

**On the gamma-ray
emission from the core of
the Sagittarius dwarf
galaxy**

TeVPA 2023

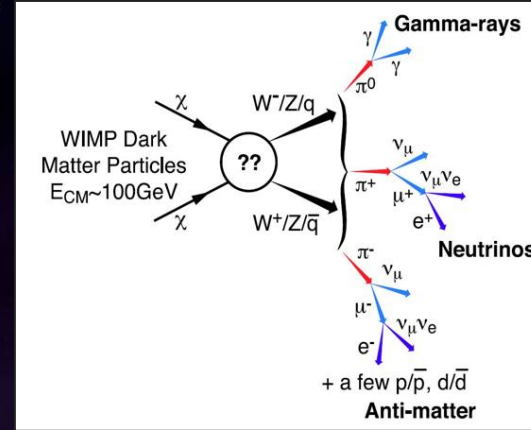
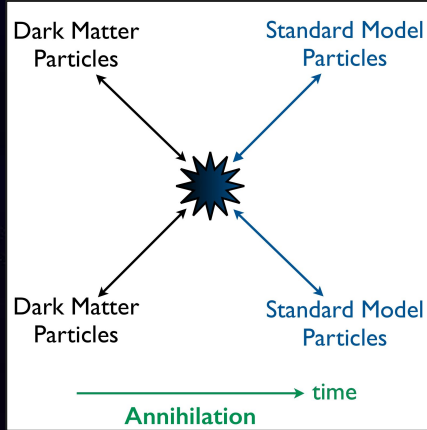
Addy J. Evans

Texas A&M University

with Louis Strigari, Pat Harding, Andrea Albert, Tim Linden, Dan
Hooper, Oskar Svenborn, and Andrew Pace

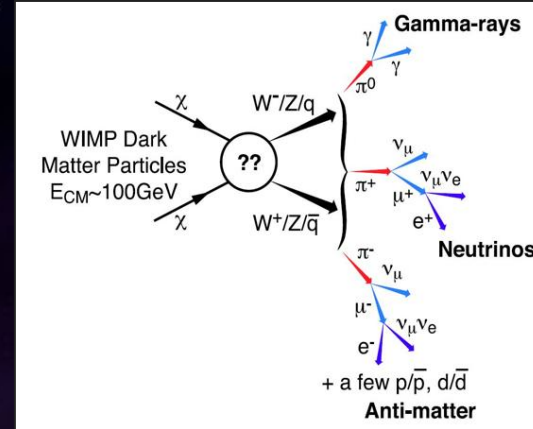
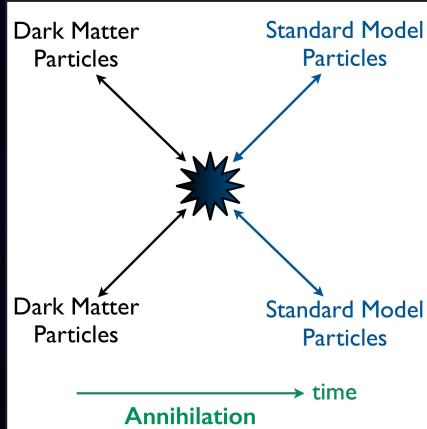
Astrophysical probes of dark matter: annihilation products

Assuming dark matter is a WIMP that self-annihilates:



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The goal is to find high-energy Standard Model particles that cannot be attributed to other astrophysical sources.

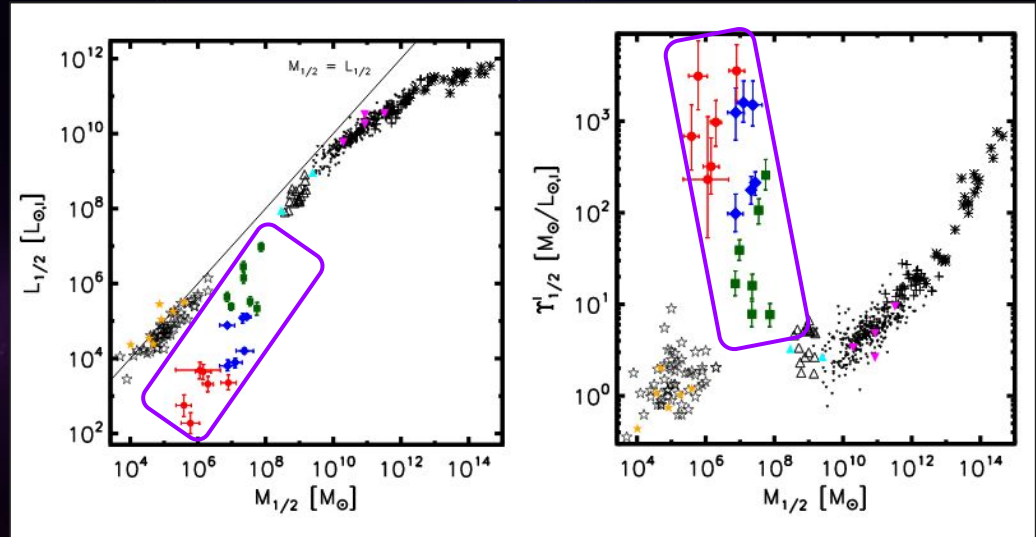
Indirect detection with dwarf galaxies

Dwarf galaxies: the most dark matter dominated galaxies in the Universe

→ Very high mass to light ratios

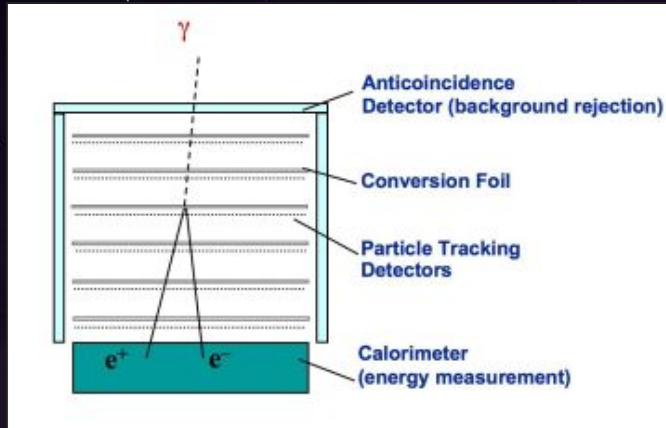
→ Fewer astrophysical gamma-ray sources

→ A great place to go looking for dark matter annihilation products!



