

- The **Astroparticle Physics experiments** at LNS cover a wide range of **physics goals** in a long-range programme of science
- Members of CSN-2 at LNS are constantly **growing** in number and responsibilities
- Role of LNS in international experiments is solid and respected

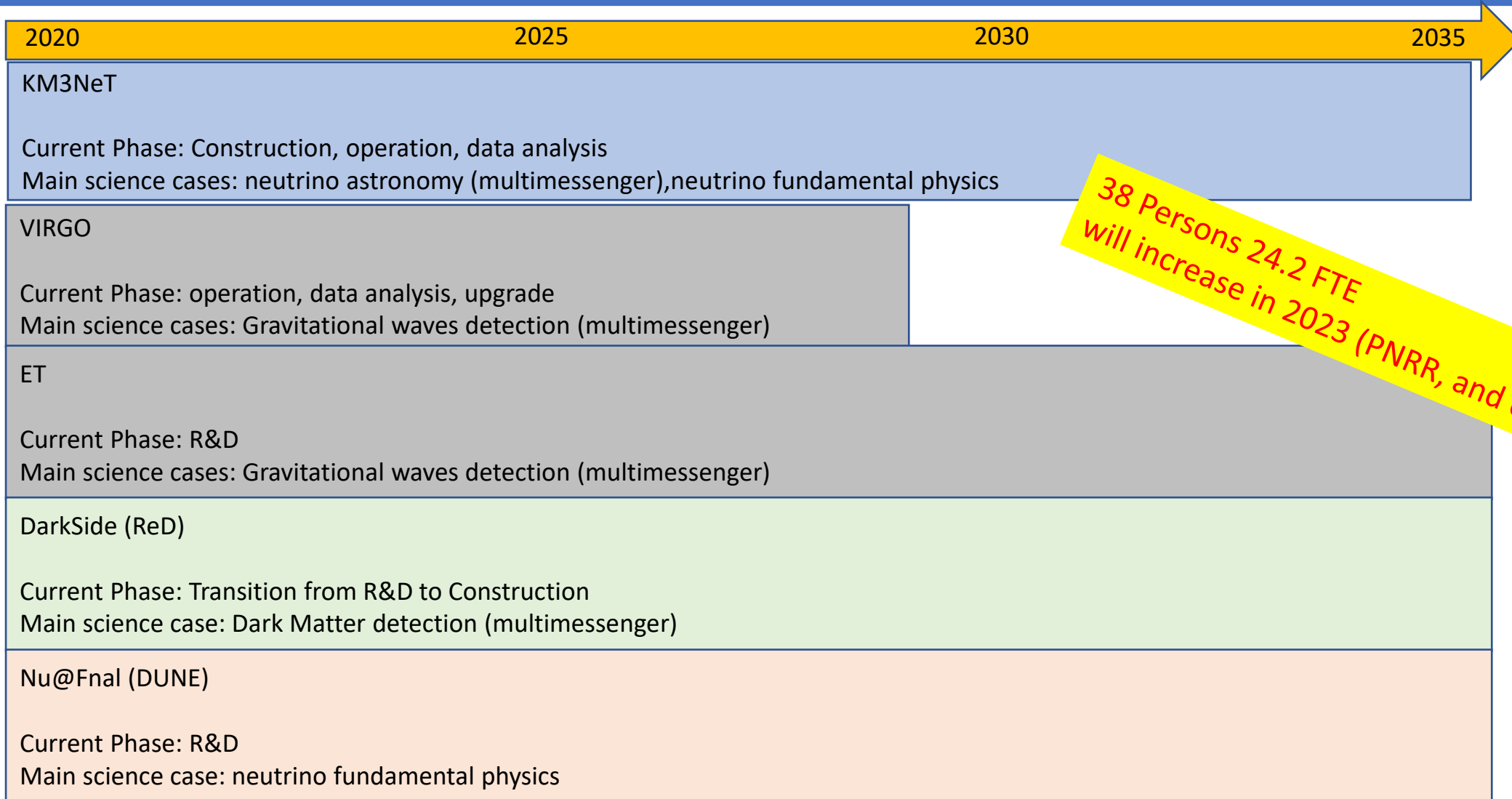
DarkSide /ReD

KM3NeT



DUNE

Virgo / ET





Network of cabled observatories located in deep waters of the Med Sea

**Neutrino astronomy: ARCA @ Capo Passero**

3500 m water depth, 100 km offshore

2 building blocks

230 units (700 m Height)

**Neutrino oscillations and Mass Hierarchy: ORCA @ Toulon**

2500 m water depth, 40 km offshore

1 building block

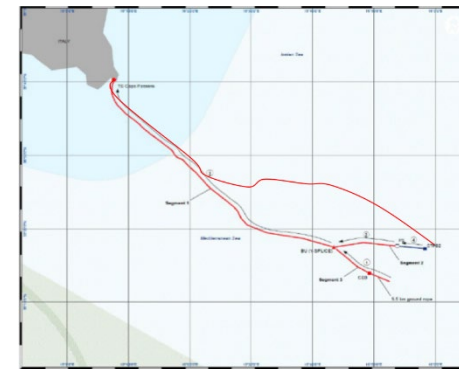
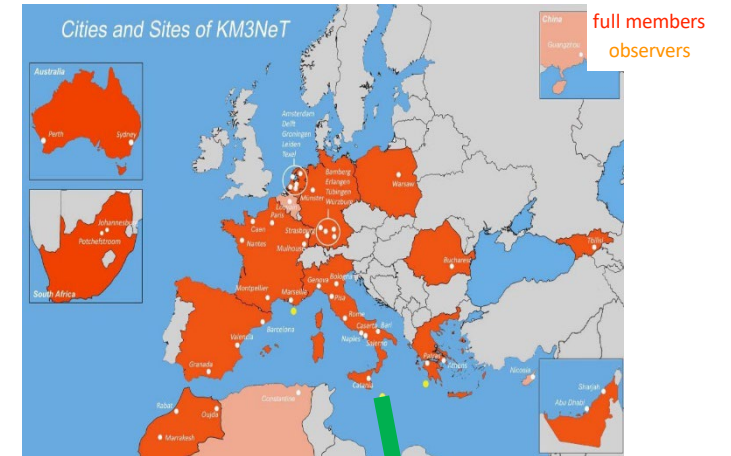
115 units (250 m Height)

**First undersea neutrino detector: ANTARES @ Toulon**

2500 m water depth, 40 km offshore

13 units (300 m Height)

**Decommissioned in 2022!**



ARCA Capo Passero:  
21 DUs in operation

Two electro-optical  
cable and power/optic  
distribution system  
onshore ready for ARCA  
230 DUs

## PNRR-ITINERIS

Seafloor data “e-highway”: optical data link and power connection for EMSO-ERIC stations in Capo Passero  
Seafloor acoustic data hub at LNS.

