

The broad physics case of a muon collider

Nathaniel Craig **UCSB**



Vision circa 1984

ECFA 84/85
CERN 84-10
5 September 1984

Satisfied with these successes, we have now to face deeper questions such as:

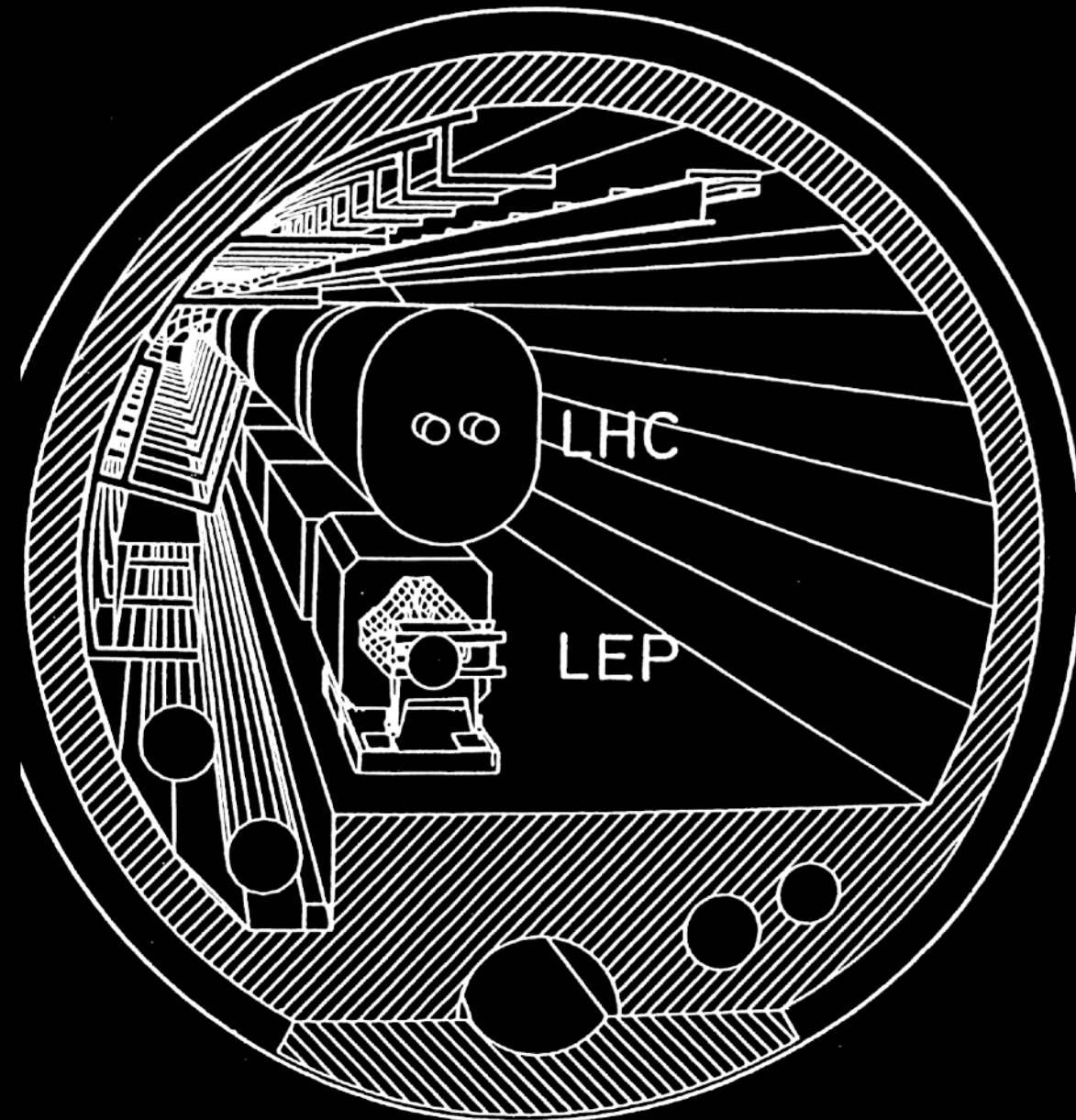
what is the origin of mass?

what kind of unification may exist beyond the standard model?

what is the origin of flavour?

is there a deeper reason for gauge symmetry?

We have simply too many a priori plausible hypotheses concerning the nature of symmetry breaking in the standard model. Experimentation in the TeV range at the constituent level is bound to provide most essential clues, and the present successes of the $p\bar{p}$ collider are a very strong encouragement to go to higher energies and to higher luminosities in hadron-hadron collisions.



LARGE HADRON COLLIDER
IN THE LEP TUNNEL

Vol. I

PROCEEDINGS OF THE ECFA-CERN WORKSHOP

held at Lausanne and Geneva,
21-27 March 1984

Vision circa 2023

Vision circa 2023

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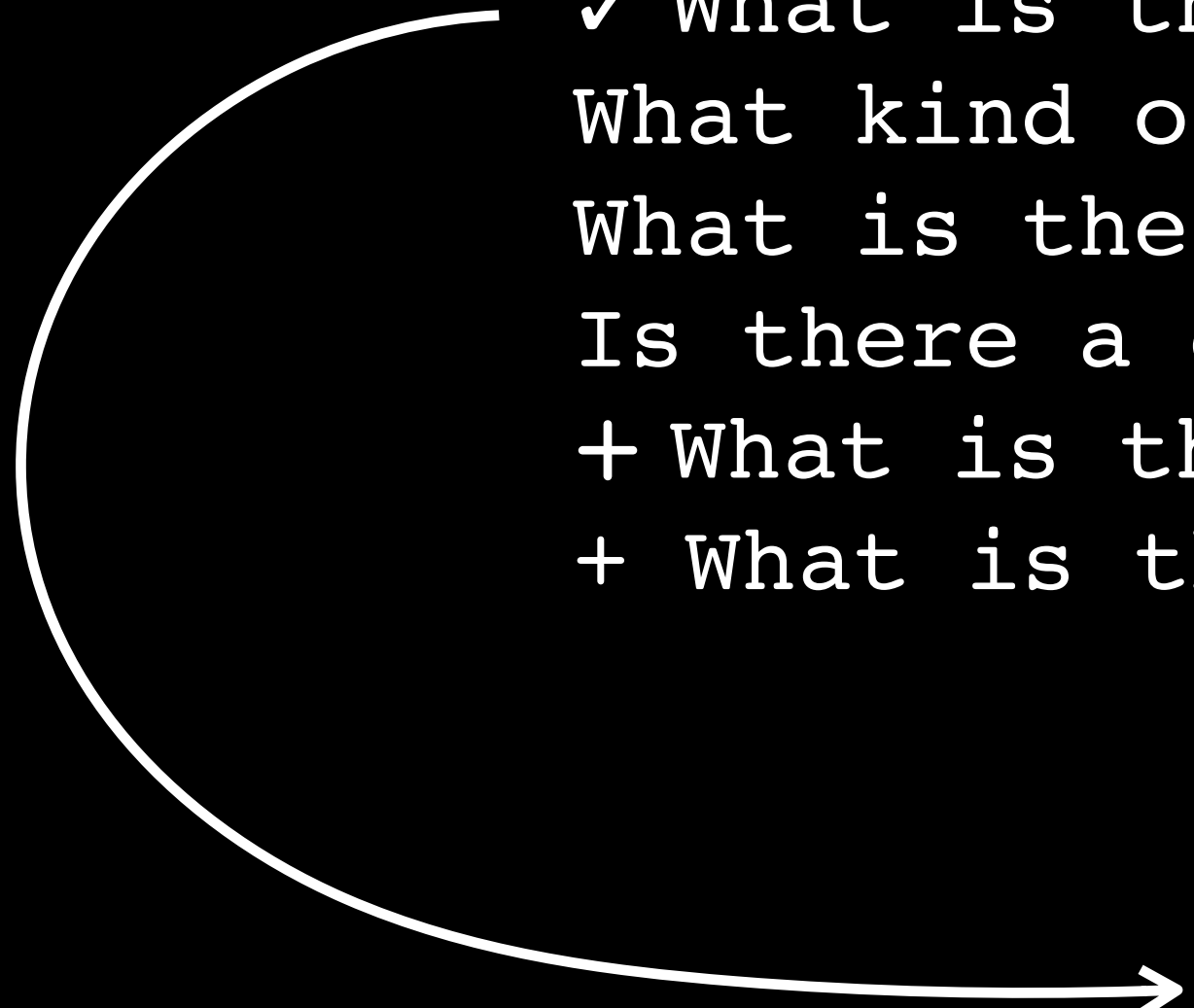
Vision circa 2023

- ✓ What is the origin of mass?
- What kind of unification may exist?
- What is the origin of flavor?
- Is there a deeper reason for gauge symmetry?
- + What is the nature of dark matter?

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A Higgs! Yet:

“The more ambitious goal...is to identify and understand the nature of electroweak symmetry breaking, the asymmetry that is key to the material universe. The Higgs boson is but its herald.”

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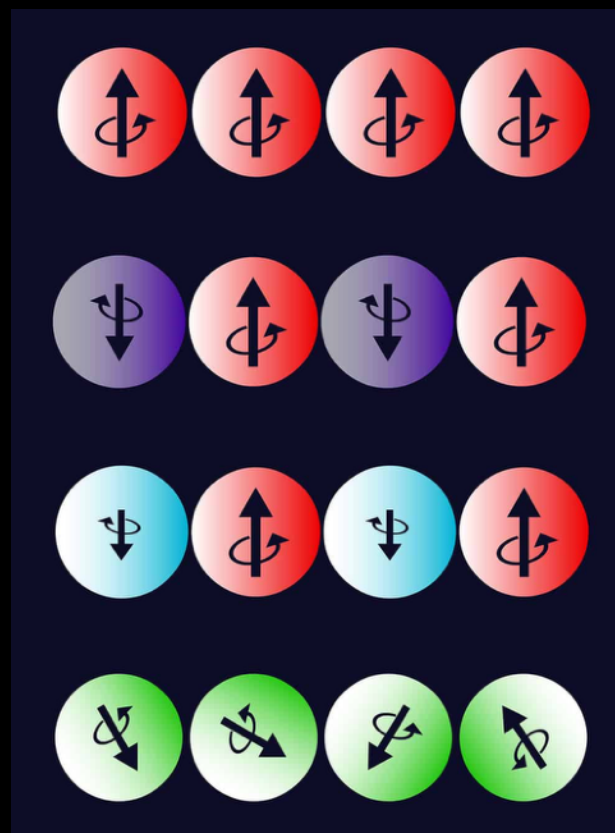
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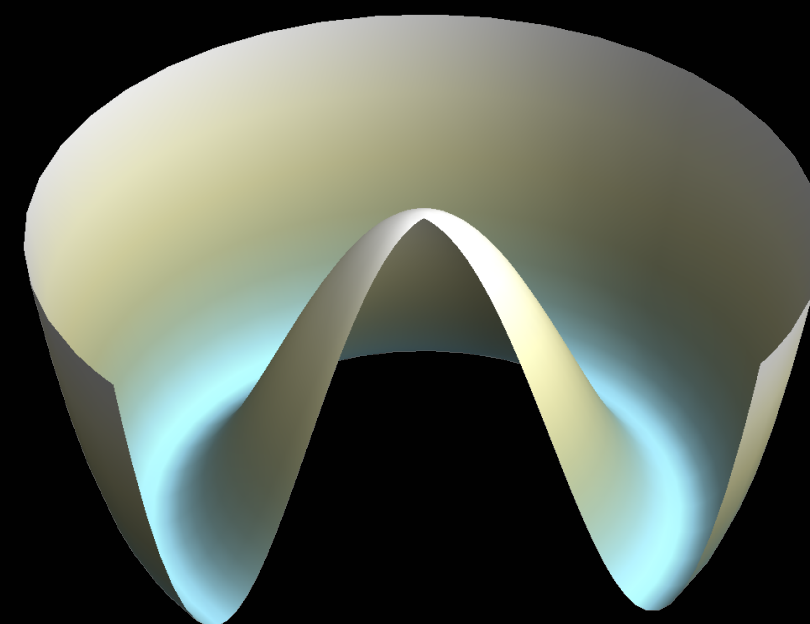
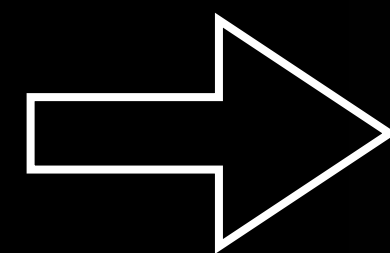
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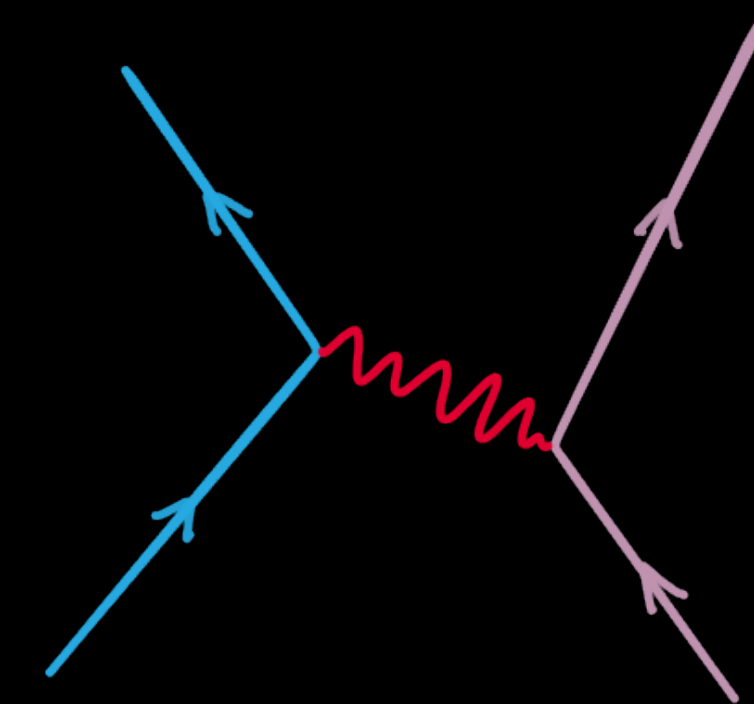
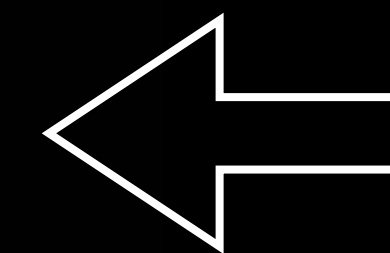
A superconducting analogy:



High-Tc Superconductors

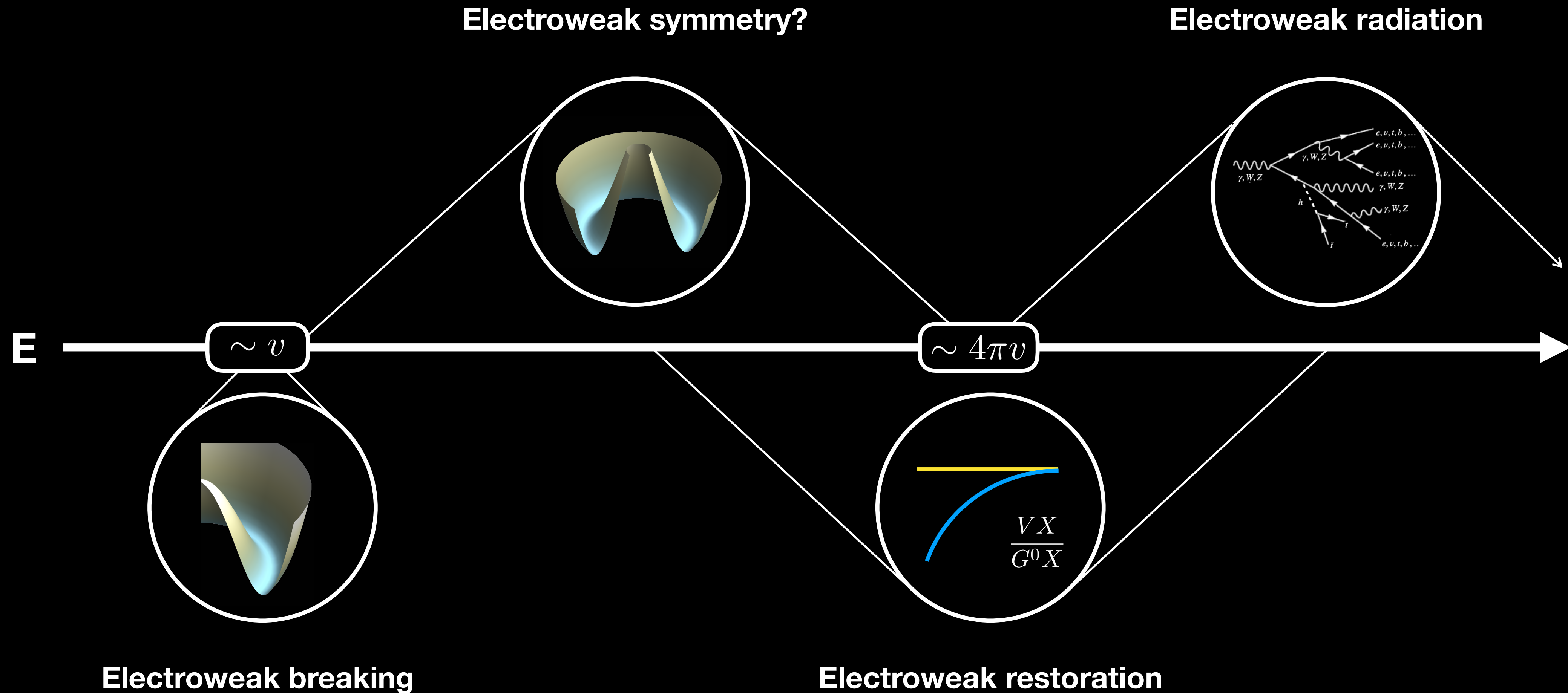


Ginzburg-Landau Theory

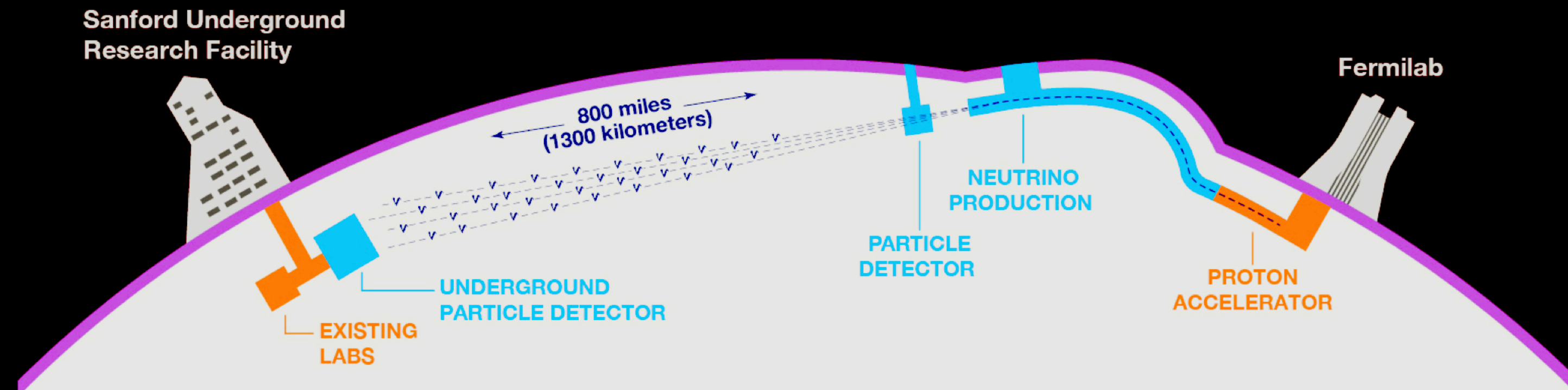


Low-Tc Superconductors (BCS)

The Electroweak Era



The Electroweak Era



Neutrinos: the purely weak frontier of the Standard Model

Dark sectors: portals are weak interactions & SM electroweak states

weak portal

$$g(W_\mu)^i_j \mathcal{O}_j$$

kinetic mixing portal

$$F^{\mu\nu} \mathcal{O}_{\mu\nu}$$

neutrino portal

$$HL^\alpha \mathcal{O}_\alpha$$

Higgs portal

$$|H|^2 \mathcal{O}$$

The Electroweak Era

...and exploration after the LHC

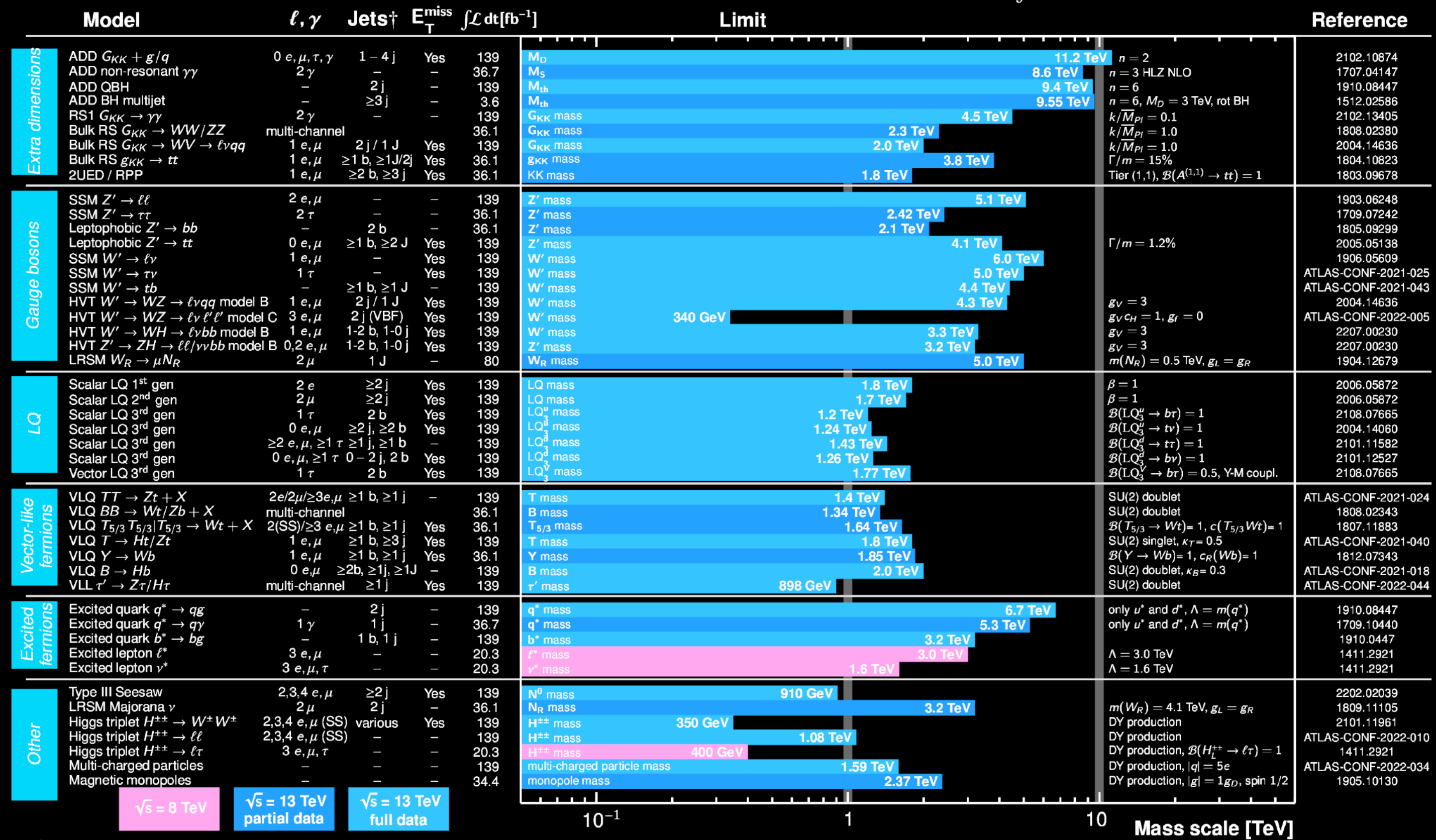
ATLAS Heavy Particle Searches* - 95% CL Upper Exclusion Limits

