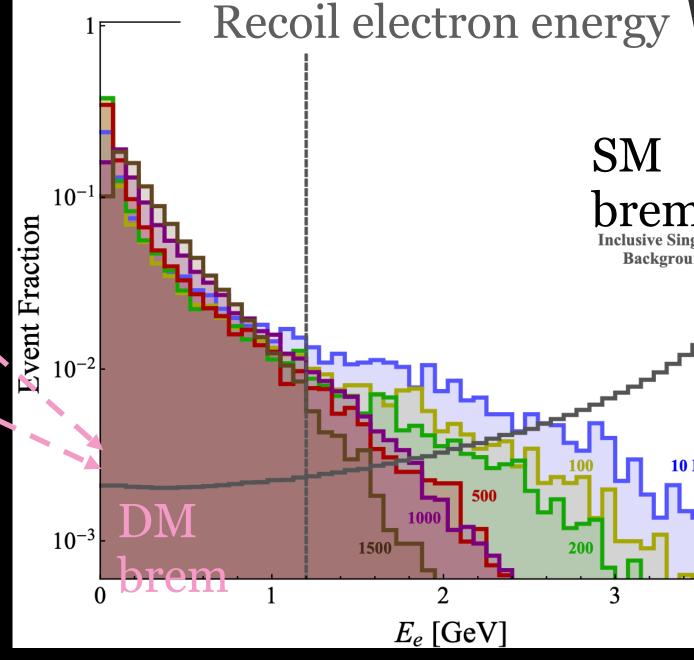
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accelerator-basel

appearance experi

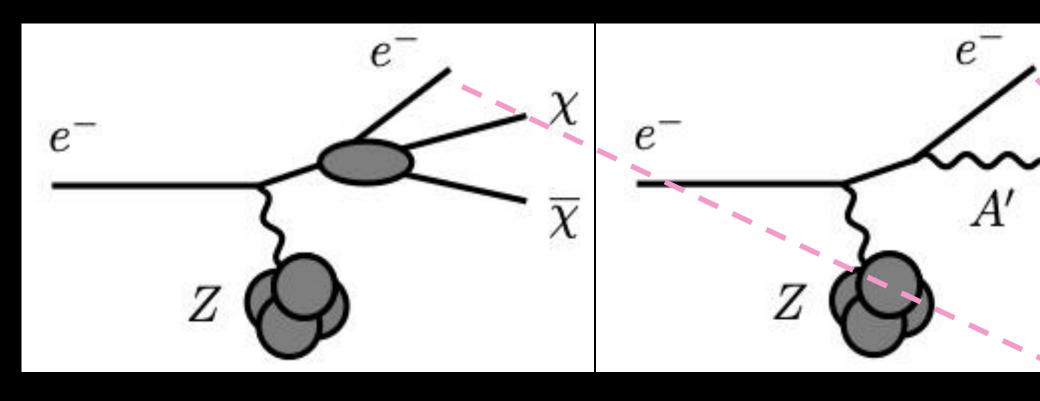






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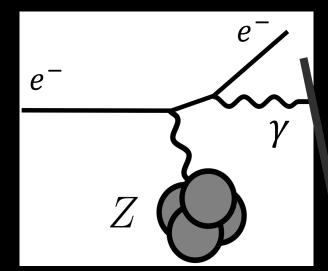
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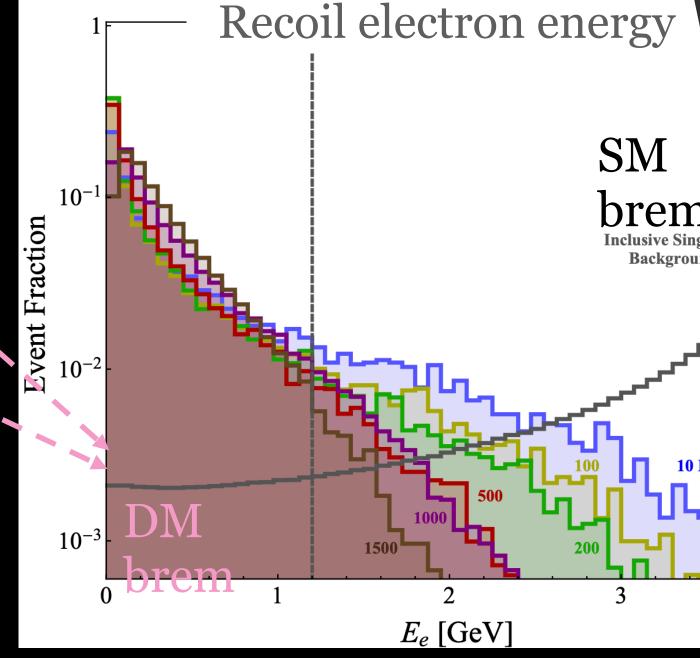
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XMG



accelerator-basel

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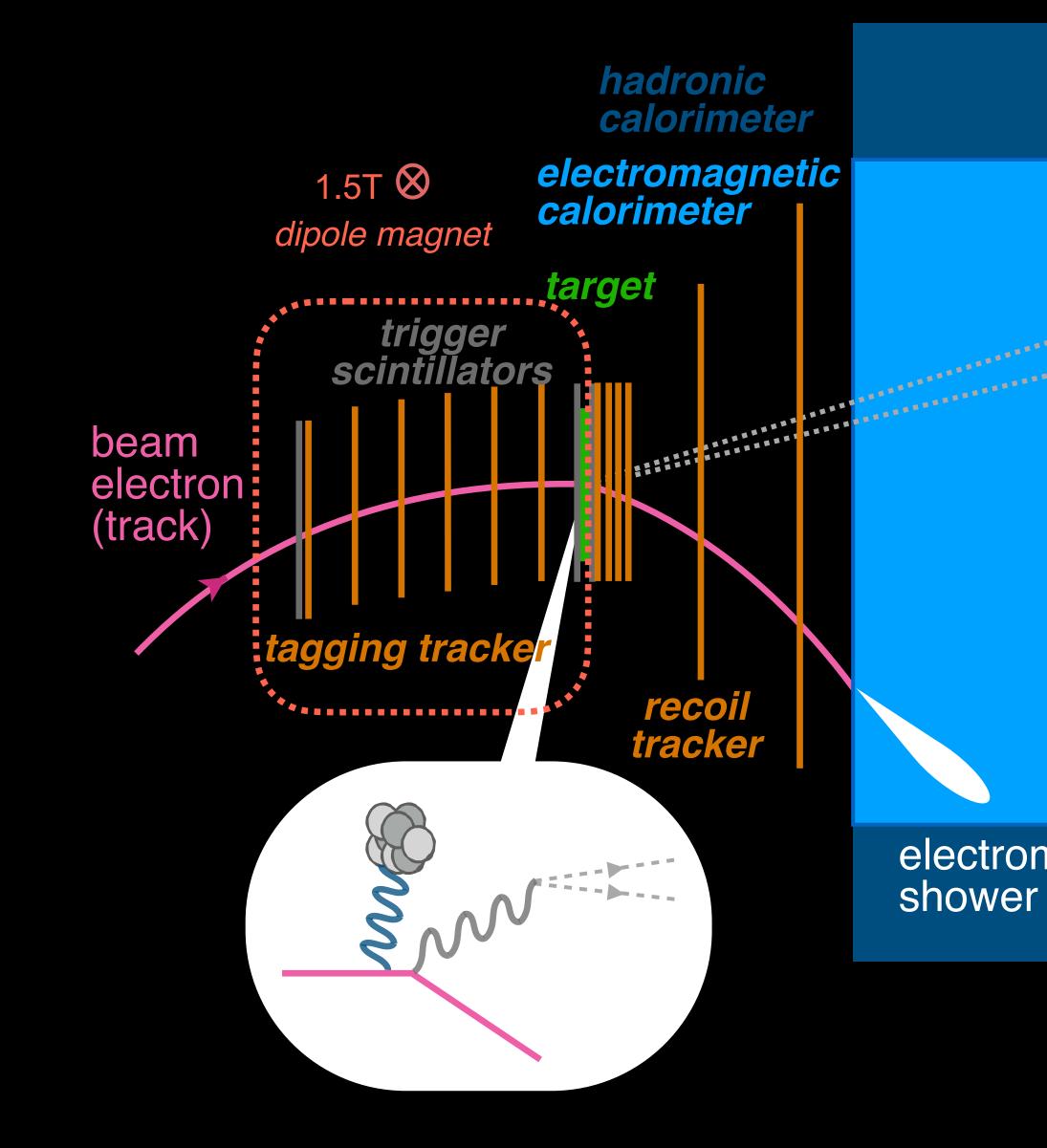


Massive photon  $\rightarrow$  momentum kick; missing momentum experiment





# LDMX detector concept



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Signal cartoon (not to scale)

escaping dark matter

\*

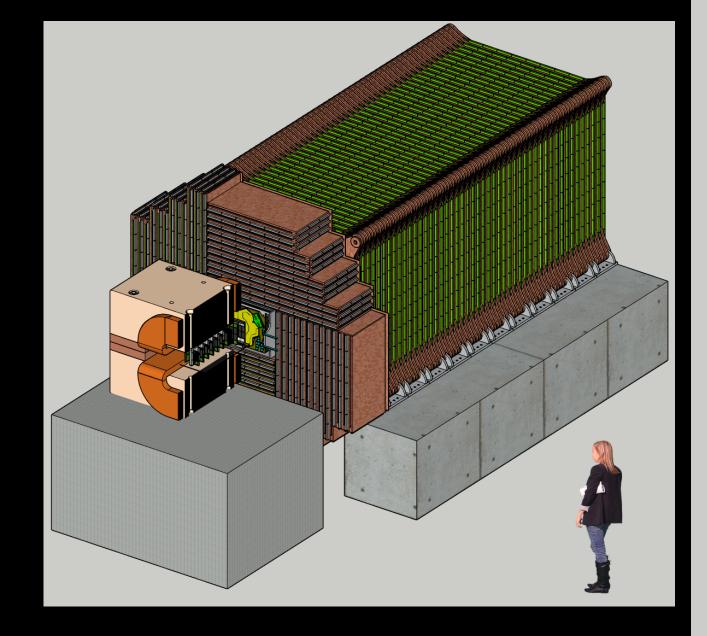
Measure (missing) energy in calorimeters, momentum in trackers

electromagnetic





# Current design in more detail

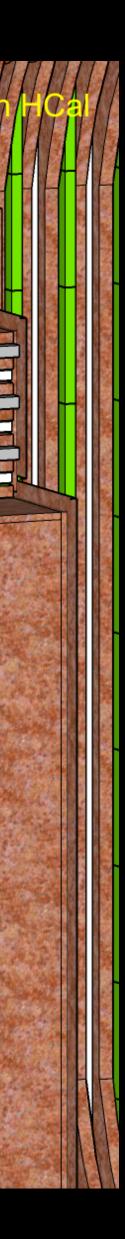


#### Leverage technologies from CMS, HPS, MINOS/Mu2e to minimize R&D

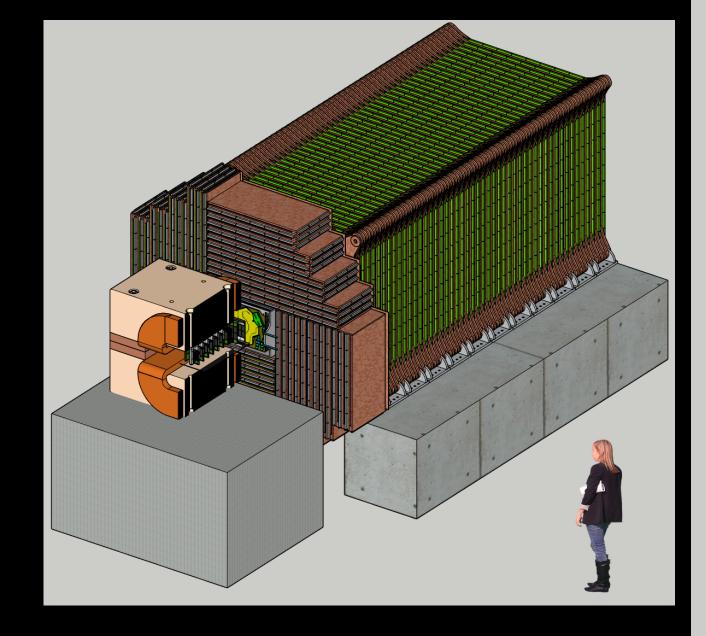








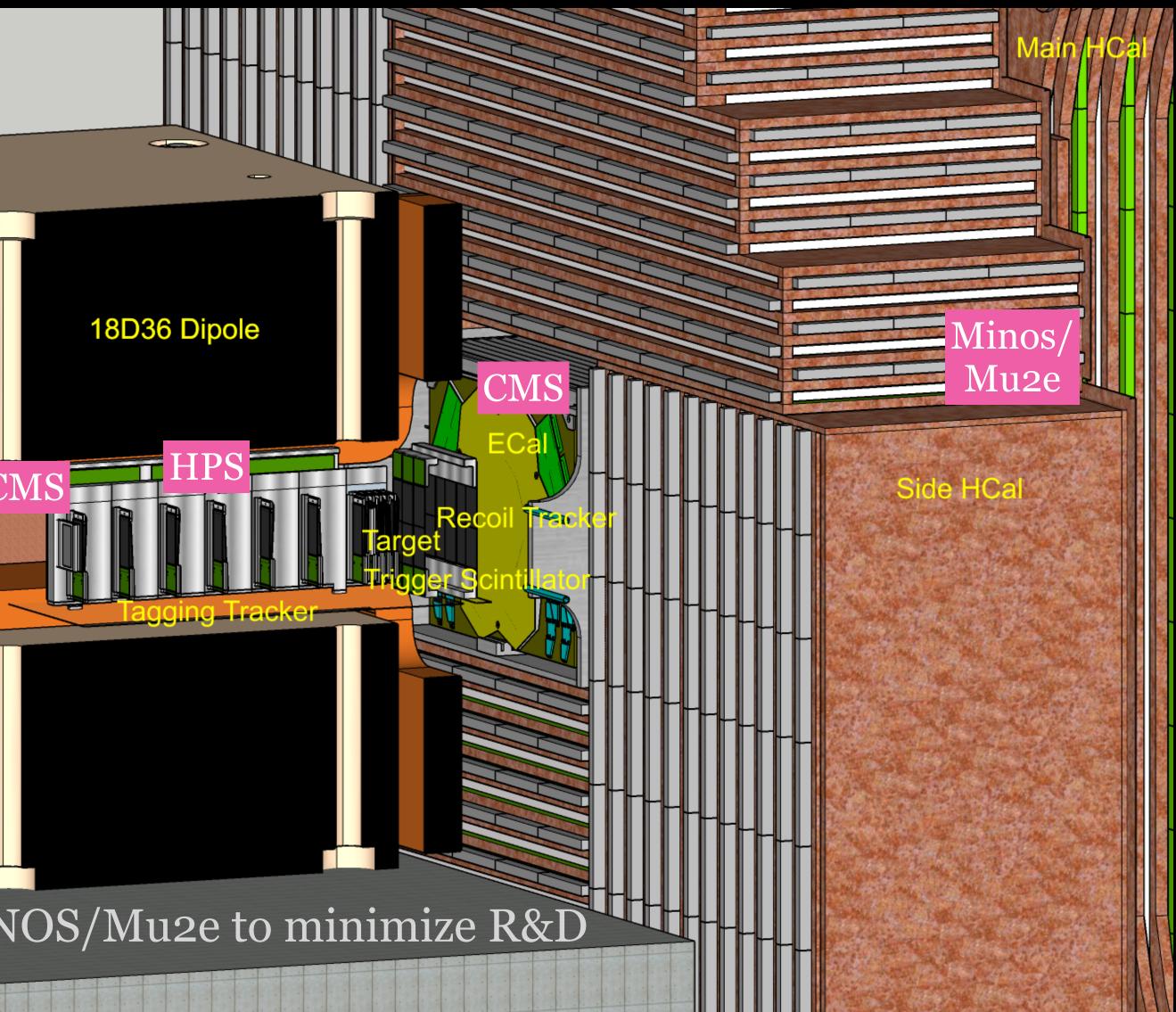
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XMC





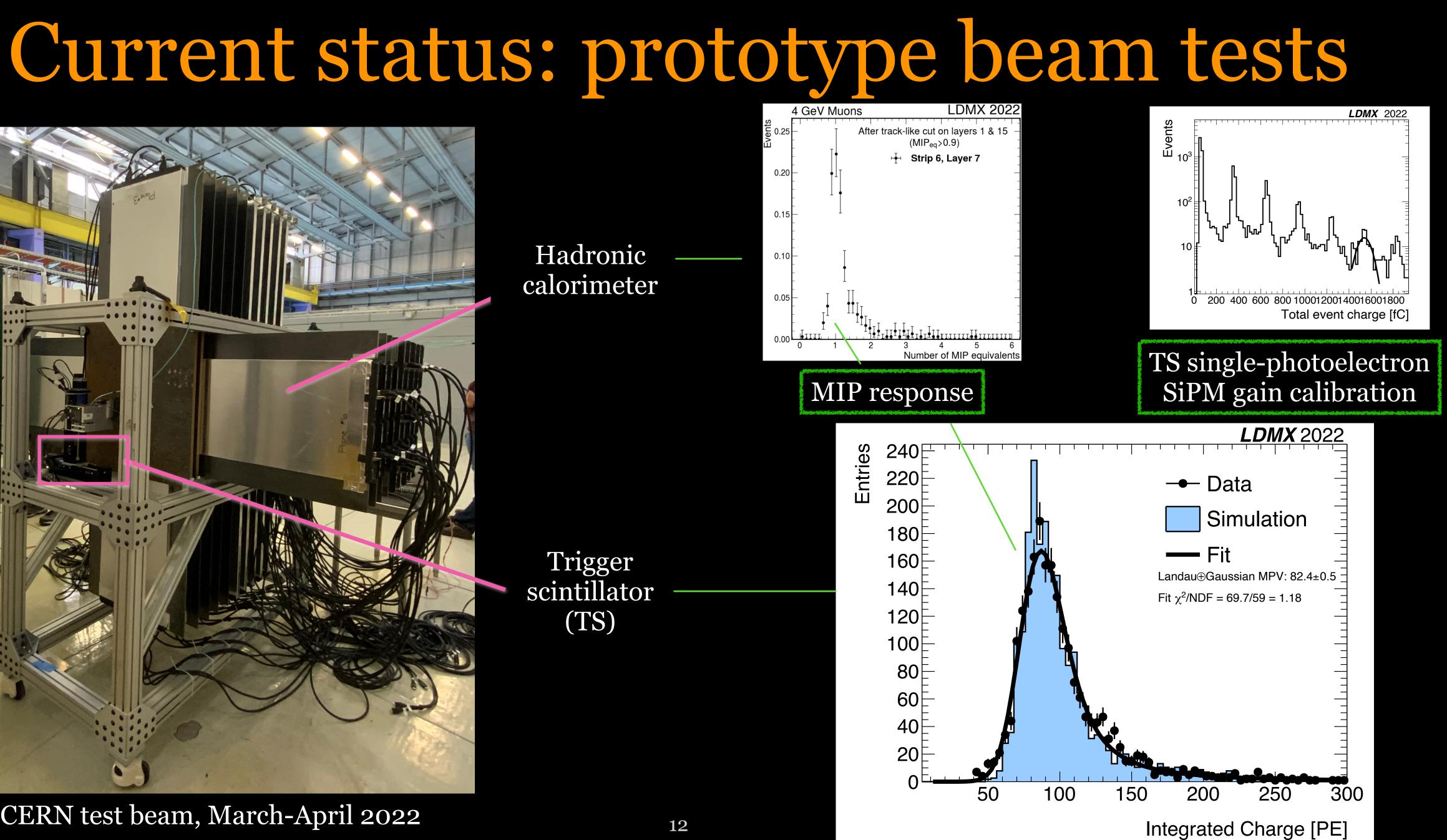
#### Hadronic calorimeter

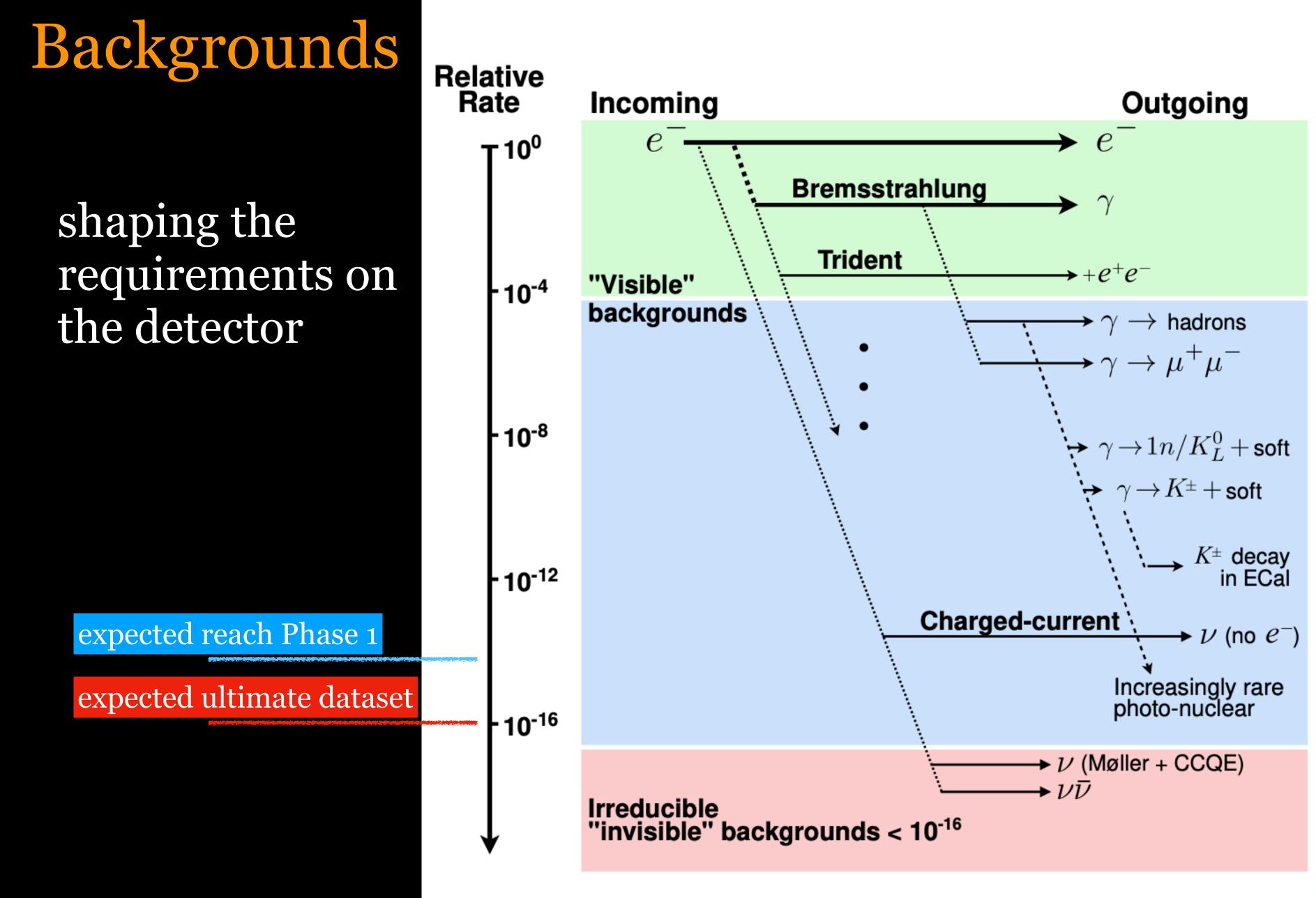
Trigger scintillator (TS)

CERN test beam, March-April 2022



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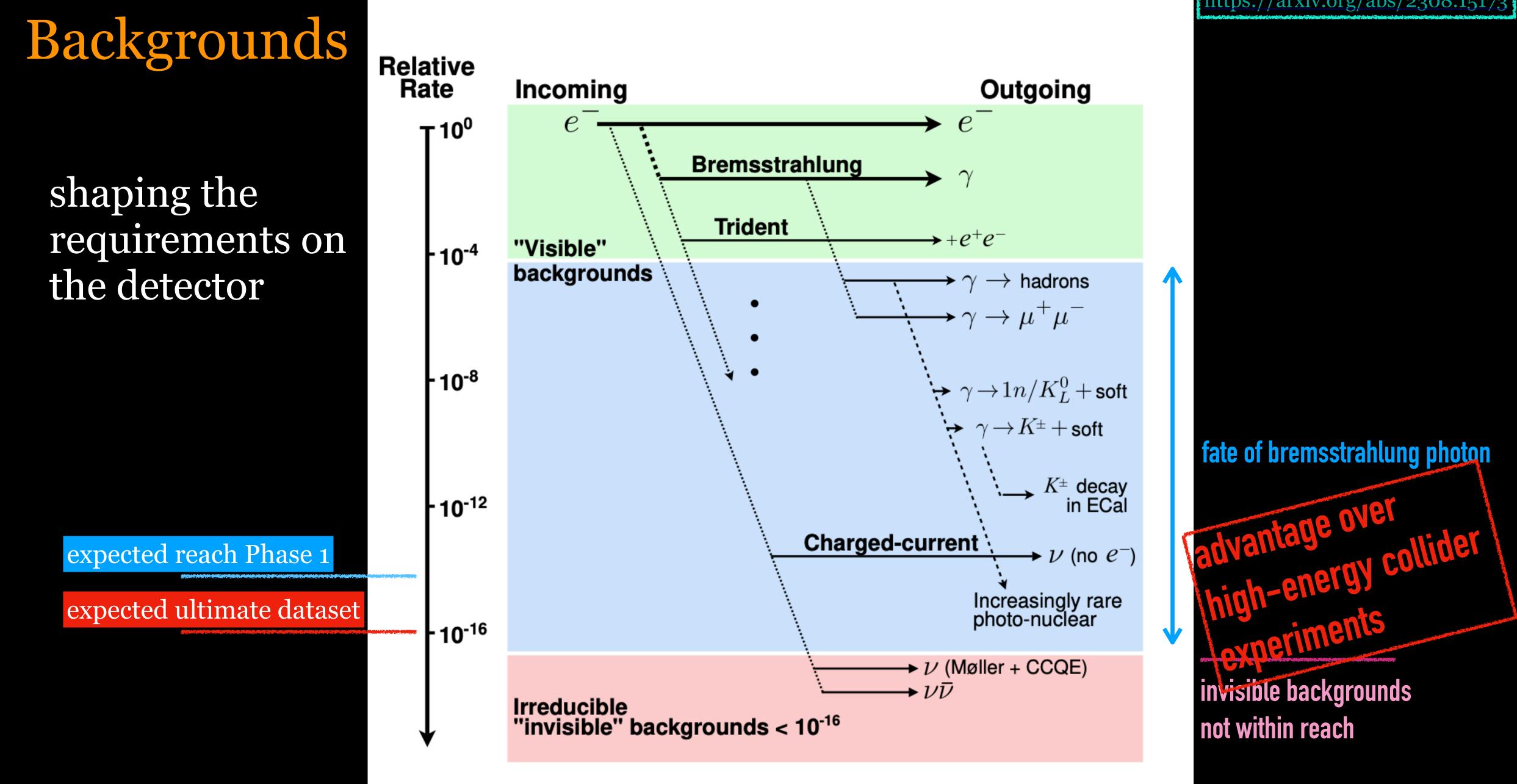
#### https://arxiv.org/abs/2308.15173

fate of bremsstrahlung photon

invisible backgrounds not within reach







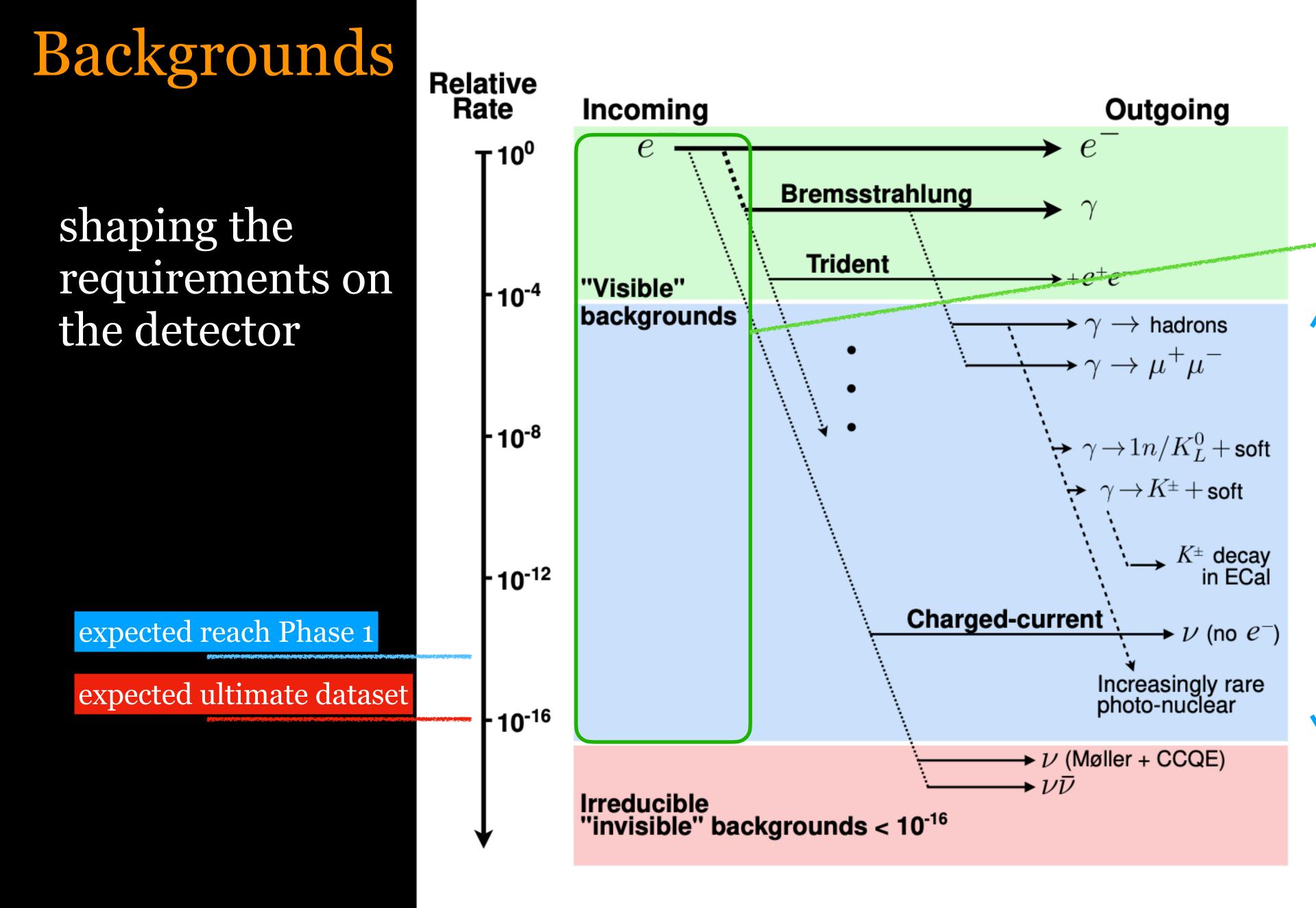
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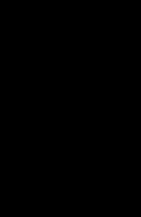
make visible with tracker and highgranularity calorimetry