## ESMO-ERIC: European Seafloor and water column multidisciplinary observatories

Based at Port of Catania Laboratory and at Capo Passero:  
Hub for seafloor geophysics and biology observatories and data centre

## Marine Hazard

Multidisciplinary seafloor station in Capo Passero

## ECCSEL-ERIC: IPANEMA:

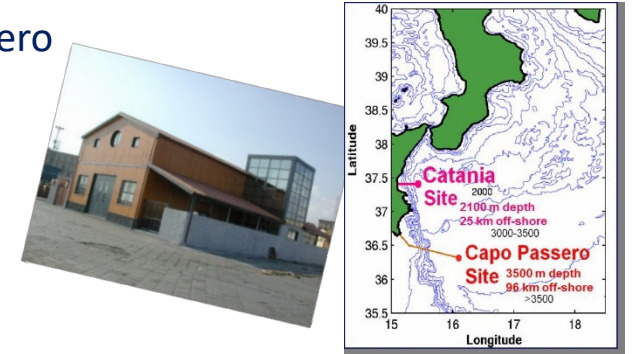
Monitoring of CO<sub>2</sub> with acoustic antennas and chemical sensors  
two stations: in Panarea (20 m depth) and Catania (2100 m depth)

## FOCUS-ERC

Monitoring of seismic phenomena with Brillouin Optical time reflectometry using the fibers of the subsea cable of the Port of Catania Test Site

## Geo-Inquire EU (K. Fleming PI, GFZ, G. Riccobene PI INFN) – Ends in 2026

Transnational access activity to use optical fibers of the subsea cable at the port of Catania for seismic and volcanic monitoring  
56 k€ funded

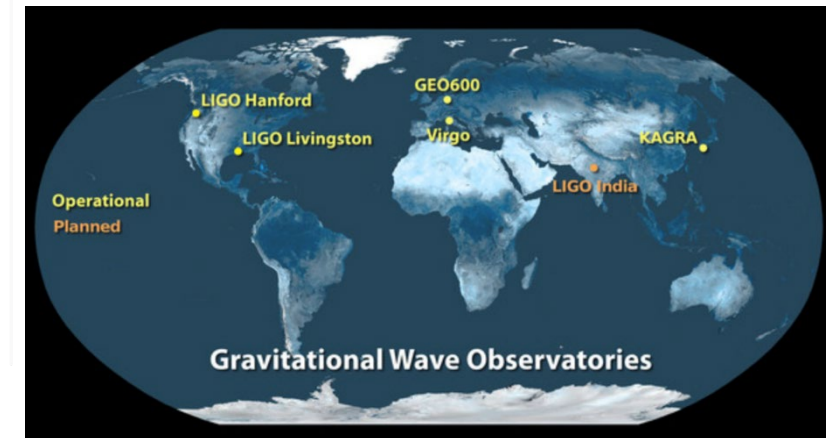


small impact and good reward

test new technologies  
multidisciplinary plan  
and new personnel

## The VIRGO Collaboration

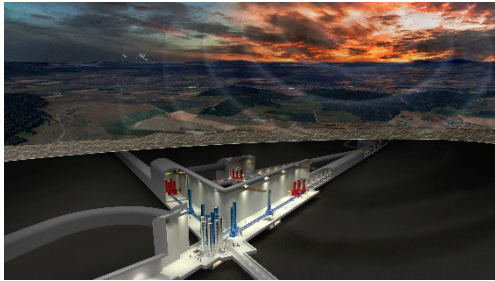
- ~770 members, ~450 authors, 131 institutions from 15 countries
- 34 Groups:
  - 32 full members
  - 2 in the first year (L2I Toulouse, KU Leuven)
- 9 countries represented in the VSC



Based in **Cascina (Pisa)** at the **European Gravitational Observatory**

new director Massimo Carpinelli (Uni Sassari - on leave to Milano Bicocca- and LNS)

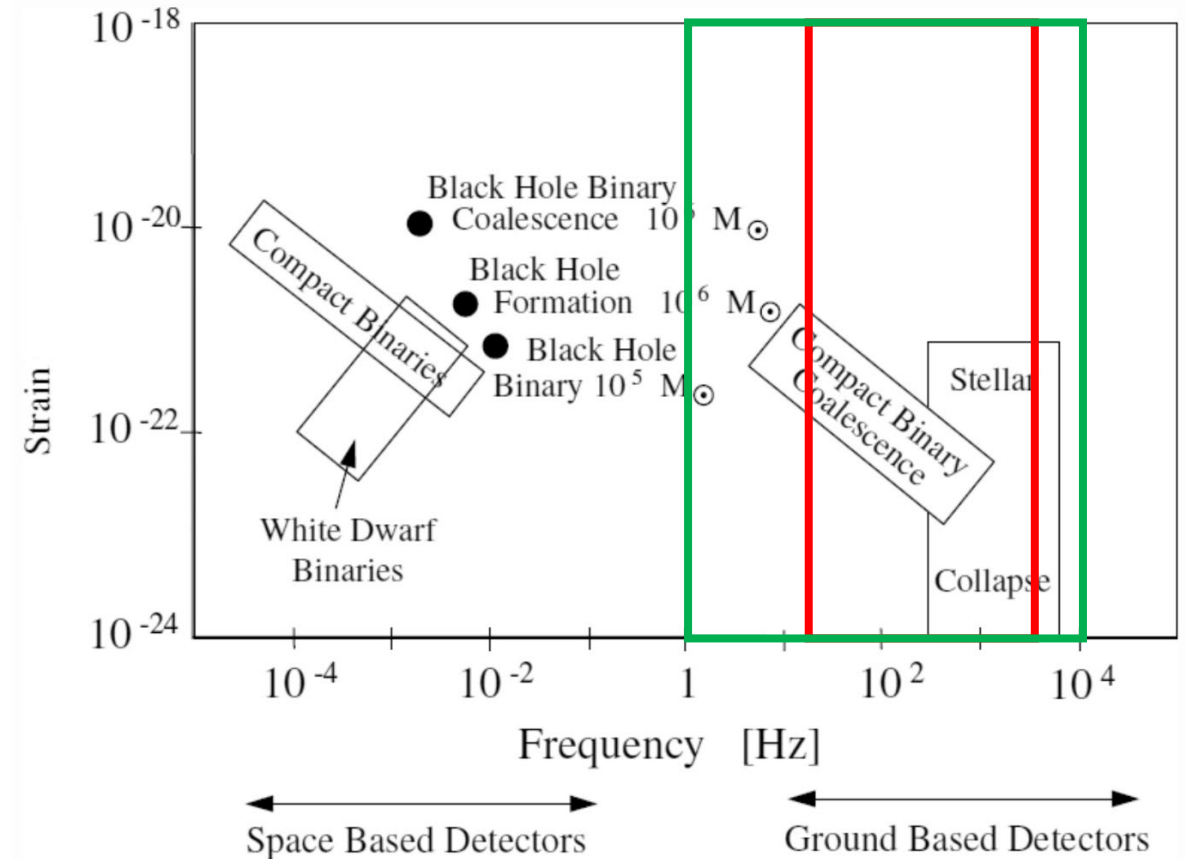
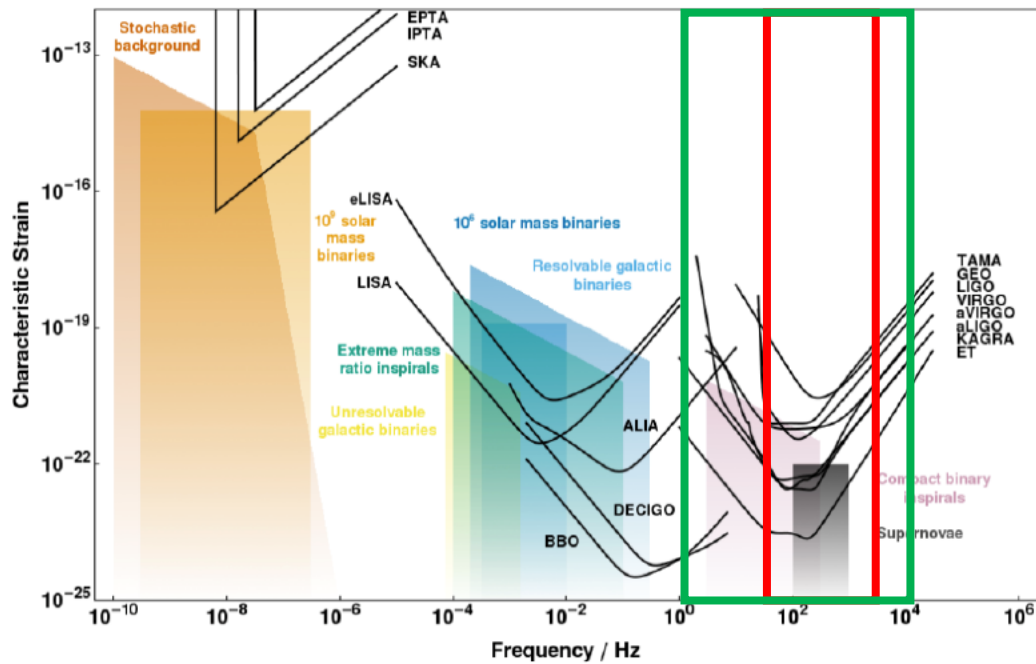
# VIRGO and Einstein Telescope



