Measurements of proton-induced reactions on ruthenium-96 in the Experimental Storage Ring(ESR) at GSI

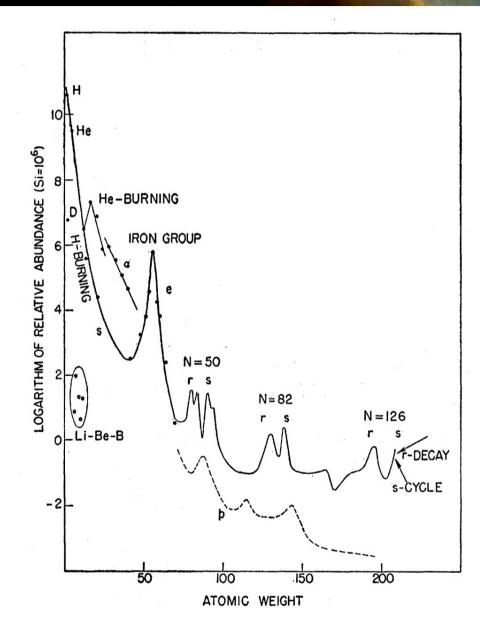
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Outlook

- •p-process
- Experimental setup
- •Monte-Carlo simulations with Geant4 code
- Analysis results
- Summary

Introduction

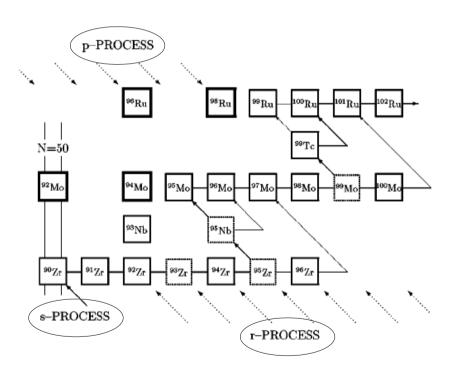


- •most of the elements heavier than iron are built in stars via neutron capture in s- or r-processes;
- •s(slow)-process with neutron densities in the order of $10^6 \div 10^{11}$ cm⁻³, neutron capture rates slower than β -decay rates;
- •r(rapid)-process with neutron density > 10²⁰ cm⁻³, stable nuclides reached through beta decay chains after the neutron capture;
- •35 nuclei from the proton-rich side of the valley of stability are produced by a combination of the (p,γ) , (γ, n) , (γ,p) and (γ,α) reactions on existing s- or r-nuclei at very high temperatures.

p-process

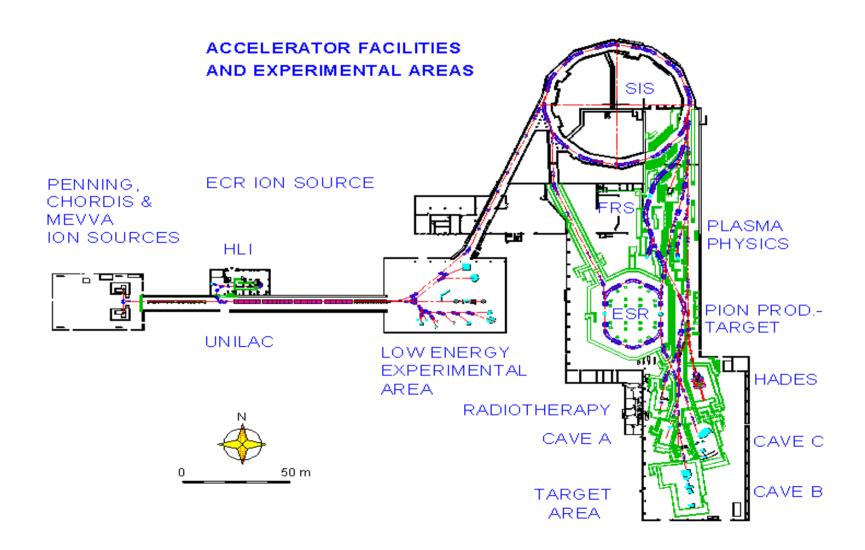


- •to adequately describe pprocess need reliable information on the thousands of reaction rates involved;
- •the reaction rates based on detailed balance theorem to predict a scenario of pprocess;
- •p-process appears in supernovae, where (p,γ) and (γ,n) reactions on s- and r-nuclei occurred at high temperatures.