Status: July 2022

ATLAS Preliminary

$\int \mathcal{L} dt = (3.6 - 139) \text{ fb}^{-1}$

$\sqrt{s} = 8, 13 \text{ TeV}$



The Electroweak Era

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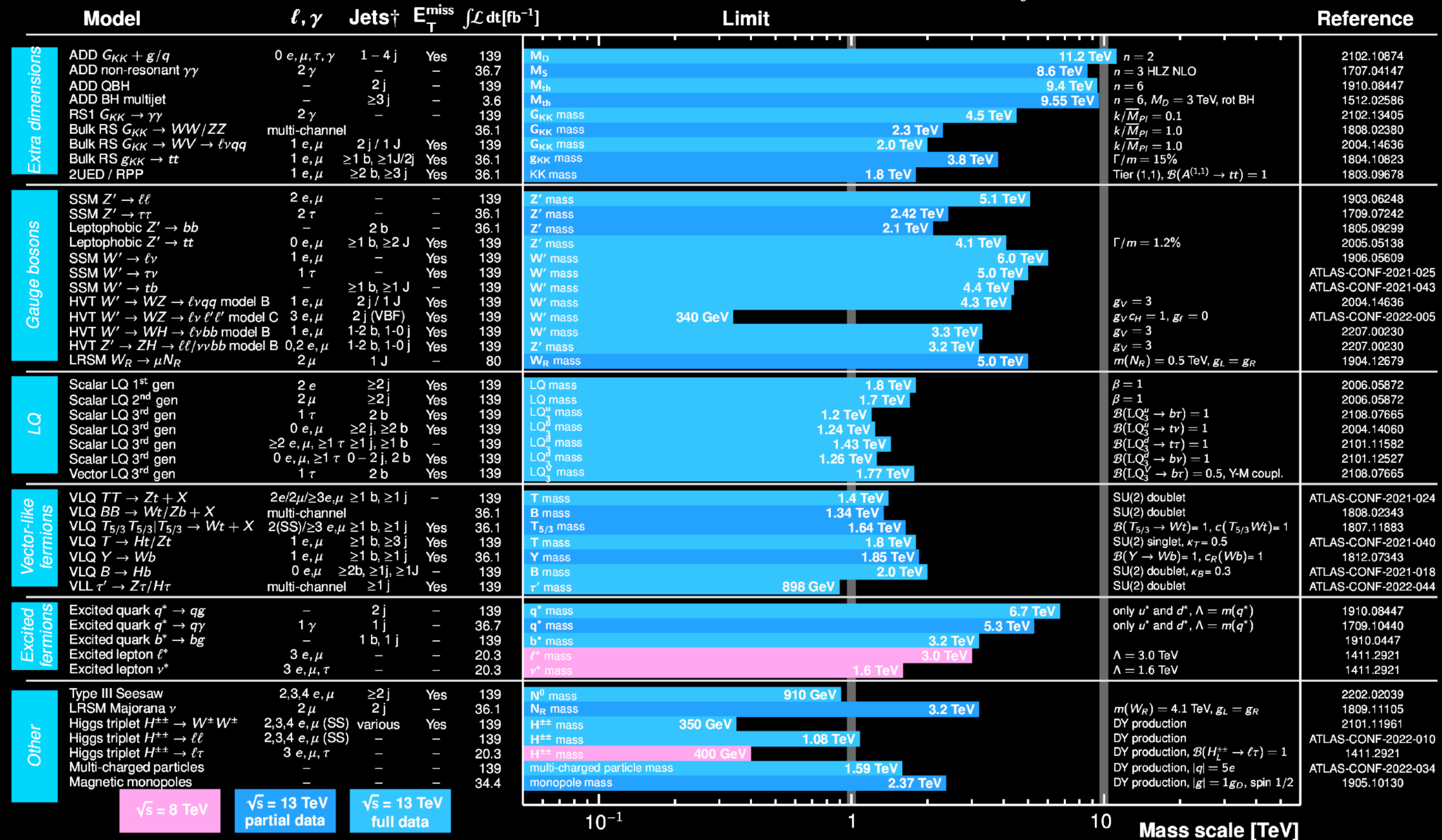
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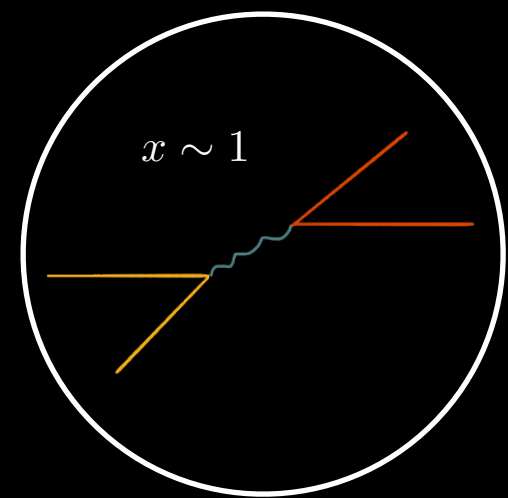


$\sqrt{s} = 8 \text{ TeV}$ $\sqrt{s} = 13 \text{ TeV}$ partial data $\sqrt{s} = 13 \text{ TeV}$ full data

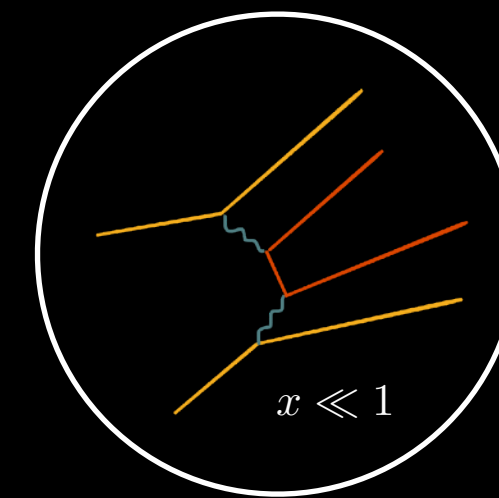
Singlet scalar $(1, 1)_0$
 Doublet scalar $(1, 2)_{1/2}$
 Doublet scalar $(1, 2)_{3/2}$

Muon Colliders for the EWK Era

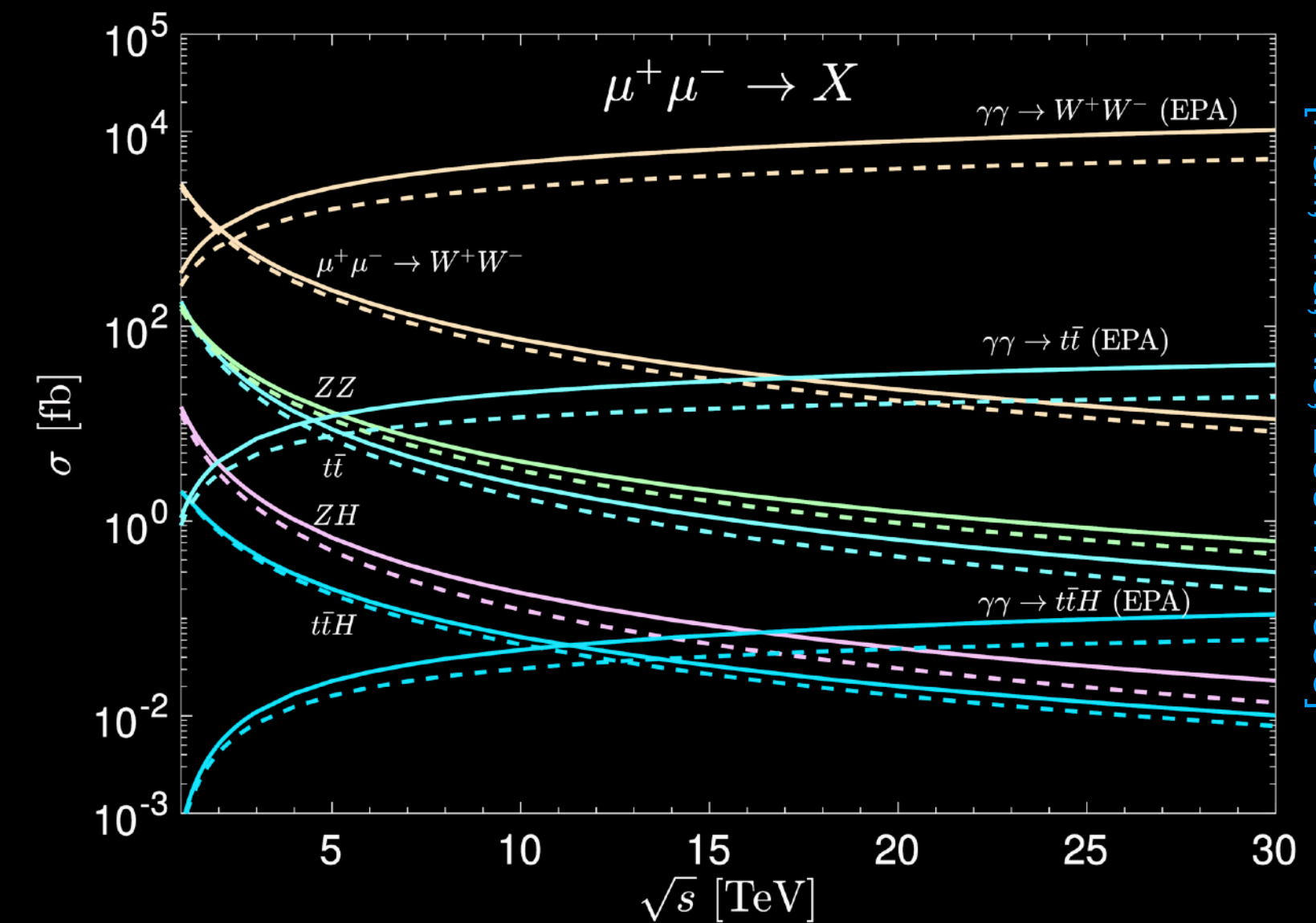
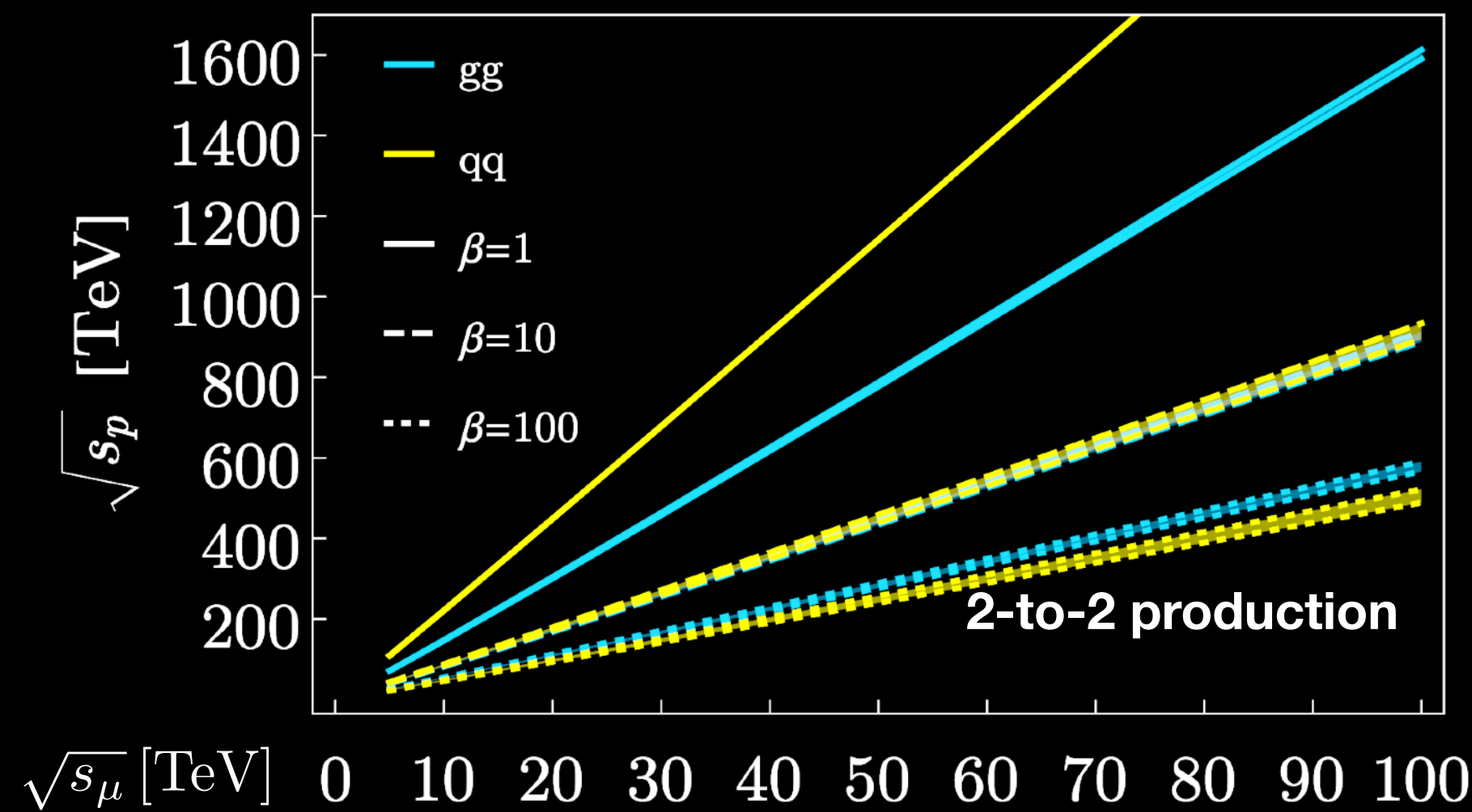
Muon colliders blur the dichotomy of probing microscopic phenomena with **precision** or **energy**.



Muon annihilation
deploys the entire
energy of the collider



Vector boson fusion
leverages the muon's
virtual boson content

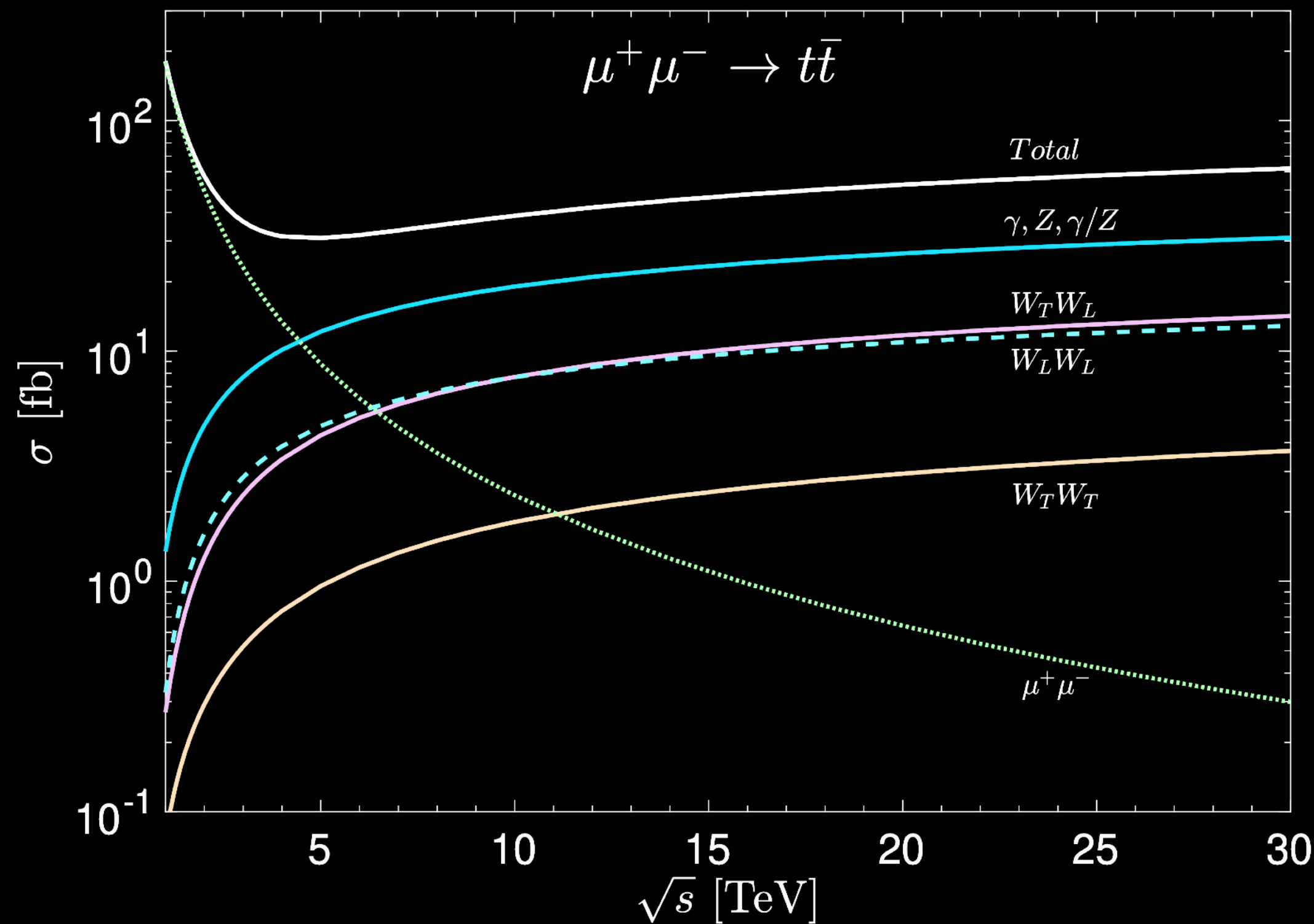


[Han, Ma, Xie, 2007.14300]

Moreover, muon collider energy in a (relatively) clean environment provides **precision from energy**.