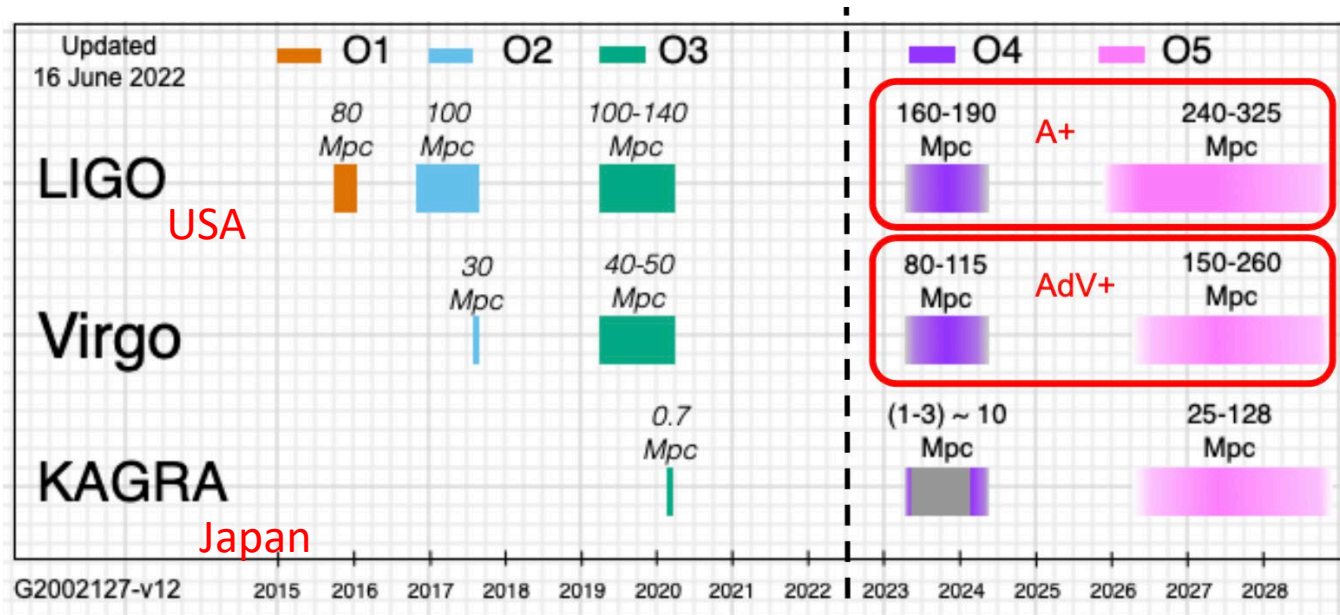
Virgo: 20-2000 Hz  
compact binary inspirals (BBH, BNS and BH-NS), supernovae and bursts.



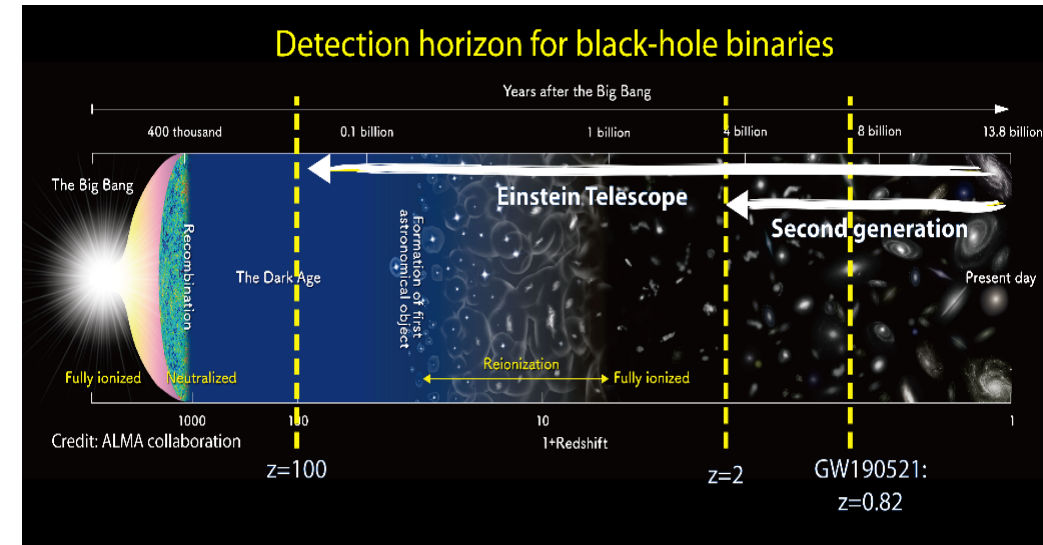
Einstein Telescope: 1-10000 Hz  
Precision measurements **AND** new discoveries