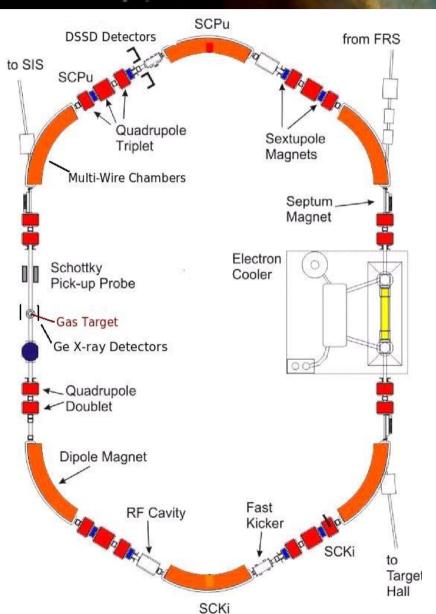
- most p-process nuclides are not known with optimum accuracy;
- •the light p-nuclei ⁹⁶Ru are underproduced in supernovae within a factor of 10, but have rather large isotopic abundance 5.52%.

- •Only a few measurements of the cross-sections are performed, mostly in activation experiments.
- •Limitation of activation method: need radioactive reaction products and stable samples.
- •New experiments in the Experimental Storage Ring (ESR) with circulating heavy ion beams.
- •Not limited as experiments using an activation technique.
- •The purpose is to compare results measured with activation method and to explore the possibilities of a new method for measuring the same reactions with a short-lived isotopes and stable reaction products.
- •Test case: 96 Ru(p, γ) 97 Rh.

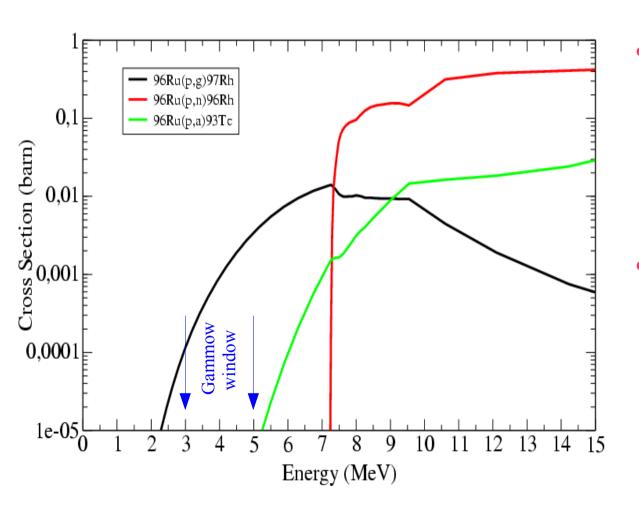


Electron Storage Ring at GSI

Experimental Astrophysics



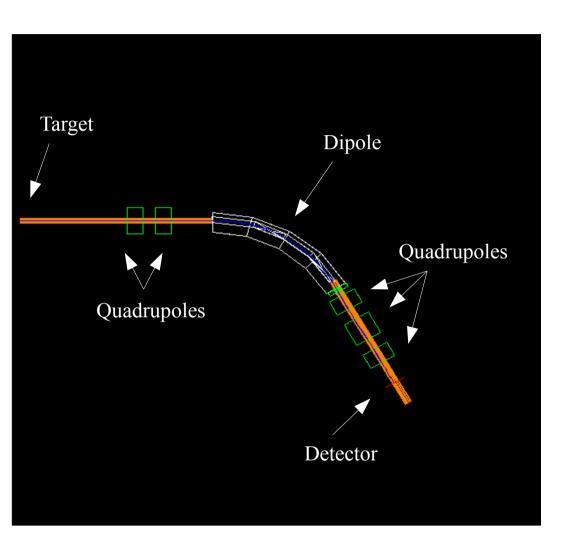
- •experiment with ⁹⁶Ru beam at 9-11 MeV and a hydrogen target;
- •products of the ⁹⁶Ru(p,γ) reaction were detected by 2 Double Sided Silicon Strip Detectors (DSSD);
- •Multiwire Proportional Chamber detected electron pick-up reaction;
- •X-ray detectors detected the emitted X-rays.