fate of bremsstrahlung photon

advantage over advantage over tigh-energy collider ments

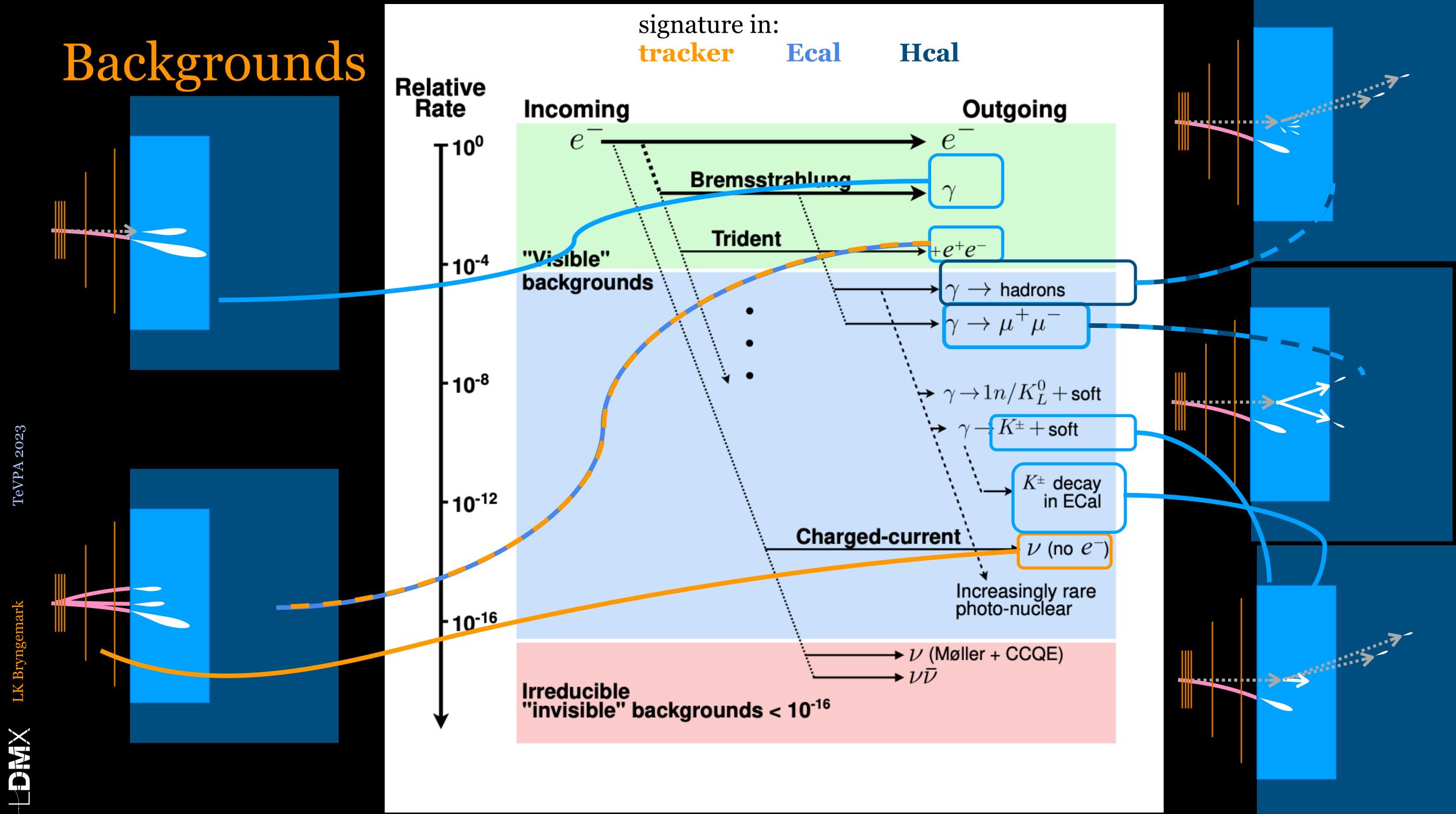
invisible backgrounds not within reach











# Backgrounds

#### Relative Rate

#### Incoming

\*\*\*\*\*

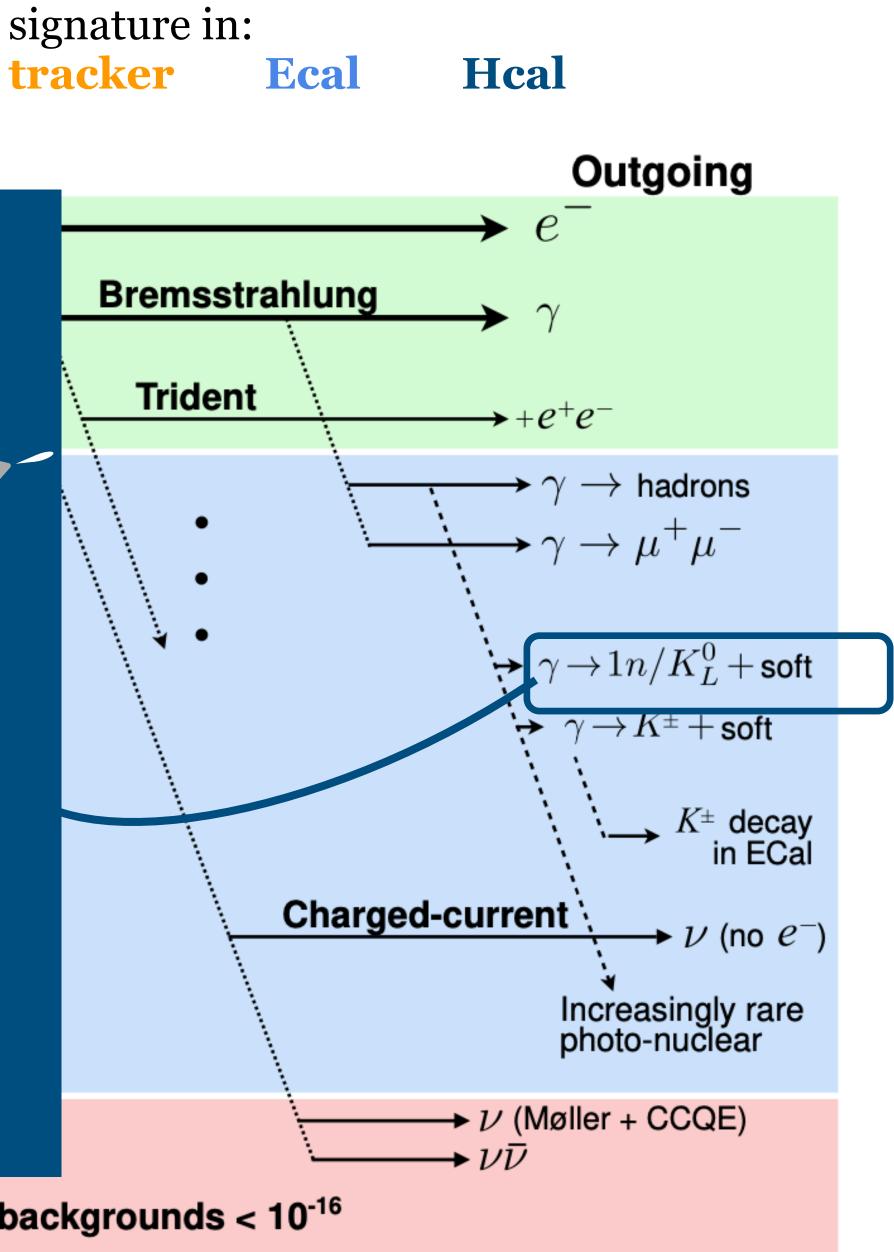
neutral hadrons

Irreducible "invisible" backgrounds < 10<sup>-16</sup>

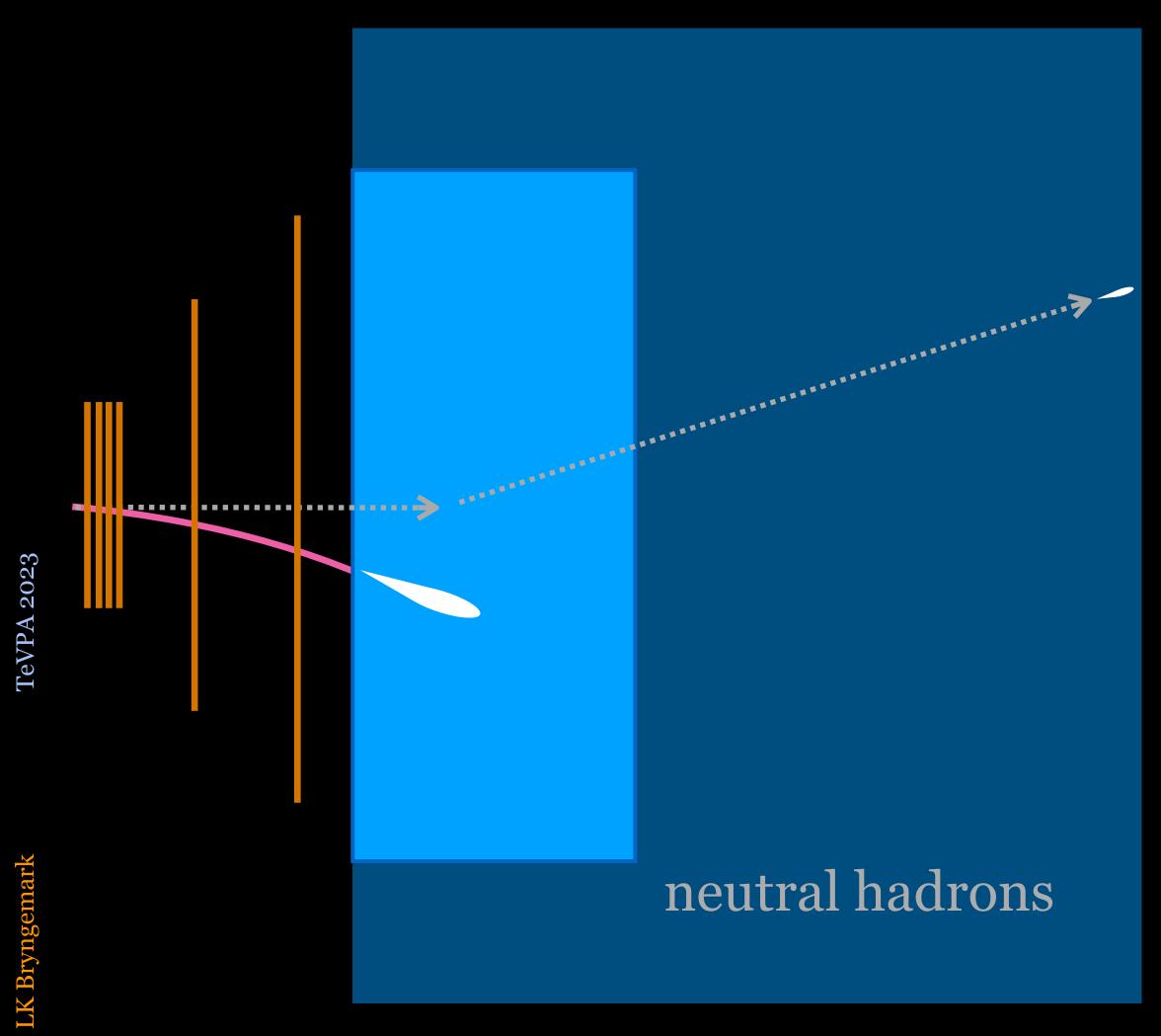
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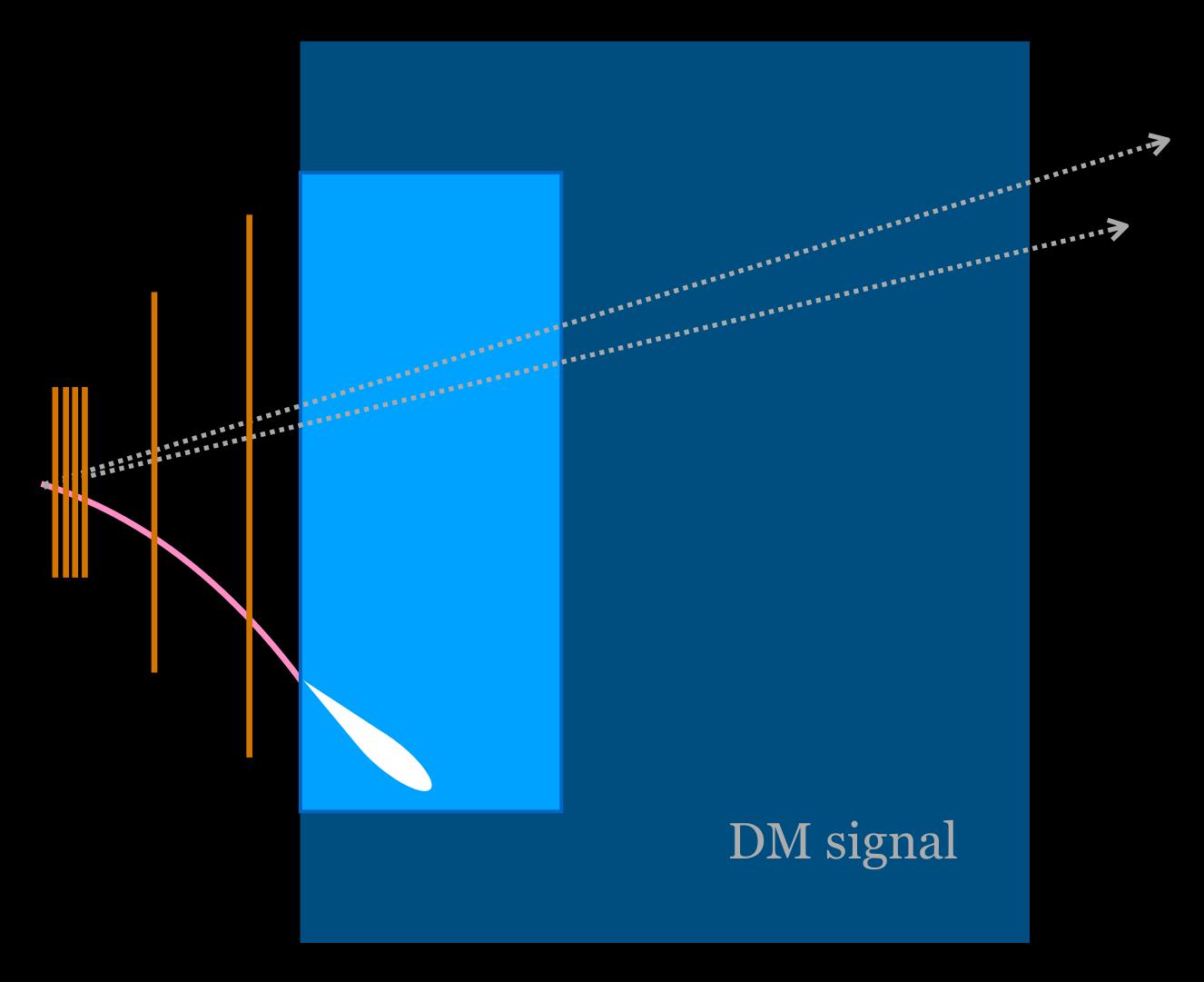
# Backgrounds



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yngemark

XMG





# Goal: no background from ~10<sup>14</sup> electrons on target

# First handle: recoiling electron

- Count incoming electrons, measure in Ecal, calculate missing energy
  - trigger if consistent with one electron losing significant energy
- veto if number of outgoing tracks > incoming tracks
- veto if no soft recoil electron track

# Exploit high-granularity Ecal features

- BDT trained to reject photonuclear events in Ecal
- MIP tracking powerful on sparse events

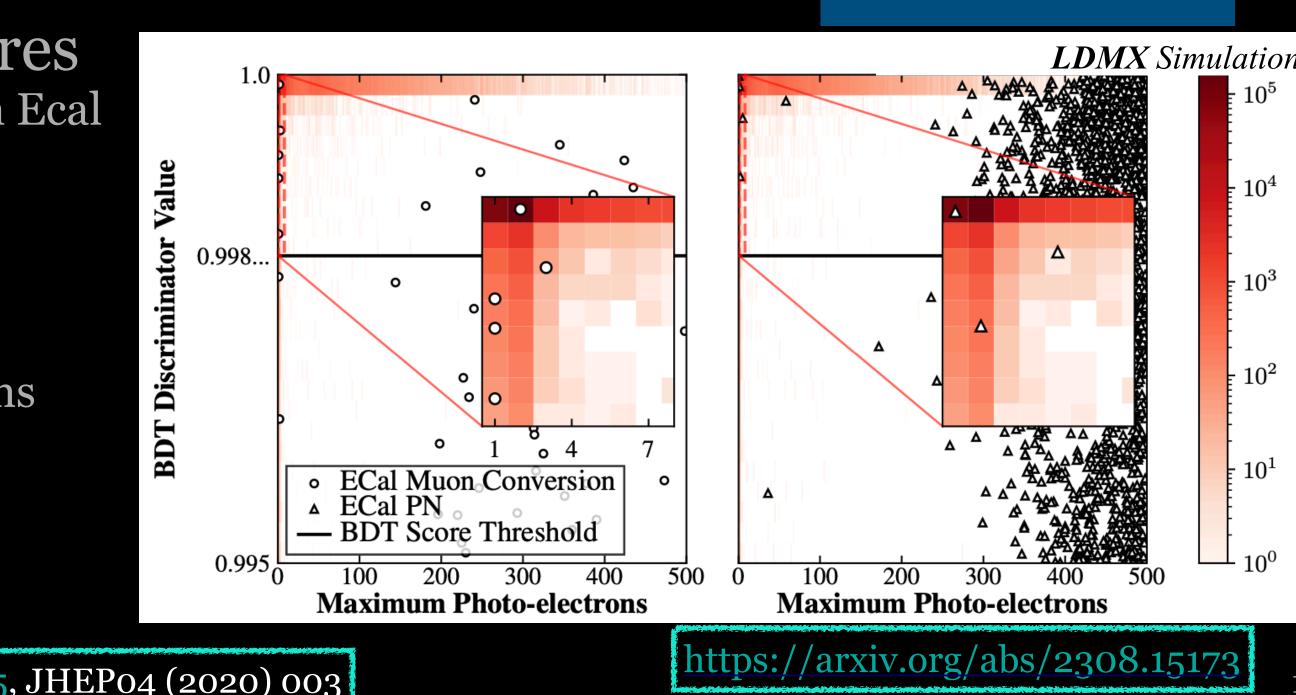
# Veto on Hcal activity

- allow no activity above detector noise
- deep enough to tease out even single neutrons

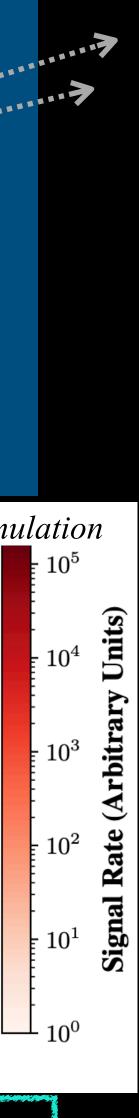
LK

extensively studied for 4 GeV beam in <u>https://arxiv.org/abs/1912.05535</u>, JHEP04 (2020) 003

calculate missing energy nificant energy g tracks



THE



16

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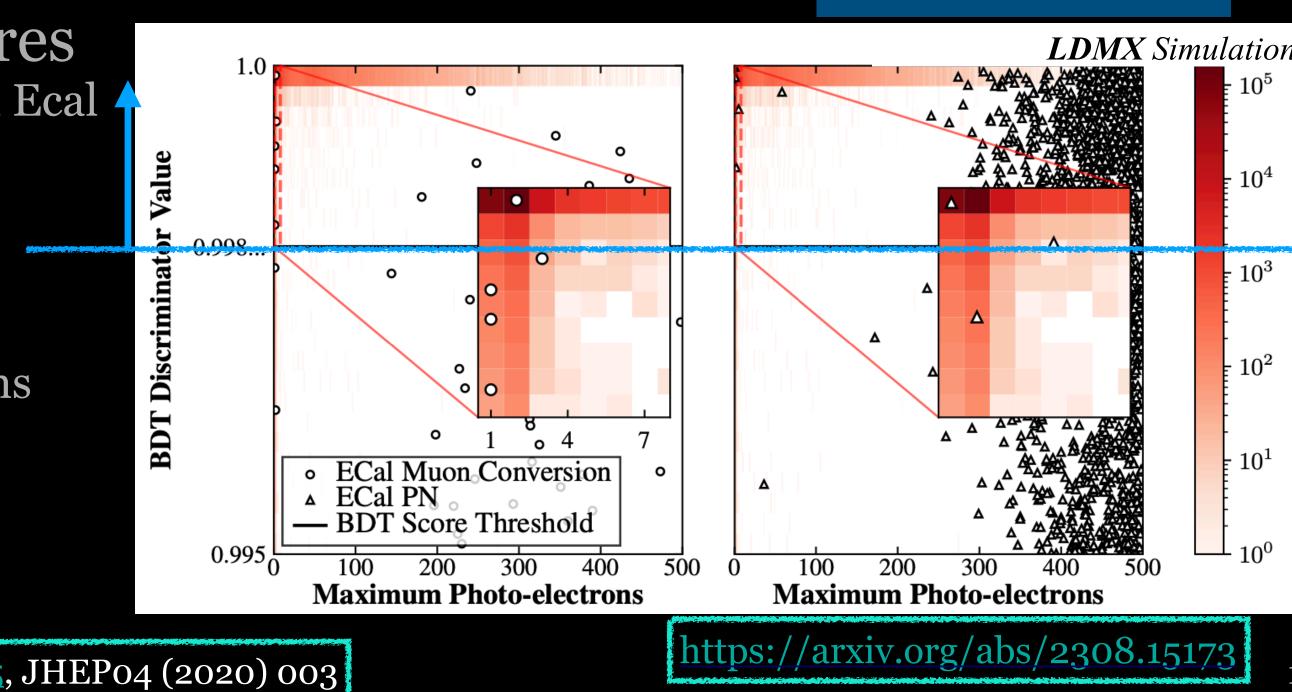
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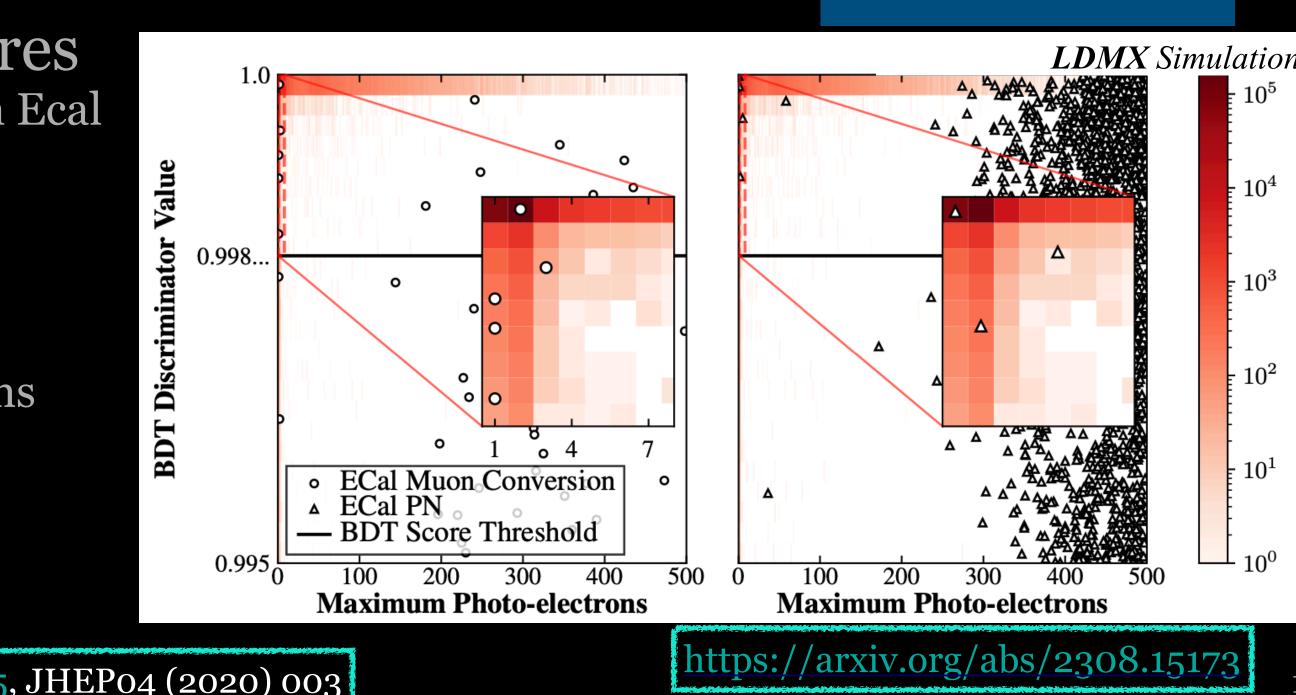
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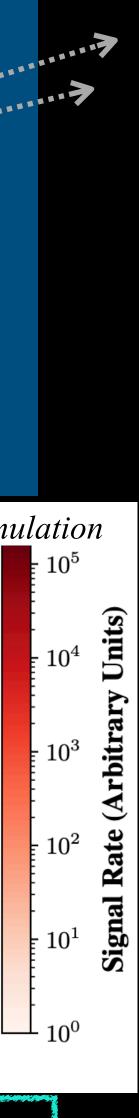
LK

extensively studied for 4 GeV beam in <u>https://arxiv.org/abs/1912.05535</u>, JHEP04 (2020) 003

calculate missing energy nificant energy g tracks



THE



16

# Goal: no background from ~10<sup>14</sup> electrons on target

# First handle: recoiling electron

- Count incoming electrons, measure in Ecal, calculate missing energy
  - trigger if consistent with one electron losing significant energy
- veto if number of outgoing tracks > incoming tracks
- veto if no soft recoil electron track