# From VIRGO to Einstein Telescope



constant increase of observation horizon



Collaboration birth in 2022: 1200 scientists from USA, Europe and Japan

ET will be a new discovery machine:

will explore *almost the entire Universe* listening the gravitational waves emitted by black hole

ET will be a precision measurement observatory:

will detect, with high SNR, *hundreds of thousands* coalescences of binary systems of Neutron Stars per year

## Requirements

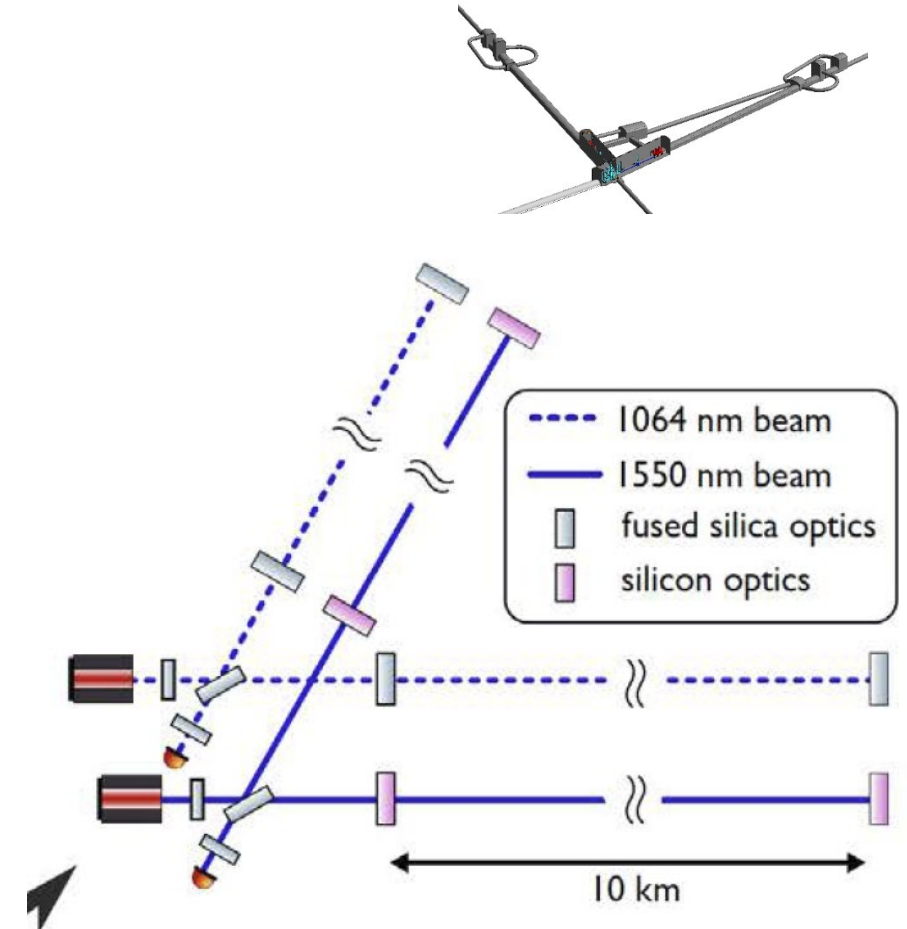
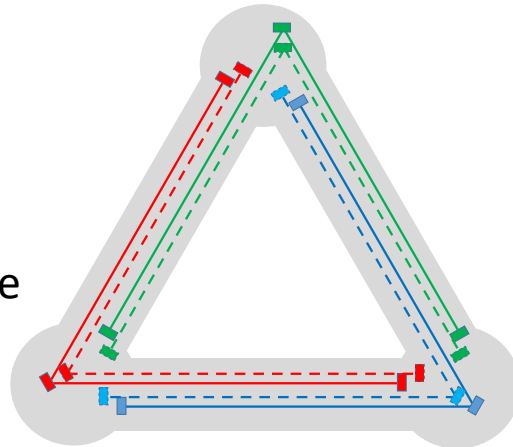
- Wide frequency range
- Massive black holes (LF focus)
- Localisation capability
- (more) Uniform sky coverage
- Polarisation disentanglement
- High Reliability (high duty cycle)
- High SNR



## Design Specifications

- Xylophone (multi-interferometer) Design
- **Underground**
- **Cryogenic**
- **Triangular shape**
- Multi-detector design
- Longer arms

6 arms: 1 LF and 1 HF arm in each side





Coordination of the **Sardinian Site Characterization** Activity:

Data analysis and ET sensitivity estimation

**Infrastructure** design and feasibility study

Sensor Instrumentation and maintenance

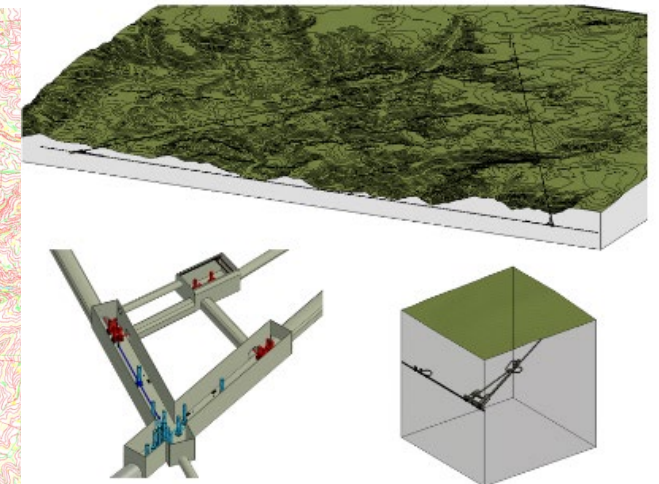
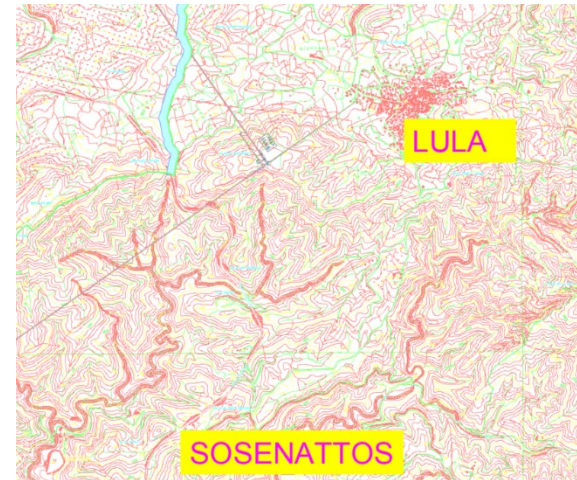
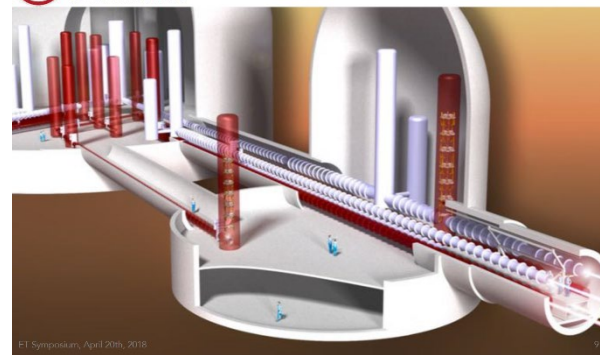
Acquisition system and mechanics of the novel “Archimedes” tiltmeter