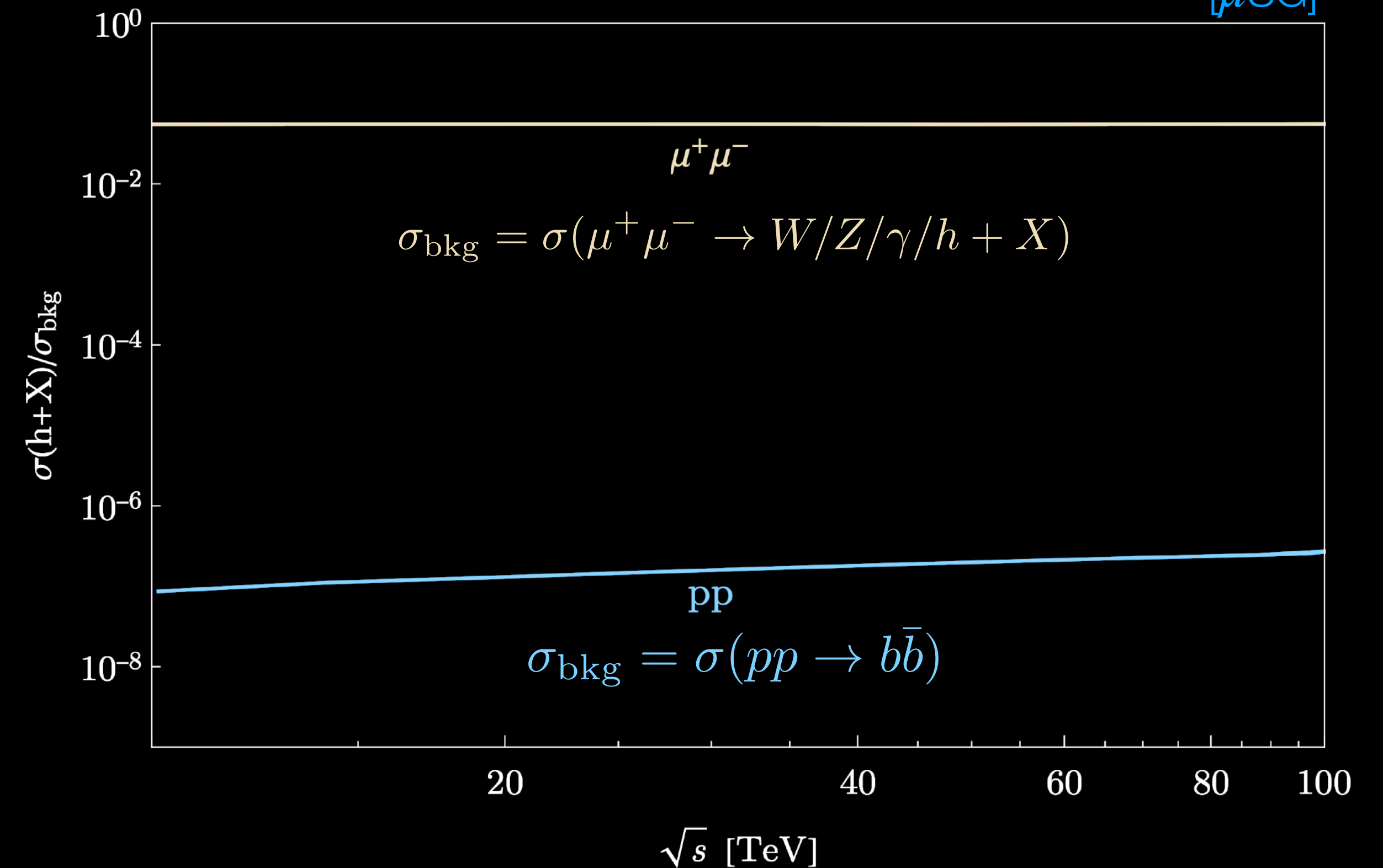
Muon Colliders as EWK Laboratories

[Han, Ma, Xie, 2007.14300]



Longitudinal polarizations play a key role, making an extraordinary laboratory for EWSB

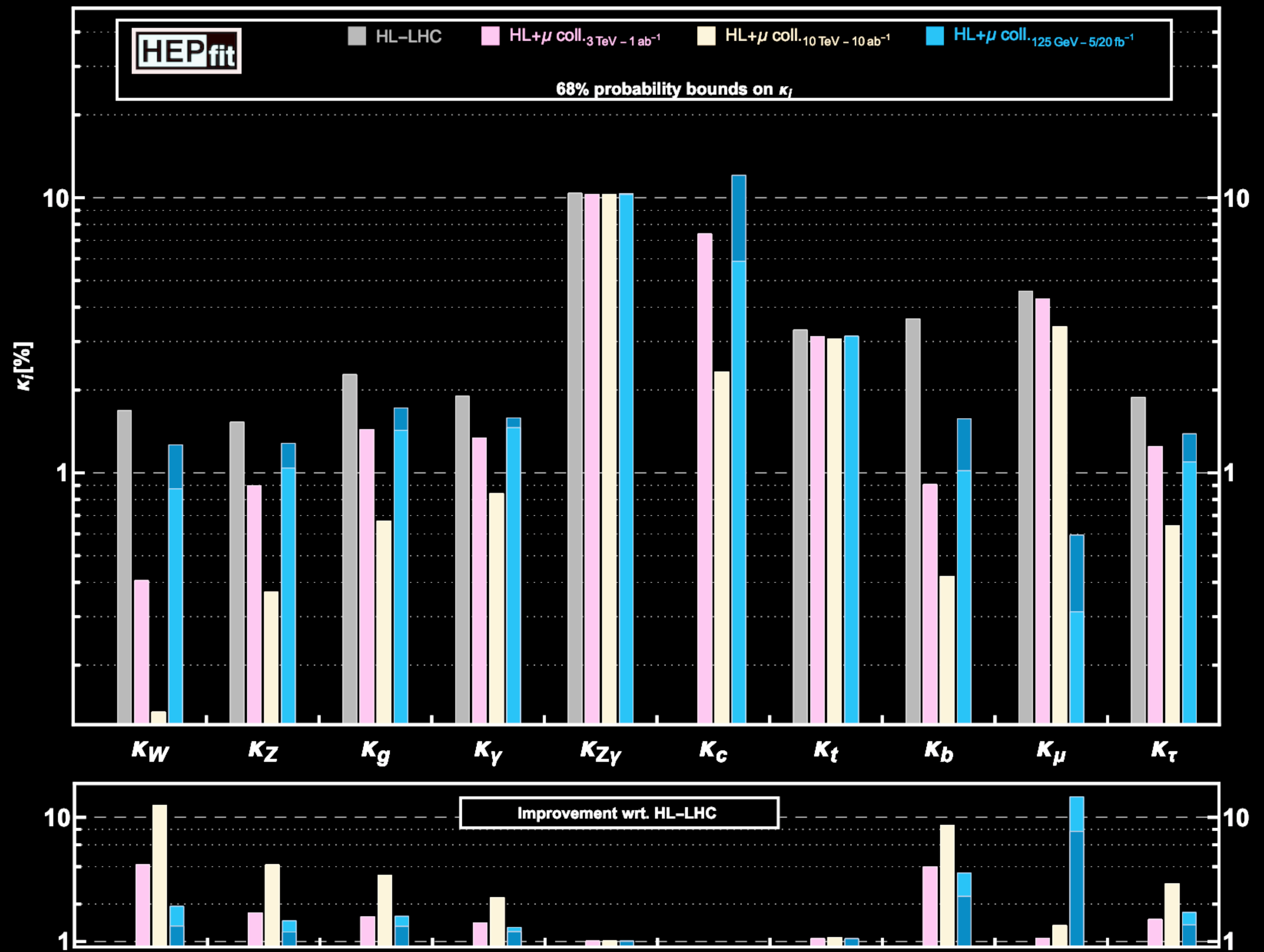
[μ SG]



Dominant signals and backgrounds both have electroweak cross sections

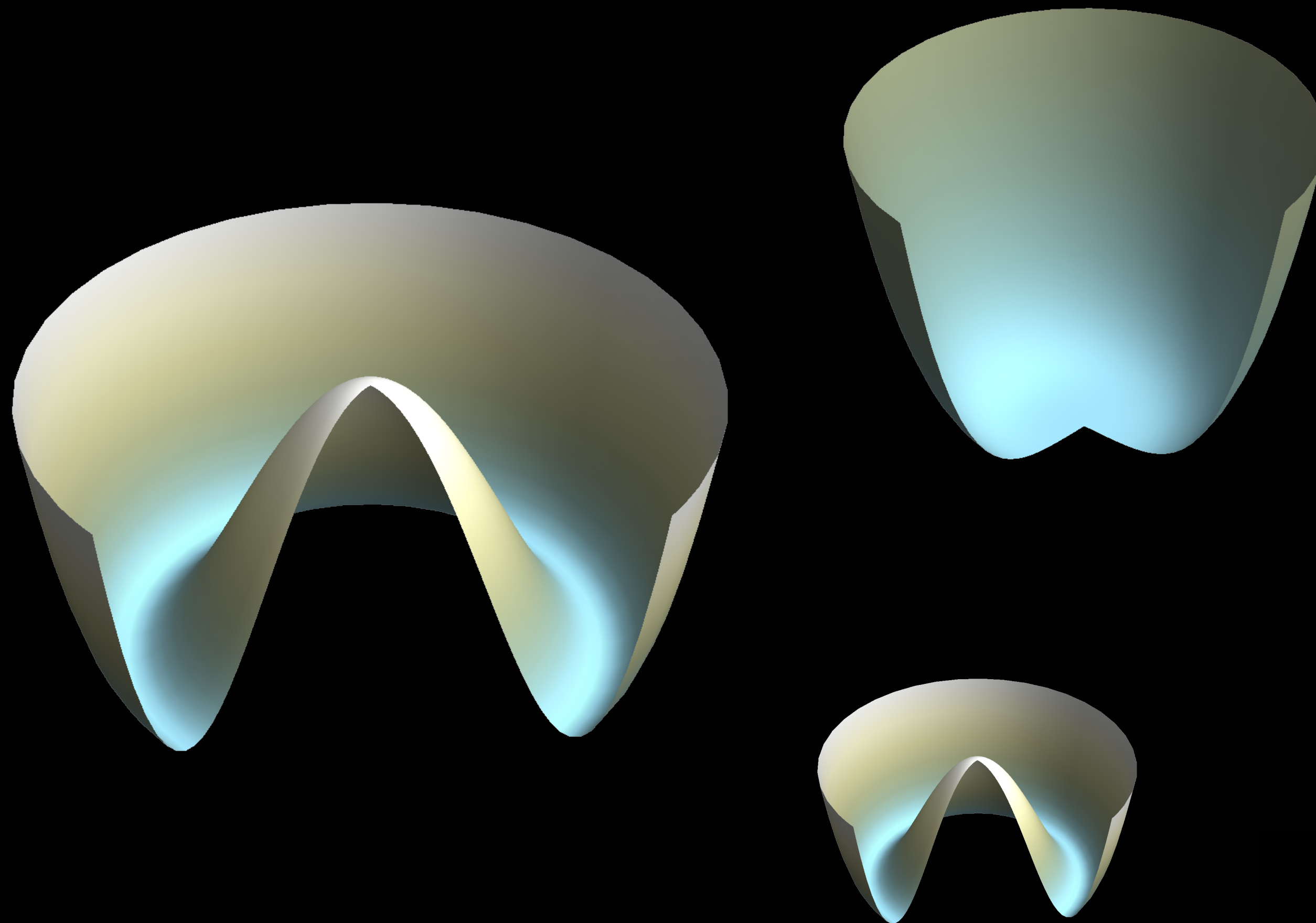
**How do muons illuminate the
physics vision?**

What is the Origin of Mass ?

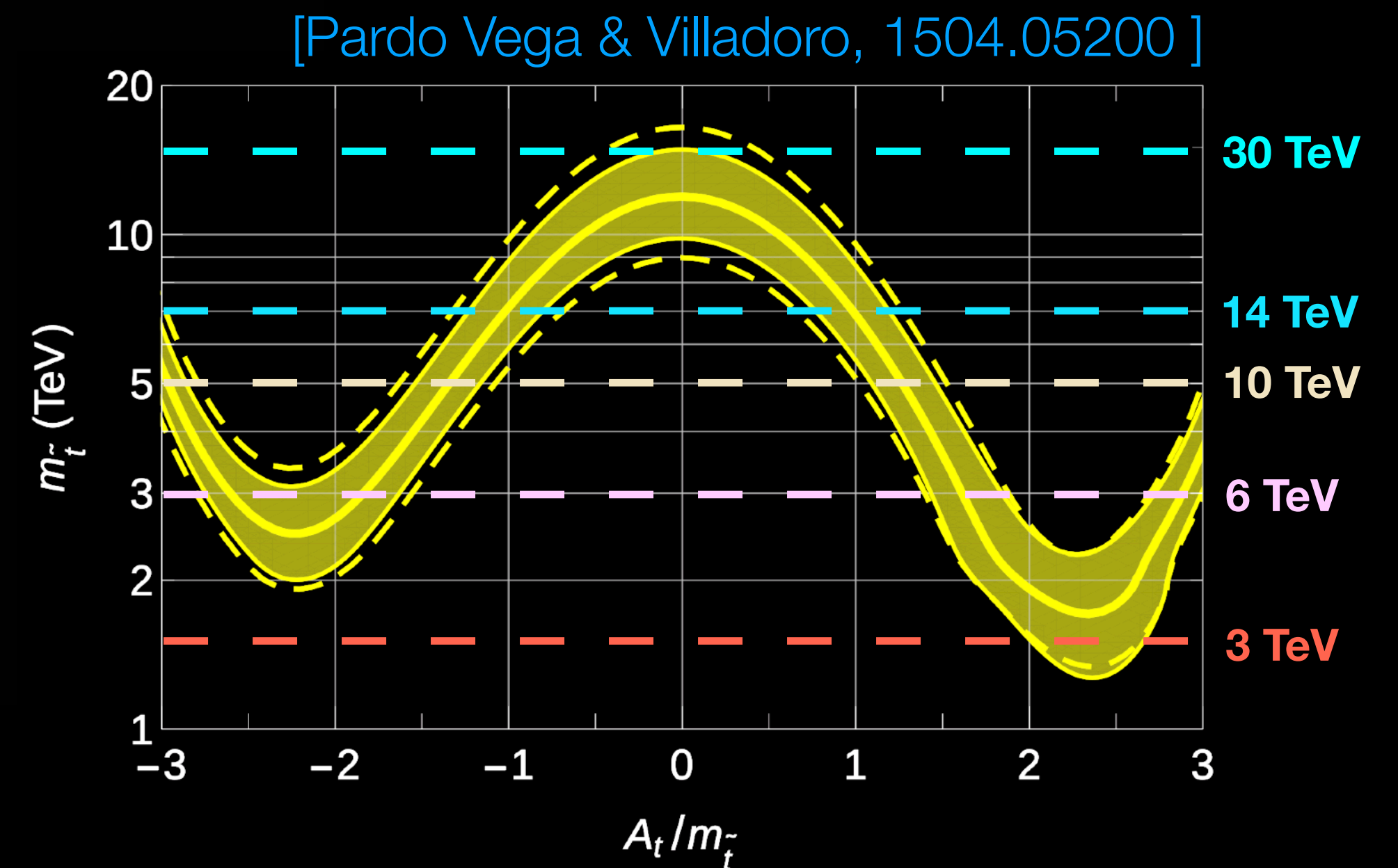


What is the BCS Theory of EWSB?

Theories that predict the Higgs mass & EWSB provide sharp targets for new physics.



**Direct targets set by the
observed Higgs mass
(e.g. supersymmetry)**



What is the BCS Theory of EWSB?

Theories that predict the Higgs mass & EWSB provide sharp targets for new physics.

Compositeness leaves fingerprints in EFT:

$$\mathcal{O}_{2W} = (D_\mu W^{\mu\nu,a})^2$$

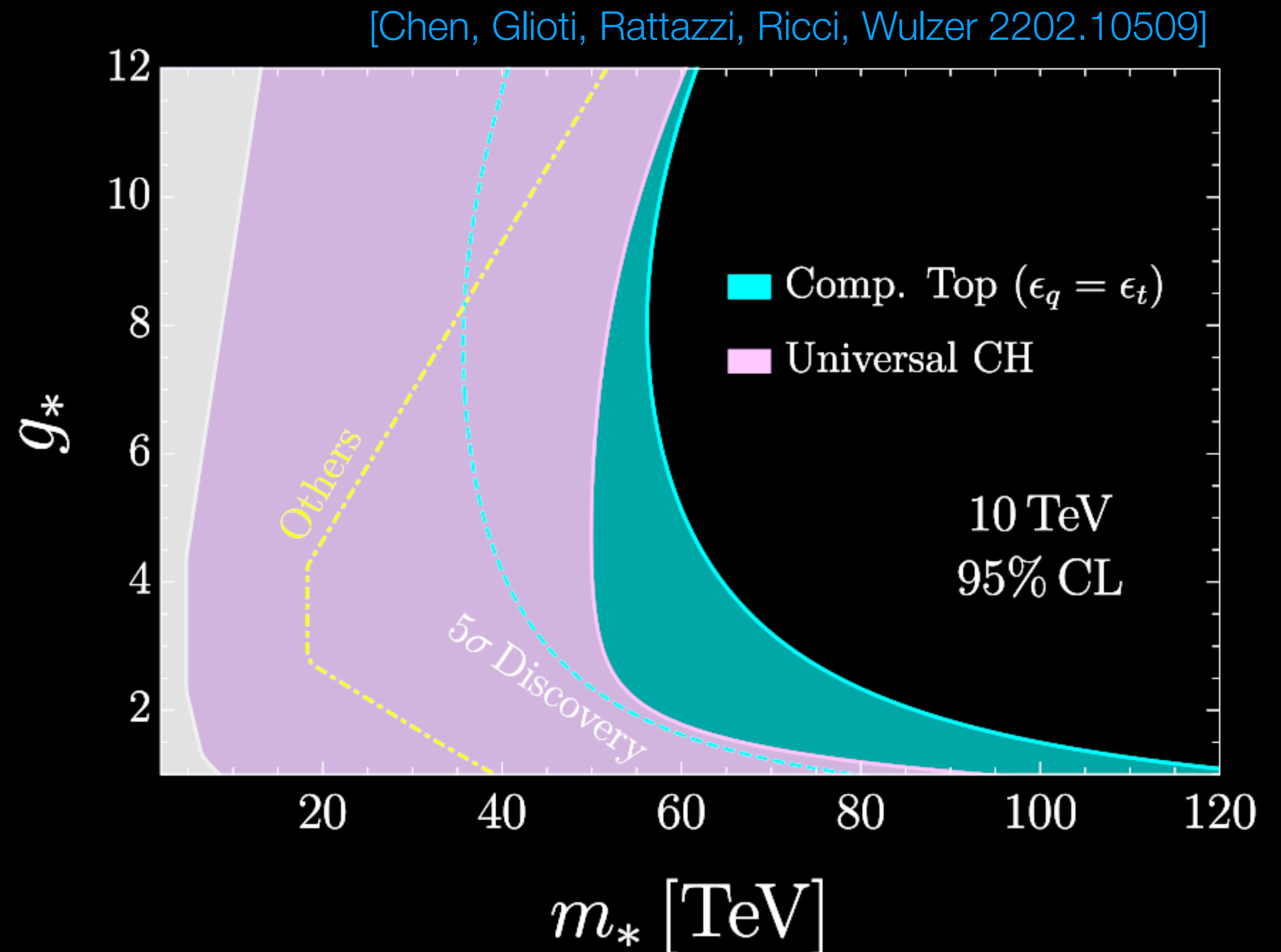
$$\mathcal{O}_{2B} = (\partial_\mu B^{\mu\nu})^2$$

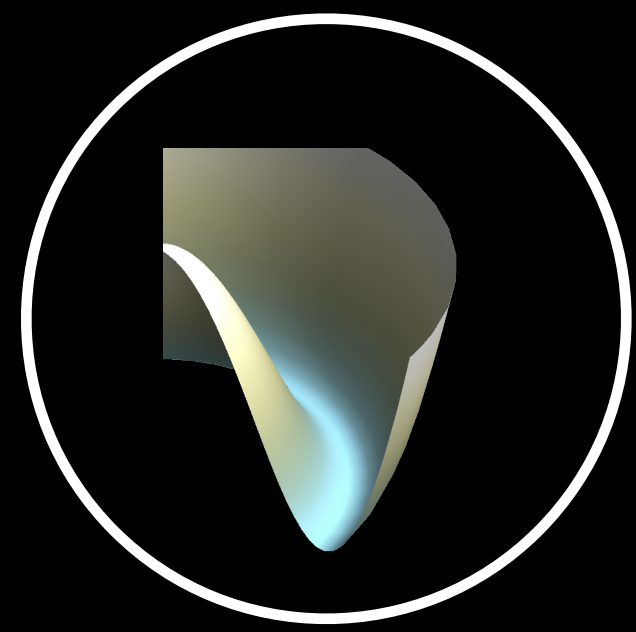
$$\mathcal{O}_W = ig(H^\dagger \sigma^a D_\mu H) D^\nu W_{\mu\nu}^a$$