# Exploit high-granularity Ecal features

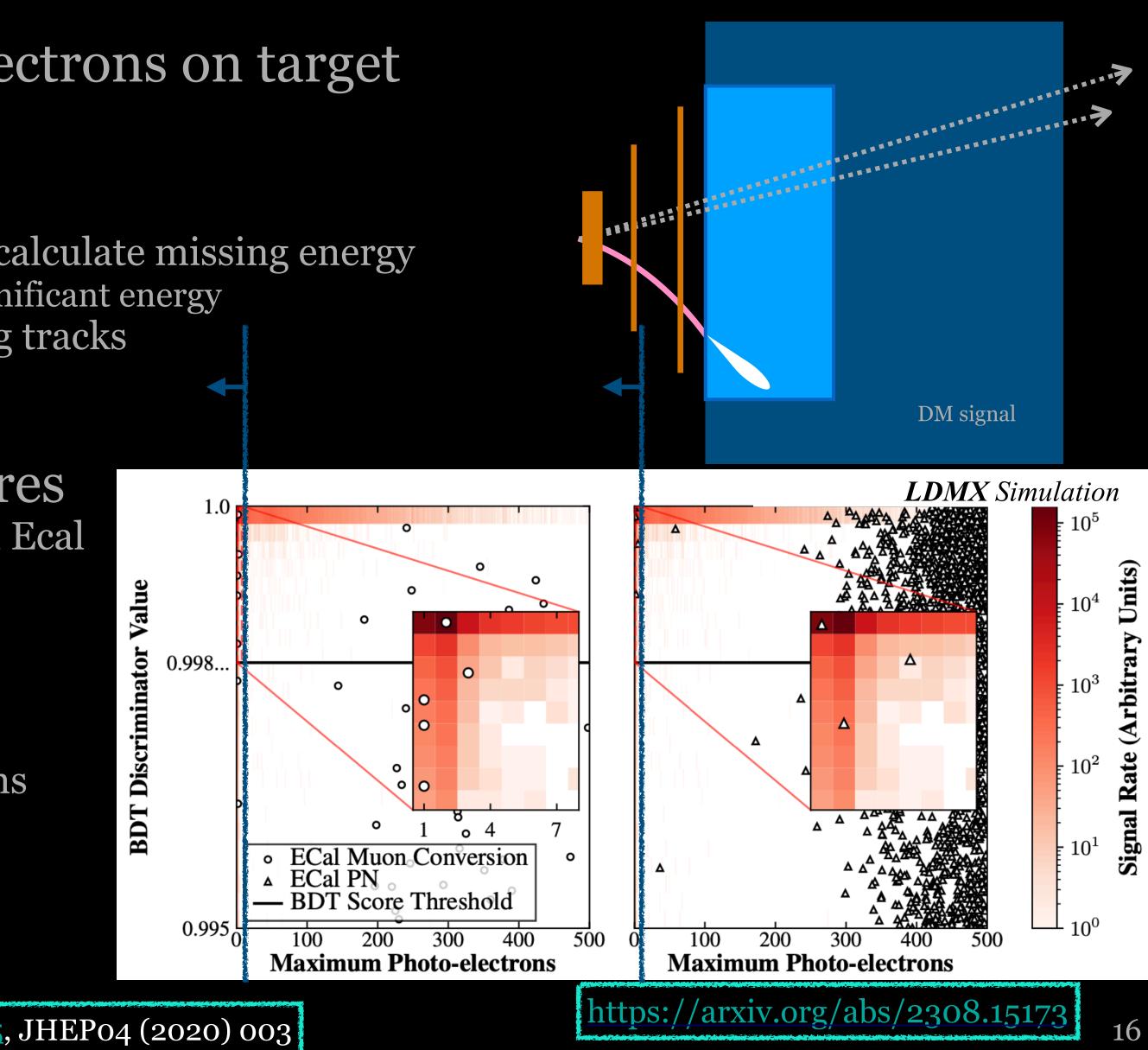
- BDT trained to reject photonuclear events in Ecal
- MIP tracking powerful on sparse events

# Veto on Hcal activity

- allow no activity above detector noise
- deep enough to tease out even single neutrons

LK

extensively studied for 4 GeV beam in <u>https://arxiv.org/abs/1912.05535</u>, JHEP04 (2020) 003



# Goal: no background from ~10<sup>14</sup> electrons on target

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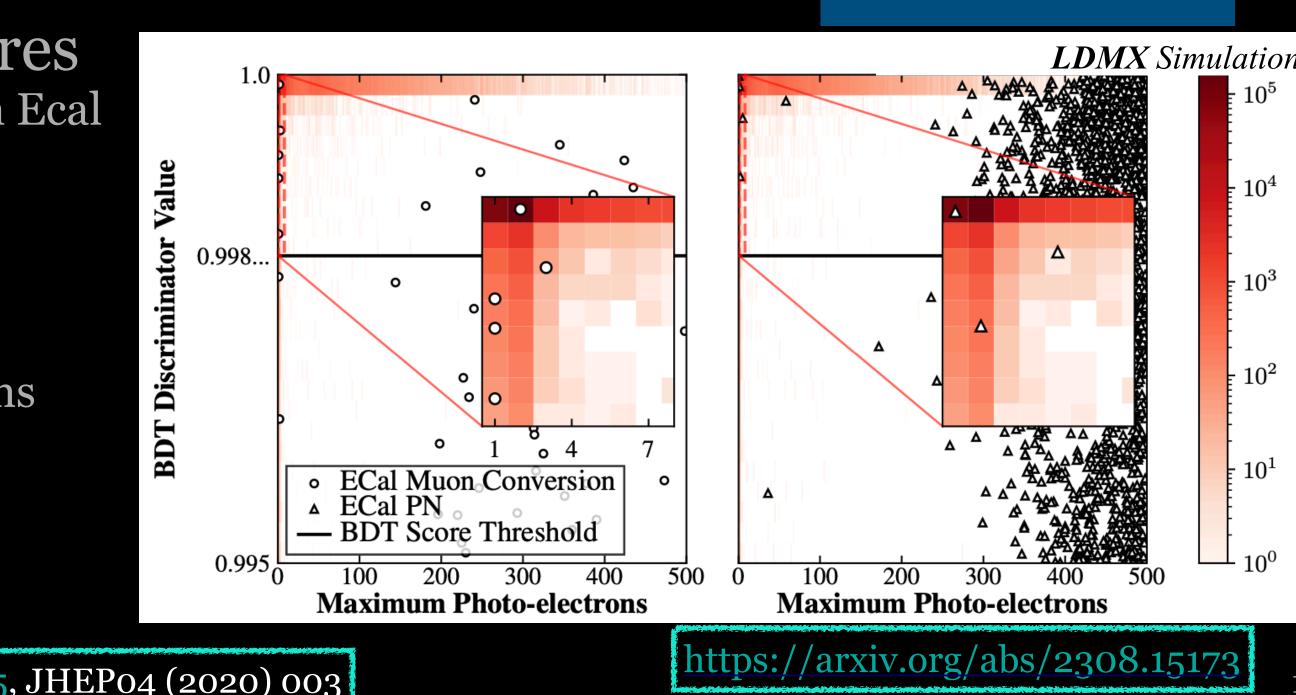
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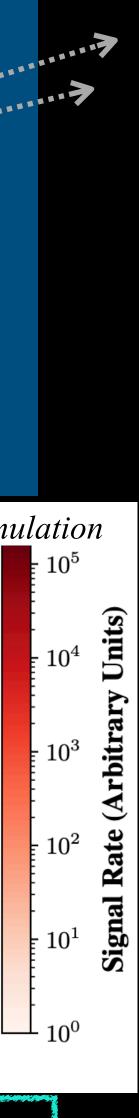
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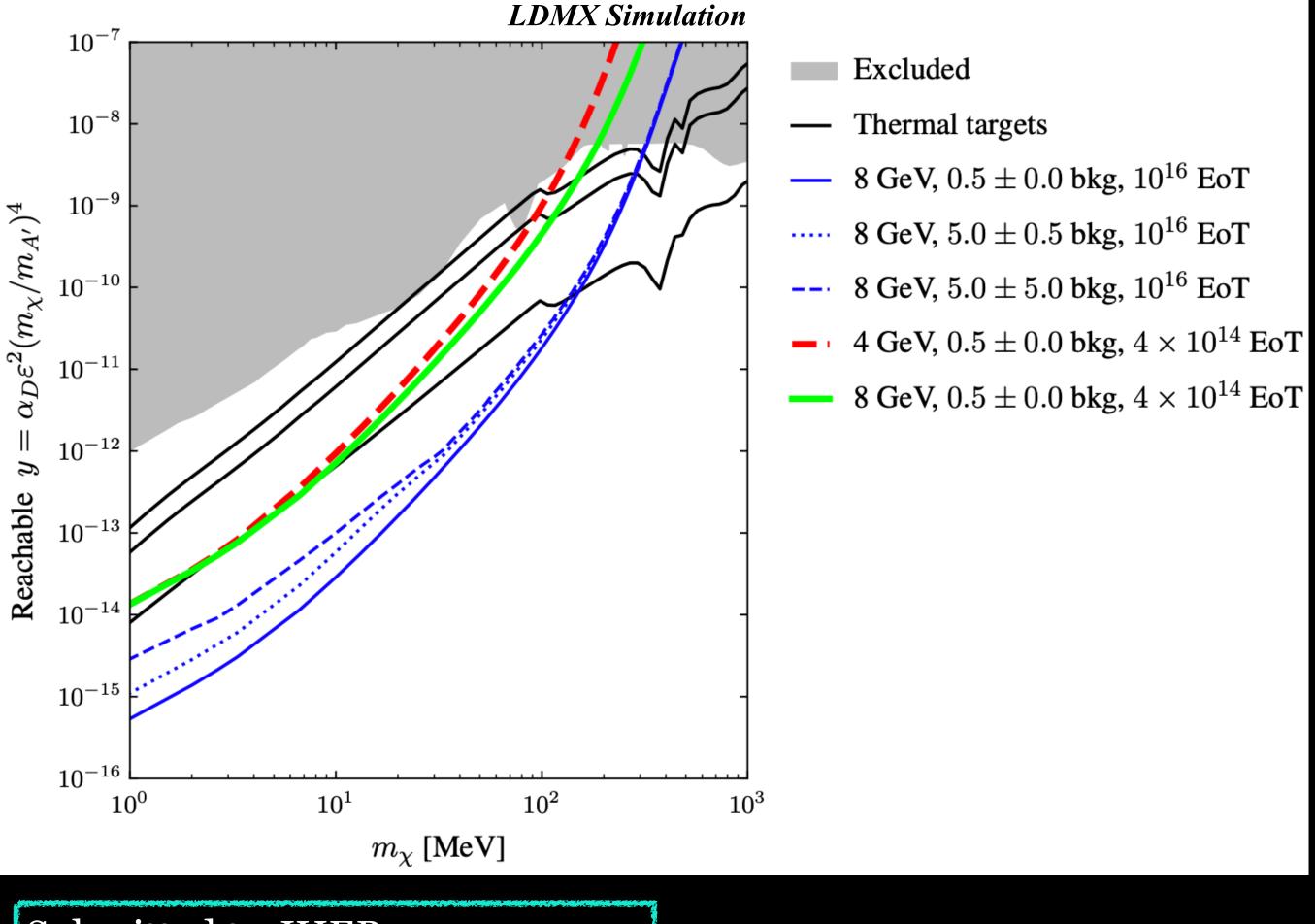
calculate missing energy nificant energy g tracks



THE



16



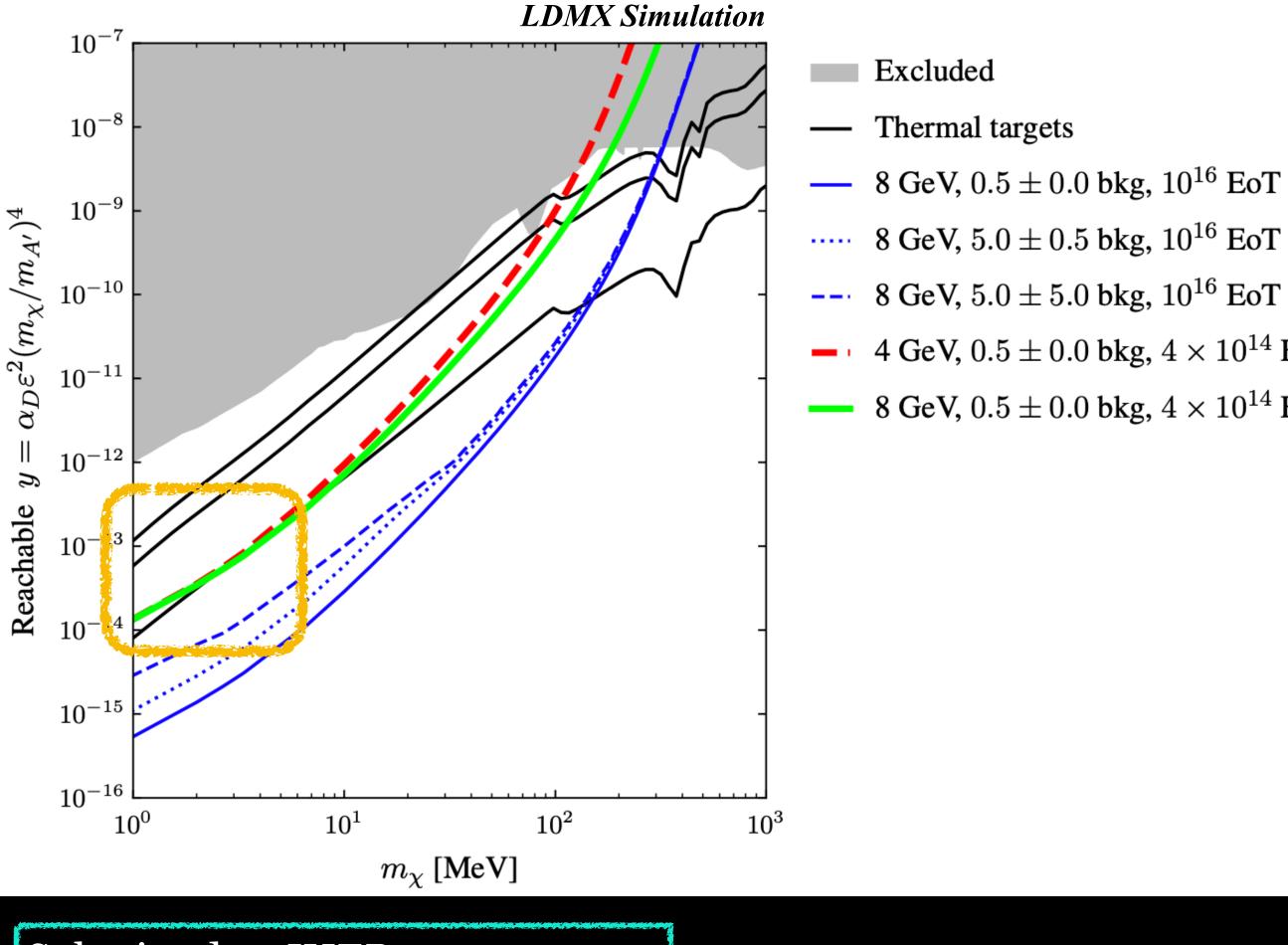
Submitted to JHEP https://arxiv.org/abs/2308.15173

TeVPA 2023



XMO

8 GeV,  $5.0 \pm 5.0$  bkg,  $10^{16}$  EoT 4 GeV,  $0.5 \pm 0.0$  bkg,  $4 \times 10^{14}$  EoT



Submitted to JHEP https://arxiv.org/abs/2308.15173

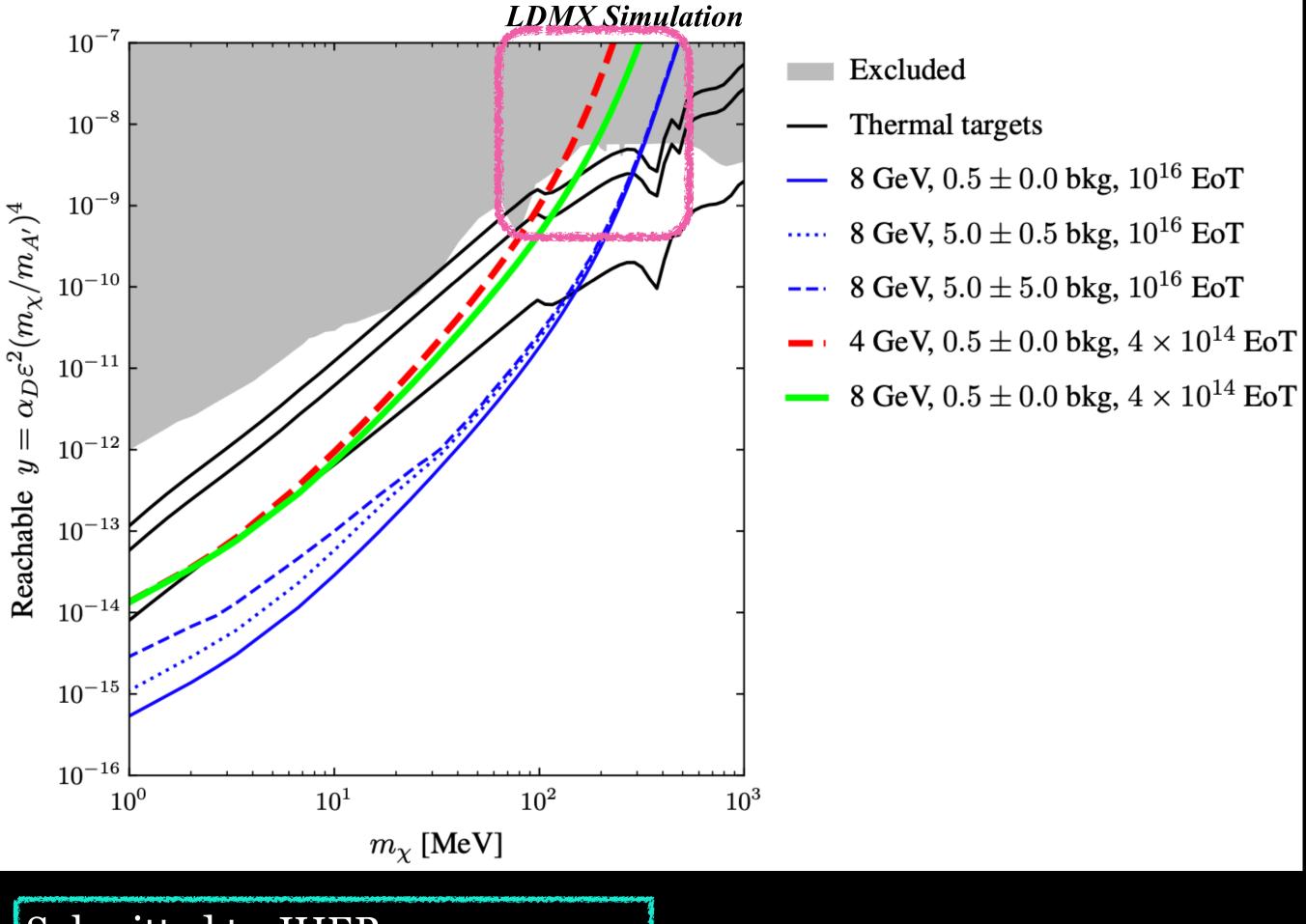
TeVPA 2023



XMO

Same sensitivity as 4 GeV at low masses

- 4 GeV,  $0.5 \pm 0.0$  bkg,  $4 \times 10^{14}$  EoT
- 8 GeV,  $0.5 \pm 0.0$  bkg,  $4 \times 10^{14}$  EoT



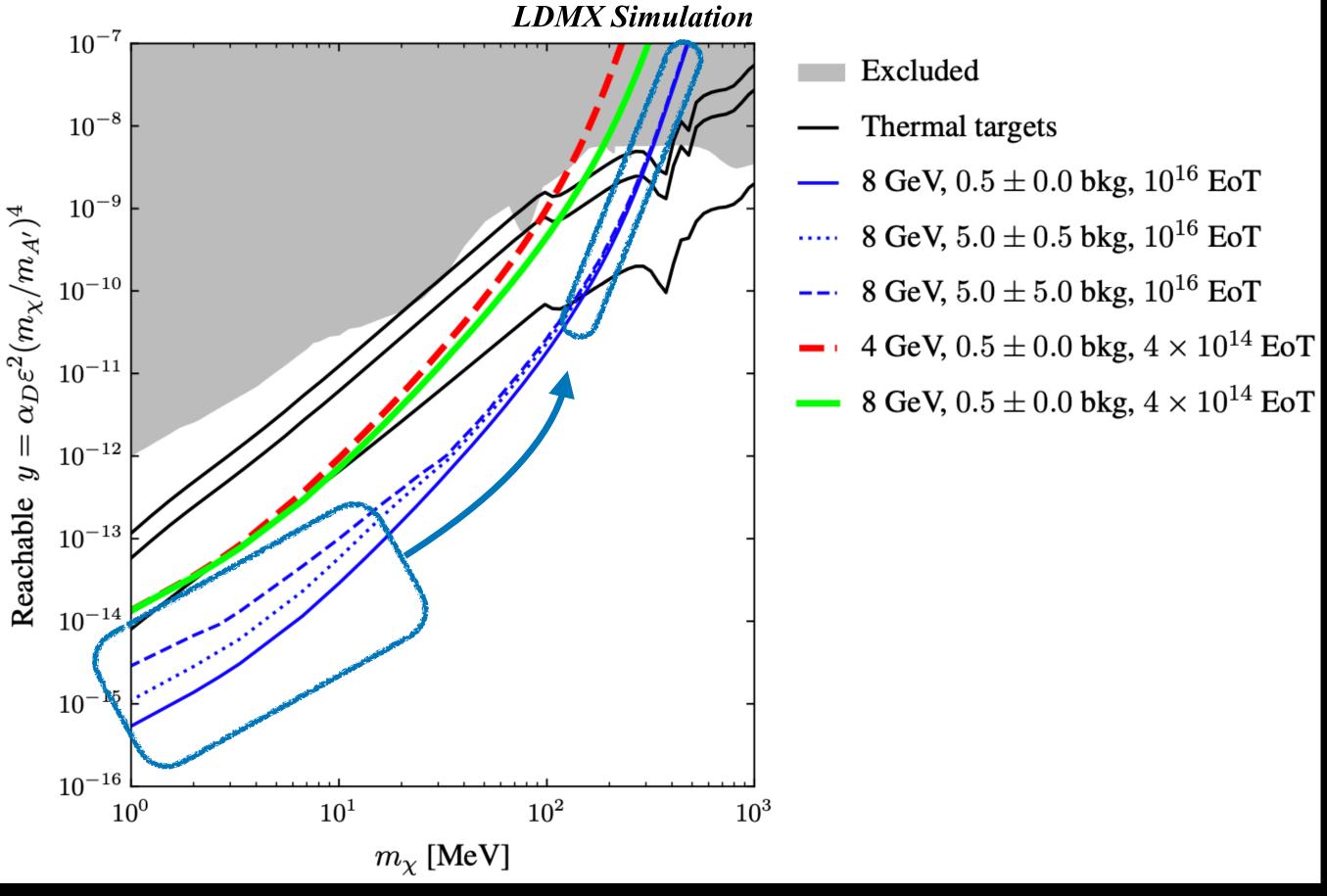
Submitted to JHEP https://arxiv.org/abs/2308.15173

TeVPA 2023



XMC

- Same sensitivity as 4 GeV at low masses
  - Increased reach at higher masses
- 4 GeV,  $0.5 \pm 0.0$  bkg,  $4 \times 10^{14}$  EoT



Submitted to JHEP https://arxiv.org/abs/2308.15173

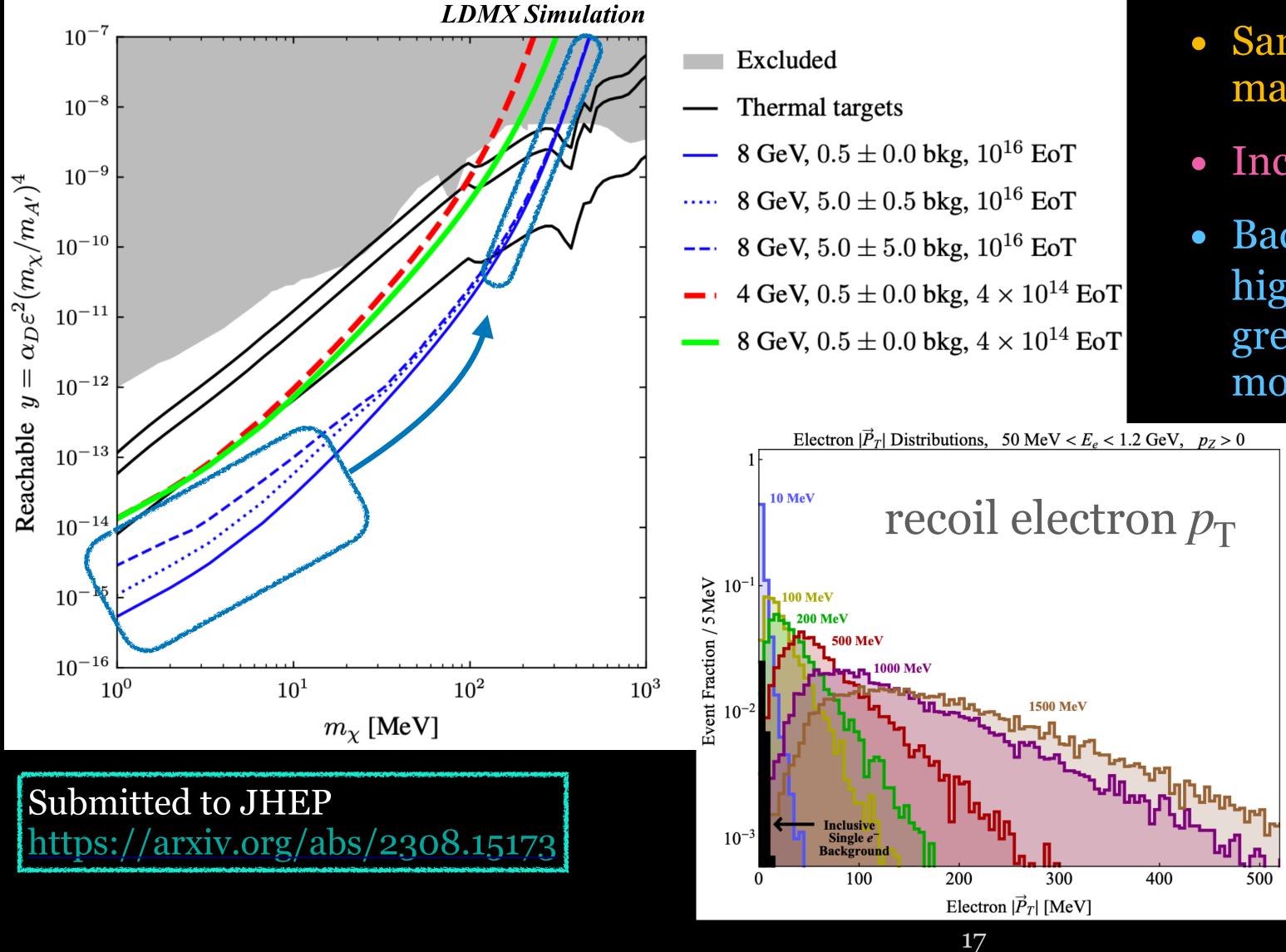
TeVPA 2023





- Same sensitivity as 4 GeV at low masses
  - Increased reach at higher masses
  - Background uncertainty reduced at higher mass using  $p_{\rm T}$  information great advantage of missing momentum experiment!