Sardinian site shows extremely low seismic noise in worldwide data

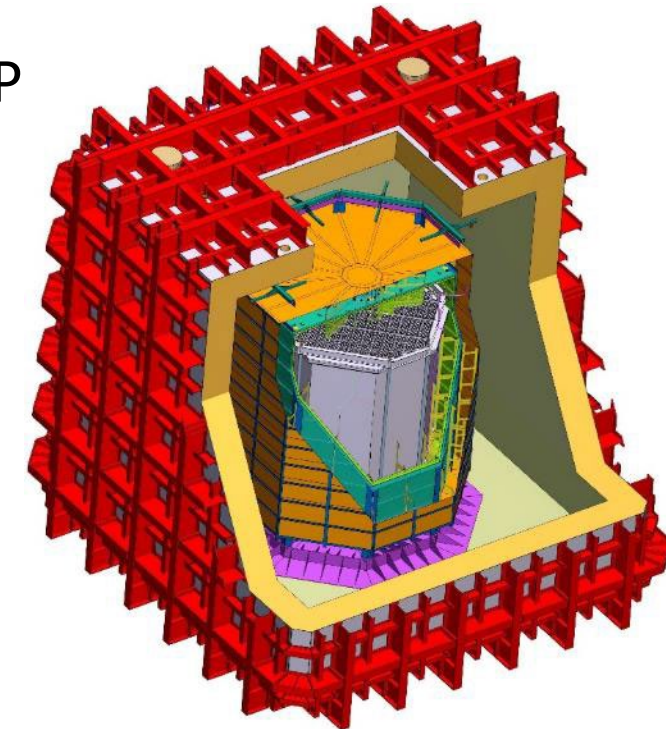
**Role of LNS empowered within PNRR ETIC**

Site characterization and infrastructure design

17.6 M€ at LNS over 50 M€ funding



- Search for **dark matter** in the form of Weakly Interacting Massive Particles (**WIMPs**)
- Signature: **low energy (< 100 keV) nuclear recoil** produced by WIMP elastic scattering
- **DarkSide** at Gran Sasso Laboratory, WIMPs search using a **dual-phase TPC** with **low-radioactivity LAr**
  - Operated a **50 kg TPC** (DarkSide-50) with **UAr**
  - Next step: 30 ton LAr **TPC** (DarkSide-20k)
    - Novel light **readout** with **SiPM**
    - Getting ready for data in **2026**, exposure O(100) ton yr
    - Expected sensitivity  $10^{-47} \text{ cm}^2 @ M_W = 1 \text{ TeV}/c^2$
  - Next-next step: global worldwide effort (ARGO, 300 ton LAr) → **Global Argon Dark Matter Collaboration (GADMC)**



**darkside**



- **ReD** is a project within the GADMC, three-fold goal:
  - check if a **dual phase LAr TPC** has sensitivity to the **direction** of Ar recoil
  - characterize the **response** of the LAr TPC to **very low-energy recoils** (< few keV) → **hot topic (S2-only)**
  - act as a **test bench** of the technical solutions for DarkSide-20k TPC
- Major involvement of LNS
- Phase 1 (Napoli) – **Characterization** of the TPC
  - **Stability** over time scale of months
  - Check that **performance** OK for phase 2
- Phase 2 (LNS) – **Directionality run**
  - **TANDEM** beam **LNS** (February 2020)
  - **No indication** for directionality
- Phase 3 (Sez. Catania) – **Low-energy run** with  $^{252}\text{Cf}$ 
  - **In preparation** (up to 2-5 keV)
  - Complemented with low-energy ERs

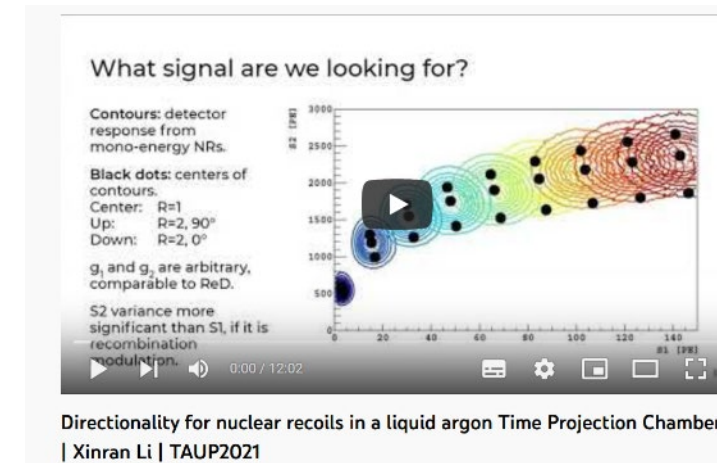
Eur. Phys. J. C (2021) 81:1014  
https://doi.org/10.1140/epjc/s10052-021-09801-6

THE EUROPEAN  
PHYSICAL JOURNAL C

Regular Article - Experimental Physics

Performance of the ReD TPC, a novel double-phase LAr detector with silicon photomultiplier readout

P. Agnes<sup>1</sup>, S. Albergo<sup>2,3</sup>, I. Albuquerque<sup>4</sup>, M. Arba<sup>5</sup>, M. Ave<sup>1</sup>, A. Boiano<sup>6</sup>, W. M. Bonivento<sup>5</sup>, B. Bottino<sup>7,8</sup>, S. Bussino<sup>9,10</sup>, M. Cadeddu<sup>2</sup>, A. Caminata<sup>1</sup>, N. Canci<sup>11</sup>, G. Cappello<sup>2,3</sup>, M. Caravati<sup>12</sup>, M. Cariello<sup>7</sup>,