$$\mathcal{O}_B = ig'(H^\dagger D_\mu H) \partial^\nu B_{\mu\nu}$$

$$\mathcal{O}_{tD} = (\bar{t}\gamma^\mu t)(\partial^\nu B_{\mu\nu})$$

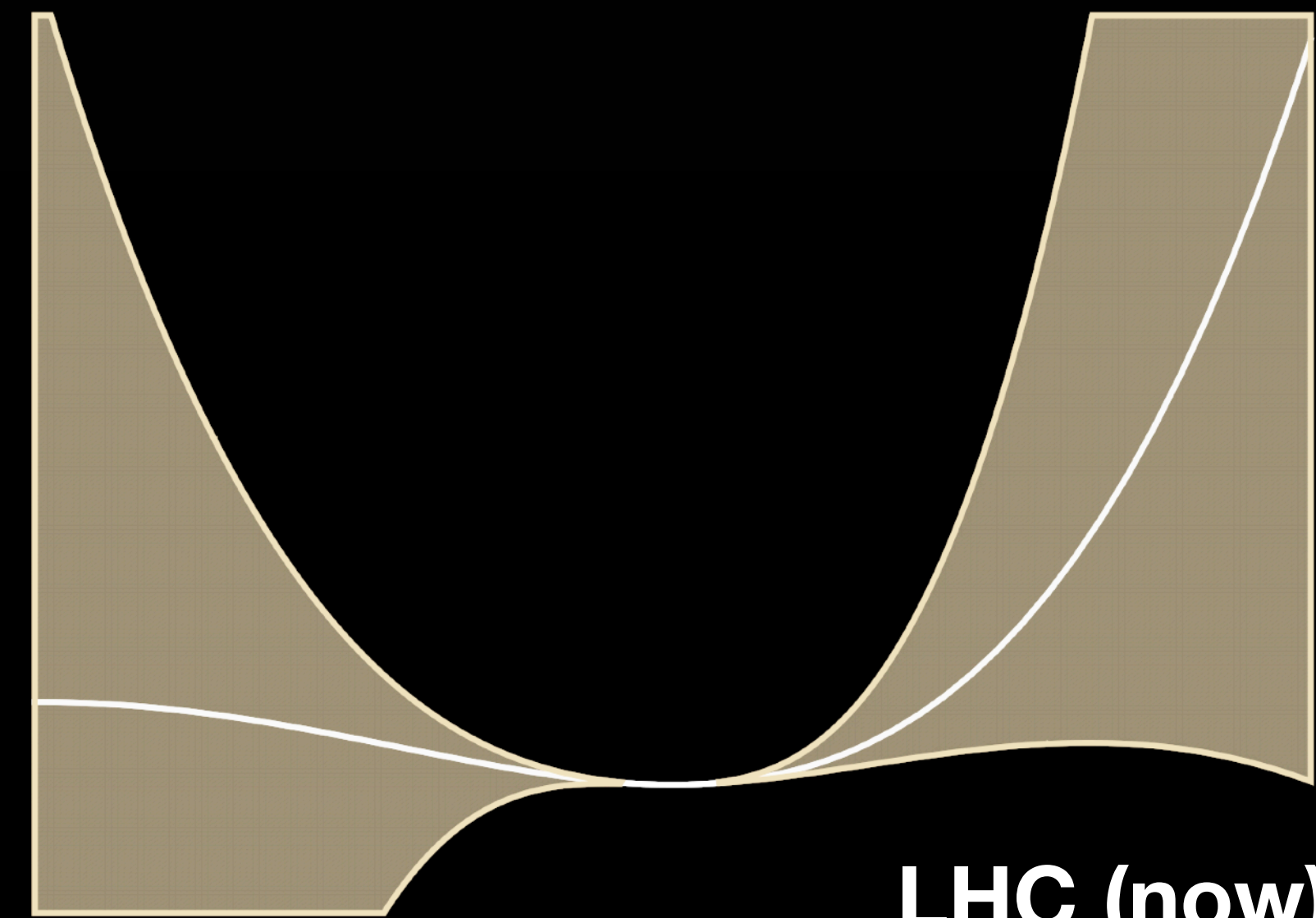
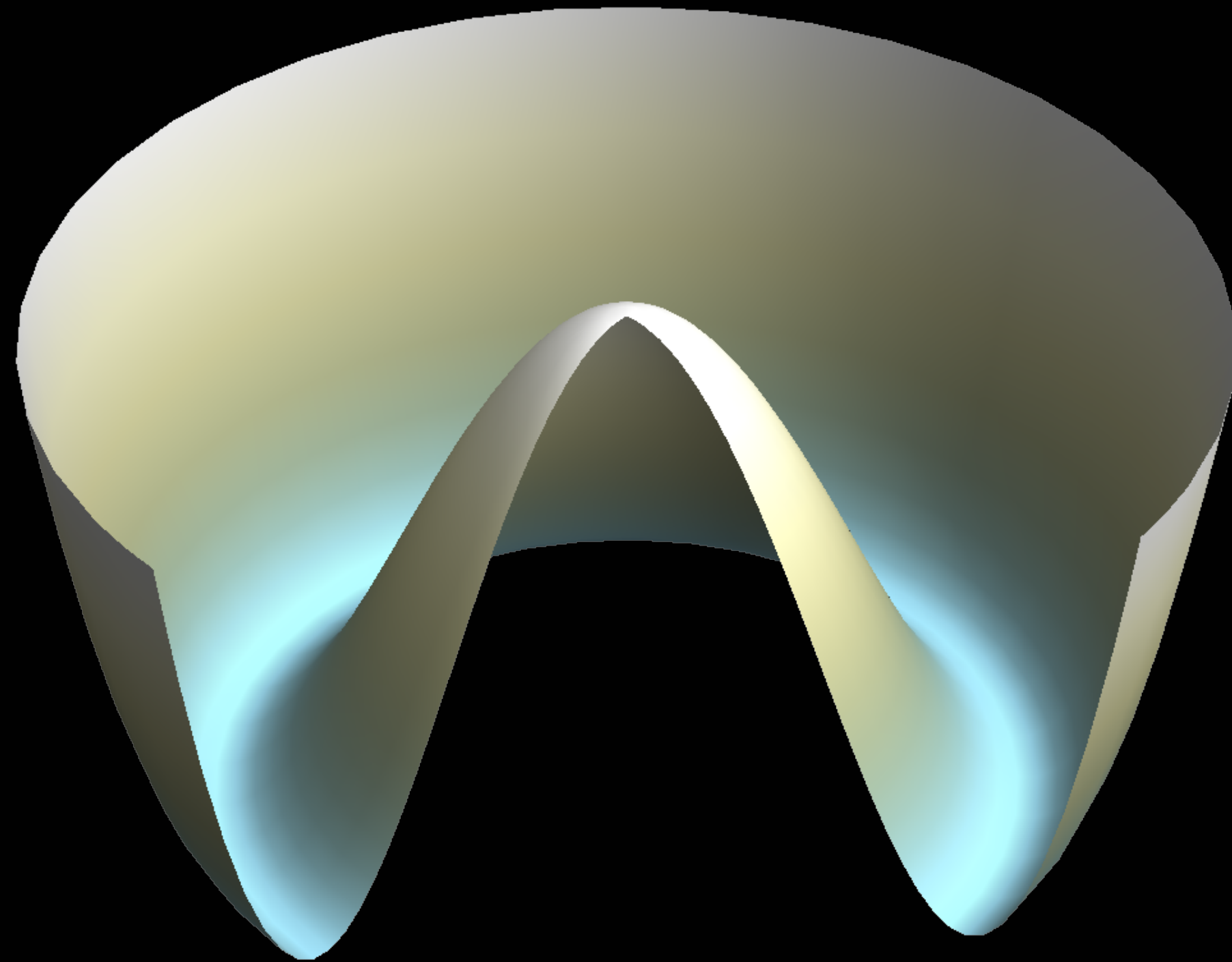
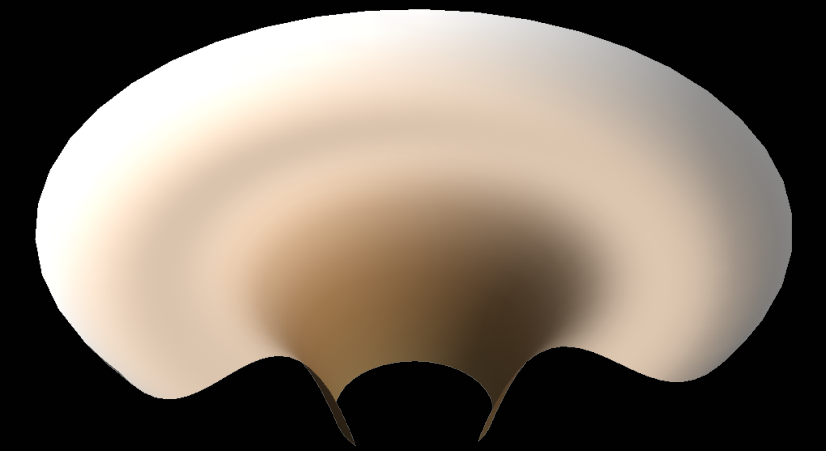
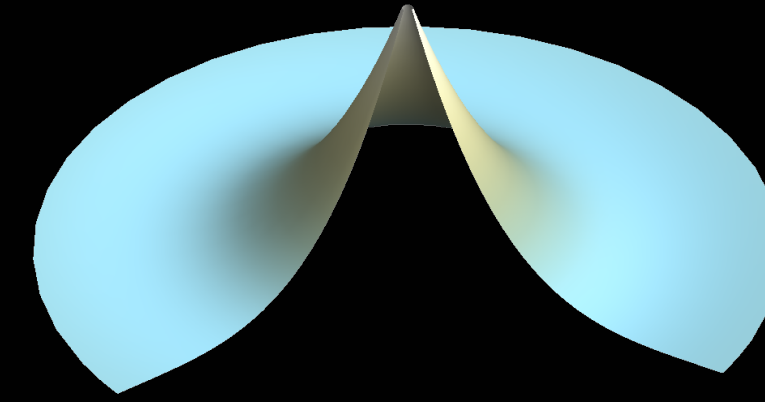
Splendid example of precision from energy...
(See also [Buttazzo, Franceschini, Wulzer 2012.11555])





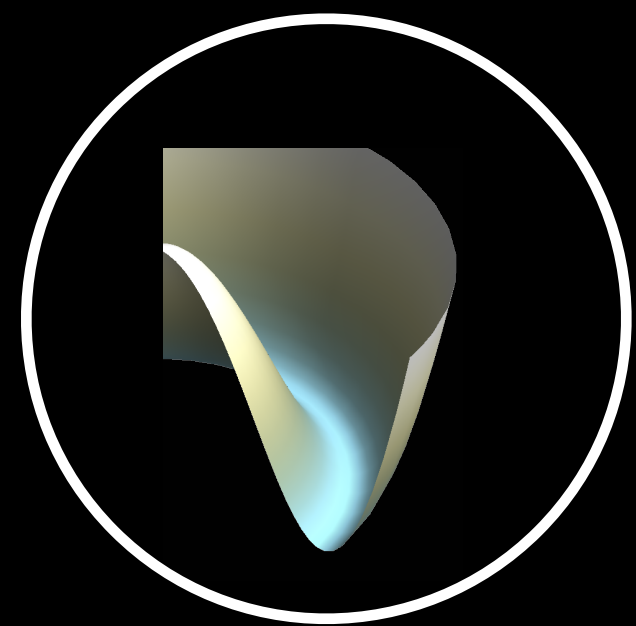
Electroweak Breaking

Our current experimental knowledge of the Higgs potential is far from the cartoon we usually show...



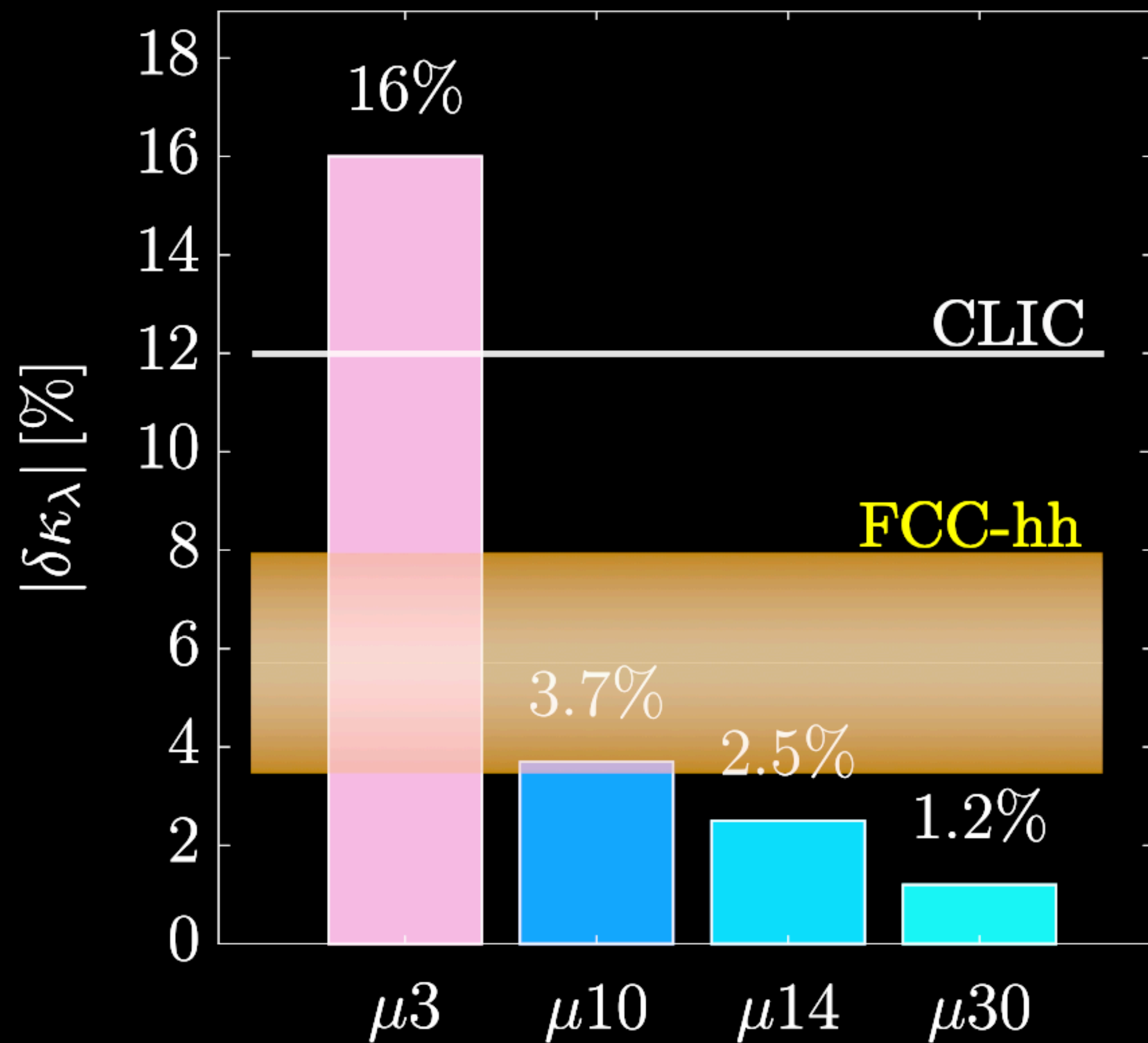
LHC (now)

(If you like this way of presenting Higgs self-coupling precision, feel free to use it w/ credit to R. Petrossian-Byrne.)

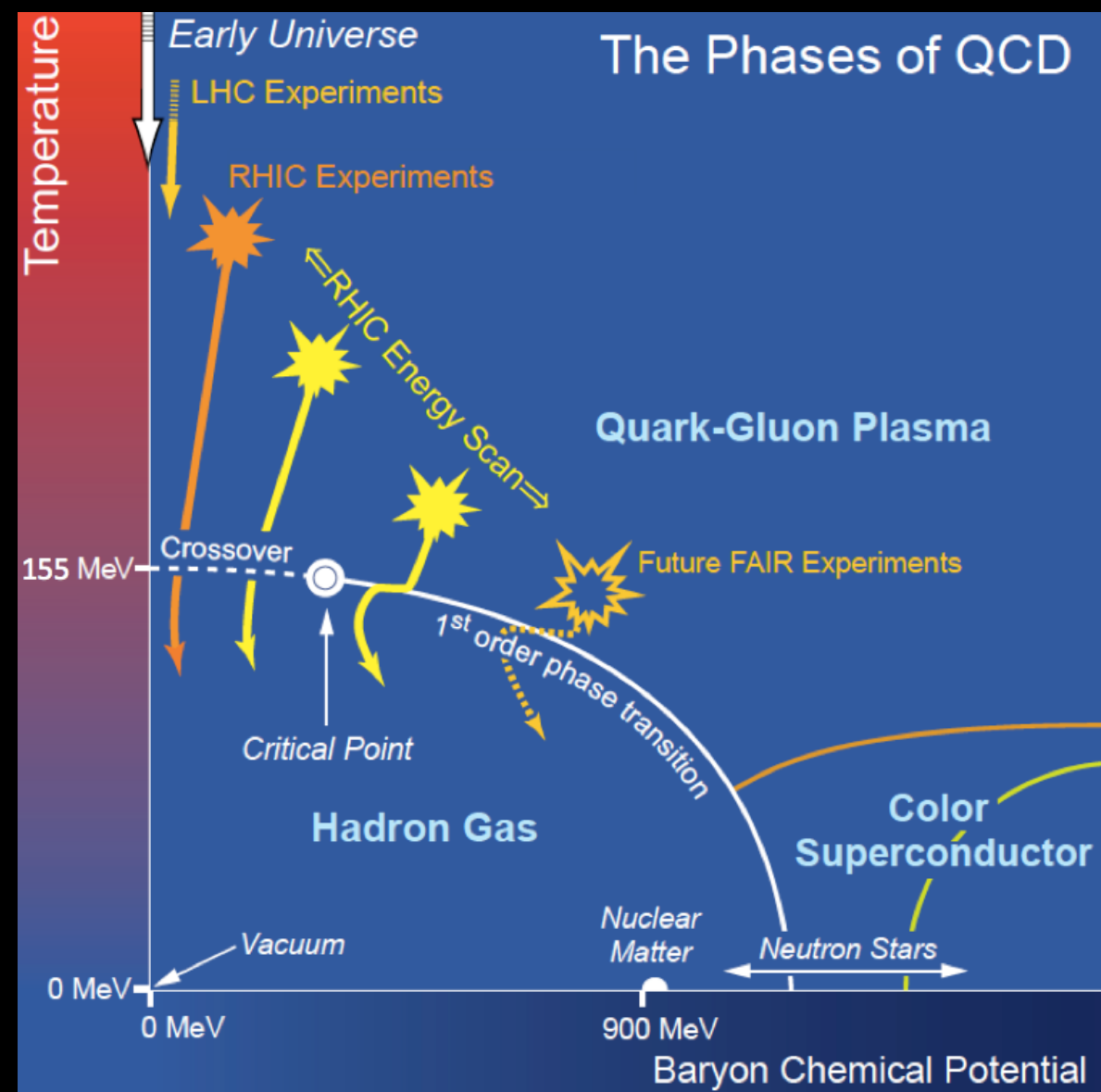


Electroweak Breaking

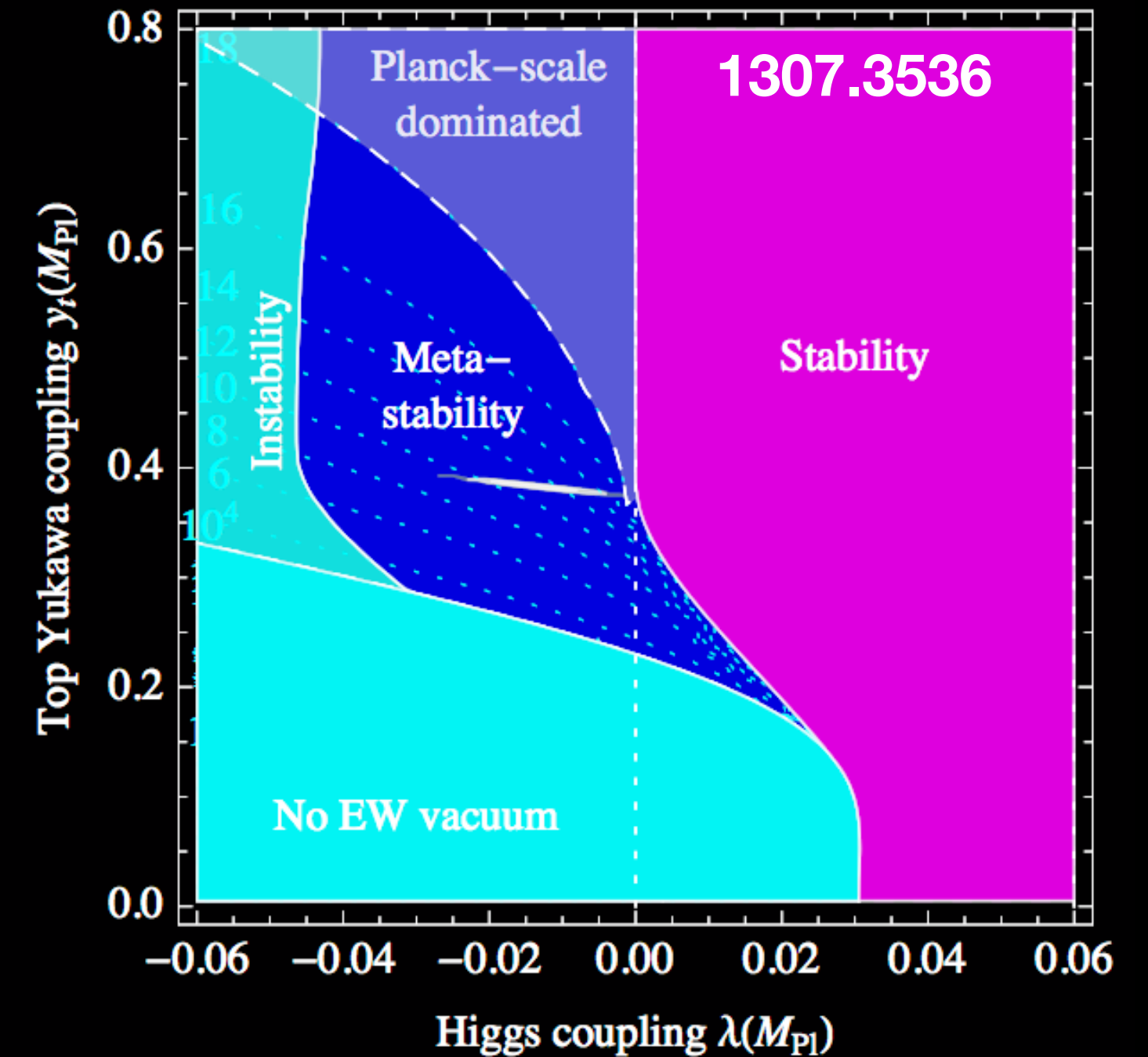
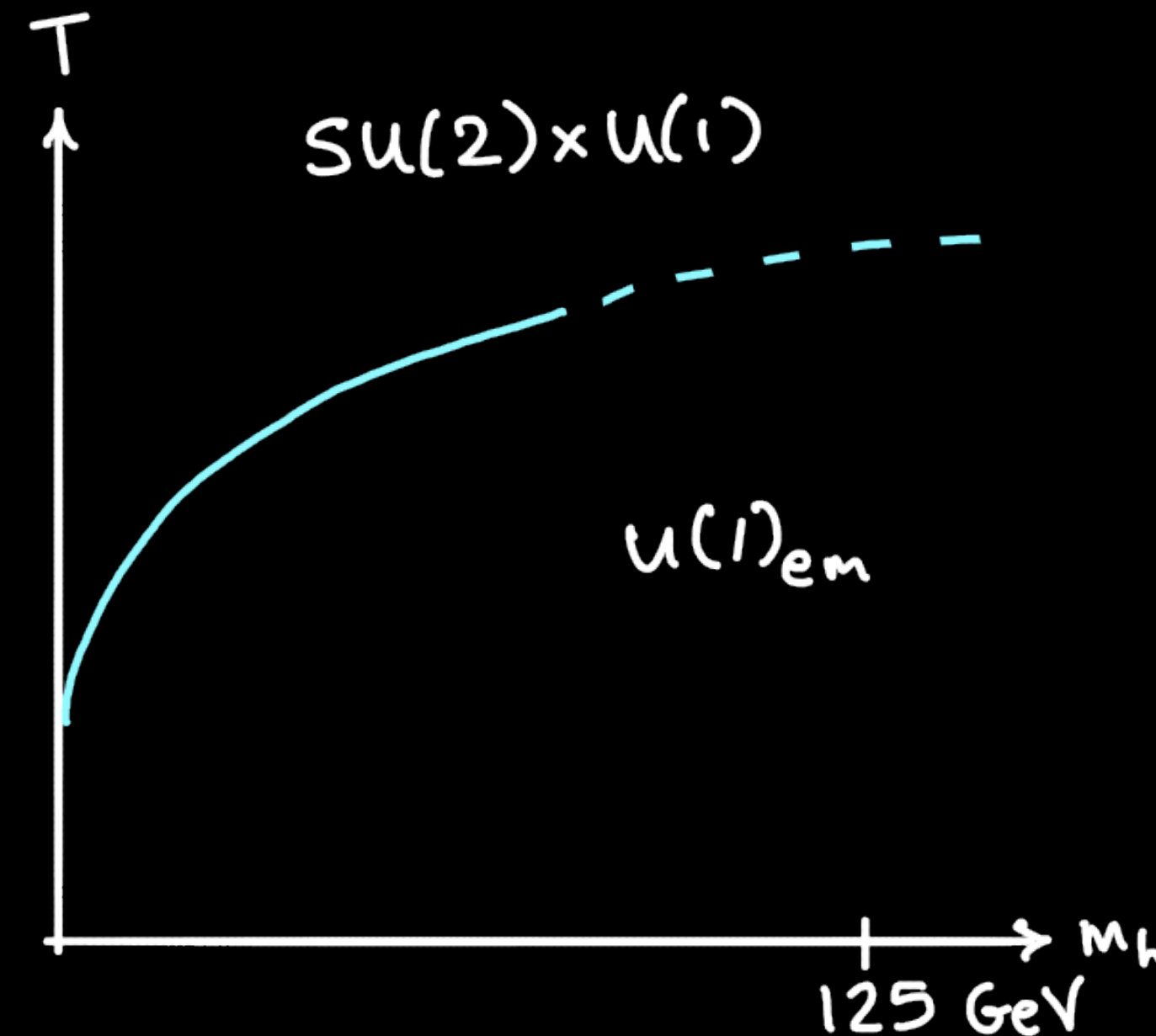
[Accettura et al. 2303.08533]



The birth and death of the Universe?