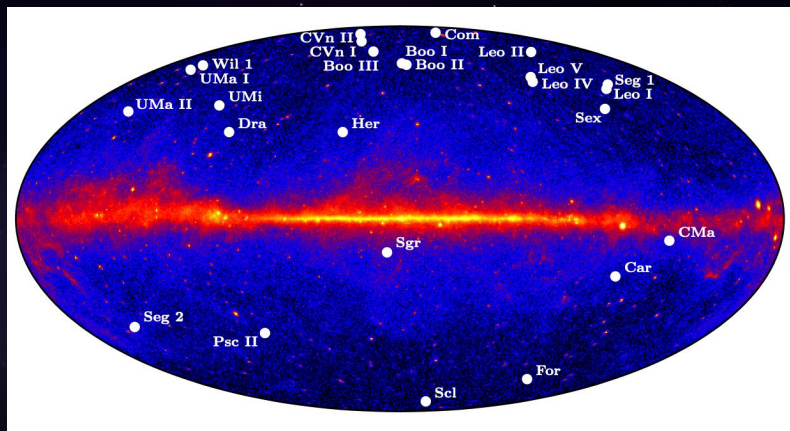
The Fermi Large Area Telescope

→ The Fermi-LAT is a wide field-of-view high energy gamma-ray space telescope sensitive to the energy range of hundreds of MeV to thousands of GeV.

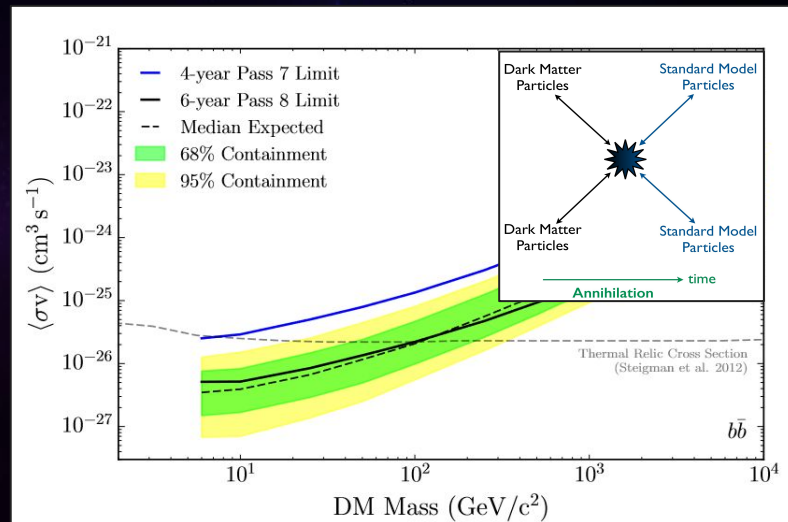


Fermi-LAT's search for dark matter

→ There have been many studies of dwarf galaxies with the Fermi-LAT



Drlica-Wagner, 2013

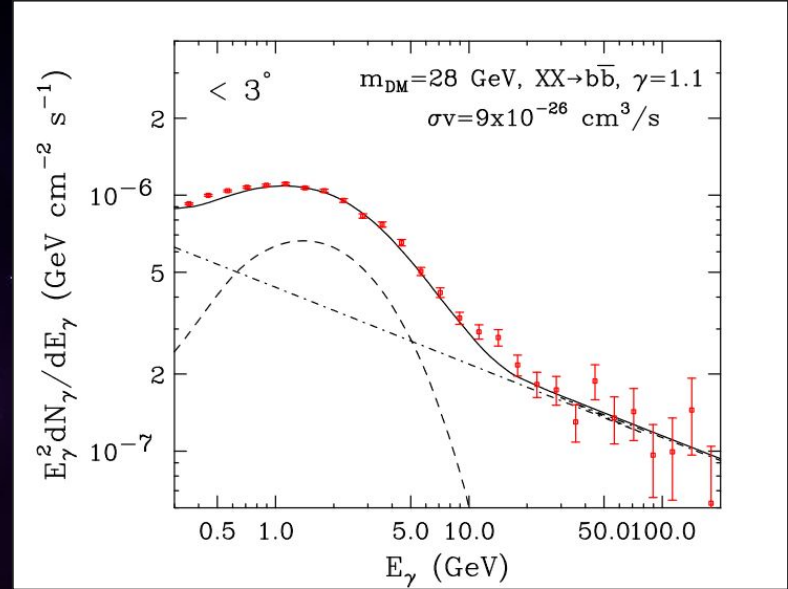


Fermi-LAT Collaboration, 2014

The GeV Excess at the Galactic Center

→ In 2009, Goodenough & Hooper discovered an excess of GeV emission from the Galactic Center consistent with annihilating dark matter.

→ The source is extended, spherically symmetric, and gives a mass and cross-section consistent with a thermal WIMP.

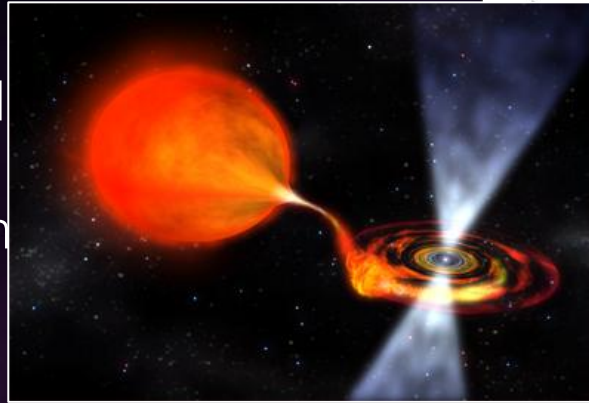


Goodenough & Hooper
2009

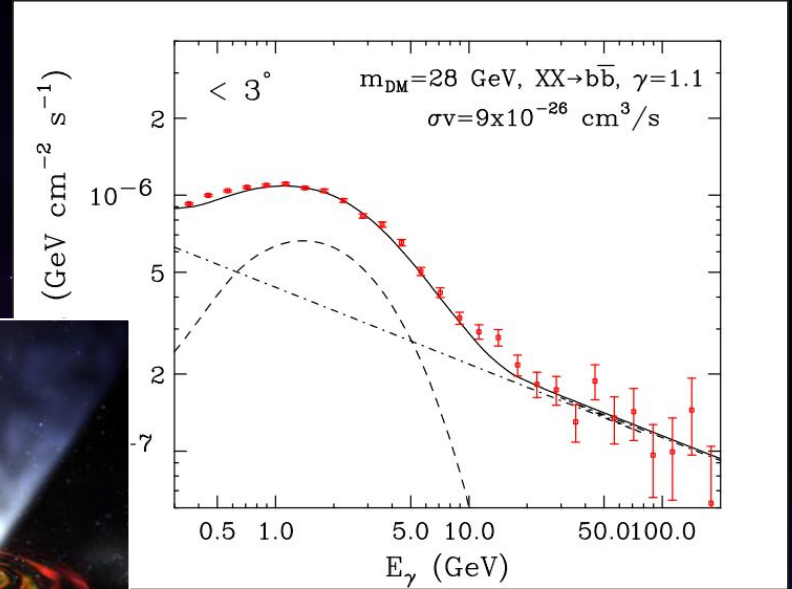
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Could this be a population of millisecond pulsars instead?