EPJ C 81 (2021) 1014

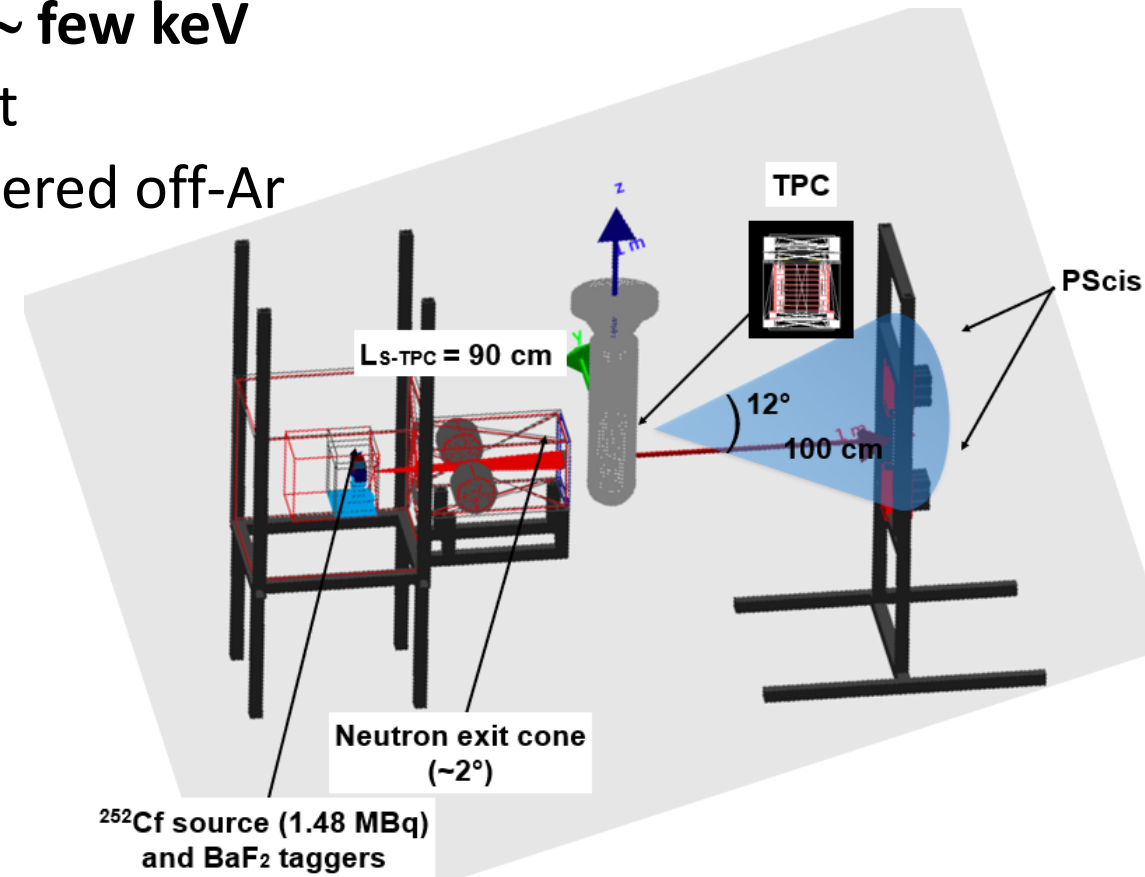
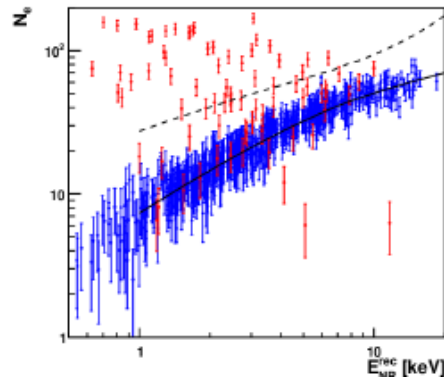
Talk at TAUP 2021

- **Low-energy recoil measurements** ( $< \text{few keV}$ ) by using neutrons from a  **$^{252}\text{Cf}$  fission source**

- Neutrons **O(2 MeV)**, more appropriate for  $E_{\text{rec}} \sim \text{few keV}$
- Use close fission tagger ( $\text{BaF}_2$ ) and time of flight
- Neutron spectrometer to detect neutrons scattered off-Ar
  - Use 1-inch plastic scintillators

- Sensitivity down to

**2-5 keV<sub>NR</sub>**



- The **URANIA plant** will extract and purify the **underground Ar** (low in  $^{39}\text{Ar}$ ) from the  $\text{CO}_2$  wells at the Kinder Morgan Doe Canyon Facility, Colorado
  - Plant **built & commissioned** at the Company site
  - Ready for **shipment** to Colorado
- Expected production: 50 tons
  - To be purified and further depleted by distillation in the ARIA facility
- LNS actively involved in the **design and construction** of the **plant**

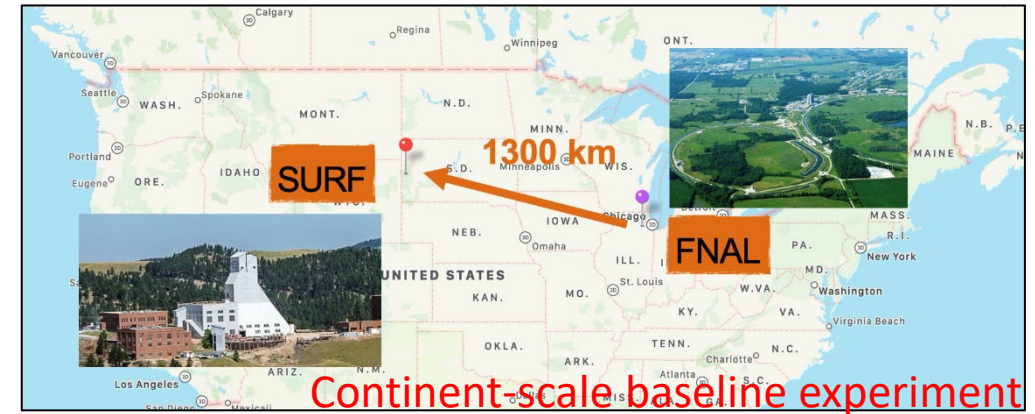


The Deep Underground Neutrino Experiment

1300 scientists and engineers

190 institutes

37 countries + CERN

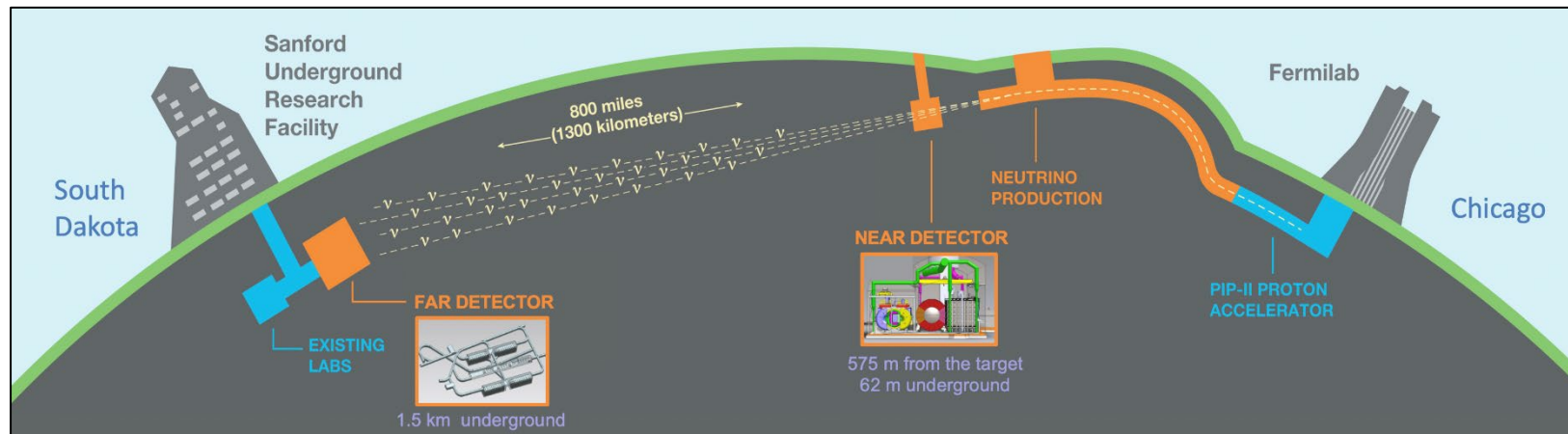


Continent-scale baseline experiment

LNS has responsibilities in **hardware construction and testing** (Far and near detector) and **software** (Geasigen for DUNE)