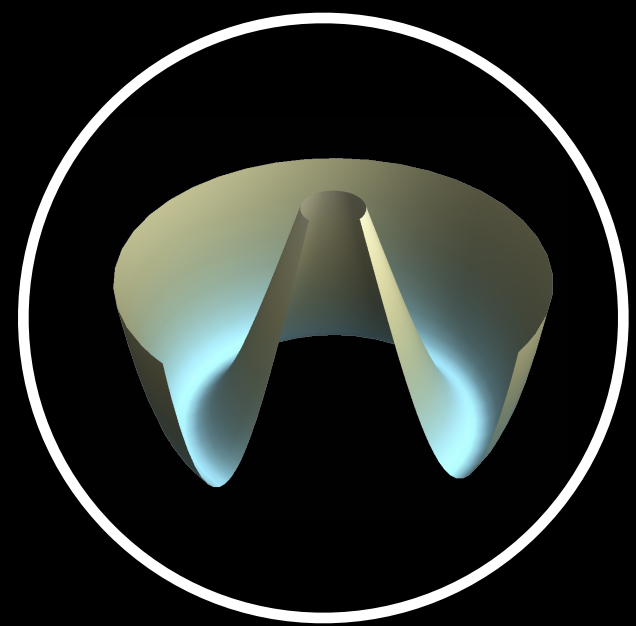


[2015 NSAC LRP / 1501.06477]



First-order electroweak phase transition?

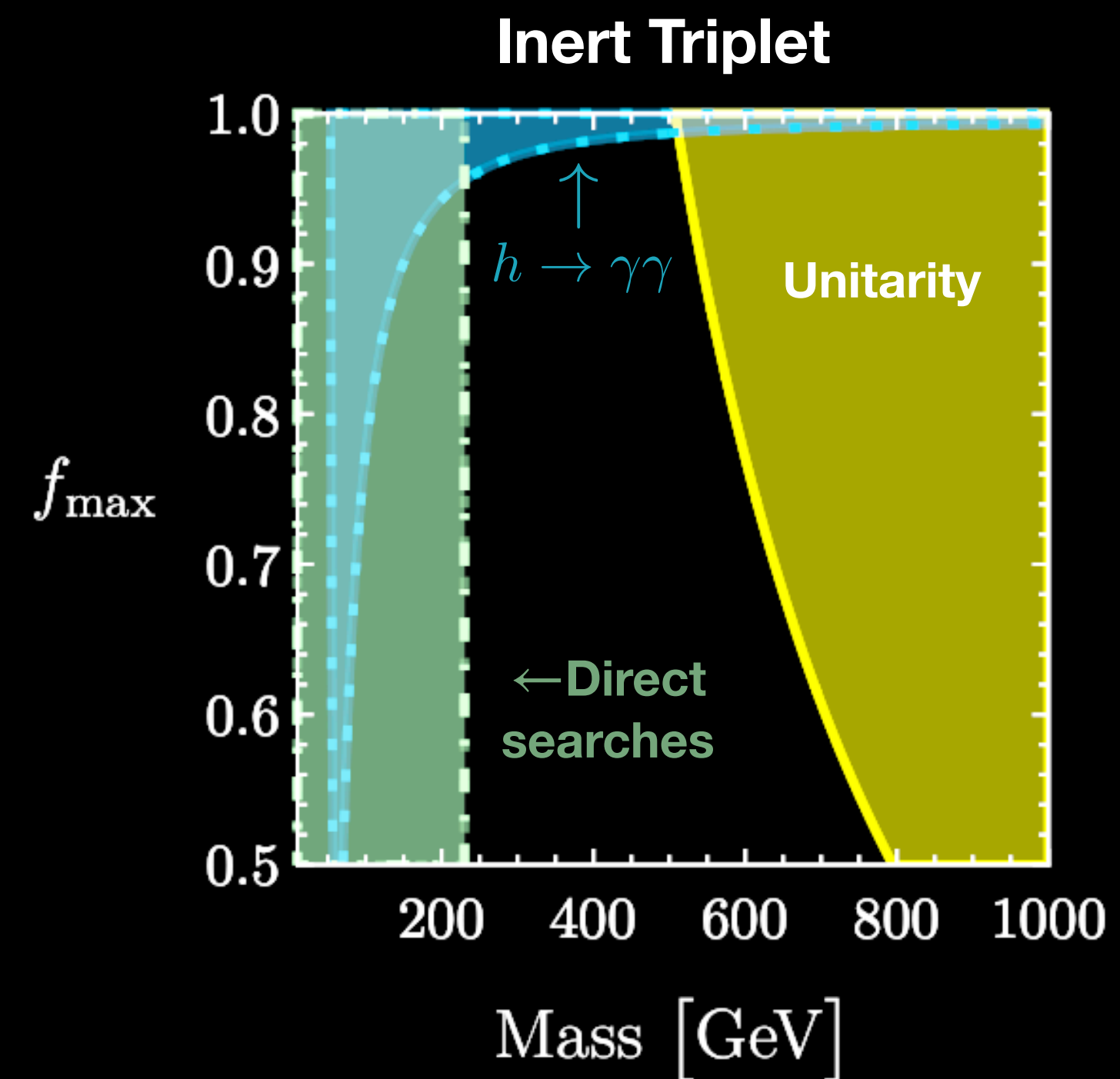
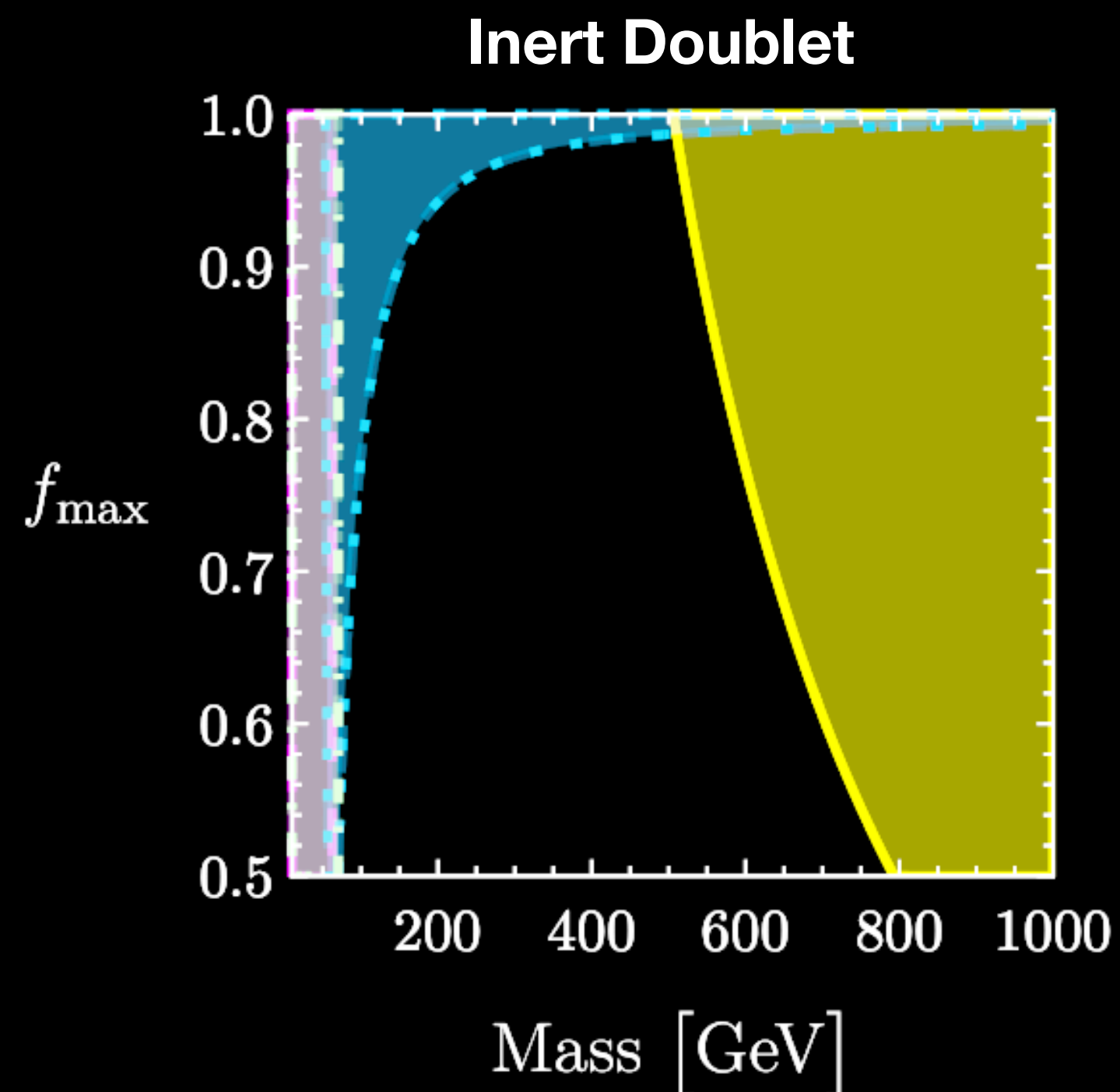
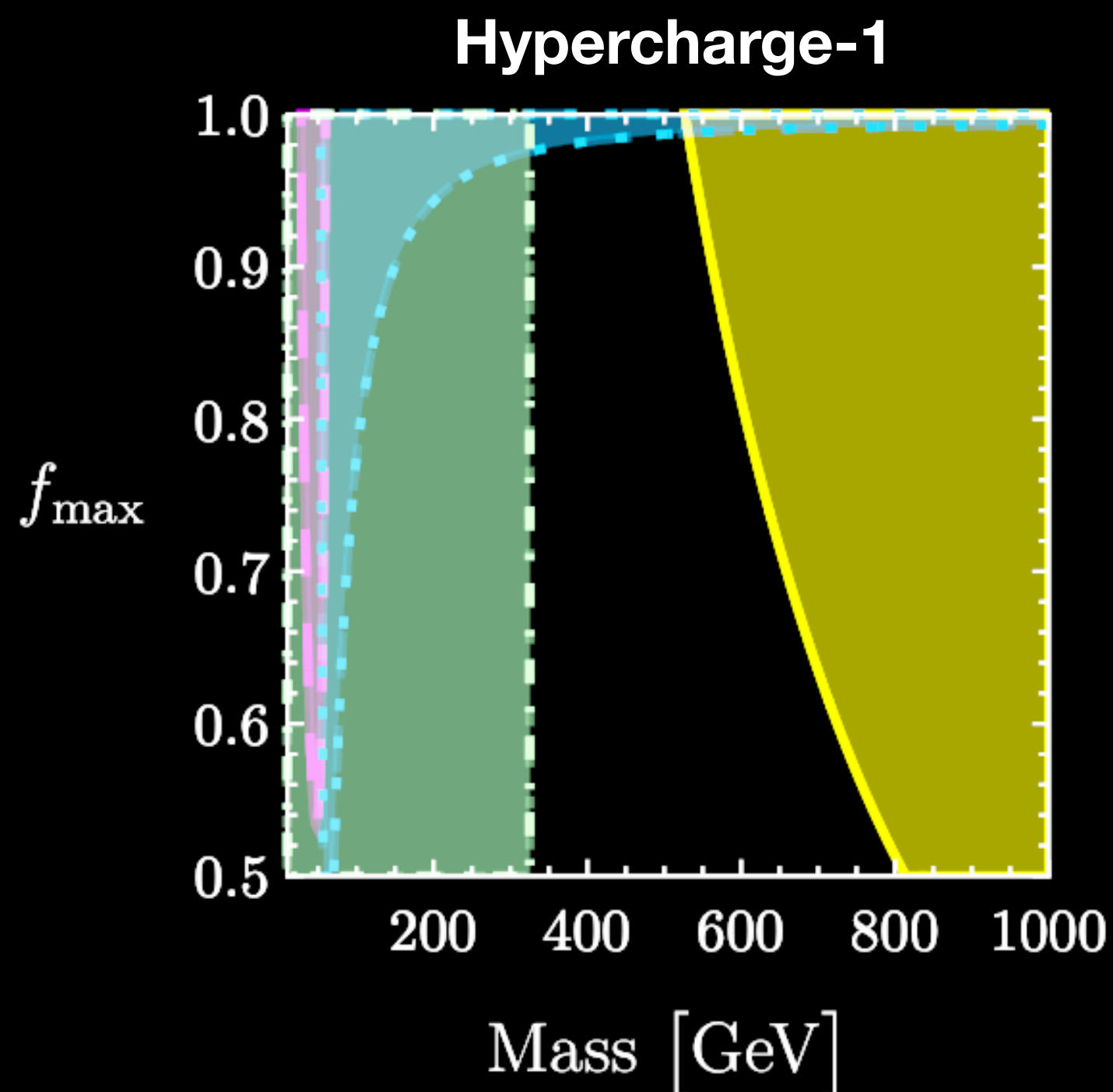
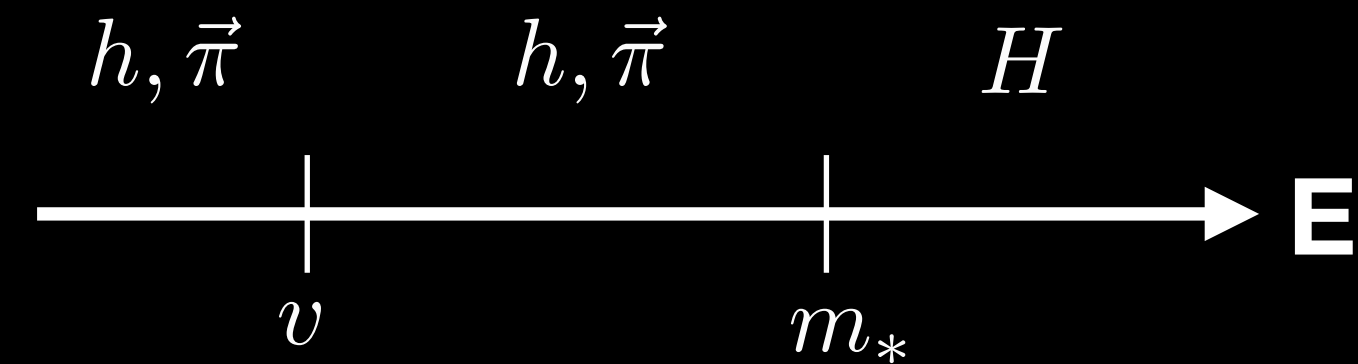
Vacuum stability?

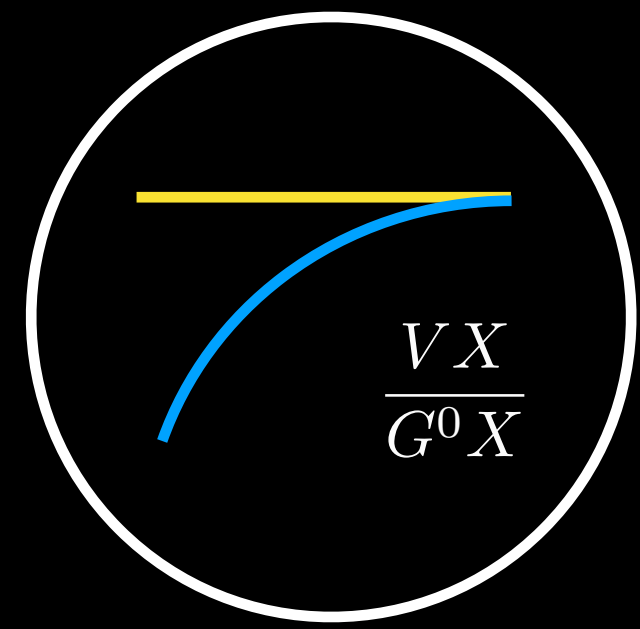


Electroweak Symmetry?

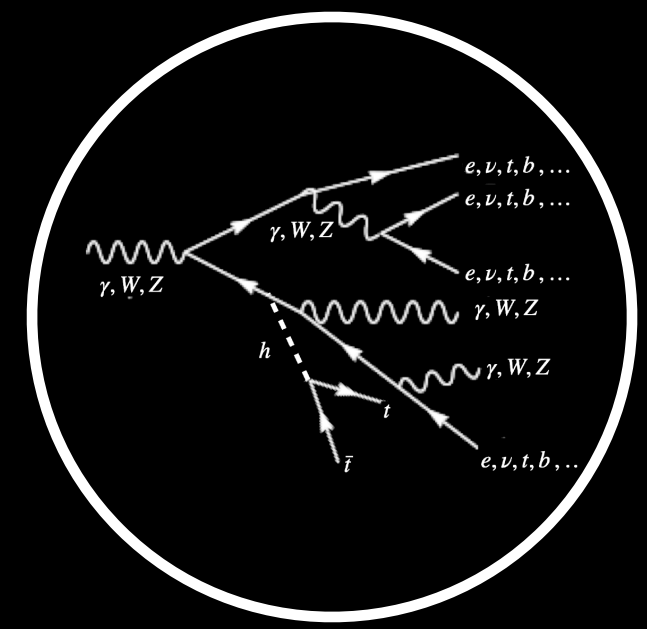
Local EFT of the Higgs does not have electroweak symmetry given e.g. extra EWSB or heavy particles acquiring $>1/2$ mass from Higgs. Many examples viable, fully covered by 10 TeV μ C.

[Banta, Cohen, NC, Lu, Sutherland, 2110.02967]



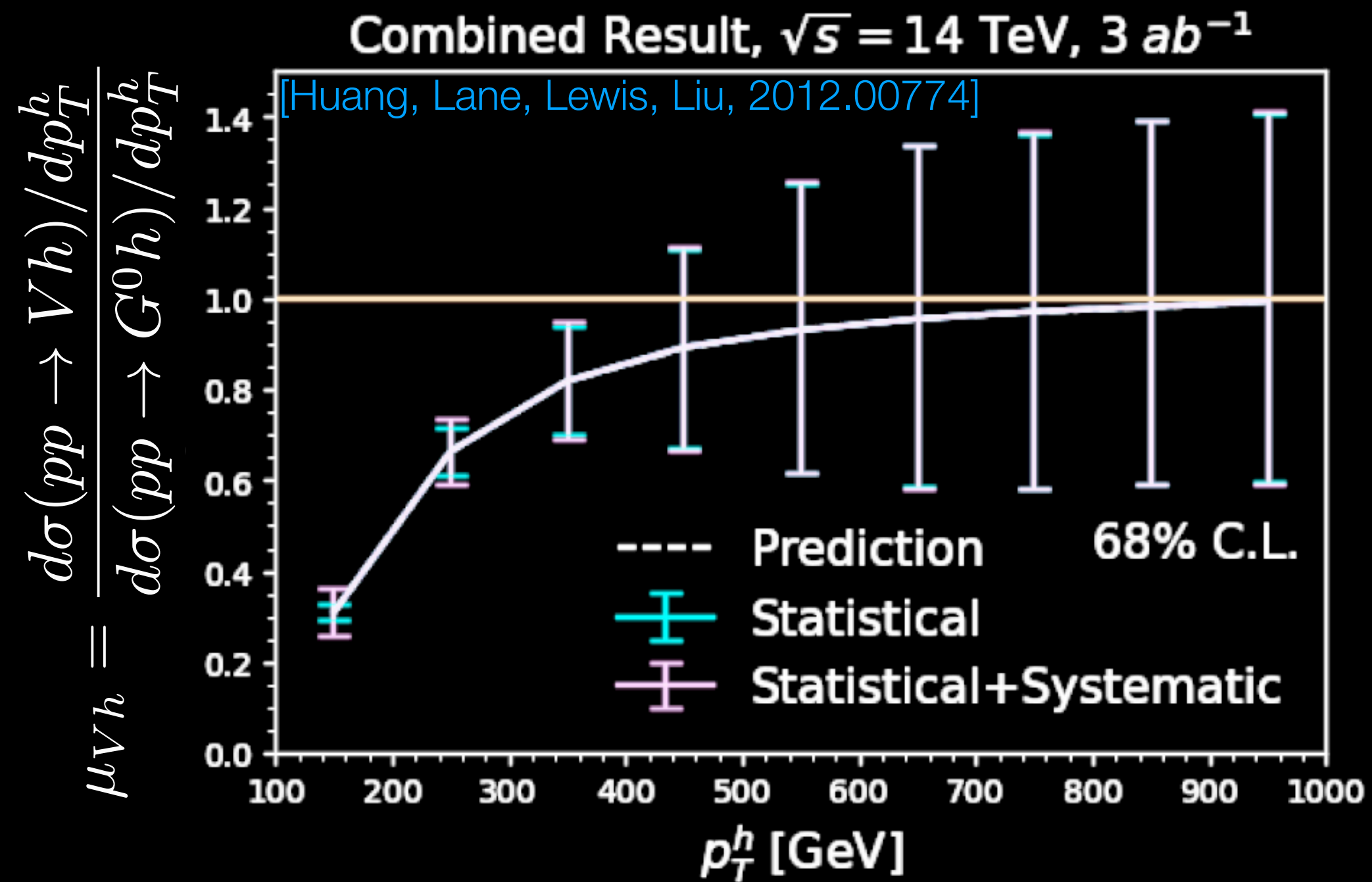


Electroweak Restoration, Electroweak Radiation

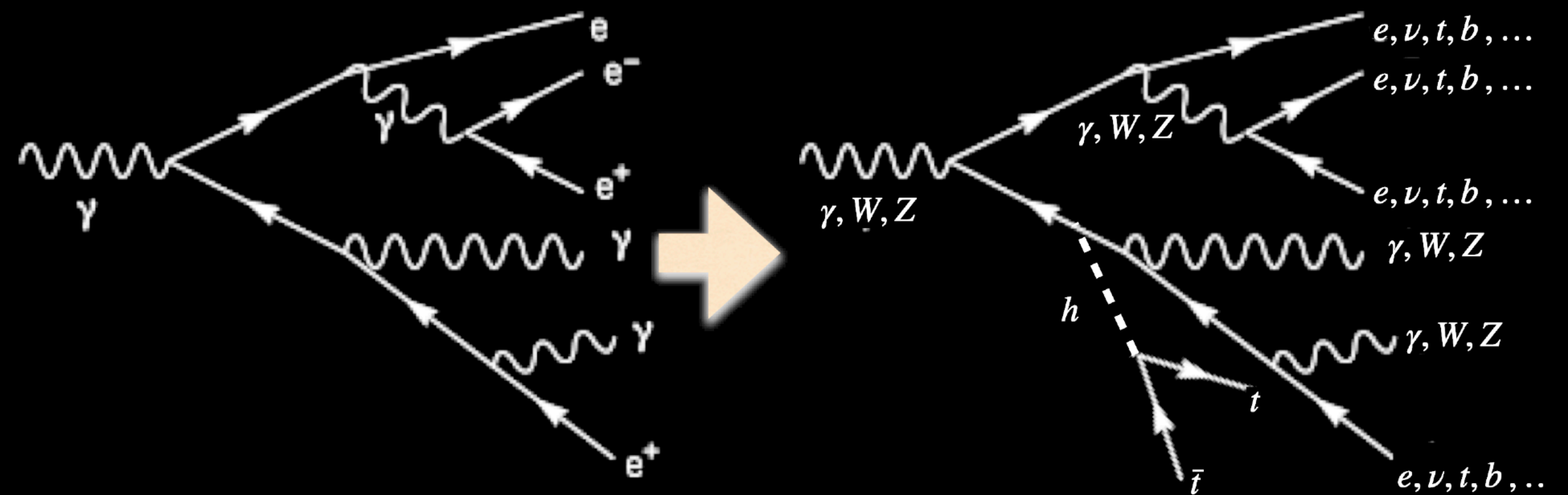


Experimentally demonstrate
Goldstone equivalence

Electroweak radiation:
Sudakov suppression of non-emission
probability becomes significant @ 10 TeV...