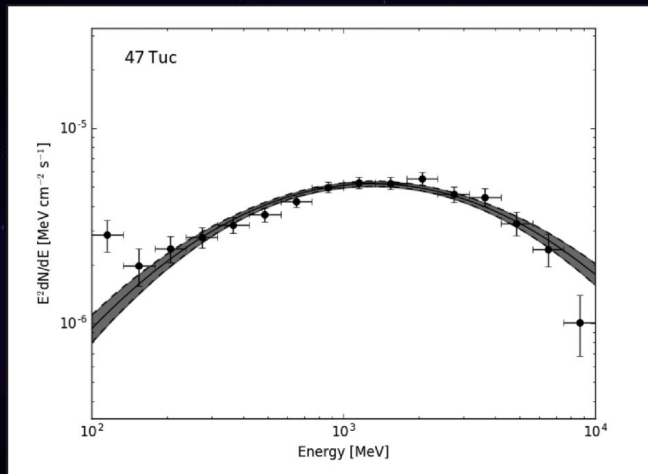


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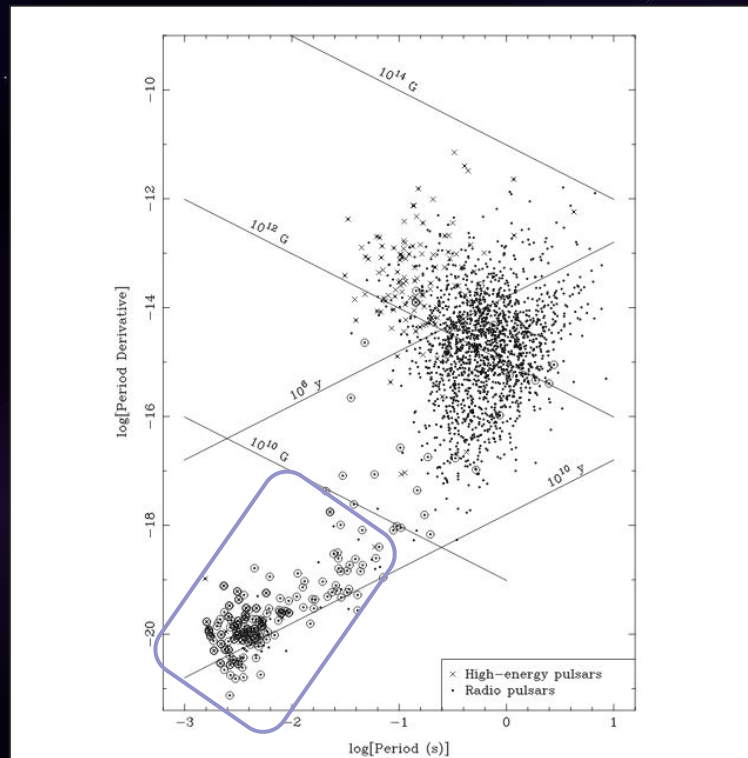
The gamma-ray emission of millisecond pulsars and globular clusters

→ Populations of recycled pulsars contribute to the total gamma-ray flux of globular clusters

→ Out of ~180 in the Galaxy, there are ~35 known gamma-ray bright globular clusters



S.J. Lloyd et al., 2018

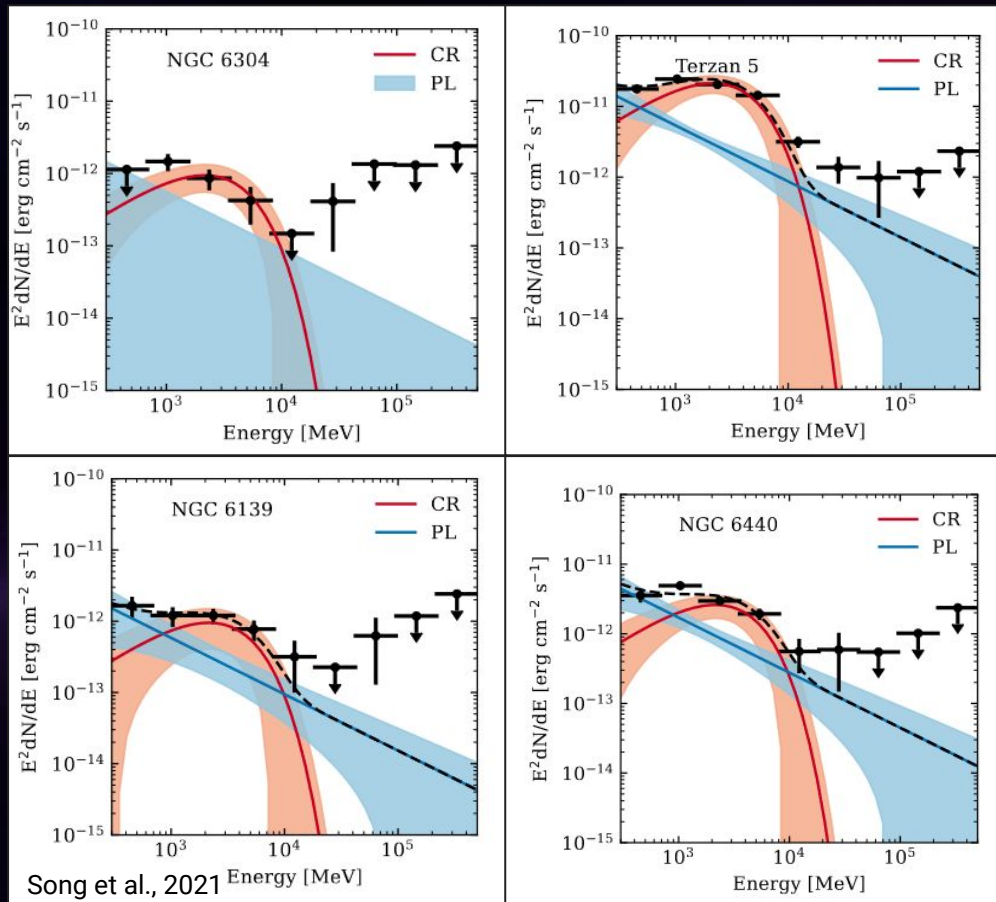


R.N. Manchester, 2017

The source of γ -ray emission in globular clusters

Two possible sources of emission:

1. Curvature or prompt radiation from millisecond pulsar magnetospheres (Harding et al., 2005)
2. Inverse Compton emission from leptons injected by millisecond pulsars into the surrounding environment (Bednarek & Sitarek, 2007)