### 10 MeV -10 GeV nu physics:

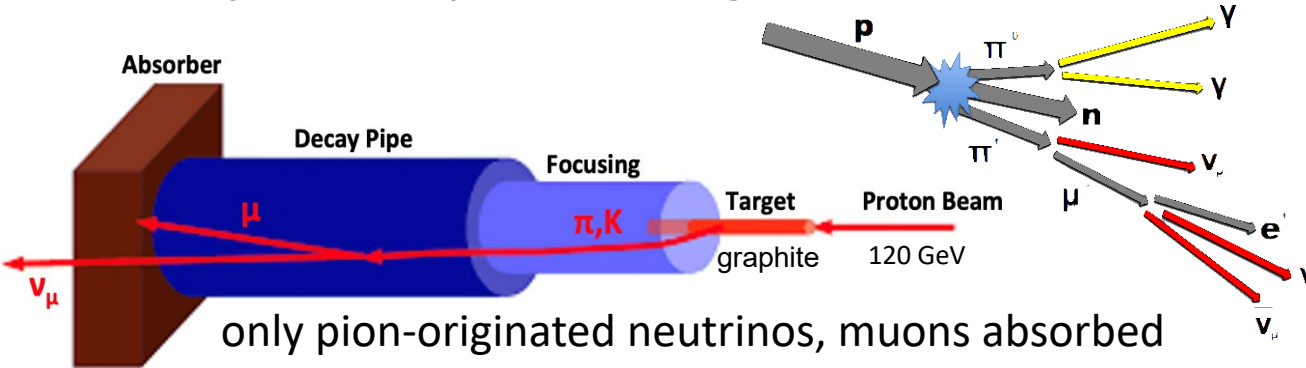
- Neutrino Physics  
CPV in the leptonic sector  
Mass Hierarchy and Oscillation Physics
- Nucleon Decay
- Supernova physics & astrophysics  
→ **multi-messenger**  
Galactic core collapse supernova
- Other topics  
Atmospheric neutrinos  
Neutrino interaction physics



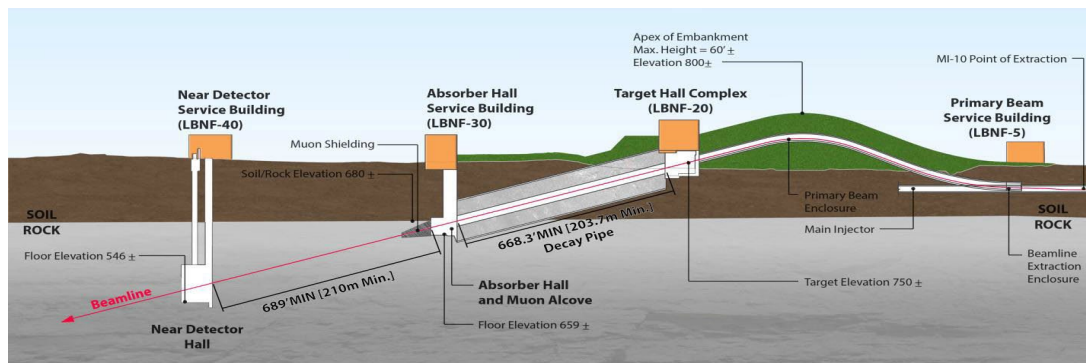
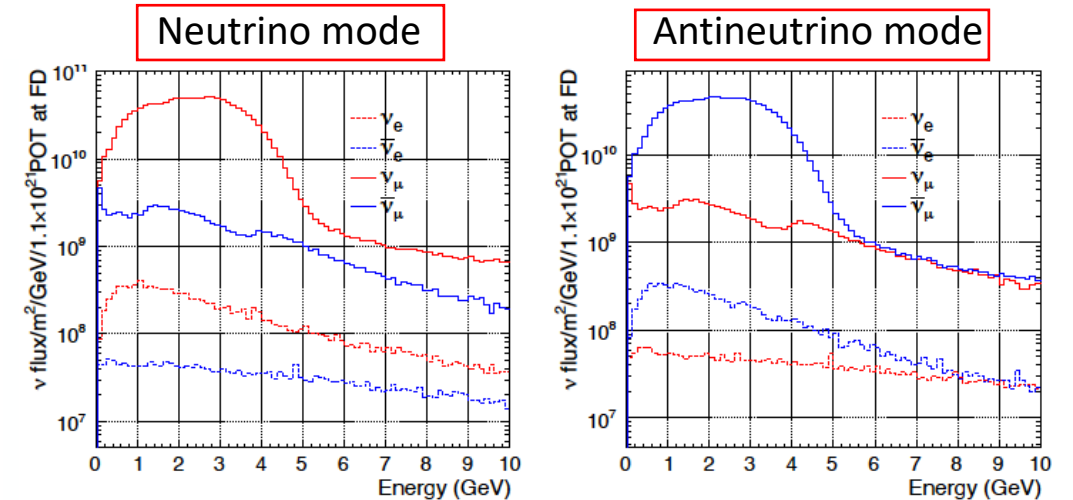
PIP-II accelerator: proton beam 1.2 MW (2028)

(PIP-III upgrade to 2.4 MW)

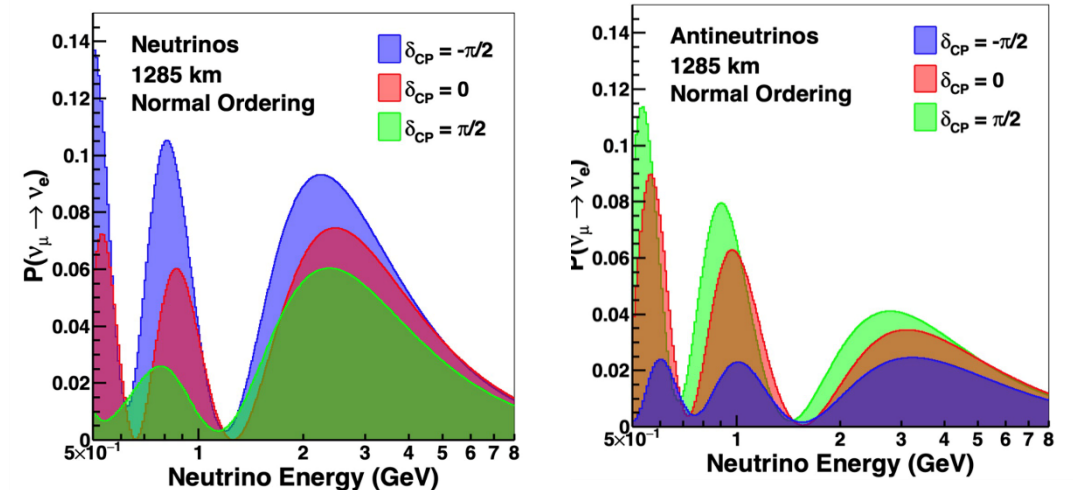
$10^{21}$  POT/year (POT: proton on target)