$$\exp \left[-\text{Casimir} \times \frac{g^2}{16\pi^2} \log^2 \left(\frac{E_{\text{cm}}^2}{m_W^2} \right) \right] \approx \exp[-1]$$



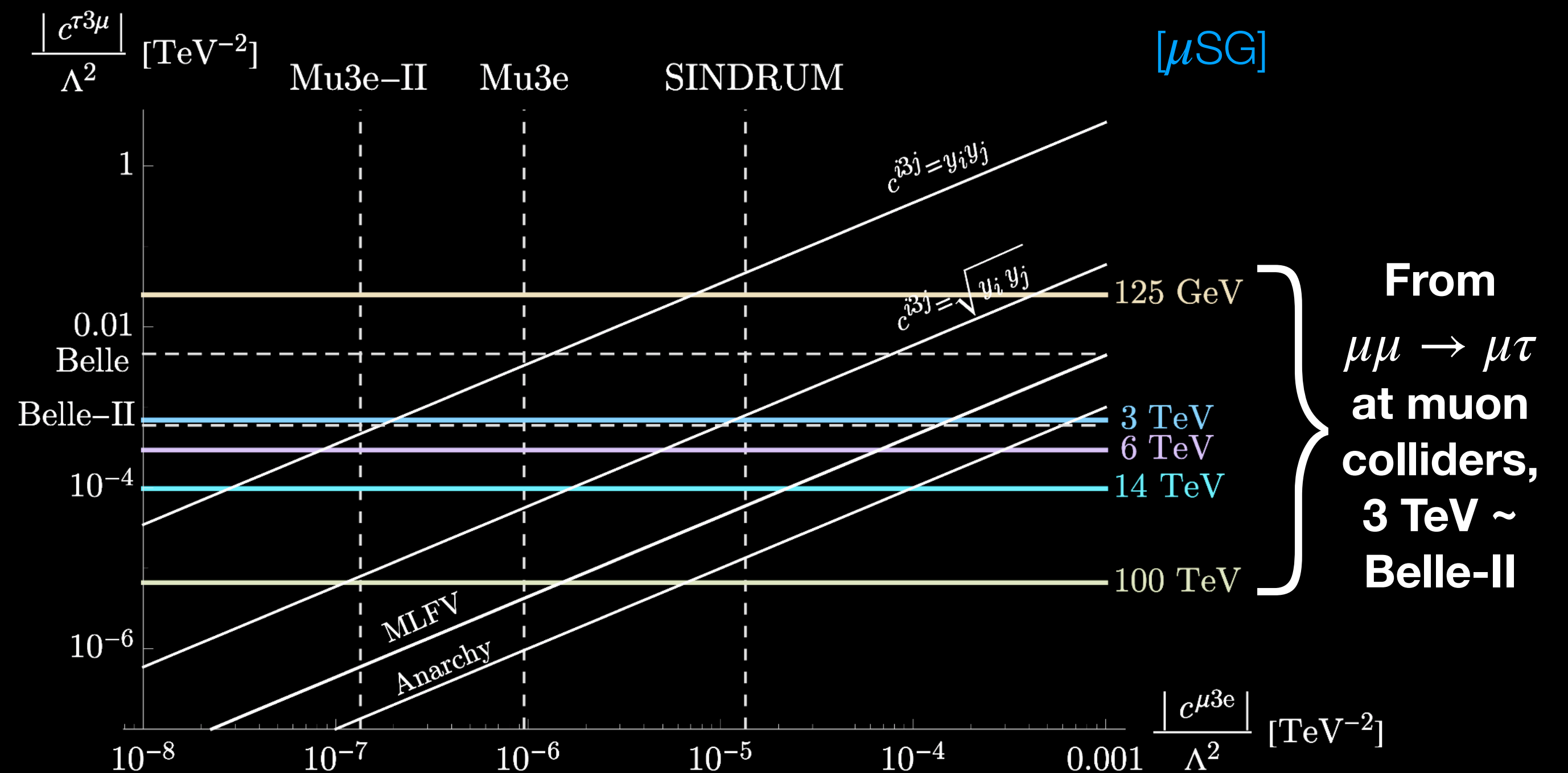
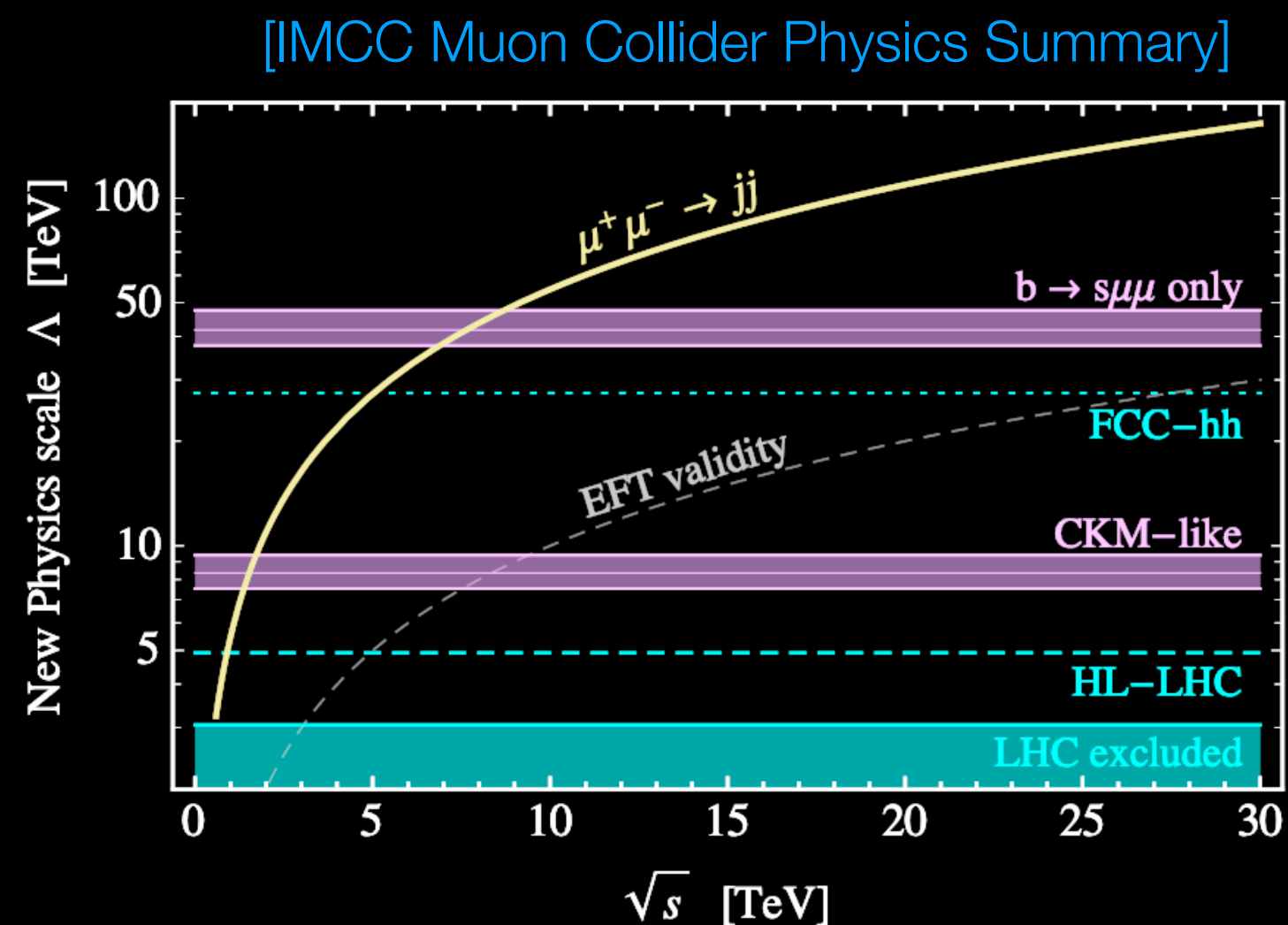
[Wulzer; Chen, Glioti, Rattazzi, Ricci, Wulzer 2202.10509]

What is the origin of flavor?

First high-energy accelerator to primarily collide second-generation fermions.

Direct access to hypothetical new particles associated with flavor structure

Indirect access to flavor structure via lepton flavor violating operators



An outstanding probe of explanations for B flavor anomalies, indicative of complementarity w/ future signs of flavor violation.

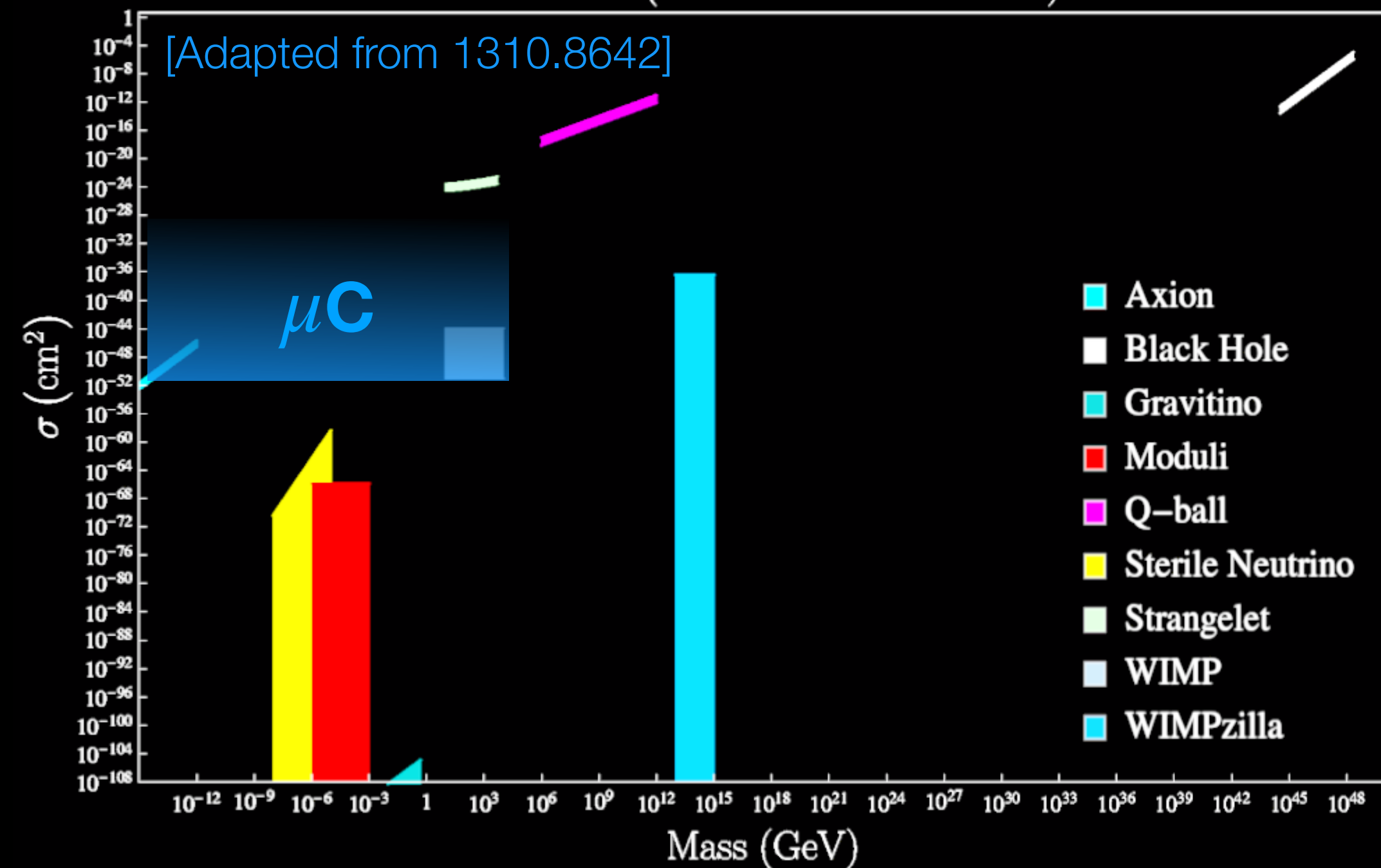
[Huang, Queiroz, Rodejohann, 2101.04956; Huang, Sana, Queiroz, Rodejohann, 2103.01617, Asadi, Capdevilla, Cesarotti, Homiller 2104.05720]

What is the nature of dark matter?

An ideal laboratory for producing & detecting weakly-interacting dark matter.

We know DM is there; coincidence of Ω_b , Ω_{dm} suggests interactions beyond gravitational

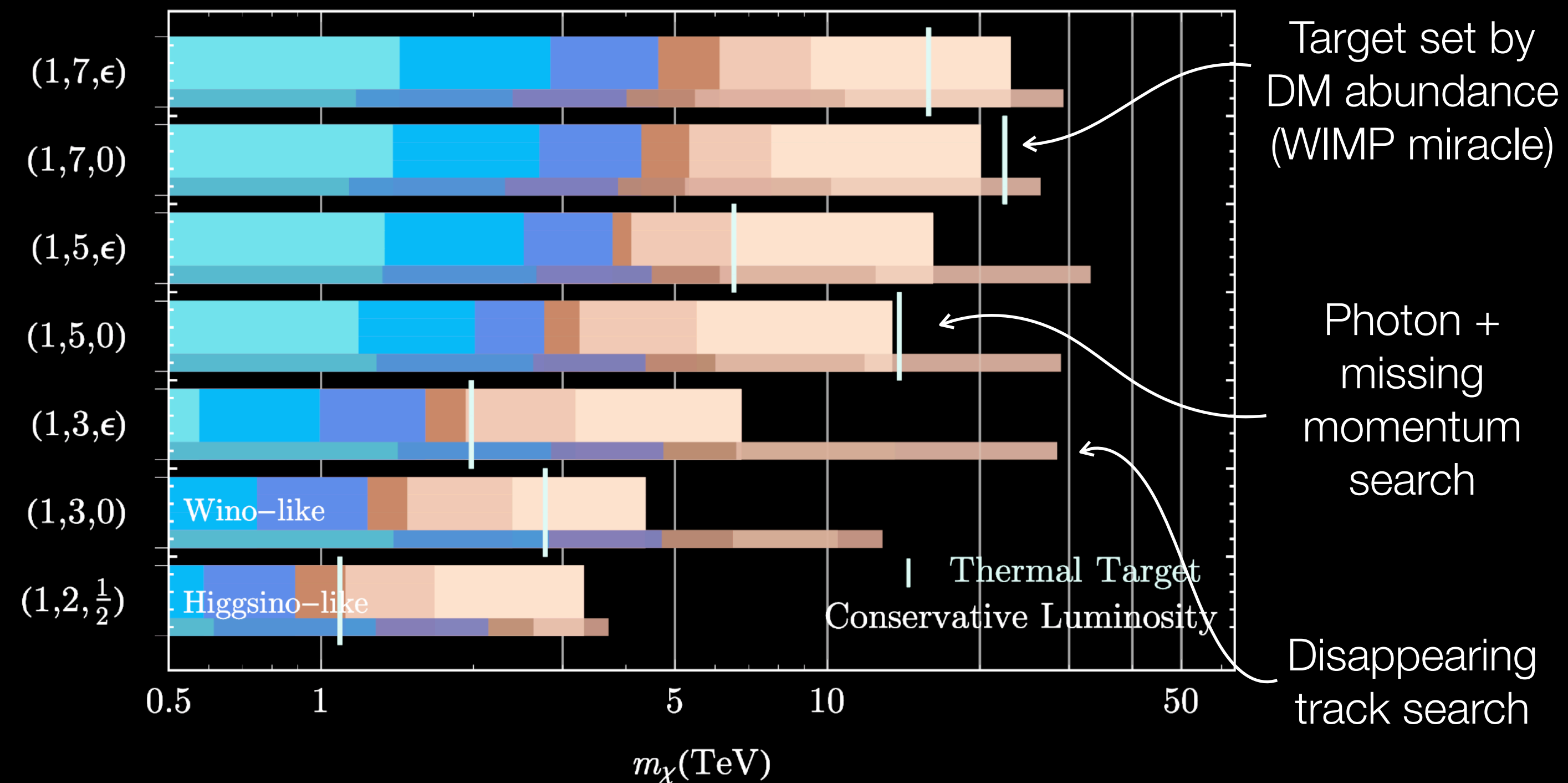
Cross Section (Xenon for Reference)



“Minimal dark matter”

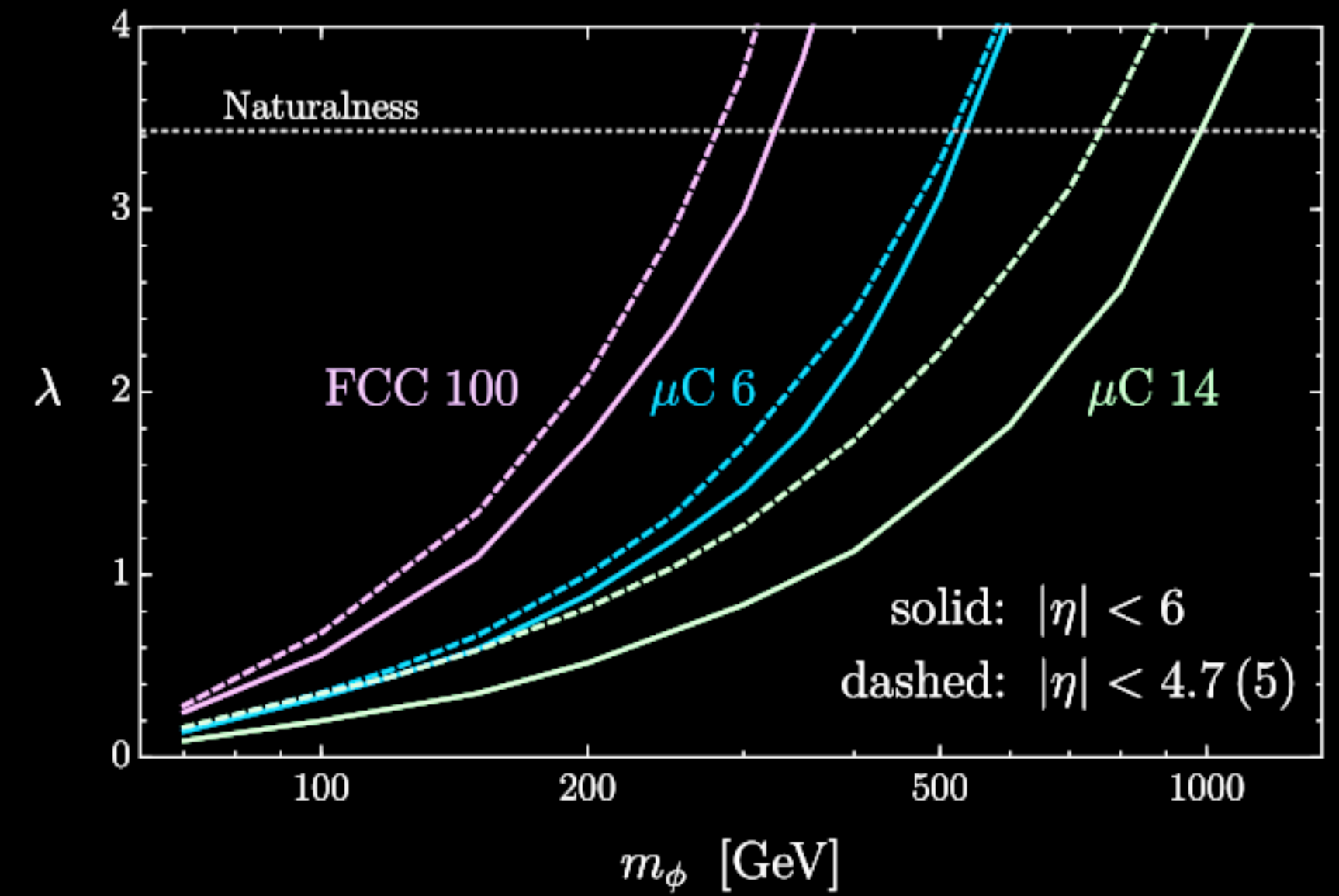
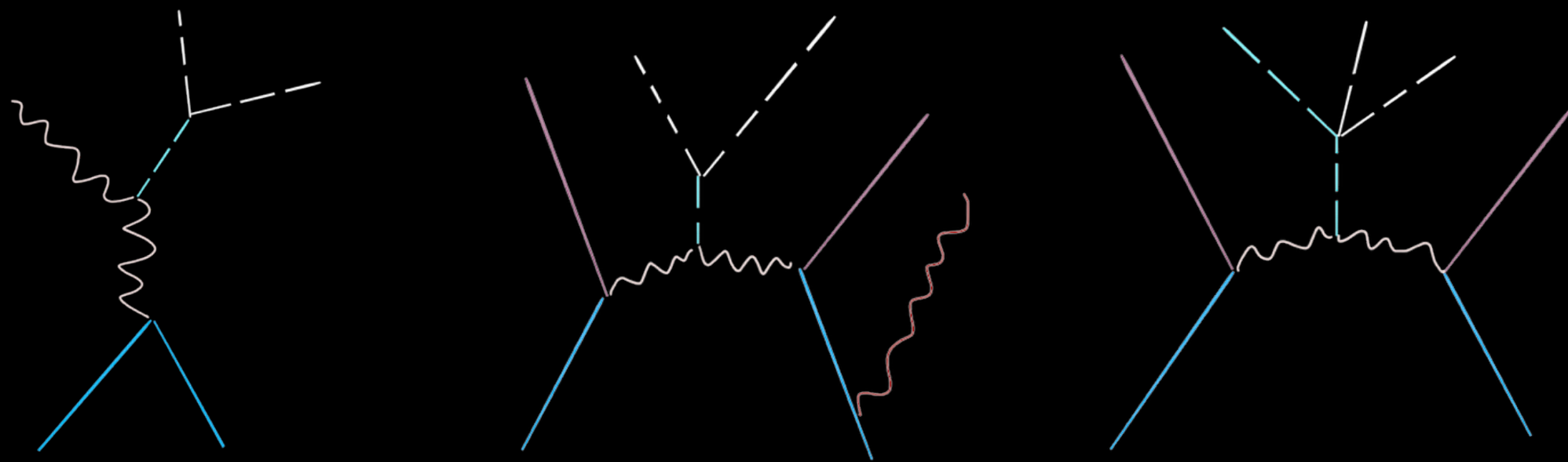
(Electroweak multiplet w/ neutral lightest particle)