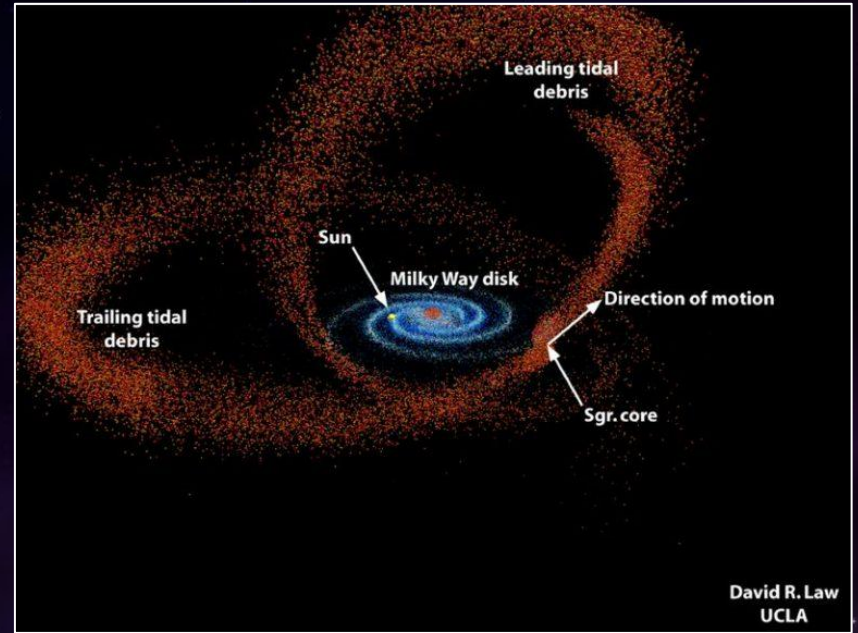
The Sagittarius dwarf spheroidal galaxy

→ Near the Galactic center & overlapping with structure in the Fermi bubbles

→ Undergoing tidal disruption

→ ~ 26 kpc away

→ Population of globular clusters, including recently identified ones (Minniti et al, 2022)



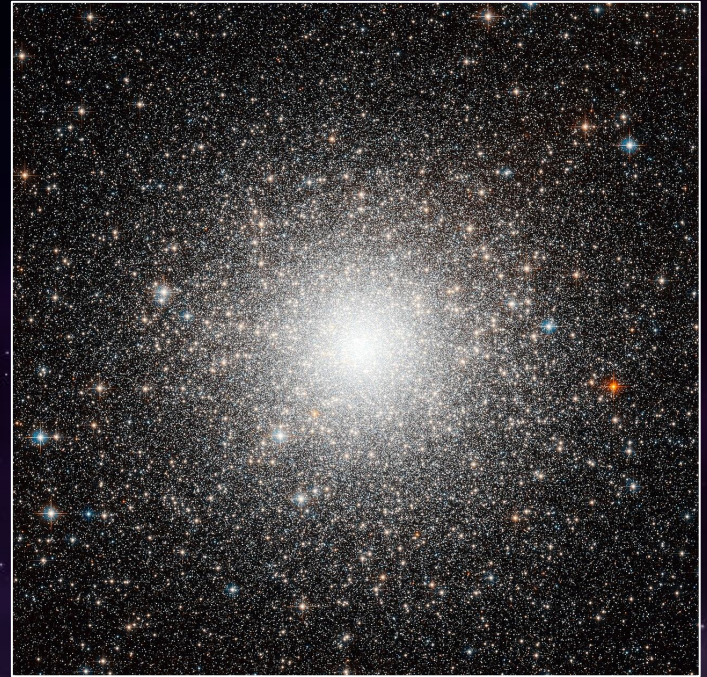
Fermi detections of the Sgr/M54 system

→ Recently detected by the Fermi Collaboration (4FGL-DR3)

→ Significance of ~ 5

→ At the core of Sagittarius

→ Coincident with the globular cluster M54



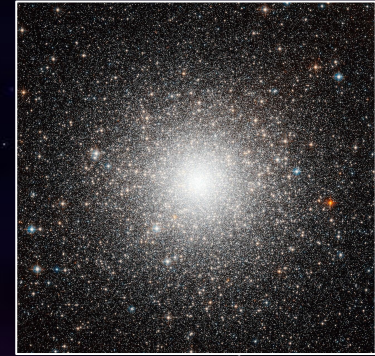
M54 (credit: Hubble/NASA)

Fermi-LAT Collaboration, 4FGL-DR3 2022

Possible sources of gamma-ray emission in Sgr

→ Point-like

- a. Gamma-rays from millisecond pulsar population in M54 and other globular clusters
 - i. Inverse compton scattering component to the flux?



M54 (Hubble/NASA)

→ Extended

- a. Dark matter annihilation in the dense DM core of Sgr
- b. Emission from an isotropically distributed population of millisecond pulsars



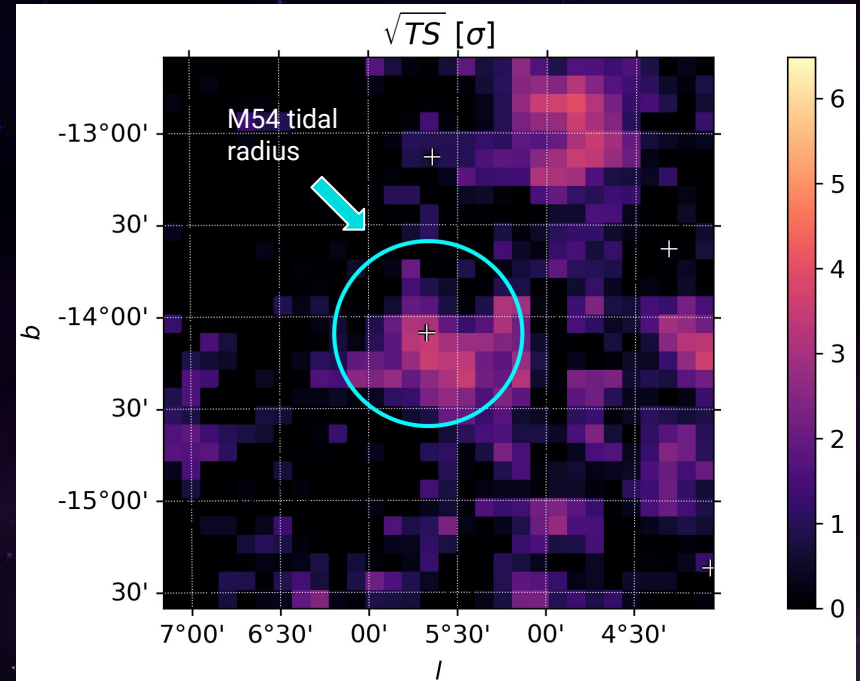
Dark matter substructure (Springel et al., 2008)

This work: the Sgr/M54 system in gamma-rays

→ We find point-like emission from the core of the Sgr galaxy, in agreement with the 4FGL-DR3 catalog