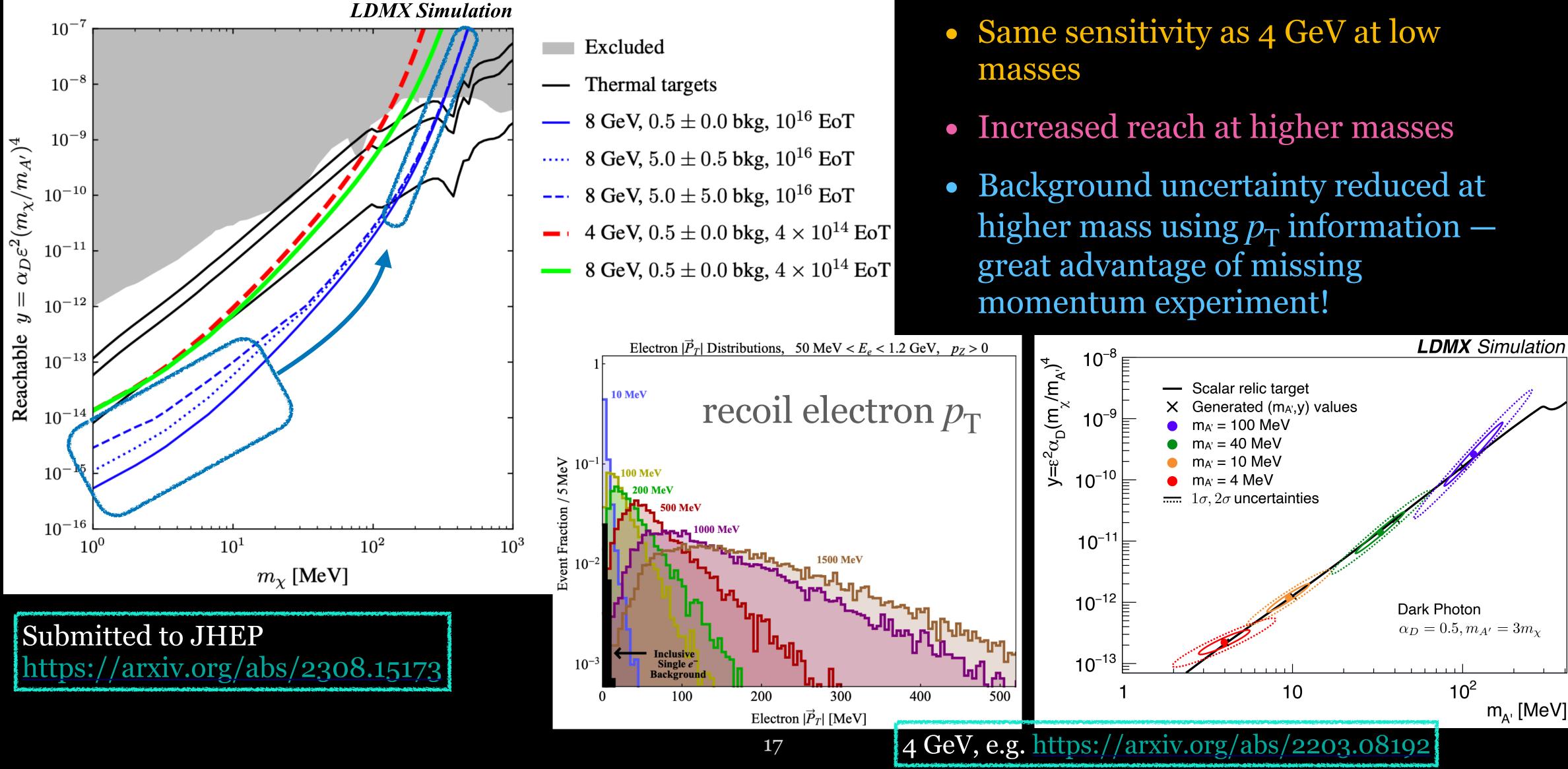
- 8 GeV,  $0.5 \pm 0.0$  bkg,  $4 \times 10^{14}$  EoT





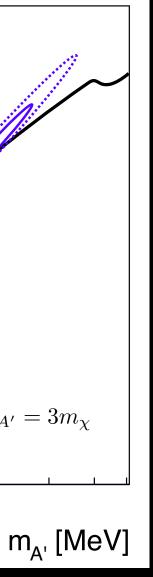
XMQ

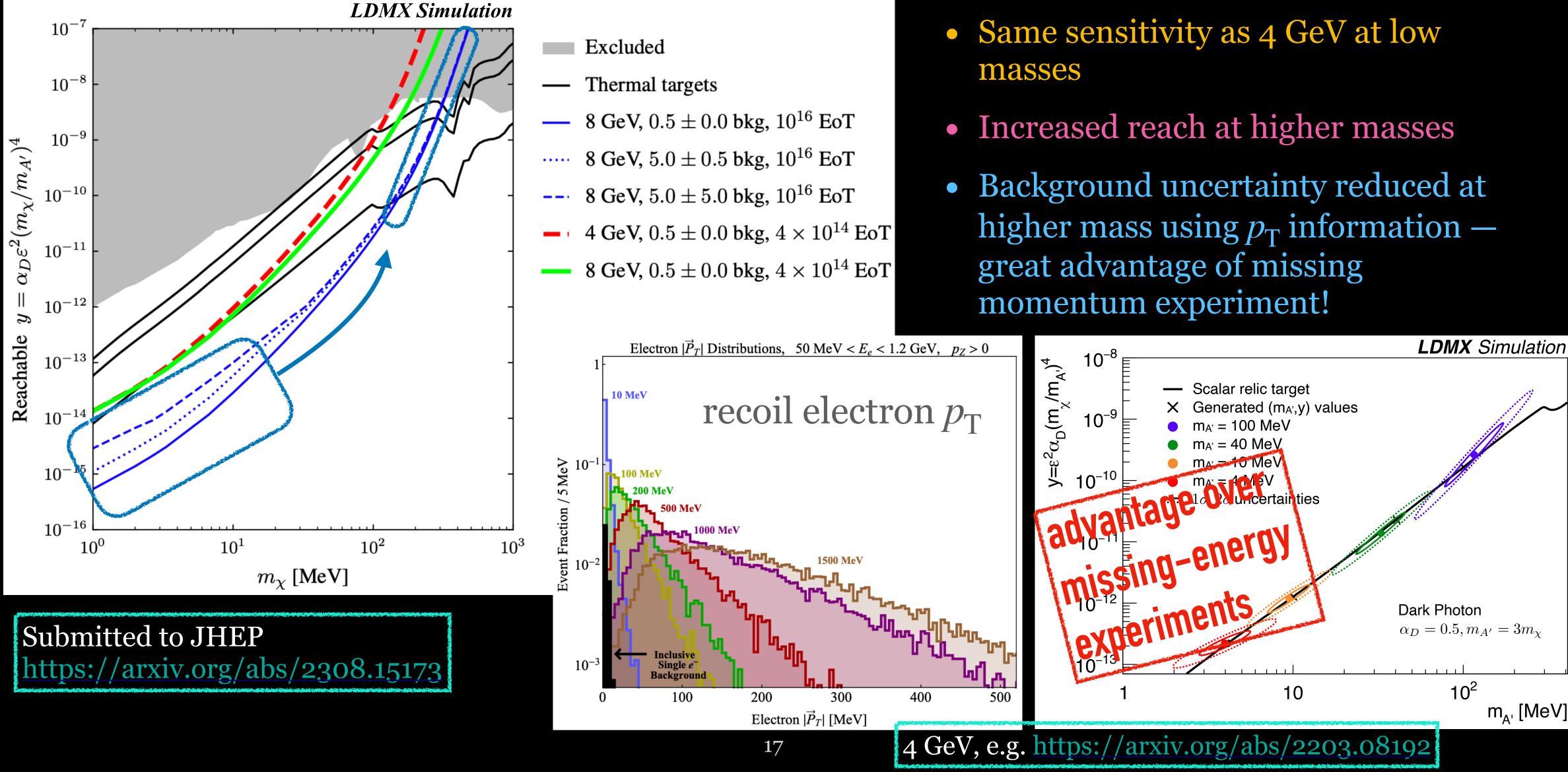
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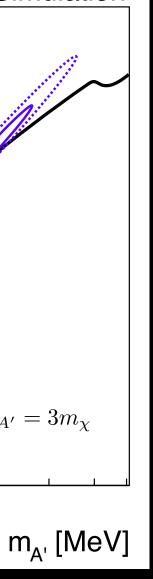
XMO



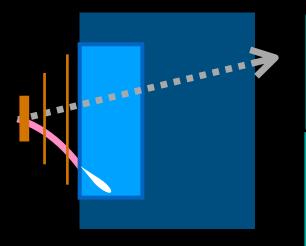




XMG

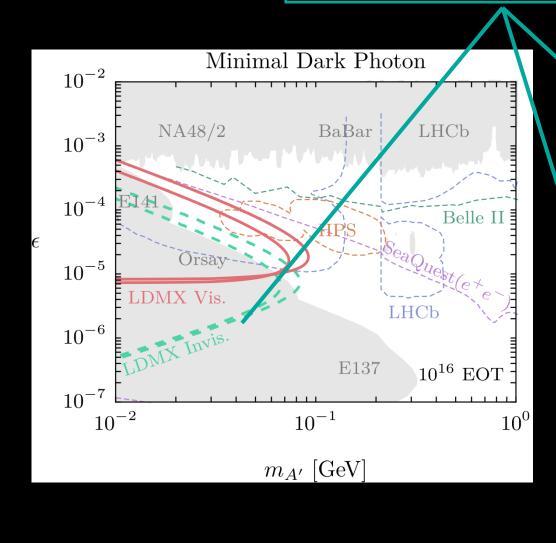


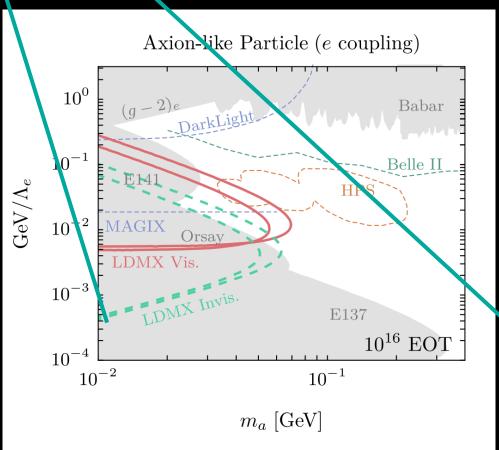
# A rich physics potential



#### Interpret missing momentum measurements

Secluded DM models, millicharge particles, invisibly decaying dark photons or mesons, axion-like particles, ...



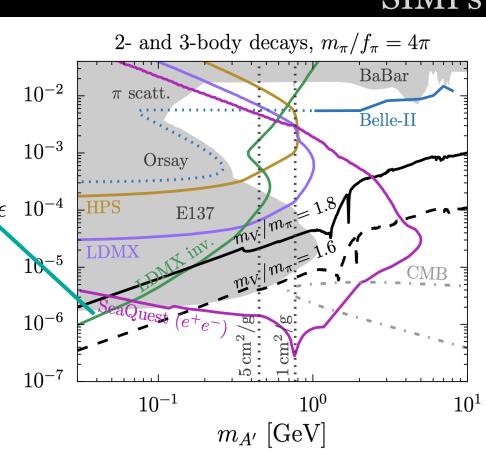




LK

see also: vector meson decays, Phys. Rev. D 105, 035036 (2022), https://arxiv.org/abs/2112.02104 and spin-1 DM models (submitted to JCAP), https://arxiv.org/abs/2307.0220

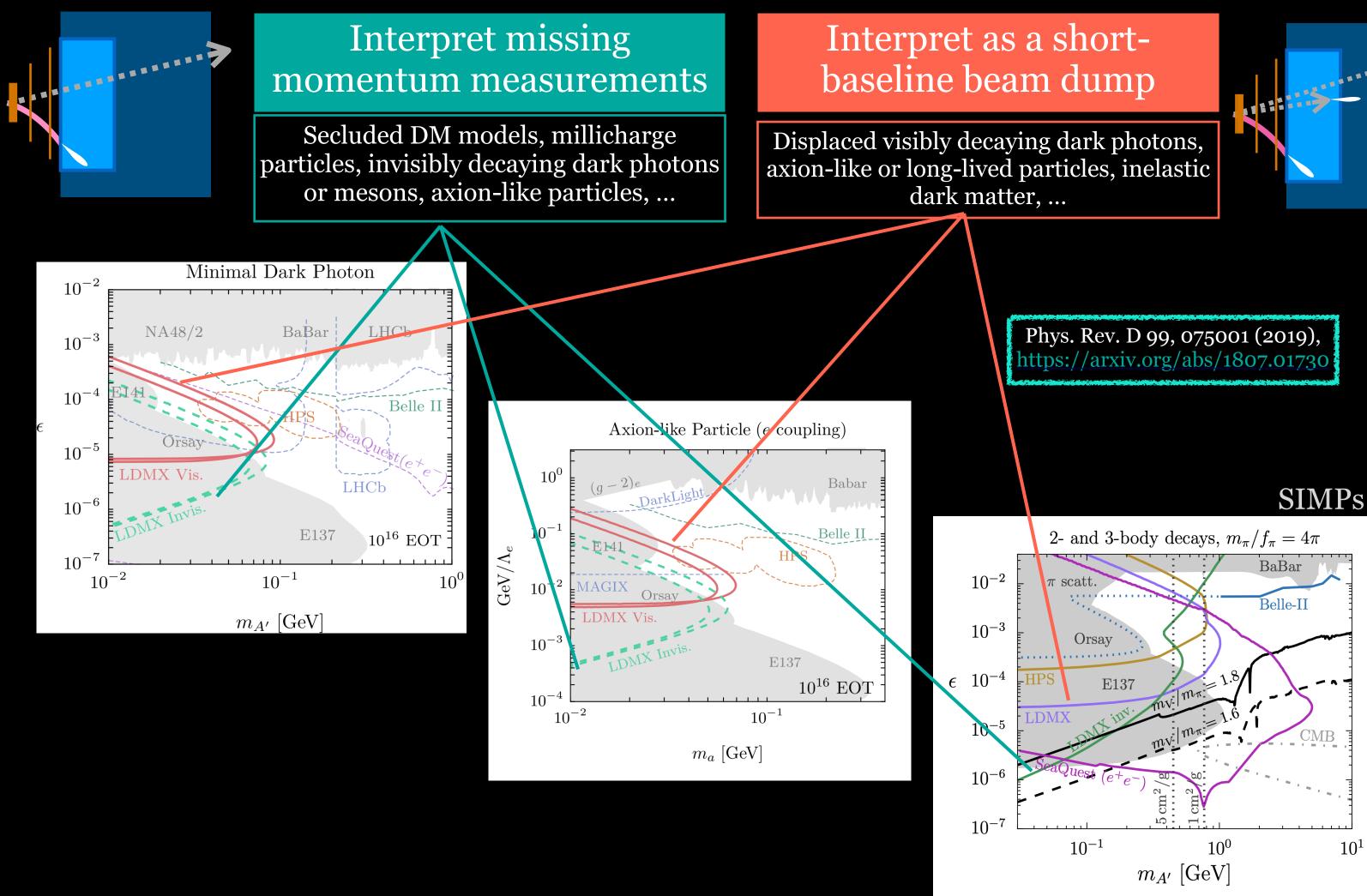
Phys. Rev. D 99, 075001 (2019), https://arxiv.org/abs/1807.01730



#### SIMPs



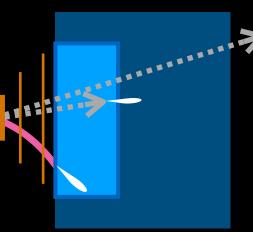
# A rich physics potential





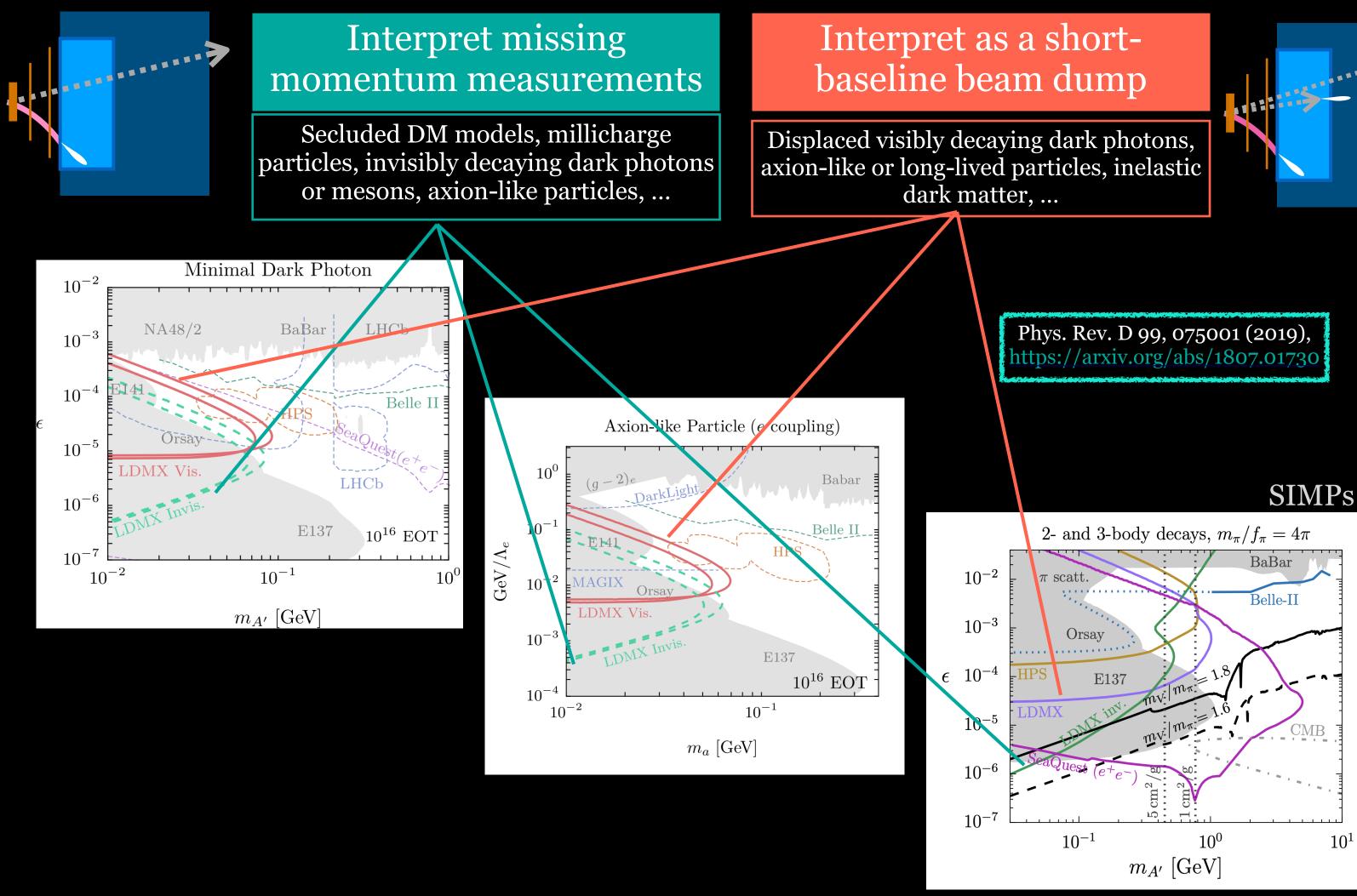
LK

see also: vector meson decays, Phys. Rev. D 105, 035036 (2022), https://arxiv.org/abs/2112.02104 and spin-1 DM models (submitted to JCAP), https://arxiv.org/abs/2307.0220





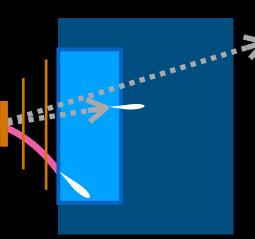
# A rich physics potential





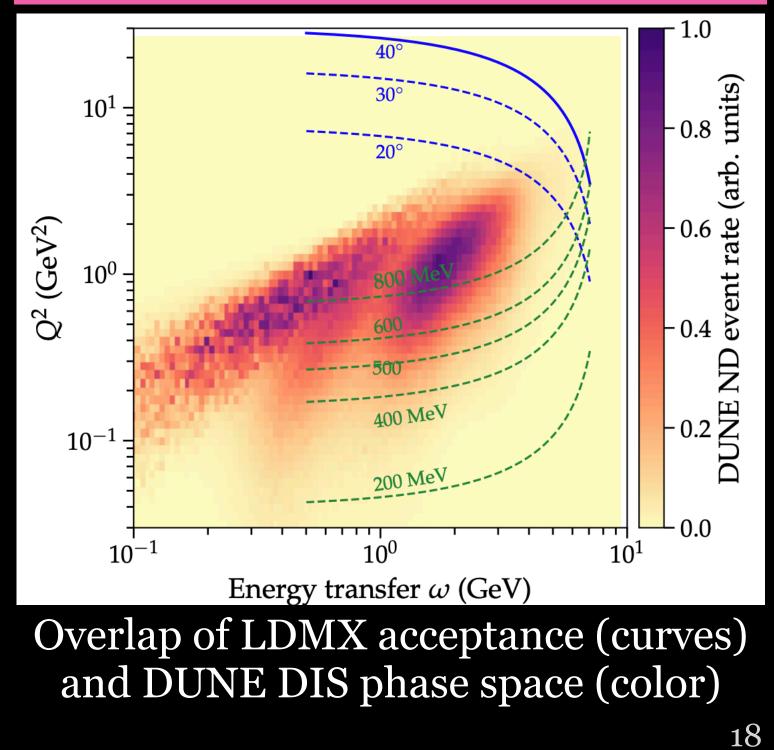
LK

see also: vector meson decays, Phys. Rev. D 105, 035036 (2022), https://arxiv.org/abs/2112.02104 and spin-1 DM models (submitted to JCAP), https://arxiv.org/abs/2307.0220



#### Phys. Rev. D 101, 053004, https://arxiv.org/abs/1912.06140

Electro-nuclear scattering particle spectra constrain theoretical uncertainties in neutrino experiments







# Searching for DM with an electron beam at the GeV scale is a wellmotivated and worthwhile effort.

- LDMX can conclusively probe the stable matter mass range for DM of thermal origin, and has a broad physics potential beyond this goal • Simulation and hardware understanding rapidly maturing  $\rightarrow$  design
- is being finalized
  - can reach required background rejection confirmed in new 8 GeV sensitivity study submitted to JHEP!
  - prototype beam tests performed at CERN, paper in preparation































# Searching for DM with an electron beam at the GeV scale is a wellmotivated and worthwhile effort.





LK

TeVPA















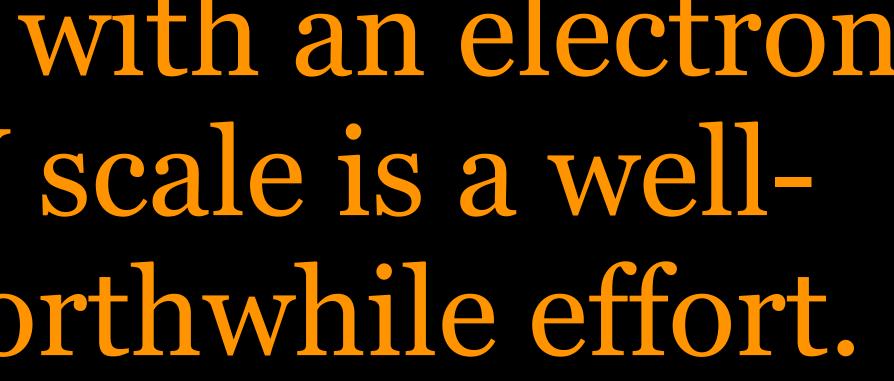
















Measurements of photo- and electro-nuclear processes for neutrino experiments: Lepton-Nucleus Cross Section Measurements for DUNE with the LDMX Detector ,Phys. Rev. D 101, 053004, https://arxiv.org/abs/1912.06140

Background rejection to < 1 event at 4 GeV: A High Efficiency Photon Veto for the Light Dark Matter eXperiment, J. High Energ. Phys. 2020, 3 (2020) https://arxiv.org/abs/1912.05535, at 8 GeV: Photon-rejection Power of the Light Dark Matter eXperiment in an 8 GeV Beam, submitted to JHEP (SLAC-PUB-17550, FERMILAB-PUB-23-433-PPD-T) https://arxiv.org/abs/2308.15173

Current Status and Future Prospects for the Light Dark Matter eXperiment, LDMX Snowmass 2021 contribution, https://arxiv.org/abs/2203.08192

The SLAC Linac to ESA (LESA) Beamline for Dark Sector Searches and Test Beams, Snowmass 2021 contribution, https://arxiv.org/abs/2205.13215

Schuster, Toro, Zhou, Probing Invisible Vector Meson Decays with NA64 and LDMX, Phys. Rev. D 105, 035036 (2022) https://arxiv.org/abs/2112.02104

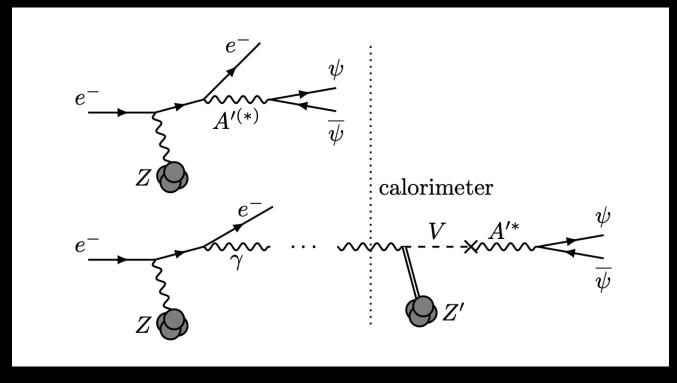
Building a Distributed Computing System for LDMX: Challenges of creating and operating a lightweight e-infrastructure for small-to-medium size accelerator experiments, vCHEP2021 proceedings, https://arxiv.org/abs/2105.02977

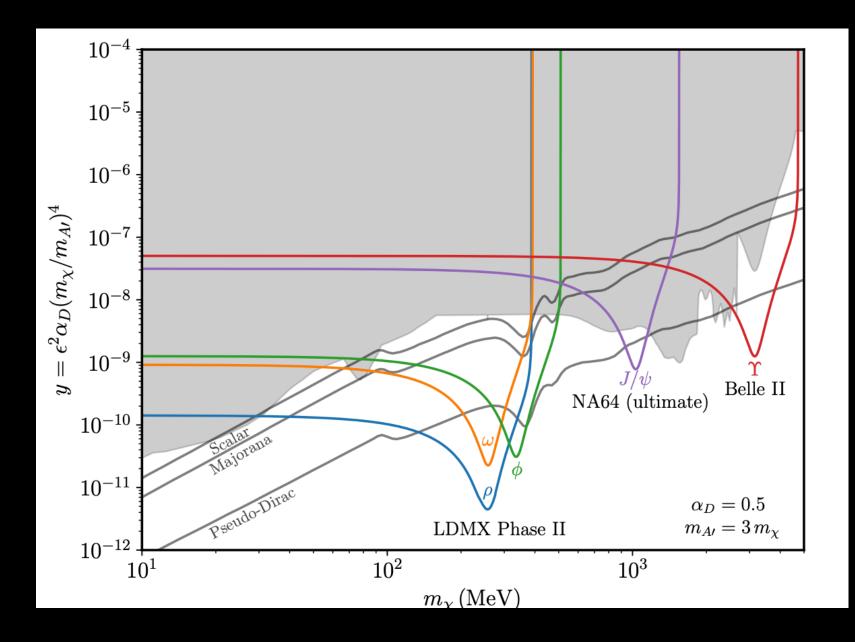
Light Dark Matter eXperiment (LDMX), https://arxiv.org/abs/1808.05219



# references

# Probing invisible meson decays



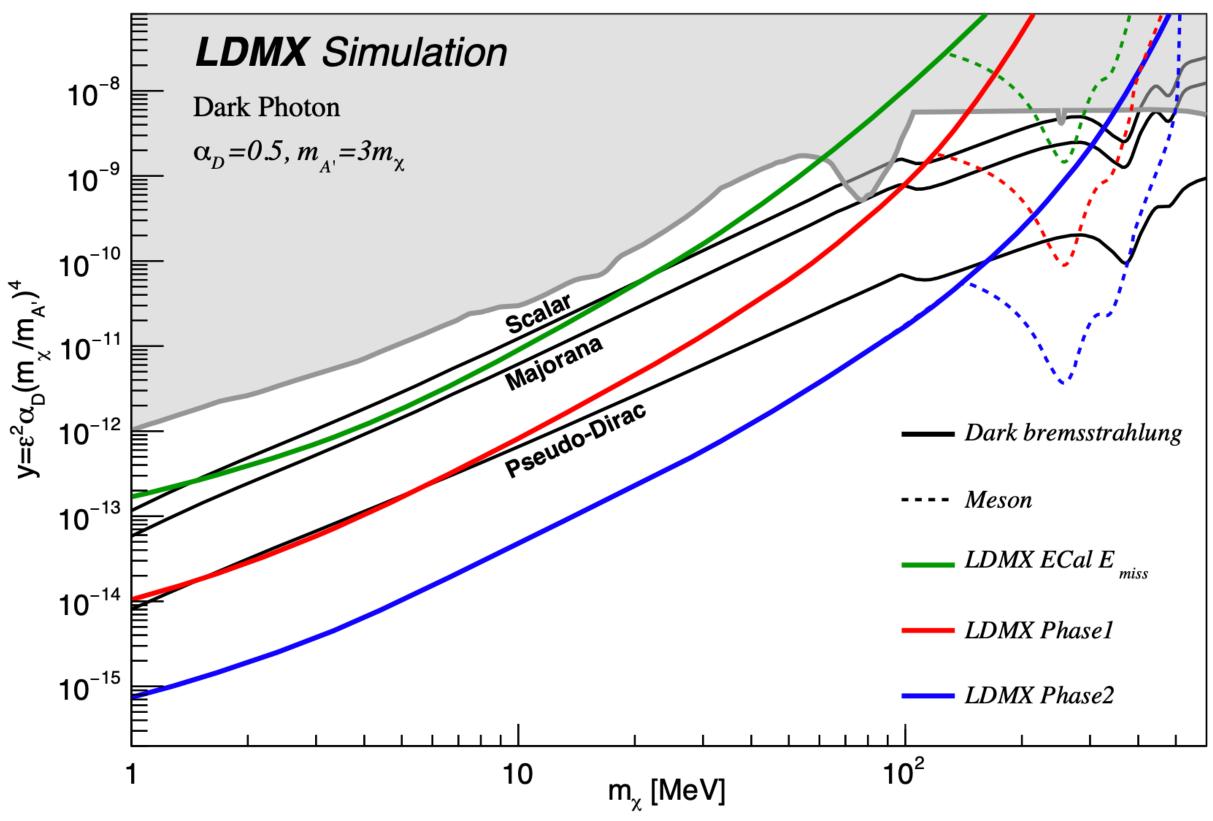


https://arxiv.org/pdf/2112.02104.pdf



TeVPA 2023





LDMX Snowmass 2021 contribution https://arxiv.org/abs/2203.08192



# Detectors pre-target: knowing what's coming in



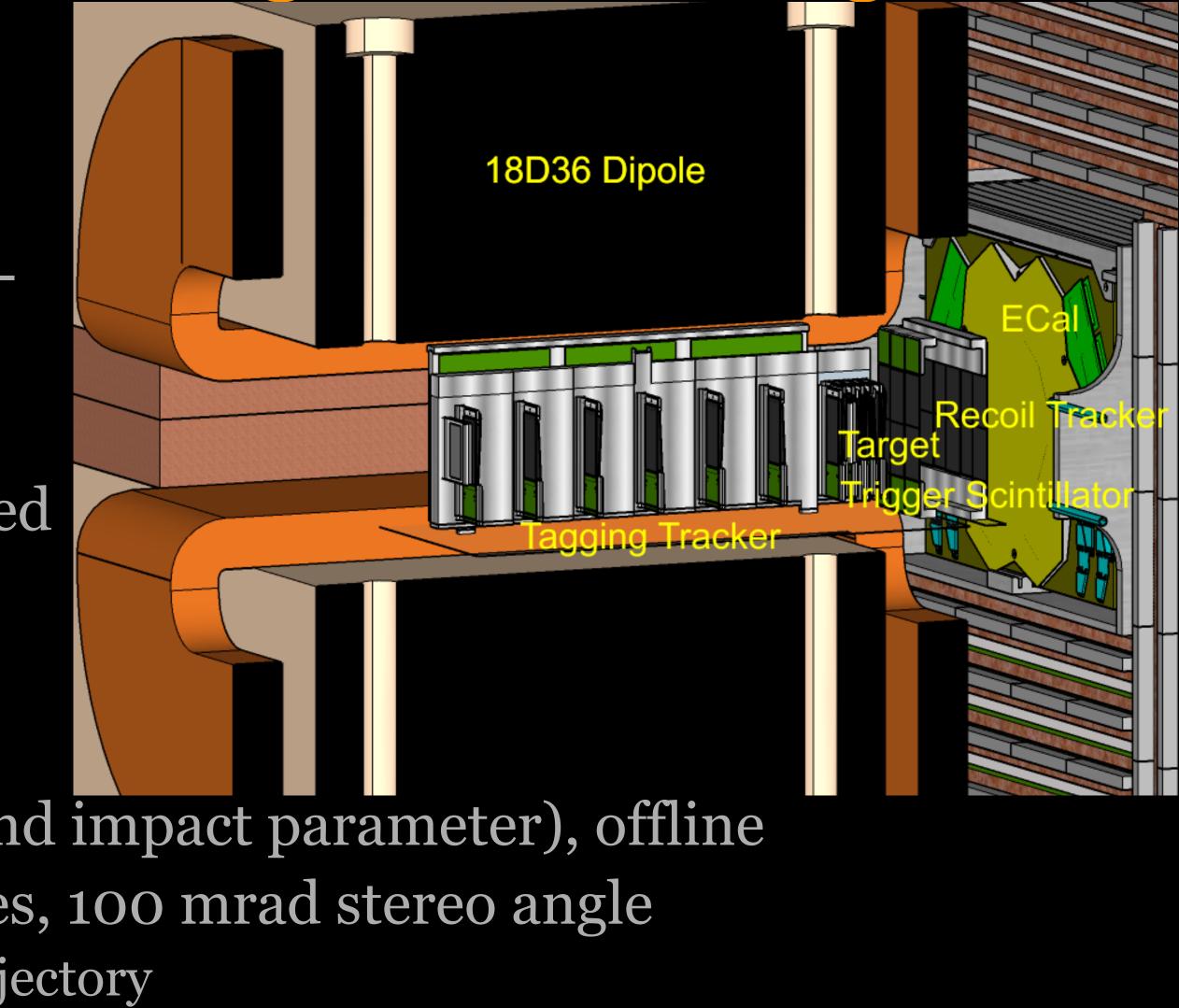
Trigger scintillators:

- Trigger-level electron counting, input to missingenergy trigger
- 3 arrays of horizontal scintillating bars, positioned along incoming beam

Tagging tracker (based on HPS tracker design):

- Tag beam electrons (momentum and impact parameter), offline
- 7 double-sided silicon strip modules, 100 mrad stereo angle
  - 10 cm apart, along incoming beam trajectory
  - vertically oriented strips for optimal momentum resolution





23

# After target: recoil tracker, Ecal and Hcal

Recoil tracker:

optimized for momentum resolution and acceptance at 1-2 GeV

Electromagnetic calorimeter (based on CMS HGCal upgrade):

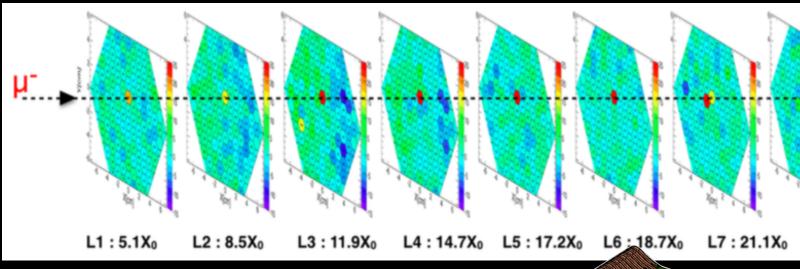
- high-granularity, capable of MIP tracking
- $40 X_0$  Si-W calorimeter, radiation hard

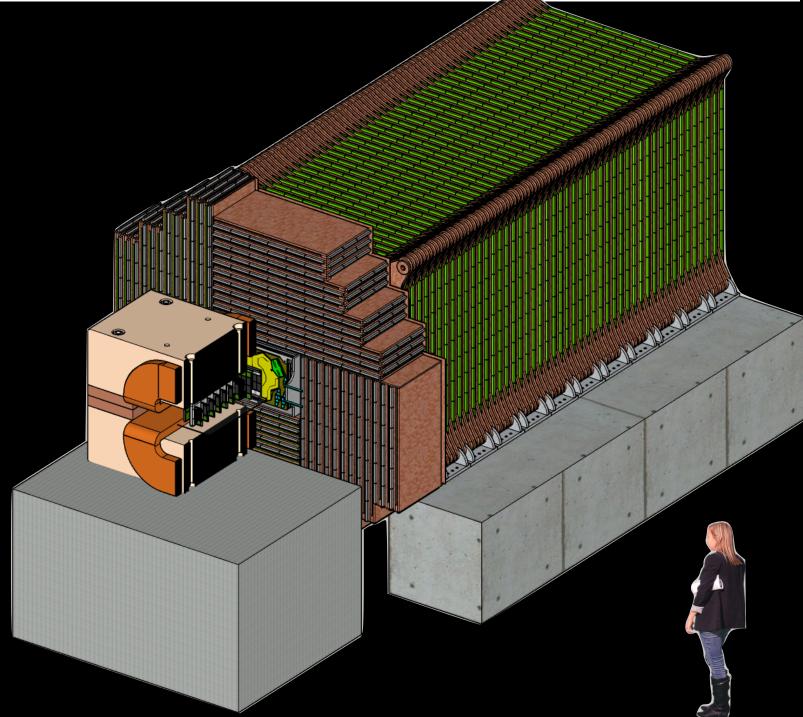
Hadronic calorimeter (adapted from Mu2e cosmic ray veto):

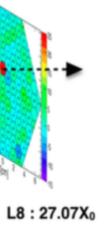
- on all sides and behind Ecal for wide and deep coverage
  - sampling calorimeter, steel absorber,  $17\lambda$  deep
- readout in scintillator bars of alternating x/yorientation

XMQ









### Goal: no background from ~10<sup>14</sup> electrons on target

# First handle: recoiling electron

- trigger on Ecal E < 1.5 GeV (1e);
- select recoil electron momentum < 1.2 GeV
- veto if number of outgoing tracks > incoming tracks

# Exploit high-granularity Ecal features

- BDT trained to reject photonuclear events in Ecal
- MIP tracking powerful on sparse events

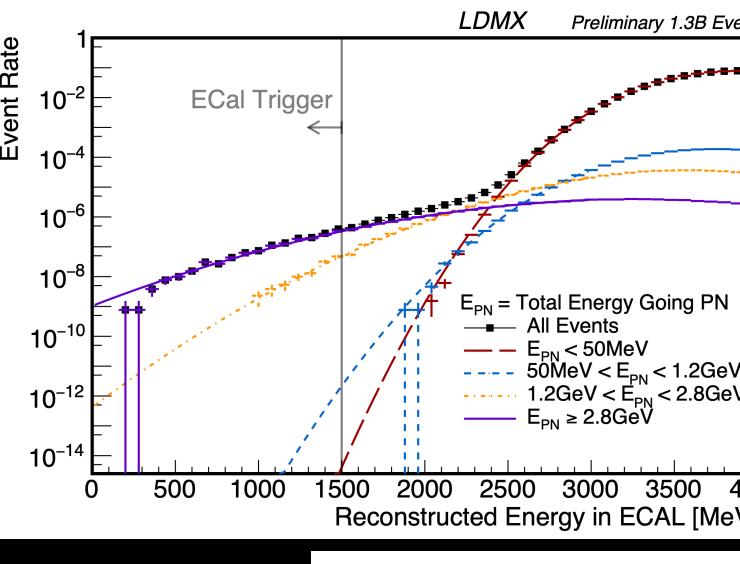
# Veto on Hcal activity

- allow only a few photoelectrons
- deep enough to tease out even single neutrons









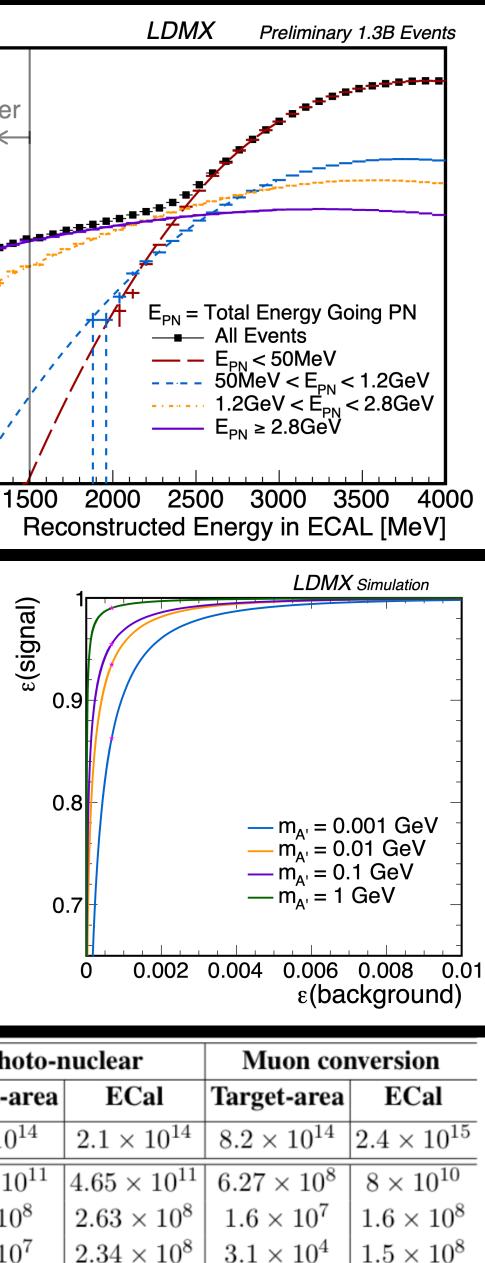


	Photo-nuclear		Muon conversion	
	Target-area	ECal	Target-area	ECal
EoT equivalent	$4 \times 10^{14}$	$2.1 \times 10^{14}$	$8.2 \times 10^{14}$	$2.4 \times 10$
Total events simulated	$8.8\times10^{11}$	$4.65\times 10^{11}$	$6.27 \times 10^8$	$8 \times 10^1$
Trigger, ECal total energy $< 1.5~{\rm GeV}$	$1 \times 10^8$	$2.63\times 10^8$	$1.6  imes 10^7$	$1.6 \times 10^{-1}$
Single track with $p < 1.2 \mathrm{GeV}$	$2 \times 10^7$	$2.34\times 10^8$	$3.1 \times 10^4$	$1.5 \times 10^{-1}$
ECal BDT ( $> 0.99$ )	$9.4  imes 10^5$	$1.32  imes 10^5$	< 1	< 1
HCal max $PE < 5$	< 1	10	< 1	< 1
ECal MIP tracks $= 0$	< 1	< 1	< 1	< 1

# Goal: no background from ~10<sup>14</sup> electrons on target

# First handle: recoiling electron

- trigger on Ecal E < 1.5 GeV (1e);
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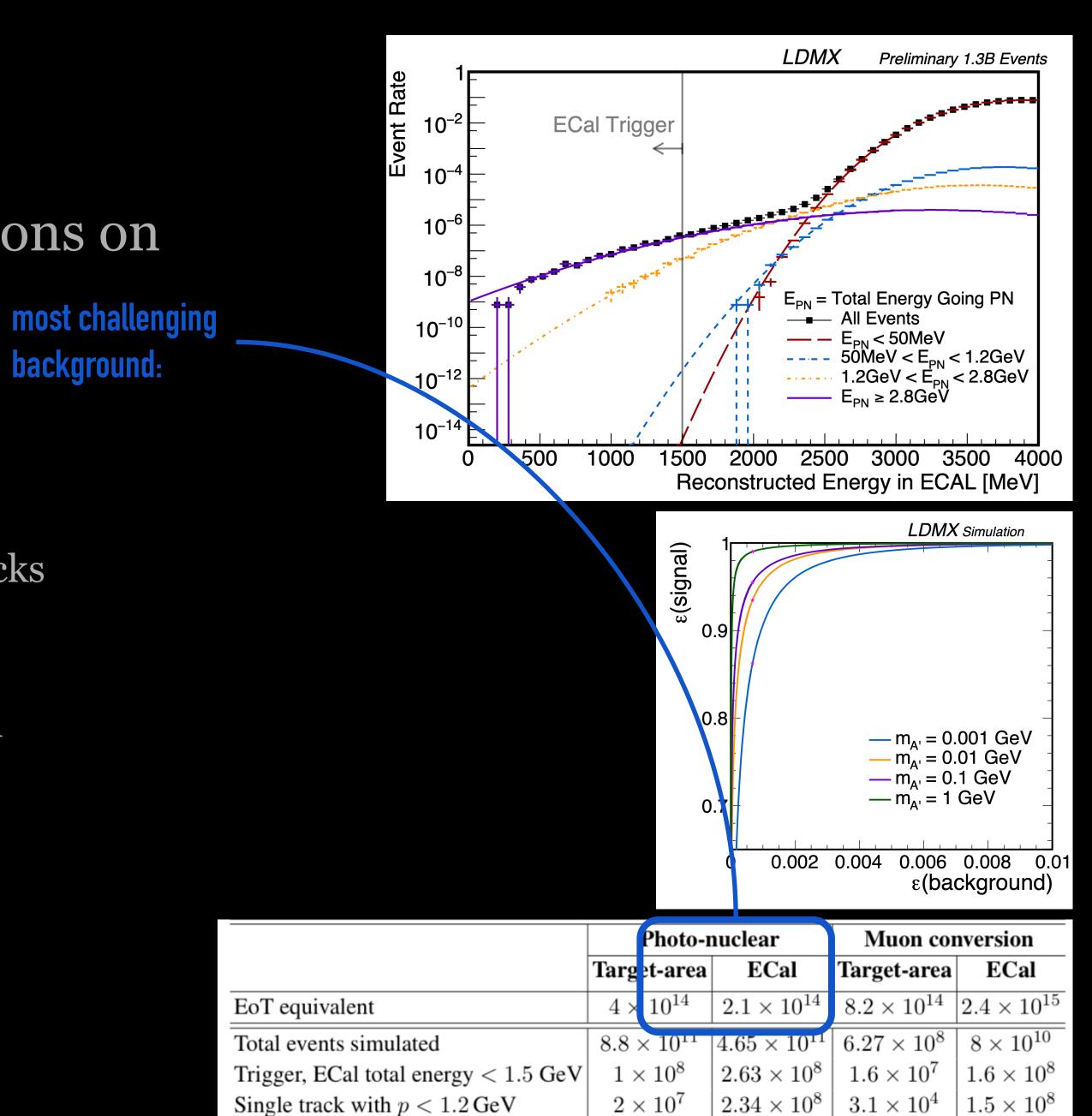
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- BDT trained to reject photonuclear events in Ecal
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- allow only a few photoelectrons
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 $9.4 \times 10^5$ 

< 1

< 1

 $1.32 \times 10^{5}$ 

10

< 1

< 1

< 1

< 1

< 1

< 1

< 1

ECal BDT (> 0.99)

ECal MIP tracks = 0

HCal max PE < 5

# Goal: no background from ~10<sup>14</sup> electrons on target

# First handle: recoiling electron

- trigger on Ecal E < 1.5 GeV (1e);
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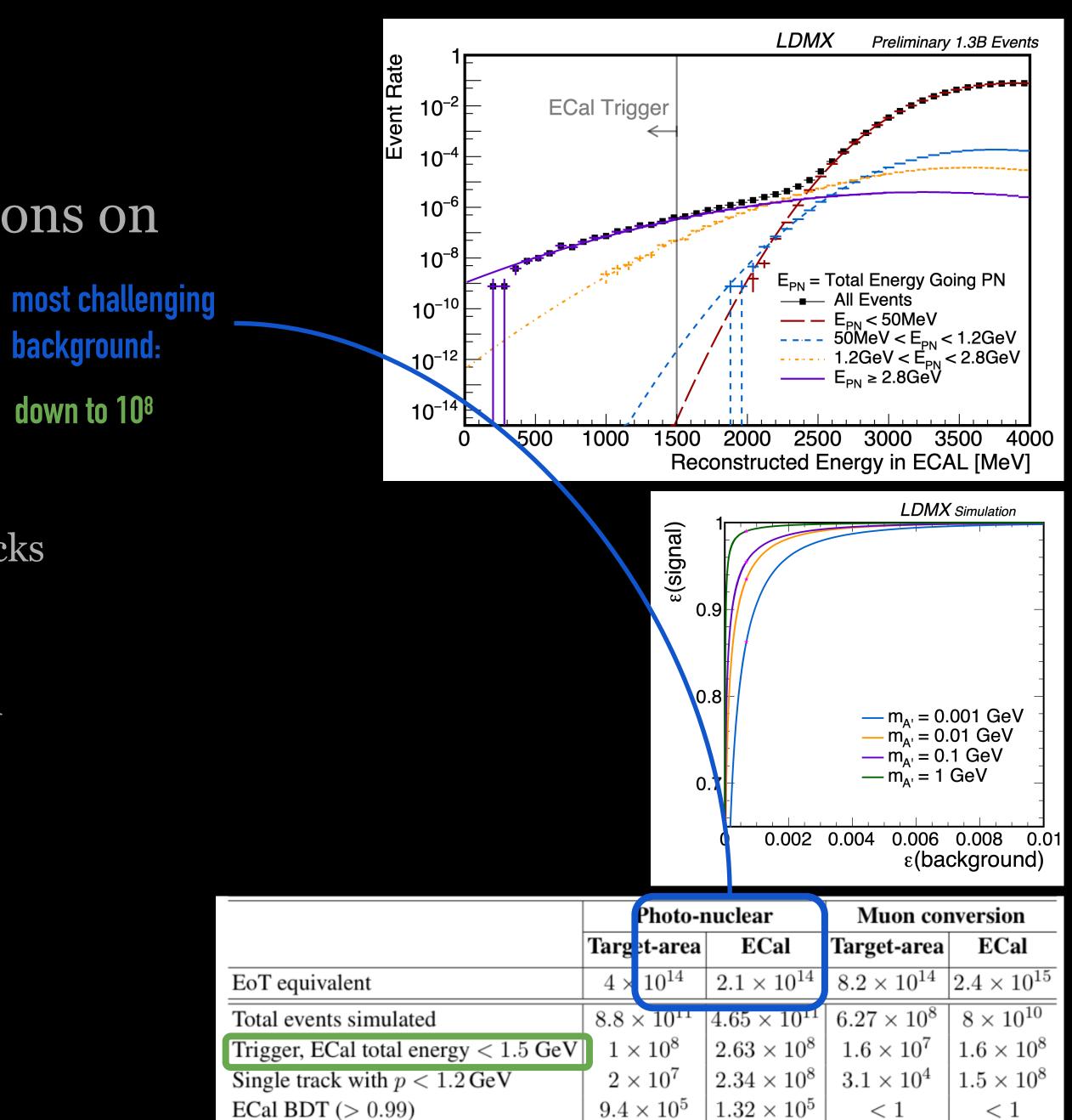
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- BDT trained to reject photonuclear events in Ecal
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- allow only a few photoelectrons
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< 1

< 1

10

< 1

< 1

< 1

< 1

< 1

HCal max PE < 5

ECal MIP tracks = 0

# Goal: no background from ~10<sup>14</sup> electrons on target

# First handle: recoiling electron

- trigger on Ecal E < 1.5 GeV (1e);
- select recoil electron momentum < 1.2 GeV
- veto if number of outgoing tracks > incoming tracks

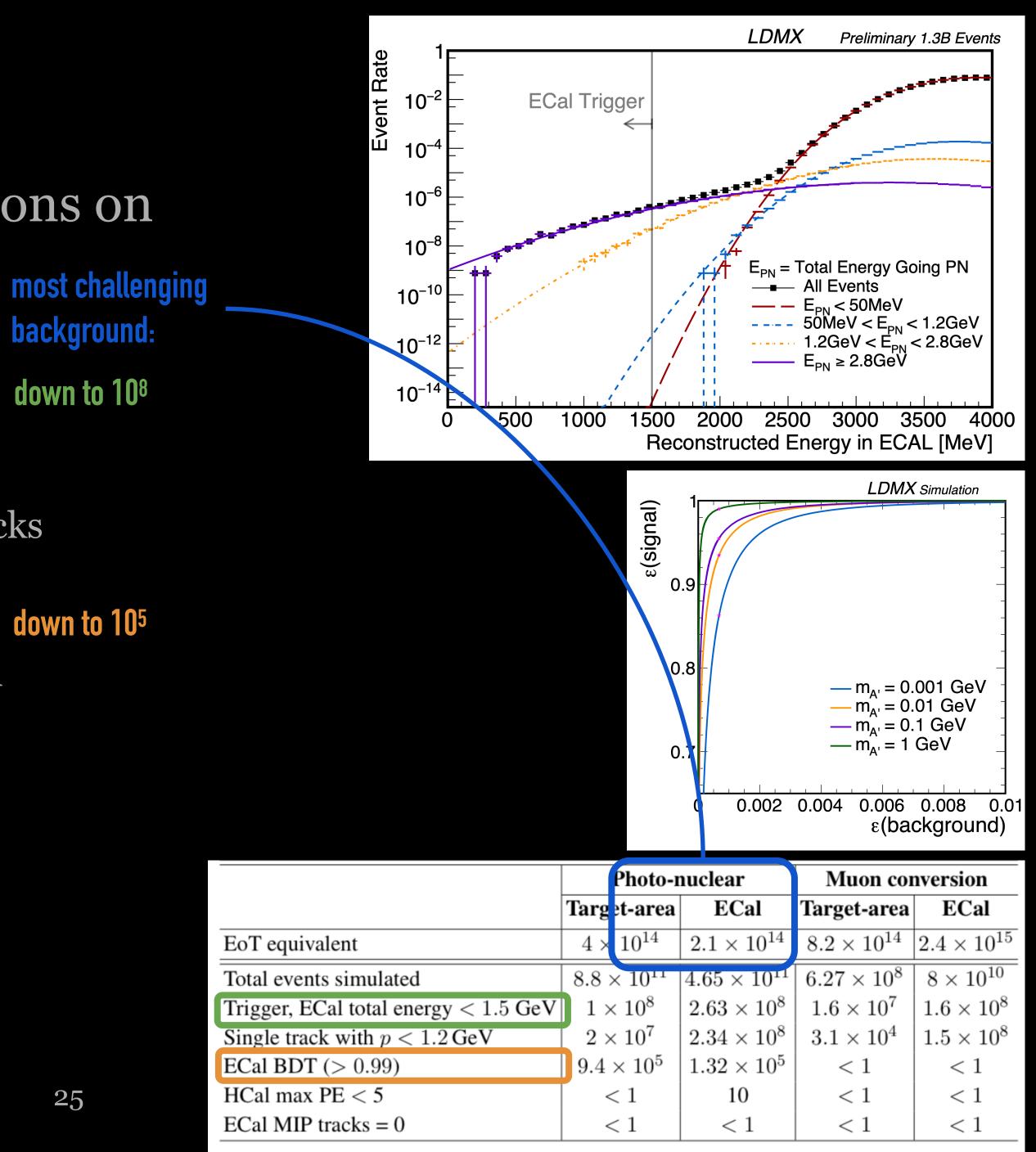
# Exploit high-granularity Ecal features down to 10<sup>5</sup>

- BDT trained to reject photonuclear events in Ecal
- MIP tracking powerful on sparse events

# Veto on Hcal activity

- allow only a few photoelectrons
- deep enough to tease out even single neutrons





### Goal: no background from ~10<sup>14</sup> electrons on target

# First handle: recoiling electron

- trigger on Ecal E < 1.5 GeV (1e);
- select recoil electron momentum < 1.2 GeV
- veto if number of outgoing tracks > incoming tracks

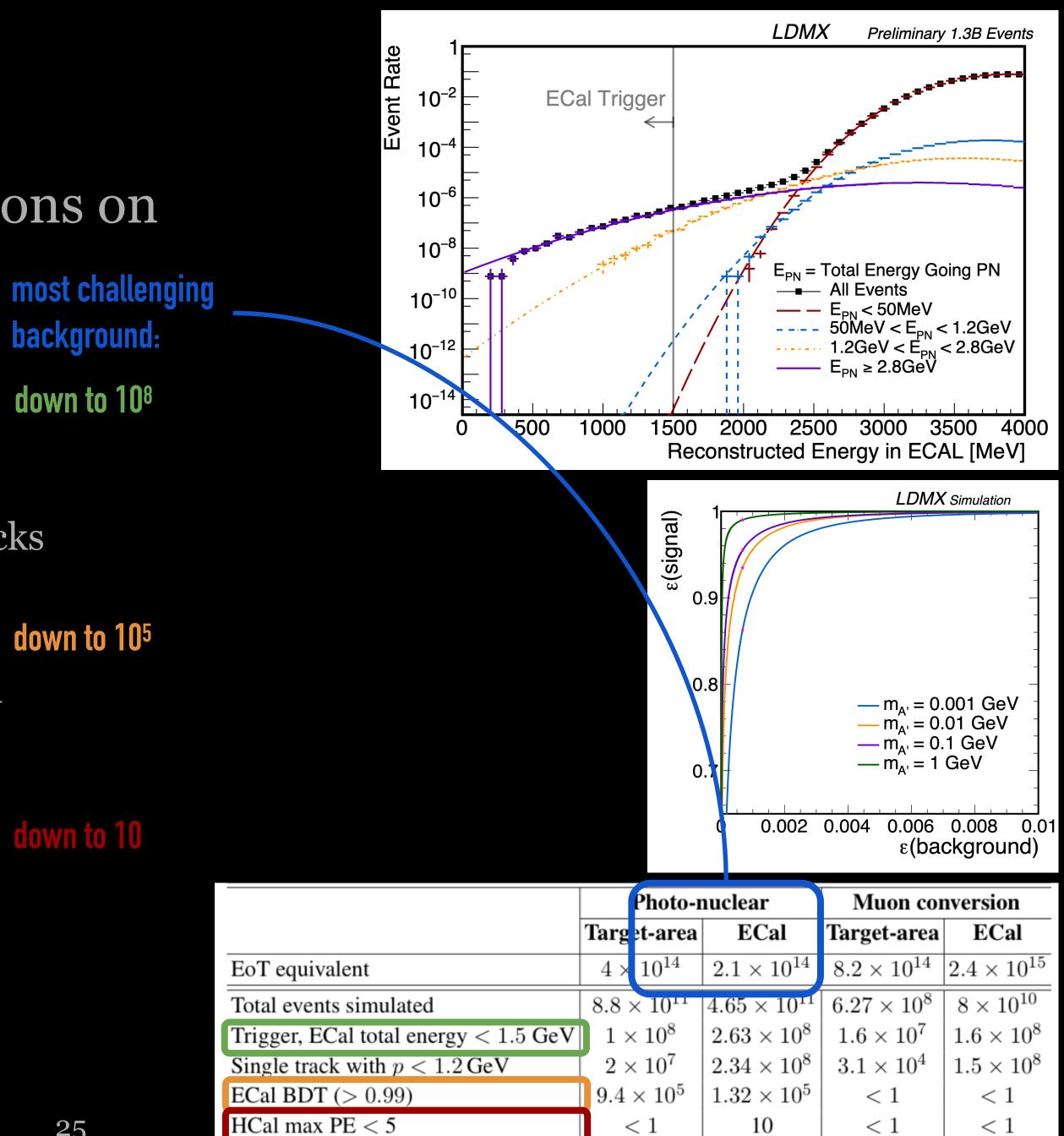
# Exploit high-granularity Ecal features down to 10<sup>5</sup>

- BDT trained to reject photonuclear events in Ecal
- MIP tracking powerful on sparse events

# Veto on Hcal activity

- allow only a few photoelectrons
- deep enough to tease out even single neutrons