- •gammow window from 3 to 5 MeV area of astrophysical interest, where the nuclear reactions occur in stars;
- •at energies 9, 10 and 11 MeV not only (p,γ) component is important, also components from (p,n), (p,α) , (p,p) reactions appear.

Monte-Carlo simulations using Geant4 code

Experimental Astrophysics

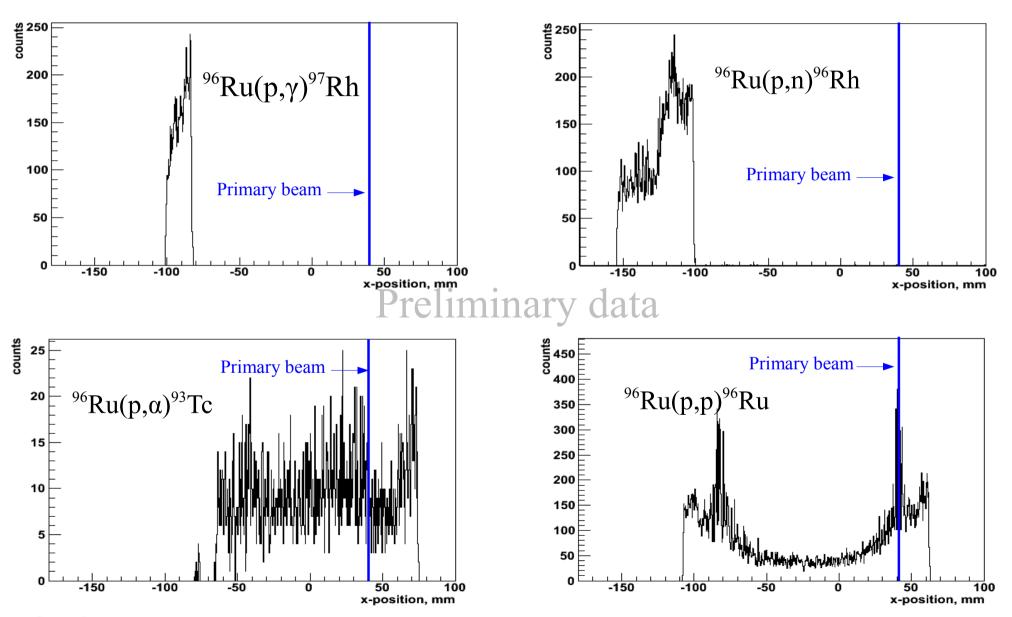


- •describe the part of ESR from H₂-target to silicon detectors;
- •simulate track of secondary particles through the experimental setup;
- •sensitive to the beam position;
- •calculate particles momentum.

Nuclear reaction channels of proton-induced reactions on ruthenium-96 isotopes

Experimental Astrophysics

Goethe-Universität Frankfurt am Mair



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10.10.2011, 8-th International Conference on Nuclear Physics at Storage Rings, Frascati, Italy

Measurement of 96 Ru(p, γ) cross section relative to the 96 Ru e- pick up cross section:

 $Z = \sigma * N * \Phi * t$

Z – quantity of nuclei, produced in reaction (peak area),

 σ - cross section,

N - quantity of target nuclei/area,

 Φ - quantity of incoming particles (flux),

t - time of measurement.

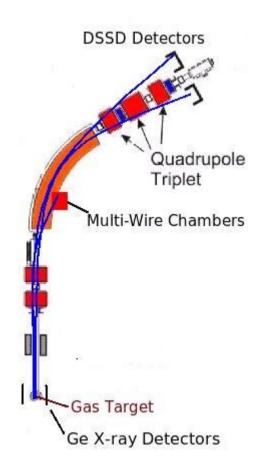
Cross section of electron capture:

$$\sigma_{ec} = Z_{ec} / (N*\Phi)*t,$$

$$N^*\Phi = Z_{ec} / (\sigma_{ec}^* t)$$

Reaction cross section:

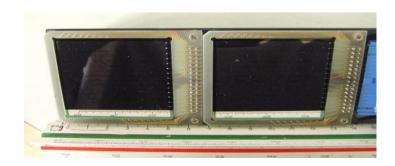
$$\sigma_{p\gamma} = (Z_{p\gamma} * \sigma_{ec})/Z_{ec}$$



Double Sided Silicon Strip Detectors

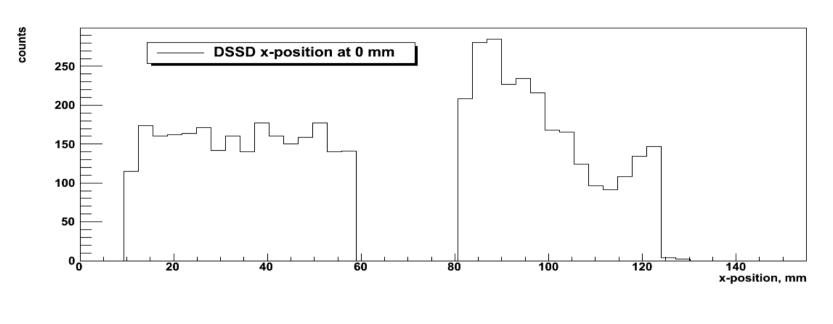
Experimental Astrophysics

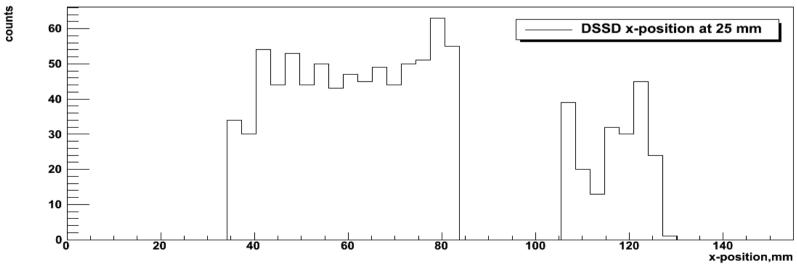
- 2 silicon detectors for measuring position of reaction products;
- the distance between detectors was 21 mm;
- each detector had 16 strips in X- and Y-direction;
- strip pitch = 3.0 mm, strip length = 49.5 mm;
- the detection efficiency for (p,γ) events was about 100%;
- possibilities to get the full view of particles position spectrum by moving detector from 0 to 25 mm from the beam-line.