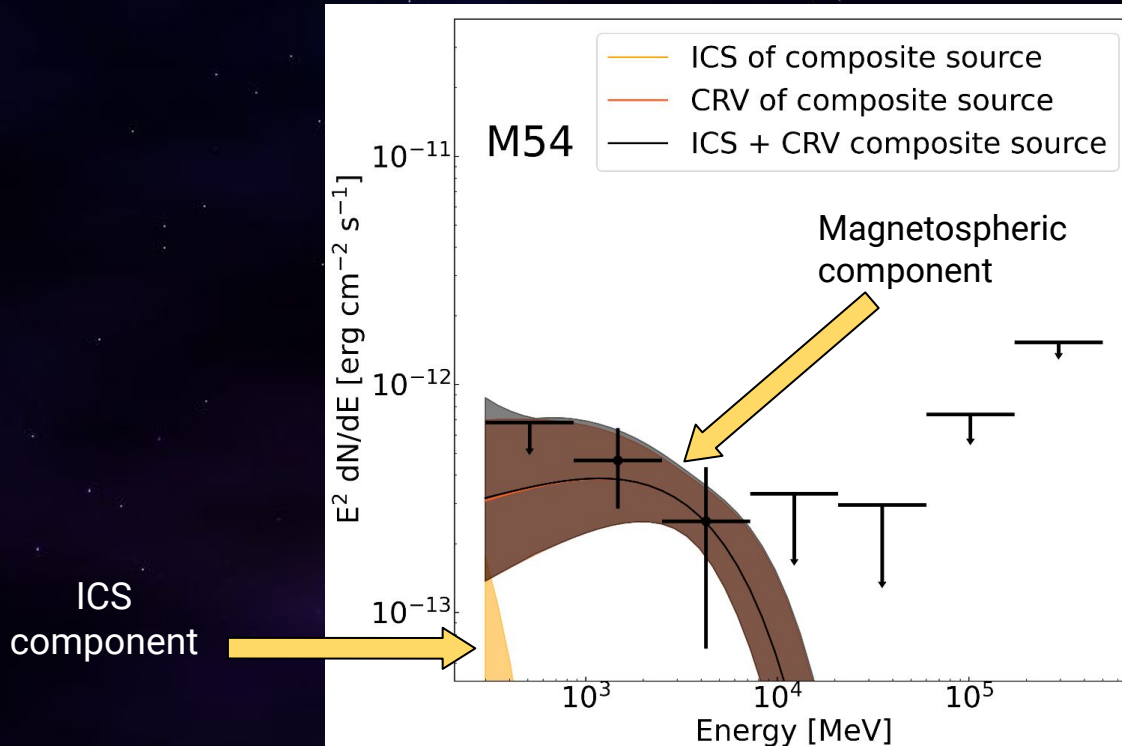
→ Source is spatially coincident with the M54 globular cluster and the core of Sgr

→ Significance of ~ 5



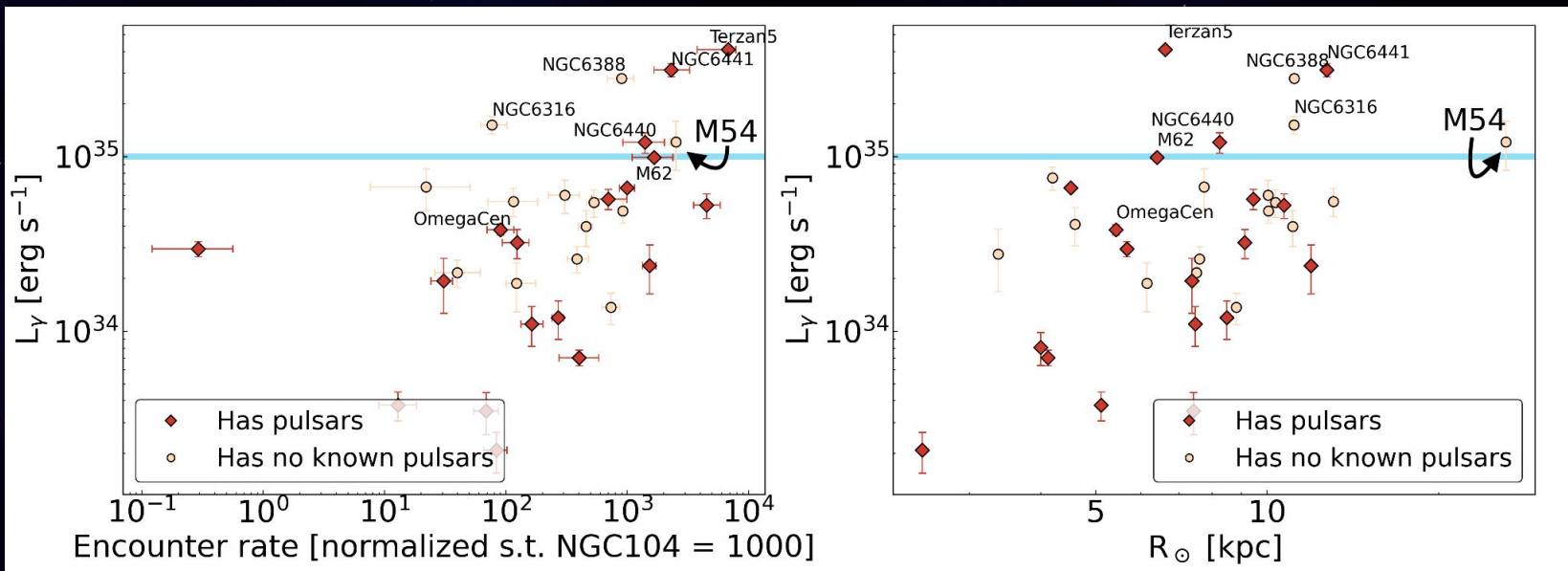
Globular cluster interpretation of the emission

→When using a two-component model, we find no evidence for a high-energy component with a negligible ICS flux compared to the magnetospheric flux