Feature: select the **polarity** of the current of the "magnetic horn" (focusing system)



5.8° inclined beam (specially created 18 m artificial hill)



The site consists of 5 main caves at a depth of 1.5 km:

- 4 detection modules

- 1 support module (cryogenics and DAQ)

Detection modules:

- Detectors based on LArTPC technologies

- LAr is both the target and the detection medium

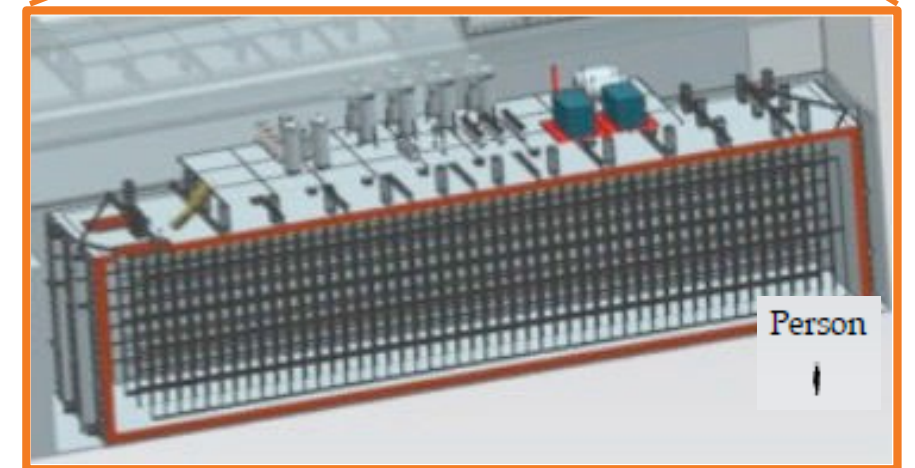
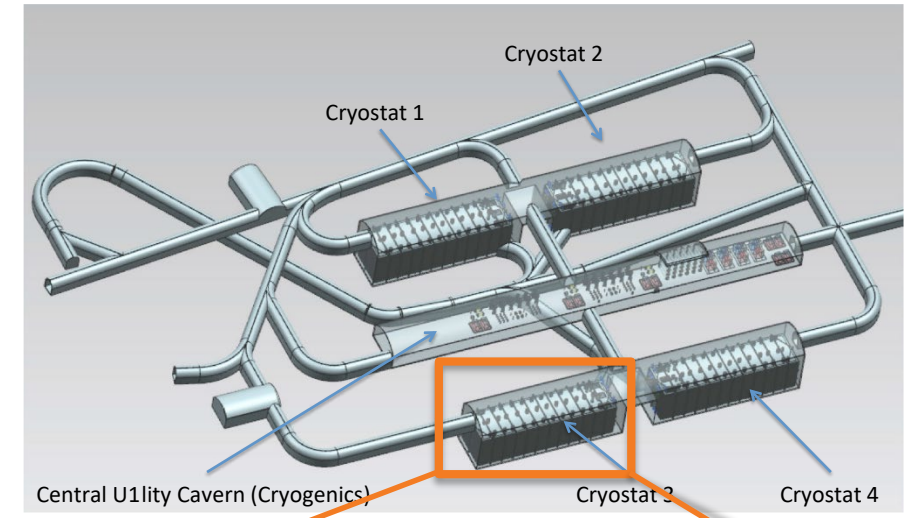
- Same cryostat (62 m x 19 m x 18 m)

- 17 kt of LAr in the whole volume (70 kt total)

- 10 kt di LAr in the fiducial volume (40 kt total)

- possibility of hosting detectors of different designs

**LNS responsible for the qualification of  
electronic components at cryogenic  
temperatures**





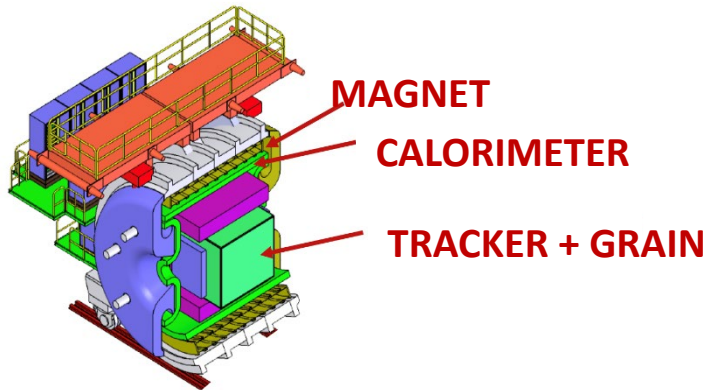
Complex of detectors based on different technologies and focused on:

- precise characterization of the neutrino beam
- reduction of uncertainties in the cross sections

## SAND

will monitor the intensity, spectrum and profile of the neutrino beam:

→ real-time variations in beam operating conditions



re-use/upgrade KLOE (LNF): →

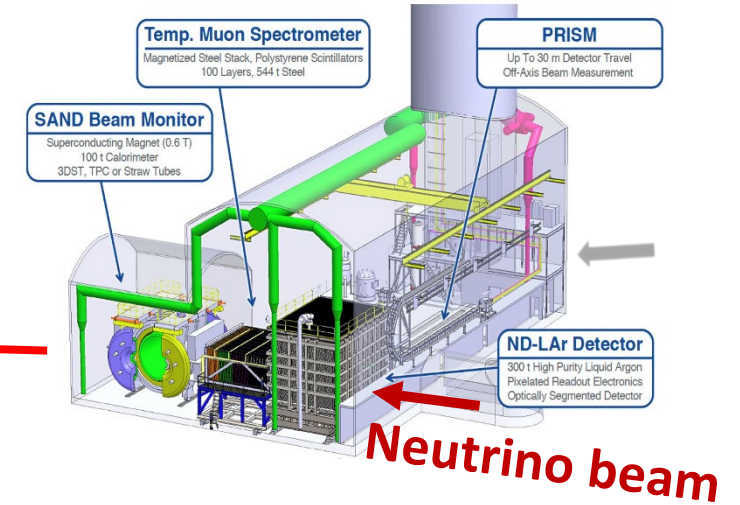
0.6 T superconduction solenoid magnet  
4pi electromagnetic calorimeter

new design:

internal straw tube tracker  
GRAIN LAr Meniscus (1 T)

**LNS involved in both activities**

SAND must be operational at first beam (2028)



Thanks for your attention!