< 1

< 1

25

ECal MIP tracks = 0

10

< 1

< 1

< 1

< 1

< 1

### Goal: no background from ~10<sup>14</sup> electrons on target

# First handle: recoiling electron

- trigger on Ecal E < 1.5 GeV (1e);
- select recoil electron momentum < 1.2 GeV
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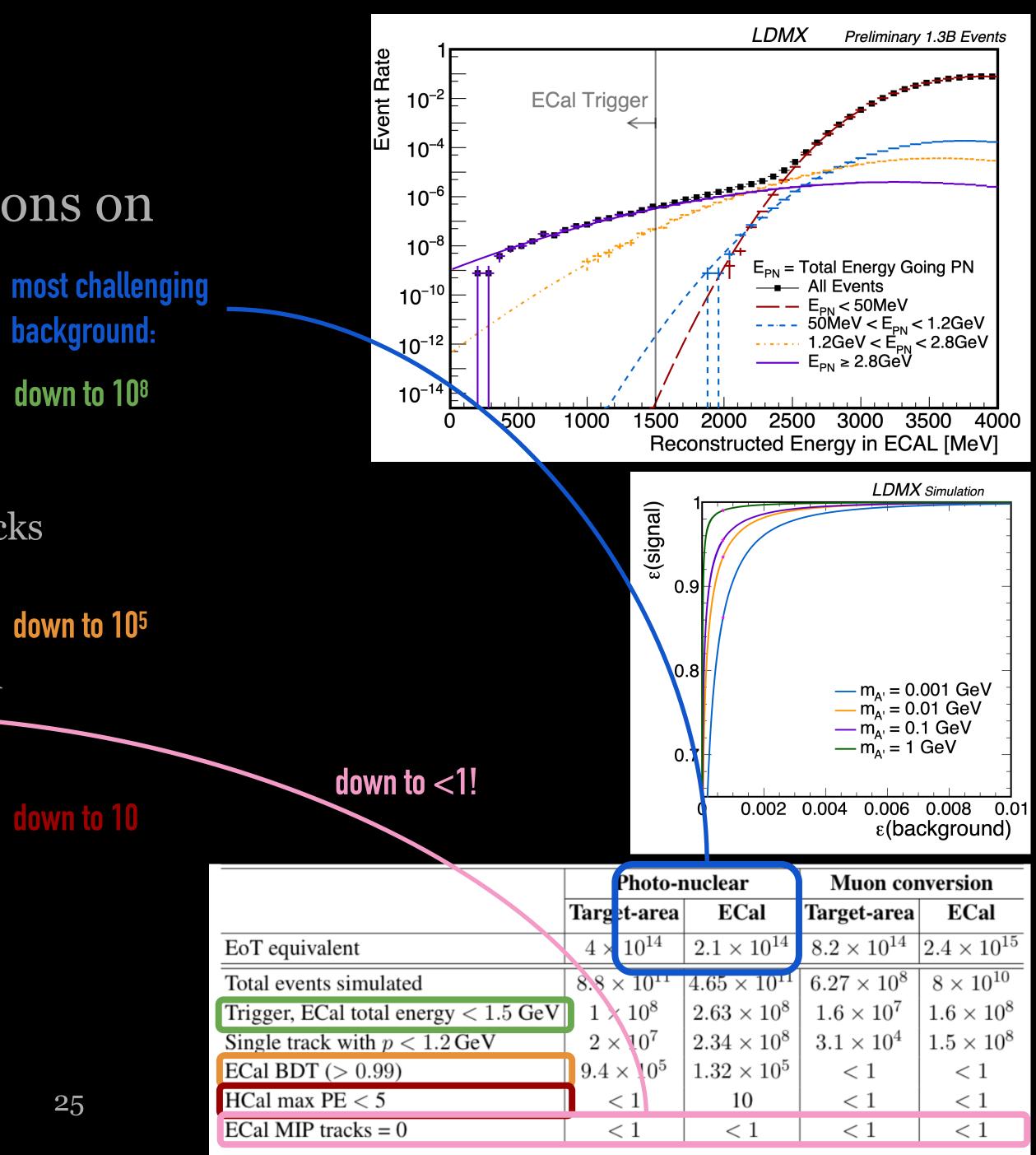
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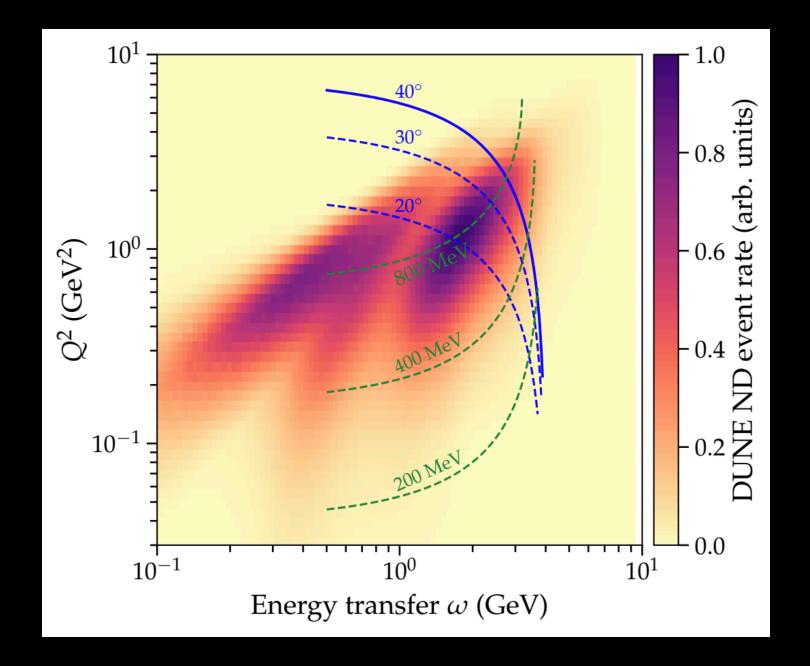
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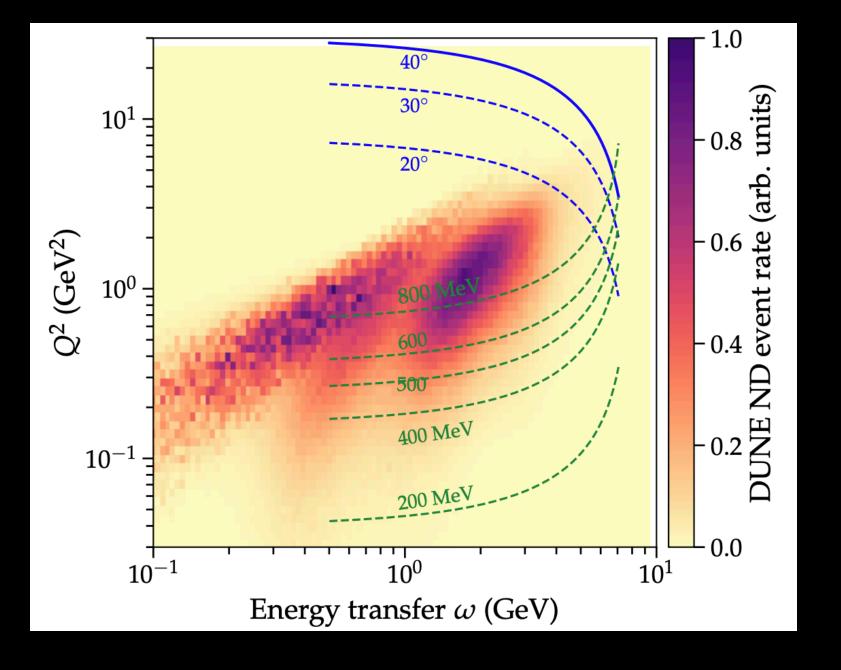
# Electronuclear measurements at 4 and 8 GeV



# Curves show LDMX acceptance, color scale represents event rate for DIS events in DUNE, indicating the relevant phase space







# Prototype at CERN Testbeam

# March-April 2022

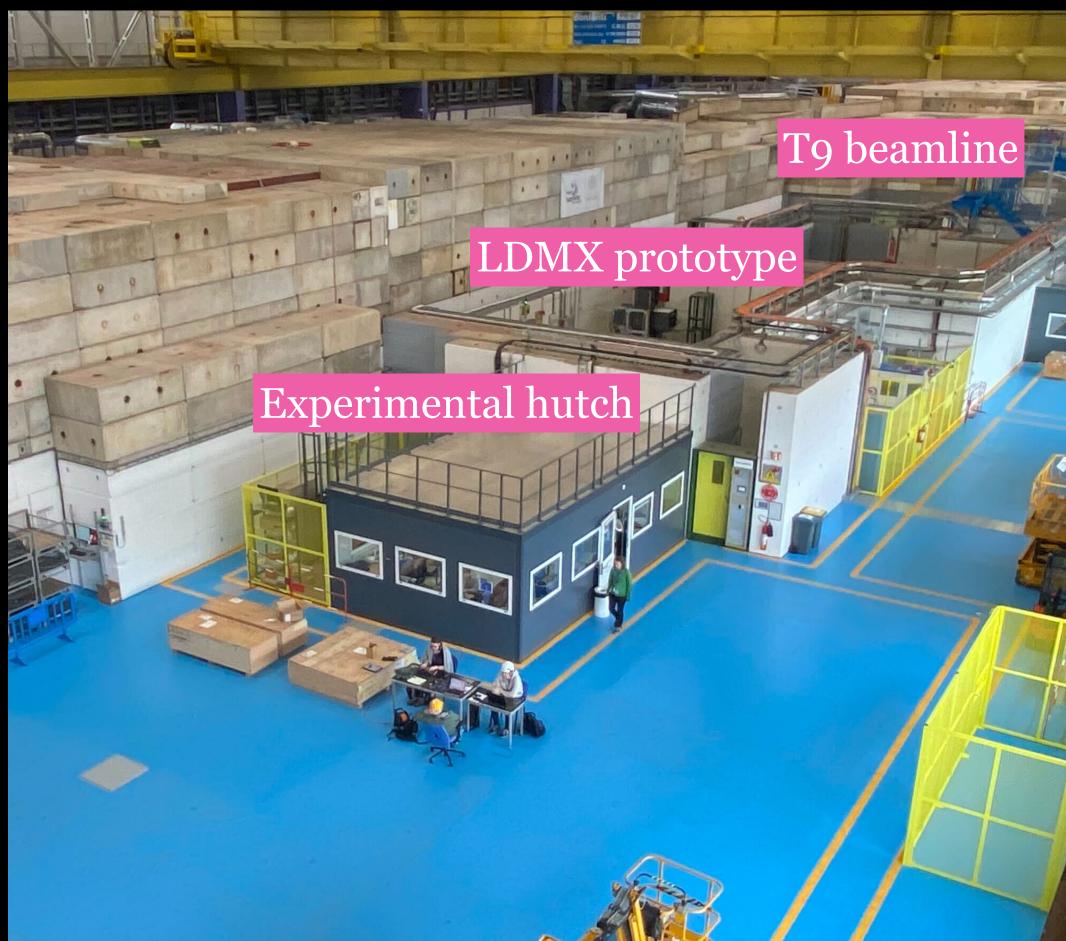


# CERN East Area: T9 beamline

- PS protons  $\rightarrow$  East Area
- Beam via North Target to T9: e,  $\mu$ ,  $\pi$ 
  - Beamline's configuration isolates final particle species from secondary beam
    - ~1k particles/spill
    - Particle ID: Cherenkov detectors
- Maximum intensity: Few million particles/spill

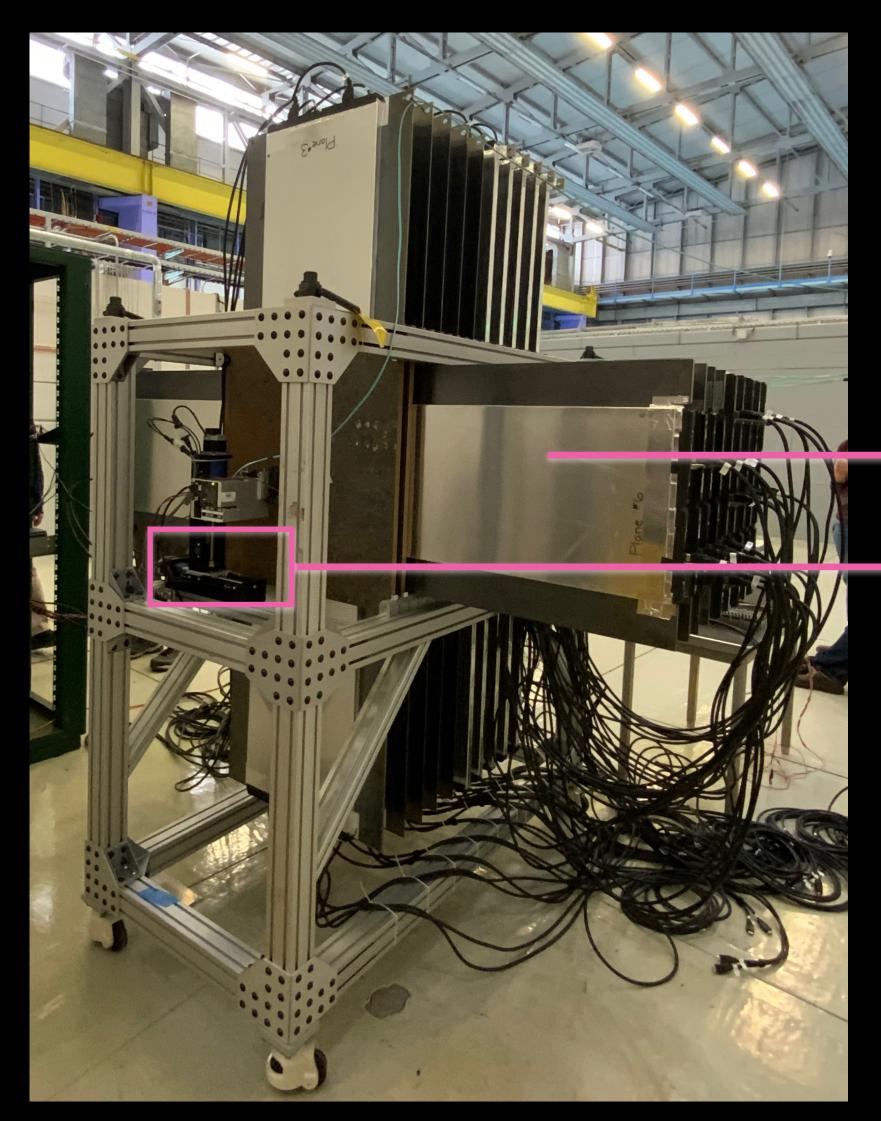








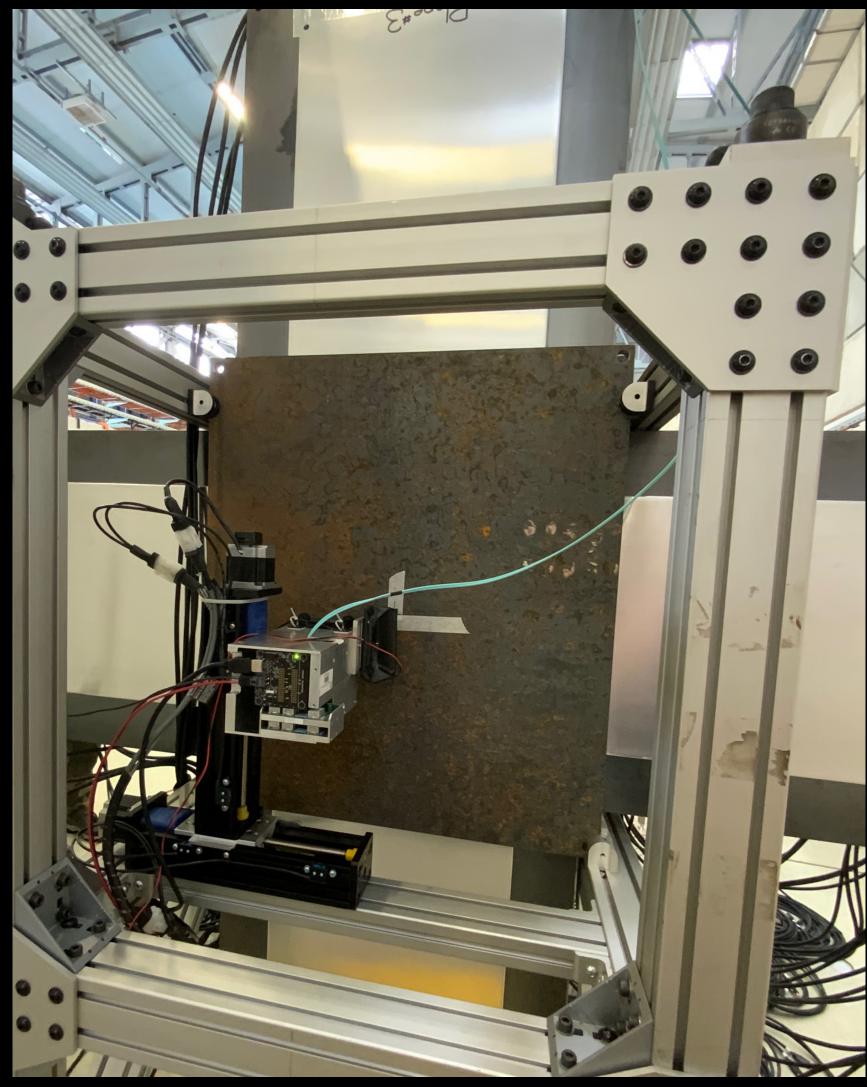
# Prototype in the beam area



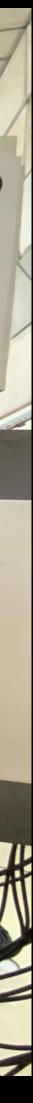
Hadronic Calorimeter (HCal) Trigger scintillator (TS)