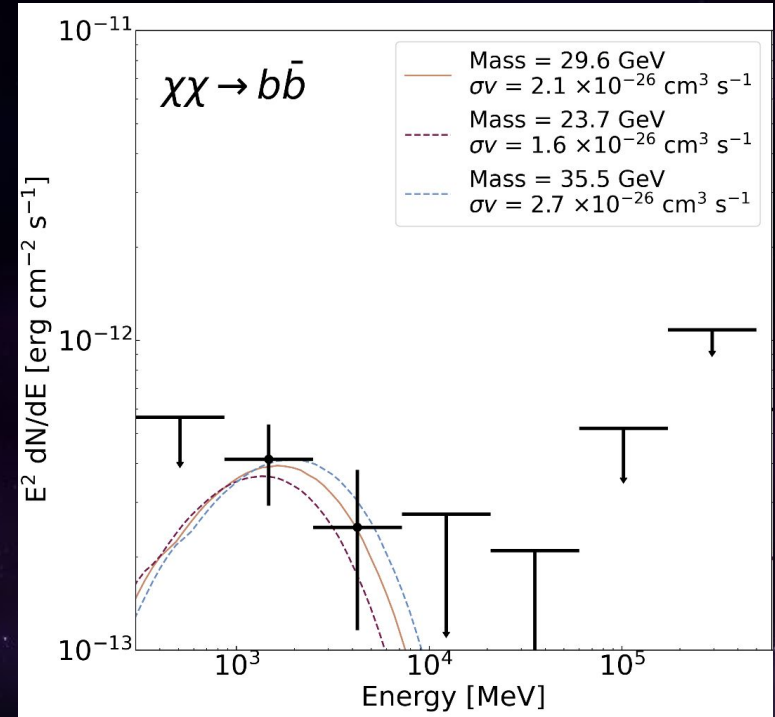
Globular cluster interpretation of the emission

→ The Sgr/M54 source shares similar properties to known Fermi-bright globular clusters



Dark matter interpretation of the emission

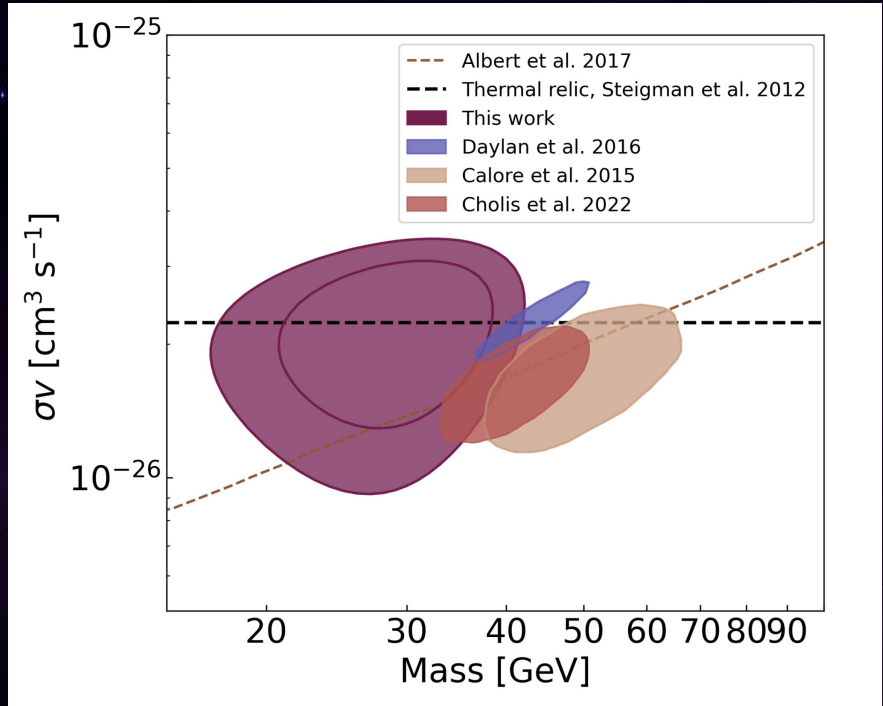
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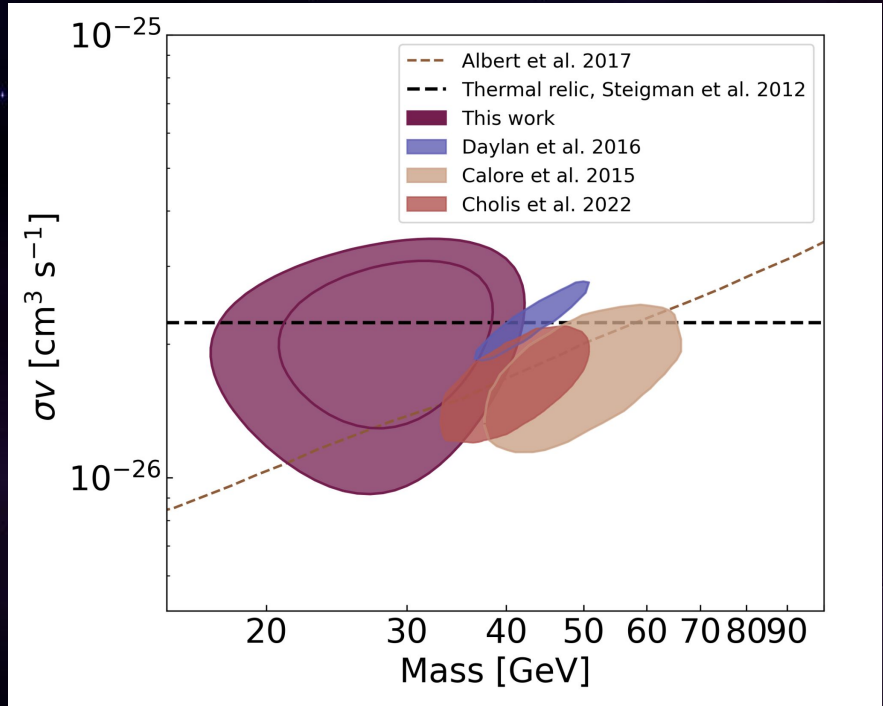


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→ However, the source remains better fit by point-like emission



Summary

Stay in touch :)
addyevans@tamu.edu

- We search the Sgr/M54 system in Fermi data and find point-like emission associated with the core of the Sgr galaxy
- This emission is possibly due to a millisecond pulsar population within the globular cluster M54
- However, a dark matter interpretation is also possible
- We find no evidence for high-energy emission from the catalog source, nor do we find any other sources possibly associated with Sgr
- We find that the source is best-fit by a point-like model rather than an extended source for several choices of background models