Muon Collider 5σ Reach ($\sqrt{s} = 3, 6, 10, 14, 30, 100$ TeV)

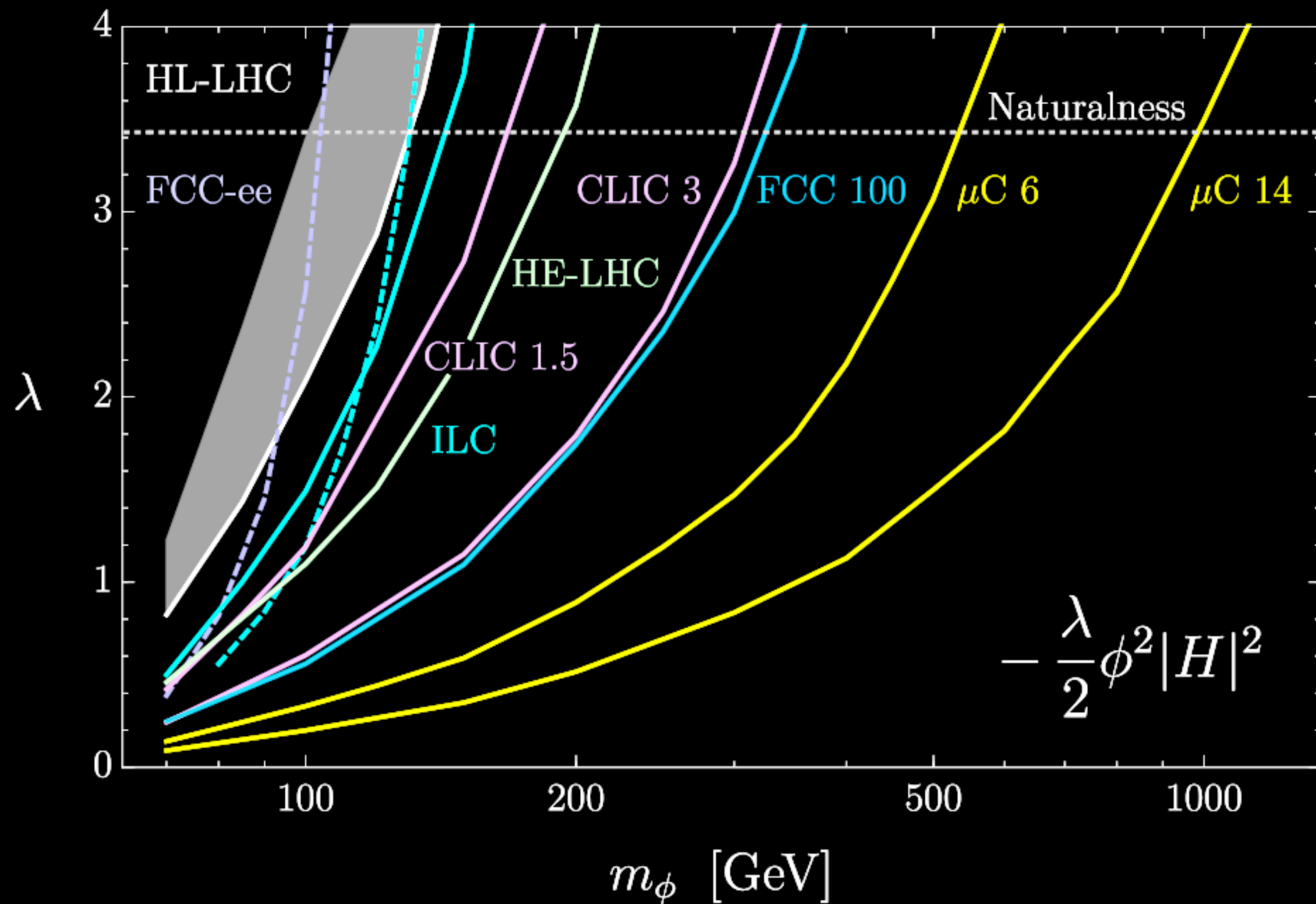


[Han, Liu, Wang, Wang, 2009.11287, lumi updated for μ SG], see also [Capdevilla, Meloni, Simoniello, Zurita 2102.11292; Bottaro, Buttazzo et al. 2107.09688 & 2205.04486]

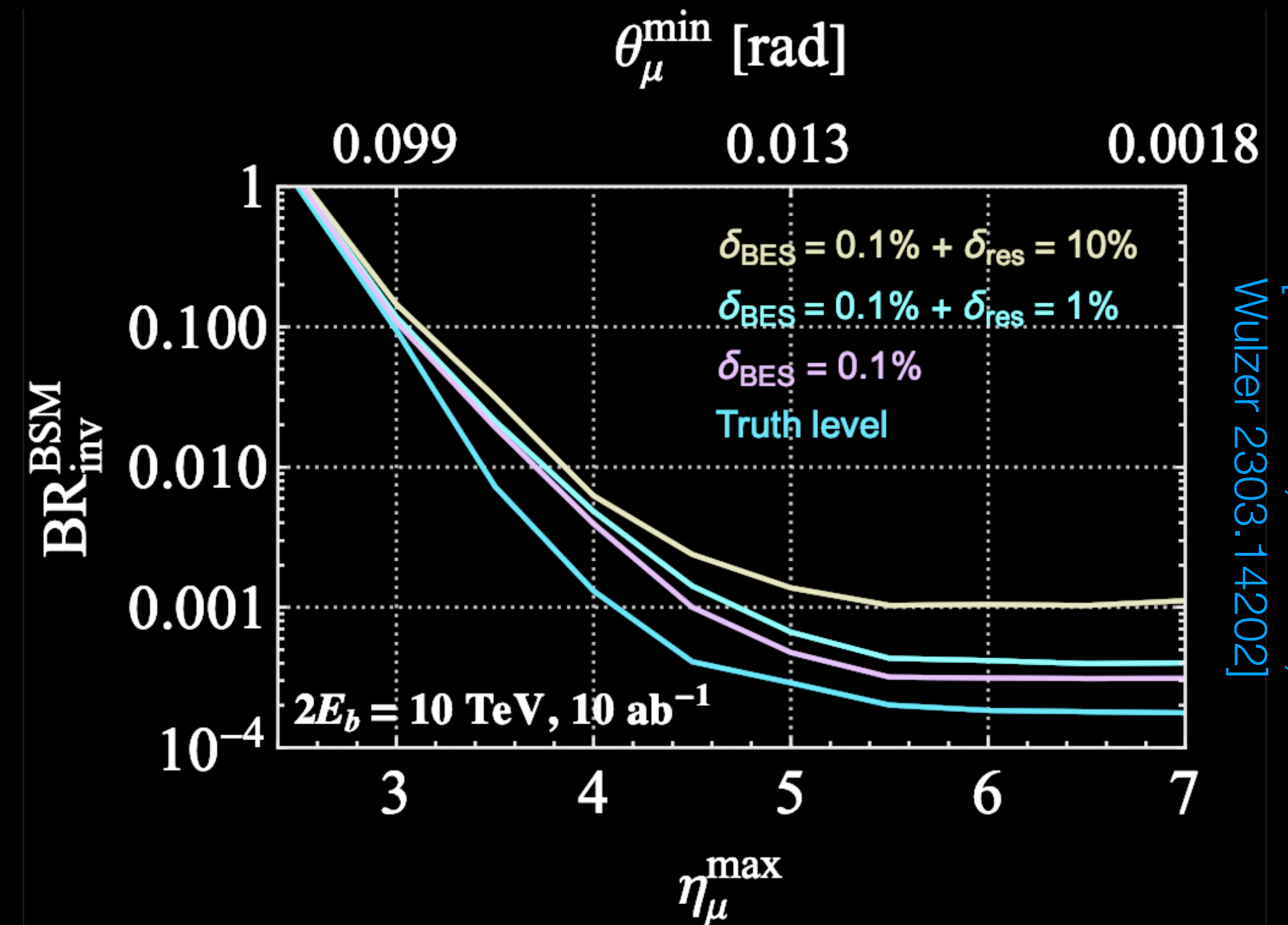
Higgs Portal DM



[1910.04170]

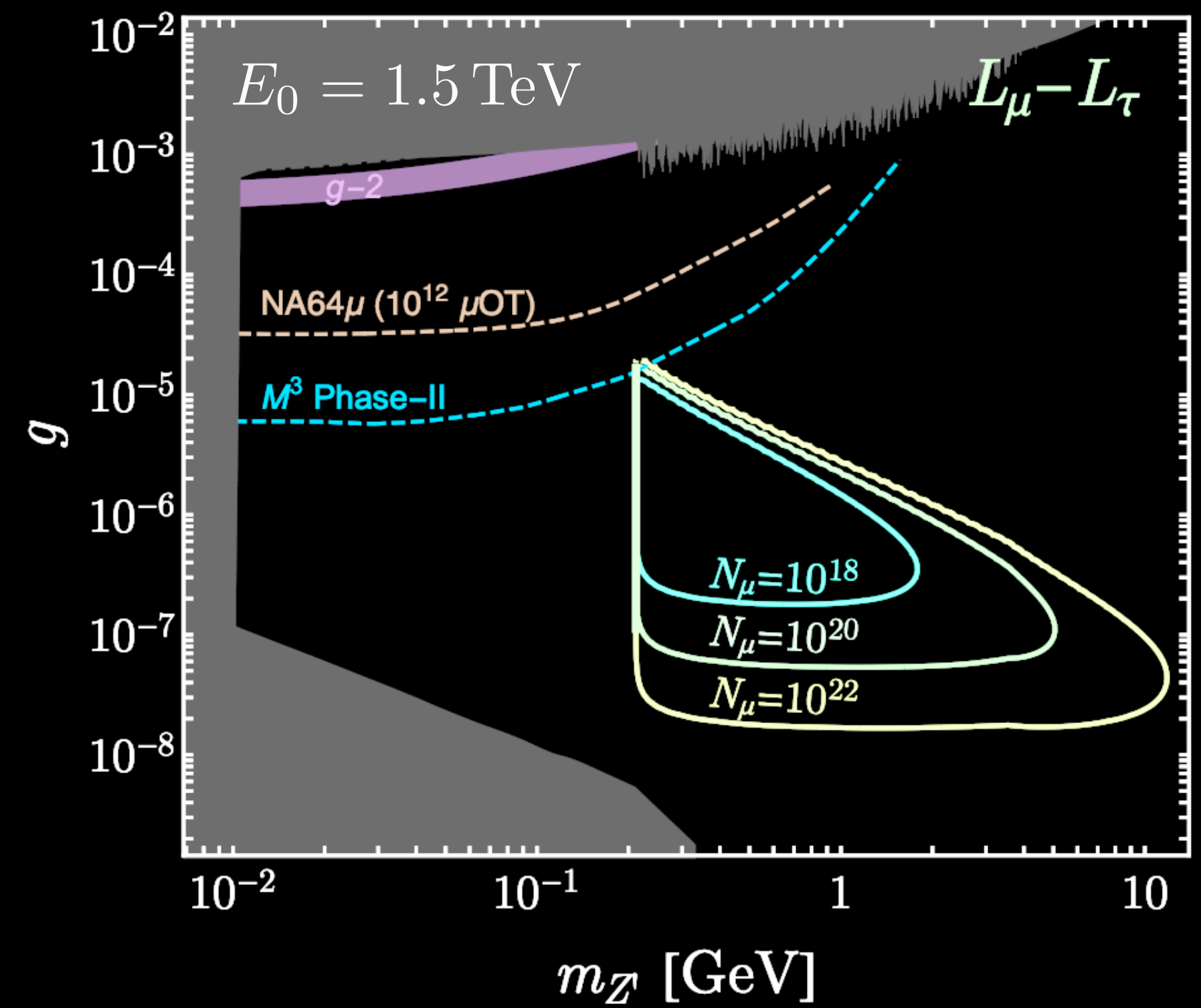
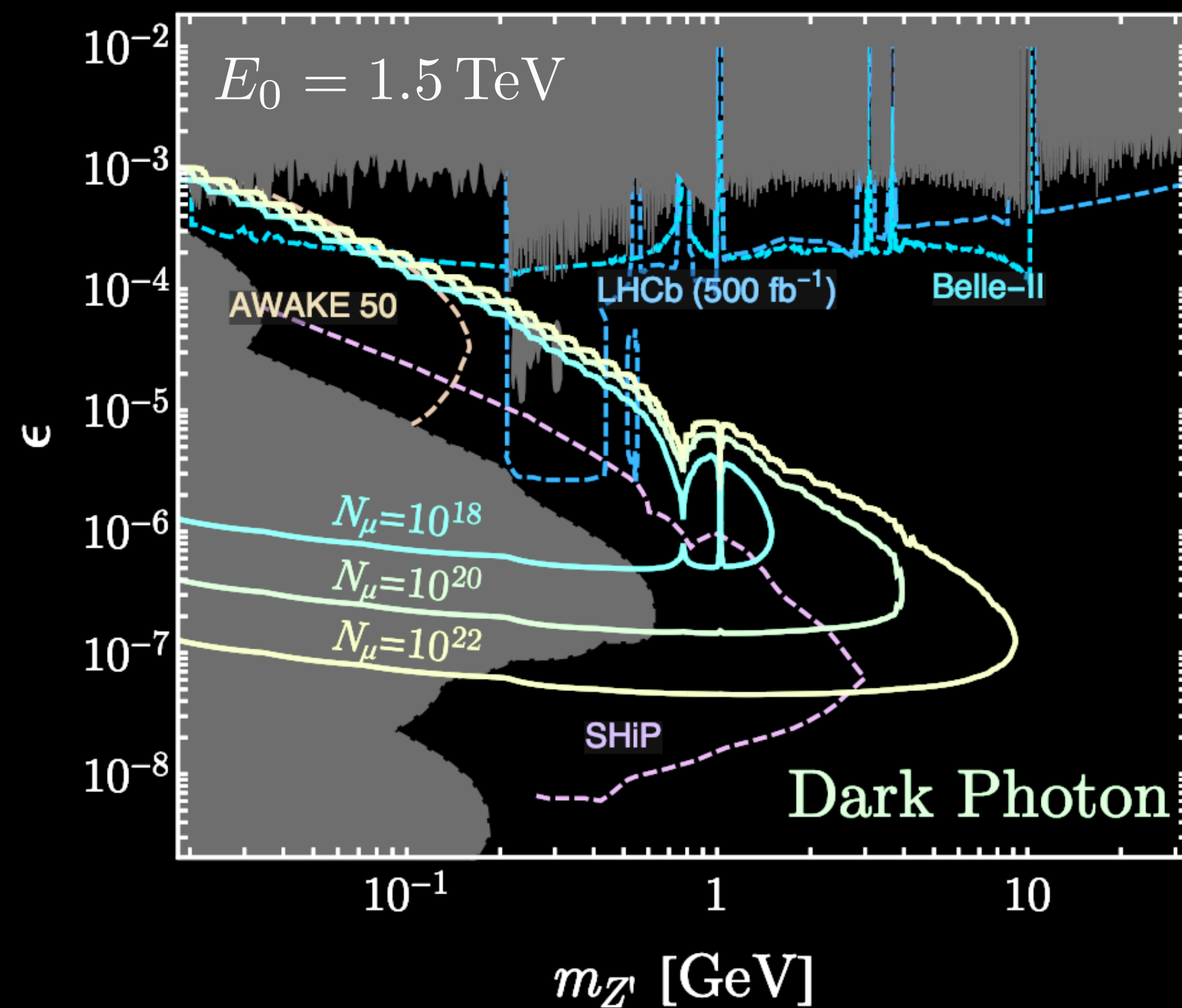
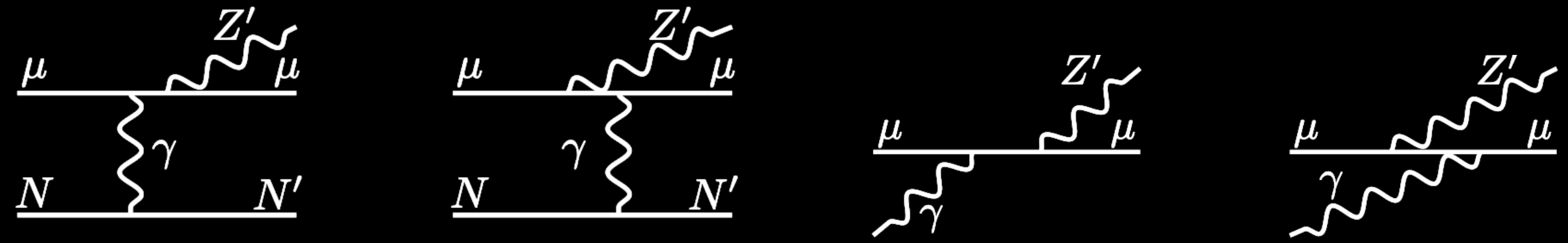
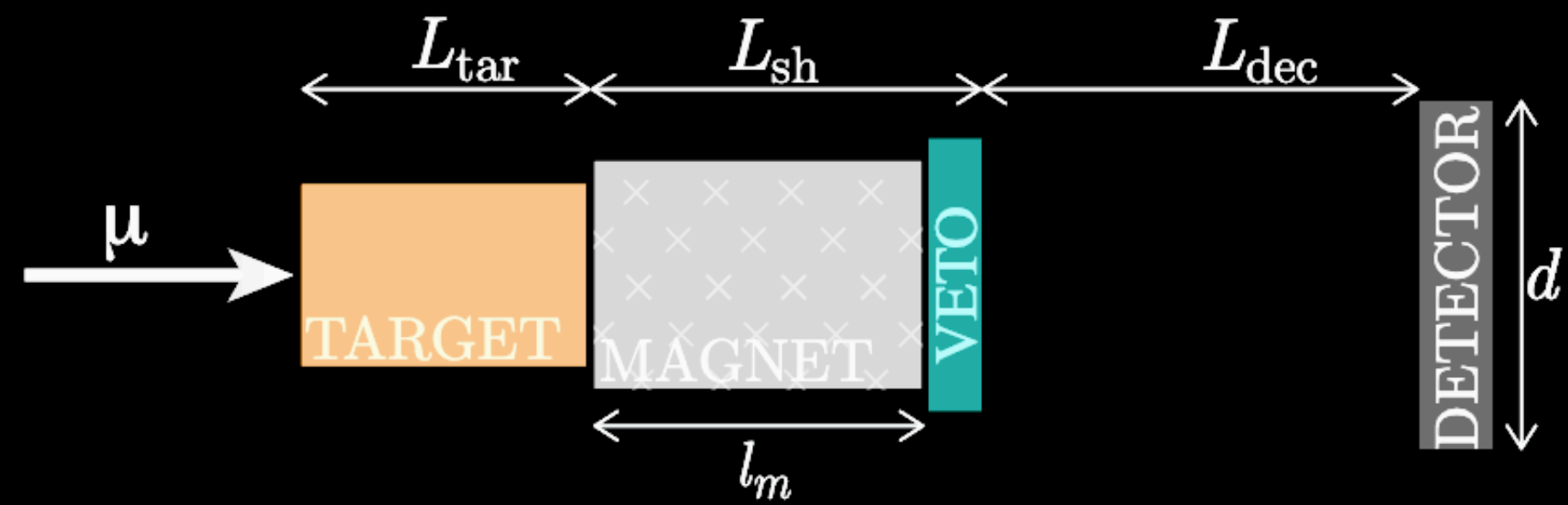


[Ruhdorfer, Salvioni, Weiler 1910.04170]