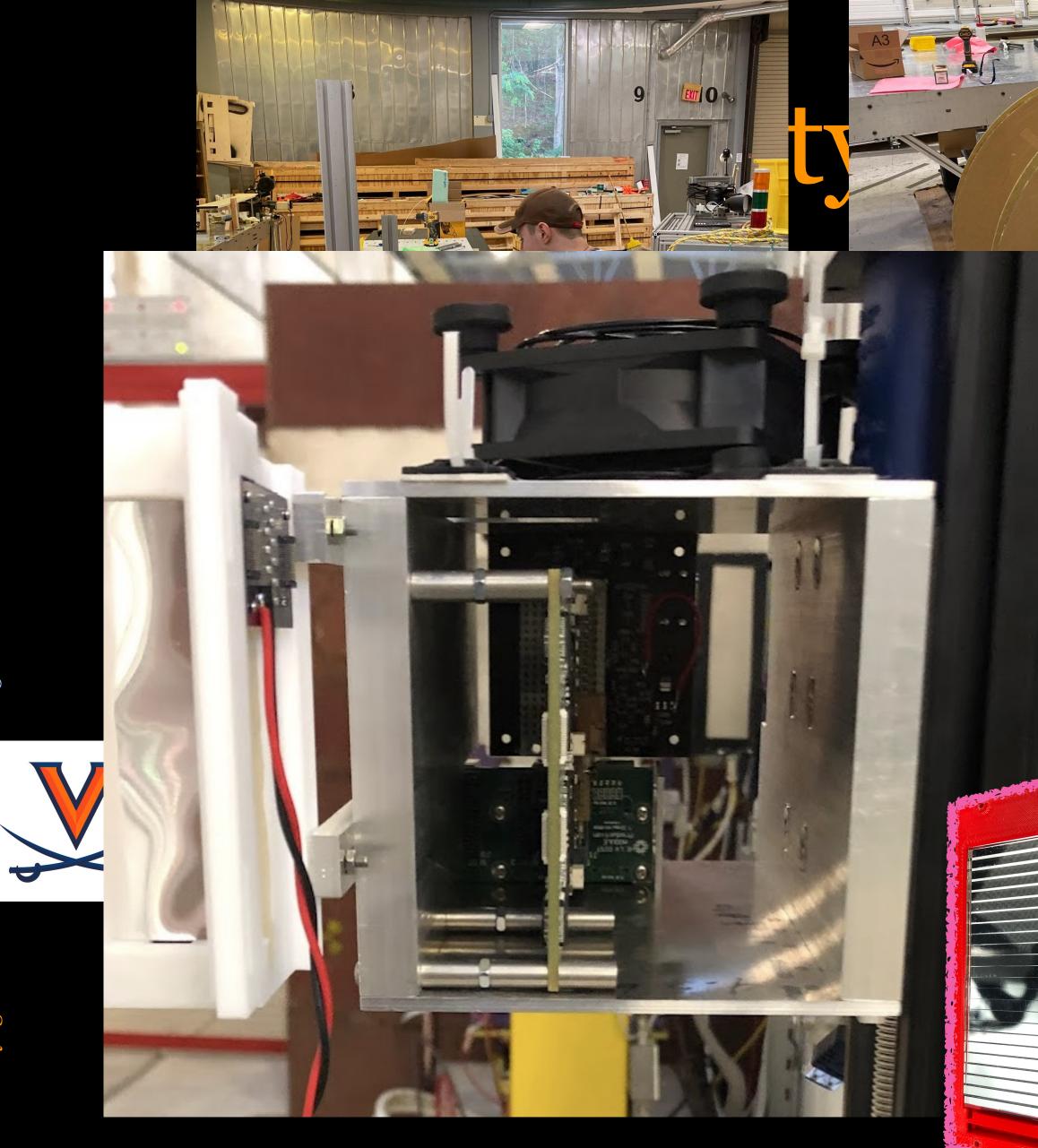






CERN test beam, March-April 2022





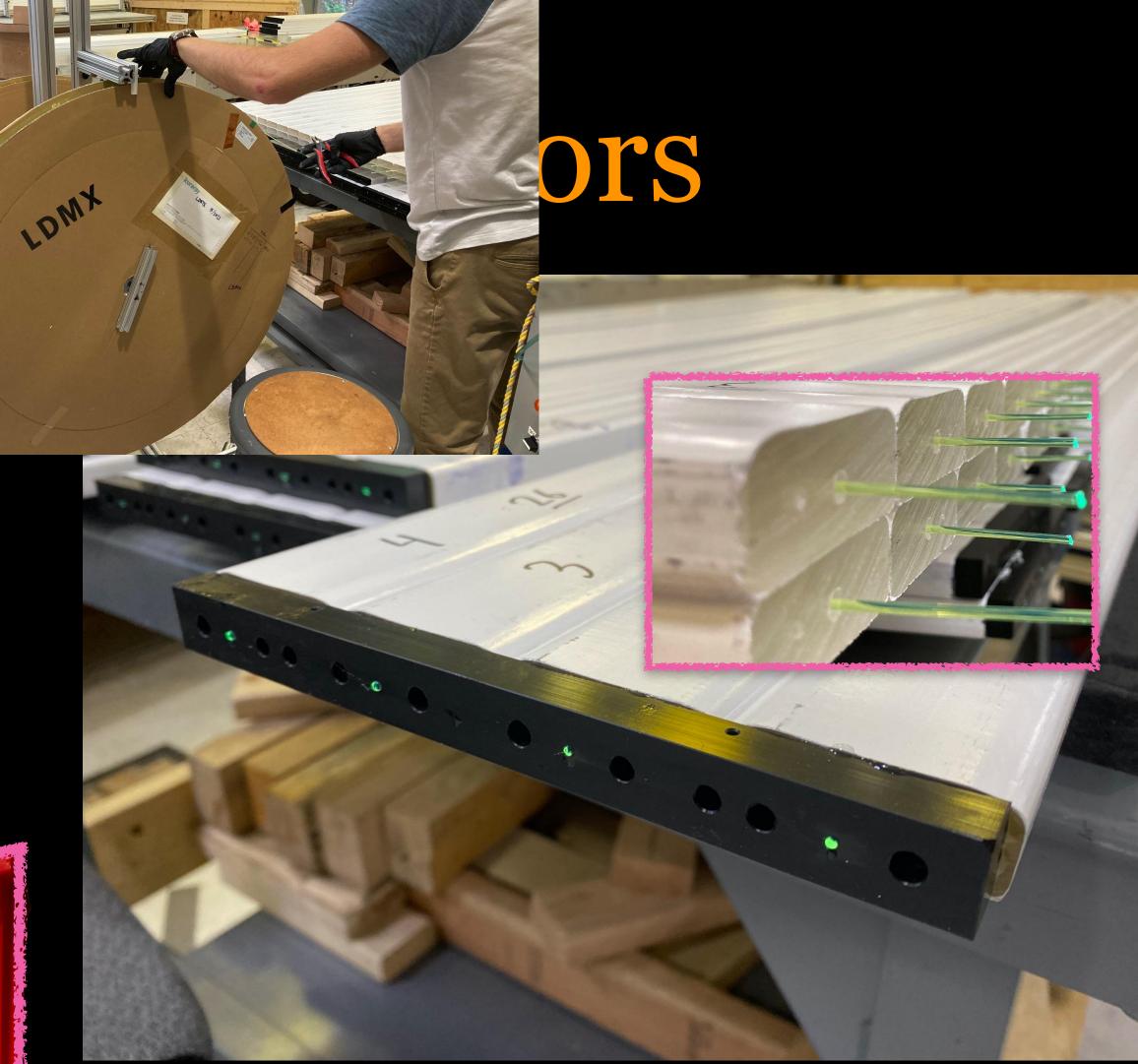
#### Trigger scintillator prototype

Inset: Aluminium-coated bare scintillator bar array in its 3D-printed holder

2023

LK

XMC

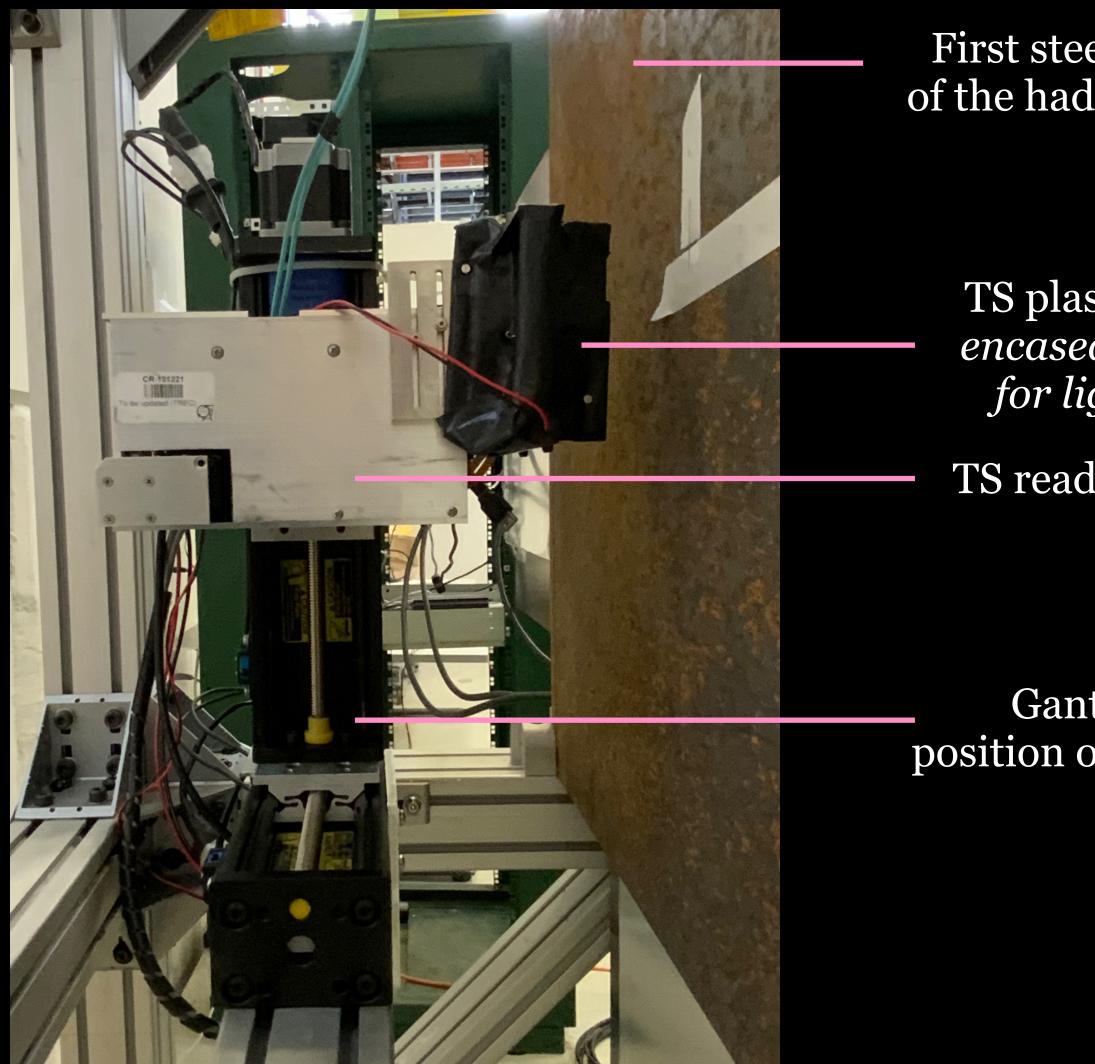


#### Hadronic calorimeter scintillator bars layer

Inset: Fibre optic cable pokes through bare scintillator bar



# Trigger scintillator (TS) prototype



TeVPA 2023

LK

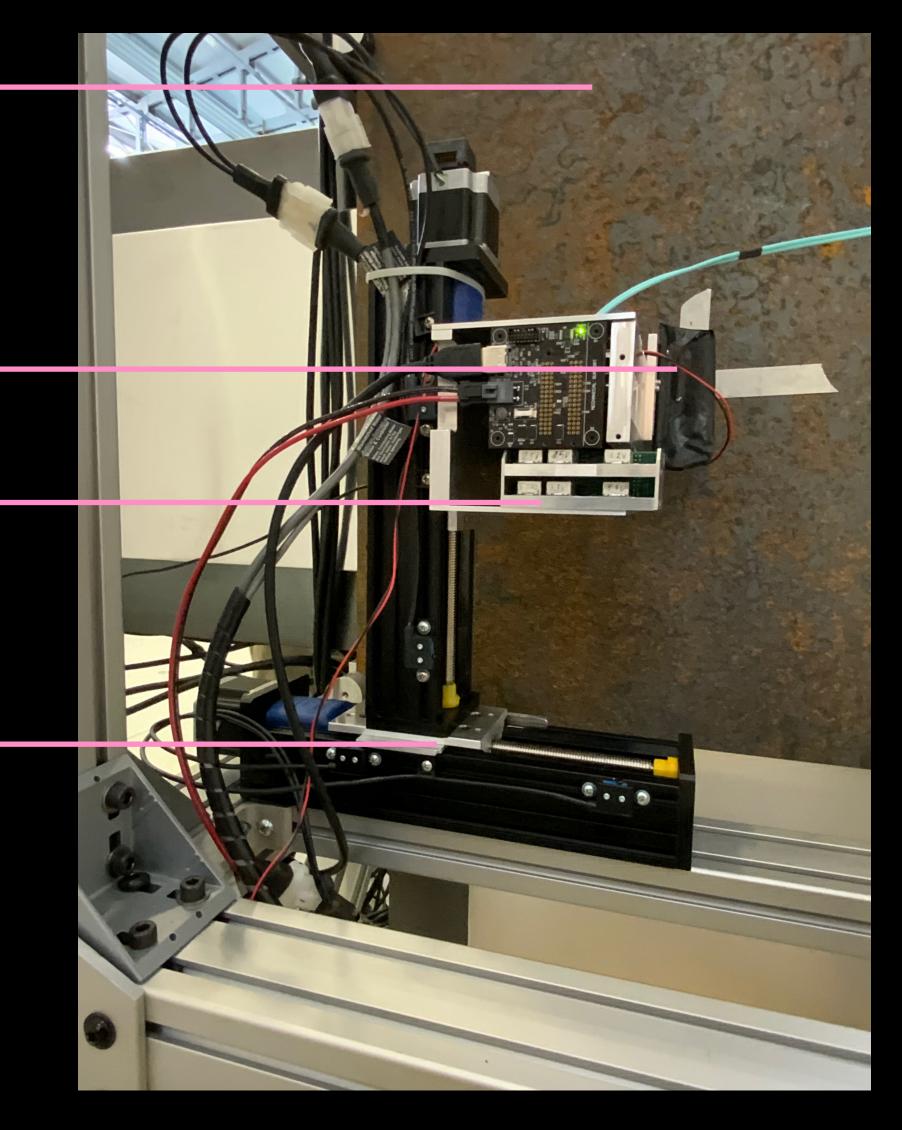
XMGT

First steel absorber layer of the hadronic calorimeter

> TS plastic scintillator encased in black tape for light tightness

TS readout electronics

Gantry to adjust position of TS in beamspot



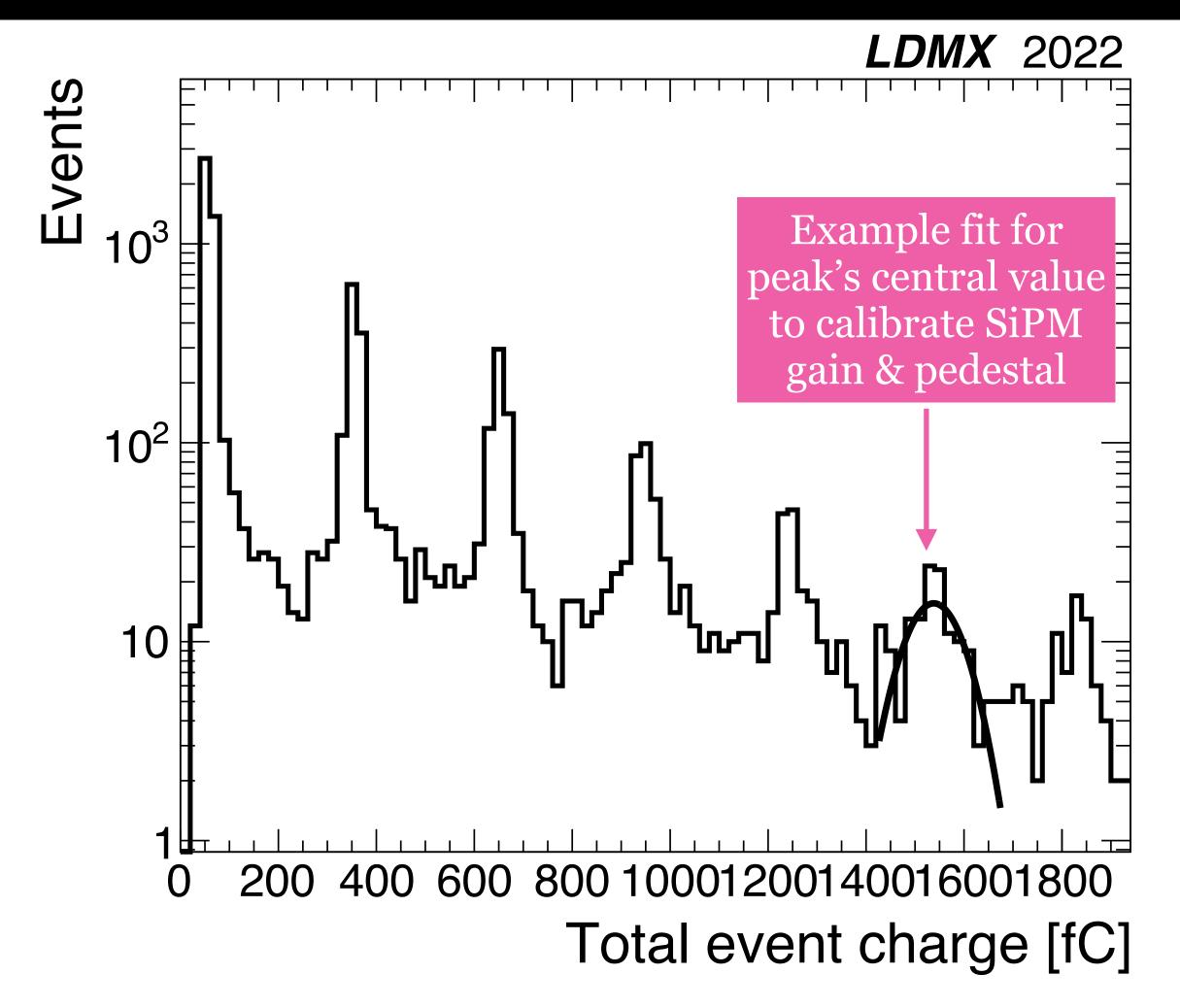
## TS: Single photoelectron spectrum

### Gain calibration

- Integrated charge/event for each TS channel
- Peaks
  - 1st: System pedestal
  - Additional: integer numbers of Si photomultiplier pixels firing







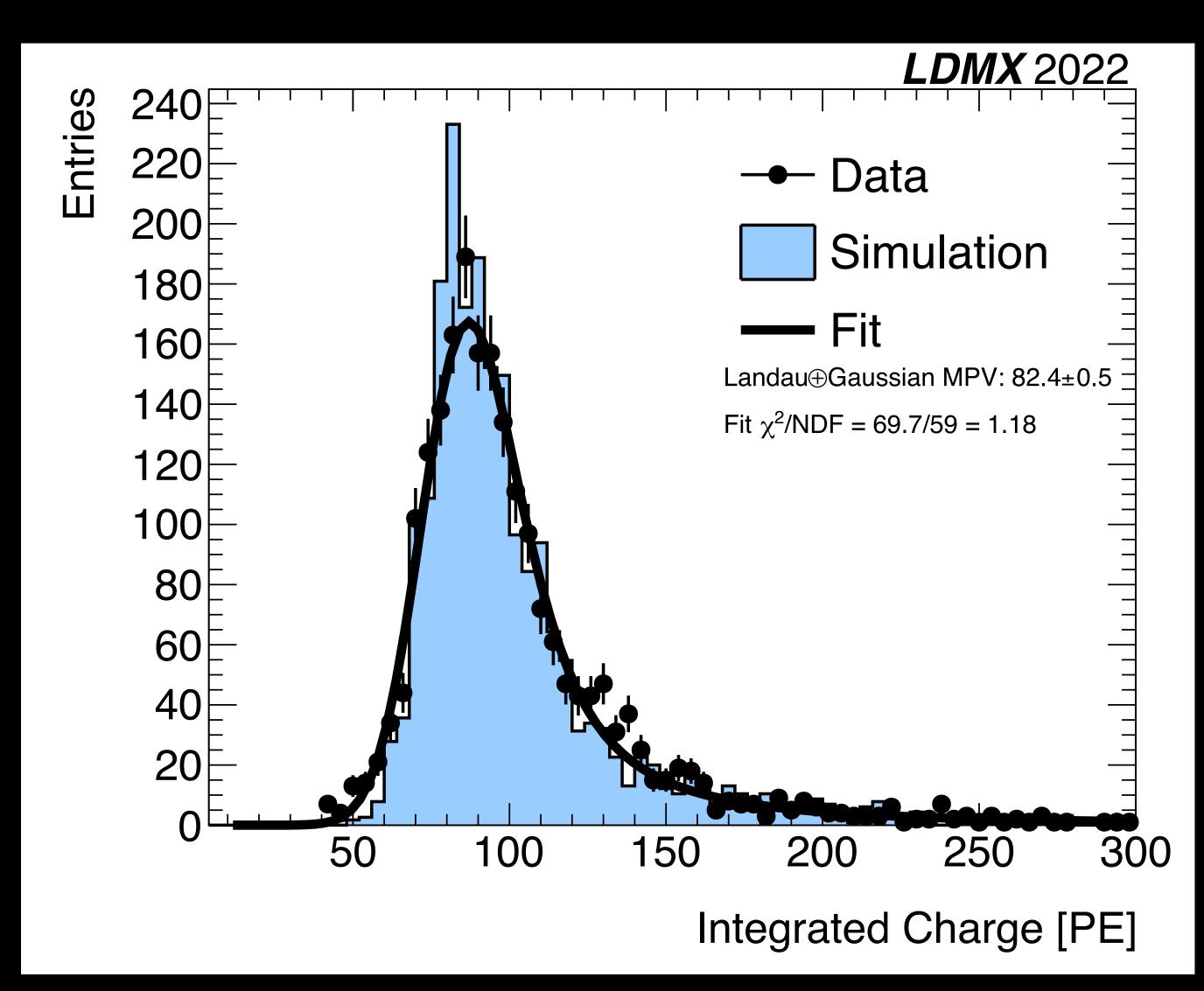
## TS: Plastic MIP response

### 4 GeV electron beam

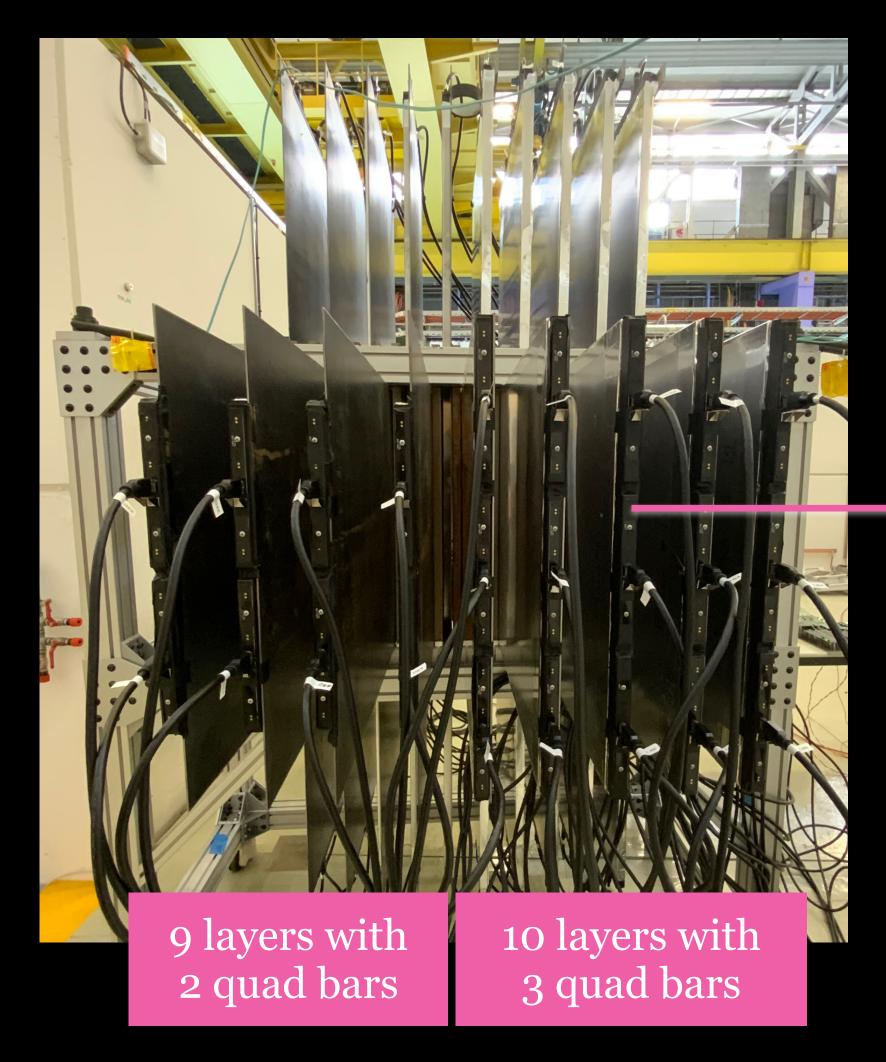
- Amplitude: Sum of charge measured over several time samples
  - Normalized to 1 photoelectron equivalent
- Most probable value of channel's response to MIP = 82 photoelectrons
  - Model: Landau + Gaussian convolution



CERN test beam, March-April 2022



## Hadronic calorimeter (HCal) prototype 19 alternating layers, usually<sup>1</sup> Al cover • scintillator bars • steel absorber plate



6 HGCRoc boards (384 total channels; 64 per board) required for readout

HDMI cable

with readout Steel absorber

Readout manifold Scintillator bars Aluminium cover



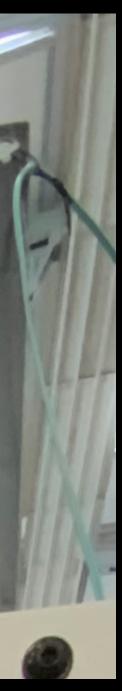




Example: Section of a vertical layer of the HCal

<sup>1</sup> First HCal layer is steel absorber, then scintillator bars CERN test beam, March-April 2022







## HCal: MIP response

### 4 GeV muon beam

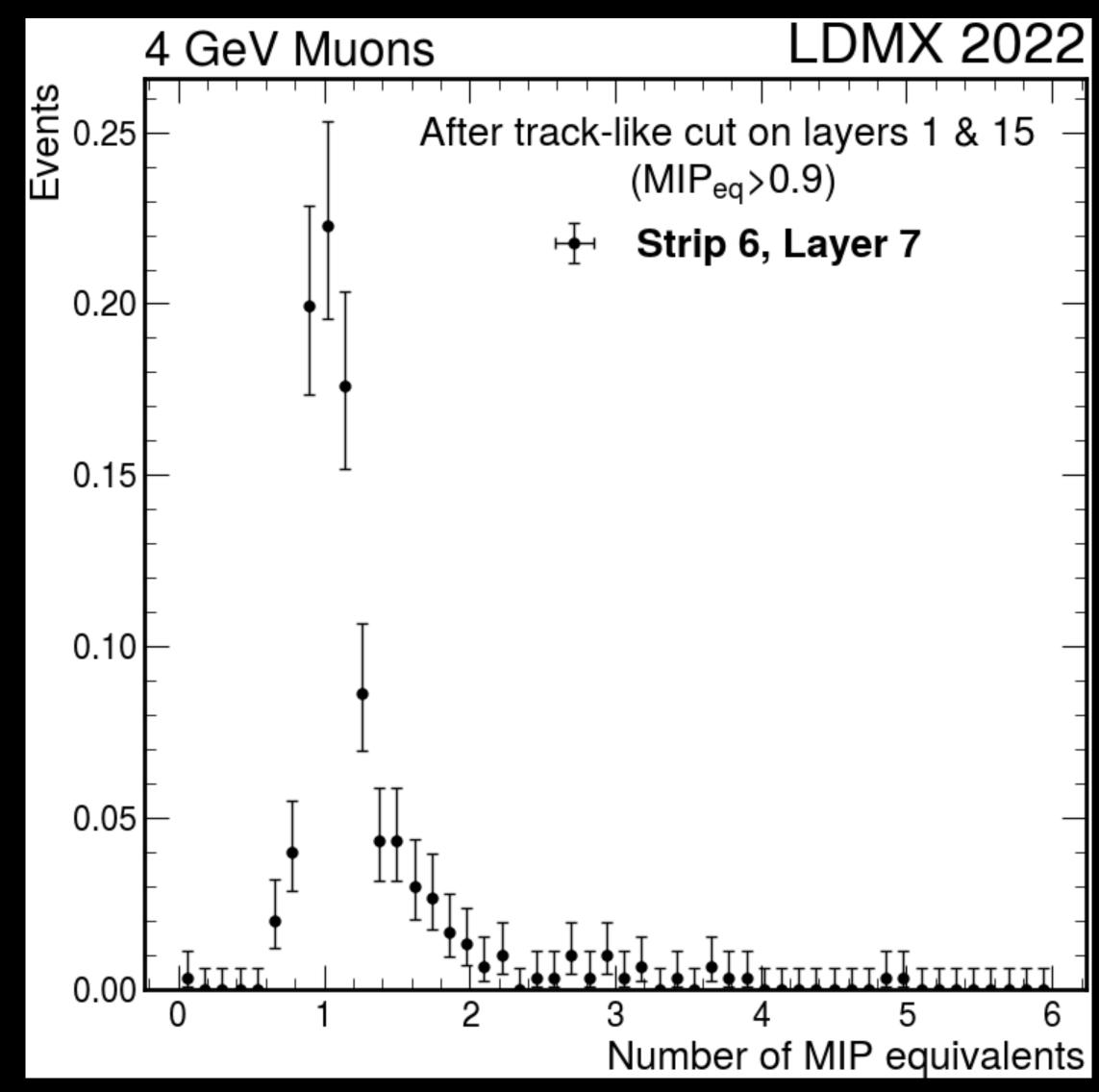
Sum of ADC counts in a single layer and strip of HCal prototype