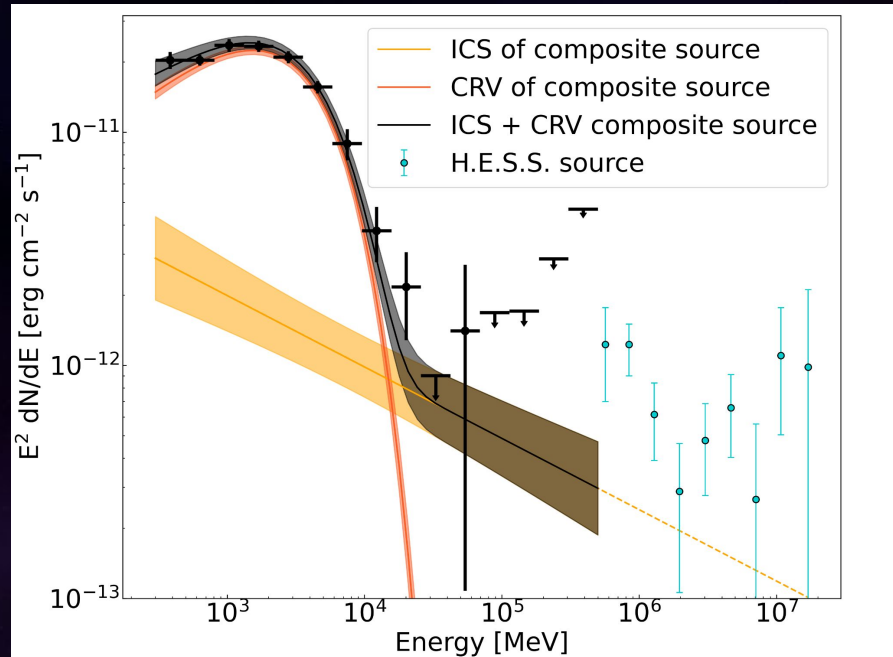
The background is a deep space scene with a dark purple and black color palette. A large, glowing planet is visible in the upper right quadrant, and numerous stars are scattered across the field of view.

Extras

Future work: Disentangling astrophysical backgrounds from dark matter

→ Differentiating emission from a population of pulsars from that of annihilating dark matter is difficult at the Fermi-LAT's energy range

→ We can explore the spectra of MSPs and DM at higher energies in an attempt to disentangle them from each other



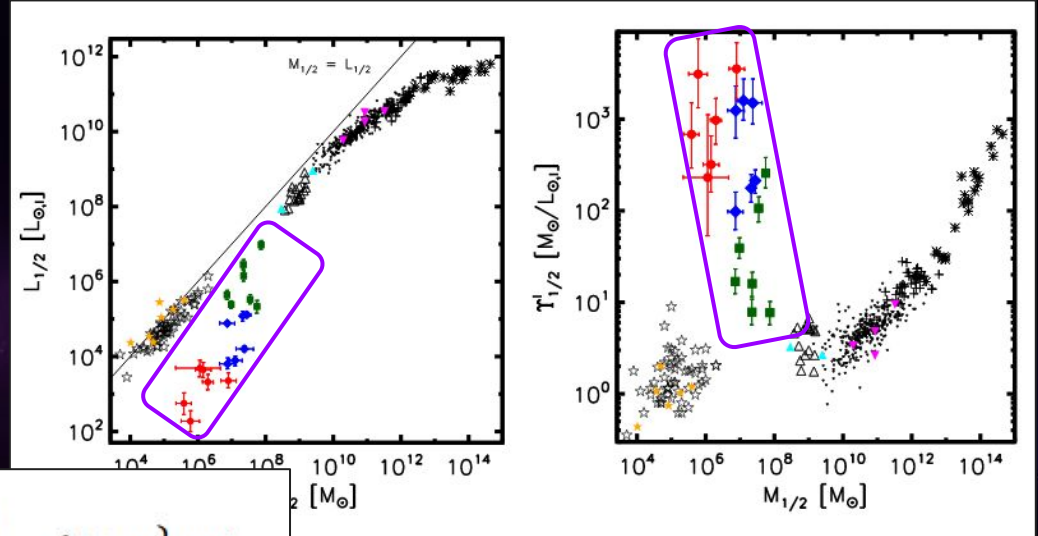
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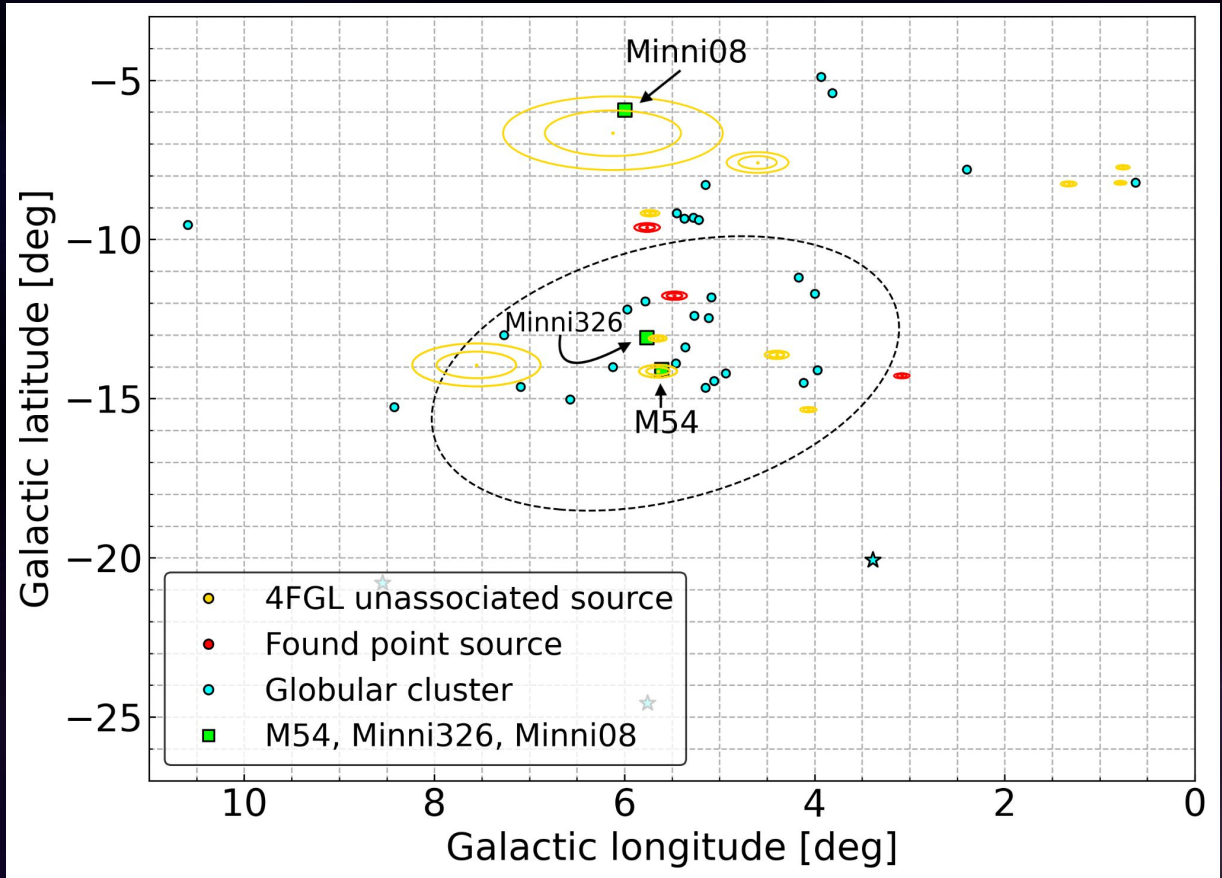
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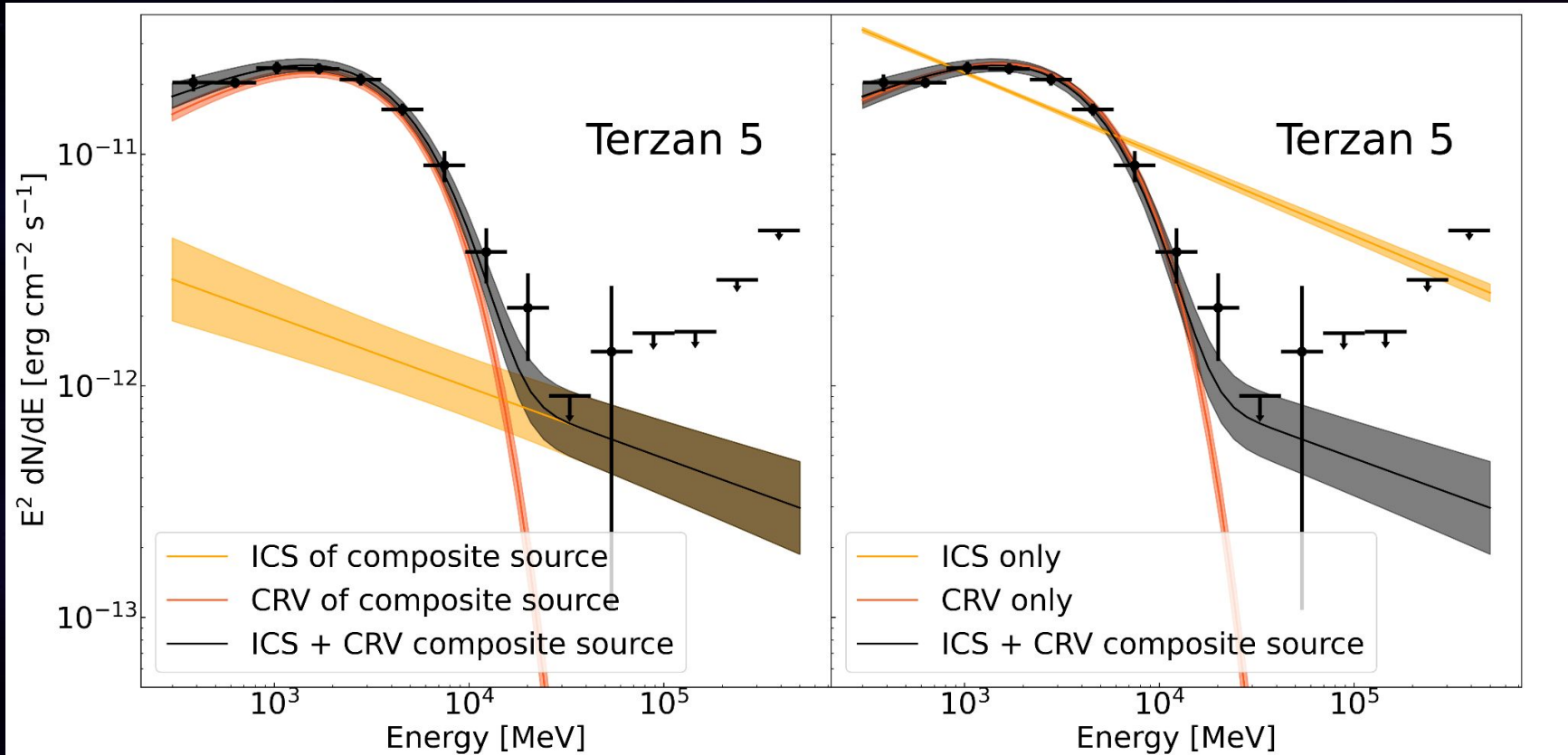
$$\phi_s(\Delta\Omega) = \underbrace{\frac{1}{4\pi} \frac{\langle\sigma v\rangle}{2m_{\text{DM}}^2} \int_{E_{\text{min}}}^{E_{\text{max}}} \frac{dN_{\gamma}}{dE_{\gamma}} dE_{\gamma}}_{\Phi_{\text{PP}}} \cdot \underbrace{\int_{\Delta\Omega} \left\{ \int_{\text{l.o.s.}} \rho^2(\mathbf{r}) dl \right\} d\Omega'}_{\text{J-factor}}$$