Position of secondary particles on silicon detectors at 0 and 25 mm

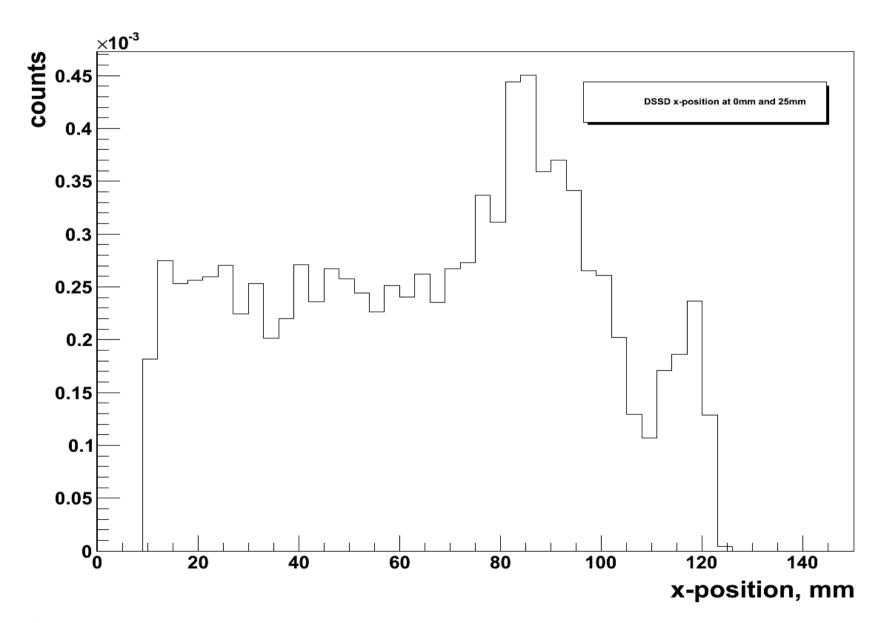
Experimental Astrophysics



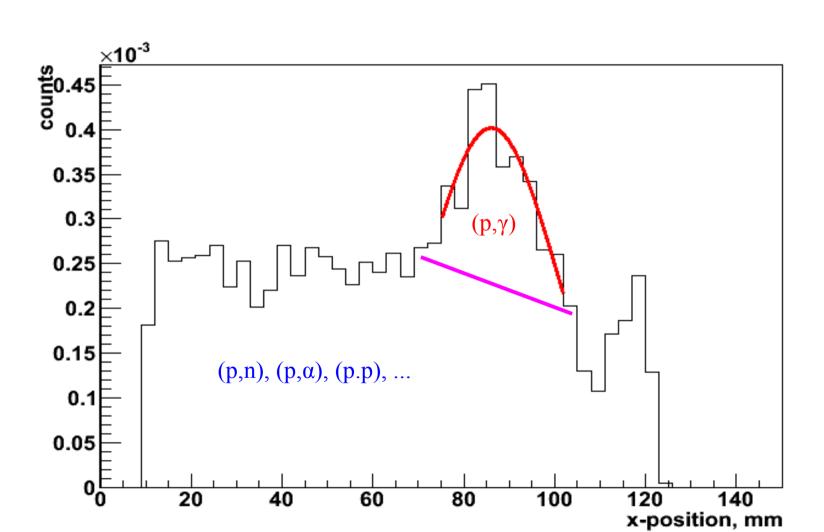


Combined DSSD position at 0 and 25 mm

Experimental Astrophysics



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Measurement of 96 Ru(p, γ) cross section relative to the 96 Ru e- pick up cross section:

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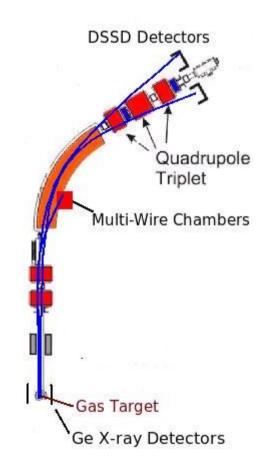
Cross section of electron capture:

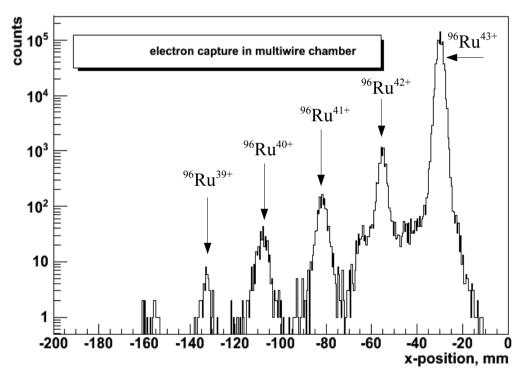
$$\sigma_{ec} = Z_{ec} / (N*\Phi)*t,$$

$$N*\Phi = Z_{ec}/(\sigma_{ec}*t)$$

Reaction cross section:

$$\sigma_{p\gamma} = (Z_{p\gamma} * \sigma_{ec})/Z_{ec}$$





- used for electron pick-up reactions;
- measured secondary particles position;
- able to handle the high flux of down charged ions from cooled heavy- ion beams;
- have efficiency close to 100%;
- particle position does not depend on the angle of incidence of the particle;
- have a good time resolution.

Preliminary result of cross-section

Experimental Astrophysics

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Preliminary result of cross section for 96 Ru(p, γ) 97 Rh reaction:

$$\sigma_{p\gamma} = 1.5 \div 3.5 \text{ mb}$$

Summary

- Simulations with Geant4 code showed particles track through the experimental setup and made a good prediction of the shape for all nuclear components.
- Comparison of simulated 96 Ru(p, γ) 97 Rh position spectrum and experimental results are in good agreement.
- Preliminary results of cross-section have been shown.
- Analysis of this experiment are in a progress.
- Future plans are to place detector at the end of the first dipole after the gas target for registration all nuclear reactions components.
- Detector installation in vacuum inside of the ESR.

Acknowledgements

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