#### $\sum$ ADC counts N<sub>MIPeq</sub> Measured value for 1 MIP

Require MIP-like signature in entrance & exit of HCal

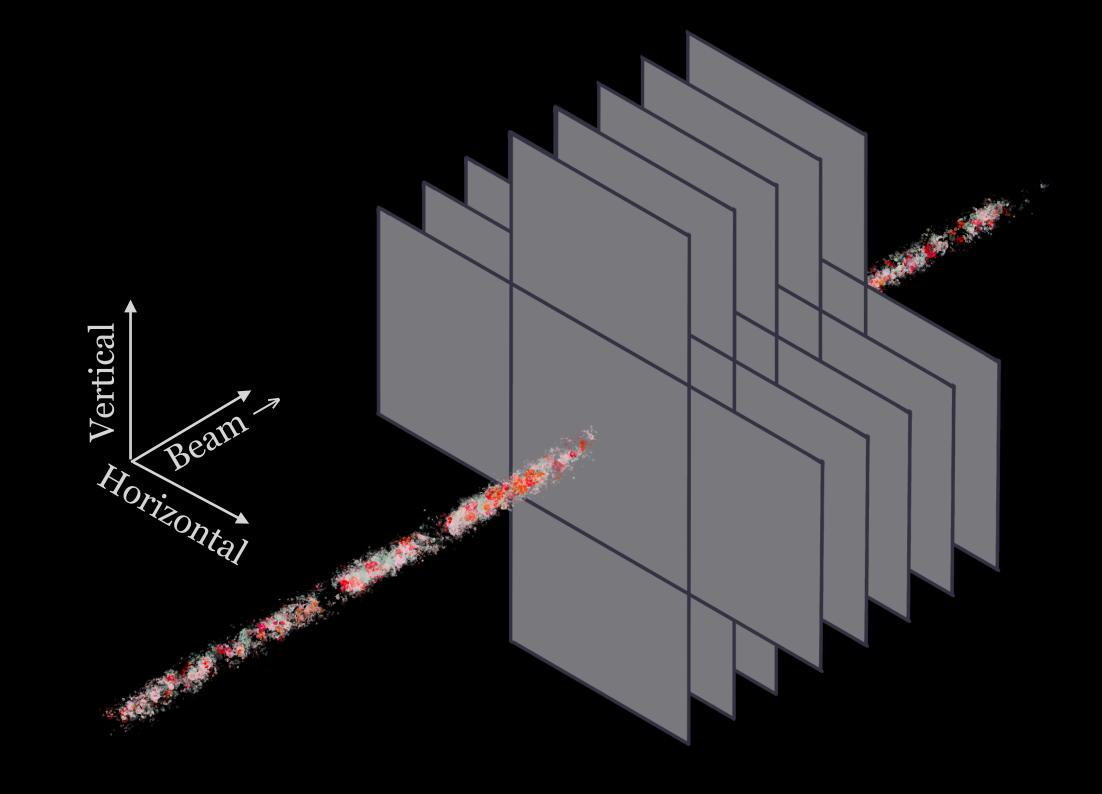






CERN test beam, March-April 2022



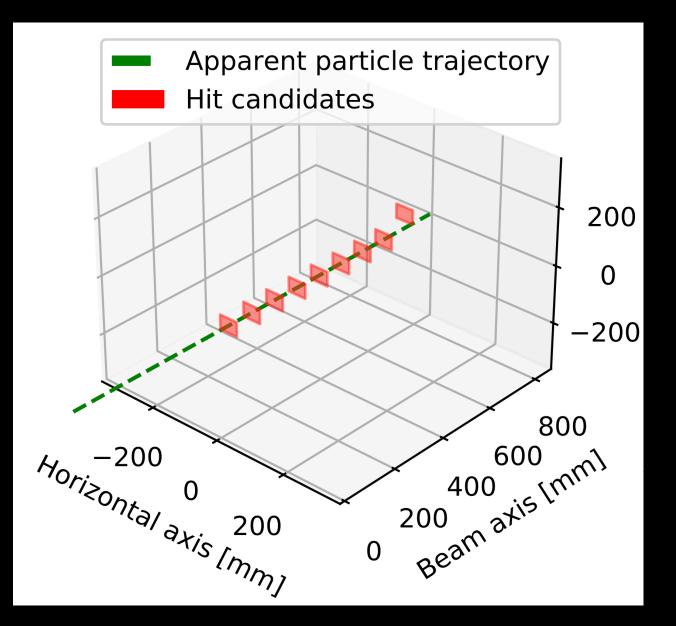


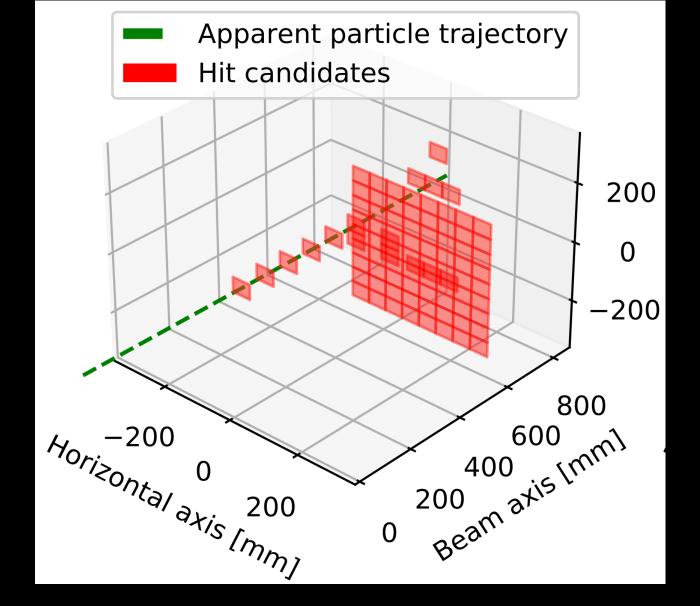
TeVPA 2023



#### MIP candidate

Sequential, crisp signature in HCal





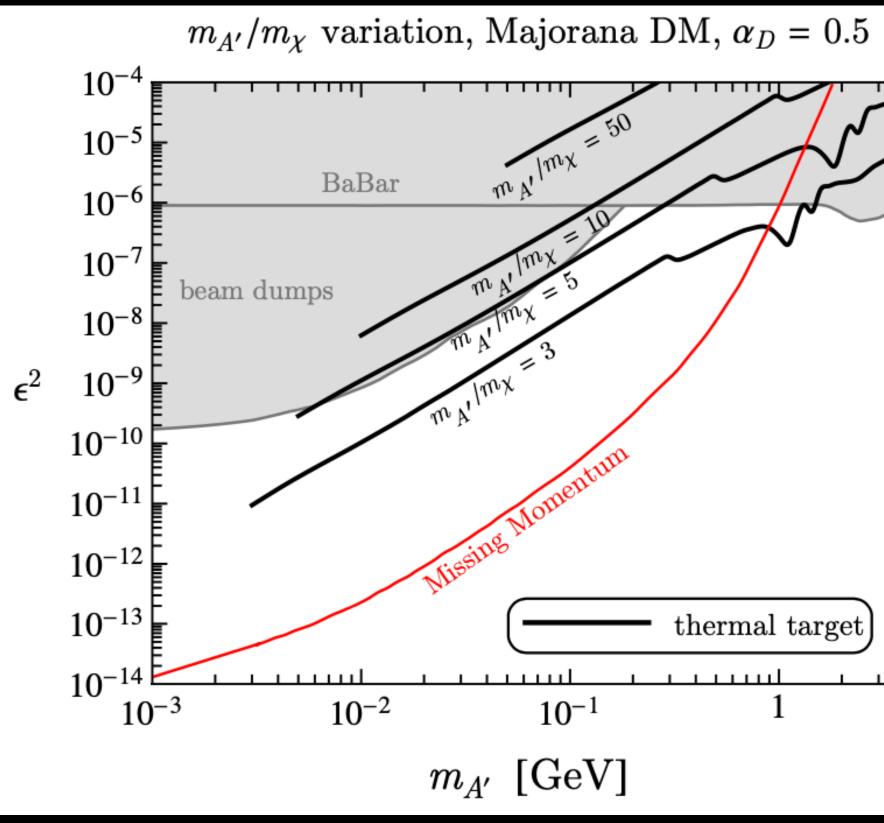
#### Pion candidate

MIP-like deposits followed by cloud in HCal







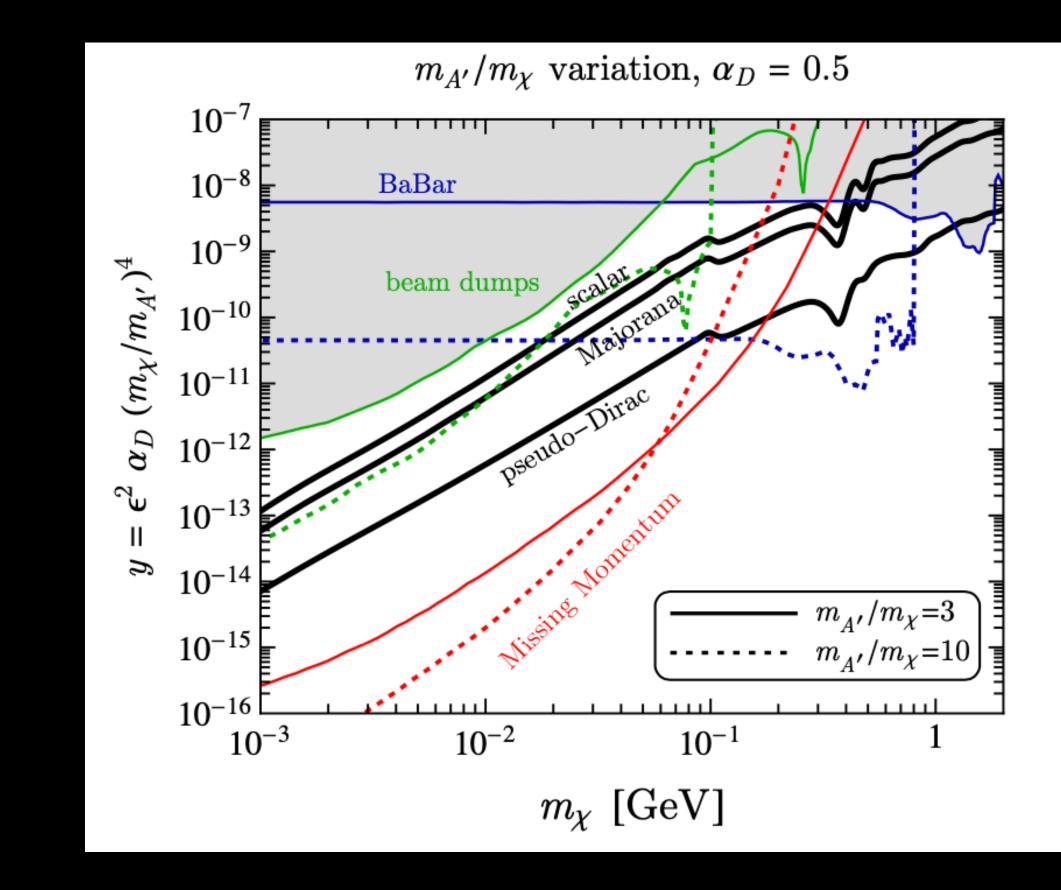


Expressed in  $\epsilon^2$ , thermal curves move, accelerator line stays



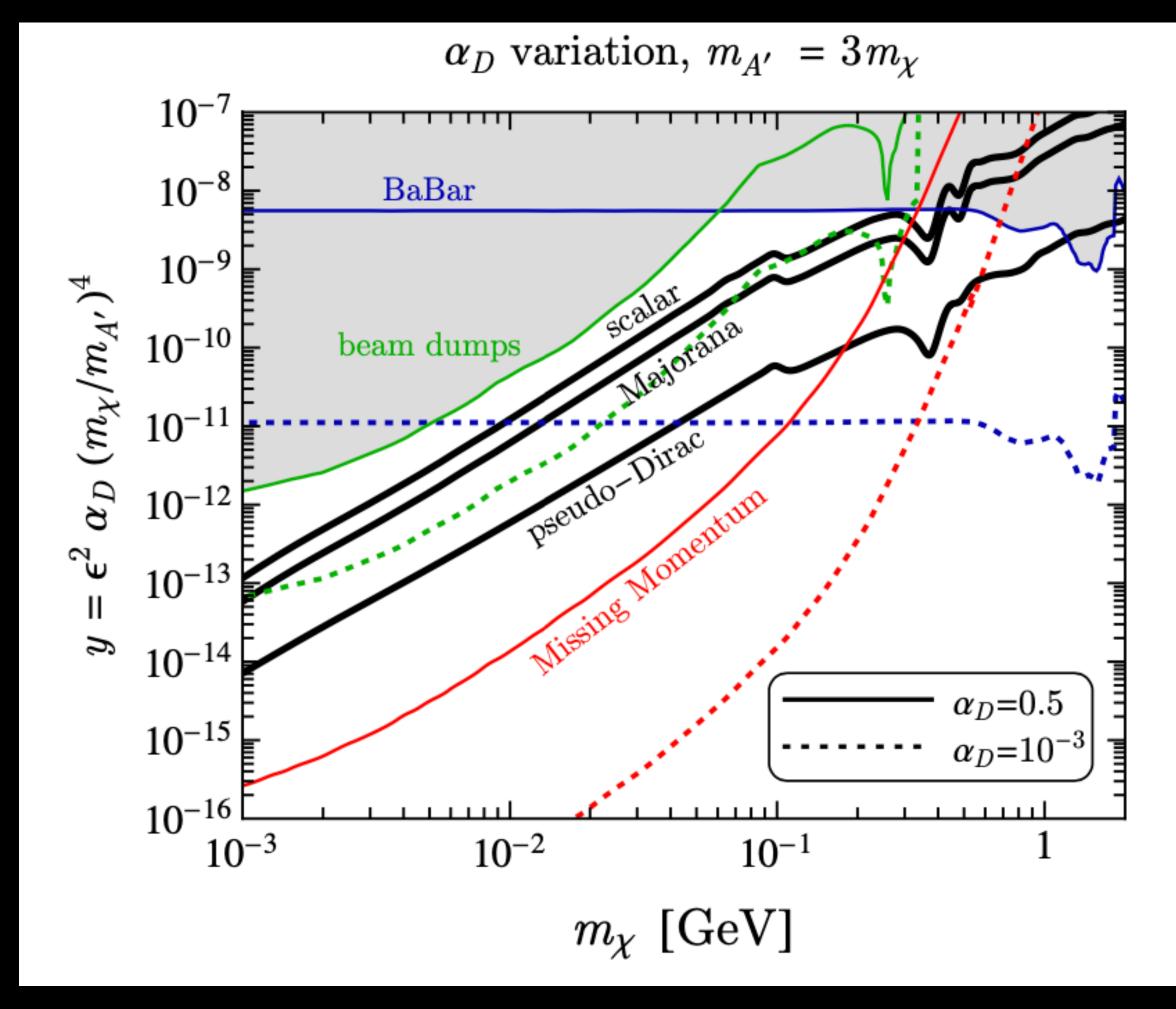


## Varying mass ratio



Expressed in *y*, thermal curves stay, accelerator lines tilt





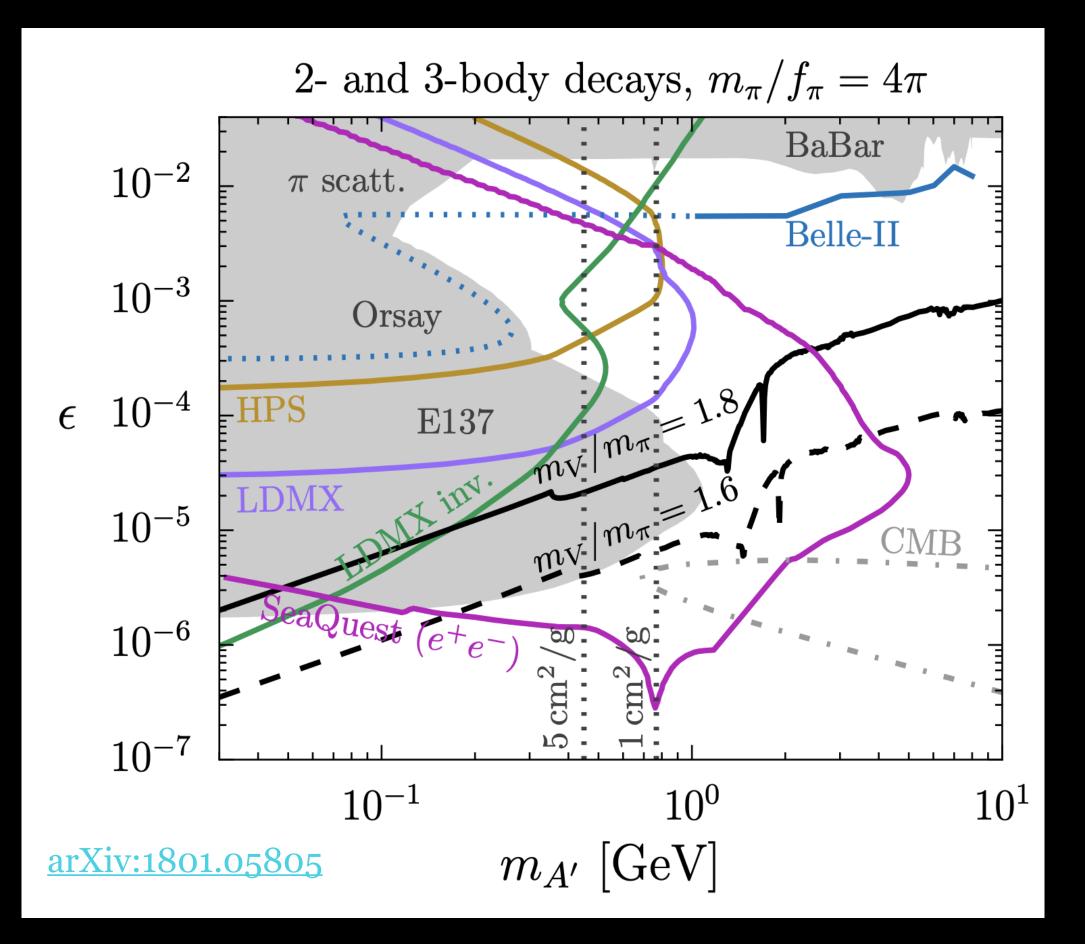
LK

XMC

## Varying coupling

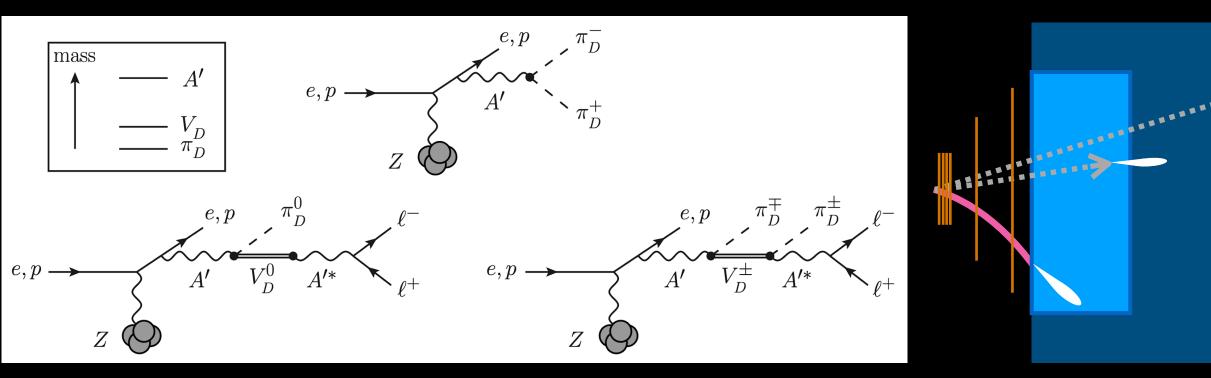
**Bottom:** parameter space in the y vs.  $m_{\chi}$  plane where the solid curves are identical to those shown in Fig. 5 (with  $\alpha_D = 0.5$ ), but the dotted curves show how the constraints and projections vary for the choice  $\alpha_D = 10^{-3}$ . For fixed values of y, a smaller  $\alpha_D$  requires a larger  $\epsilon^2$  (i.e. larger mediator coupling), which makes that parameter point *easier* to constrain. Hence, accelerator sensitivity generally improves in the yvs.  $m_{\chi}$  plane for smaller  $\alpha_D$ . Note that the thermal freeze-out curves in this plane are identical for both values of  $\alpha_D$  shown here because the thermal abundance scales with y.

## Expected sensitivity: visible signatures (example)







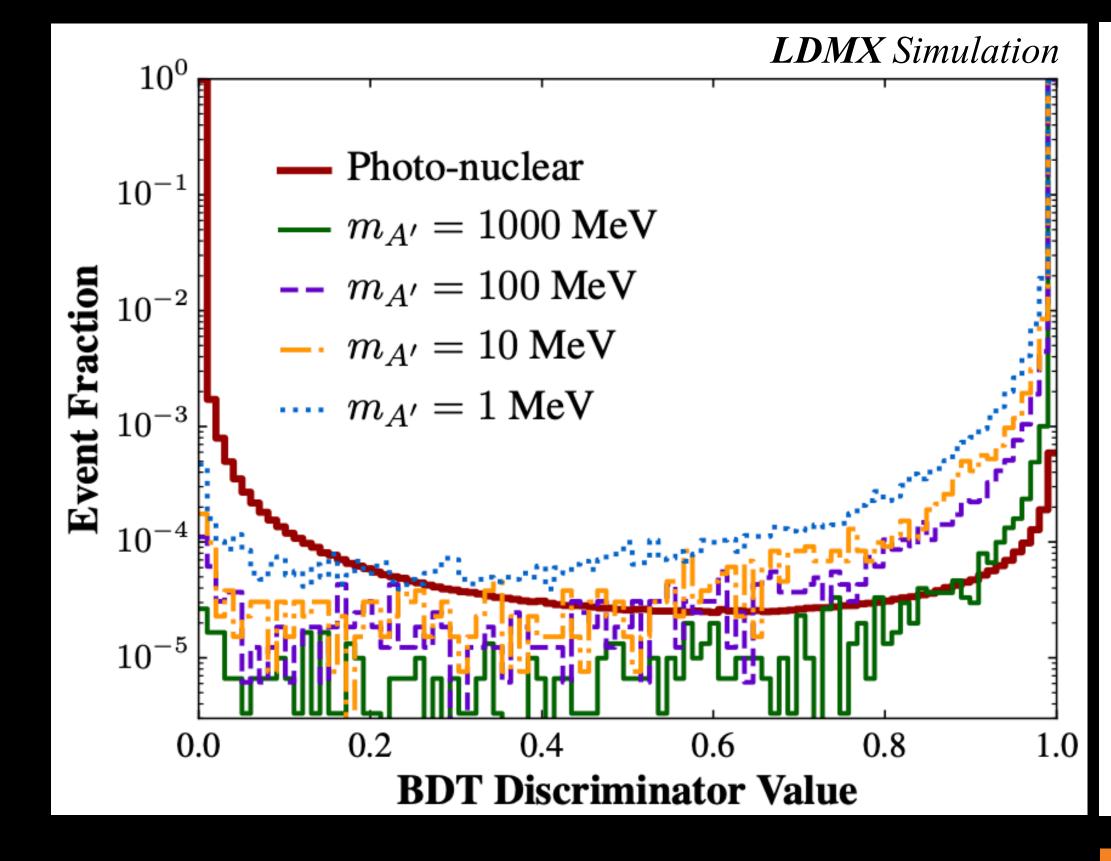


Example displaced decay model: Strongly Interacting Massive Particles (SIMPs)

- slightly more complicated (indirect) DM depletion mechanism
- *A*' decays to hidden sector pions
- delayed lepton pair production
- → missing momentum or with displaced decay activity

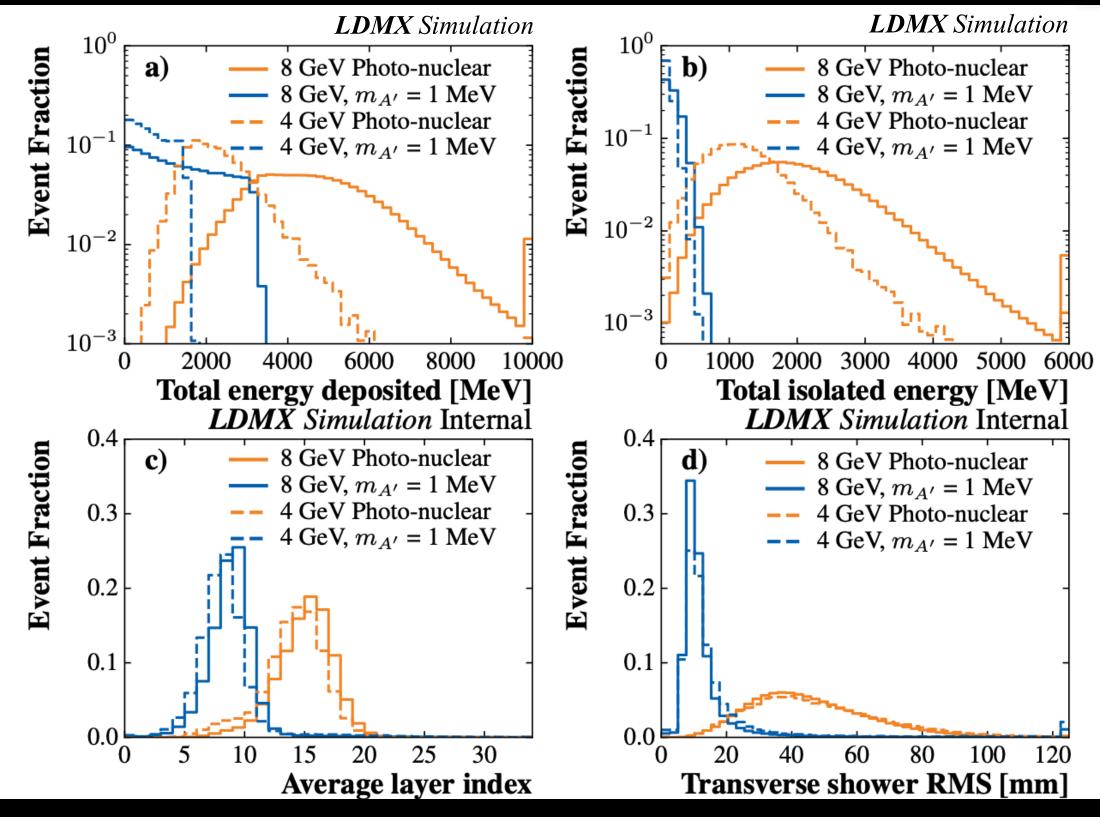


# BDT output, and input distributions for key ECal variables at 4 and 8 GeV





XMO



Photonuclear interactions Signal process

## Simulation cutflow at 4 and 8 GeV

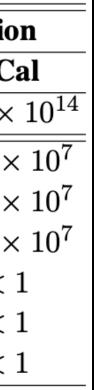
	Photo-nuclear		Muon conversion			Photo-nuclear		Muon conversion	
	Target-area	ECal	Target-area	ECal		Target-area	ECal	Target-area	ECa
EoT equivalent	$4 \times 10^{14}$	$2.1 \times 10^{14}$	$8.2 \times 10^{14}$	$2.4 \times 10^{15}$	EoT Equivalent	$2.00  imes 10^{14}$	$2.00\times10^{14}$	$2.00\times10^{14}$	$2.00 \times$
Total events simulated	$8.8\times10^{11}$	$4.65\times10^{11}$	$6.27 \times 10^{8}$	$8 \times 10^{10}$	Trigger (front ECal energy < 3160 MeV)	$7.57 \times 10^7$	$4.43 \times 10^{8}$	$2.37  imes 10^7$	$8.12 \times$
Trigger, ECal total energy $< 1.5~{\rm GeV}$	$1 \times 10^8$	$2.63 \times 10^8$	$1.6 \times 10^7$	$1.6 \times 10^8$	Total ECal energy $< 3160 \text{ MeV}$	$2.73  imes 10^7$	$7.27 \times 10^{7}$	$1.76 \times 10^7$	$6.06 \times$
Single track with $p < 1.2 \mathrm{GeV}$	$2 \times 10^7$	$2.34 \times 10^8$	$3.1 \times 10^4$	$1.5 \times 10^8$	Single track with $p < 2400$ MeV/c	$3.03  imes 10^6$	$6.64 \times 10^7$	$5.32 \times 10^4$	$5.69 \times$
ECal BDT ( $> 0.99$ )	$9.4 \times 10^5$	$1.32 \times 10^5$	< 1	< 1	ECal BDT (85% eff. $m_{A'} = 1 \text{ MeV}$ )	$1.50 \times 10^5$	$1.04 \times 10^{5}$	< 1	< 1
HCal max $PE < 5$	< 1	10	< 1	< 1	HCal max $PE < 8$	< 1	2.02	< 1	< 1
ECal MIP tracks $= 0$	< 1	< 1	< 1	< 1	ECal MIP tracks $= 0$	< 1	< 1	< 1	< 1

4-GeV beam energy

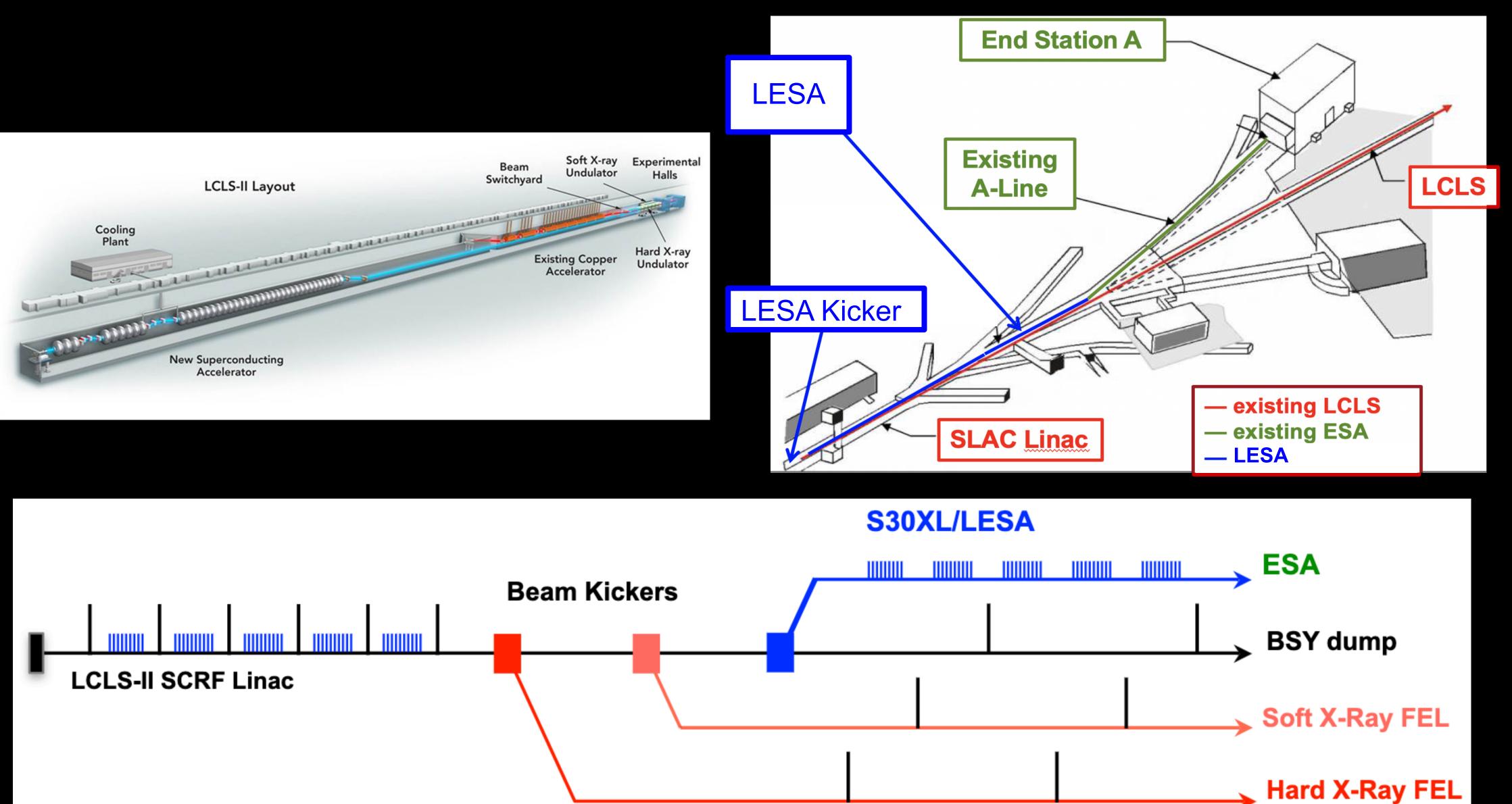


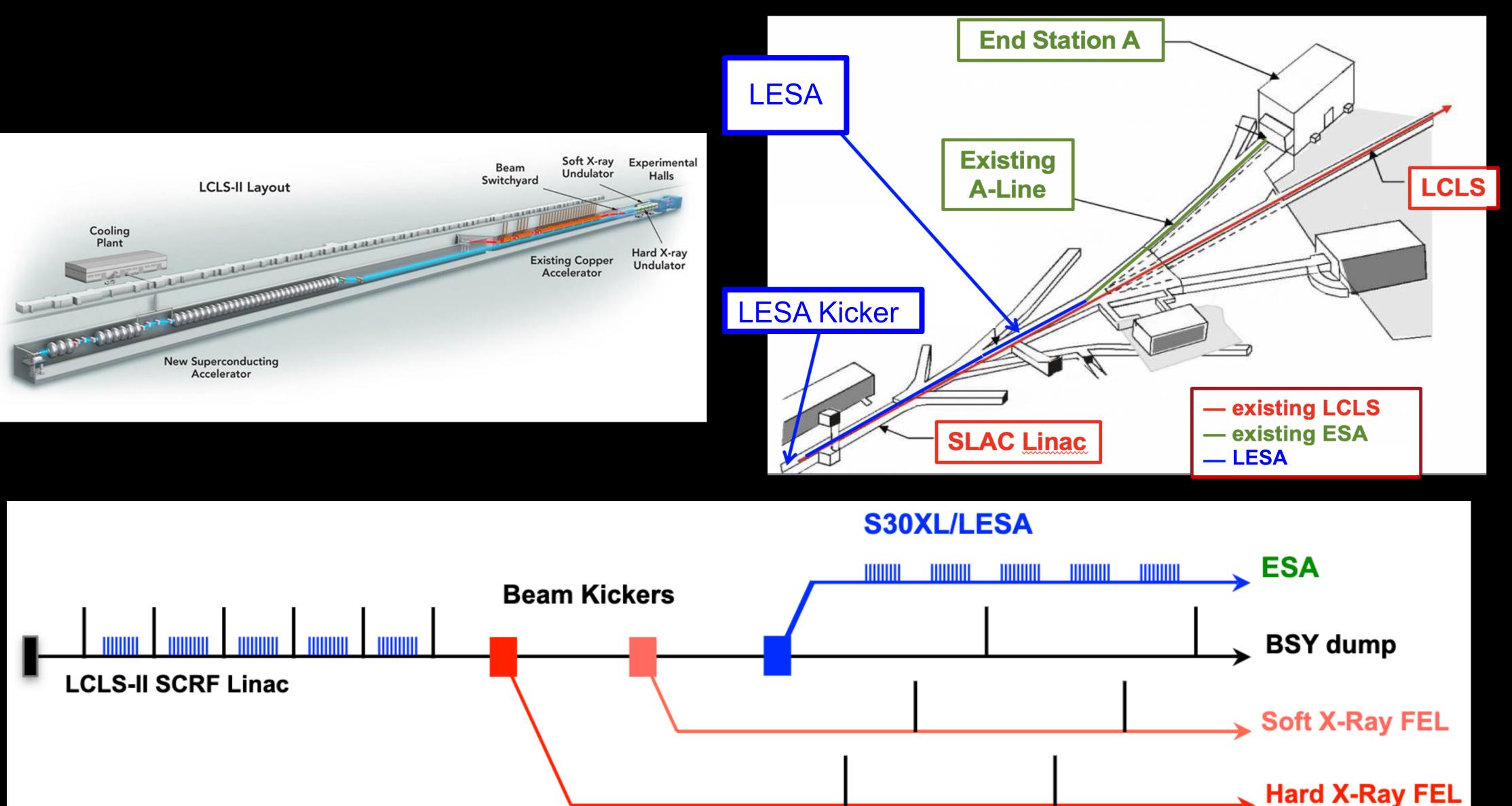


8-GeV beam energy



### GeV-scale electron beam at SLAC



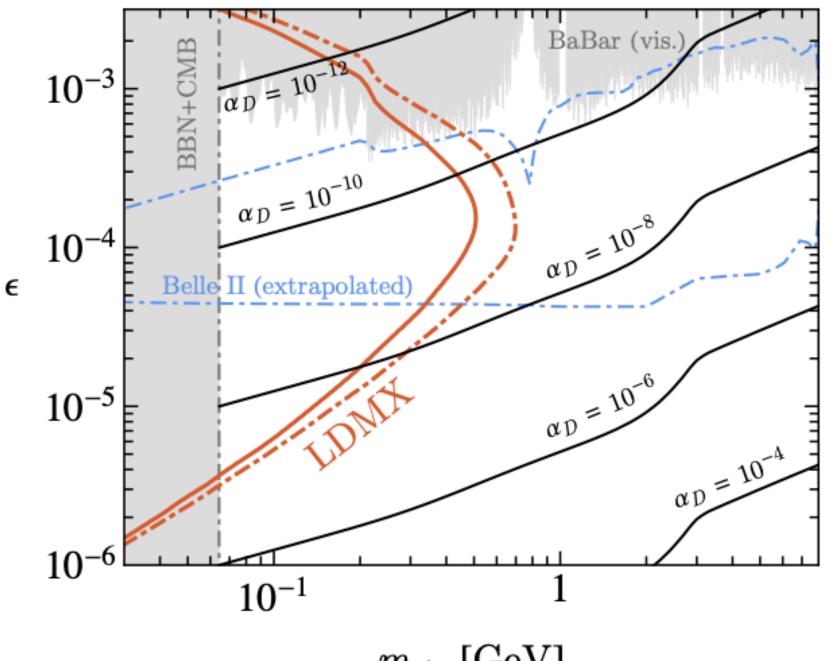












 $ab^{-1}$  assuming statistics limitation only [1, 87].





## Freeze-1n

#### Low–Reheat Freeze–In, $m_{A'}~=15~T_{\rm RH},~m_{\chi}=10~{\rm keV}$

 $m_{A'}$  [GeV]

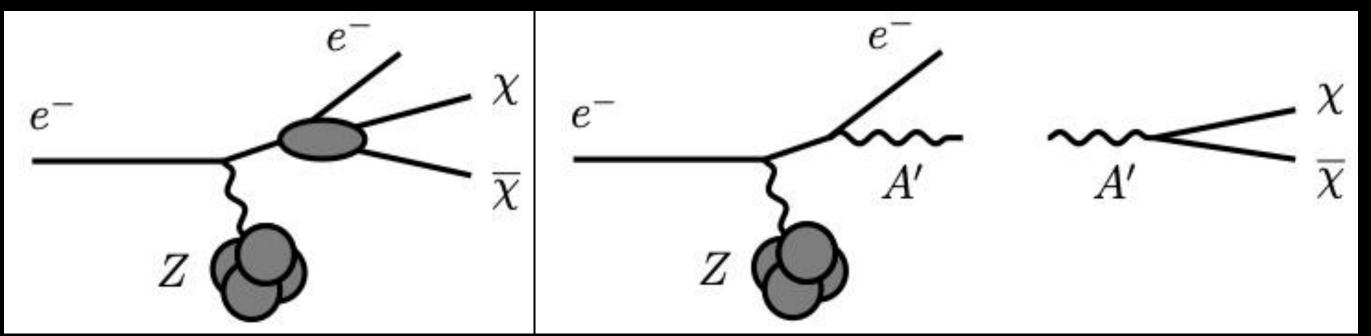
FIG. 14: LDMX sensitivity to the freeze-in scenario with a heavy dark photon and low-reheat temperature. The projected reach of LDMX is shown as the solid red (dashed-dotted red) line for a tungsten (aluminum) target and a 8 (16) GeV beam. The correct relic abundance is obtained along the black contours for different choices of  $\alpha_D$ . The gray shaded regions are excluded by the BaBar resonance search [19] and by cosmological constraints on low reheating temperatures [139]. We also show the projected sensitivity of the Belle II monophoton search (blue dot-dashed) as computed by rescaling the 20 fb<sup>-1</sup> background study up to 50

Phys. Rev. D 99, 075001 (2019), https://arxiv.org/abs/1807.01730



## LDMX: a fixed-target missing momentum experiment

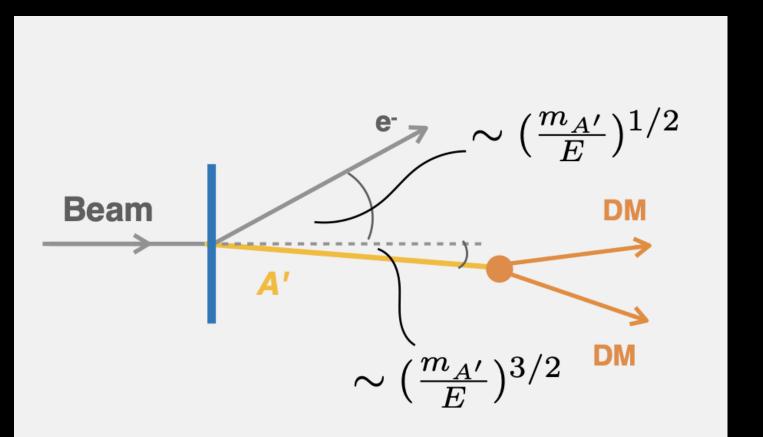
LDMX focuses on *escaping* dark matter:



- massive dark photon (*A*') bremsstrahlung in thin target
- agnostic when it comes to (invisible) fate of the A'
  - notice that energy goes missing  $\bigcirc$
- strategy: make "all" SM backgrounds appear in detector
- veto everything but low-activity events  $\bigcirc$







## What do we need from Dark Matter to produce it with a ~GeV electron beam?

- Has to be light
  - In the mass range of known stable particles can't be a classical WIMP
    - Has to interact through some other mediator
- Not too feeble to ever happen
  - Rate should respect present DM relic abundance, unitarity, ...

### And from us: a suitable beam and detection strategy

XMC