The Sgr/M54 system in gamma-rays

→ We search for other sources possibly associated with the Sgr/M54 system and find no substantial evidence for further sources



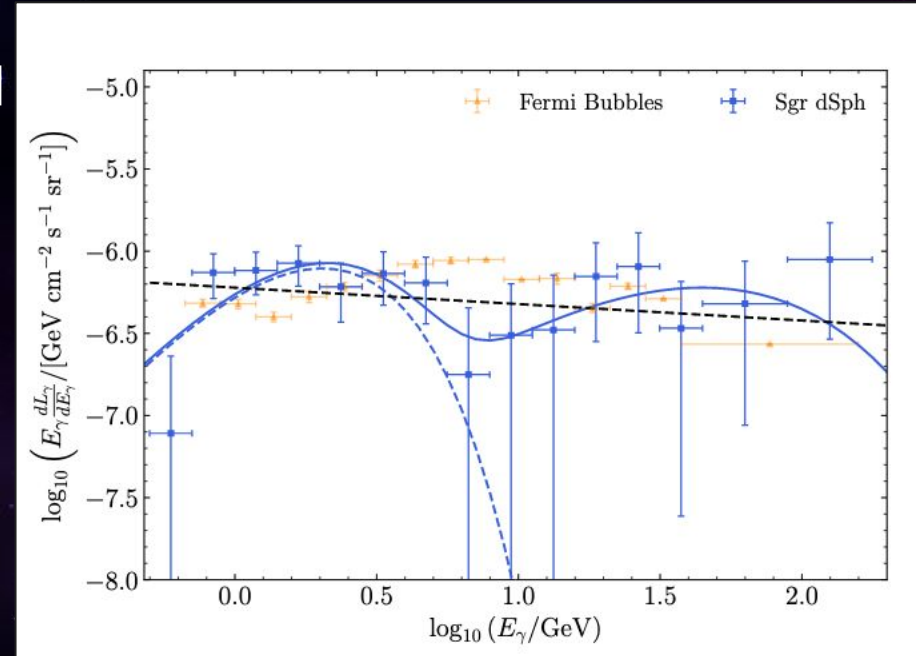
Two-component modeling of Terzan 5



The Sgr/M54 system in gamma-rays

→ Crocker et al. report a highly significant, extended source associated with Sgr

→ Attributed to the dwarf's millisecond pulsar population, with a high-energy component that is comparable to the low-energy component



Crocker et al.,
2022