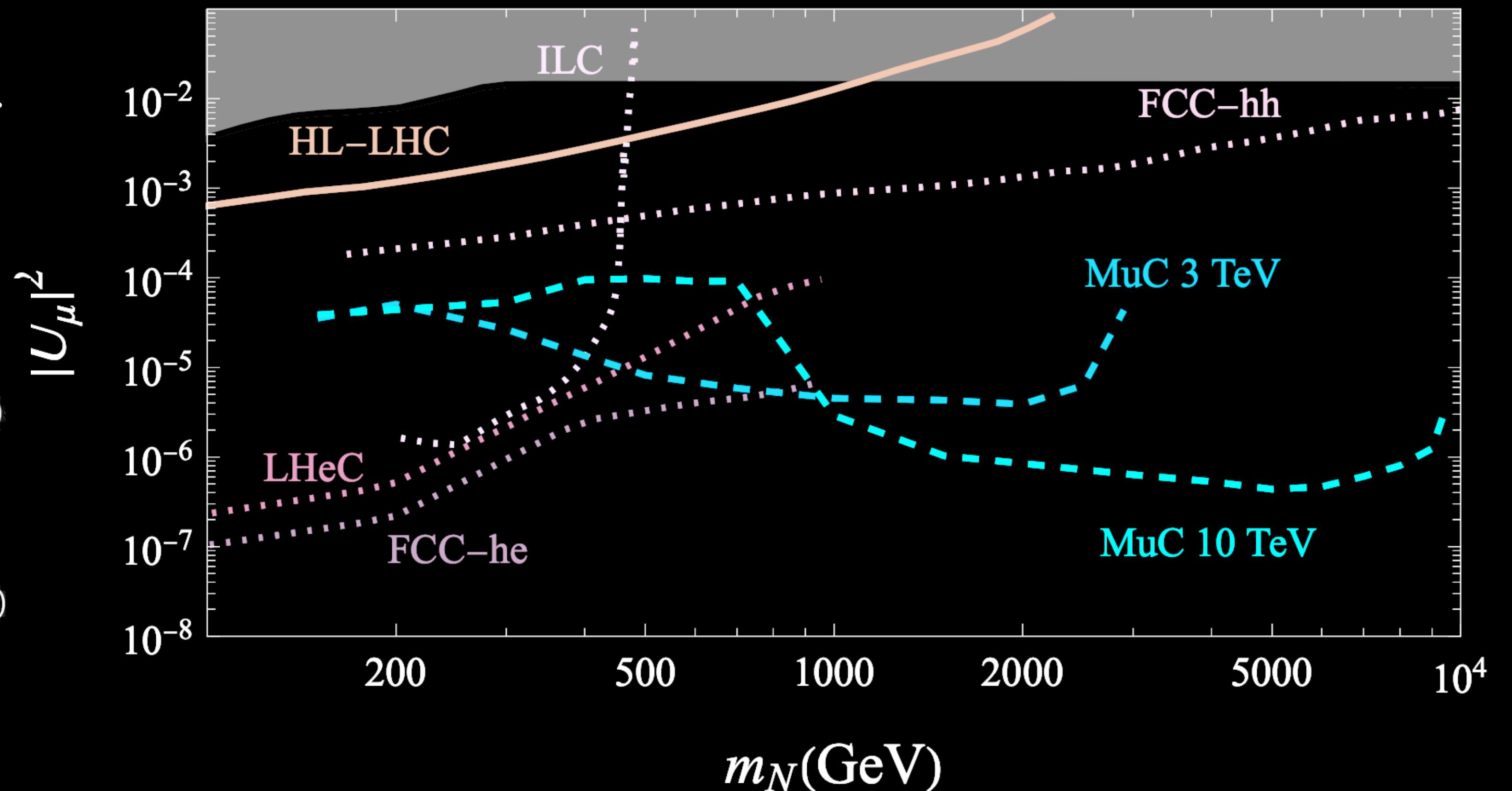
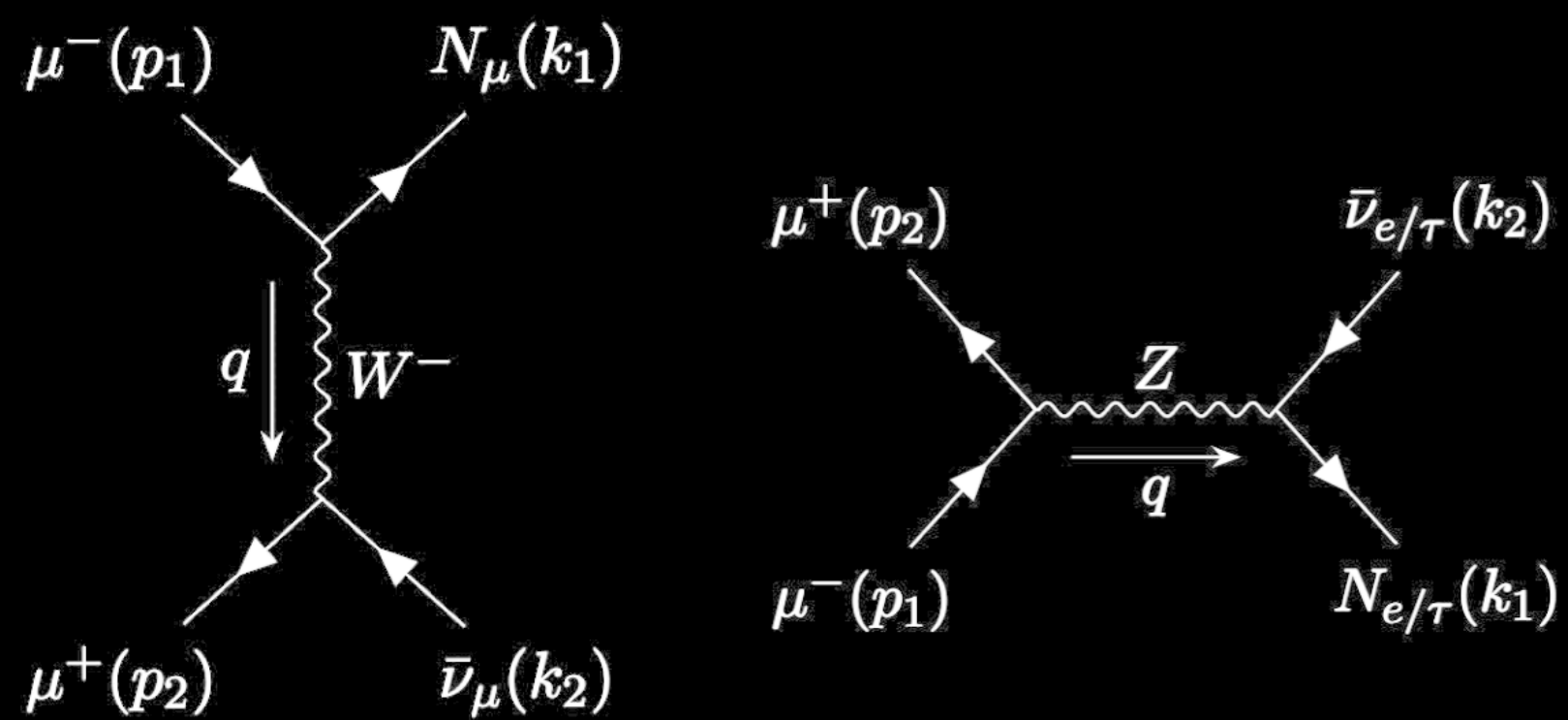
[Ruhdorfer, Salvioni, Weiler 2303.14202]

Dark Portals: μC Beam Dump?

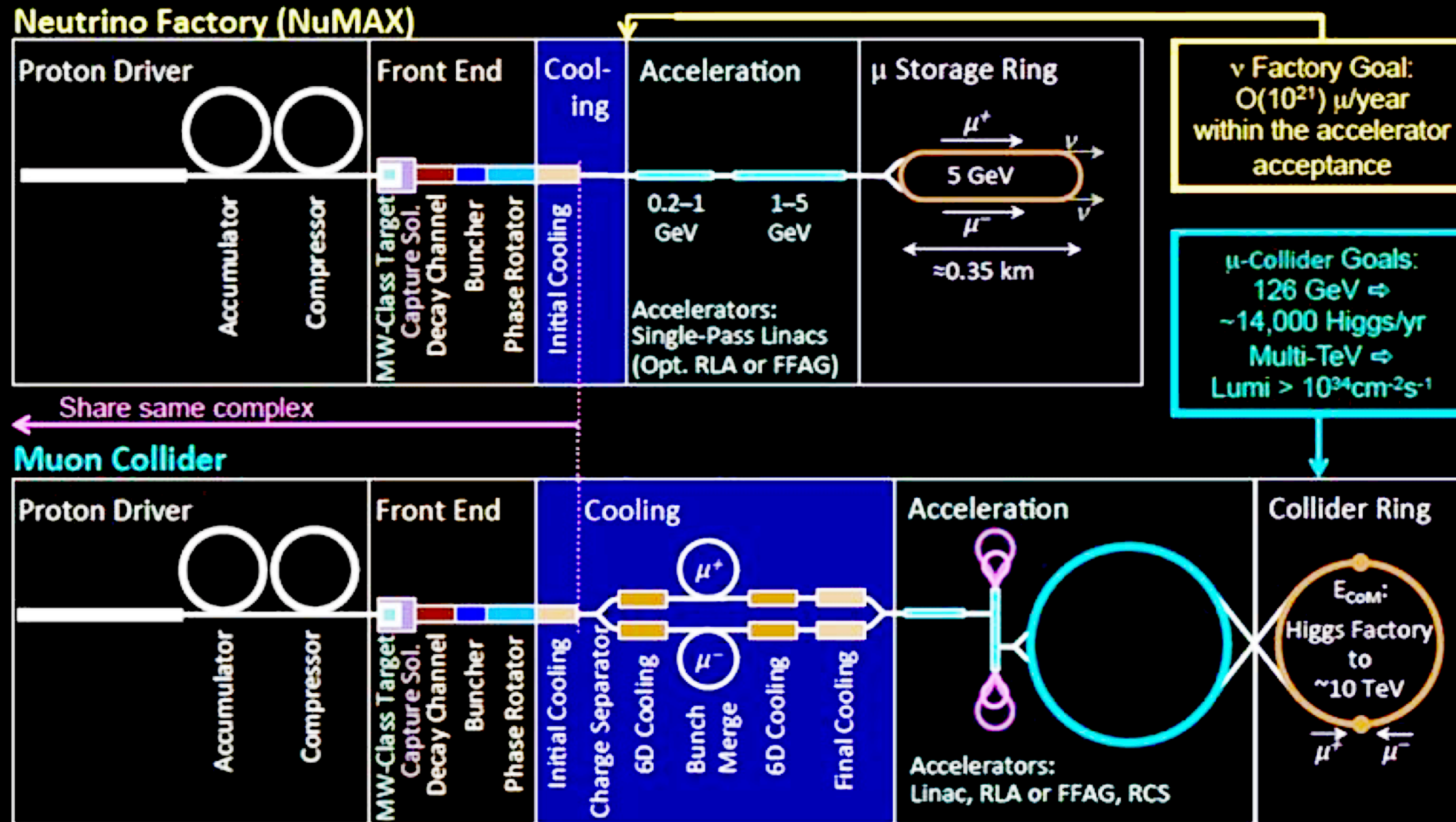


Nature of the Neutrino Sector: Heavy Neutral Leptons

$$\mathcal{L} \supset -\lambda_\nu \bar{L} \tilde{H} N - \frac{m_N}{2} \bar{N}^c N + \text{h.c.}$$

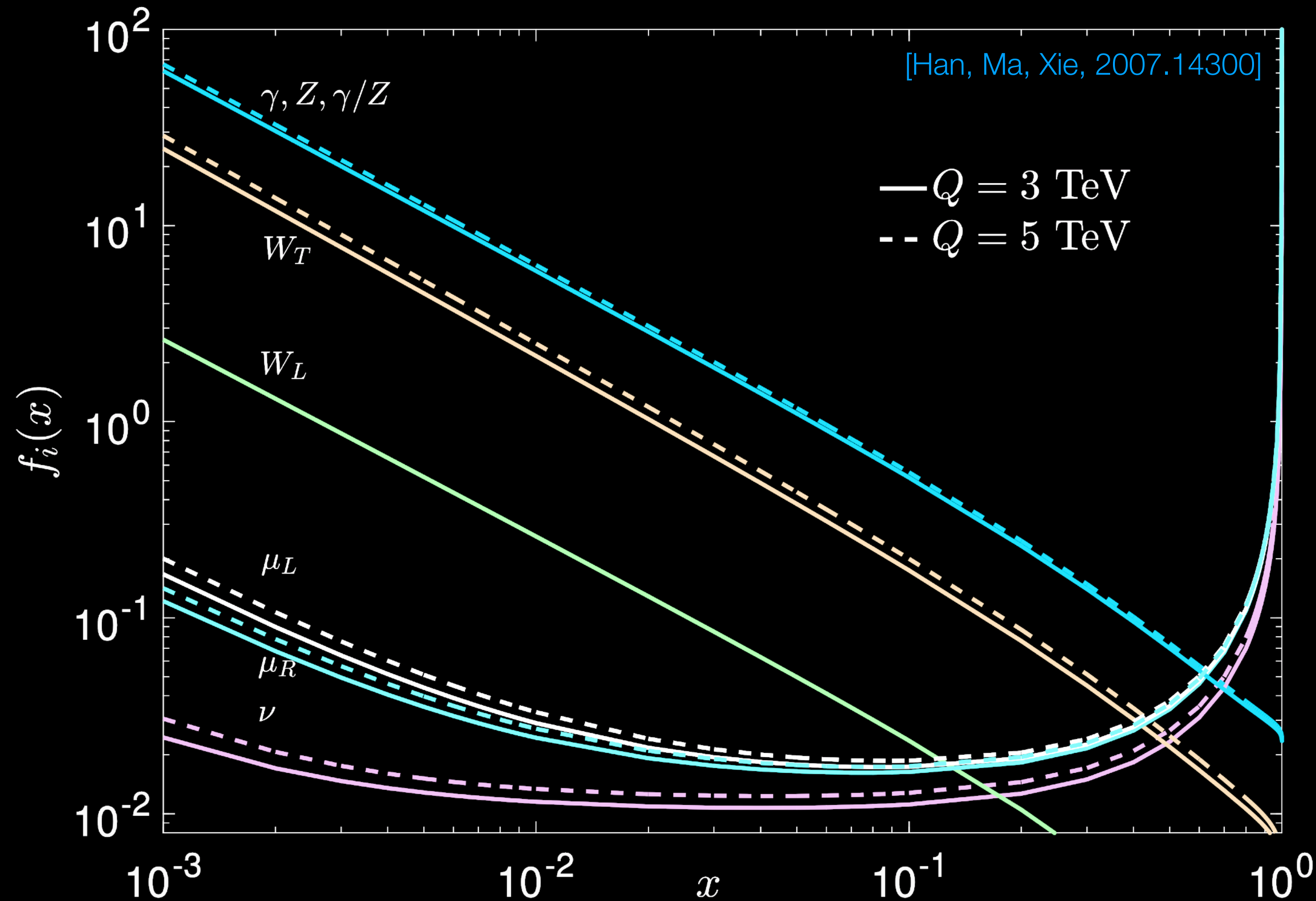


$\mu C/\nu F$ Complementarity



[Delahaye et al. 1803.07431]

High-energy Neutrinos?

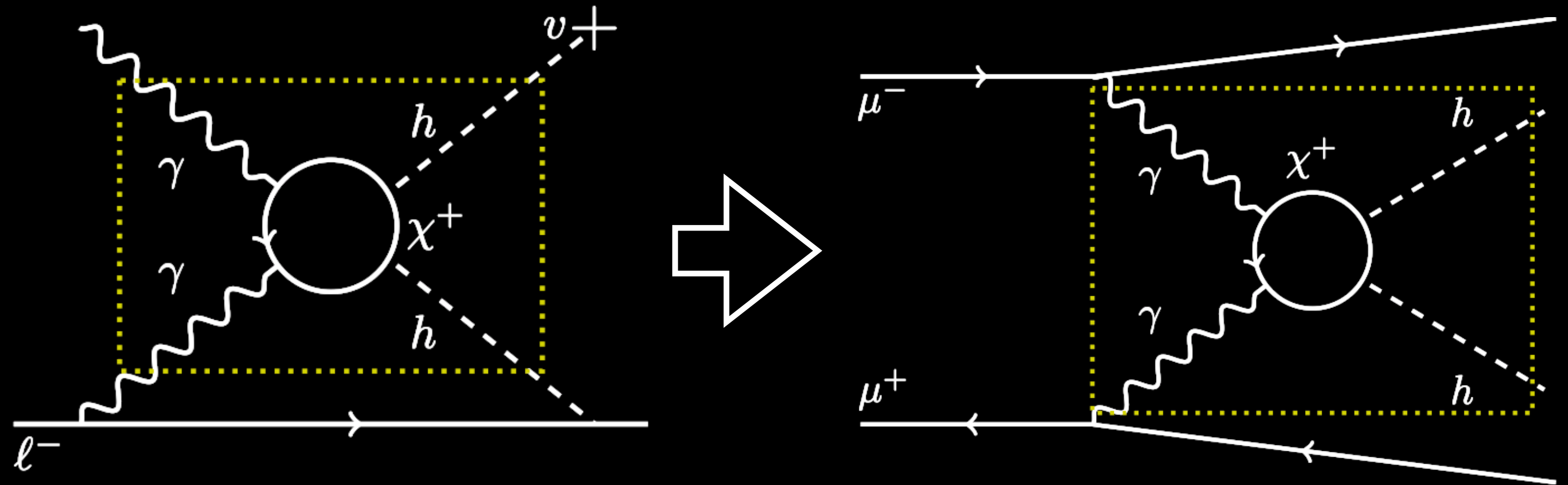


- Neutrino radiation for high-energy fixed target? Akin to FASER ν but w/ well-known neutrino flavor composition & spectrum, narrow beam.
- Neutrino-neutrino or neutrino-charged lepton collisions? Probe neutrino interactions, dark sector portals, ...

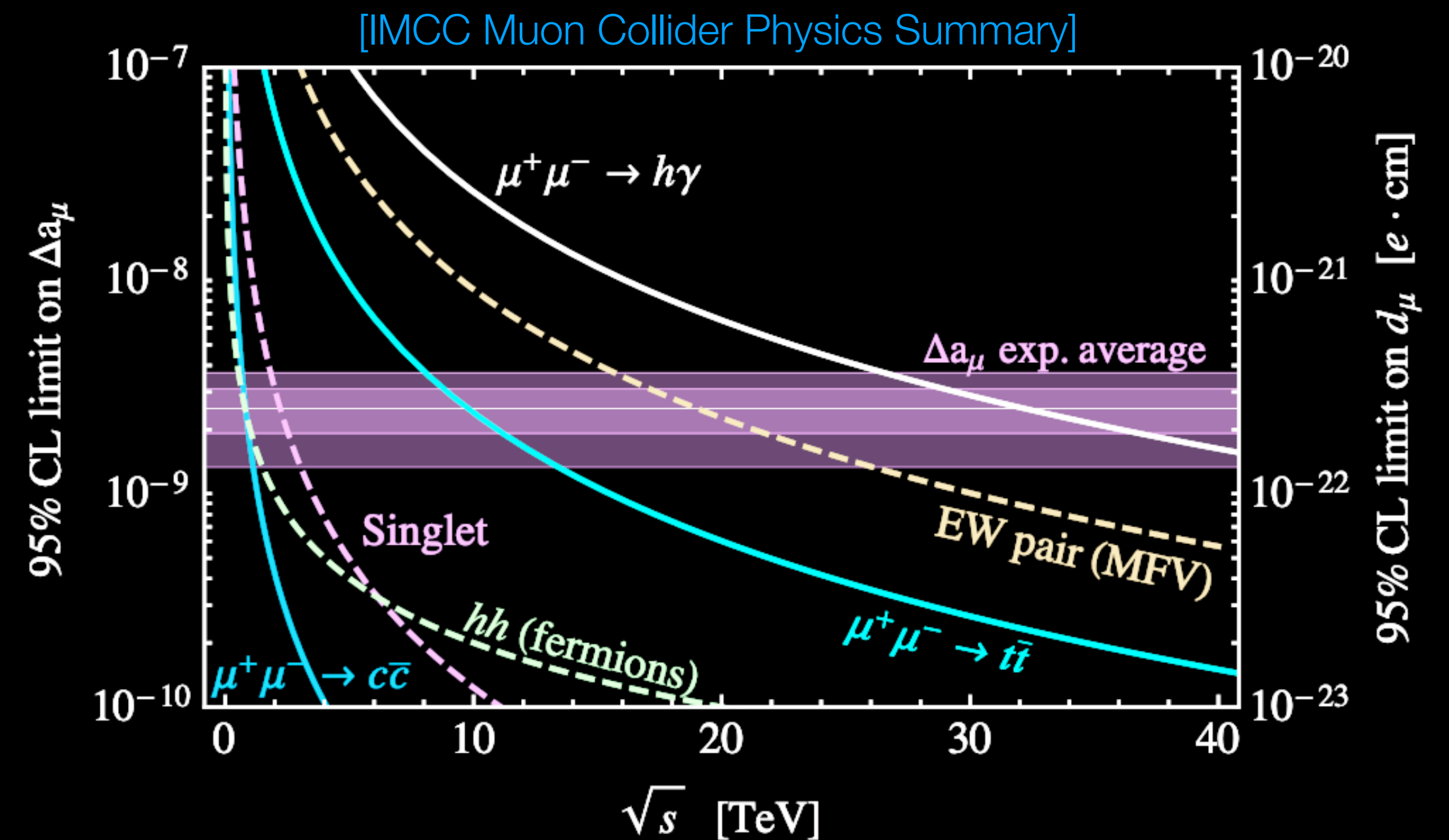
Compelling complementarity

E.g. next-gen. **electron EDM** experiments sensitive to ~ 20 TeV particles in Barr-Zee diagrams; same diagram probed in muon colliders

(See also: [Homiller, Lu, Reece 2203.08825])



Any new physics contributions to **Muon g-2** efficiently probed at muon colliders
 [Capdevilla, Curtin, Kahn, Krnjaic, 2006.16277; Buttazzo & Paradisi, 2012.02769; Capdevilla, Curtin, Kahn, Krnjaic, 2101.10334; Chen, Wang, Yao 2102.05619; Yin, Yamaguchi 2012.03928]



Conclusions



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Thank you!

